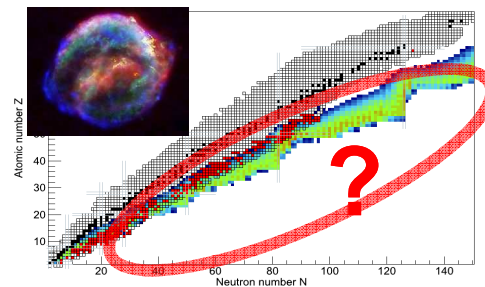
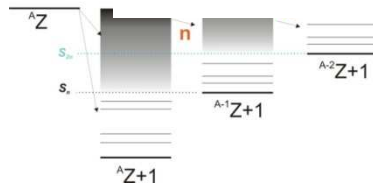
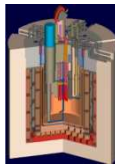
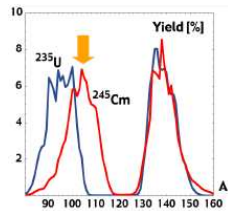
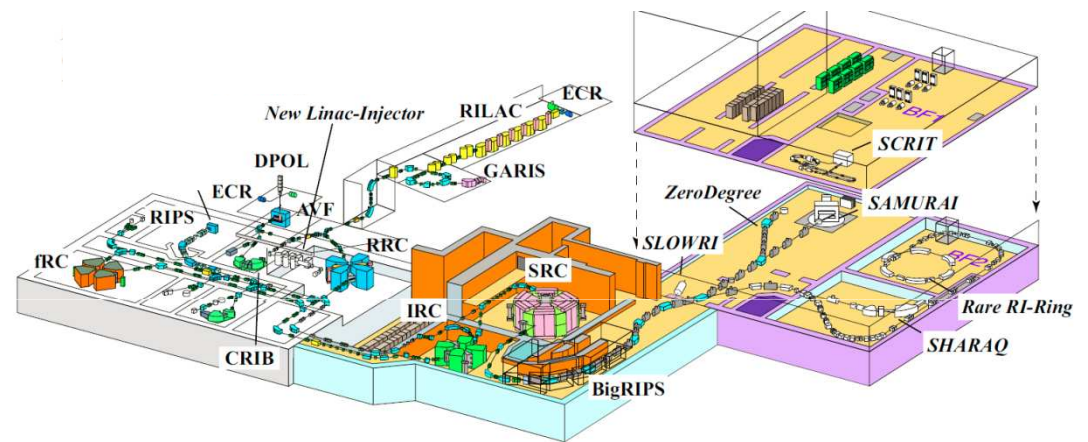
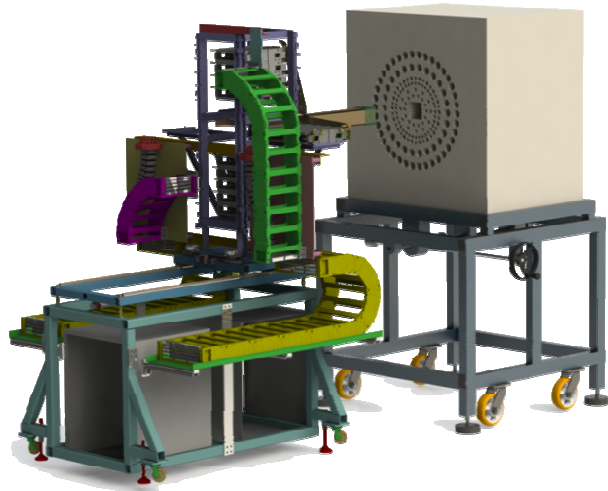


BRIKEN: β -delayed neutron measurements at RIKEN for nuclear structure, astrophysics and applications

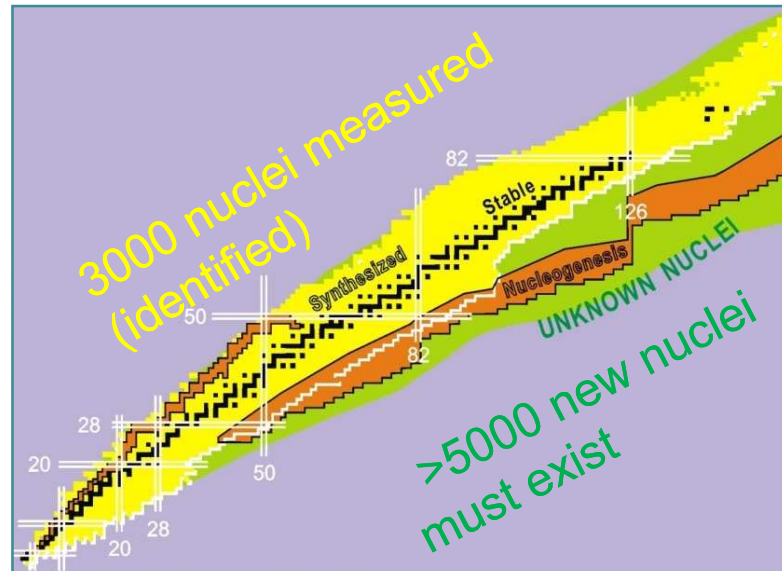


Outline

- Motivation & Introduction
- Recent examples of measurements: 2n-emitters (ORNL), βn around ^{78}Ni (JYFL) and stellar nucleosynthesis around $N=126$ (GSI).
- The BRIKEN approach:
 - BRIKEN-Collaboration
 - Detector design: a high- and flat-efficiency detection system
- Summary & Outlook

Motivation

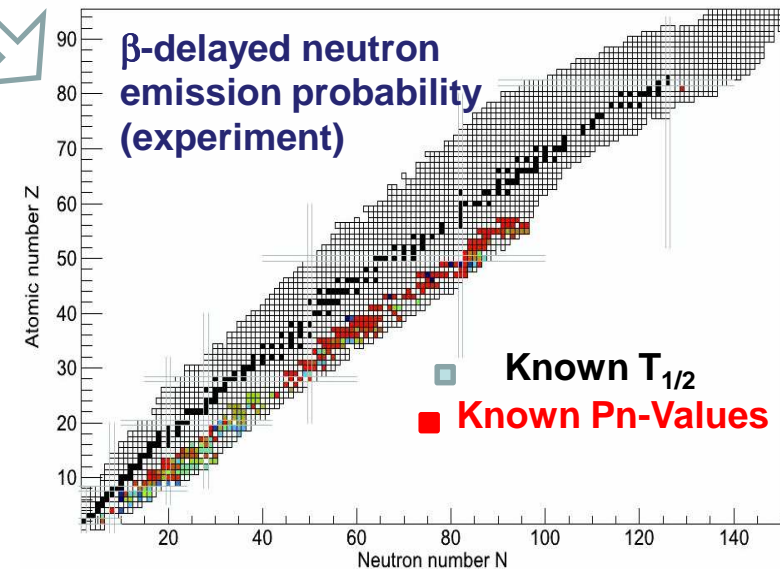
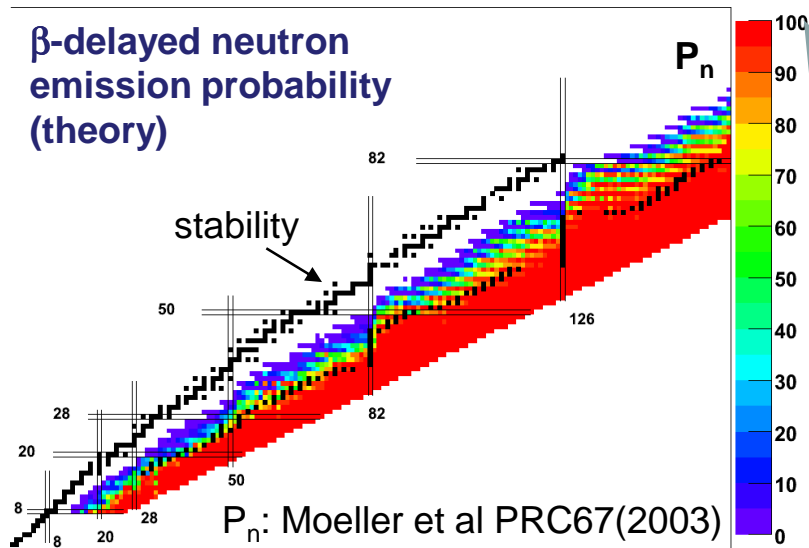
The knowledge we have on nuclear structure and dynamics is based on about 3000 nuclei, whereas still more than 5000 new nuclei must exist.



Almost all these new nuclei are expected to be neutron emitters, and hence, an understanding of this property and the involved technique becomes of pivotal importance for NS and future studies.

What we expect (theory):

What we know (experiments):



• Almost all new nuclei are expected to be n-emitters

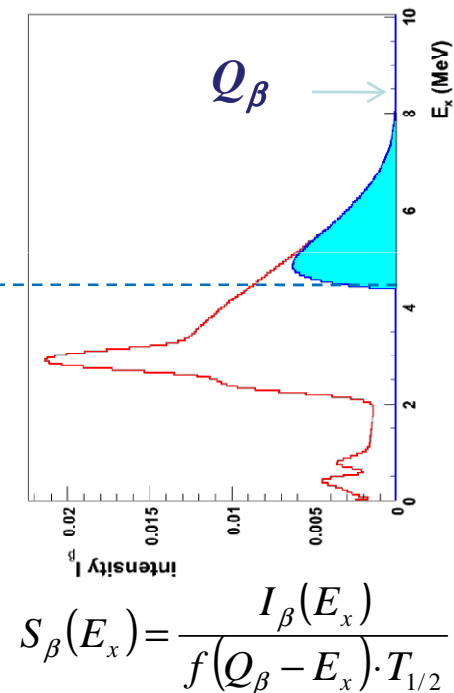
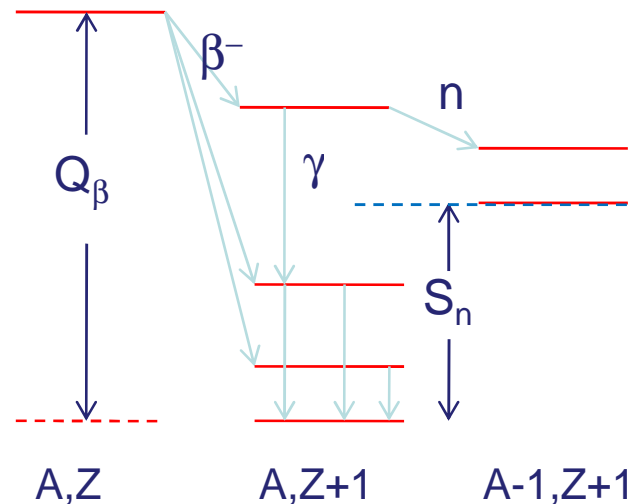
• Only about 200 n-emitters are known

Introduction

- β -delayed neutron emission may happen when the β -decay energy window Q_β exceeds the neutron separation energy S_n in the daughter nucleus. First reported by Roberts et al. in 1939.
- The half-life $T_{1/2}$ yields information on the average β -feeding of a nucleus.
- P_n yields information on the β -feeding **above the S_n**

$$\frac{1}{T_{1/2}} = \sum_0^{Q_\beta} S_\beta(E_x) \cdot f(Q_\beta - E_x)$$

$$P_n = \frac{\sum_0^{Q_\beta} S_\beta(E_x) \cdot f(Q_\beta - E_x)}{\sum_0^{Q_\beta} S_\beta(E_x) \cdot f(Q_\beta - E_x)}$$

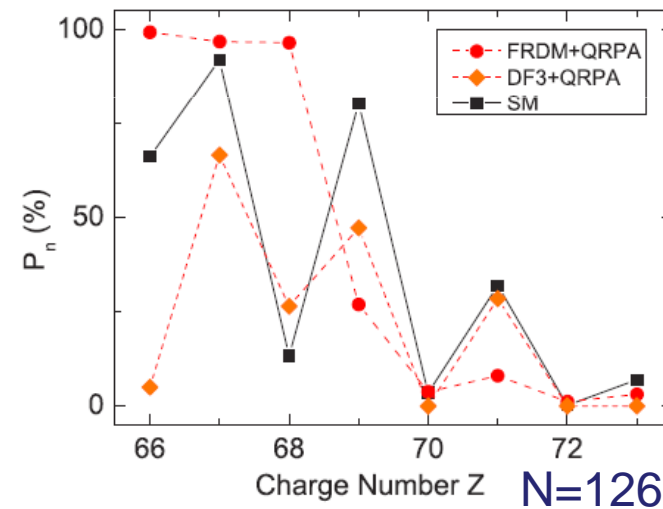
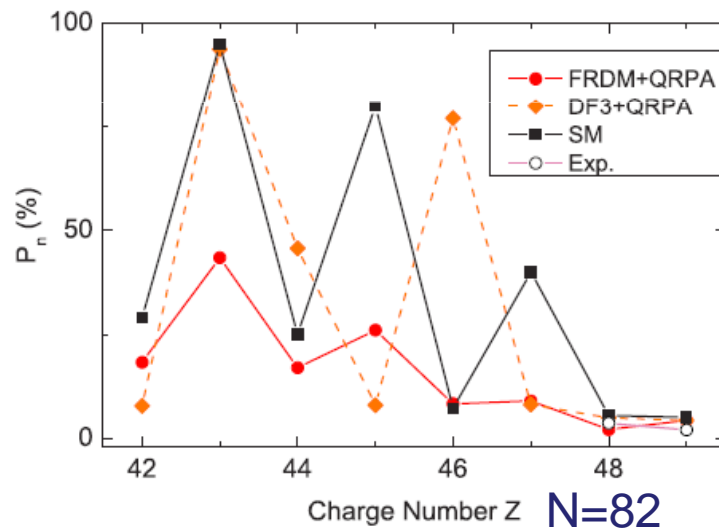


Despite of the relatively simple P_n “definition”, P_n values are rather difficult to predict theoretically, as they are reflecting the “shape” of the b-strength distribution and the underlying fine-structure of the nucleus at high excitation energy (!).

Introduction

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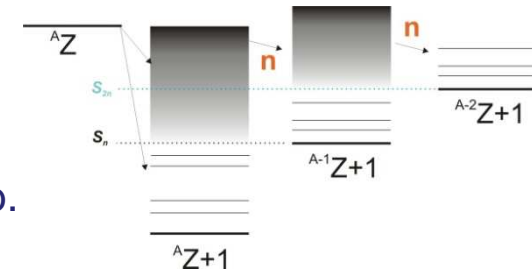
Credit: Q. Zhi et al., Phys. Rev. C 87, 2013



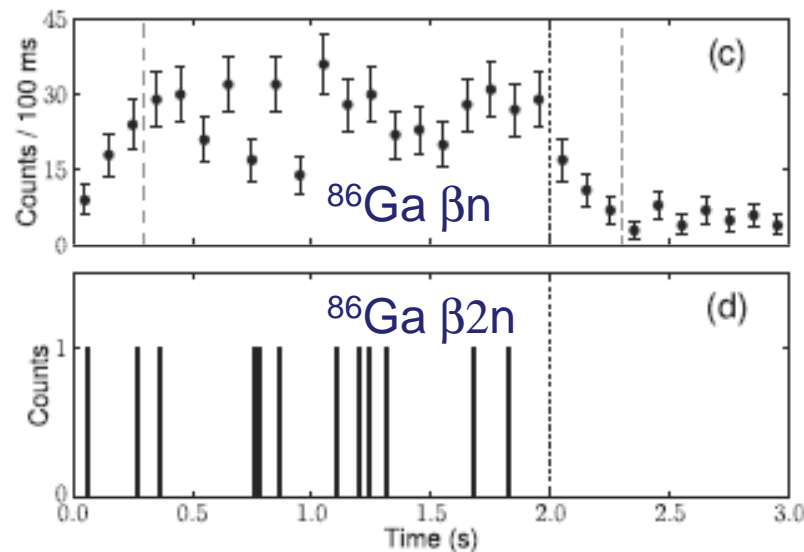
Despite of the relatively simple P_n “definition”, P_n values are rather **difficult to predict theoretically**, as they are reflecting the “shape” of the b-strength distribution and the underlying fine-structure of the nucleus at high excitation energy (!).

Competition between 1n and 2n emission

- 1n-2n competition and theoretical description
- xn emission is even more difficult to model.
- Postulated for first time in 1960s (Goldansky et al.)
- Only 18 b2n emitters, only three heavier than Fe: ^{86}Ga and $^{98,100}\text{Rb}$.



Credit: K. Miernik, et al., Phys. Rev. Lett. 111 (2013)



$$\epsilon_n = 10\%$$

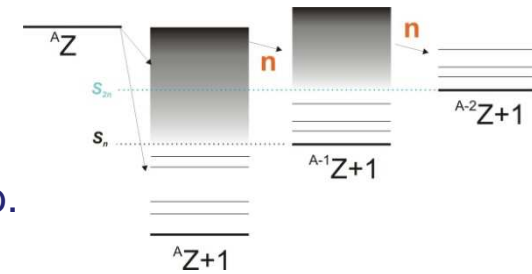
$$\epsilon_{2n} = 1\%$$



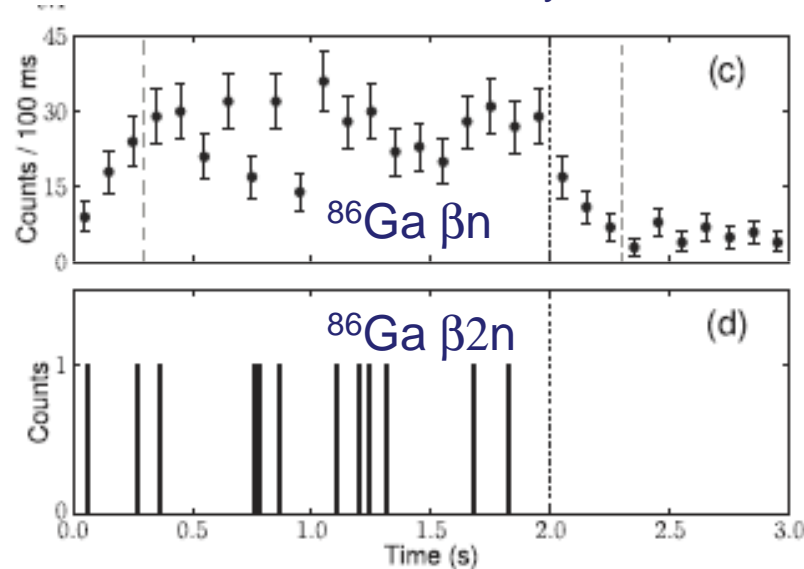
Hybrid 3He @ HRIBF-LeRIBSS:
48 ^3He Counters + 2 HPGe Clovers

Competition between 1n and 2n emission

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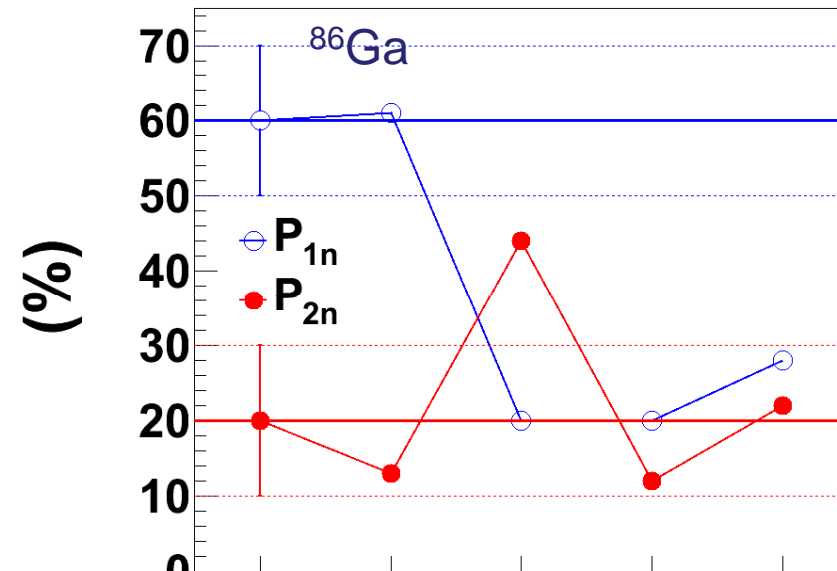


Credit: K. Miernik, et al., Phys. Rev. Lett. 111 (2013)



$$\epsilon_n = 10\%$$

$$\epsilon_{2n} = 1\%$$

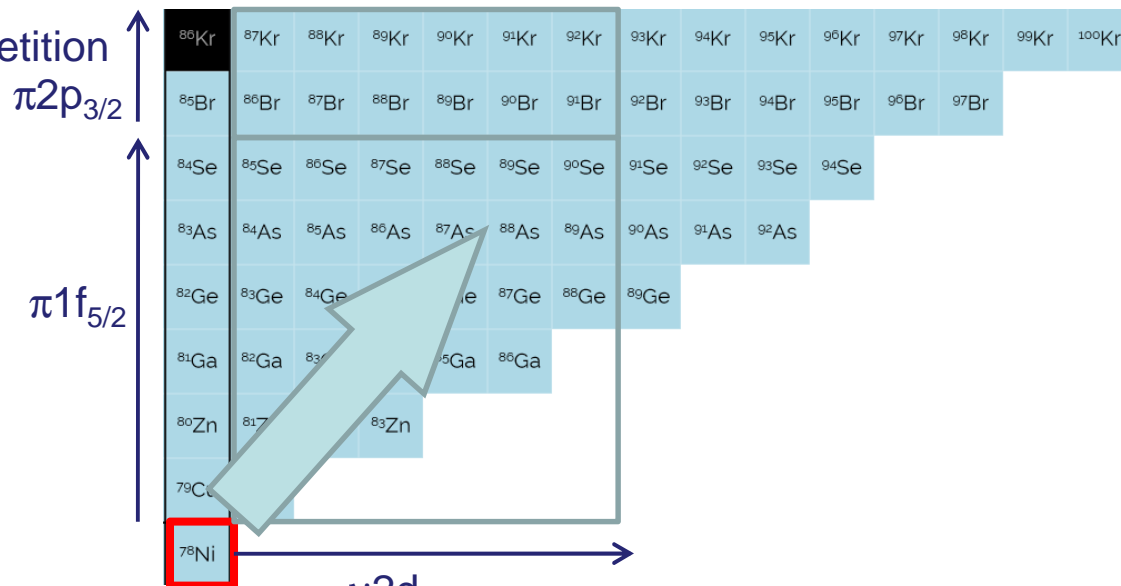
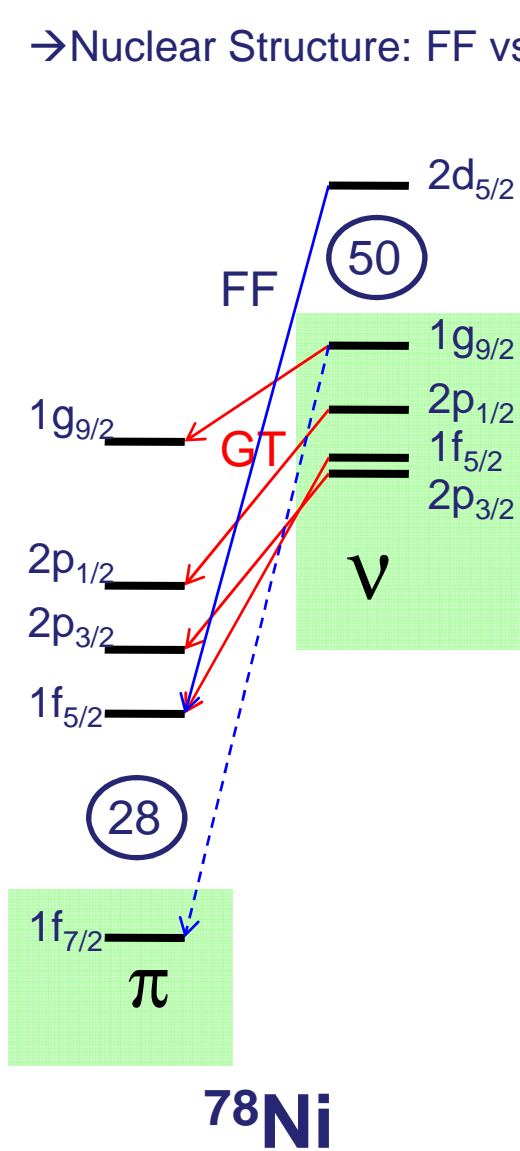


Miernik et al. (2013)
 FRDM+QRPA (1997)
 FRDM+QRPA2 (2003)
 DF3a+CQRPA (2003)
 DF3a+CQRPA (2013)

→ P_n and P_{2n} measurements are a very stringent test for theory models far-off stability!

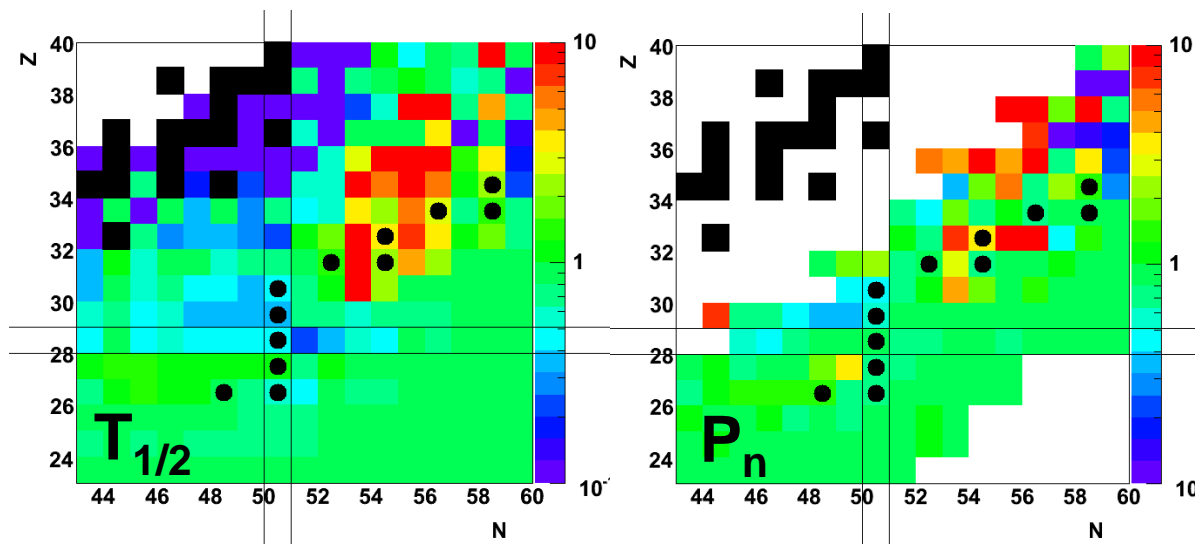
Shell structure and GT- FF- competition

→ Nuclear Structure: FF vs. GT competition



(GT+FF)/GT

Moeller et al., PRC67(2003)
Borzov, PRC71(2005)



Shell structure and GT- FF- competition

→ Nuclear Structure: FF vs. GT competition
BELEN20 @ JYFL



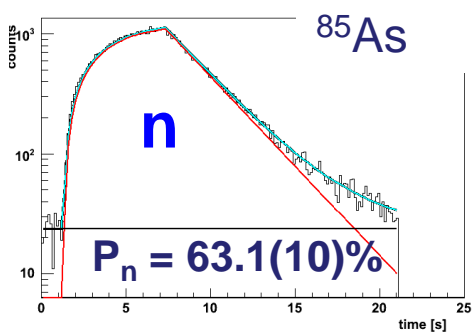
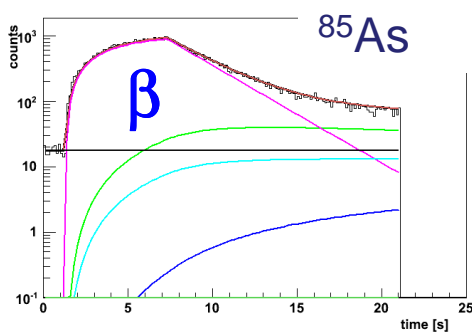
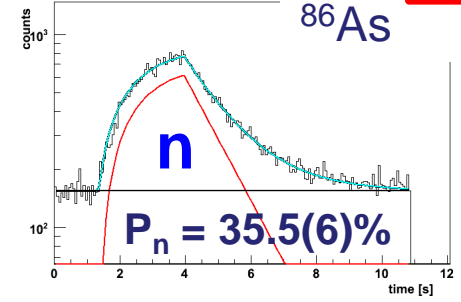
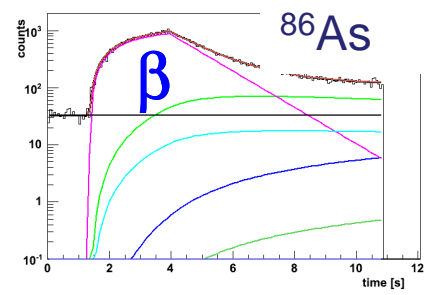
$\pi 2p_{3/2}$
 $\pi 1f_{5/2}$

⁸⁶ Kr	⁸⁷ Kr	⁸⁸ Kr	⁸⁹ Kr	⁹⁰ Kr	⁹¹ Kr	⁹² Kr	⁹³ Kr	⁹⁴ Kr	⁹⁵ Kr	⁹⁶ Kr	⁹⁷ Kr	⁹⁸ Kr	⁹⁹ Kr	¹⁰⁰ Kr
⁸⁵ Br	⁸⁶ Br	⁸⁷ Br	⁸⁸ Br	⁸⁹ Br	⁹⁰ Br	⁹¹ Br	⁹² Br	⁹³ Br	⁹⁴ Br	⁹⁵ Br	⁹⁶ Br	⁹⁷ Br		
⁸⁴ Se	⁸⁵ Se	⁸⁶ Se	⁸⁷ Se	⁸⁸ Se	⁸⁹ Se	⁹⁰ Se	⁹¹ Se	⁹² Se	⁹³ Se	⁹⁴ Se				
⁸³ As	⁸⁴ As	⁸⁵ As	⁸⁶ As	⁸⁷ As	⁸⁸ As	⁸⁹ As	⁹⁰ As	⁹¹ As	⁹² As					
⁸² Ge	⁸³ Ge	⁸⁴ Ge	⁸⁵ Ge	⁸⁶ Ge	⁸⁷ Ge	⁸⁸ Ge	⁸⁹ Ge							
⁸¹ Ga	⁸² Ga	⁸³ Ga	⁸⁴ Ga	⁸⁵ Ga	⁸⁶ Ga									
⁸⁰ Zn	⁸¹ Zn	⁸² Zn	⁸³ Zn											
⁷⁹ Cu	⁸⁰ Cu													

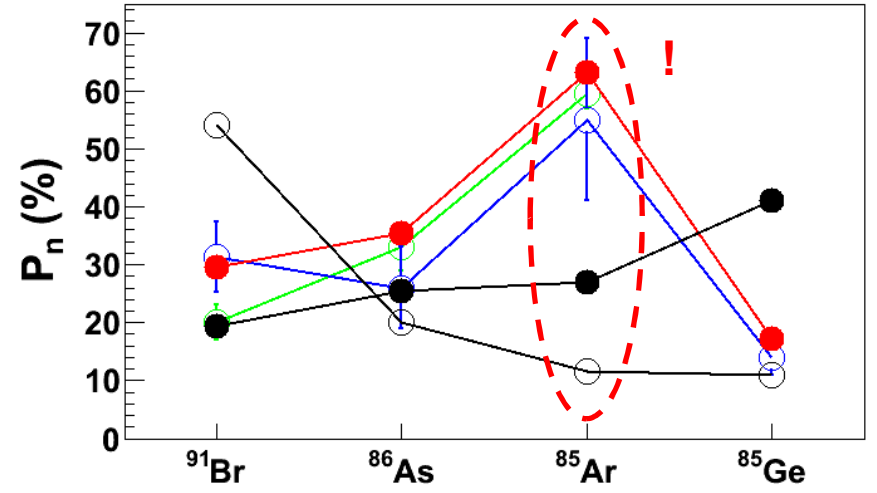
⁷⁸Ni

$v2d_{5/2}$ Nuc. Data. Conf., J.L.Tain et al. (2013)

PhD Thesis:
A. García (CIEMAT)
J. Agramunt (IFIC)

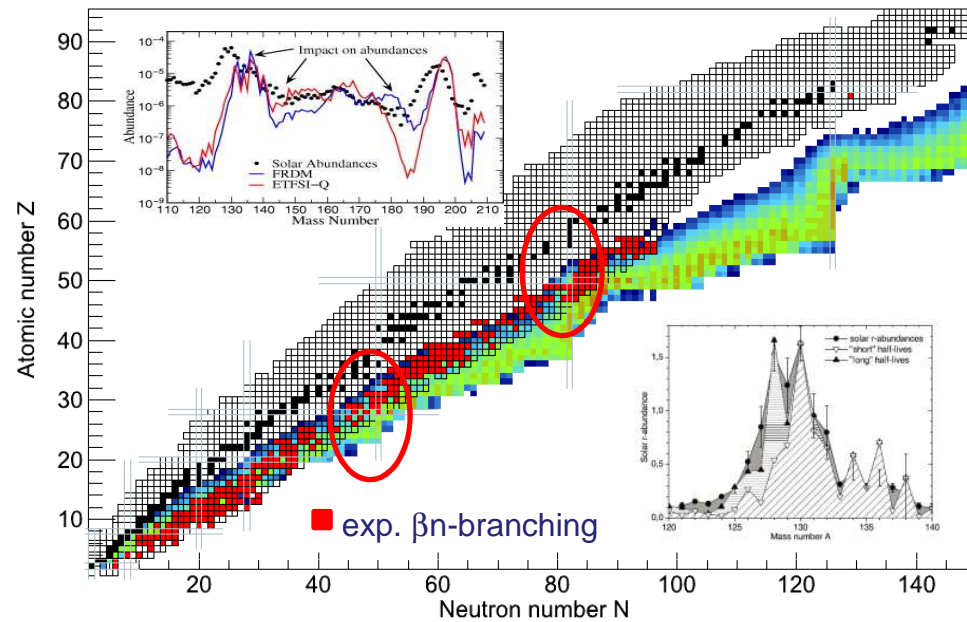


- BELEN @ JYFL, J.L.Tain et al. (2013)
- Rudstam et al., ADNDT 53 (1993)
- Pfeiffer et al., Prog.Nuc.En.41 (2002)
- DF3-cQRPA, I.Borzov (2005)
- FRDM-cQRPA, P.Moller et al. (2003)



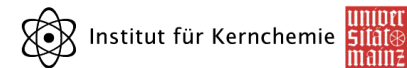
Nucleosynthesis of the heaviest elements

- Re-activation
- Shift towards lower A



F. Montes et al., Phys. Rev. C 73 (2006)
 J. Pereira et al., Phys. Rev. C 79 (2009)
 P. Hosmer et al., Phys. Rev. C 82 (2010)
 H. Schatz et al., The Astr. Jour. 579 (2002)

I. Dillmann et al., Phys. Rev. Lett 91 (2003)
 K.L. Kratz et al., Hyp. Int. 129 (2000)
 H. Ohm et al., Zeit. Phys. 296 (1980)
 K.L. Kratz et al., Phys. Lett. B 103 (1981)
 H. Gabelmann et al., Zeit. Phys. A 308 (1982)
 K.-L. Kratz et al., Zeit. Phys. A 306 (1982)
 J.C. Wang et al., Phys. Lett. B 454 (1999)
 M. Hannawald et al., Phys. Rev. C 62 (2000)



Nucleosynthesis of the heaviest elements

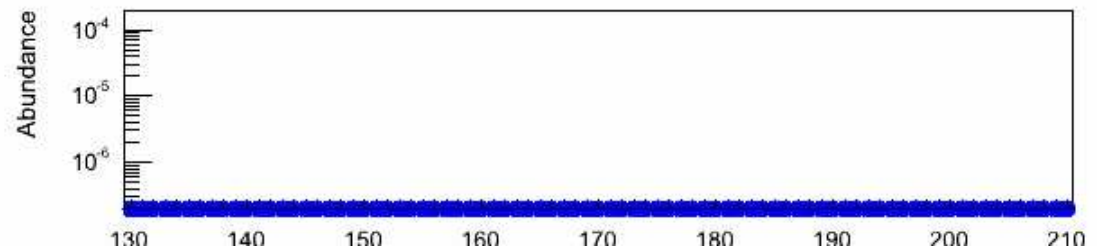
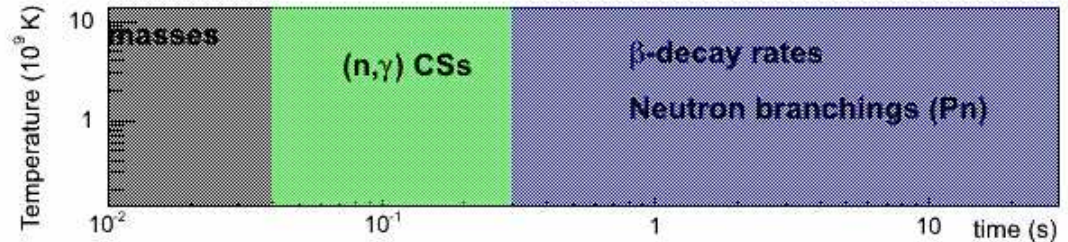
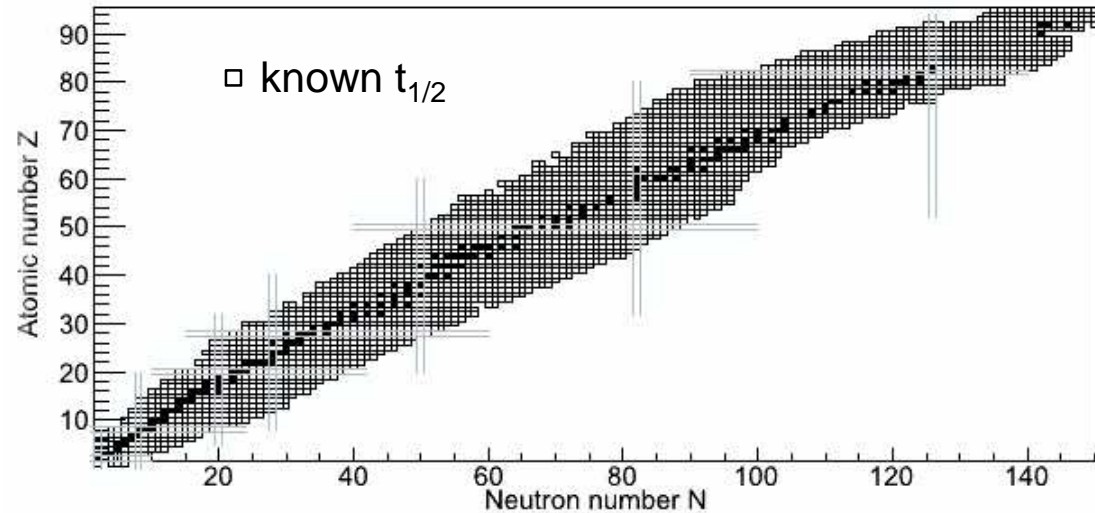
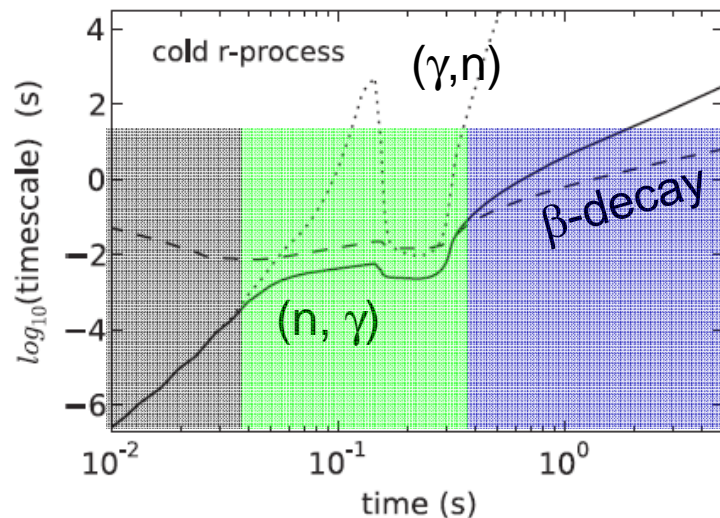
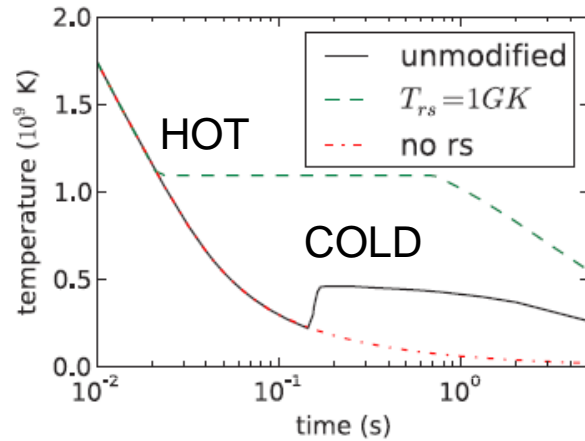
PHYSICAL REVIEW C 83, 045809 (2011)

Dynamical *r*-process studies within the neutrino-driven wind scenario and its sensitivity to the nuclear physics input

A. Arcones^{1,2,*} and G. Martínez-Pinedo²

¹Institut für Kernphysik, Technische Universität Darmstadt, D-64289 Darmstadt, Germany

²GSI Helmholtzzentrum für Schwerionenforschung, Planckstraße 1, D-64291 Darmstadt, Germany



NucNet network code, B. Meyer et al., Clemson University
FRDM+QRPA (P. Möller) + JINA Reaclib Database

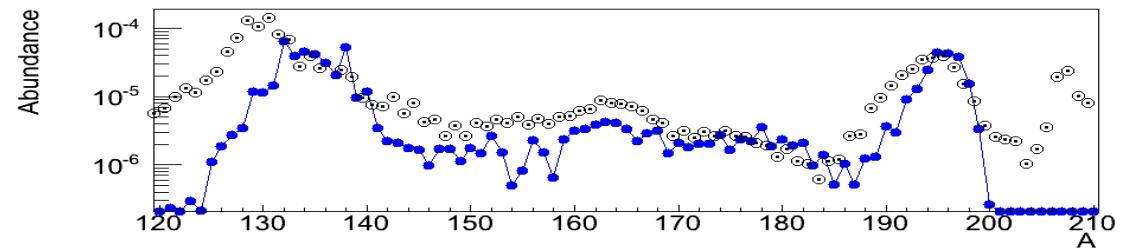
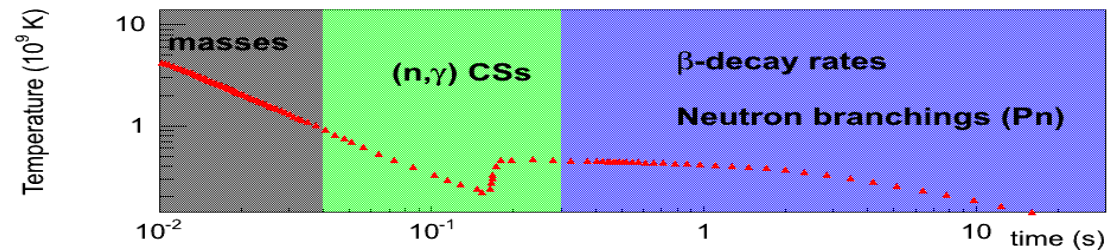
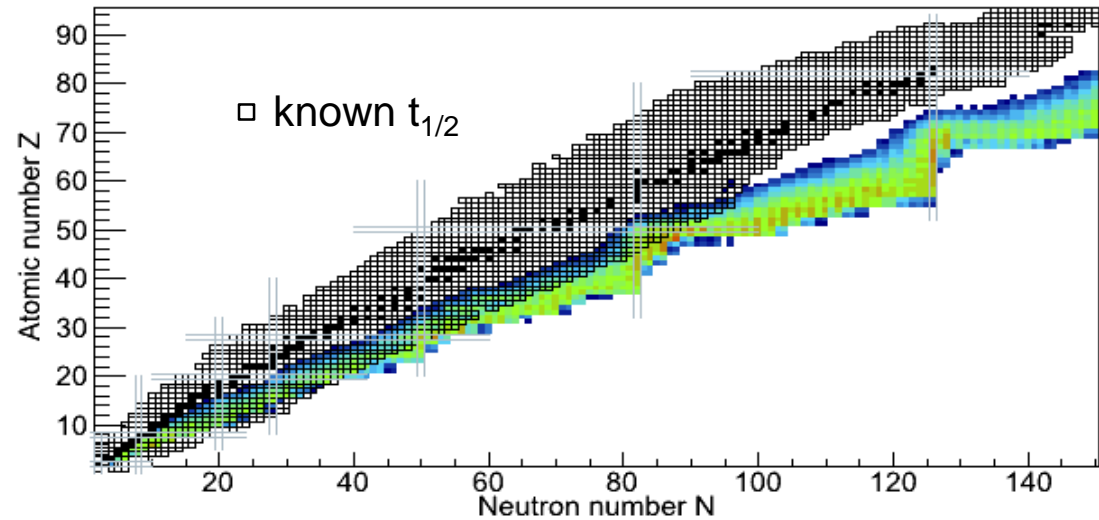
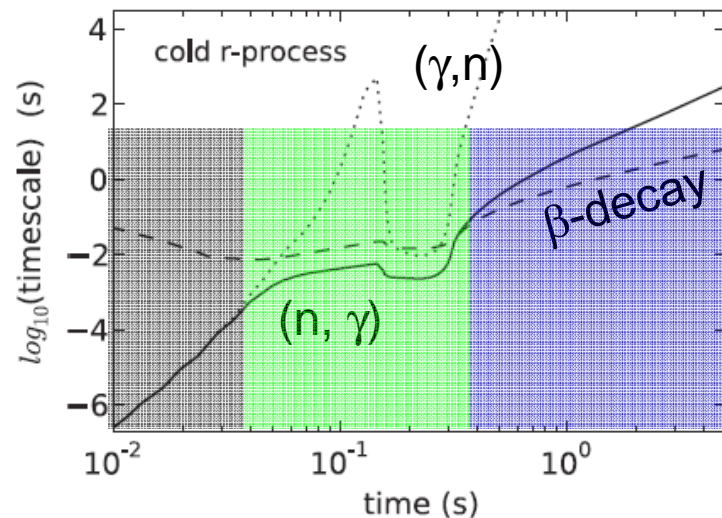
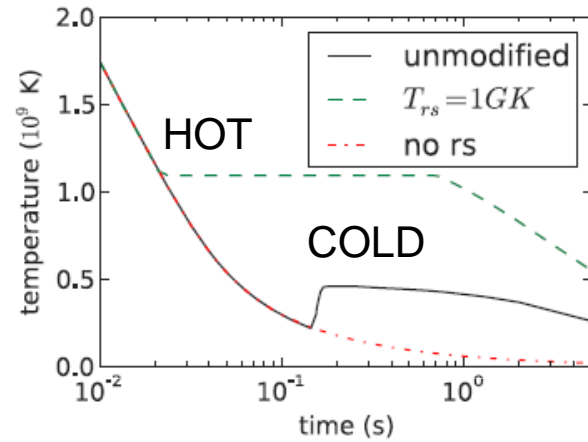
Nucleosynthesis of the heaviest elements

PHYSICAL REVIEW C 83, 045809 (2011)

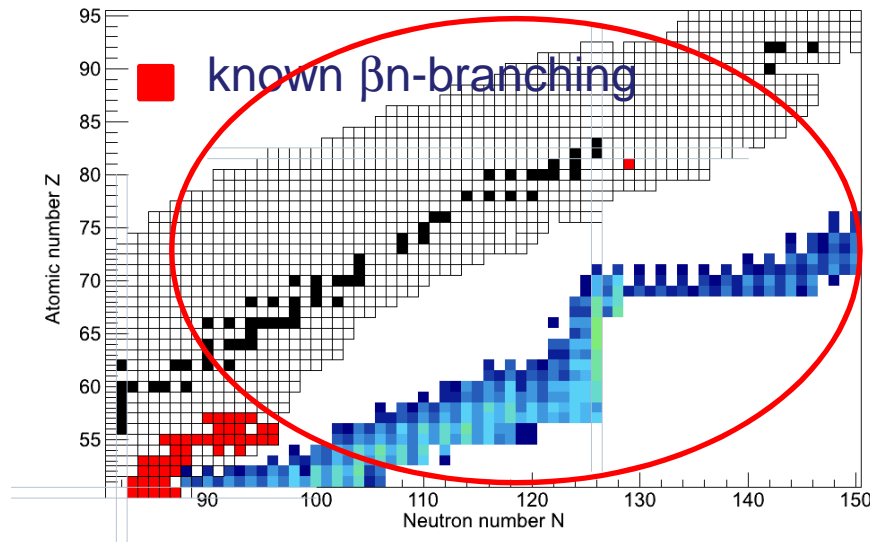
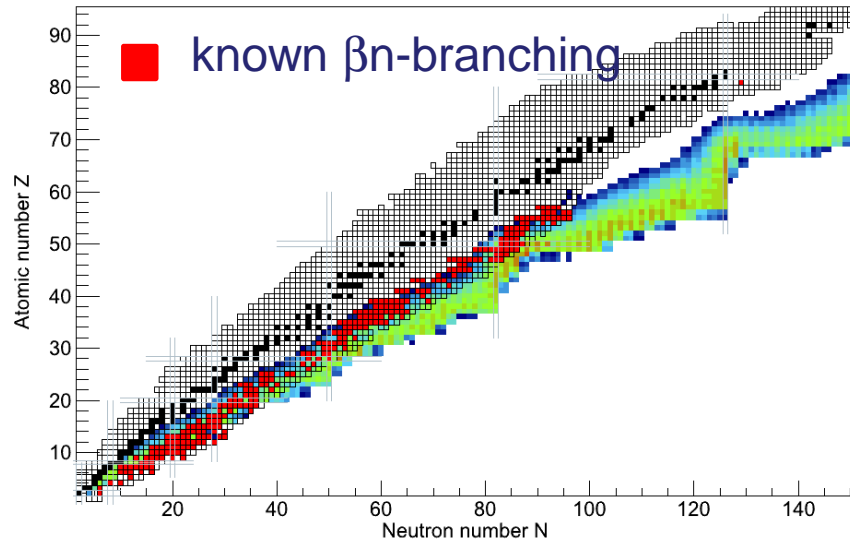
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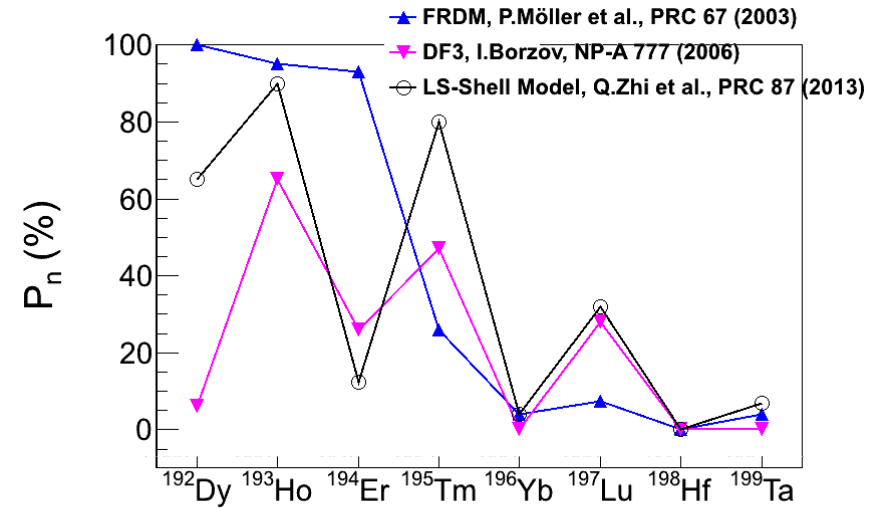
¹Institut für Kernphysik, Technische Universität Darmstadt, D-64289 Darmstadt, Germany
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Nucleosynthesis of the heaviest elements



N=126

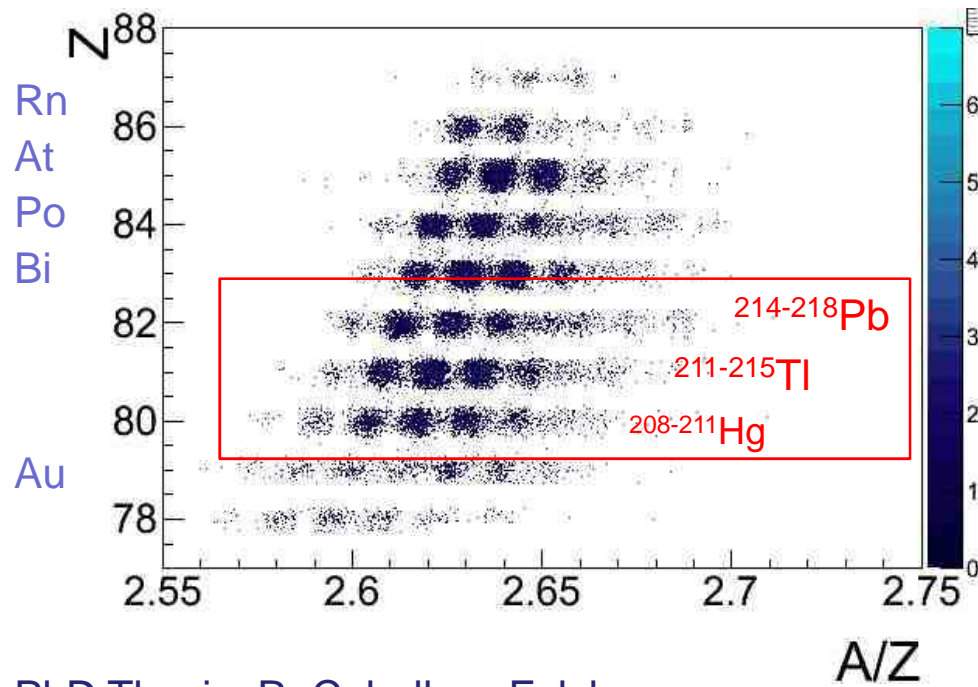
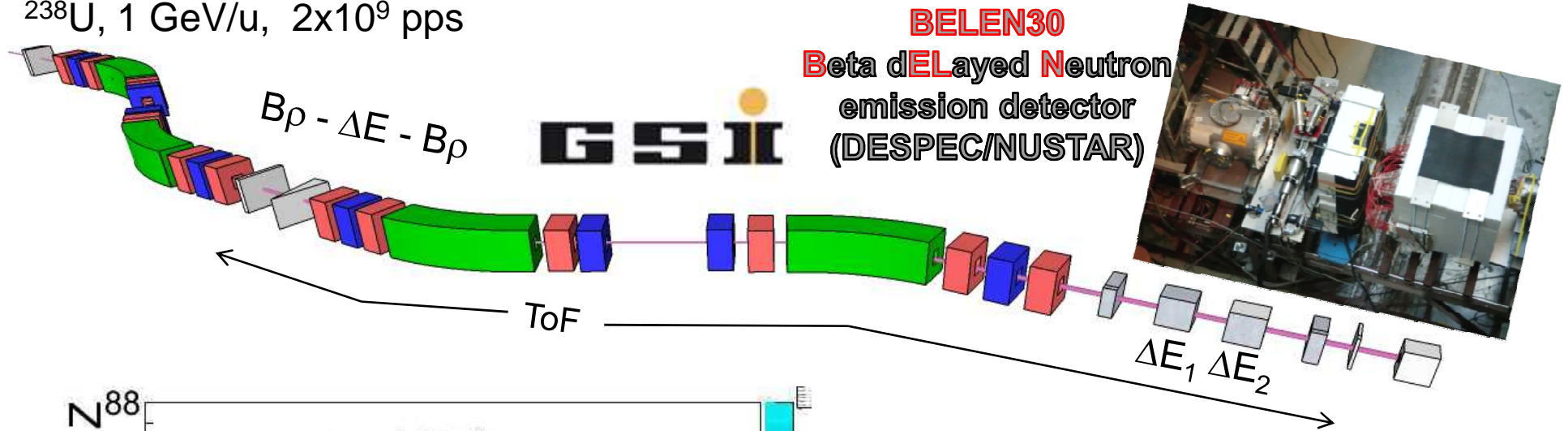


Only **one** experimental value is known: **TI-210** !!

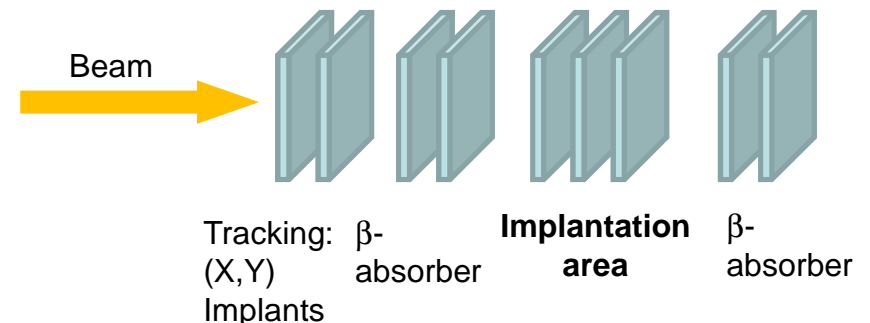
- A.V.Kogan, et al. Sov. Phys. JETP 5(1957) 365
- G.Stetter, TID-14880(1961)

Nucleosynthesis of the heaviest elements

^{238}U , 1 GeV/u, 2×10^9 pps



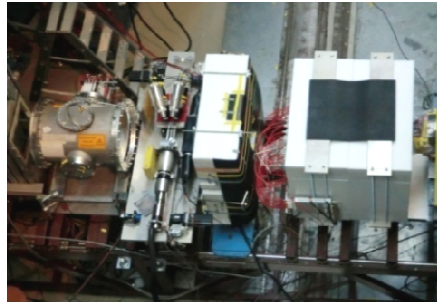
SIMBA
Silicon Implantation
Array
(TU-Munich)



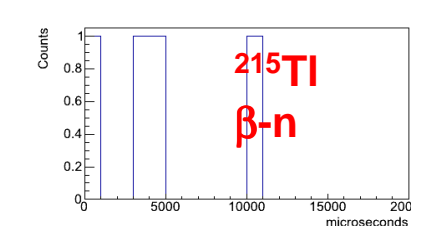
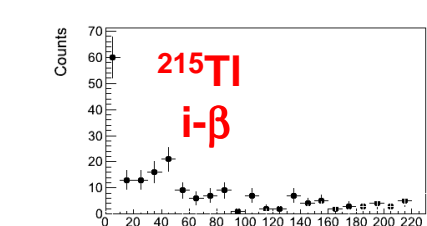
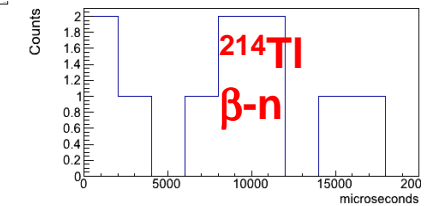
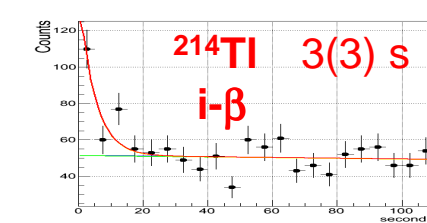
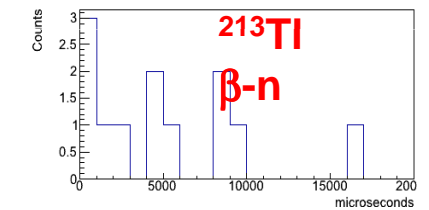
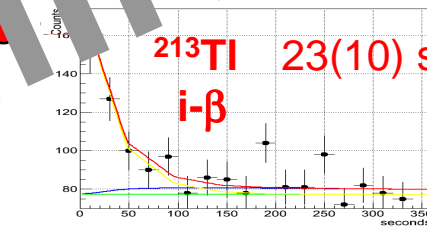
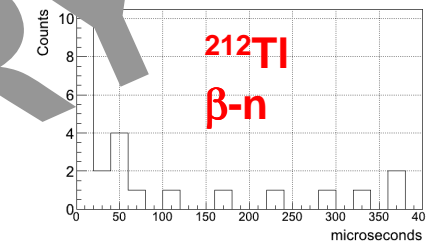
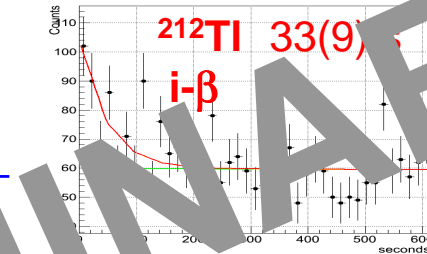
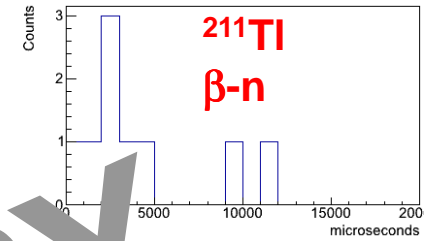
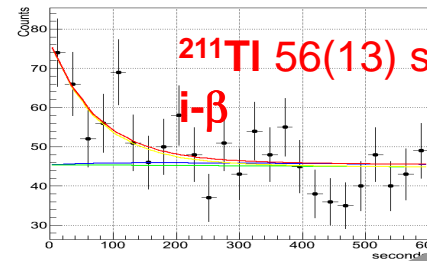
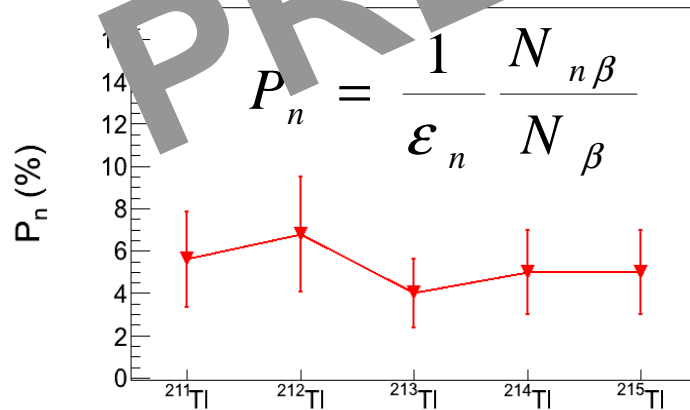
PhD Thesis: R. Caballero-Folch

Nucleosynthesis of the heaviest elements

BELEN
Beta dELayed Neutron
emission detector

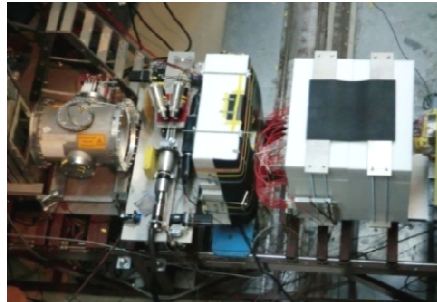


	Bi202	Bi203	Bi204	Bi205	Bi206	Bi207	Bi208	Bi209	Bi210	Bi211	Bi212	Bi213	Bi214
82	Pb201	Pb202	Pb203	Pb204	Pb205	Pb206	Pb207	Pb208	Pb209	Pb210	Pb211	Pb212	Pb213
	Tl200	Tl201	Tl202	Tl203	Tl204	Tl205	Tl206	Tl207	Tl208	Tl209	Tl210	Tl211	Tl212
80	Hg199	Hg200	Hg201	Hg202	Hg203	Hg204	Hg205	Hg206	Hg207	Hg208	Hg209	Hg210	Hg211
	Au198	Au199	Au200	Au201	Au202	Au203	Au204	Au205	Au206	Au207	Au208	Au209	Au210
78	Pt197	Pt198											
	Ir196	Ir197											
76	Os195	Os196											
	Re194												
74													



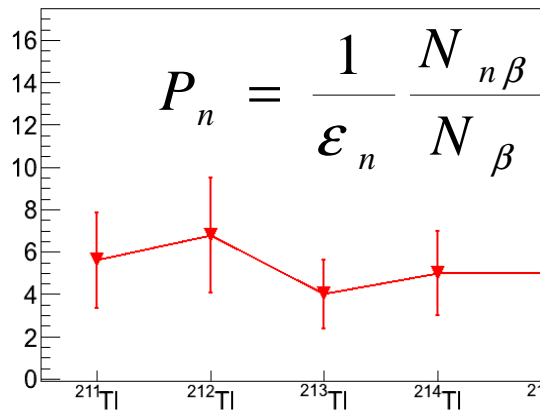
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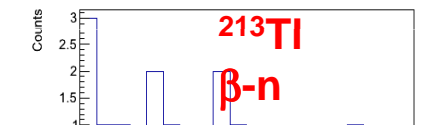
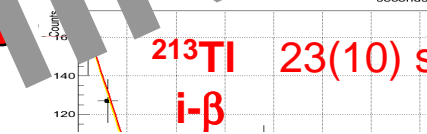
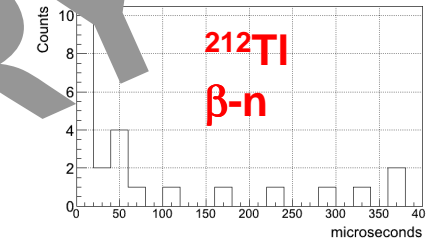
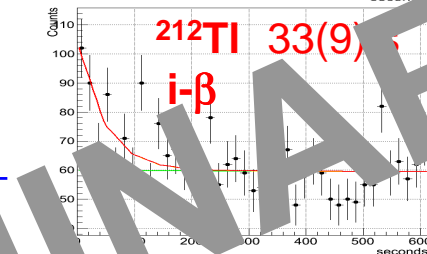
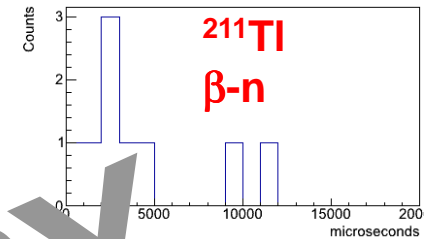
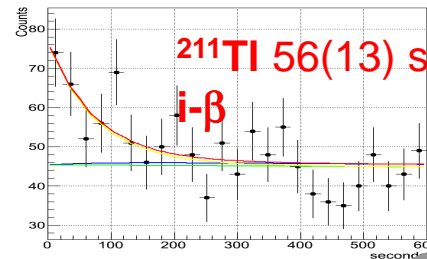
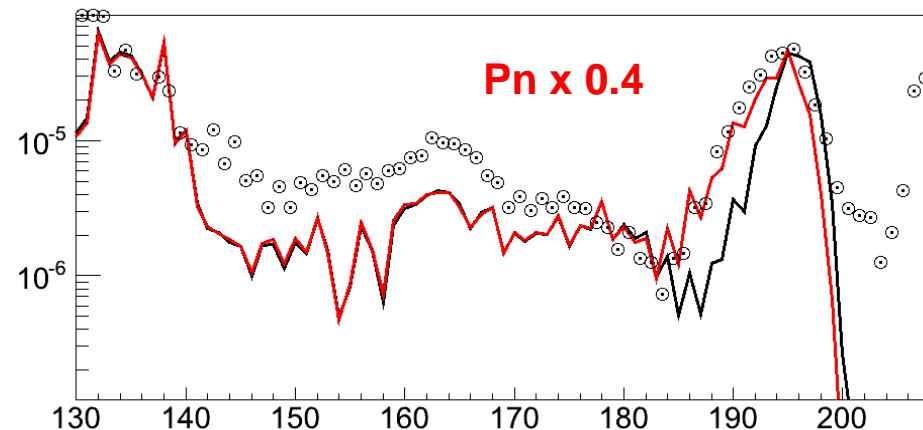


	Bi202	Bi203	Bi204	Bi205	Bi206	Bi207	Bi208	Bi209	Bi210	Bi211	Bi212	Bi213	Bi214
82	Pb201	Pb202	Pb203	Pb204	Pb205	Pb206	Pb207	Pb208	Pb209	Pb210	Pb211	Pb212	Pb213
	Tl200	Tl201	Tl202	Tl203	Tl204	Tl205	Tl206	Tl207	Tl208	Tl209	Tl210	Tl211	Tl212
80	Hg199	Hg200	Hg201	Hg202	Hg203	Hg204	Hg205	Hg206	Hg207	Hg208	Hg209	Hg210	Hg211
	Au198	Au199	Au200	Au201	Au202	Au203	Au204	Au205	Au206	Au207	Au208	Au209	Au210
78	Pt197	Pt198											
	Ir196	Ir197											
76	Os195	Os196											
	Re194												
74													

P_n (%)



Abundance

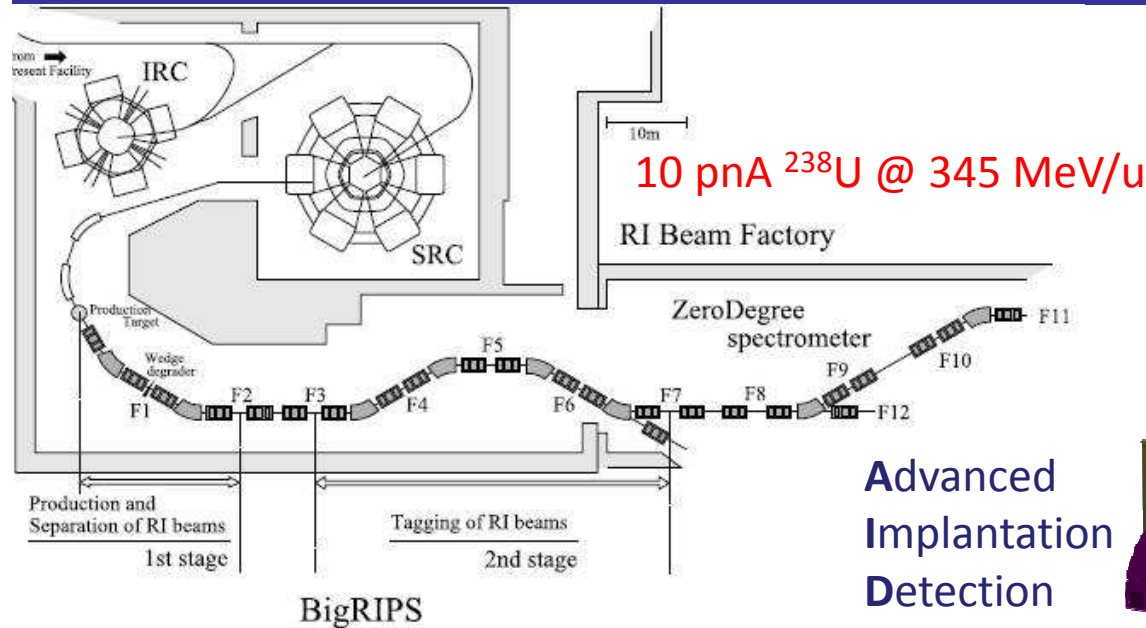


0 20 40 60 80 100 120 140 160 180 200 220

A

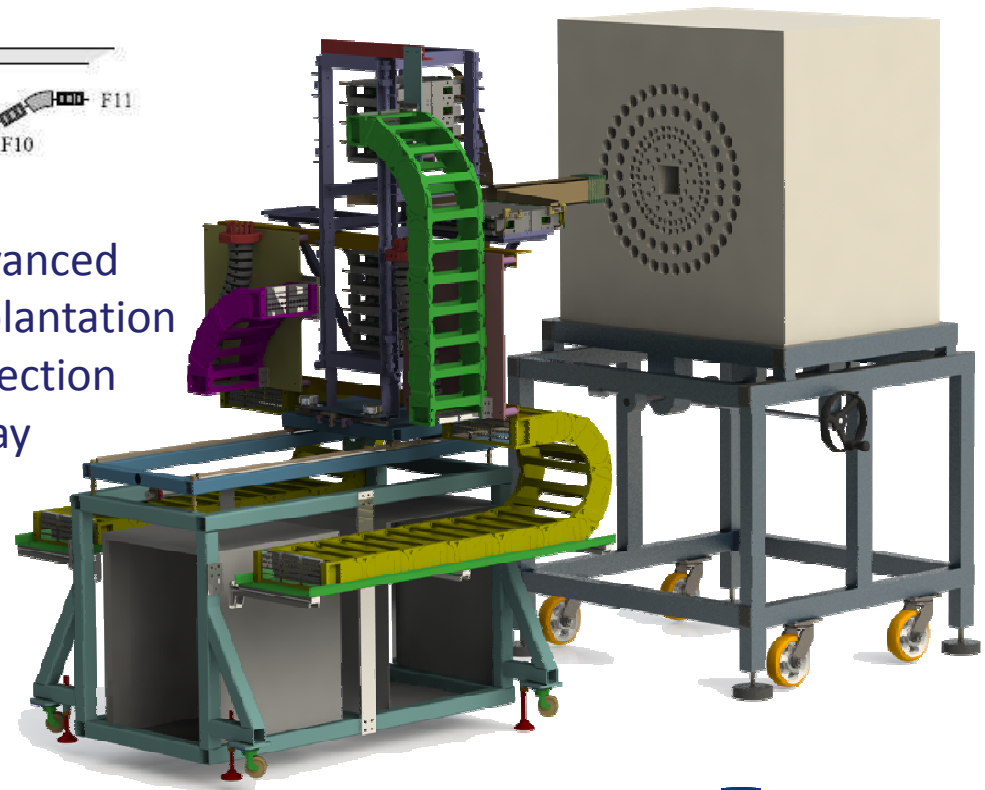
microseconds

BRIKEN: βn measurements of the most exotic nuclei



BRIKEN neutron detection set-up

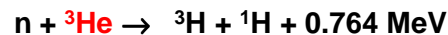
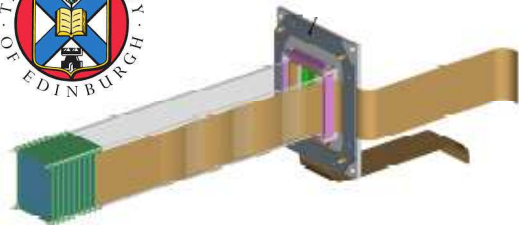
**Advanced
Implantation
Detection
Array**



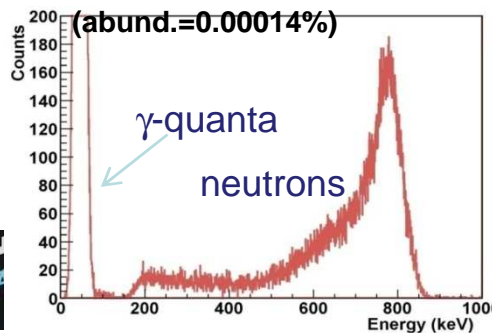
BRIKEN: βn measurements of the most exotic nuclei



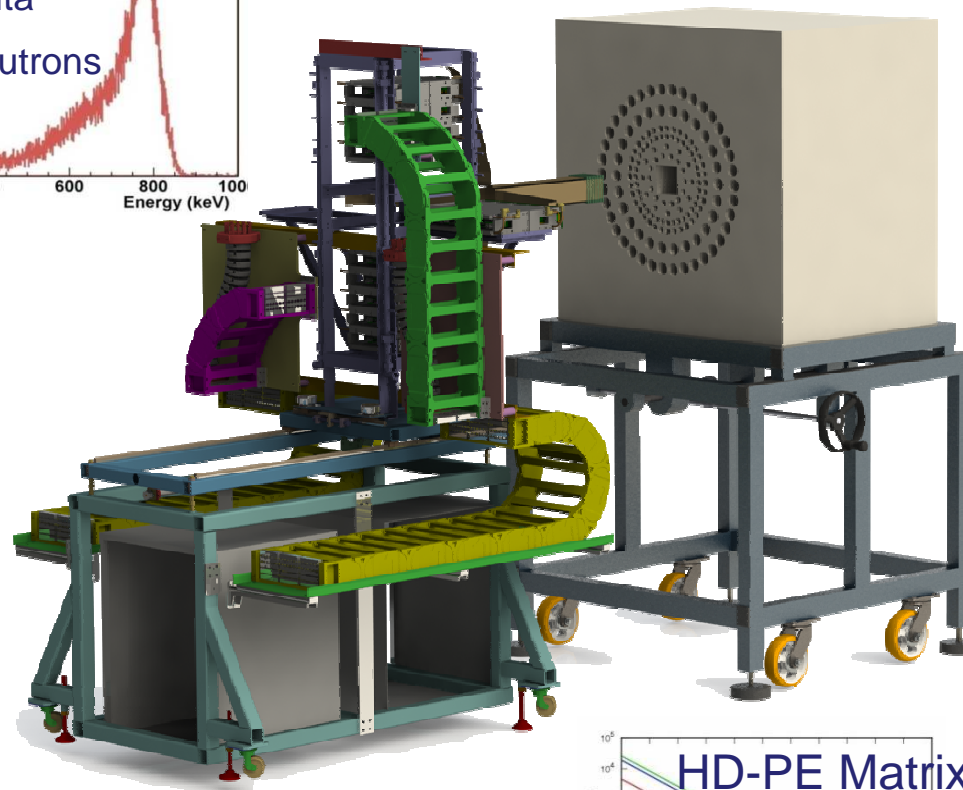
AIDA



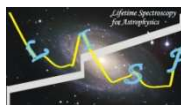
(abund. = 0.00014%)



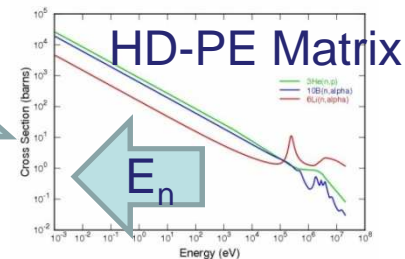
BRIKEN neutron detection set-up



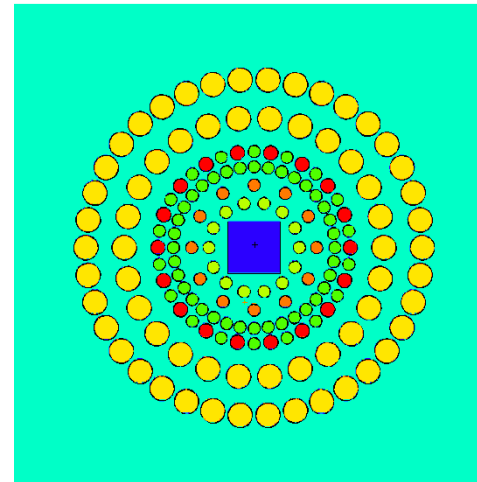
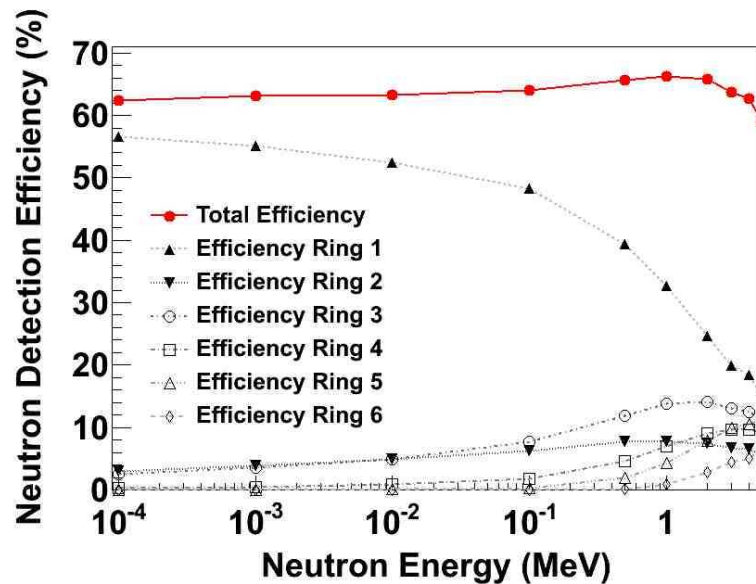
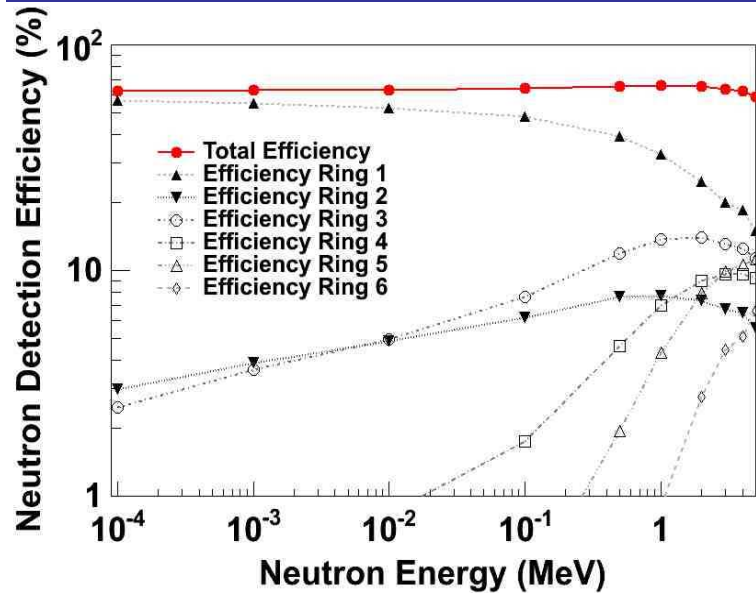
P (atm)	# Tubes	
10	10	GSI
4	20	JINR
10	67	ORNL
10	17	ORNL
5.13	26	RIKEN
8	42	UPC
Total	182	



NERO (NSCL)	3HEN (ORNL)	BELEN (FAIR)	ALTO (DESIR)
60	74	48	90



BRIKEN neutron detector array



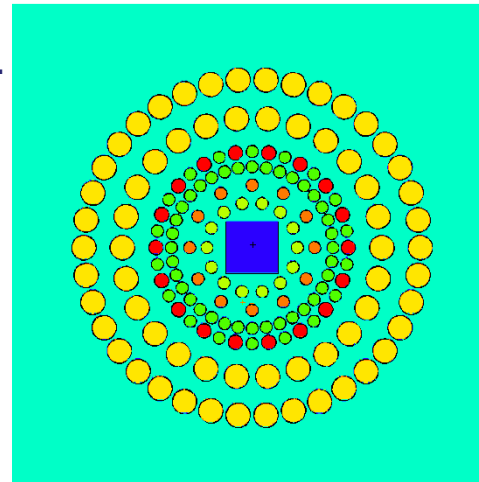
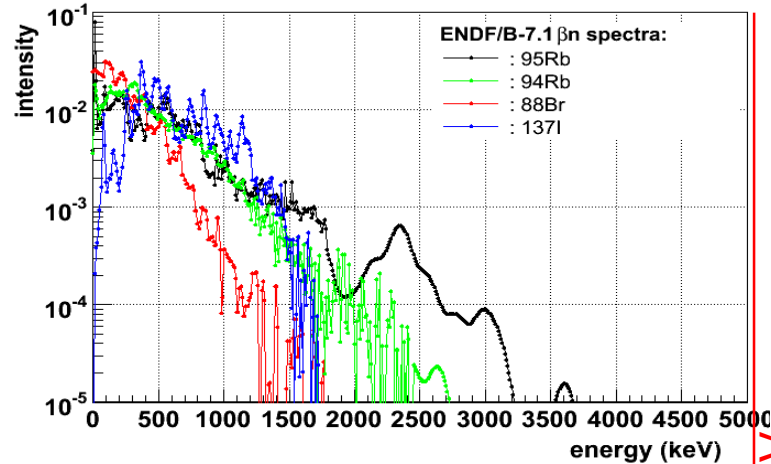
174 ^3He tubes of 6 different types:

Ring	Radius (cm)	# ^3He Tubes	Pressure (atm)	Diameter (inch)	Institute
1	9.4	14	10	1	ORNL
2	13	12+12	5.13	1	RIKEN
3	16.8	10+26	10/8	1	GSI/UPC
4	20	18+18	5/8	1.18/1	JINR/UPC
5	27	26	10	2	ORNL
6	35	38	10	2	ORNL

- High average efficiency of > 60 %
- Flat efficiency 6% up to 4 MeV, 12% up to 5 MeV.

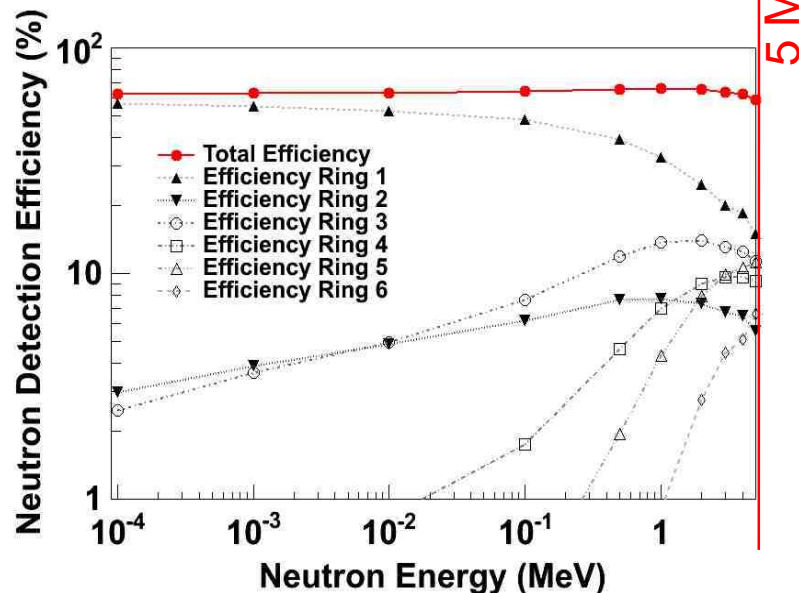
BRIKEN neutron detector array

- Eff. independent on neutron E-spectrum.
- E-spectrum can be inferred from rings info.



174 ^3He tubes of 6 different types:

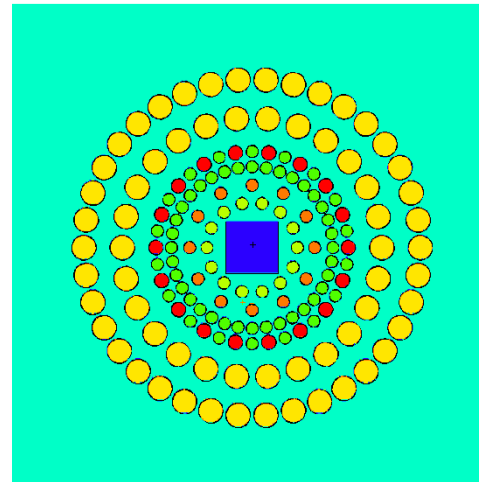
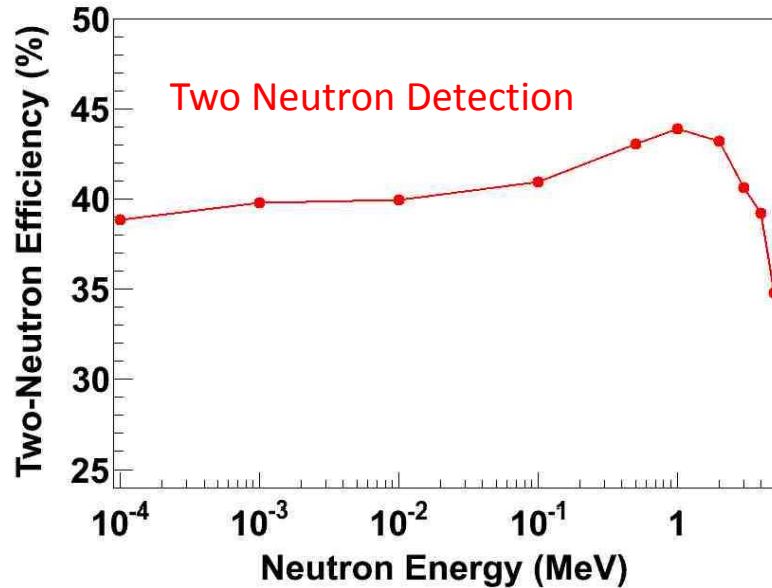
Ring	Radius (cm)	# ^3He Tubes	Pressure (atm)	Diameter (inch)	Institute
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✓ BRIKEN Construction Proposal Approved @ RIKEN-PAC
December, 2013

BRIKEN neutron detector array

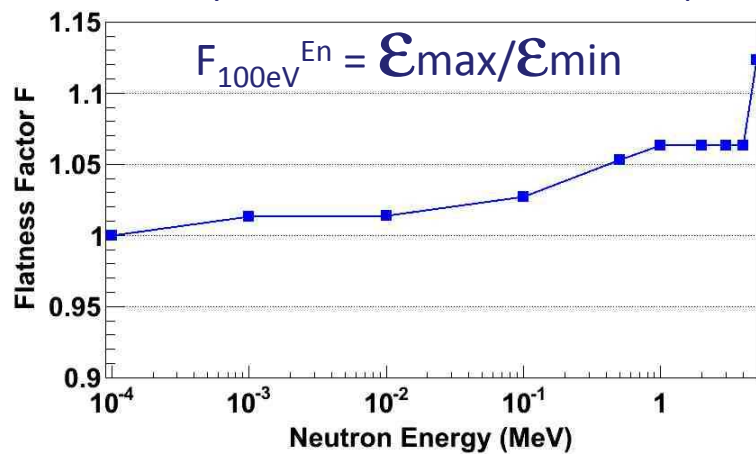
High efficiency also for two-fold neutron emission:



174 ³He tubes of 6 different types:

Ring	Radius (cm)	# ³ He Tubes	Pressure (atm)	Diameter (inch)	Institute
1	9.4	14	10	1	ORNL
2	13	12+12	5.13	1	RIKEN
3	16.8	10+26	10/8	1	GSI/UPC
4	20	18+18	5/8	1.18/1	JINR/UPC
5	27	26	10	2	ORNL
6	35	38	10	2	ORNL

Flat efficiency → Pn insensitive to neutron spectrum



✓ BRIKEN Construction Proposal Approved @ RIKEN-PAC December, 2013

BRIKEN Physics Goals for a Project Proposal in 2014

Presenter	Topic	Nuclei
S.Nishimura (RIKEN)	Below the 2nd r-process peak	^{112}Zr - ^{129}Pd
F. Montes (MSU)	2nd r-process peak	^{139}Sb
C.Domingo (IFIC)	Rare-earth r-process peak	^{151}La - ^{173}Tb
G. Lorusso (RIKEN)	2nd r-process peak	^{129}Ag - ^{142}Te , 133 - ^{134}Cd
M.Marta I.Dillmann (GSI/TRIUMF)	Multiple n emission	^{76}Co - ^{81}Cu , ^{134}Sn - ^{133}Cd
K.Rykaczewski (ORNL)	One and two n emitters above ^{78}Ni and ^{132}Sn	Ni, Cu, Zn, Ga, Ge, As, Se, In
R.Griwacz (U.Tennessee)	One and two n-emission below and at ^{78}Ni	Cl, Ar, K, Ca, Sc, Ti, Ni, Cu, Zn, Mn, Fe, Co
A Algora (IFIC)	Deformation $A \sim 110$	106 - ^{110}Zr , 110 - ^{114}Mo
B. Rubio (IFIC)	Nuclear structure $\sim ^{132}\text{Sn}$	^{130}Ag - ^{138}Sb
A.Estrade (Edinburgh)	Masses	Several
J.L. Tain (IFIC)	β -strength NE of ^{78}Ni	^{85}Ge - ^{97}Br
D.Cano-Ott (CIEMAT)	Reactor technologies	^{86}Ge , ^{96}Rb , ^{100}Rb , $^{98\text{m}}\text{Y}$, ^{131}Cd , ^{137}Sb

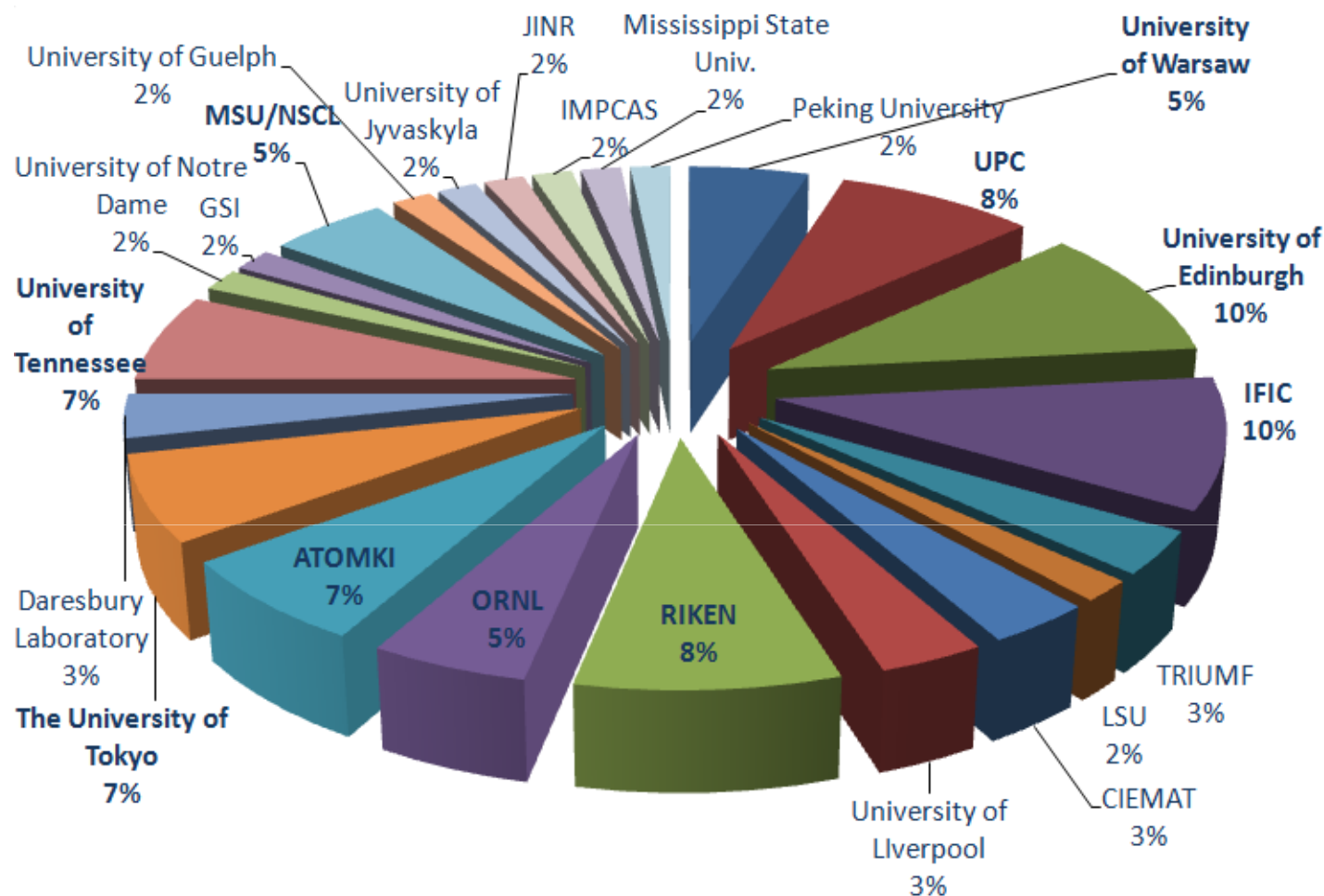
PROJECT PROPOSAL

Astrophysics:
r-process
nucleosynthesis

Nuclear
Structure:
1n, 2n-
competition
in heavy
nuclei, FF vs.
GT, etc

New reactor
technologies

BRIKEN Collaboration



>50 Scientists
>20 Research Centers

Open project, to join: briken.project@gmail.com

<http://indico.ific.uv.es/indico/event/briken> 1st BRIKEN Workshop, Valencia 17-18 Dec. 2012

<http://indico.ific.uv.es/indico/event/briken2> 2nd BRIKEN Workshop, Tokyo, 30-31 July 2013

BRIKEN: Summary & Outlook

- Beta-delayed neutrons will be one of the key Gross properties we will aiming at measuring in the next generation of RIB facilities, like NUSTAR-FAIR.
- βn measurements represent an stringent test for nuclear models far-off stability and how well the nuclear structure details (beta-strength function) are included.
- In stellar nucleosynthesis β -delayed neutron emission plays a relevant role for understanding both the observed r-process distributions and dynamical evolution.
- We intend to study these aspects in the framework of the BRIKEN Project devoted to the measurement of the most exotic nuclei at the RIB facility of RIKEN.
- BRIKEN is a joint international effort, to join instrumentation and expertise in order to build a high-performance –high+flat efficiency- neutron detector array, to be set-up and operated for an experimental campaign at RIKEN.
- Physics proposals will be submitted within the same BRIKEN “umbrella” at the next NP-PAC in June, 2014. The project is open, new collaborators are welcome to join!

Thank you for your attention!

Thanks to BRIKEN Collaborators

Agnieszka Korgul	University of Warsaw	Toshiuki Kubo	RIKEN
Albert Riego	UPC	Marc Labiche	Daresbury Laboratory
Alfredo Estrade	University of Edinburgh	Keishi Matsui	University of Tokyo
Alejandro Algora	IFIC	Megum Niikura	University of Tokyo
Anu Kankainen	University of Edinburgh	Michele Marta	GSI
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Belen Gomez	UPC	Fernando Montes	NSCL
Berta Rubio	IFIC	Nathan Brewer	University of Tennessee
Francisco Calvino	UPC	Shunji Nishimura	RIKEN
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