

Contents

- 1. Introduction
 Why we study hypernuclei?
- 2. ΛN interaction from light Λ hypernuclei
 A=4 charge symmetry breaking
 A very new result of ⁴_ΛHe γ-ray
- 4. Summary

1. Introduction

Why we study hypernuclei?

Motivations of hypernuclear physics

BB interactions

Unified understanding of BB forces by u,d ->u, d, s particularly short-range forces by quark pictures

Test lattice QCD calculations

Impurity effect in nuclear structure

Changes of size, deformation, clustering, Appearing new symmetry, Properties and behavior of baryons in nuclei • μ_ν

 μ_{Λ} in a nucleus, Single particle levels of heavy Λ hypernuclei

Clues to understand hadrons and nuclei from quarks

Cold and dense nuclear matter with strangeness

Importance of BB interactions

YN, YY interactions in free space

YN scattering

YY correlation in HIC

Light Hypernuclei (Few body systems)

YN, YY interactions in nuclear matter

= (YNN force / ρ dependence of YN, YN in neutron-rich matter)

Heavy hypernuclei

YN, YY interactions from high density matter in HIC?

 \Rightarrow Necessary to understand the baryonic matter EOS (=neutron star matter) for $\rho \sim 0$ -- $5\rho_0$, with strangeness

Few-body A hypernuclei

Exact calculations possible

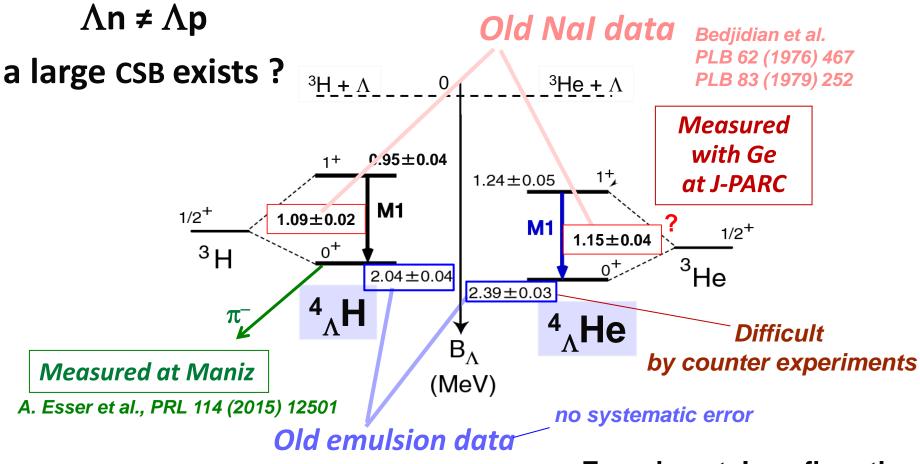
A=3
$$^{3}_{\Lambda}$$
H $^{3}_{\Lambda}$ =130 keV => stringent test for Λ N interaction τ < 200 ps (by HIC): why so short?? Weak decay branches not measured well nn Λ bound? $^{3}_{\Lambda}$ H $^{4}_{\Lambda}$ He $^{5}_{\Lambda}$ He $^{5}_{\Lambda}$ He $^{6}_{\Lambda}$ H

A=5 $^{5}_{\Lambda}$ He
Overbinding problem
=> B_{Λ} (A=3,4,5) explained well with $\Lambda N-\Sigma N$ interaction

2. ΛN interaction from Light Λ hypernuclei

Charge Symmetry Breaking in A=4

Charge Symmetry Breaking puzzle in A=4



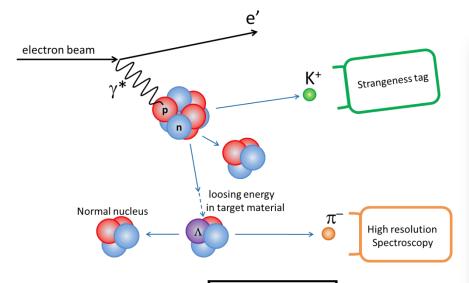
M. Juric et al. NPB 52 (1973) 1

Experimental confirmation of CSB also necessary

Origin is unkown.

 Λ N- Σ N coupling? But 4-body calc's with Λ - Σ mixing using Nijmegen interactions give Δ B < 100keV => Long standing puzzle

Pion decay spectroscopy @Mainz



Primary Beam

Energy 1.5 GeV

Target

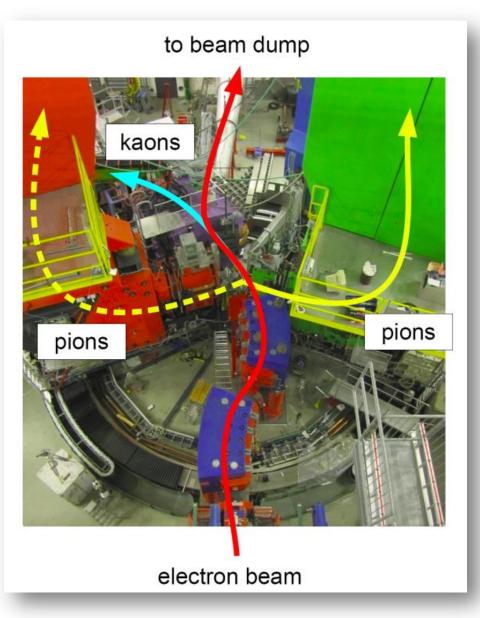
)	
Material	⁹ Be
Thickness	125 μm
Tilt angle	54 deg

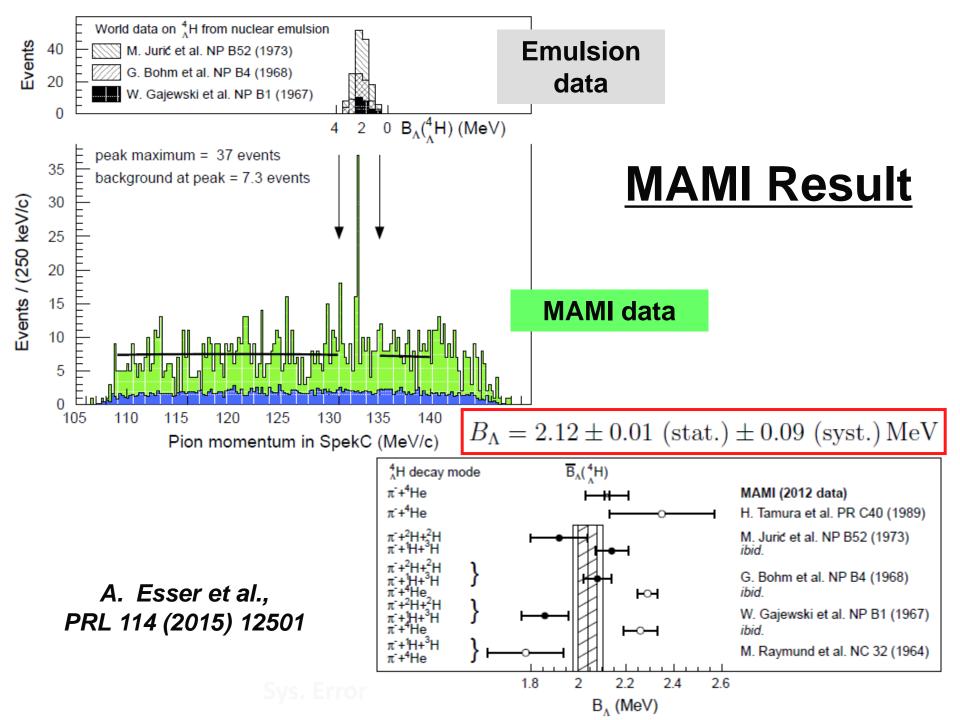
Kaos

Cent. Mom	+900 MeV/c
Detector	MWPC, TOF, AC

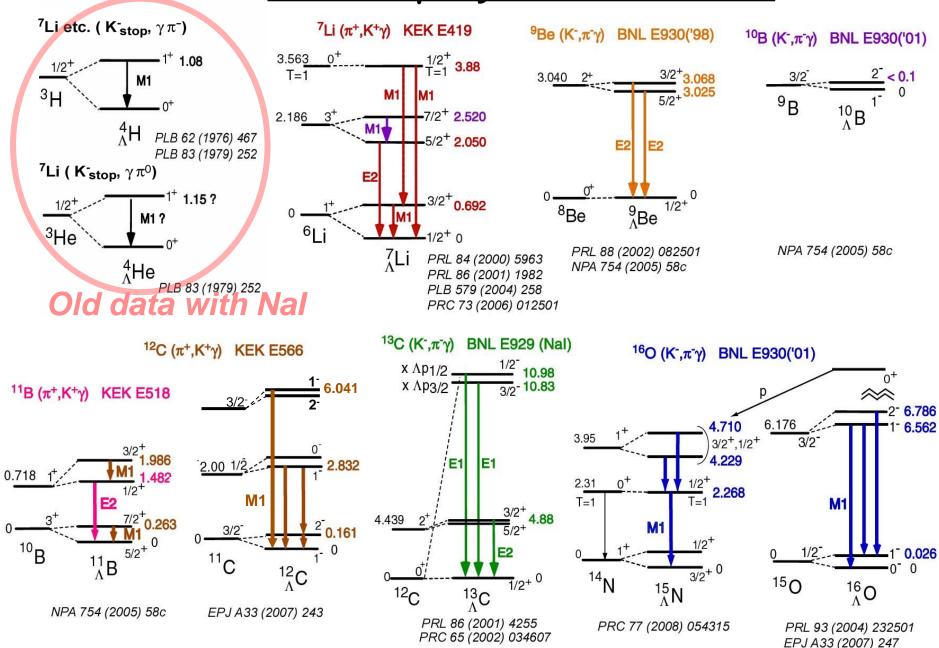
Spek-A, C

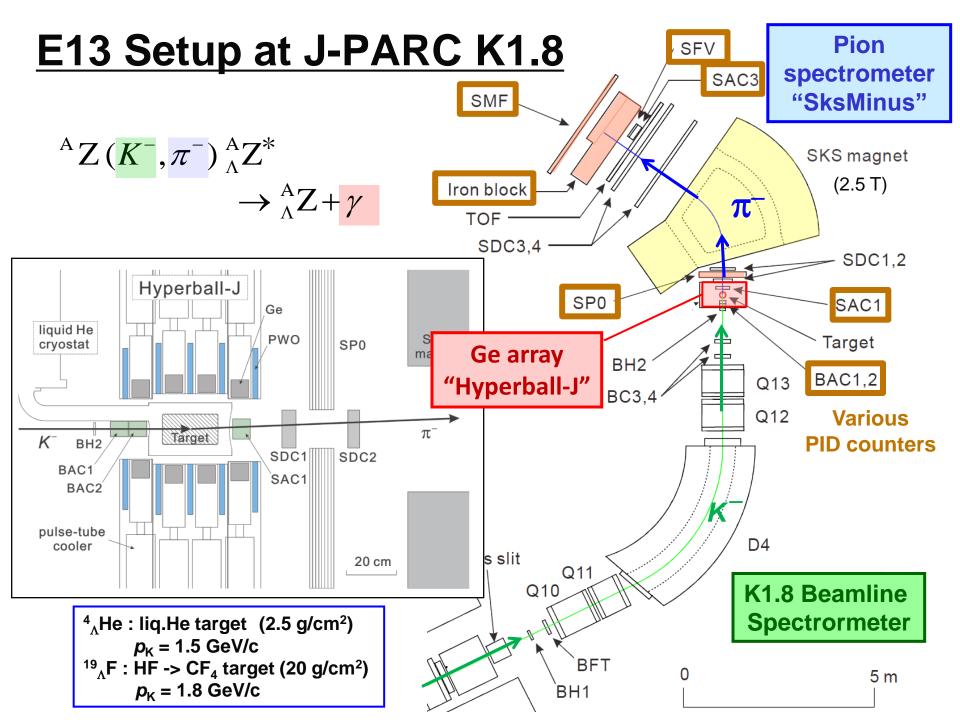
_	
Cent. Mom	- 115/ -125 MeV/c
Detector	DC, TOF, GC

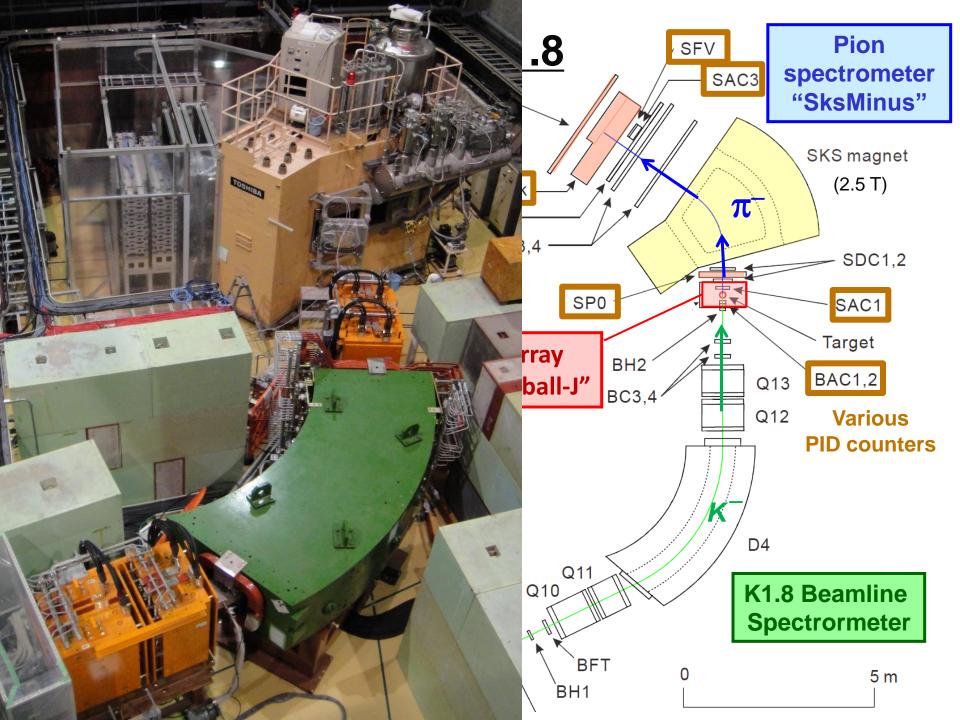


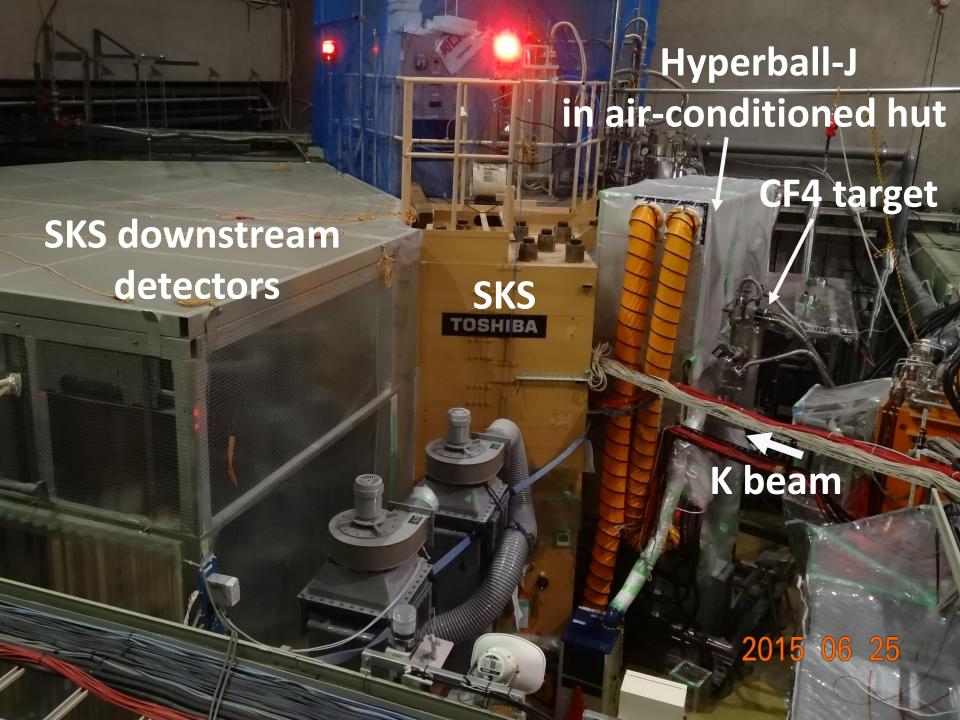


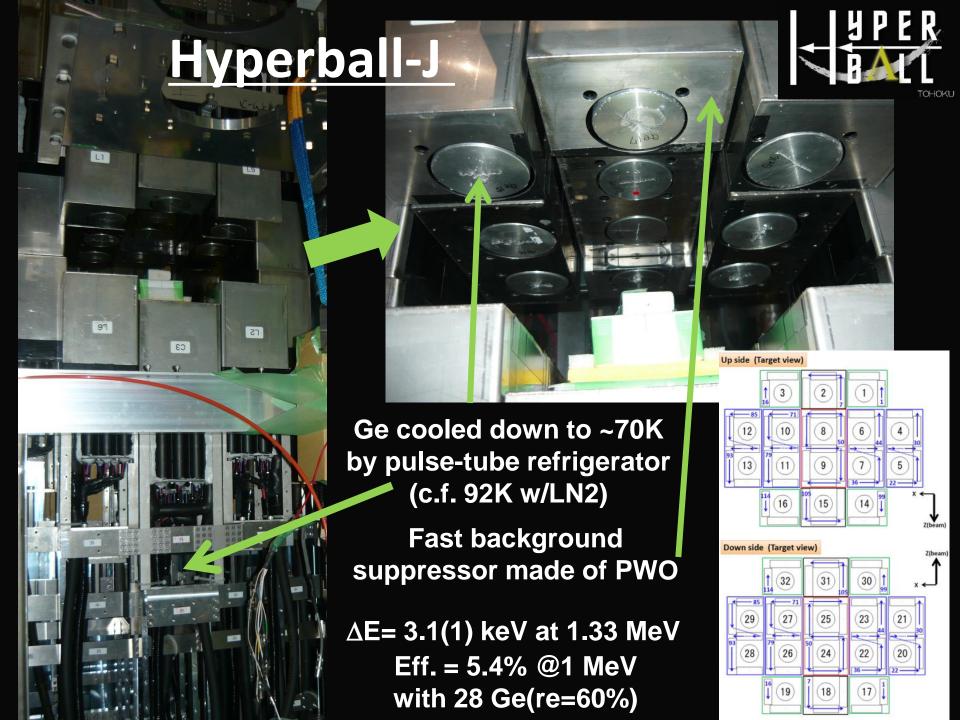
Hypernuclear γ-ray data as of 2014











E13 Missing mass: ${}^{4}\text{He}(K^{-},\pi^{-}) {}^{4}_{\Lambda}\text{He}$

 $p_{\rm K}$ =1.5 GeV/c

highly unbound region

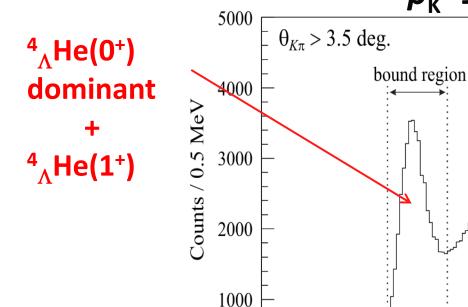
empty target

30

20

10

Excitation energy [MeV]



Peak width = resolution ~5.4 MeV (FWHM)

Byproduct: Spectrum for ${}^4_{\Sigma}$ He (p_{K} =1.5 GeV/c) was also successfully taken.

-20

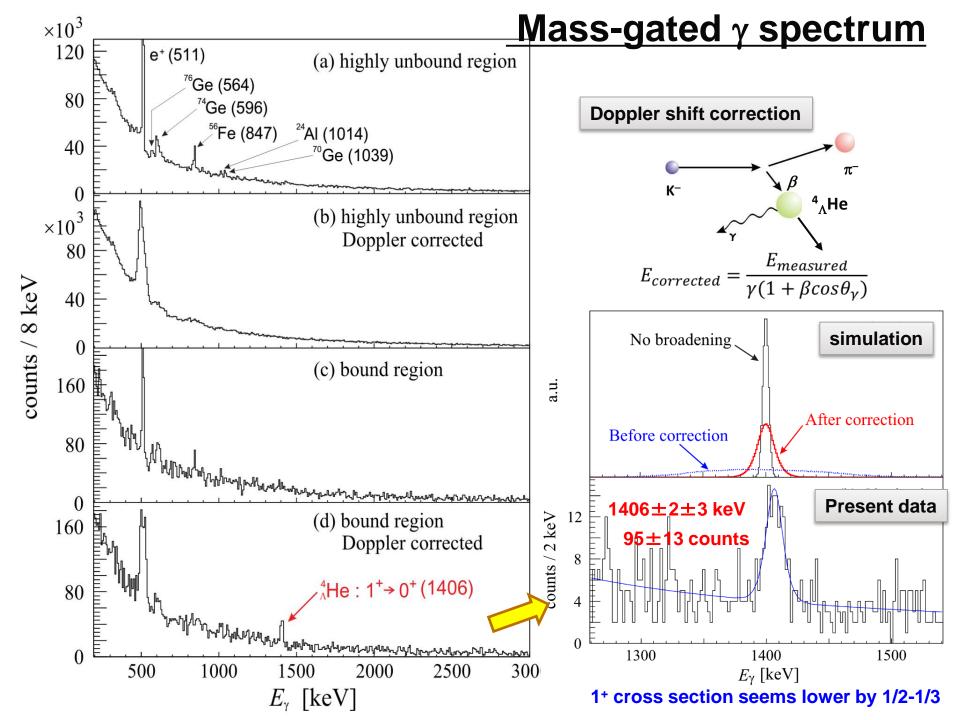
-10

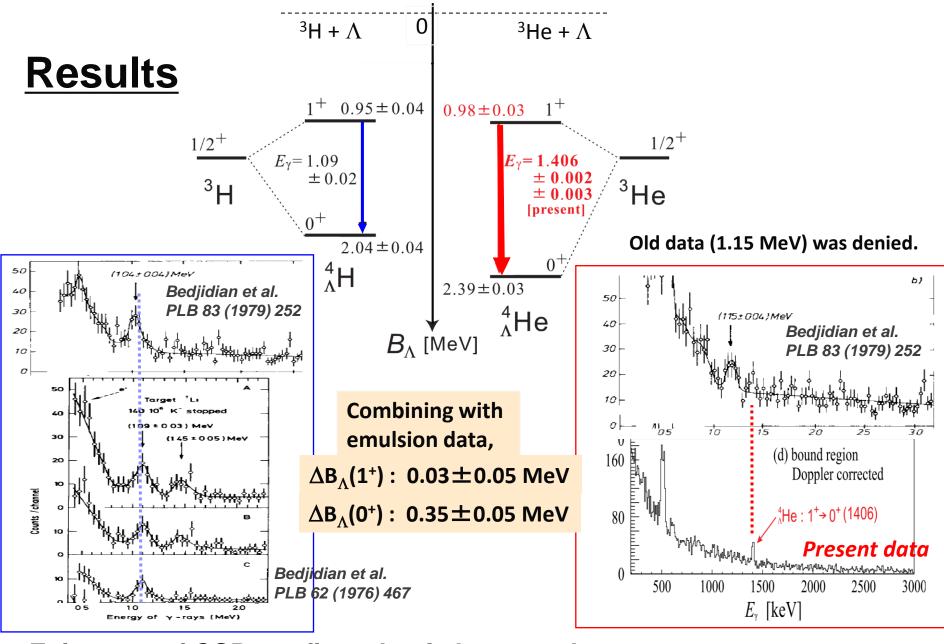
Black: with liq. He (physics run)

Blue: empty target

40

(w/o liq. He, w/target cell)





- Existence of CSB confirmed only by γ-ray data
- Large spin dependence in CSB found by combining with emulsion data

From Akaishi's slide $^4_{\Lambda}$ H 0+ S=1 pairs $\Lambda \mathbf{p} \Leftrightarrow -\sqrt{\frac{1}{3}} \Sigma^{0} \mathbf{p} + \sqrt{\frac{2}{3}} \Sigma^{+} \mathbf{n} \begin{bmatrix} \mathbf{s}_{3} = 1 \\ \mathbf{s}_{3} = 0 \\ \mathbf{s}_{3} = -1 \end{bmatrix}$ $\Lambda \mathbf{n} \Leftrightarrow \sqrt{\frac{1}{3}} \Sigma^{0} \mathbf{n} - \sqrt{\frac{2}{3}} \Sigma^{-} \mathbf{p} \begin{bmatrix} \mathbf{s}_{3} = 1 \\ \mathbf{s}_{3} = 0 \end{bmatrix}$ -1/3 Cancel +1/2 +1/3 +1/2 +1/2 added +1/2 Contribution to $U_{\Sigma\Lambda}$ 1/2 3/2 Λ -Σ coupling energy 1 9

<u>ΛN-ΣN coupling</u>

Our result strongly suggests that $\Lambda N - \Sigma N$ coupling is responsible for CSB, because $\Lambda N - \Sigma N$ coupling gives by one order smaller energy shift to 1+ state than to 0+ state.

Y. Akaishi et al. PRL 84 (2000) 3539

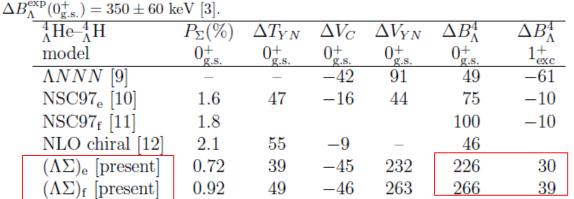
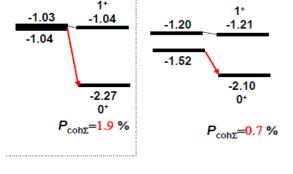


Table 2: Calculated CSB contributions to $\Delta B_{\Lambda}^4(0_{\rm g.s.}^+)$ and total values of $\Delta B_{\Lambda}^4(0_{\rm g.s.}^+)$ and

 $\Delta B_{\Lambda}^{4}(1_{\rm exc}^{+})$, in keV, from several model calculations of the A=4 hypernuclei. Recall that

"D2" potential A. Gal, PLB 744 (2015) 352



 $\Lambda N - \Sigma N$ D2 SC97e(S) central tensor-dominated

Theoretical studies will elucidate the origin of CSB and the Λ N - Σ N interaction.

3. Toward Strange matter in neutron stars

Under preparation (Partly) took data

Status of Strangeness NP @J-PARC

toward neutron star matter

- ----- S=-1 -----• n-rich Λ hypernuclei by (π^-, K^+) E10
 - γ spectroscopy of Λ hypernuclei
 - -> Λ N, Λ N- Σ N (Λ NN) int.
 - => Fraction of Λ in n-rich matter
 - K-pp by 3 He(K-,n) _{E15} K-pp by d(π +,K+) _{E27}

E13

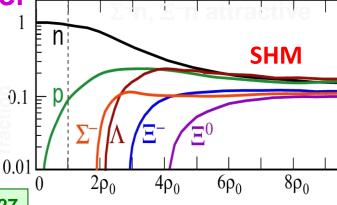
- -> K^{bar}N int. in matter => K condensation in n star? **Property of high density nuclear systems**
- $\triangleright \Sigma^{\pm}$ p scattering E40
 - -> Σ^- n (= Σ^+ p) (Quark Pauli effect), Σ^- p-> Λ N int. => Σ^- exists in n-star?

----- S=-2 -----

- **E07** ◆ ΛΛ hypernuclei
 - -> $\Lambda\Lambda$ interaction , $\Lambda\Lambda$ correlation?
- $\Rightarrow \Lambda$ fraction in Strange **Hadronic Matter**



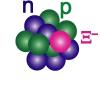
- -> **ΞN** interaction => **Ξ**⁻ exists in n-star?
- H dibaryon search from H-> $\Lambda\Lambda$, $\Lambda p\pi^-$ E42
 - -> Short-range BB force (Color magnetic int.)











First observation of a **E-nuclear** bound state

KEK E373 "Kiso event" K. Nakazawa et al. PTEP 2015, 033D02

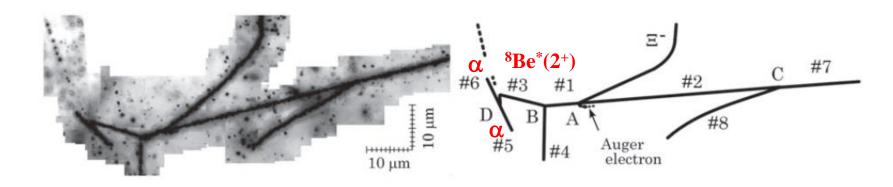


Fig. 1. A superimposed image from photographs and a schematic drawing of the KISO event.

$$\Xi^- + {}^{14}\mathrm{N} \rightarrow {}^{10}_{\Lambda}\mathrm{Be} + {}^{5}_{\Lambda}\mathrm{He}$$

$$B_{\Xi^{-}}=$$
 4.38 \pm 0.25 MeV $-$ 1.11 \pm 0.25 MeV $>>$ E(3D) $=$ 0.17 MeV $^{10}_{\Lambda}$ Be in g.s. $^{10}_{\Lambda}$ Be in highest excited state

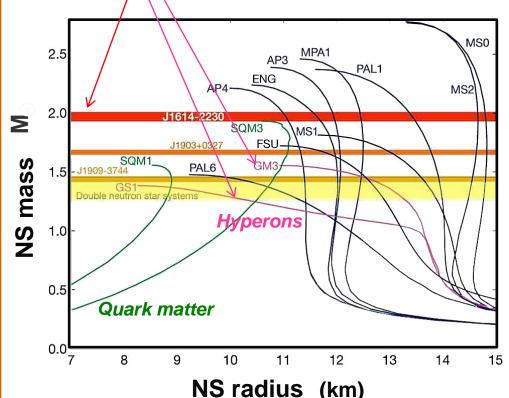
(Ehime pot.) $U_{\Xi} \sim 20 \text{ MeV} \implies B_{\Xi} (2\text{p state}) = 1.1 \text{ MeV}$

"The heavy neutron star puzzle"

- Hyperons must appear at $\rho = 2^3 \rho_0$
- EOS's with hyperons (or kaons) too soft -> can support M < 1.5 M_{sun}

PSR J1614-2230 (2010) $1.97\pm0.04~M_{sun}$ PSR J0348-0432 (2013) $2.01\pm0.04~M_{sun}$

We do not know BB interaction in high density nuclear matter!



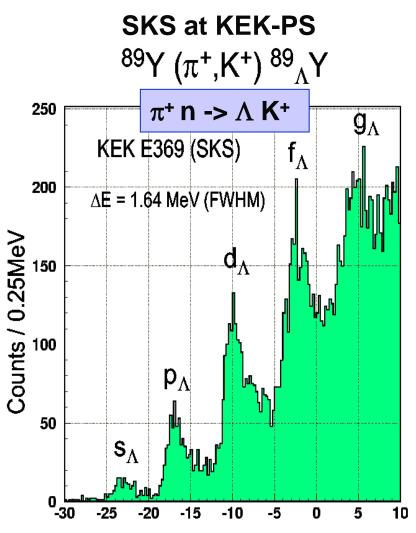
Unknown repulsion at high ρ exists

- Strong repulsion in three-body force including hyperons are necessary. (NNN, YNN, YYN, YYY)
- ■Phase transition to quark matter ? (quark star or hybrid star)

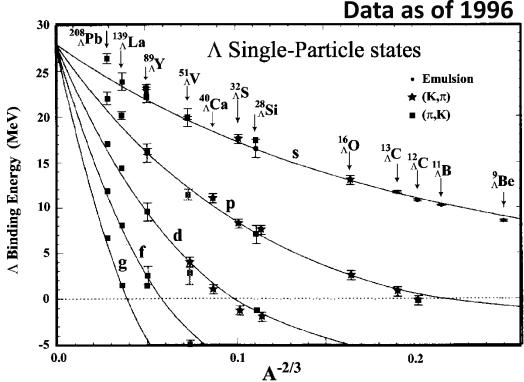
But we have no data on BBB force at high ρ nuclear matter, except for indirect info. in HI collisions.

-> Rijken, Pagliara,...

Previous (π^+ ,K $^+$) data and ΛN interaction



Mass of hypernucleus -B $_{\Lambda}$ (MeV) Hotchi et al., PRC 64 (2001) 044302



 U_{Λ} = -30 MeV (< U_{N} = -50 MeV) established better resolution n-rich hypernuclei



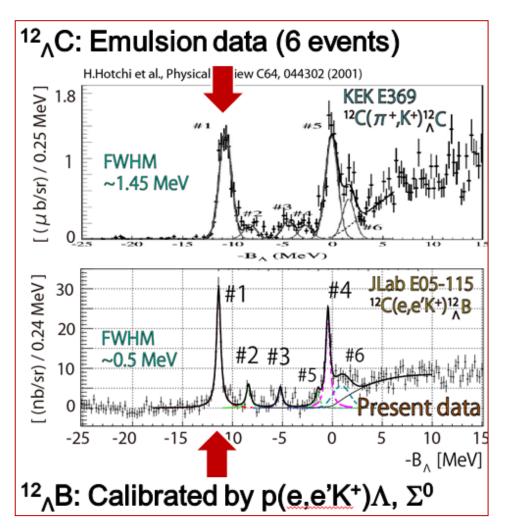
<u>(e,e'K+) at JLab</u>

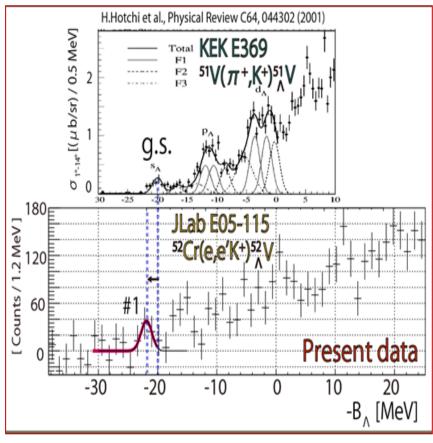
on ΛN int.

 γ spectroscopy and (π^- ,K+) at J-PARC

Slide by Nakamura

(e,e'K+) data at JLab



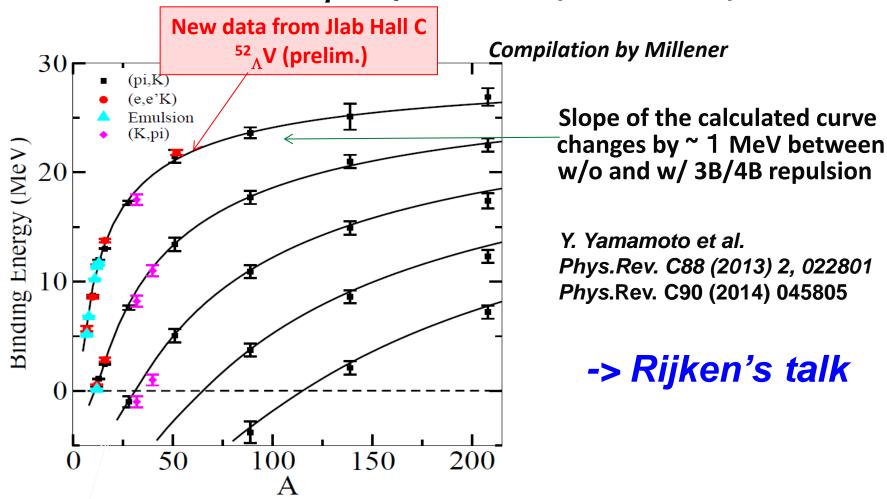


Going to heavier hypernuclei

Resolution ~0.5 MeV (FWHM) Absolute accurary ~0.1 keV

⁴⁰Ca/ ⁴⁸Ca target runs conditionally approved

Precise Hypernuclear B_{Λ} data provide information on ρ dependence (Λ NN force)?



 (π^+, K^+) , (K^-, π^-) systematic error ~1 MeV

(e,e'K'): systematic error \sim 0.1 MeV --- B_{\wedge} will be measured at JLab

J-PARC will also measure B_{Λ} by high resolution (π^+ , K^+) with 0.2 MeV resolution

Summary

- Light and heavier hypernuclear data provide information on YN, YY interactions in free space and in nuclear medium.
- New data on CSB in A=4 hypernuclei:
 - B_{Λ} of ${}^{4}_{\Lambda}H(0^{+})$ measured via pion decay spectroscopy
 - ---- Consistent with old emulsion data
 - 4 _{Λ}He(1+->0+) γ -ray measured to be 1.406 MeV

⇔1.09 MeV for ⁴_∧H

- -- A large CSB effect in ΛN interaction confirmed.
- --CSB has a spin dependence
- Ξ nuclear bound system (Kiso event) was observed for the first time -> Ξ potential is attractive
- To solve the hyperon puzzle, Λ NN 3-body force should be studied via precise B $_{\Lambda}$ measurements at JLab (+J-PARC).