## Overview of hypernuclei What are exciting now?

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4. Summary

## 1. Introduction

## Why we study hypernuclei?

## **Motivations of hypernuclear physics**

#### **BB** interactions

Unified understanding of BB forces by u,d ->u, d, s particularly short-range forces by quark pictures Test lattice QCD calculations

#### **Impurity effect**

#### in nuclear structure

Changes of size, deformation, clustering, Appearing new symmetry, Properties and behavior of baryons

#### <u>in nuclei</u>

μ<sub>Λ</sub>

 $\mu_{\Lambda}$  in a nucleus, Single particle levels of heavy  $\Lambda$  hypernuclei

Clues to understand hadrons and nuclei from quarks Cold and dense nuclear matter with strangeness



## **Importance of BB interactions**

#### YN, YY interactions in free space

- **YN** scattering
- YY correlation in HIC
- Light Hypernuclei (Few body systems)

#### YN, YY interactions in nuclear matter

= (YNN force /  $\rho$  dependence of YN, YN in neutron-rich matter) Heavy hypernuclei

YN, YY interactions from high density matter in HIC?

 $\Rightarrow \mbox{Necessary to understand} \\ \mbox{the baryonic matter EOS (=neutron star matter)} \\ \mbox{for } \rho \sim 0 - 5 \rho_0 \,, \mbox{ with strangeness} \end{cases}$ 

## Few-body ∧ hypernuclei

#### Exact calculations possible

A=3  ${}^{3}_{\Lambda}H$  B<sub> $\Lambda$ </sub> =130 keV => stringent test for  $\Lambda N$  interaction  $\tau < 200$  ps (by HIC): why so short?? Weak decay branches not measured well  $nn\Lambda$  bound?  $^{3}_{\Lambda}$ H  $^{4}_{\Lambda}$ He <sup>5</sup>He  $pn\Lambda(T=1)$ ? 0.0 --0.13 MeV 1<sup>+</sup> -1.24 MeV A=4  ${}^{4}_{\Lambda}H/{}^{4}_{\Lambda}He$  $0+/1+\Lambda N$  spin-spin,  $\Lambda N-\Sigma N$  interaction Exp. 0<sup>+</sup> -2.39 MeV Charge symmetry breaking? -3.12 MeV A=5  ${}^{5}$  He Overbinding problem =>  $B_{\Lambda}$  (A=3,4,5) explained well with  $\Lambda N-\Sigma N$  interaction

# 2. ΛN interaction from Light Λ hypernuclei

**Charge Symmetry Breaking in A=4** 

## Charge Symmetry Breaking puzzle in A=4



## Pion decay spectroscopy @Mainz



|--|

	,			
	Material	<sup>9</sup> Be		
	Thickness	125 μm		
	Tilt angle	54 deg		

Kaos					
Cent. Mom	+900 MeV/c				
Detector	MWPC, TOF, AC				

Spek-A, C	
Cent. Mom	- 115/ -125 MeV/c
Detector	DC, TOF, GC



#### electron beam



#### Hypernuclear γ-ray data as of 2014







## Hyperball-J in air-conditioned hut

CF4 target

## SKS downstream detectors





2015 06 25

## **Hyperball-J**

L2

C3

97



Ge cooled down to ~70K by pulse-tube refrigerator (c.f. 92K w/LN2)

Fast background suppressor made of PWO

∆E= 3.1(1) keV at 1.33 MeV Eff. = 5.4% @1 MeV with 28 Ge(re=60%)



Up side (Target view)





was also successfully taken.





Existence of CSB confirmed <u>only by γ-ray data</u>

Large spin dependence in CSB found by combining with emulsion data



$^{4}_{\Lambda}\mathrm{He}{}^{-4}_{\Lambda}\mathrm{H}$	$P_{\Sigma}(\%)$	$\Delta T_{YN}$	$\Delta V_C$	$\Delta V_{YN}$	$\Delta B_{\Lambda}^4$	$\Delta B_{\Lambda}^4$
model	$0_{g.s.}^{+}$	$0_{g.s.}^{+}$	$0_{g.s.}^{+}$	$0_{g.s.}^{+}$	$0_{g.s.}^{+}$	$1_{\rm exc}^+$
$\Lambda NNN$ [9]	_	_	-42	91	49	-61
$NSC97_{e}$ [10]	1.6	47	-16	44	75	-10
NSC97 <sub>f</sub> [11]	1.8				100	-10
NLO chiral [12]	2.1	55	-9	_	46	
$(\Lambda \Sigma)_{\rm e}$ [present]	0.72	39	-45	232	226	30
$(\Lambda \Sigma)_{\rm f}$ [present]	0.92	49	-46	263	266	39

"D2" potential

A. Gal, PLB 744 (2015) 352

Theoretical studies will elucidate the origin of CSB and the  $\Lambda$  N - $\Sigma$  N interaction.

central

tensor-dominated

## 3. Toward Strange matter in neutron stars



## **First observation of a** *Ξ***-nuclear bound state**

#### KEK E373 "Kiso event" K. Nakazawa et al. PTEP 2015, 033D02



Fig. 1. A superimposed image from photographs and a schematic drawing of the KISO event.

$$\Xi^- + {}^{14}N \rightarrow {}^{10}_{\Lambda}Be + {}^{5}_{\Lambda}He$$

 $B_{\Xi} = 4.38 \pm 0.25 \text{ MeV} - 1.11 \pm 0.25 \text{ MeV} >> E(3D) = 0.17 \text{ MeV}$   $^{10}_{\Lambda}\text{Be in g.s.} \qquad ^{10}_{\Lambda}\text{Be in highest excited state}$   $(Ehime pot.) \quad U_{\Xi} \sim 20 \text{ MeV} \implies B_{\Xi}(2p \text{ state}) = 1.1 \text{ MeV}$ 

## "The heavy neutron star puzzle"

Hyperons must appear at  $\rho = 2^3 \rho_0$ 

EOS's with hyperons (or kaons) too soft -> can support M < 1.5 M<sub>sun</sub>

PSR J1614-2230 (2010)  $1.97 \pm 0.04 M_{sun}$ PSR J0348-0432 (2013)  $2.01 \pm 0.04 M_{sun}$ 

We do not know BB interaction in high density nuclear matter!



Unknown repulsion at high  $\rho$  exists

- Strong repulsion in three-body force including hyperons are necessary. (NNN, YNN, YYN, YYY)
- Phase transition to quark matter ? (quark star or hybrid star)

But we have no data on BBB force at high  $\rho$  nuclear matter, except for indirect info. in HI collisions.

-> Rijken, Pagliara,..

#### **Previous (** $\pi^+$ ,K<sup>+</sup>) data and $\Lambda N$ interaction



## (e,e'K+) data at JLab

## Lab Slide by Nakamura



Resolution ~0.5 MeV (FWHM) Absolute accurary ~0.1 keV <sup>40</sup>Ca/ <sup>48</sup>Ca target runs conditionally approved

#### <u>Precise Hypernuclear B<sub> $\Lambda$ </sub> data provide</u> <u>information on $\rho$ dependence (ANN force)?</u>



Compilation by Millener

Slope of the calculated curve changes by ~ 1 MeV between w/o and w/ 3B/4B repulsion

Y. Yamamoto et al. Phys.Rev. C88 (2013) 2, 022801 Phys.Rev. C90 (2014) 045805

-> Rijken's talk

( $\pi^+, K^+$ ), ( $K^-, \pi^-$ ) systematic error ~1 MeV (e,e'K<sup>+</sup>): systematic error ~ 0.1 MeV --- B<sub>A</sub> will be measured at JLab J-PARC will also measure B<sub>A</sub> by high resolution ( $\pi^+, K^+$ ) with 0.2 MeV resolution

## **Summary**

- Light and heavier hypernuclear data provide information on YN, YY interactions in free space and in nuclear medium.
- New data on CSB in A=4 hypernuclei:

 $B_{\Lambda}$  of  ${}^4_{\Lambda}H(0^{\scriptscriptstyle +})$  measured via pion decay spectroscopy

---- Consistent with old emulsion data

 ${}^{4}_{\Lambda}$ He(1+->0+)  $\gamma$ -ray measured to be 1.406 MeV

 $\Leftrightarrow$  1.09 MeV for  ${}^{4}_{\Lambda}$ H

-- A large CSB effect in  $\Lambda N$  interaction confirmed.

--CSB has a spin dependence

- E nuclear bound system (Kiso event) was observed for the first time -> E potential is attractive
- To solve the hyperon puzzle,  $\Lambda$ NN 3-body force should be studied via precise B<sub> $\Lambda$ </sub> measurements at JLab (+J-PARC).