

# Unraveling stellar alchemy with NUSTAR

## From ideas to in-kind

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*FAIR*

2<sup>nd</sup> In-kind Contributions Workshop (IKCW 2015)  
GSI/FAIR Darmstadt, November 4, 2015



## What are the limits for existence of nuclei?

Where are the proton and neutron drip lines situated?

Where does the nuclear chart end?

## How does the nuclear force depend on varying proton-to-neutron ratios?

What is the isospin dependence of the spin-orbit force?

How does shell structure change far away from stability?

## How to explain collective phenomena from individual motion?

What are the phases, relevant degrees of freedom, and symmetries of the nuclear many-body system?

## How are complex nuclei built from their basic constituents?

What is the effective nucleon-nucleon interaction?

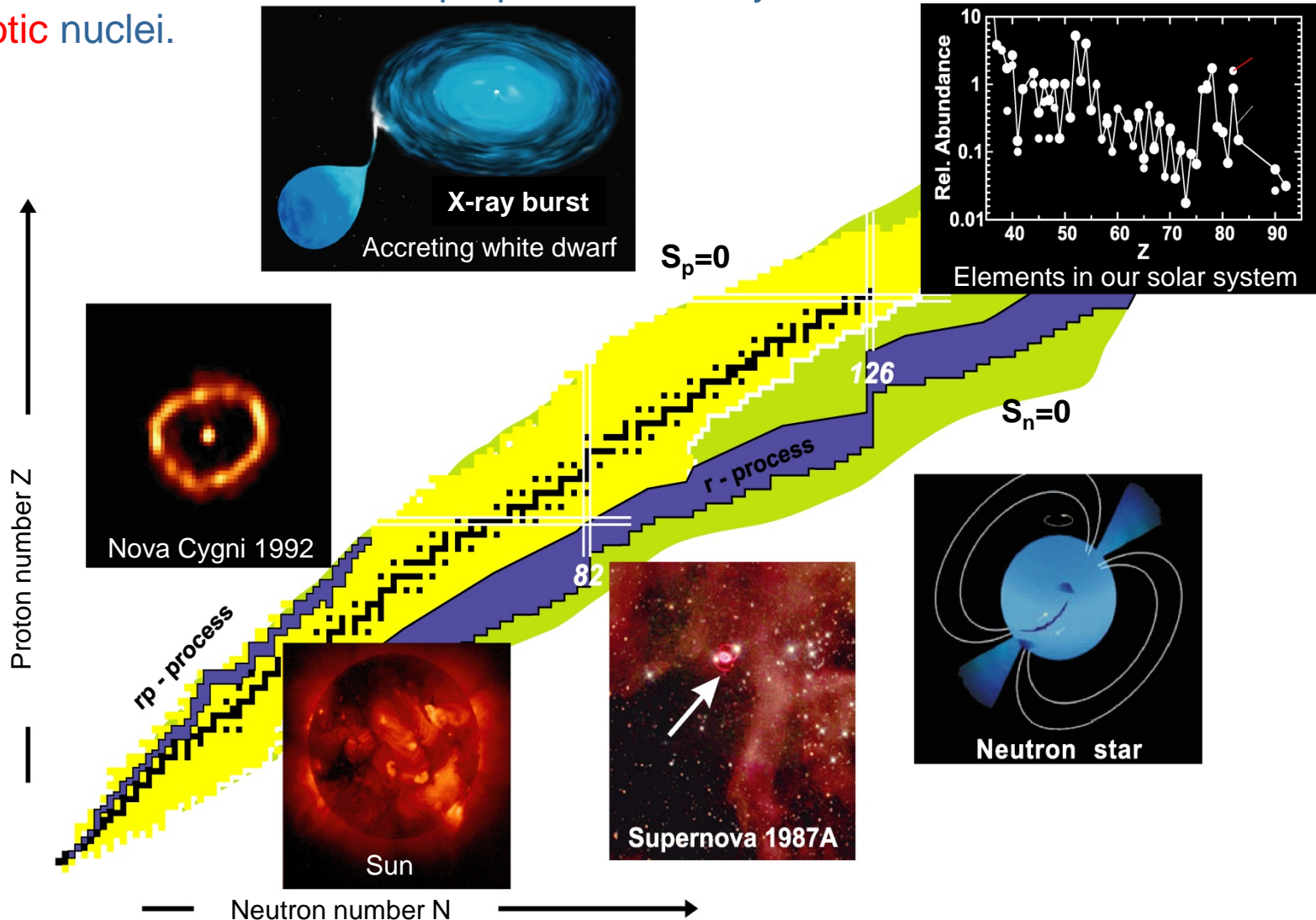
How does QCD constrain its parameters?

## Which are the nuclei relevant for astrophysical processes and what are their properties?

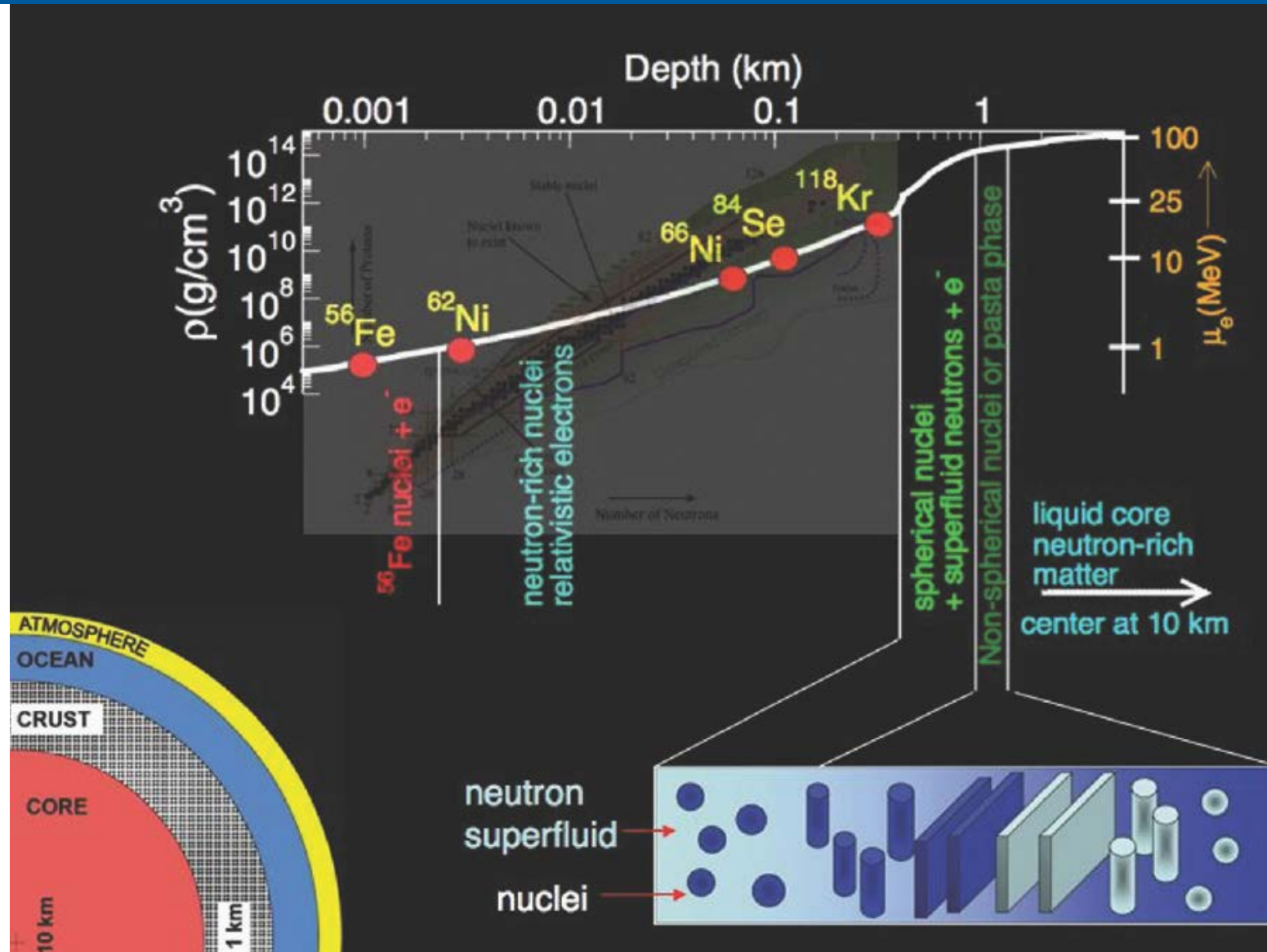
What is the origin of the heavy elements?

# Nuclear physics in the universe

For the understanding of nucleosynthesis and stellar dynamics we need to know properties of many **exotic** nuclei.



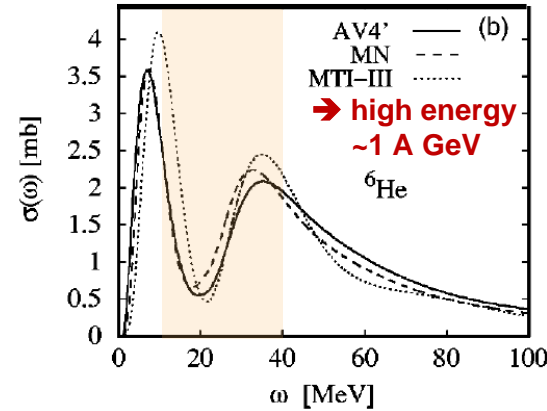
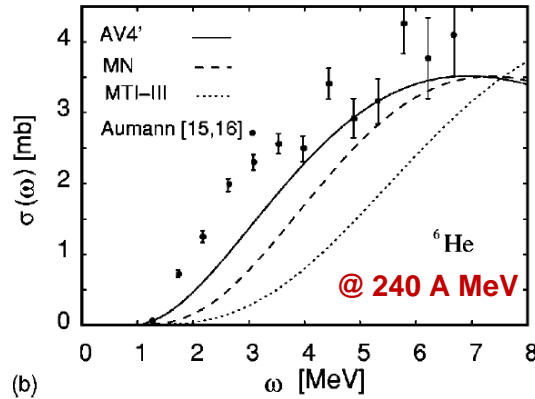
# Nuclei in neutron stars





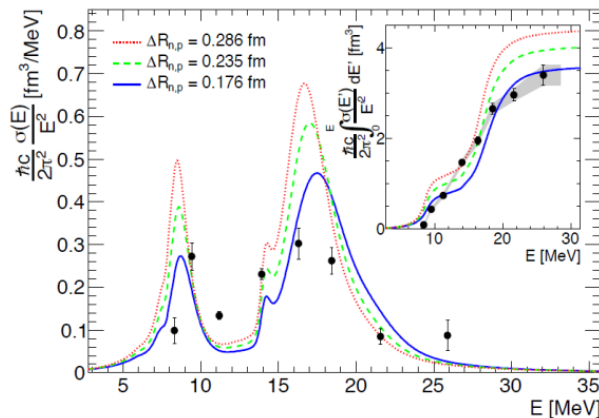
core vs. neutron skins & halos → density / asymmetry

to be measured  
with R<sup>3</sup>B



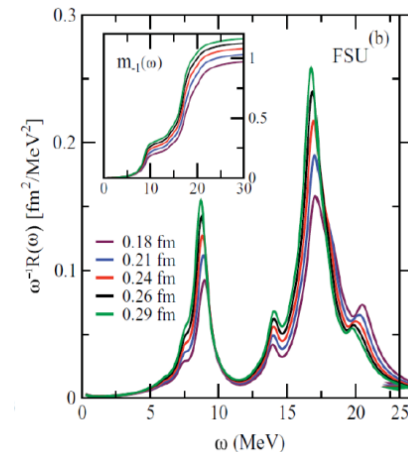
S. Bacca et al.  
PRL **89** (2002) 052502  
PRC **69** (2004) 057001

access to EoS (e.g. neutron star) & low lying E1 strength (r-process)



D. Rossi et al.  
PRL **111** (2013) 242503

skin thickness <sup>68</sup>Ni  
0.175(21) fm



Pb chain &  
N=126 isotones

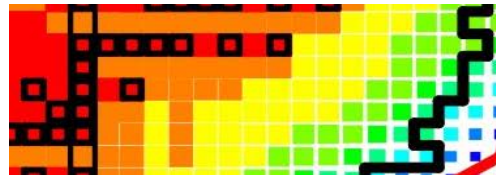
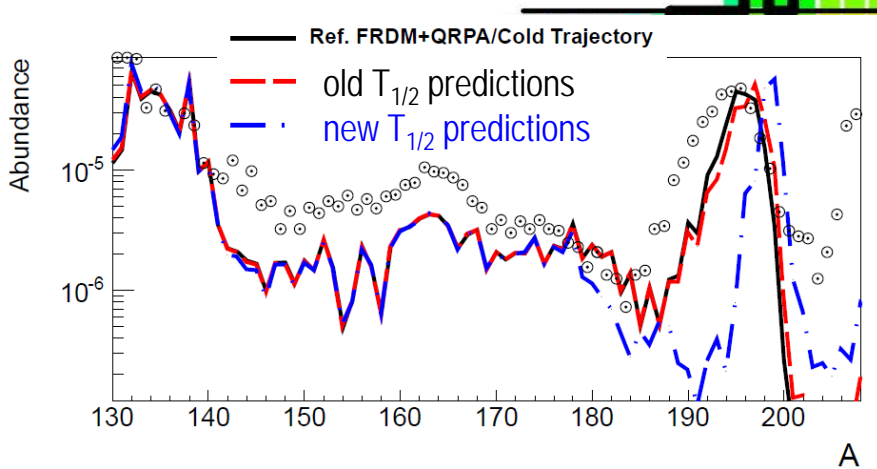
~1 A GeV →  
bare ions  
Fragment  
identification

$$\alpha_D = \frac{\hbar c}{2\pi^2} \int_0^\infty \frac{\sigma(E)}{E^2} dE$$

J. Piekarewicz, PRC **83** (2011) 034319

# The N=126 physics case

Previous GSI measurements  
contradict earlier lifetime predictions!  
→ Mass abundance not understood!

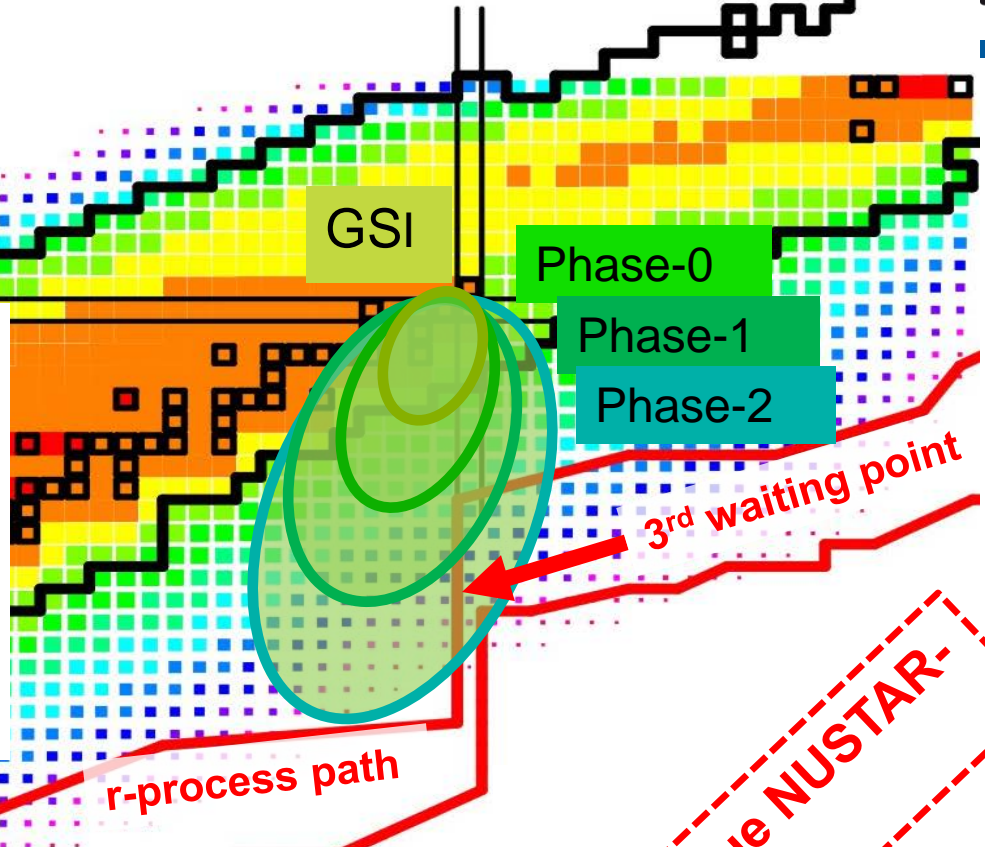


Mass abundance depends  
on the detailed structure  
of N=126 nuclei around the  
3<sup>rd</sup> r-process waiting point

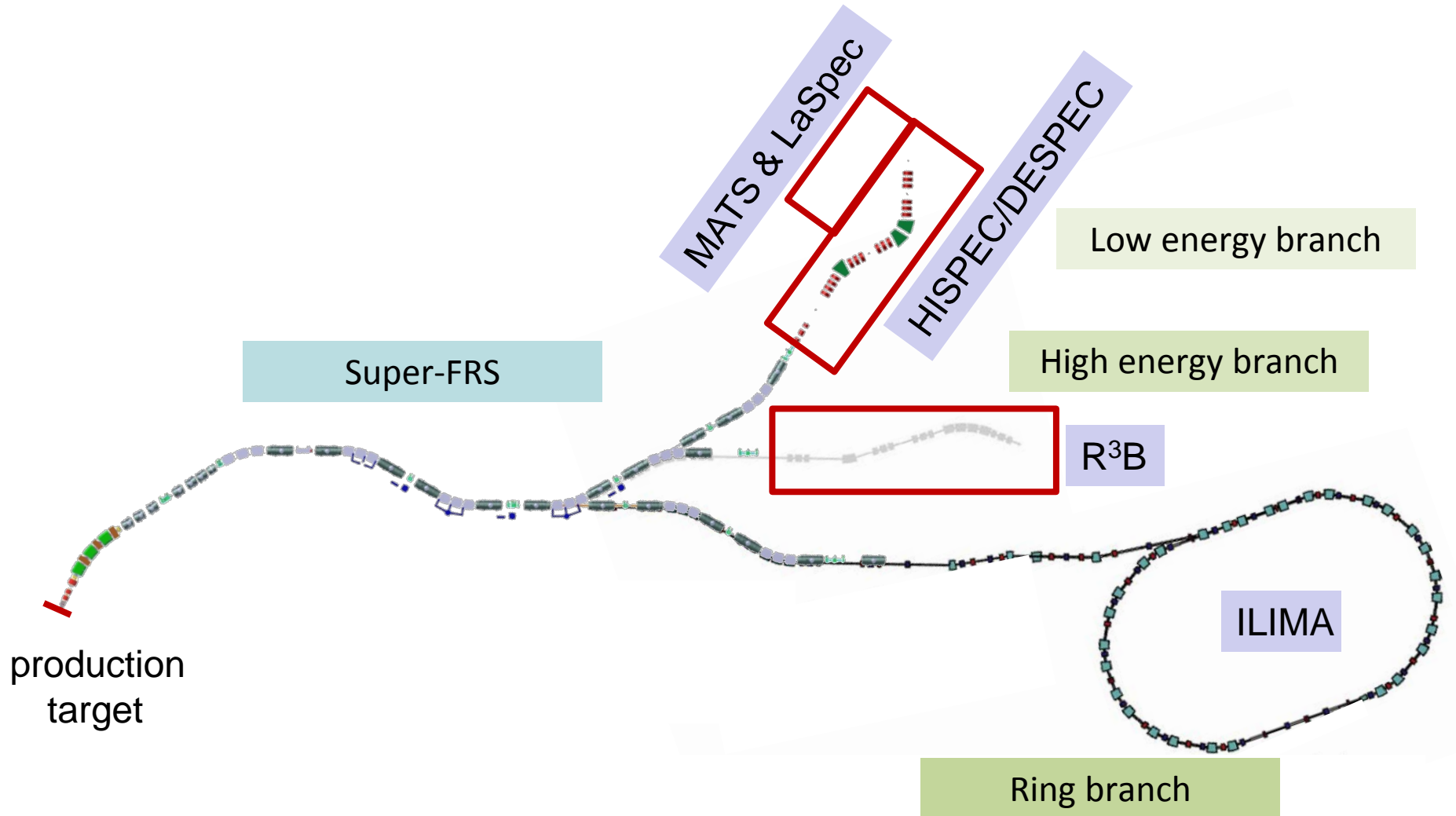
NUSTAR aims to measure:

- masses
- $\beta$ -lifetimes
- neutron-branchings
- strength distributions
- level structure

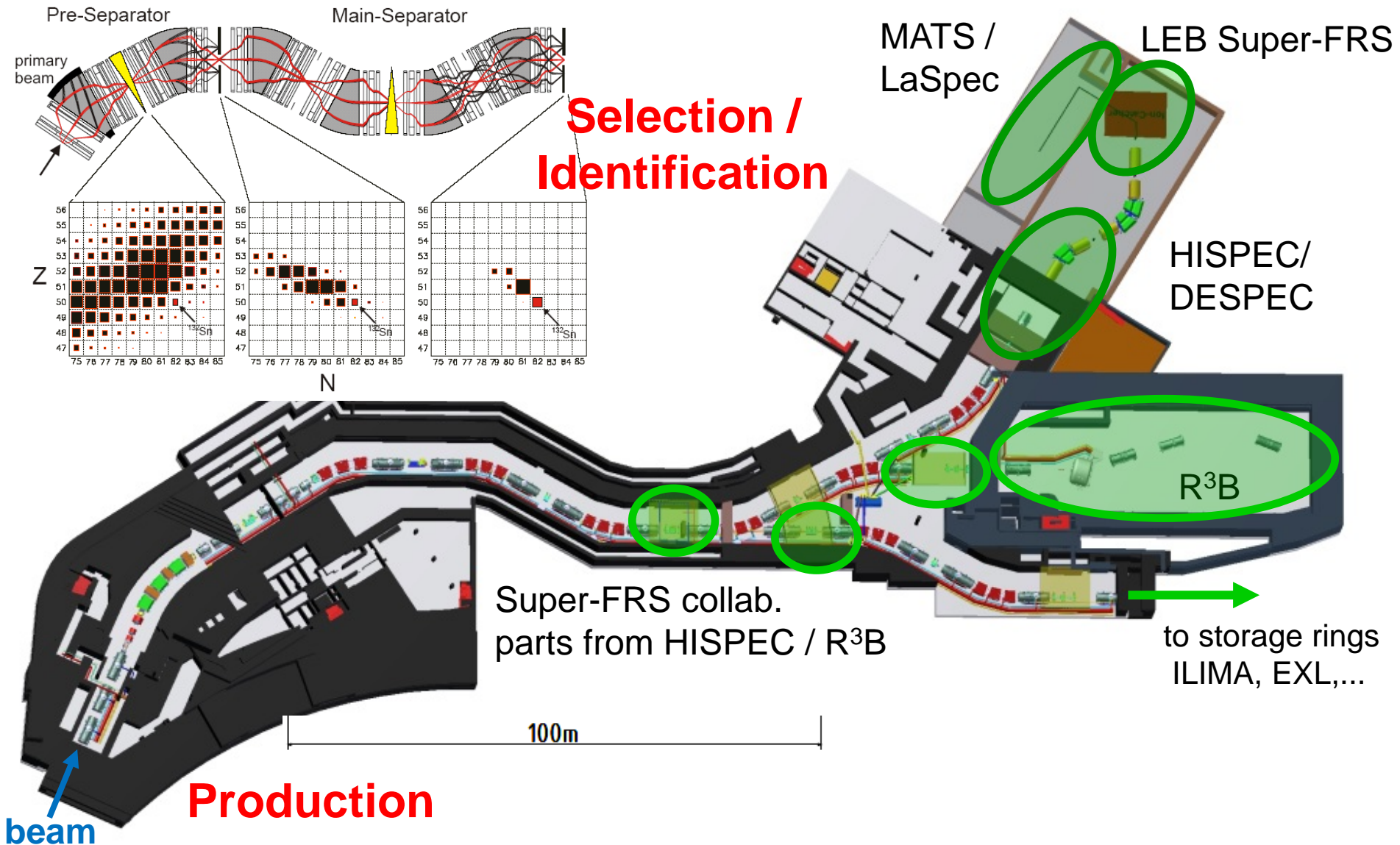
Important unique NUSTAR-  
LEB experiment



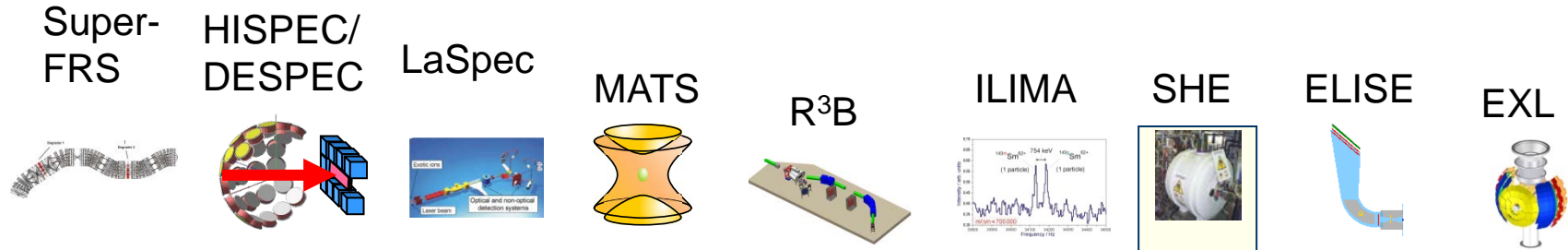








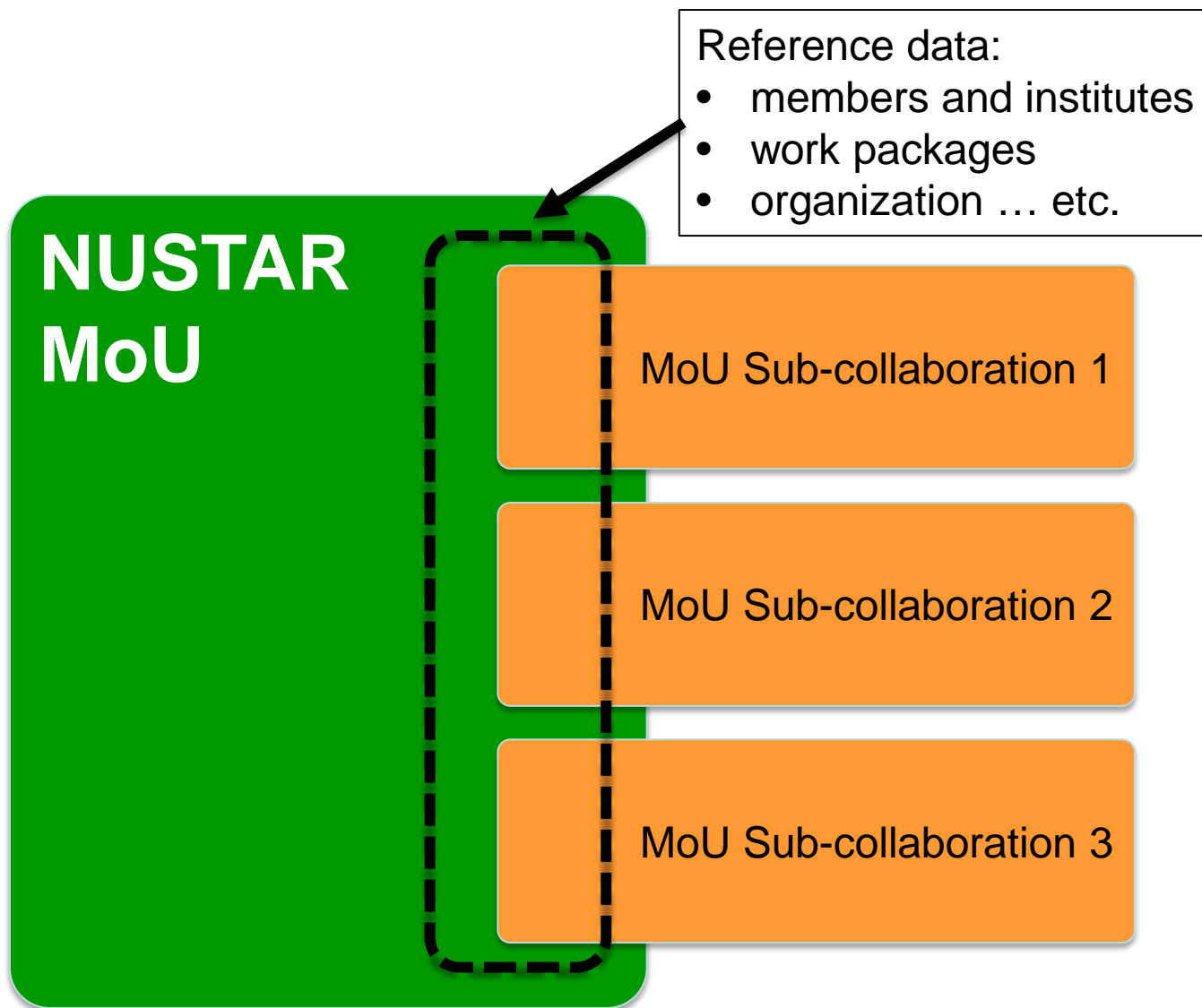
# Complementarity of NUSTAR experiments



	Super-FRS	HISPEC/DESPEC	LASPEC	MATS	R3B	ILIMA	SHE	ELISE	EXL
<b>Masses</b>		Q-values, isomers		dressed ions, highest precision	unbound nuclei	bare ions, mapping study	precision mass of SHEs		
<b>Half-lives</b>	ps...ns-range	dressed ions, $\mu$ s...s			resonance width, decay up to 100ns	bare ions, ms...years	$\mu$ s...days		
<b>Matter radii</b>	interaction x-section				interaction x-section				matter density distribution
<b>Charge radii</b>	charge-changing cross sections		mean square radii		charge-changing cross sections			charge density distribution	
<b>Single-particle structure</b>	high resolution, angular momentum	high-resolution particle and $\gamma$ -ray spectroscopy	magnetic moments, nucl. spins	evolution of shell str., pairing int., valence nucl.	quasi-free knockout, short-range and tensor	evolution of shell closures, pairing corr.	shell structure of SHEs		low momentum transfers
<b>Collective behavior</b>		electromag. transitions	quadrupole moments	halo structure	dipole response	changes in deformation		electromag. transitions	monopole resonance
<b>EoS</b>					polarizability, neutron skin			neutron skin $\rightarrow$	neutron skin, Compressibility
<b>Exotic Systems</b>	bound mesons, hypernuclei, nucleon res.								



- # NUSTAR Week 2015, Warsaw





Technical proposal



Technical design report (TDR)



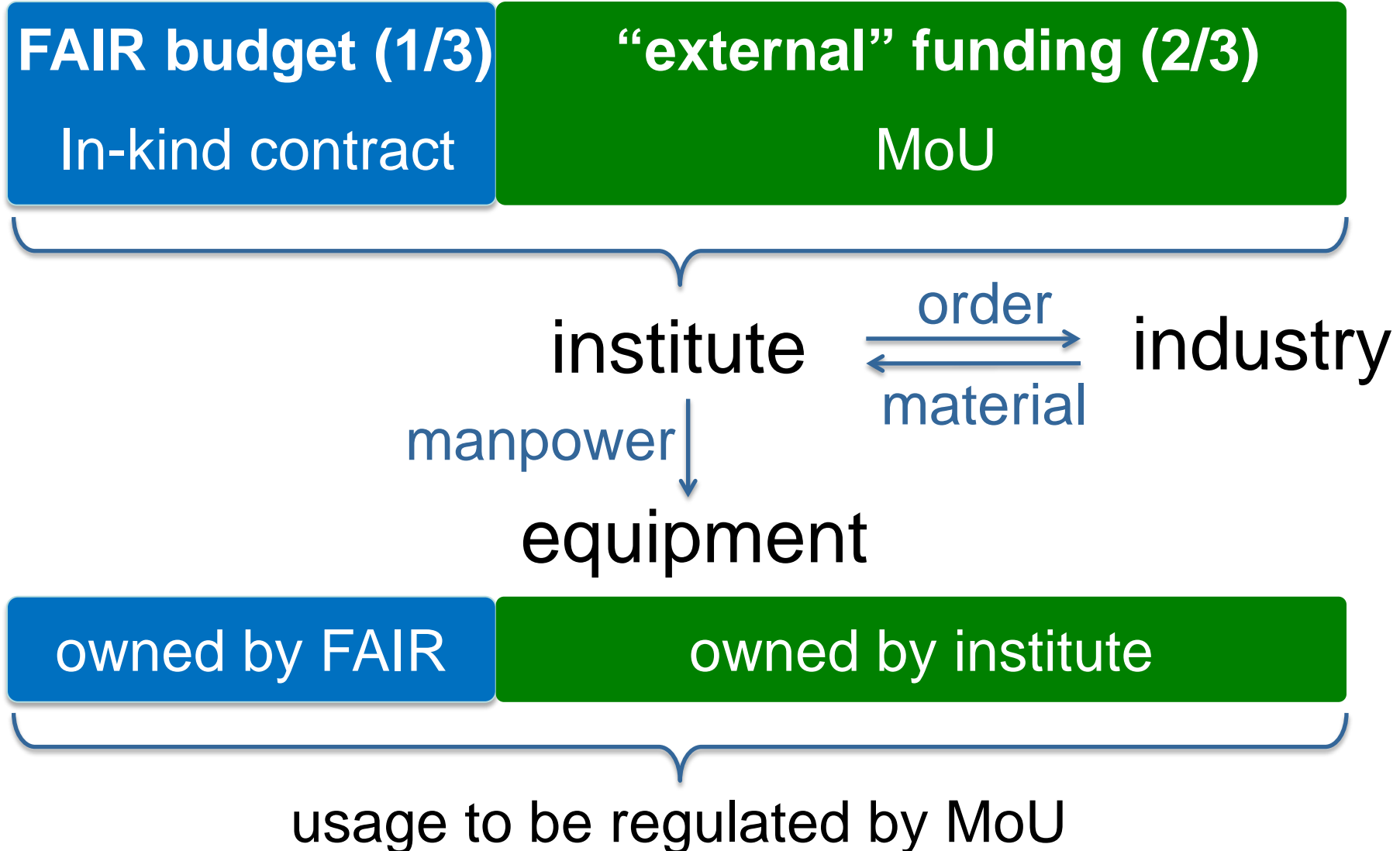
Identification of partner institutes

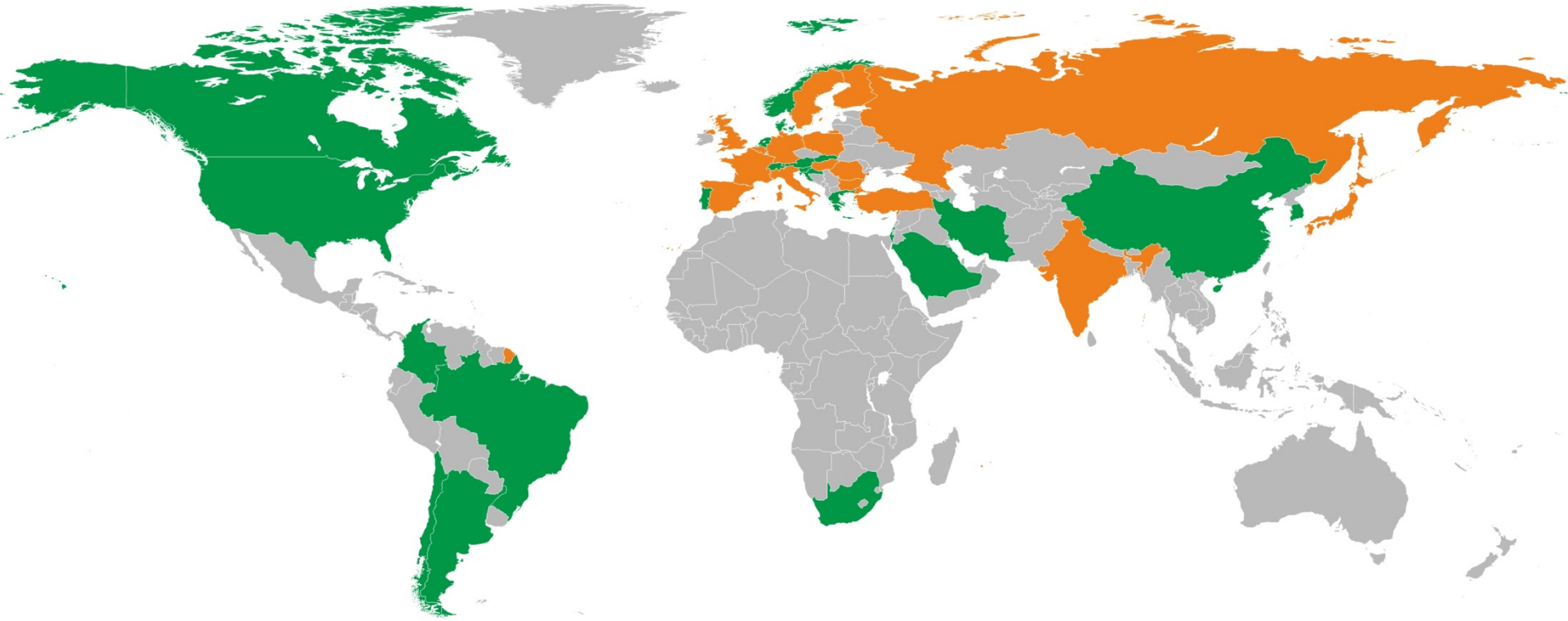


Funding requests



Delivery of in-kind contributions

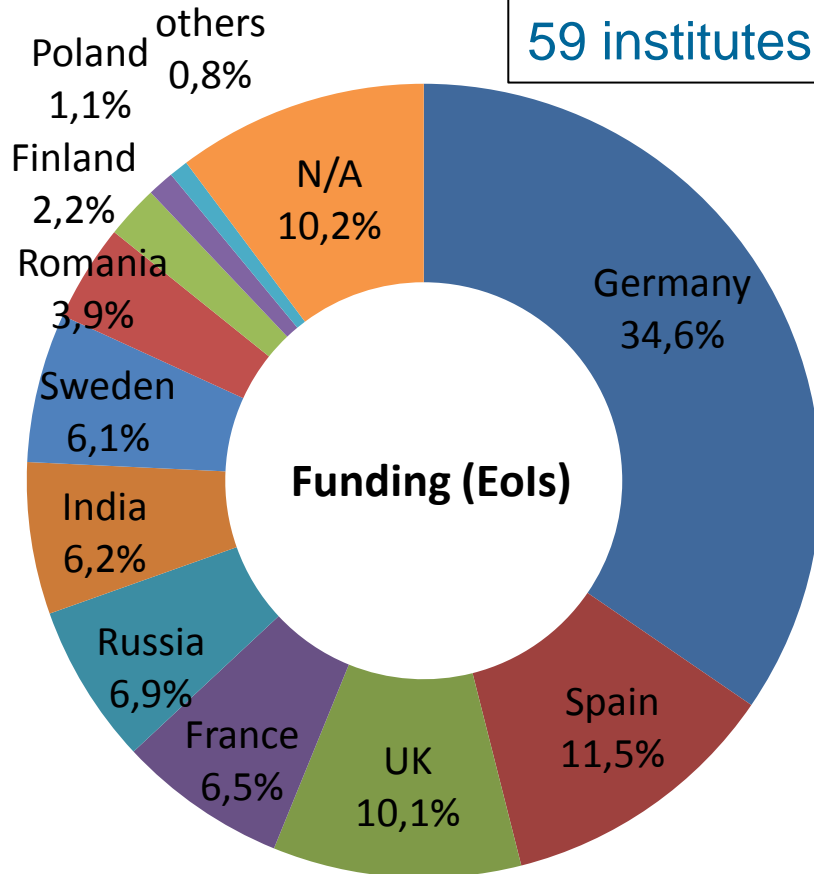




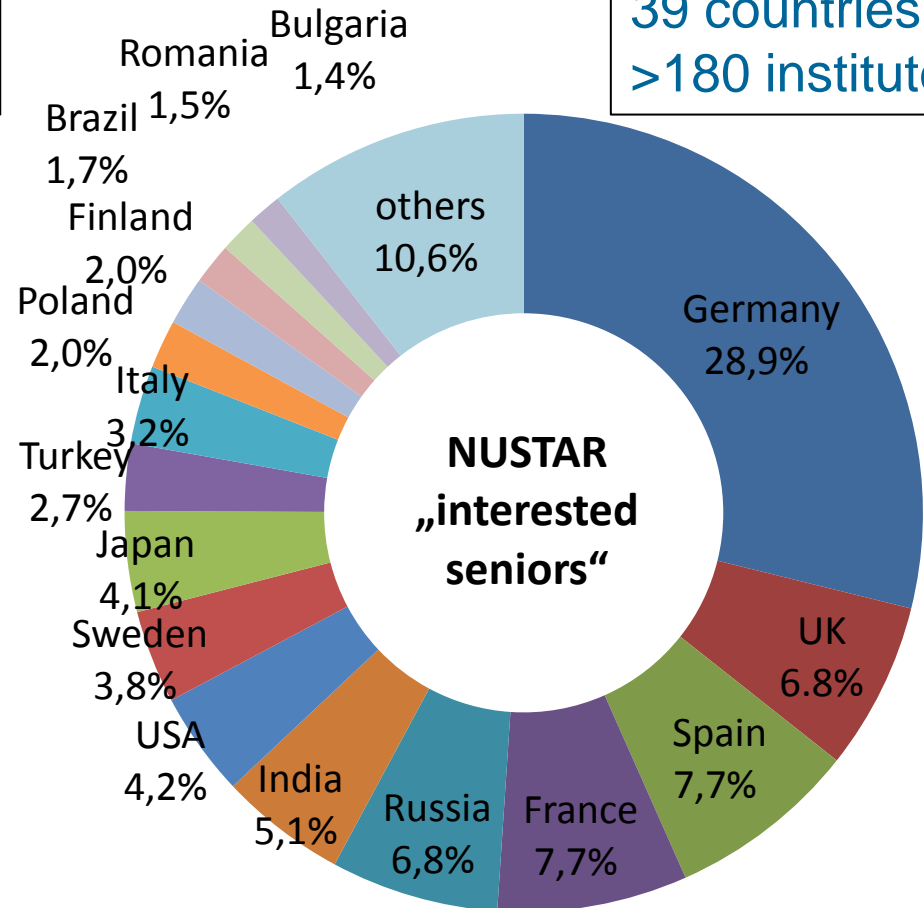
- Secured funding and expression of interest in funding  
(status: July 2015)  
16 countries (incl. 9 FAIR partner countries)

# Comparison: funding vs. senior scientists

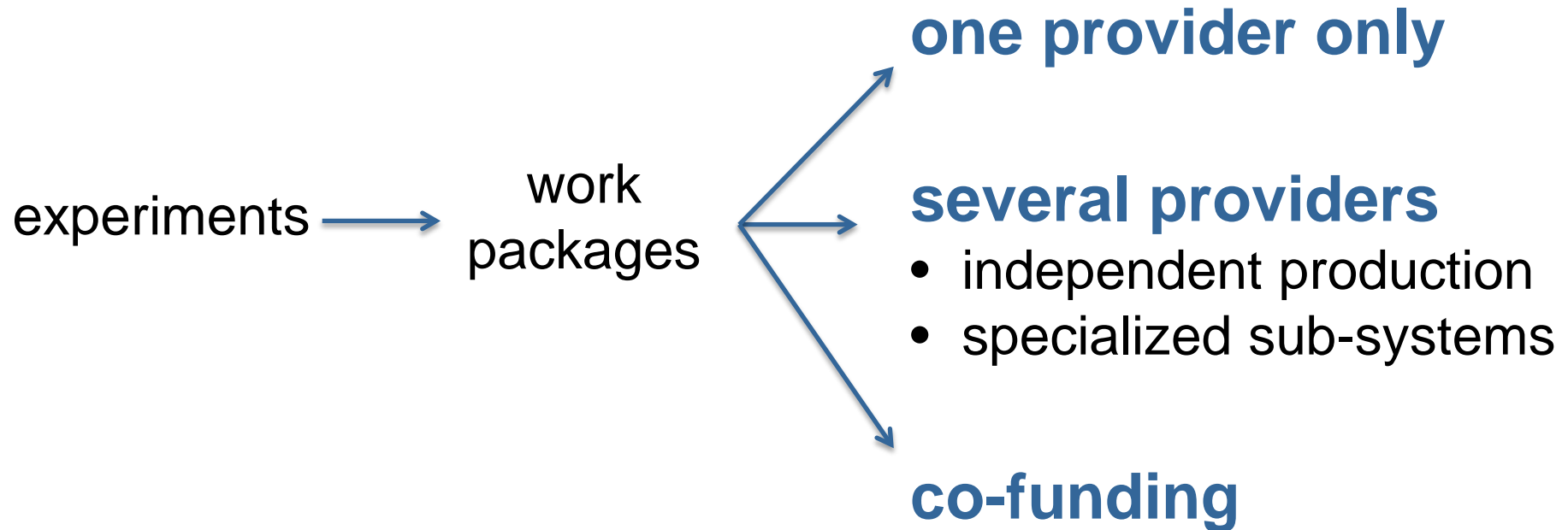
16 countries  
59 institutes



39 countries  
>180 institutes







## Modularized Start Version (MSV)

1.2.1	LEB Super-FRS
1.2.2	HISPEC/DESPEC
1.2.3	MATS
1.2.4	LaSpec
1.2.5	R <sup>3</sup> B
1.2.6	ILIMA

## Extending or beyond MSV

1.2.8	ELISe
1.2.9	EXL

## New experiments

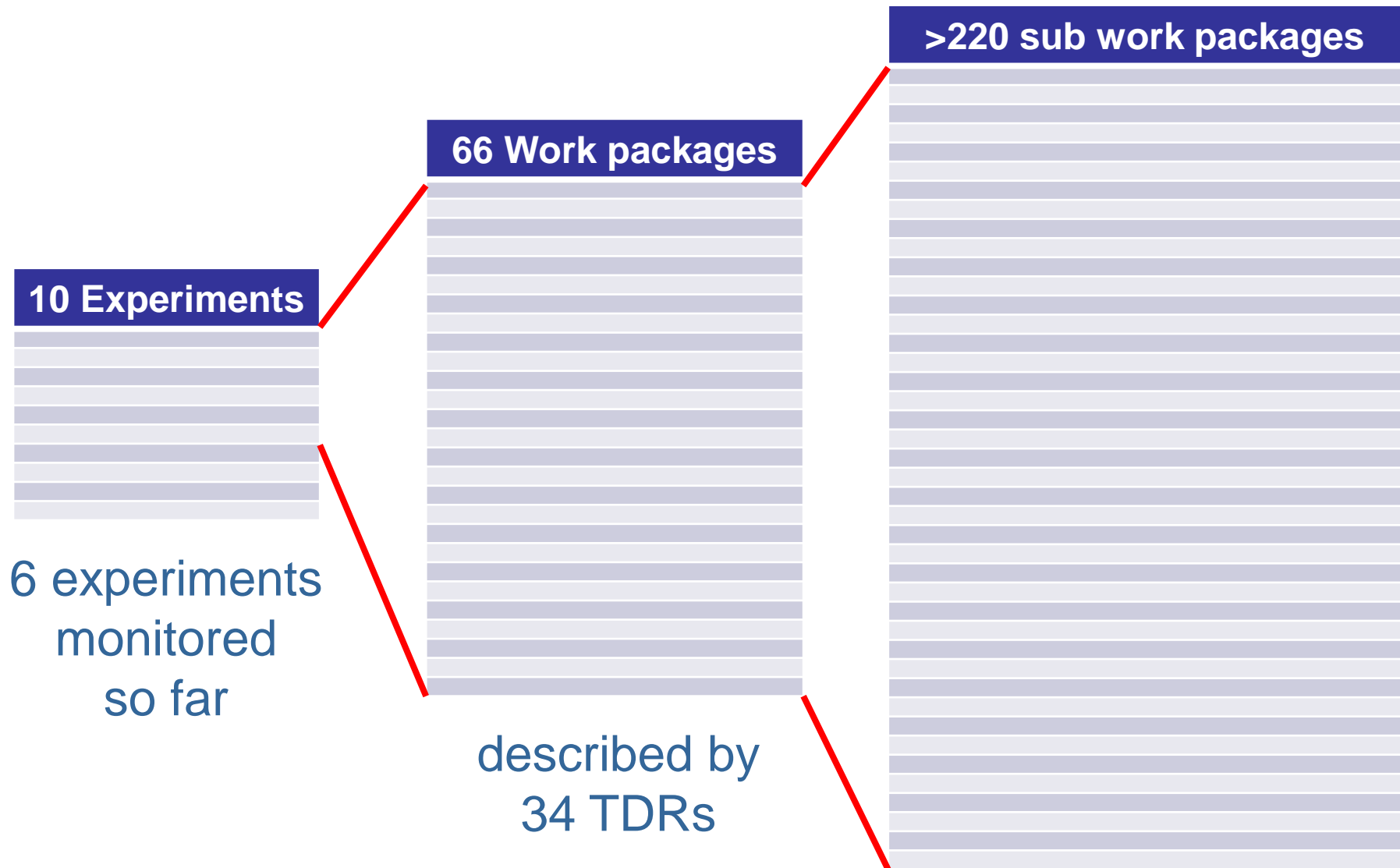
1.2.10	Super-FRS Experiments
1.2.11	SHE

## 1.2 NUSTAR

**initially NESR required** –  
alternative/intermediate „operation“  
within MSV under discussion

„operation“ within MSV planned

# Work packages of NUSTAR



# NUSTAR work packages (63 with TDR)



Described by 34 TDRs



## 1.2.5.1 R<sup>3</sup>B (stage I)

1.2.5.1.1  
Magnets

1.2.5.1.2 Detectors

1.2.5.1.3  
Vacuum

1.2.5.1.4  
DAQ

1.2.5.1.5  
Infra-  
structure

GLAD

Multiplet

Tracking

**CALIFA**

Si tracker  
(target  
recoils)

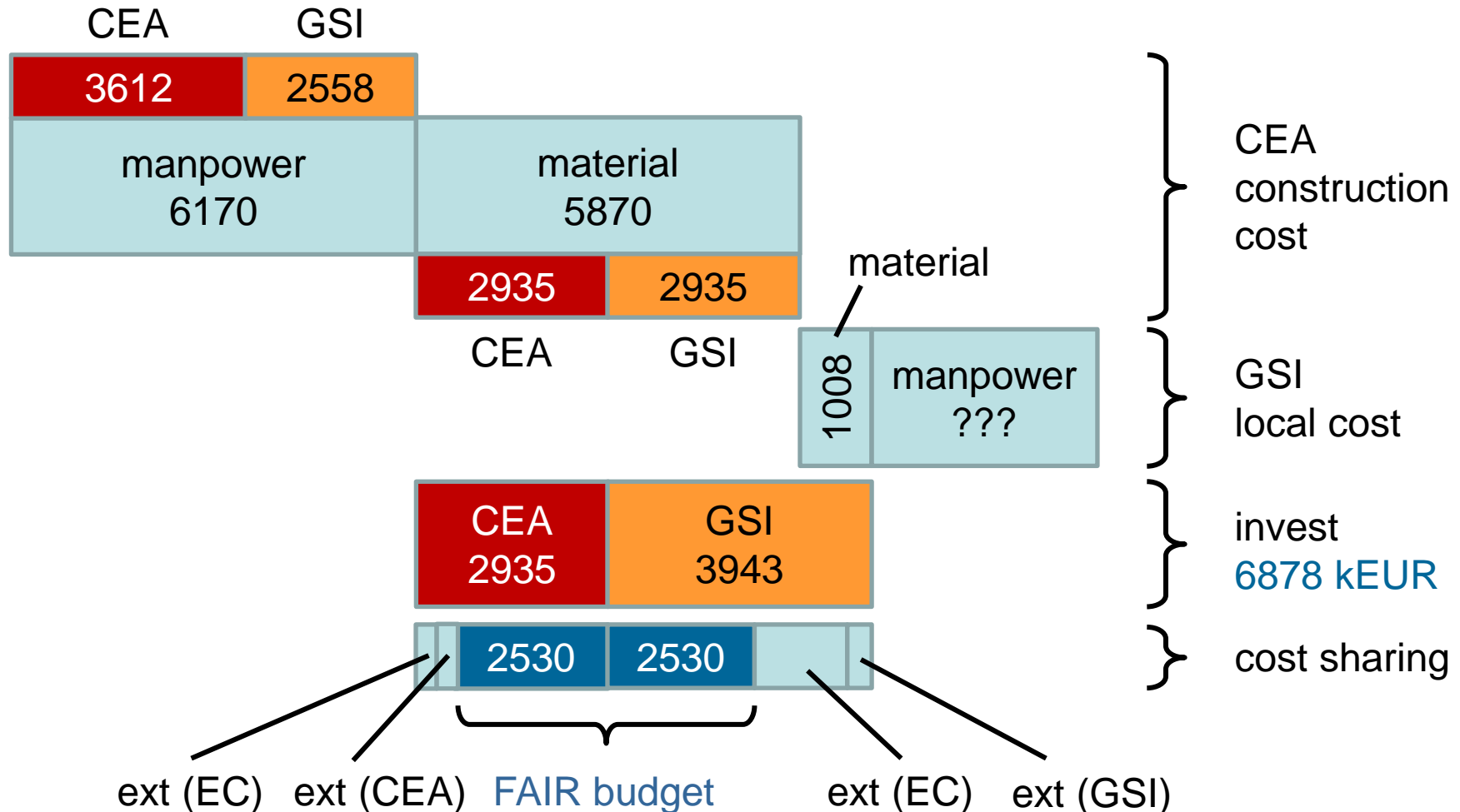
**NeuLAND**

↓ ↓  
**Staged construction**



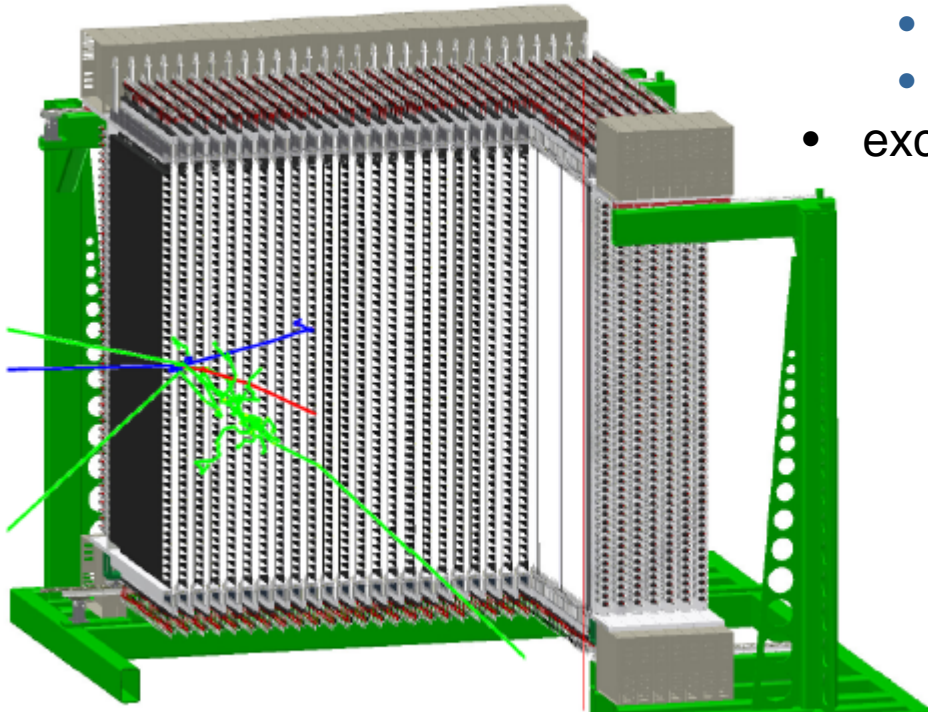
- GLAD magnet construction cost (July 2015)

- Detailed cost overview (CEA Saclay + GSI local cost) in kEUR 2005

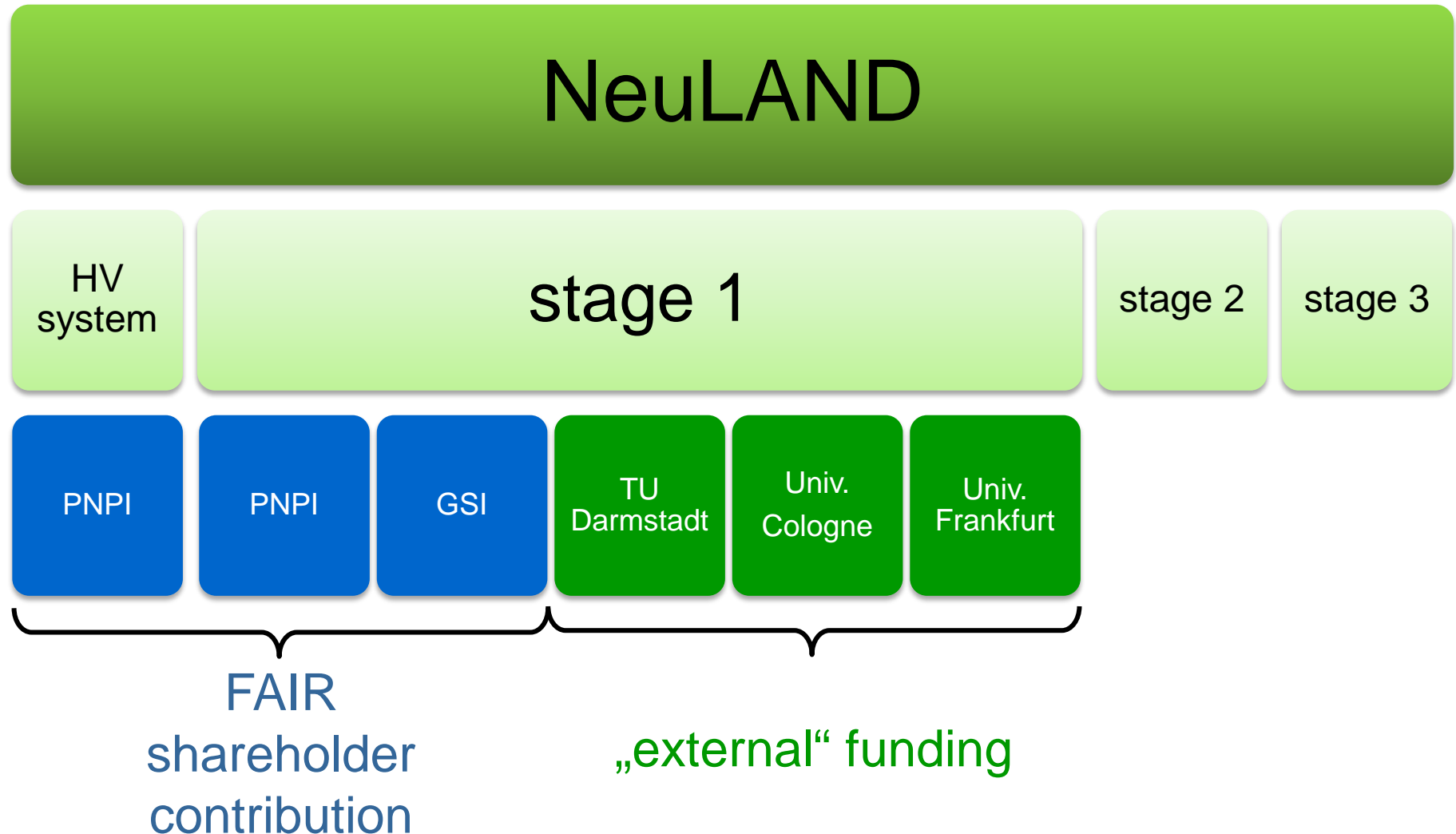




- 30 double planes
  - 100 x 100 scintillator bars x 30 planes
  - 6000 PMTs
- excellent multi-neutron capability



		200 MeV generated							600 MeV generated								1000 MeV generated						
		%	1n	2n	3n	4n	5n			%	1n	2n	3n	4n	5n			%	1n	2n	3n	4n	5n
detected	1n	88	31	6	1	0		detected	1n	92	22	2	0	0		detected	1n	89	12	1	0	0	
	2n	2	62	37	10	2			2n	2	71	32	7	1			2n	7	78	23	3	0	
	3n	0	5	49	38	14			3n	0	6	55	32	9			3n	0	8	63	26	5	
	4n	0	0	8	48	54			4n	0	0	10	57	50			4n	0	0	12	63	40	
	5n	0	0	0	3	26			5n	0	1	1	4	35			5n	0	0	0	7	46	
	6n	0	0	0	0	3			6n	0	0	0	0	5			6n	0	0	0	0	8	





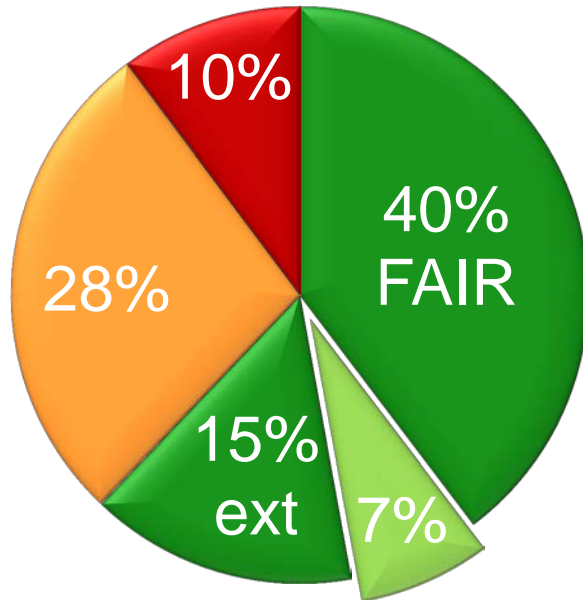
- **Phase 0**
  - R&D and experiments to be carried out with **present facilities** and FAIR/NUSTAR equipment
- **Phase 1**
  - Core detectors and subsystems completed
  - First measurements with FAIR/Super-FRS beams
    - **Carry out experiments with highest visibility as part of the core program and within the FAIR MSV**
- **Phase 2**
  - FAIR evolving towards full power
  - Completion of experiments within MSV
    - **Essentially the full program of MSV can be performed**
- **Phase 3**
  - Moderate projects, which have been initiated on the way (**outside MSV**) can be included (e.g. experiments related to return line for rings)

# Phases of HISPEC/DESPEC – funding



phase		Description	TDR	PSP code
0		Beam tracking and identification detectors		1.2.2.1
1	X	HISPEC/DESPEC Beamline		1.2.2.2
2	X	Mechanics + installation		1.2.2.3
3	X	Common EDAQ		1.2.2.4
	X	Safety		1.2.2.5
	X	Cabling and related (HISPEC/AGATA)		1.2.2.6
		Active target (MINOS)		1.2.2.7.1
	X	Active target (India)		1.2.2.7.2
	X	AGATA		1.2.2.8
	X	HYDE		1.2.2.9
		LYCCA		1.2.2.10
	X	Plunger		1.2.2.11
		AIDA		1.2.2.13
	X	DEGAS		1.2.2.14
		FATIMA		1.2.2.15
		BELEN		1.2.2.16.1
	X	MONSTER		1.2.2.16.2
	X	NEDA		1.2.2.16.3
		DTAS		1.2.2.17
	X	Isomeric Moments		1.2.2.18

(status: July 2015)



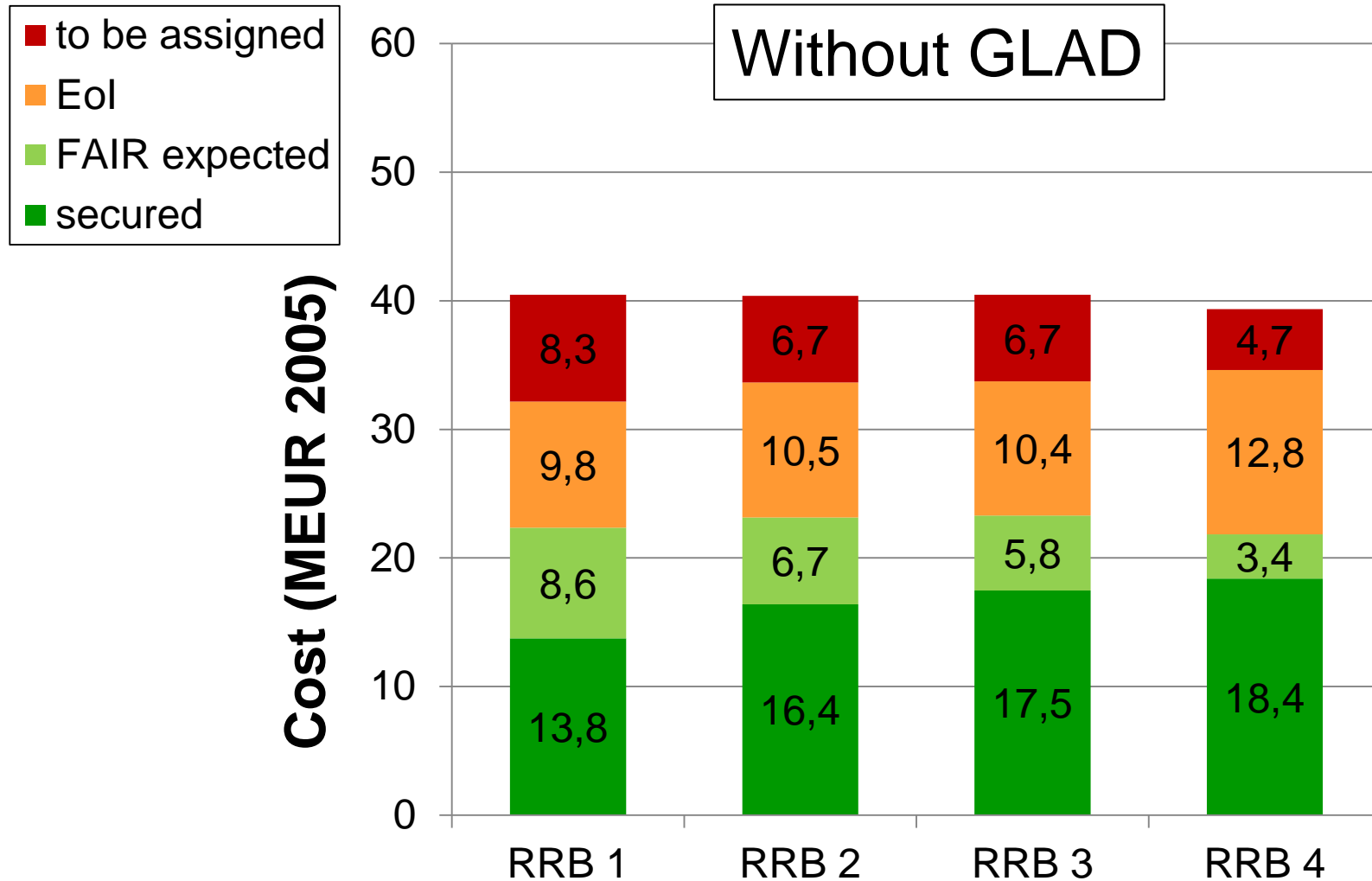
- secured FAIR
- expected from FAIR
- secured external
- Eol
- to be assigned

- FAIR shareholders and associates

- Finland
- France
- Germany
- India
- Poland
- Romania
- Russia
- Sweden
- UK

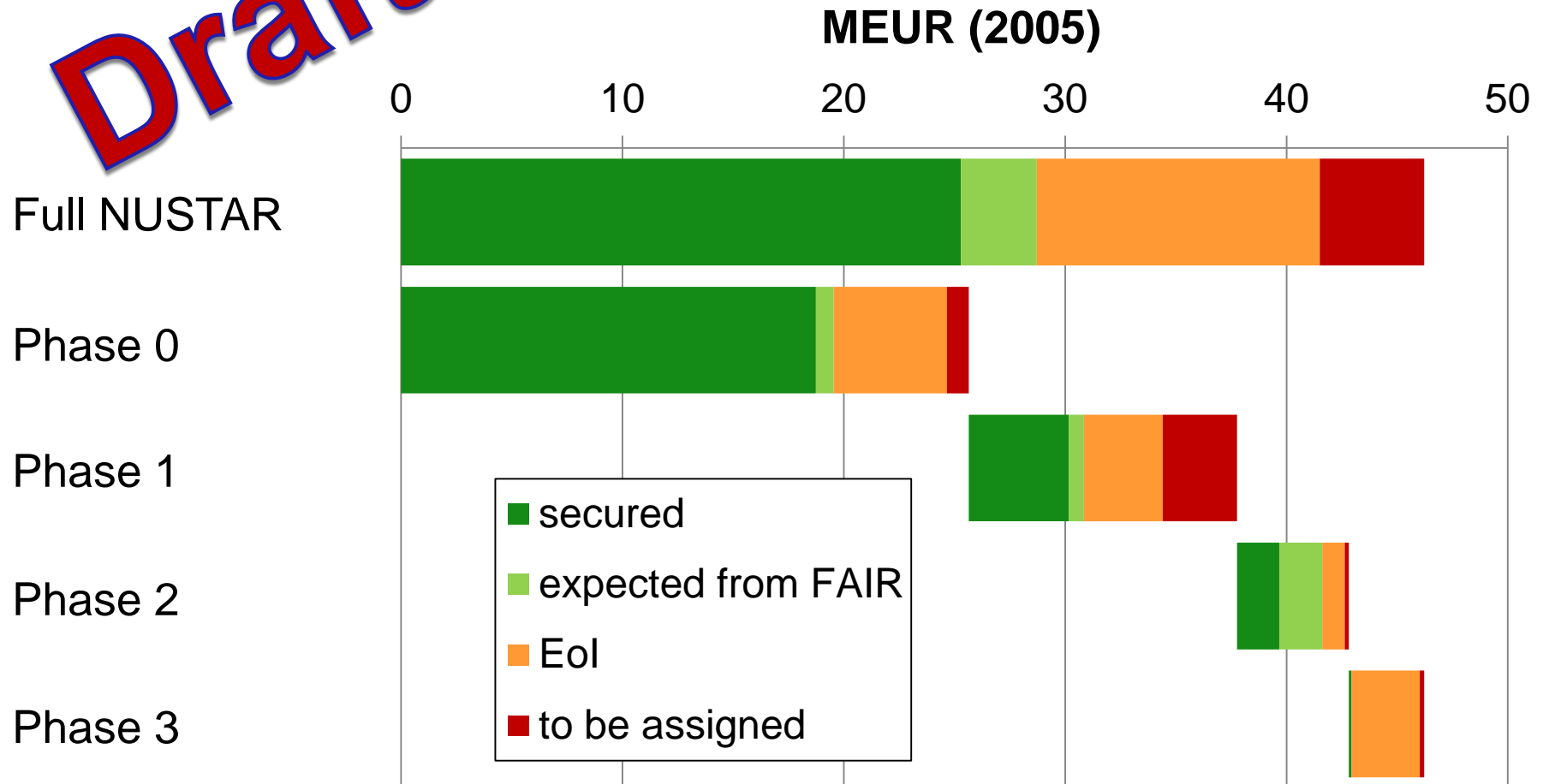
- Additional funding from:

- Belgium
- Bulgaria
- Hungary
- Israel
- Italy
- Japan
- Spain
- Turkey



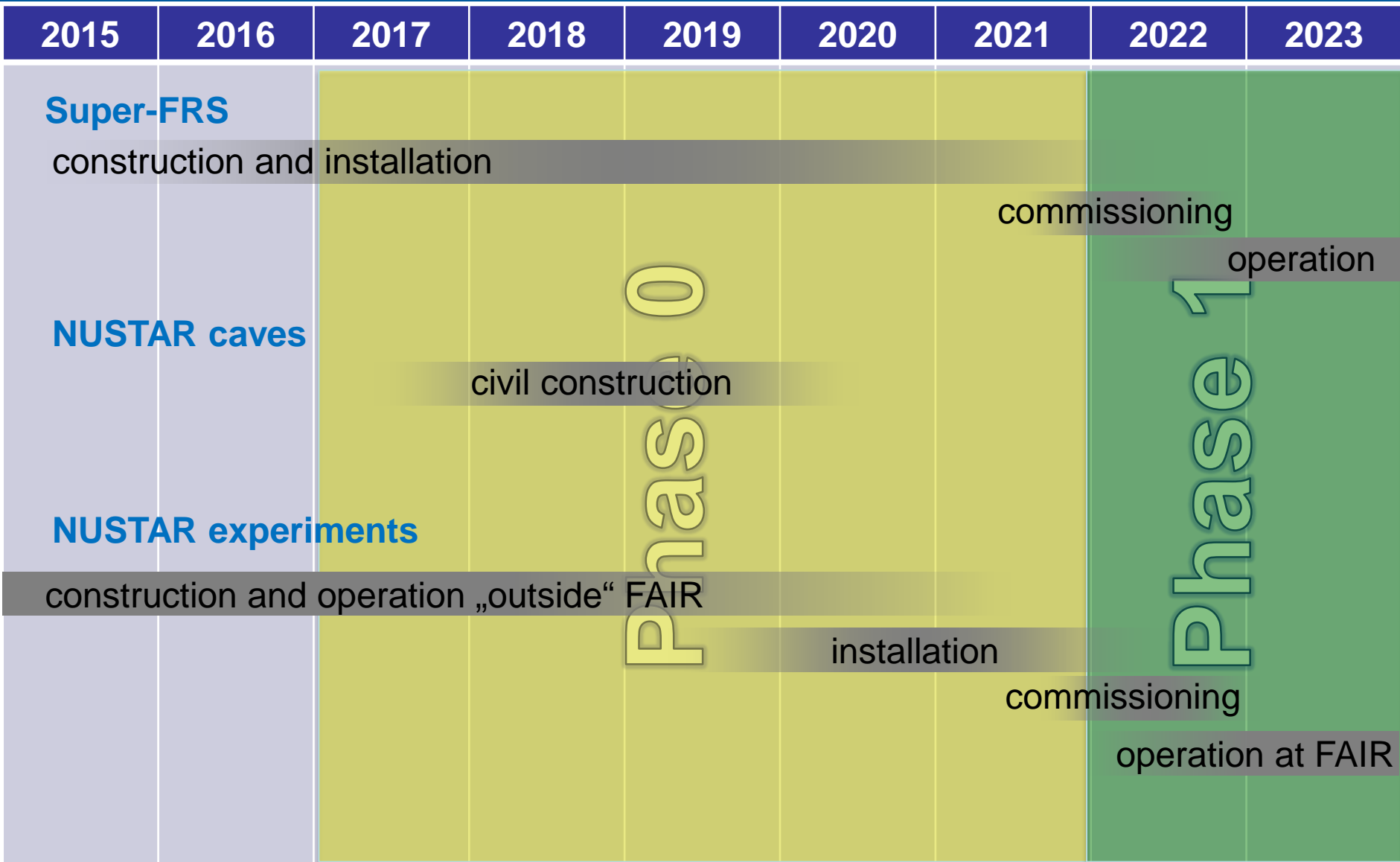
# NUSTAR experiment funding – phases

Draft



July 2015 - iteration within NUSTAR Collaboration ongoing

# Scenario for phase 0 and phase 1 operation



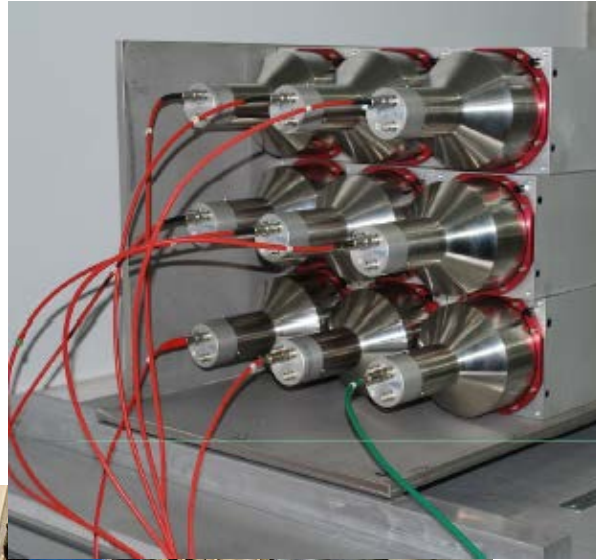


# HISPEC/DESPEC – ready for operation

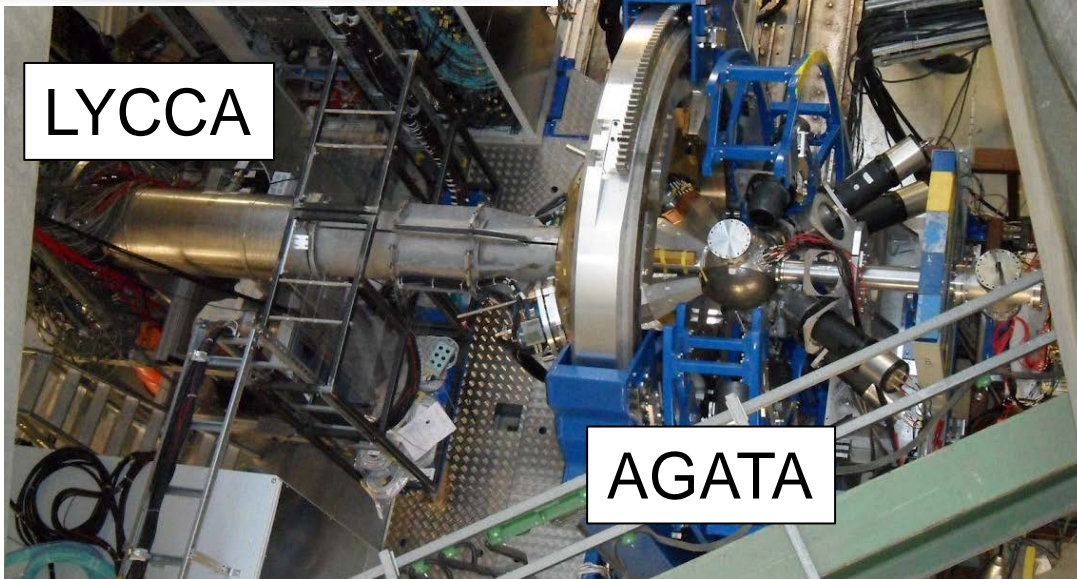
BELEN



DTAS

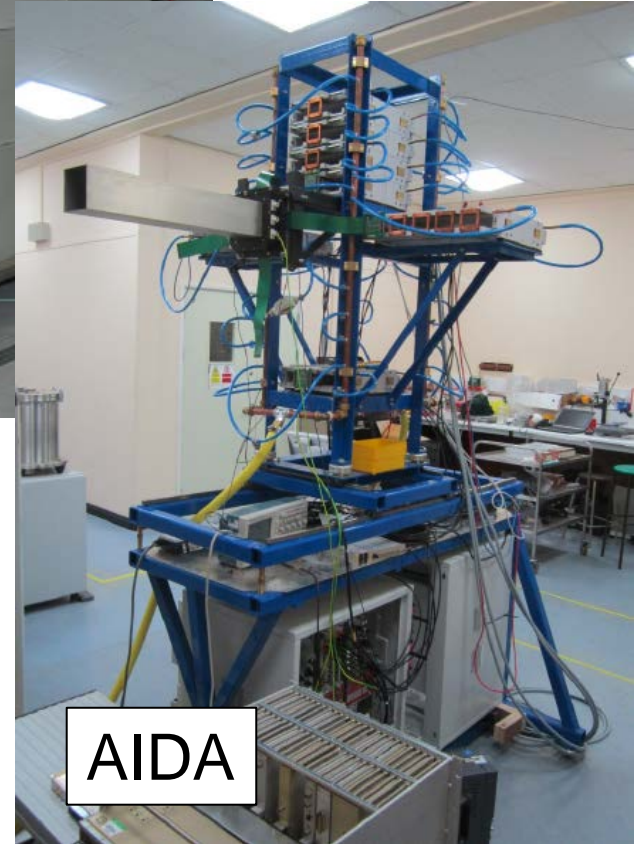


LYCCA



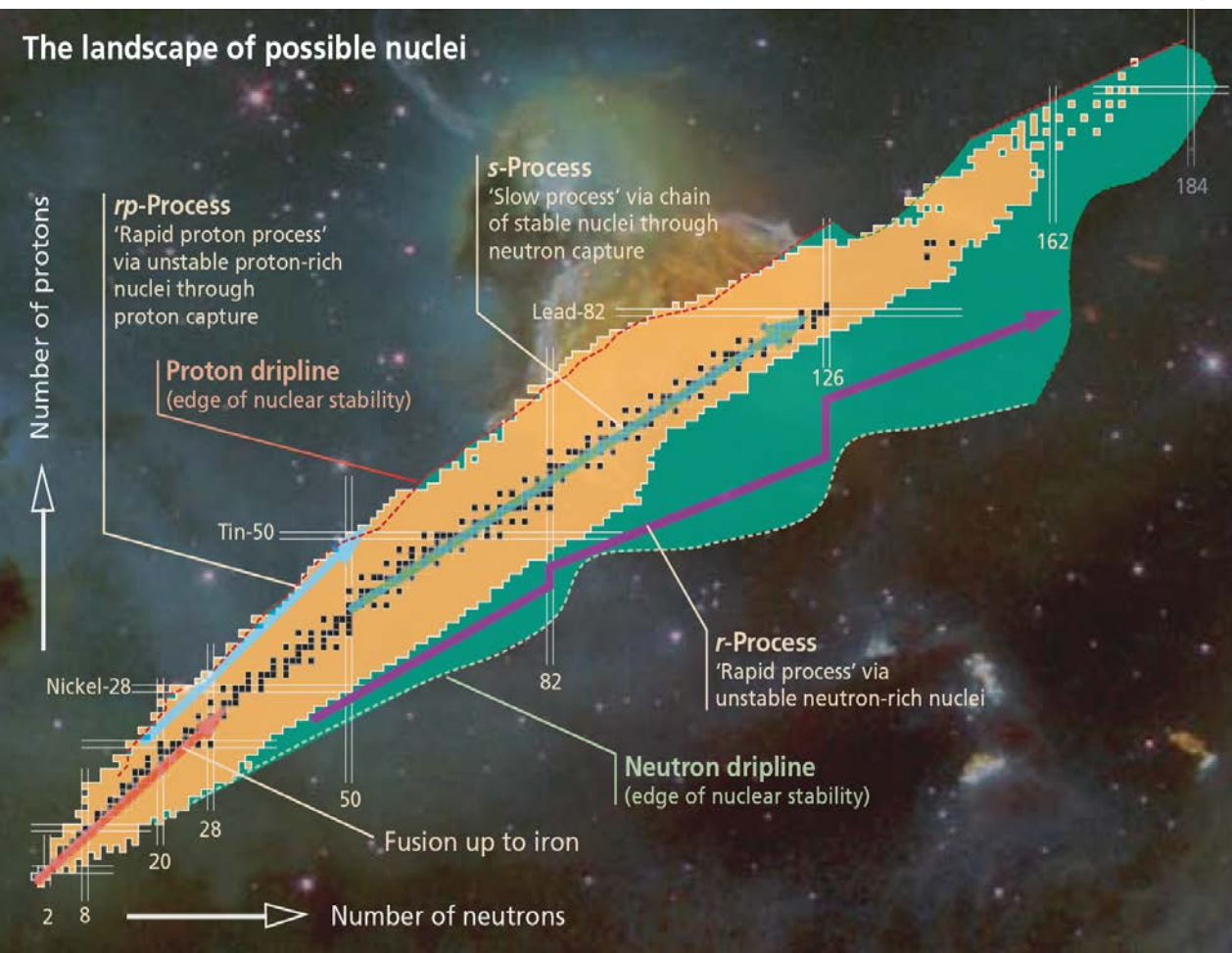
AGATA

AIDA



With in-kind from international partners ...

... the NUSTAR project aims at studying exotic nuclei ...



### Physics subject

*to understand the formation of the elements and to finally describe the atomic nucleus*

### Instrumentation

*a multitude of novel particle and radiation detectors with sophisticated electronics and data acquisition systems*

### Applications

*many new devices and methods for medicine, security, industry other research areas*

... and is on a good way!