



Susceptibilities from a black hole engineered EoS with a critical point

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Lattice QCD

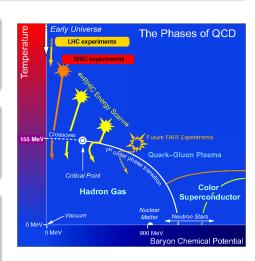
Perform calculations at $\mu_B = 0$, and extrapolate via Taylor expansion to finite μ_B

Black Hole Engineering

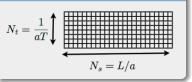
Based on Lattice data at $\mu_B = 0$, allows us to calculate observables at finite density.

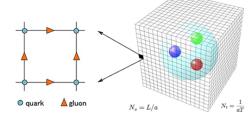
Susceptibilities of Conserved Charges

Are sensitive to the critical point and can be measured









- Study QCD from first principles in the non-perturbative region (Performs path integral using Monte-Carlo technique).
- Calculate equilibrium properties at $\mu_B = 0$ or at imaginary- μ_B (sign problem!). Calculations can be extrapolated to a small regime of μ_B .
- Has technical difficulties to compute transport properties

Black Hole Engineering

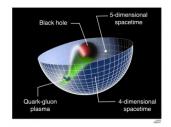


Holography (gauge/string duality at Strong Coupling)

Quantum Field Theory in 4- dimensions



- Coupling >> 1 in QFT \rightarrow vanishing string coupling
- (T, μ_B) in QFT \rightarrow black hole solution
- Holography → Near Perfect fluidity



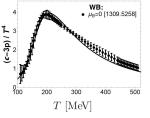
J M Maldacena 1999 Int. J. Theor. Phys. (38) 1113

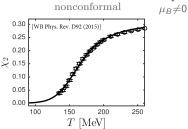


Non-conformal holographic gravity dual in 5 dimensions

$$\Rightarrow$$
 Black Hole Solution

$$S = \frac{1}{16\pi G_5} \int \mathrm{d}x^5 \sqrt{-g} \left[\mathcal{R} - \frac{1}{2} (\partial_M \phi)^2 - \underbrace{V(\phi)}_{g} \right] - \frac{1}{4} \underbrace{f(\phi) F_{MN}^2}_{g}$$





- Input parameters are fixed by lattice QCD results at $\mu_B = 0$
- Finite T and $\mu_B \to \text{Predictions}$

O DeWolfe, S S Gubser, and C Rosen, Phys. Rev. D 83, (2011)

R Rougemont, A Ficnar, S Finazzo and J Noronha, JHEP (2016) 102.

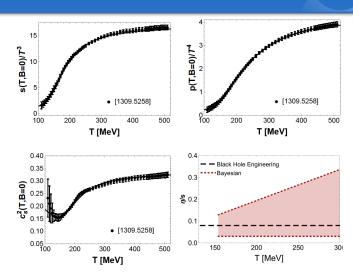
Black Hole Model

- Input parameters are fixed by Lattice data at $(T, \mu_B = 0)$
- Non-conformal Equation of State
 - \blacksquare at finite T and finite μ_B
 - with a critical end point
 - lacksquare agrees with lattice data at small μ_B
- Near perfect fluidity
 - Ability to compute transport coefficients near the crossover and at large μ_B



Model Predictions at $\mu_B = 0$





R Rougemont, A Ficnar, S Finazzo and J Noronha, JHEP (2016) 102.

A system in thermal equilibrium is characterized by

$$Z = Tr \left[-\frac{H - \sum_{i} \mu_{i} Q_{i}}{T} \right]$$

The Pressure

$$P = \frac{T}{V} \ln Z$$

■ The Baryonic Susceptibilities χ_n^B are defined as

$$\chi_n^{\scriptscriptstyle B}(T,\mu_{\scriptscriptstyle B}) = \frac{\partial^n}{\partial (\mu_{\scriptscriptstyle B}/T)^n} \left(\frac{P}{T^4}\right)$$



Susceptibilities of Conserved Charges



- The susceptibilities $\chi_n = \chi_n^B(T, \mu_B)$ are related directly to the moments of the distribution.
- The volume-independent ratios are useful quantities to compare to experimental data.

mean: $M = \chi_1$

variance : $\sigma^2 = \chi_2$

skewness : $S = \chi_3/\chi_2^{3/2}$

kurtosis : $\kappa = \chi_4/\chi_2^2$

$$M/\sigma^{2} = \chi_{1}/\chi_{2}$$

$$S\sigma = \chi_{3}/\chi_{2}$$

$$\kappa\sigma^{2} = \chi_{4}/\chi_{2}$$

$$S\sigma^{3}/M = \chi_{3}/\chi_{1}$$

[STAR] Phys. Rev. Lett. 112 (2014) 032302

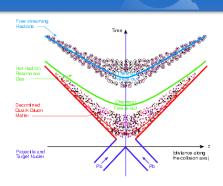
Karsch Central Eur J. Phys. 10 (2012)



Evolution of heavy ion collisions



- Chemical freeze-out: all inelastic interactions cease.
 The chemical composition of the system is fixed
- Kinetic freeze-out: all elastic interactions cease: the spectra of the particles are fixed

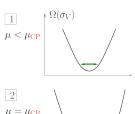


- We want to study the chemical freeze-out
- Observables: susceptibilities of conserved charges
 - They are fixed at the freeze-out
 - They can be measured and calculated
 - They are sensitive to the critical point



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Fluctuations at the Critical End Point







The probability distribution for the order parameter $P[\sigma] \sim \exp\{-\Omega[\sigma]/T\}$

$$\Omega = \int d^3x \left[\frac{1}{2} (\nabla \sigma)^2 + \frac{m_\sigma^2}{2} \sigma_2 + \frac{\lambda_3}{3} \sigma^3 + \cdots \right]$$

The correlation length $(\xi = m_{\sigma}^{-1})$ $\xi \sim |T - T_c|^{-\nu}$ where $\nu > 0$

$$\chi_{2} = VT\xi^{2}$$

$$\chi_{3} = 2VT^{3/2}\hat{\lambda}_{3}\xi^{9/2}$$

$$\chi_{4} = 6VT^{2}[2\hat{\lambda}_{3}^{2} - \hat{\lambda}_{4}]\xi^{7}$$

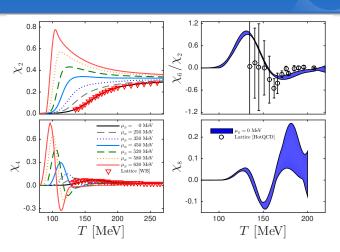
M. A. Stephanov, Phys. Rev. Lett. 102 (2009) 032301





Black Hole Susceptibilities





BH curves: R. Critelli, I. P. et al., to appear.

Lattice results: [WB] Phys. Rev. D, **92**, 114505 (2015).

[HotQCD] Phys. Rev. D, **95**, 054504 (2017).

Observables at finite $\mu_{\it B}$



Taylor expansion of observables in terms of susceptibilities

$$\chi_n = \chi_n^B(T, \mu_B = 0)$$

Pressure

$$\frac{p(T, \mu_B) - p(T, \mu_B = 0)}{T^4} = \sum_{n=1}^{\infty} \frac{\chi_{2n}}{(2n!)} \left(\frac{\mu_B}{T}\right)^{2n}$$

Baryonic density

$$\frac{\rho_B(T, \mu_B)}{T^3} = \sum_{n=1}^{\infty} \frac{\chi_{2n}}{(2n-1)!} \left(\frac{\mu_B}{T}\right)^{2n-1}$$

■ Susceptibilities χ_2 and χ_4

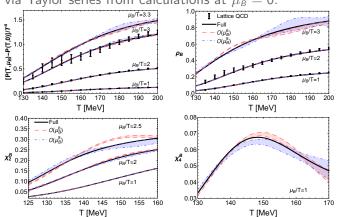
$$\chi_{2}(T, \mu_{B}) = \sum_{n=0}^{\infty} \frac{\chi_{2n+2}}{(2n!)} \left(\frac{\mu_{B}}{T}\right)^{2n} \quad \chi_{4}(T, \mu_{B}) = \sum_{n=0}^{\infty} \frac{\chi_{2n+4}}{(2n!)} \left(\frac{\mu_{B}}{T}\right)^{2n}$$



Taylor Reconstruction up to $\mathcal{O}(\mu_B^8)$



Reconstruction of thermodynamic quantities at different values of μ_B/T via Taylor series from calculations at $\mu_B=0$.

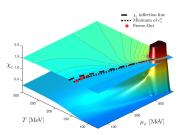


BH curves: R. Critelli, I. P. et al., to appear.

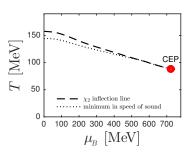
Lattice results: [HotQCD] Phys. Rev. D, 95, 054504 (2017).

The black hole model contains a critical end point at

$$\mu_{\rm B} = 723 \pm 36 \; {\rm MeV}$$



$$T = 89 \pm 11 \text{ MeV}$$



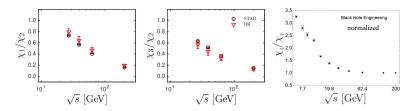
R. Critelli, I. P. et al., to appear.



Connection to Experiment



- We compare the baryonic BH susceptibilities ratios with the net-proton moments measured at STAR
- Freeze-out parameters are extracted by fitting the experimental values for χ_1/χ_2 and χ_3/χ_2
- $= \chi_4/\chi_2$ predicted at the minimum in speed of sound



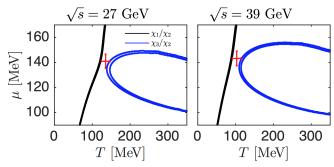
R. Critelli, I. P. et al., to appear. [STAR] Phys. Rev. Lett. 112 (2014)



Freeze out parameters from the Black Hole model



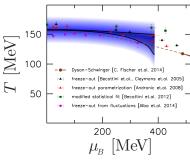
Trajectories in the $[T - \mu]$ plane that satisfy the experimental values



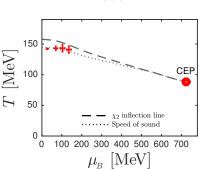
■ Freeze out points $[T - \mu_B]$ are extracted from the line made by the closest points between χ_1/χ_2 and χ_3/χ_2







BH Model

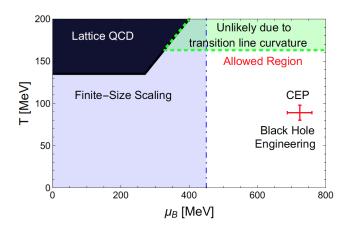


[WB:] R. Bellwied et. al.,

Phys. Lett. B 751 (2015) 053

R. Critelli, I. P. et al., to appear.





R. Critelli, I. P. et al., to appear.

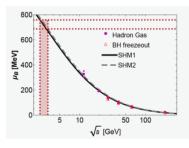


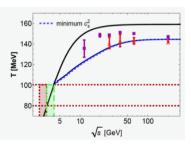
Collision Energy Estimates



We estimate a collision energy needed to hit the CEP

$$\sqrt{s} = 2.5 - 4.1 \text{ GeV}$$





 The collision energy is reachable by the next generation of colliders

[BH] R. Critelli, I. P. et al., to appear.

[HRG] Paolo Alba et al. Phys. Lett. B738 (2014),

[SHM1] A. Andronic et al. Phys. Lett. B673 (2009). [SHM2] J. Cleymans et al. Phys. Rev. C73 (2006).



The holographic Black Hole Model

- Reproduces lattice data at $\mu_B = 0$
- Contains a critical end point at $\mu_{\rm B}\!=\!723\pm36$ MeV and $T\!=\!89\pm11$ MeV
- Allows us to compute baryonic susceptibilities, and extract freeze-out parameters
- Estimates that the collision energy needed to hit the CEP should be $\sqrt{s} = 2.5 4.1 \text{ GeV}$