EMMI Rapid Reaction Task Force The systematic treatment of the Coulomb interaction in few-body systems, Second meeting. May 31-June 3, 2016

Updates on solar proton-proton fusion from  $\pi$ EFT



האוניברסיטה העברית בירושלים THE HEBREW UNIVERSITY OF JERUSALEM

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Widely believed:



#### $\checkmark$ Weak proton-proton fusion in the Sun – theory standards

SFII – Adelberger et al., Rev. Mod. Phys. 83, 195 (2011)

 $3.99(1 \pm 0.030) \times 10^{-25}$  MeV b pionless EFT.

SFII recommended value (2011):  $S_{11}(0) = 4.01(1 \pm 0.009) \times 10^{-25}$  MeV b.

<u>" $\chi$ EFT" calculation by Marcucci et al., Phys. Rev. Lett. (2013)</u>: Use consistent <sup>3</sup>H decay-rate to constrain consistently axial MEC (DG, Quaglioni, Navratil, PRL 2009), and predict pp-fusion rate.

$$S(0) = (4.030 \pm 0.006) \times 10^{-23} \text{ MeV fm}^2$$

Including: p-wave contribution (+0.005%), full EM (-0.0025-(-0.0075)%), difference between 500 and 600 MeV cutoff and potential models.

Recently Archaya et al (1603.01593)  $\chi$ EFT:  $S(0) = (4.081^{+0.024}_{-0.032}) \times 10^{-23} \text{ MeV fm}^2$ 



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Can we reach precision physics with  $\pi EFT$ ?





Role of  $\pi$ EFT: Coherent and sysytematic (theoretical) uncertainty quantification. Big question: is precision physics a possible frontier of  $\pi$ EFT?

We revisit the pp-fusion problem within pionless EFT, fixing the unknown LEC using triton decay. May 31, 2016



Advantages of  $\pi$ EFT for proton-proton fusion:

Small number of parameters.
 Two NLO π EFT arrangements.
 A "cheat-sheet" in the electromagnetic sector.
 Cutoff independence up to infinity.

#### A fully perturbative pionless EFT A=2, 3 calculation @NLO

- LO Parameters:
  - nn and 2-np Scattering lengths: <sup>3</sup>S<sub>1</sub>, <sup>1</sup>S<sub>0</sub>.
  - pp scattering length.
  - Fine structure constant.
  - Three body force strength to prevent Thomas collapse.
- NLO parameters:
  - 2 effective ranges.
  - Renormalizations of pp and 3NF.
  - (isospin dependent 3NF to prevent logarithmic divergence in the binding energy of <sup>3</sup>He).
- Weak Interaction: LO (g<sub>A</sub> 1 body), NLO (L<sub>1A</sub> 2 body)
- EM Interaction: LO  $(\kappa_s, \kappa_v) 1$  body), NLO  $(L_1, L_2 2$  body)



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# The role of the deuteron tail

Many low energy reactions depend on deuteron normalization.

$$Z_d^{-1} = i \frac{\partial}{\partial_{p_0}} \frac{1}{i \mathcal{D}_t(p_0, p)} \Big|_{p_0 = \frac{\gamma_t^2}{M_N}, p = 0}$$

• One has a choice of rearranging the expansion:



Z-parameterization has quicker convergence, especially for observables sensitive to the deuteron tail.

> Phillips, Rupak, Savage, Phys. Lett. **B473**, 209 (2000) Grießhammer, Nucl. Phys. A744, 192 (2004)

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#### ▲ A=3 magnetic moments calculations:



- All NLO contributions of the same order of magnitude 5-10% –
  Natural NLO contributions useful for theoretical error estimates!
- No effect due to Zed-Rho parameterizations.
- Cutoff independence.
- When L<sub>1</sub> and L<sub>2</sub> are fixed **from A=2 observables**:

LO:	$\mu_{^{3}\rm{H}}^{^{LO}}=3.09\pm_{Z_d}0.01$	$\mu_{^{3}\text{He}}^{LO} = -2.455 \pm_{Z_{d}} 0.005$
NLO:	$\mu_{{}^{3}\mathrm{H}}^{NLO} = 3.005 \pm_{Z_d} 0.01$	$\mu_{{}^{3}\text{He}}^{NLO} = -2.13 \pm_{Z_d} 0.01$
exp:	$\mu_{{}^{3}{}_{H}}^{exp} = 2.9789$	$\mu_{{}^{3}\mathrm{He}}^{\mathrm{exp}} = -2.1276$

#### ▲ A=3 magnetic moments calculations:



- All NLO contributions of the same order of magnitude 5-10% –
  Natural NLO contributions useful for theoretical error estimates!
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- Cutoff independence.
- When L<sub>1</sub> and L<sub>2</sub> are fixed **from A=3 magnetic moments**:

LO:
$$\mu_d^{LO} = 0.8798$$
 $\sigma_{np}^{LO} = 298.2 \,\mathrm{mb}$ NLO: $\mu_d^{NLO} = 0.8617 \pm_{Z_d} 0.0002$  $\sigma_{np}^{NLO} = 335(Z_d) - 320(\rho)$ exp: $\mu_d^{exp} = 0.8574...$  $\sigma_{np}^{exp} = 334.2 \pm 0.5 \,\mathrm{mb...}$ 





## Lattice QCD calculation of $l_1$

PRL 115, 132001 (2015)

PHYSICAL REVIEW LETTERS

week ending 25 SEPTEMBER 2015

Ab initio Calculation of the  $np \rightarrow d\gamma$  Radiative Capture Process

Silas R. Beane,<sup>1</sup> Emmanuel Chang,<sup>2</sup> William Detmold,<sup>3</sup> Kostas Orginos,<sup>4,5</sup> Assumpta Parreño,<sup>6</sup> Martin J. Savage,<sup>2</sup> and Brian C. Tiburzi<sup>7,8,9</sup>

(NPLQCD Collaboration)





$$\tilde{X}_{M1} = \frac{Z_d}{-\frac{1}{a_1} + \frac{1}{2}r_1|\mathbf{p}|^2 - i|\mathbf{p}|} \times \left[\frac{\kappa_1\gamma_0^2}{\gamma_0^2 + |\mathbf{p}|^2} \left(\gamma_0 - \frac{1}{a_1} + \frac{1}{2}r_1|\mathbf{p}|^2\right) + \frac{\gamma_0^2}{2}l_1\right]$$





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**Rho-parameterization** 





K

"Empirical" extraction of GT (using calculated F strength)





Adding the NLO 1-body contributions



All NLO contributions are of the same order (1-2%), one can estimate higher order effects as the NLO contribution.





2<sup>nd</sup> estimate of theoretical uncertainty: difference between Zed and Rho Paramerizations.



#### Translates to $\pm 2\%$ difference in pp fusion

2<sup>nd</sup> estimate of theoretical uncertainty: difference between Zed and Rho Paramerizations.

#### So... is 3% too big to be called precision physics?

![](_page_33_Figure_1.jpeg)

i.e., theoretical uncertainty of the same order of systematic experimental error encapsulated in  $g_A$  and <sup>3</sup>H half life (2% total).

![](_page_33_Figure_3.jpeg)

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# Summary

- Pionless EFT reproduces low-energy electroweak observables to a very good precision (~1%), even at NLO, and allows reliable uncertainty estimates.
- Theoretical uncertainty estimated from:
  - (Natural) Size of NLO contribution (all NLO contributions are of the same order of magnitude).
  - Difference between Zed and Rho parameterizations.
  - Both error estimates lead to about 2% uncertainty.
- EM sector confirms calculation procedure.
- Lattice QCD for nuclei is a new front for  $\pi$  EFT
- Based on the EM sector, a theoretical prediction for pp fusion:

$$S_{pp}(g_A = 1.2701) = 4.01 \pm_{theory} 0.08 \pm_{g_A(1\sigma)} 0.07 \pm_{^{3}\text{H half life}} 0.04$$
$$S_{pp}(g_A = 1.275) = 4.12 \pm_{theory} 0.08 \pm_{g_A(1\sigma)} 0.07 \pm_{^{3}\text{H half life}} 0.04$$

- Better determination of g<sub>A</sub> is necessary!
- (<sup>3</sup>H half life is also an open exp. issue).