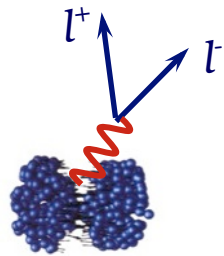


Isolation of excess pairs an experimental approach

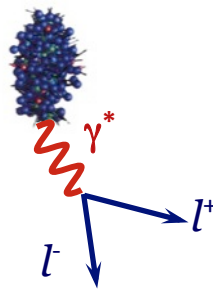
EMMI RRTF
10 October 2013
Tetyana Galatyuk

Probing EM structure of dense/hot hadronic matter

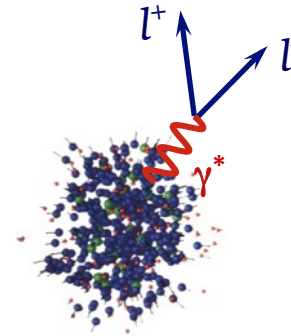
Sources of lepton pairs



first-chance collisions



hot/dense fireball



freeze-out

Characteristics of dilepton rates (cocktail)



Characteristic features of dilepton invariant mass spectra

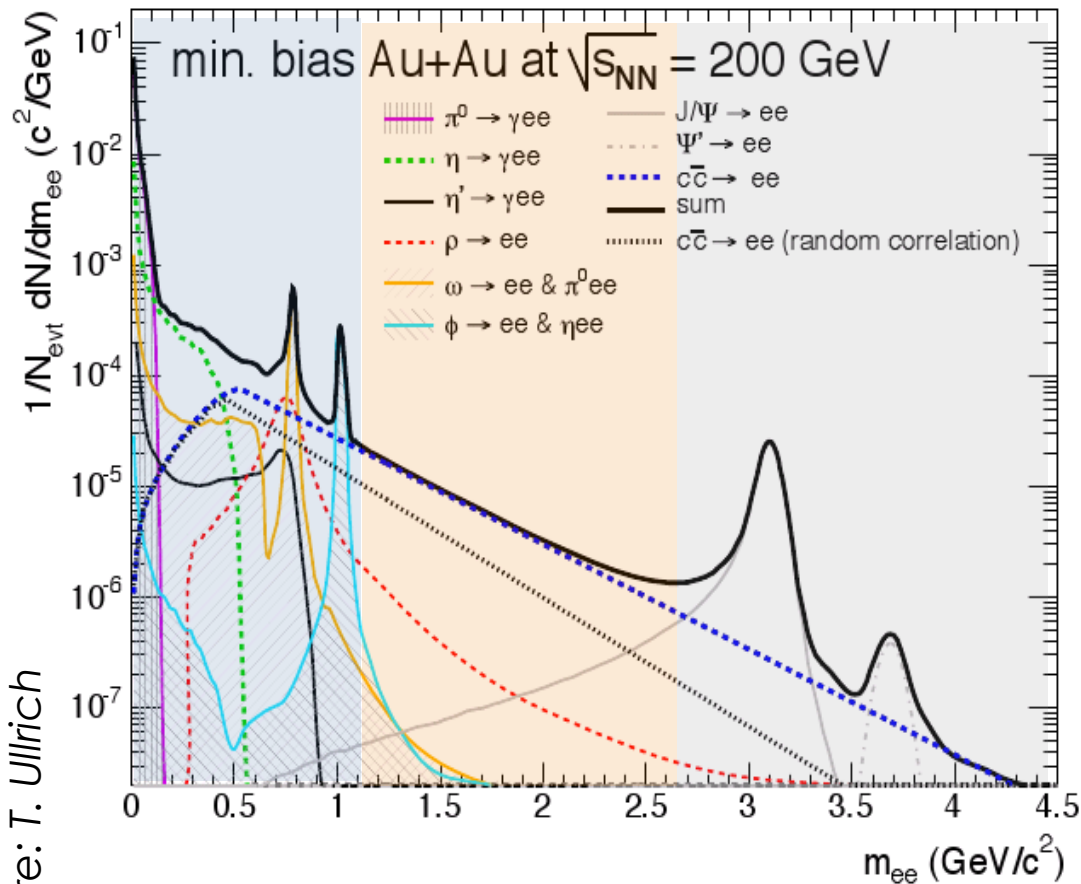


Figure: T. Ullrich



Low mass:

- continuum enhancement?
- in-medium modification of vector mesons \rightarrow link to the chiral symmetry restoration?

Intermediate mass:

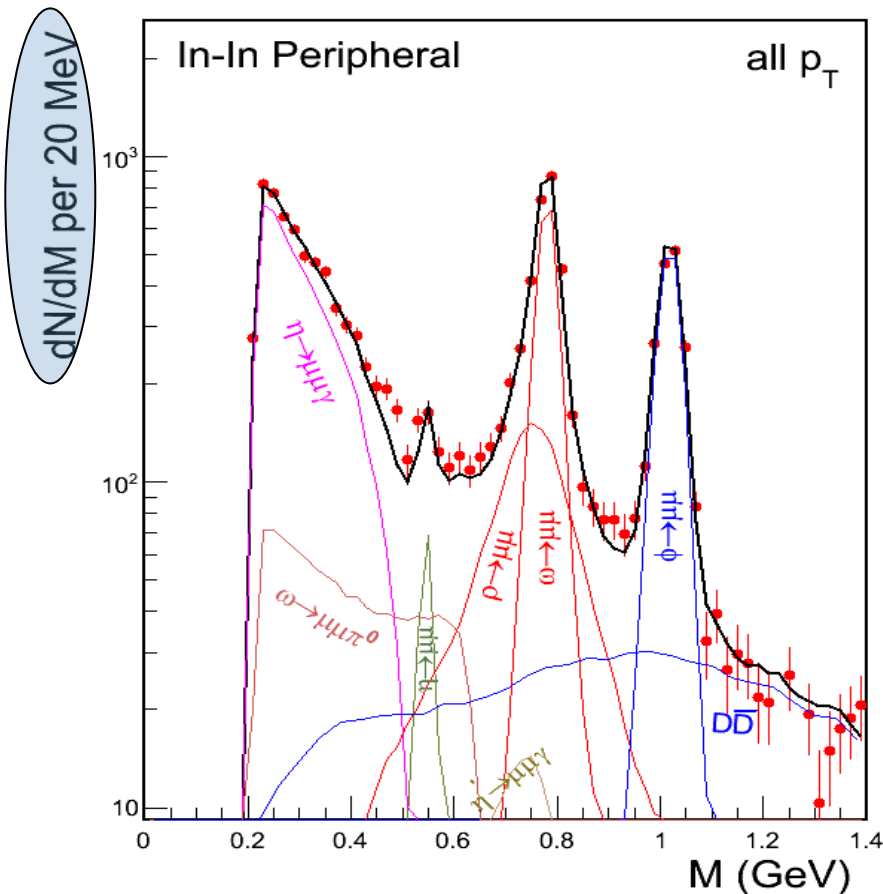
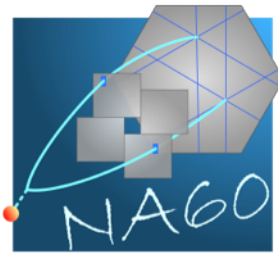
- QGP thermal radiation?
- heavy-flavor modification?

High mass:

- J/ψ suppression ?
enhancement ?
- Drell-Yan, primordial emission

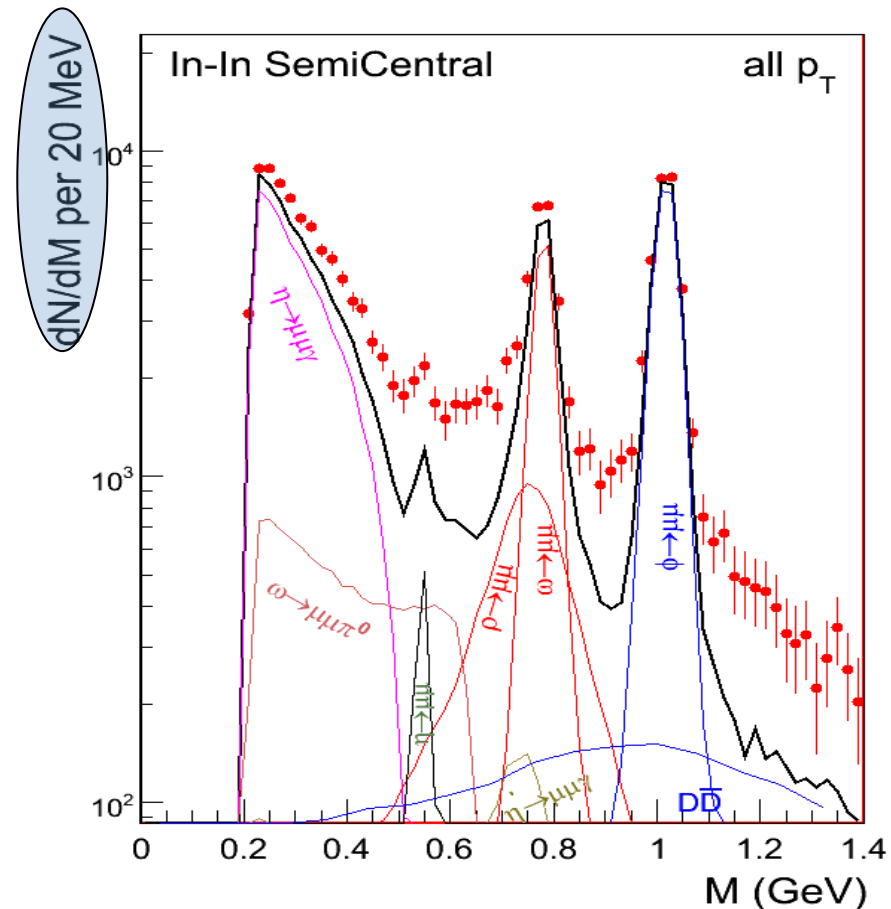
Understanding the mass spectra

S. Damjanovic, Trento 2010



Peripheral data

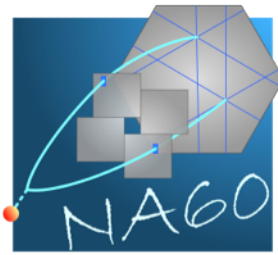
well described by meson decay 'cocktail' (η , η' , ρ , ω , ϕ) and $D\bar{D}$



More central data

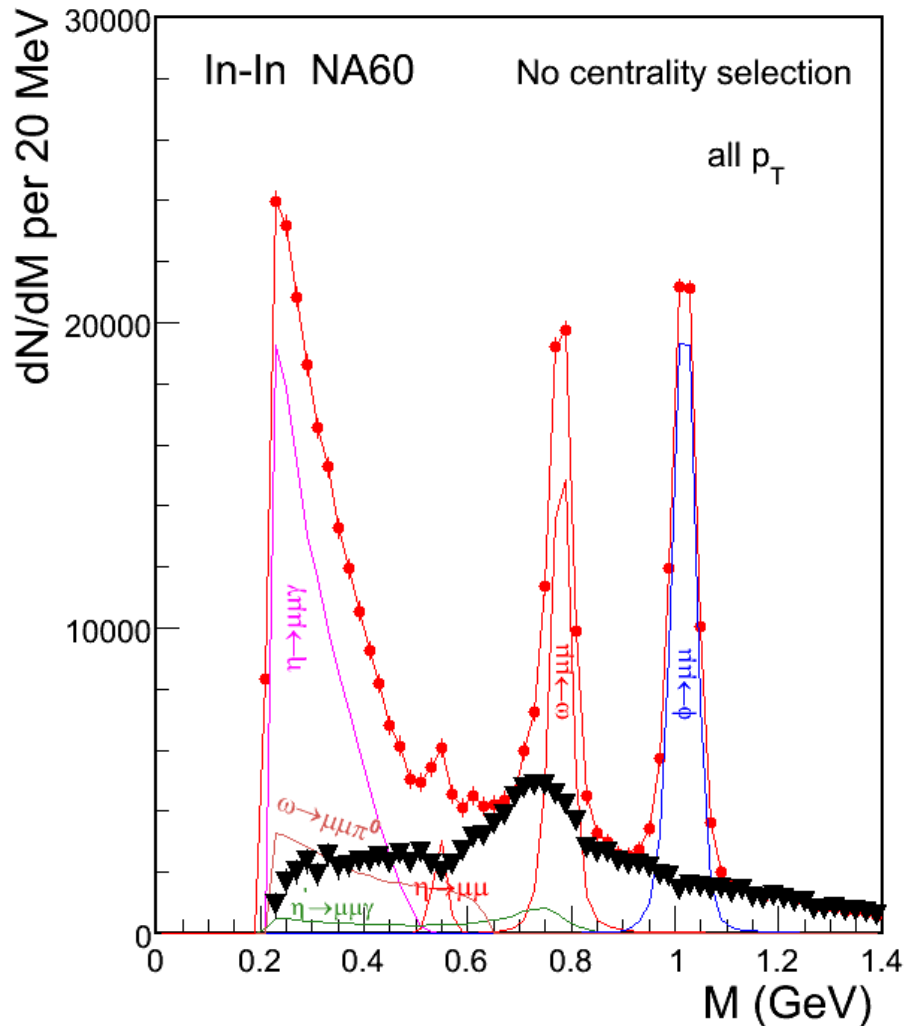
Clear excess of data above decay 'cocktail'. But, what is the spectral shape of the excess?

Isolation of excess dimuons



S. Damjanovic, Trento 2010

Phys. Rev. Lett. 96 (2006) 162302



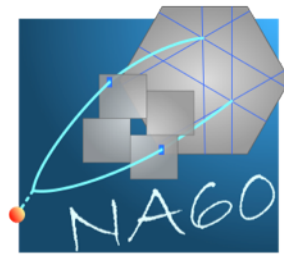
ω and f : fix yields such as to get, after subtraction, a smooth underlying continuum

η : fix yield at $p_T > 1$ GeV profiting from the very high sensitivity of the spectral shape of the Dalitz decay to any underlying admixture from other sources; lower limit from peripheral data

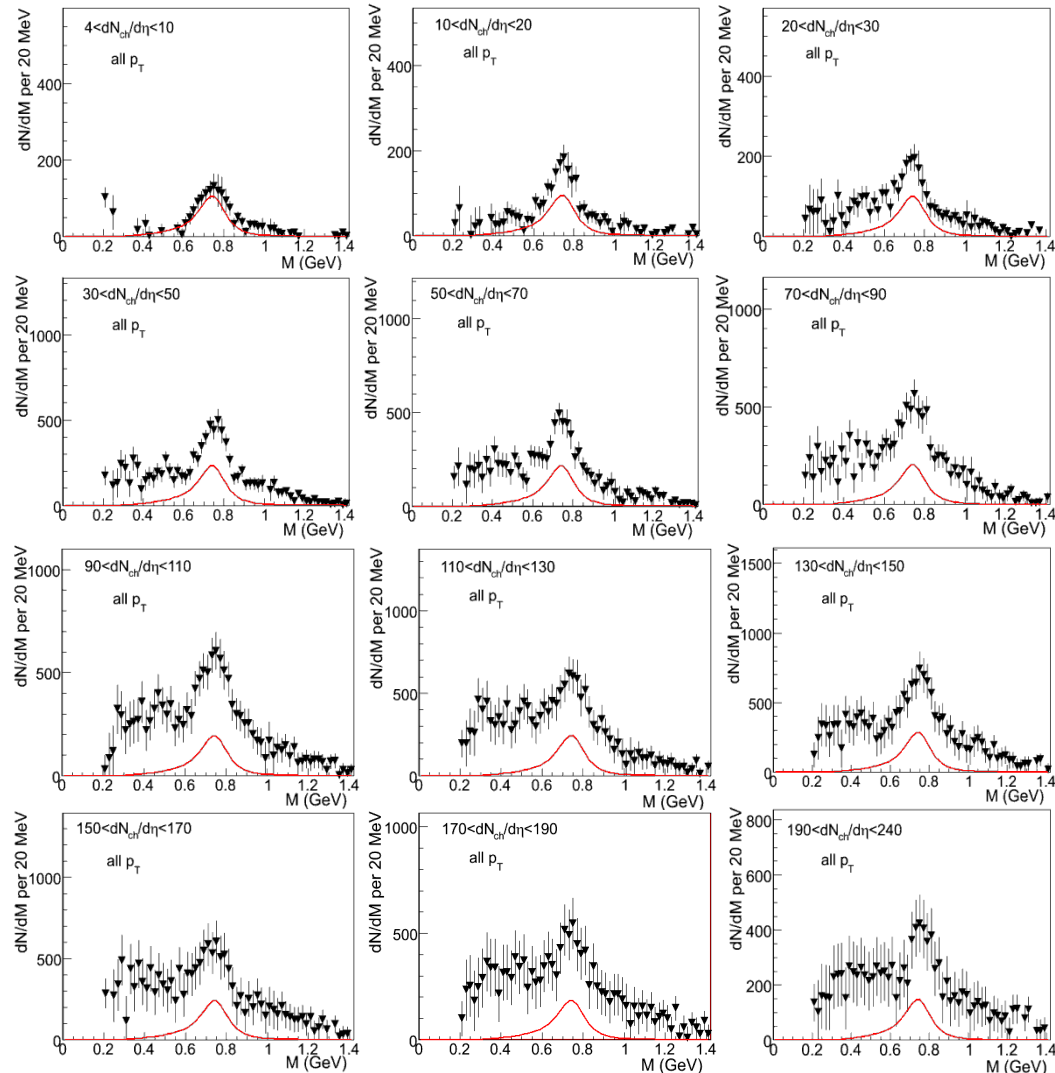
accuracy 2-3%, but results robust to mistakes even at the 10% level

Excess mass spectra in 12 centrality bins

S. Damjanovic, Trento 2010

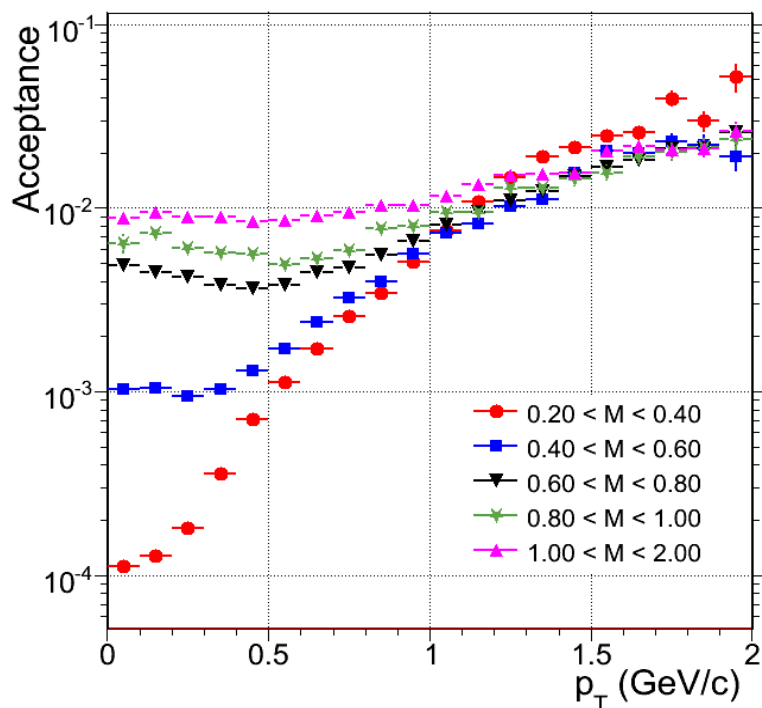
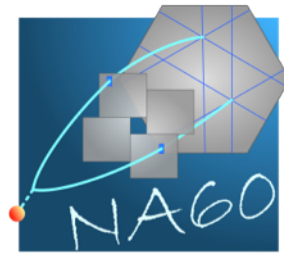


Eur.Phys.J.C 49 (2007) 235



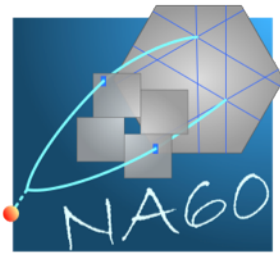
Acceptance corrections

S. Damjanovic, Trento 2010

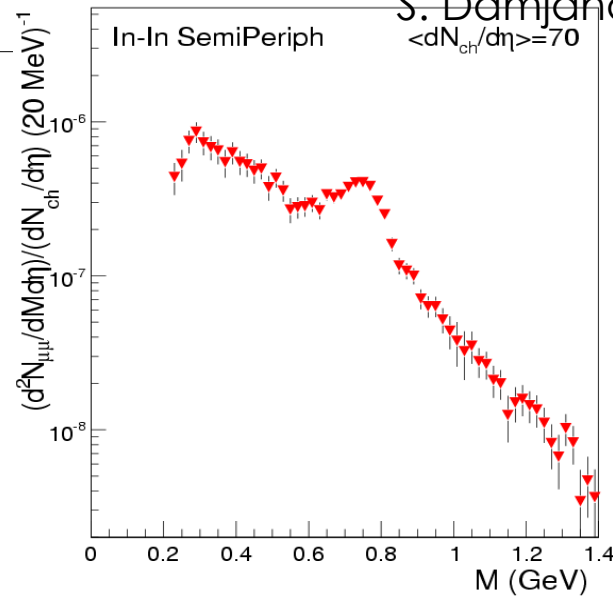
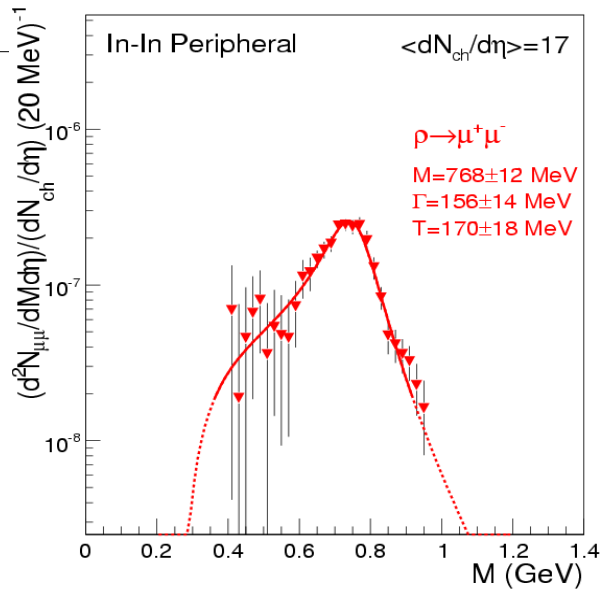


- reduce 4-dimensional acceptance correction in M - p_T - y - $\cos\Theta_{CS}$ to (mostly) 2-dimensional corrections in pairs of variables. Example M - p_T , using **measured** y distributions and **measured** $\cos\Theta_{CS}$ distributions as an input; same for other pairs (iteration)
- requires separate treatment of the excess and the other sources, due to differences in the y and the $\cos\Theta_{CS}$ distribution
- acceptance vs. M , p_T , y , and $\cos\Theta$ understood to within $<10\%$, based on a detailed study of the peripheral data

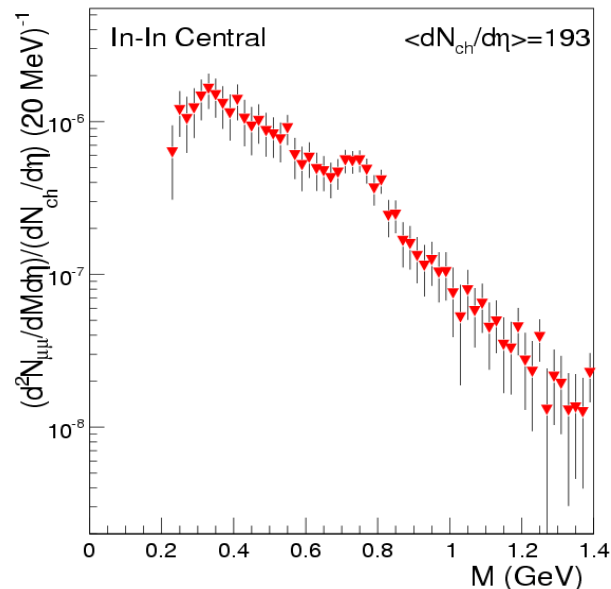
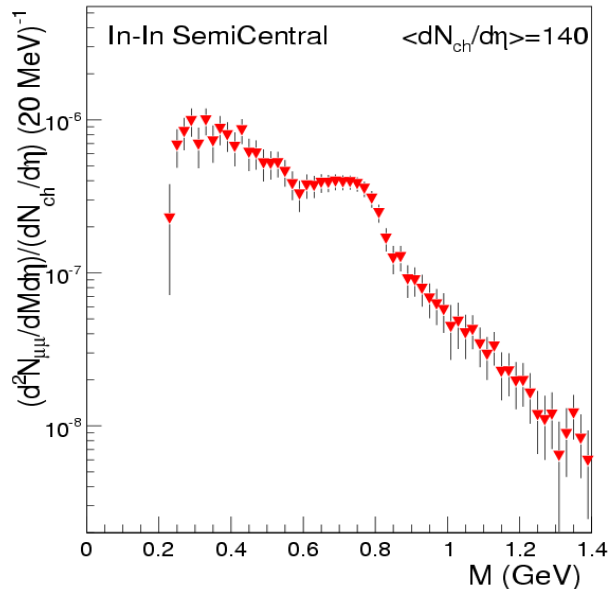
Centrality dependence of excess mass spectra



S. Damjanovic, Trento 2010

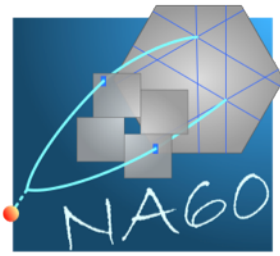


very fast evolution from vacuum ρ to Planck-like spectra



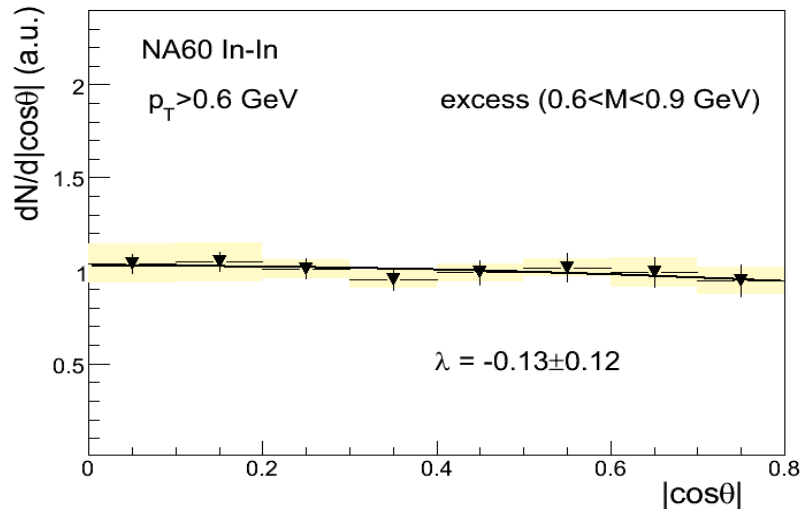
increasing masking of residual freeze-out ρ

Angular distributions



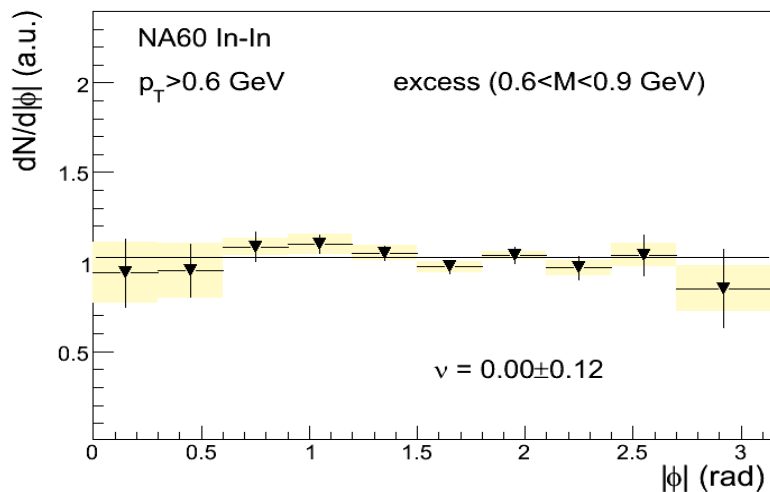
S. Damjanovic, Trento 2010

Phys. Rev. Lett. 102 (2009) 222301

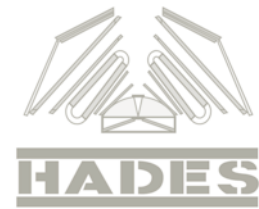


$$\frac{1}{\sigma} \frac{d\sigma}{d \cos \theta d\phi} = \left(1 + \lambda \cos^2 \theta + \mu \sin 2\theta \cos \phi + \frac{\nu}{2} \sin^2 \theta \cos 2\phi \right)$$

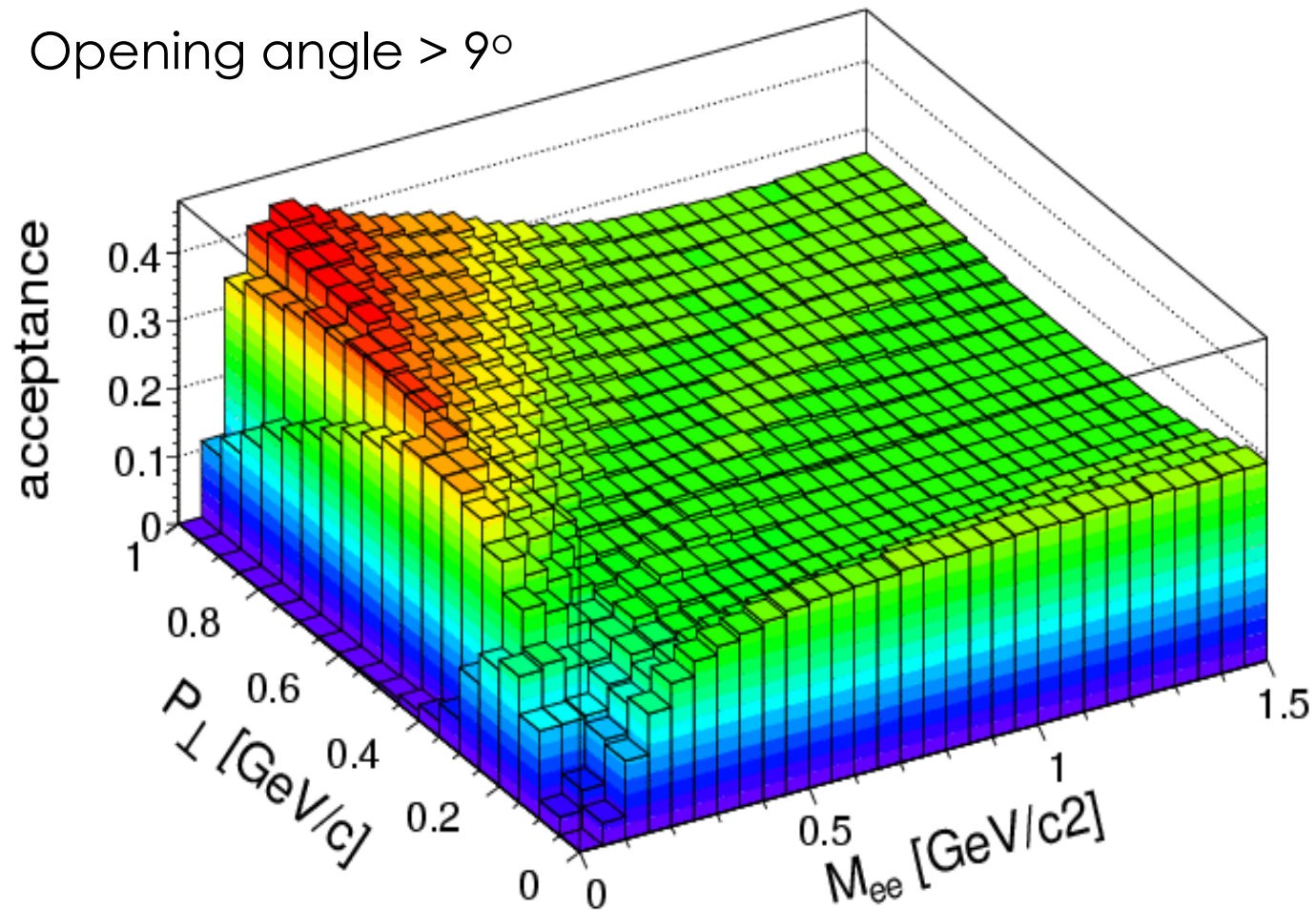
Zero polarization within errors



HADES pair acceptance



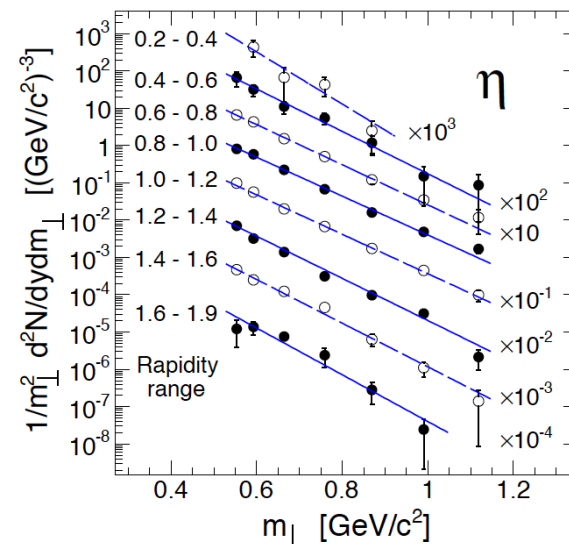
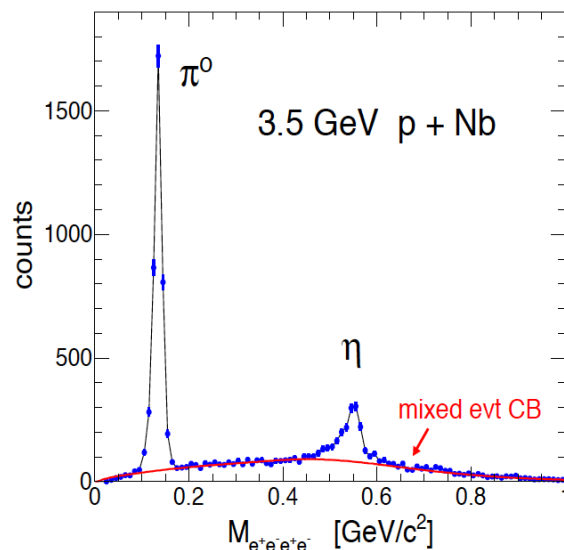
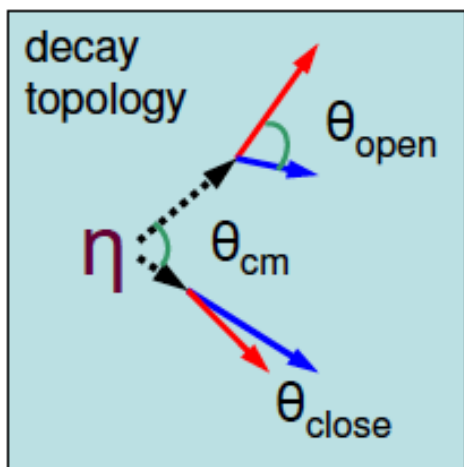
Opening angle $> 90^\circ$



Reconstruction of π^0 and η decays through conversion

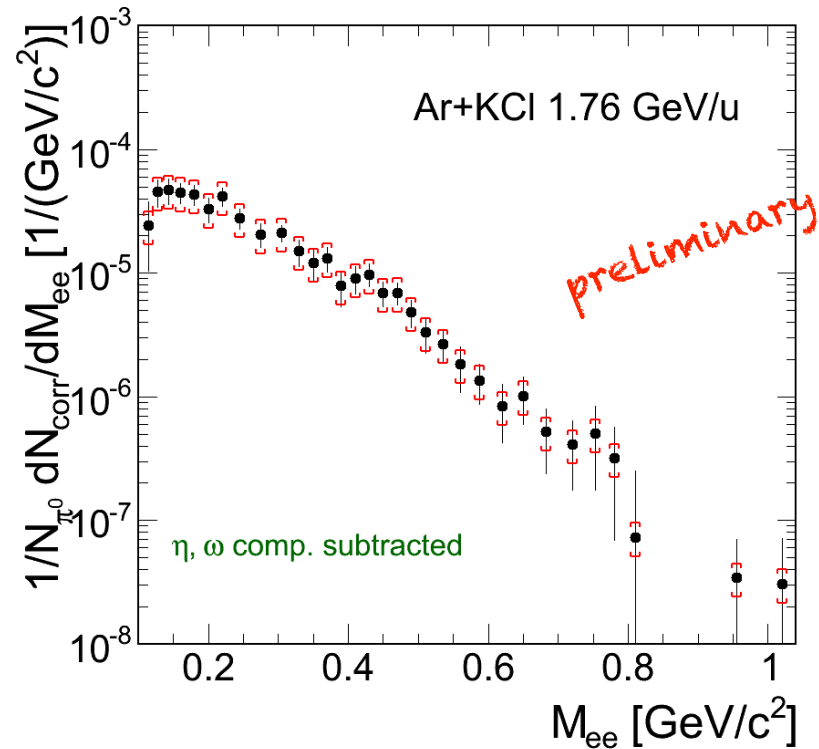
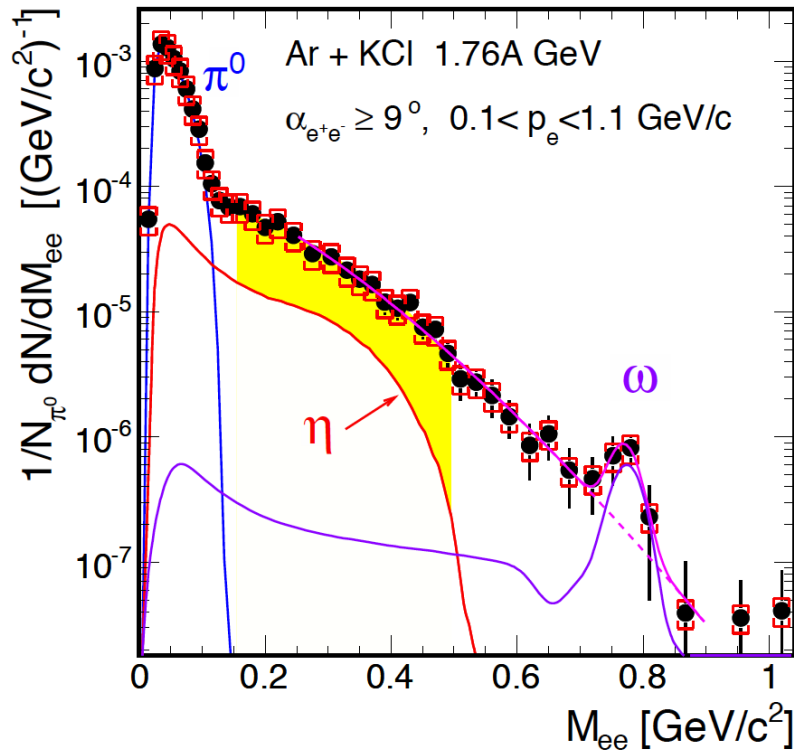
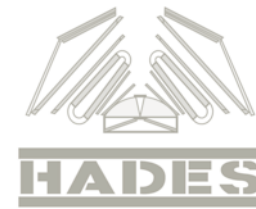
September 2008 run: 3.5 GeV p+Nb: $\rightarrow 9.2 \cdot 10^9$ LVL1 events:

- Meson $\rightarrow \gamma e^+e^-$ + pair conversion of photon: meson $\rightarrow e^+e^- e^+e^-$
- Meson $\rightarrow \gamma\gamma$ + conversion of both photons: meson $\rightarrow e^+e^- e^+e^-$



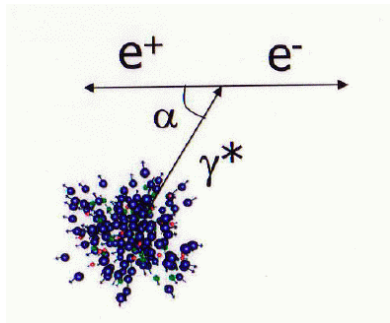
Nearly full phase-space coverage for η detection!

Ar+KCl data, $E_{\text{kin}}=1.76$ GeV/u

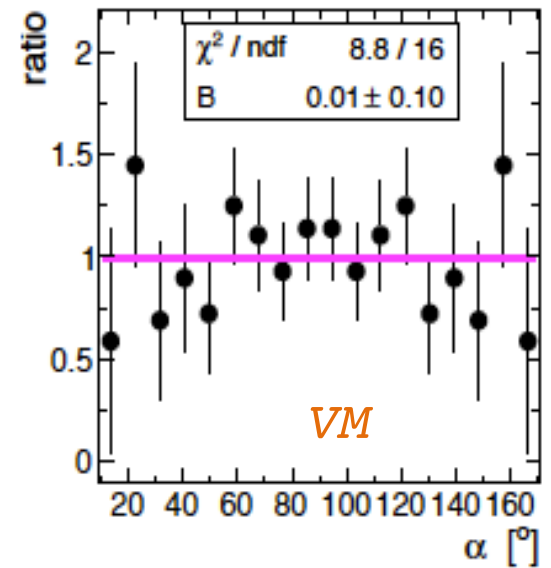
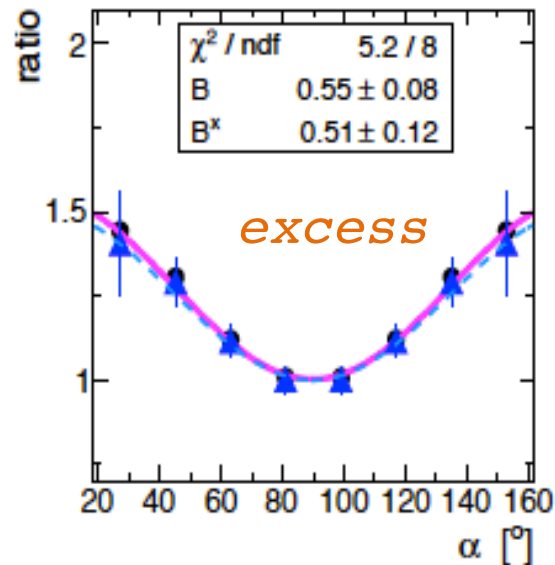
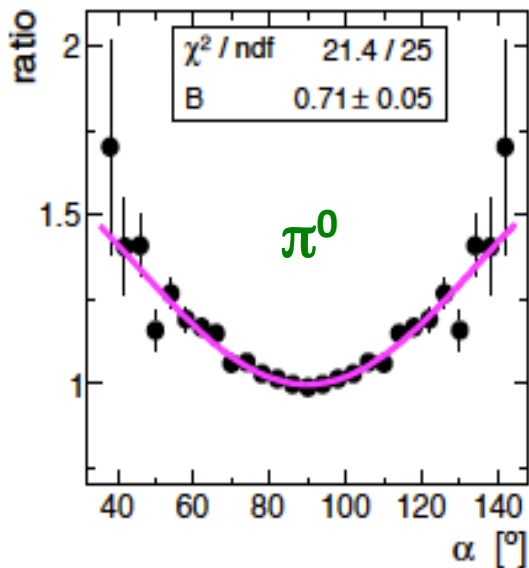
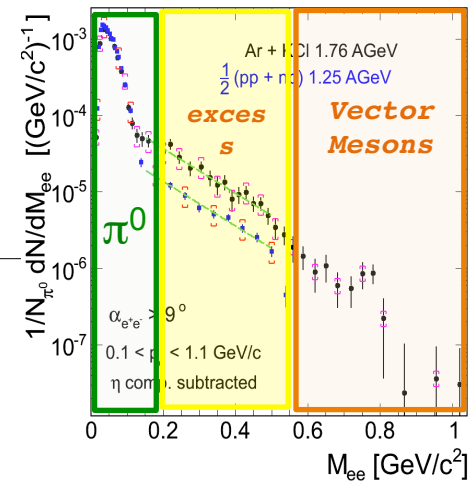


Phys.Rev.C84:014902,2011

e+e- angular distributions



$$\frac{dN}{d\alpha} = A(1 + B \cos^2 \alpha)$$

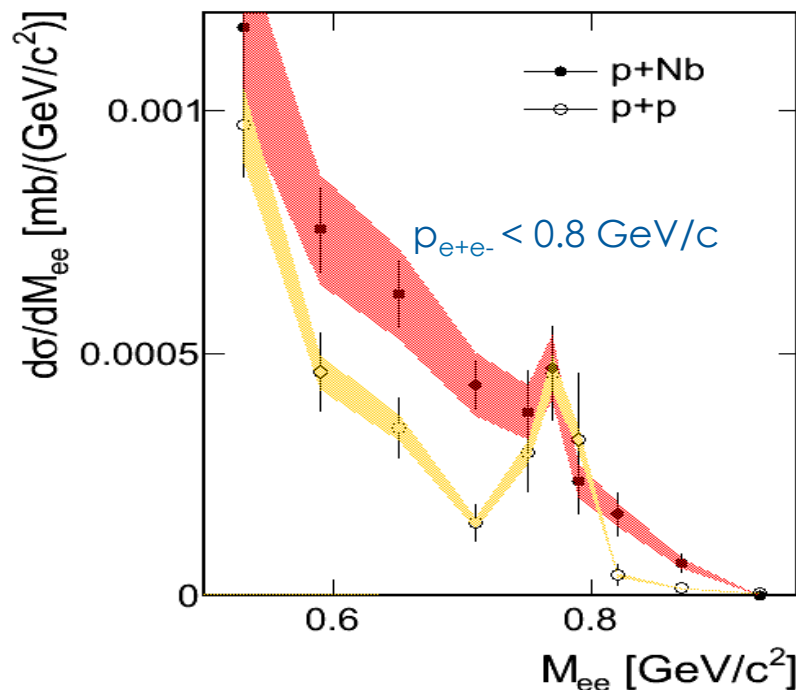
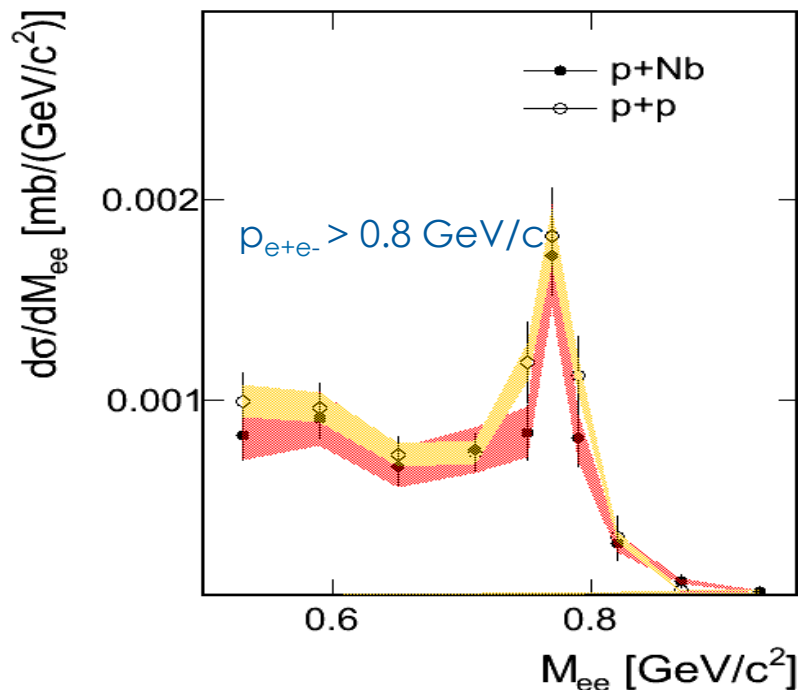


Excess has polarization consistent with Δ

Not clear \rightarrow need more statistic
($B = 0$ at SPS)

Omega in cold nuclear matter

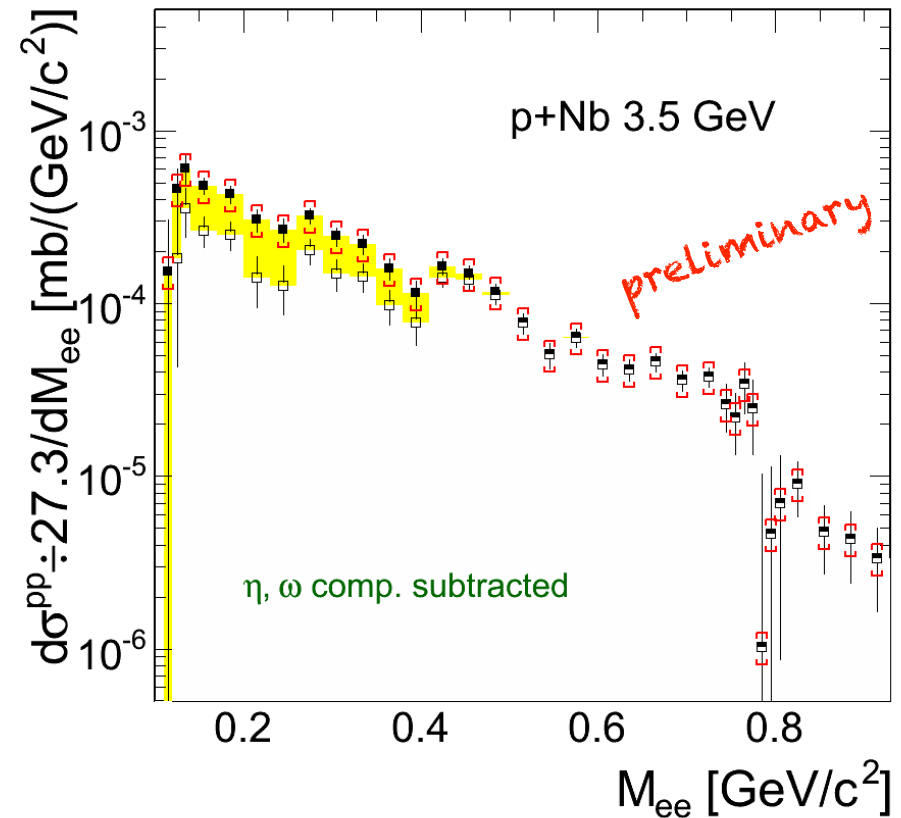
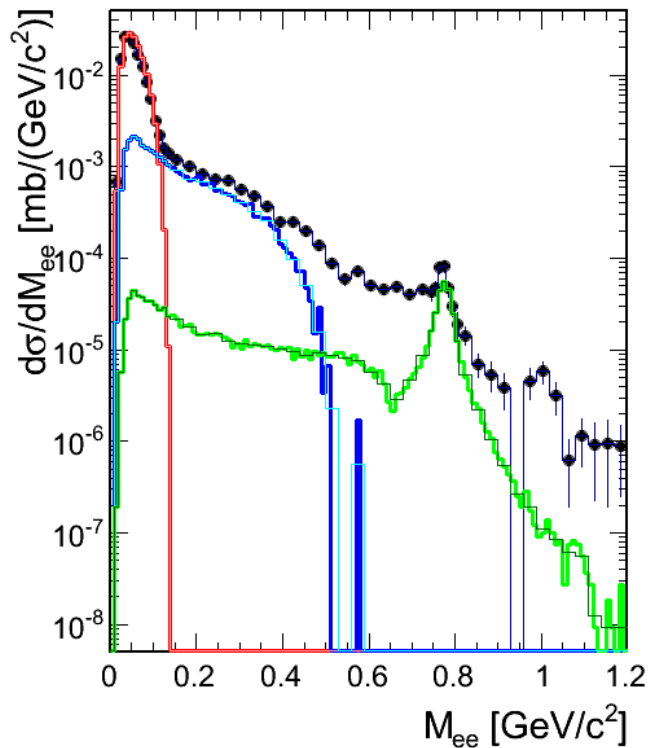
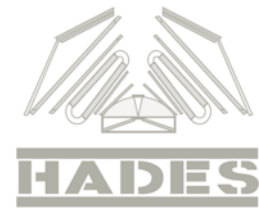
Phys.Lett. B715 (2012) 304-309



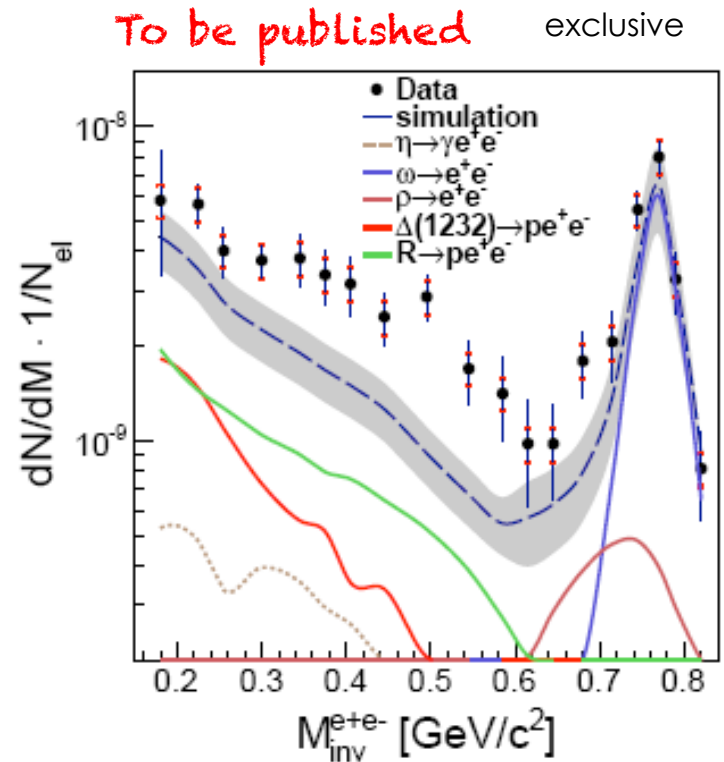
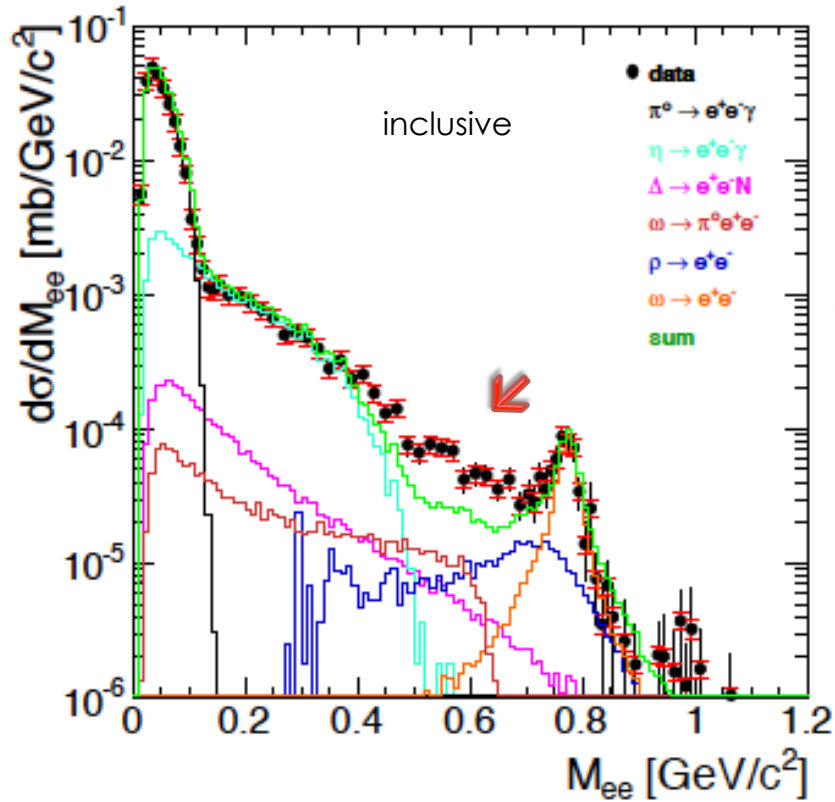
Selection on momentum of the decaying ω meson

- First measurement of in-medium vector meson decays in the relevant momentum region (high-momentum ω mesons “decouple” from the medium)
- ω suppressed, in-medium decays buried under ρ -like contribution

p+Nb data, $E_{\text{kin}}=3.5$ GeV



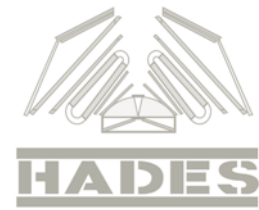
pp reference



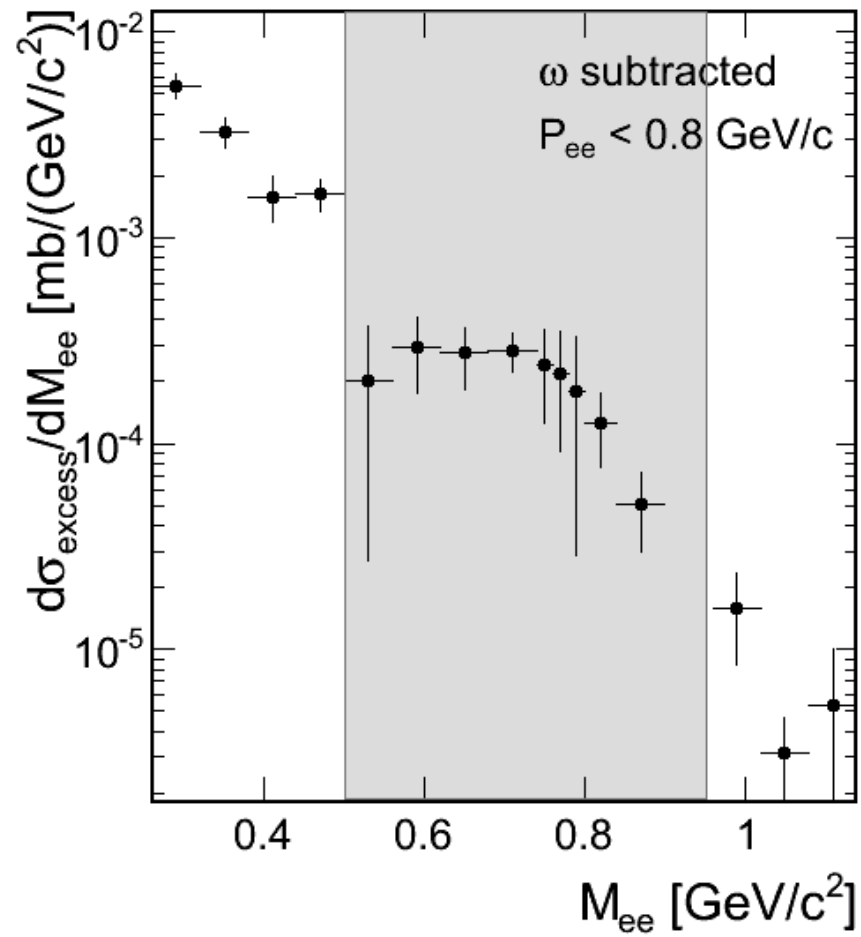
PDG Entry 2012
 $BR(h \rightarrow e^+e^-) < 5.6 \times 10^{-6}$ (90% CL)

- Coupling of ρ to baryonic resonances
 → Cross check with hadronic final states needed!

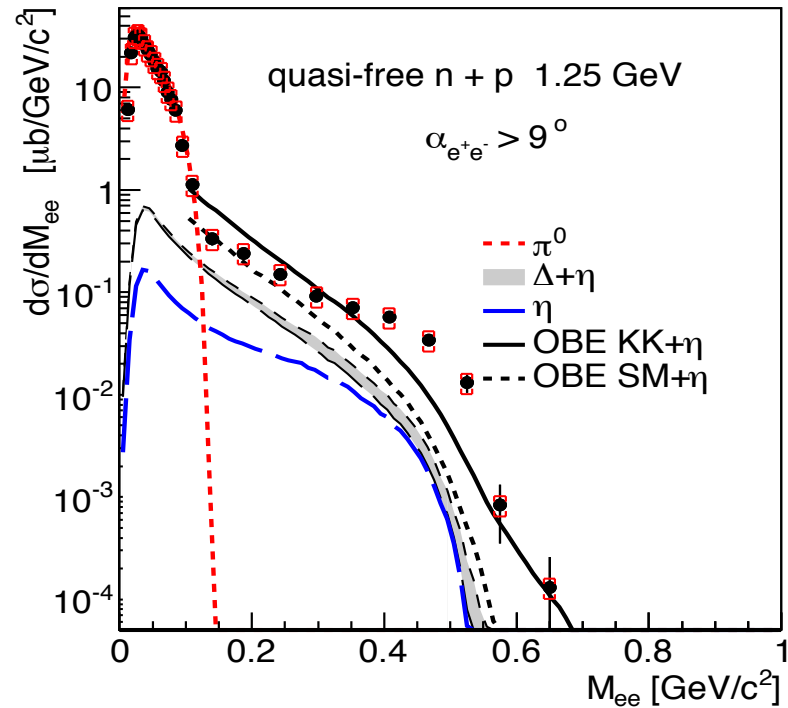
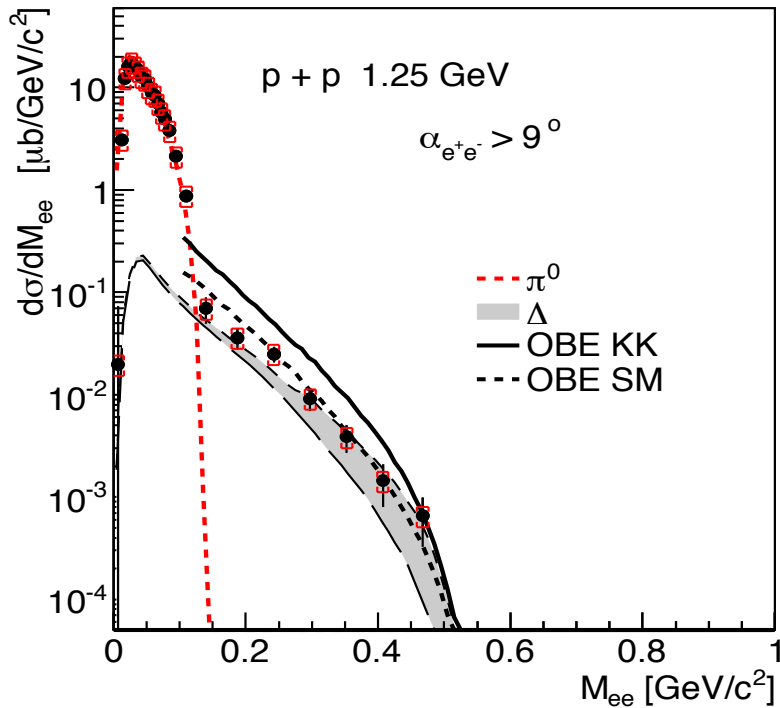
pNb – pp – ω



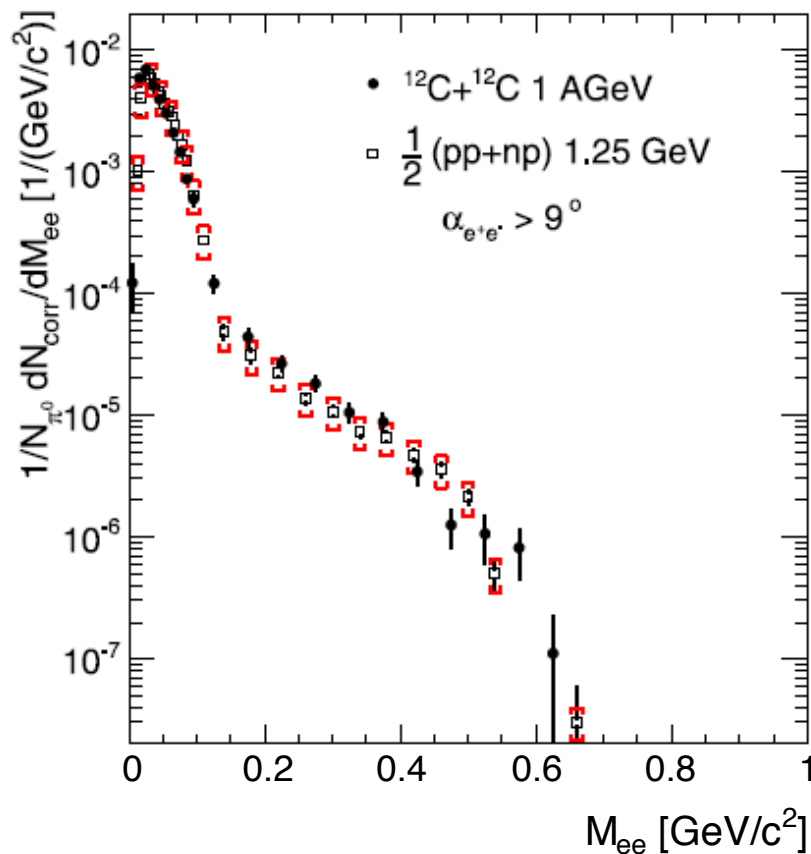
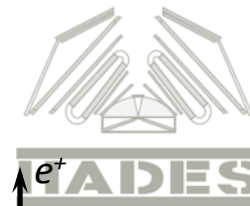
Phys.Lett. B715 (2012) 304-309



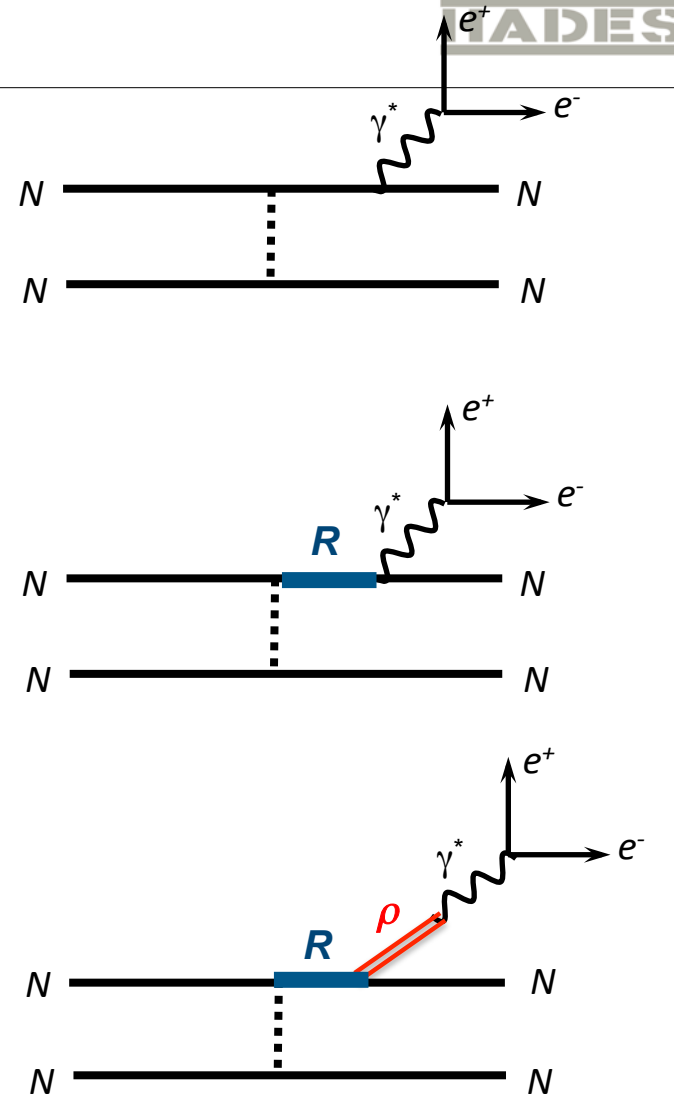
NN reference



Origin of the low-mass pair excess in C+C collisions

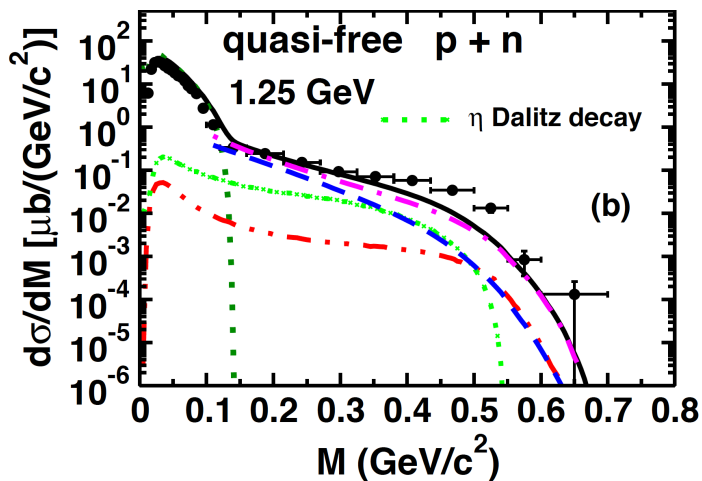
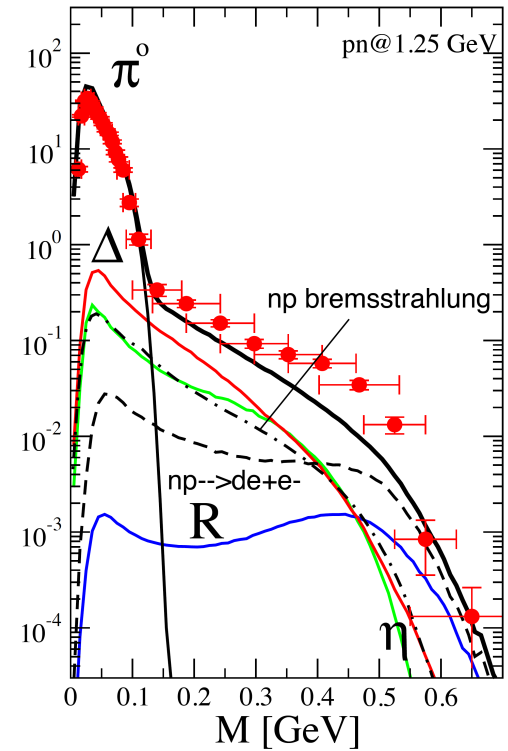
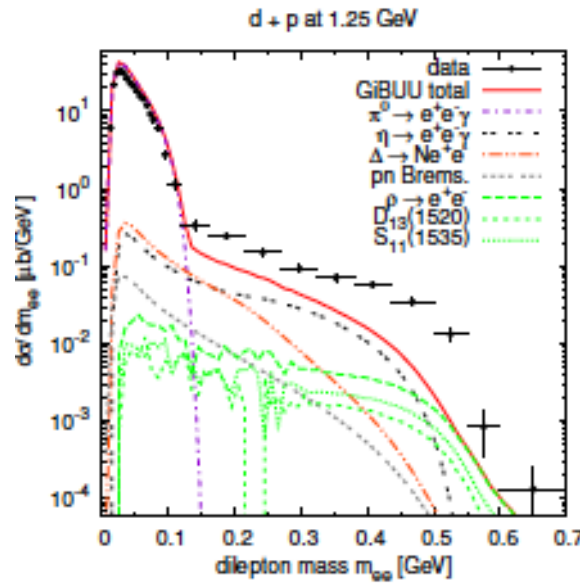
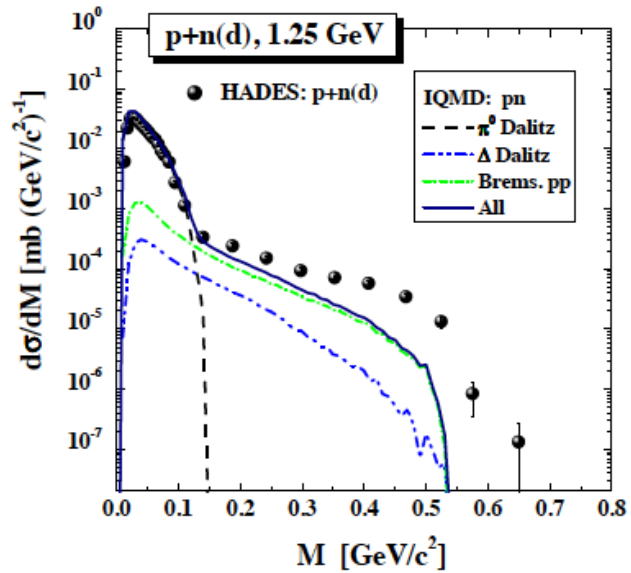


HADES: Phys. Lett. B 690 (2010) 118

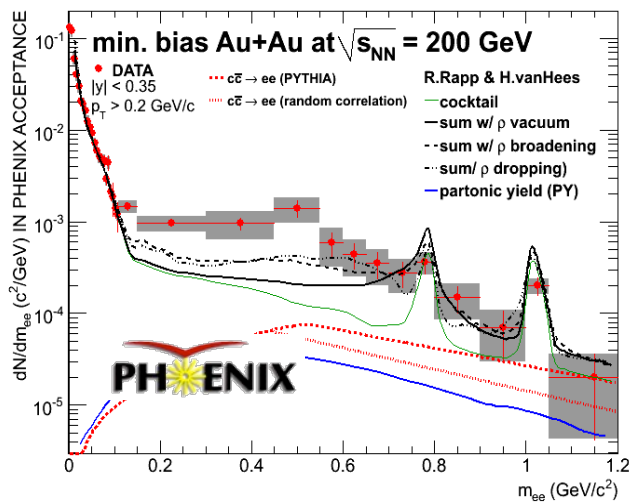
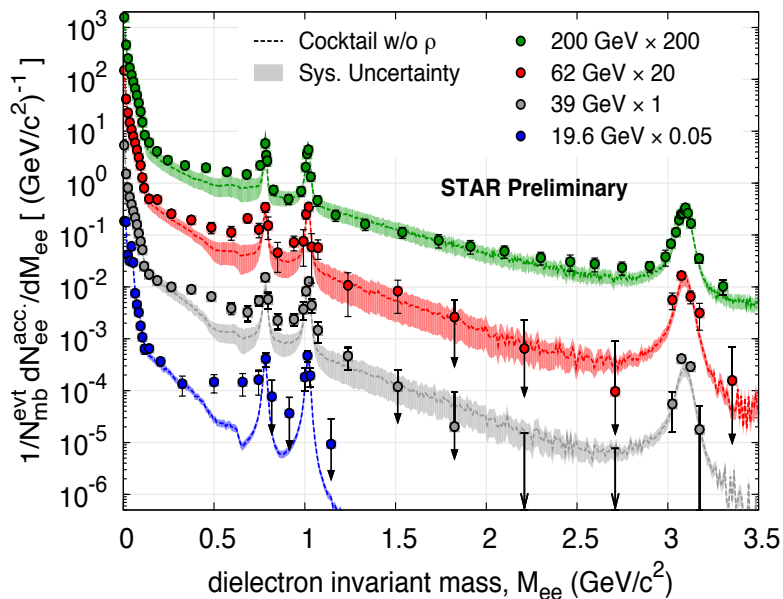
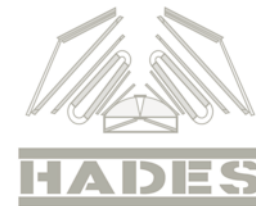


Baryonic contributions from NN "reference"

pn collisions

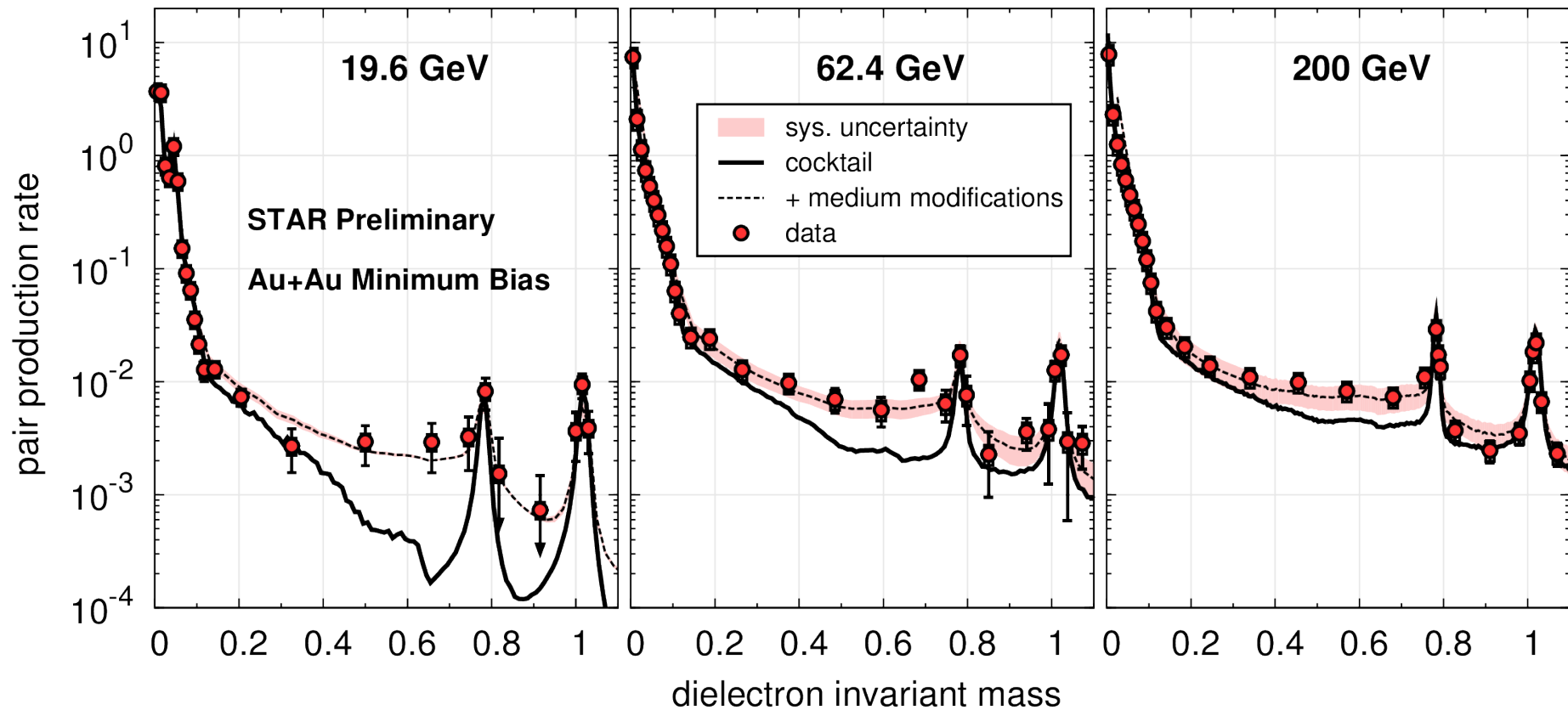


Dileptons at RHIC and LHC

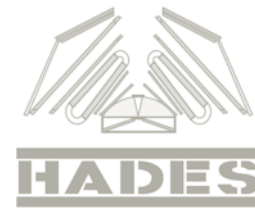


- SPS $\sqrt{s_{NN}} \leq 17.2$ GeV
 - net-baryon density: $\mu_B \sim 250$ MeV (at $T_{ch} \approx 160$ MeV)
- RHIC $\sqrt{s_{NN}} = 200$ GeV
 - $\mu_B \ll T$, i.e. vanishing net-baryon density
 - **Lattice QCD computations are most powerful!**
- Precision measurement from $\sqrt{s_{NN}} = 200$ GeV is absolutely needed to “calibrate” EM rates
- Wish: excess mass spectrum a-la NA60, i.e. subtracting all know sources, at $\sqrt{s_{NN}} = 200$ GeV
 - High precision measurements of charm
 - Statistic!

STAR Beam Energy Scan



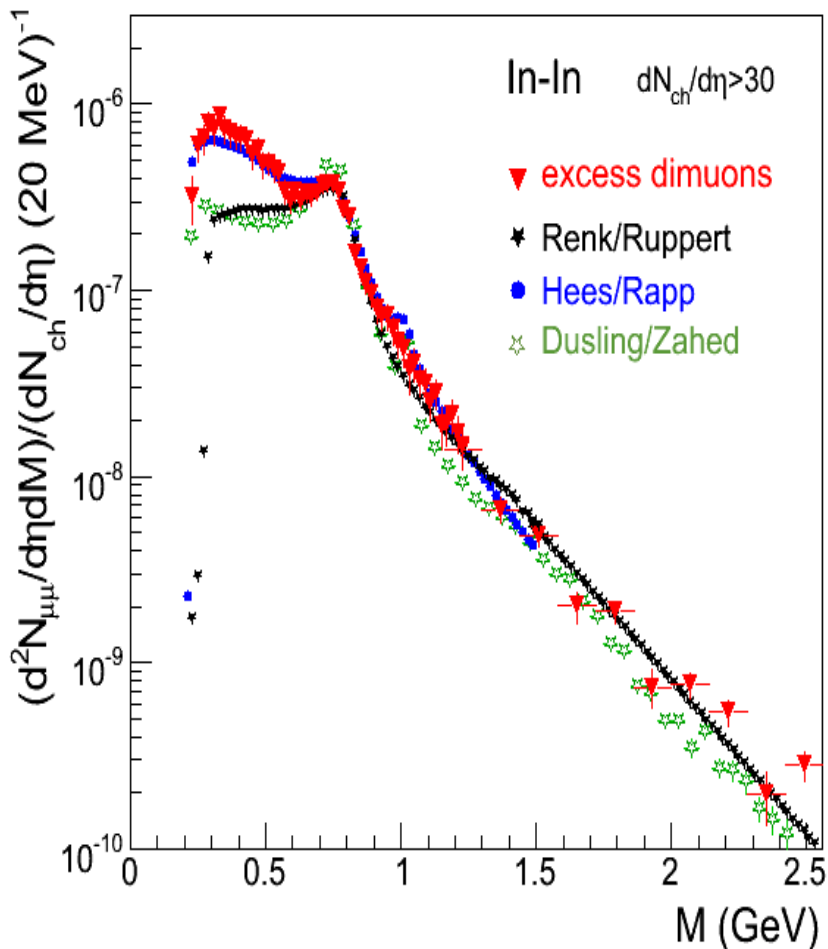
Acceptance corrected excess mass spectrum



[Eur. Phys. J. C 59 (2009) 607-623]

CERN Courier 11/ 2009, 31-35

Chiral 2010 , AIP Conf.Proc. 1322 (2010) 1-10



M > 1 GeV

~ exponential fall-off → 'Planck-like'

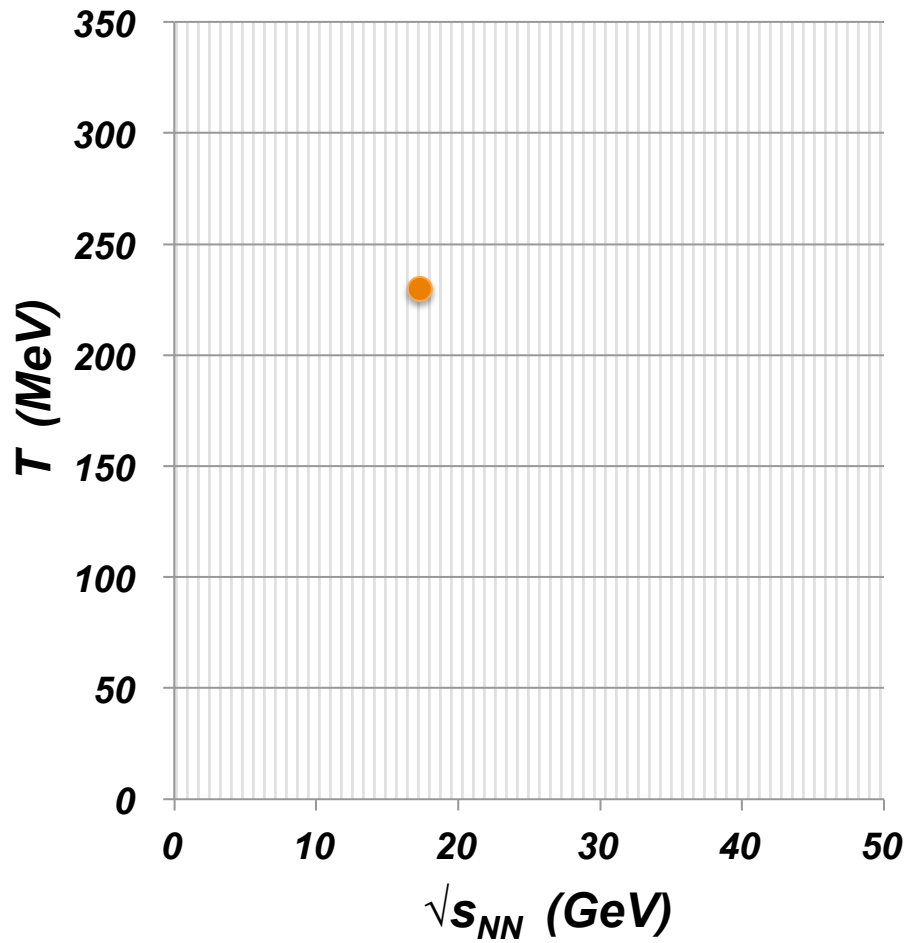
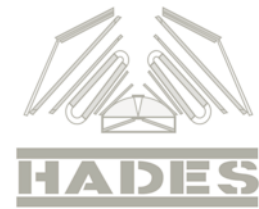
fit to $dN / dM \propto M^{3/2} \times \exp(-M / T)$

range 1.1-2.0 GeV: $T = 205 \pm 12 \text{ MeV}$

1.1-2.4 GeV: $T = 230 \pm 10 \text{ MeV}$

described by R/R, D/Z and H/R models

Future explorations



- Look for non-monotonic behavior of T as a function of $\sqrt{s_{NN}}$? (Note, $T < T_{\text{initial}}$)
- Appearance (disappearance) of QGP radiation
- What would pp and pA say?