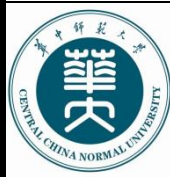


# QCD matter physics at the future Facility for Antiproton and Ion Research

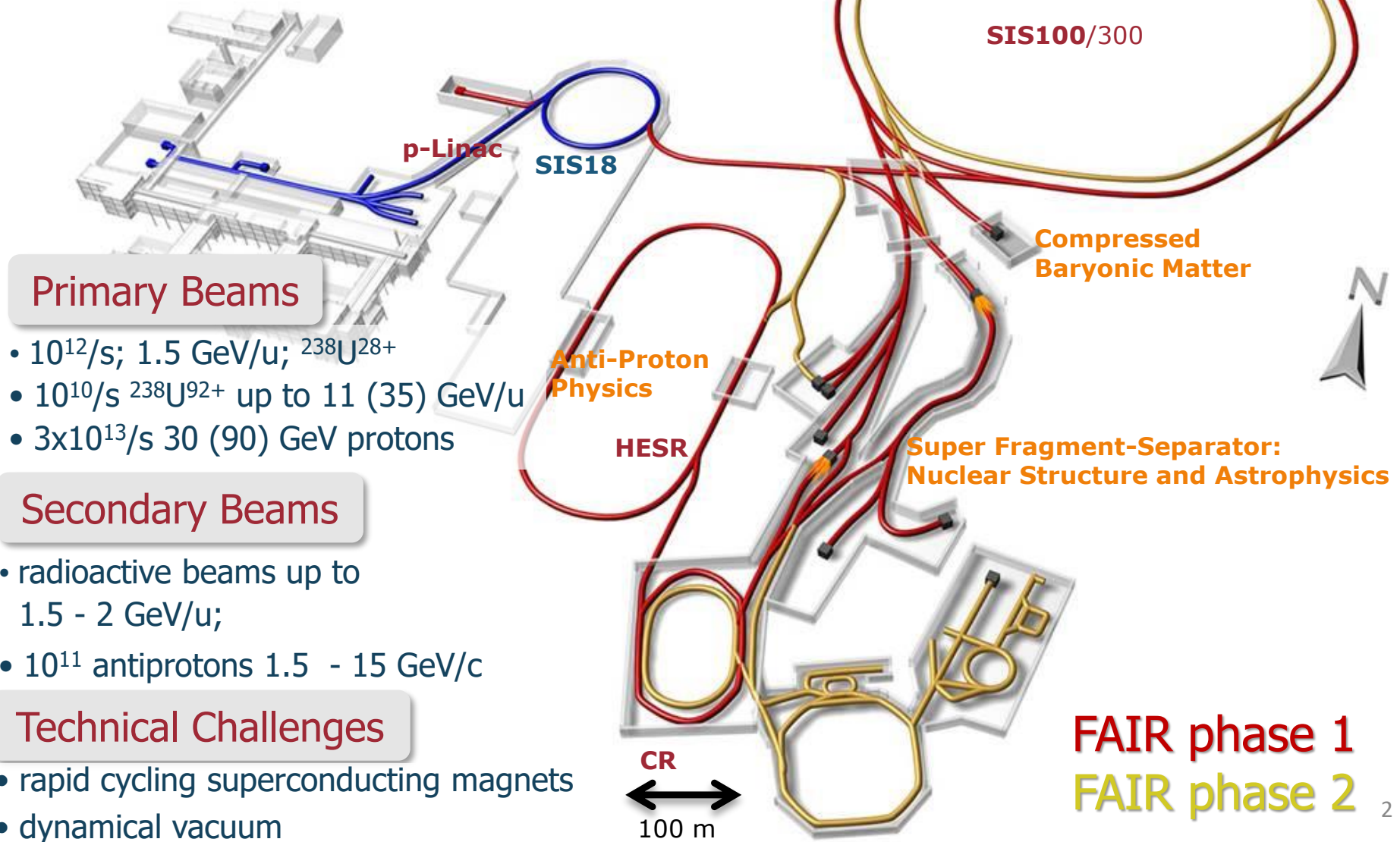
## Outline:

- The Facility on Antiproton and Ion Research
- Exploring cosmic matter in the laboratory:
  - the high-density nuclear matter equation-of-state
  - the QCD phase diagram
- The Compressed Baryonic Matter (CBM) experiment

Peter Senger



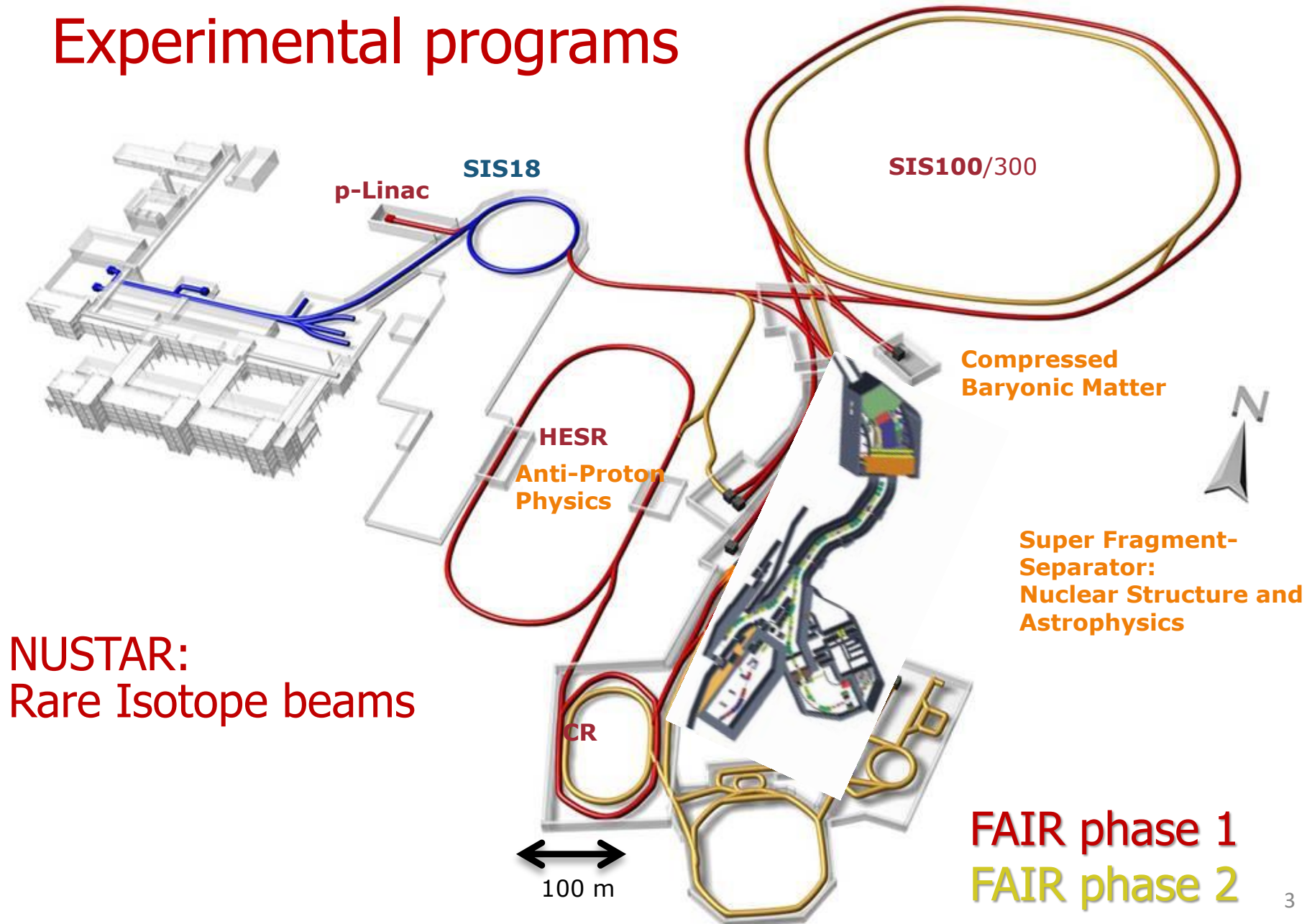
# Facility for Antiproton & Ion Research



**FAIR phase 1**  
**FAIR phase 2**

# Facility for Antiproton & Ion Research

## Experimental programs



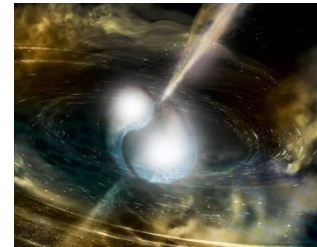
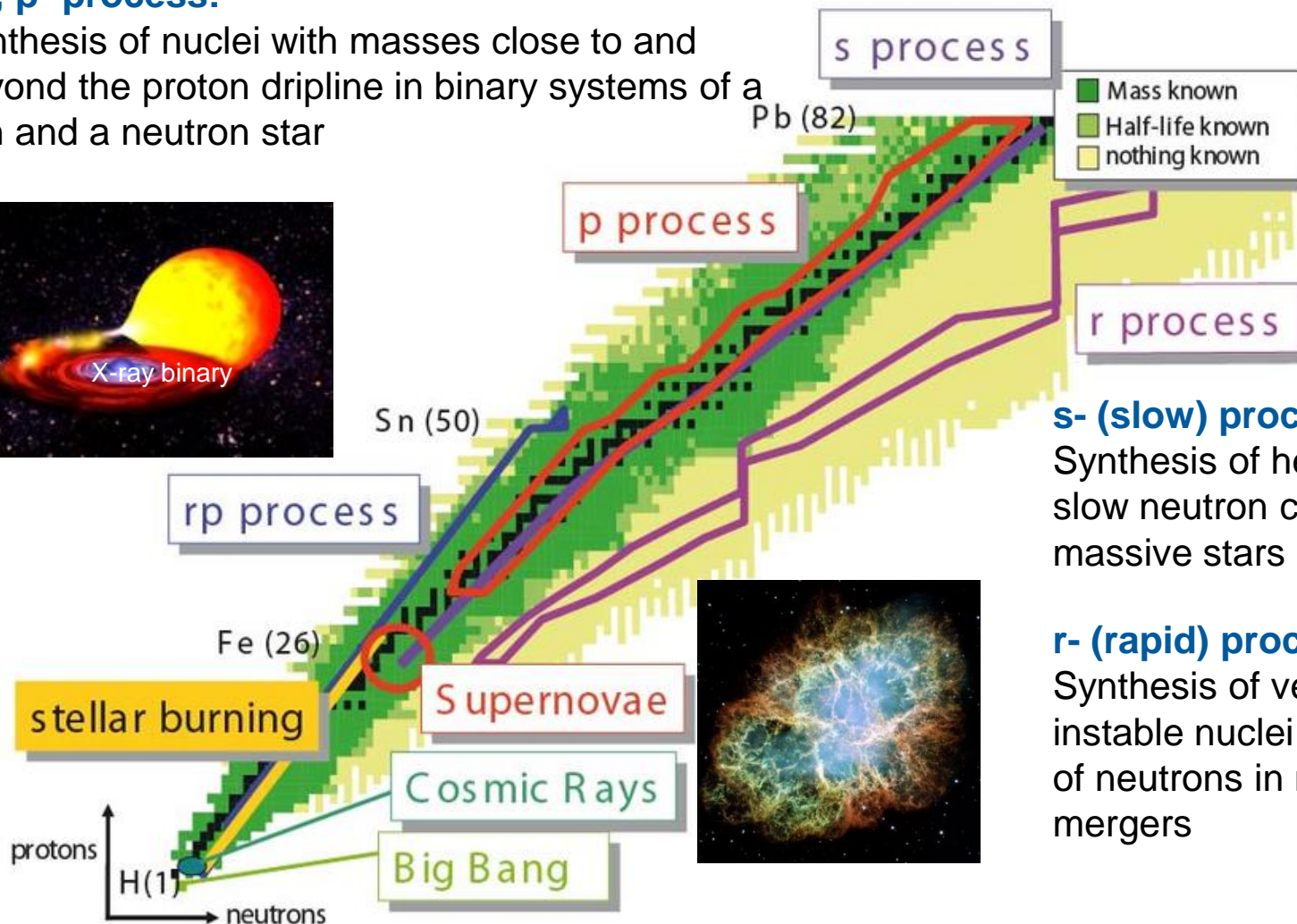


# Nuclear astrophysics: The origin of elements

Measurements in the laboratory: Mass, lifetime, decay channels, structure of very rare instable (neutron or proton rich) nuclei

## rp-, p- process:

Synthesis of nuclei with masses close to and beyond the proton dripline in binary systems of a sun and a neutron star

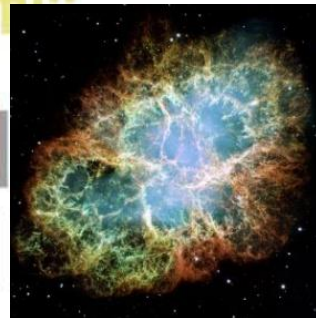


## s- (slow) process:

Synthesis of heavy nuclei via slow neutron capture in very massive stars

## r- (rapid) process:

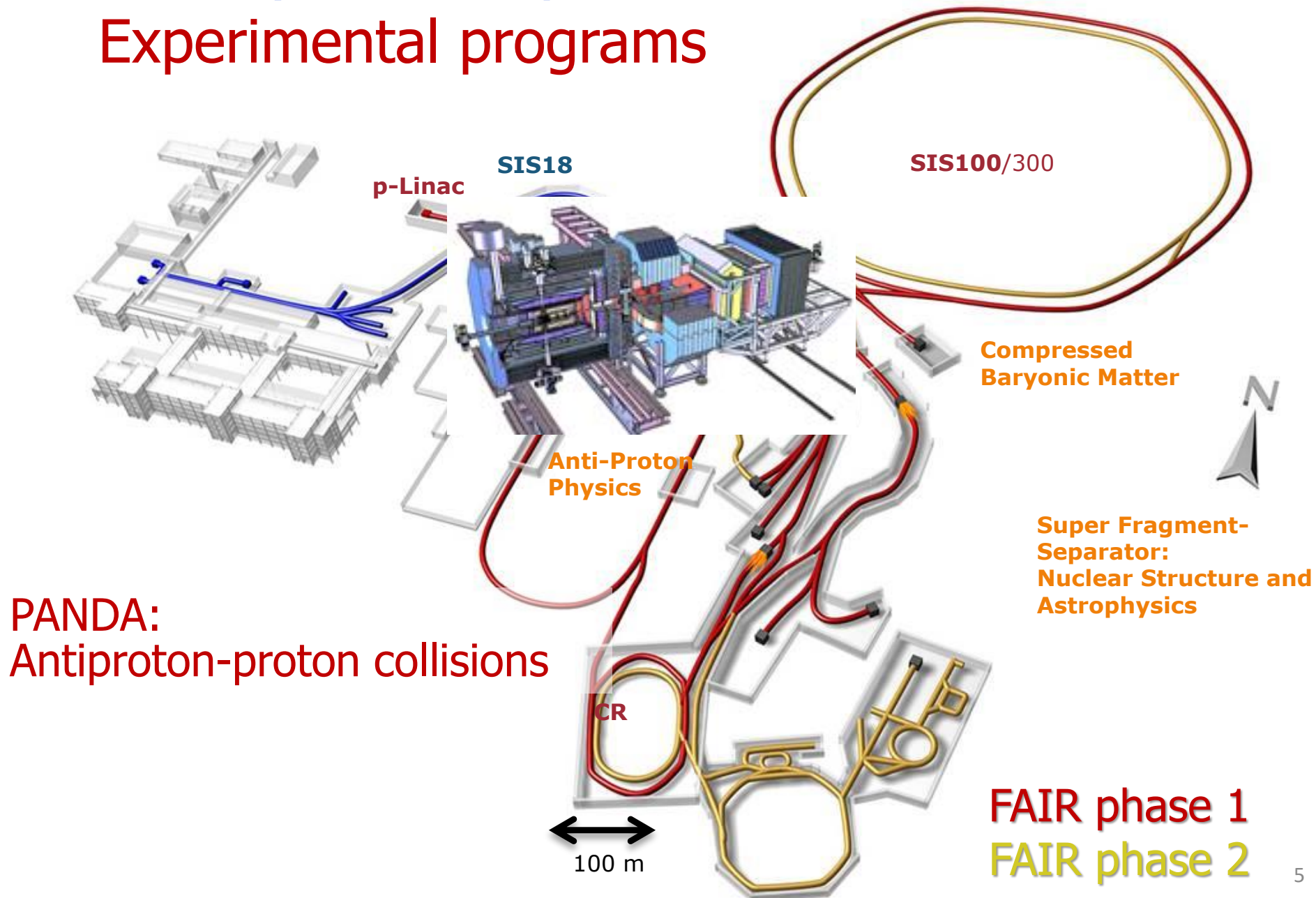
Synthesis of very neutron-rich instable nuclei via rapid capture of neutrons in neutron star mergers





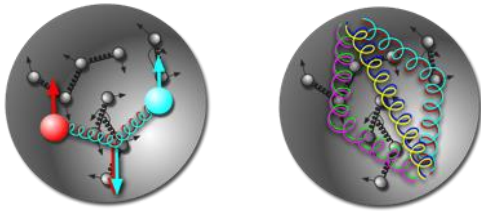
# Facility for Antiproton & Ion Research

## Experimental programs

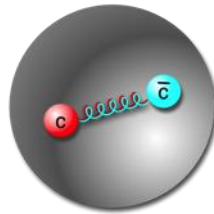


# Hadron Physics with antiprotons at FAIR

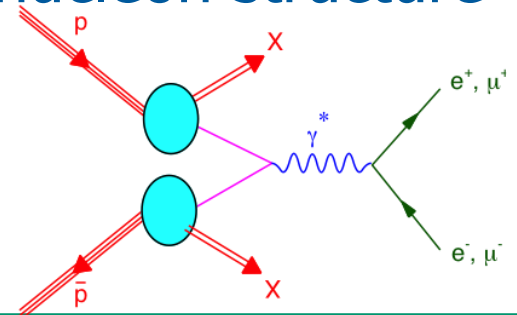
Gluonic excitations:  
Hybrids, glueballs



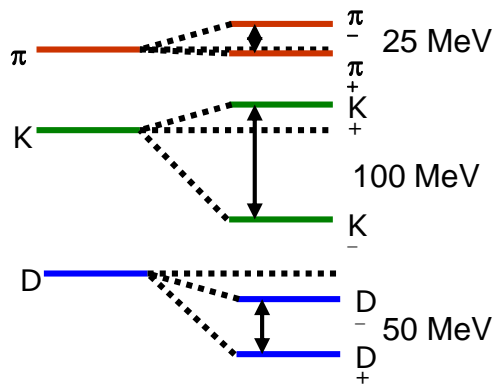
Charmonium states:  
Precision  
spectroscopy



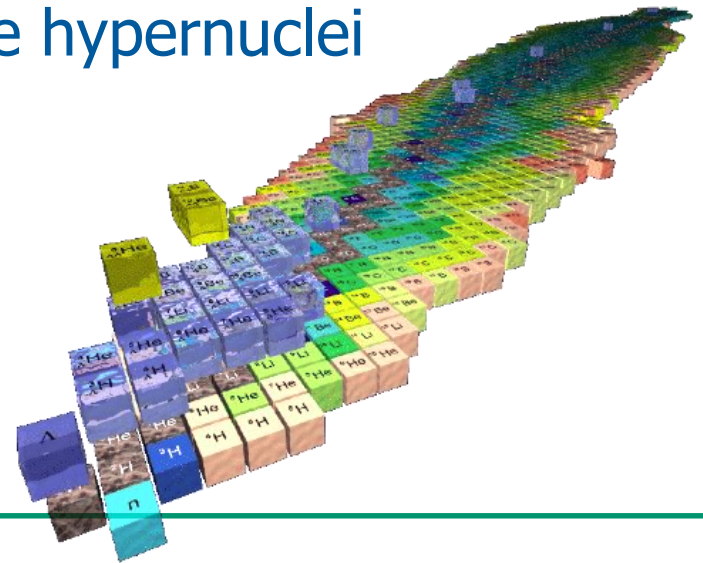
Time-like form factors,  
nucleon structure



In medium mass modifications:  
Extension to the charm sector

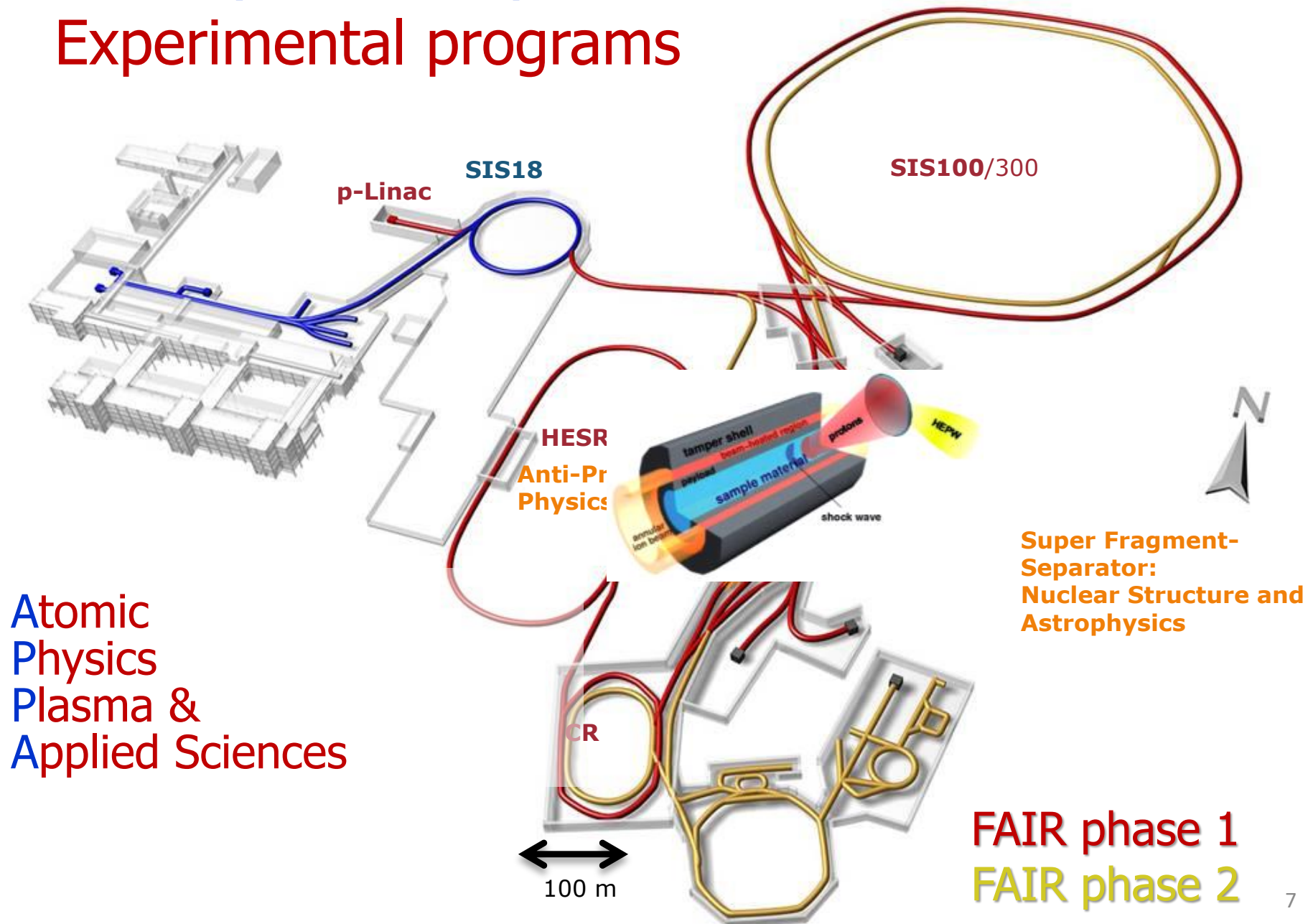


Extension of nuclear chart:  
Double hypernuclei



# Facility for Antiproton & Ion Research

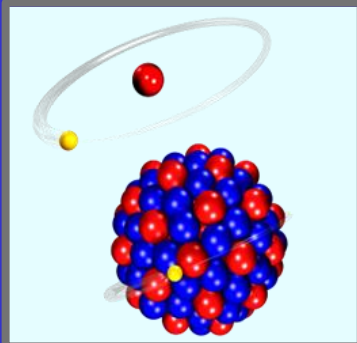
## Experimental programs





# Atomic Physics, Plasma and Applied Sciences

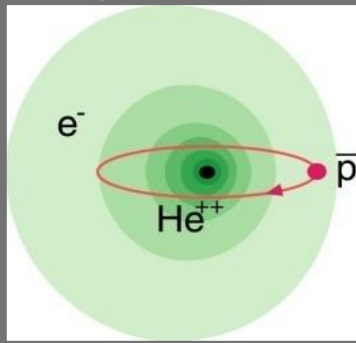
## Atomic Physics



### SPARC

#### strong field research

... probing of fundamental laws of physics

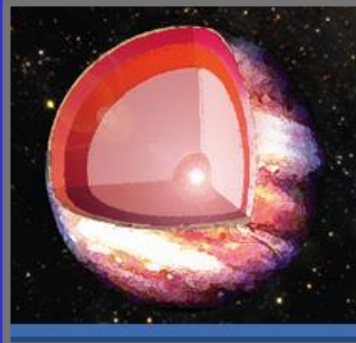


### FLAIR

#### anti-matter

... matter / anti-matter asymmetry

## Plasma

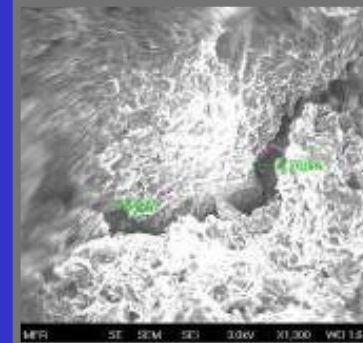


### HEDgeHOB/WDM

#### planetary interiors

... states of matter common in astrophysical objects

## Materials

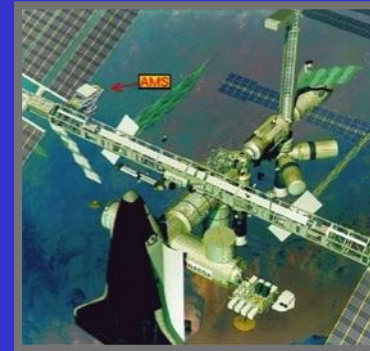


### MAT/BIOMAT

#### extreme conditions

... radiation hardness and modification of materials

## Bio



### BIO/BIOMAT

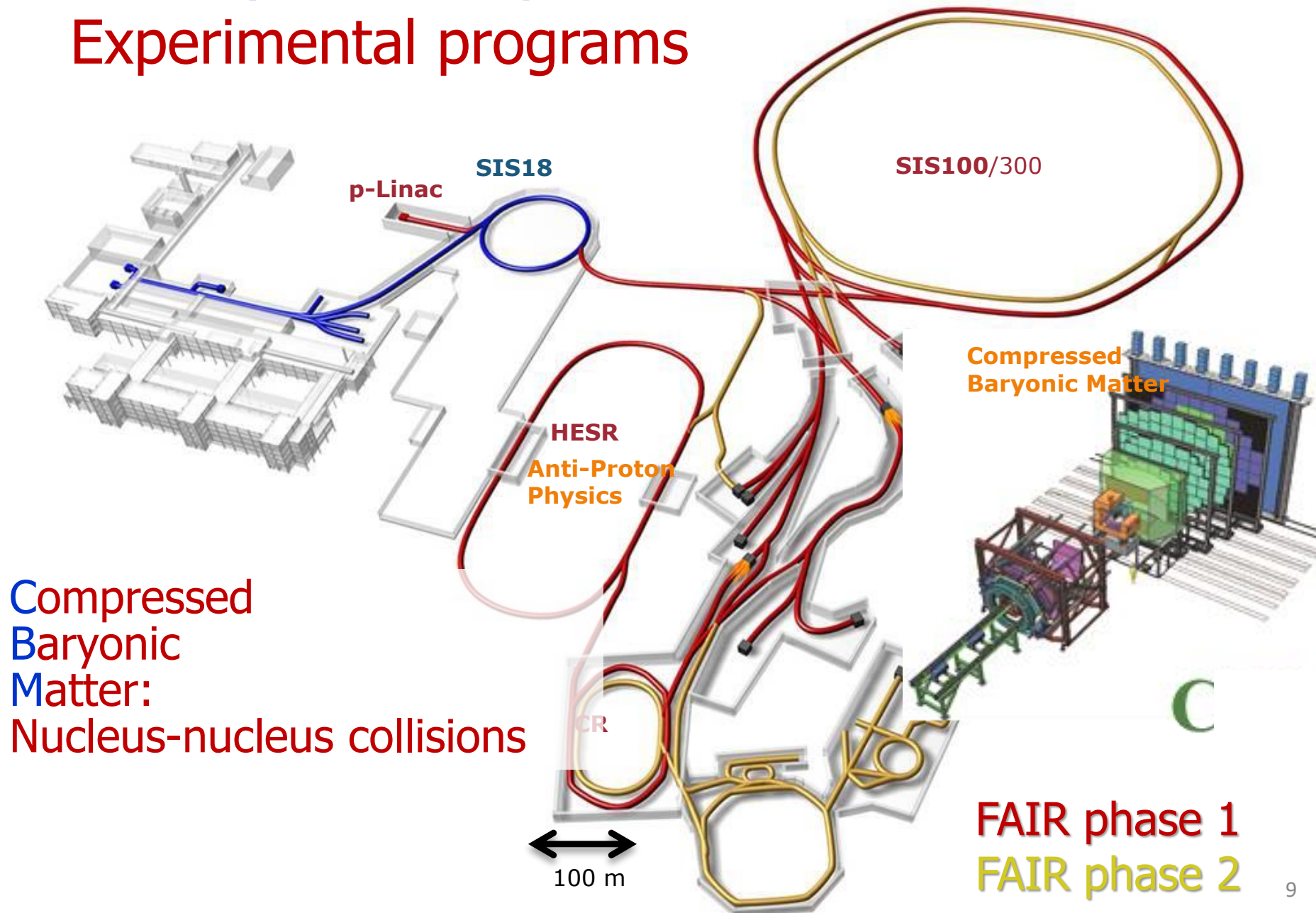
#### aerospace engineering

... radiation shielding of cosmic radiation

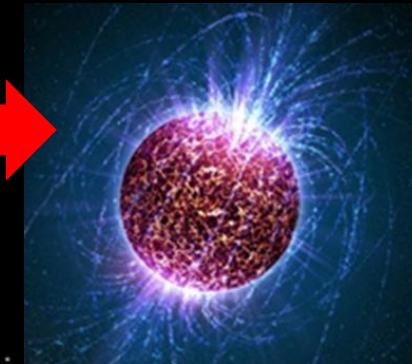
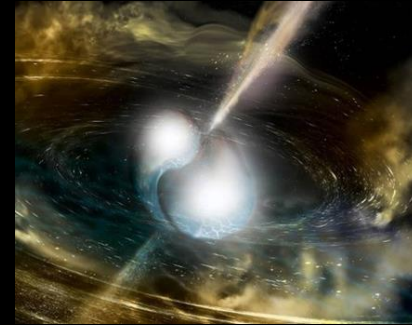
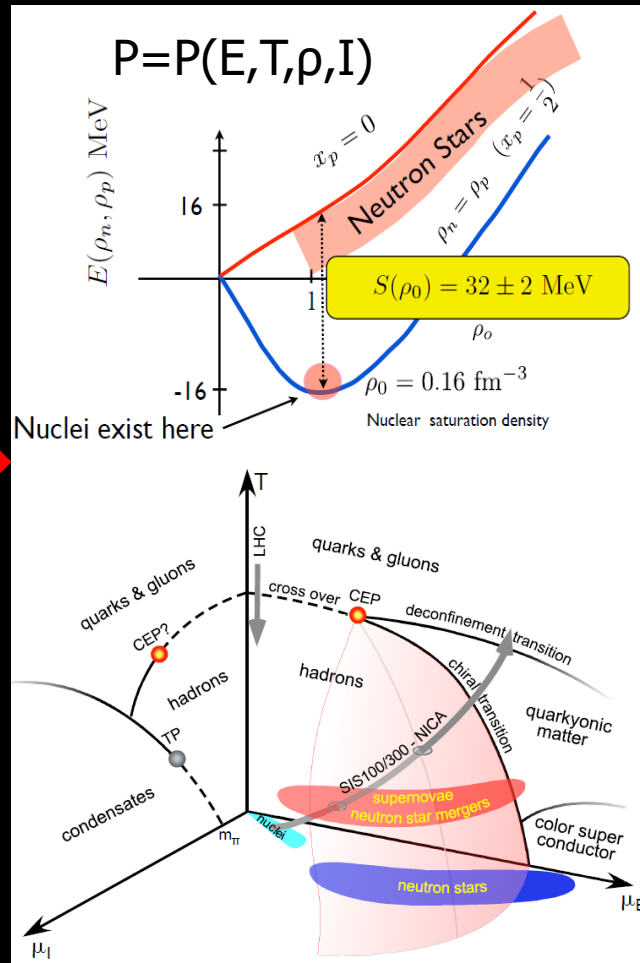
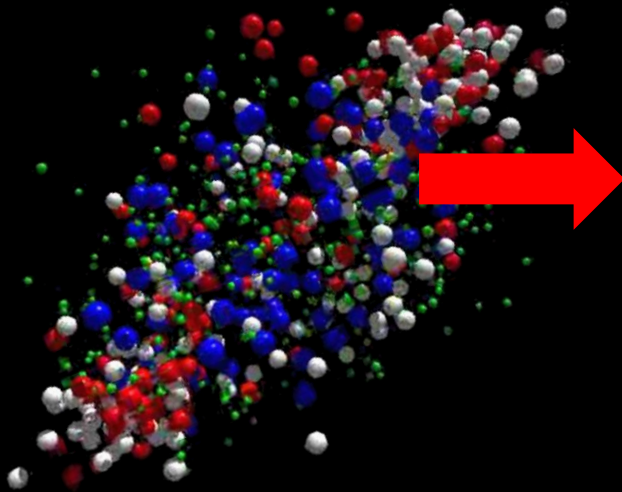
- Highest Charge States: **Extreme Static Fields**
- Relativistic Energies: **Extreme Dynamical Fields and Ultrashort Pulses**
- High Intensities: **Very High Energy Densities and Pressures**
- High Charge at Low Velocity: **Large Energy Deposition**
- Low-Energy Anti-Protons: **Antimatter Research**

# Facility for Antiproton & Ion Research

## Experimental programs



# QCD matter physics

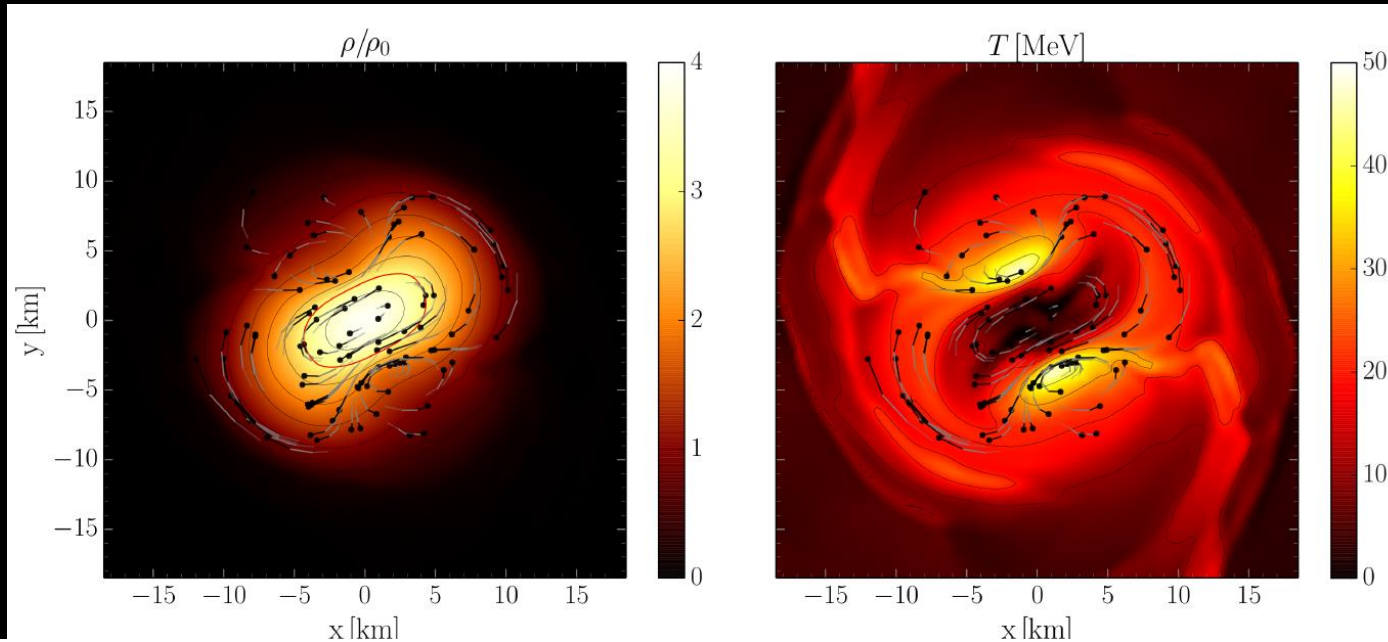




# Neutron star mergers and heavy-ion collisions

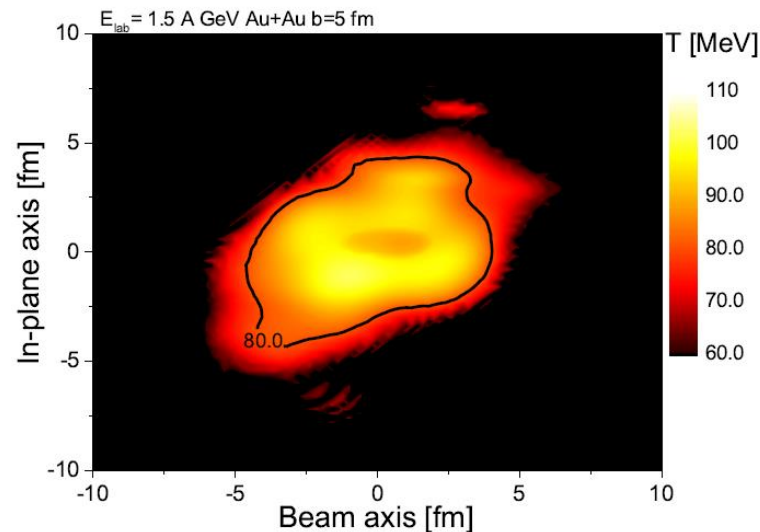
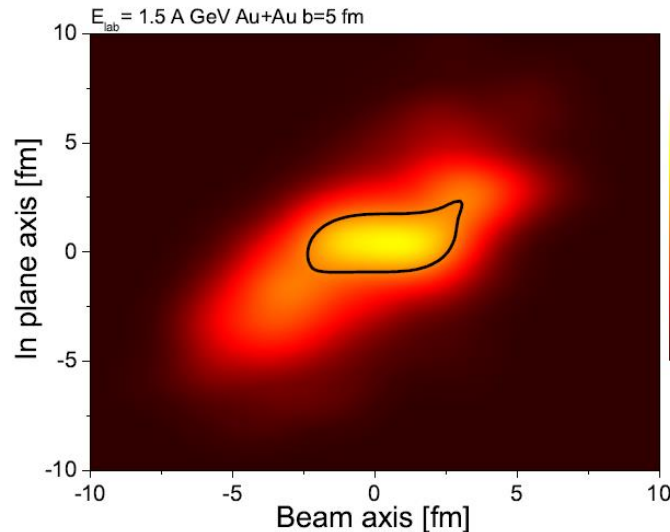
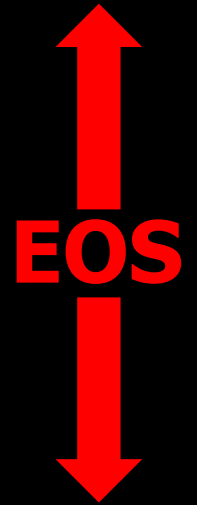
density

temperature



M. Hanauske et al.,  
J. Phys.: Conf. Ser.  
878 012031

n-star merger



Au +Au  
1.5A GeV

# The nuclear matter equation-of-state

The nuclear matter equation of state (EOS) describes the relation between density, pressure, temperature, energy, and isospin asymmetry

$$P = \left. \delta E / \delta V \right|_{T=\text{const}}$$

$$V = A/\rho$$

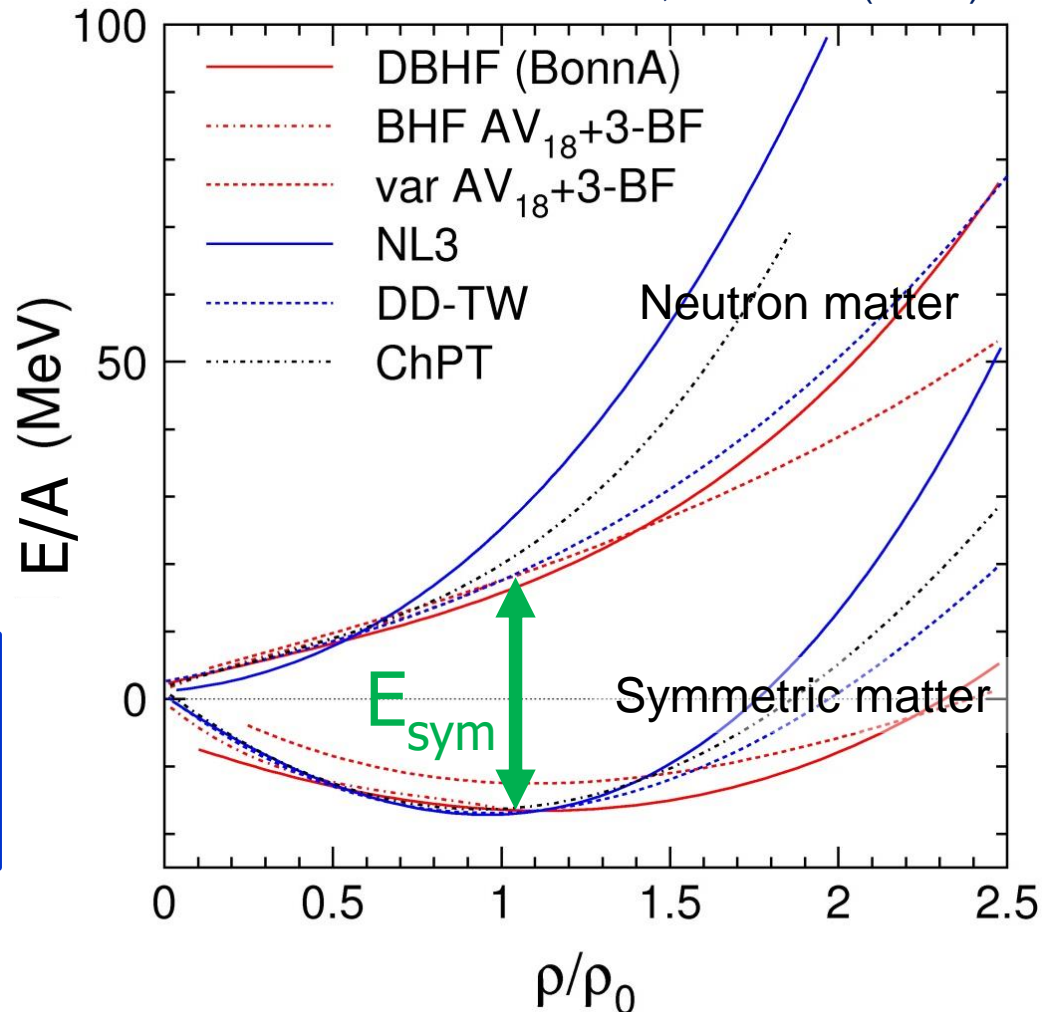
$$\delta V / \delta \rho = -A/\rho^2$$

$$P = \rho^2 \left. \delta(E/A) / \delta \rho \right|_{T=\text{const}}$$

$$E_A(\rho, \delta) = E_A(\rho, 0) + E_{\text{sym}}(\rho) \cdot \delta^2$$

$$\text{with } \delta = (\rho_n - \rho_p) / \rho$$

Ch. Fuchs and H.H. Wolter, EPJA 30 (2006) 5



# The EOS of (symmetric) nuclear matter

$$E_A(\rho, \delta) = E_A(\rho, 0) + E_{\text{sym}}(\rho) \cdot \delta^2 + O(\delta^4)$$

$T=0$ :  $E/A = 1/\rho \int U(\rho) d\rho$

Effective NN-potential:

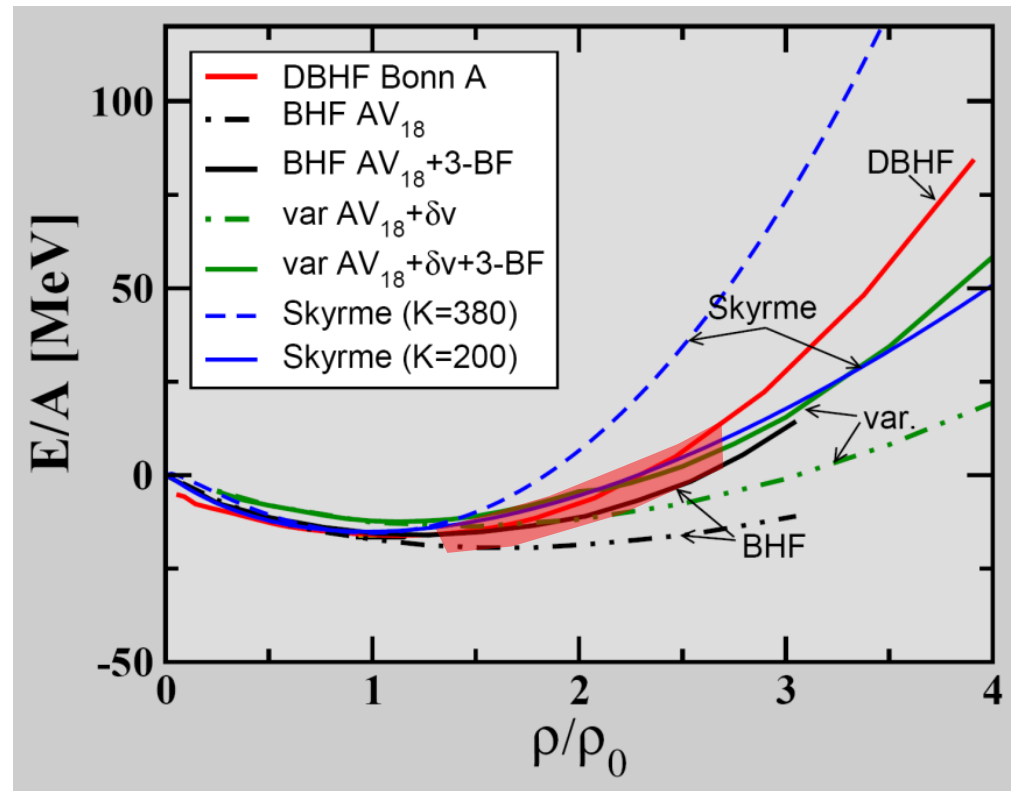
$$U(\rho) = \alpha\rho + \beta\rho^\gamma$$

- $E/A(\rho_0) = -16 \text{ MeV}$
- slope  $\delta(E/A)(\rho_0)/\delta\rho = 0$
- curvature  $K_{\text{nm}} = 9\rho_0^2 \delta^2(E/A)/\delta\rho^2$   
(nuclear incompressibility)

Measurements at GSI SIS18:

- elliptic flow of light fragments
- subthreshold kaon production

C. Fuchs, Prog. Part. Nucl. Phys. 56 (2006) 1



$K_{\text{nm}} = 220 \text{ MeV}$ : "soft" EOS

$K_{\text{nm}} = 380 \text{ MeV}$ : "stiff" EOS

A. Le Fevre et al., Nucl. Phys. A945 (2016) 112

C. Sturm et al., (KaoS Collaboration) Phys. Rev. Lett. 86 (2001) 39

Ch. Fuchs et al., Phys. Rev. Lett. 86 (2001) 1974



# The nuclear symmetry energy

$$E_A(\rho, \delta) = E_A(\rho, 0) + E_{\text{sym}}(\rho) \cdot \delta^2$$

$$E_{\text{sym}}(\rho) = E_{\text{sym}}(\rho_0) + \frac{L}{3} \left( \frac{\rho - \rho_0}{\rho_0} \right) + \frac{K_{\text{sym}}}{18} \left( \frac{\rho - \rho_0}{\rho_0} \right)^2$$

Empirical value  $E_{\text{sym}}(\rho_0) \approx 30 \text{ MeV}$

slope 
$$L = 3\rho_0 \left. \frac{\partial E_{\text{sym}}(\rho)}{\partial \rho} \right|_{\rho=\rho_0}$$

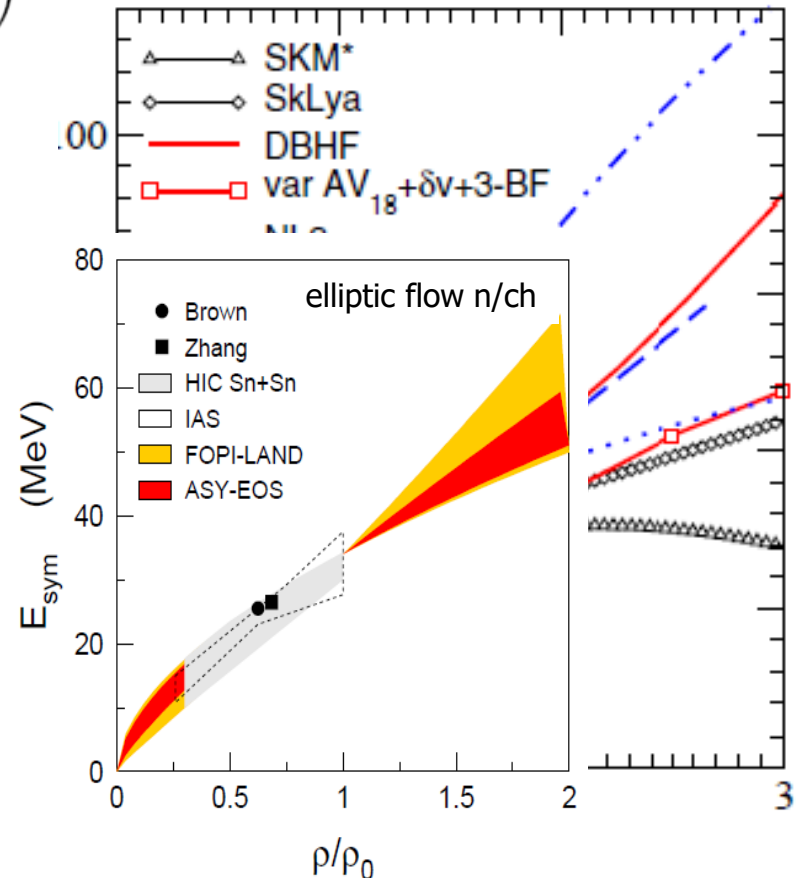
theoretical value  $L(\rho_0) \approx 60 \text{ MeV}$

B.A. Li and X. Han, Phys. Lett. B 727 (2013) 276

curvature 
$$K_{\text{sym}} = 9\rho_0^2 \left. \frac{\partial^2 E_{\text{sym}}(\rho)}{\partial^2 \rho} \right|_{\rho=\rho_0}$$

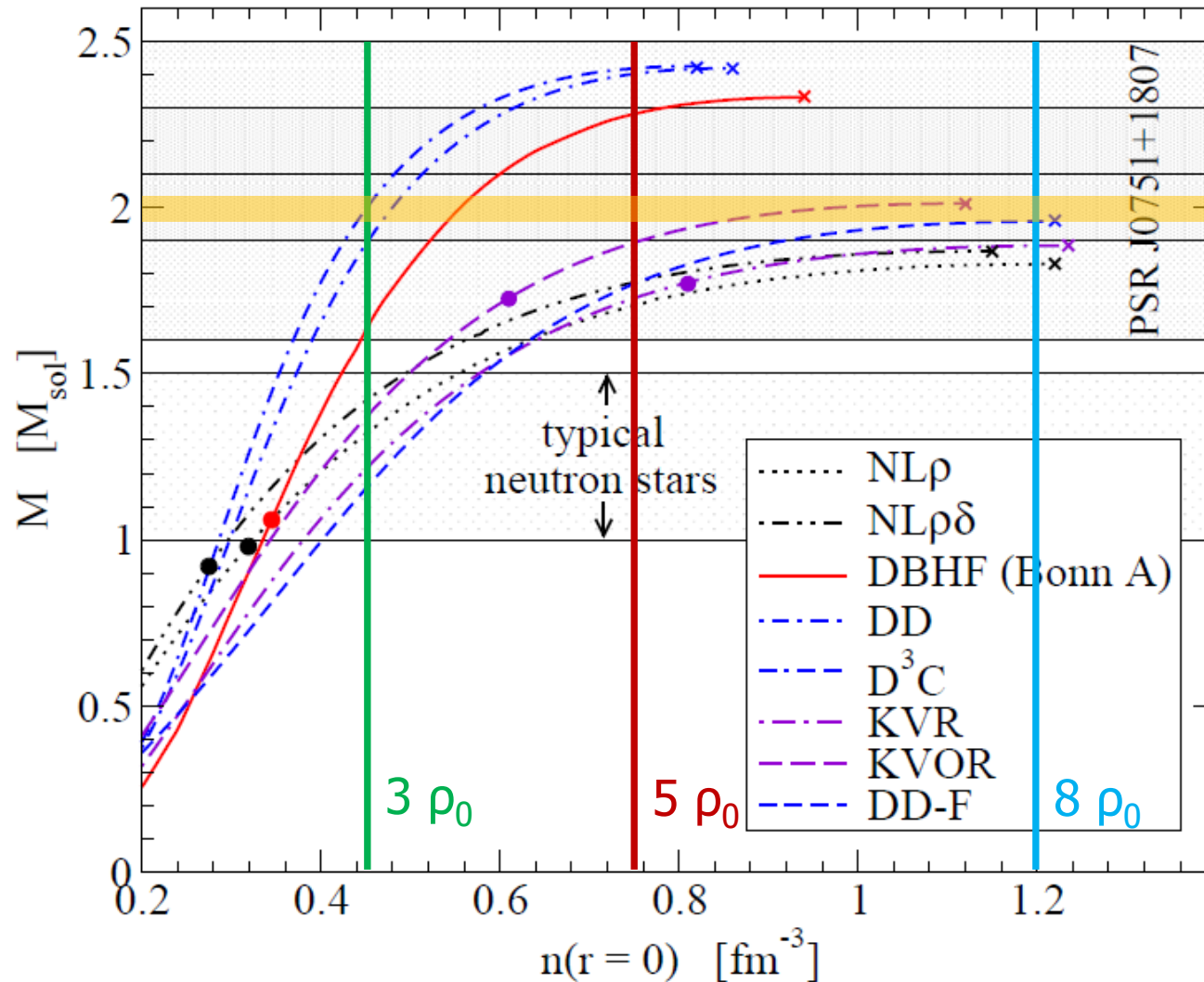
theoretical value  $K_{\text{sym}} = -700 \text{ to } 470 \text{ MeV}$

Ch. Fuchs and H.H. Wolter, EPJA30 (2006) 5



P. Russotto et al., Phys. Rev. C 94, 034608 (2016)

# Mass-density relation of neutron stars for different EOS



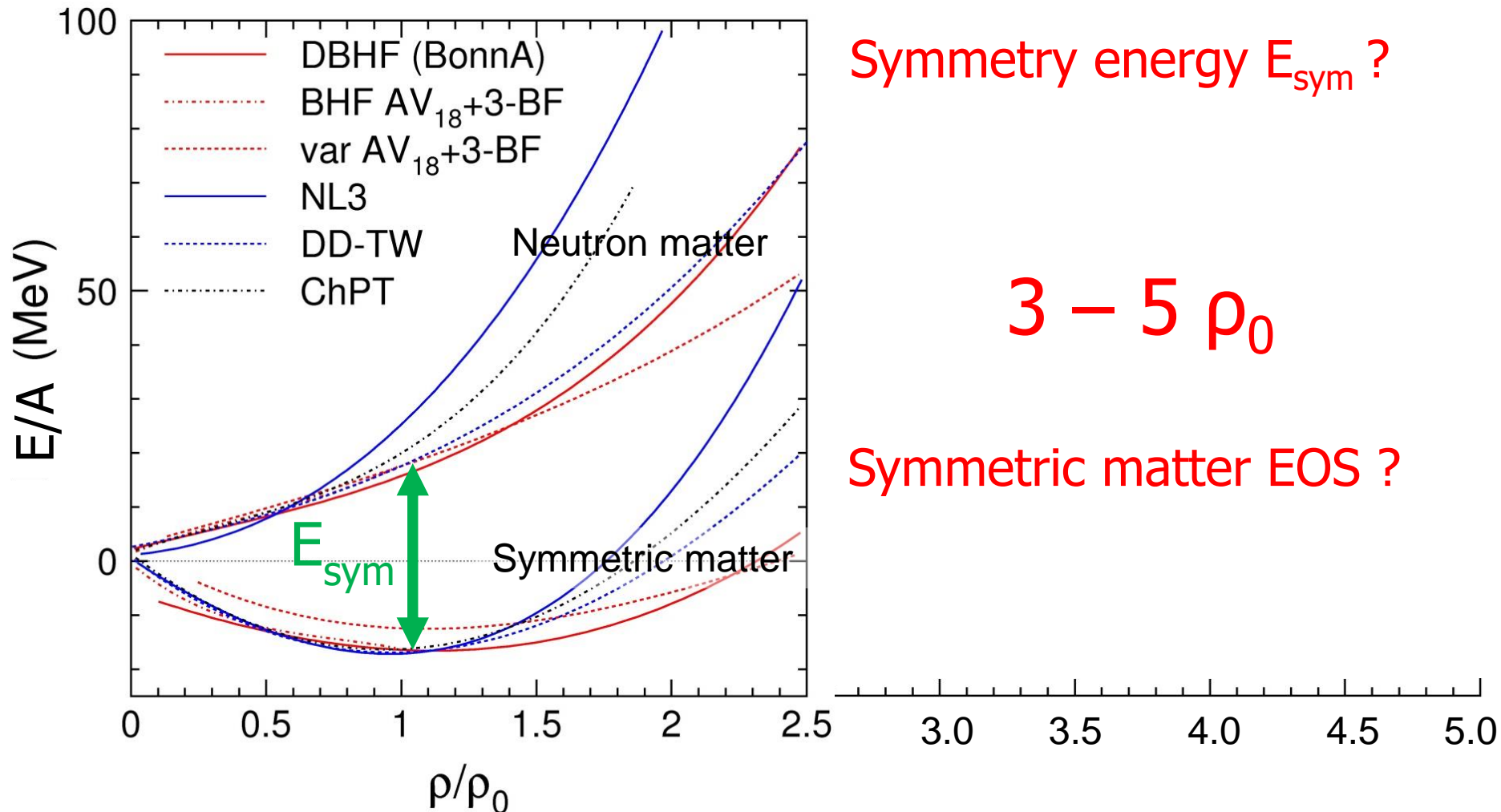
PSR J1614-2230  
 $M = 1.97 \pm 0.04 M_{\text{sun}}$

P. Demorest et al.,  
*Nature* 467, 1081 (2010)

PSR J0348+0432  
 $M = 2.01 \pm 0.04 M_{\text{sun}}$

J. Antoniadis et al.,  
*Science* **340**, 6131 (2013)

# The high-density nuclear matter equation-of-state

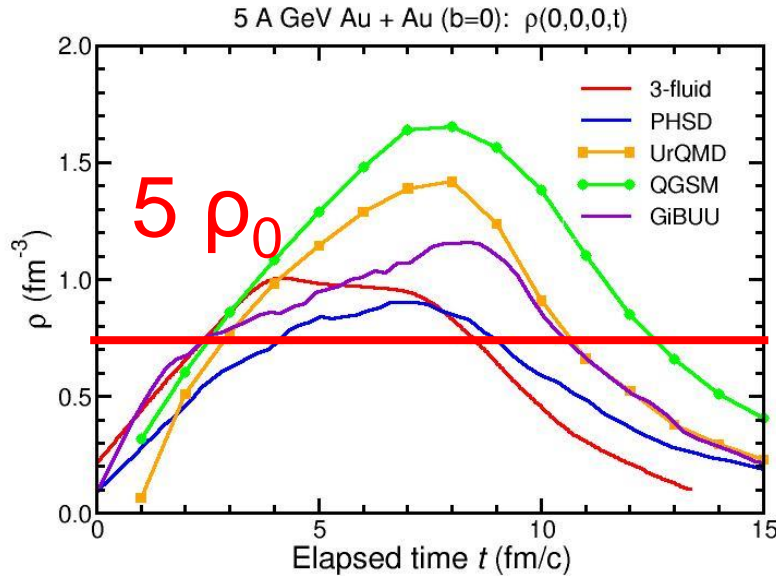




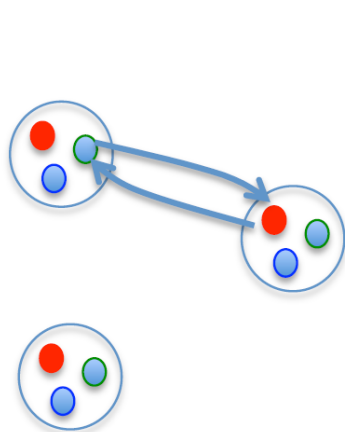
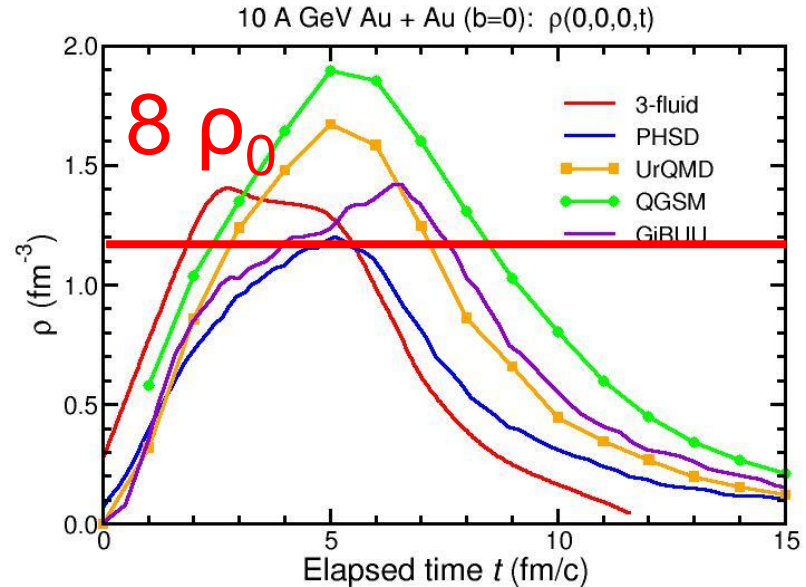
# Baryon densities in central Au+Au collisions

I.C. Arsene et al., Phys. Rev. C 75, 24902 (2007)

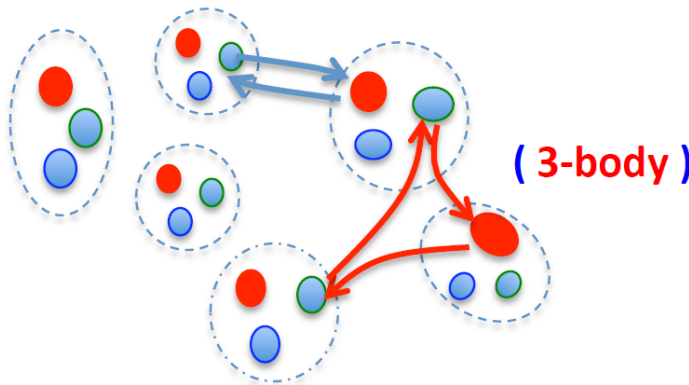
**5 A GeV**



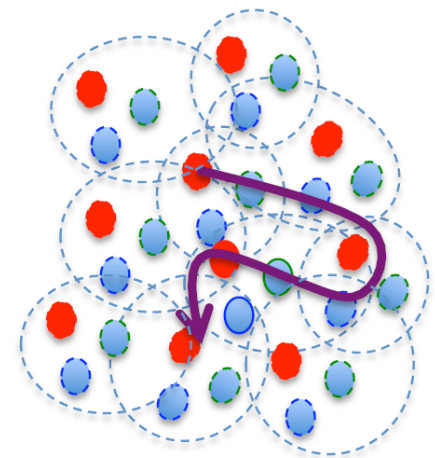
**10 A GeV**



$\sim 2 \rho_0$



$\sim 5 \rho_0$



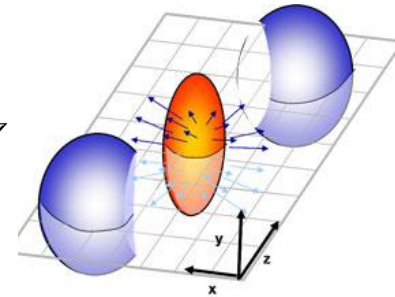
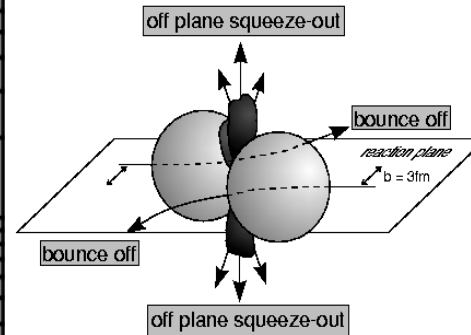
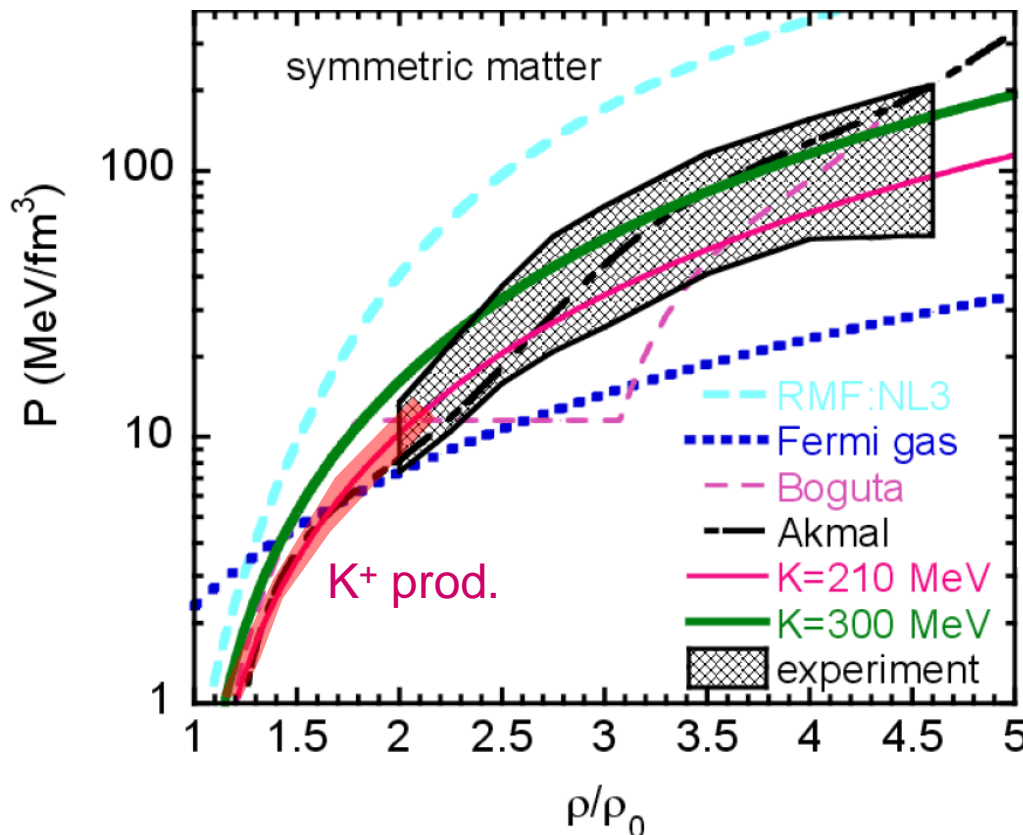
courtesy Toru Kojo (CCNU)

# CBM physics case and observables

The QCD matter equation-of-state at neutron star core densities

- collective flow of identified particles ( $n, K, p, \Lambda, \Xi, \Omega, \dots$ ) driven by the pressure gradient in the early fireball

EOS of symmetric matter extracted from proton flow in Au+Au collisions measured at AGS for beam energies from 2 to 11A GeV.



Azimuthal angle distribution:

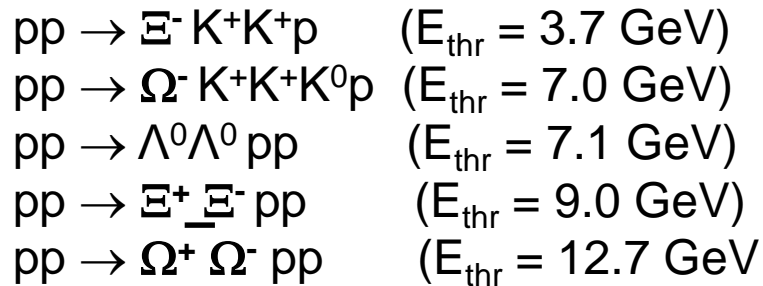
$$dN/d\phi = C (1 + v_1 \cos(\phi) + v_2 \cos(2\phi) + \dots)$$

# CBM physics case and observables

The QCD matter equation-of-state at neutron star core densities

- particle production at (sub)threshold energies via multi-step processes (multi-strange hyperons, charm)

Direct multi-strange hyperon production:

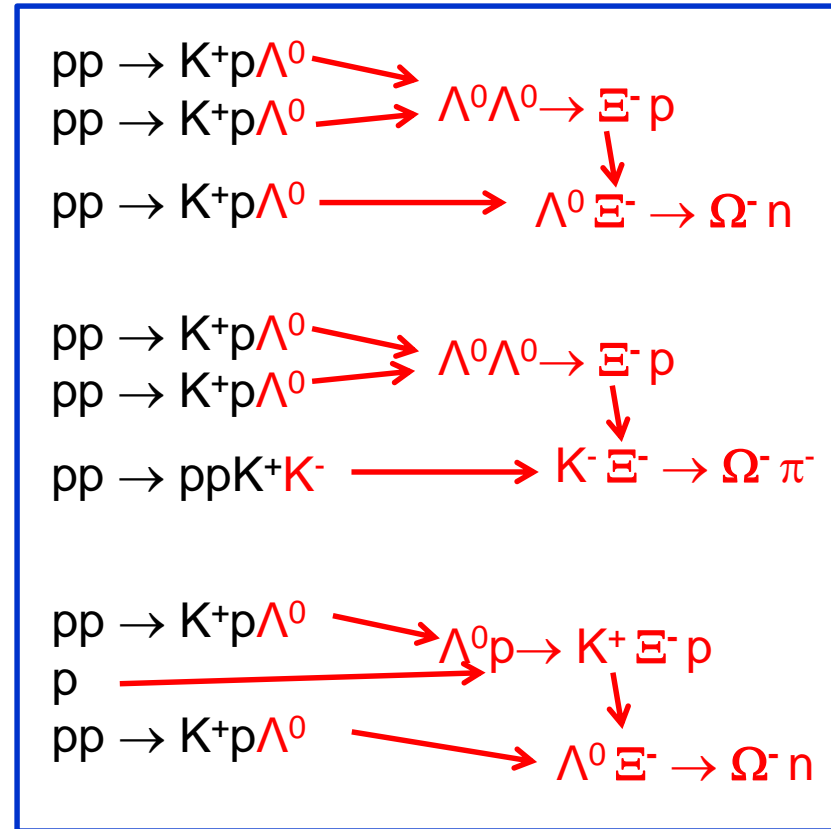


Hyperon production via multiple collisions

1.  $pp \rightarrow K^+ \Lambda^0 p$ ,  $pp \rightarrow K^+ K^- pp$ ,
2.  $p \Lambda^0 \rightarrow K^+ \Xi^- p$ ,  $\pi \Lambda^0 \rightarrow K^+ \Xi^- \pi$ ,  
 $\Lambda^0 \Lambda^0 \rightarrow \Xi^- p$ ,  $\Lambda^0 K^- \rightarrow \Xi^- \pi^0$
3.  $\Lambda^0 \Xi^- \rightarrow \Omega^- n$ ,  $\Xi^- K^- \rightarrow \Omega^- \pi^-$

Antihyperons

1.  $\Lambda^0 K^+ \rightarrow \Xi^+ \pi^0$ ,
2.  $\Xi^+ K^+ \rightarrow \Omega^+ \pi^+$ .



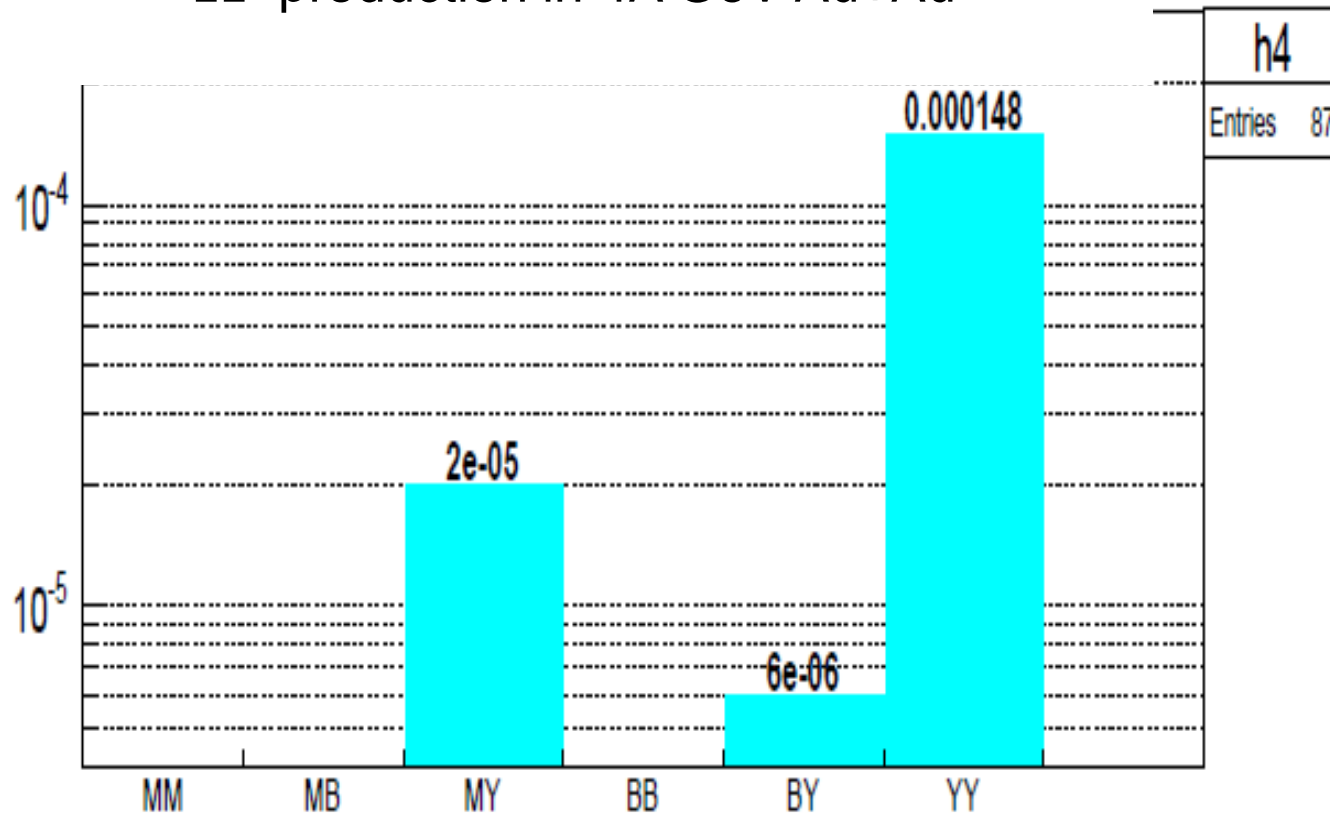
➡ Hyperon yield ~ multi-step collisions ~ density → EOS

# CBM physics case and observables

The QCD matter equation-of-state at neutron star core densities

- particle production at (sub)threshold energies via multi-step processes (multi-strange hyperons, charm)

$\Omega^-$  production in 4A GeV Au+Au



HYPQGSM calculations, K. Gudima, Y. Murin et al. , priv. comm.

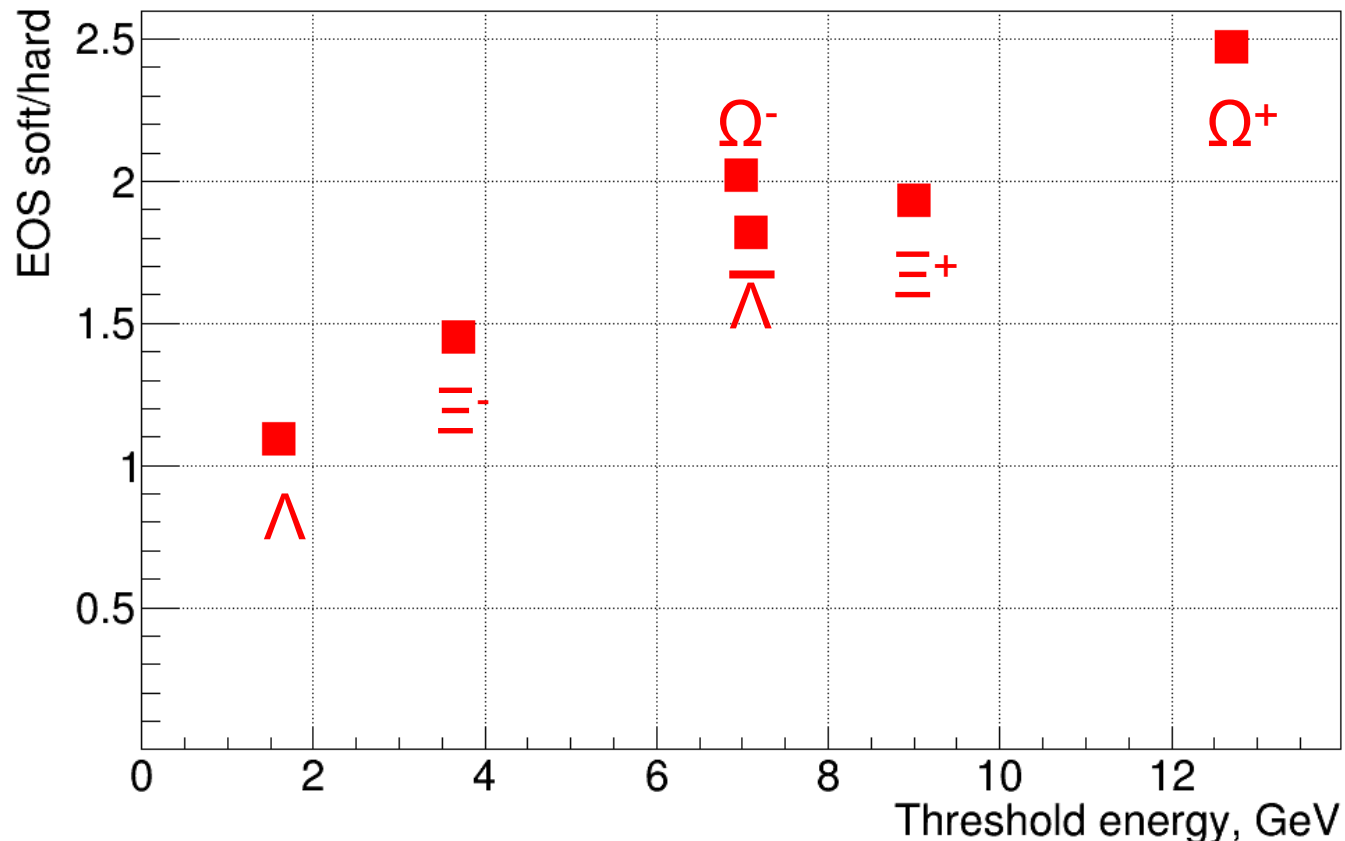


# CBM physics case and observables

The QCD matter equation-of-state at neutron star core densities

- particle production at (sub)threshold energies via multi-step processes (multi-strange hyperons, charm)

Hyperon yield in 4A GeV Au+Au:  
soft EOS (K=240 MeV) / hard EOS (K=350) MeV

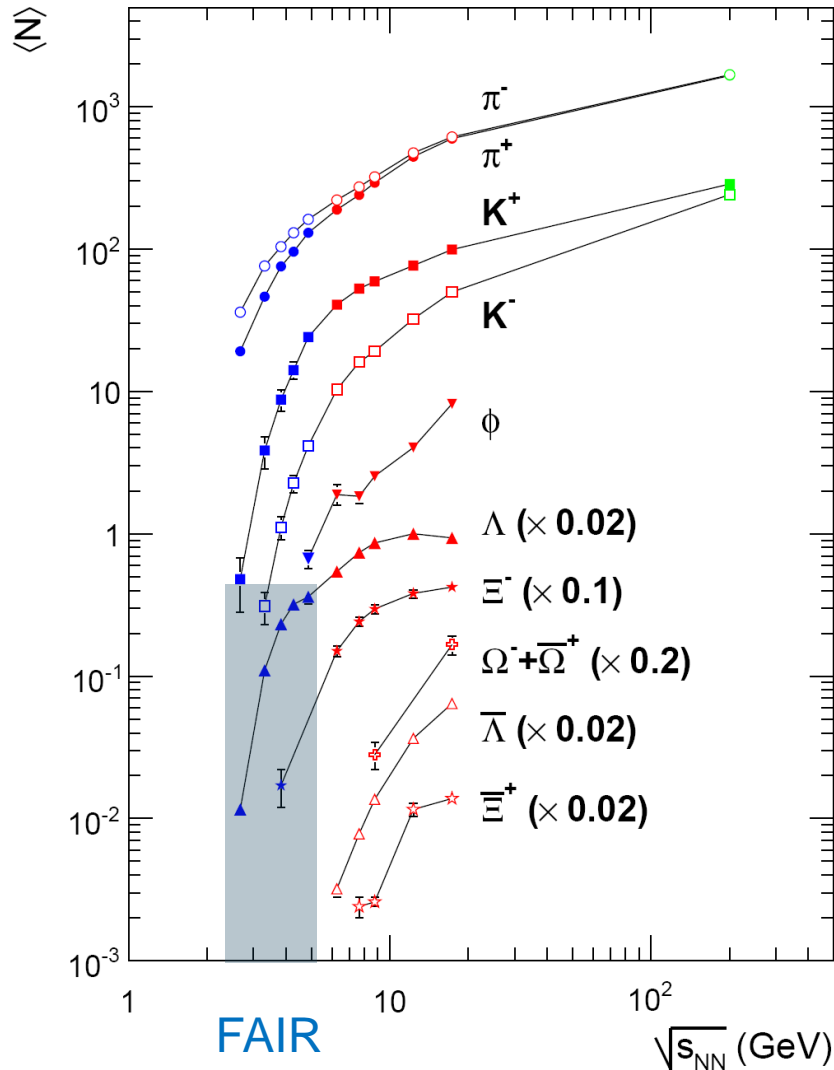


PHQMD calculations , V. Kireyeu et al., priv. comm.

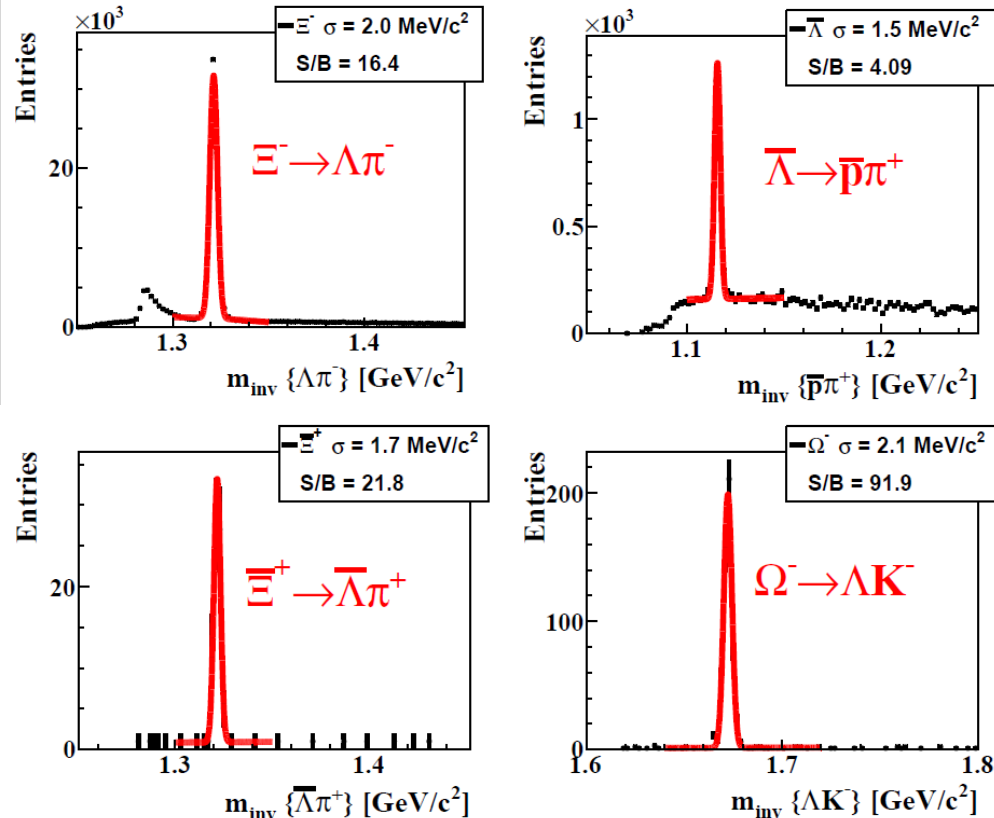
# CBM physics case and observables

The QCD matter equation-of-state at neutron star core densities

- particle production at (sub)threshold energies via multi-step processes (multi-strange hyperons, charm)



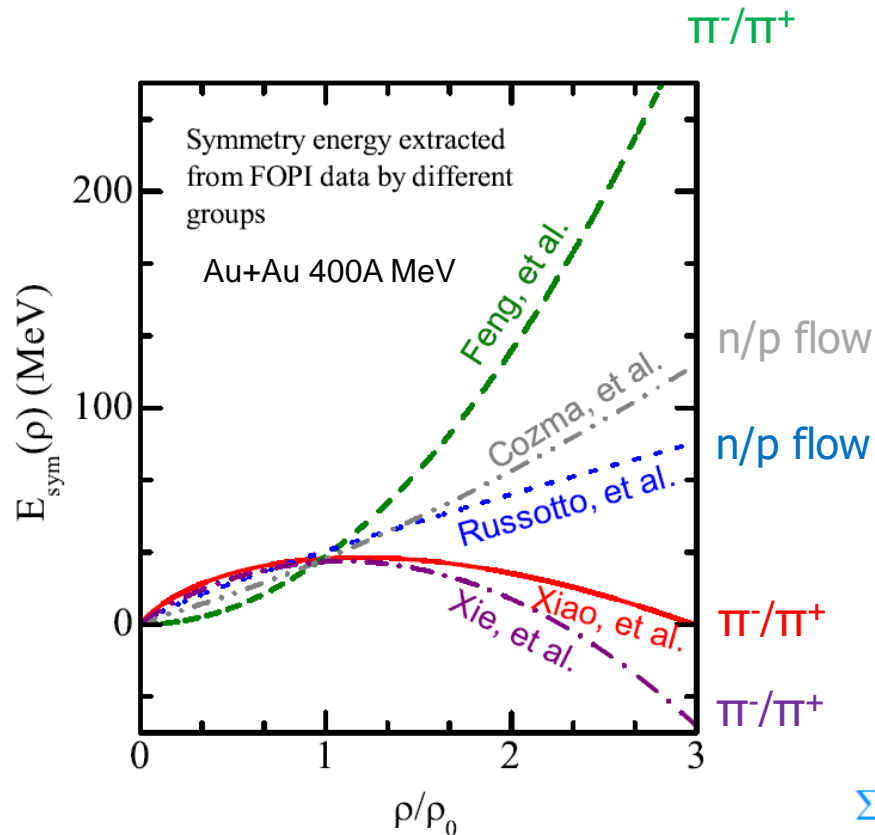
Simulations using the UrQMD event generator for central Au+Au collisions 10A GeV based on realistic detector responses



# CBM physics case and observables

The symmetry energy  $E_{\text{sym}}$  at high density

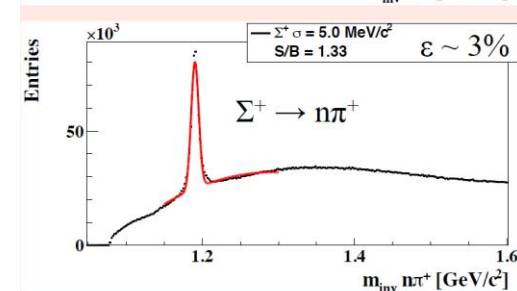
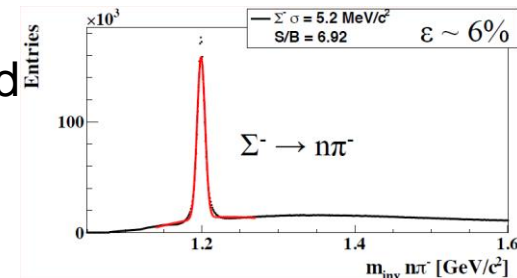
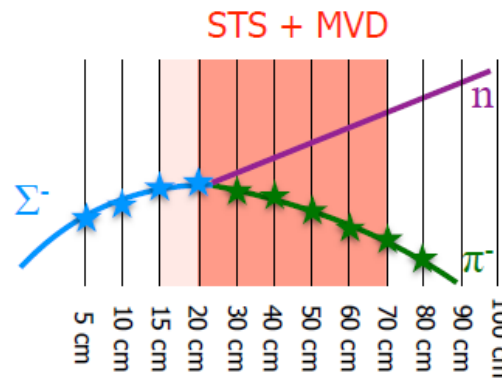
- Elliptic flow neutrons/protons (upgrade option)
- Particles with opposite isospin



W.-M. Guo et al., Phys. Lett. B738 (2014) 397

| $I_3$ | particle        | production  | $E_{\text{thr}}$<br>GeV | decay  |
|-------|-----------------|---|-------------------------|--|
| +1    | $\Sigma^+(uus)$ | $pp \rightarrow \Sigma^+ K^+ n$<br>$pp \rightarrow \Sigma^+ K^0 p$<br>$pn \rightarrow \Sigma^+ K^0 n$ | 1.8                     | $\Sigma^+ \rightarrow p \pi^0$<br>$\Sigma^+ \rightarrow n \pi^+$ |
| -1    | $\Sigma^-(dds)$ | $pn \rightarrow \Sigma^- K^+ p$<br>$nn \rightarrow \Sigma^- K^+ n$                                    | 1.8                     | $\Sigma^- \rightarrow n \pi^-$                                   |

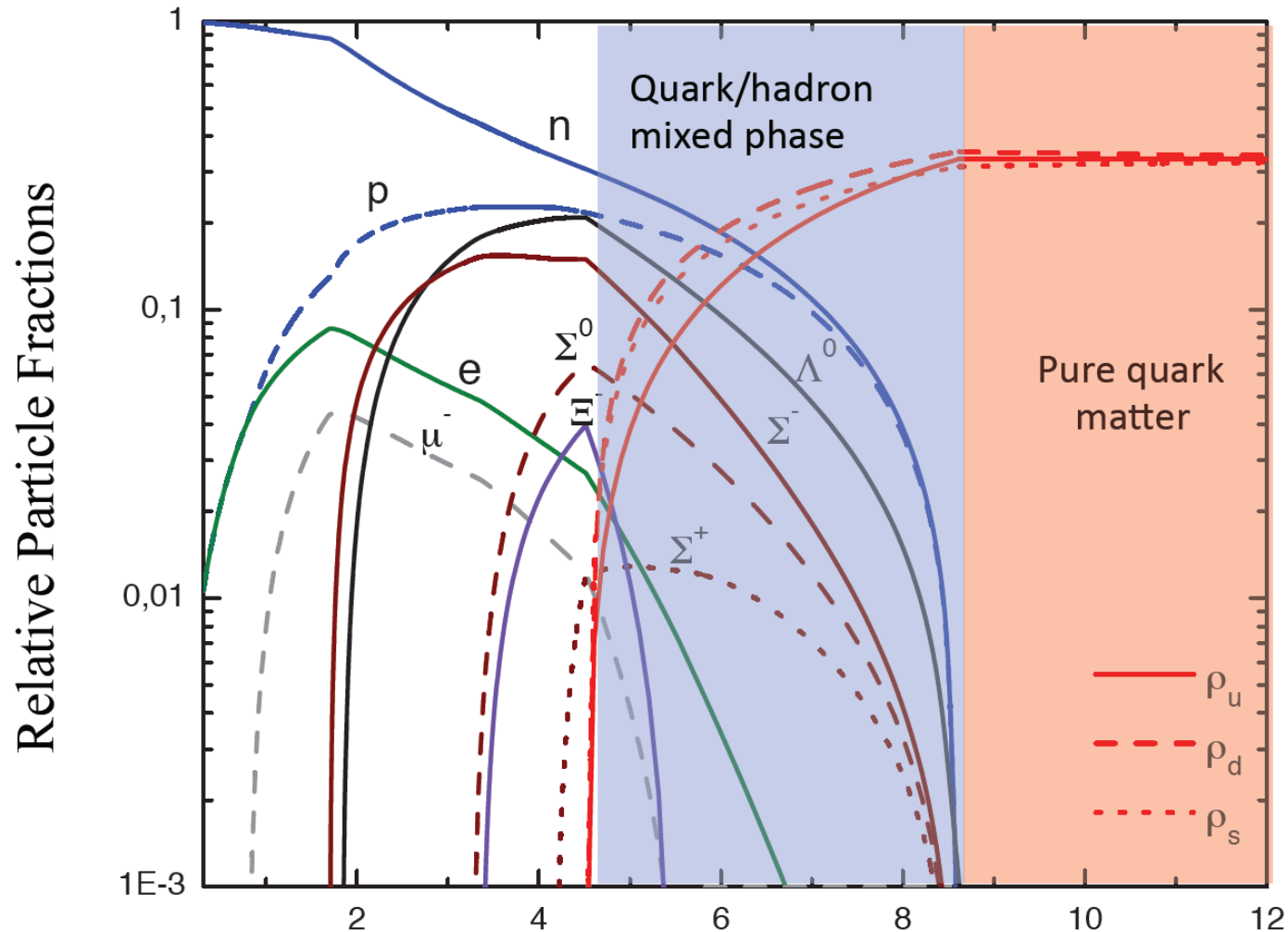
Missing mass method



# Quark matter in massive neutron stars?

M. Orsaria, H. Rodrigues, F. Weber, G.A. Contrera, arXiv:1308.1657

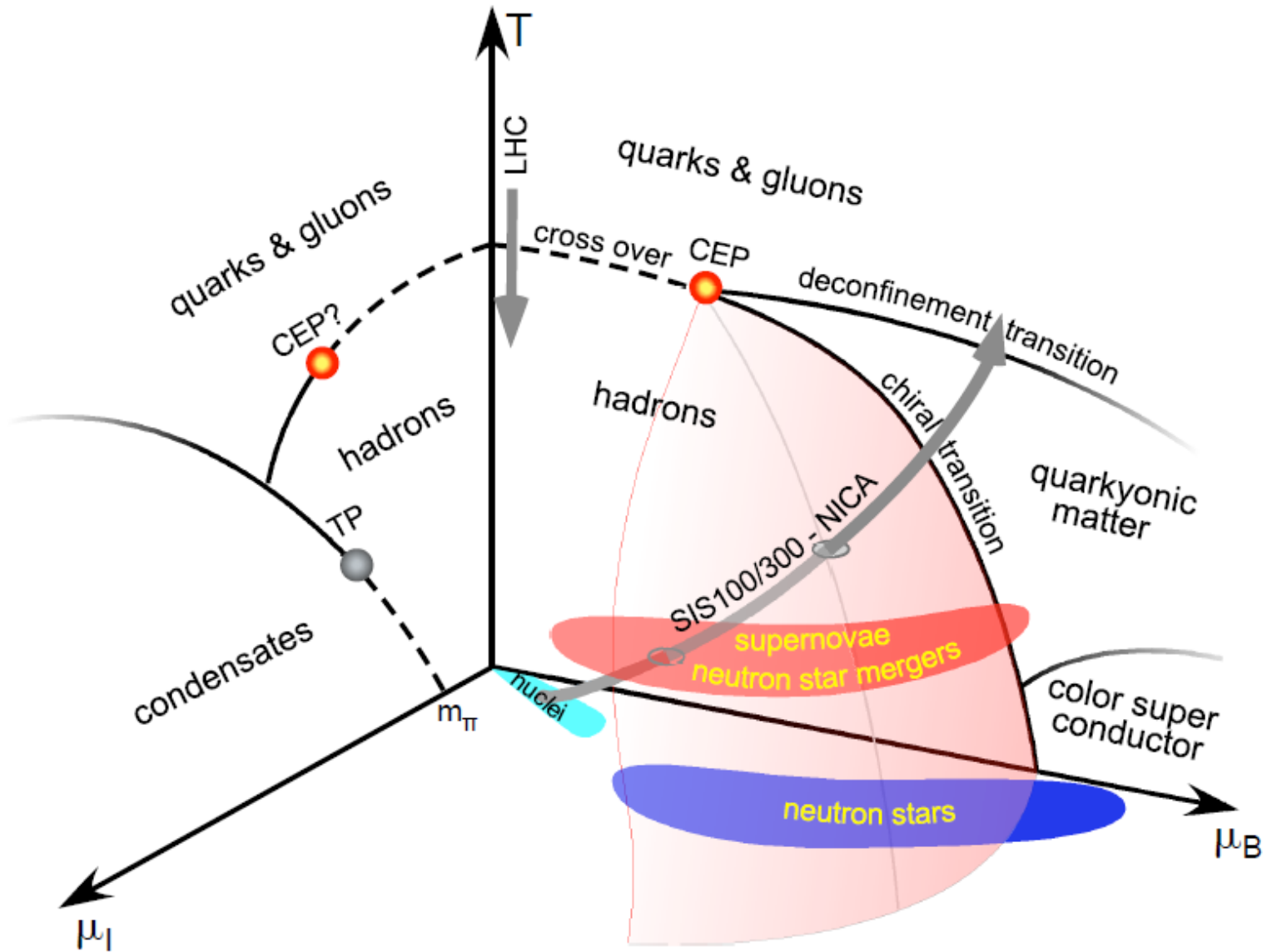
Phys. Rev. C 89, 015806, 2014



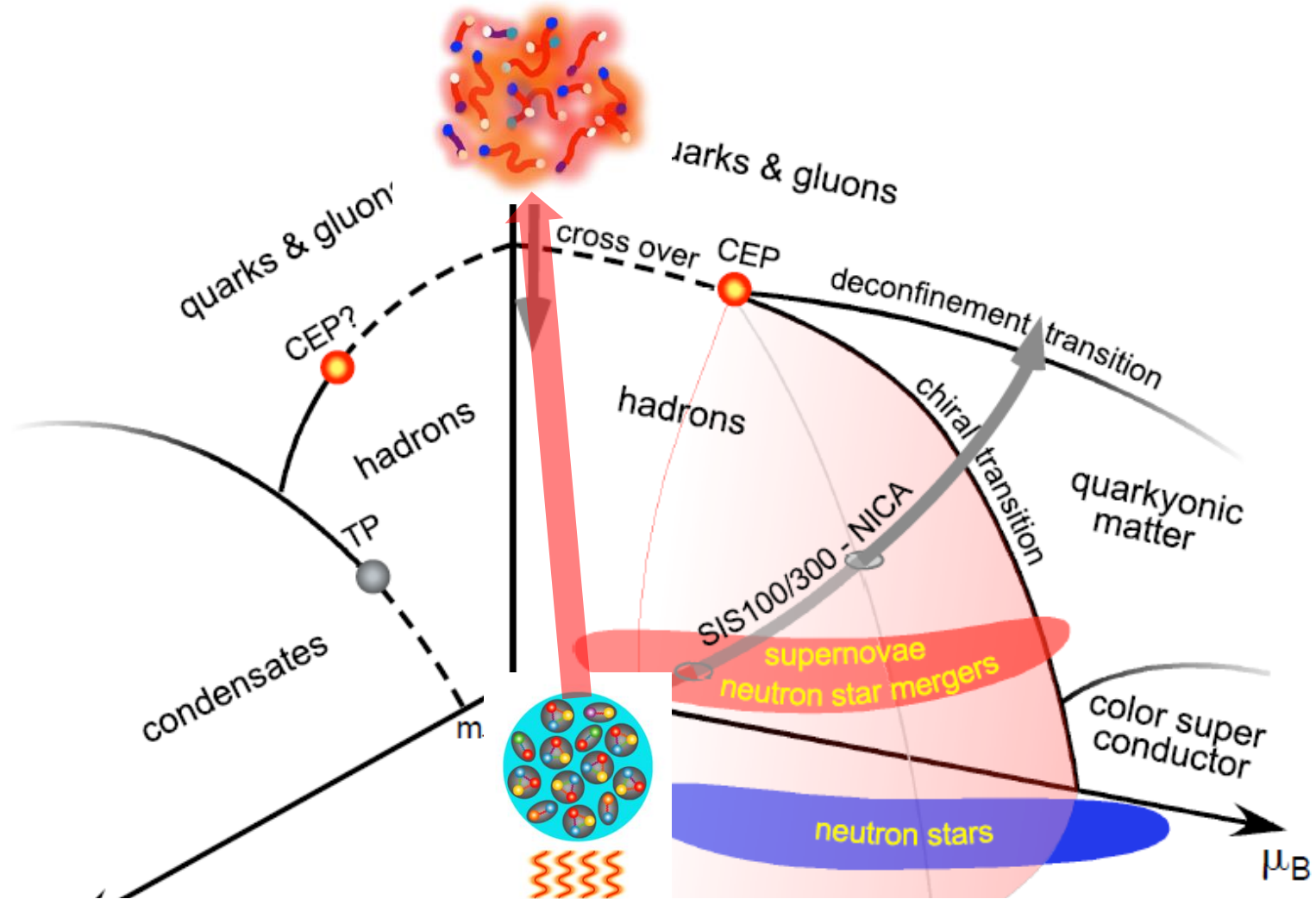
→ QCD phase diagram



# Exploring the QCD phase diagram



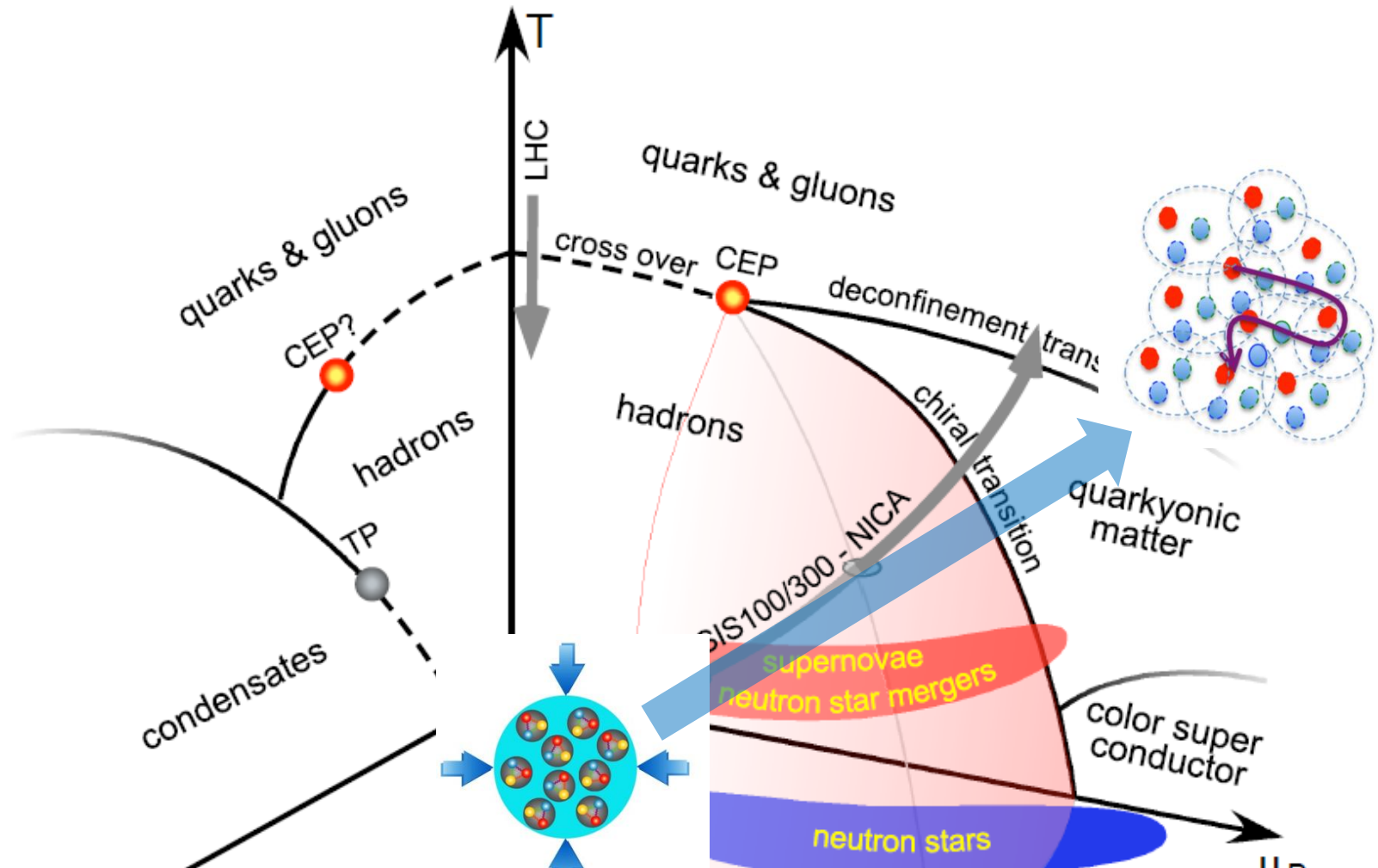
# Exploring the QCD phase diagram



At very high temperature:

- $N$  of baryons  $\approx N$  of antibaryons, situation similar to early universe
- L-QCD finds crossover transition between hadronic matter and Quark-Gluon Plasma
- Experiments: [ALICE](#), [ATLAS](#), [CMS](#) at LHC, [STAR](#), [PHENIX](#) at RHIC

# Exploring the QCD phase diagram



At high baryon density:

- N of baryons  $\gg$  N of antibaryons , densities like in neutron star cores
- L-QCD not (yet) applicable, models predict phase transitions and exotic phases
- Experiments:  
BES at RHIC, NA61 at CERN SPS, CBM at FAIR, MPD/BM@N at NICA, CEE at HIAF

# Linking hadron production in heavy-ion collisions to the QCD phase diagram: Statistical Hadronization Model

## Particle density:

$$n_i = N_i/V = -\frac{T}{V} \frac{\partial \ln Z_i}{\partial \mu} = \frac{g_i}{2\pi^2} \int_0^\infty \frac{p^2 dp}{\exp[(E_i - \mu_i)/T] \pm 1},$$

$N_i$ : number of particles  $i$

$V$ : volume

$T$ : temperature

$Z_i$ : partition function

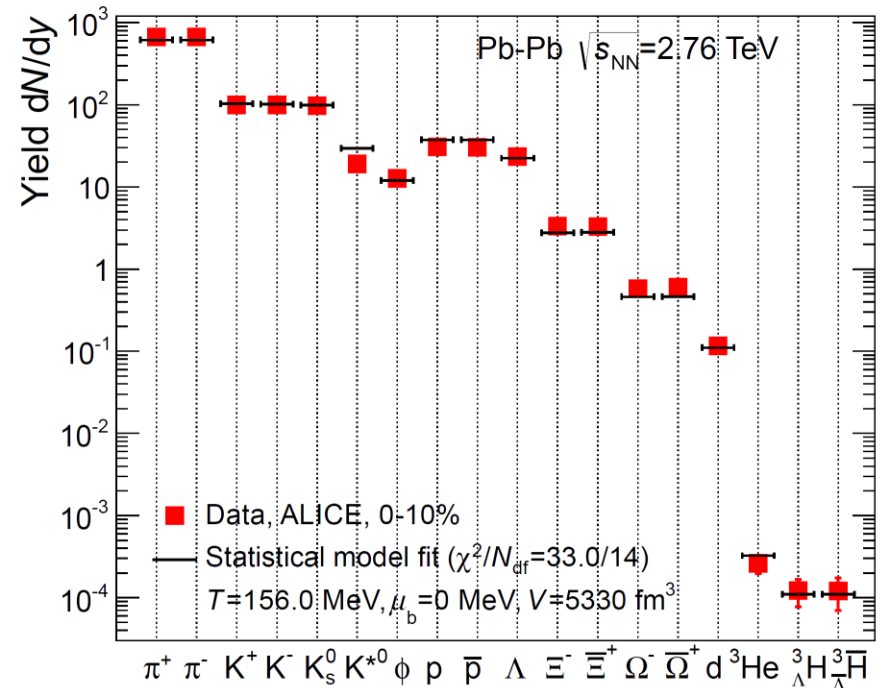
$\mu$ : chemical potential

$g_i = (2J_i + 1)$ : spin degeneracy factor

$E_i = \sqrt{(p^2 + m_i^2)}$ : total energy

## Assumptions:

- chemical equilibrium
- simultaneous freeze-out of all particles





# Linking hadron production in heavy-ion collisions to the QCD phase diagram: Statistical Hadronization Model

## Particle density:

$$n_i = N_i/V = -\frac{T}{V} \frac{\partial \ln Z_i}{\partial \mu} = \frac{g_i}{2\pi^2} \int_0^\infty \frac{p^2 dp}{\exp[(E_i - \mu_i)/T] \pm 1},$$

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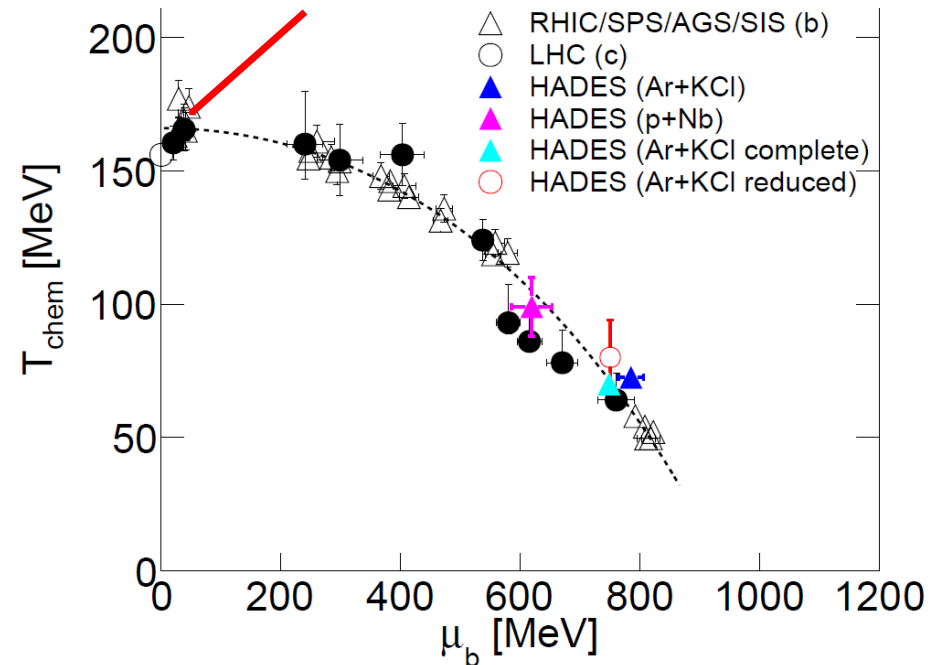
$g_i = (2J_i + 1)$ : spin degeneracy factor

$E_i = \sqrt{p^2 + m_i^2}$ : total energy

## Assumptions:

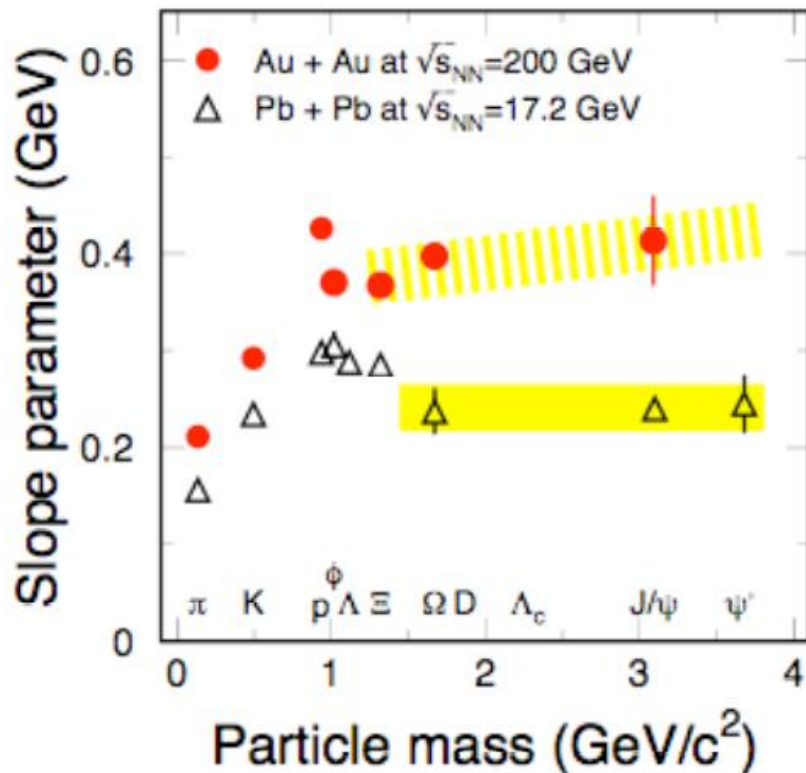
- chemical equilibrium
- simultaneous freeze-out of all particles

Cross-over transition at  $\mu_B \approx 0$  with pseudo-critical temperature of  $T_c = 156.5 \pm 1.5$  MeV

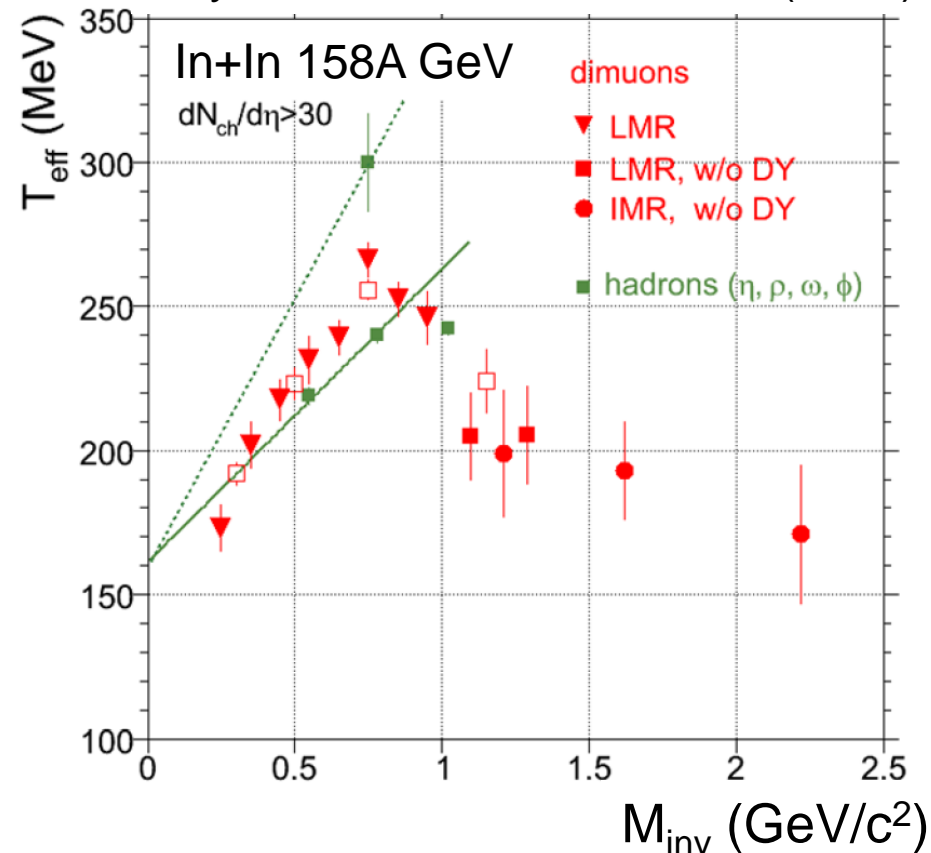


# Slope of transverse mass distributions of particles and lepton pairs

Nu Xu priv. comm.



R. Arnaldi et. al., NA60 Collaboration  
Phys. Rev. Lett. 100, 022302 (2008)



$M < 1 \text{ GeV}/c^2$ : radial flow generated in the late hadronic phase

$M > 1 \text{ GeV}/c^2$ : messengers from the early partonic phase ?

# CBM physics case and observables

## Phase transitions from partonic to hadronic matter

- excitation function of yields and slope parameters of hadrons (from  $\pi$  to  $\Omega$ ) and of lepton pairs (from  $M_{\text{inv}} = 0.2$  to  $\sim 3$  GeV)  $\rightarrow$  separation of radial flow generated in late hadronic phase and early partonic phase

# CBM physics case and observables

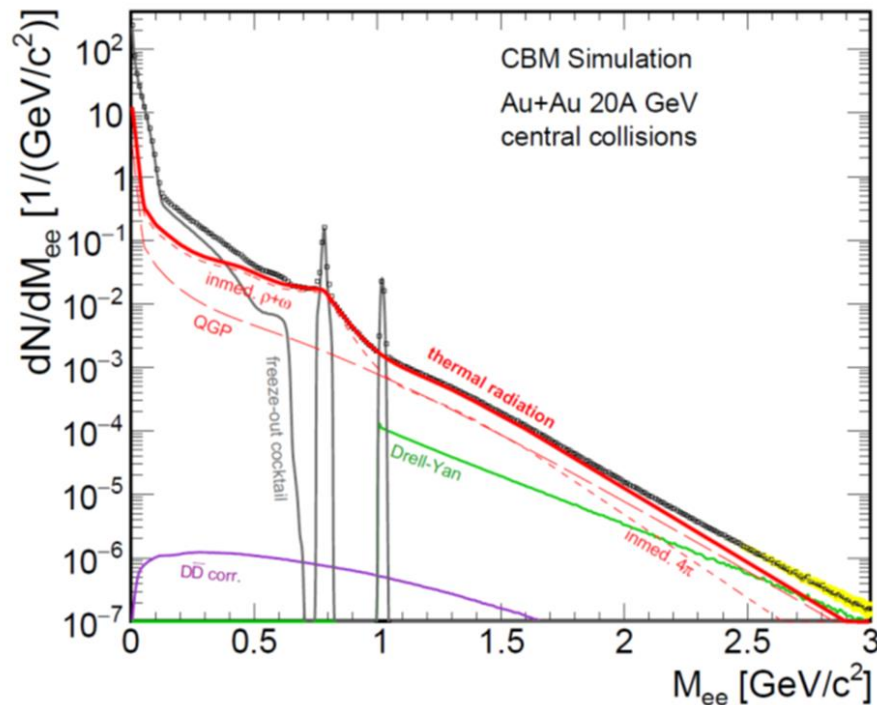
## Phase transitions from partonic to hadronic matter

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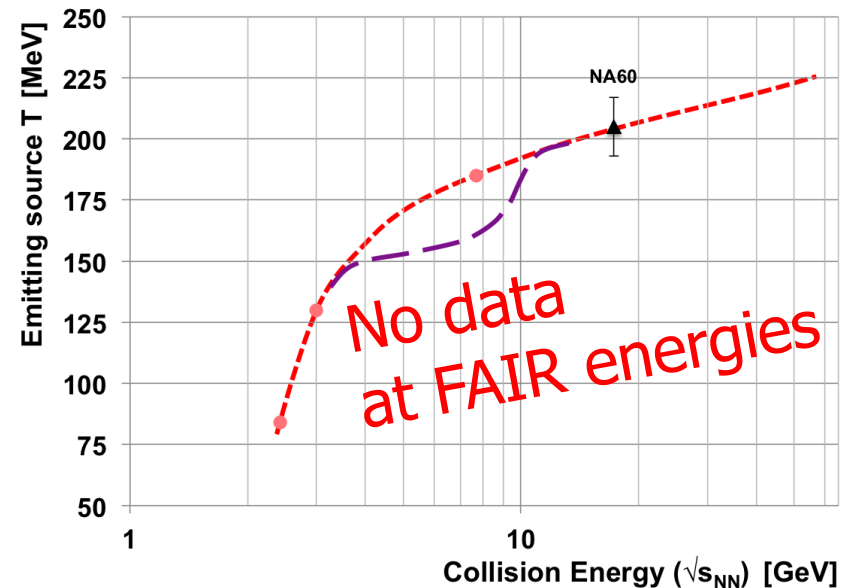
## Phase coexistence

- excitation function (invariant mass) of lepton pairs: thermal radiation from QGP, caloric curve

Invariant mass distribution of lepton pairs



Slope of dilepton invariant mass spectrum  
 $1 \text{ GeV}/c^2 < M_{\text{inv}} < 2.5 \text{ GeV}/c^2$





# CBM physics case and observables

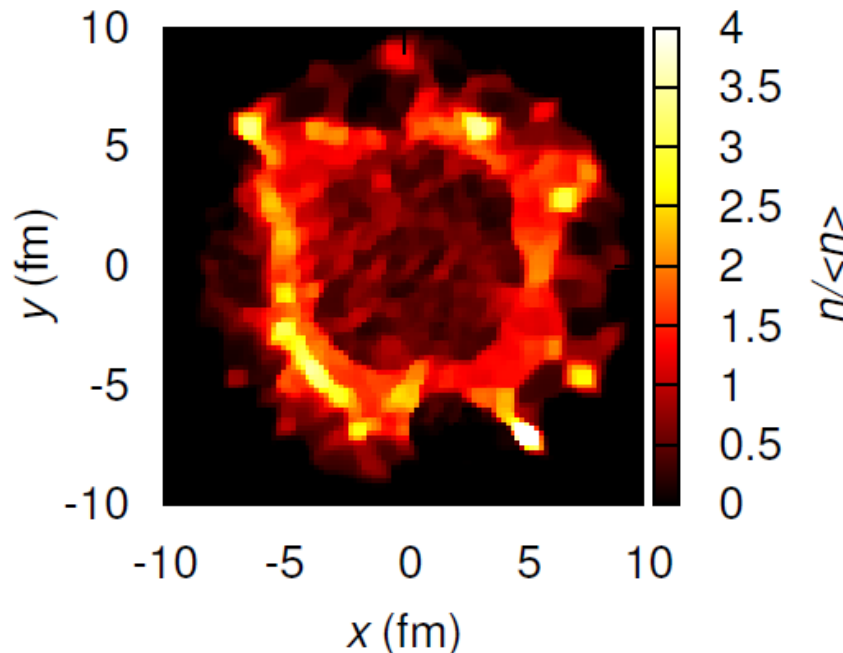
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## Phase coexistence

- excitation function (invariant mass) of lepton pairs: thermal radiation from QGP, caloric curve
- anisotropic azimuthal angle distributions: “spinodal decomposition”

Spinodal decomposition of the mixed phase: net baryon number density fluctuations



C. Herold, M. Nahrgang, I. Mishustin, M. Bleicher  
Nuclear Physics A 925 (2014) 14

Jan Steinheimer, Jorgen Randrup  
Phys. Rev. C 87, 054903 (2013)  
Eur. Phys. J. A (2016) 52: 239

No data  
at FAIR energies

# CBM physics case and observables

## Phase transitions from partonic to hadronic matter

- excitation function of yields and slope parameters of hadrons (from  $\pi$  to  $\Omega$ ) and of lepton pairs (from  $M_{\text{inv}} = 0.2$  to  $\sim 3$  GeV)  $\rightarrow$  separation of radial flow generated in late hadronic phase and early partonic phase

## Phase coexistence

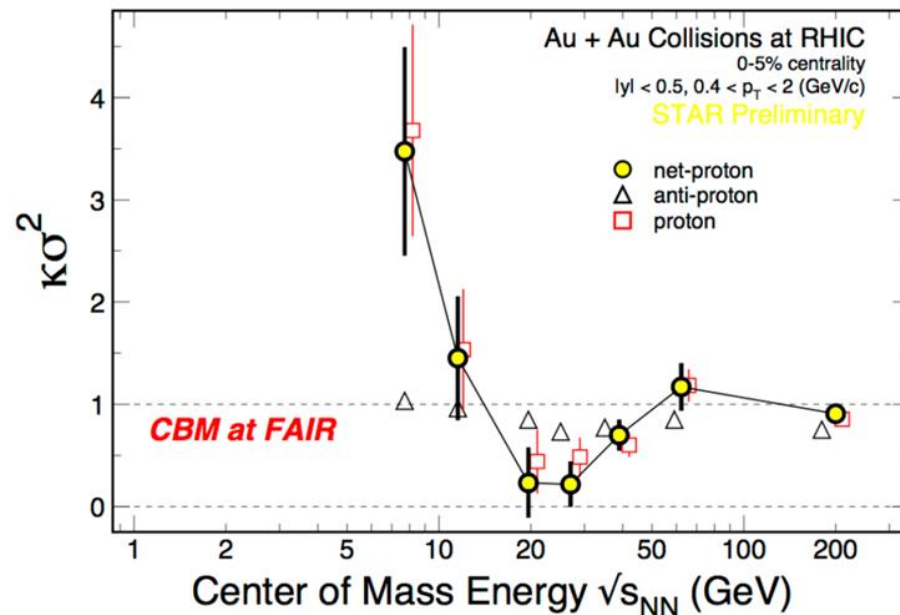
- excitation function (invariant mass) of lepton pairs: thermal radiation from QGP, caloric curve
- anisotropic azimuthal angle distributions: “spinodal decomposition”

## Critical point

- event-by-event fluctuations of conserved quantities (B,S,Q)  
“critical opalescence”

4<sup>th</sup> moment of net-proton  
multiplicity distribution:  
critical fluctuations

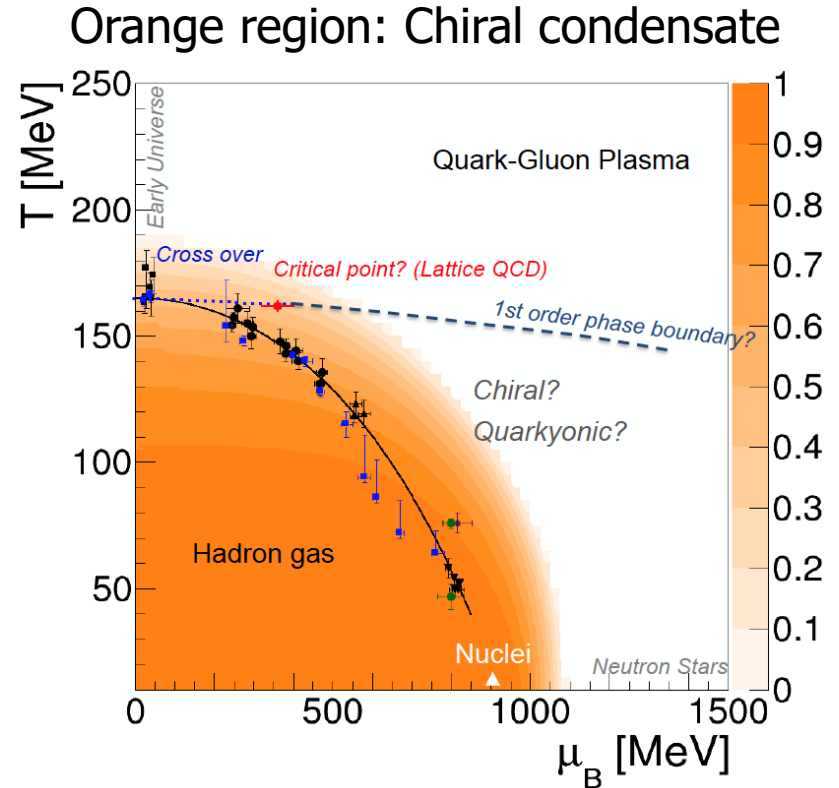
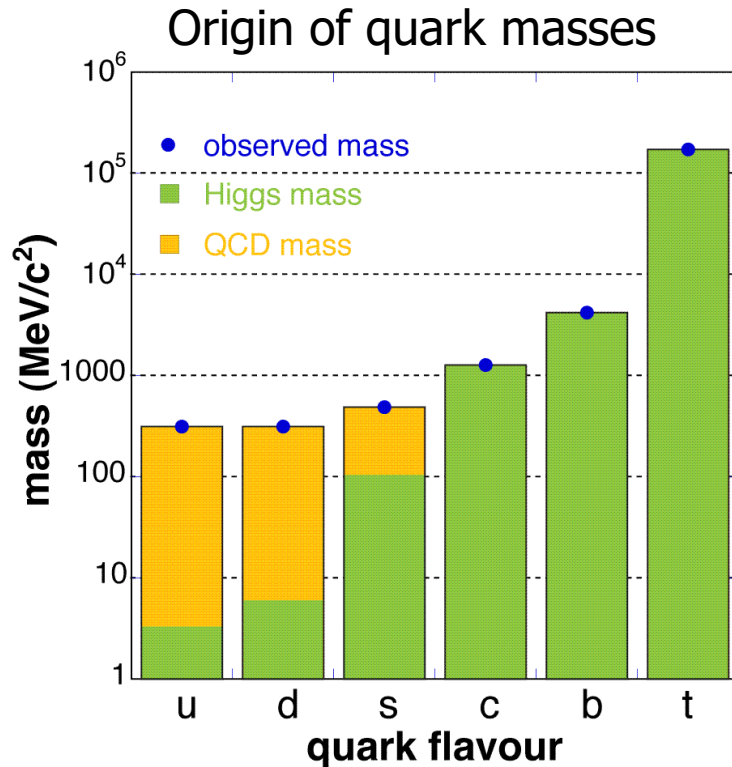
No data  
at FAIR energies



# CBM physics case and observables

Onset of chiral symmetry restoration at high  $\rho_B$

- in-medium modifications of hadrons:  $\rho, \omega, \phi \rightarrow e^+e^-(\mu^+\mu^-)$
- dileptons at intermediate invariant masses:  $4\pi \rightarrow \rho\text{-}a_1$  chiral mixing



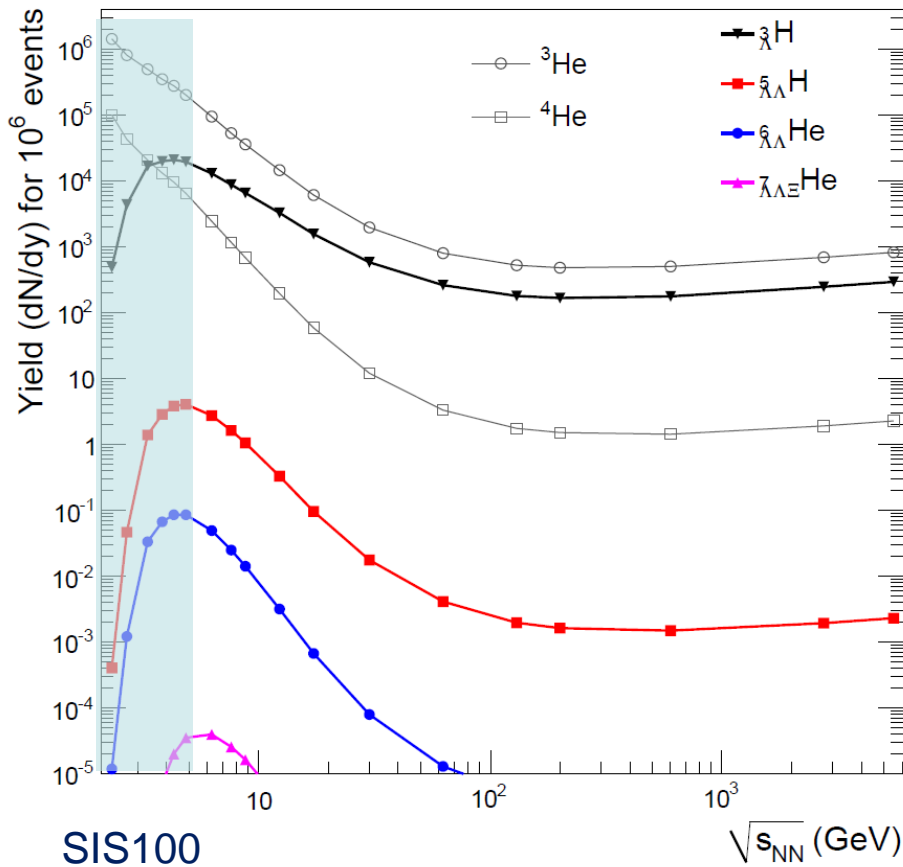
QCD mass generation by spontaneous/dynamical Chiral Symmetry breaking:  
Hadrons acquire mass by coupling to the virtual quark-antiquark pairs of the chiral condensate

# CBM physics case and observables

N- $\Lambda$ ,  $\Lambda$ - $\Lambda$  interaction, strange matter?

- (double-) lambda hypernuclei
- meta-stable objects (e.g. strange dibaryons)

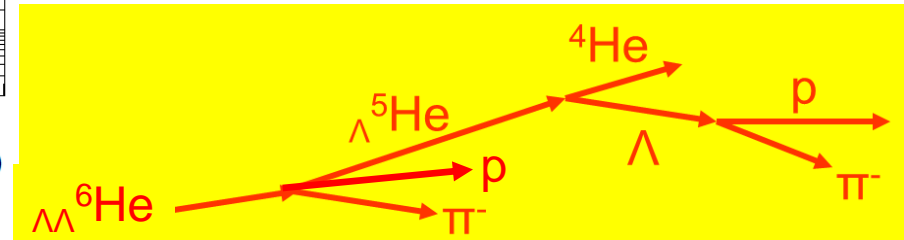
No data  
at FAIR energies



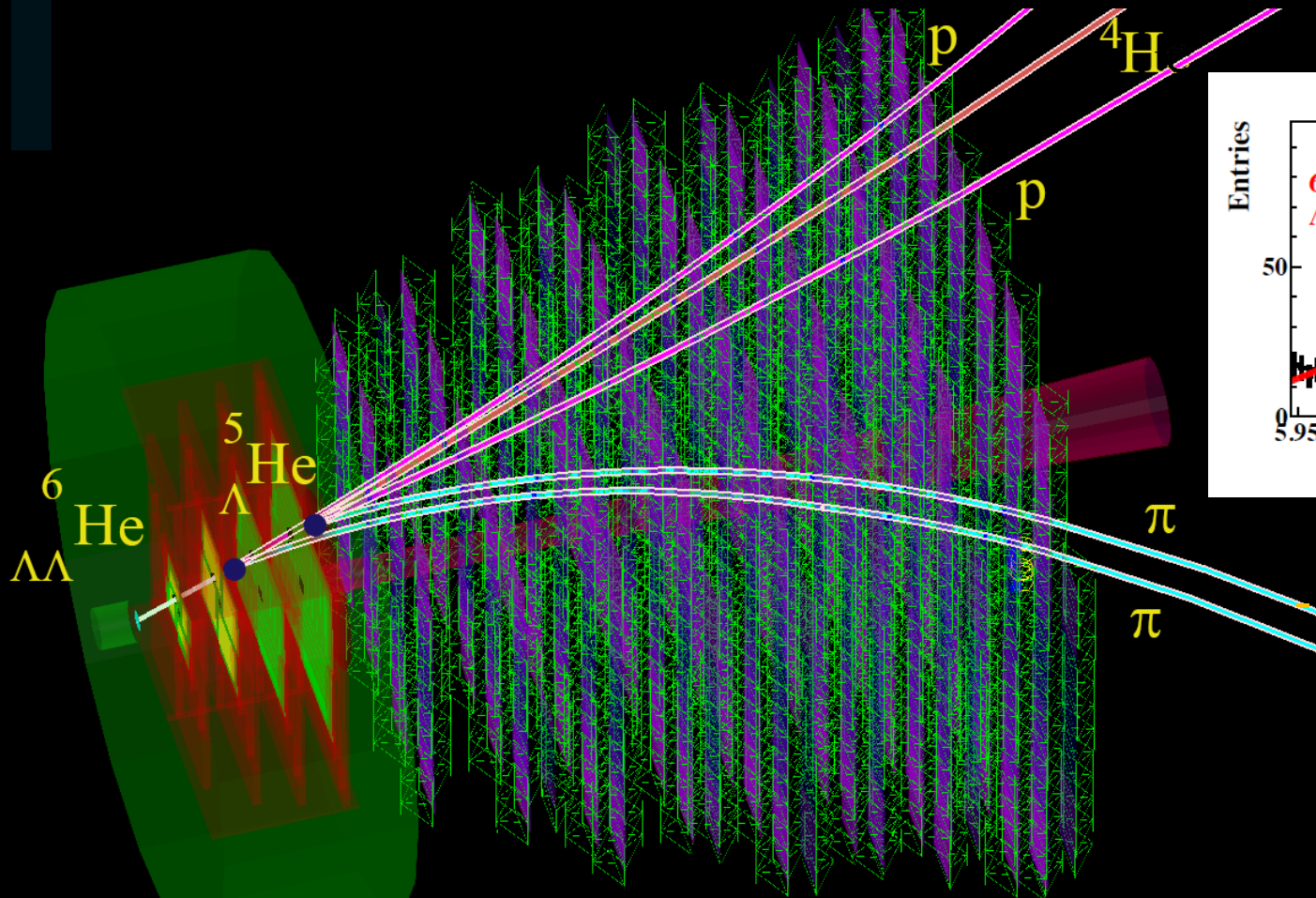
Central Au+Au collisions at 10 A GeV:

|                            | Multiplicity      | Yield in 1 week |
|----------------------------|-------------------|-----------------|
| $5\Lambda\Lambda\text{H}$  | $5 \cdot 10^{-6}$ | 3000            |
| $6\Lambda\Lambda\text{He}$ | $1 \cdot 10^{-7}$ | 60              |

Reaction Rate 1 MHz, BR 10%, efficiency 1%

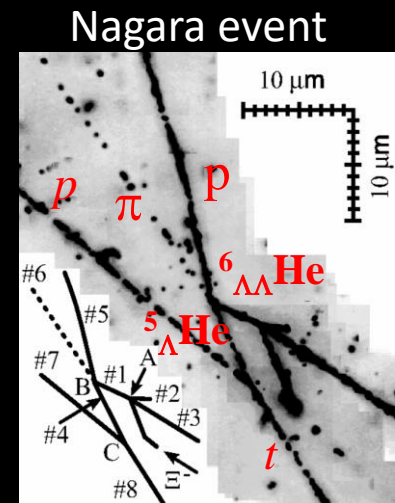
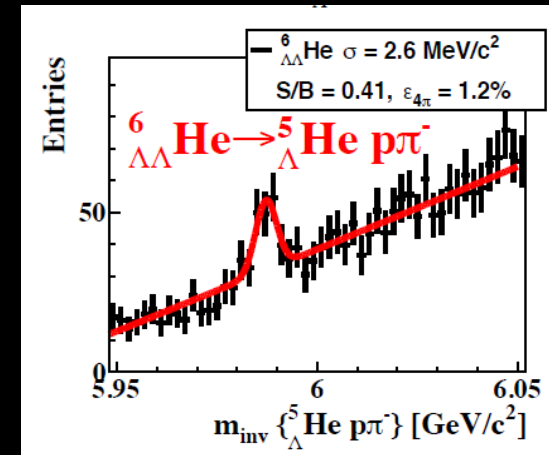


# $\Lambda\Lambda$ ${}^6\text{He}$ reconstruction in CBM detector



**Micro Vertex Detector**  
(4 layers of  
Monolithic Active  
Pixel Sensors)

**Silicon Tracking System**  
(8 layers of double-sided  
micro-strip sensors)

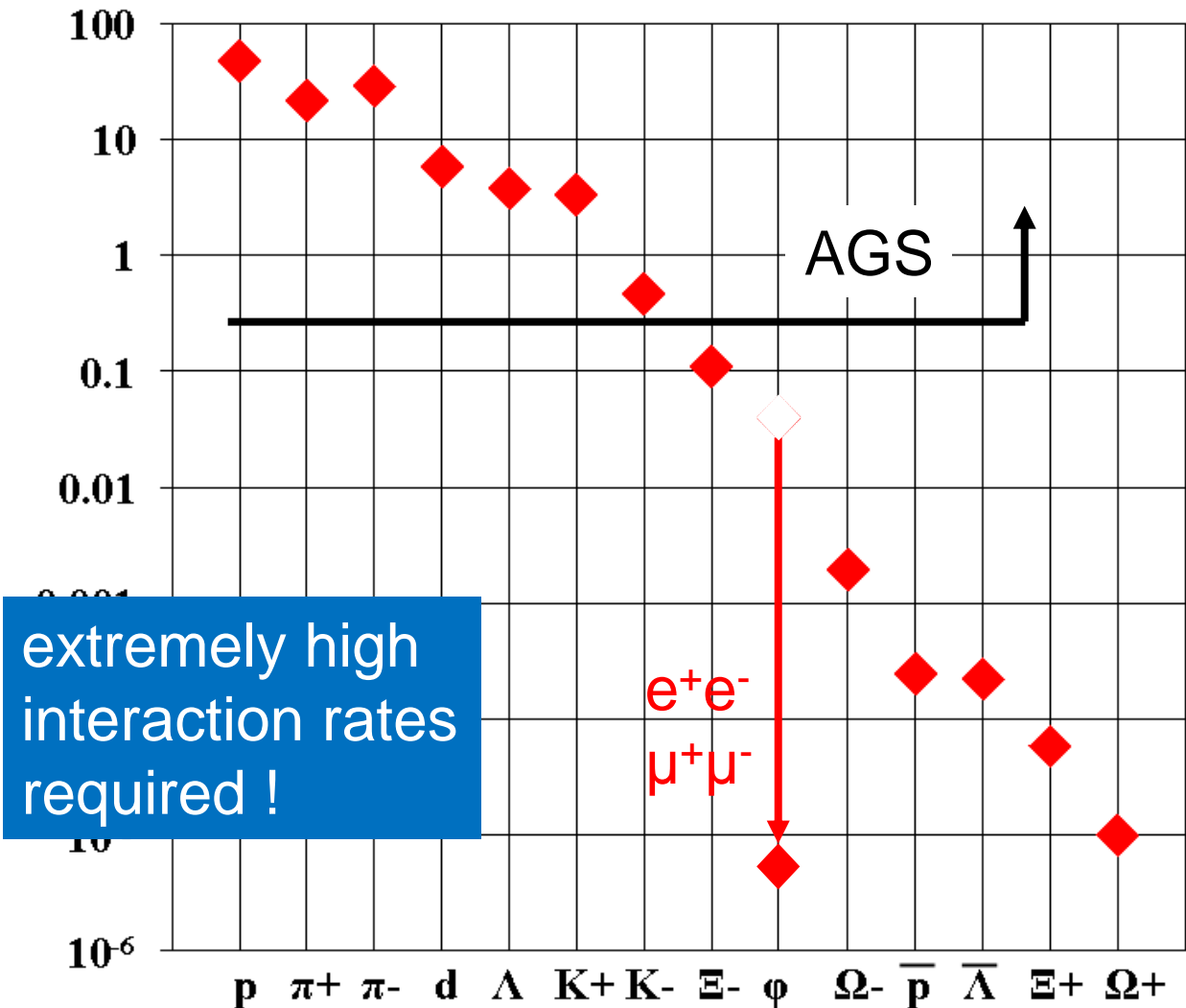




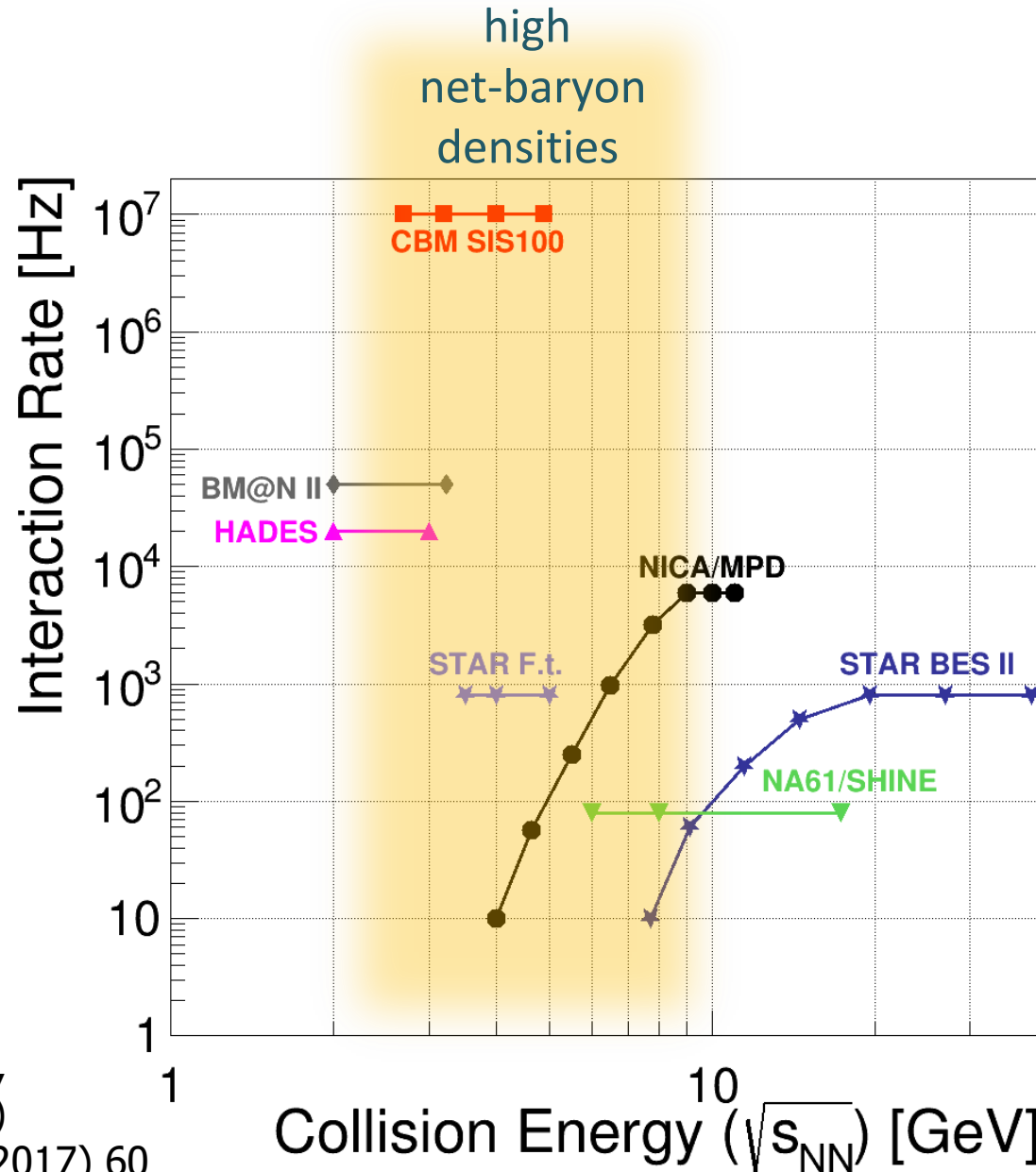
# Experimental challenges

## Particle yields in central Au+Au 4 A GeV

Multiplicity Statistical model, A. Andronic, priv. com.



# Experiments exploring dense QCD matter





# Experimental requirements

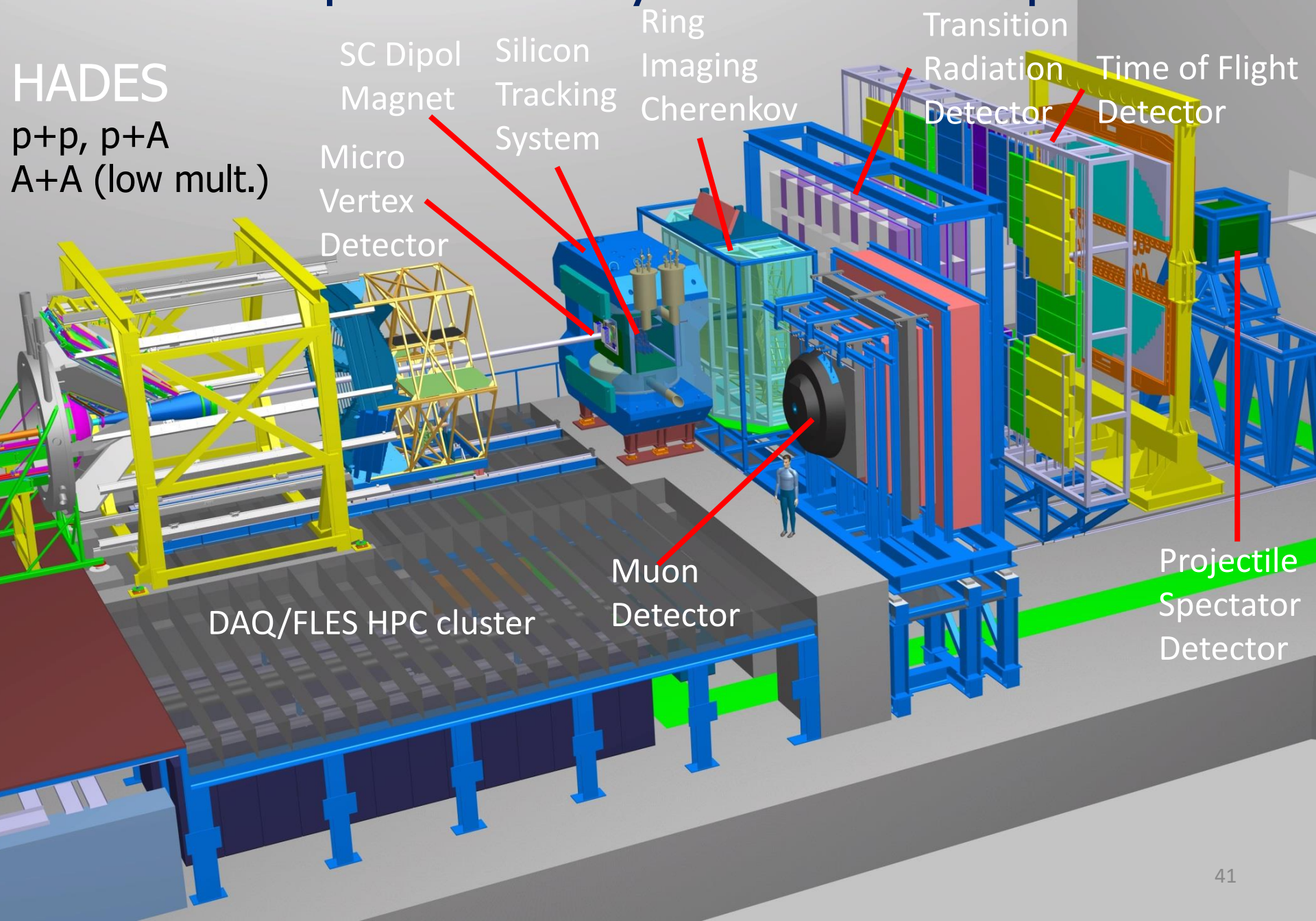
- $10^5 - 10^7$  Au+Au reactions/sec
- determination of displaced vertices ( $\sigma \approx 50 \mu\text{m}$ )
- identification of leptons and hadrons
- fast and radiation hard detectors and FEE
- free-streaming readout electronics
- high speed data acquisition and high performance computer farm for online event selection
- 4-D event reconstruction



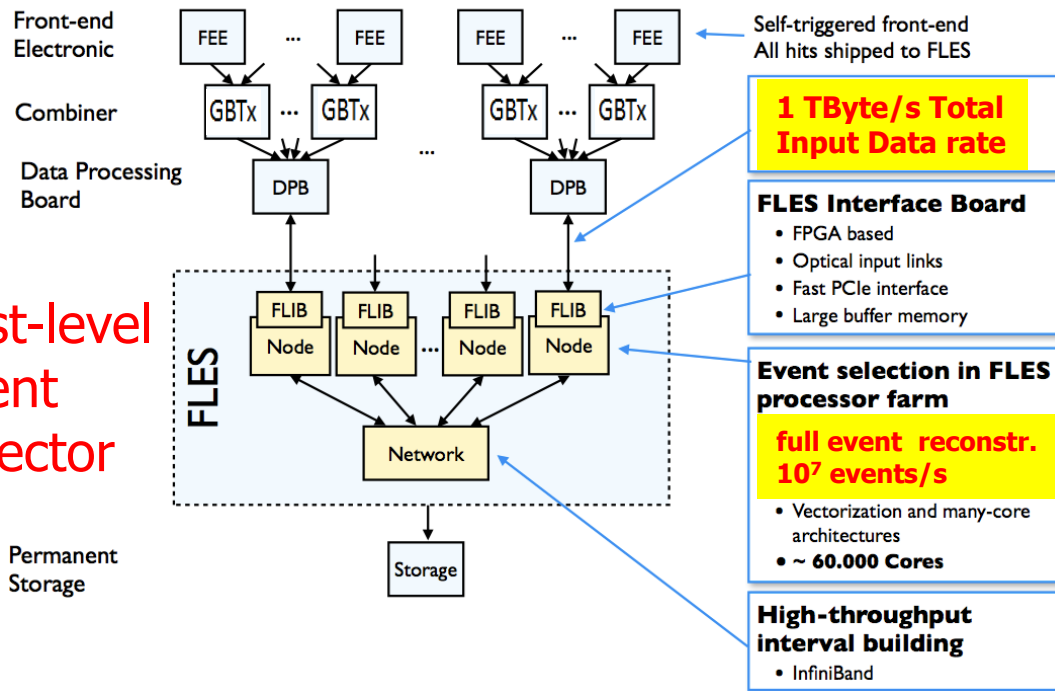
# The Compressed Baryonic Matter Experiment

HADES

p+p, p+A  
A+A (low mult.)



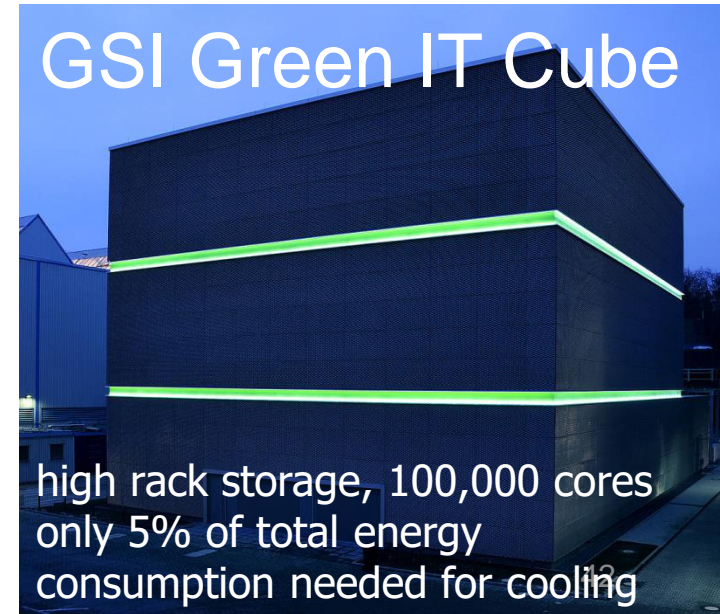
# CBM DAQ and online event selection



## Novel readout system:

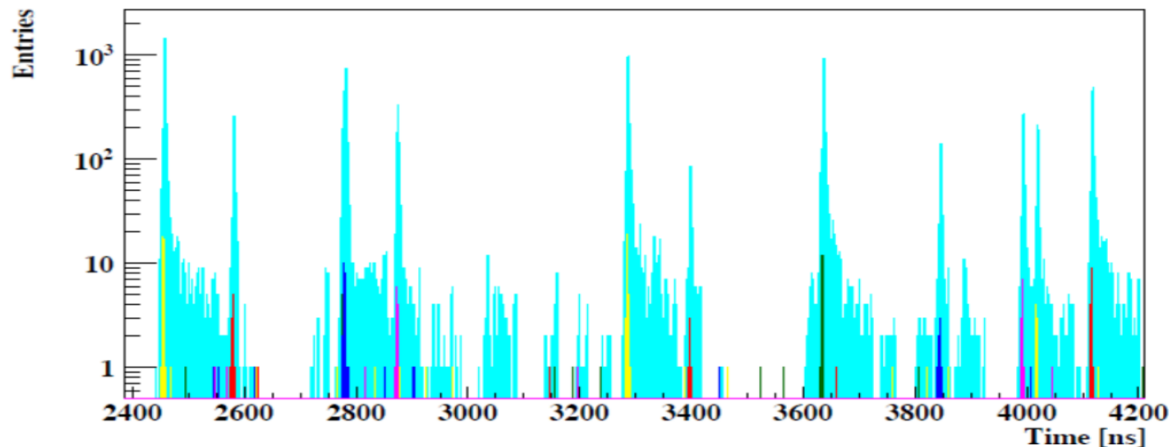
- no hardware trigger on events
- detector hits with time stamps
- full online 4-D track and event reconstruction.

## GSI Green IT Cube



high rack storage, 100,000 cores  
only 5% of total energy  
consumption needed for cooling

Hit and track time distribution for Au+Au 10A GeV collisions at 10 MHz (UrQMD)



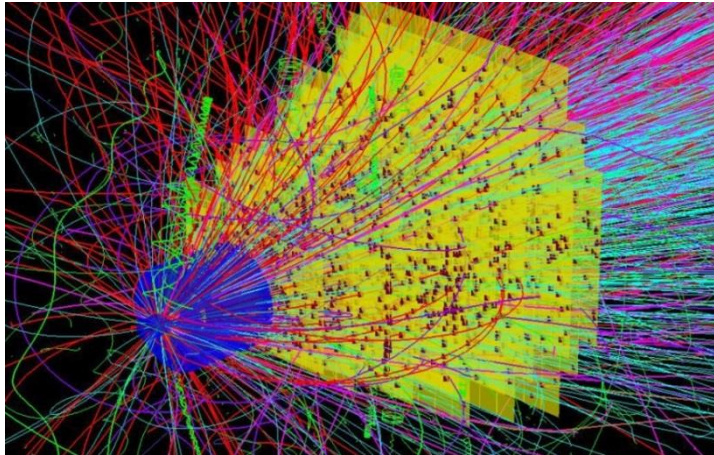


# Simulation and reconstruction

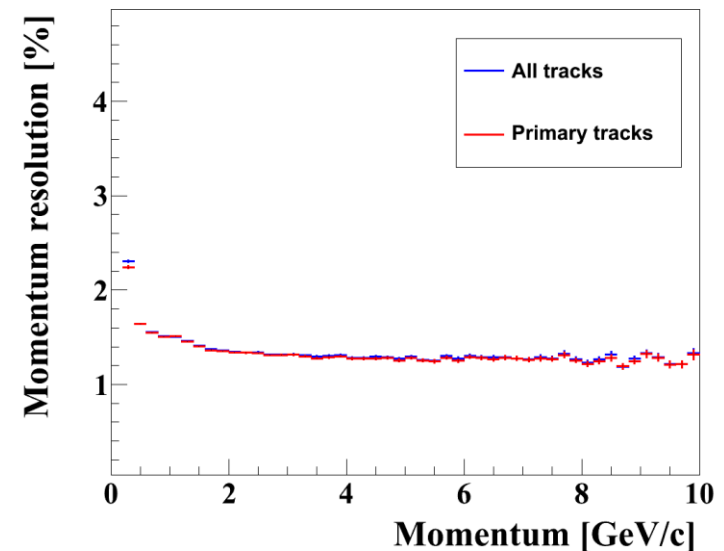
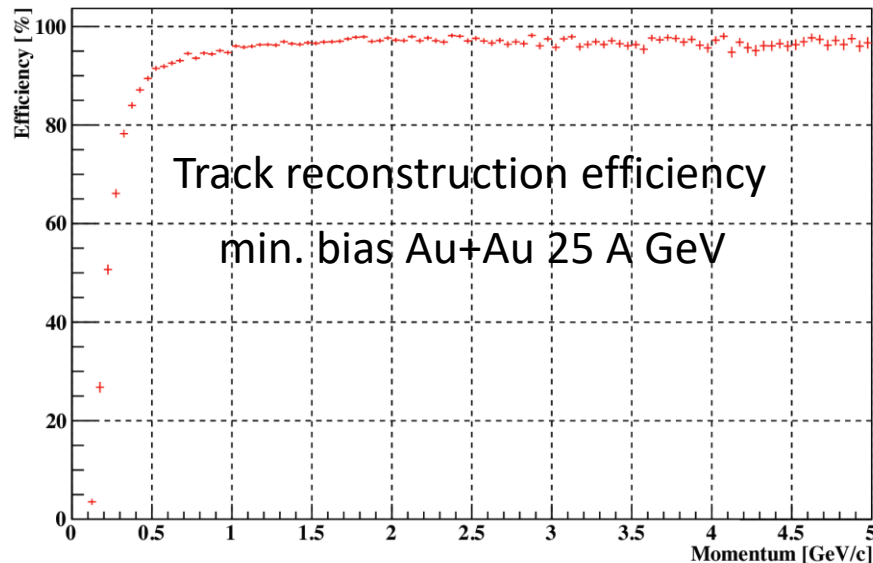
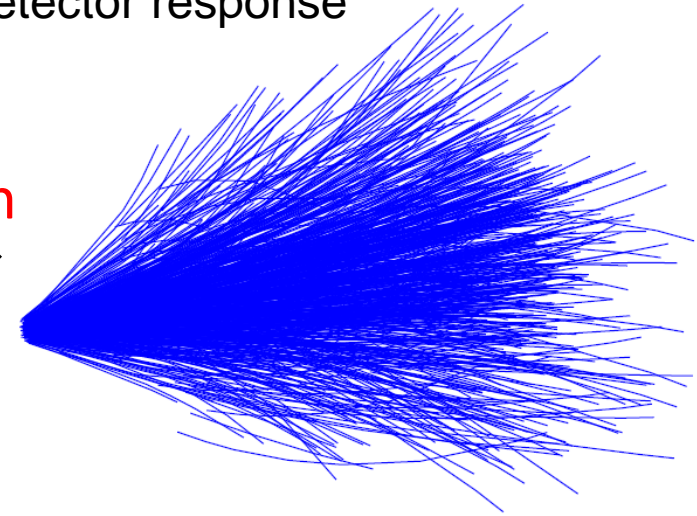
Event generators UrQMD 3.3

Transport code GEANT3, FLUKA

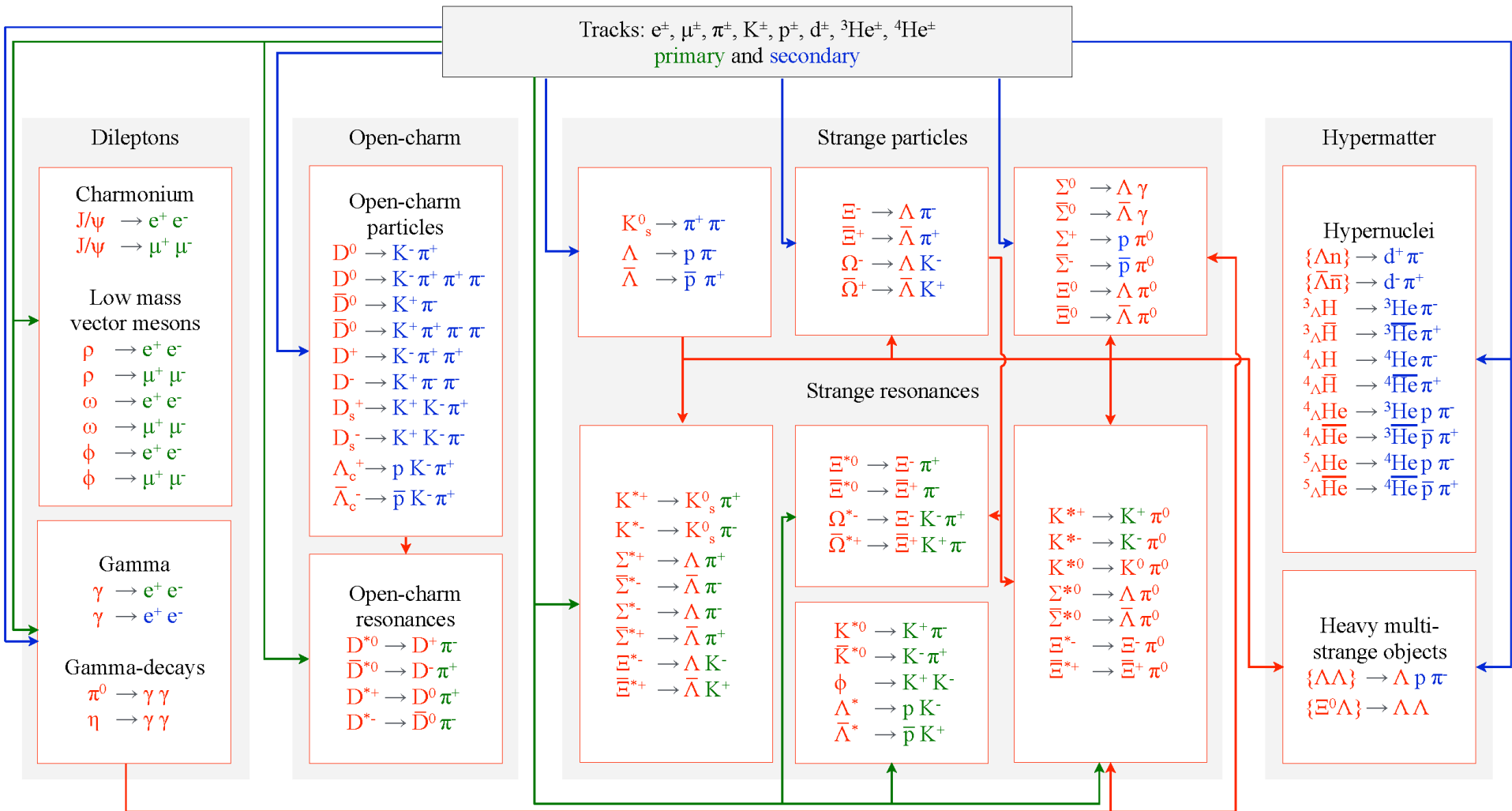
Realistic detector geometries, material budget and detector response



reconstruction



# Online particle identification in CBM: The KF Particle Finder



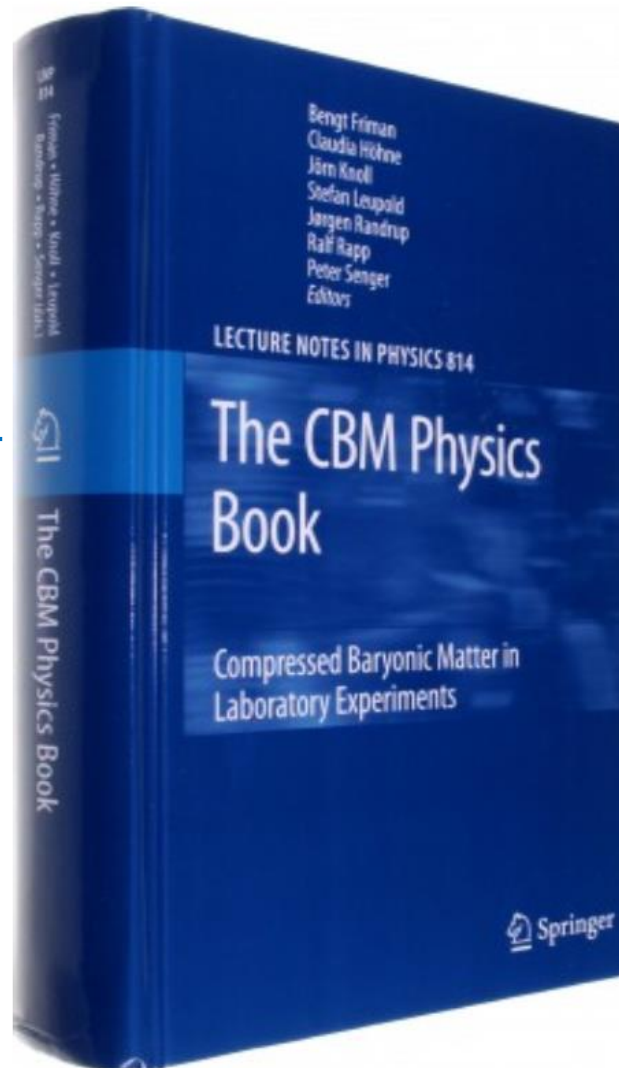
successfully used online in the STAR experiment at RHIC

# For further reading ...

## The CBM Physics Book

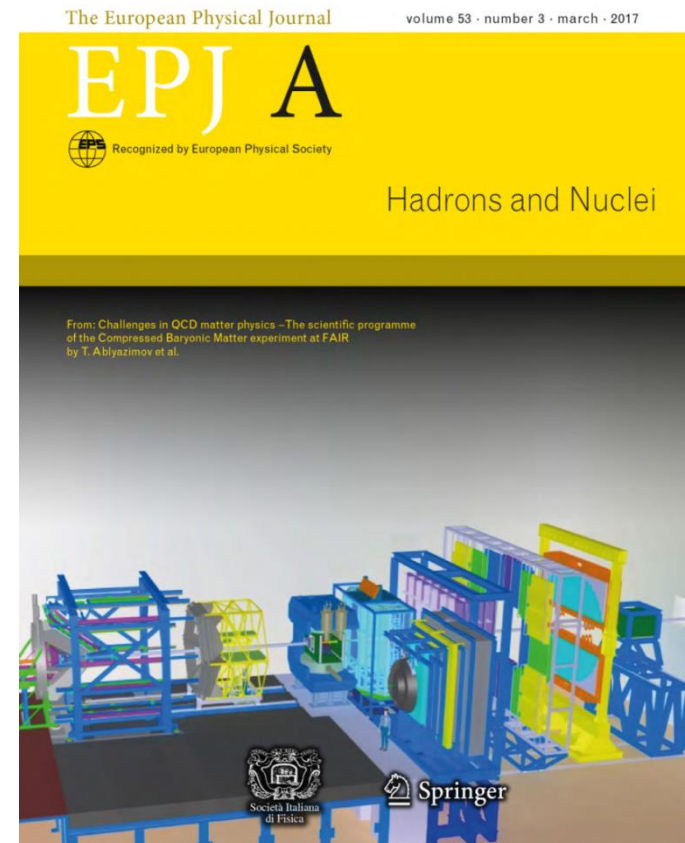
Foreword by  
Frank Wilczek

Springer Series:  
Lecture Notes in  
Physics, Vol. 814  
1<sup>st</sup> Edition.,  
2011, 960 p.,



Electronic Authors  
version:

<http://www.gsi.de/documents/DOC-2009-Sep-120-1.pdf>

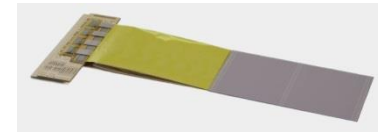
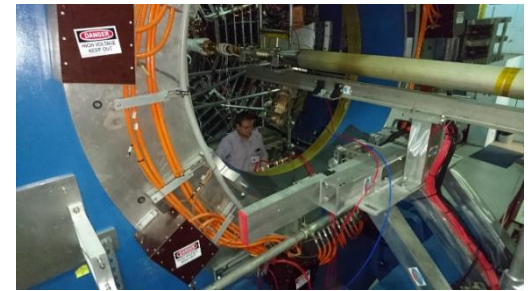
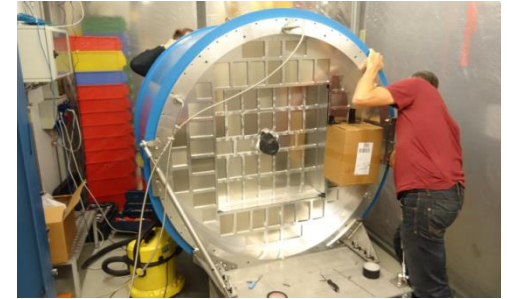


“Challenges in QCD Matter Physics – the scientific programme of the Compressed Baryonic Matter Experiment at FAIR”

Ablyazimov, et al. Eur. Phys. J. A (2017) 53

# FAIR phase 0 experiments on dense QCD matter

1. Install, commission and use 430 out of 1100 CBM RICH multi-anode photo-multipliers (MAPMT) in HADES RICH photon detector
2. Install, commission and use 10% of the CBM TOF modules including read-out chain at STAR/RHIC (BES II 2019/2020)
3. Install, commission and use 4 Silicon Tracking Stations and the Project Spectator Detector in the BM@N experiment at the Nuclotron in JINR/Dubna (start 2019 with Au-beams up to 4.5 A GeV)
4. Build miniCBM at GSI/SIS18 for a full system test with high-rate nucleus-nucleus collisions from 2018 - 2021





# The CBM Collaboration: 56 institutions, > 460 members

## China:

CCNU Wuhan  
Tsinghua Univ.  
USTC Hefei  
CTGU Yichang  
Chongqing Univ.

## Czech Republic:

CAS, Rez  
Techn. Univ. Prague

## France:

IPHC Strasbourg

## Hungary:

KFKI Budapest  
Eötvös Univ.

## Germany:

Darmstadt TU  
FAIR  
Frankfurt Univ. IKF  
Frankfurt Univ. FIAS  
Frankfurt Univ. ICS  
GSI Darmstadt  
Giessen Univ.  
Heidelberg Univ. P.I.  
Heidelberg Univ. ZITI  
HZ Dresden-Rossendorf  
KIT Karlsruhe  
Münster Univ.  
Tübingen Univ.  
Wuppertal Univ.  
ZIB Berlin

## India:

Aligarh Muslim Univ.  
Bose Inst. Kolkata  
Panjab Univ.  
Rajasthan Univ.  
Univ. of Jammu  
Univ. of Kashmir  
Univ. of Calcutta  
B.H. Univ. Varanasi  
VECC Kolkata  
IOP Bhubaneswar  
IIT Kharagpur  
IIT Indore  
Gauhati Univ.

## Korea:

Pusan Nat. Univ.

## Poland:

AGH Krakow  
Jag. Univ. Krakow  
Warsaw Univ.  
Warsaw Univ. Tech.

## Romania:

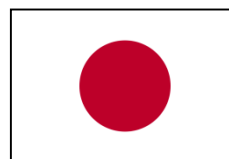
NIPNE Bucharest  
Univ. Bucharest

## Russia:

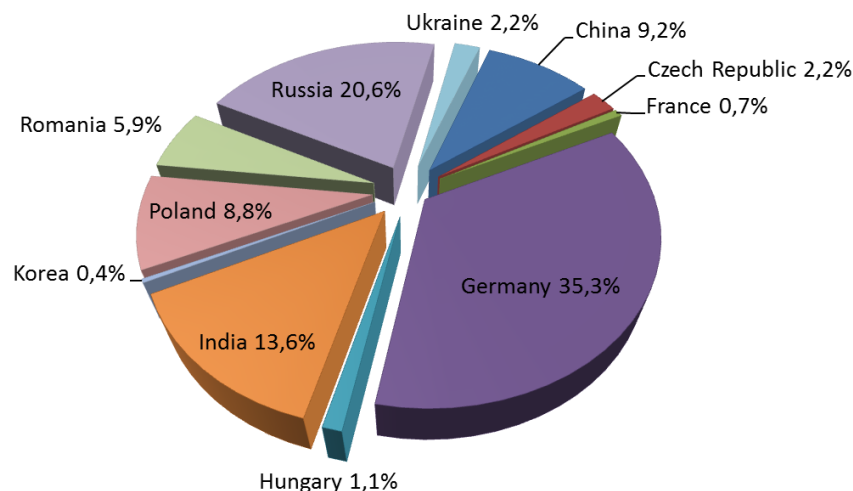
IHEP Protvino  
INR Troitzk  
ITEP Moscow  
Kurchatov Inst., Moscow  
VBLHEP, JINR Dubna  
LIT, JINR Dubna  
MEPHI Moscow  
PNPI Gatchina  
SINP MSU, Moscow

## Ukraine:

T. Shevchenko Univ. Kiev  
Kiev Inst. Nucl. Research



## CBM Scientists





# FAIR Project Status 2018



- Comprehensive civil construction plan: completion of all buildings by 2022
- Full integrated planning for construction and commissioning of the entire project: Completion of the full FAIR facility by 2025.
- Civil construction as well as procurement of accelerators and realization of experiments are progressing well.



**2014: 1350 pillars 60 m deep**



**Ground breaking - 4 July 2017**







# FAIR tunnel excavation August 2018





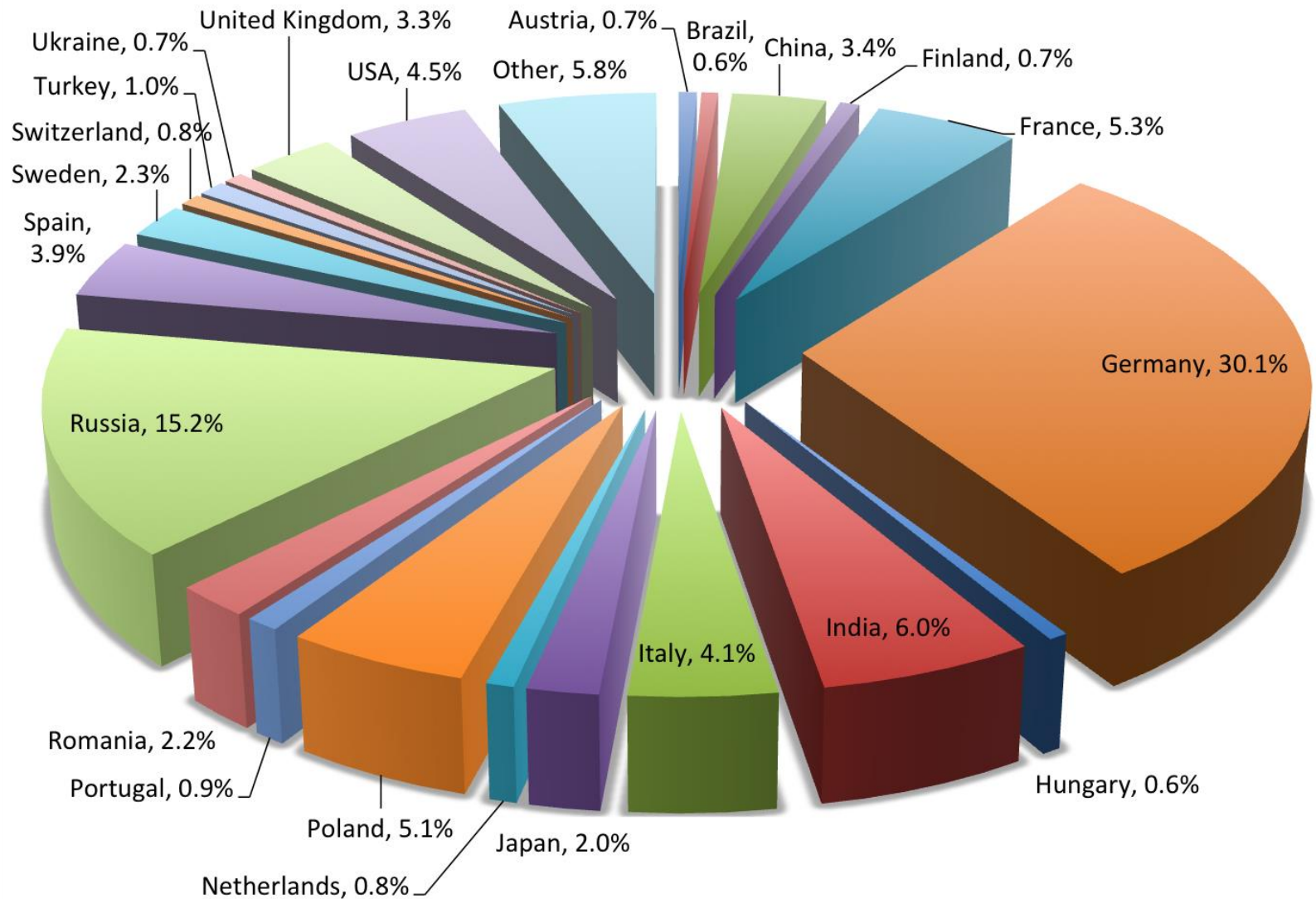
# FAIR civil construction status Sept. 2018



EXCAVATION PIT FOR  
CBM EXPERIMENT



# FAIR Collaboration Members by Country





# Instead of a summary: NuPECC recommendation

Key Summary Recommendation of the NuPECC Long Range Plan 2017 presented in Brussels on Nov 27<sup>th</sup> :

**Complete urgently the construction of the ESFRI\* flagship FAIR and develop and bring into operation the experimental program of its four scientific pillars APPA, CBM, NUSTAR and PANDA.**

FAIR is a European flagship facility for the coming decades. Worldwide unique it will allow for a large variety of unprecedented fore-front research in physics and applied science. It focuses on the structure and evolution of matter. Its multi-faceted research opens a new era in our understanding of the fundamental building blocks of matter and the forces as well as of the evolution of our Universe: the new possibilities for research in Darmstadt are unique and are expected to produce ground breaking new insights for nuclear research.

\*European Strategy Forum on Research Infrastructures

