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EMMI RRTF

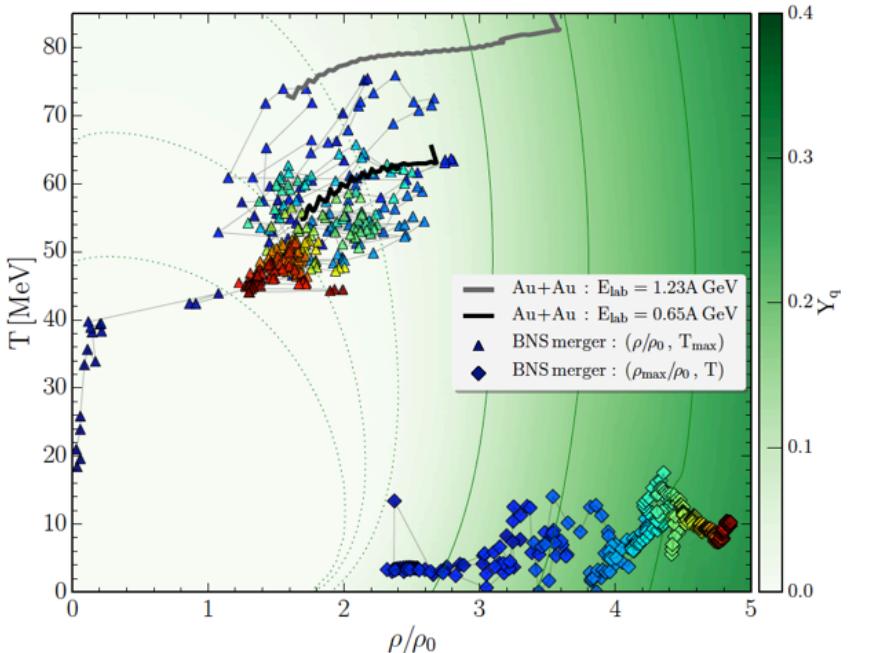
The physics of neutron star mergers at GSI/FAIR

Heavy-ion collisions - Discussion session III

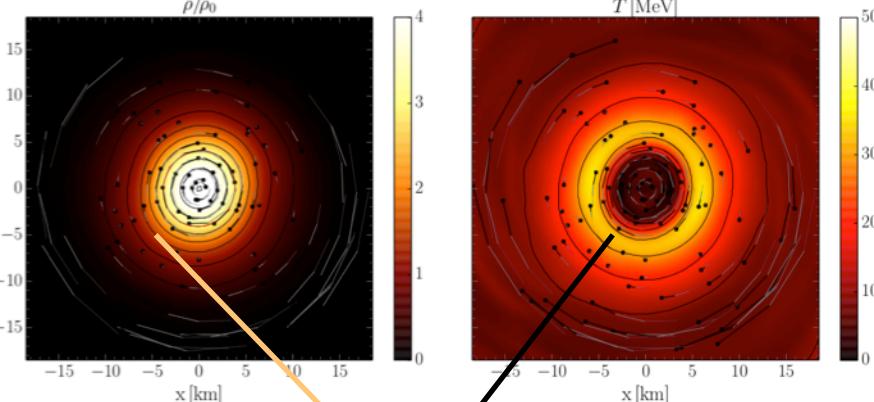
Tetyana Galatyuk TU Darmstadt / GSI

Low beam energy HIC compared to NS merger simulations

- Chiral Mean Field Model enables to treat heavy-ion collisions and NS mergers on the same footing



A dense and cold core with a hot hadronic corona



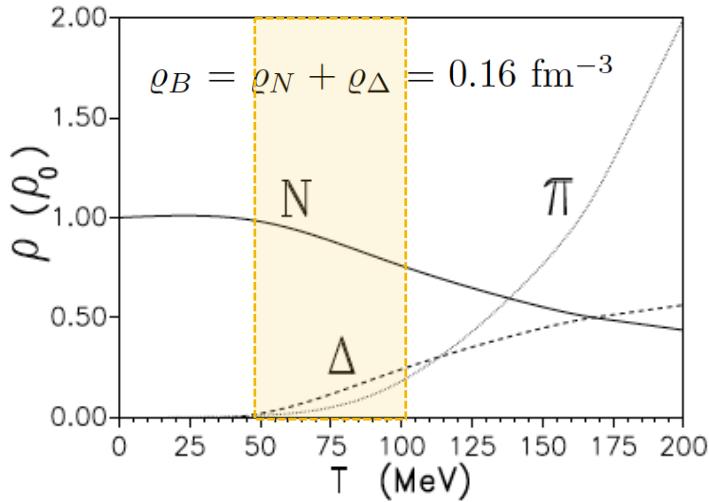
NS mergers and HIC at $E_{\text{lab}} < 2A$ GeV probe the same region in the phase diagram

Can HIC help to constrain the EoS?

See J. Steinheimer, Mo 11 Jun

Baryonic matter at 1-2A GeV beam energy

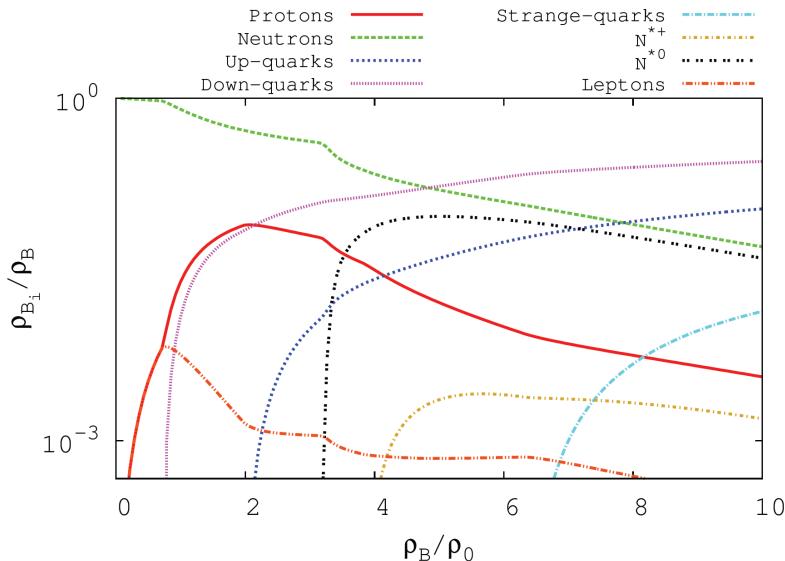
Composition of a hot $\pi\Delta N$ gas



Rapp, Wambach, Adv.Nucl.Phys. 25 (2000)

Isospin?

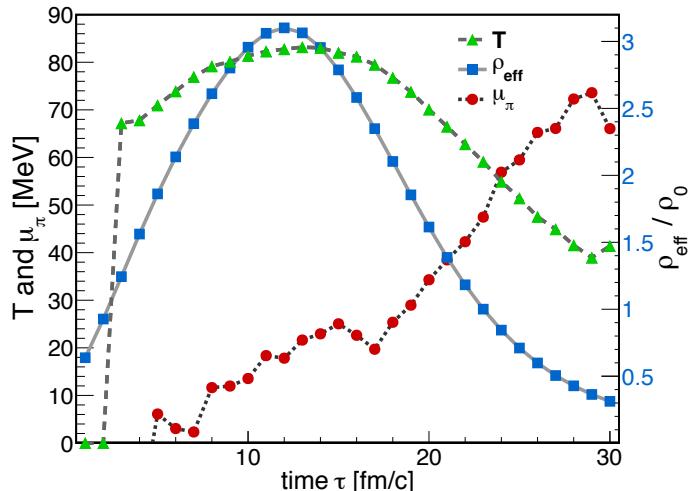
Chemical composition of neutron star matter
A smooth transition from hadrons to quarks



A. Mukherjee, S. Schramm, J. Steinheimer and V. Dexheimer,
Astron. Astrophys. 608, A110 (2017)

Evolution of HIC and NS merger

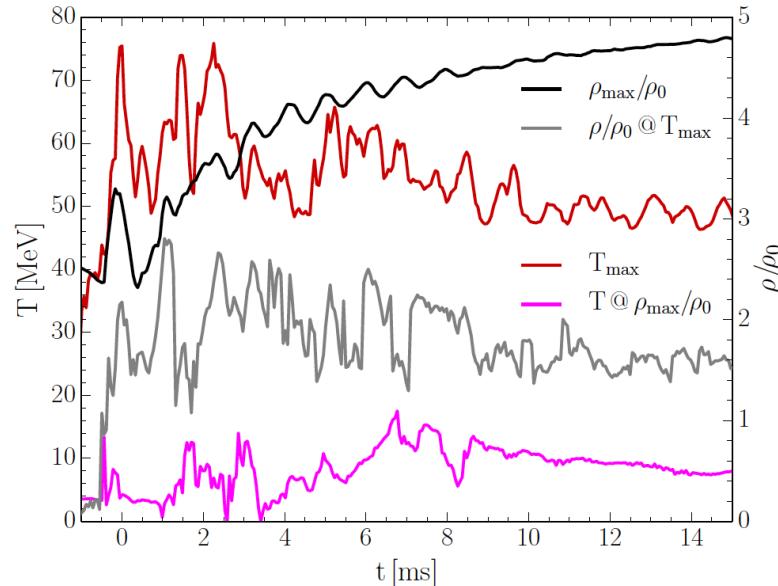
Central cell (3x3x3 fm³) thermodynamic properties from coarse graining UrQMD



TG, F. Seck, R. Rapp, J. Stroth, Eur. Phys. J. A 52 (2016) 131

Thermalization?

Evolution of the central region in a binary NS merger for CMF equation-of-state (see s.1)



M. G. Alford, L. Bovard, M. Hanuske, L. Rezzolla, K. Schwenzer.,
Phys. Rev. Lett. 120, 041101 and Phys. Rev. D 96, 124005,
M. Hanuske (priv. com.)

Observables

- Elliptic flow of p, d, t from 0.25 to 1.5A GeV
 - Compressibility (model dependent way) → "soft" EoS
 - [UrQMD3.4 P. Hillmann, J. Steinheimer, M. Bleicher, arXiv:1802.01951](#) → "hard" EoS
- Differential elliptic flow v_2 of n/p
 - Symmetry energy at supra-normal densities (towards model invariance: tested stability with different models)
- Particle production (below NN production threshold), i.e. K^+ (Au/C), K_s^0/K^+
 - Compressibility (model dependent way) → "soft" EoS
- Particle spectra
 - Transverse momentum distribution of i.e. K_s^0 → support in-medium repulsive vector K^0 potential ~40 MeV
- YN, YNN interactions ($Y = \Lambda, \Xi$)
 - Femtoscopy (ΛN correlation function, HADES collab., PRC 94 (2016) no.2, 025201)
- In-medium characteristics of particles
 - effective masses, spectral functions, decay widths



Measurable in pp, pA, AA

- Compare updated and new models against all available sets
- Apply machine learning methods to determine parameters in the model(s) consistently?

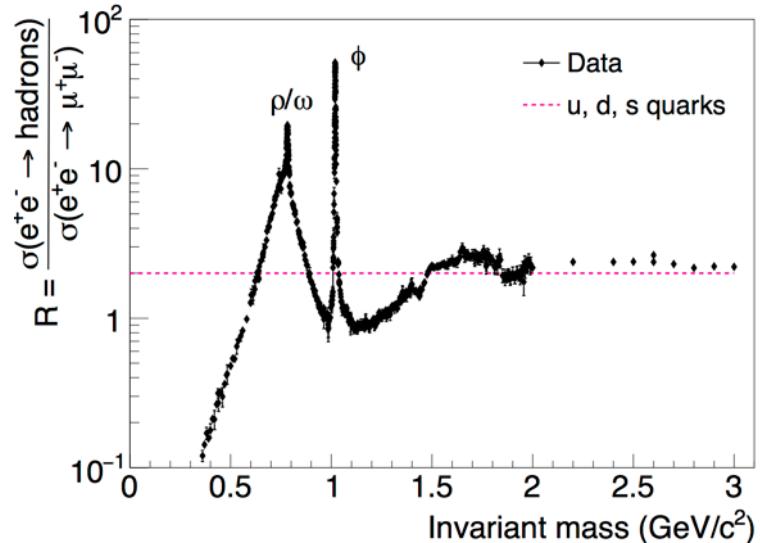
Electromagnetic correlator in vacuum

McLerran-Toimela formula

$$\frac{dN_{ll}}{d^4x d^4q} = \frac{-\alpha_{EM}^2}{\pi^3 M^2} f^B(q \cdot u; T) \text{Im} \Pi_{EM}^{\mu\nu}(M, q; u_B, T)$$

$$R = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)} \propto \frac{\text{Im} \Pi_{EM}^{\text{vac}}}{M^2}$$

- If $\frac{\text{Im} \Pi_{EM}^{\mu\nu}}{M^2} \sim \text{const.}$ thermal dilepton emission will follow a Bose distribution
- Dileptons as thermometer



$$\text{Im} \Pi_{EM}^{\text{vac}}(M \leq 1.5 \text{ GeV}) = \sum_{\nu=\rho,\omega,\phi} \left(\frac{m_\nu^2}{g_\nu} \right)^2 \text{Im} D_\nu^{\text{vac}}(M)$$



Vector Meson Dominance: $J^\rho = 1^-$ for both γ^* and VM (with ρ playing a dominant role)

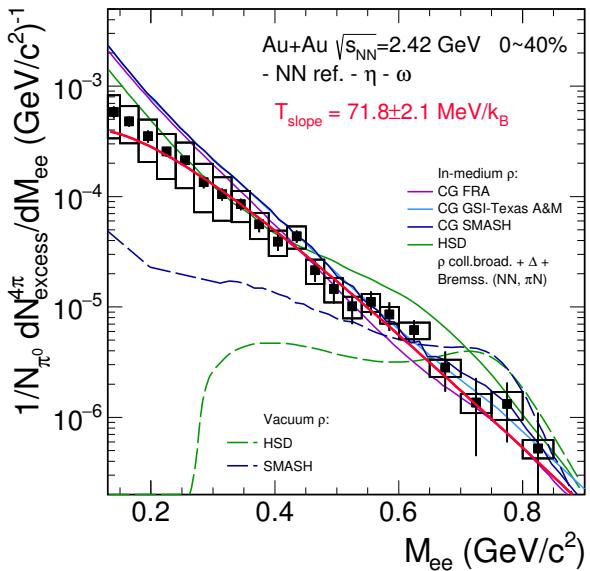
$$\text{Im} \Pi_{EM}^{\text{vac}}(M > 1.5 \text{ GeV}) = -\frac{M^2}{12\pi} \left(1 + \frac{\alpha_s(M)}{\pi} + \dots \right) N_c \sum_{q=u,d,s} (e_q)^2$$



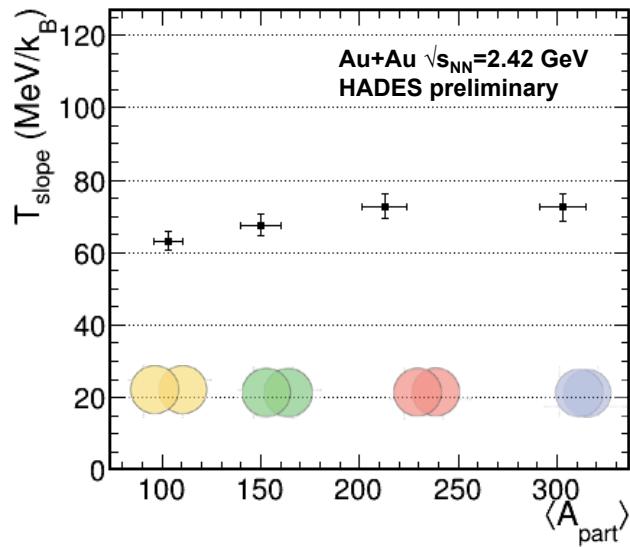
pQCD continuum

Results from HADES at SIS18, GSI

Emissivity of matter created in Au+Au collisions at 1.23A GeV

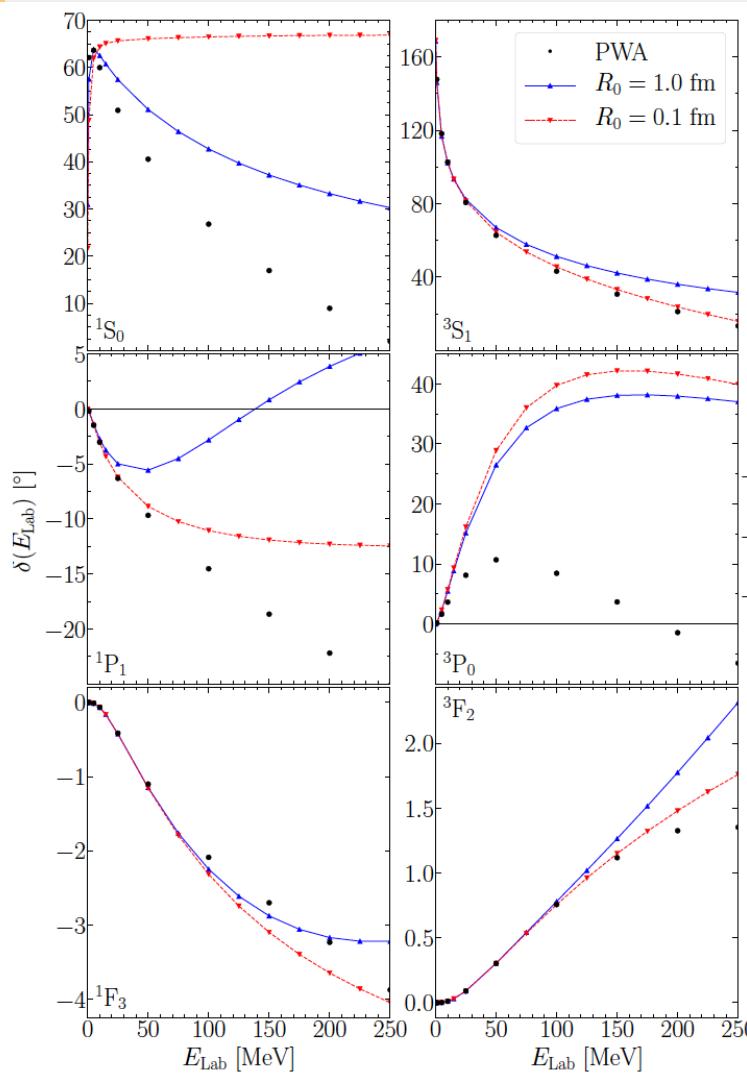


- Mass spectrum falling exponentially
→ "Planck-like" → $T = 72 \pm 2 \text{ MeV}$
- Agreement with coarse-grained approach
 $\rho/\rho_0 = 2 - 3$



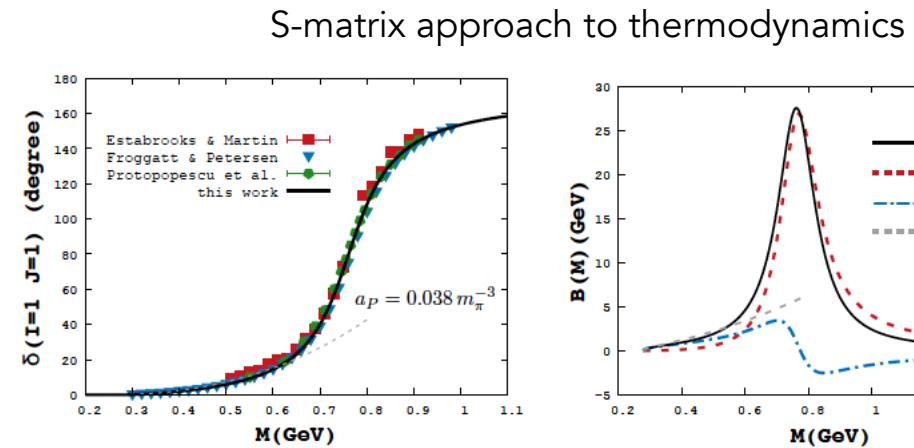
- Control T, density and pressure by varying centrality, system size, energy
- **Measure γ^* radiation in Au+Au collisions at 0.6 – 0.8A GeV**

→ Start from comparing spectral distributions
(i.e. @ $T=40 \text{ MeV}$, $\rho/\rho_0 = 1.5-2$)
→ go for EoS



Phase shifts as function of the laboratory energy

I. Tews et al.,
arXiv: 1806.00233v1 [nucl-th]

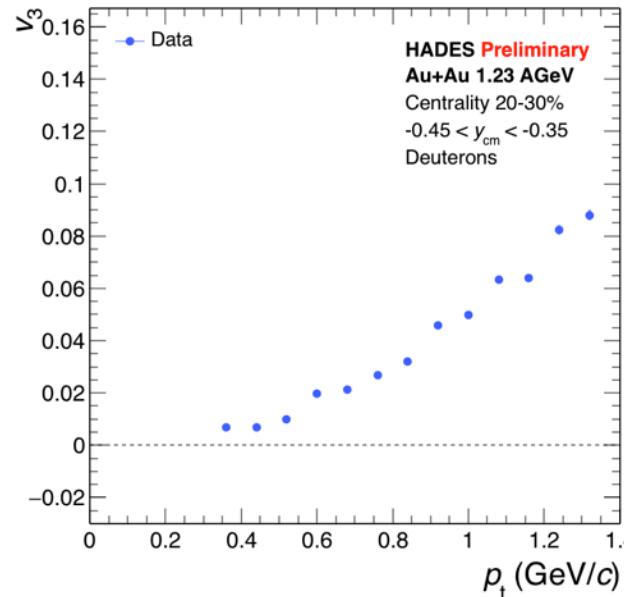
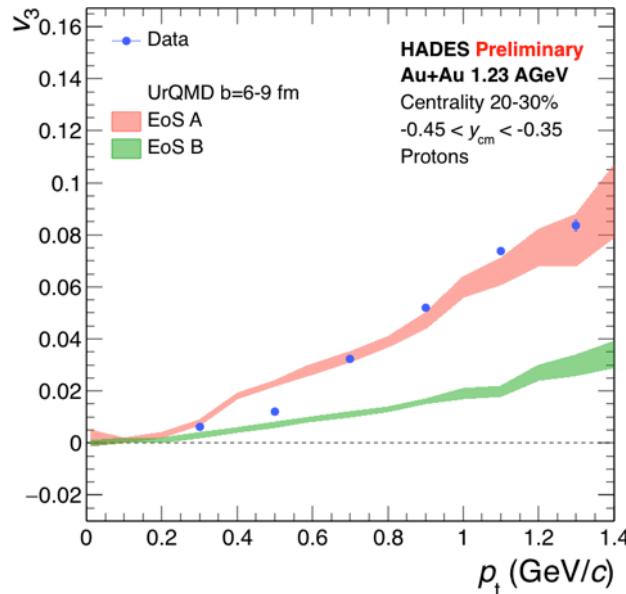


Pok Man Lo, arXiv:1707.04490v2 [hep-ph]

Bonus slides

Triangular flow $v_3\{\psi_{RP}\}$ and EoS at SIS18

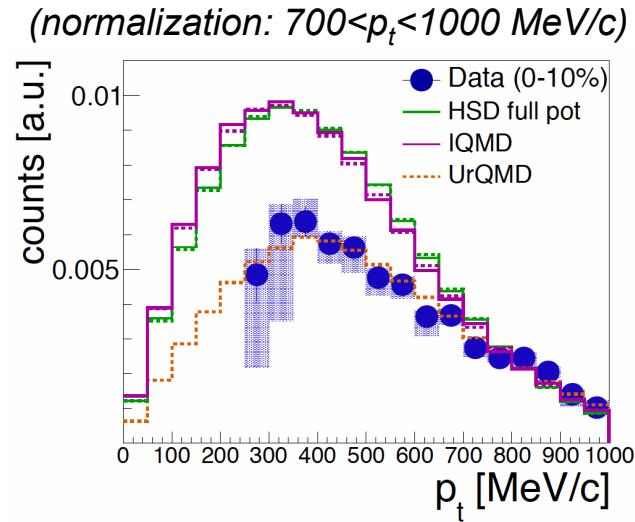
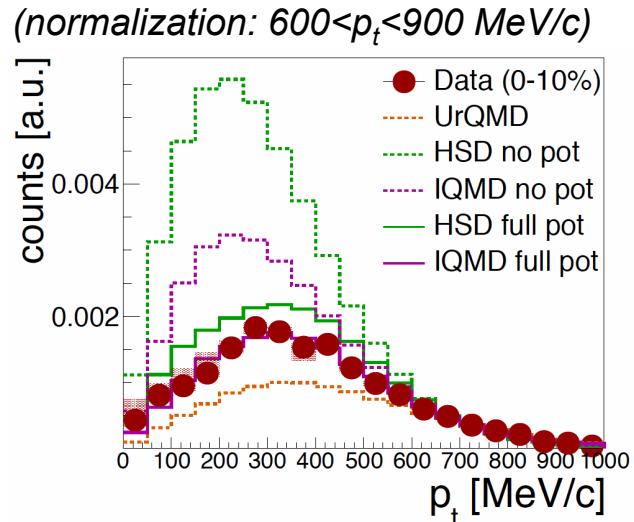
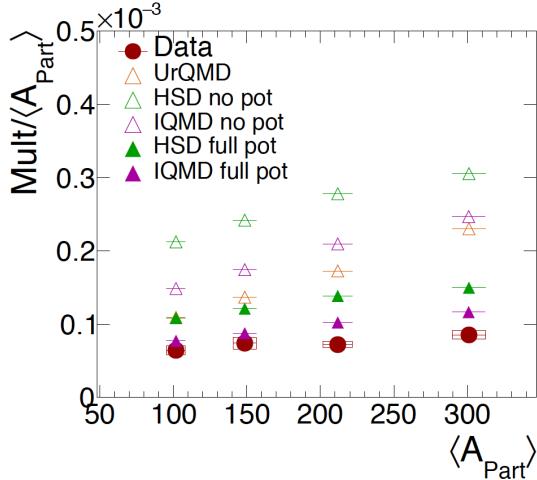
UrQMD3.4 P. Hillmann, J. Steinheimer, M. Bleicher, arXiv:1802.01951



UrQMD predicts high sensitivity of v_3 to EoS

Microscopic Description

Λ, K_S^0 production compared to transport models



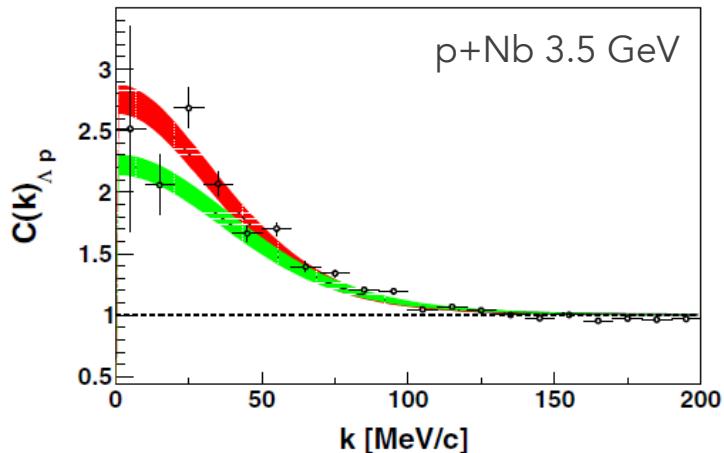
- “no pot” - state-of-the-art transport model calculations overestimate yield and low p_t
- “full pot” - inclusion of repulsive KN potential ($U_{\text{opt}}=+40 \text{ MeV} \cdot p/p_0$) → describes trends better

- Λp_t spectrum best described by UrQMD (no potential!)
- No simultaneous description of K and Λ results

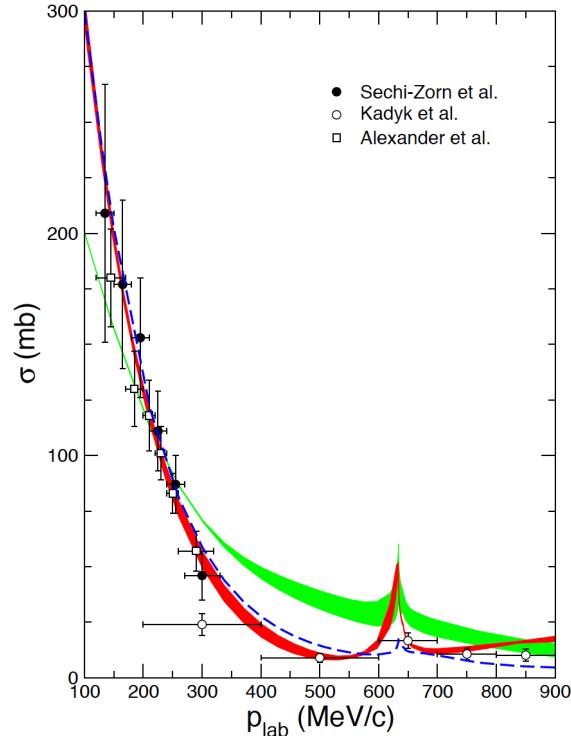
YN, YNN interactions

J. Haidenbauer et al., arXiv:1304.5339v1 [nucl-th]
 $\Lambda p \rightarrow \Lambda p$

ΛN correlation function

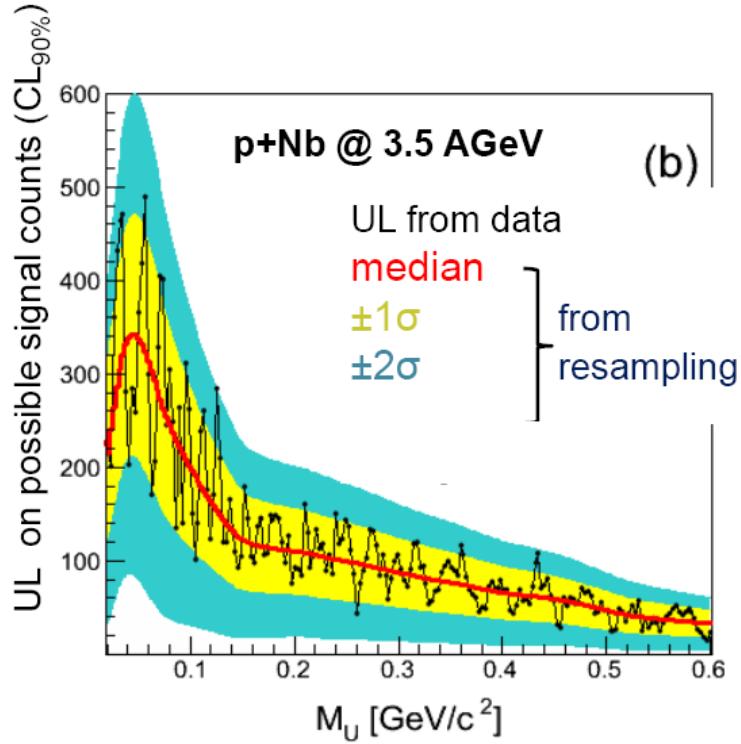


HADES collab., PRC 94 (2016) no.2, 025201

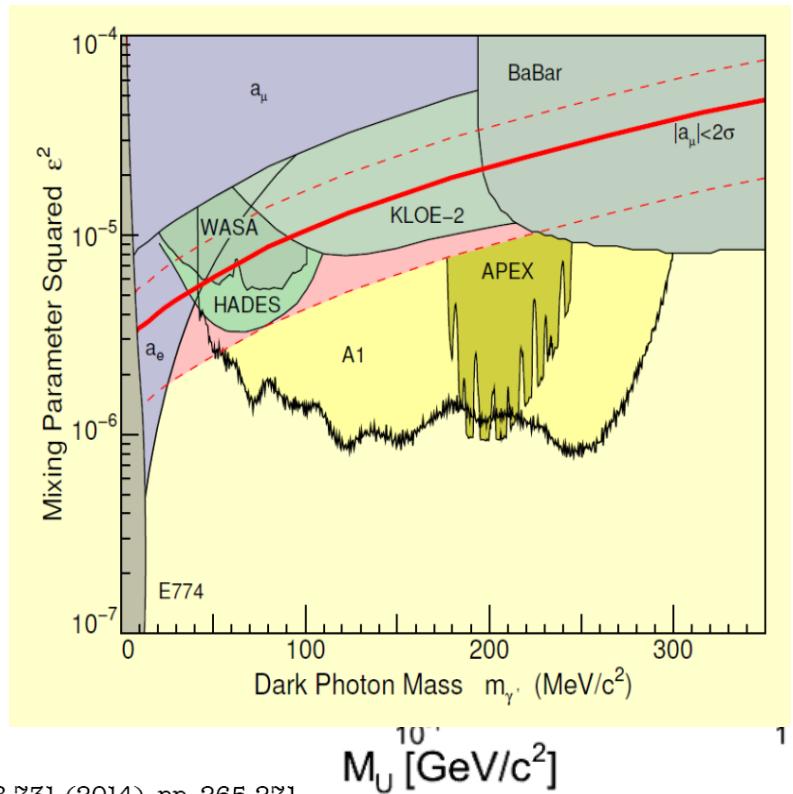


The red band shows the chiral EFT results to NLO, the green band are results to LO

Dark Photons



New A1 results



Phys. Lett. B 731 (2014), pp. 265-271