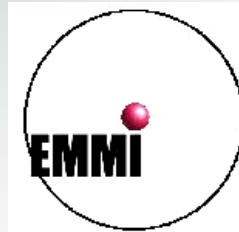


# 2<sup>nd</sup> EMMI Workshop on Plasma Physics with Intense Laser and Heavy Ion Beams, May 14-15, Moscow



## Investigations on warm dense plasma with PHELIX facility

**S.A. Pikuz Jr., I.Yu. Skobelev, A.Ya. Faenov,  
T.A. Pikuz, S.V. Gasilov**

**Joint institute for High Temperatures RAS**

# Application of high-resolution x-ray spectroscopy and radiography for investigations of plasma created by PHELIX laser.

- X-ray diagnostics of temperature, density, ionization state of **high-temperature** plasma created under interaction of high intense PHELIX laser pulse with **structured and homogeneous** targets.
- Observations of the **fast ions** generated in the PHELIX laser-produced plasma by X-Ray spectroscopy methods.
- X-ray spectroscopy and radiography of **warm dense matter**.
- Diagnostics of **MG-magnetic fields** generated in the laser-produced plasma by observations of X-Ray plasma satellites and Zeeman splitting of X-Ray spectral lines.

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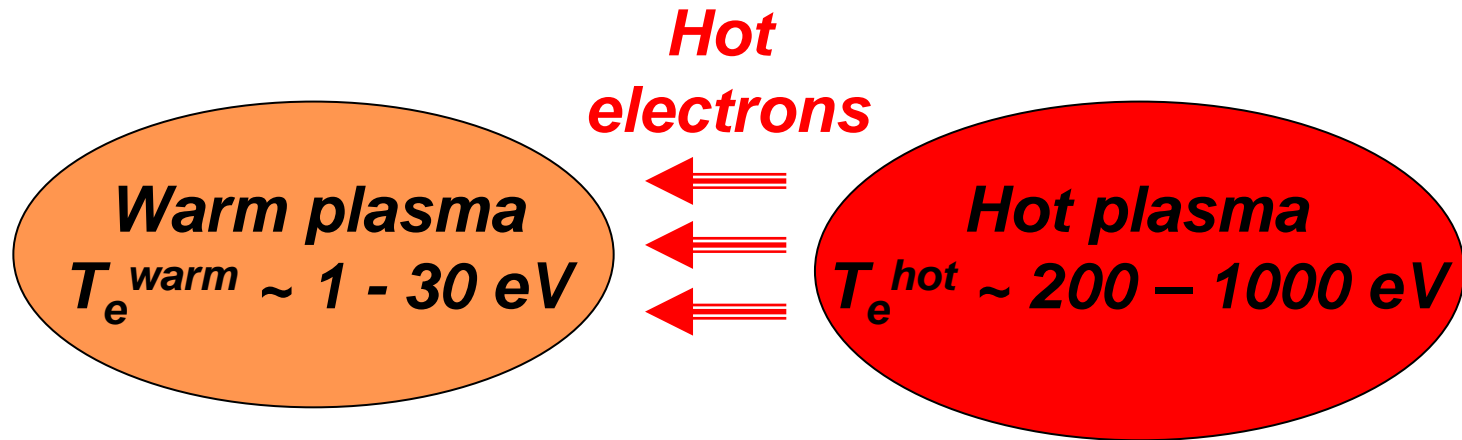
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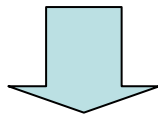
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# High resolution X-ray spectroscopy for WDM

# The way to produce and investigate WDM



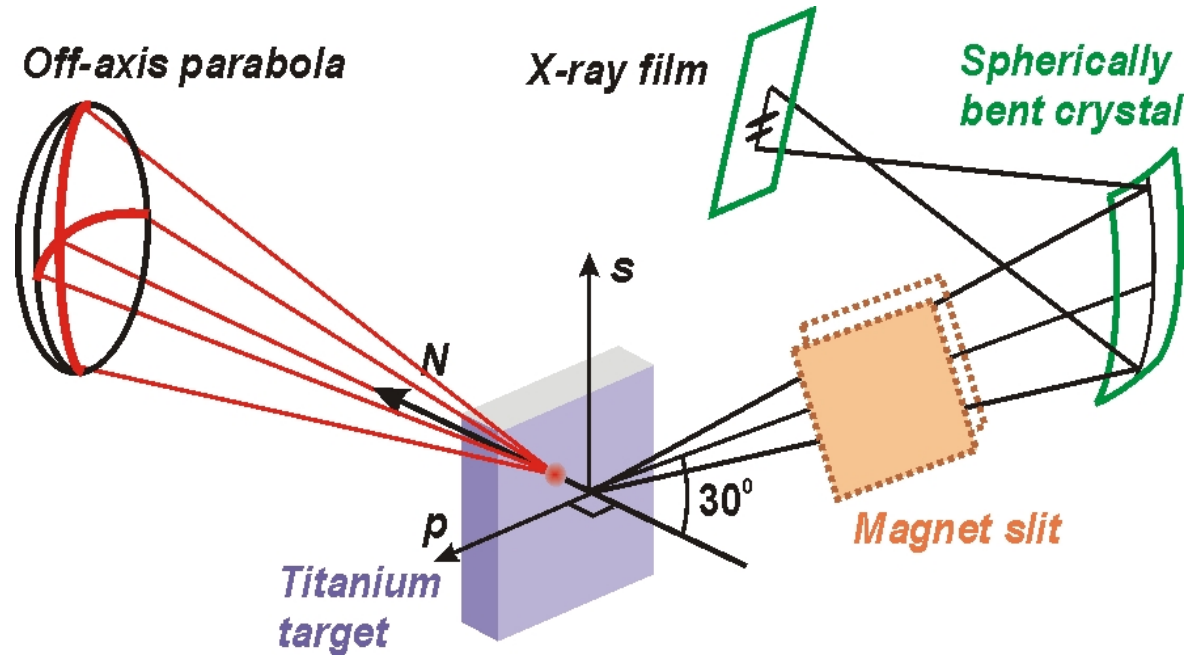
$K_\alpha$  lines  
of ions with small charges



to spectrometer

Directly measured value— plasma ionization state;  
 $T_e^{warm}$  can be determined

# X-ray spectrometers and experimental setup

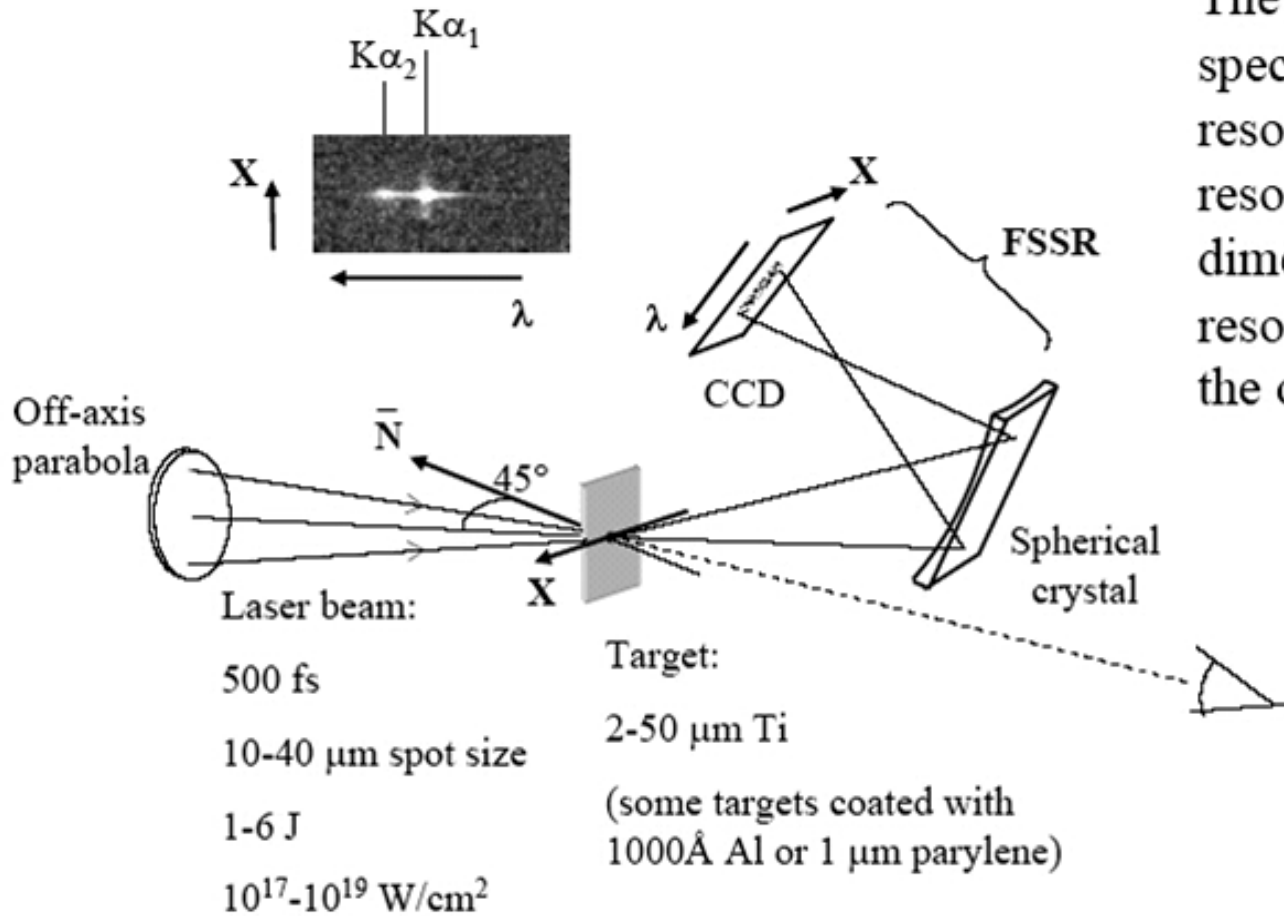


Hot plasma ( $T_e \sim 1$  keV) on the front side of foil target

WDM – on the rare side of the foil target ( $T_e \sim 10$  eV,  $N_e \sim 10^{23}$  cm $^{-3}$ )

FSSR spectrometers provide high spectral and spatial resolution, simultaneously

# Experiments on COMET LLNL facility

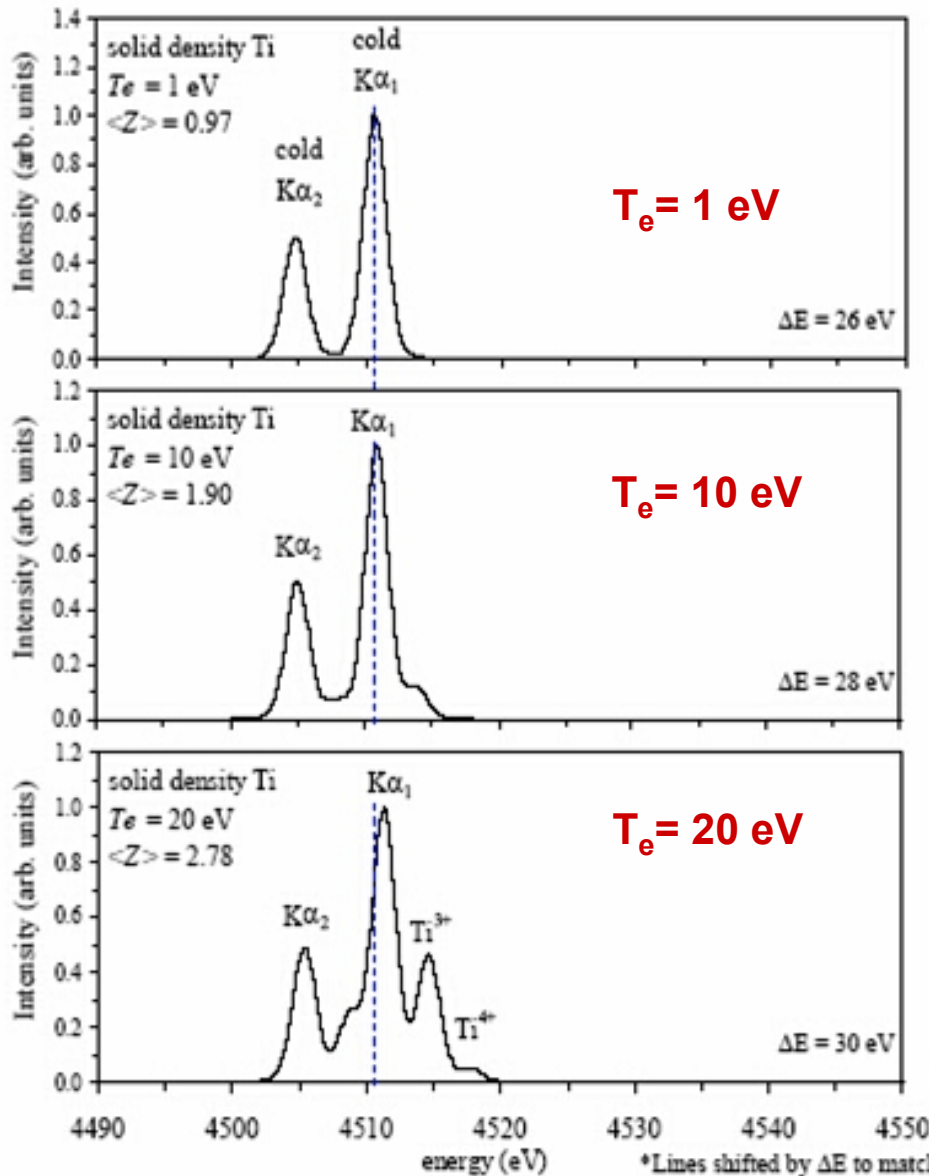


The FSSR-2D (focusing spectrometer with spatial resolution) has spatial resolution of 48  $\mu\text{m}$  in one dimension and spectral resolution  $\lambda/\delta\lambda \sim 3800$  in the other.

Electron and photon counting detectors ( $\lambda/\delta\lambda = 42$ ) also view the back side of the targets.



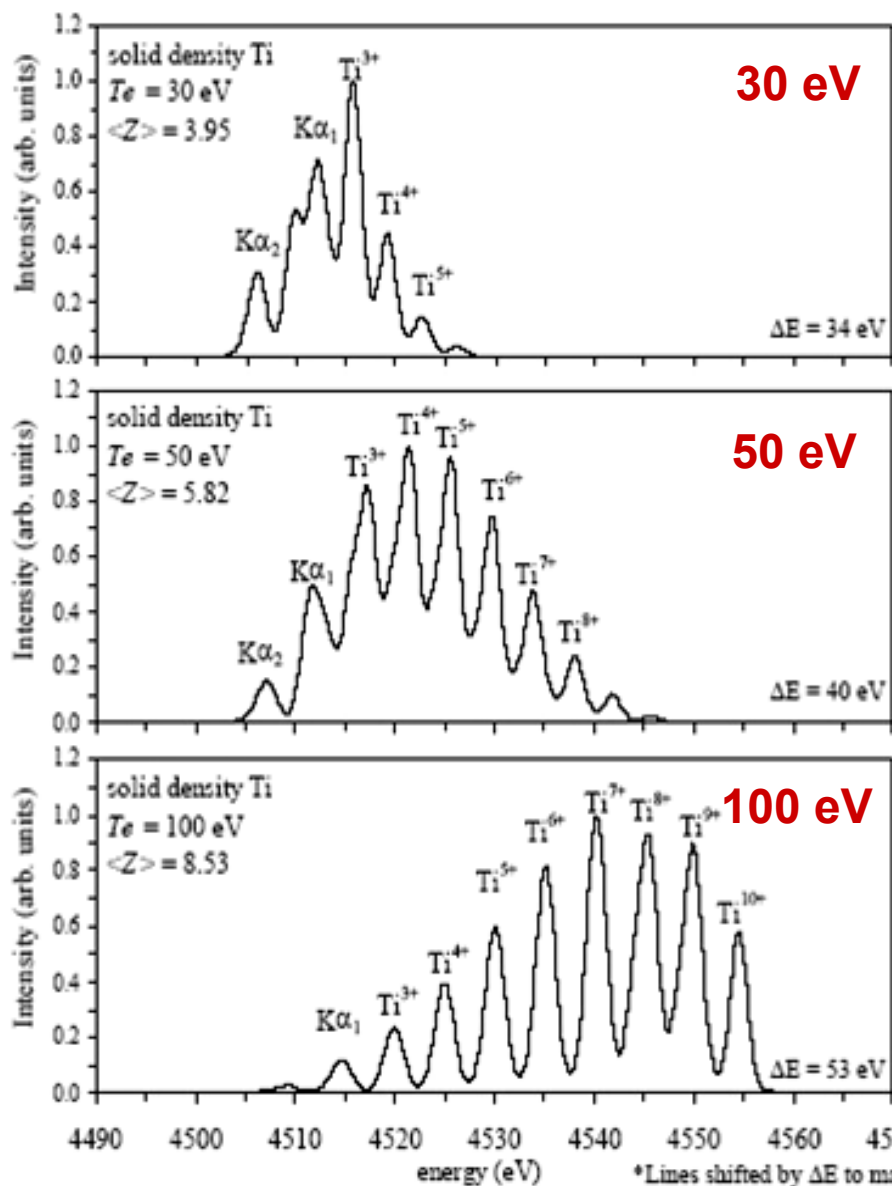
# Inner-shell $K\alpha$ of solid Ti, $T_e = 1 - 20$ eV



As the temperature of the solid increases, ionizing the Ti atoms, the characteristic K-shell lines shift to higher energies. Removing the 4s electrons ( $Ti^{1+}$  and  $Ti^{2+}$ ) produces a shift of only  $\sim 1$  eV, while removing M-shell (3l) electrons ( $Ti^{3+} - Ti^{10+}$ ) produces a  $\sim 4$  eV shift for each charge state.

Between 10 and 20 eV, satellites are evident on the blue sides of the cold  $K\alpha$  lines.

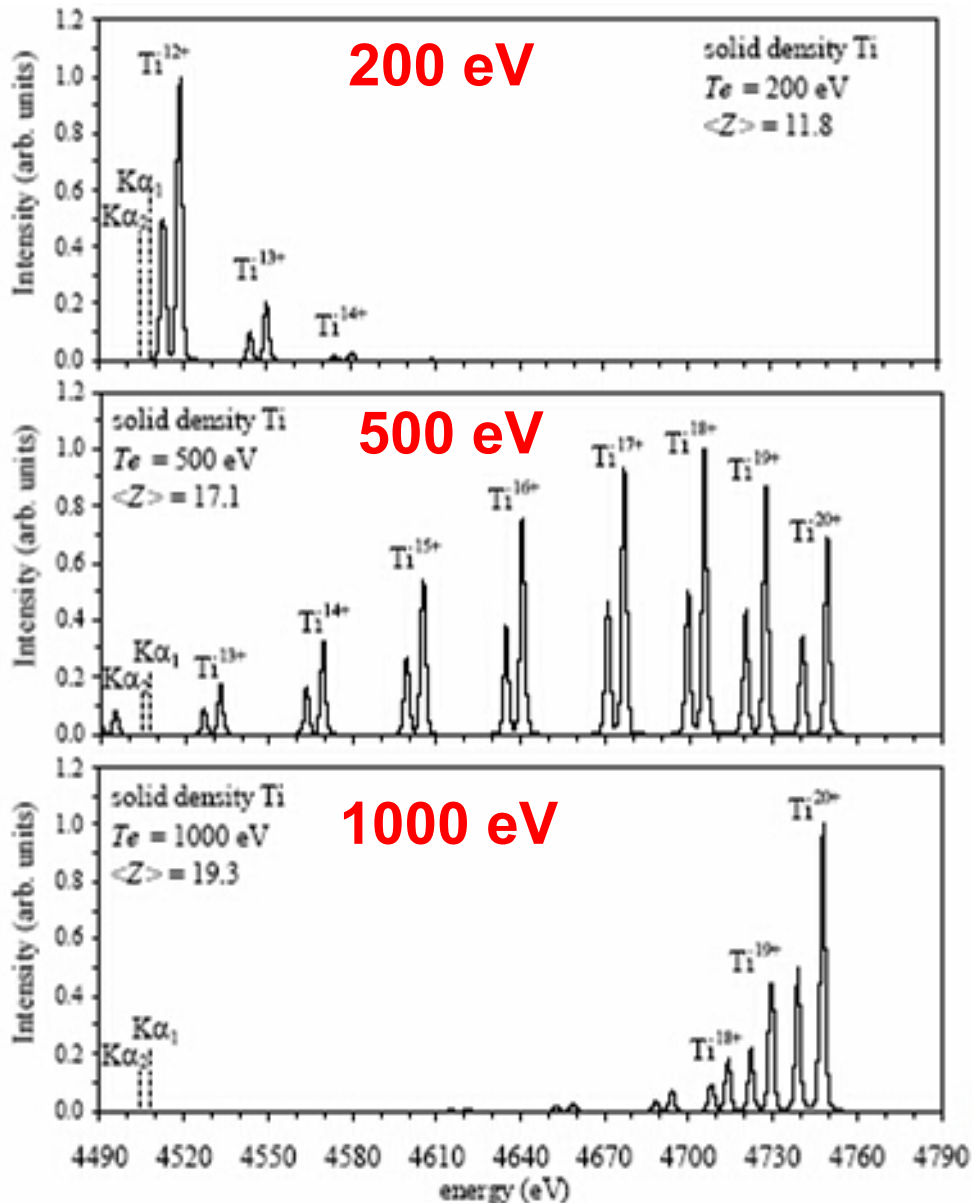
# Inner-shell $K\alpha$ of solid Ti, $T_e = 30 - 100$ eV



With further increases of temperature, Ti burns through the M-shell ions, producing  $K\alpha$  emission shifted by about 4 eV for each charge state. This shift is very near the energy difference between the  $K\alpha_1$  ( $2p_{3/2} - 1s_{1/2}$ ) and  $K\alpha_2$  ( $2p_{1/2} - 1s_{1/2}$ ) lines.

Between 30 and 100 eV, the mean energy of the  $K\alpha$  emission feature shifts by 30 – 50 eV.

# Inner-shell $K\alpha$ of solid Ti, $T_e = 200 - 1000$ eV



As Ti burns through the L-shell ions, the inner-shell  $K\alpha$  emission is shifted by about 30 eV for each charge state; each such shift is comparable to the total shift in mean energy from neutral Ti to  $Ti^{10+}$ .

Above about 1 keV, K-shell ionization stages are reached:

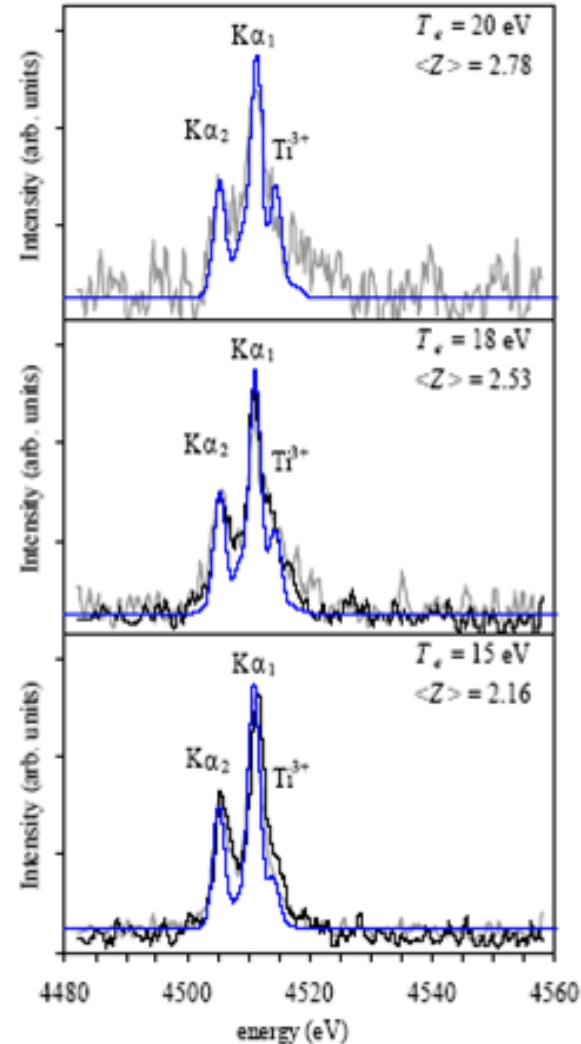
He-like ( $Ti^{20+}$ ): 4750 eV

H-like ( $Ti^{21+}$ ): 4970 eV

# Effect of laser intensity

shot	energy (J)	focus	intensity (W/cm <sup>2</sup> )	K $\alpha_2$ $\Delta x$ ( $\mu$ m)	film images
22-11	4.42	1	$3 \times 10^{18}$	190	
22-16	4.19	1/10	$3 \times 10^{17}$	170	
22-17	4.30	1/10	$3 \times 10^{17}$	220	
22-18	4.19	1/100	$3 \times 10^{16}$	190	
22-19	4.78	1/100	$3 \times 10^{16}$	240	

$T_e$  and  $\langle Z \rangle$  increase with laser intensity

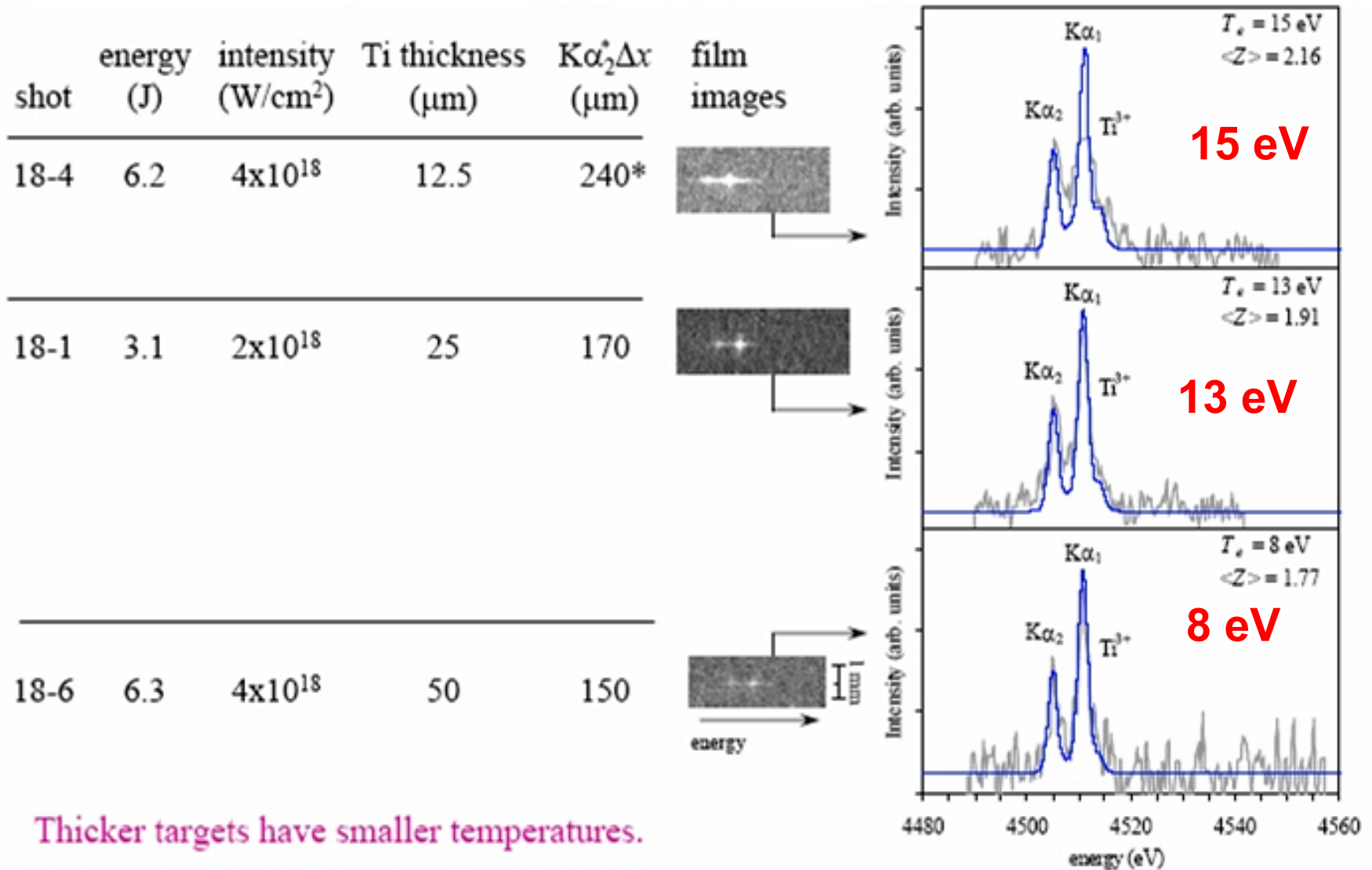


$T_e = 20 \text{ eV}$

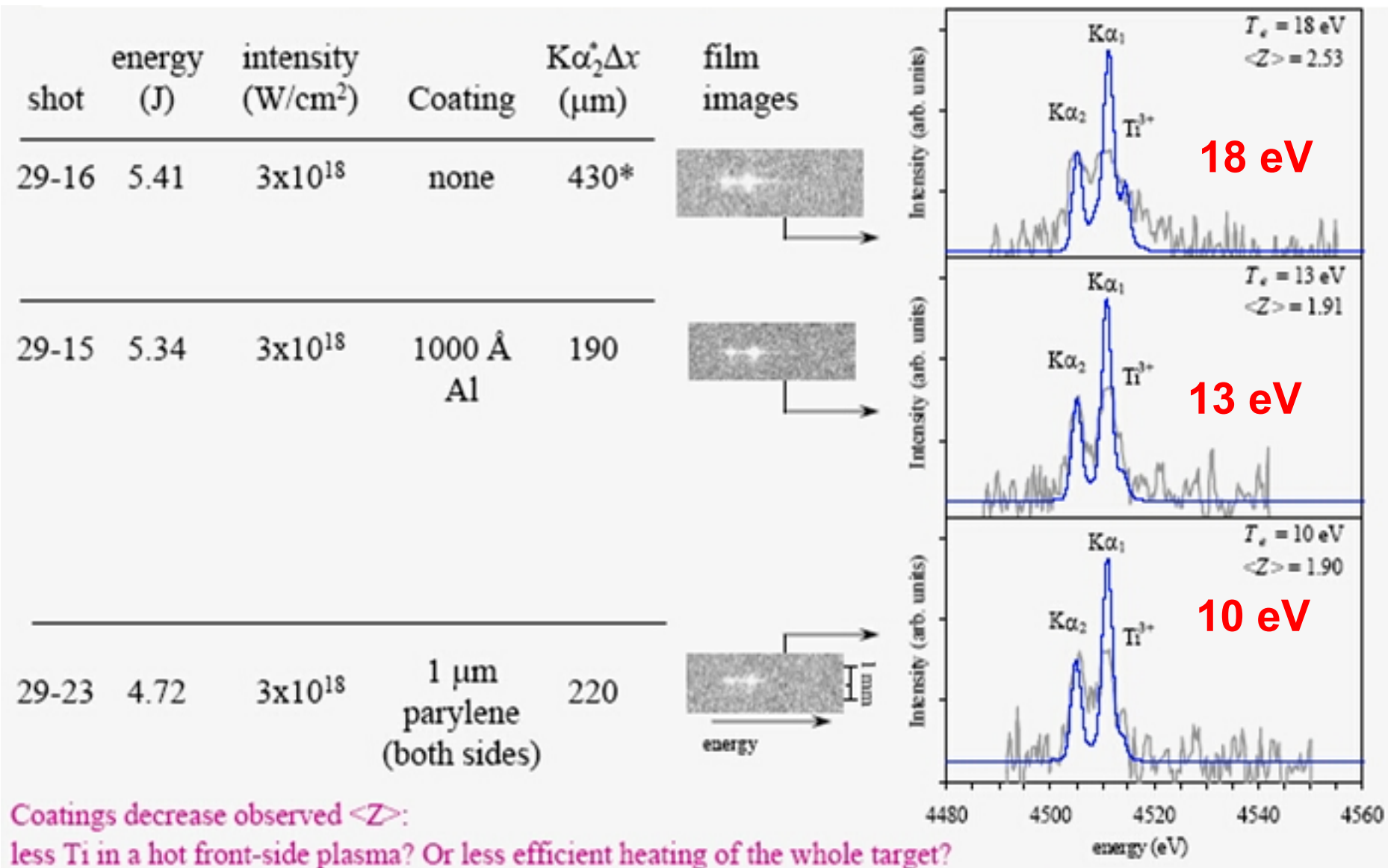
$T_e = 18 \text{ eV}$

$T_e = 15 \text{ eV}$

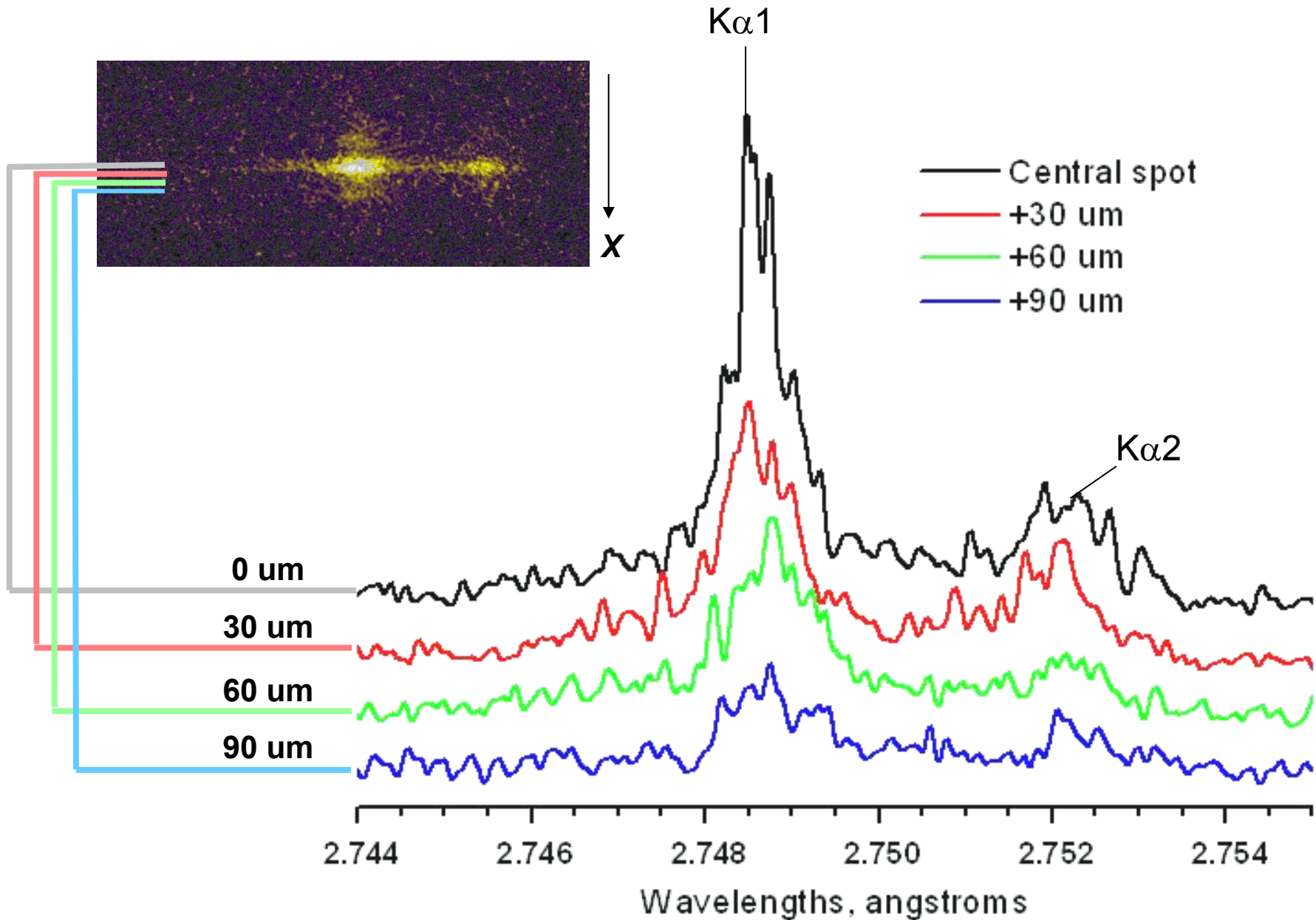
# Effect of Ti foil thickness (Ti +0.1 um Al)



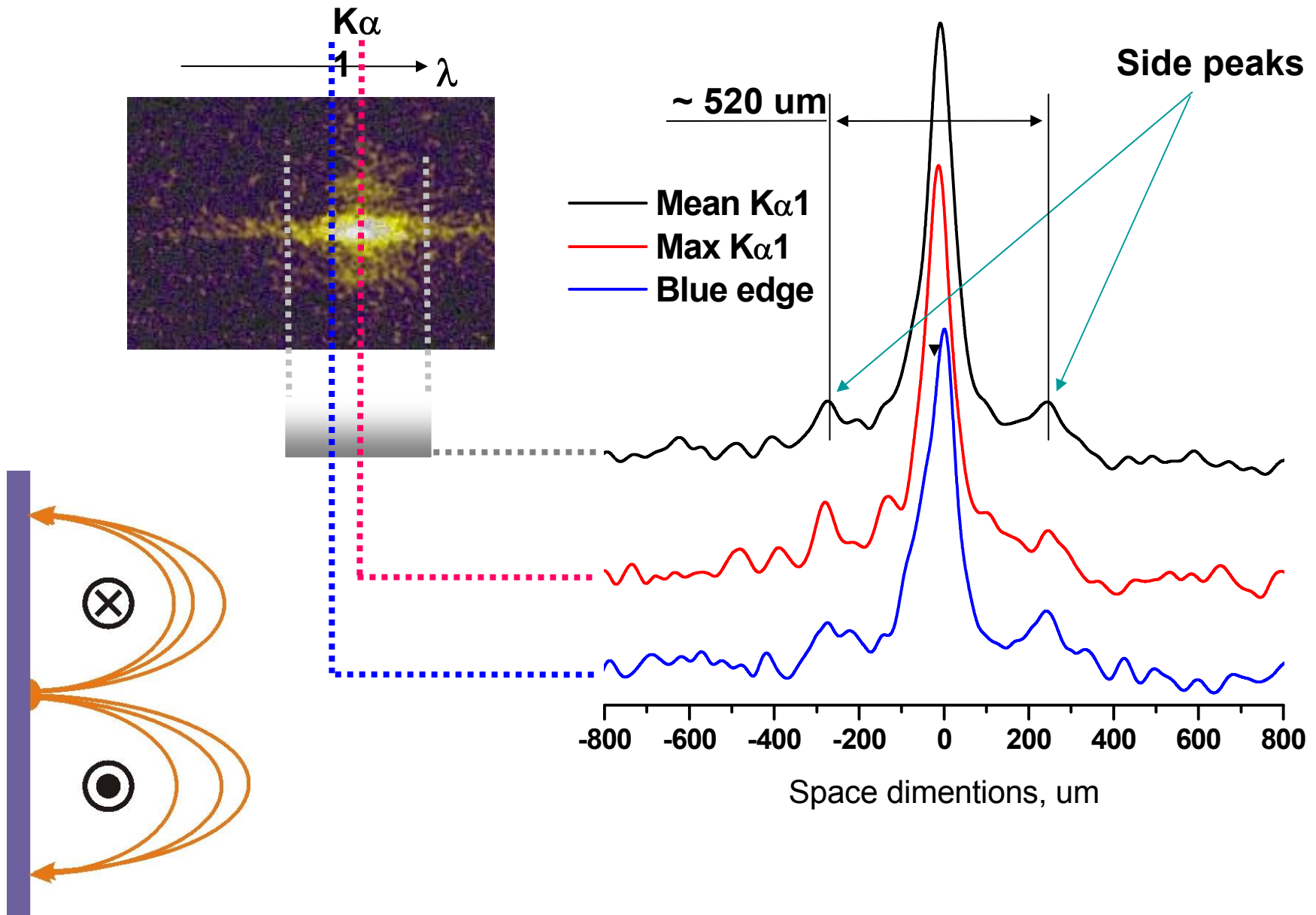
# Effect of coating on 25 um Ti foil



# Spatial resolution along target plane



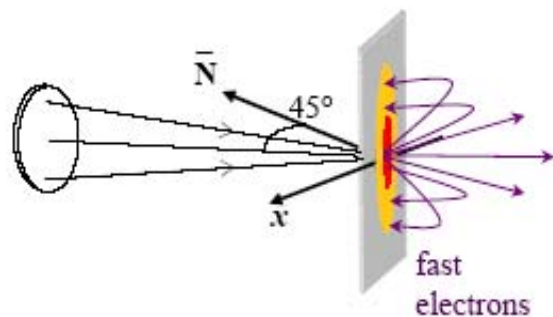
# Spatial resolution and "fountain effect"





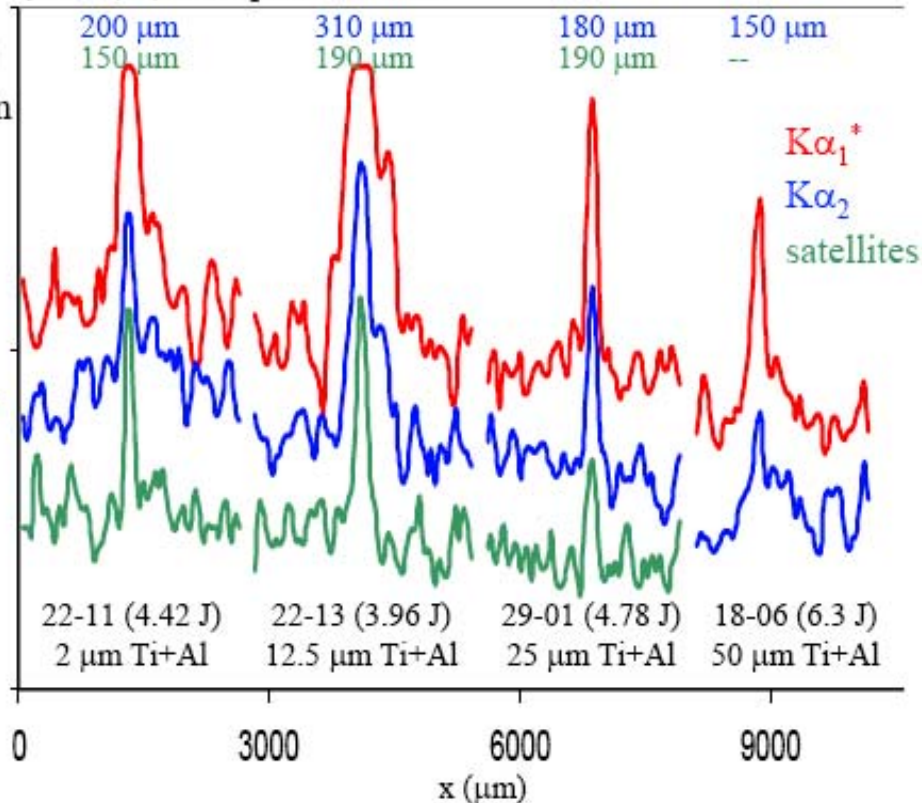
# Line widths change with energy and target

Satellite emission (4515-4520 eV) is narrower than  $K\alpha$  emission in thin targets: evidence for a "fountain effect?"



Or simply an indication of front-side plasma gradients?

*fwhm* ( $\Delta x$ ) of  $K\alpha_2$  and satellites:



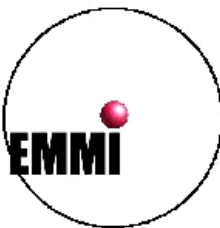
\* Pixel bleeding can occur in  $x$  where lines are saturated

# WDM investigations – suggestions

*We suggest in the frame of the EMMI project:*

*Using ultrahigh laser fluxes  $\sim 10^{19}$ - $10^{20}$  W/cm<sup>2</sup> to increase spectral brightness and find which of excited levels can exist in strongly coupled plasma by mean of registration  $K_{\alpha}$ ,  $K_{\beta}$ ,  $K_{\gamma}$ ,... spectral lines*

*To provide higher energy of hot electrons on front target surface, so the investigations of WDM for large  $Z$  materials will be available.*



# X-ray and ion radiography of WDM

# Laser generated plasma as a source of probe radiation

Intense laser pulse is unique and useful tool to generate a flux of X-ray photons as well as energetic charge particles, simultaneously

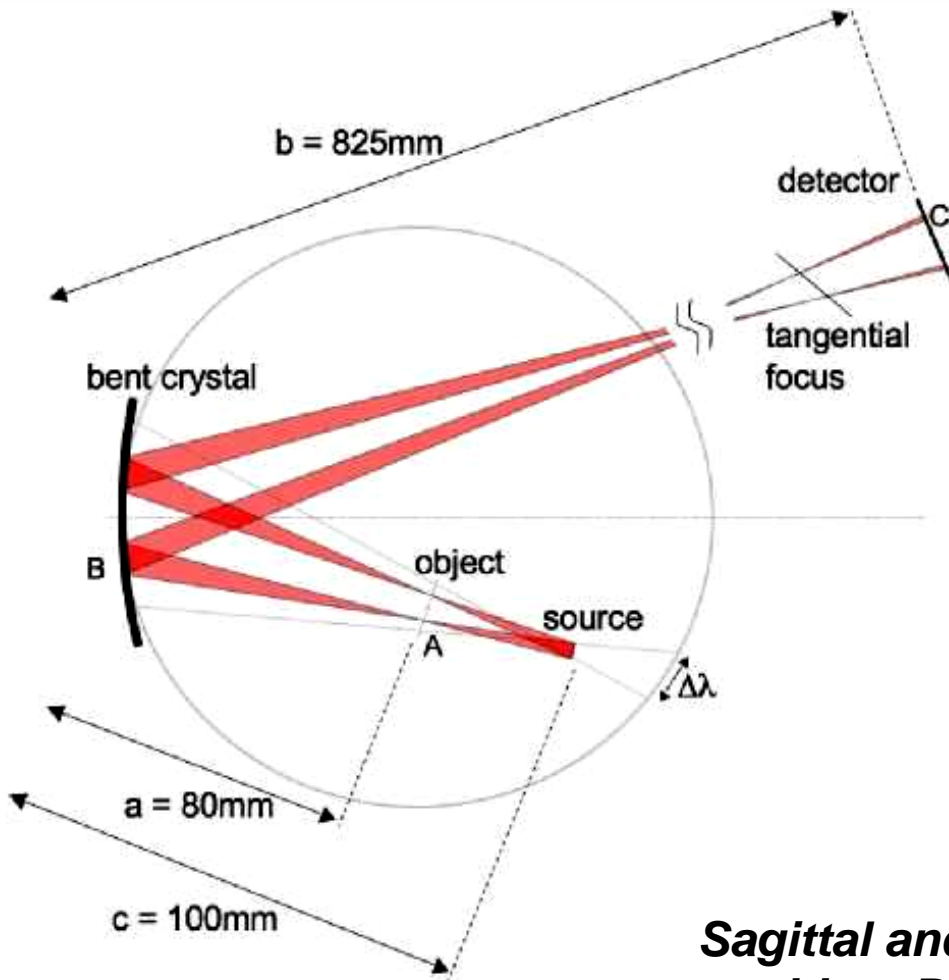
- Both fluxes can be used for shadowgraphy imaging and absorption measurements of different objects and processes including low contrast structures in plasma
- Following diagnostic opportunities can be realized
  - Ion and **proton radiography**
  - **X-ray absorption imaging** and spectroscopy
  - X-ray phase contrast imaging

**Source size** is important to provide spatial and spectral resolution, spatial coherency

**Energy of probing particles** and **monochromatization** are important to provide the sensitivity on density and chemical staff of an object

**Source brightness** and **scheme luminosity** are important to provide wider dynamic range, precise measurements and image quality

# X-ray monochromatic backlighting scheme



## Features:

- Backlighter and object are inside Rowland circle

= High luminosity

= Higher spatial resolution

- Angle of radiation incidence is far from normal one

= Wide spectral range

= Spatial resolution tunable

$$f_t = \frac{aR \sin \theta}{2a - R \sin \theta}, M_t = \frac{c}{c - a} \left[ \left( \frac{2}{R \sin \theta} - \frac{1}{c} \right) b - 1 \right],$$

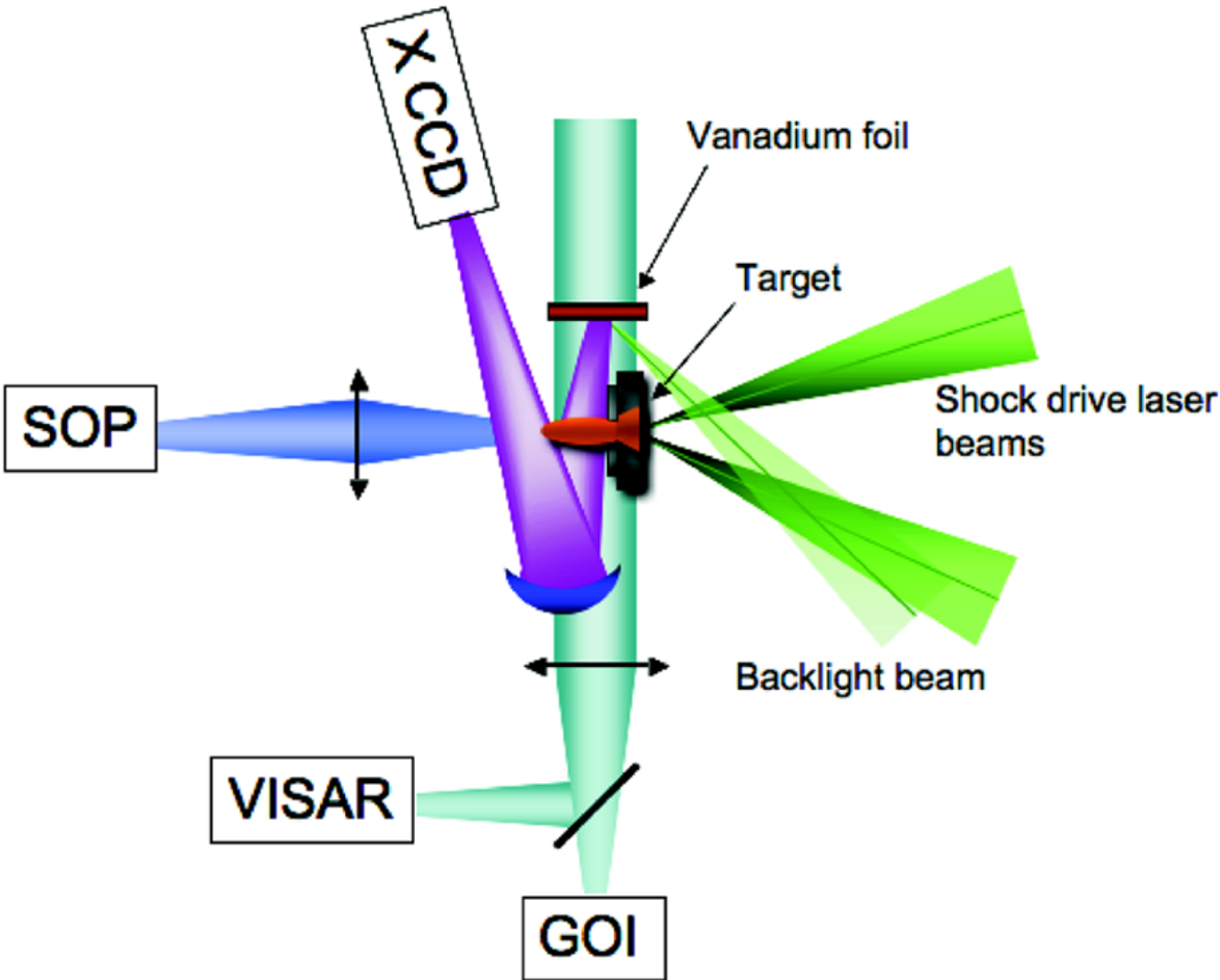
$$f_s = \frac{aR}{2a \sin \theta - R}, M_s = \frac{c}{c - a} \left[ \left( \frac{2 \sin \theta}{R} - \frac{1}{c} \right) b - 1 \right].$$

***Sagittal and tangential focuses are not on the same position. By choosing detector position the spatial resolution along each direction can be tuned up.***



Monochromatic scheme allows to measure object density or chemical staff in precise according to probe radiation absorption

# Experimental setup



## Main laser:

1 ns,  $2\omega$ ,  
300 J,  $d = 400 \mu\text{m}$   
 $I = 5 \cdot 10^{13} \text{ W/cm}^2$

## Backlighting laser:

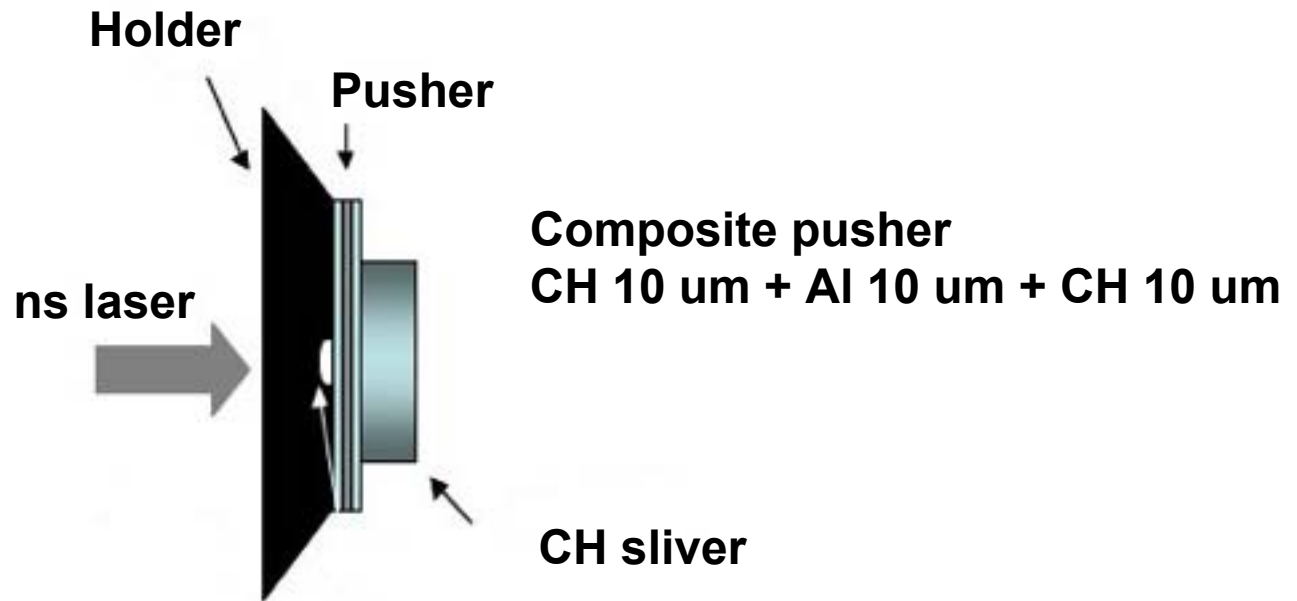
1 ns,  $2\omega$ ,  
250 J,  $d = 150 \mu\text{m}$ ,  
 $I = 10^{15} \text{ W/cm}^2$

## Spherical crystal:

Quartz,  $2d = 4.912 \text{ \AA}$   
II reflection order  
Curvature 150 mm  
Magnification 10

# Target design

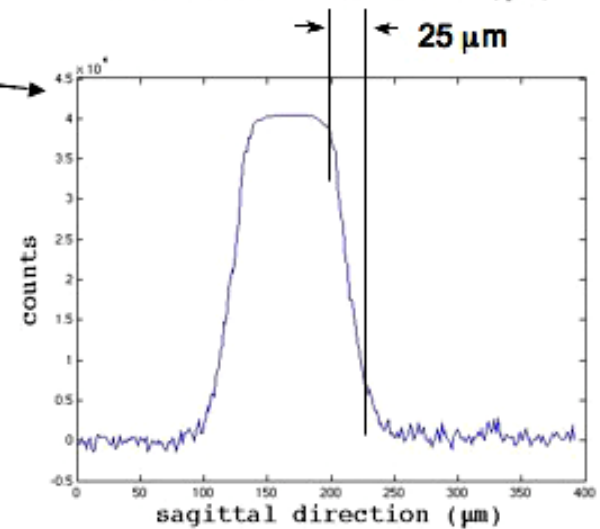
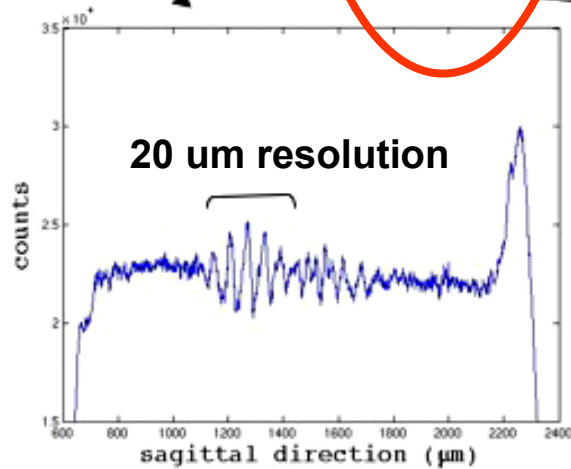
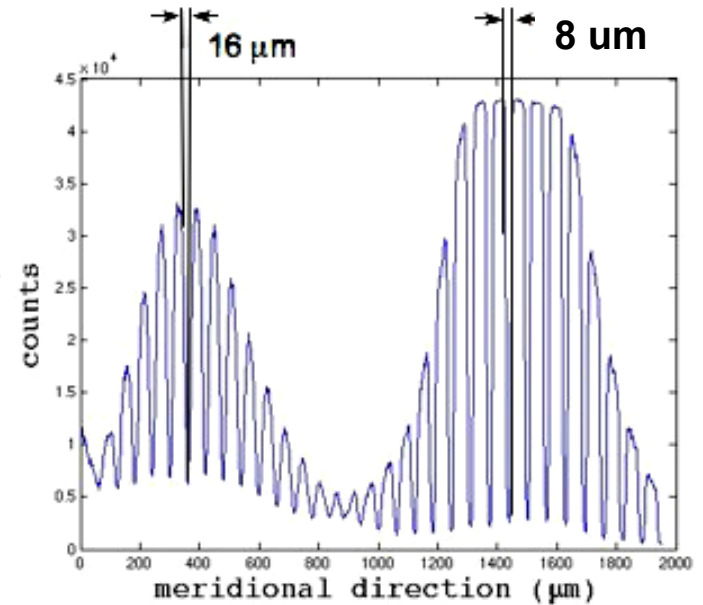
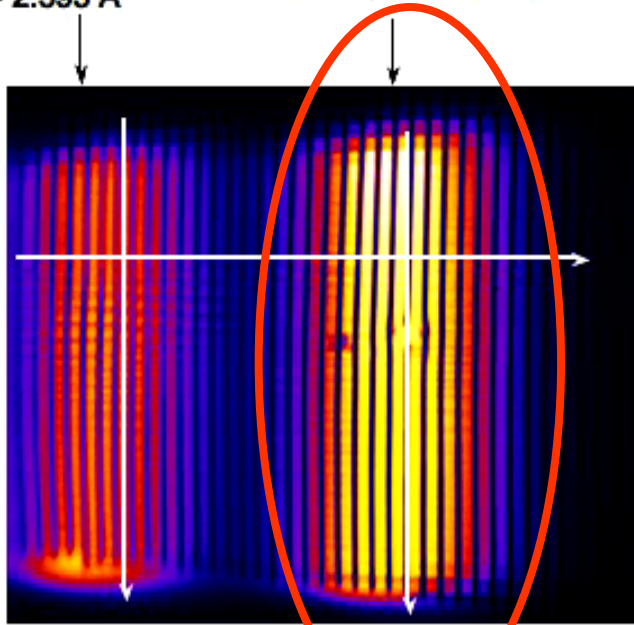
**Shock wave  
generation in  
CH sliver**



# Calibration results

V He<sub>α</sub> Intercombinaison  
line - 2.393 Å

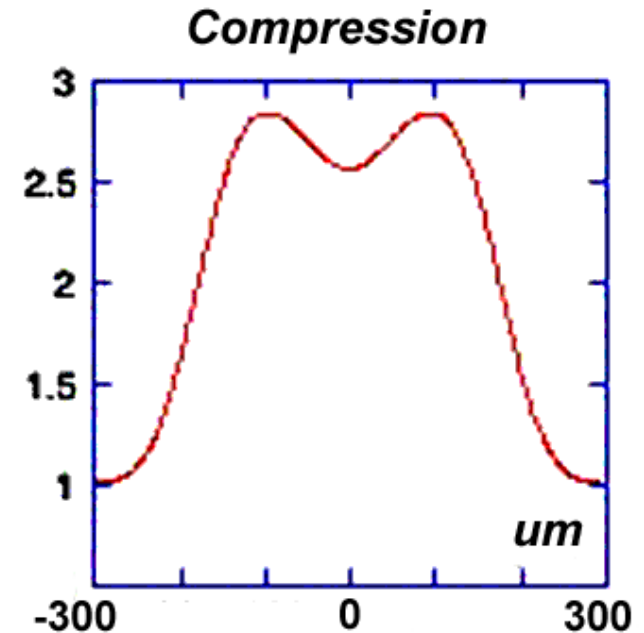
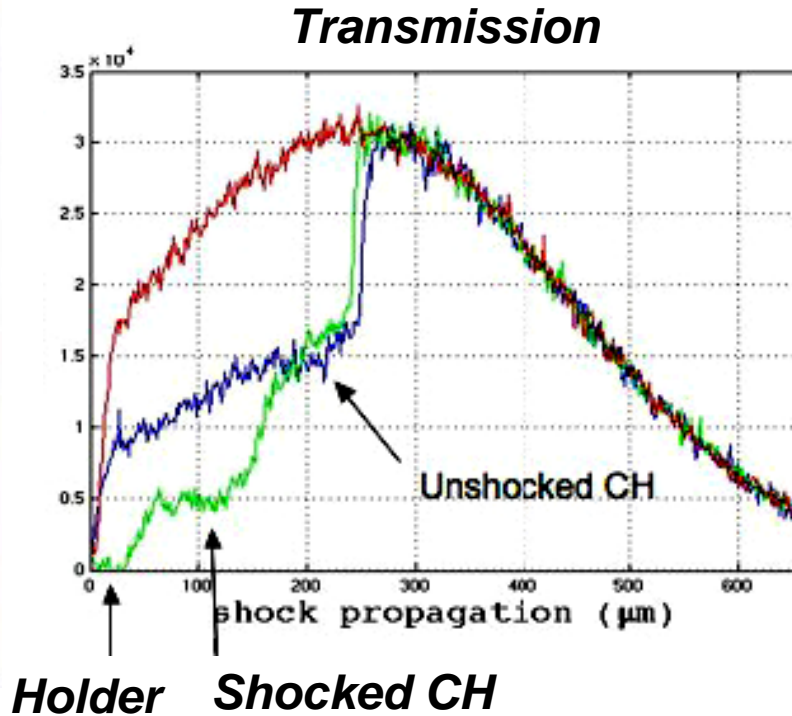
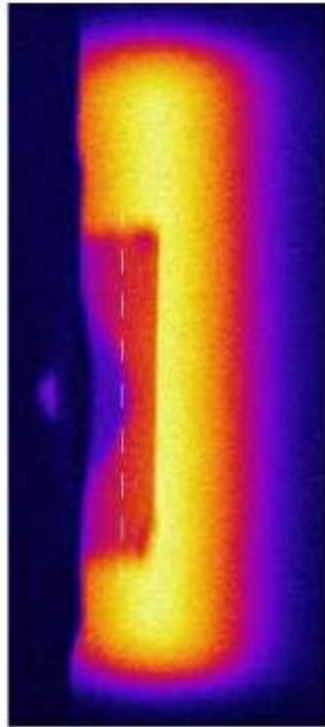
V He<sub>α</sub> satellite line - 2.414 Å





# Shock wave density measurements

X-ray backlighting image in V He $\alpha$  spectral line

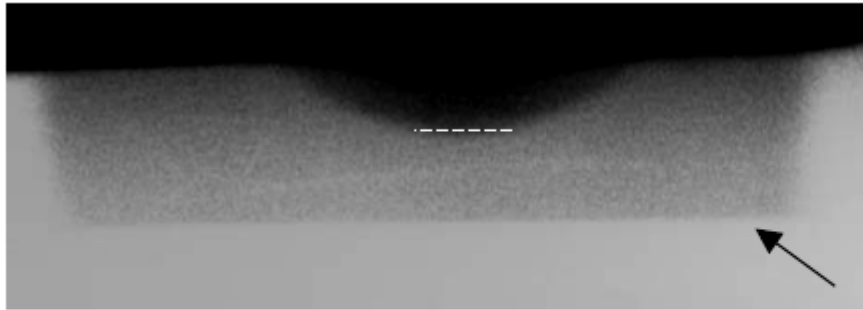


Due to spectral selectivity of backlighting image the density profile of the object can be measured in precise. with spatial resolution along shock front.

The maximum compression of **2.85** is obtained for shock wave in CH sliver

# Shock wave evolution and velocity

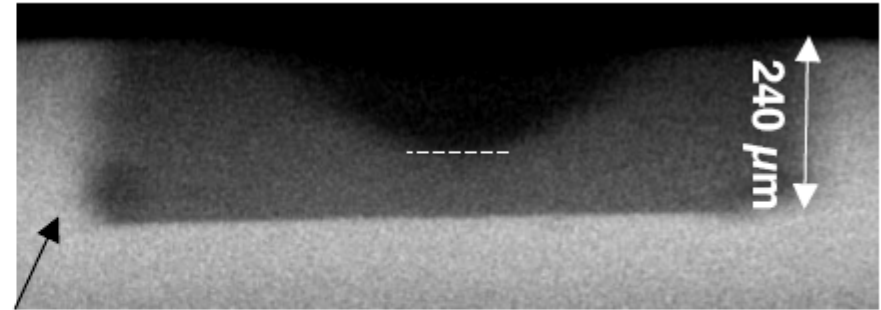
In order to resolve the evolution of imaging object the delay between main and backlighter laser pulses can be aligned



Shock position  $\sim 100 \mu\text{m}$   
at  $t = 5 \text{ ns}$

$t = 5 \text{ ns}$

CH  
sliver

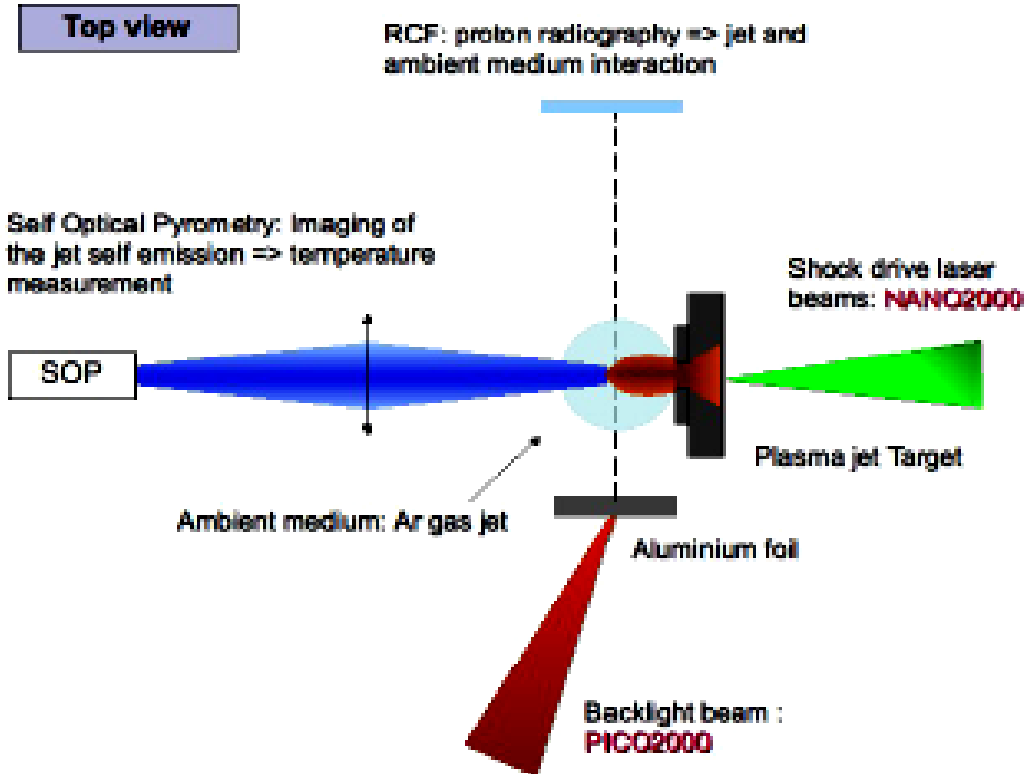


Shock position  $\sim 175 \mu\text{m}$   
at  $t = 9 \text{ ns}$

$t = 9 \text{ ns}$

Shock front velocity is measured to be equal  $19 \text{ km/s}$   
During the propagation the cylindrical symmetry of shock front is stable

# Diagnostic



## Obstacle

Brightness of x-ray source preferably depends on laser energy not by pulse intensity

On upgraded LULI facility **50 J** in **1 ps** is available now instead of **500 J** in **1 ns** before. X-ray source has brightness not enough for monochromatic shadowgraphy imaging.

## Temporary solution:

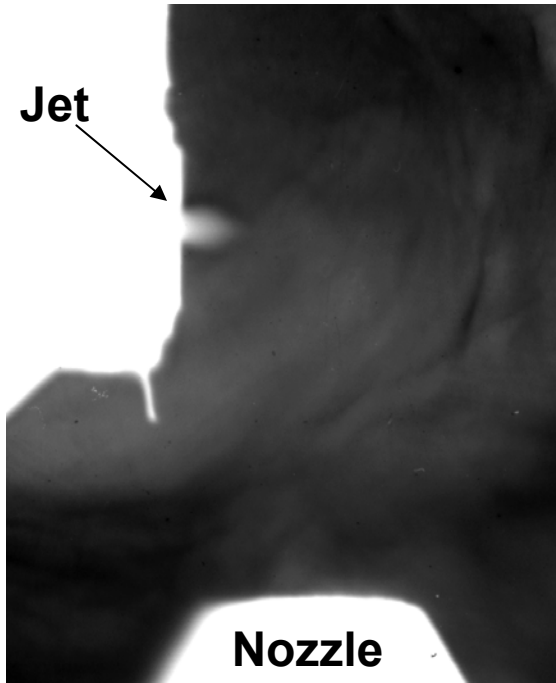
Change to proton radiography diagnostics in quite the same experimental setup.  
RCF film instead of X-ray crystal

## Solution:

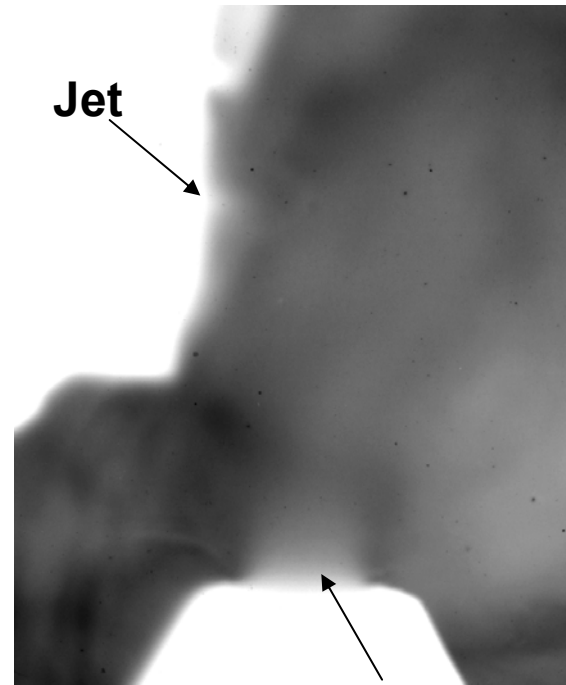
Special design for backlighting target.  
Increase of detector sensitivity

# Proton radiography imaging results

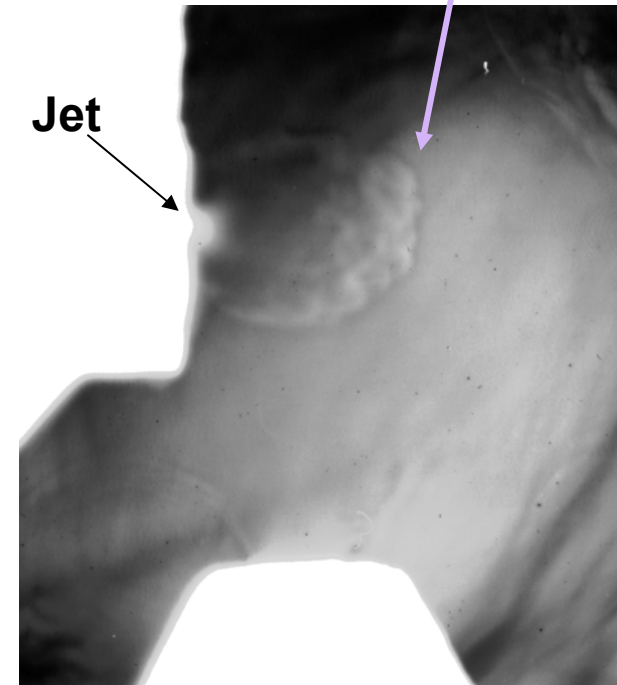
No ambient gas



Ar 20 bar



Secondary structure



Probe proton energy  $\sim 5$  MeV



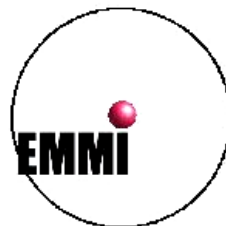
Proton radiography is perfect to image the evolution of plasma structures but not for quantitative measurements on their densities. Further development of soft X-ray diagnostics is on demand.

# WDM investigations – suggestions 2, 3

*We suggest in the frame of the EMMI project:*

*To study shock compression in WDM  
for higher Z materials, metals, by means of x-ray  
monochromatic shadowgraphy and ion radiography.*

*To develop hard X-ray backlighting imaging using fs  
PHELIX beam to produce bright source of probe  
radiation.*



# Conclusion

## Available devices and methods

High resolution spectrometers  
Spherically bent crystals  
Modeling of spectra  
Backlighting schemes  
Absorption and phase contrast

## Forthcoming tasks

Create and study of  
Super dense plasma  
Hot dense plasma  
Warm dense plasma  
Effective sources for  
ion- and proton-radiography  
Images of nanoscale foils and  
biological samples  
Measurement of MG magnetic fields  
Monochromatic imaging of  
shock waves and plasma jets  
Modeling of astrophysical  
phenomena in laboratory  
Investigation of new spectroscopy  
phenomena

