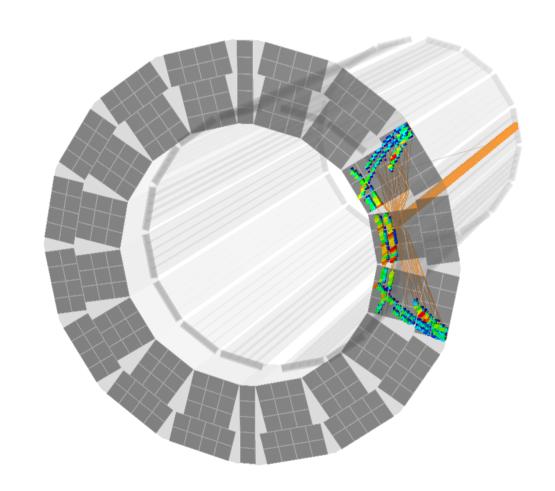
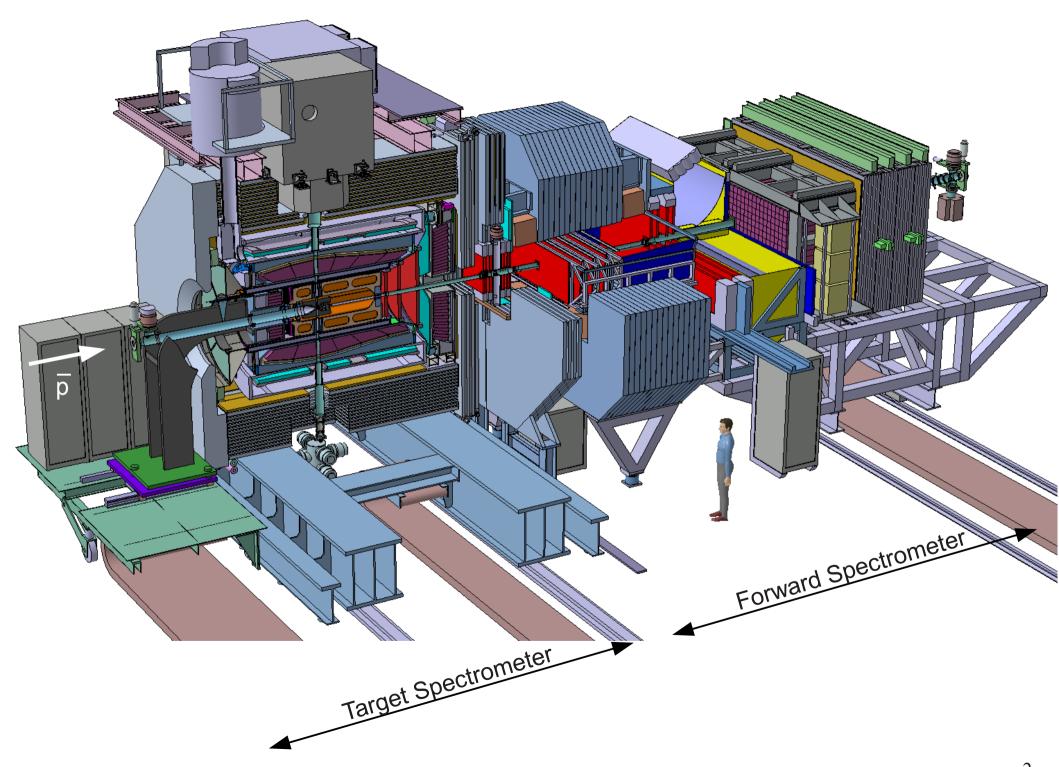
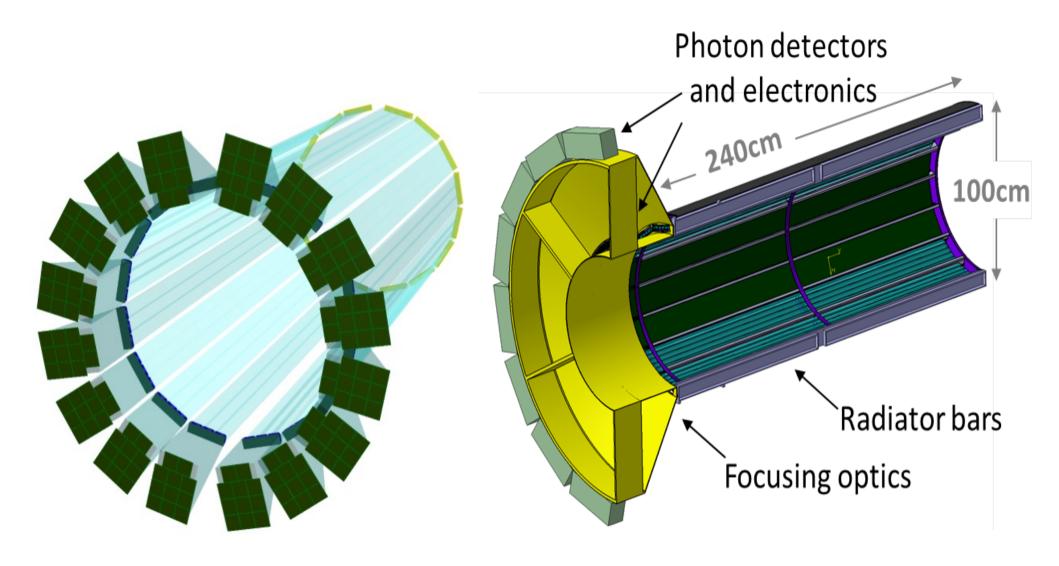
# Readout of the Barrel DIRC system

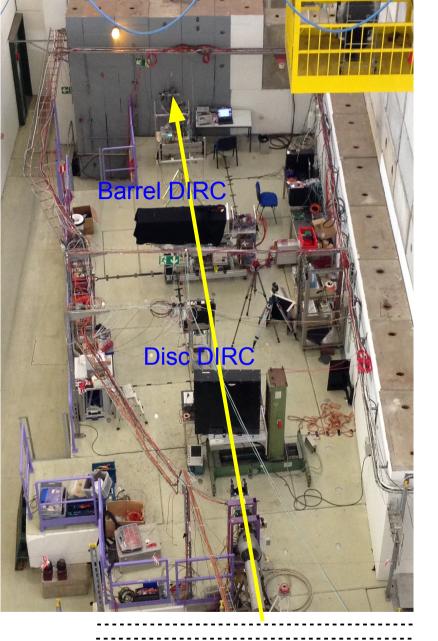
## Carsten Schwarz, G 55 it

- Introduction
- Beam experiments
- ●PADIWA3 + TRB3
- Outlook DIRICH

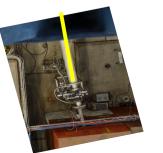








**29 m TOF** 



Experiments,

2008 - GSI

2009 - GSI

2011 - GSI, CERN

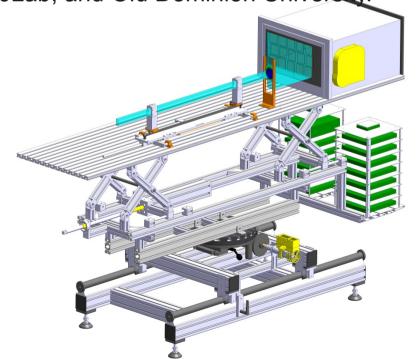
2012 - CERN

2013 - Mainz

2014 - GSI

2015 - CERN

2015: Joint effort of groups from GSI, Uni Mainz, Uni Giessen, Uni Erlangen, JLab, and Old Dominion University.

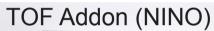






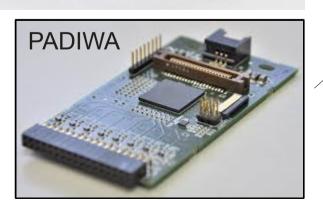
TRB 2

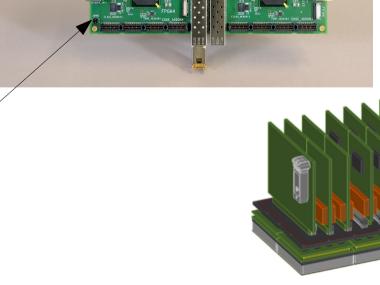




ades Discriminator

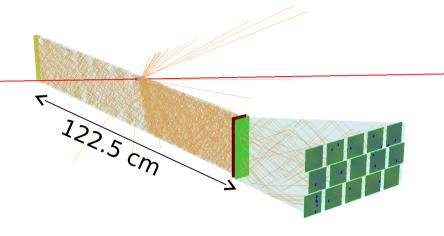








## Photon detector





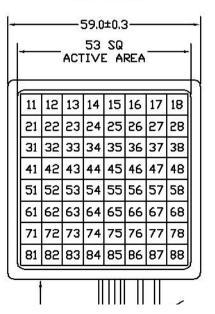
5 x 3 Planacon MCP-PMT (XP85012/A1-Q, Photonis) 960 pixels (in total >1200 readout channels)

with pixel size 6.5 x 6.5 mm<sup>2</sup>

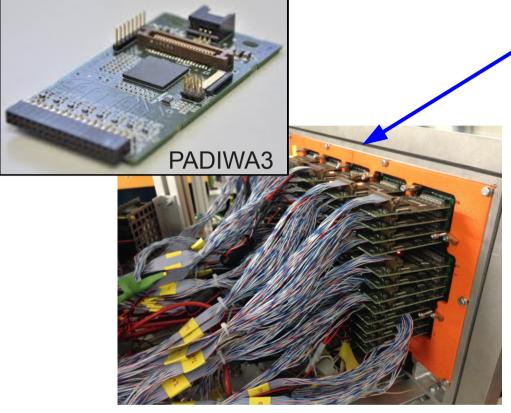
Work in 1T magnetic field

Survive **10 years** of PANDA (ageing)

#### TOP VIEW

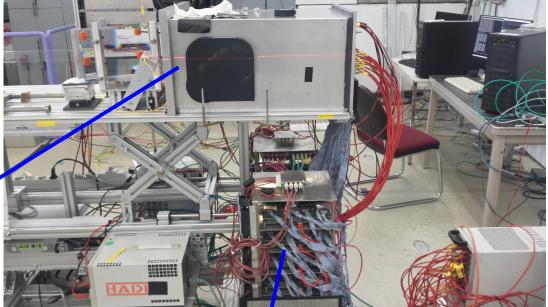


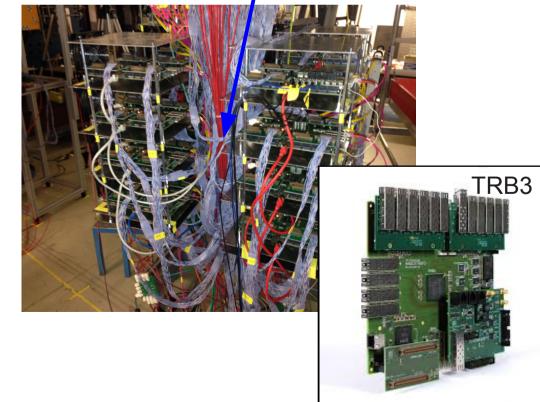
## Readout chain



PADIWA3 discriminator Keep It Small & Simple = KISS Amplifier + LVDS discriminator

TRB3 TDC board Leading edge → timing (~10ps) Trailing edge → TOT → walk correction

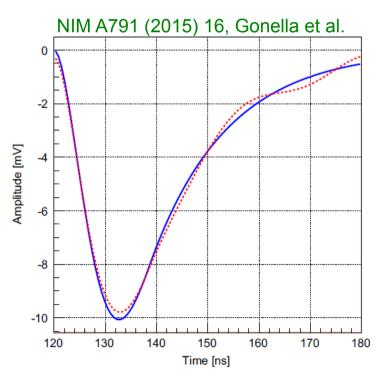




# Walk correction by TOT

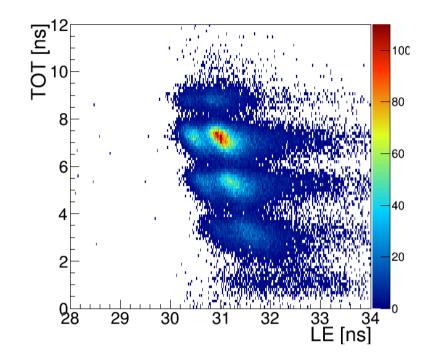
Time spectra (Leading Edge) show modulation which we ignore for **walk correction** 

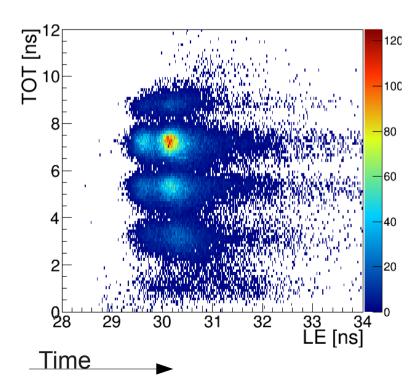
Understood: TOT + small high frequency noise



Amplitude

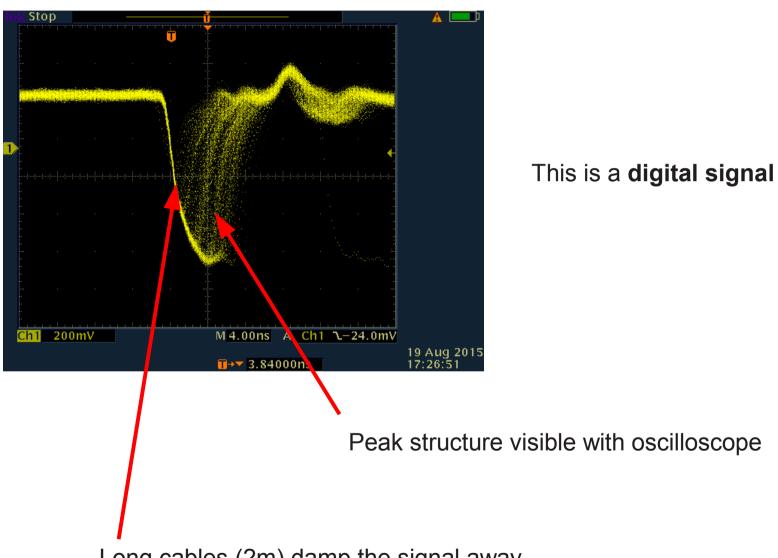
Fig. 5. Simulated shape of the output signal of the system lead-glass block – PMT without (solid curve) and with addition of  $300\,\mu V$  noise at  $40\,MHz$  frequency (dashed curve).





## Peak structure of TOT

#### LVDS output (PADIWA)



Long cables (2m) damp the signal away. Larger effective thresholds? (No, we see all photons.) Better skip cables...

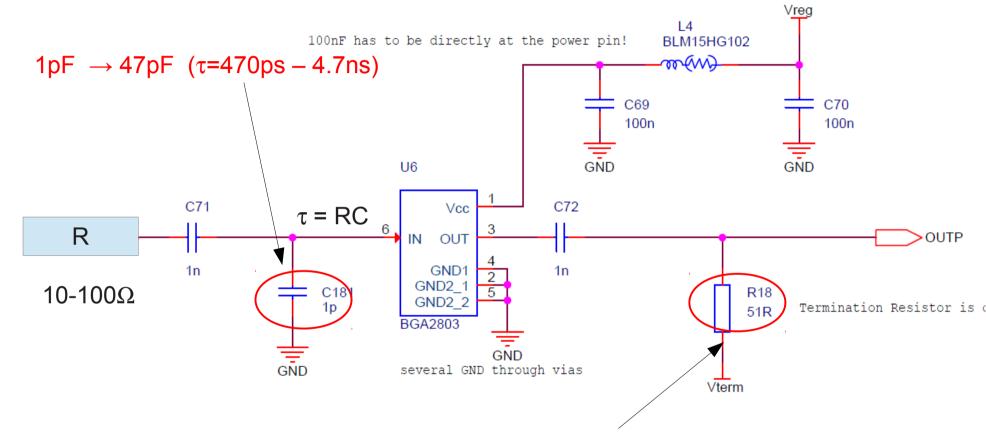
- The FEE is stable in the lab in small setups
- In a larger setup (GSI 2014 beam) a high frequency feedback from the FEE to the MCP-PMTs and back to other channels of the FEE has been observed
- Result: high frequency oscillation forced the use of unreasonable high thresholds
- Only solution available:

attenuate high frequency noise at the input of the 3GHz amplifiers

disadvantage: slower rise time of the signal + smaller amplitude needs higher amplification and is more affected to lower frequency noise

## Padiwa modification

Bandwidth reduced PADIWA → prevent oscillations observed @ GSI 2014 beam time



Regain amplification by 51R → 200R

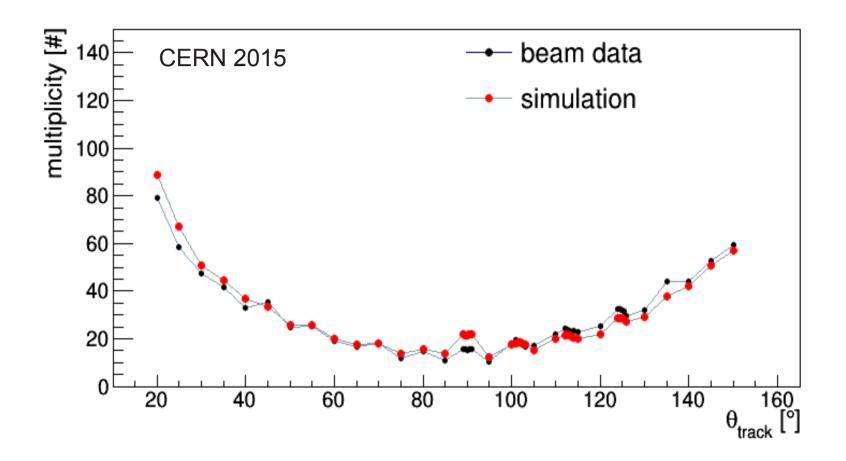
After this modification

The system of 60 PADIWA did not oscillate and low thresholds <1mV were used.

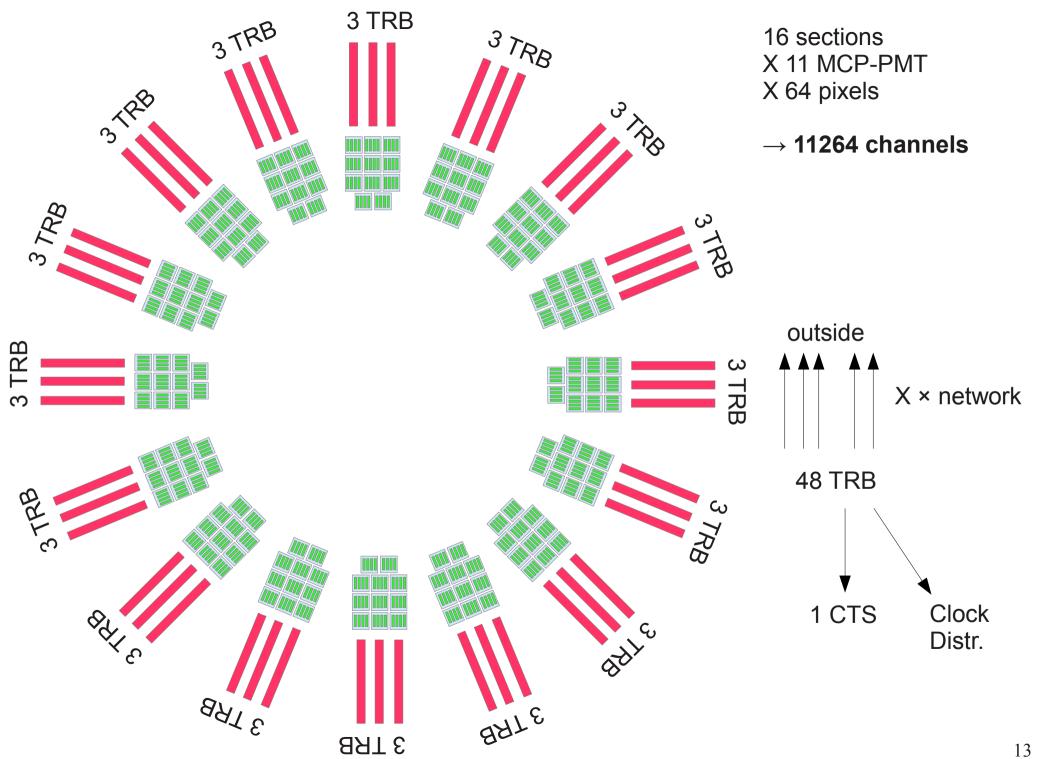
The timing resolution became worse > 200ps (CERN2015 beam time)

→ investigated in electronic lab. (Mainz, GSI)

Beside of electronic issues, measurements have been successful.



Number of measured photons as function of polar angle of charged particle. Good agreement between data and simulation.



### 20 Mhz interaction rate

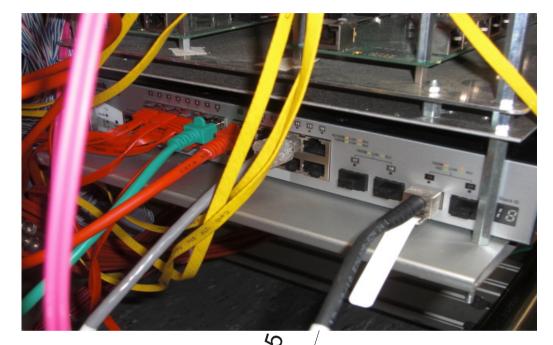
- x 2 charged particles (barrel)
- x 50 detected Cherenkov photons
- $\rightarrow$  2 10<sup>9</sup> hits/s

Leading edge (32 bit)

- + trailing edge (32 bit)
- + overhead (32 bit)

for  $TOT \rightarrow 96$  bits (12 bytes)

- $\rightarrow$  200 Gbit/s
  - 1,2,4,... copper Ethernet links (routing geometry to outside)



#### **PADIWA**

80 mW / channel

11264 \* 80 mW

= 900W

 $= 4.4 \lor x 200 A$ 

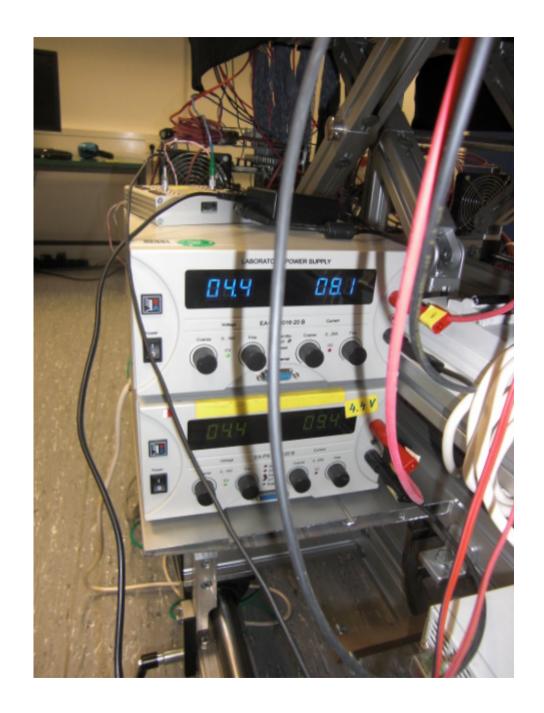
Free copper cable 2 x 10mm diameter Ampacity @ 75 degree → 200 A

https://en.wikipedia.org/wiki/American\_wire\_gauge

#### **TRB**

48 x 20 W = 960 W

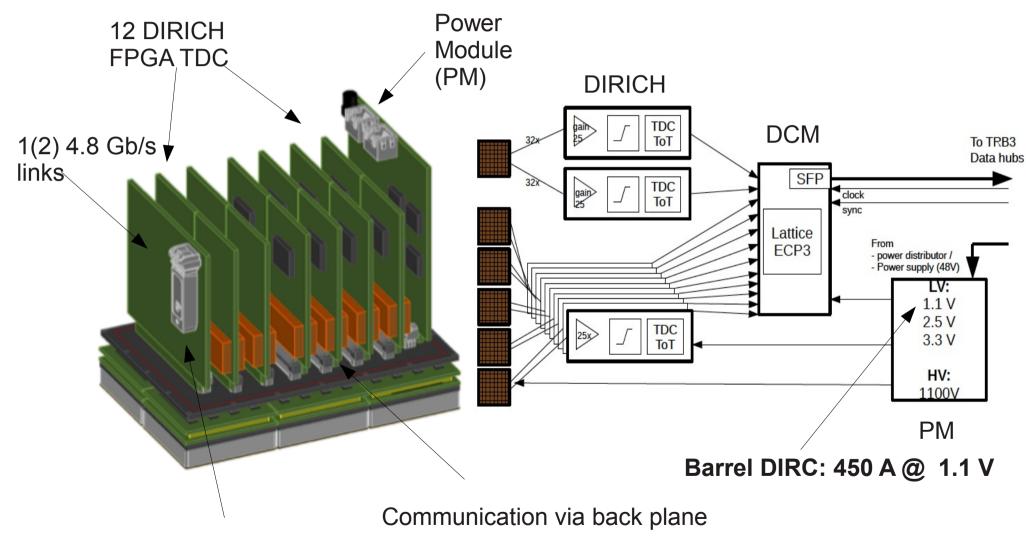
About the same as above...



## DIRICH

#### PANDA-DIRC/CBM-RICH and HADES-RICH

In design phase: DIRICH, Successor of PADIWA3 + TRB3



Data Combiner Module (DCM)

## Summary

## Barrel DIRC FEE

- based on PADIWA3 + TRB3
- CERN 2015 readout of 960 pixels 1500 channels total
- Results agree with simulation
- Option DIRICH