

# **Probing the occurrence of material structures at subnuclear densities with a dynamical self-consistent description**

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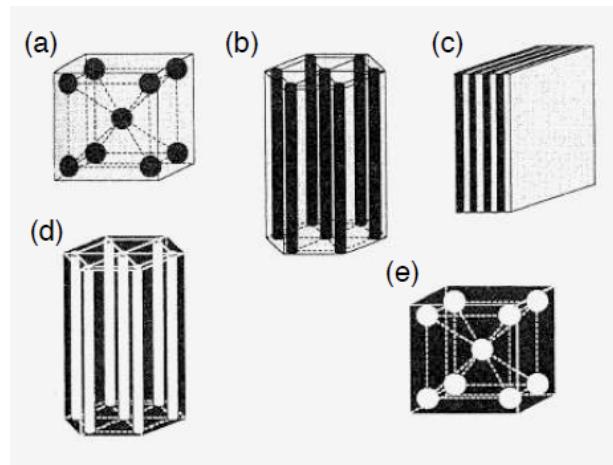


# Neutron Star Crusts: Exotic Shapes

$\rho \leq 10^{14} \text{ g/cm}^3$

at sub-nuclear densities: frustration

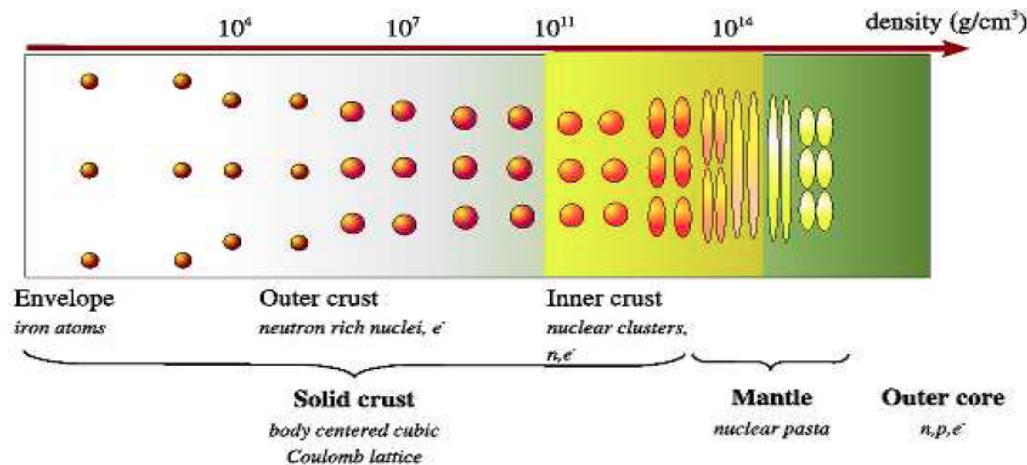
$T \approx 0.1 \text{ MeV}$



« Nuclear Pasta »

- a) spheres
- b) rods
- c) slabs
- d) cylindrical holes
- e) spherical holes

K. Oyamatsu, Nucl. Phys. **A561**, 431 (1993)



N. Chamel & P. Hænsel arXiv:0812.3955[astro-ph]

# Pasta formation

## astrophysical consequences:

- \* neutrino scattering
- \* electron scattering
- \* elastic properties of the crust

## connection with nuclear collisions

- \* fragment formation
- \* EoS sensitivity, asymmetry energy

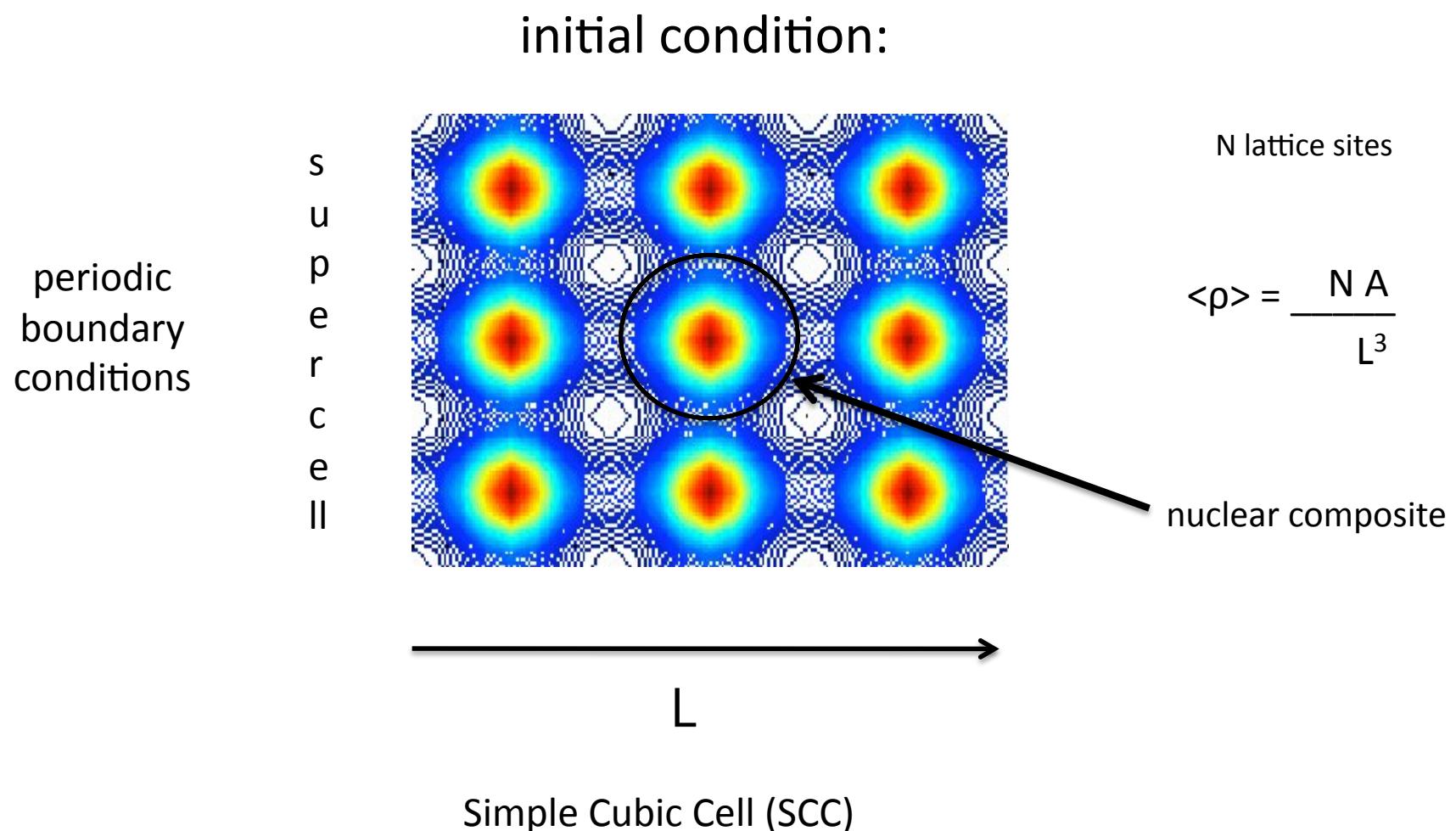
# The DYnamical WAvelets in Nuclei Model

F. Sébille, V. de la Mota, S. Figerou

Heavy Ion Collisions     $\longleftrightarrow$     Neutron Stars

**simulation of the dynamical processes  
in inhomogeneous nuclear matter using  
a large number of nucleons without any  
assumptions on the structure of nuclear  
matter.**

# Studying the behavior of matter in the crust: interacting nucleons in a uniform electron background



nuclear composites prepared self-consistently

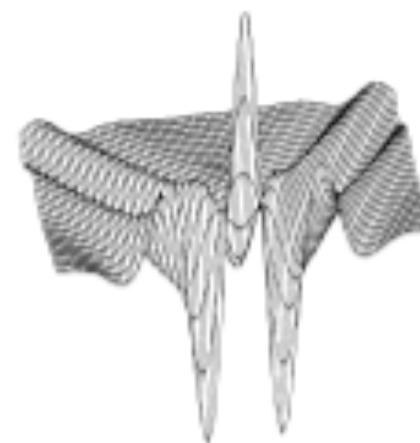
$$[\mathbf{h}(\rho), \rho] = 0$$

$$\mathbf{h} = \frac{\mathbf{p}^2}{2m} + \mathbf{V}^{HF}(\rho)$$

$$\mathbf{V}^{HF} = Tr_2\{\mathbf{V}^A(1, 2)\rho(2)\}$$

$$\rho = \sum_{\lambda=0}^N \sum_i n_i^\lambda |\alpha_i^\lambda\rangle\langle\alpha_i^\lambda|$$

$\{\alpha(\vec{r}, t)\}$  : moving basis



$$\alpha(\vec{r}) = \textcolor{brown}{\alpha_x(x)} \alpha_y(y) \alpha_z(z)$$

$$\alpha_x(x)=\mathcal{N}\exp\{-a(\chi,\phi)(x-\langle x\rangle)^2+i\frac{\langle p_x\rangle}{\hbar}(x-\langle x\rangle)\}$$

$$\{\langle x\rangle,\langle p_x\rangle,\chi,\gamma\}$$

$$\begin{array}{ll} \chi=\langle (x-\langle x\rangle)^2\rangle & \sigma=\langle [(x-\langle x\rangle),(p_x-\langle p_x\rangle)]_+\rangle \\ \phi=\langle (p_x-\langle p_x\rangle)^2\rangle & \gamma=\dfrac{\sigma^2}{2\chi} \end{array}$$

$$\Delta=\chi\phi-\sigma^2=\frac{\hbar}{4}$$

dynamical evolution:

$$i\hbar\dot{\rho} = [\mathbf{h}, \rho]$$

$$\mathbf{h} = \frac{\mathbf{p}^2}{2m} + V^{HF}(\rho(t))$$

$$\rho = \sum_{\lambda=0}^{\infty} \sum_i n_i^\lambda |\alpha_i^\lambda\rangle\langle\alpha_i^\lambda|$$

$$i\hbar \frac{\partial |\alpha_k^\lambda(t)\rangle}{\partial t} = \mathbf{h} |\alpha_k^\lambda(t)\rangle$$

$$\mathcal{A} = \int_{t_1}^{t_2} \langle \alpha | i\hbar \frac{\partial}{\partial t} - \mathbf{h} | \alpha \rangle$$

TDHF equation

time dependent one-body Hamiltonian

c.s. expansion of the one-body density matrix

equation of motion for the basis elements

variational principle

evolution of wavelet centroids and widths:

$$\frac{d\langle x \rangle}{dt} = \frac{\langle p_x \rangle}{m}$$

$$\frac{d\langle p_x \rangle}{dt} = -\frac{\partial \langle V^{HF} \rangle}{\partial \langle x \rangle}$$

$$\frac{d\chi}{dt} = \frac{4\gamma\chi}{m}$$

$$\frac{d\gamma}{dt} = \frac{\hbar^2}{8m\chi^2} - \frac{2\gamma^2}{m} - \frac{\partial \langle V^{HF} \rangle}{\partial \chi}$$

$$\langle \quad \rangle \equiv \langle \alpha | \quad | \alpha \rangle$$

effective force:

$$V_q^{HF}(\rho, \xi) = \frac{t_0}{\rho_\infty} \rho + \frac{t_3}{\rho_\infty^{\nu+1}} \rho^{\nu+1} + \frac{c}{\rho_\infty^2} \xi^2 + \frac{4qc}{\rho_\infty^2} \rho \xi + \frac{\Omega}{3\rho_\infty^2} \xi^2 + \frac{4q\Omega}{3\rho_\infty^2} (\rho - \rho_\infty) \xi + V_q^C$$

$$\rho = \rho_n + \rho_p$$

$$\xi = \rho_n - \rho_p$$

$$q = \begin{cases} \frac{1}{2} \text{ neutrons} \\ -\frac{1}{2} \text{ protons} \end{cases}$$

$$t_0 = -356/\rho_\infty (\text{MeV}/\text{fm}^{-3}) \quad \rho_\infty = 0.145 \text{ fm}^{-3}$$

$$t_3 = 303/\rho_\infty^{\nu+1} (\text{MeV}/\text{fm}^{-3}) \quad \nu = 1/6$$

$$c = 20 \text{ MeV} \ (J=31.5 \text{ MeV})$$

$V_q^C$  lattice calculation with Ewald summation technique

## macroscopic properties of nuclear matter

$$\omega = \frac{\int V_q^{HF} \, d\rho}{\rho} + \omega_{kin} \quad \delta = \frac{\xi}{\rho} \quad \omega_\delta = \frac{1}{2} \lim_{\delta \rightarrow 0} \frac{\partial^2 \omega}{\partial \delta^2}$$

$$\omega(\rho, \delta) = \frac{t_0}{2\rho_\infty} \rho + \frac{t_3}{(\nu+2)\rho_\infty^{\nu+1}} \rho^{\nu+1} + \frac{c}{\rho_\infty^2} \delta^2 \rho^2 + \frac{\Omega}{3\rho_\infty^2} (\rho - \rho_\infty) \rho \delta^2 + \omega_{kin}$$

$$\boxed{\omega = \omega_0 + \frac{K_0}{18\rho_\infty^2} (\rho - \rho_\infty)^2 + \left[ J + \frac{L}{3\rho_\infty} (\rho - \rho_\infty) \right] \delta^2}$$

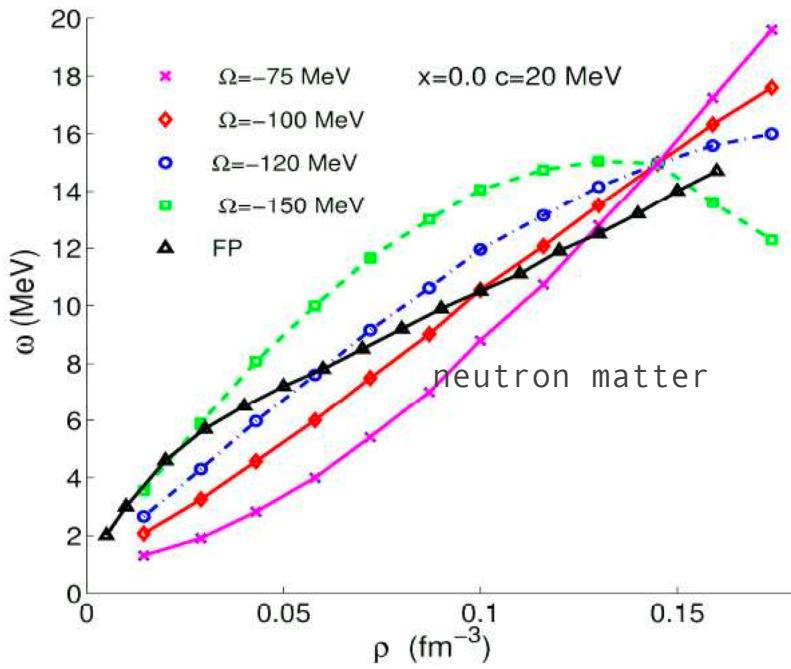
J.M. Lattimer, Ann. Rev. Nucl. Part. Sci. 31  
 (1981) 337  
 K. Oyamatsu, Pr. Th. Phys. 109 (2003) 631

$$J = \omega_\delta(\rho_\infty) \quad L = 3\rho \frac{\partial \omega_\delta}{\partial \rho} \Big|_{\rho=\rho_\infty} \quad L_{kin} = 3\rho_\infty \left[ \frac{d\omega_{kin}(\rho)}{d\rho} \right]_{\rho_\infty}$$

$$K_{sym} = 9\rho^2 \frac{\partial^2 \omega_\delta}{\partial \rho^2} \Big|_{\rho=\rho_\infty} \quad K_0 = K_{sym}(\delta = 0)$$

$$L = 6c + \Omega + L_{kin}$$

# neutron matter properties

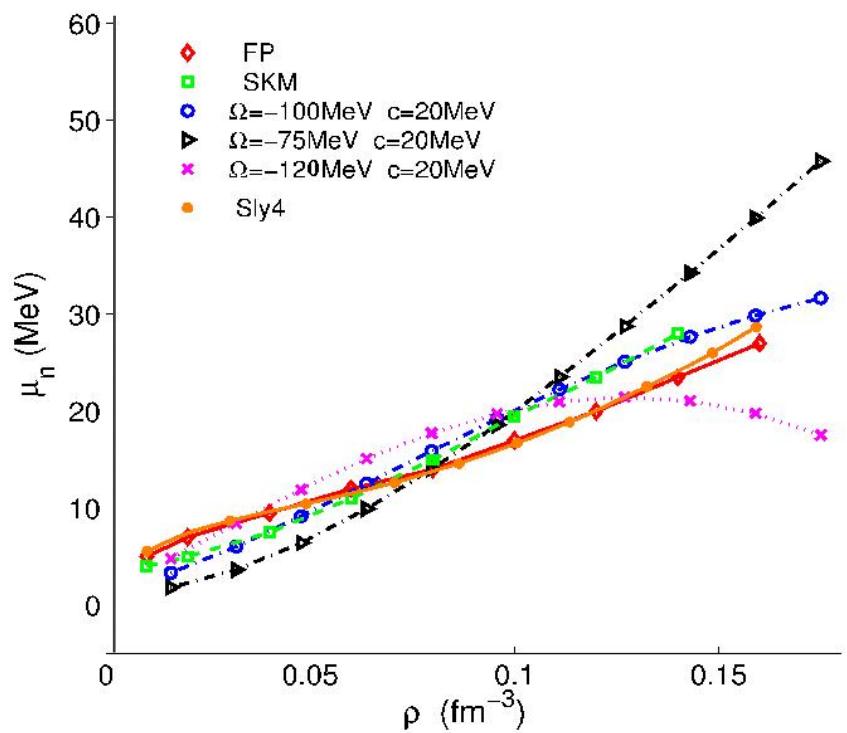


**FP** B. Friedman et B.R. Pandharipande, Nucl. Phys. 361(1981) 501

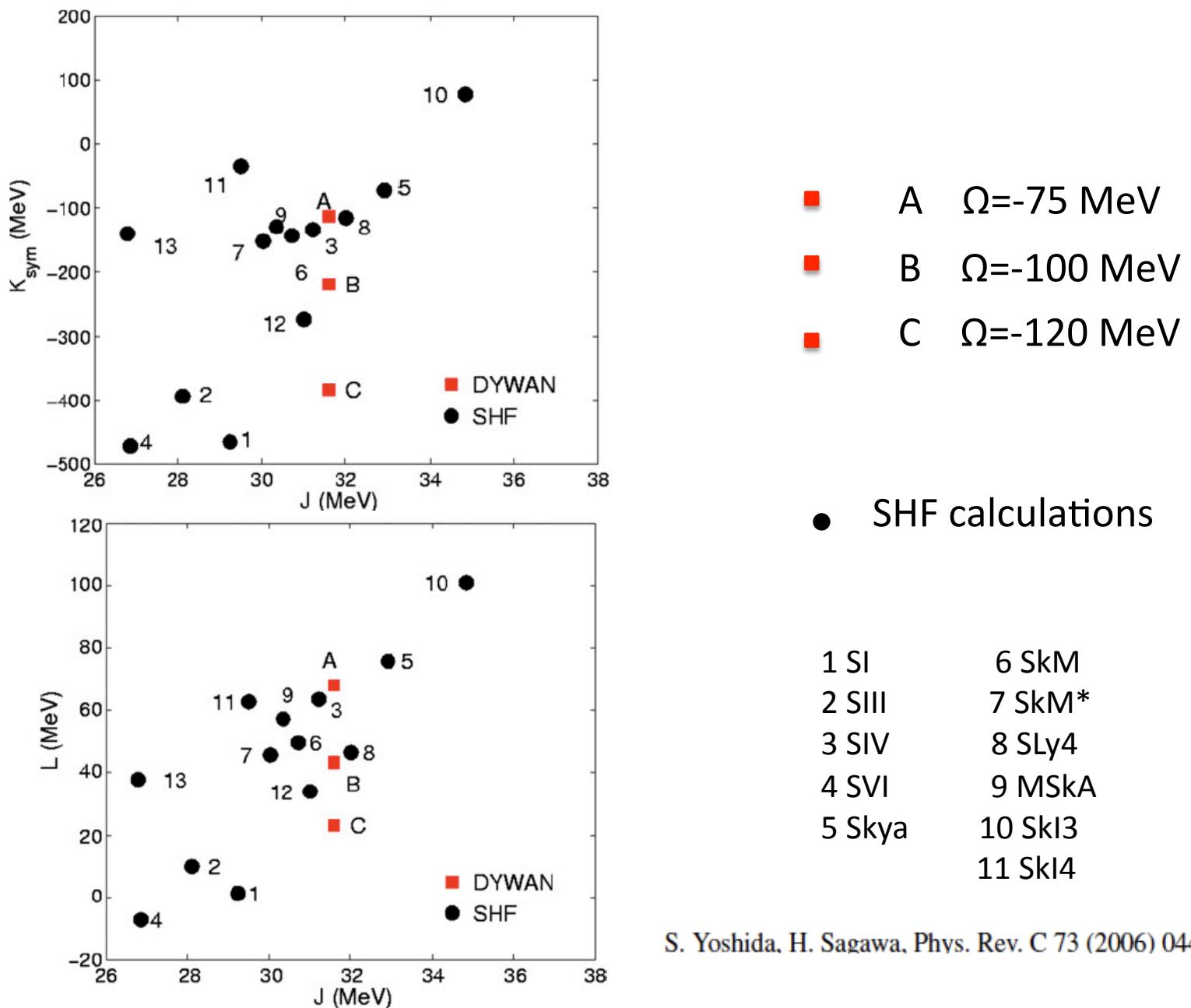
**SKM** H. Krivine et al, Nucl. Phys. A 336 (1980) 155

**Sly4** F. Douchin et al, Phys. Lett. B 435 (2000) 107

$$\mu_i = \frac{\partial \varepsilon}{\partial \rho_i} \quad i = n, p$$

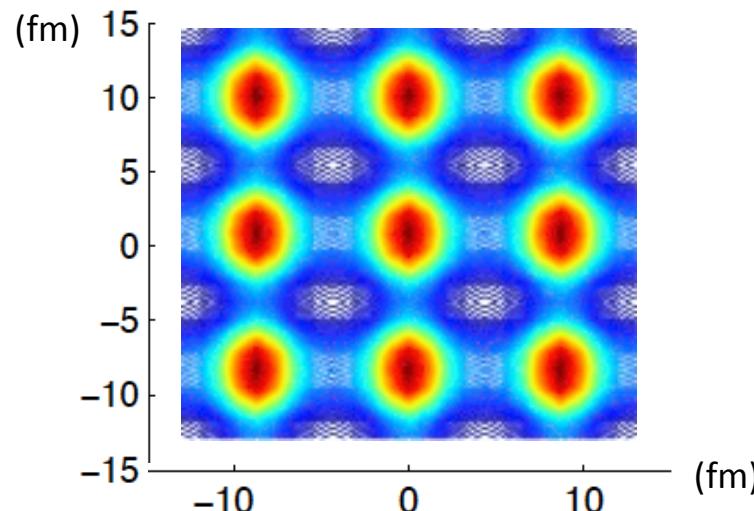


## correlations between macroscopic properties



dynamical evolution

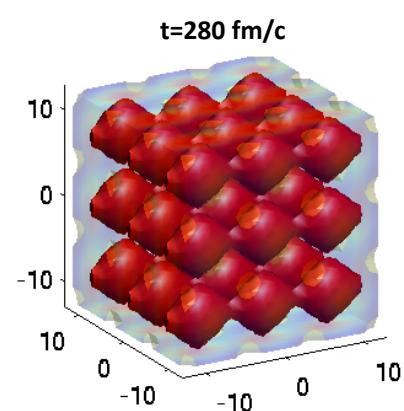
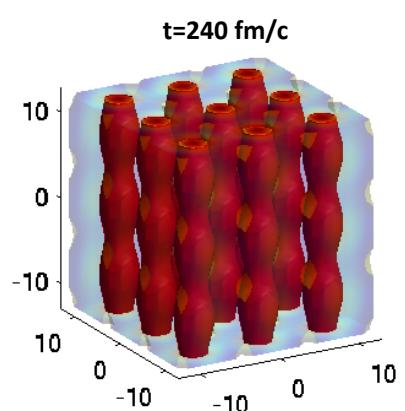
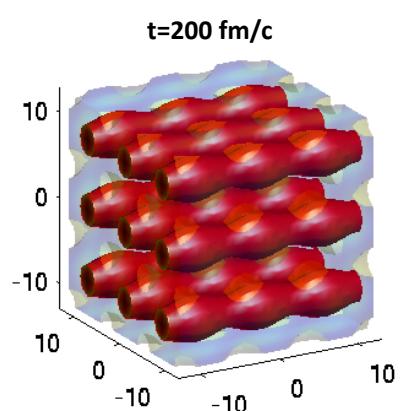
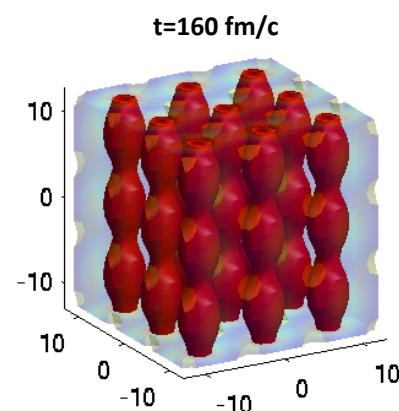
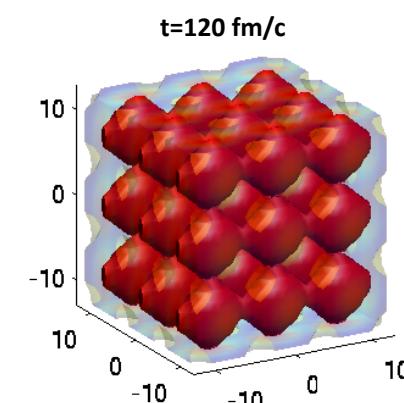
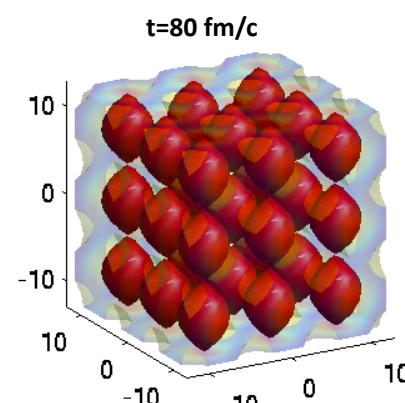
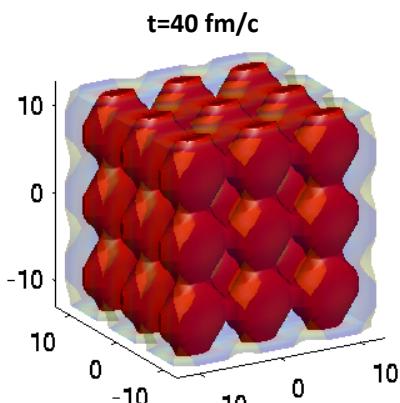
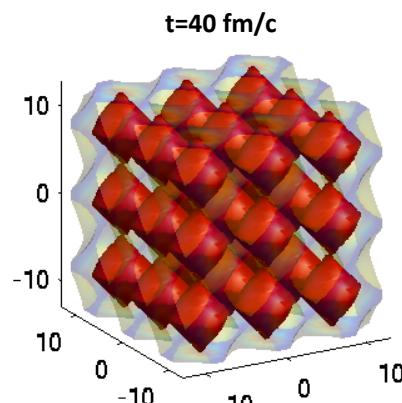
$\Omega = -100 \text{ MeV}$



Oxygen  $x=0.2$

$$\langle \rho \rangle = 0.072 \text{ fm}^{-3}$$

$$\rho_t = 0.08 \text{ fm}^{-3}$$

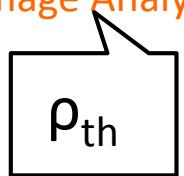


# neutron phase diagrams

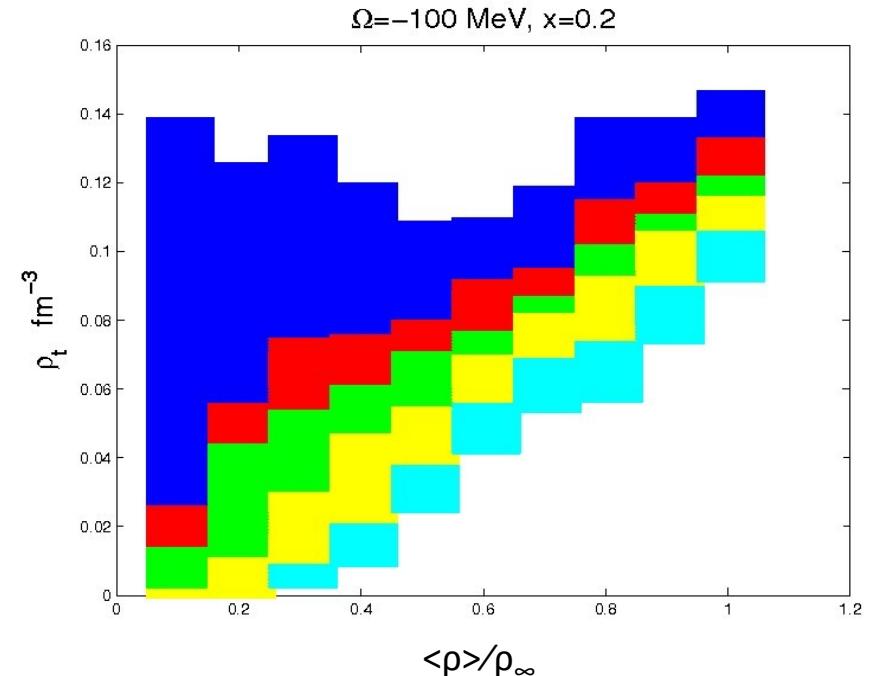
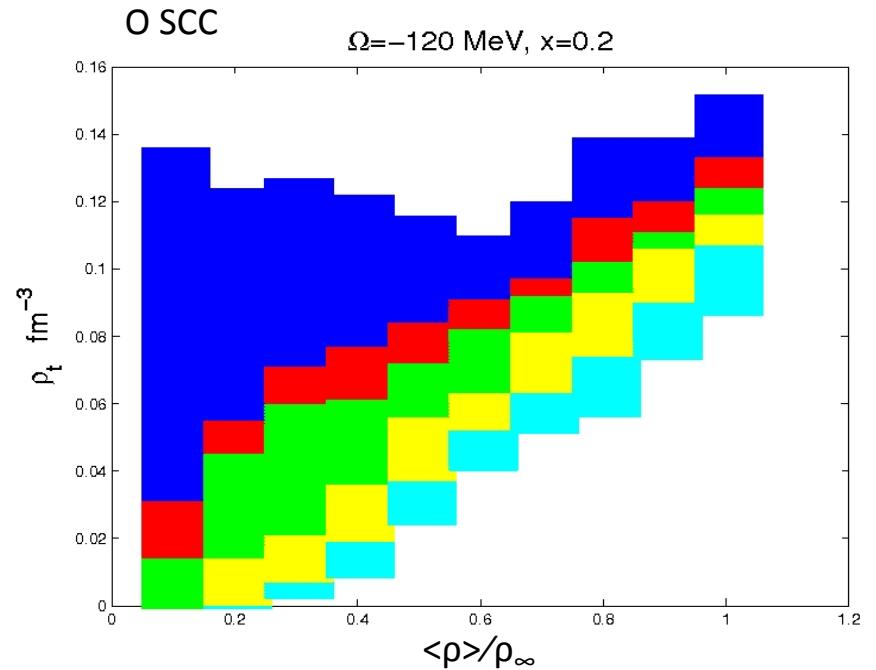
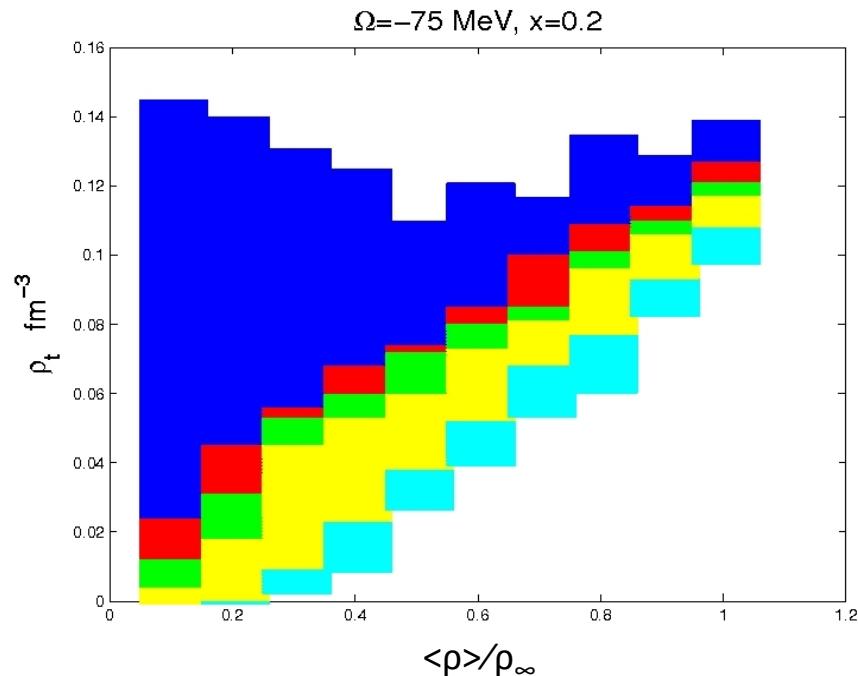
Structure patterns in the  $\rho_t$  -  $\langle \rho \rangle / \rho_\infty$  plane  
for different values of  $\Omega$

structures characterization

Morphological  
Image Analysis



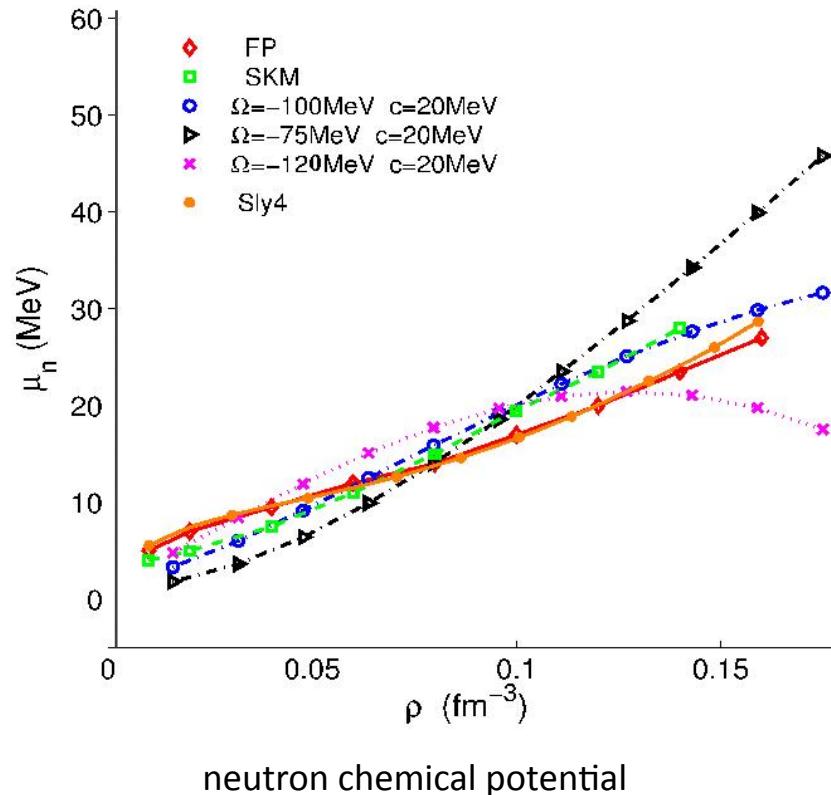
- █ spheres
- █ rods
- █ slabs
- █ sponges
- █ bubbles



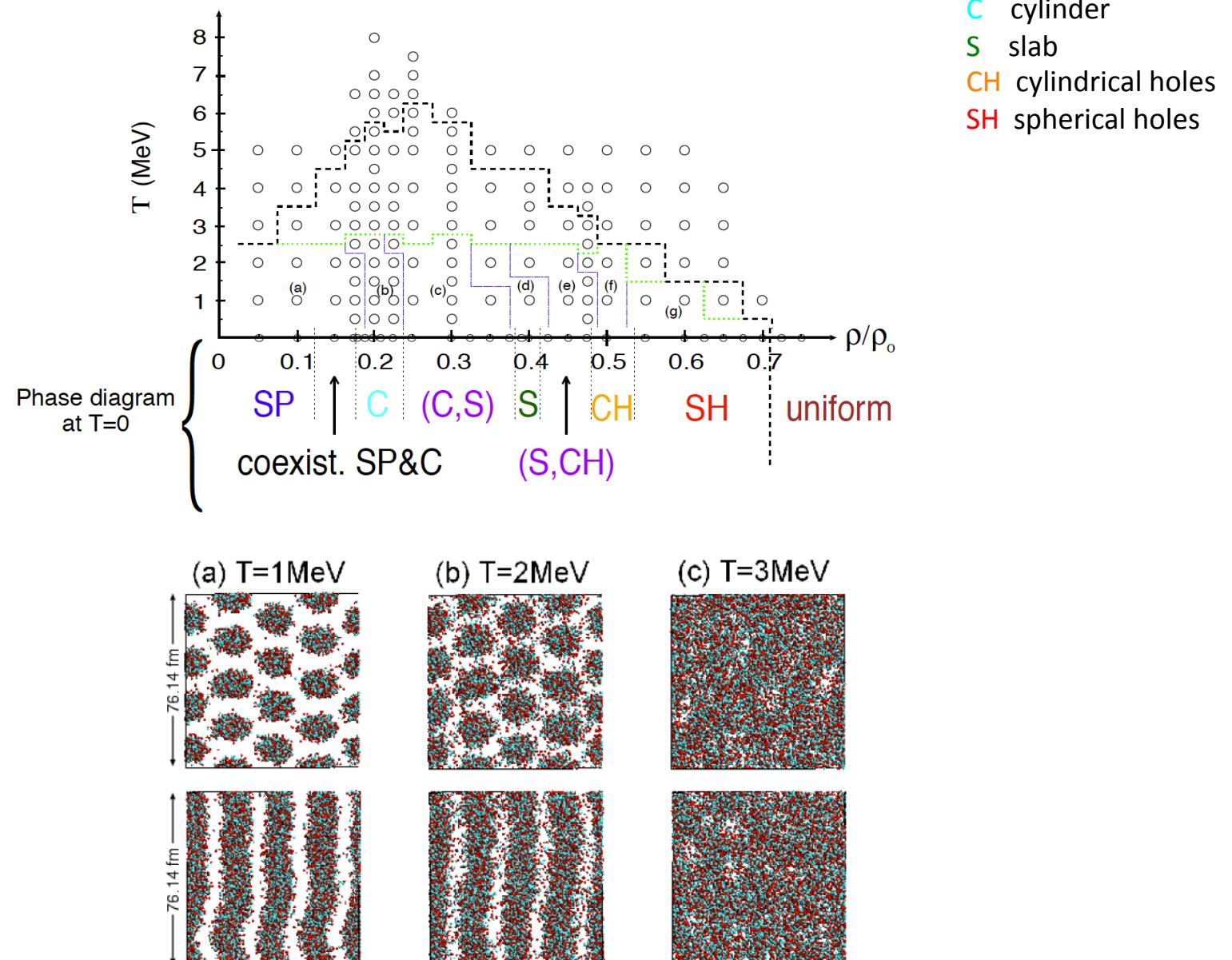
# influence of $\Omega$

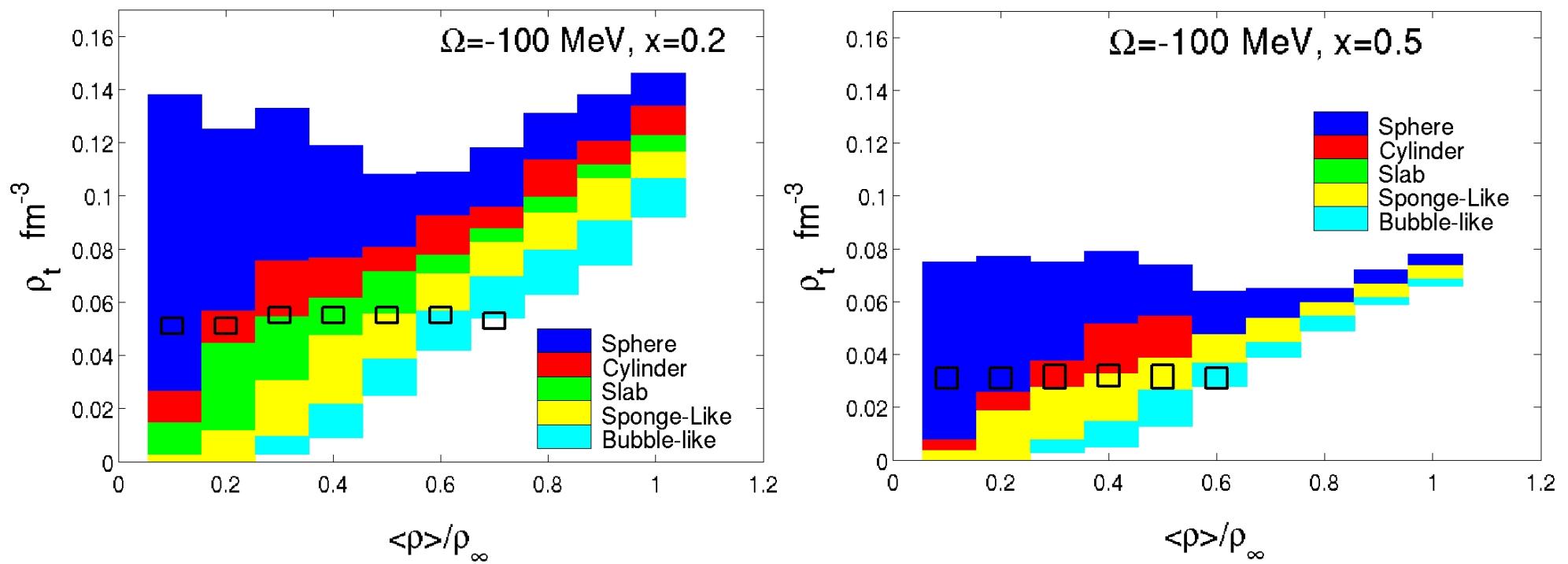
**stiffer** force

favours spherical & sponge-like structures  
restrains cylinder & slabs.



# Quantum Molecular Dynamics





$$0.03 \text{ fm}^{-3} < \rho_t^{\text{ref}} < 0.08 \text{ fm}^{-3}$$

B. Schuetrumpf et al. PRC 87 055805 (2013)

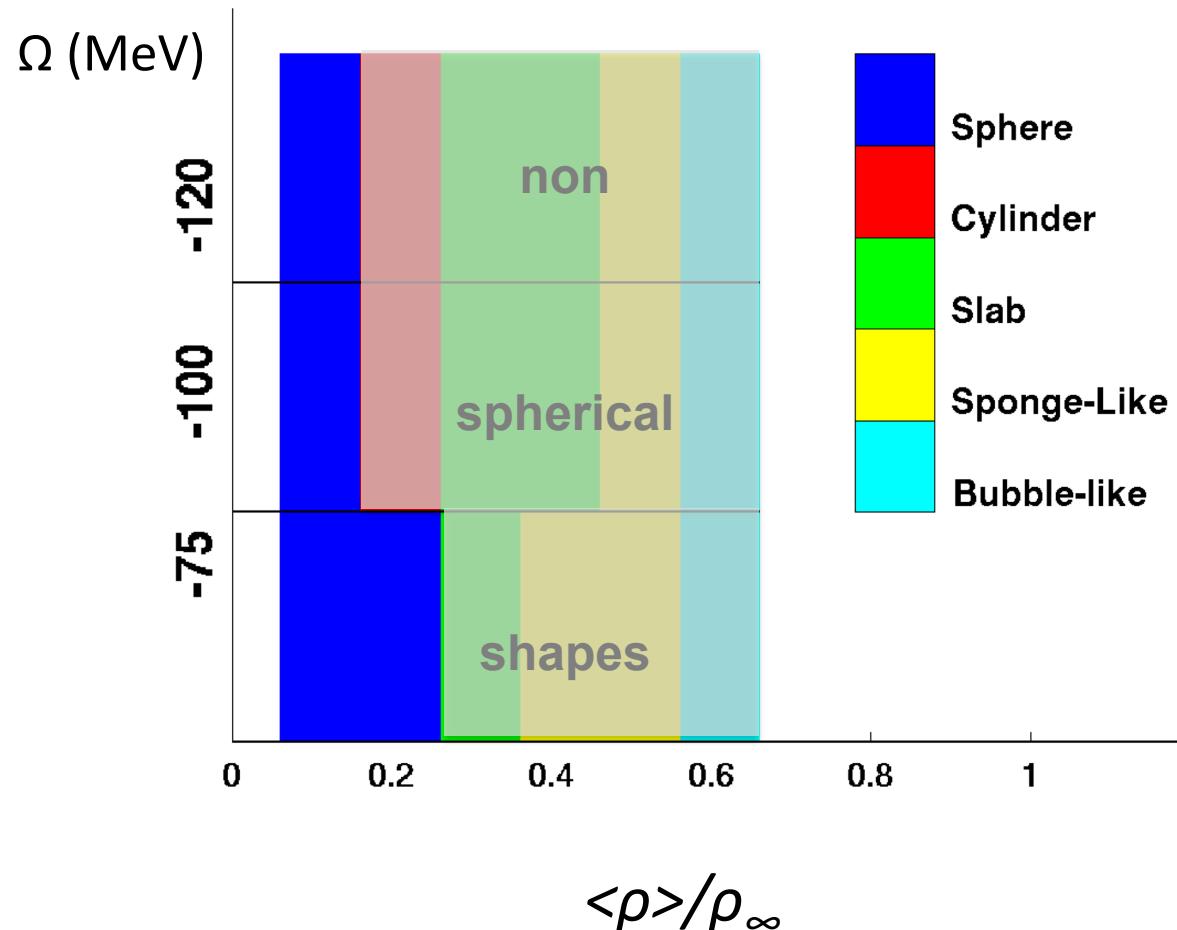
◻  $\rho_t^{\text{ref}}$  satisfies the condition:

$$\left( \frac{d^2(\mathcal{V}/\mathcal{A})}{d\rho_t^2} \right)_{\rho_t^{\text{ref}}} = 0$$

G. Watanabe et al.  
PRC 66 012801 (R) (2002)

- ◆  $\rho_t^{\text{ref}} \sim$  independent of  $\langle \rho \rangle$
- ◆ classical sorting of structures
- ◆ different structures exist for the same  $\langle \rho \rangle$  value: embedded
- ◆ sensitivity to isotope composition

synopsis: shapes vs  $\Omega$  at  $\rho_t^{\text{ref}}$



the density region  
with non spherical  
shapes  
is narrow for  
the stiffer force

agreement with Sonoda et al.  
PRC 77 035806 (2008)

# Summary and Perspectives

self-consistent dynamical description of matter has been performed

a small initial deposited energy permits to explore a landscape of meta-stable structures with constant energy preserving the initial symmetries

the ordered standard types of pasta emerge naturally whatever the threshold density is

sensitivity to the asymmetry dependent term of the potential: the stiffer force favours the occurrence spherical and sponge-like structures, restraining rods and slabs

effective force drawbacks: non-locality effects, spin-orbit, pairing

temperature: effects of varying T

inclusion of residual interactions in the dynamical description

dissipation: influence on pasta formation

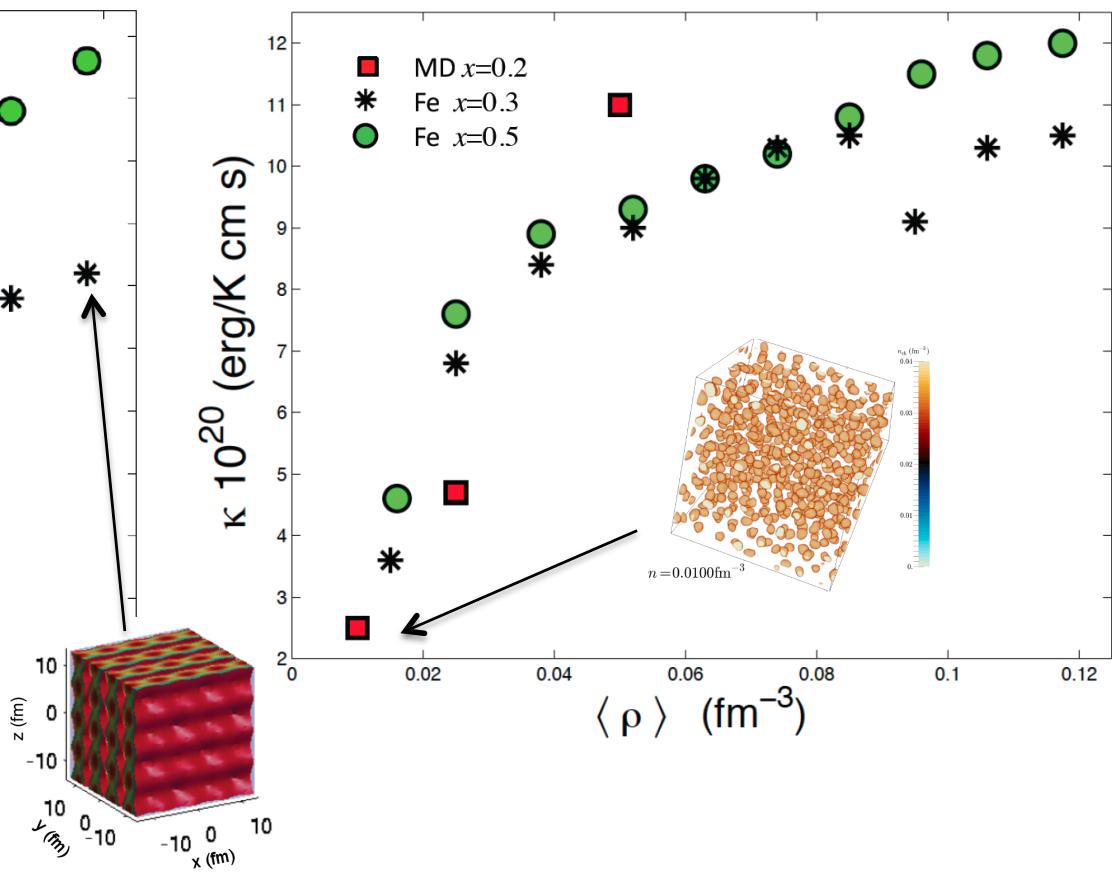
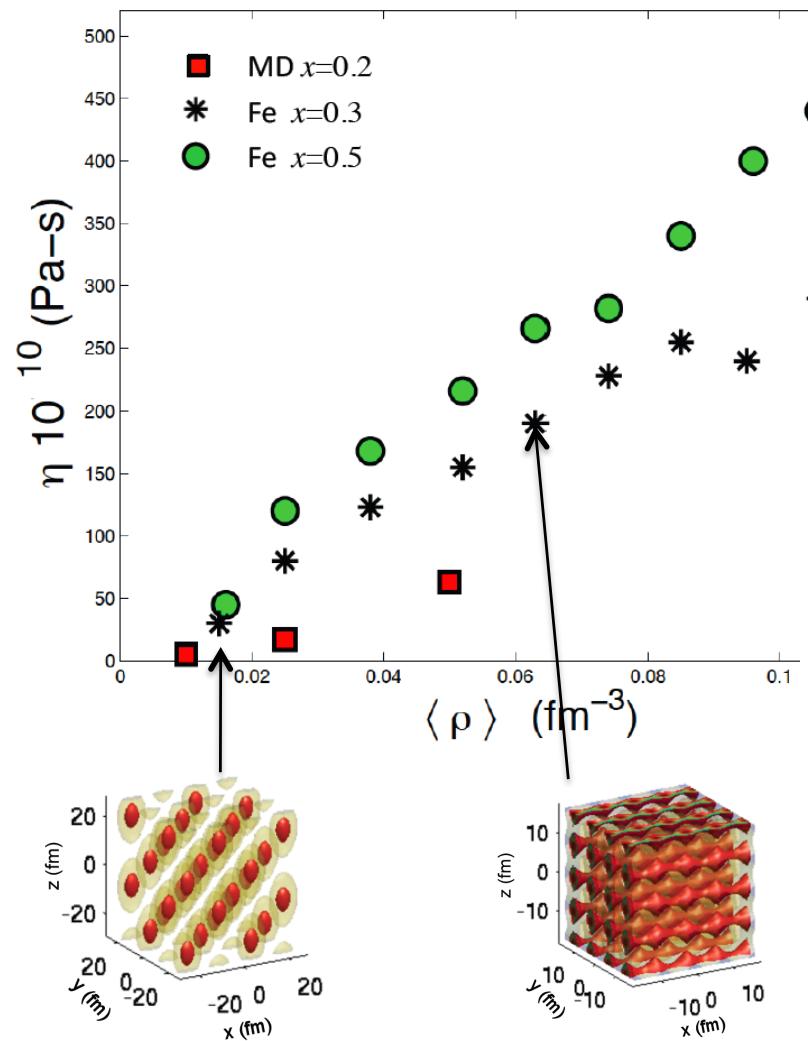


work in progress

bonus ...

## shear viscosity

## thermal conductivity



■ C.J. Horowitz and D.K. Berry, Phys. Rev. C 78 (2008) 035806  
● \* DYWAN calculations Fe FCC

# The Coulomb potential

protons

wavelet expansion of the density

$$\rho(r) \sim \sum_i \beta_i g_i(r)$$

charged particles

protons  
electrons

electrons

$$\rho_{\text{electron}} = Q/\mathcal{V}$$

$\rho_T = \rho_{\text{proton}} + \rho_{\text{electron}}$

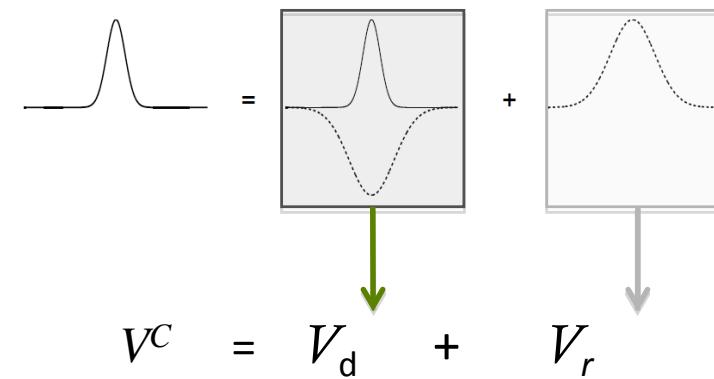
total density decomposition



potential decomposition

$V_d$ : direct term  $\rightarrow$  « cut-off » analytic

$V_r$ : reciprocal term  $\rightarrow$  solution of the Poisson equation with a FFT



[1]

## MORPHOLOGICAL IMAGE ANALYSIS (MIA)

convex ensemble  $K$       → Minkowski Functionals:

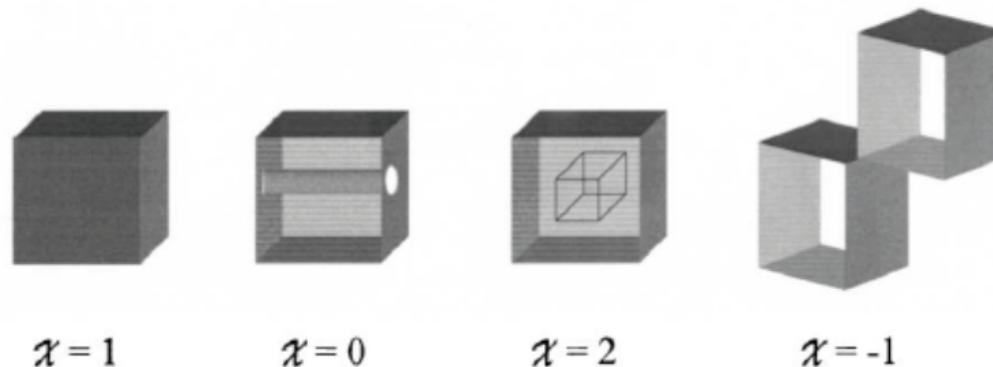
**additive:**  $K = \bigcup K_i$ ,  $K_i$ =unitary

$\mathcal{V}$  volume  
 $\mathcal{S}$  surface  
 $\mathcal{B}$  curvature  
 $\chi$  Euler char.

they provide a complete measure of  $K$  morphology

e  
x  
a  
m  
p  
l  
e

3D



$$\chi = N_{\text{elements}} - N_{\text{tunnels}} + N_{\text{cavities}}$$

[1] INTEGRAL-GEOMETRY MORPHOLOGICAL IMAGE ANALYSIS

*K. Michielsen, H. De Raedt / Physics Reports 347 (2001) 461}538 461*