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Thermal Evolution of Hybrid Stars modeled with an SU(3) non-linear Sigma Model

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The thermal evolution of hybrid stars is investigated. The structure and composition of these objects are obtained by means of an extended hadronic and quark SU(3) non-linear sigma model. Within this model the degrees of freedom of the system change naturally from quarks to hadrons, allowing a more natural description of hybrid stars. In this work we will focus on the thermal evolution of these objects. Furthermore special attention will be given to the possible effects that spin-down may have on the cooling of these stars.

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