



- Neutron Stars and what we learn from Femtoscopy
 - The Hyperon Puzzle
 - How can Femtoscopy help?
- Correlation functions
 - Theoretical description of the correlation function
 - Experimental correlation functions
- Results of the Analysis in the p-p collision system at 7 TeV
 - pp and Λp Correlations
 - ΛΛ Correlations
 - Future prospects with CATS
- Summary and Outlook







Observation from X-Ray measurements of binary Systems:

> Neutron star masses $m > m_{\odot}$ and radii between 5 – 15 km

• From Shapiro Delay Measurements of binary

Systems: (P.Demorest et. Al, Nature 467 (2010), pp. 1081-1083)

> Neutron star $m = (1.97 \pm 0.04) m_{\odot}$

• Large nucleon chemical potential ($ho >
ho_0$)



Hyperon production becomes energetically possible? e.g. Λ, Ξ, Σ

Ozel et al. 2015 (arXiv:1505.05155v2)







- Interaction potential between hyperons and neutron matter governs the onset of production and behavior of the EoS
- Quantum Monte Carlo study of pure nuclear matter and hyper neutron matter
- Attractive ΛN interaction softens EoS
- Repulsive ΛNN interaction stiffens EoS



D. Lonardoni, A. Lovato, S. Gandolfi, F. Pederiva Phys. Rev. Lett. 114, 092301 (2015)





Mass – Radius Relation of Neutron Stars

- EoS allows to solve the Tolman-Oppenheimer-Volkoff Equations
 Mass Radius Relation
- Repulsive 3 body interaction shifts onset of hyperon production to larger densities and allows for higher masses
 - Not well constrained
- To understand the role of hyperons in neutron stars more constraints on the hyperon-neutron force are needed



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ALICE



Global Proton-A Scattering Data

- Data from scattering experiments from 1968 and 1971 in bubble chambers
 - $K^- + p \rightarrow \Sigma^0 + \pi^0, \Sigma^0 \rightarrow \Lambda + \gamma$
 - Production threshold for $\Lambda 's:p~\gtrsim 100~MeV$
- Different type of measurement needed to obtain constraints at low momentum
- Can we use Femtoscopic measurements?



LO: H. Polinder, J.H., U. Meiβner, NPA 779 (2006) 244 NLO: J.Haidenbauer., N.Kaiser, et al., NPA 915 (2013) 24







































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Theoretical formulation of the Correlation Function:

$$C(\overrightarrow{p_{a}},\overrightarrow{p_{b}}) = \frac{P(\overrightarrow{p_{a}},\overrightarrow{p_{b}})}{P(\overrightarrow{p_{a}})P(\overrightarrow{p_{b}})} = C(k) = \int S(\overrightarrow{r},k)|\psi(\overrightarrow{r},k)|^{2} d\overrightarrow{r}$$

Source Wave Function

If we know about the **source**, we can learn about the **interaction**

Fix the parametrization of the source by fitting the pp and Λp Correlation Function simultaneously to test different models of the p Λ interaction











The pp Correlation Function

- Governed by:
 - Coulomb Interaction
 - Strong Interaction
 - Quantum Statistics
- Koonin Fit Function
 - Assumes a Gaussian source of size R_G

$$C(k) = \int dr^3 \phi_{rel}^2(r,k) \exp\left(-\frac{r^2}{4R_g^2}\right)$$

S. E. Koonin, Phys. Lett. B 70 (1977) 43 S. Pratt et al., Nucl. Phys. A 566 (1994) 103c

φ_{rel} from solving the Schrödinger
 Equation with the known potentials
 for the Coulomb and Strong interaction







- Governed by:
 - Strong Interaction
 - No Coulomb Interaction
- Lednický model
 - Assumes a **Gaussian source** of size R_G
 - Based on the effective Range expansion
 - The interaction is modeled using the scattering length (f₀) and the effective range (d₀)

R. Lednicky and V. L. Lyuboshits, Sov. J. Nucl. Phys. 35, 770 (1982), [Yad. Fiz.35,1316(1981)].



$$C(k) = 1 + \sum_{S} \rho_{S} \left[\frac{1}{2} \left| \frac{f^{S}(k)}{R_{G}^{\Lambda p}} \right|^{2} \left(1 - \frac{d_{0}^{S}}{2\sqrt{\pi}R_{G}^{\Lambda p}} \right) + 2 \frac{\mathcal{R}f^{S}(k)}{\sqrt{\pi}R_{G}^{\Lambda p}} F_{1}(QR_{G}^{\Lambda p}) - \frac{\mathcal{I}f^{S}(k)}{R_{G}^{\Lambda p}} F_{2}(QR_{G}^{\Lambda p}) \right]$$





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An expression containing f_0



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- Discrimination of the two models not possible due to the lack of data
- NLO predicts a repulsive core of the interaction potential

	NLO						LO
Λ [MeV]	450	500	550	600	650	700	600
$a_s^{\Lambda p}$	-2.90	-2.91	-2.91	-2.91	-2.90	-2.90	-1.91
$r_s^{\Lambda p}$	2.64	2.86	2.84	2.78	2.65	2.56	1.40
$a_t^{\Lambda p}$	-1.70	-1.61	-1.52	-1.54	-1.51	-1.48	-1.23
$r_t^{\Lambda p}$	3.44	3.05	2.83	2.72	2.64	2.62	2.13

Dark Matter





A Large Ion Collider Experiment



pp 7 TeV



- Tracking and Particle Identification (PID) of charged particles
 - |η| < 0.9
 - 2π coverage in the azimuth
- Inner Tracking System (ITS)
- Time Projection Chamber (TPC)
- Time of Flight (TOF) System





$$C(\overrightarrow{\boldsymbol{p}_{a}}, \overrightarrow{\boldsymbol{p}_{b}}) = \frac{P(\overrightarrow{p_{a}}, \overrightarrow{p_{b}})}{P(\overrightarrow{p_{a}})P(\overrightarrow{p_{b}})} = \mathcal{N}\frac{N_{SE}(k)}{N_{ME}(k)}$$













λ Parameters can be related to measured single-particle quantities, e.g. Purity $\mathcal P$ or feed-down fractions f

(PhD Thesis of O. Arnold - https://www.das.ktas.ph.tum.de/DasDocs/Public/PhD_Theses/arnold-oliver_thesis.pdf)





Decomposition of the Correlation Functions

$$\begin{split} \{pp\} &= pp + p_{\Lambda}p + p_{\Lambda}p_{\Lambda} + p_{\Sigma^{+}}p + p_{\Sigma^{+}}p_{\Sigma^{+}} + p_{\Lambda}p_{\Sigma^{+}} + \tilde{p}p + \tilde{p}\tilde{p} \ . \\ \{p\Lambda\} &= p\Lambda + p\Lambda_{\Xi^{-}} + p\Lambda_{\Xi^{0}} + p\Lambda_{\Sigma^{0}} + p_{\Lambda}\Lambda + p_{\Lambda}\Lambda_{\Xi^{-}} \\ &+ p_{\Lambda}\Lambda_{\Xi^{0}} + p_{\Lambda}\Lambda_{\Sigma^{0}} + p_{\Sigma^{+}}\Lambda + p_{\Sigma^{+}}\Lambda_{\Xi^{-}} + p_{\Sigma^{+}}\Lambda_{\Xi^{0}} + p_{\Sigma^{+}}\Lambda_{\Sigma^{0}} \\ &+ \tilde{p}\Lambda + p\tilde{\Lambda} + \tilde{p}\tilde{\Lambda} \end{split}$$

Secondary contributions from measurements

Pair	Percentage %
pp	75
$p_{\Lambda}p$	16
$p_{\Lambda}p_{\Lambda}$	1
$p_{\Sigma^+}p$	6
$p_{\Sigma^+}p_{\Sigma^+}$	0
$p_{\Sigma^+} p_{\Lambda}$	0
$\widetilde{p}p$	2
$\widetilde{p}\widetilde{p}$	0

Pair	Percentage %
$p\Lambda$	49
$p\Lambda_{\Xi^-}$	10
$p\Lambda_{\Xi^0}$	10
$p\Lambda_{\Sigma^0}$	16
$p_{\Lambda}\Lambda$	5
$p_{\Lambda}\Lambda_{\Xi^-}$	1
$p_{\Lambda}\Lambda_{\Xi^0}$	1
$p_{\Lambda}\Lambda_{\Sigma^0}$	2
$p_{\Sigma^+}\Lambda$	0
$p_{\Sigma^+}\Lambda_{\Xi^-}$	0
$p_{\Sigma^+}\Lambda_{\Xi^0}$	1
$p_{\Sigma^+}\Lambda_{\Sigma^0}$	2
$\widetilde{p}\Lambda$	1
$p\widetilde{\Lambda}$	2
$\widetilde{p}\widetilde{\Lambda}$	0





• Mini Jet Background not present for baryon baryon pairs

(see J.Adam et al., Europ. Phys. Jour. C, Aug. 17, 77:569)

- Evolution of the system is better understood compared to Pb-Pb
 - Same freeze out times and source size for all particles
- Small source sizes
 - Stronger signal
 - Sensitivity to the shape of the potential









The pp and Λp Correlation Function



• Other collision systems e.g. p-Pb









? Any Sensitivity to the repulsive core of the potential? Included in pp not in Λp currently.















- Analysis of the $\Lambda\Lambda$ correlation function in Au-Au Collisions (STAR):
 - Fit of the Lednický Model with an additional residual term as a free parameter
 - Results in a slightly repulsive interaction ($a_0 < 0 \text{ fm}$)
- Modeled correlation function does not describe ALICE Data with an adjusted source









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• Reanalysis of the data (K. Morita et al., T. Furumoto, AO PRC91('15) 024916) finds an attractive interaction $(a_0 > 1.25 \text{ fm})$

- An attractive interaction allows to describe both STAR and ALICE data
- Larger statistics of Run 2 could yield a conclusive answer





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CATS – Correlation Analysis Tool Using the Schrödinger Equation



http://www.denseandstrange.ph.tum.de/index.php?id=78

$$C(k) = \int S(\vec{r},k) |\psi(\vec{r},k)|^2 d\vec{r} \xrightarrow{k \to \infty} 1$$











- Using CATS phase shifts δ_l can be extracted for any given potential
 - Comparison to global data









- Λp cross sections for any given potential can be extracted
 - Comparison to global data





Comparison CATS to Lednicky



• CATS:

- Usmani Potential for the interaction
- Gaussian Source
- Lednicky:
 - Scattering parameters and effective range obtained from the Usmani Potential

Wang and Pratt, Phys. Rev. Lett. 83, 3138, 1999







Comparison CATS to Lednicky



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- Agreement within 4% for a source size R_G = 1.30 fm
 - Corrections for small source sizes in the Lednicky Model play a role







Messenger



- CATS allows to describe experimental data of our ALICE analysis
- Input1:

ALICE

- Argonne V18 Potential
- Gaussian Source function with R_G = 1.30 fm
- <u>Input 2:</u>
 - Argonne V18 Potential
 - Source Distribution from EPOS
- EPOS not suitable to describe the source

Argonne V18 potential and data: Phys. Rev. C 51, 38









core

____35

r* [fm]

of the potential • Femtoscopy: pairs are also produced in the region inside the

> Access to the shape of the potential?

- Scattering experiments:
 intermediate region outside the core
- Reduce degrees of freedom by fitting different collision systems (pp 7 and 13 TeV and pPb at 5 TeV) at different energies at the same time



Gaussian Source Function ($R_{c} = 1.5$ fm)



Summary and Outlook

- Femtoscopy in small systems is feasible
- New method to calculate different contributions to the total correlation function based on single particle properties
- Modelling of the correlation function with CATS
- > Analysis of Run 2 Data in p-p at 13 TeV and p-Pb Collisions at 5 TeV ongoing
 - \succ Additionally obtain the Σ and Ξ Correlation Function
- <u>Universal and Robust Femto Analysis Tool</u>
 - Fit the correlation function of various systems simultaneously in combination with CATS





Tool





Thank you for your attention!











Considered Shapes



Study of the theoretical prediction for the $p\Sigma^0$ correlation in





A. Kiesel et al Phys. Rev. C89 (2014) no.5, 054916 Workshop on anti-matter, hyper-matter and exotica production at the LHC



Pythia low k*









Pythia high k*









Minijets Background



orrelations:



