

Collective modes in chiral relativistic plasmas

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**The 16th International
Workshop on QCD in
Extreme Conditions**



Chiral plasmas ($\vec{B}, \vec{\omega}$)

- **Early Universe, e.g.,**

[Boyarsky, Frohlich, Ruchayskiy, Phys.Rev.Lett. 108, 031301 (2012)]

- **Heavy-ion collisions, e.g.,**

[Kharzeev, Liao, Voloshin, Wang, Prog.Part.Nucl.Phys. 88, 1 (2016)]

- **Super-dense matter in compact stars, e.g.,**

[Yamamoto, Phys.Rev. D93, 065017 (2016)]

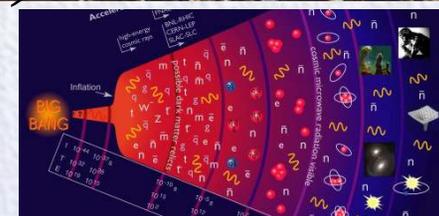
- **Ultra-relativistic jets from black holes**

- **Dirac/Weyl (semi-)metals, e.g.,**

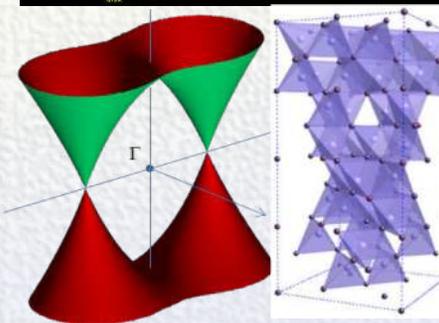
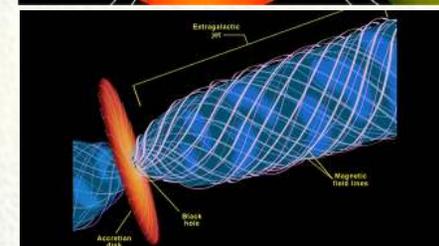
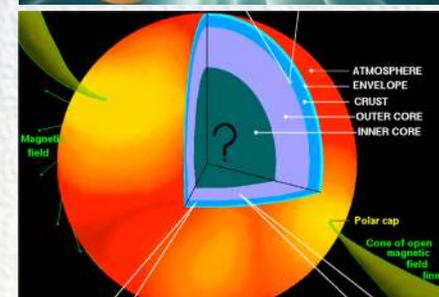
[Li et. al. Nature Phys. 12, 550 (2016)]

- **Superfluid $^3\text{He-A}$, e.g.,**

[Volovik, JETP Lett. 105, 34 (2017)]



Credit: Brookhaven National Laboratory



- *Massless* Dirac fermions:

$$\left(\gamma^0 p_0 - \vec{\gamma} \cdot \vec{p}\right) \Psi = 0 \quad \Rightarrow \quad \frac{\vec{\Sigma} \cdot \vec{p}}{|\vec{p}|} \Psi = \text{sign}(p_0) \gamma^5 \Psi$$

For particles ($p_0 > 0$): chirality = helicity

For antiparticles ($p_0 < 0$): chirality = - helicity

- Massive Dirac fermions in *ultrarelativistic* regime

– High temperature: $T \gg m$

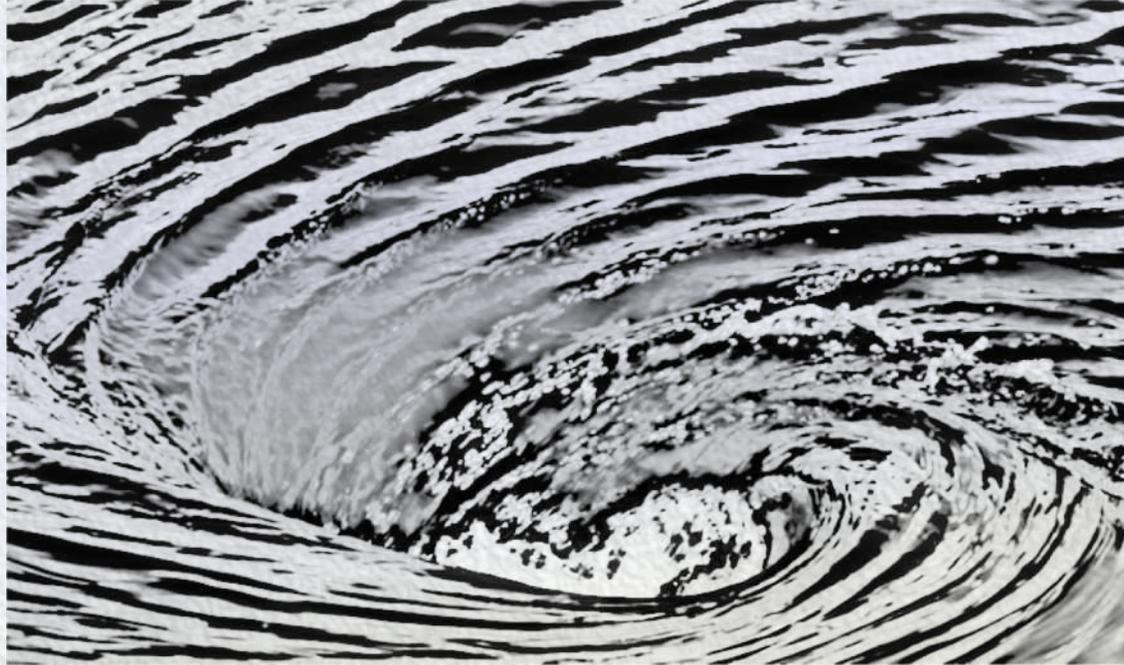
– High density: $\mu \gg m$

- Matter made of chiral fermions with $n_L \neq n_R$
- Unlike the electric charge ($n_R + n_L$), the chiral charge ($n_R - n_L$) is **not** conserved

$$\frac{\partial(n_R + n_L)}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$$

$$\frac{\partial(n_R - n_L)}{\partial t} + \vec{\nabla} \cdot \vec{j}_5 = \frac{e^2 \vec{E} \cdot \vec{B}}{2\pi^2 c}$$

- The chiral symmetry is anomalous in quantum theory



CHIRAL HYDRODYNAMICS

- Continuity equations: [Son, Surowka, Phys. Rev. Lett. 103, 191601 (2009)]
[Neiman and Oz, JHEP 03, 023 (2011)]

$$\partial_\mu j^\mu = 0$$

$$\partial_\mu j_5^\mu = -\frac{e^2}{2\pi^2 \hbar^2} E^\mu B_\mu$$

$$\partial_\nu T^{\mu\nu} = e F^{\mu\nu} j_\nu$$

together with the constitutive relations:

$$j^\mu = n u^\mu + \nu^\mu$$

$$j_5^\mu = n_5 u^\mu + \nu_5^\mu$$

$$T^{\mu\nu} = \epsilon u^\mu u^\nu - \Delta^{\mu\nu} P + (h^\mu u^\nu + u^\mu h^\nu) + \pi^{\mu\nu}$$

- Currents included new non-dissipative terms:

$$j^\mu = nu^\mu + \sigma_\omega \omega^\mu + \sigma_B B^\mu$$

$$j_5^\mu = n_5 u^\mu + \sigma_\omega^5 \omega^\mu + \sigma_B^5 B^\mu$$

where the anomalous coefficients are

$$\sigma_\omega = \frac{\mu\mu_5}{\pi^2 \hbar^2}, \quad \sigma_B = \frac{e\mu_5}{2\pi^2 \hbar^2}$$

$$\sigma_\omega^5 = \frac{3(\mu^2 + \mu_5^2) + \pi^2 T^2}{6\pi^2 \hbar^2}, \quad \sigma_B^5 = \frac{e\mu}{2\pi^2 \hbar^2}$$

The Question

- **Can one derive chiral hydrodynamics from first principles?**
- Chiral kinetic theory (CKT) is a good starting point
- Note: CKT can be “derived” from field theory
- Original versions of CKT had several limitations:
 - No explicit Lorentz covariance
 - Collisions are tricky

[Son and Yamamoto, Phys. Rev. D **87**, 085016 (2013)]
 [Stephanov and Yin, Phys. Rev. Lett. **109**, 162001 (2012)]

- Kinetic equation:

$$\frac{\partial f_\lambda}{\partial t} + \frac{\left[e\tilde{\mathbf{E}}_\lambda + \frac{e}{c}(\mathbf{v} \times \mathbf{B}_\lambda) + \frac{e^2}{c}(\tilde{\mathbf{E}}_\lambda \cdot \mathbf{B}_\lambda)\boldsymbol{\Omega}_\lambda \right] \cdot \nabla_{\mathbf{p}} f_\lambda}{1 + \frac{e}{c}(\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda)} + \frac{\left[\mathbf{v} + e(\tilde{\mathbf{E}}_\lambda \times \boldsymbol{\Omega}_\lambda) + \frac{e}{c}(\mathbf{v} \cdot \boldsymbol{\Omega}_\lambda)\mathbf{B}_\lambda \right] \cdot \nabla_{\mathbf{r}} f_\lambda}{1 + \frac{e}{c}(\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda)} = 0$$

where $\tilde{\mathbf{E}}_\lambda = \mathbf{E}_\lambda - (1/e)\nabla_{\mathbf{r}}\epsilon_{\mathbf{p}}$, $\mathbf{v} = \nabla_{\mathbf{p}}\epsilon_{\mathbf{p}}$,

$$\epsilon_{\mathbf{p}} = v_F p \left[1 - \frac{e}{c}(\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda) \right]$$

and $\boldsymbol{\Omega}_\lambda = \lambda\hbar \frac{\hat{\mathbf{p}}}{2p^2}$ is the Berry curvature

- Lorentz covariant formulation of CKT:

[Hidaka, Pu, Yang, Phys. Rev. D 95, 091901 (2017); Phys. Rev. D 97, 016004 (2018)]

$$\mathcal{D}_\mu W^\mu(p, x) = \delta(p^2) p \cdot C + \lambda \hbar e \tilde{F}^{\mu\nu} C_\mu p_\nu \delta'(p^2)$$

where $\mathcal{D}^\mu = \partial/\partial x^\mu - e F^{\mu\nu} \partial/\partial p^\nu$

$S^{\mu\nu} = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} p_\alpha u_\beta / (p \cdot u)$ is the spin tensor

C^μ is the collision operator

- Quasi-classical solution:

$$W^\mu(p, x) \equiv \underbrace{p^\mu \delta(p^2) f}_{\mathcal{O}(1)} + \underbrace{\lambda \hbar S^{\mu\nu} \delta(p^2) (D_\nu f - C_\nu) + \lambda \hbar e \tilde{F}^{\mu\nu} p_\nu \delta'(p^2) f}_{\mathcal{O}(\hbar)}$$

Approximations

- Relaxation time approximation:

$$\mathcal{D}_\mu W^\mu = -\frac{u_\mu (W^\mu - W_{\text{eq}}^\mu)}{\tau}$$

Note, the Wigner function W^μ is expressed in terms of distribution function, e.g.,

$$f_{\text{eq}}(p, x) = \frac{1}{1 + e^{\text{sign}(p_0)(\varepsilon_p - \mu_\lambda)/T}}$$

where $\mu_\lambda \equiv \mu + \lambda\mu_5$, $\varepsilon_p = \underbrace{u_\mu p^\mu}_{O(1)} + \underbrace{\frac{\lambda\hbar}{2} \frac{p \cdot \omega}{p \cdot u}}_{O(\hbar)}$

and $\omega^\mu = \frac{1}{2}\epsilon^{\mu\nu\alpha\beta} u_\nu \partial_\alpha u_\beta$ is the vorticity

Constitutive relations

- Conserved currents in CKT are moments of W^μ :

$$j^\mu = 2 \sum_\lambda \int \frac{d^4 p}{(2\pi)^3} W^\mu$$

$$j_5^\mu = 2 \sum_\lambda \lambda \int \frac{d^4 p}{(2\pi)^3} W^\mu$$

$$T^{\mu\nu} = \sum_\lambda \int \frac{d^4 p}{(2\pi)^3} (W^\mu p^\nu + p^\mu W^\nu)$$

However, when using the CKT equation, we get

$$\partial_\mu j^\mu = -\frac{1}{\tau} (n - n_{\text{eq}})$$

$$\partial_\mu j_5^\mu + \frac{e^2}{2\pi^2 \hbar^2} E^\mu B_\mu = -\frac{1}{\tau} (n_5 - n_{5,\text{eq}})$$

$$\underbrace{\partial_\nu T^{\mu\nu} - e F^{\mu\nu} j_\nu}_{\text{good}} = \underbrace{-\frac{u^\mu}{\tau} (\epsilon - \epsilon_{\text{eq}} + \dots) - \frac{1}{\tau} (h^\mu - h_{\text{eq}}^\mu + \dots)}_{\text{bad}}$$

Constitutive relations

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$$\partial_\nu T^{\mu\nu} - e F^{\mu\nu} j_\nu = 0$$

6 constraints, defining
local equilibrium
parameters T, μ, μ_5, u^μ

good

bad

- Dissipative terms (first-order):

$$\nu^\mu = \nu_{\text{eq}}^\mu + \frac{\tau}{3} \nabla^\mu n - \tau \dot{u}^\mu n + \sigma_E E^\mu$$

$$\nu_5^\mu = \nu_{5,\text{eq}}^\mu + \frac{\tau}{3} \nabla^\mu n_5 - \tau \dot{u}^\mu n_5 + \sigma_E^5 E^\mu$$

$$\pi^{\mu\nu} = \frac{8\tau\epsilon}{15} \Delta_{\alpha\beta}^{\mu\nu} (\partial^\alpha u^\beta)$$

where

$$\sigma_E = \tau e \frac{3(\mu^2 + \mu_5^2) + \pi^2 T^2}{9\pi^2 \hbar^3}$$

$$\sigma_E^5 = \tau e \frac{2\mu\mu_5}{3\pi^2 \hbar^3}$$

- At this order, the constitutive relations are differential equations,

[Gorbar, Rybalka, Shovkovy, Phys. Rev. D 95, 096010 (2017)]

$$\dot{i}^{\langle\mu\rangle} + \frac{\nu^\mu - \nu_{\text{eq}}^\mu}{\tau} = -\dot{i}^\mu n + \frac{1}{3} \nabla^\mu n - \frac{n}{\epsilon + P} \Delta^{\mu\nu} \partial^\rho \pi_{\rho\nu} - \nu_\rho \omega^{\rho\mu} - (\partial \cdot u) \nu^\mu - \frac{9}{5} (\partial^{\langle\mu} u^{\rho\rangle}) \nu_\rho + \dots$$

$$\dot{i}_5^{\langle\mu\rangle} + \frac{\nu_5^\mu - \nu_{5,\text{eq}}^\mu}{\tau} = -\dot{i}^\mu n_5 + \frac{1}{3} \nabla^\mu n_5 - \frac{n_5}{\epsilon + P} \Delta^{\mu\nu} \partial^\rho \pi_{\rho\nu} - \nu_{5,\rho} \omega^{\rho\mu} - (\partial \cdot u) \nu_5^\mu - \frac{9}{5} (\partial^{\langle\mu} u^{\rho\rangle}) \nu_{5,\rho} + \dots$$

$$\dot{\pi}^{\langle\mu\nu\rangle} + \frac{\pi^{\mu\nu}}{\tau} = -2h^{\langle\mu} \dot{i}^{\nu\rangle} + 2\pi_\rho^{\langle\mu} \omega^{\nu\rangle\rho} - \frac{10}{7} \pi_\rho^{\langle\mu} \sigma^{\nu\rangle\rho} - \frac{4}{3} \pi^{\mu\nu} \partial_\alpha u^\alpha + \frac{8}{15} (\partial^{\langle\mu} u^{\nu\rangle}) \epsilon + \dots$$

- Causality is Ok
- Stability is (probably) Ok

- Sound waves

$$\Omega = \pm \frac{k_z}{\sqrt{3}} + \frac{3}{8} \hbar \bar{\omega} \frac{n_{5,\text{eq}}}{\epsilon_{\text{eq}}} k_z + \frac{2}{15} i \tau k_z^2$$

- Chiral vortical waves

$$\Omega = \hbar \bar{\omega} v_1 k_z - \frac{1}{3} i \tau k_z^2, \quad \Omega = \hbar \bar{\omega} v_2 k_z - \frac{1}{3} i \tau k_z^2$$

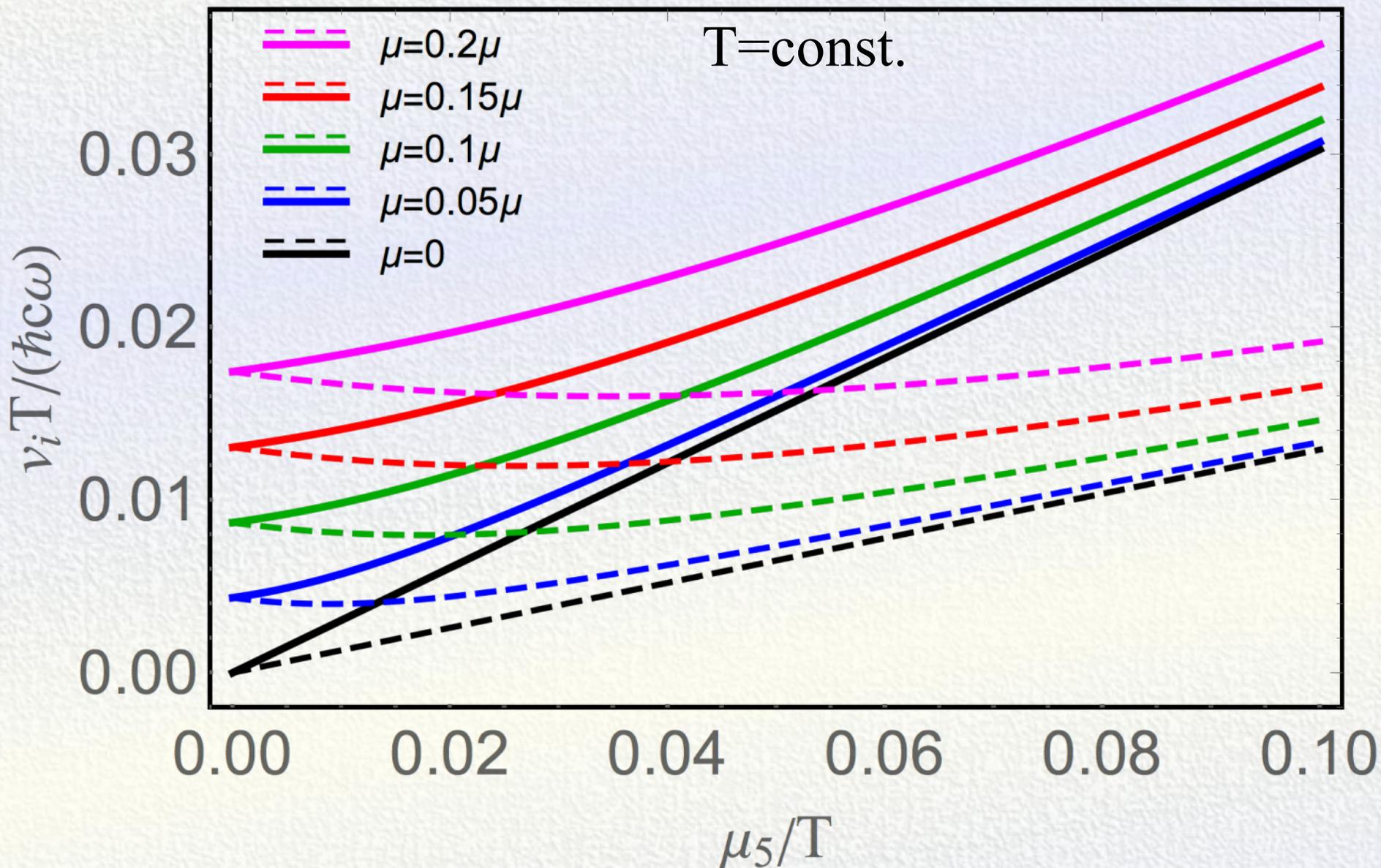
where $v_1 \neq v_2$ (along/against $\vec{\omega}$ direction)

- Oscillations of all thermodynamic parameters are important:

$$\delta\mu \neq 0, \quad \delta\mu_5 \neq 0, \quad \delta T \neq 0, \quad \delta u^\mu \neq 0$$

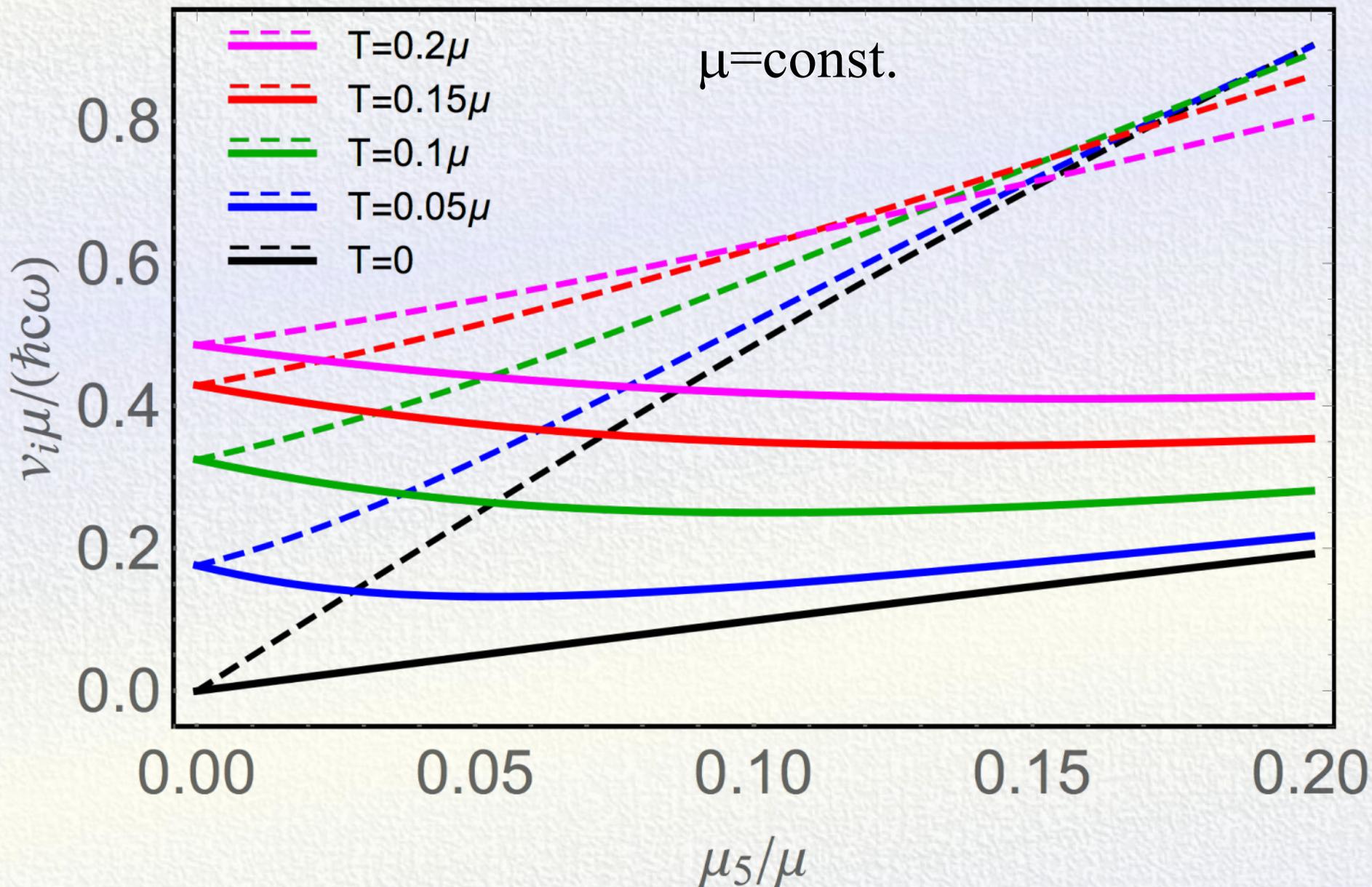
[Gorbar, Rybalka, Shovkovy, Phys. Rev. D 95, 096010 (2017)]

CVW velocities @ high T

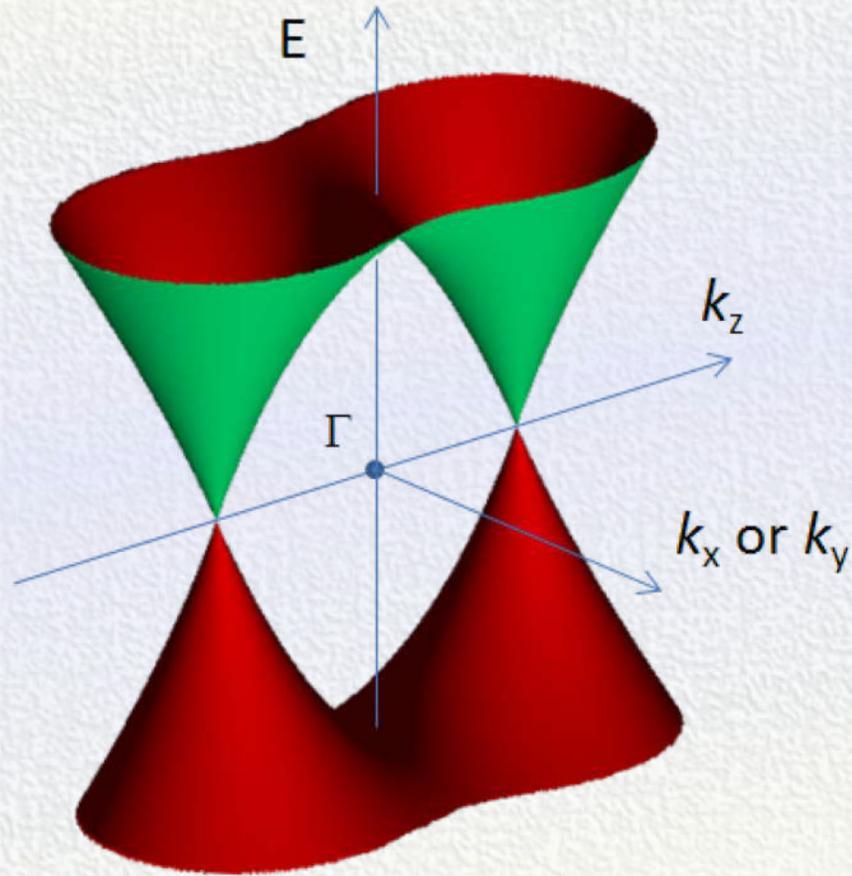


[Gorbar, Rybalka, Shovkovy, Phys. Rev. D 95, 096010 (2017)]

CVW velocities @ large μ



[Gorbar, Rybalka, Shovkovy, Phys. Rev. D 95, 096010 (2017)]



Credits: Borisenko et al., Phys. Rev. Lett. 113, 027603 (2014)

DIRAC & WEYL MATERIALS

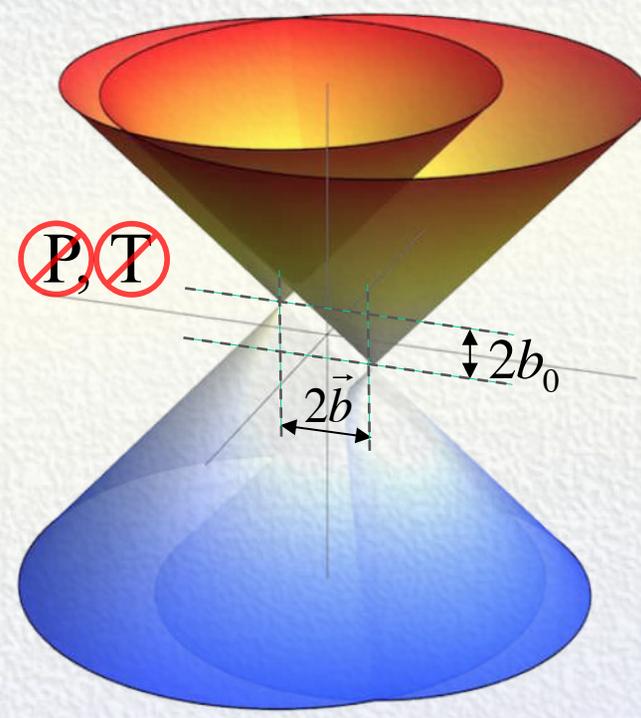
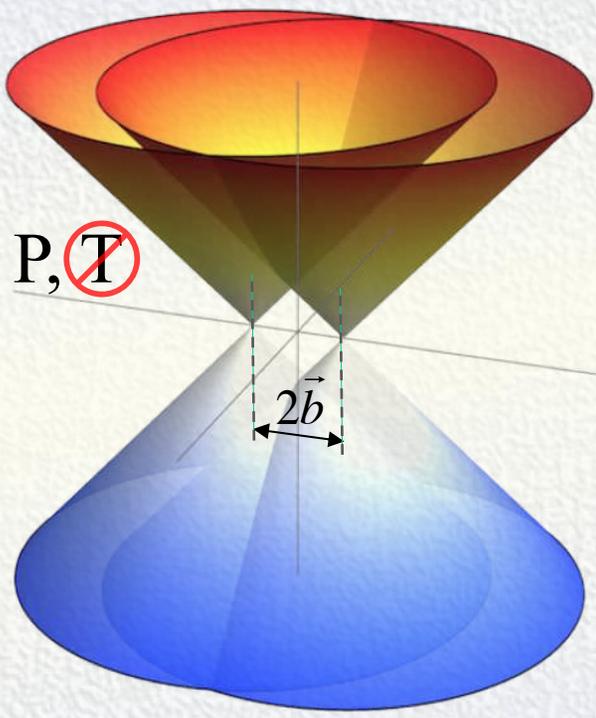
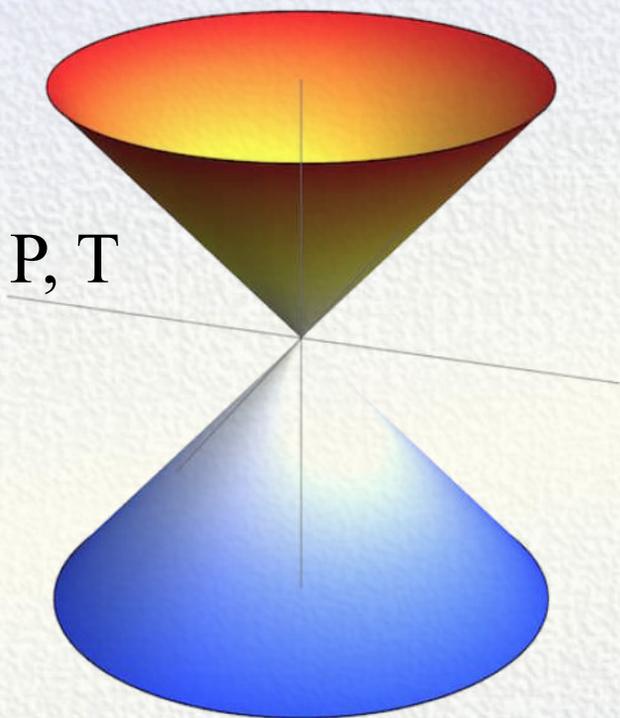
Dirac vs. Weyl materials

- Low-energy Hamiltonian of a Dirac/Weyl material

$$H = \int d^3\mathbf{r} \bar{\psi} \left[-iv_F (\vec{\gamma} \cdot \vec{\mathbf{p}}) - \underbrace{(\vec{b} \cdot \vec{\gamma}) \gamma^5}_{\cancel{T}} + \underbrace{b_0 \gamma^0 \gamma^5}_{\cancel{P}} \right] \psi$$

Dirac (e.g., Na₃Bi, Cd₃As₂, ZrTe₅)

Weyl (e.g., TaAs, NbAs, TaP, NbP, WTe₂)



Strain in Weyl materials

- Strains in the low-energy effective Weyl Hamiltonian

$$H = \int d^3\mathbf{r} \bar{\psi} \left[-iv_F (\vec{\gamma} \cdot \vec{\mathbf{p}}) - (\vec{b} + \vec{A}_5) \cdot \vec{\gamma} \gamma^5 + (b_0 + A_{5,0}) \gamma^0 \gamma^5 \right] \psi$$

where the chiral gauge fields are

$$A_{5,0} \propto b_0 \left| \vec{b} \right| \partial_{||} u_{||}$$

$$A_{5,\perp} \propto \left| \vec{b} \right| \partial_{||} u_{\perp}$$

$$A_{5,||} \propto \alpha \left| \vec{b} \right|^2 \partial_{||} u_{||} + \beta \sum_i \partial_i u_i$$

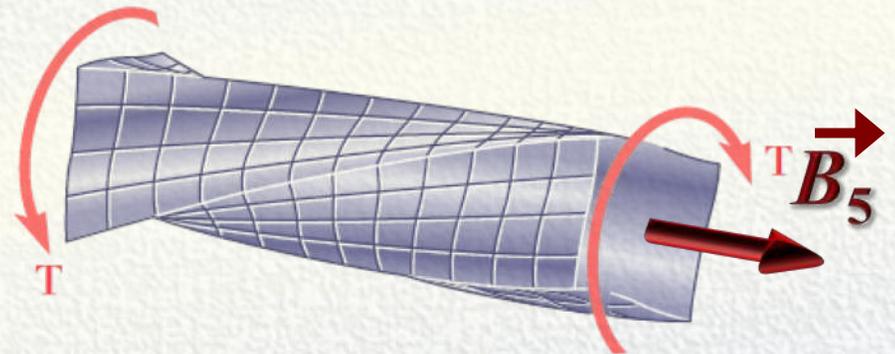
[Zubkov, Annals Phys. **360**, 655 (2015)]

[Cortijo, Ferreira, Landsteiner, Vozmediano. PRL **115**, 177202 (2015)]

[Pikulin, Chen, Franz, PRX **6**, 041021 (2016)]

[Grushin, Venderbos, Vishwanath, Ilan, PRX **6**, 041046 (2016)]

[Cortijo, Kharzeev, Landsteiner, Vozmediano, PRB **94**, 241405 (2016)]



leading to the pseudo-EM fields

$$\vec{B}_5 = \vec{\nabla} \times \vec{A}_5 \quad \text{and} \quad \vec{E}_5 = -\vec{\nabla} A_0 - \partial_t \vec{A}_5$$

- The definitions of density and current are

$$\rho_\lambda = e \int \frac{d^3 p}{(2\pi\hbar)^3} \left[1 + \frac{e}{c} (\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda) \right] f_\lambda,$$

$$\mathbf{j}_\lambda = e \int \frac{d^3 p}{(2\pi\hbar)^3} \left[\mathbf{v} + \frac{e}{c} (\mathbf{v} \cdot \boldsymbol{\Omega}_\lambda) \mathbf{B}_\lambda + e (\tilde{\mathbf{E}}_\lambda \times \boldsymbol{\Omega}_\lambda) \right] f_\lambda$$

$$+ e \nabla \times \int \frac{d^3 p}{(2\pi\hbar)^3} f_\lambda \epsilon_{\mathbf{p}} \boldsymbol{\Omega}_\lambda,$$

They satisfy the following anomalous relations:

$$\frac{\partial \rho_5}{\partial t} + \nabla \cdot \mathbf{j}_5 = \frac{e^3}{2\pi^2 \hbar^2 c} \left[(\mathbf{E} \cdot \mathbf{B}) + (\mathbf{E}_5 \cdot \mathbf{B}_5) \right] \quad \checkmark$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{j} = \frac{e^3}{2\pi^2 \hbar^2 c} \left[(\mathbf{E} \cdot \mathbf{B}_5) + (\mathbf{E}_5 \cdot \mathbf{B}) \right] \quad \times$$

- Additional Bardeen-Zumino term is needed,

$$\delta j^\mu = \frac{e^3}{4\pi^2 \hbar^2 c} \epsilon^{\mu\nu\rho\lambda} A_\nu^5 F_{\rho\lambda}$$

- i.e.,

[Gorbar, Miransky, Shovkovy, Sukhachov, PRL 118, 127601 (2017)]

$$\delta \rho = \frac{e^3}{2\pi^2 \hbar^2 c^2} (\mathbf{A}^5 \cdot \mathbf{B})$$

$$\delta \mathbf{j} = \frac{e^3}{2\pi^2 \hbar^2 c} A_0^5 \mathbf{B} - \frac{e^3}{2\pi^2 \hbar^2 c} (\mathbf{A}^5 \times \mathbf{E})$$

- Its role and implications:

- Electric charge is conserved locally ($\partial_\mu J^\mu = 0$)
- Anomalous Hall effect is reproduced
- CME vanishes in equilibrium ($\mu_5 = -eb_0$)

ASU Hydrodynamics in Weyl metals

The Euler equation

$$\frac{1}{v_F} \partial_t \left(\frac{\epsilon + P}{v_F} \mathbf{u} + \sigma^{(\epsilon, B)} \mathbf{B} \right) = -en \left(\mathbf{E} + \frac{1}{c} [\mathbf{u} \times \mathbf{B}] \right) + \frac{\sigma^{(B)} (\mathbf{E} \cdot \mathbf{B})}{3v_F^2} \mathbf{u} - \frac{\epsilon + P}{\tau v_F^2} \mathbf{u} + O(\nabla_{\mathbf{r}})$$

The energy conservation

$$\partial_t \epsilon = -\mathbf{E} \cdot (en\mathbf{u} - \sigma^{(B)} \mathbf{B}) + O(\nabla_{\mathbf{r}})$$

plus the Maxwell equations that include the Chern-Simons (Bardeen-Zumino) terms, i.e.,

$$\rho_{\text{CS}} = -\frac{e^3 (\mathbf{b} \cdot \mathbf{B})}{2\pi^2 \hbar^2 c^2}$$
$$\mathbf{J}_{\text{CS}} = -\frac{e^3 b_0 \mathbf{B}}{2\pi^2 \hbar^2 c} + \frac{e^3 [\mathbf{b} \times \mathbf{E}]}{2\pi^2 \hbar^2 c}$$

[Gorbar, Miransky, Shovkovy, Sukhachov, PRB 97, 121105(R) (2018)]

- Magneto-acoustic wave ($\rho = 0$):

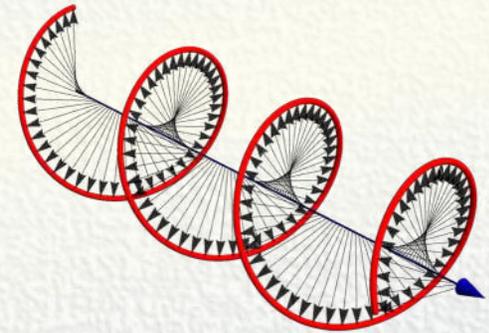
$$\omega_{s,\pm} = -\frac{i}{2\tau} \pm \frac{i}{2\tau} \sqrt{1 - 4\tau^2 v_F^2 \frac{|\mathbf{k}|^2 w_0 - \sigma^{(\epsilon,u)} [2|\mathbf{k}|^2 B_0^2 - (\mathbf{k} \cdot \mathbf{B}_0)]}{3w_0}}$$

- *Gapped* chiral magnetic wave ($\rho = 0$):

$$\omega_{\text{gCMW},\pm} = \pm \frac{eB_0 \sqrt{3v_F^3 \left(4\pi e^2 T_0^2 + 3\varepsilon_e \hbar^3 v_F^3 k_{\parallel}^2\right)}}{2\pi^2 T_0^2 c \sqrt{\varepsilon_e \hbar}}$$

- Helicons ($\rho \neq 0$):

$$\omega_{h,\pm} \approx \mp \frac{ck_{\parallel}^2 B_0}{4\pi\mu_m \rho_0} - \frac{i}{\tau} \frac{c^2 k_{\parallel}^2 w_0}{4\pi\mu_m v_F^2 \rho_0^2} + O(k_{\parallel}^3)$$



- New anomalous Hall waves at $\vec{b} \neq 0$, etc.

[Gorbar, Miransky, Shovkovy, Sukhachov, arXiv:1802.10110]

- Lorentz covariant form of consistent chiral kinetic (CKT) theory is formulated for (pseudo-)relativistic plasmas
- Topological terms play a critical role in Dirac and Weyl (semi-)metals
- Dissipative chiral hydrodynamics is derived from CKT
- Anomalous terms also affect dissipative dynamics
- Hydrodynamic modes in (pseudo-)relativistic plasmas with vorticity/magnetic field are studied