# Exploring the symmetry energy in <sup>40,48</sup>Ca+ <sup>40,48</sup>Ca systems at E/A = 35 MeV

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- Introduction to the nuclear EOS
- Analysis of the experiment <sup>40,48</sup>Ca+<sup>40,48</sup>Ca @ E/A=35 MeV
- Extraction of the symmetry energy term of EOS
- Conclusions

#### Time evolution of central collisions at intermediate energies

M. Colonna, A. Ono and J. Rizzo PRC82, 054613 (2010) Typical event of Xe+Sn @ E/A =50 MeV



- EOS and transport properties paly a fundamental role in the understanding of many a aspects of nuclear physics and astrophysics.
- Explore the EOS under laboratory controlled conditions
- HIC is Femtonovae which mimic Supernovae

ECT\* 2014 : Simulating the Supernova Neutrinosphere with Heavy Ion Collisions

### Experimental probe of nuclear stopping



#### Experimental probe of collective flow



#### Radial flow:

Influence of radial collective flow on multif partitions in central Xe+Sn & Ta+Zn



Amount of the radial collective energy governs the degree of fragmentation of hot nuclei

E. Bonnet et al., Nucl. Phys. A816, 1 (2009) Phys. Rev. Lett. 105 (2010) J.D. Frankland et al., in prep

# Production of fragments and secondary decay





- Primary fragments are excited :
- S. Hudan et al., PRC67, 064613 (2003).

• E\*/A ≈ 3 MeV

$$E(\rho,\delta)/A = E(\rho,\delta=0)/A + S(\rho) \cdot \left(\frac{\rho_n - \rho_p}{\rho_n + \rho_p}\right)^2$$

- Information on the EOS: need to study fragments at freeze-out, BEFORE secondary decays : PRIMARY FRAGMENTS
- Observables : Isotopic distribution of fragments and isoscaling
- But : observables measured after secondary decay (need back-tracing)

# Accessing the symmetry energy



# Effects of secondary decays





- After decay : Cannot distinguish between the two interactions (soft/stiff)
- Secondary decays need to be taken into account for comparison to experimental data
- Statistical model
- Or/and : experimentally provide the primary distributions



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# Isotopic distributions of PLF



- Broad A<sub>PLF</sub> distributions (more than 13 isotopes)
- Sensitive to the n-richness of the system
- N/Z up to 1.58 (11% N/Z <sup>48</sup>Ca) very exotic

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# Reconstruction of primary fragments

- Up to 20 isotopes
- Average value and  $\sigma$  increases with  $Z_{\rm pr}$
- Small differences for light Z<sub>pr</sub>
- Strong dependence on the n-richness of the system for heavy fragments
- small dependence on the n-richness of the target



Can be used as an observable

Evaluation of the effect of neutron emission on the width : AMD+GEMINI It will be used as correction to the data.

# Staggering effect for PLF and primary frag



The staggering is determined by secondary evaporative stage No sensitivity to dynamics.

Validity of method: peripheral collisions and Z>12





Consistent with the values of saturation density The method is validated and should be applied to more dissipative collisions

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4.00 - 4.50

4.50 - 5.00

2.68

2.07

2.17

1.16

### Surface to volume contribution



#### Surface effect is important No big difference between the two interactions

#### Production of exotic nuclei beyond the drip lines 50 F 10<sup>3</sup> 45 E 40<sup>F</sup> PLF <sup>48</sup>Ca+<sup>48</sup>Ca 10<sup>2</sup> Ξ 35 10 30 Z=28 25 20 Z=20 50E 10<sup>3</sup> 45 15 primary 40 10 10<sup>2</sup> N=28 Ξ 35 10 30 30 10 20 Z=28 25 N=50 1 20 Z=20 10<sup>-1</sup> 15 10 N=28 10<sup>-2</sup> N-20 10<sup>-3</sup> 20 30 40 50 60 10

NuSym2014, Liverpool

# Spectroscopy of <sup>29</sup>Si



- Experimental reconstruction (correlation techniques)
- Study excited states of exotic nuclei
- Study cluster states in nuclei (in progress)

# Signals of boson condensation and fermion quenching



- VAMOS to select QP + INDRA to reconstruct its identity and E\*
- Study of fluctuations of multiplicity and quadrupole momentum accounting for the QUANTAL nature of the probe particle (A. Bonasera et al)
- Determine Excitation energy, density, temperature event-by-event

# Signals of boson condensation and fermion quenching in the dilute phase?

(P. Marini et al., in preparation)

# **Summary and Conclusions**

- Exploration of  $E_{sym}(\rho)$  with HI-Collisions (<sup>48,40</sup>Ca+<sup>48,40</sup>Ca)
- Observables : isotopic distribution & isoscaling
- Accessing the symmetry energy
  - Take into account the secondary effects
  - Primary experimental isotopic distributions
    - Zprimay distributions were reconstructed experimentally
    - Aprimary neutrons distributions reconstructed exp. but need to take into account the effect of neutron emission on the Apr neutrons distributions
    - Staggering effects are washed with this reconstruction
- Both methods (isoscaling and isotopic distributions) can be used to extract the symmetry energy term if applied for primary quantities
- Esym was extracted for peripheral collisions, the values obtained are consistent with the value at normal density :
- work is in progress for central collisions

A. Chbihi, G. Verde, E. Bonnet, J.D. Frankland, M. Boisjoli, P. Marini, J. Moisan, B. Sorgunlu, F. Rejmund, M. Rejmund, J.P. Wieleczko, Sarmishtha Bhattacharya, P. Napolitani GANIL, CEA, IN2P3-CNRS, FRANCE INFN, Catania, ITALY B. Borderie, E. Galichet, N. Le Neindre, M.F. Rivet IPN Orsay, IN2P3-CNRS, FRANCE R. Dayras, L. Nalpas, C. Volant DAPNIA/SPhN, CEA Saclay, FRANCE P. Marini, M. Boisjoli, P. Wigg, Catex, FRANCE Institut de Physique N. A. Ono Department of Physics, Tohoku University, Sendai, JAPAN R. Roy Université de Laval, Quebec, CANADA W. Trautmann, J. Lukasik GSI, D-64291 Darmstadt, GERMANY E. Rosato, M. Vigilante Dipartimento di Scienze Fisiche, Un. Federico II, Napoli, ITALY M. Bruno, M. D'Agostino, E. Geraci, G. Vannini INFN and Dipartimento di Fisica, Bologna ITALY L. Bardelli, G. Casini, A. Olmi, S. Piantelli, G. Poggi INFN and Dipartimento di Fisica, Firenze, ITALY F. Gramegna, G. Montagnoli INFN, Laboratori Nazionali di Legnaro, ITALY U. Abbondanno INFN, Trieste, ITALY NUS Vm2021,4G LTveppool 20 A. Chbihi National Institute for Physics and Nuclear Engineering, Bucarest-Maguerele, ROMANIA

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