
TOPICS IN HALO EFT

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OUTLINE

- Subtractive renormalization
 - Seduced by the $\Lambda \rightarrow \infty$ limit
 - Some wins for perturbative power counting
 - Coulomb energies and isospin symmetry in Halo EFT
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SUBTRACTIVE RENORMALIZATION

Afnan & Phillips, PRC, 2004

Yang, Elster, Phillips, PRC, 2008

- Consider the zero-energy amplitude resulting from a long-range potential v , and a contact term, C

$$T(p, 0; 0) = [v(p, 0) + C] + \int_0^\Lambda dp' p'^2 [v(p, p') + C] G_0(p'; 0) T(p', 0; 0)$$

$$G_0(q; E) = \frac{1}{E - q^2 / (2m_R)}$$

- Take difference of $T(p, 0; 0)$ and $T(0, 0; 0)$

$$T(p, 0; 0) = T(0, 0; 0) + [v(p, 0) - v(0, 0)] + \int_0^\Lambda dp' p'^2 [v(p, p') - v(0, p')] G_0(p'; 0) T(p', 0; 0)$$

Hammer & Mehen, NPA, 2001

- Same trick suffices to compute $T(p, p'; 0)$:

$$T(p, p'; 0) = T(0, p'; 0) + [v(p, p') - v(0, p')] + \int_0^\Lambda dq q^2 [v(p, q) - v(0, q)] G_0(q; 0) T(q, p'; 0)$$

WHAT DO WE LEARN?

- Nothing here that can't be done in original formulation
 - Avoiding computation of $C(\Lambda)$ allows $\Lambda \rightarrow \infty$ limit to be straightforwardly taken
 - Off-shell behavior of zero-energy amplitude entirely determined by scattering length and differences of v
 - Kernel is negative definite, therefore equation can be used, together with RG for T with Λ , to show that effects of cutoff in T are of relative order p^3
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FINITE ENERGIES

- Off-shell behavior at one energy suffices to get T for all energies

$$T(E) = [v + C] + [v + C]G_0(E)T(E)$$

$$T(E) = T(0) + T(0)[G_0(E) - G_0(0)]T(E)$$

“First-resolvent method”

- For small E , high-momentum behavior will be as for $T(0)$

- Extensions:

- Higher partial waves

Yang, Elster, Phillips, PRC, 2009

- Energy-dependent potential

Afnan, Phillips, PRC, 2004

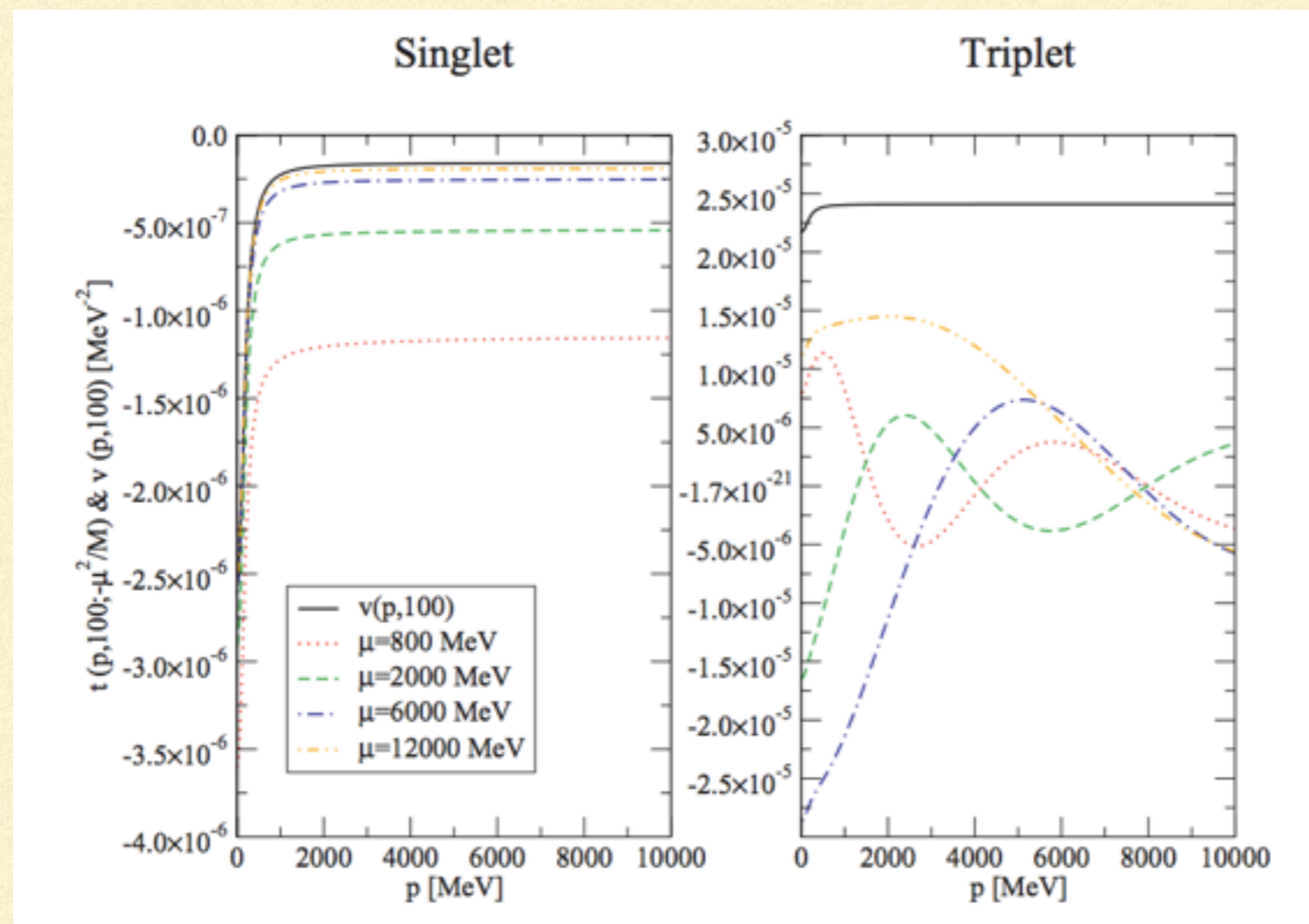
- Contact terms $\sim p^2$

Yang Elster, Phillips, PRC, 2009

NOT THAT SUBTRACTION METHOD

- Cf. Frederico et al., who perform difference with $T(-\mu^2/2m_R)$ and then assume Born approximation valid for latter

Frederico, Timoteo, Tomio, PRC, 2007



Yang, Elster, Phillips,
PRC, 2007

Born approximation never holds for singular potential

APPLICATION: ATOM-DIMER SCATTERING AT N²LO

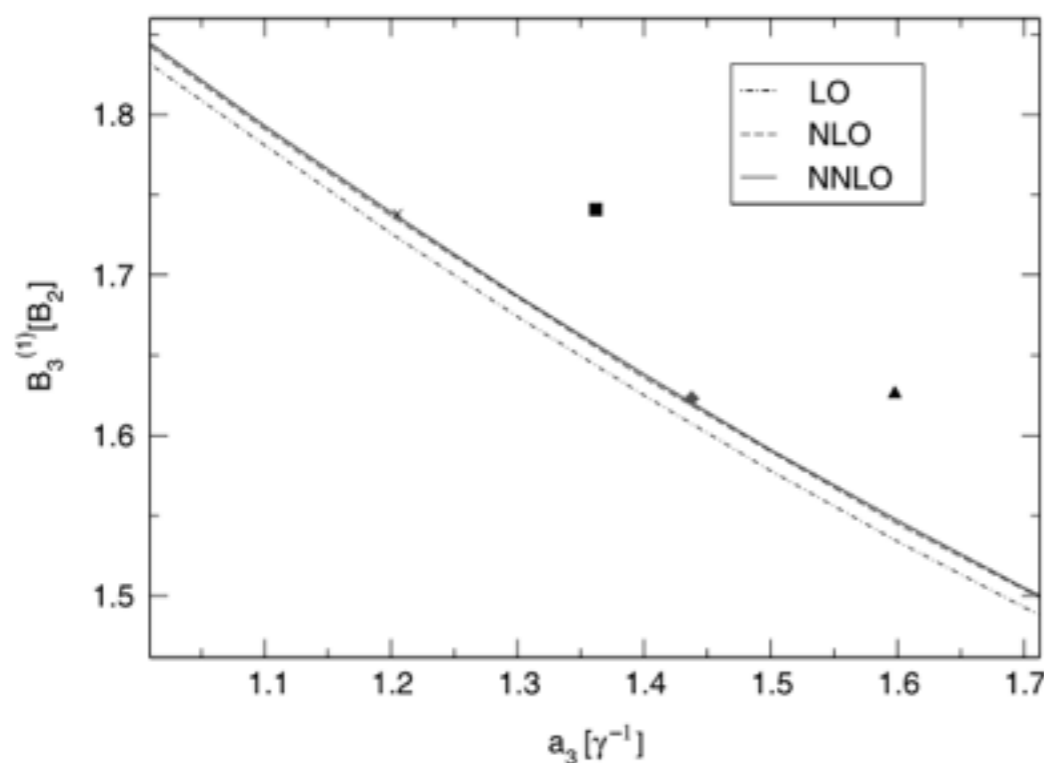
Platter, Phillips, FBS, 2006

- Employ “partial resummation”: take nth-order kernel in three-body integral equation, and solve for amplitude

“Weinberg counting”

Bedquae, Griesshammer, Hammer, Rupak, NPA, 2003

$$K^{(n)}(p, 0; -B_2) = \frac{2m}{p^2 + \gamma^2} + \frac{4m^2}{3} \int_0^\Lambda dq \frac{1}{pq} \log \left(\frac{p^2 + q^2 + pq + \gamma^2}{p^2 + q^2 - pq + \gamma^2} \right) \\ \times S^{(n)}(-B_2; q) K^{(n)}(q, 0; -B_2),$$

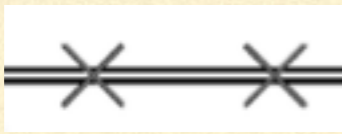


$S^{(n)} \sim q^{n+1}$ for large q

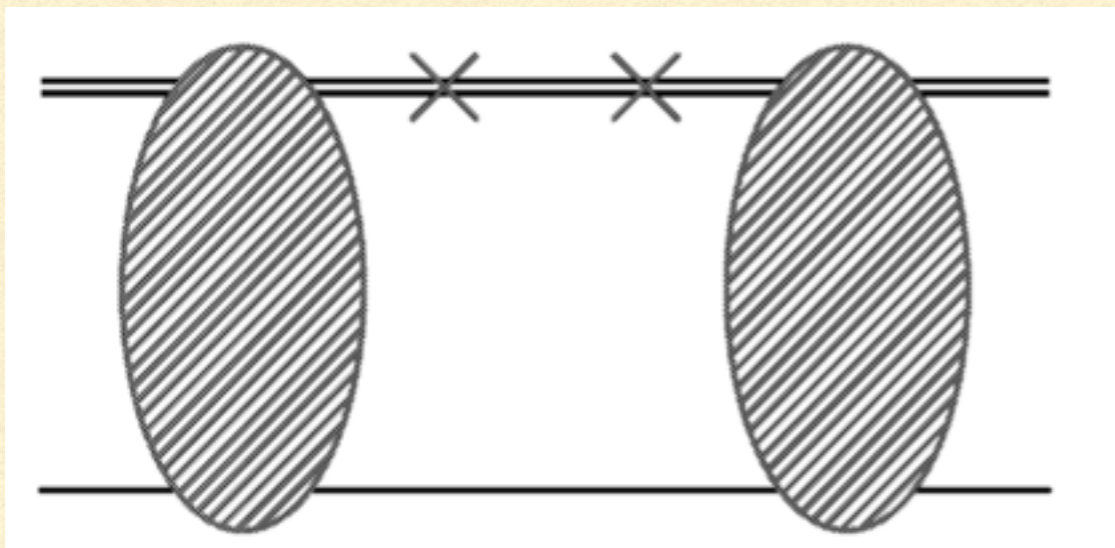
- After subtraction, can be solved numerically for $\Lambda \rightarrow \infty$
- Cutoff independent results at N²LO

THE H₂ CONFLICT

- This suggests no additional three-body force is needed at N²LO
- In contradistinction to the findings of BGHR



$$\sim \frac{2\pi}{m_R} \frac{1}{\gamma + ip} \left(\frac{\rho}{2} \frac{\gamma^2 + p^2}{\gamma + ip} \right)^2 \sim p \text{ as } p \rightarrow \infty$$



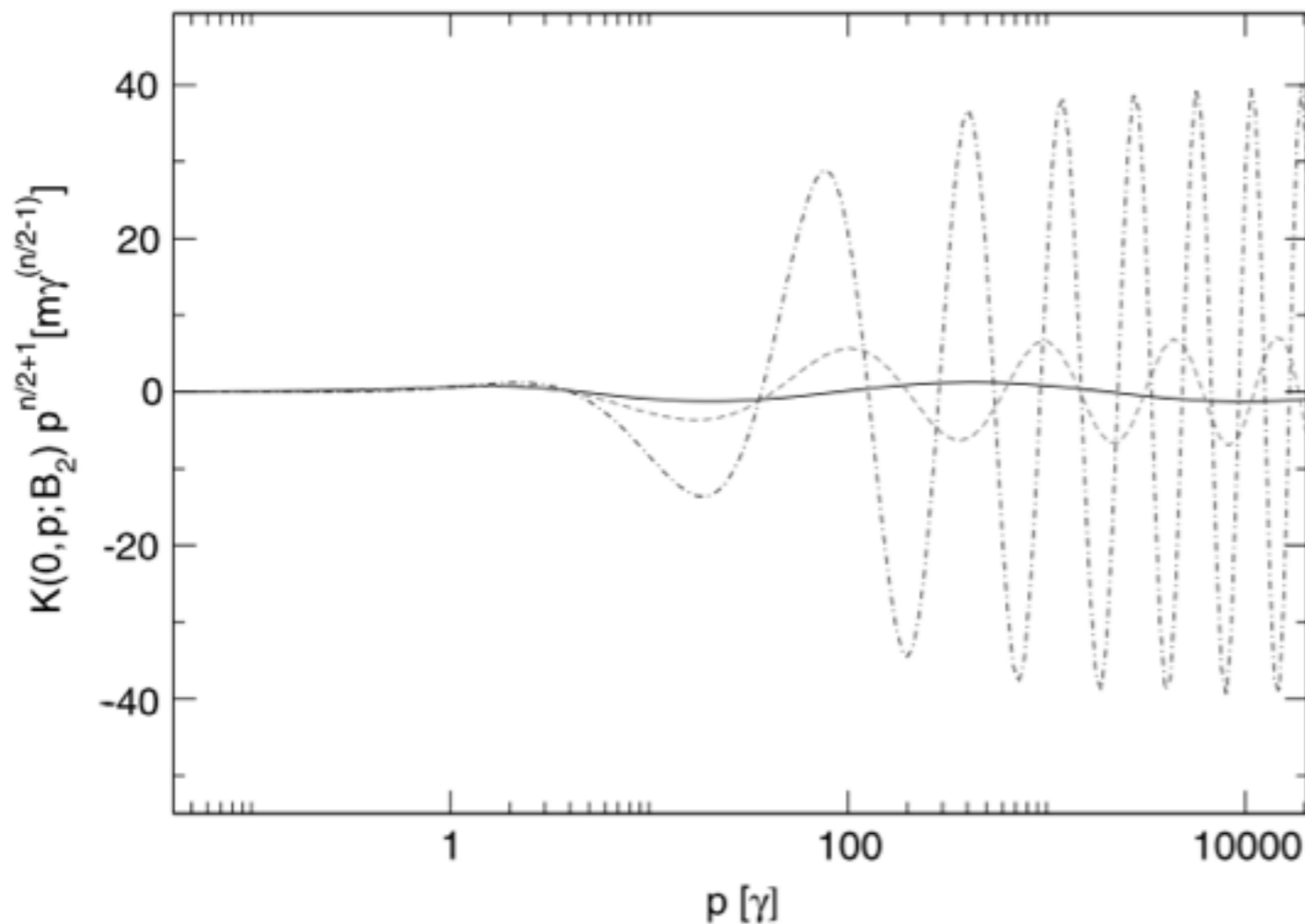
$$\sim \frac{1}{p^{1+is_0}} p \frac{1}{p^{1+is_0}} p^3$$

Anticipate $\Lambda^2 + m_R E \ln(\Lambda)$

- Buttressed by slope of Lepage plot at low cutoffs
- Confirmed in subsequent analysis

Chen, Phillips, FBS, 2013
Vnnasse, PRC, 2014

CONFLICT RESOLUTION



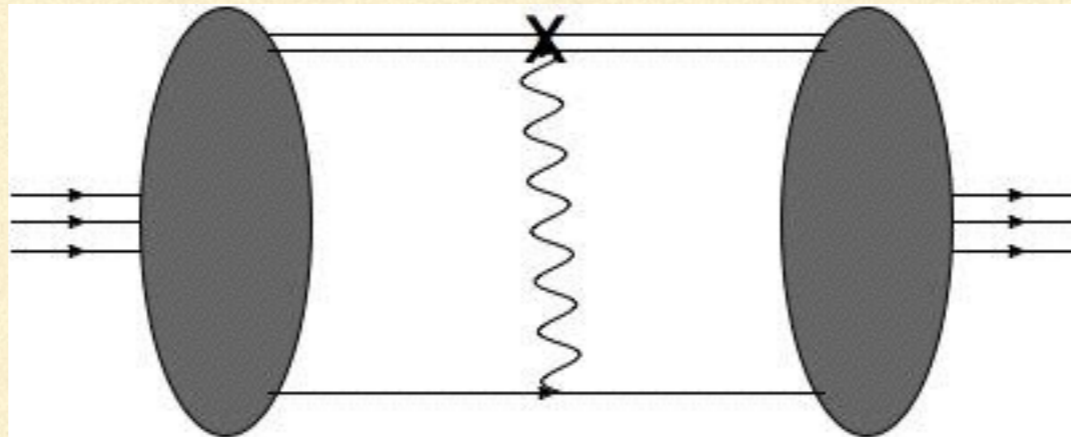
Non-perturbative
treatment of kernel
modifies high-p amplitude

$$T^{(n)} \sim 1/p^{1+n/2} \text{ (Oscillatory)} \\ \text{as } p \rightarrow \infty$$

- Harder kernel \Rightarrow softer amplitude (cf. NN scattering)
- Platter-Phillips results emerge only if $|a| \gg |r| \gg \ell$

PERTURBATIVE COULOMB

- NLO graph:



$$\text{Graph} \sim \frac{1}{p^{1+is_0}} \frac{1}{p} \frac{1}{(p-p')^2} \frac{1}{p'} \frac{1}{p'^{1+is_0}} p^3 p'^3$$

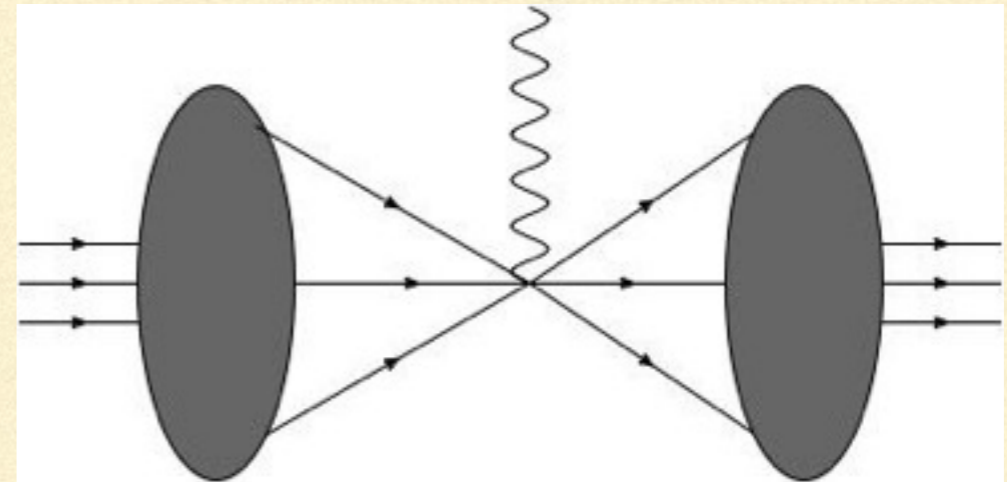
- In co-ordinate space

$$\text{Graph} \sim \int_R^\infty dr_{pd} r_{pd}^2 \frac{1}{r_{pd}} \frac{\alpha_{em}}{r_{pd}} \frac{1}{r_{pd}} (\text{Oscillatory in } r_{pd})$$

- Anticipate $\alpha_{em} \ln(R)$ divergence at NLO

IMPLICATIONS FOR EM OPERATORS

- Short-distance contribution to tri-nucleon form factors



- Unitary limit, analog of reduced radial wave function

$$f_n(\rho) = 2\kappa_n \sqrt{\frac{\sinh(\pi s_0)}{\pi s_0}} \rho^{1/2} K_{is_0}(\sqrt{2}\kappa_n \rho)$$
$$\langle \rho^2 \rangle = \kappa^2 \int_R^\infty K_{is_0}(\sqrt{2}\kappa \rho) \rho^3 K_{is_0}(\sqrt{2}\kappa \rho) d\rho$$

Short-distance contribution $\sim \kappa^2 R^4$ cf. $1/\kappa^2$ at LO

Vanasse result for ${}^3\text{H}$: $\langle r^2 \rangle^{1/2} = 1.13 + 0.46 + 0.27 \pm 0.07$ fm

THE STRONG PC SCATTERING LENGTH

- ‘Strong’ proton-core scattering length is defined as the proton-core scattering length when the Coulomb potential is off

- In Halo EFT with PDS it is $\frac{1}{a_{pc}} = \frac{2\pi}{C_0(\mu)m_R} + \mu$

- Relationship to observable, a_C :

$$\frac{1}{a_{pc}^{MS}(\mu)} = \frac{1}{a_C} + 2k_C \left[\ln \left(\frac{\sqrt{\pi}\mu}{2k_C} \right) + 1 - \frac{3C_E}{2} \right]$$

Kong & Ravndal, 1998; Gegelia, 2001; Higa, Hammer, van Kolck, 2008; Ryberg et al., 2014

- Scheme and scale dependent
 - Includes effects of photons “above μ ”: but dressed by strong interactions, “Coulomb-nuclear interference”
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BINDING ENERGY SHIFTS DUE TO COULOMB

$$H_{\text{strong}}|\psi_s\rangle = E_{\text{strong}}|\psi_s\rangle;$$

$$(H_{\text{strong}} + V_C)|\psi\rangle = E|\psi\rangle$$

- Coulomb energy then defined as $E - E_{\text{strong}}$
 - Coulomb energy of a proton halo is scheme and scale dependent
 - Thomas-Ehrman shift?
 - Recent evaluation of ${}^3\text{He}$ - ${}^3\text{H}$ binding-energy difference in expansion about unitary limit: $B({}^3\text{He}) - B({}^3\text{H}) = -0.86 \pm 0.17 \text{ MeV}$

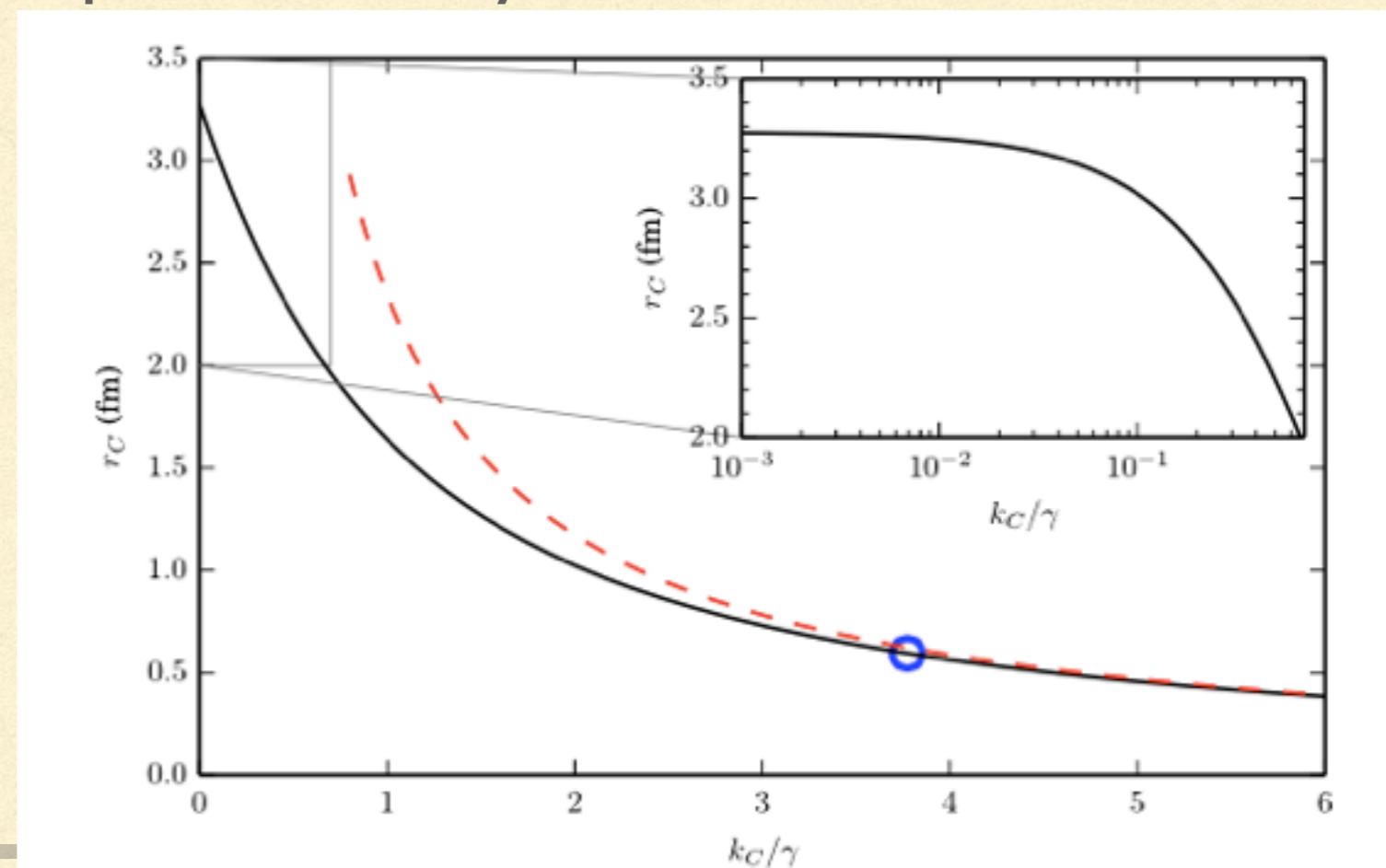
Koenig, Griesshammer, Hammer, van Kolck, JPG, 2016
 - Possible because Coulomb does not require additional renormalization at LO in α_{em} in pionless EFT
 - Coulomb energy of 2p halos? “Three-body Thomas-Ehrman shift” (${}^{16}\text{Ne}/{}^{16}\text{C}$)?
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PROTON HALOS FROM NEUTRON HALOS?

- Can we predict the energy of a proton halo from its isospin mirror?
- First problem: neutron halo already fine tuned, require another fine tuning to also have proton halo bound or nearly bound
- Note that if neutron is bound by little enough to be in a halo then Coulomb must be treated non-perturbatively

Theory seems to work for $^{17}\text{F}^*$, but this state is in the deep Coulomb regime, with $k_c r \sim 1$

Ryberg et al., Ann. Phys., 2016



ISOSPIN SYMMETRY IN HALO EFT

- Second problem: Halo EFT is typically asserted to have isospin symmetry, but does it?
- Simplest case: pp vs. nn systems
- $C_{0,nn}$ and $C_{0,pp}$ differ in their μ -dependence, due to $C_{0,pp}$ having to account for Coulomb interactions
- So at what scale does isospin apply? $\mu=m_\pi$? Predicts:

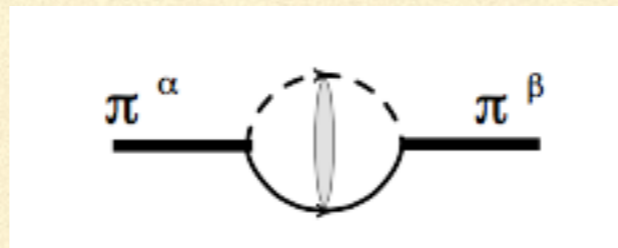
$$\frac{1}{a_{C,pp}} = \frac{1}{a_{nn}} - 2k_C \left[\ln \left(\frac{\sqrt{\pi} m_\pi}{2k_C} \right) + 1 - \frac{3C_E}{2} \right]$$

For $a_{nn} = -18.6$ fm, predicts $a_{C,pp} = -6.45$ fm cf. $a_{C,pp} = -7.8063(26)$ fm

AND IT GETS WORSE...

	${}^7\text{Li-n}$	${}^7\text{Be-p}$
γ_1 (MeV)	57.8	15
r_1 (fm $^{-1}$)	-1.43	-0.34

- Consider a p-wave proton-core system, need to calculate:



- Results for effective-range parameters in terms of Lagrangian

$$\frac{(-)}{a_1} \equiv (-)6\pi \frac{M_R \Delta}{h_{Pt}^2} + 2k_C^3 \left(\frac{1}{3-d} + \ln \frac{\mu\sqrt{\pi}}{k_C} - \frac{3}{2}C_E + \frac{4}{3} \right) - 3\mu k_C^2 \left(1 + \frac{\pi^2}{3} \right),$$
$$\frac{r_1}{2} \equiv (-)\frac{3\pi}{h_{Pt}^2} + 2k_C \left(\frac{1}{3-d} + \ln \frac{\mu\sqrt{\pi}}{k_C} - \frac{3}{2}C_E + \frac{4}{3} \right) - 3\mu,$$

CONCLUSION

- Subtractive renormalization illuminates aspects of the three-body problem in pionless EFT/halo EFT, but is not a silver bullet
 - Higher-order corrections should **not** be iterated at arbitrarily large cutoffs: they change the asymptotic behavior of the amplitude. This tends to produce erroneous conclusions about the order at which counterterms are needed.
 - Perturbative analysis should permit extraction of order at which a particular effect becomes sensitive to short-distance pieces of the 3B wave function; unitary limit wf in hyperspherical co-ordinates can be useful for this
 - UV piece of Coulomb-nuclear interference is associated with non-observability of Coulomb energies
 - Complicates implementation of isospin symmetry in Halo EFT: how does isospin relate, e.g. ${}^3\text{He}({}^4\text{He}, \gamma){}^7\text{Be}$ and ${}^3\text{H}({}^4\text{He}, \gamma){}^7\text{Li}$
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