



Erice: INTERNATIONAL SCHOOL OF  
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# SEARCH FOR $CE\nu NS$ WITH CRYOGENIC SAPPHIRE DETECTORS AT MINER

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Rupak Mahapatra**

**(on behalf of MINER collaboration)**

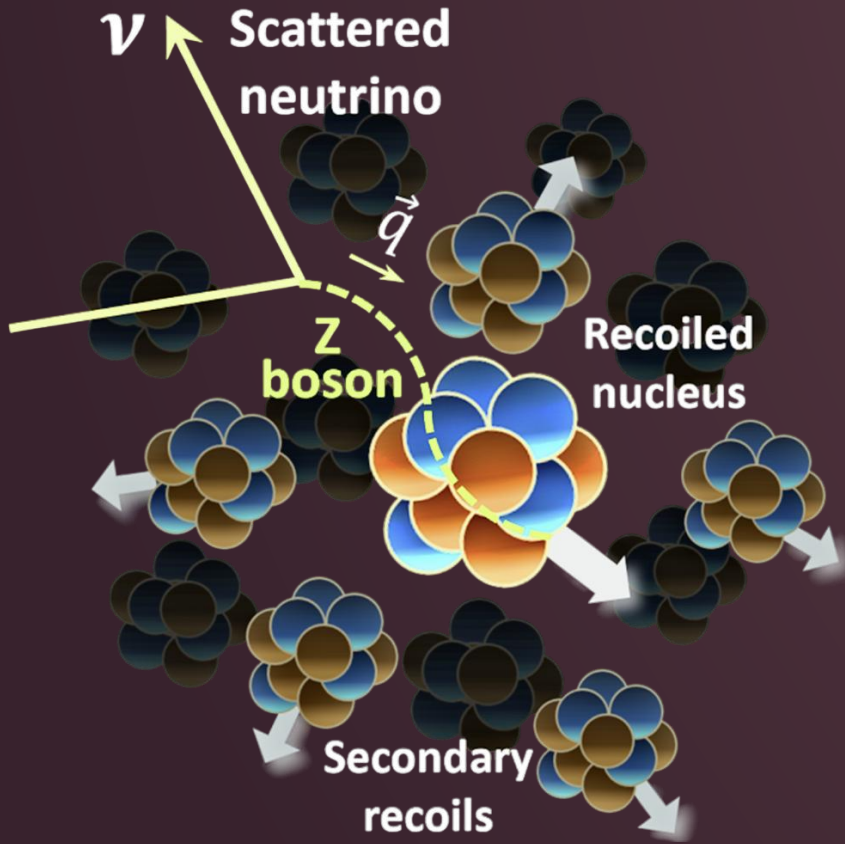
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- Introduction
- MINER experiment
- Low energy excess
- Prospects



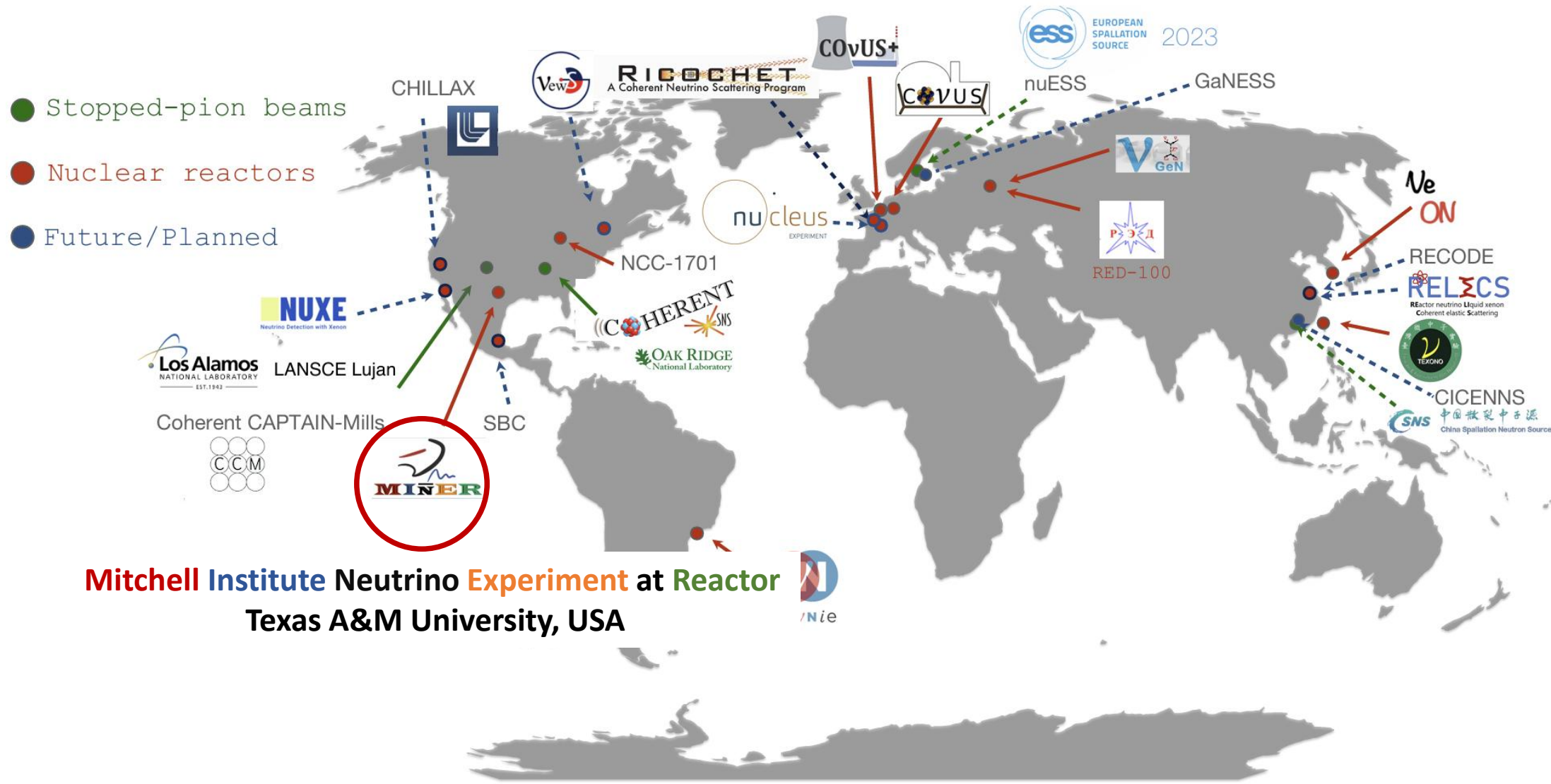
## CE $\nu$ NS: Coherent Elastic Neutrino Nucleus Scattering



- First predicted in 1974 by D. Freedman
- Neutrino scatters off the entire nucleus
- Low momentum transfer  $|\vec{q}| \leq 1/R_{\text{nucleus}}$
- Flavor independent neutral current process
- The SM cross-section scales as  $N^2$ ,  
 $N \rightarrow$  neutron number
  - Enhanced for heavy nuclei but recoil energy reduces
- Recoil energy  $\mathcal{O}(10 - 1000 \text{ eV})$
- Offers sensitivity to both SM and BSM physics.



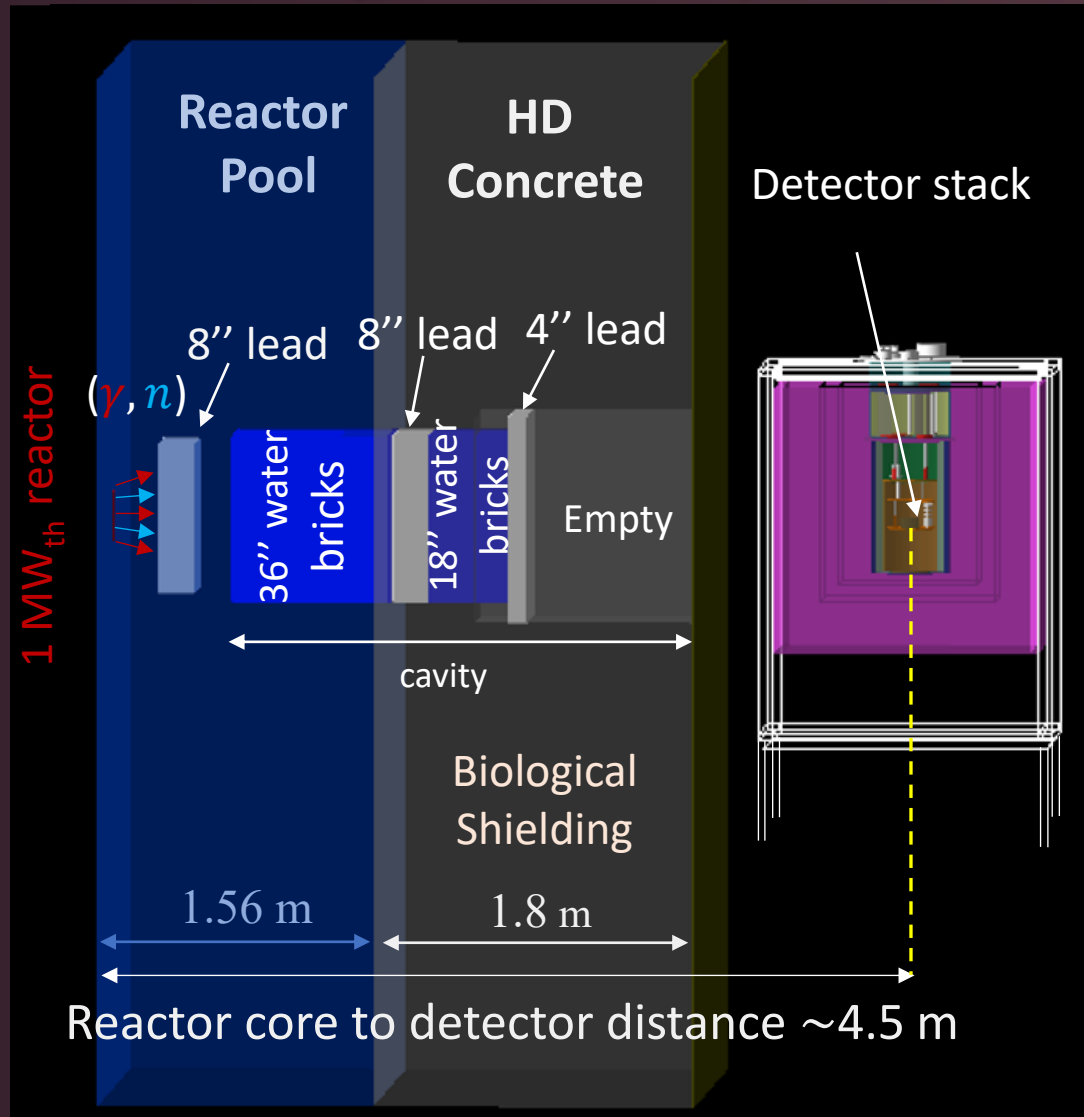
Adapted from C. Bonifazi, Neutrino 2022



**Mitchell Institute Neutrino Experiment at Reactor**  
**Texas A&M University, USA**

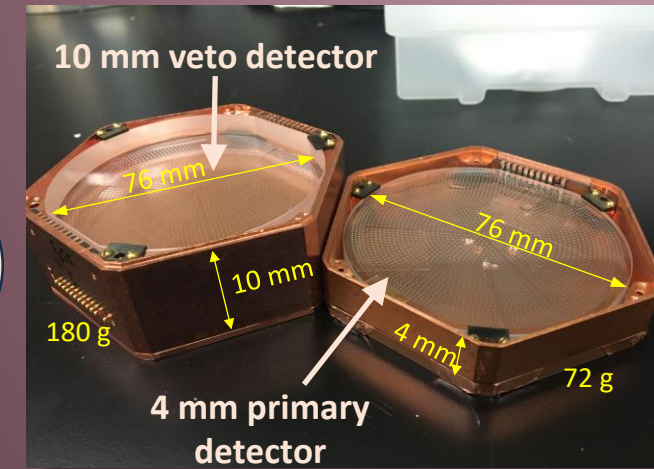
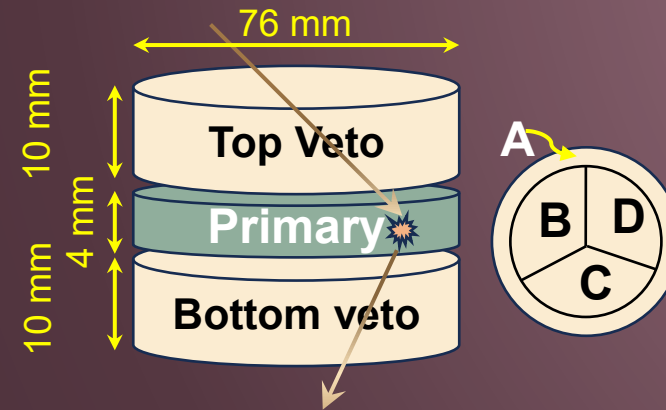
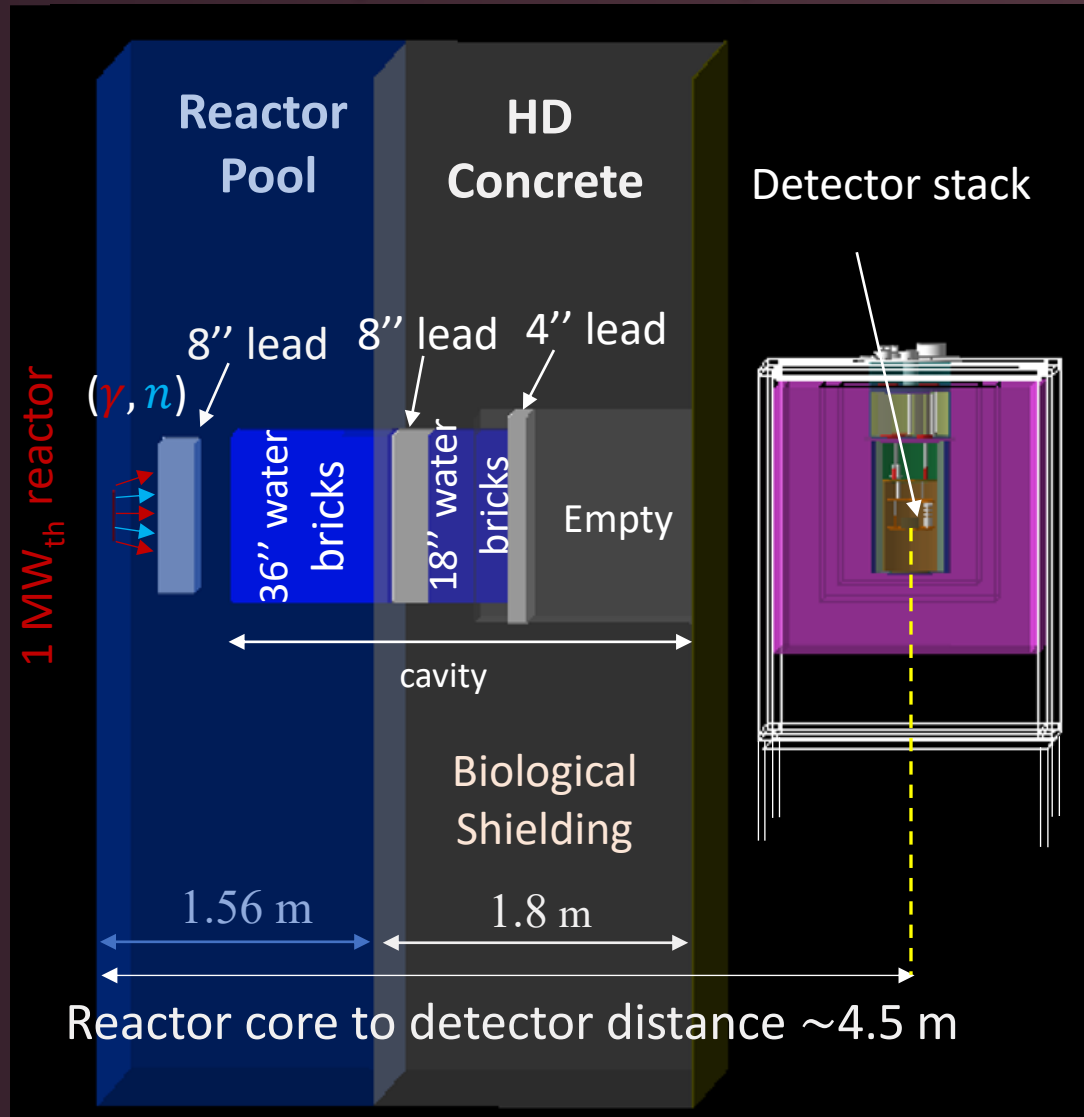
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## Experimental setup



- Reactor power: 1 MW<sub>th</sub>
- Baseline distance: 4.5 meters

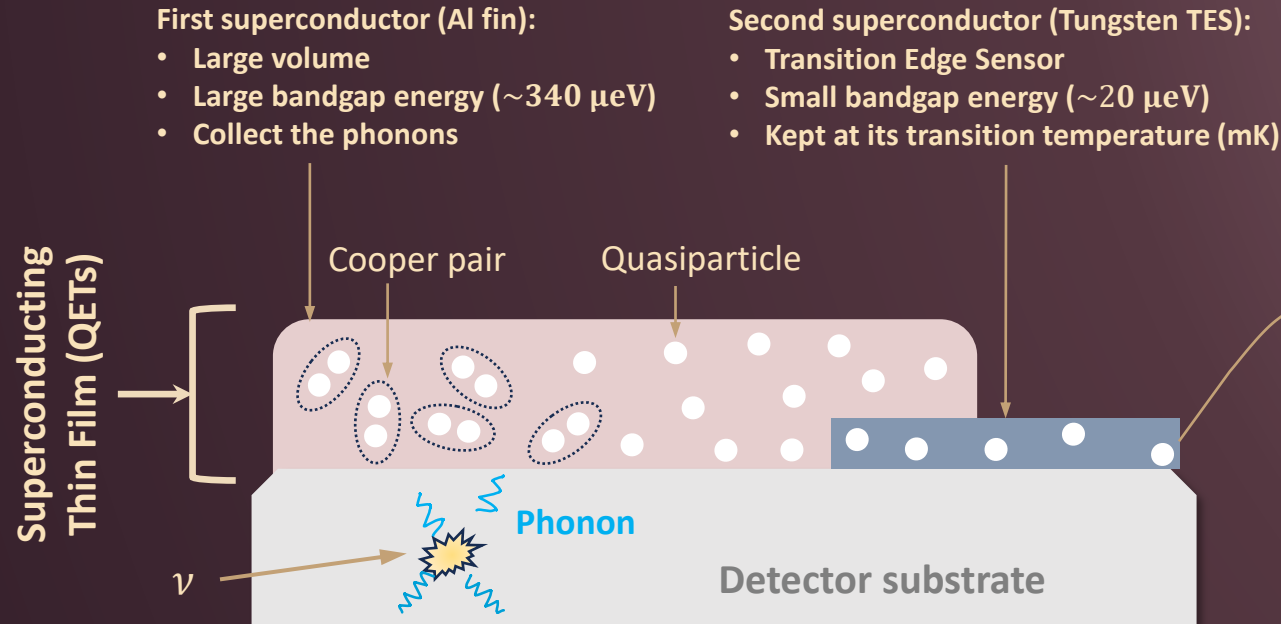
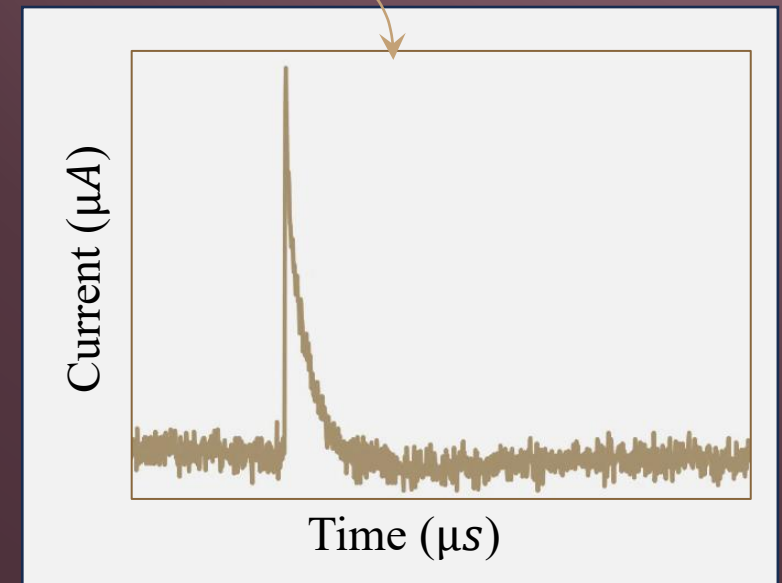
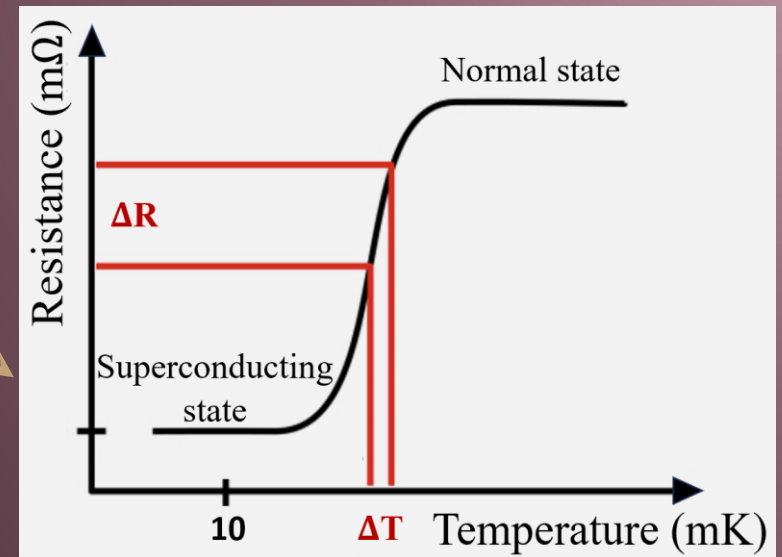
## Experimental setup



| Detector | Baseline resolution [eV] |
|----------|--------------------------|
| Primary  | 40                       |
| Veto     | 120                      |

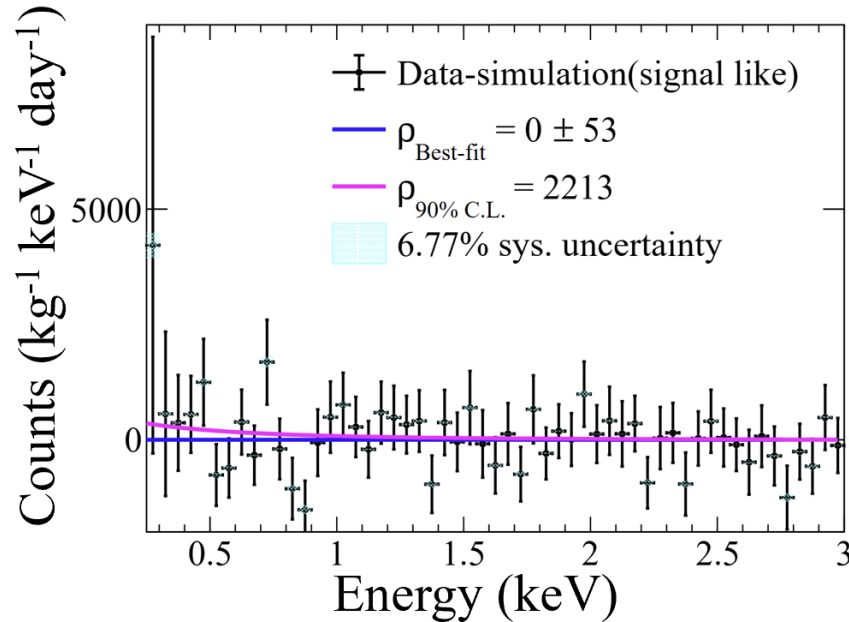
- Reactor power: 1 MW<sub>th</sub>
- Baseline distance: 4.5 meters
- Region of interest: 250 eV – 3 keV

## Transition Edge Sensor graph



- Suitable for low recoil energy detection
- Provides excellent baseline resolution

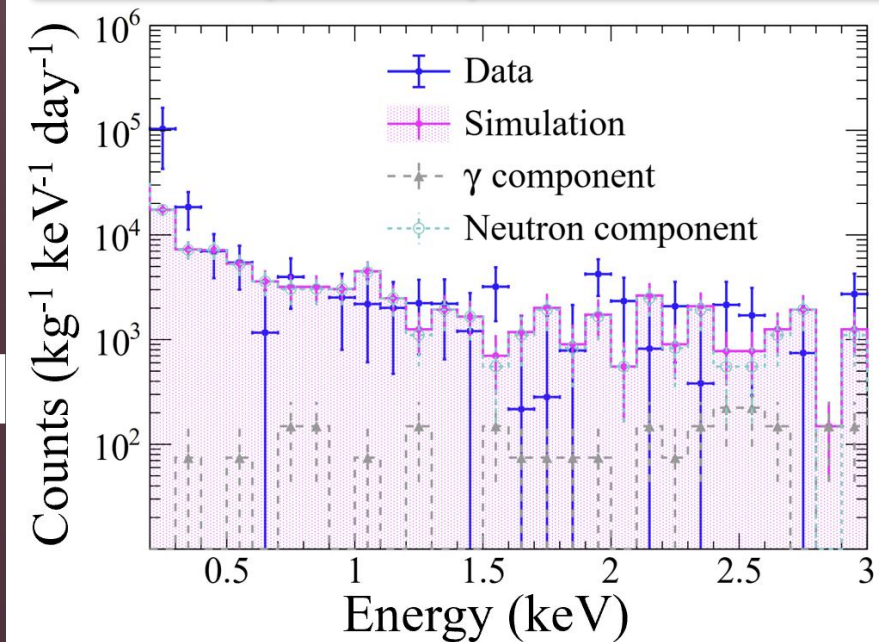
## $\chi^2$ - minimization



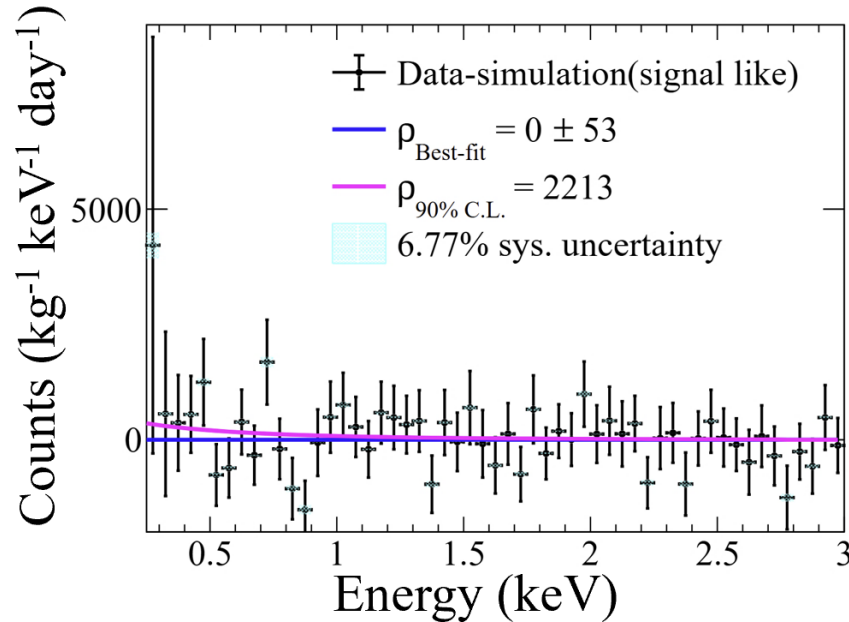
- $\rho \rightarrow$  ratio between data to SM cross section
- Upper limit set on  $\rho$
- Theory prediction  $\sim 0.14$  events/day/kg
- No hint of CEvNS event observed

- Reactor ON-OFF spectra agrees with simulation
- Data is dominated by reactor-related background
- Neutron background dominates over  $\gamma$  background
- Mismatch between data & simulation in 200-400 eV

## Data (ON-OFF) vs simulation



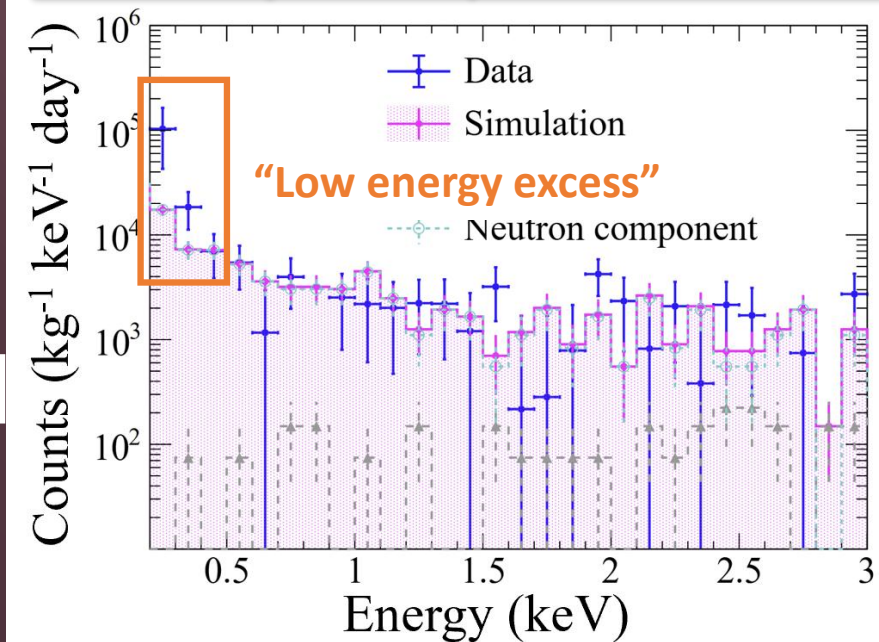
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## Data (ON-OFF) vs simulation



Background in  
cryogenic rare event  
search experiment

Cosmogenic

Environmental

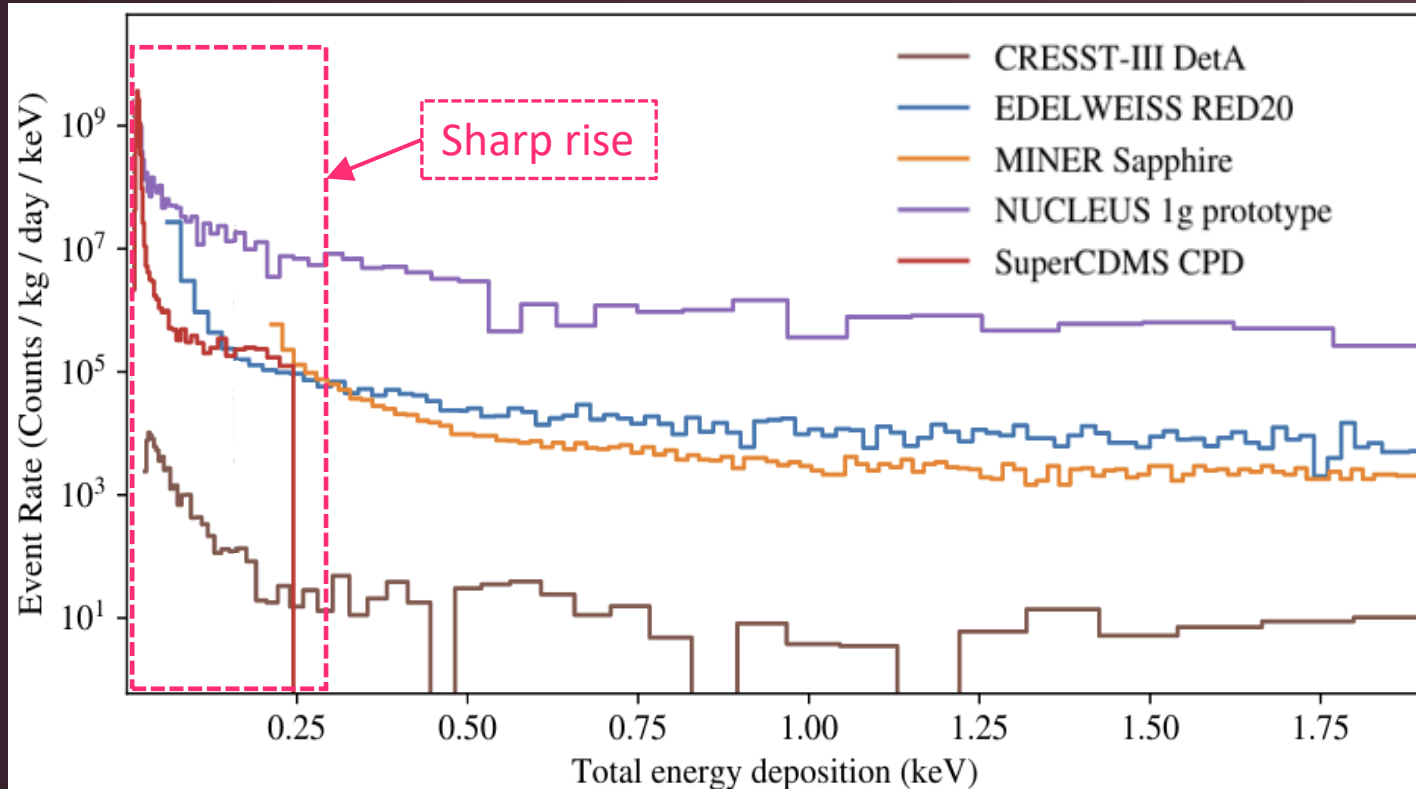
Intrinsic

Known

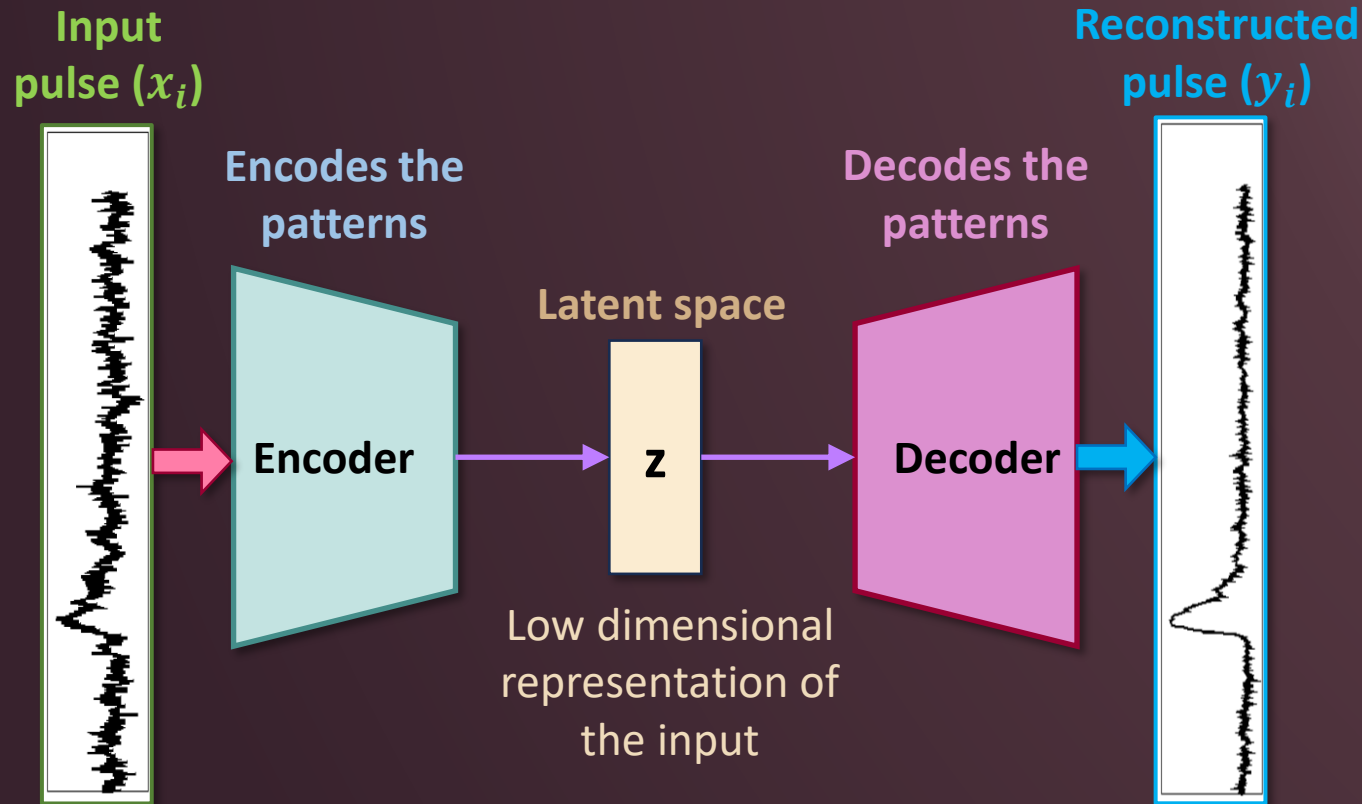
## LEE: Low-Energy Excess

- Event rate rises sharply toward the detector threshold
- Unexplained by known background sources
- Rate decreases with time
- Reappears following a warm-up and recooling cycle
- Limits sensitivity to rare low-energy signals such as  $\text{CE}\nu\text{NS}$  and Dark Matter searches

*SciPost Phys. Proc. 9, 001 (2022)*



## CVAE: Convolutional Variational Autoencoder

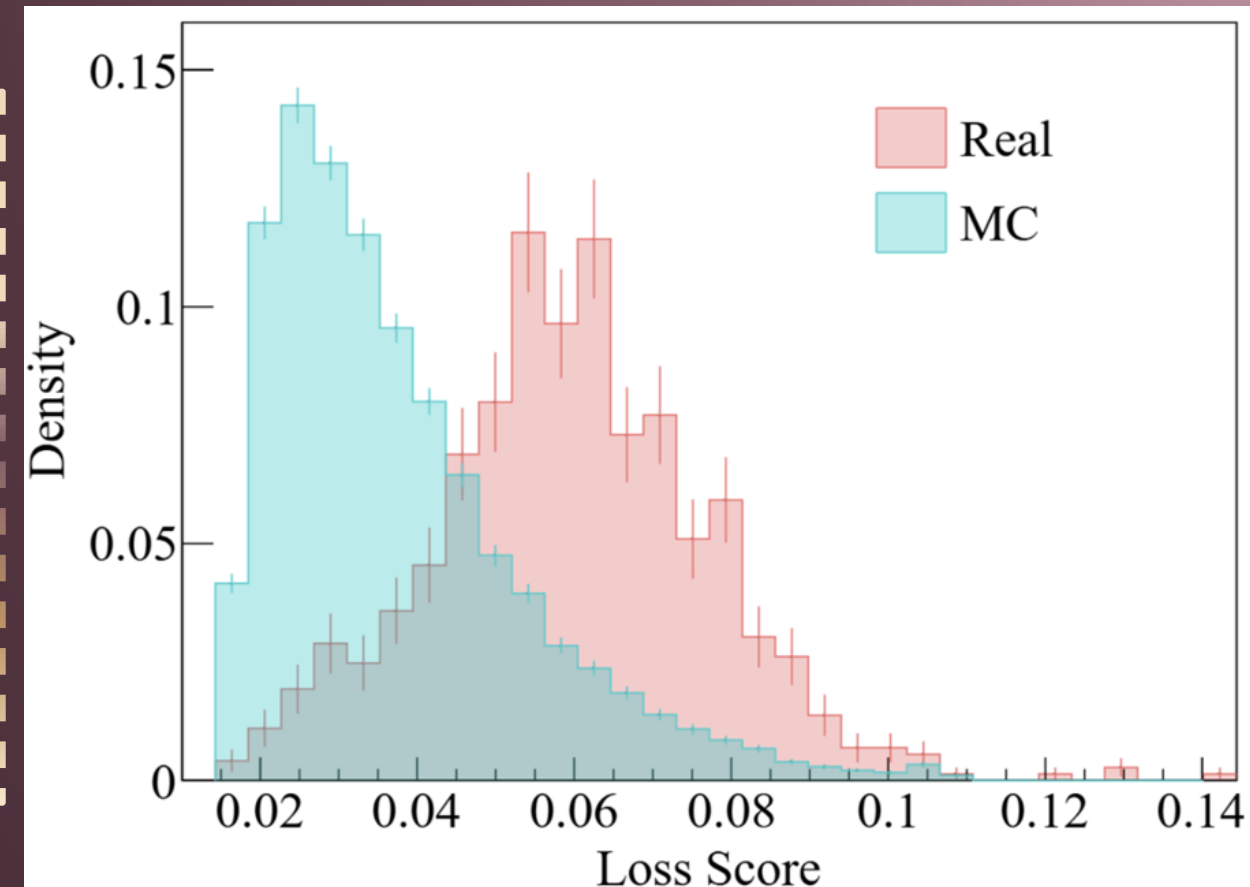


## Reconstruction loss

$$\left( \frac{1}{N} \sum_{i=0}^N (x_i - y_i)^2 + \beta D_{KL} \right)$$

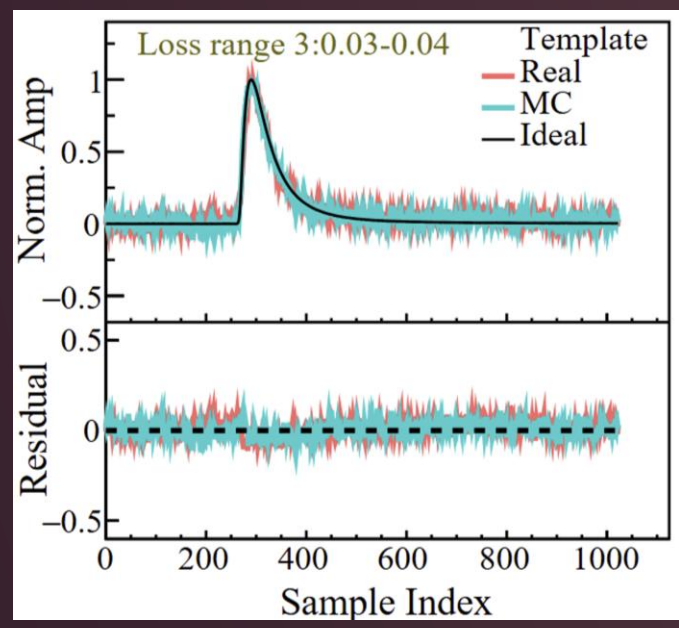
- $x_i \rightarrow$  input (pulse)
- $y_i \rightarrow$  CVAE output (reconstructed)
- $N \rightarrow$  No. of sample points (1024)
- $\beta \rightarrow$  constant weightfactor
- $D_{KL} \rightarrow$  Regularization term

- Most MC events populate the low-loss region.
- Real data exhibits a significant population at higher loss values.
- An overlap region remains where signal-like and LEE events are not fully separated.

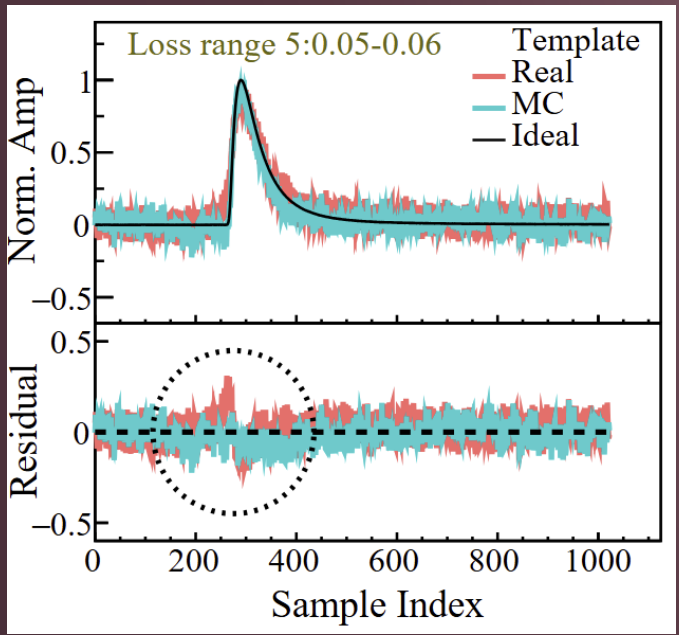


## Pulse-shape evolution across loss-score intervals

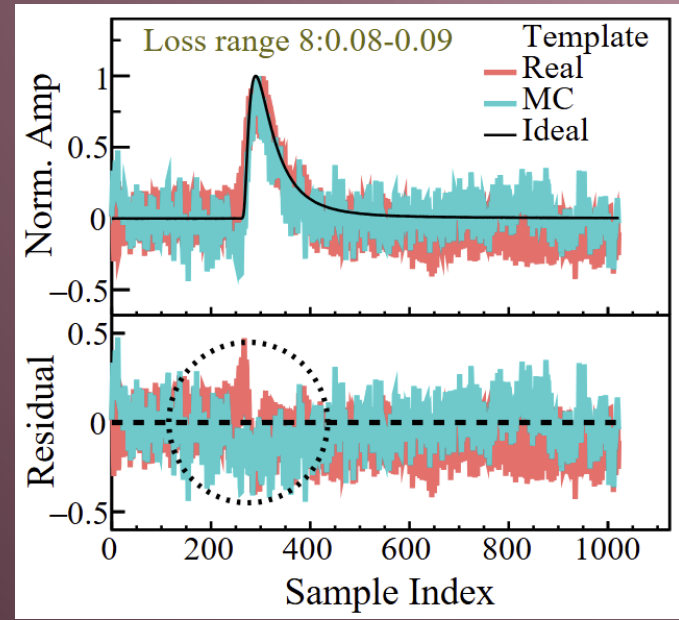
Low loss score



Mid loss score



High loss score

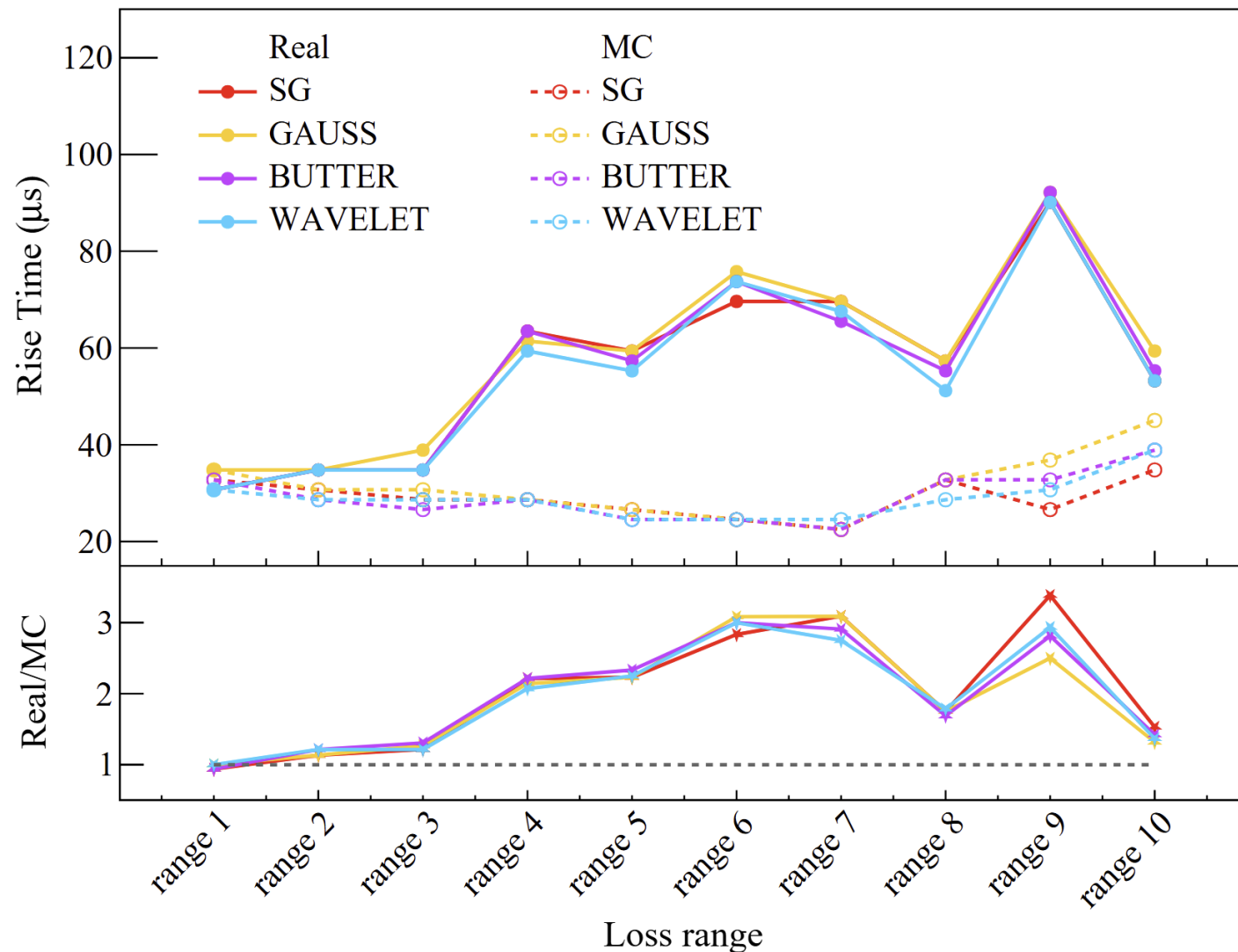


- Loss scores divided into 10 intervals (0–0.11).
- Residual mismatch appears for loss > 0.04.
- A slow-rise component emerges in real-data templates at higher loss scores.



Real-template show slower rise times compared to MC templates.

# TEMPLATE LEVEL RISE-TIME ESTIMATION



Multiple filtering algorithms tested to avoid bias.



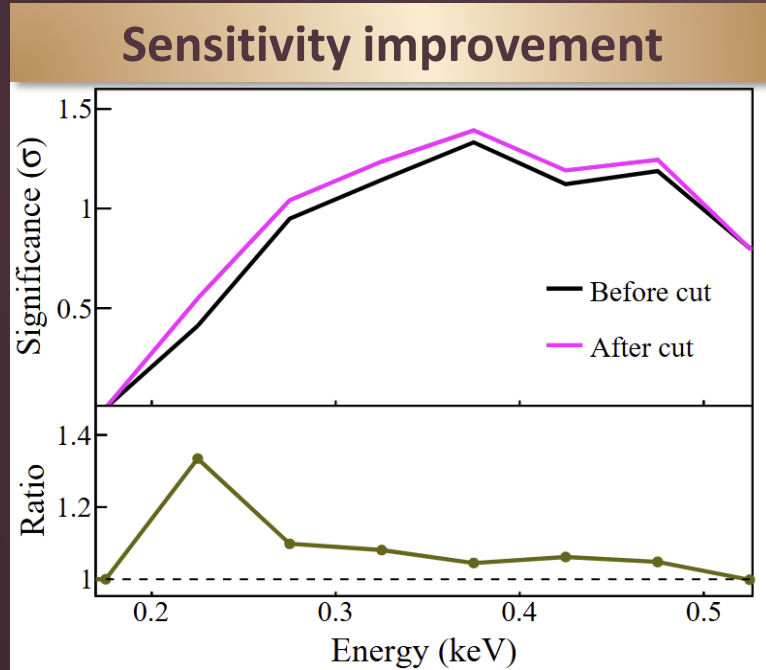
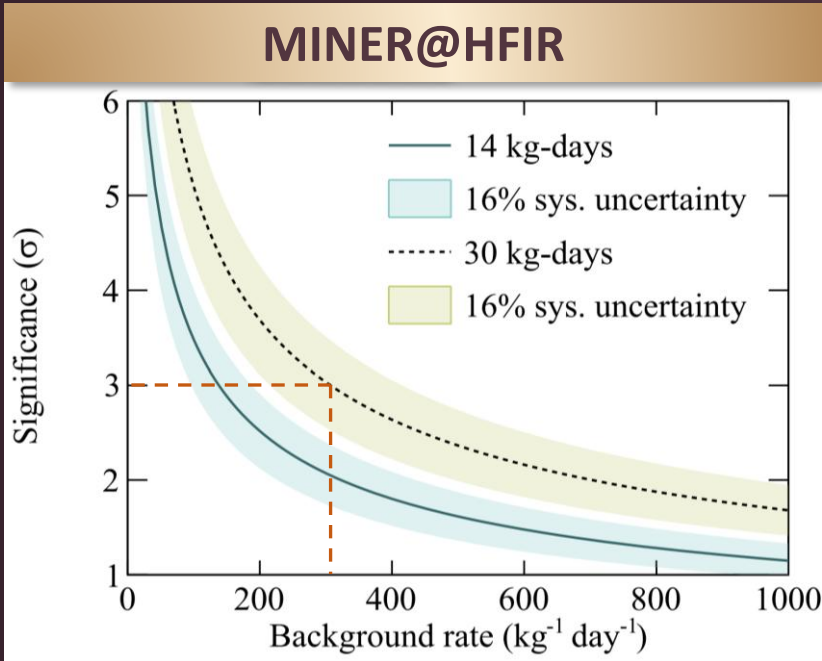
- MC templates' rise time: 20 – 40  $\mu s$
- Rise time for real data templates reaches 3 times of the MC between 0.06 – 0.08.



**Developed a rise-time threshold to estimated the improvement in detection sensitivity.**

| Experiment | Power [MW <sub>th</sub> ] | Baseline [m] | flux [cm <sup>-2</sup> s <sup>-1</sup> ] | Event rate [kg <sup>-1</sup> day <sup>-1</sup> ] |
|------------|---------------------------|--------------|------------------------------------------|--------------------------------------------------|
| MINER      | 1                         | 4.5          | $\sim 7 \times 10^{10}$                  | $\sim 0.14$                                      |
| MINER@HFIR | 85                        | 5            | $\sim 5 \times 10^{12}$                  | $\sim 10$                                        |

- $3\sigma$  can be achieved in 30 kg. day of data taking with  $\sim 300/\text{kg/day}$  background level
- $Z = \frac{s}{\sqrt{s+b}}$
- 40% improvement in the lowest-energy bin
- Overall significance:  $2.8\sigma$  to  $3.1\sigma$  ( $\sim 10\%$  improvement)



- MINER demonstrates ultra-low-threshold  $\text{CE}\nu\text{NS}$  detection capability using cryogenic sapphire detectors, achieving a baseline energy resolution of 40 eV.
- No statistically significant  $\text{CE}\nu\text{NS}$  excess was observed in MINER at TRIGA
- Pulse-shape analysis identifies an anomalous slow-rise pulse population associated with the low-energy excess and enables rejection of a substantial fraction of these events.
- The move to HFIR, combined with improved background rejection, provides significantly enhanced prospects for a first  $\text{CE}\nu\text{NS}$  measurement with MINER.

*Stay tuned !*

Thank you!!



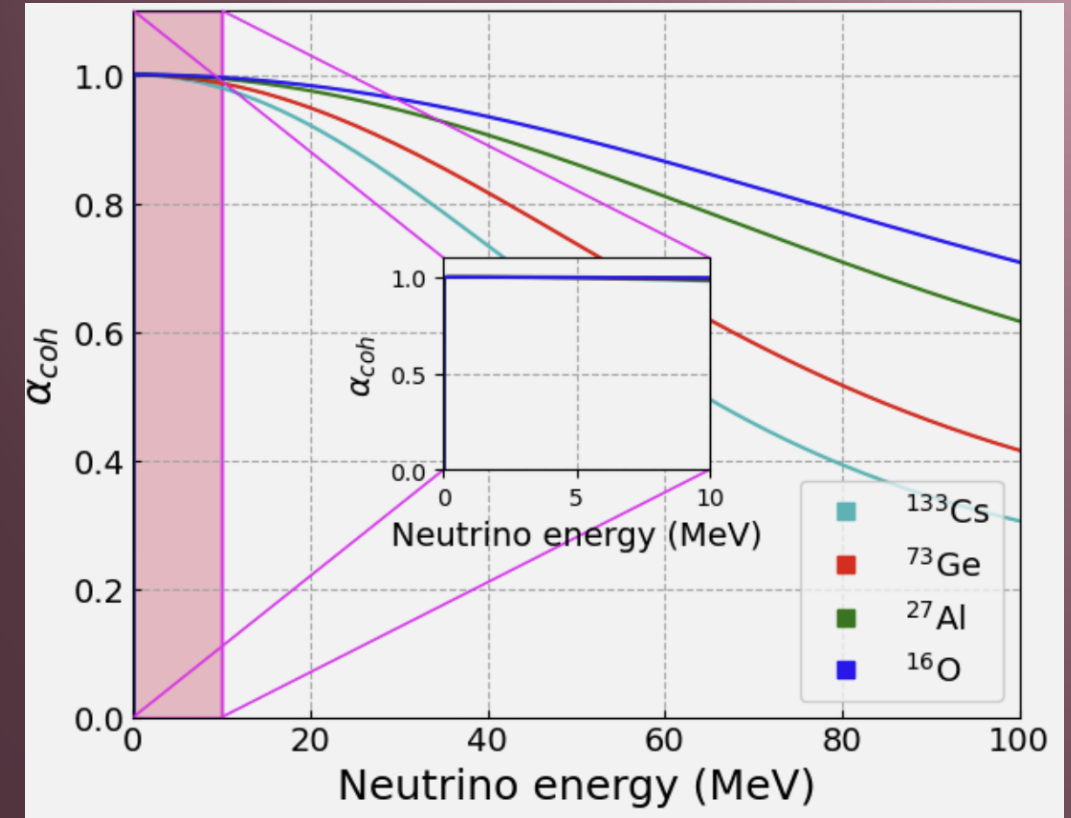
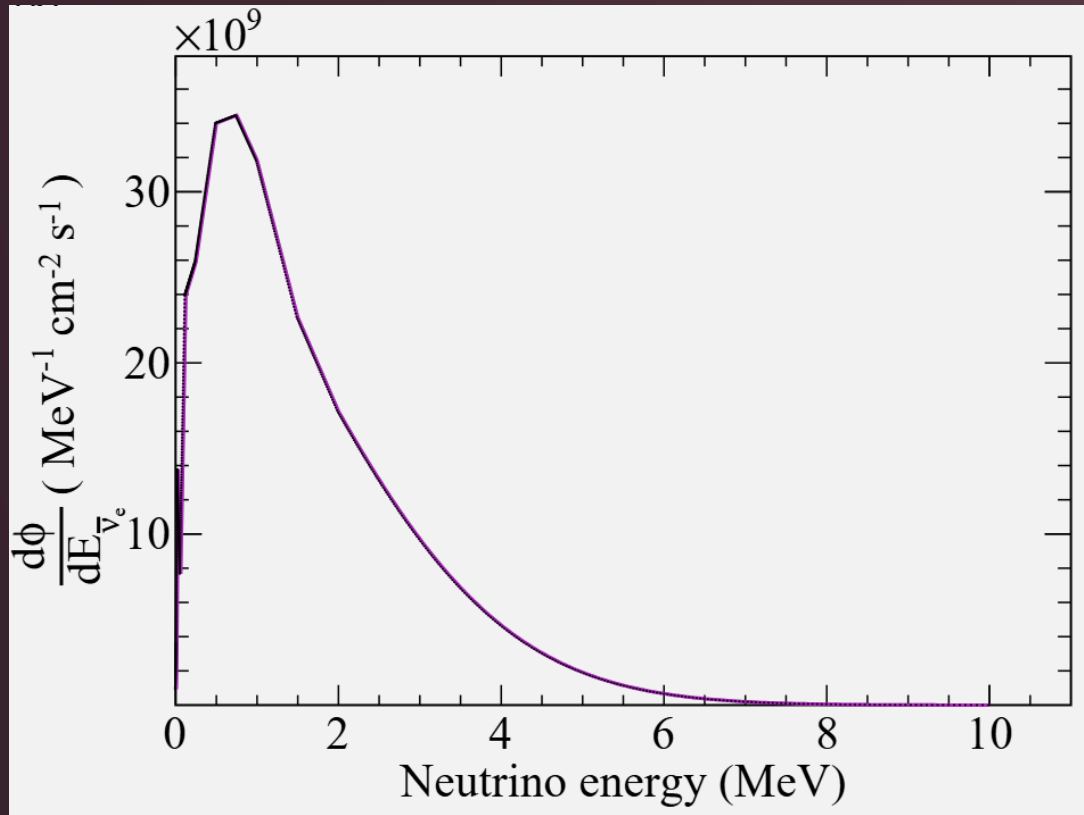
*BACKUPS*



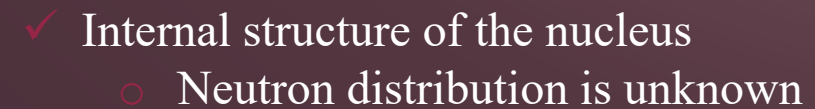
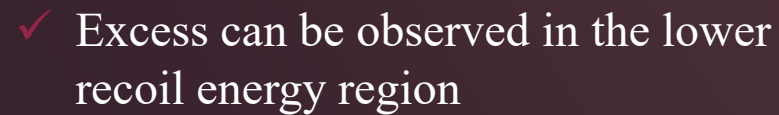
$$\frac{d\sigma(E_\nu, T)}{dT} = \frac{G_F^2 M}{\pi} \left( 1 - \frac{MT}{2E_\nu^2} \right) \times [\mathcal{Q}_{\mathcal{W}}^{\mathcal{N}}(|\vec{q}|)]^2$$

$$\sin^2\vartheta_W \approx 0.239$$



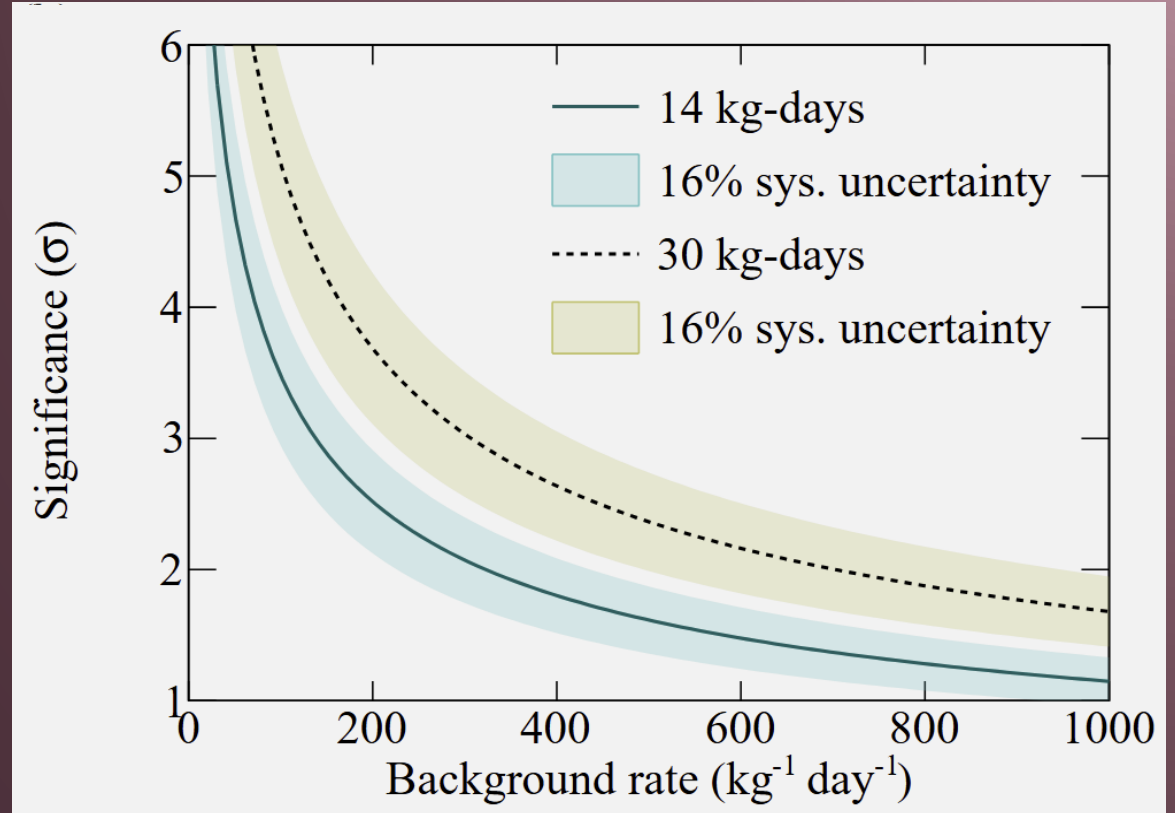


- ✓ Coherency almost 100% up to 10 MeV



$$\chi^2(\delta, \beta, \zeta; \rho) = \sum_i \frac{[R_i^{ON} - \rho(1 + \delta)R_i^{SM} - (1 + \beta)R_i^{sim} - (1 + \zeta)R_i^{OFF}]^2}{\sigma_{i,ON}^2 + (1 + \zeta)\sigma_{i,OFF}^2} + \sum_{k=\delta, \beta, \zeta} \left(\frac{k}{\sigma_k}\right)^2$$

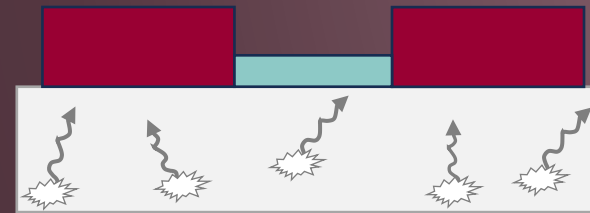
- Pull terms & corresponding systematic uncertainties  $\rightarrow \delta(\sigma_\delta = 16\%), \beta(\sigma_\beta = 10\%) \text{ and } \zeta(\sigma_\zeta = 6.77\%)$
- $R_i^{ON} \rightarrow$  reactor ON event rate
- $R_i^{SM} \rightarrow$  theoretical CEvNS rate from SM
- $R_i^{sim} \rightarrow$  simulated background model
- $R_i^{OFF} \rightarrow$  reactor OFF event rate (background)
- $\sigma_{i,ON} \rightarrow$  statistical uncertainty for reactor ON spectrum
- $\sigma_{i,OFF} \rightarrow$  statistical uncertainty for reactor OFF spectrum







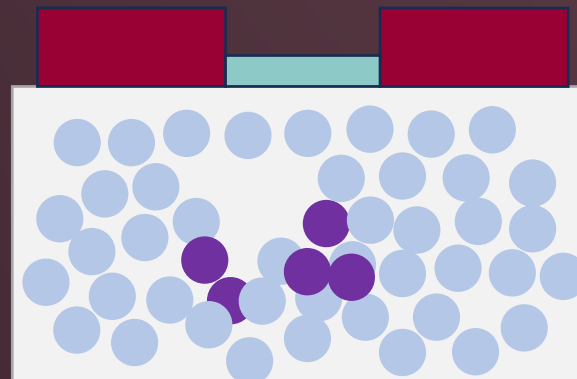
## (b) Bulk related



Detector substrate  
Al phonon collection fins  
Tungsten TES

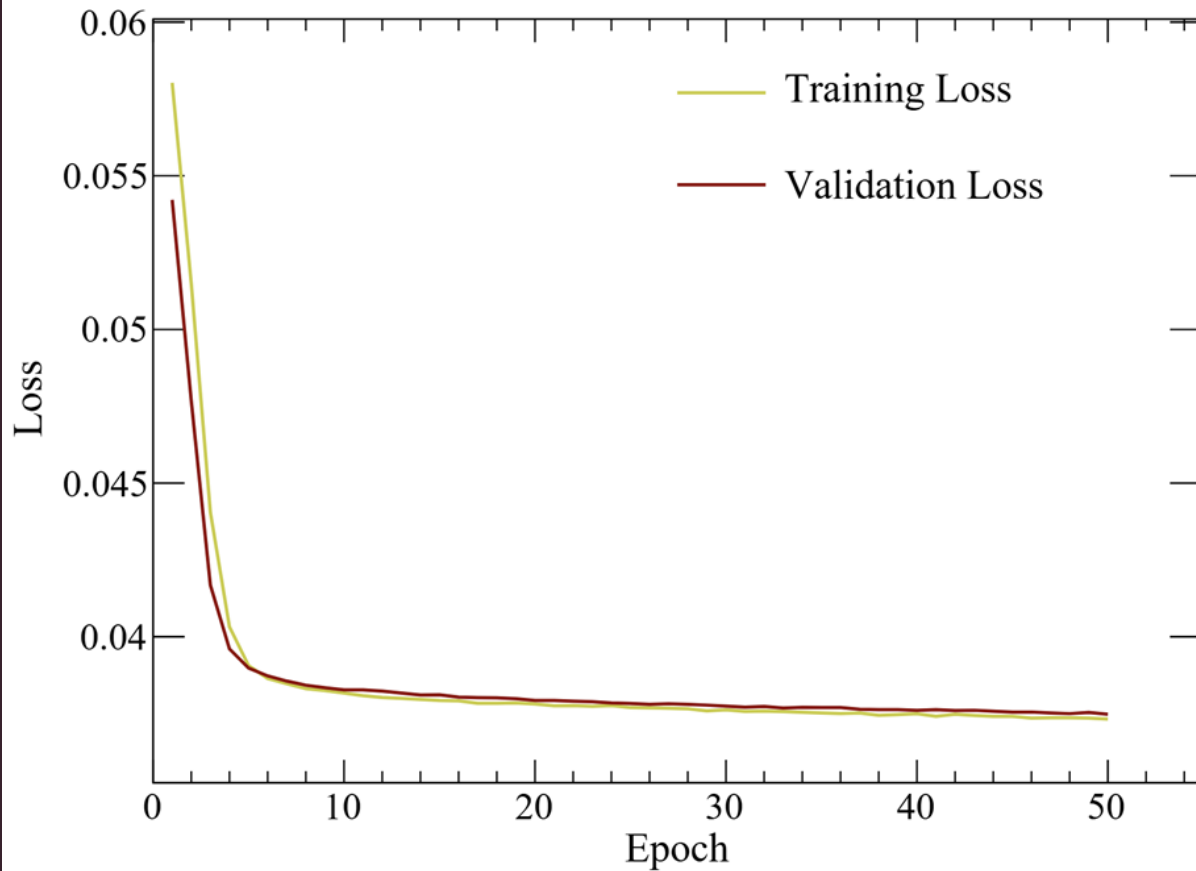
**Cosmic ray shower**

**Fast neutron**



## Defect relaxation

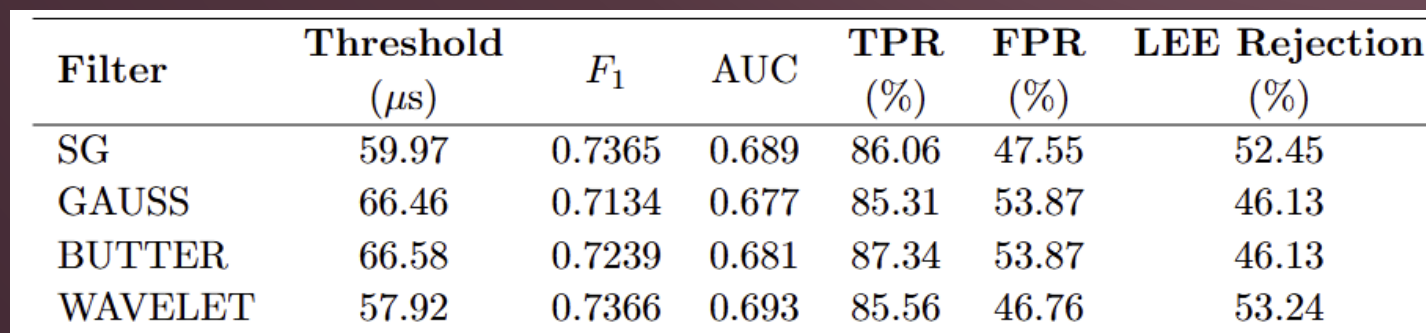




- Loss quantifies the difference between the input and reconstructed pulse
- Training and validation losses converge
- Model generalizes well to unseen events



- True positive rate (TPR)
- False negative rate (FNR)
- False positive rate (FPR)



- ○ ○ ○ ○ ○ ○ ○ ○ ○ ○ ○ ○ ○

