

### Questions

- 1. Can HIC help constrain the EOS at NS core densities?
- 2. Signatures of phase transitions in QCD matter at large baryon densities?
- 3. What are suitable observables?

# **EOS from NSs**





"While terrestrial experiments are able to test and constrain the EOS at densities below the saturation density of nuclei  $\rho_{nuc} = 2.8 \times 10^{14}$  g cm<sup>-3</sup> (see e.g. [14–17] for a review), currently they cannot probe the extreme conditions in the core of NSs"

#### EMMI RRTF, GSI, Darmstadt

### **EOS determination in HIC**



## Elliptic flow at energies 0.25-1.5 AGeV













- heavy ion reactions rapidly evolving, transient state
- influenced by
  - nuclear matter equation of state
    - momentum dependence
  - in-medium cross sections
    - Pauli blocking
  - in-medium characteristics of particles
    - effective masses/potentials, spectral functions
    - decay widths
- microscopic transport models needed
- convincing conclusions on basic nuclear properties imply a successful simulation:
  - of the full set of experimental observables
  - with the same code
  - using the same physical and technical parameters.



# Symmetry energy from neutron/proton flow





for details see Y. Leifels, EMMI – RRTF symposium

 $E_{sym} = E_{sym}^{pot} + E_{sym}^{kin} = 22 \text{MeV} \cdot (\rho/\rho_0)^{\gamma} + 12 \text{MeV} \cdot (\rho/\rho_0)^{2/3}$ Density dependence:  $\gamma = 0.72 \pm 0.19$ Pressure contribution at  $\rho = \rho_0$ :  $\rho_0 = 3.4 \pm 0.7 \text{ MeV/fm}^3$ 

### EOS at center of largest NSs ?





#### AP4 incomplete, strangeness DoF missing.

# **Excitation function of collective flow**



P. Danielewicz, R. Lacey, W.G. Lynch, Science 298 (2002) 1592



Large sensitivity for high densities in the (lower) AGS/SIS100 beam energy range.

### **Phases of QCD matter ?**







### **CBM – Goals**





### **Mission:**

Systematically explore QCD matter at large baryon densities with high accuracy and rare probes.

### **Dense Baryonic Matter**





### **Neutron stars**

Temperature T < 20 MeV

Density  $\rho < 10 \rho_0$ Lifetime



### **Neutron star merger**

Temperature T < 70 MeV

Density  $\rho < 2 - 6 \rho_0$ 

Reaction time (GW170817) T ~ 10 ms

### Heavy ion collisions at SIS100



#### Compressed Baryonic Matter

Temperature T < 120 MeV

Density  $\rho < 8\rho_0$ 

Reaction time  $t \sim 10^{-23} s$ 

# Baryon densities in central Au+Au collisions





I.C. Arsene et al., Phys. Rev. C 75, 24902 (2007)

### Reminder: Subthreshold Kaon – measurements (KAOS at SIS18)







# **Final state particle abundance**



#### Particle yields from central Au + Au collisions



Strange and charmed particle production thresholds in pp - collisions

reaction	$\sqrt{s}$ (GeV)	T <sub>lab</sub> (GeV)
$pp \to K^+ \Lambda p$	2.548	1.6
$pp \rightarrow K^+ K^- pp$	2.864	2.5
$pp \to K^+ K^+ \Xi^- p$	3.247	3.7
$pp \to K^+ K^+ K^+ \Omega^- n$	4.092	7.0
$pp \rightarrow \Lambda \bar{\Lambda} pp$	4.108	7.1
$pp \rightarrow \Xi^- \overline{\Xi}^+ pp$	4.520	9.0
$pp \rightarrow \Omega^- \bar{\Omega}^+ pp$	5.222	12.7
$pp \rightarrow J/\Psi pp$	4.973	12.2

# **CBM – Strategy**





# **CBM experimental setup (day-1)**

![](_page_16_Picture_1.jpeg)

![](_page_16_Figure_2.jpeg)

- Tracking acceptance:  $2^{\circ} < \theta_{lab} < 25^{\circ}$
- Free streaming DAQ
- R<sub>int</sub> = 10 MHz (Au+Au)

$$\begin{split} R_{int} &\approx 0.5 \; MHz \\ \text{full bandwith:} \\ & \text{Det.} - \text{Entry nodes} \\ \text{reduced bandwidth} \\ & \text{Entry nodes} - \text{Comp. farm} \end{split}$$

with R<sub>int</sub> (MVD)=0.1 MHz

Software based event selection

Day-1 funding: ~ 90% secured

# **CBM data processing system**

![](_page_17_Picture_1.jpeg)

![](_page_17_Picture_2.jpeg)

Reaction rate Au + Au:

10<sup>7</sup> collisions per second

Data rate:

~ 1 TB/s

![](_page_17_Picture_7.jpeg)

![](_page_17_Figure_8.jpeg)

### Main features:

- radiation tolerant detectors and front-end electronics
- free streaming (triggerless) data with time stamps,
- software based event selection

N.Herrmann, June 11, 2018

#### EMMI RRTF, GSI, Darmstadt

# **CBM** physics and observables

### QCD equation-of-state

- collective flow of identified particles
- particle production at threshold energies

### Phase transition

- excitation function of hyperons
- excitation function of LM lepton pairs

### **Critical point**

event-by-event fluctuations of conserved quantities

### Chiral symmetry restoration at large $\rho_B$

- in-medium modifications of hadrons
- dileptons at intermediate invariant masses

### Strange matter

- (double-) lambda hypernuclei
- Search for meta-stable objects (e.g. strange dibaryons)

### Heavy flavour in cold and dense matter

excitation function of charm production

![](_page_18_Picture_18.jpeg)

![](_page_18_Picture_19.jpeg)

![](_page_18_Picture_20.jpeg)

volume 53 + number 3 + march + 2011

![](_page_18_Picture_21.jpeg)

 $\rightarrow$  V. Friese

The European Physical Journal

# CBM day-1 – program

![](_page_19_Picture_1.jpeg)

### Observables: Strangeness and Dileptons

Excitation function of yields and phase-space distributions of multi-strange hyperons and lepton pairs in AA (C+C, Au+Au) collisions from 2-11 A GeV. Search for hypernuclei (no data available in this energy range).

![](_page_19_Figure_4.jpeg)

## **Chemical Freeze-out data**

![](_page_20_Picture_1.jpeg)

Analyses in framework of Statistical Hadronisation Model

![](_page_20_Figure_3.jpeg)

High energies:

grandcanonical ensemble

$$n_i(\mu,T) = \frac{N_i}{V} = -\frac{T}{V} \frac{\partial \ln Z_i}{\partial \mu} = \frac{g_i}{2\pi^2} \int \frac{p^2 dp}{e^{\frac{E_i - \mu_i}{T}} \pm 1}$$
$$\mu_i = \mu_B B_i + \mu_S S_i + \mu_{I_3} I_{3,i}$$

Lower energies / small systems: canonical ensemble, strangeness suppression factor  $\gamma_s$ 

Equilibrium achieved in small systems?

Equilibrium as signature for phase transition?

Freeze-out line at large baryon densities as phase boundary to quarkyonic matter?

![](_page_20_Figure_11.jpeg)

A. Andronic et al., Nucl. Phys. A837 (2010) 65

# **HADES:** Sub-threshold $\Xi^-$ - production

![](_page_21_Picture_1.jpeg)

### Ar+KCI reactions at 1.76A GeV

•  $\Xi^{-}$  yield by appr. factor 25 higher than thermal yield

![](_page_21_Figure_4.jpeg)

G. Agakishiev et al. (HADES), PRL103, 132301, (2009)

## Hyperons as probes of dense matter

![](_page_22_Picture_1.jpeg)

### **PHSD** interpretation of $\Xi^-$ - production

A. Palmese et al. Phys.Rev. C94 (2016) no.4, 044912

![](_page_22_Figure_4.jpeg)

Predicted sensitivities of production yields:

strong dependence on Chiral Symmetry Restoration (CSR)

Measurable dependence on Equation of State (NL1, NL3)

Alternative explanation (URQMD): Tuned resonance parameter J. Steinheimer, M. Bleicher, J.Phys. G43 (2016), 015104

### **Prediction of PHSD transport model**

(E. Bratkovskaya, W. Cassing)

I. Vassiliev, CBM, private communication

![](_page_23_Figure_4.jpeg)

# **CBM experimental setup (MSV)**

![](_page_24_Picture_1.jpeg)

![](_page_24_Figure_2.jpeg)

- Tracking acceptance:  $2^{\circ} < \theta_{lab} < 25^{\circ}$
- Free streaming DAQ

 $R_{int} = 10 \text{ MHz} (Au+Au)$ 

with  $R_{int}$  (MVD) = 0.1 MHz

 Software based event selection

# **Equation of State & Neutron stars**

![](_page_25_Picture_1.jpeg)

![](_page_25_Figure_2.jpeg)

Soft EOS (Skyrme, K = 200 MeV) is not repulsive enough to allow for a neutron star with 2 solar masses.

DBHF BONN A corresponds to AP4, however, does not contain strange baryons.

Stiffening of EOS must occur in the range of densities up to 4  $\rho_0$  (SIS100 energy range).N.Herrmann, June 11, 2018EMMI RRTF, GSI, Darmstadt26

# Dileptons as probes for dense matter (Day 1)

![](_page_26_Picture_1.jpeg)

![](_page_26_Figure_2.jpeg)

- LMR:  $\rho$  chiral symmetry restoration fireball space time extension
- IMR: access to fireball temperature  $\rho$ -a<sub>1</sub> chiral mixing

Measurement program:

e.g. excitation function of IMR – slope full performance, uses MVD (100 kHz)

![](_page_26_Figure_7.jpeg)

Collision Energy (√s<sub>NN</sub>) [GeV]

# **CBM Day 1 – further unique measurements**

![](_page_27_Picture_1.jpeg)

1.9

 $\Omega^{-} \rightarrow \Lambda K^{-}$ 

1.7

eff = 5.1%

S/Bg = 1.0

1.8

 ${}^{4}_{\Lambda}\text{He} \rightarrow {}^{3}\text{He+p+}\pi^{-}$ 

4.1

m [GeV/c<sup>2</sup>]

 $m_{inv} K^{-} \Lambda (GeV/c^2)$ 

#### Hyperon measurements:

Au+Au at 10A GeV, $\varepsilon_{dutv}$ = 50	0%, R=100kHz
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Particle	Multi- plicity	BR	ε (%)	yield (s⁻¹)	yield in 1 week
Ω <sup>-</sup> (1672)	5.6·10 <sup>-3</sup>	0.68	5	1.64	5•10⁵
<sup>4</sup> ∧He (3930)	1.9·10 <sup>-3</sup>	0.32	14.7	0.87	3∙10⁵

### Hypernuclei measurement:

#### **Di-Muon**

### LM measurement at 8A GeV

complementary measurement to e<sup>+</sup>e<sup>-</sup> with different systematic errors

![](_page_27_Figure_9.jpeg)

Entries / (8MeV/c<sup>2</sup>)

Entries

100

**50** 

0 3.9

800

600

400

200

1.6

# **Indian contribution: Muon Detector and Physics**

![](_page_28_Picture_1.jpeg)

![](_page_28_Figure_2.jpeg)

EMMI RRTF, GSI, Darmstaat

### Reference data for $\Lambda$ – production

![](_page_29_Picture_1.jpeg)

M. Merschmeyer et al. (FOPI), PRC 76, 024906 (2007)

![](_page_29_Figure_3.jpeg)

Reaction:

<sup>58</sup>Ni + <sup>58</sup>Ni at 1.93 AGeV

Centrality: 350 mb (most central)  $\frac{\sigma_{cen}}{\sigma_{geo}} \le 0.13$ 

Data taking period: 17.1.2003 – 3.2.2003

Statistics:

~ 60.000 reconstructed  $\Lambda$ 

Derived quantities: slope parameter integrated yield

# Physics of the benchmark observable

![](_page_30_Figure_1.jpeg)

![](_page_30_Figure_2.jpeg)

![](_page_30_Figure_3.jpeg)

- smaller than proton
- not explained by transport models
- reason unclear:
  - rescattering cross section
  - repulsive potential