

# BES II and Future Dilepton Measurements at STAR

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#### Outline



- Motivation for dilepton physics at STAR
- Dileptons results in Beam Energy Scan Phase I
- Beam Energy Scan Phase II
  - Physics goals
  - Detector upgrades
- Near future dilepton studies at STAR
  - Measurement of  $\mu^+\mu^-$  continuum at STAR
  - Studies of correlated charm in IMR
  - Low  $p_T$  dilepton pairs in Isobar collisions
- Summary

# Motivation for Dilepton Physics at STAR

## Dileptons provide an excellent penetrating probe of the medium

- Leptons have low interaction cross section with QGP medium – they carry information to the final state
- Dilepton pairs are created throughout the entire lifetime of the system





#### Low Mass Range (LMR) $M_{\parallel} < 1.1 \text{ GeV/c^2}$

- In-medium modification of vector mesons
- Link to chiral symmetry restoration

#### Intermediate Mass Range (IMR) $1.1 < M_{\parallel} < 3.0 \text{ GeV/c^2}$

- Dominant contribution from semi-leptonic correlated charm decays
- QGP thermal radiation

#### High Mass Range (HMR) $M_{\parallel} > 3.0 \text{ GeV/c^2}$

• Primordial emission, Drell-Yan, J/ $\psi$  and Upsilon suppression

#### NA60 : Low Mass Range



#### PRL 96 (2006) 162302



NA60's dimuon measurement:

- Compare  $\rho$  broadening/melting models
- Data favors Rapp/Wambach model - ρ broadening through interactions with hadronic medium
- Link to chiral symmetry restoration?

#### NA60 : Intermediate Mass Region





PRL 100 (2008) 022302

Measurment of inverse slope parameter ( $T_{eff}$ ) from  $m_T$ distribution

- $T_{eff}$  = 205  $\pm$  12 MeV
- Indicative of thermal radiation from partonic medium
- No mass dependence visible in IMR

Challenge to disentangle correlated charm from QGP thermal radiation

 STAR upgrades (Muon Telescope Detector & Heavy Flavor Tracker) will help – more later



# STAR Dielectron Results in Beam Energy Scan Phase I

## STAR Dielectrons in BES I

- BES I beam energies :  $\sqrt{s_{NN}} = 7.7, 11.5, 14.5, 19.6, 27.0, 39.0, and 62.4 GeV$
- Insufficient statistics available below 19.6 GeV for dielectron analyses
- Probe a wide range of initial conditions
  Approximately constant total baryon
- Approximately constant total baryon density in the  $\sqrt{S_{NN}}$  range from 19.6 to 200 GeV

How does LMR excess yield evolve?

- Are in-medium modification effects visible at all energies down to  $\sqrt{s_{NN}}$  = 19.6 GeV?
- Is the enhancement consistent with model predictions at these energies?





STAR Dielectrons in BES I





8

0.8

#### STAR Acceptance Corrected Excess Spectra Acceptance-corrected spectra for Au+Au at 27, 39,

- Au+Au at 19.6 and 200 GeV
- Normalized to  $(dN_{ch}/dy)_{y=0}$  to cancel out volume effects
- 17.3 GeV (NA60) and 19.6 GeV consistent



Hohler & Rapp: "Is p-meson melting compatible with chiral restoration?"

d<sup>2</sup>N/dydM)/(dN<sub>ch</sub>/dy) (20 MeV)<sup>-1</sup>

10<sup>-5</sup> .

0<sup>-6</sup>

10-7

10<sup>-8</sup>

10<sup>-9</sup>

10<sup>-10</sup>

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# Link to Chiral Symmetry Restoration

Strong Experimental evidence that  $\rho$  meson "melts"

- Input for Phenomenological approach:
- Vector meson SF from phenomenological model, verified against experimental data
- T dependence of condensate, from lattice QCD

QCD Sum Rules : constrain vector / axial vector SFs individually

$$\frac{1}{M^2} \int ds \frac{\rho_{V/A}(s)}{s} e^{-s/M^2} = \sum_n C_n \langle O_n \rangle$$

Weinberg sum rules : difference between vector & axial vector SFs  $\int ds(\rho_V - \rho_A)s^n = f_n$ 



Quantatively compatible with (approach to ) chiral restoration

Hohler, Rapp *PLB 731 (2014) 103* But, still need microscopic calculations of  $a_1(1260)$  ... (Massive Yang-Mills)



# Link to Chiral Symmetry Restoration



**Effective Chiral Lagrangian** 

- Gauge  $\rho$  and  $\mathbf{a}_1$  into chiral pion Lagrangian
  - Massive Yang Mills in hot pion gas starting point for evaluation of chiral restoration in medium





## Beam Energy Scan Phase II 2019 - 2020

#### RHIC Beam Energy Scan Phase II



Purpose: Refine our understanding of the phase structures of QCD matter

- Beam time : 2019 2020
- Revisit lower BES I energies (19.6 GeV and below), possibly add 9.1 GeV
- Gain significantly more statistics necessary for dilepton measurements!
- Systematically study dielectron continuum from  $\sqrt{s_{NN}}$  = 7.7 to 19.6 GeV

Table 2. Event statistics (in millions)	needed for	Beam Energy	gy Scan Pha	se-II for vari	ous observables.
Collision Energy (GeV)	7.7	9.1	11.5	14.5	19.6
$u_B$ (MeV) in 0-5% central collisions	420	370	315	260	205
Observables					
$\overline{R_{CP}}$ up to $p_T = 5 \text{ GeV}/c$	_		160	125	92
Elliptic Flow (\$\$ mesons)	100	150	200	200	400
Chiral Magnetic Effect	50	50	50	50	50
Directed Flow (protons)	50	75	100	100	200
Azimuthal Femtoscopy (protons)	35	40	50	65	80
Net-Proton Kurtosis	80	100	120	200	400
Dileptons	100	160	230	300	400
Required Number of Events	100	160	230	300	400

STAR Whitepaper for BES II



## Dileptons in BES II

- BES I :  $\sqrt{s_{NN}}$  = 19.6 to 62.4 GeV
- Dilepton emission dominant in T<sub>c</sub> region
- ~constant total baryon density
- Emission proportional to lifetime

BES II :  $\sqrt{s_{NN}}$  = 7.7 to 19.6 GeV

- Average temperature of medium ~constant
- Probe life time + baryon density dependence of the ρ-meson spectral function



#### **Collision energy**

# Dileptons in BES II

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BES II :  $\sqrt{s_{NN}}$  = 7.7 to 19.6 GeV

- Average temperature of medium ~constant
- Probe life time + baryon density dependence of the ρ-meson spectral function
- Total baryon density expected to increase by ~2x
- Close to QCD critical point?
  - Increase in correlation length
  - Critical slow down anomalous increase in fireball lifetime
- Linking top RHIC energies with SPS, FAIR energies
- Comparison with NA60, HADES, CBM
- At FAIR energies, probe lifetime + total baryon density + temperature

NA60+ proposed at SPS – overlap with RHIC and FAIR



#### Collision energy

# BES II Detector Upgrades : iTPC

#### STAR proposal to upgrade the inner sector of its TPC

w/ run11 acceptance

w/ cocktail mas

0.8

M<sub>ee</sub> GeV/c<sup>2</sup>

w/ flat mass

- Increase rapidity coverage
- Lowers  $p_{\tau}$  threshold •
- Improves dE/dx (PID) resolution

#### Substantial improve dielectron systematics

- Systematics expected to be reduced by ~2x
- Improved PID  $\rightarrow$  increased electron purity
- Lower  $p_T$  threshold  $\rightarrow$  better acceptance in LMR ٠







## BES II Detector Upgrades : eTOF

#### STAR proposal to install CBM TOF

- Endcap TOF (eTOF) mounted on east pole tip
- Increases PID rapidity coverage
- Takes advantage of large iTPC coverage

**Dielectron Measurement:** 

• Allows rapidity differential

- LMR excess expected to depend on total baryon density
- π yields drop by 2x from y=0 to y=1.2
   → "baryon density" drops by factor of two.
- Analysis at y=1.2 is equivalent to lowering the beam energy





# Near Future STAR Dilepton Measurements current - 2018

## Measurements of $\mu^+\mu^-$ Continuum at STAR



• STAR's Muon Telescope Detector upgrade (completed 2014) allows for new studies of the dimuon continuum at RHIC energies with STAR.

Species	Energy (GeV)	~Sampled Luminosity	$22 \times 10^{\circ}$ STAR Preliminary $20 = 2015 \text{ p+p at } \sqrt{s} = 200 \text{ GeV}$
p+p (63% MTD)	500	28 pb <sup>-1</sup>	$ \begin{array}{c}                                     $
p+p	200	122 pb <sup>-1</sup>	Image: Sign and
p+Au	200	409 nb <sup>-1</sup>	10 Signal 10 Signal 10 Signal ↓ ■
d+Au	200	94 nb <sup>-1</sup>	
Au+Au 2014	200	14 nb <sup>-1</sup>	4 ψ(2S)
Au+Au 2016	200	12 nb <sup>-1</sup>	
			$M_{\rm uu}$ (GeV/c <sup>2</sup> )

- High quality p+p data set : provides baseline for Au+Au analyses
  - Clear  $\omega$ ,  $\rho$ , J/ $\psi$ , and  $\psi$ (2S) peaks visible in p+p collisions @  $\sqrt{s_{NN}}$  = 200 GeV
- High statistic Au+Au data samples triggered by dedicated dimuon trigger
- Dimuon channel fewer background source in LMR compared to  $e^+e^-$  channel
- Analysis of data from p+p and Au+Au @  $\sqrt{s_{NN}}$  = 200 GeV is ongoing

# Disentangling Correlated Charm in IMR

Heavy Flavor Tracker:

- Provides precise tracking
- Allows charm to be better distinguished via secondary vertex reconstruction
- Analyses ongoing with 2014 & 2016 data





Muon Telescope Detector:

- Allows for a dedicated  $e-\mu$  trigger
- Gain handle on charm contribution via  $e-\mu$  correlation
- Study possible charm modification in medium



# Isobaric Collisions : Low $p_T$ Excess

- STAR and ALICE have observed significant excess production of J/ $\psi$  in peripheral A+A collisions at low p<sub>T</sub> ( p<sub>T</sub> < 300 MeV/c)  $d^{4}$  10<sup>2</sup>
- Two potential sources:
  - Photo-nuclear  $\propto Z^2$
  - Photon-Photon  $\propto Z^4$

Investigate Z-dependence of low- $p_T$  dielectron excess:

- Measure & compare in A=96 isobars:
  - ${}^{96}_{44}$ Ru +  ${}^{96}_{44}$ Ru
  - $\frac{96}{40}$ Zr +  $\frac{96}{40}$ Zr

Proposed for 2018



# Summary

- At SPS, NA60 and CERES demonstrated the physics potential of accurate dilepton measurements
- STAR has developed a strong dilepton program
  - Detailed study of dielectron production in Au+Au and U+U @ top RHIC energy
  - BES I Allowed measurement of low mass excess down to 19.6 GeV, compare with SPS energy
- Progress in thermal field theory computations of dilepton production in heavy-ion collisions has led to robust description of LMR excess
- BES Phase II (2019-2020): systematic dilepton measurements down to  $\sqrt{s_{NN}}$  = 7.7 GeV
  - Probe dependence on total baryon density
  - Measure  $p_T$  distribution's inverse slope parameter ( $T_{eff}$ )
  - Look for anomalous increase in yield  $\rightarrow$  indicative of critical behavior
  - Utilize iTPC and eTOF upgrades to reduce systematic uncertainties & add rapidity differential measurements
- Near Future dilepton measurements at STAR
  - New STAR measurements of  $\mu^+\mu^-$  continuum thanks to MTD upgrade
  - New methods of distinguishing correlated charm contributions (MTD and HFT)
  - Investigation of low  $p_{\rm T}$  excess dilepton yield through isobaric collisions

## Thank you





AuAu@200 GeV, 19.6 GeV STAR, PLB 750 (2015) 64

#### Isobaric collisions at RHIC in 2018

Quantities	Systems				
	U+U	Ru+Ru	Zr+Zr		
Centrality	60-80	47-75	47-75		
Multiplicity ( $ \eta  < 0.5$ )	12-52	12-52	12-52		
N <sub>part</sub>	$21 \pm 1$	$21 \pm 1$	$21 \pm 1$		
$B^{2}(fm^{-4})$	$30.8 \pm 0.1$	$30.1 \pm 0.1$	$26.2 \pm 0.1$		
$B^4$ (fm <sup>-8</sup> )	$1984 \pm 4$	$2121 \pm 4$	$1672 \pm 4$		

source: STAR Note 657 RHIC Beam Use Request For Runs 17 and 18 https://drupal.star.bnl.gov/STAR/starnotes/public/sn0657

Table 5.2: Results obtained from the Glauber model calculations [87] for different colliding systems. The estimations of the magnetic fields are done at the time of the collisions (t=0) and at the center of the participant zone. The multiplicity is obtained using two-component model that is tuned to fit Au+Au data.

Physics process	47-75% Zr+Zr (data/cocktail)	47-75% Ru+Ru (data/cocktail)	Difference between Zr+Zr and Ru+Ru
Photonuclear	$14.3 \pm 0.4$	$16.1 \pm 0.4$	$1.8 \pm 0.6 (3.0 \sigma)$
Two-photon	$14.2 \pm 0.4$	$17.4 \pm 0.4$	$3.2 \pm 0.6 (5.3 \sigma)$

Table 5.3: The expected di-electron data over cocktail ratios in the mass region 0.4-0.76  $\text{GeV/c}^2$  for  $p_T < 0.15 \text{ GeV/c}$  with 1.2 billion minimum-bias isobar collisions and the projected differences for the two physics scenarios in Zr+Zr and Ru+Ru collisions.

Physics process	47-75% Zr+Zr (data/cocktail)	47-75% Ru+Ru (data/cocktail)	Differences between Zr+Zr and Ru+Ru
Photonuclear	$17.5 \pm 1.7$	$20.0 \pm 1.7$	$2.5 \pm 2.4 (1.0 \sigma)$
Two-photon	$17.3 \pm 1.7$	$21.8 \pm 1.7$	$4.5 \pm 2.4 (1.9 \sigma)$

Table 5.4: The expected di-electron data over cocktail ratios in the mass region 3.0-3.2 GeV/c<sup>2</sup> for  $p_T < 0.15$  GeV/c with 1.2 billion minimum-bias isobar collisions and the projected differences for the two physics scenarios in Zr+Zr and Ru+Ru collisions.

## Excess yield and Medium Lifetime

- Normalized excess yields in LMR proportional to medium life time (QGP+HG) for Vs<sub>NN</sub>=17.3–200 GeV
  - nearly constant total baryon density
  - emission rates dominated around T<sub>c</sub>
- Yields in U+U@193GeV and Au+Au@200GeV
  - higher yields in central than in lower energies
  - observe increase from peripheral to central

Indications of longer medium lifetime in central UU@193GeV and central AuAu@200GeV



#### First Measurements of Dielectron V<sub>2</sub> STAR, PRC 90 (2014) 64904

#### $\succ$ challenge: isolate v<sub>2</sub> of excess dielectrons

$$v_2^{\text{total}}(m_{ee}) = v_2^{\text{signal}} \left[ \frac{N_S}{N_B + N_S} \right] (m_{ee}) + v_2^{\text{background}} \left[ 1 - \frac{N_S}{N_B + N_S} \right] (m_{ee})$$

cocktail simulations based on published light-hadron v<sub>2</sub> measurements



v<sub>2</sub> from π<sup>0</sup> Dalitz decay consistent with simulations based on published π v<sub>2</sub>



# Dielectron v<sub>2</sub> : proof of principle

#### PRC 90 (2014) 64904

- based on combined Run 10 and 11 data (760M events)
- $p_T$  integrated  $v_2$  of dielectrons in STAR acceptance



# What to expect at RHIC

#### **Opportunities**:

- expect significant increase of partonic source contribution
- Beam Energy Scan provides unique opportunities to
  - systematically study in-medium ρ broadening
  - on-set of QGP thermal radiation





#### Challenges:

- increased particle multiplicities at higher Vs<sub>NN</sub> lead to significant increase in combinatorial backgrounds
- STAR at 200 GeV for  $M_{ee}^{\sim}$  0.5 GeV/ $c^2$ 
  - p+p: S/B~1/10
  - Au+Au: S/B ~1/250

30

# $\mu^+\mu^-$ in Run14 Au+Au @ $\sqrt{s_{NN}}$ = 200 GeV



Data Triggered by dedicated Dimuon Trigger

In 60-80% Au+Au :

 $\rightarrow$  Clear  $\phi$  and J/ $\psi$  peaks

→S/B > ~1/10 (~1/100 to 1/250 in  $e^+e^-$ )

Significantly more data in semicentral and central collisions



 $J/\psi R_{AA}$ 



#### Muon Telescope Detector (MTD)

- |η|< 0.5
- Azimuthal coverage ~ 45%
- Precise timing  $\sigma$ ~100 picoseconds
- Precise spatial resolution ~1 cm

- →Excess is consistent in Au+Au and U+U
- →Excess cannot be explained by hadronic contributions modified by medium





# STAR Dielectron Results in Au+Au & U+U Collisions @ RHIC Top Energies

# Dielectrons in Au+Au @ $\sqrt{s_{NN}}$ =200 GeV



R. Rapp, Phys.Rev. C 63 (2001) 054907 O. Linnyk et al., Phys. Rev. C 85 024910 (2012)

#### Low Mass Range

- Significant enhancement w.r.t. hadronic cocktail  ${\rm w/o}~\rho$ 

#### Intermediate Mass Range

- Dominant contribution from correlated charm decay
- Consistent with cocktail within uncertainties

#### Heavy Flavor Tracker Upgrade:

#### Data from 2014+2016

- Includes Heavy Flavor Tracker upgrade for precise secondary vertex reconstruction
- Will help disentangle contributions in IMR (Charm modification, QGP thermal radiation, etc.)
- Analyses are ongoing

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- Data/Cocktail "enhancement factor" in LMR does not show strong dependence on centrality or  $p_{T}$
- Model calculations (Rapp & PHSD) give a reasonably good description across  $p_T$  and centrality differentials.

#### Daniel Brandenburg | Rice University

#### 10 Expected to have ~20% higher energy density compared to Au+Au @ 200 GeV

- Kikola et al., PRC84 054907 2011
- Expect longer medium lifetime
- Expect higher excess yield in LMR

#### Significant excess yield in $\rho$ -like mass region ( $300 < M_{ee} < 760 \text{ MeV/c}^2$ )

- Data/cocktail = 2.1 ± 0.1 (stat) ± 0.2 (syst) ± 0.3 (cocktail)
- Large contribution (~48%) from charm in this range
  - $\sigma_{cc}$ =797 µb,  $\sigma_{bb}$ =3.7 µb,  $\sigma_{DY}$ =42nb
- Model vs. data shows good agreement

Model simulation:

3/18/17

R. Rapp – Adv. High energy Phys. (2013) 148253 Cocktail simulations: STAR, PRC 92 (2015) 024912







### Dielctrons in Au+Au and U+U



What have we learned?

- Vacuum  $\rho$  disfavored by data
- Both Rapp & PHSD models consistently describes the data
- LMR excess shows little dependence on  $p_T$  or centrality





### Dielctrons in Au+Au and U+U



What have we learned?

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Strong motivation to further explore production dynamics + models' ability to describe data

• measure excess yield in  $\rho$ -like mass region vs.  $\sqrt{s_{NN}}$ 

