

# *Direct Capture and the $r$ -process*

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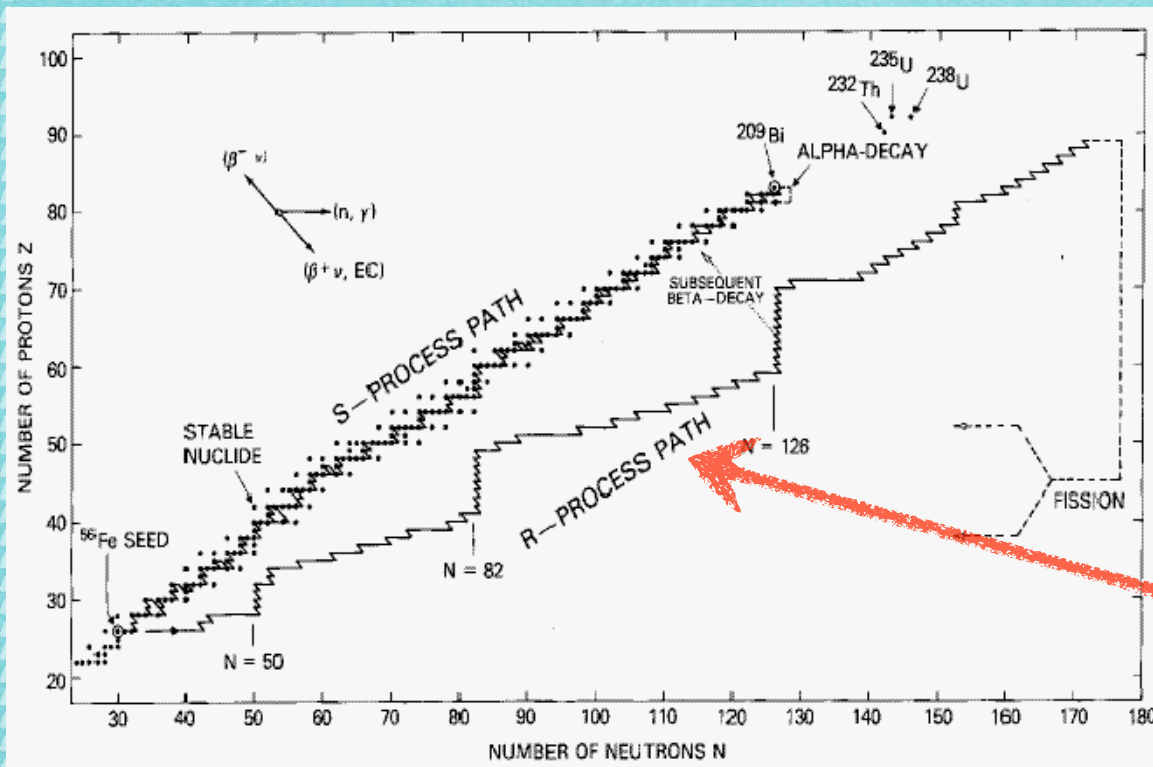
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**2nd EMMI-EFES WS at RIKEN  
June 16-18 (2010)**



# r-process nucleosynthesis



Astrophysical site(s) for the r-process is still unknown.

r-process path is not always in  $(n, \gamma)$ - $(\gamma, n)$  equilibrium.

beta-decay,  $(n, \gamma)$  rates are also important!



# Direct capture

Neutron-capture process

---> compound process + direct capture process

Mathews et al. (1983)

- Importance of direct capture on r-process  
(HF method is invalid for nuclei on r-process path)
- Compare the reaction rates of Compound process and direct capture of Cd

Goriely (1997)

- Calculate direct capture reactions for all nuclei
- discuss the role of direct capture with canonical model.

→ how direct capture works in dynamical calculations?

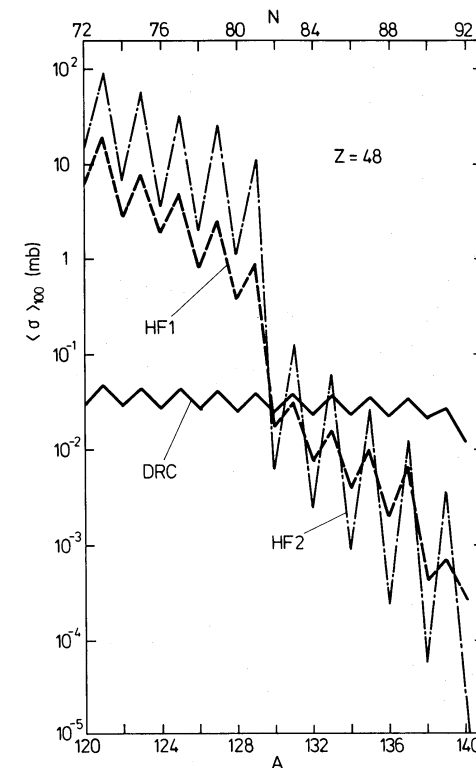


FIG. 1.—A comparison of Maxwellian averaged neutron capture cross sections at  $kT = 100$  keV for Cd isotopes through the  $r$ -process path at  $A = 130$ . DRC indicates the direct radiative capture component. HF1 is the statistical model calculation of Thielemann (1980). HF2 is the analytical statistical model calculation of Woosley *et al.* (1975). Here it can be seen that, beyond the neutron closed shell at  $N = 82$ , the direct radiative capture component dominates the neutron capture rate.



# r-process calculation

## Dynamical full network code

based on Meyer et al. 1994,  
Modified by Terasawa & Orito(2001), Otsuki(2003)

$$\dot{Y}_p = \sum_{r,s} \frac{c_p}{c_r!c_s!} Y_r^{c_r} Y_s^{c_s} R_{rs} - \sum_q \frac{c_p}{c_p!c_q!} Y_p^{c_p} Y_q^{c_q} R_{pq}$$

~7000 nuclide, ~10000 reactions for each time step

MASS...FRDM+AME03 (Möller et al. 1995)

(n, $\gamma$ )...based on FRDM (Rauscher et al. 2000, Panov et al. 2009)

$\beta$ -decay...based on FRDM (Möller et al. 2002)

charged current reaction...NACRE+Reaclib

Physical condition

Spherical steady state flow



# Direct capture reaction rates

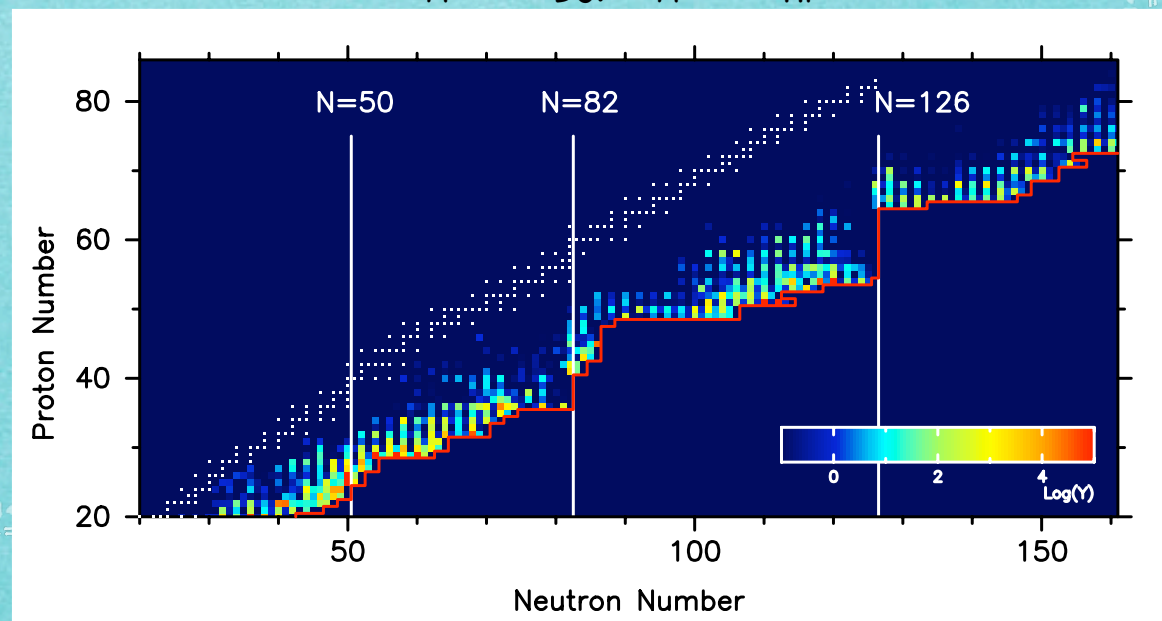
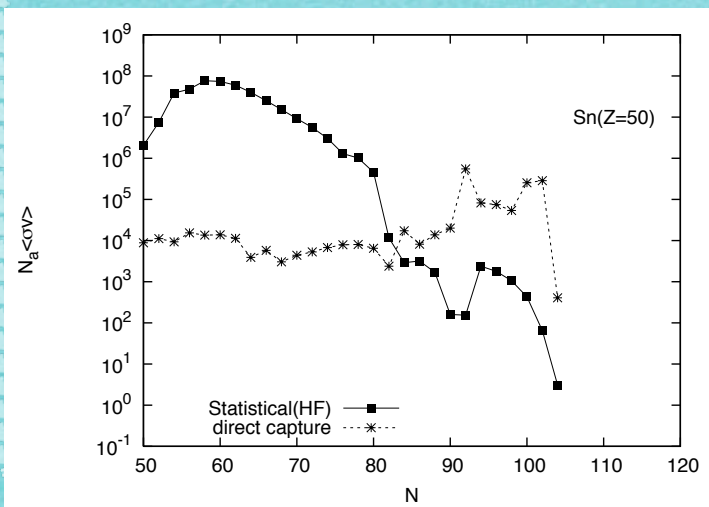
semi-analytic model (Typel et al. 2005)

- define a pattern of single-particle level using experimental data and scale it for each nuclei.
- considering deformation
- deformation and Sn from FRDM

Comparison between reaction rates of compound process(HF method, Rauscher et al. 2000) and direct capture rates(this work)

Sn(Z=50)

$$N_A \langle \sigma v \rangle_{DC} / N_A \langle \sigma v \rangle_{HF}$$

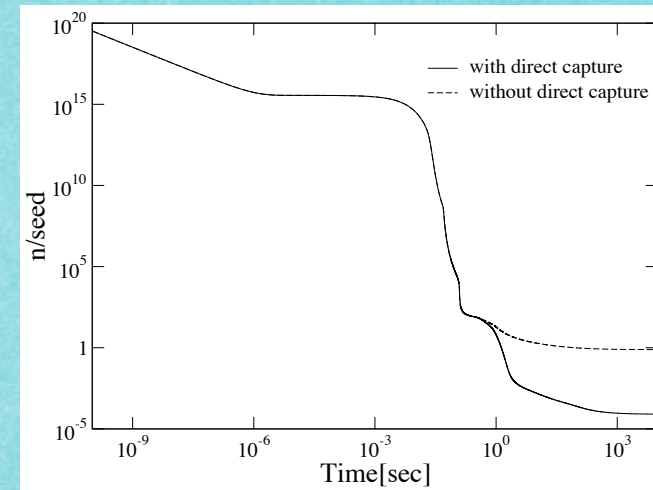




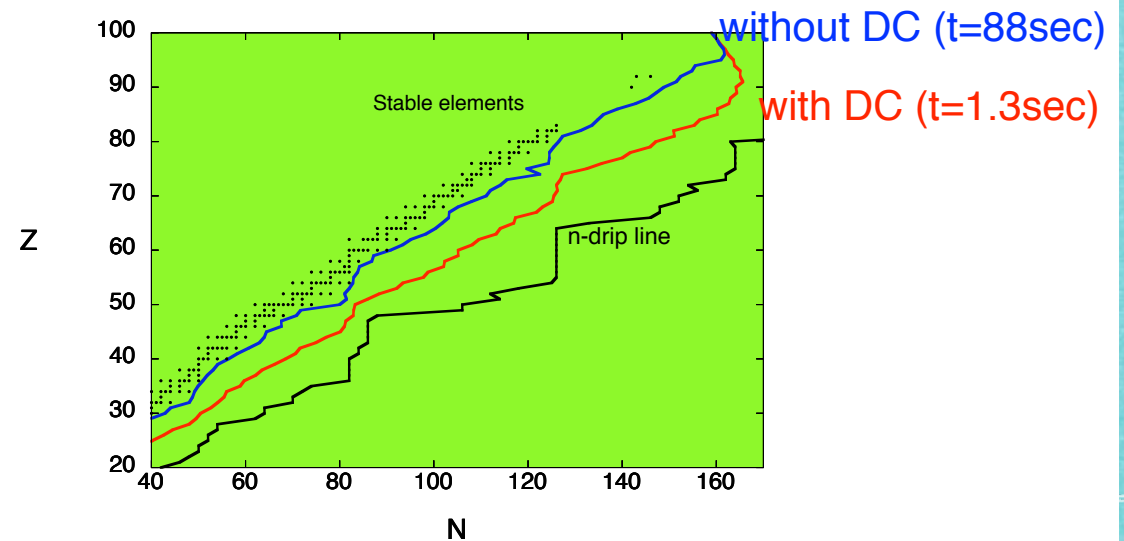
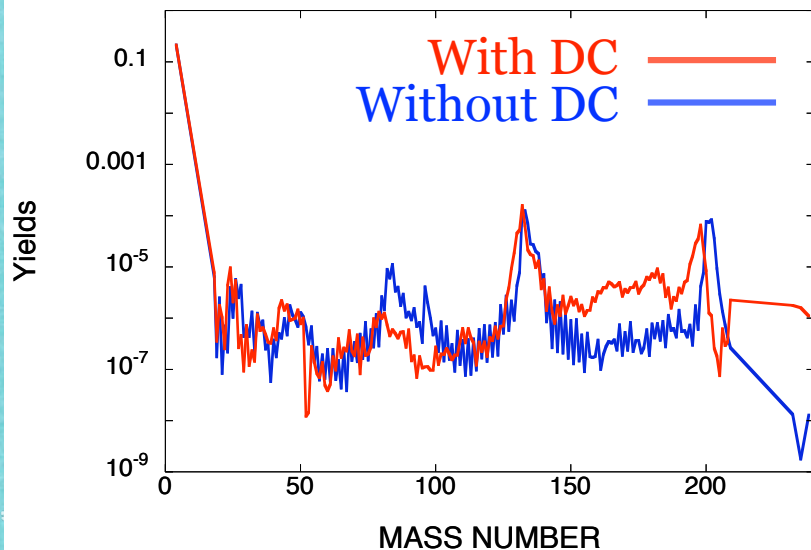
# Effect of direct capture reaction

the direct capture furthers r-process and make freeze out earlier.

---> Change final abundance



Wind solution with  $S=250$



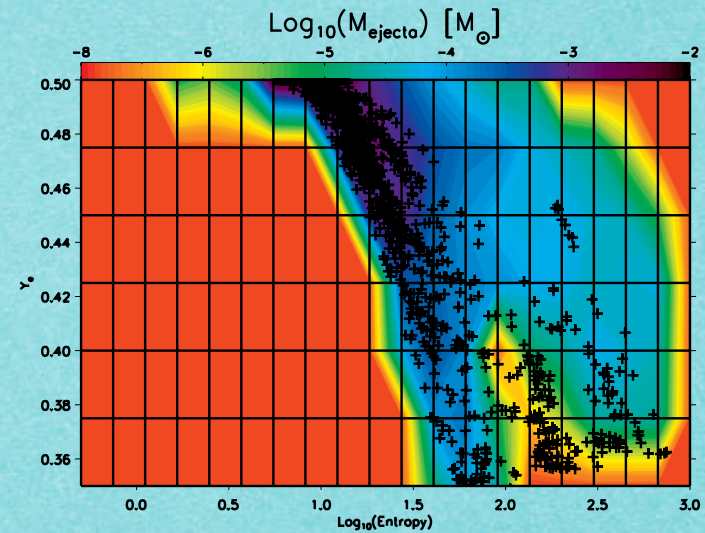
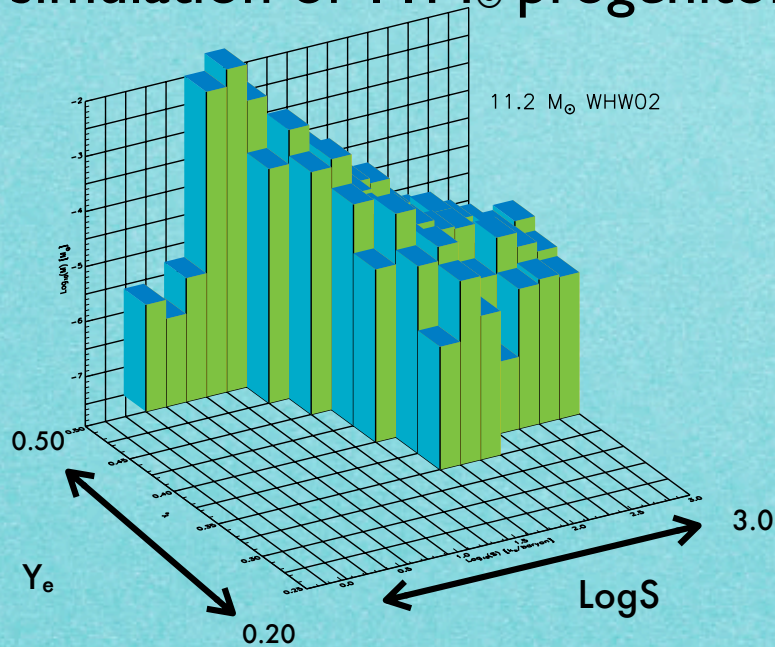


# Direct capture and r-process in supernova

accretion cause core oscillation

Burrows et al. (2006)

simulation of  $11 M_{\odot}$  progenitor



	NDW	Burrows
S	<500	$9 \sim 10^5$
$Y_e$	$\sim 0.45$	$0.33 \sim 0.49$

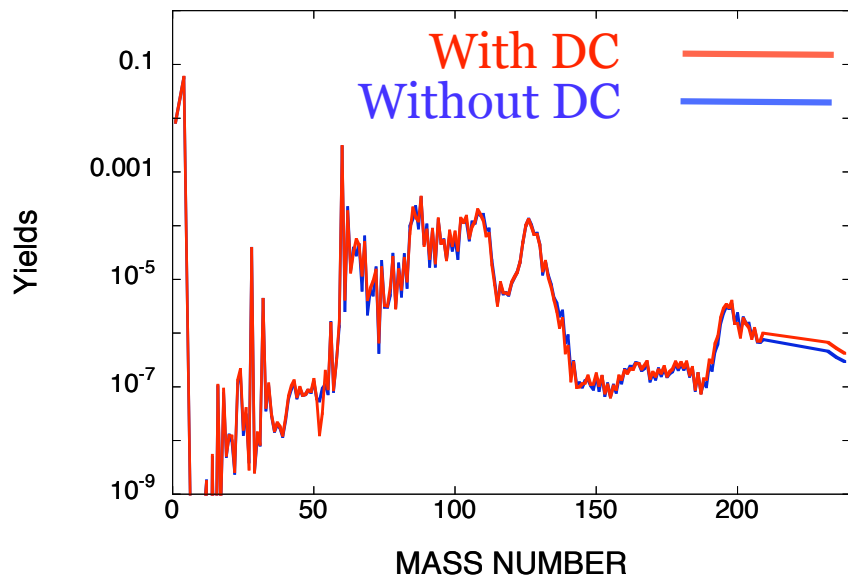
S,  $V_{\text{rad}}$ ,  $Y_e$  of 9086 points at  $r=1000\text{km}$   
 assume constant  $V_{\text{rad}}$  &  $S, Y_e(1000\text{km})=Y_{ei}$

r-process calculation of 366 representative

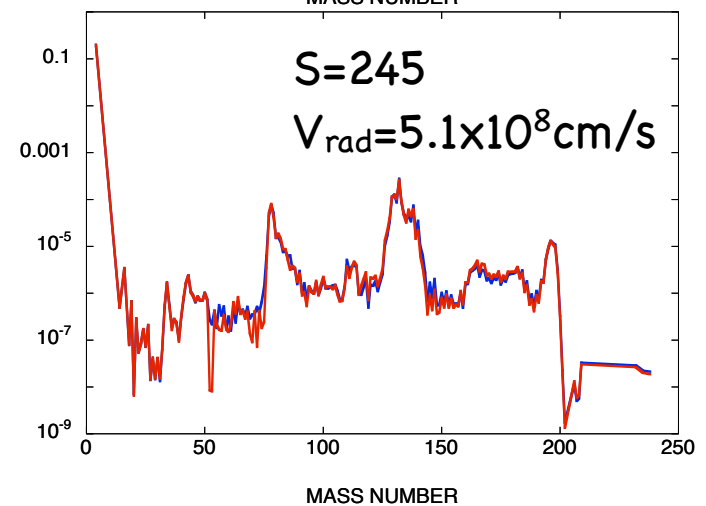
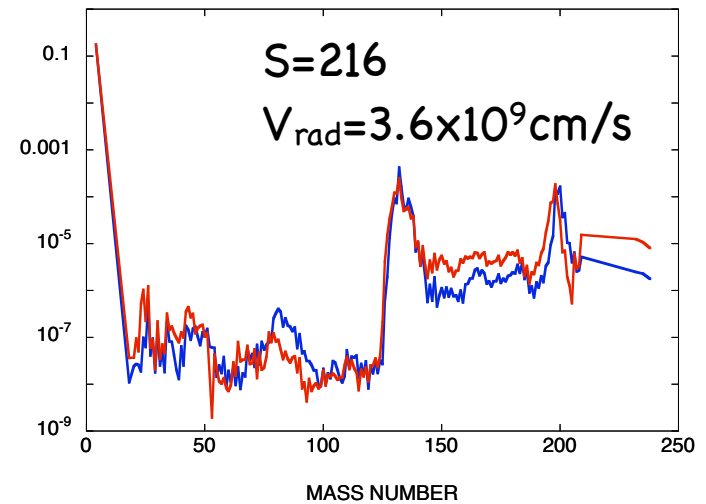


# Direct capture and r-process in supernova

Total yield with and without direct capture



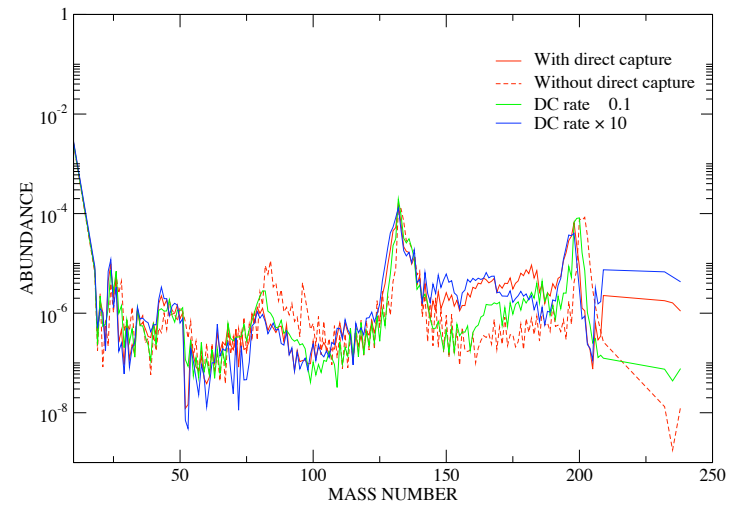
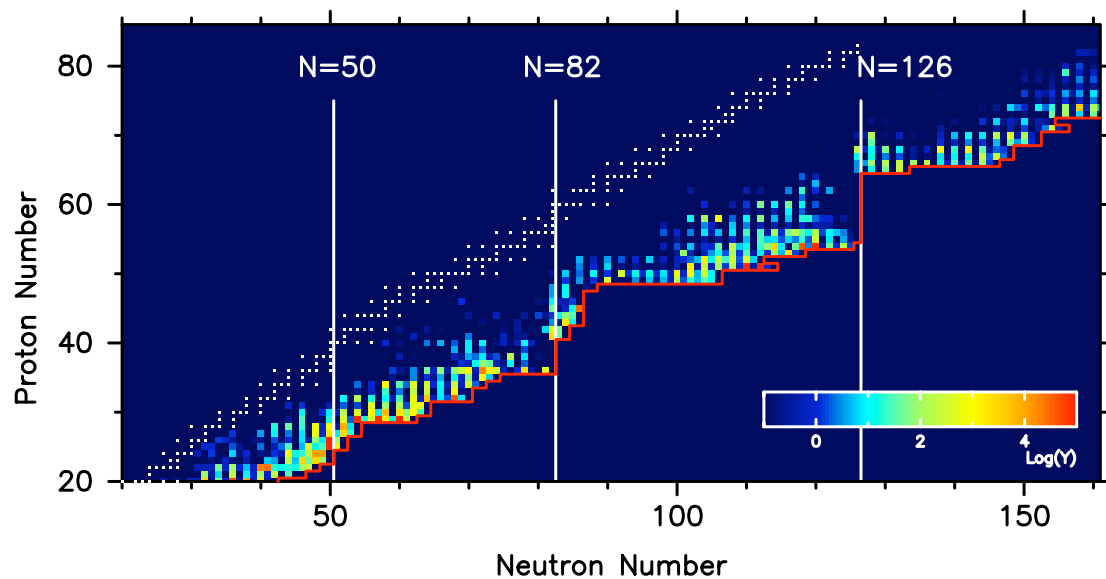
Direct capture increases actinide slightly,  
but did not change the position of the 3rd peak





# Direct capture reaction rates

In our model, direct capture of  $N \sim 50$  nuclide are significant



An order of magnitude difference changes final yield significantly.



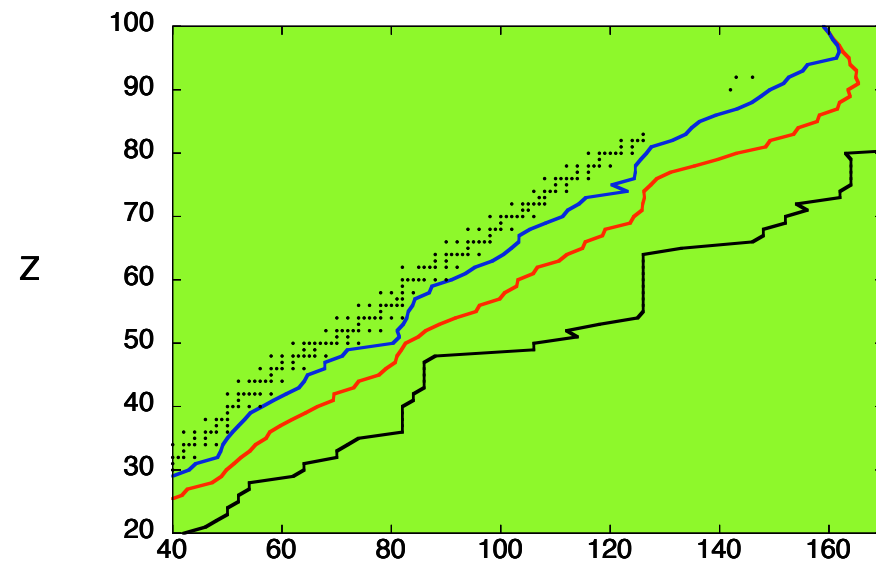
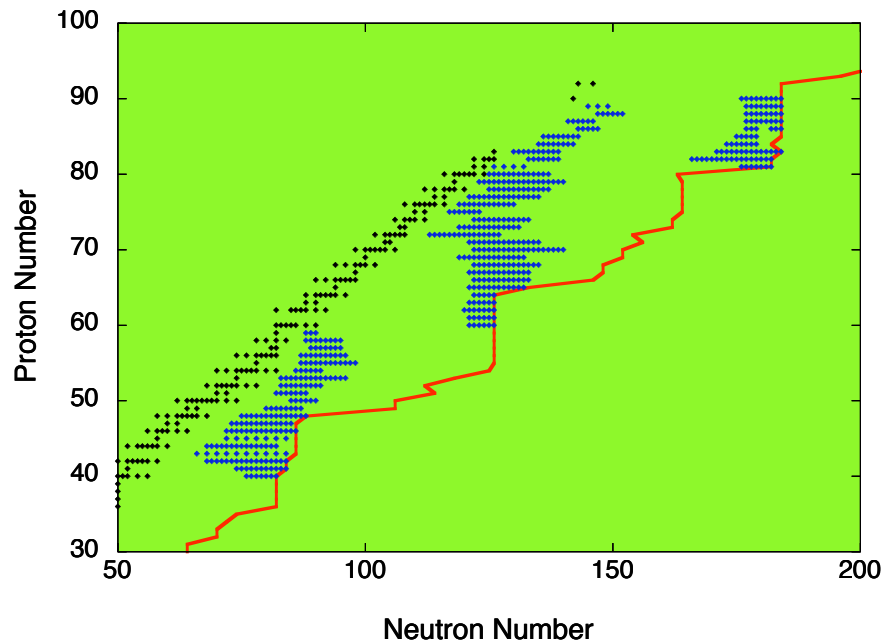
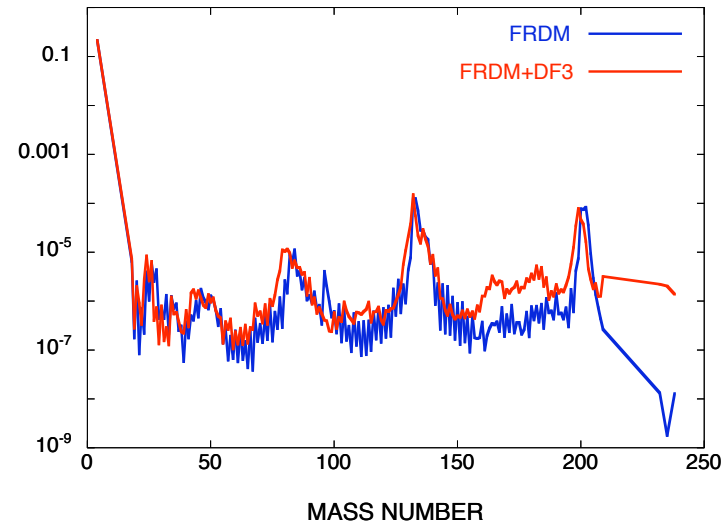
# Effect of $\beta$ decay

Beta-decay rates By I.Borzov

DF theory + CQRPA

$\beta_{1/2} \sim 1/2$  compare to FRDM

Short  $\beta$  half-lives also make freeze-out earlier.





# Summary

- ▶ We have studied the role of direct neutron capture reaction in r-process nucleosynthesis
- ▶ The direct capture furthers r-process and make freeze-out earlier. It changes the final yield drastically.
- ▶ Short  $\beta$  half lives also make freeze-out earlier.
- ▶ neutron-capture reaction rates and beta-decay are important for the r-process.
- ▶ Experimental measurement of neutron-capture reaction rates around  $N=50$  nuclei are strongly desired.