The energy of the ²²⁹Th clock transition

Adriana Pálffy Max Planck Institute for Nuclear Physics, Heidelberg





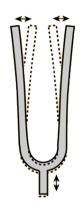


EMMI Featured Talk/GSI Colloquium 19 November 2019



Clocks

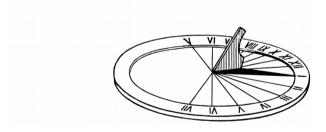
involve a periodic event (oscillation) and a counting device.

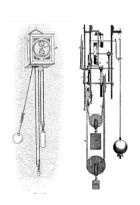












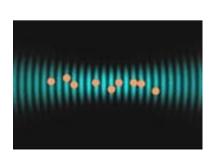


Did you know?

- The earliest clocks to display seconds appeared during the last half of the 16th century.
- The first pendulum clock was invented in 1656 by the Dutch scientist C. Huygens.
- In 1960, the SI second was defined as the fraction of 1/31,556,925.9747 of the tropical year for 1900 January 0th at 12 hours ephemeris this lasted only until 1967.
- The SI second is close to 1/86,400 of a mean solar day in the mid-19th century. In earlier centuries, the mean solar day was shorter than 86,400 SI seconds, and in more recent centuries it is longer than 86,400 seconds.

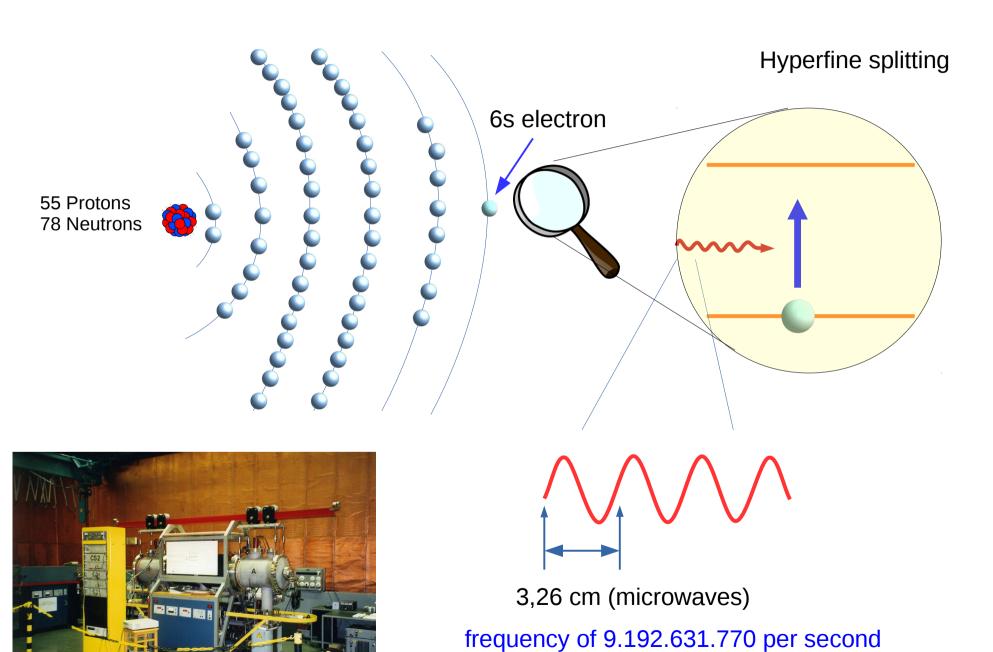






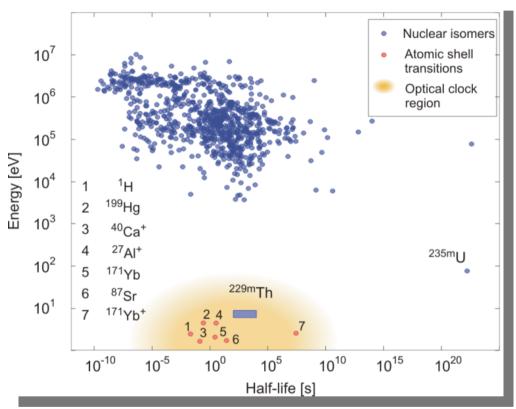


Atomic clock transition in ¹³³**Cs**



A possible nuclear clock

→ lowest excitation energy of all ca. 176,000 presently known excited nuclear states



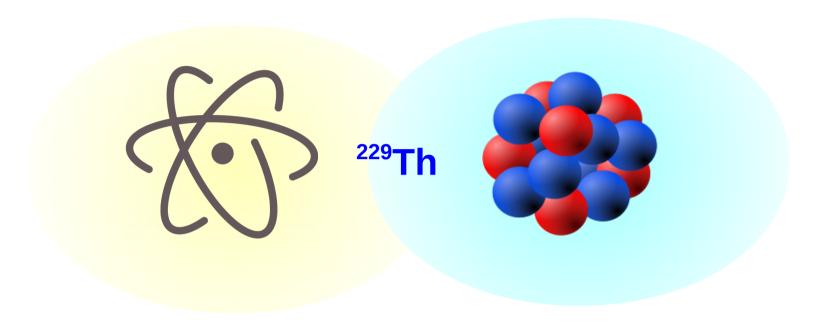
L. v.d. Wense, et al., Nature 533, 47-53 (2016)



E ~ 8.28 eV, λ ~ 150 nm M1 (+E2) transition

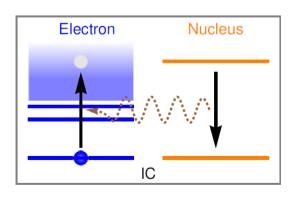
Insensitivity of ^{229m}Th nuclear transition frequency to external perturbations small nuclear moments largely immune to systematic frequency shifts:

Between atomic and nuclear physics



Nuclear decay processes that involve atomic electrons – internal conversion

Optical, UV, VUV lasers open control possibilities – direct excitation, electronic bridge



Possible applications

- Enhanced GPS precision for cm positioning
- lacktriangle Variation of fundamental constants, not only $\,\dotlpha/lpha\,$ but also $\,m_q/\Lambda_{QCD}$
- Tests for Lorentz invariance
- ▶ 3D gravity sensors
- Search for new physics, topological dark matter, etc.

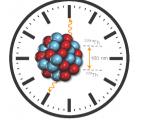
temporal variation in transition energy of ^{229m}Th may provide enhanced sensitivity by (10² – 10⁵) for fine structure constant $\dot{\alpha}/\alpha$ $\frac{\delta\omega}{\omega}=K\frac{\delta\alpha}{\alpha}$ and strong interaction parameter m_q/Λ_{QCD}

$$\dot{\alpha}/\alpha = (-1.6 \pm 2.3) \times 10^{-17} {
m yr}^{-1}$$
 Rosenband et al., Science 319, 1808 (2008)

Outline

I. What we know about ²²⁹Th

Short history
Experimental achievements
Coupling to atomic shell



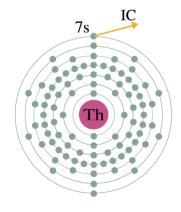
II. Nuclear structure perspective

Nuclear model Excitation probabilities Magnetic moments



III. New energy measurement

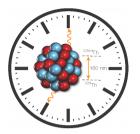
Internal conversion Struggle with excited states Extracted new energy value



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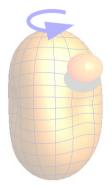
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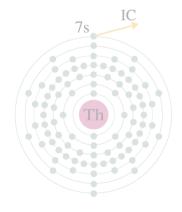
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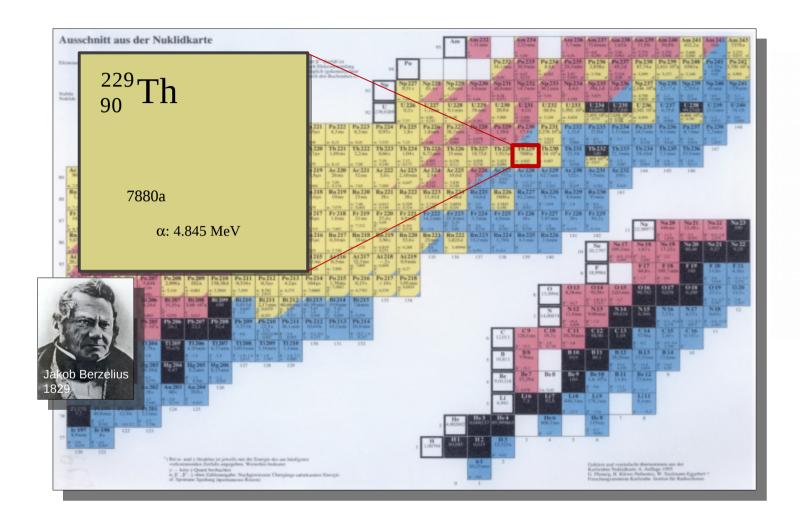


III. New energy measurement

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The ²²⁹Th nucleus





Thor

1828 discovered by Morten Thrane Esmark on the Norwegian island of Løvøya.

Chasing a phantom

For 40 years nuclear physicists seemed to chase a phantom:

presumed low-lying isomer in ²²⁹Th:

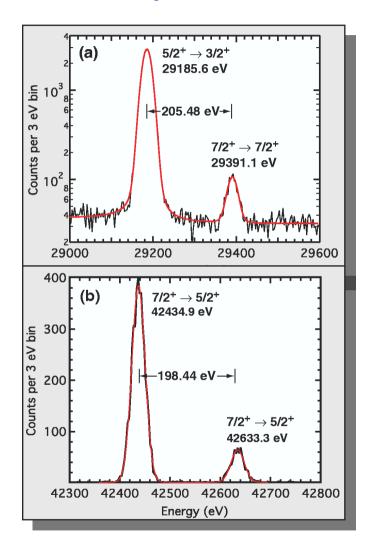
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1976 (Kroger, Reich): < 100 eV 1990 (Reich, Helmer): -1 \pm 4 eV 1994 (Reich, Helmer): 3.5 \pm 1 eV 2005 (Guimaraes-Filho, Helene): 5.5 \pm 1 eV 2007 (Beck et al.): 7.6 \pm 0.5 eV
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energy: 7.8(5) eV (indirect evidence) wavelength: 159(10) nm radiative \tau \approx 10^4 s \rightarrow \Delta E/E \approx 10^{-20} \rightarrow \Delta V \sim 10^{-4} Hz
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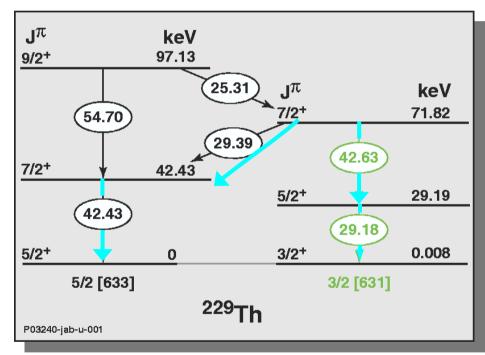
B.R. Beck et al., LLNL-PROC-415170 (2009)

Indirect evidence

data from X-ray micro-calorimeter:



$$E(^{229m}Th) = (7.8 \pm 0.5) eV$$

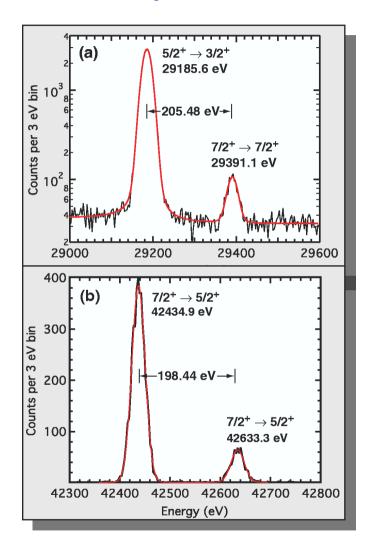


B.R. Beck et al., PRL 98 (2007) 142501 LLNL-PROC-415170 (2009)

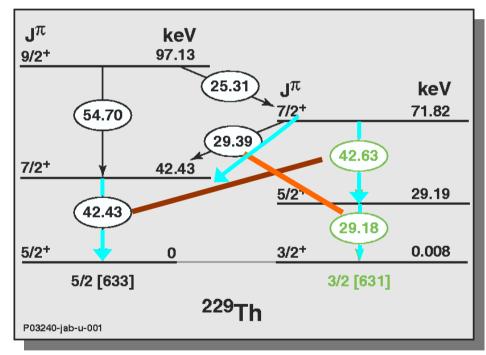
from energy difference of high-energy transitions → low isomer excitation energy

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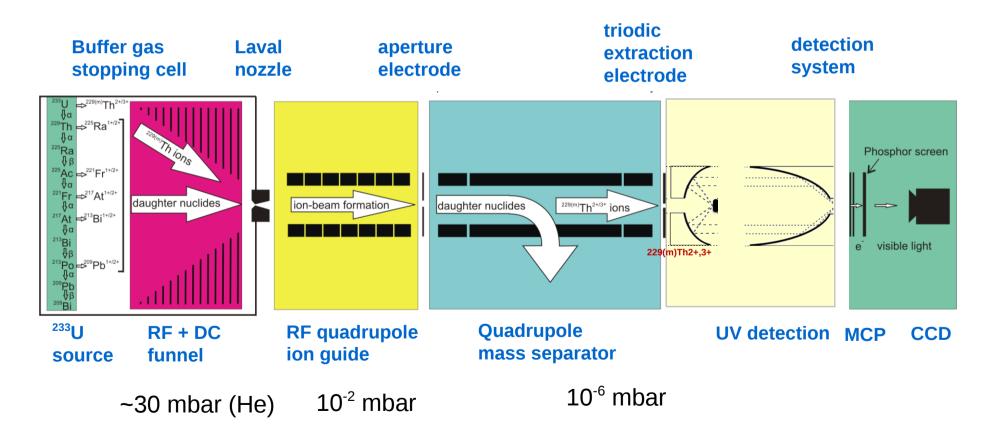
from energy difference of high-energy transitions → low isomer excitation energy

The direct proof

Group of Peter Thirolf @ LMU München

concept:

- populate the isomeric state via 2% decay branch in the lpha decay of 233 U
- spatially decouple ^{229(m)}Th recoils from the ²³³U source
- detect the subsequently occurring isomeric decay



Munich apparatus

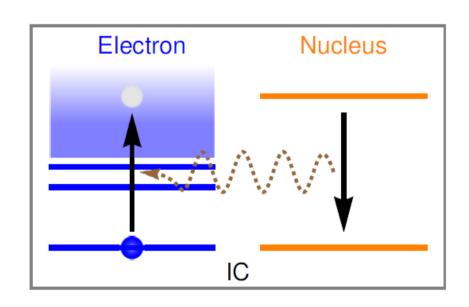


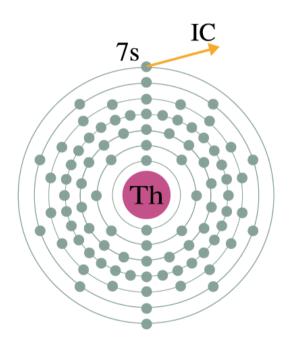
located at Maier-Leibnitz Laboratory, Garching

No VUV photons!

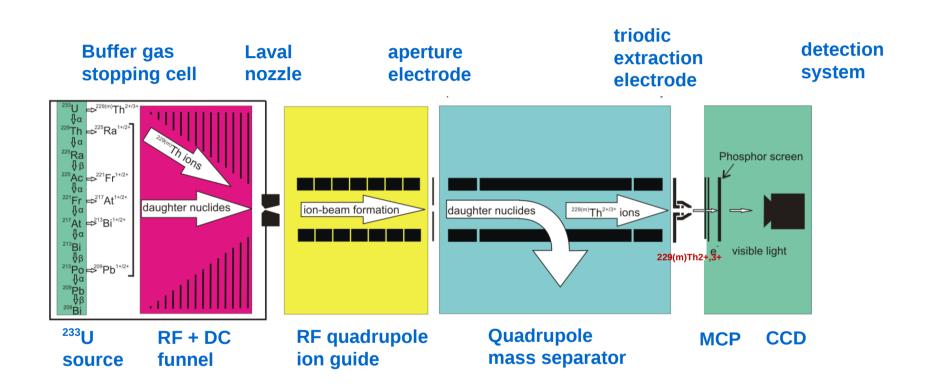
What else could have happened?

For low-energy transition, internal conversion is the preferred decay channel!





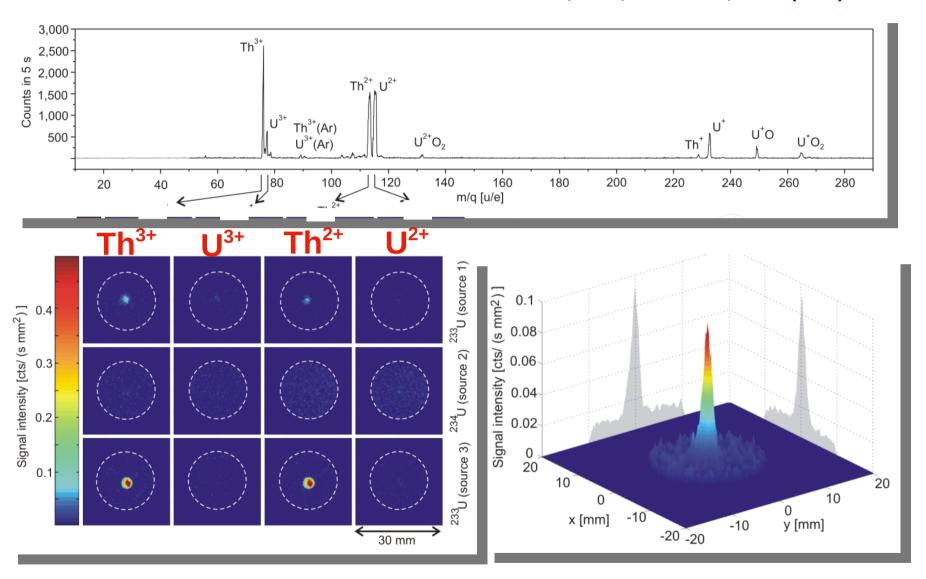
Detection of IC electron



accumulate ^{229(m)}Th ions directly onto MCP surface

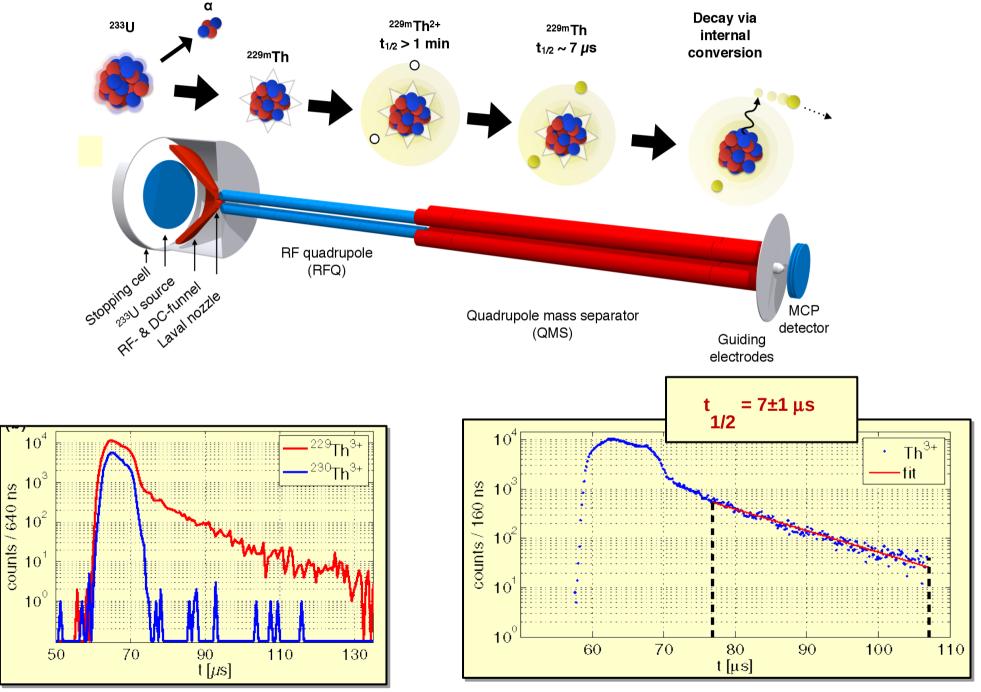
Direct IC signal

L. v.d. Wense, et al., Nature 533, 47-53 (2016)



clear signal from Th³⁺, Th²⁺ no signal from U³⁺, U²⁺

IC lifetime measurement



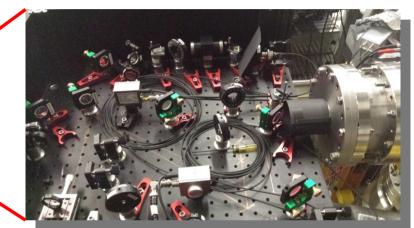
Courtesy P. Thirolf

Collinear laser spectroscopy on ^{229m}Th

PTB Braunschweig, group of Ekkehard Peik

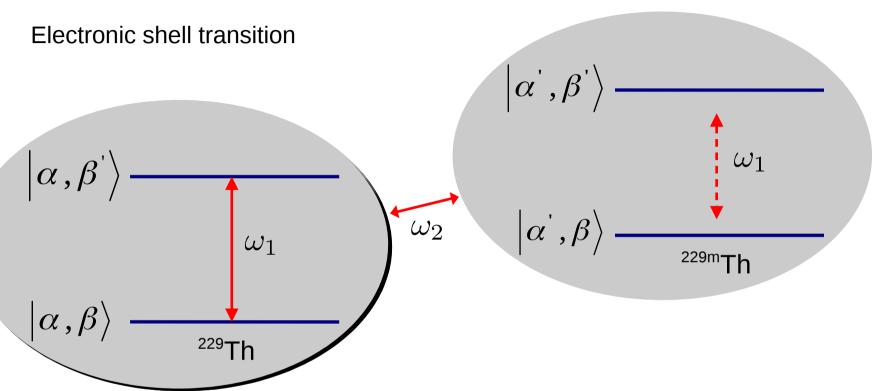
isomer beam (LMU) + laser system (PTB) - resolve hyperfine structure of ^{229m}Th²⁺





- laser excitation of ^{229(m)}Th²⁺ ions behind QMS:
 3 external-cavity diode lasers
 co- and counter-propagating laser beams
- preparatory experiments on ²²⁹Th at PTB Paul trap

Laser spectroscopy on hyperfine structure



 α : nuclear, β : electronic

after nuclear transition

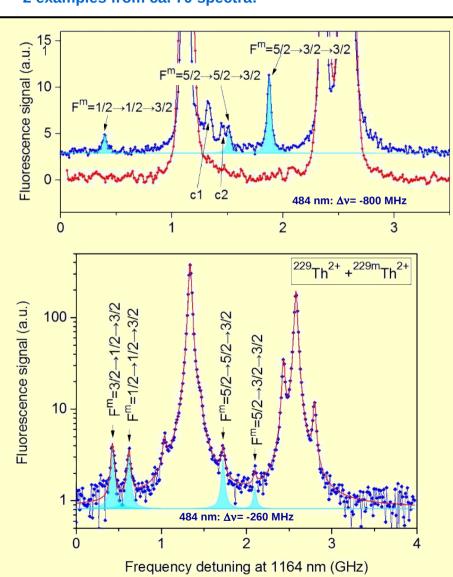
change of nuclear moments, spin change of hyperfine splitting, total angular momenta $\omega_{_1}$ out of resonance (~GHz)

drop in resonance fluorescence

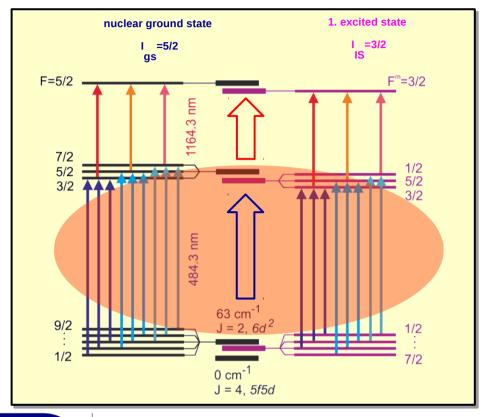
Peik, Tamm, Eur. Phys. Lett. 61 (2003) 181

Hyperfine structure of ^{229m}**Th**

2 examples from ca. 70 spectra:



J. Thielking et al., Nature 556, 321-325 (2018)



ground state: (I=5/2): 9 transitions **isomeric state:** (I=3/2): 8 transitions

Determine the hyperfine constants A and B!

Outline

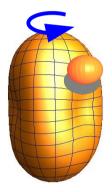
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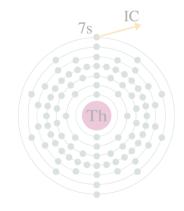
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Nuclear model Excitation probabilities Magnetic moments

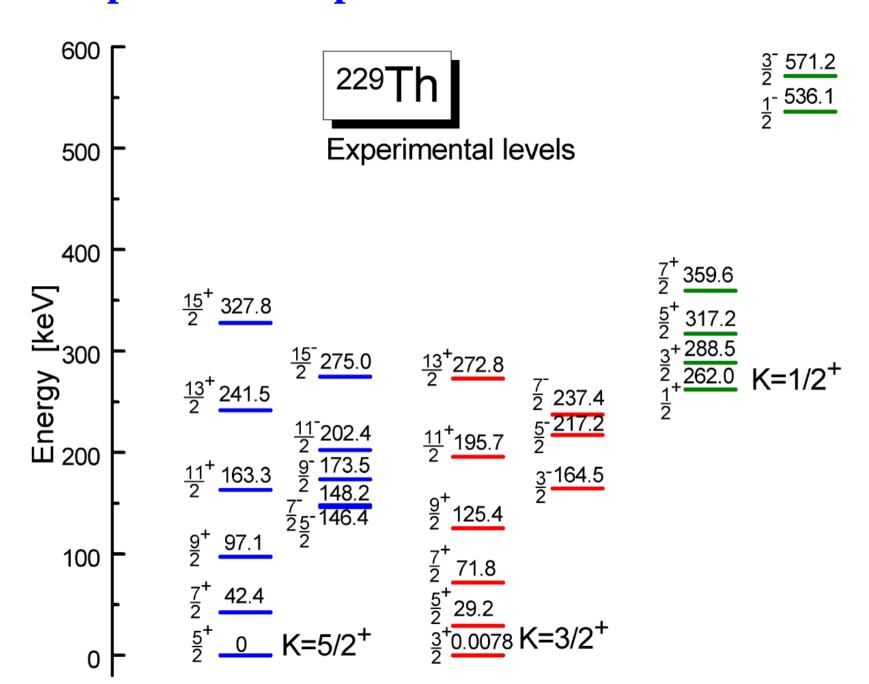


III. New energy measurement

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²²⁹Th experimental spectrum



Nuclear structure model

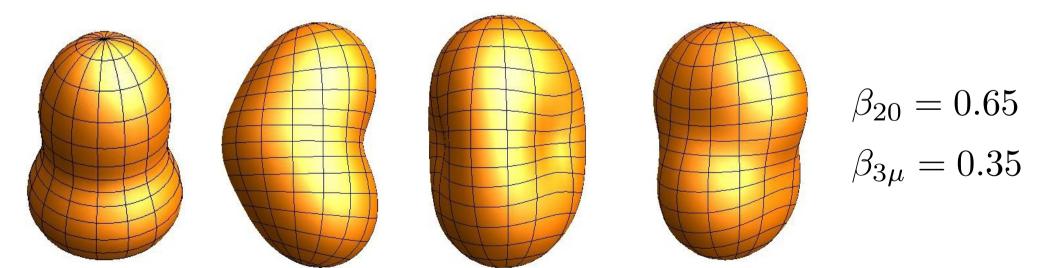
90 protons + 128 neutrons Even – even core

+ one single neutron

Additional single particle

+ Coriolis interaction

Quadrupole – octupole deformation for core



Nuclear structure model

90 protons + 128 neutrons

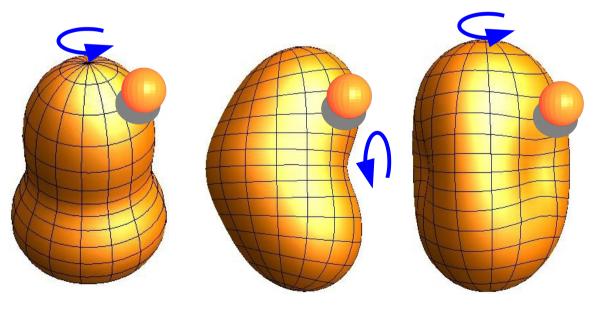
Even – even core

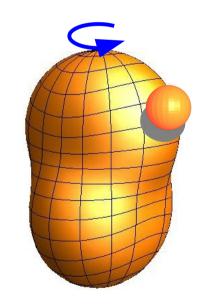
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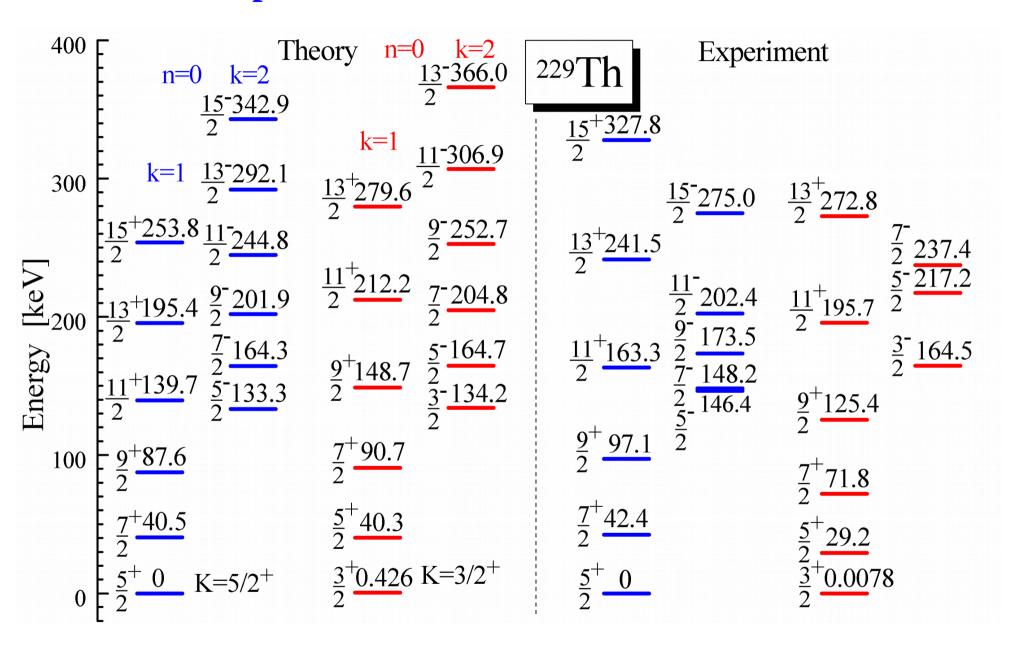




$$\beta_{20} = 0.65$$

$$\beta_{20} = 0.65$$
$$\beta_{3\mu} = 0.35$$

Theoretical predictions



Theoretical B(M1) and B(E2) values

Type/Mult	Transition	Th1[Th2] (W.u.)	Exp (W.u.)
E2	$7/2_{ m yrs}^+ ightarrow 5/2_{ m yrs}^+$	252 [267]	300 (±16)
E2	$9/2_{yrs}^+ \rightarrow 5/2_{yrs}^+$	82 [85]	65 (±7)
E2	$9/2_{yrs}^+ \rightarrow 7/2_{yrs}^+$	213 [224]	170 (\pm 30)
E2	$9/2_{yrs}^{+} o 5/2_{ex1}^{+}$	19.98 [17.37]	$6.2~(\pm 0.8)$
E2	$3/2_{ ext{ex}1}^+ \; o \; 5/2_{ ext{yrs}}^+$	27.04 [23.05]	?
M1	$7/2_{yrs}^+ ightarrow 5/2_{yrs}^+$	0.0093 [0.0085]	$0.0110~(\pm 0.0040)$
M1	$9/2_{yrs}^+ \rightarrow 7/2_{yrs}^+$	0.0178 [0.0157]	$0.0076~(\pm 0.0012)$
M1	$9/2_{yrs}^{+} \to 7/2_{ex1}^{+}$	0.0151 [0.0130]	$0.0117~(\pm 0.0014)$
M1	$3/2_{ ext{ex}1}^{ ext{+}} \; o \; 5/2_{ ext{yrs}}^{ ext{+}}$	0.0076 [0.0061]	?

Th1
$$\rightarrow E(3/2^+) = 0.4263 \text{ keV}$$

Th2 $\rightarrow E(3/2^+) = 0.0078 \text{ keV}$

Smaller than the previously used Theoretical value of 0.048 W.u.

Magnetic moment

Model Hamiltonian

$$H = H_{\text{qo}} + H_{\text{s.p.}} + H_{\text{pair}} + H_{\text{Coriol}}$$

Core plus particle magnetic dipole operator

$$\hat{M}1 = \sqrt{3/(4\pi)}\mu_N \left[g_R(\hat{I} - \hat{j}) + g_s \,\hat{s} + g_l \,\hat{l} \right]$$

Magnetic moment is the matrix element

$$\mu = \sqrt{\frac{4\pi}{3}} \langle \widetilde{\Psi}_{IIK_b} | \hat{M} 1_0 | \widetilde{\Psi}_{IIK_b} \rangle$$

Magnetic moment

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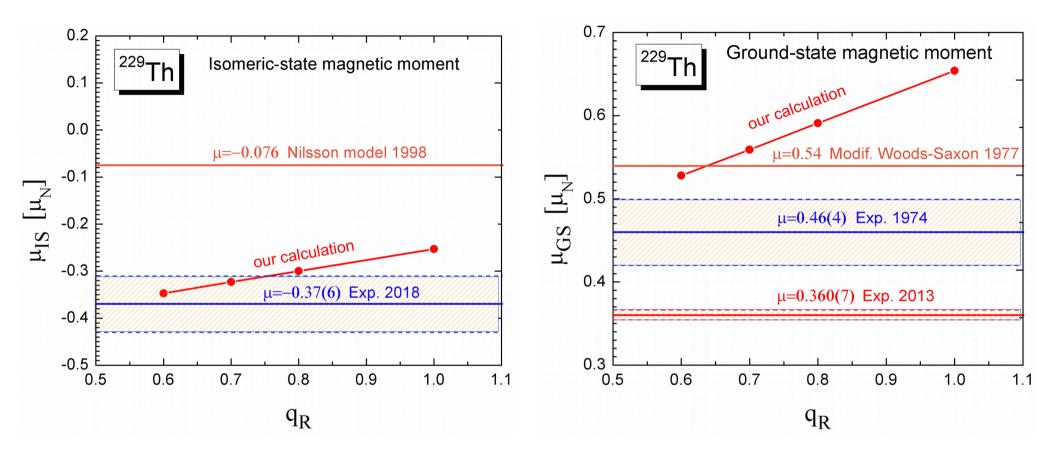
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Magnetic moment is the matrix element

$$\mu = \sqrt{\frac{4\pi}{3}} \langle \widetilde{\Psi}_{IIK_b} | \hat{M} 1_0 | \widetilde{\Psi}_{IIK_b} \rangle$$
 phenomenological Z/A introduce quenching factor for deformed nuclei (20%-30% less)

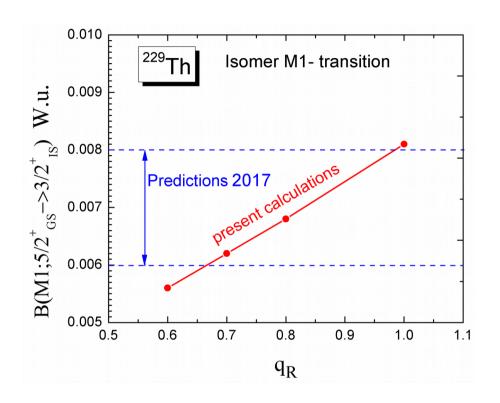
collective gyromagnetic factor

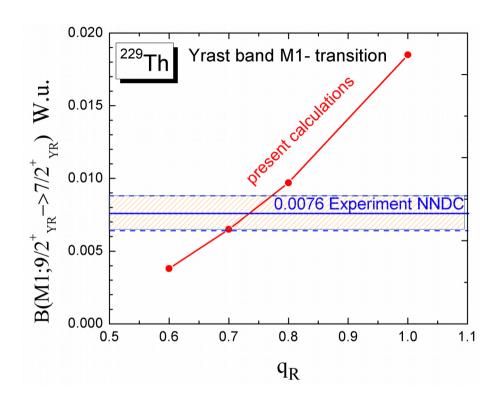
Results



Model parameters from our previous work on B(M1) chosen to reproduce the low-lying energy spectrum

For B(M1)





It appears likely that B(M1) is even lower than previously estimated!

Summary

- Collective CQOM + microscopic DSM+BCS + Coriolis interaction
- Quenching of collective gyromagnetic factor
- Predictions for magnetic moments $\mu_{GS} = 0.528 \mu_N \qquad \text{too large}$ $\mu_{IS} = -0.347 \mu_N \quad \text{surprisingly good}$
- B(M1): will very likely be smaller than previously predicted, 0.0056 W.u.
- New search through model parameters planed.
 Interestingly, Coriolis mixing factor affects only ground state magnetic moment!

Outline

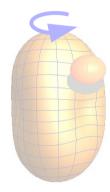
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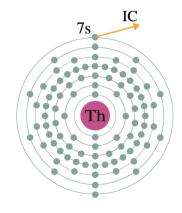
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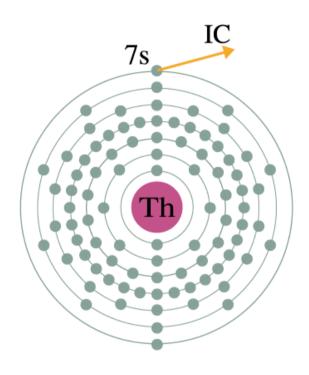


III. New energy measurement

Internal conversion Struggle with excited states Extracted new energy value



Internal conversion



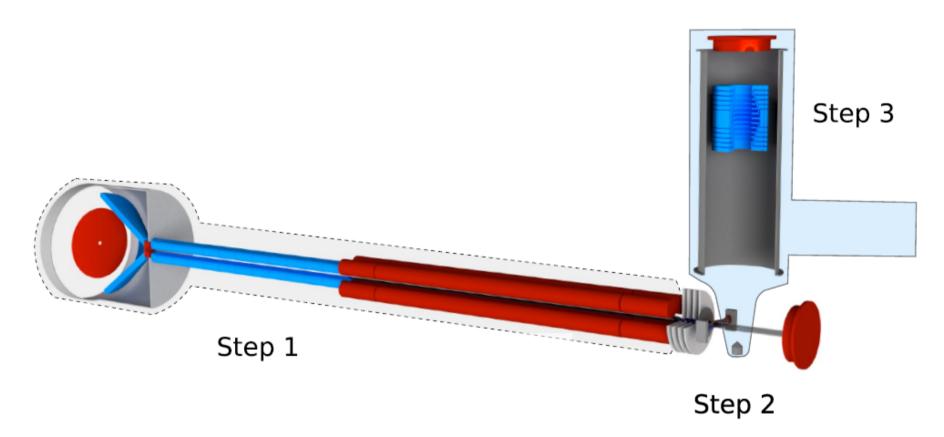
IC coefficient $\alpha = \frac{\text{IC rate}}{\gamma\text{-decay rate}}$

For M1-transition at $E \simeq 10 \text{ eV}$ γ -decay rate $\propto E^3$ is small $\Rightarrow \alpha$ is large

In neutral atom of $^{229}{\rm Th}~\alpha \simeq 10^9$ F. F. Karpeshin *et al.* Phys. Rev. C **76**, 054313 (2007)

lon charge	0	1+	2+	3+	4+
Ion. threshold (eV)	6.3	12.1	20.0	28.7	58

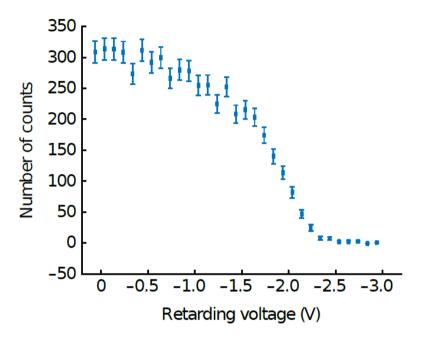
Experimental setup



LMU Munich
Group of Peter Thirolf

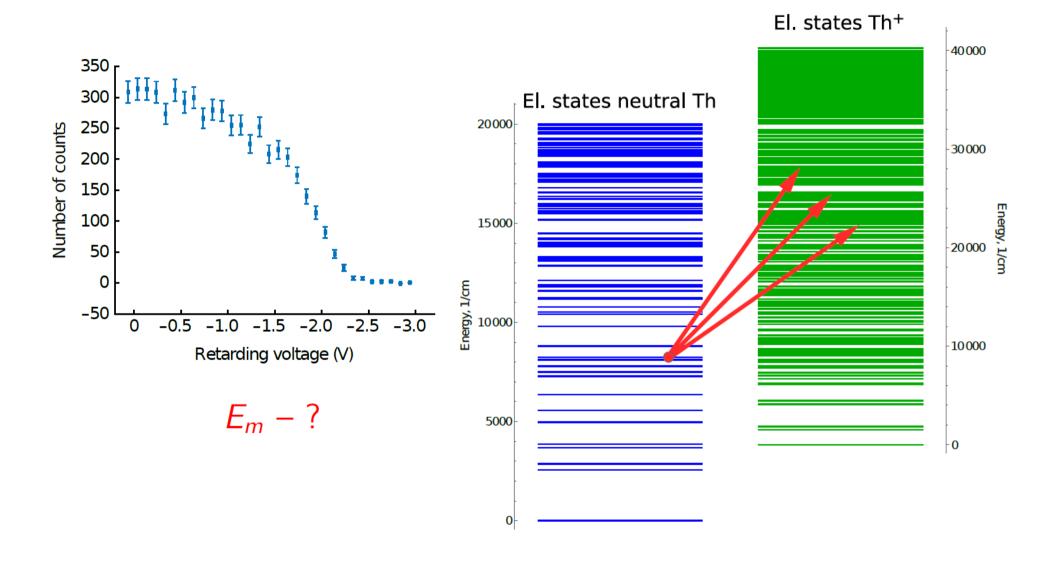
Experimental setup MCP detector Electric Solenoid retarding-field unit coil Magnetic-bottle-type retarding-field spectrometer Charged particles Internal conversion electrons Bending electrode Neutral atoms Focusing electrodes Permanent Graphene magnet Neutralization

Spectrum

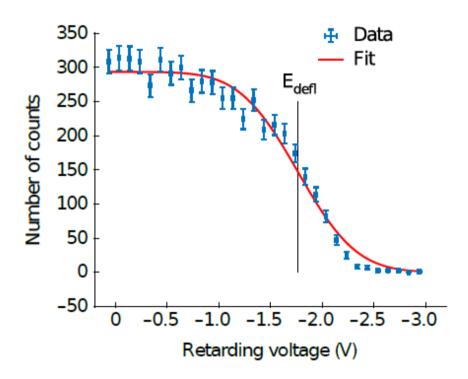


 E_m – ?

Spectrum

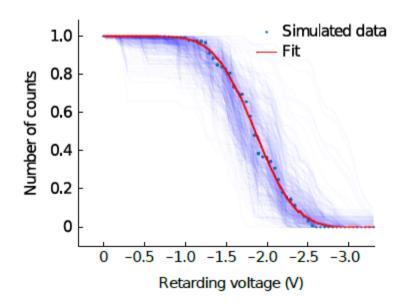


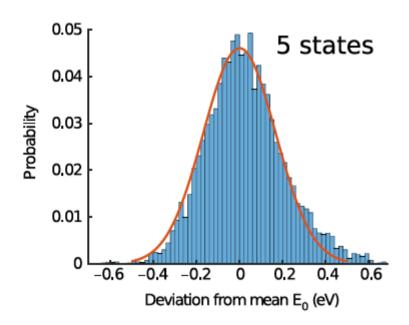
The isomer energy



$$f(U) = a \left\{ 1 - \operatorname{erf}\left[rac{U - E_{\operatorname{defl}}}{b}
ight]
ight\}$$
 $E_{\operatorname{defl}} = 1.77 \pm 0.03 \; \operatorname{eV}$
 $E_{\operatorname{m}} = E_0 + E_{\operatorname{defl}}$

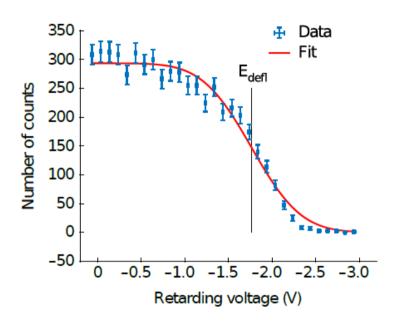
The isomer energy





Number of excited states (n)	E_0	Width of the distribution σ		
2	$6.56{\pm}0.02$	0.22		
4	6.517 ± 0.003	0.18		
5 *	6.506 ± 0.001	0.16		
10	6.504 ± 0.001	0.12		
20	6.492 ± 0.001	0.09		
30	6.494 ± 0.001	0.06		
40	6.493 ± 0.001	0.05		

The isomer energy

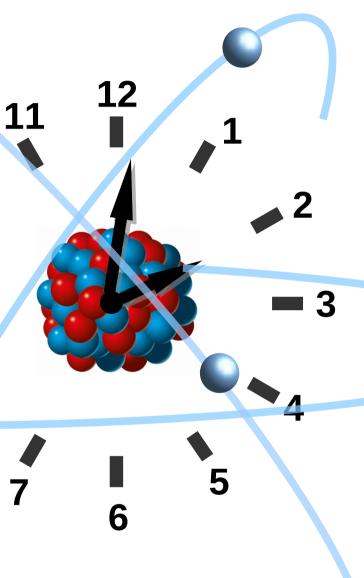


IC calculations to extract the energy value from the measured spectrum

$$E_m = 8.28 \pm 0.17 \,\mathrm{eV}$$



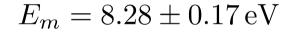
Conclusions



²²⁹Th could provide a unique chance to develop a nuclear frequency standard, stable, compact, with increased accuracy.

Appealing applications:

GPS variation of fundamental constants tests for new physics



Optical (VUV) control of the transition requires more time - "needle in the haystack" + laser development

"but we know much better how this needle should look like!"

Peter Thirolf



Thanks

Atomic and nuclear quantum dynamics

Brenden Nickerson Pavlo Bilous Wen-Te Liao Nikolay Minkov



Jose Crespo

Peter Thirolf Lars von der Wense Benedict Seifere



Thorsten Schumm



