



EMMI Workshop: Anti-matter, hyper-matter and exotica production at the LHC

20-22 July 2015

CERN

Europe/Berlin timezone



Hypernuclei – Hyperon-hyperon interactions

Hypernuclei

Few-body Λ hypernuclei

Exact calculations possible

$A=3$ ${}^3_{\Lambda}\text{H}$

$B_{\Lambda}=130$ keV \Rightarrow stringent test for ΛN interaction

$\tau < 200$ ps (by HIC): why so short??

Weak decay branches not measured well

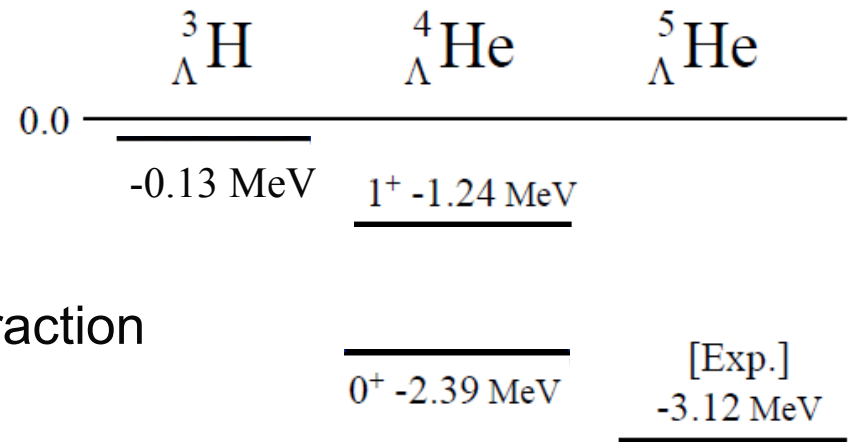
$nn\Lambda$ bound ?

$pn\Lambda(T=1)$?

$A=4$ ${}^4_{\Lambda}\text{H}/{}^4_{\Lambda}\text{He}$

$0+/1+$ ΛN spin-spin, $\Lambda\text{N}-\Sigma\text{N}$ interaction

Charge symmetry breaking ?



$A=5$ ${}^5_{\Lambda}\text{He}$

Overbinding problem

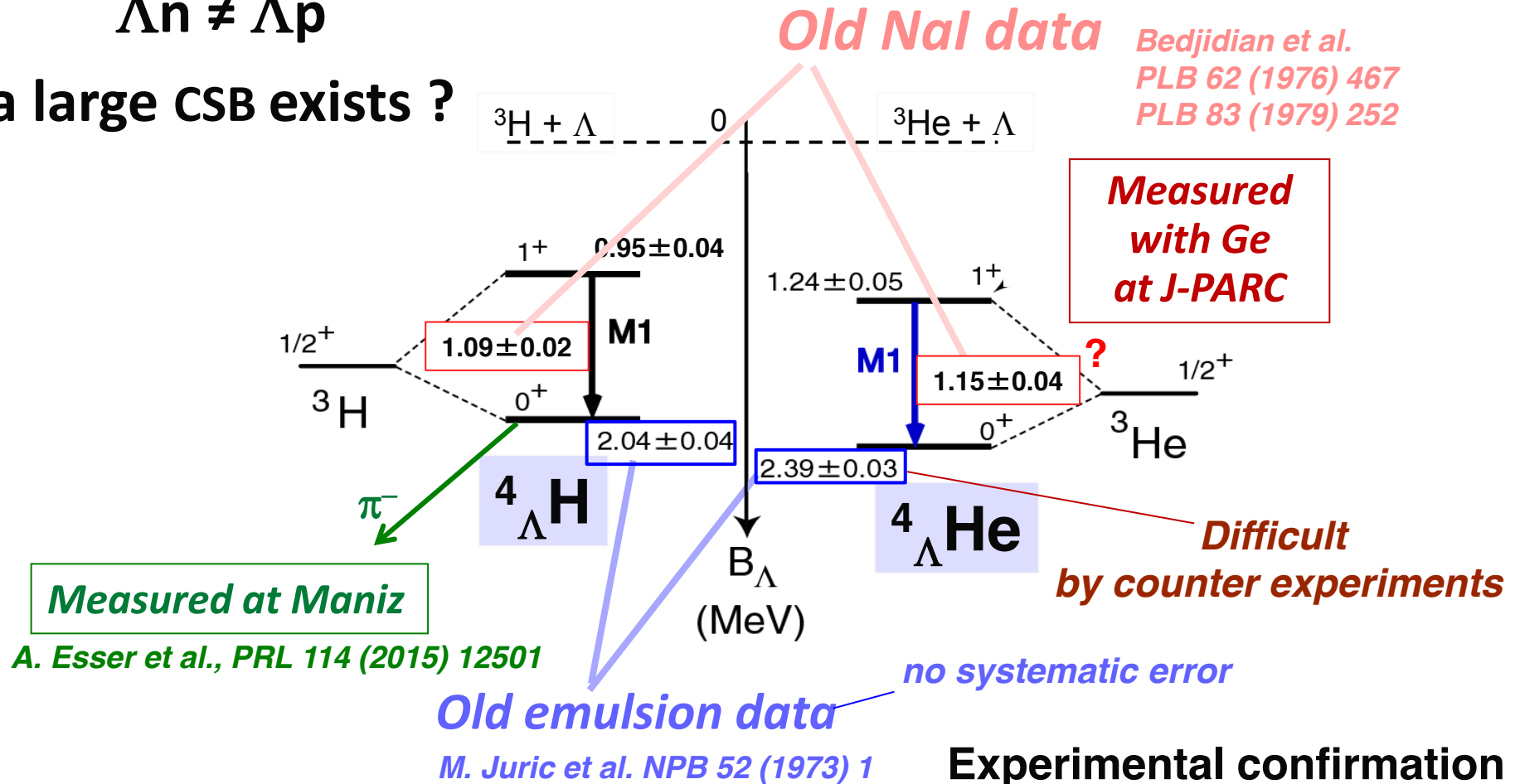
$\Rightarrow B_{\Lambda}(A=3,4,5)$ explained well with $\Lambda\text{N}-\Sigma\text{N}$ interaction

C. Rappold et al. PLB 728 (2014) 543

Charge Symmetry Breaking puzzle in A=4

$$\Lambda n \neq \Lambda p$$

a large CSB exists ?

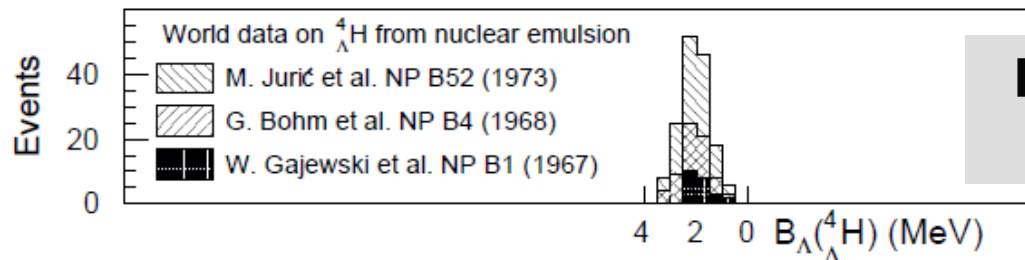


Origin is unknown.

ΛN - ΣN coupling? But 4-body calc's with Λ - Σ mixing using Nijmegen interactions give $\Delta B < 100\text{keV}$

=> Long standing puzzle

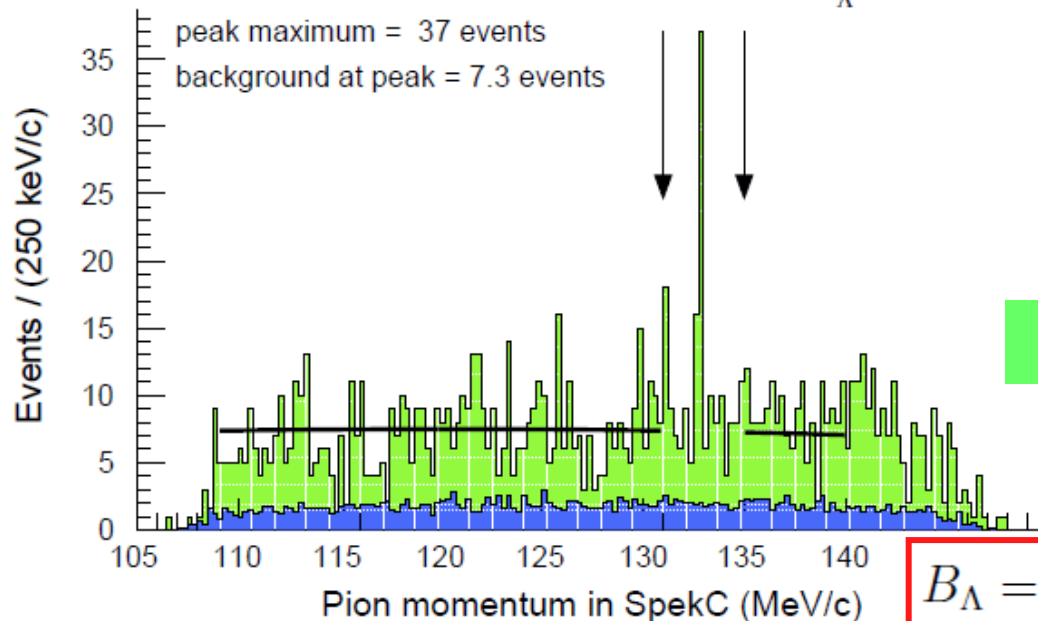
Experimental confirmation of CSB also necessary



**Emulsion
data**

electroproduction

MAMI Result



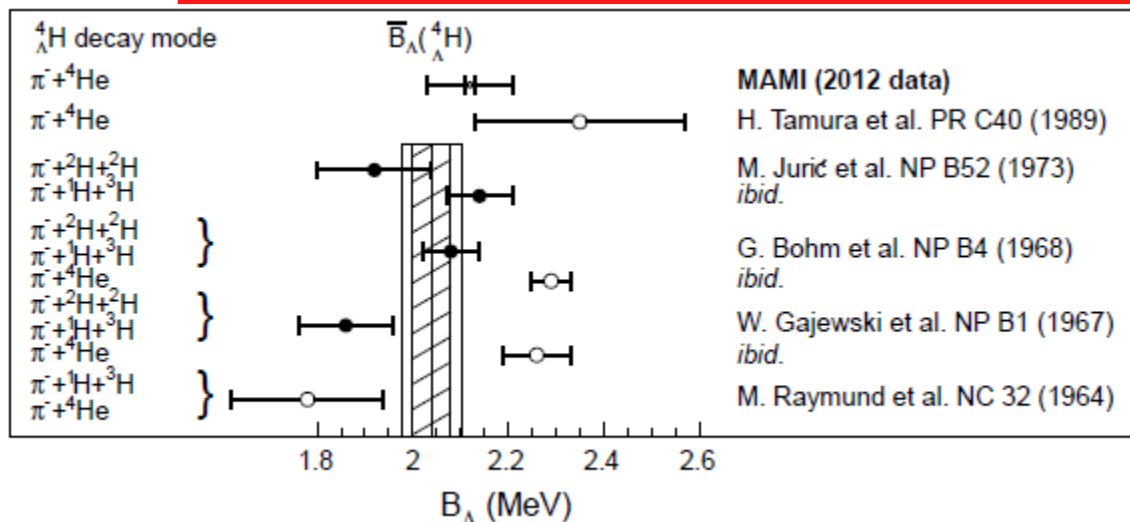
MAMI data

$$B_{\Lambda} = 2.12 \pm 0.01 \text{ (stat.)} \pm 0.09 \text{ (syst.) MeV}$$

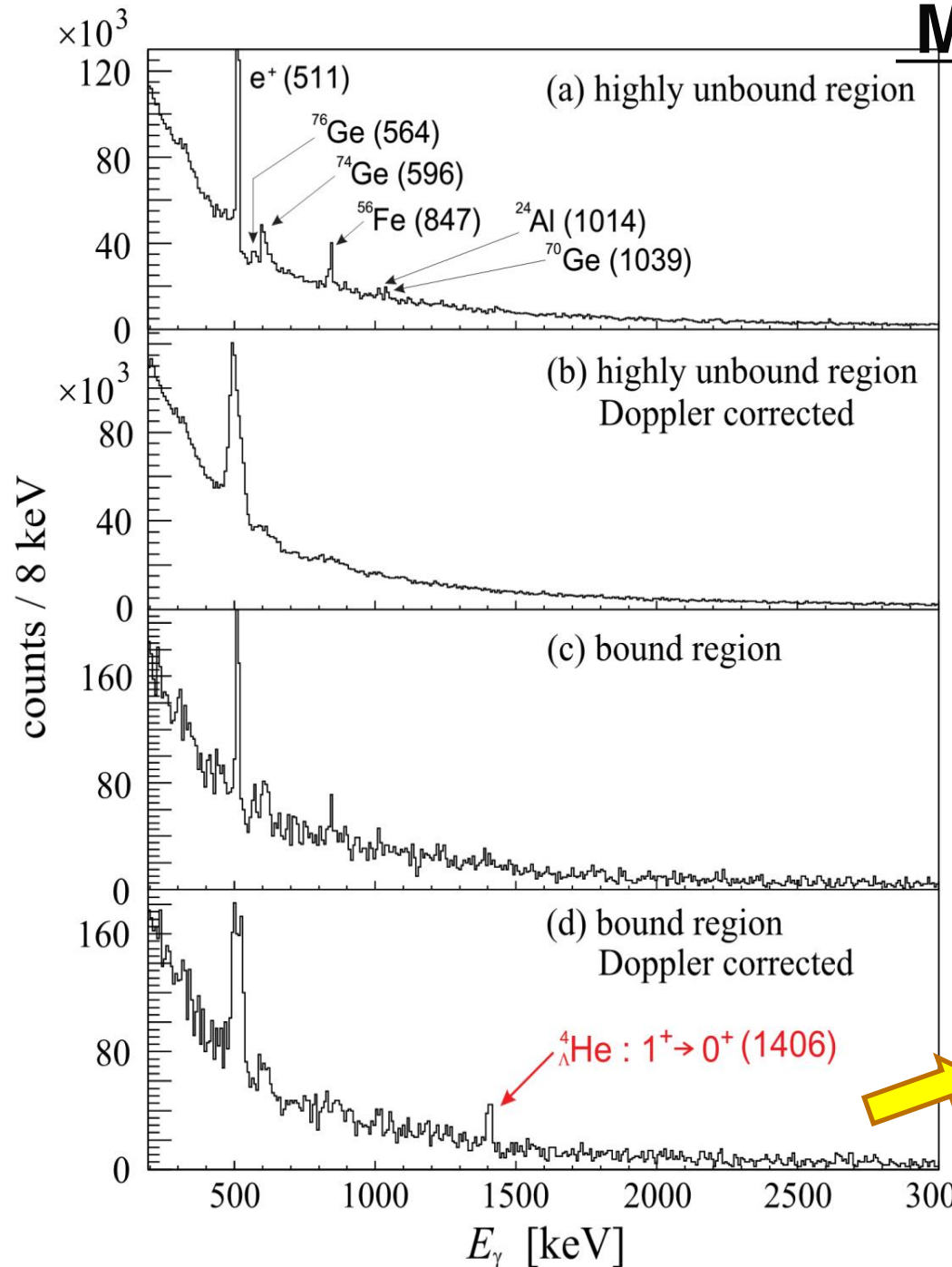
**A. Esser et al.,
PRL 114 (2015) 12501**

A1 Collaboration

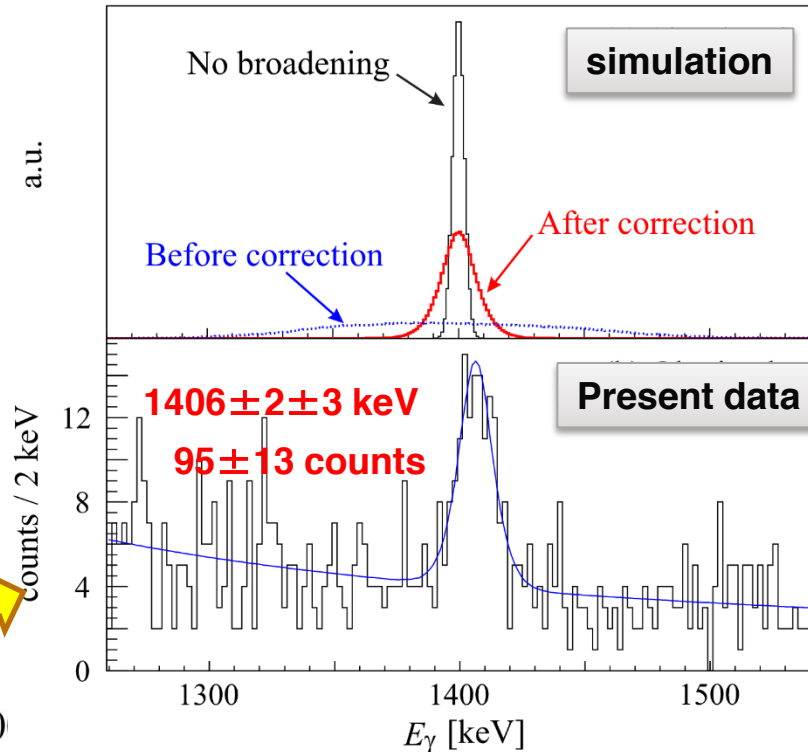
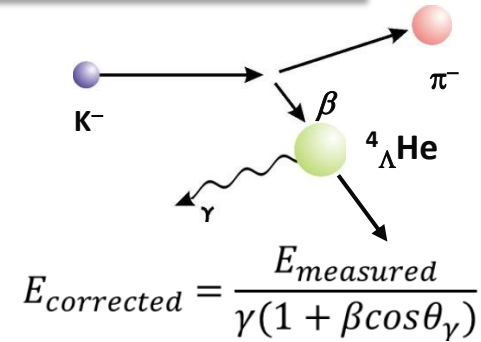
Sys. Error



Mass-gated γ spectrum

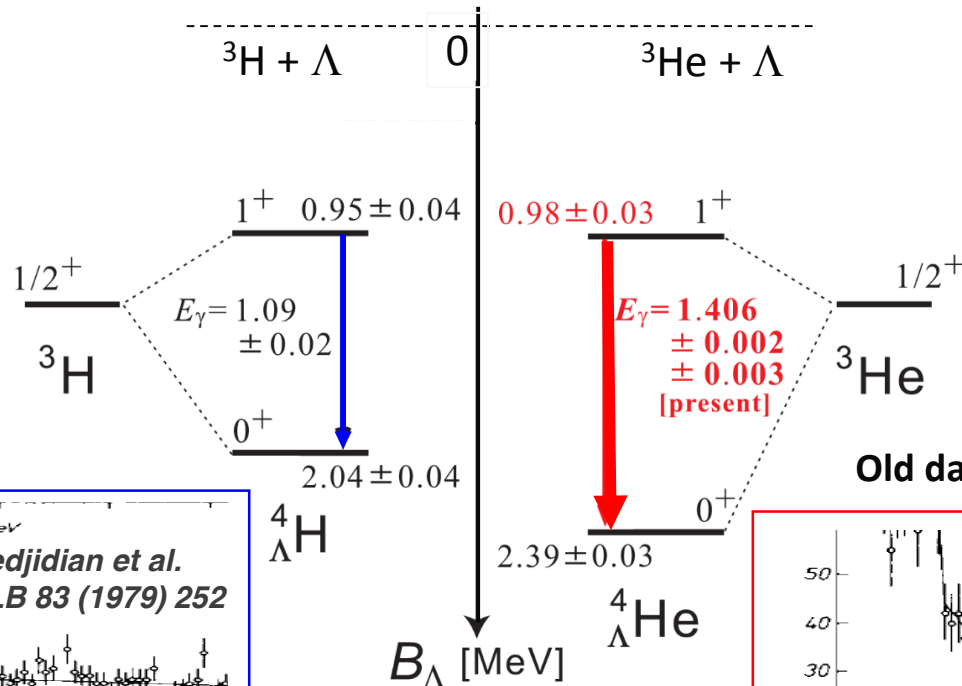


Doppler shift correction



1^+ cross section seems lower by 1/2-1/3

Results

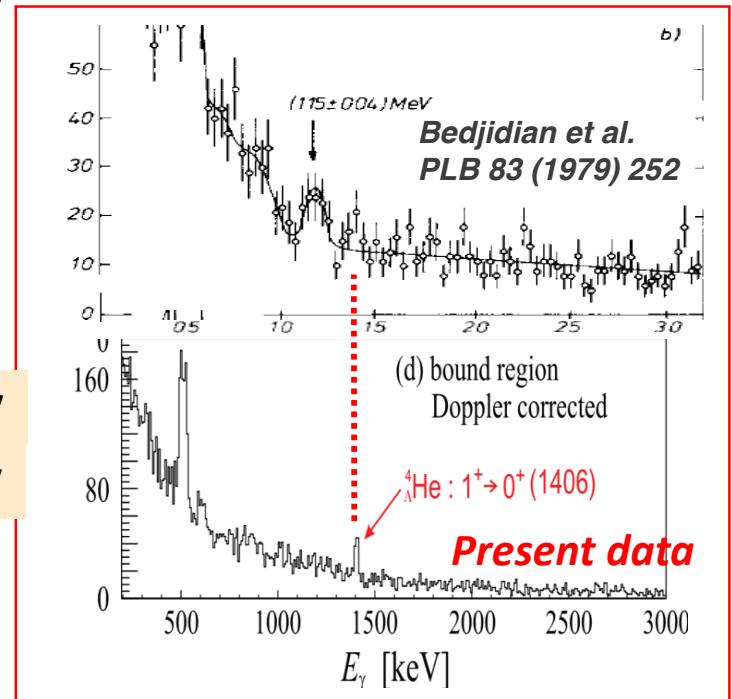
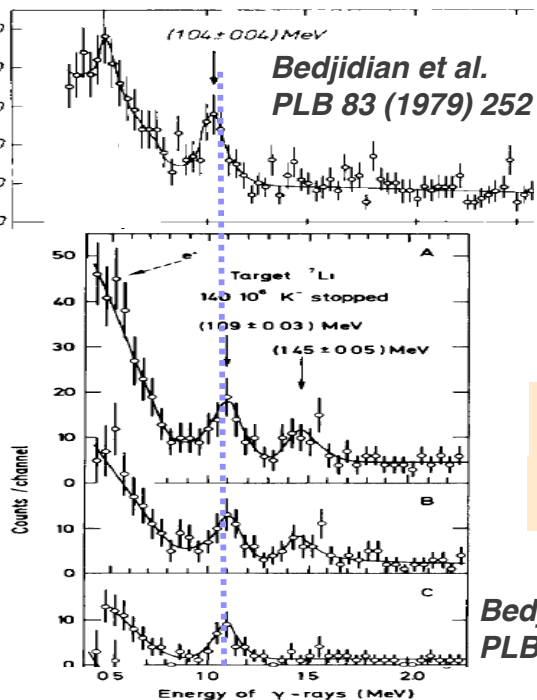


Old data (1.15 MeV) was denied.

Combining with
emulsion data,

$$\Delta B_\Lambda(1^+) : 0.03 \pm 0.05 \text{ MeV}$$

$$\Delta B_\Lambda(0^+) : 0.35 \pm 0.05 \text{ MeV}$$



■ Existence of CSB confirmed only by γ -ray data

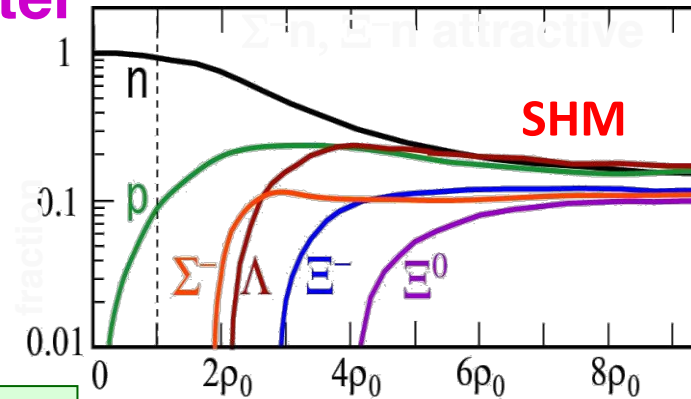
■ Large spin dependence in CSB found by combining with emulsion data

- New data on CSB in $A=4$ hypernuclei:
 - B_{Λ} of ${}^4_{\Lambda}\text{H}(0^+)$ measured via pion decay spectroscopy
 - Consistent with old emulsion data
 - ${}^4_{\Lambda}\text{He}(1^+ \rightarrow 0^+) \gamma$ -ray measured to be 1.406 MeV
 - $\Leftrightarrow 1.09$ MeV for ${}^4_{\Lambda}\text{H}$
 - A large CSB effect in ΛN interaction confirmed.
 - CSB has a spin dependence
- Ξ nuclear bound system (Kiso event) was observed for the first time $\rightarrow \Xi$ potential is attractive

Under preparation
(Partly) took data

Status of Strangeness NP @J-PARC

toward neutron star matter



----- S=-1 -----

- ◆ n-rich Λ hypernuclei by (π^-, K^+) E10
- ◆ γ spectroscopy of Λ hypernuclei E13



-> ΛN , ΛN - ΣN (ΛNN) int.

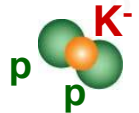
=> Fraction of Λ in n-rich matter

- ◆ K^-pp by $^3\text{He}(K^-, n)$ E15
- ◆ K^-pp by $d(\pi^+, K^+)$ E27

-> $K^{\text{bar}}N$ int. in matter => K condensation in n star?

- ◆ $\Sigma^\pm p$ scattering E40

Property of high density nuclear systems



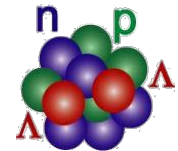
-> $\Sigma^- n$ ($= \Sigma^+ p$) (Quark Pauli effect) , $\Sigma^- p \rightarrow \Lambda N$ int. => Σ^- exists in n-star?

----- S=-2 -----

- ◆ $\Lambda\Lambda$ hypernuclei E07

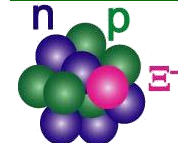
-> $\Lambda\Lambda$ interaction , $\Lambda\Lambda$ correlation?

=> Λ fraction in Strange Hadronic Matter



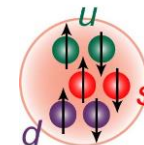
- ◆ Ξ hypernuclear spectroscopy E05
- ◆ Ξ atomic X rays E03, E07

-> ΞN interaction => Ξ^- exists in n-star?



- ◆ H dibaryon search from $H \rightarrow \Lambda\Lambda, \Lambda p \pi^-$ E42

-> Short-range BB force (Color magnetic int.)



First observation of a Ξ -nuclear bound state

KEK E373 “Kiso event”

K. Nakazawa et al. PTEP 2015, 033D02

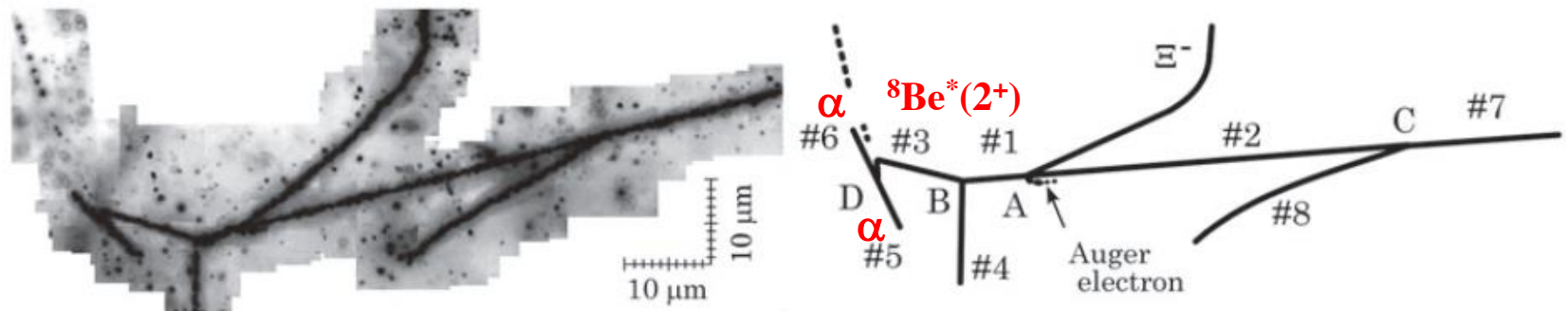
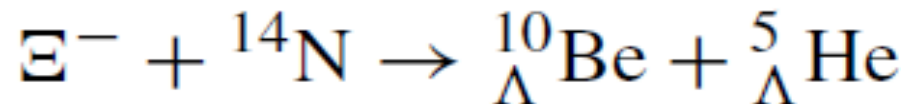


Fig. 1. A superimposed image from photographs and a schematic drawing of the KISO event.



$$B_{\Xi^-} = \boxed{4.38 \pm 0.25 \text{ MeV}} - \boxed{1.11 \pm 0.25 \text{ MeV}} \gg E(3D) = 0.17 \text{ MeV}$$

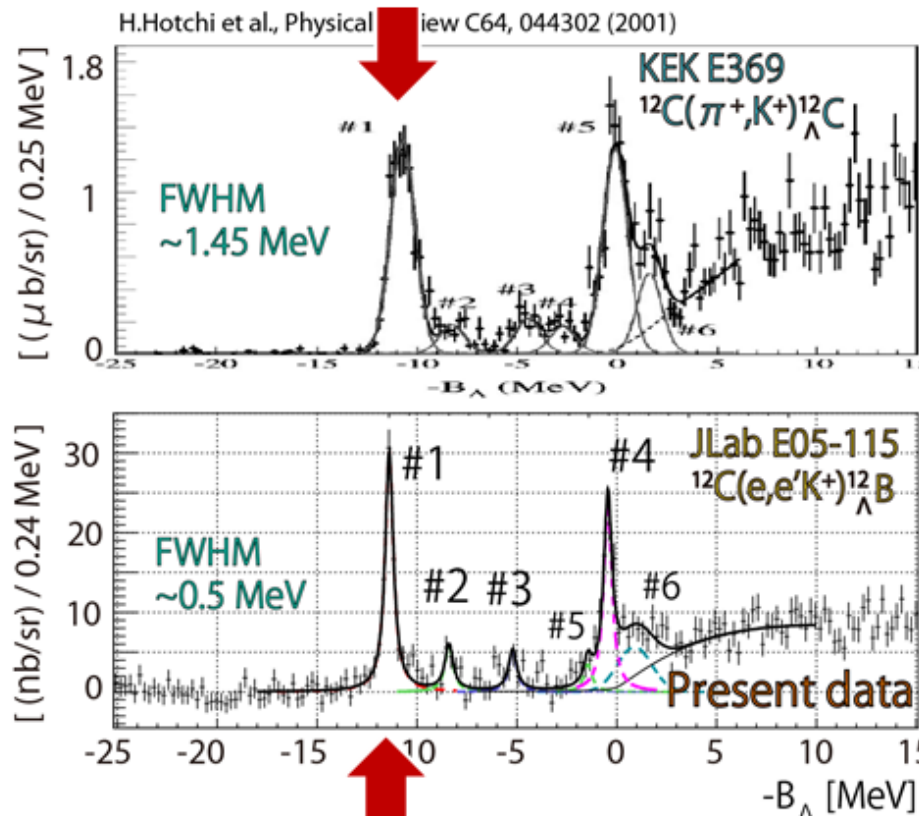
${}_{\Lambda}^{10}\text{Be}$ in g.s. ${}_{\Lambda}^{10}\text{Be}$ in highest excited state

$$(E_{\text{hime pot.}}) \quad U_{\Xi} \sim 20 \text{ MeV} \Rightarrow B_{\Xi}(2p \text{ state}) = 1.1 \text{ MeV}$$

(e,e'K⁺) data at JLab

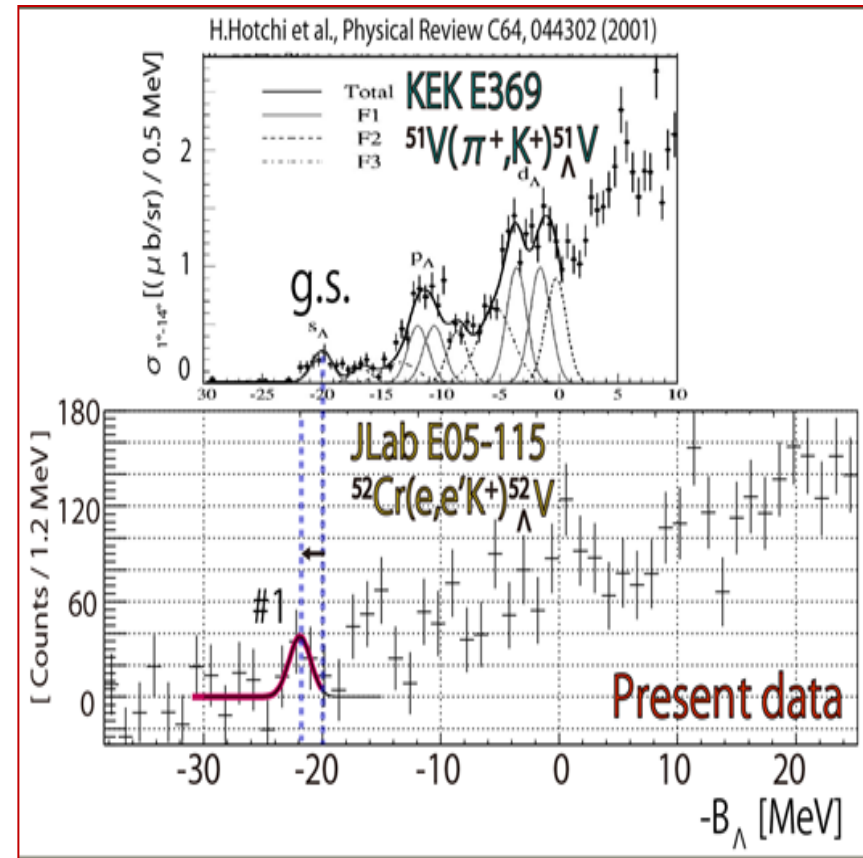
Slide by Nakamura

¹²_ΛC: Emulsion data (6 events)



¹²_ΛB: Calibrated by p(e,e'K⁺)_Λ, Σ⁰

Resolution ~0.5 MeV (FWHM)
Absolute accuracy ~0.1 keV

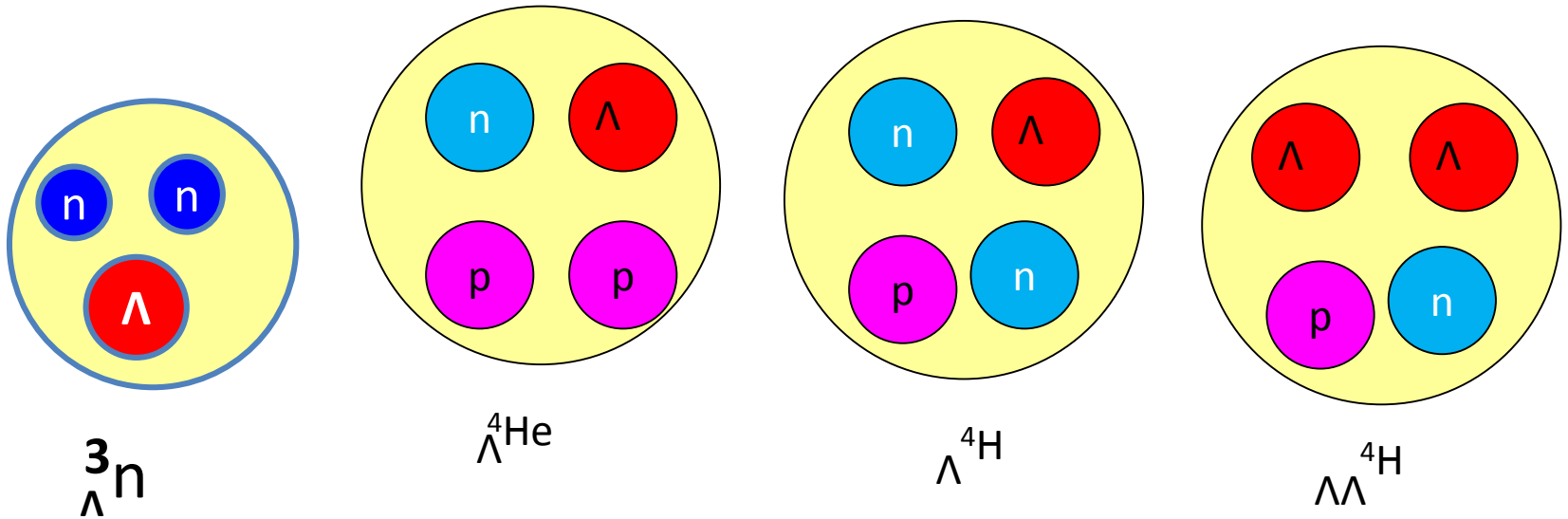


Going to heavier hypernuclei

**⁴⁰Ca/ ⁴⁸Ca target runs
conditionally approved**

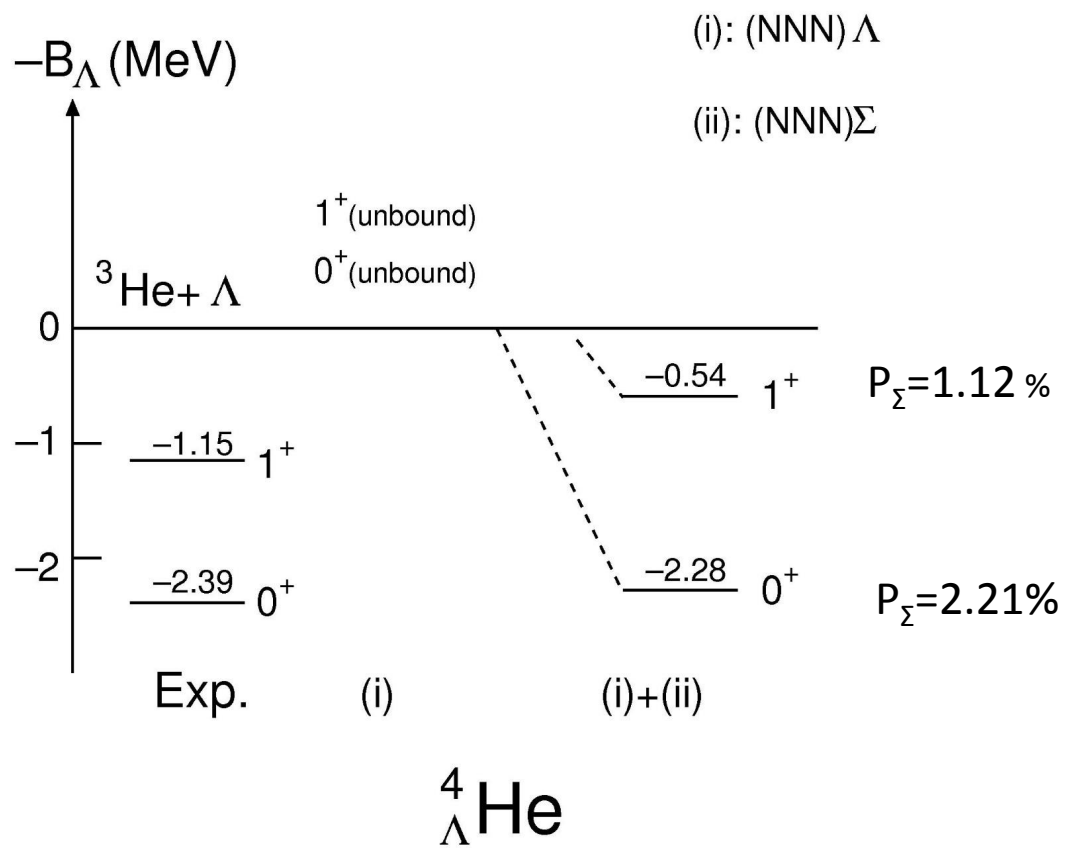
Experiment – theory

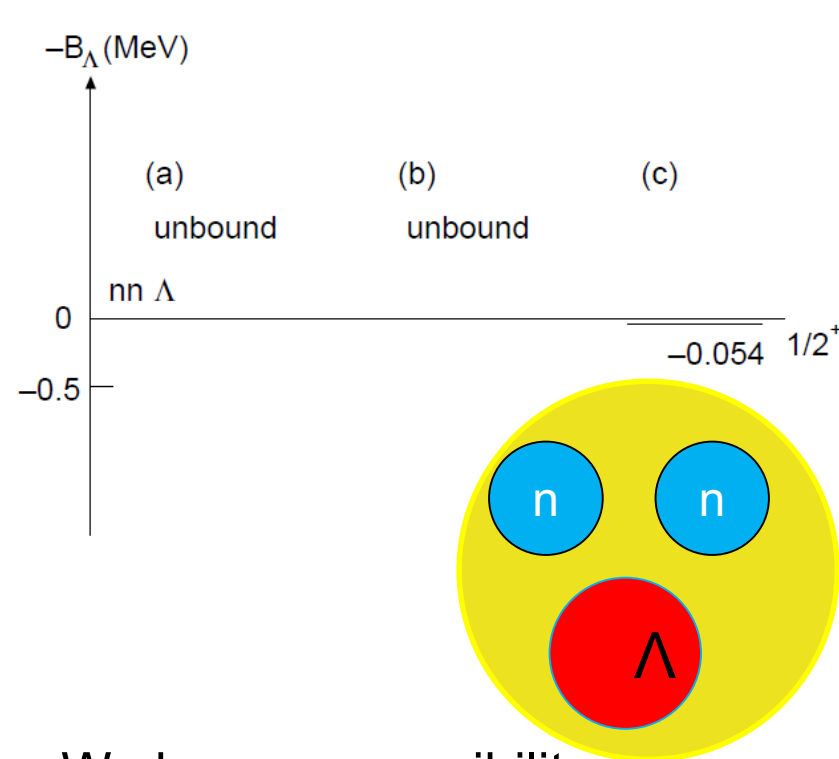
One of the important and interesting subjects :
to study three- and four-baryon systems



C. Rappold et al.,
HypHI collaboration
Phys. Rev. C 88,
041001 (R) (2013)

A. Esser, PRL114,
232501(2015)





We have no possibility to have a bound state in $nn\Lambda$ system.

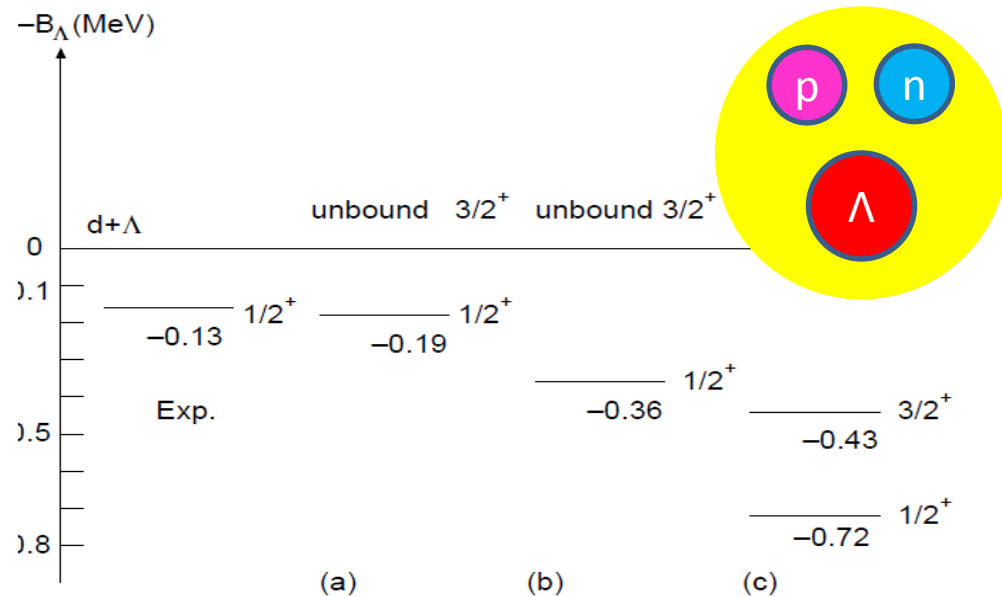
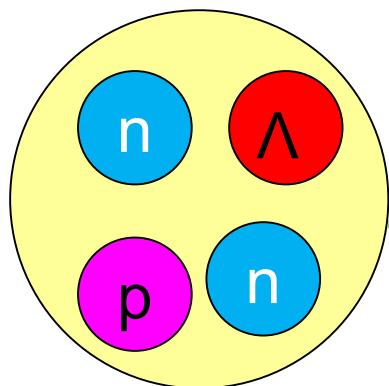
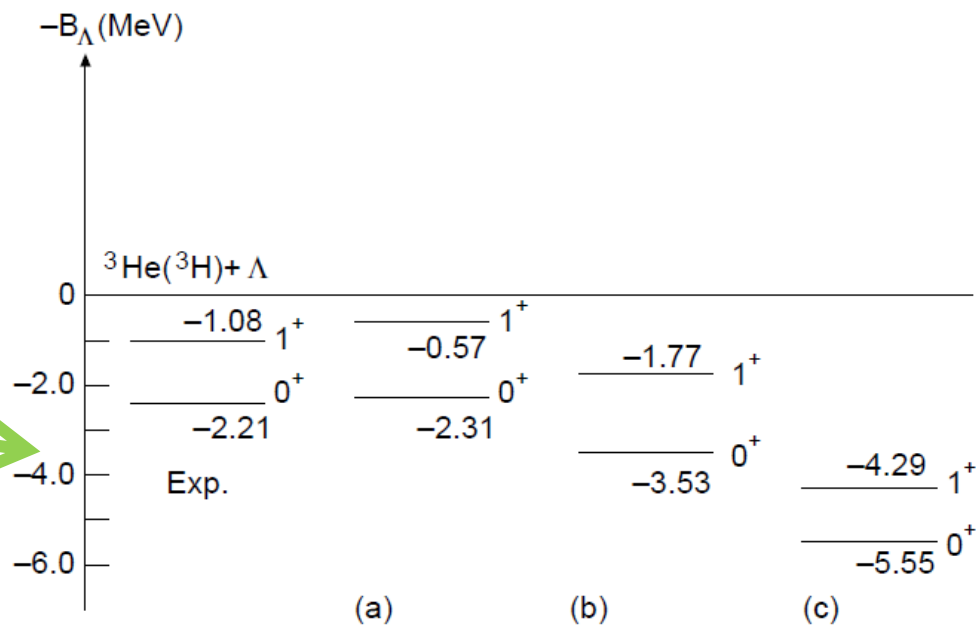


FIG. 3: Calculated Λ -separation energy for ${}^3_{\Lambda}\text{H}$ with (a) ${}^3V_{N\Lambda-N\Sigma}^T \times 1.00$, (b) ${}^3V_{N\Lambda-N\Sigma}^T \times 1.10$, and (c) ${}^3V_{N\Lambda-N\Sigma}^T \times$



$$\mathbf{H} = \Delta \mathbf{M} + \mathbf{T}_{\text{int}} + \mathbf{V}_{\text{NN}} + \mathbf{V}_{\text{3N}} + \mathbf{V}_{\text{YN}}$$

■ NN: chiral N³LO

Entem & Machleidt

Phys. Rev. C **68**, 041001(R) (2003)

$$\Lambda_{\text{NN}} = 500 \text{ MeV}$$

■ 3N: chiral N²LO

Navrátil

Few-Body Syst. **41**, 117 (2007)

$$\Lambda_{\text{3N}} = 500 \text{ MeV}$$

■ YN: chiral LO

Polinder, Haidenbauer & Meißner

Nucl. Phys. A **779**, 244 (2006)

$$\Lambda_{\text{YN}} = 600 \text{ MeV}, 700 \text{ MeV}$$

OR Jülich'04

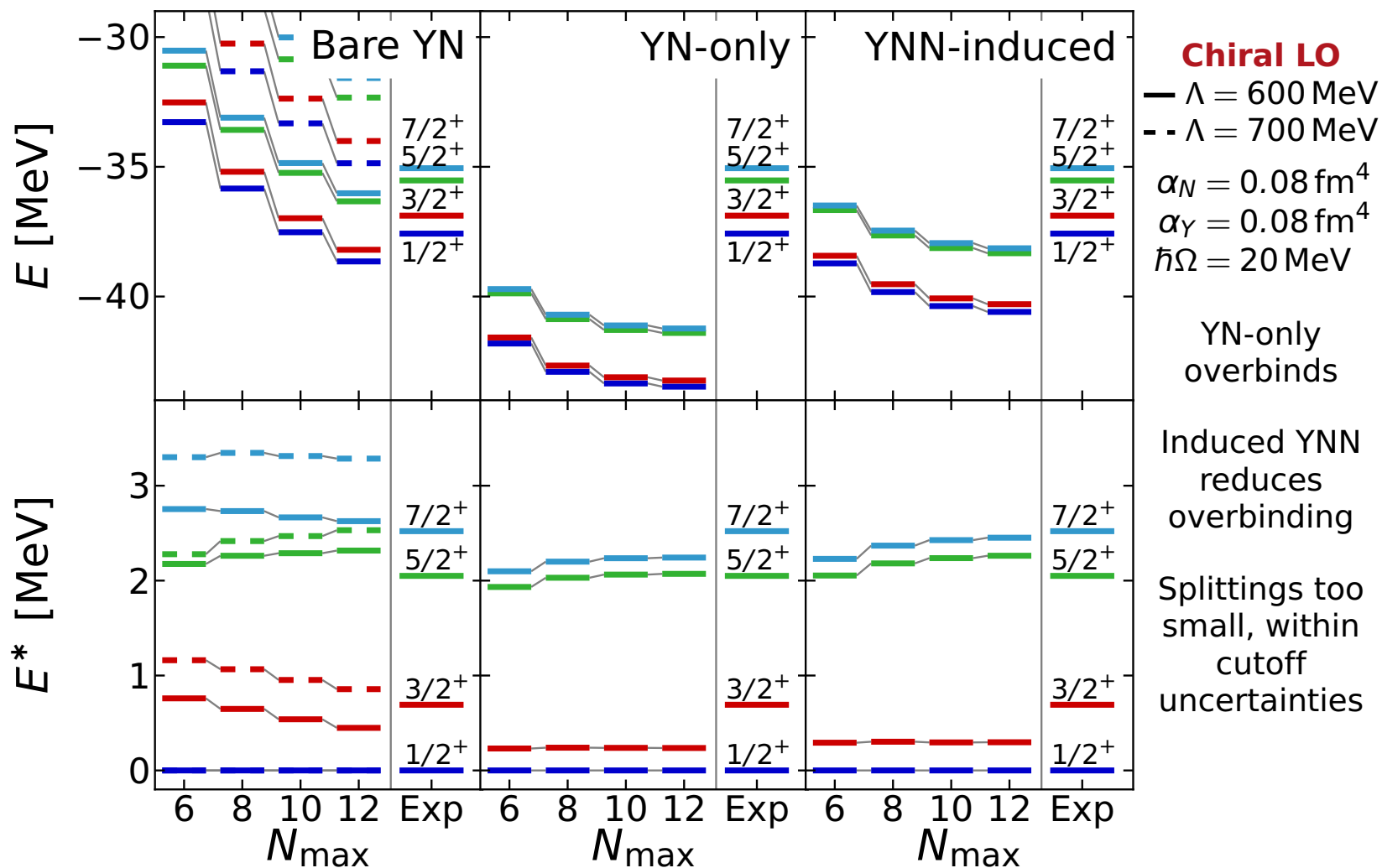
Haidenbauer & Meißner

Phys. Rev. C **72**, 044005 (2005)

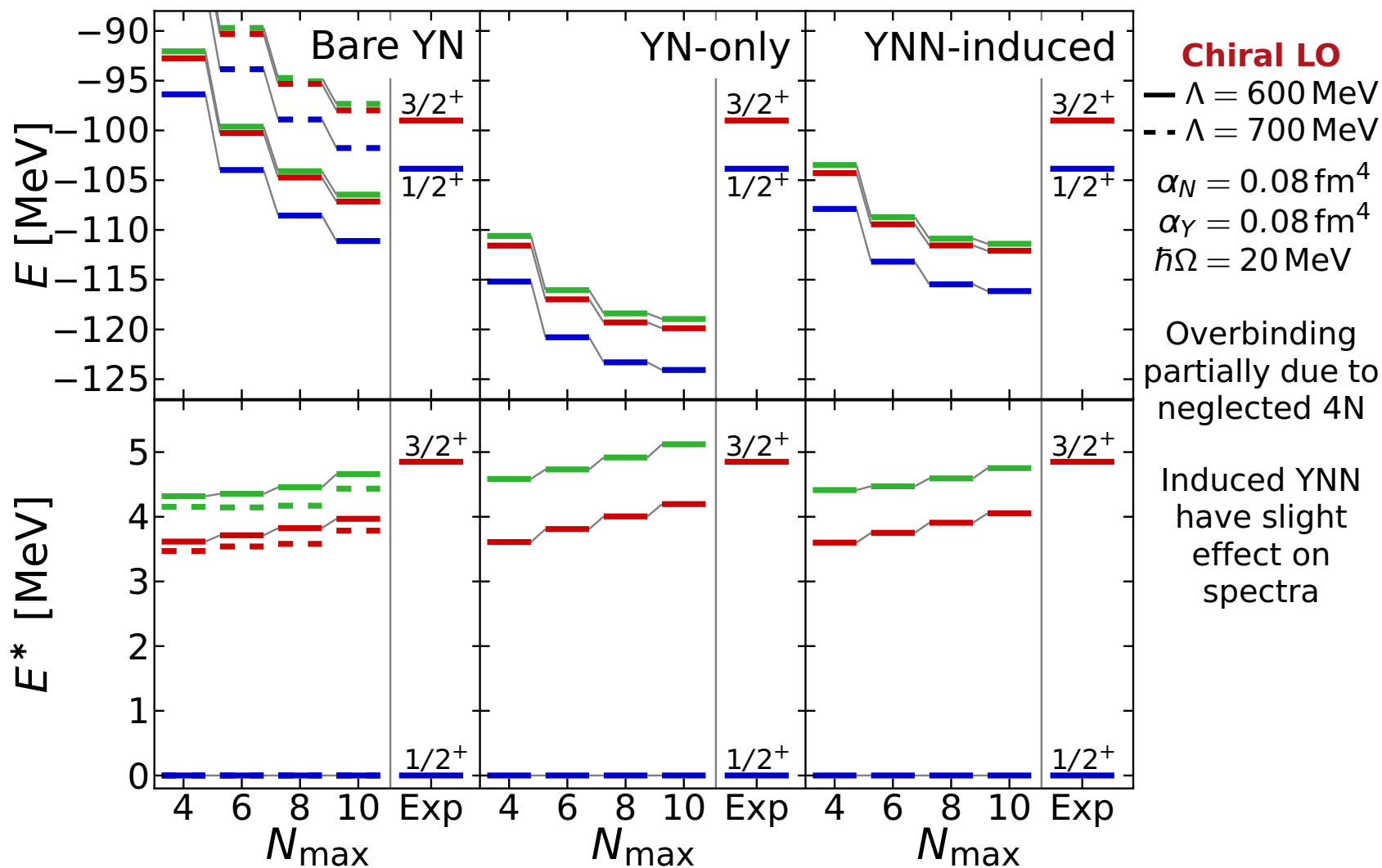
meson-exchange

NN+3N yields quantitative description of *p*-shell nuclei

${}^7_\Lambda\text{Li}$ — Effect of Induced YNN

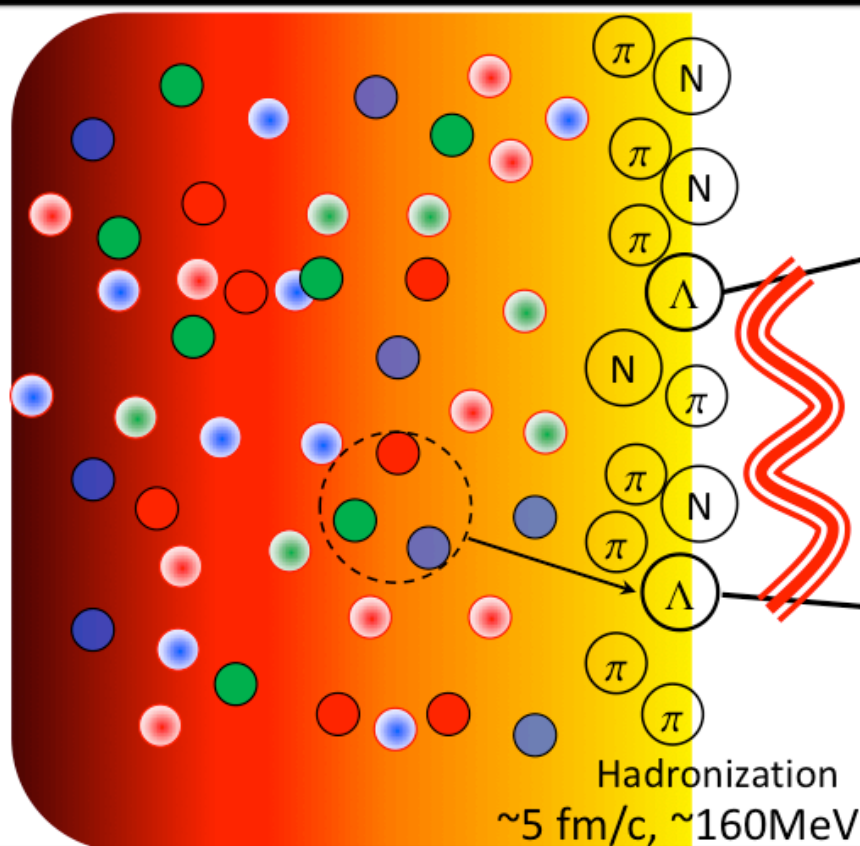


$^{13}_{\Lambda}\text{C}$ — Effect of Induced YNN



Hyperon-hyperon interaction

EMMI Workshop on anti-matter, hyper-matter and exotica production at the LHC


$$\frac{dN_{\Lambda}}{dy} \simeq 0.6 - 13 \quad \times 10^8 \text{ events}$$

(60-80%) (0-5%)

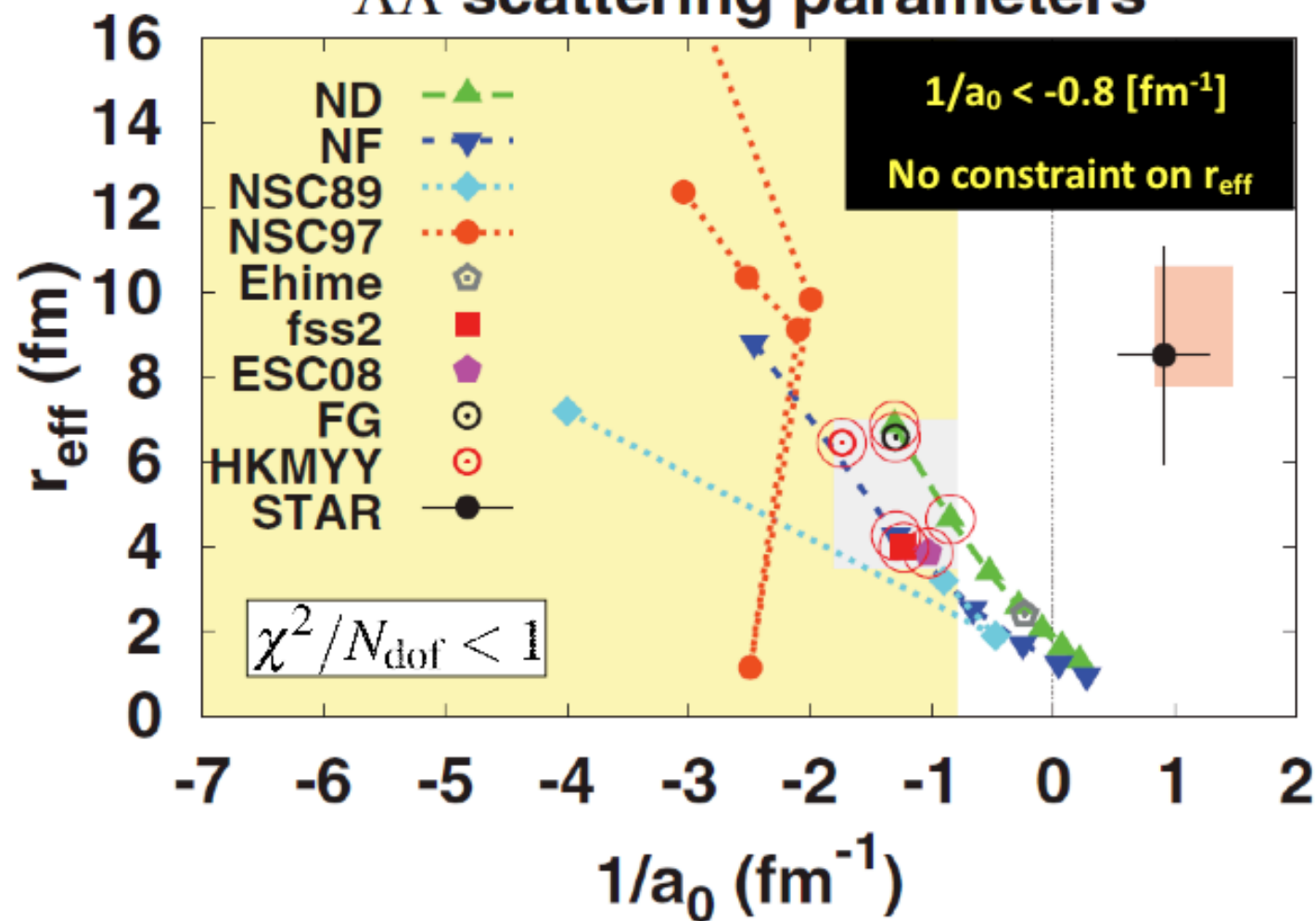
$$C(Q = p_1 - p_2) = \frac{N_2(p_1, p_2)}{N(p_1)N(p_2)}$$

 $V(r)$ p_1 p_2

time

Constraints on a_0 and r_{eff}

$\Lambda\Lambda$ scattering parameters



Interests of $S=-2$ multi-baryon system

H-dibaryon

- The flavor singlet state with $J=0$ predicted by R.L. Jaffe.
 - Strongly attractive color magnetic interaction.
 - No quark Pauli principle for flavor singlet state.

Double- Λ hypernucleus

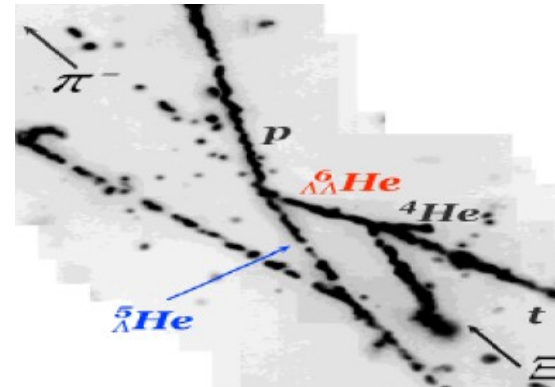
- Conclusions of the “NAGARA Event”

K.Nakazawa and KEK-E176 & E373 Collaborators

Λ -N attraction

Λ - Λ weak attraction

$$m_H \geq 2m_\Lambda - 6.9\text{MeV}$$

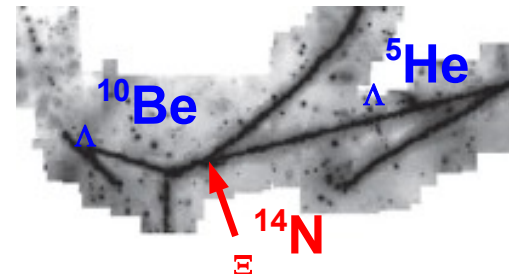


Ξ hypernucleus

- Conclusions of the “KISO Event”

K.Nakazawa and KEK-E373 Collaborators

Ξ -N attraction

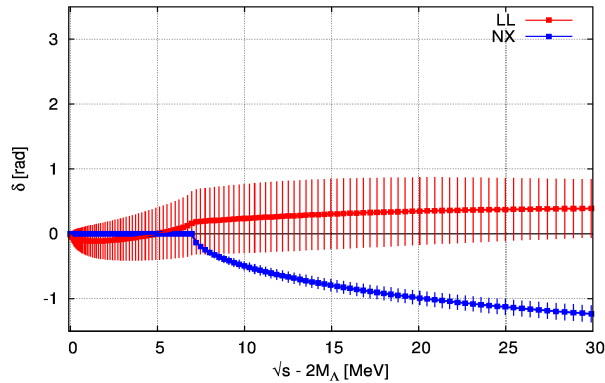


$\Lambda\Lambda$ and $N\Xi$ phase shifts

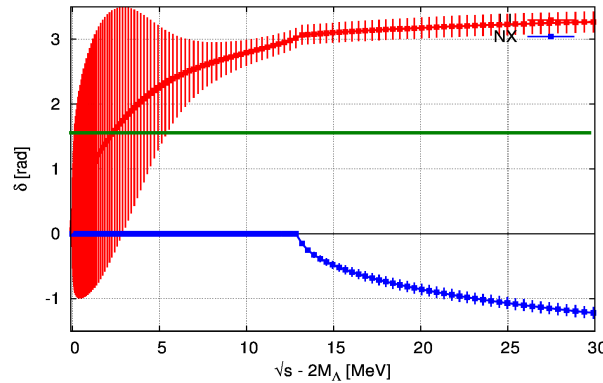
► $N_f = 2+1$ full QCD with $L = 2.9\text{fm}$

Preliminary!

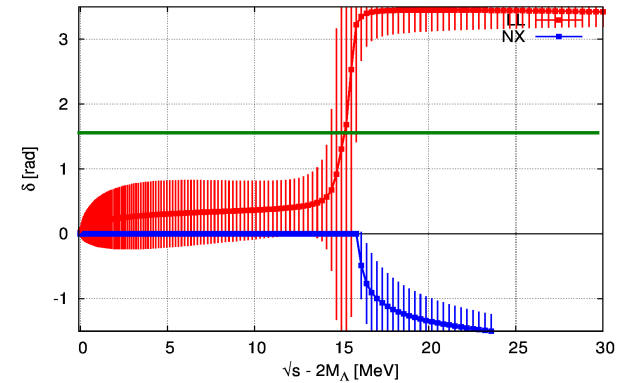
$m_\pi = 700\text{ MeV}$



$m_\pi = 570\text{ MeV}$

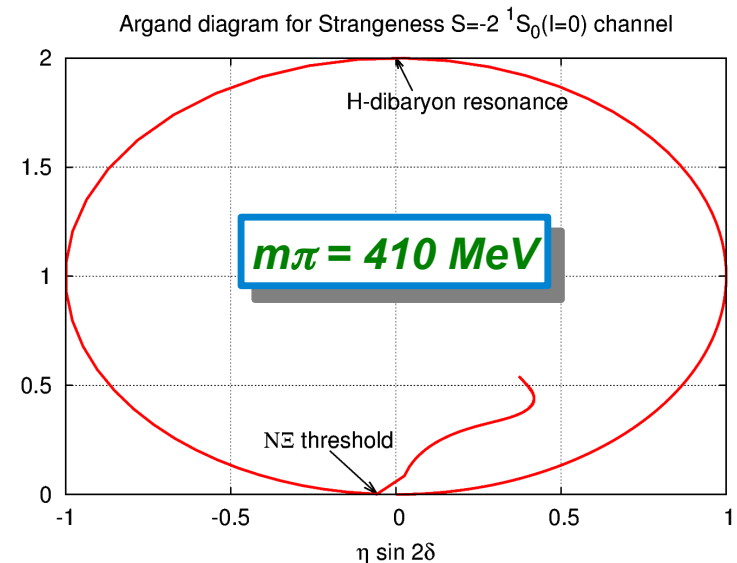


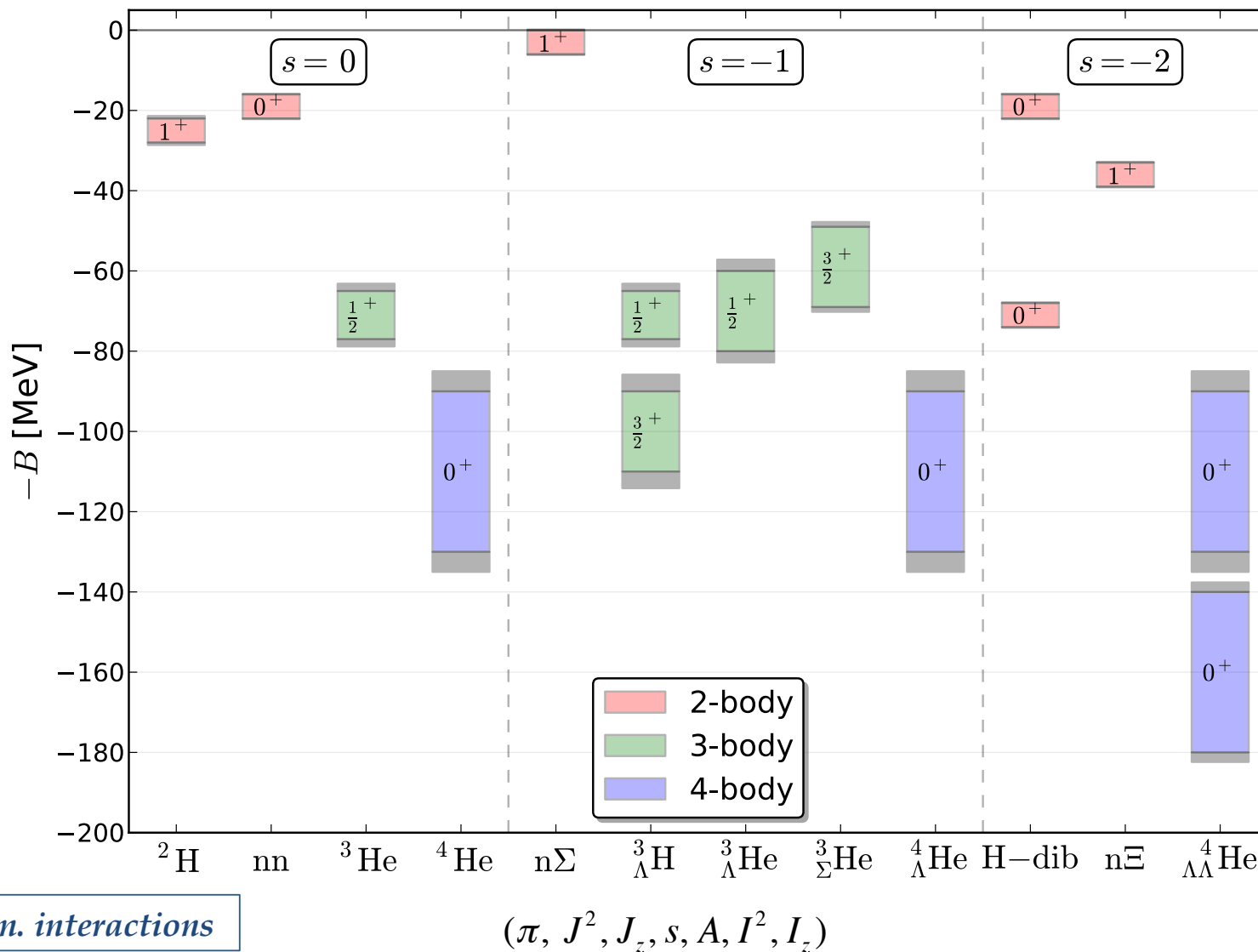
$m_\pi = 410\text{ MeV}$



- $m_\pi = 700\text{ MeV}$: bound state
- $m_\pi = 570\text{ MeV}$: resonance near $\Lambda\Lambda$ threshold
- $m_\pi = 410\text{ MeV}$: resonance near $N\Xi$ threshold

H-dibaryon is unlikely bound state

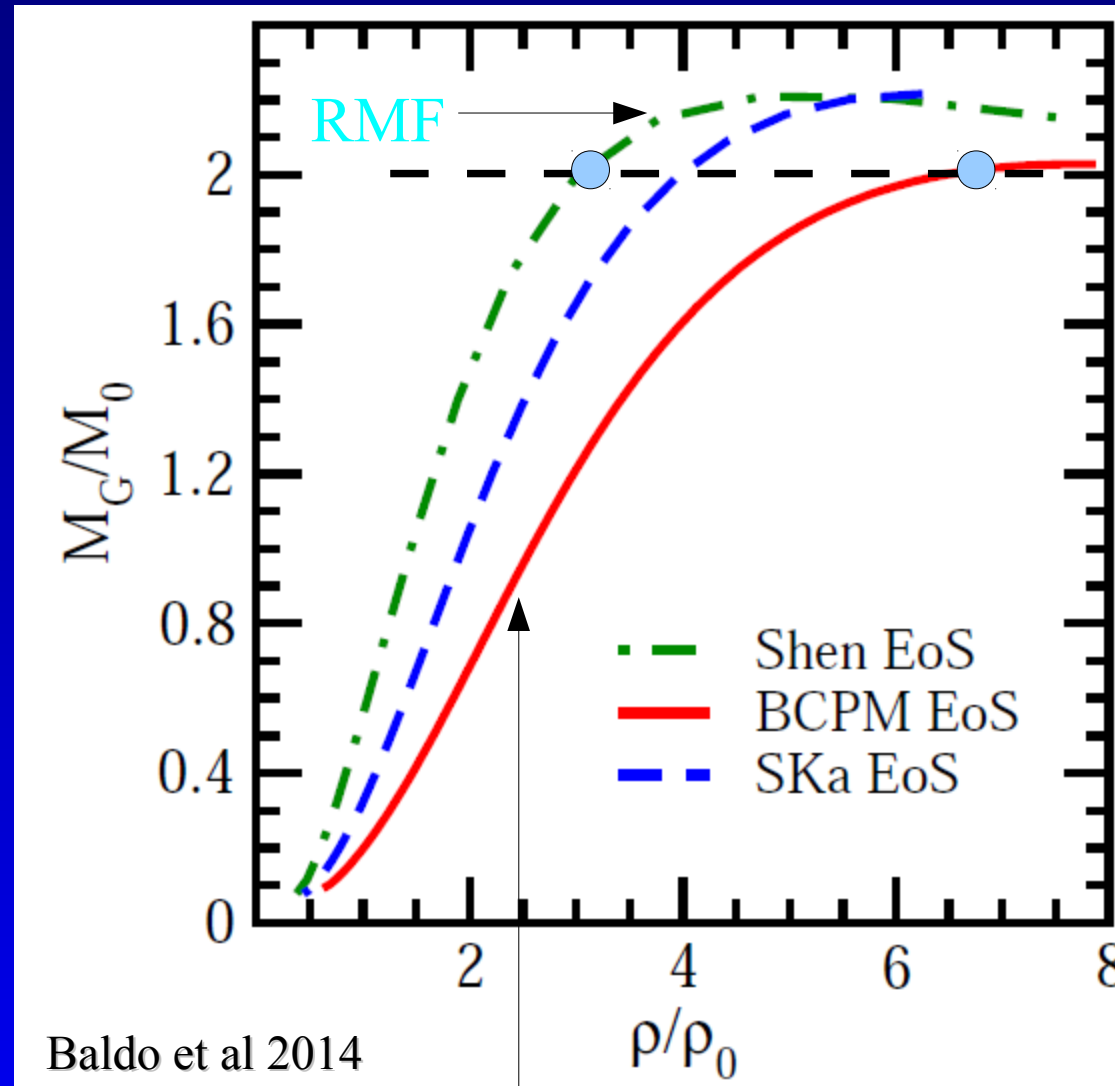




What does a $2M_{\text{sun}}$ star mean?

“Standard” neutron stars, just nucleons and electrons.

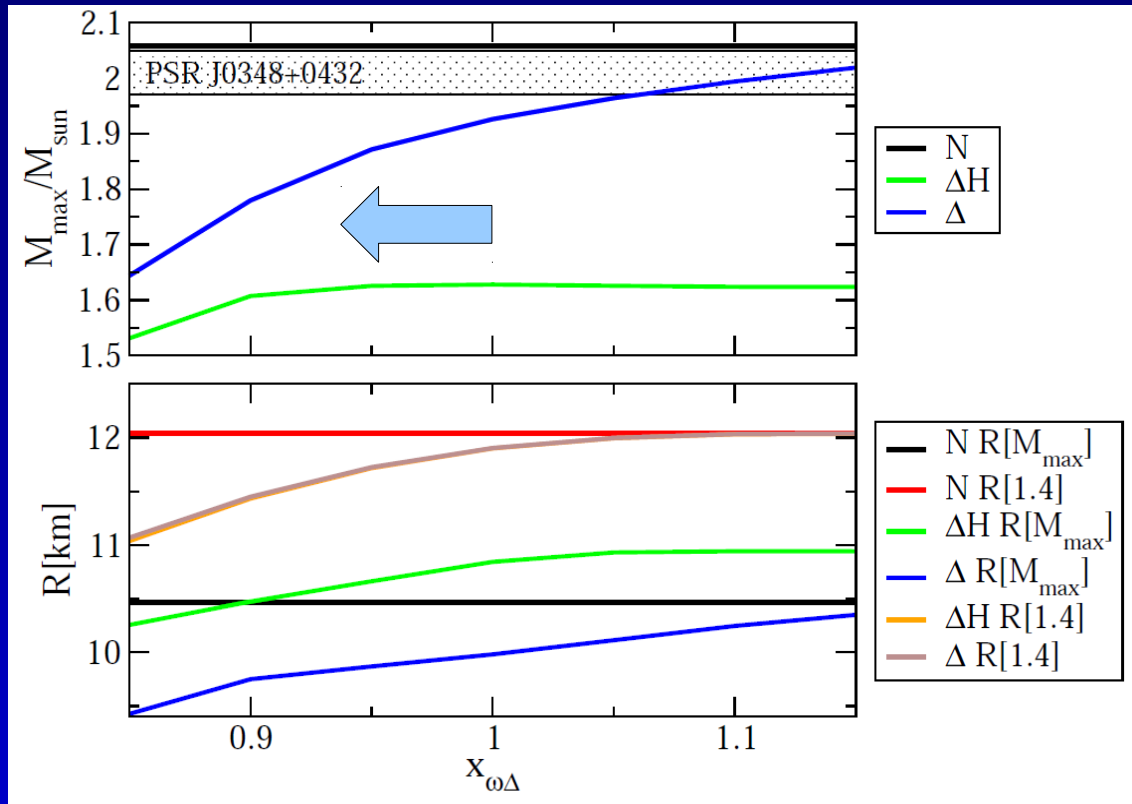
Central baryon densities of a $2M_{\text{sun}}$ star 3-7 times nuclear saturation density. Are there really just nucleons? Hyperons & Δ ?



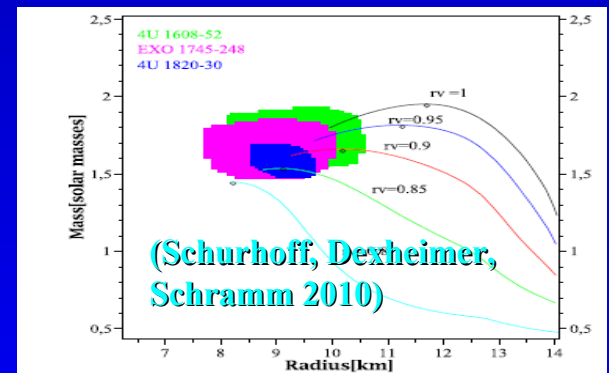
Baldo et al 2014

Microscopic calculation: nucleon nucleon potential and three body forces

Maximum mass and radii: the maximum mass is significantly smaller than the measured ones. Also, very compact stellar configurations are possible.



See also:



Punchline2: beside the “hyperon puzzle” is there also a “delta isobars puzzle”?

-) **New masses and radii measurements challenge nuclear physics: tension between high mass and small radii. 2.4 Msun candidates already exist.**
-) **Hyperons and delta puzzles**
-) **NICER mission, with a precision of 1km in radii measurements, could hopefully solve the problem**
-) **Possible existence of two families of compact stars (high mass – quark stars, low mass – hadronic stars). Rich phenomenology: frequency and mass distributions, explosive events, strangelets**

Hyperon puzzle in Neutron Stars

Massive $2M_{\odot}$ neutron stars:

2010 PSR J1614-2230 (1.97 ± 0.04) M_{\odot}

2013 PSR J0348-0432 (2.01 ± 0.04) M_{\odot}

- **Problem:** Softening EoS by hyperon mixing (e.g. Schulze et al)
- **Conclusion:** [Yamamoto et al PRC88 (2014), EPJA (2015)]

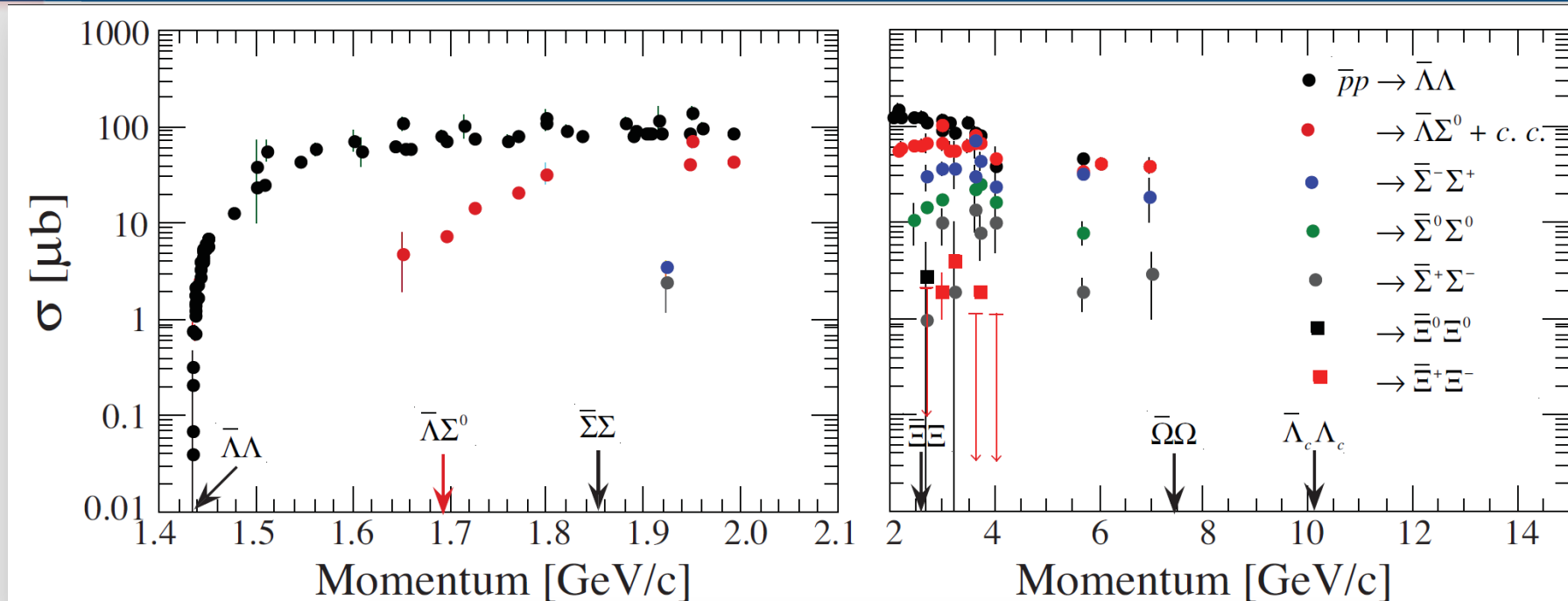
The puzzle can be solved by a
Universal Three-Baryon Repulsion
on the basis of heavy-ion data

Collaborators: Y. Yamamoto, T. Furumoto, N. Yasutake,
H.-J. Schulze, and M.M. Nagels.

1. High-quality Simultaneous Fit/Description $NN \oplus YN$,
OBE, TME, MPE meson-exchange dynamics.
 $SU_f(3)$ -symmetry, (Non-linear) chiral-symmetry.
2. Scalar-meson nonet structure \Leftrightarrow Nagara $\Delta B_{\Lambda\Lambda}$ values.
3. **NO S=-1 Bound-States, NO $\Lambda\Lambda$ -Bound-State,**
4. Prediction: $D_{\Xi N} = \Xi N(I = 1, {}^3S_1)$ **B.S.!**, $D_{\Xi\Xi} = \Xi\Xi(I = 1, {}^1S_0)$ **B.S. ??!**
5. Similar role **tensor-force** in 3S_1 NN-, $\Lambda/\Sigma N$ -, ΞN -, and $\Lambda/\Sigma\Xi$ -channels.

G-matrix and EOS of the ESC YN/YY-interactions:

- a. ESC08: Excellent G-matrix predictions for the $U_{\Lambda}, U_{\Sigma}, U_{\Xi}$ well-depth's,
 ΛN spin-spin and spin-orbit, and Nagara-event okay.
- b. Neutron Star mass $M/M_{\odot} = 1.44 - 2.10 \Leftrightarrow$ Universal Multi-Pomeron
Repulsion, including Λ, Σ, Ξ



PANDA

Production Rates (1-2 (fb)⁻¹/y)

<u>Final State</u>	<u>cross section</u>	<u># reconstr. events/y</u>
Meson resonance + anything	100 μb	10^{10}
$\Lambda\bar{\Lambda}$	50 μb	10^{10}
$\Xi\bar{\Xi} (\rightarrow \Lambda\Lambda A)$	2 μb	$10^8 (10^5)$
$D\bar{D}$	250nb	10^7
$J/\psi (\rightarrow e^+e^-, \mu^+\mu^-)$	630nb	10^9
$\chi_2 (\rightarrow J/\psi + \gamma)$	3.7nb	10^7
$\Lambda_c\bar{\Lambda}_c$	20nb	10^7
$\Omega_c\bar{\Omega}_c$	0.1nb	10^5

Perspectives: Hypernuclei

J-PARC

MAINZ

CERN

BNL

JLab

FAIR (?)