# A-A Correlation in Relativistic Heavy Ion Collisions

Kenji Morita

(YITP, Kyoto)

in Collaboration with

Akira Ohnishi (YITP)

Takenori Furumoto (Ichinoseki National Col.)

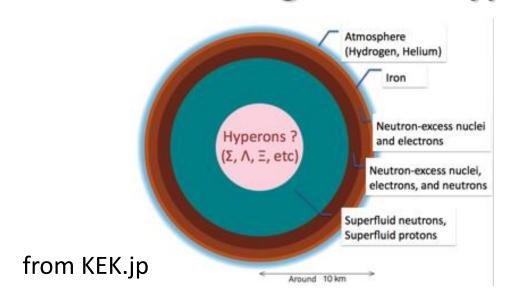
Ref.) Phys. Rev. C91, 024916 (2015)





### Role of AA Interaction

### Possible Emergence of Hyperons in NS core



To understand EoS,

Information on

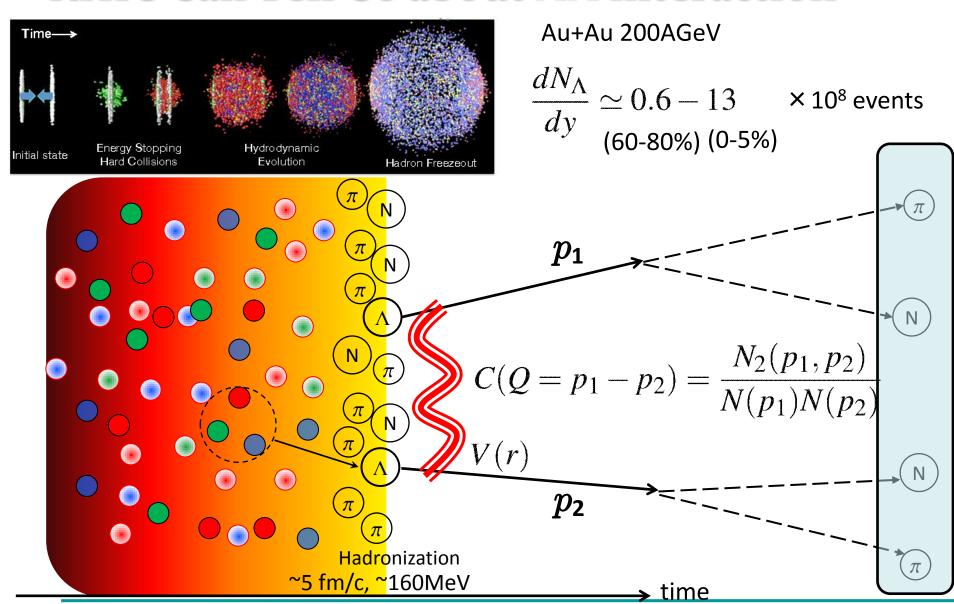
**Hyperon-Hyperon** 

Interaction is indispensable

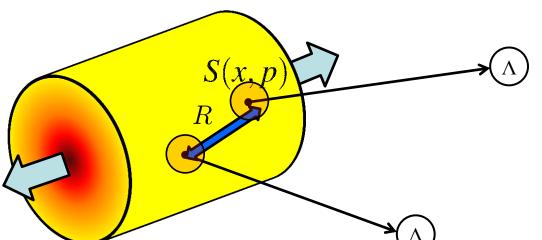
- H-dibaryon (uuddss)?

  - Resonance?

#### RHIC Can Tell Us about $\Lambda\Lambda$ interaction



#### **AA Correlation in HIC**



Independent (Chaotic) emission (←Thermal Source)

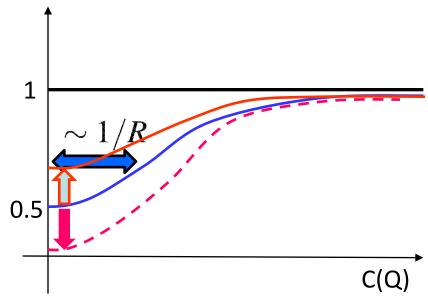
Identical particle correlation from quantum statistics (HBT effect)

C(Q): effective source size

#### $\Lambda\Lambda$ Interaction : No Coulomb!

Affects C(Q) when effective range r<sub>eff</sub> is comparable with R

Different results for repulsive and attractive interaction



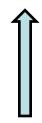
## Approach: Thermal Source + $V_{\Lambda\Lambda}$

#### Formula (Gong et al., '91)

$$C_2(Q,K) = \frac{\int d^4x_1 d^4x_2 S(x_1,K) S(x_2,K) |\Psi_{12}(Q,x_1-x_2-(t_2-t_1)K/m)|^2}{\int d^4x_1 d^4x_2 S(x_1,k_1) S(x_1,k_2)}$$

**Emission source function** 

- = direct emission + decay daughters
- = (direct + short-lived resonances)



+ Long-lived resonances

EM + Weak

Thermal source model (Mimic hydro)

 $\Lambda\Lambda$  relative wave function

Modification of S-wave by interaction

Various potentials (via 2 or 3 range Gaussian Fit)

Meson exchange models (Nijmegen model D, F, Soft Core89/97, ESC08)

<u>Phenomenological</u> (Ehime) <u>Quark model</u> (fss2)

<u>Fit to  $_{\Lambda\Lambda}$ </u> <sup>6</sup>He(Nagara) Filikhin-Gal (FG) Hiyama et al. (HKMYY)

#### **ΛΛ Wave Function**

$$|\Psi|^2 = \frac{1}{4}|\Psi_s|^2 + \frac{3}{4}|\Psi_t|^2 \longrightarrow \operatorname{Sp}_{N_t}$$

Spin: anti-symmetric

Spatially symmetric

**Modification in S-wave** 

No S-wave

$$\Psi_t = \frac{1}{\sqrt{2}} e^{2iK \cdot X} (e^{iQ \cdot r} - e^{-iQ \cdot r})$$

$$\Psi_s = \sqrt{2}e^{iK\cdot X}\left[\cos(Q\cdot r/2) + \chi_Q(r) - j_0(Qr/2)\right]$$

$$\left[-\frac{1}{m_{\Lambda}}\frac{d^2}{dr^2} + V(r)\right] \left[r\chi_Q(r)\right] = \frac{Q^2}{4m_{\Lambda}} \left[r\chi_Q(r)\right]$$

Schrödinger Eq.

## Potential, Wave func., / Correlation

#### Correlation Function for the Static Source

$$C_{\rm stat}(Q) = 1 - \frac{1}{2} e^{-Q^2 R^2} + \frac{1}{4\sqrt{\pi}R^3} \int\limits_0^\infty dr \, r^2 e^{-\frac{r^2}{4R^2}} \left[ \left[ \chi_Q(r) \right]^2 - \left[ j_0 \left( Qr/2 \right) \right]^2 \right] \\ \text{HBT (size R)} \\ \text{Interaction : deviation from free w.f.} \\ \frac{0.1}{NSC89-820} \\ \frac{NSC89-820}{FS2} \\ \frac{FS2}{FS} \\ \frac{FS}{NSC89-820} \\ \frac{FS2}{FS} \\ \frac{FS}{NSC89-820} \\ \frac{FS2}{NSC89-820} \\ \frac{FS2}{NSC89-820}$$

-0.1

0.6

0.8

1.2

r [fm]

1.4

-0.5

2

3

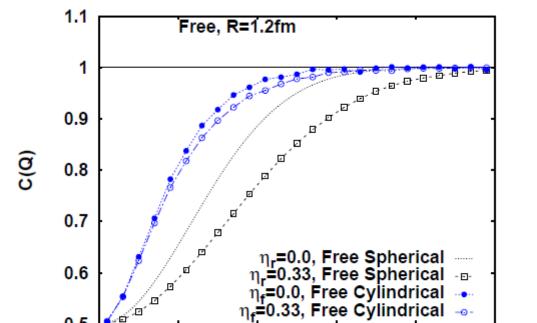
r/R

## **Expanding Source Model**

S.Chapman et al., '95

$$S(x,k) \propto m_T \cosh(y - Y_L) n_F (u \cdot k/T) \exp\left[-\frac{(\tau - \tau_0)}{2(\Delta \tau)^2} - \frac{x^2 + y^2}{2R^2}\right]$$

$$\sim \exp\left[-\frac{\gamma_T M_T}{T} \cosh(y - Y_L)\right] \exp\left[\frac{\gamma v_T k_T \cos \phi}{T}\right]$$
Give a finite longitudinal extent  $R_L \sim \tau_0 \sqrt{\frac{T}{M_T}}$ 



0.2

0.3

Q [GeV/c]

0.4

Width of C(Q)  $(Q=|p_1-p_2|)$ : effective 3d size

Large R<sub>L</sub>: narrow width

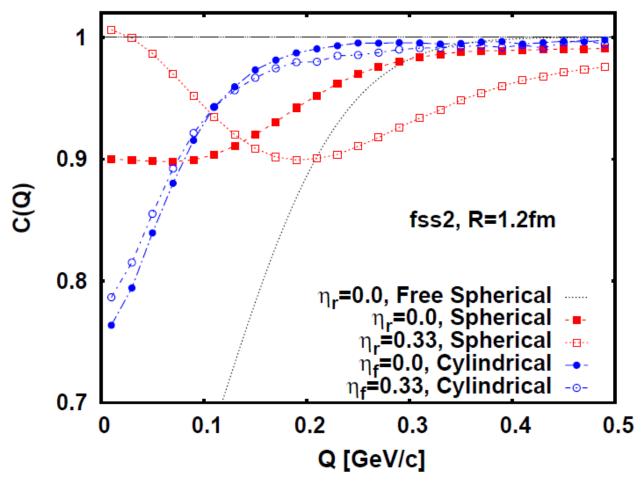
Transverse flow: fit to p<sub>t</sub> distribution of  $\Lambda$  (STAR '12)

0.5

0.1

0.5

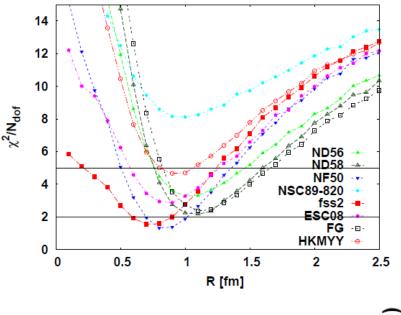
#### Results from Boost-invariant Source



Effect of  $\eta_f$  is rather small : longitudinal expansion dominates C(Q)

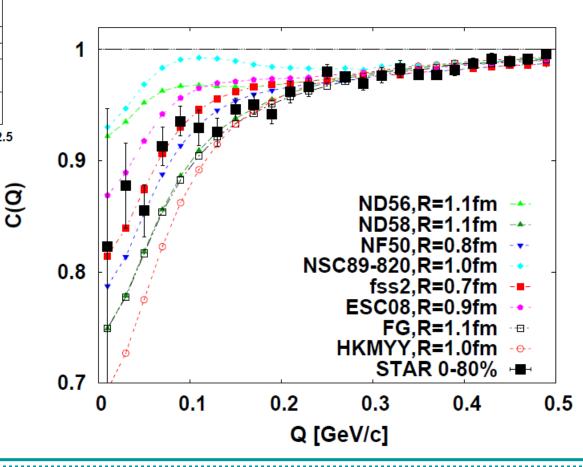
Behavior at small Q is different from the static source!

## $V_{\Lambda\Lambda}$ from Expanding Source Model



Minimize  $\chi^2$  against R

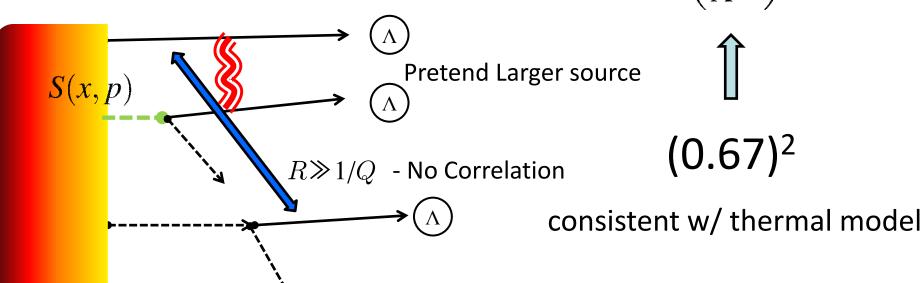
Low Q region sensitive to  $V_{\Lambda\Lambda}$ 



#### **Feed-Down Contribution**

- Short-Lived ( $\Sigma^*$ , N\* etc) : large R,  $\tau$  and  $\delta \tau$
- $\blacksquare \Xi \rightarrow \Lambda + \pi$  partly subtracted

$$C(Q) o 1 + \left(\frac{\Lambda^{\text{dir}}}{\Lambda^{\text{tot}}}\right)^2 (C(Q) - 1)$$



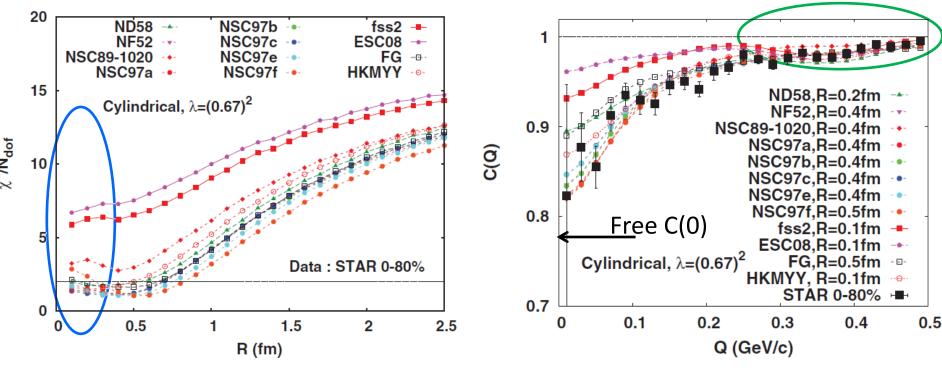
 $(0.52 if including <math>\Xi)$ 

Using  $\Sigma^0/\Lambda$  =0.278 (p+Be data) and  $\Xi/\Lambda$  = 0.15 (RHIC),  $\Lambda^{\rm dir}/\Lambda^{\rm tot}$  = 0.67

## Long Tail in C(Q)

$$C(Q) \rightarrow 1 + (0.67)^2 (C(Q) - 1)$$

## Sensitivity at low Q is reduced



Unphysically small size is preferred;

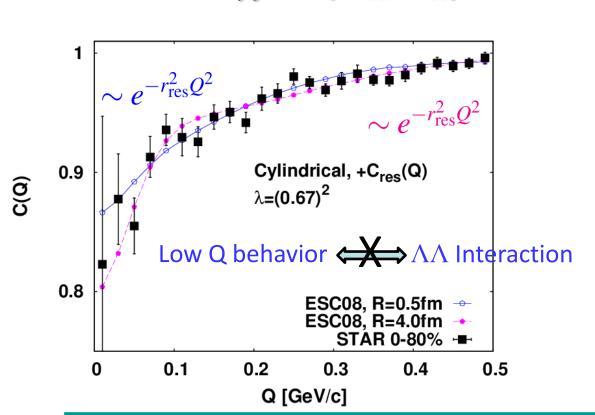
due to the long tail in C(Q) which cannot be included in the present framework

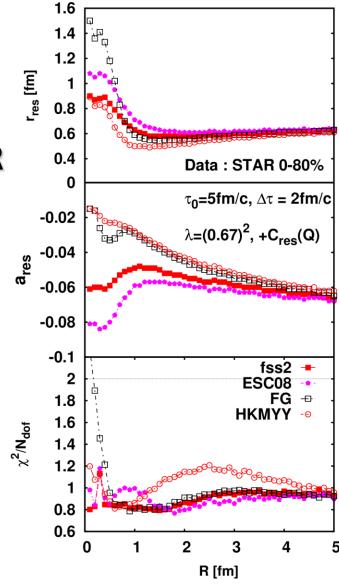
## Long Tail: Residual Correlation?



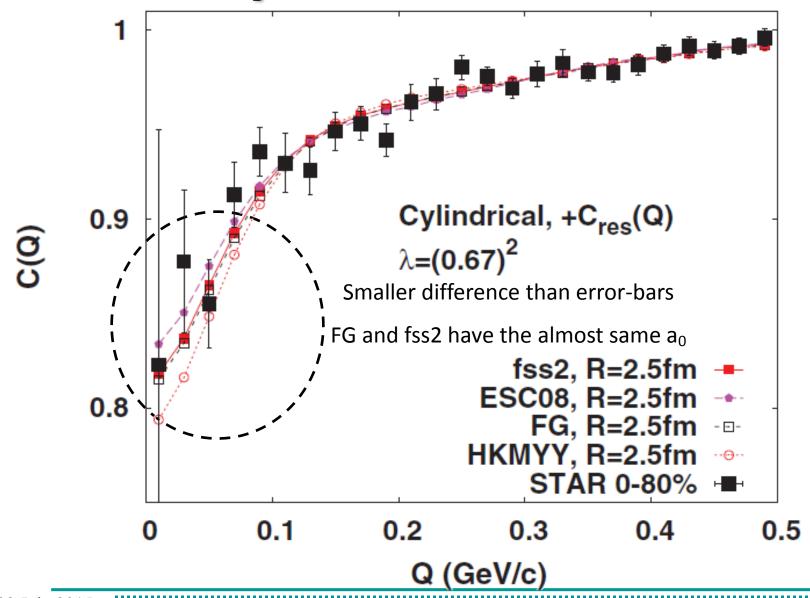
$$C(Q) \rightarrow C(Q) + a_{\text{res}}e^{-r_{\text{res}}^2Q^2}$$

 $\bullet$  minimize  $\chi^2$  in (a<sub>res</sub>, r<sub>res</sub>) for each R

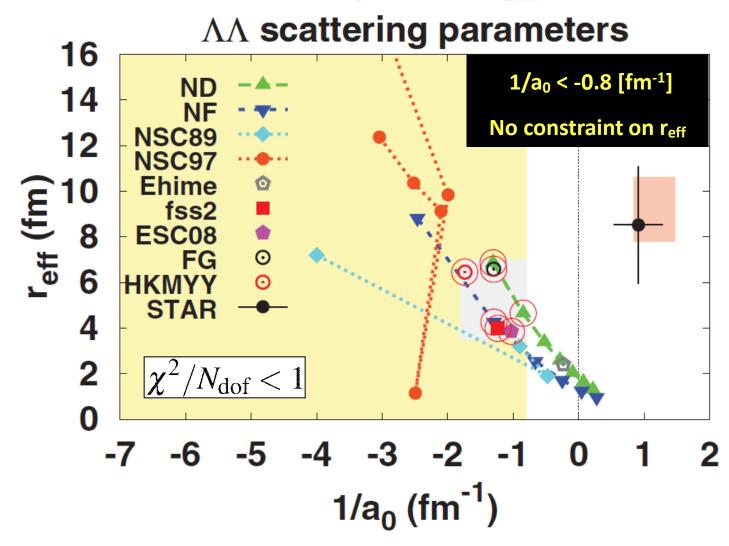




## Sensitivity to Interaction Remains



## Constraints on ao and reff



## Summary and Outlook

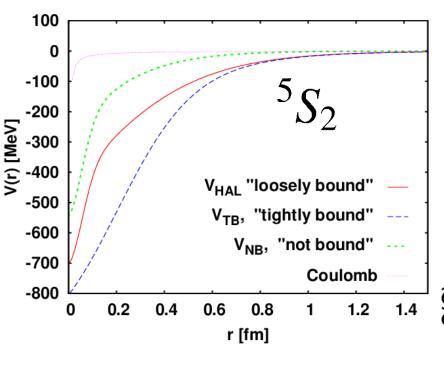
- $\blacksquare$  HIC have potential to determine  $\Lambda\Lambda$  interaction
- Ideal measurements (i.e., decay contribution is subtracted) will give stronger constraints
  - Feed-down effect reduces resolving power
- Long-tail in STAR data needs to be subtracted and its origin needs to be understood
- Scattering length  $1/a_0 < -0.8$  fm<sup>-1</sup>
  - Weakly attracting Implies no bound H-dibaryon
- Applicable to other systems
  - HIC as strange hadron factory!

#### $\Omega N$ Correlation

No Pauli Blocking: bound state candidate



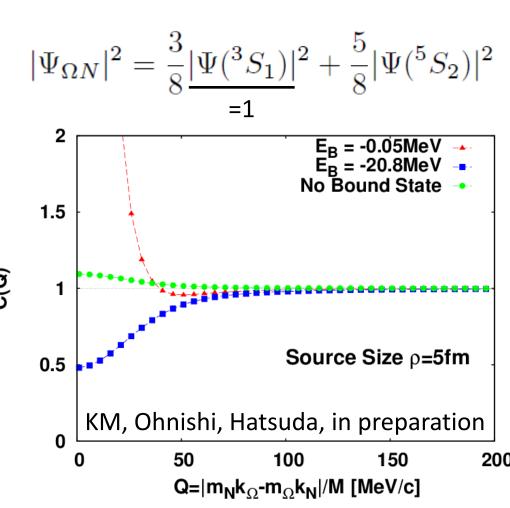
#### Potential from HAL QCD (Etminen et al., NPA'14)





Keep  $a_0$  and  $r_{\rm eff}$ 

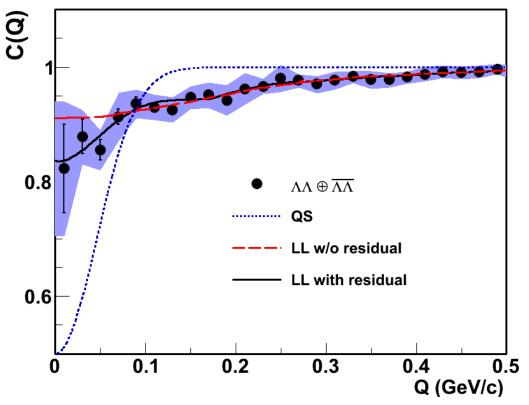
Tune parameters



## Backup

## Analysis by STAR Coll.

arXiv:1408.4360



Fit w/ Lednicky-Lyuboshitz model

Data: long tail (→ Small source size)

Introduce "residual correlation" for a better fit to data

$$a_0 = -1.10 \pm 0.37^{+0.68}_{-0.08} \text{ fm}$$

$$r_{\text{eff}} = 8.52 \pm 2.56^{+2.09}_{-0.74} \text{ fm}$$

$$\chi^2/N_{\text{dof}} = 0.56$$

$$C_{\rm fit}(Q) = N \left[ 1 + \lambda \left\{ -\frac{1}{2} e^{-r_0^2 Q^2} + \chi(a_0, r_{\rm eff}) \right\} + a_{\rm res} \exp(-r_{\rm res}^2 Q^2) \right]$$
 HBT (size r<sub>0</sub>) Interaction Residual correlation

0.18 1.006

2.96 fm

Residual correlation Interaction

-0.044

0.43 fm

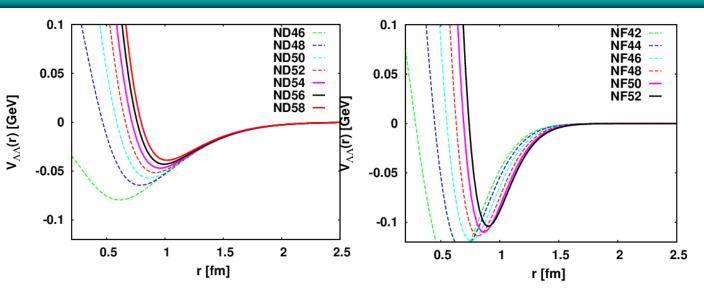
TABLE I:  $\Lambda\Lambda$  potentials. The scattering length  $(a_0)$  and effective range  $(r_{\text{eff}})$  are fitted using a two-range gaussian potential,  $V_{\Lambda\Lambda}(r) = V_1 \exp(-r^2/\mu_1^2) + V_2 \exp(-r^2/\mu_2^2)$ .

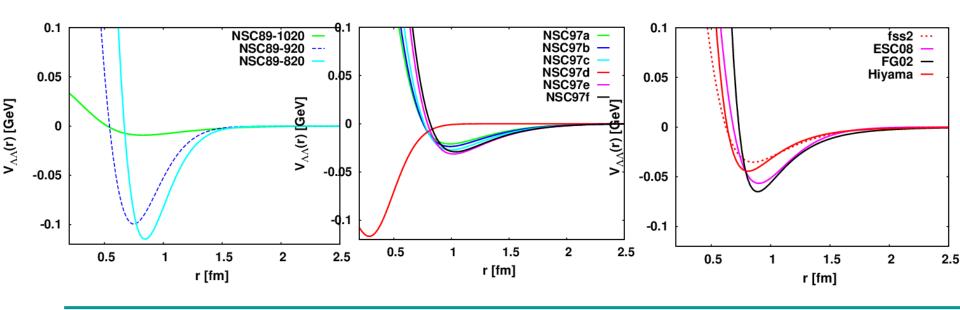
Model	$a_0$ (fm)	$r_{\rm eff}~({\rm fm})$	$\mu_1$ (fm)	$V_1 \; ({ m MeV})$	$\mu_2 \text{ (fm)}$	$V_2 \; ({ m MeV})$	Ref.
ND46	4.621	1.300	1.0	-144.89	0.45	127.87	[18] $r_c = 0.46 \text{ fm}$
ND48	14.394	1.633	1.0	-150.83	0.45	355.09	[18] $r_c = 0.48 \text{ fm}$
ND50	-10.629	2.042	1.0	-151.54	0.45	587.21	[18] $r_c = 0.50 \text{ fm}$
ND52	-3.483	2.592	1.0	-150.29	0.45	840.55	[18] $r_c = 0.52 \text{ fm}$
ND54	-1.893	3.389	1.0	-147.65	0.45	1114.72	[18] $r_c = 0.54 \text{ fm}$
ND56	-1.179	4.656	1.0	-144.26	0.45	1413.75	[18] $r_c = 0.56 \text{ fm}$
ND58	-0.764	6.863	1.0	-137.74	0.45	1666.78	[18] $r_c = 0.58 \text{ fm}$
NF42	3.659	0.975	0.6	-878.97	0.45	1048.58	[19] $r_c = 0.42 \text{ fm}$
NF44	23.956	1.258	0.6	-1066.98	0.45	1646.65	[19] $r_c = 0.44 \text{ fm}$
NF46	-3.960	1.721	0.6	-1327.26	0.45	2561.56	[19] $r_c = 0.46 \text{ fm}$
NF48	-1.511	2.549	0.6	-1647.40	0.45	3888.96	[19] $r_c = 0.48 \text{ fm}$
NF50	-0.772	4.271	0.6	-2007.35	0.45	5678.97	[19] $r_c = 0.50 \text{ fm}$
NF52	-0.406	8.828	0.6	-2276.73	0.45		[19] $r_c = 0.52 \text{ fm}$
NSC89-1020	-0.250	7.200	1.0	-22.89	0.45	67.45	$[20] m_{\text{cut}} = 1020 \text{ MeV}$
NSC89-920	-2.100	1.900	0.6	-1080.35	0.45	2039.54	$[20] m_{\text{cut}} = 920 \text{ MeV}$
NSC89-820	-1.110	3.200	0.6	-1904.41	0.45	4996.93	$[20] m_{\text{cut}} = 820 \text{ MeV}$
NSC97a	-0.329	12.370	1.0	-69.45	0.45	653.86	[21]
NSC97b	-0.397	10.360	1.0	-78.42	0.45	741.76	[21]
NSC97c	-0.476	9.130	1.0	-91.80	0.45	914.67	[21]
NSC97d	-0.401	1.150	0.4	-445.77	0.30	373.64	[21]
NSC97e	-0.501	9.840	1.0	-110.45	0.45	1309.55	[21]
NSC97f	-0.350	16.330	1.0	-106.53	0.45	1469.33	[21]
Ehime	-4.21	2.41	1.0	-146.6	0.45	720.9	[23]
fss2	-0.81	3.99	0.92	-103.9	0.41	658.2	[25]
ESC08	-0.97	3.86	0.80	-293.66	0.45	1429.27	[22]

TABLE II:  $\Lambda\Lambda$  potentials from Nagara event. The scattering length  $(a_0)$  and effective range  $(r_{\rm eff})$  are fitted using a three-range gaussian potential,  $V_{\Lambda\Lambda}(r) = V_1 \exp(-r^2/\mu_1^2) + V_2 \exp(-r^2/\mu_2^2) + V_3 \exp(-r^2/\mu_3^2)$ .

Model	$a_0$ (fm)	$r_{ m eff}~({ m fm})$	$\mu_1$ (fm)	$V_1 \; ({ m MeV})$	$\mu_2 \text{ (fm)}$	$V_2 \; ({ m MeV})$	$\mu_3$ (fm)	$V_3~({ m MeV})$	Ref.
HKMYY	-0.575	6.45	1.342	-10.96	0.777	-141.75	0.35	2136.6	[3]
FG	-0.77	6.59	1.342	-21.49	0.777	-250.13	0.35	9324.0	[2]

#### Kenji Morita (YITP, Kyoto)

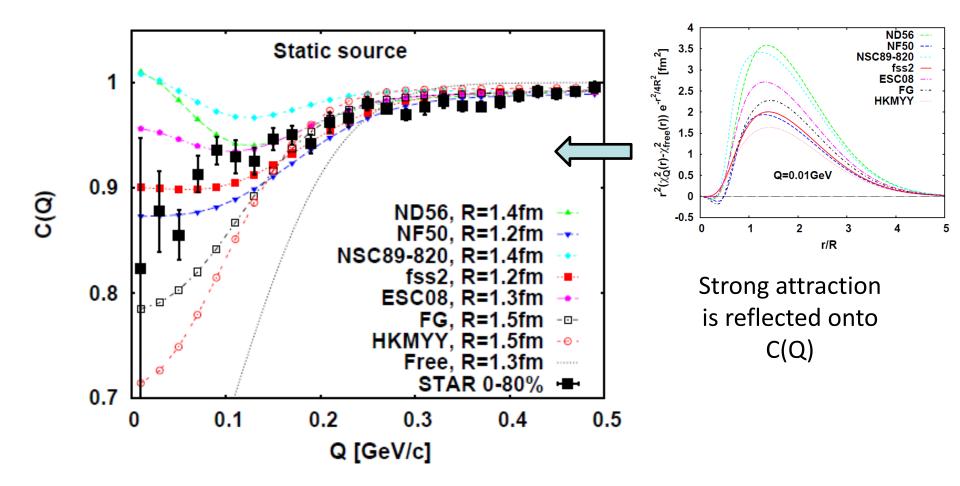




22 July 2015

EMMI Workshop on anti-matter, hyper-matter and exotica production at the LHC

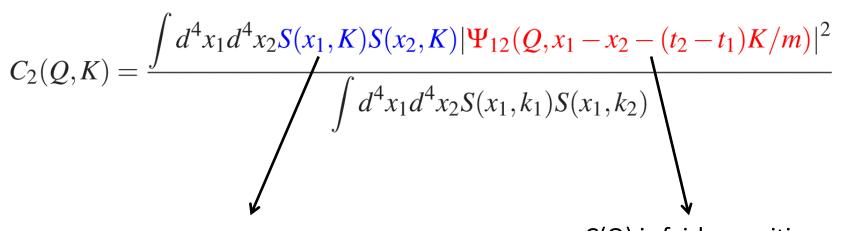
#### Results from the Static Source



Larger variation among potentials than data error-bars

Size : determined from min.  $\chi^2$ 

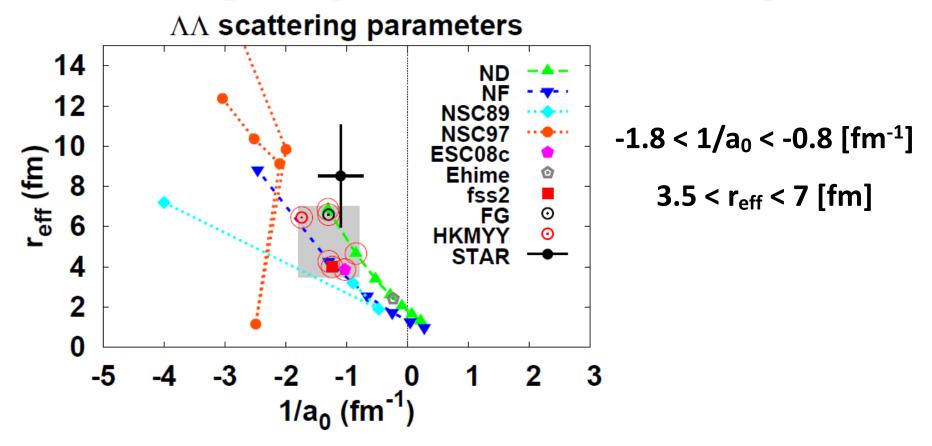
## **Collectivity Deforms Source Function**



Influence on the best-fit potentials?

C(Q) is fairly sensitive to interaction

## Scattering Length and Effective Range



## Approach: Thermal Source + $V_{\Lambda\Lambda}$

#### Formula (Gong et al., '91)

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Emission source func.



Thermal source model (Mimic hydro)

- Static Spherically Symmetric
- Spherically Symmetric + Hubble Flow
- Cylindrically Symmetric + Boostinvariance + Transverse flow

 $\Lambda\Lambda$  relative S-wave func.



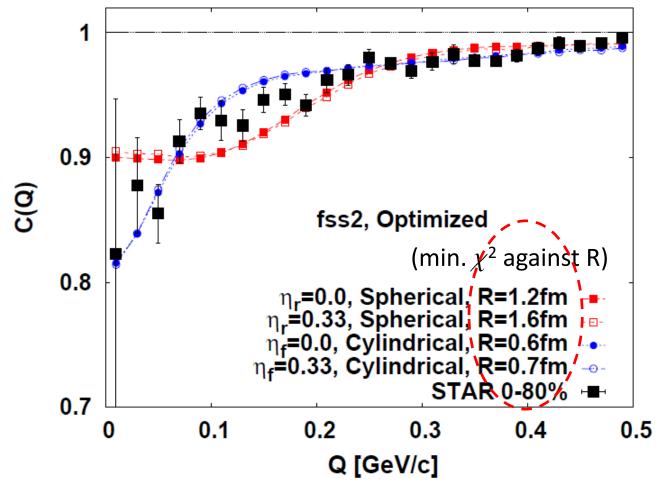
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<u>Fit to  $_{\Lambda\Lambda}$ </u> <sup>6</sup>He(Nagara) Filikhin-Gal (FG) Hiyama et al. (HKMYY)

## Combined effects from flow and $V_{\Lambda\Lambda}$



Effect of  $\eta_f$  is absorbed into change of  $R_{opt}$ 

Longitudinal expansion gives another type of "best fit"