Evgeny Epelbaum, Ruhr-University Bochum

KHuK Jahrestagung, Bad Honnef, 5. Dezember 2015

Aktuelle Entwicklungen in der Theorie: Ab initio Beschreibung der Kerne

Low-energy pion-nucleon dynamics Chiral nuclear forces in the precision era Nuclear lattice simulations: recent highlights Medium-mass nuclei and nuclear matter Nuclear systems from lattice QCD Summary

Disclaimer: apologies for not being able to show other results...

Low-energy pion-nucleon dynamics

Long-range nuclear forces are completely determined by the χ -symmetry of QCD + experimental information on π N scattering



Pion-nucleon Roy-Steiner equations

M. Hoferichter, J. Ruiz de Elvira, B. Kubis, U.-G. Meißner, PRL 115 (2015) 092301; arXiv:1510.06039 [hep-ph]

Integral equations in the form of dispersion relations which incorporate constraints from analyticity, unitarity & crossing symmetry

Input: S-,P-waves at high energy, inelasticities, ⁵/₅ D- and higher waves + scatt. lengths from hadronic atoms

Output: reliable results for S-,P-waves with systematic uncertainties; subthreshold coefficients, determination of the σ -term:

 $\sigma_{\pi N} = 59.1 \pm 3.5 \,\mathrm{MeV}$

ChPT for πN , $\pi N \rightarrow \pi \pi N$ with/without Δ (1232) Siemens et al., PRC 89 (2014) 065211; to appear

Baryon ChPT beyond the low-energy region EE, J. Gegelia, U.-G. Meißner, D.-L. Yao, EPJ C75 (2015) 499



Chiral nuclear forces



Why is it necessary/interesting to extend the χ -expansion of the NN potential to Q⁵?

- no additional parameters (except for 1 IB term) → testing the theory (long-range physics)
- there is evidence that χ -expansion for the 3NF is not yet converged at Q⁴
- understanding fine details of the 3NF requires accurate and precise NN forces

Chiral nuclear forces

New state-of-the-art chiral NN potentials up to N⁴LO



EE, Krebs, Meißner, EPJA 51 (2015) 53; PRL 115 (2015) 122301

- improved UV regulator maintains the analytic structure of the amplitude
- all LECs in the long-range part are taken from πN scattering, no fine tuning!
- coupled with the novel approach for uncertainty quantification, provides the tool for next-generation precision ab initio studies

+ 1 IB LEC

Quality of the reproduction of the Nijmegen PWA (,, χ^{2}_{datum} ")

LO $[Q^0]$	NLO [Q ²]	$N^{2}LO [Q^{3}]$	$N^{3}LO[Q^{4}]$	N ⁴ LO [Q
bhase shifts				
360	$_{31}$ no new LECs	4.5	0.7 1 LEC (¹ S ₀)	0.3
480	63	21	0.7	0.3
ase shifts				
5750	102	15	0.8 no new LECs	0.3
9150	560	130	0.7	0.6
	LO [Q ⁰] hase shifts 360 480 ase shifts 5750 9150	$\begin{array}{c c} LO \ [Q^0] & NLO \ [Q^2] \\ \hline hase shifts \\ 360 \\ 480 \\ \hline case shifts \\ 5750 \\ 9150 \\ \hline case 560 \\ \hline case shifts \\ 560 \\ \hline case shifts \\ 5750 \\ \hline case shifts \\ \hline case shifts \\ case shift$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	LO $[Q^0]$ NLO $[Q^2]$ N ² LO $[Q^3]$ N ³ LO $[Q^4]$ hase shifts 360 31 no new LECs 4.5 0.7 1 LEC (1S_0) 480 63 21 0.7 1 LEC (1S_0) 1 ase shifts 5750 102 15 0.8 no new LECs 0.7 9150 560 130 0.7 102 15 0.8 102 15

+ 7 LECs + 2 IB LECs

2 LECs

+ 15 LECs

Uncertainty quantification

A simple algorithm for estimating uncertainty from the truncation of the chiral expansion: EE, Krebs, Meißner, EPJA 51 (2015) 53

For any observable: $X^{(i)}(p) = X^{(0)} + \Delta X^{(2)} + \dots + \Delta X^{(i)} + \dots + \Delta X^{(i)} + \dots + \Delta X^{(i)}$ $\sim Q^2 X^{(0)} + \dots + \Delta X^{(i)} + \dots +$

Use the explicitly calculated $\Delta X^{(i)}$ to estimate the uncertainty $\delta X^{(i)}$ at order Qⁱ:

 $\delta X^{(0)} = Q^2 |X^{(0)}|,$

 $\delta X^{(i)} = \max_{2 \leq j \leq i} \left(Q^{i+1} | X^{(0)} |, \, Q^{i+1-j} | \Delta X^{(j)} |
ight)$

subject to the additional constraint

 $\delta X^{(i)} \, \geq \, \max_{j,k} ig(|X^{(j\geq i)} - X^{(k\geq i)}| ig).$

- no reliance on the cutoff variation (not reliable)
- easily applicable to any observable (scattering, bound states, 3N, ...)
- error bars found to be consistent with 68% degree-of-belief intervals
 Furnstahl et al., PRC 92 (2015) 024005

proton-neutron scattering observables at Elab=143 MeV



3N force studies

With these tools, we are well equipped to tackle the 3N force problem.

Is there any clear evidence for missing 3N forces effects? Yes! Binder et al., arXiv: 1505.07218 [nucl-th]



total cross section in nD scattering calculated without 3NF

3NF contributions matches well the estimated size of N²LO terms

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DARMSTAD

→ consistent with the chiral power counting!

Discrepancies between theory

of quantified uncertainties

and data well outside the range

→ unambiguous evidence

The magnitude of the required

for missing 3NF effects!





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LENPIC: Low Energy Nuclear Physics International Collaboration

IÜLICH

Kyutech

3N force studies



from Norbert Kaiser, TUM Baryonic forces from SU(3) chiral effective field theory

J. Haidenbauer (FZJ), N. Kaiser (TUM), U.-G. Meißner (Bonn), S. Petschauer (TUM), W. Weise (ECT*)

 \blacktriangleright extend successful SU(2) $\chi {\rm EFT}$ approach to nuclear forces to three flavors

▶ hyperon-nucleon interaction from SU(3) χ EFT at NLO (one- and two-meson exchange + contact terms)



good description of available
 YN scattering data
 [Nucl.Phys. A915 (2013) 24-58]

 G-matrix calculation of hyperon potentials in isospin-(a)symmetric nuclear matter

- weak Λ-nuclear spin-orbit
- repulsive Σ-nuclear potential [arXiv:1507.08808]





- construction of leading three-baryon forces (e.g. ΛNN) [arXiv:1511.02095]
- constants estimated via decuplet saturation (e.g. Σ^* for ΛNN)

Nuclear systems from lattice χEFT

Nuclear lattice simulations:

A new ab initio approach to nuclei and nuclear reactions D. Lee, EE, H. Krebs, T. Lähde, T. Luu, U.-G. Meißner, G. Rupak, ...

Some recent highlights:

Ab initio calculation of the Hoyle state EE, H. Krebs, D. Lee, U.-G. Meißner, PRL 106 (11) 192501; EE, H. Krebs, T.A.Lähde, D. Lee, U.-G. Meißner, PRL 109 (12) 252501

Viability of Carbon-based life as a function of light quark masses EE, H. Krebs, T. A. Lähde, D. Lee, U.-G. Meißner, PRL 110 (13) 112502; EPJA 49 (13) 82

Ab initio calculation of the spectrum and structure of ¹⁶O EE, H. Krebs, T. A. Lähde, D. Lee, U.-G. Meißner, G. Rupak, PRL 112 (14) 102501

Lattice EFT for medium-mass nuclei ("triangulation" method for Euclidean-time extrapol.)

T. A. Lähde, EE, H. Krebs, D. Lee, U.-G. Meißner, G. Rupak, PLB 732 (14) 110

 $E_{8Be, old} = -55(2) \text{ MeV}$ $E_{8Be, new} = -56.4(2) \text{ MeV}$

Symmetry-sign extrapol.

T.A. Lähde, T. Luu, D. Lee, U.-G. Meißner, EE, H. Krebs, G. Rupak, EPJ A51 (15) 92





Nuclear systems from lattice χEFT

Lab

nature

Ab initio alpha-alpha scattering

 $Serdar Elhatisari^1, Dean Lee^2, Gautam Rupak^3, Evgeny Epelbaum^4, Hermann Krebs^4, Timo A. Lähde^5, Thomas Luu^{1.5} \& Ulf-G. Meißner^{1.5,6}$

Nature 528, 111–114 (03 December 2015) | doi:10.1038/nature16067 Received 12 June 2015 | Accepted 30 September 2015 | Published online 02 December 2015

First ab initio calculation of alpha-alpha scattering!

Used lattice EFT to extract the effective Hamiltonian for two interacting α-clusters (adiabatic projection method [A. Rokash et al., PRC 92 (15) 054612])



Phase shifts obtained $[m_{T}] = [N_{\tau}^{-1/2}H_{\tau}N_{\tau}^{-1/2}]_{R,R}^{\ell,\ell_{z}}$ loying a hard spherical wall boundary at asymptotically large distances

Promising scaling with respect to the number of particles as $\sim (A_1 + A_2)^2$



6

E_{Lab} (MeV)

8

10

12

0

2

Frontier of ab initio calculations from Achim Schwenk, TU Darmstadt

 First NN+3N prediction of the neutron skin, weak form factor, dipole polarizability of ⁴⁸Ca





Neutron and weak-charge distributions of the ⁴⁸Ca nucleus

G. Hagen^{1,2,*}, A. Ekström^{1,2}, C. Forssén^{1,2,3}, G. R. Jansen^{1,2}, W. Nazarewicz^{1,4,5}, T. Papenbrock^{1,2}, K. A. Wendt^{1,2}, S. Bacca^{6,7}, N. Barnea⁸, B. Carlsson³, C. Drischler³¹⁰, K. Hebeler^{9,10}, M. Hjorth-Jensen^{4,11}, M. Miorelli^{6,12}, G. Orlandini^{13,14}, A. Schwenk^{9,10} and J. Simonis^{9,10}



Neutron skin smaller than previously thought! Hagen et al., Nature Phys.

In-Medium Similarity Renormalization Group

First nonperturbative derivation of shell-model interactions from NN+3N interactions Bogner et al., PRL 113, 142501 (2014)

First ab initio description of deformed nuclei Stroberg et al., 1511.02802

• Quantum Monte Carlo with local chiral 3N Lynn, Tews et al., 1507.05561, 1509.0347



Neutron-Star Matter



no "exotic" matter needed (or excluded)

from Norbert Kaiser, TUM T. Hell, N. Kaiser, B. Röttgers, S. Schulteß, W. Weise PRC 90 (2014) 045801

- Very precise measurement of a neutron P.B. Demorest et al., star (PSR J1614-2230) Nature, 467, 1081 (2010) $M=(1.97\pm0.04)M_{\odot}$
- Sets new constraints on the equation of state (EoS) of nuclear matter
- Inclusion of neutron star constraints plus Chiral Effective Field Theory at lower density





from Hans-Werner Hammer, TU Darmstadt

- Delicate interplay between strong and Coulomb interaction
 - \implies two fine tunings required to obtain shallow halo states
 - \implies proton halos are rarer in nature than neutron halos
- Range corrections in proton halos

Ryberg, Forssen, Hammer, Platter, arXiv:1507.08675

- Explore universal correlations between observables
- S-factor for ${}^{7}\text{Be}(p,\gamma){}^{8}\text{B}$ and charge radius of ${}^{8}\text{B}$ are correlated



Ryberg, Forssen, Hammer, Platter, Eur. Phys. J. A 50 (2014) 170

Nuclear physics from lattice QCD

- Lattice QCD results for light nuclei start to emerge (at high M_{π})
- Controversial results: more binding [NPLQCD, Yamazaki et al.] versus no binding at all [HAL QCD]
 - → consistency checks are needed!
 - π -less EFT & extrapolations in the # of nucl. [Barnea et al.'15]
 - Low-energy theorems: long-range interactions imply correlations between coefficients in the effective range expansion
 - \bullet model-independent; no reliance on the χ -expansion
 - good predictive power in ³S₁ for physical M_π
 - can be extended to heavier M_π
 [Baru et al., PRC99 (15) 014001]

	Nijmegen PWA	5.42	1.75	0.04	0.67	-4.0
adad ta	NLO LET	input	input	0.06	0.70	-4.0
ysical W_{π}	LOLET	input	1.60	-0.05	0.82	-5.0
veical M —		$oldsymbol{a}~[{ m fm}]$	$m{r}~[{ m fm}]$	$v_2 \; [{ m fm^3}]$	$v_3 \; [{ m fm}^5]$	$v_4 \; [{ m fm}^7]$







Summary

New generation of accurate and precise chiral nuclear forces

- + Reliable approach to uncertainty quantifications
- + Exciting progress in ab initio methods (Nuclear lattice simulations, Coupled cluster, In-Medium SRG, Green Function Monte Carlo, ...)
- + Growing computational resources

Low energy nuclear theory is entering precision era: Reliable ab initio few- and many-body calculations based on chiral EFT with quantified theoretical uncertainties.