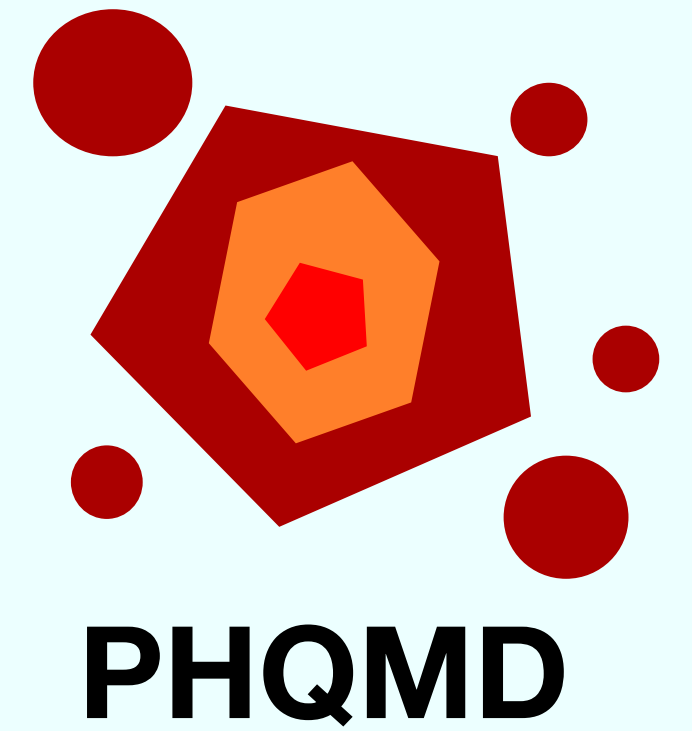


# Probing of EoS with Clusters and Hypernuclei (arXiv:2507.14255)

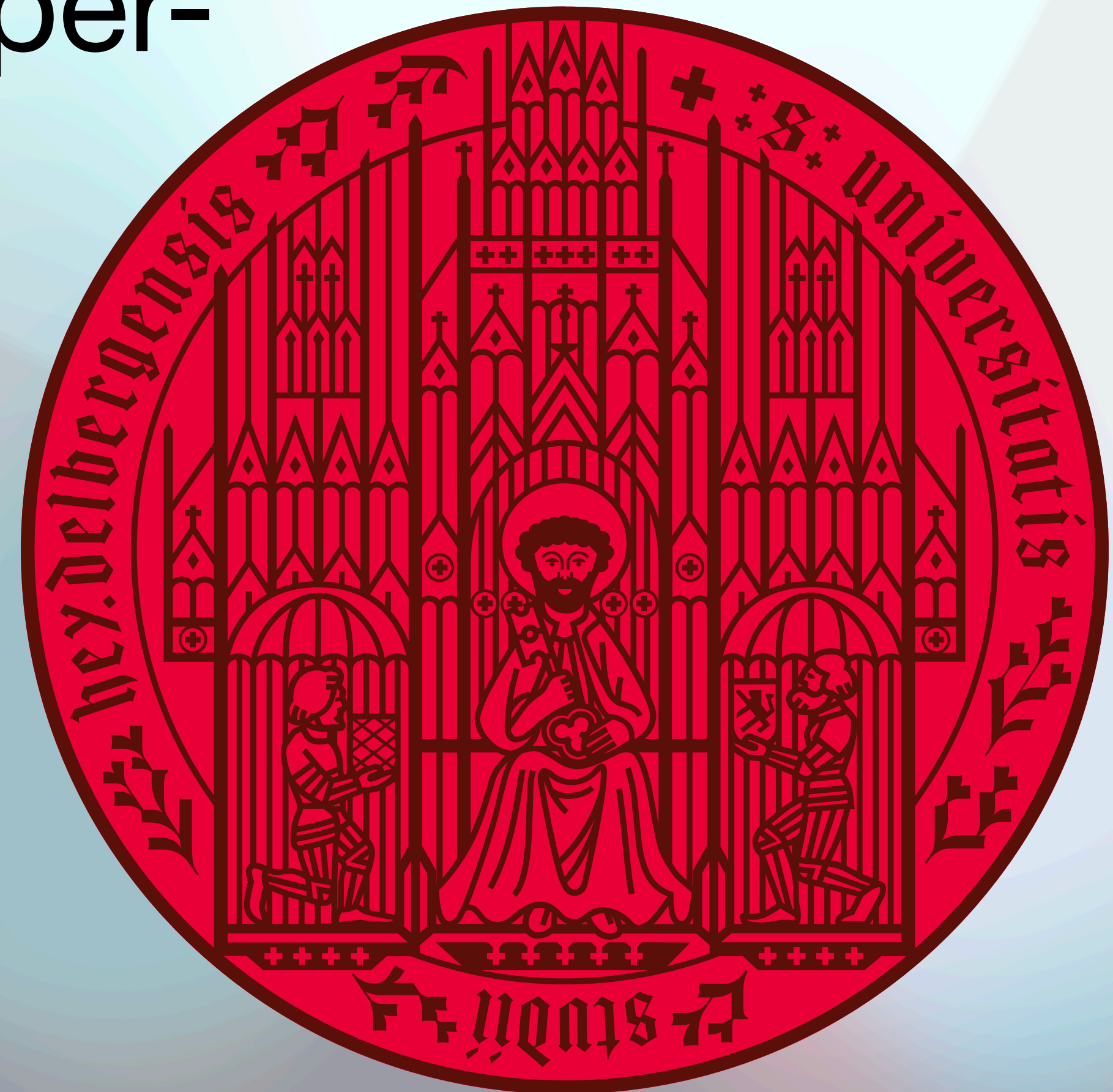


## 5<sup>th</sup> Workshop on Anti-Matter, Hyper-Matter and Exotica Production

Y. Zhou, S. Gläsel, Y. H. Leung, V. Kireyeu, J. Zhao,  
G. Coci, C. Blume, I. Vassiliev, V. Voronyuk, M. Winn,  
N. Herrmann, N. Xu, J. Aichelin, E. Bratkovskaya

University of Heidelberg

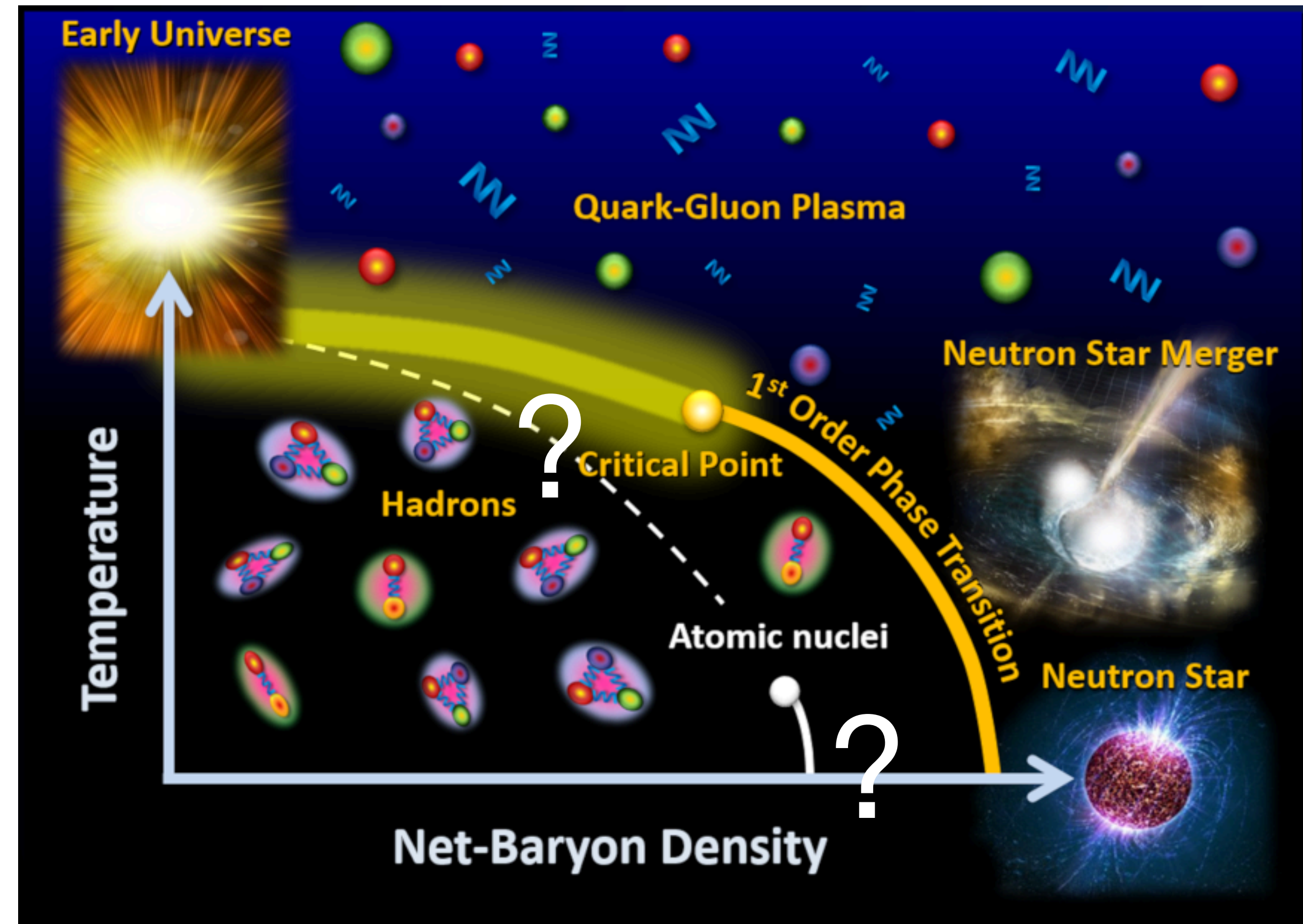
13<sup>th</sup> Nov, 2025





# Motivations

- Quest for **Equation of state (EoS)** of strongly interacting matter: a major objective of nuclear physics
- EoS at **high baryon density**: crucial for understanding the behavior of nuclear matter under extreme conditions, such as neutron star cores
- **STAR experiment** recently published new data in Au+Au collisions at  $\sqrt{s_{NN}} = 3$  GeV: highest quality dataset to date
- **Parton-Hadron-Quantum-Molecular Dynamics (PHQMD)** model is employed to interpret the new STAR data and investigate the sensitivity of various observables to different equation-of-state scenarios



# Momentum-Dependent Potential

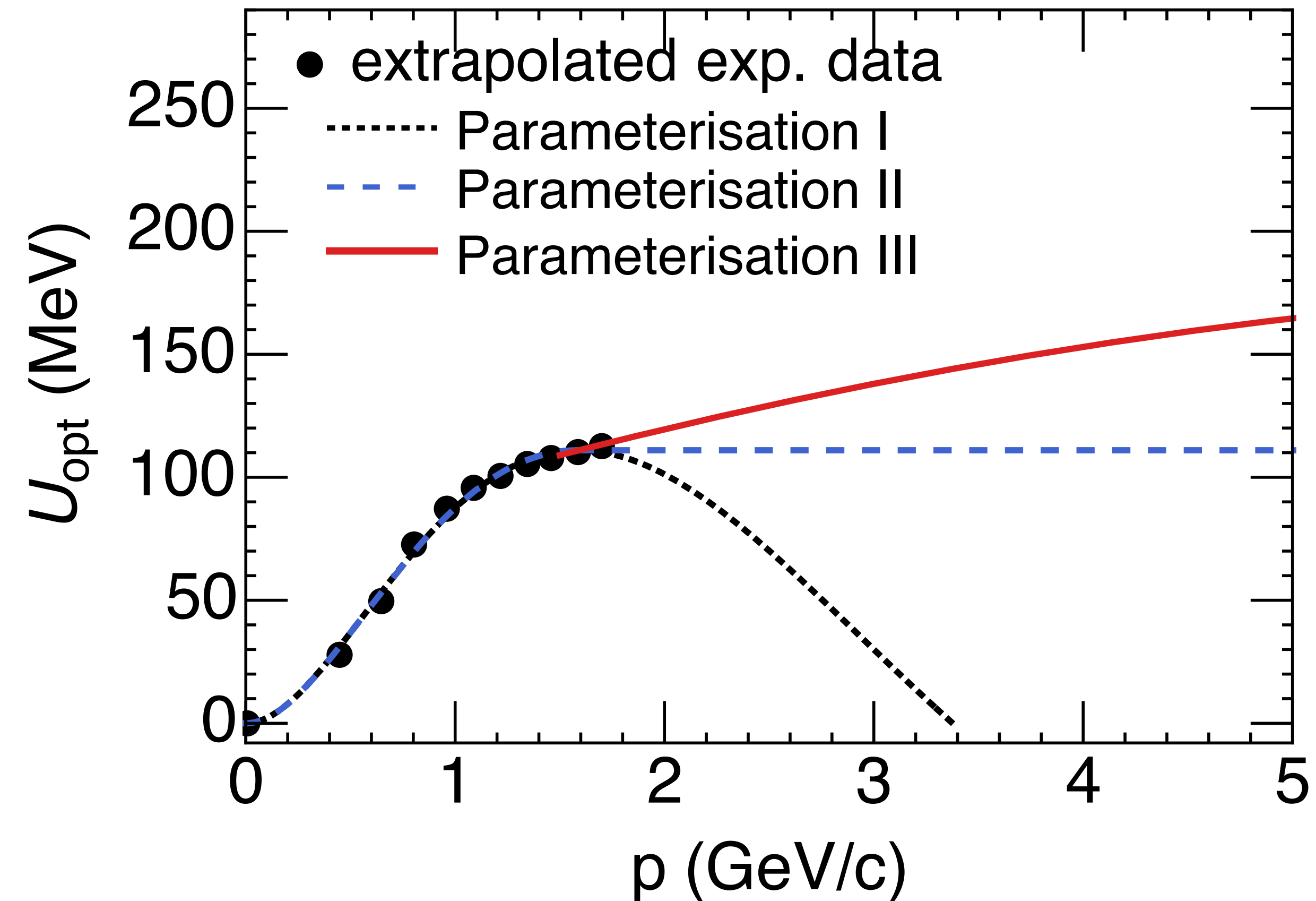
- For details on the PHQMD model, see S. Gläsel talk on 11/11 (Tue)!
- Consider two **static** EoS models, labeled “soft (**S**)” and “hard (**H**)“, which differ in compressibility modulus  $K$ , and a **momentum-dependent** soft EoS model (**SM**)

- Momentum-dependent potential

$$V(\mathbf{r}_1, \mathbf{r}_2, \mathbf{p}_{01}, \mathbf{p}_{02}) \\ = (a\Delta p + b\Delta p^2) \times \exp[-c\sqrt{\Delta p}] \delta(\mathbf{r}_1 - \mathbf{r}_2)$$

- Parameters  $a$ ,  $b$ , and  $c$ : fitted to the “optical” potential extracted from elastic scattering data in pA

$$U_{SEQ}(p) = \frac{\int^{p_F} V(\mathbf{p} - \mathbf{p}_1) dp_1^3}{\frac{4}{3}\pi p_F^3}$$





# Modeling the EoS within PHQMD

- In infinite matter a potential corresponds to an EoS for cold nuclear matter:

$$V_{Skryme,stat} = \alpha \frac{\rho}{\rho_0} + \beta \left( \frac{\rho}{\rho_0} \right)^\gamma$$

$$V_{mom} = (a\Delta p + b\Delta p^2) \exp[-c\sqrt{\Delta p}] \frac{\rho}{\rho_0}$$

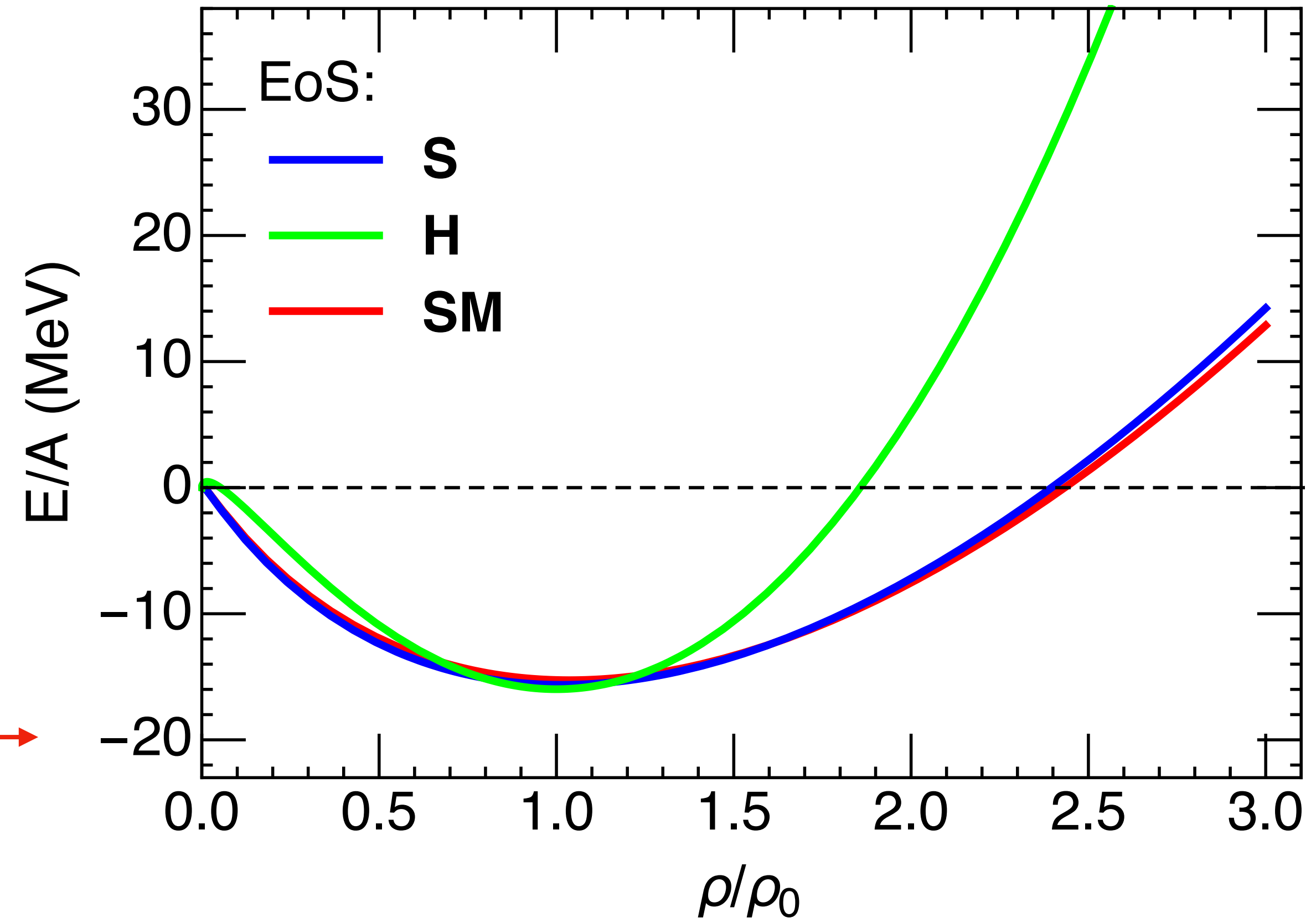
$$\frac{E}{A}(\rho) = \frac{3}{5}E_F + V_{Skryme,stat}(\rho) + V_{mom}(\rho)$$

- Parameters  $\alpha$ ,  $\beta$ , and  $\gamma$ , and compressibility modulus  $K$  is varied to construct three EoS

EoS	$\alpha$ [GeV]	$\beta$ [GeV]	$\gamma$	K [MeV]
S	-0.3835	0.3295	1.15	200
H	-0.1253	0.071	2.0	380
SM	-0.478	0.4137	1.1	200

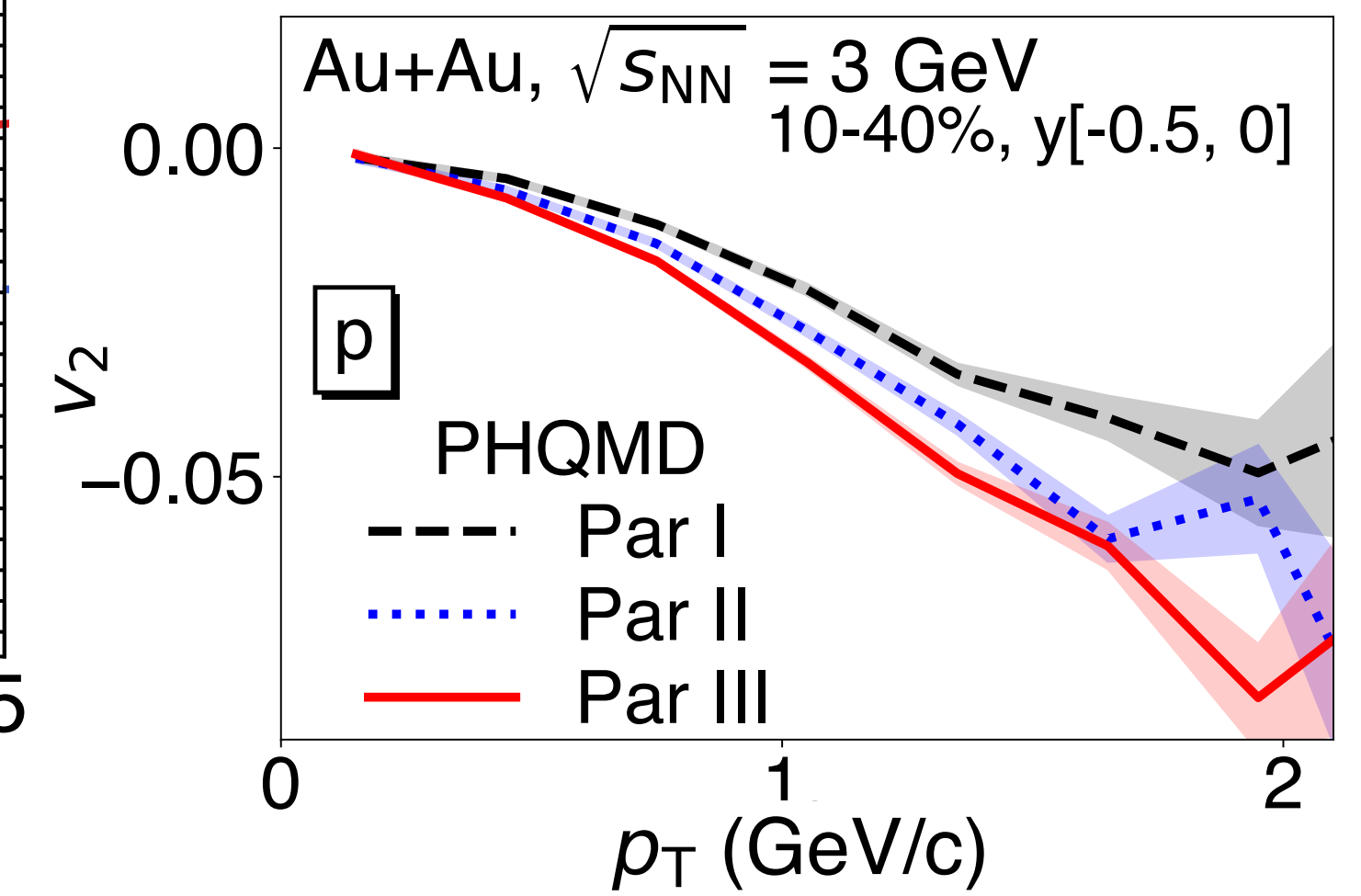
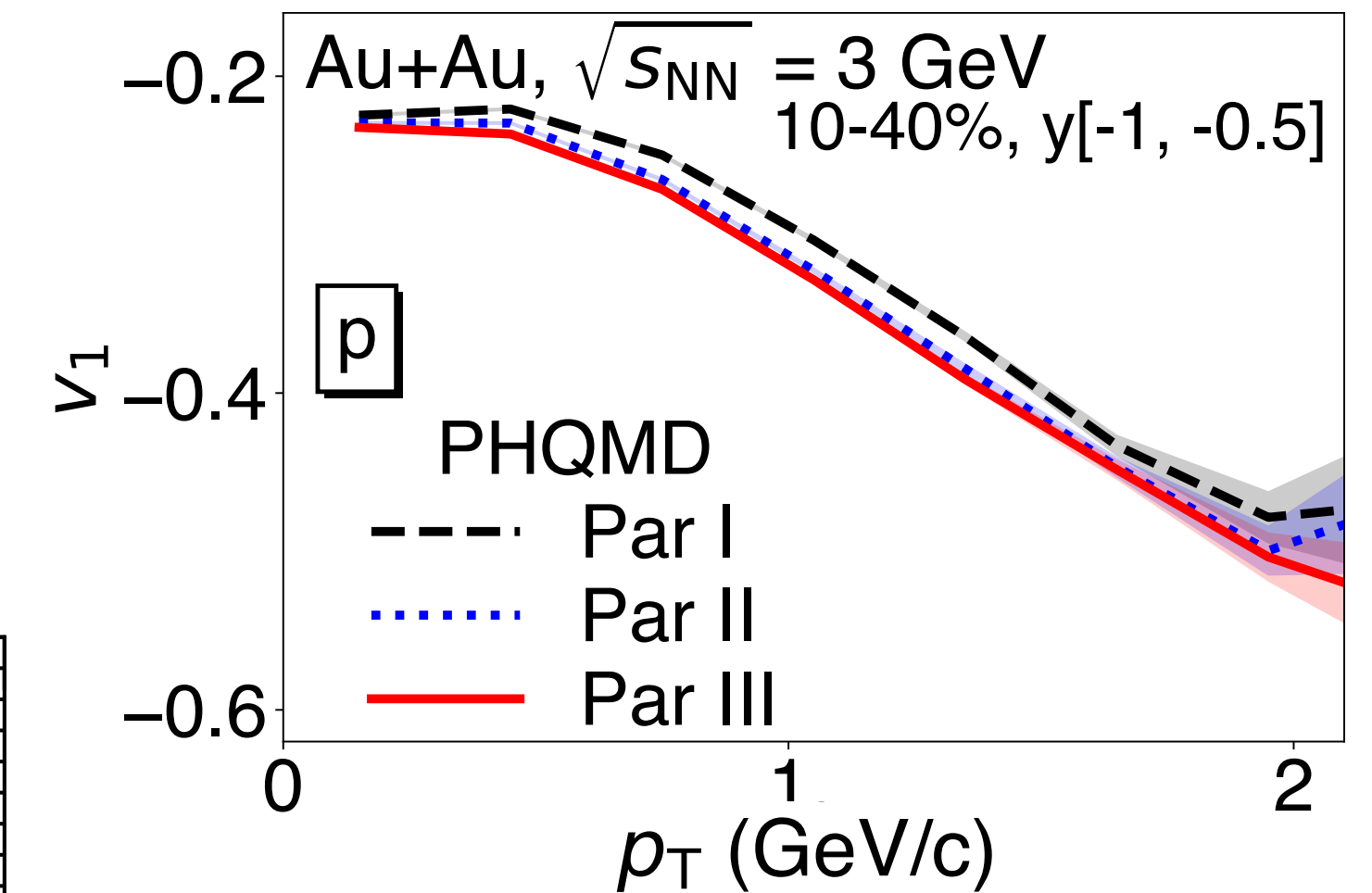
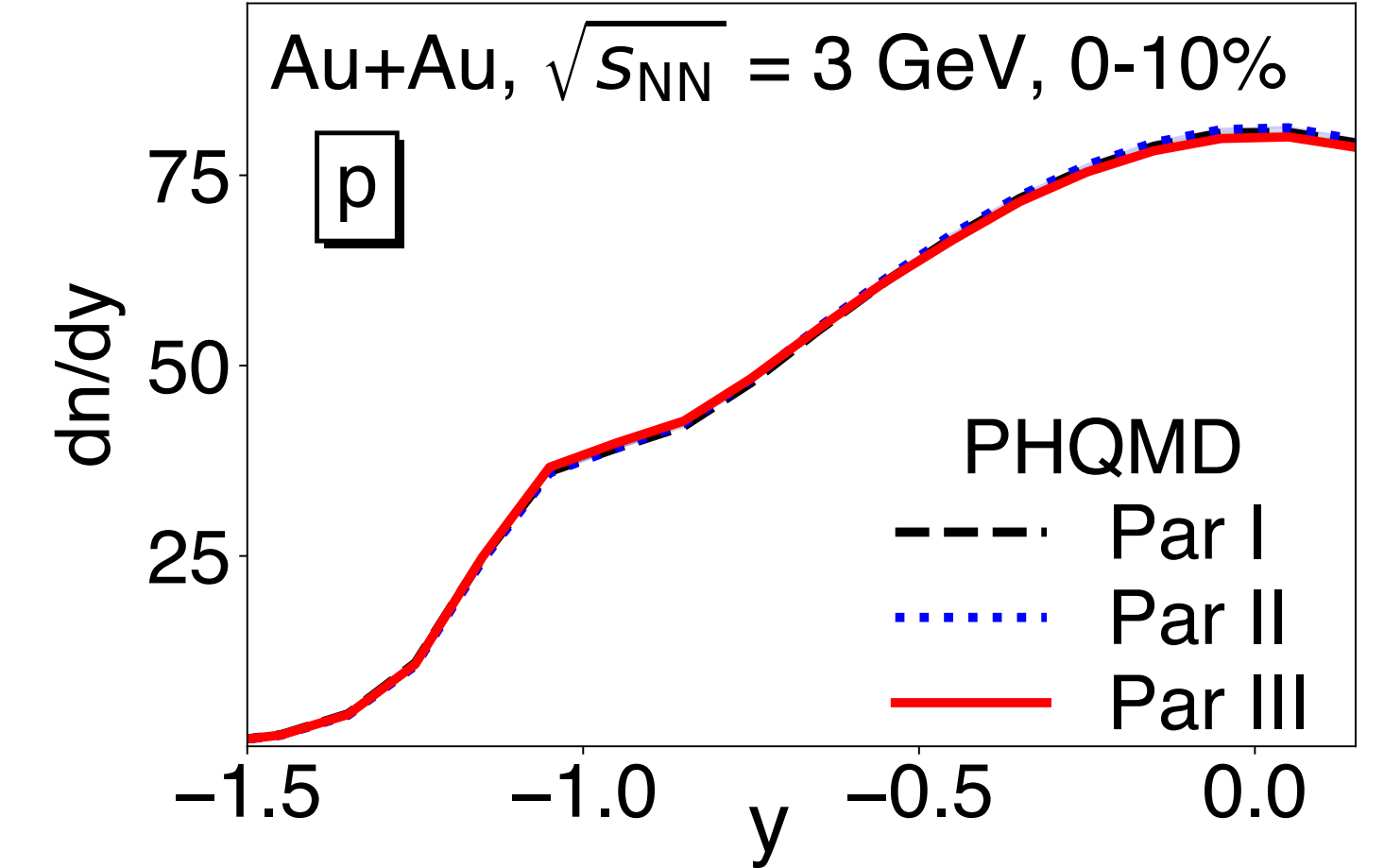
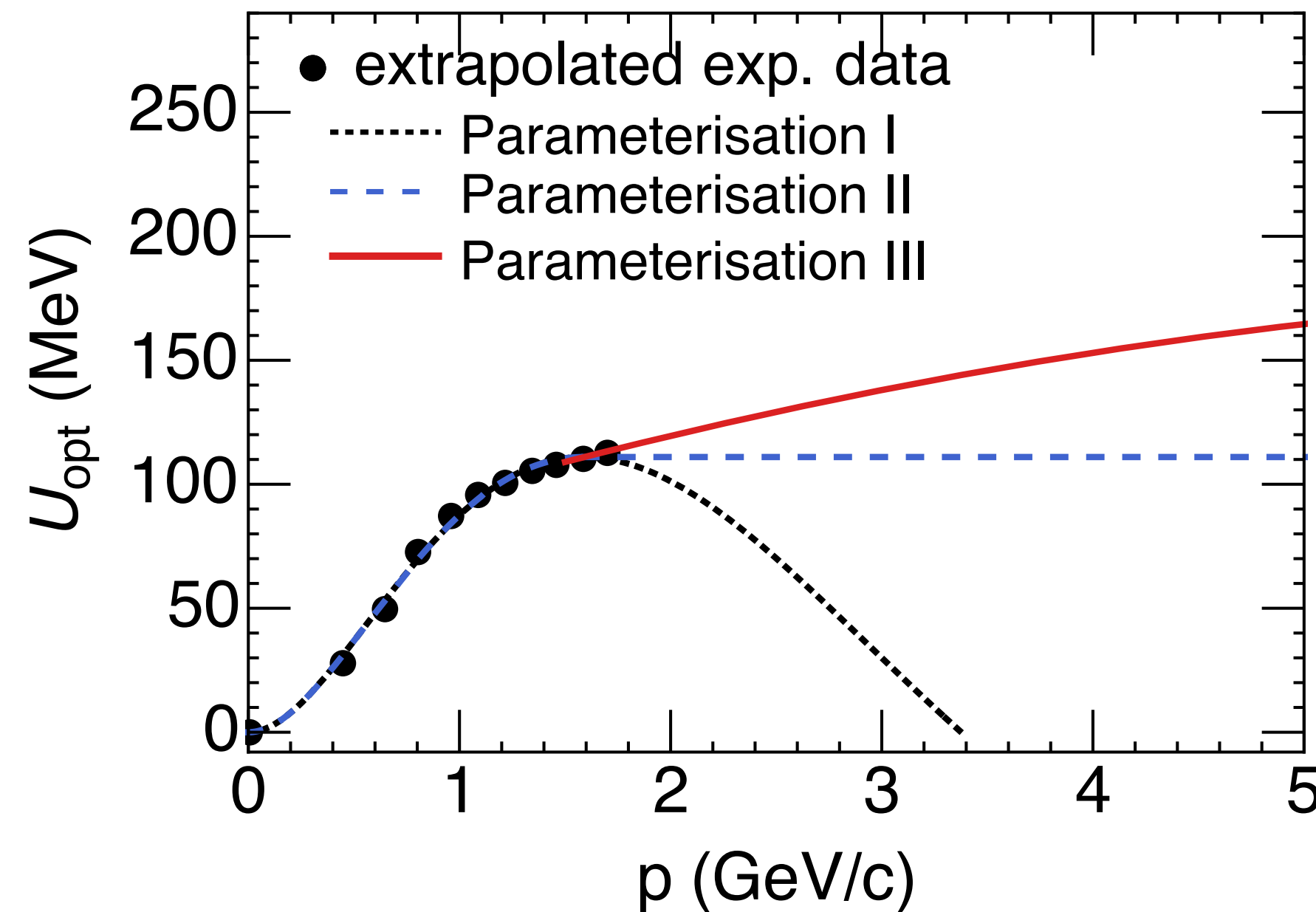
  

$a$ [GeV <sup>-1</sup> ]	$b$ [GeV <sup>-3</sup> ]	$c$ [GeV <sup>-1</sup> ]
236.326	-20.730	0.901



# Influence of $U_{opt}(p)$ on Observables

- No experimental data allow for a reliable extrapolation of  $U_{opt}(p)$  to large  $p$
- At 3 GeV,  $dN/dy$  and  $v_1$  distributions is practically identical for the three parametrizations;  $v_2$  shows mild differences at higher  $p_T$
- Collisions at higher energies (e.g. 5 GeV) will be more sensitive to this extrapolation

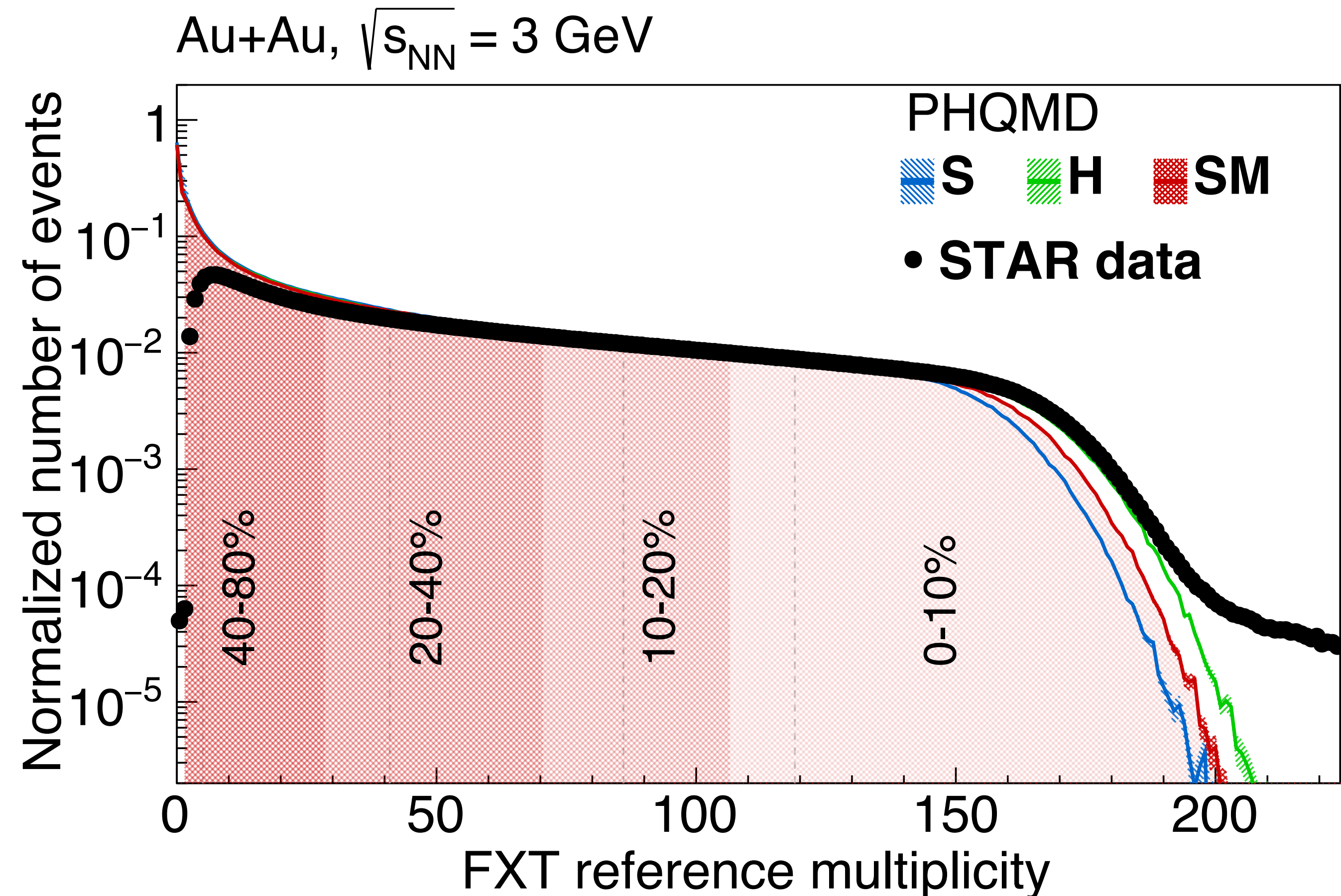




# Centrality Definition

- Experimental data: the centrality class is defined by a **Glauber Model fit to the measured charged particle multiplicity** (FXT reference multiplicity, FXTMult) within the acceptance of the STAR Time Projection Chamber (TPC) ( $-2.0 < \eta < 0$ )
- PHQMD: construct FXTMult using  $\pi$ , K, p:
  - $-2.0 < \eta < 0$
  - $p_T > 0.2$  GeV for  $\pi$ , K,  $> 0.35$  for p
- Centrality  $C_M$  defined by:

$$C_M = \frac{1}{\sigma^{AA}} \int_M^\infty dM' \frac{d\sigma}{dM'}$$



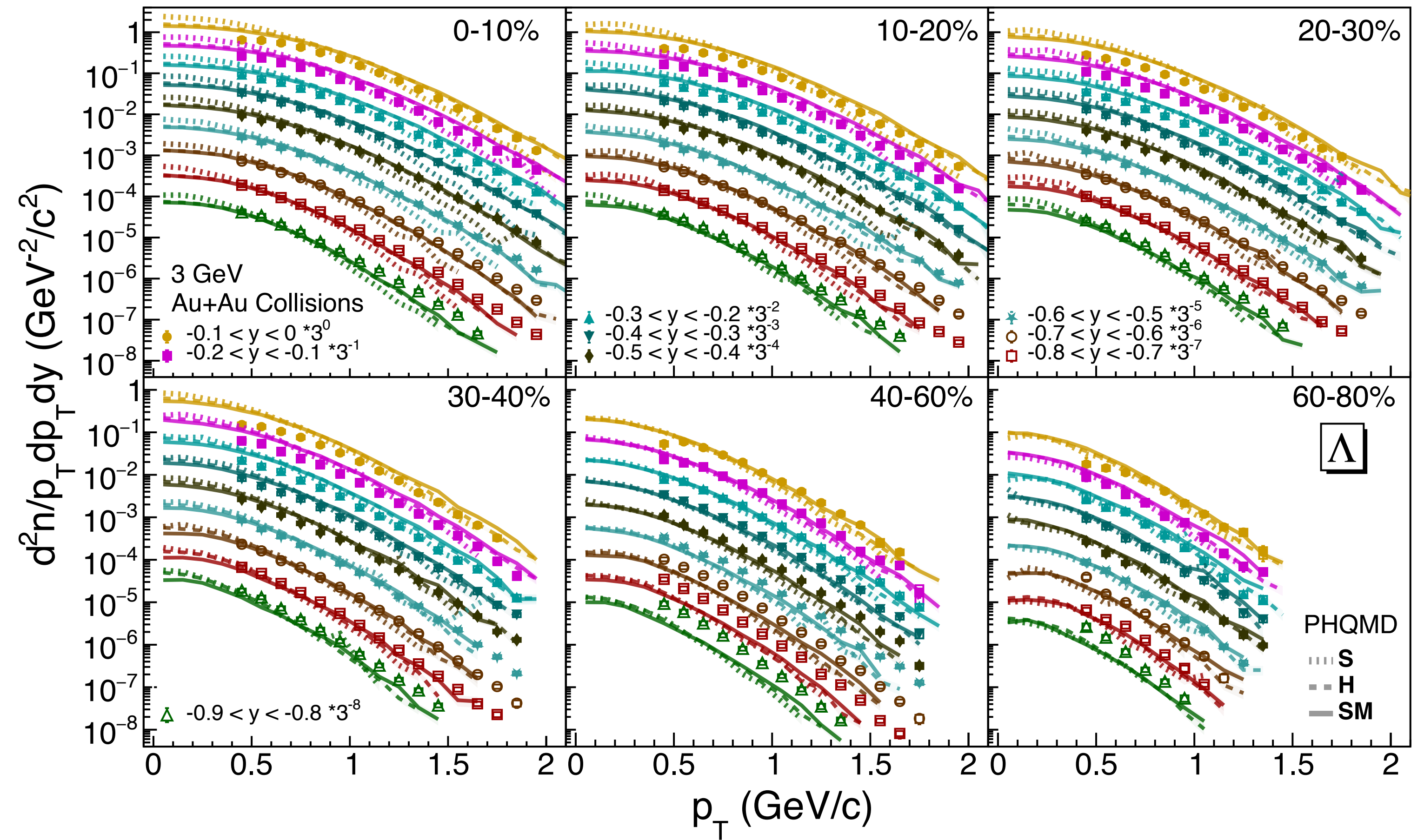
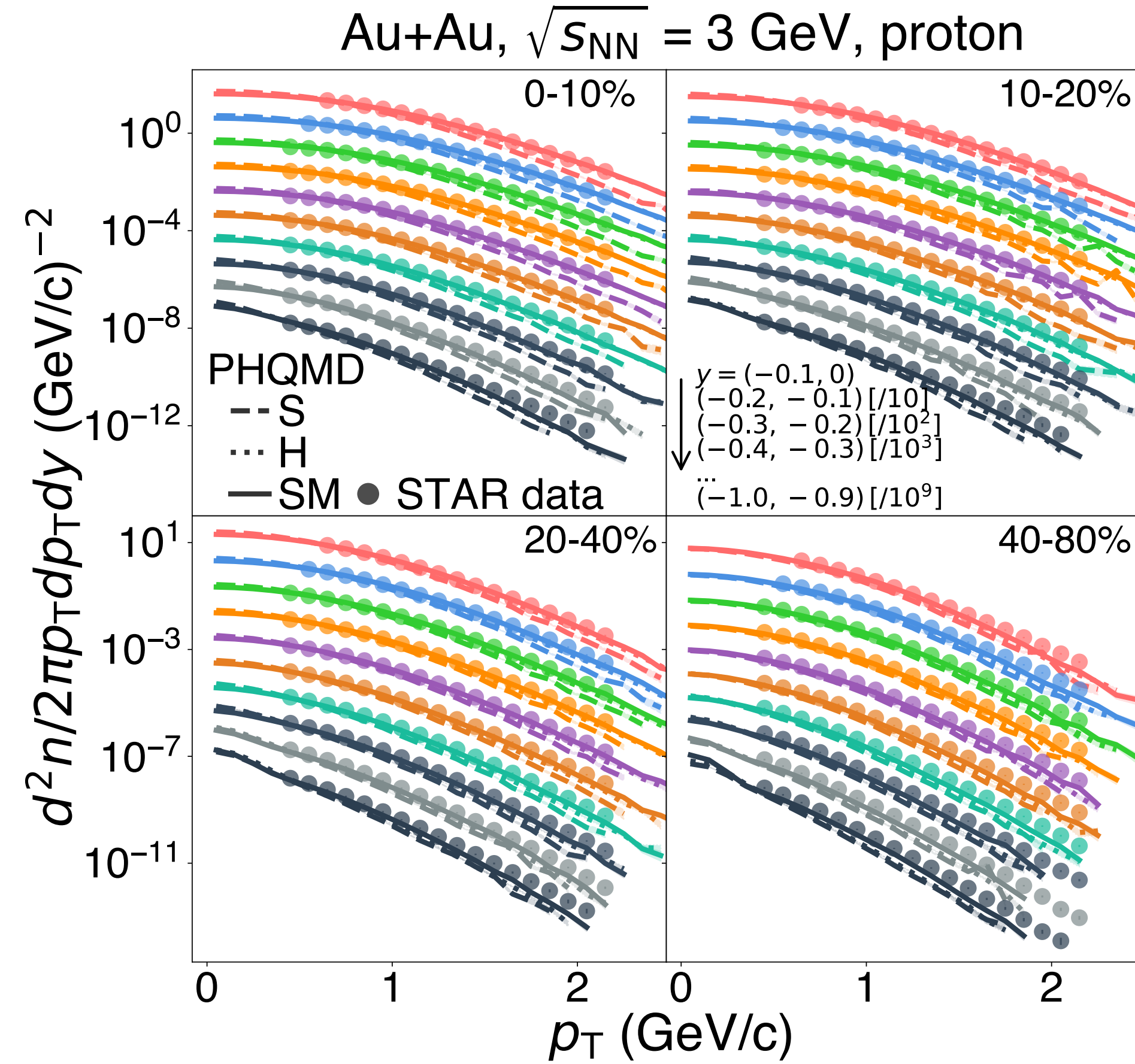
# Protons and $\Lambda$

## STAR data

- proton  $p_T$  spectra: Phys. Rev. C 110 (2024) 54911
- $\Lambda$   $p_T$  spectra: JHEP 2024 (2024) 139
- proton,  $\Lambda$   $v_1, v_2$ : Phys. Lett. B 827 (2022) 137003



# Proton and $\Lambda$ Transverse Momentum Distributions at 3 GeV

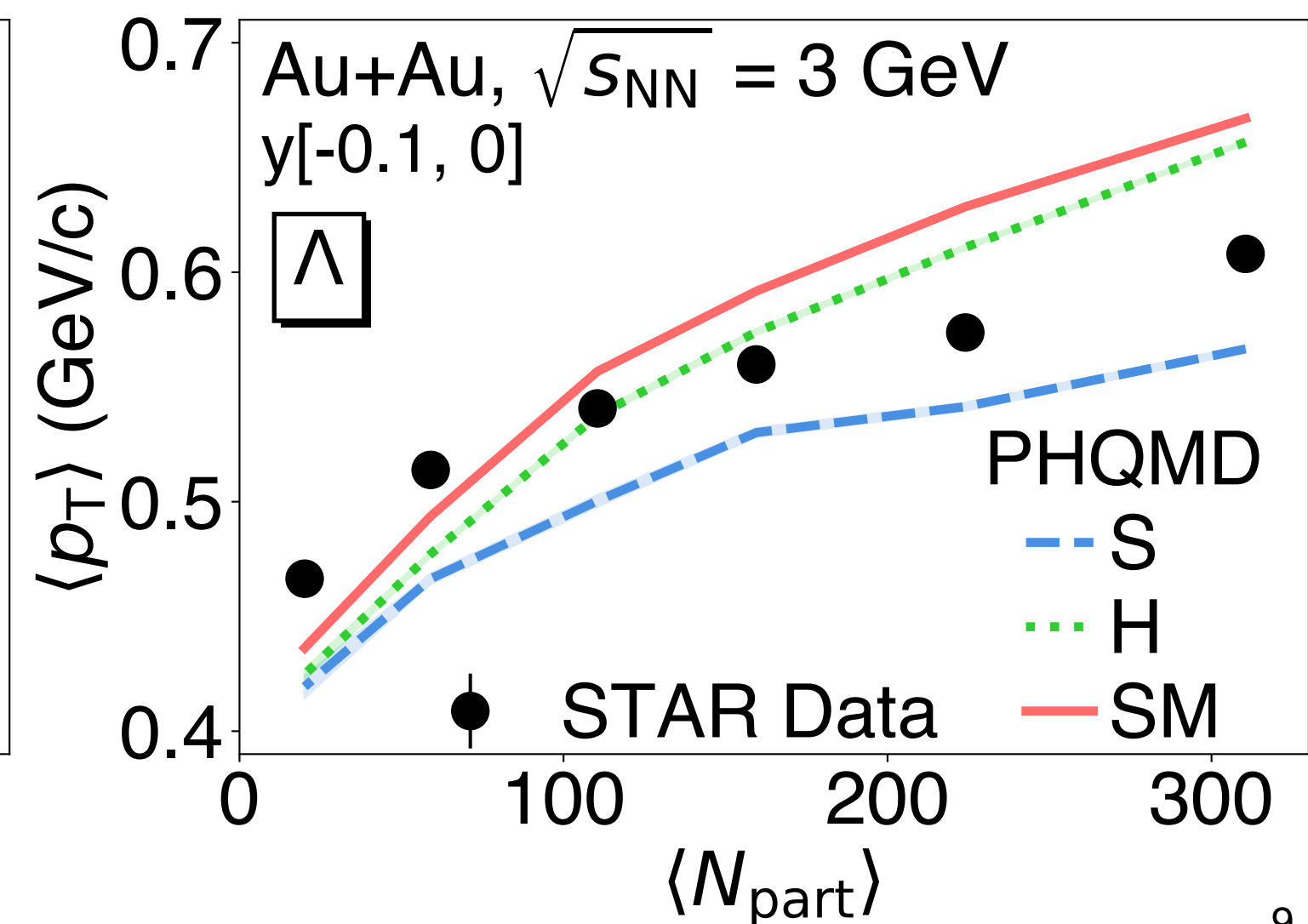
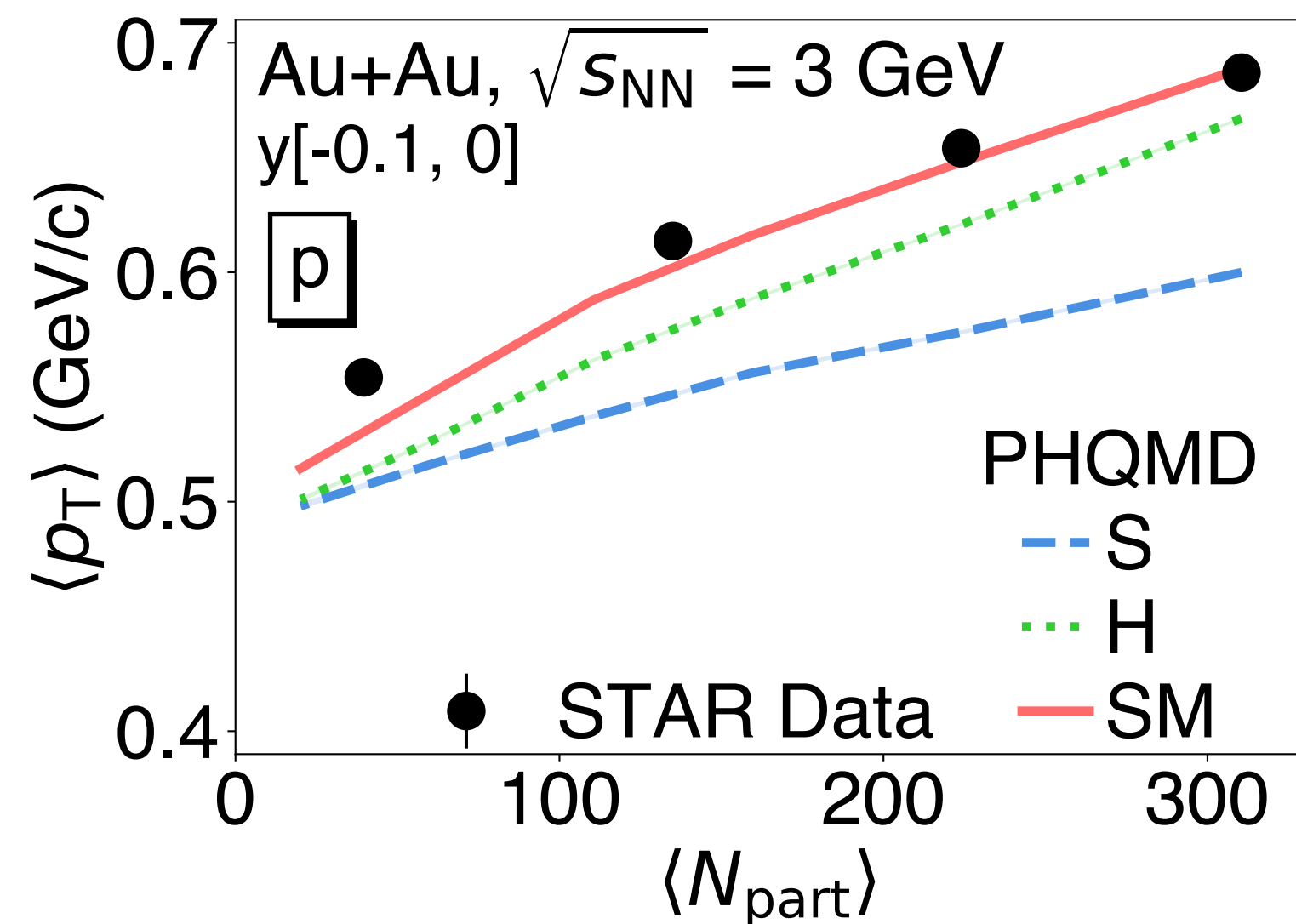
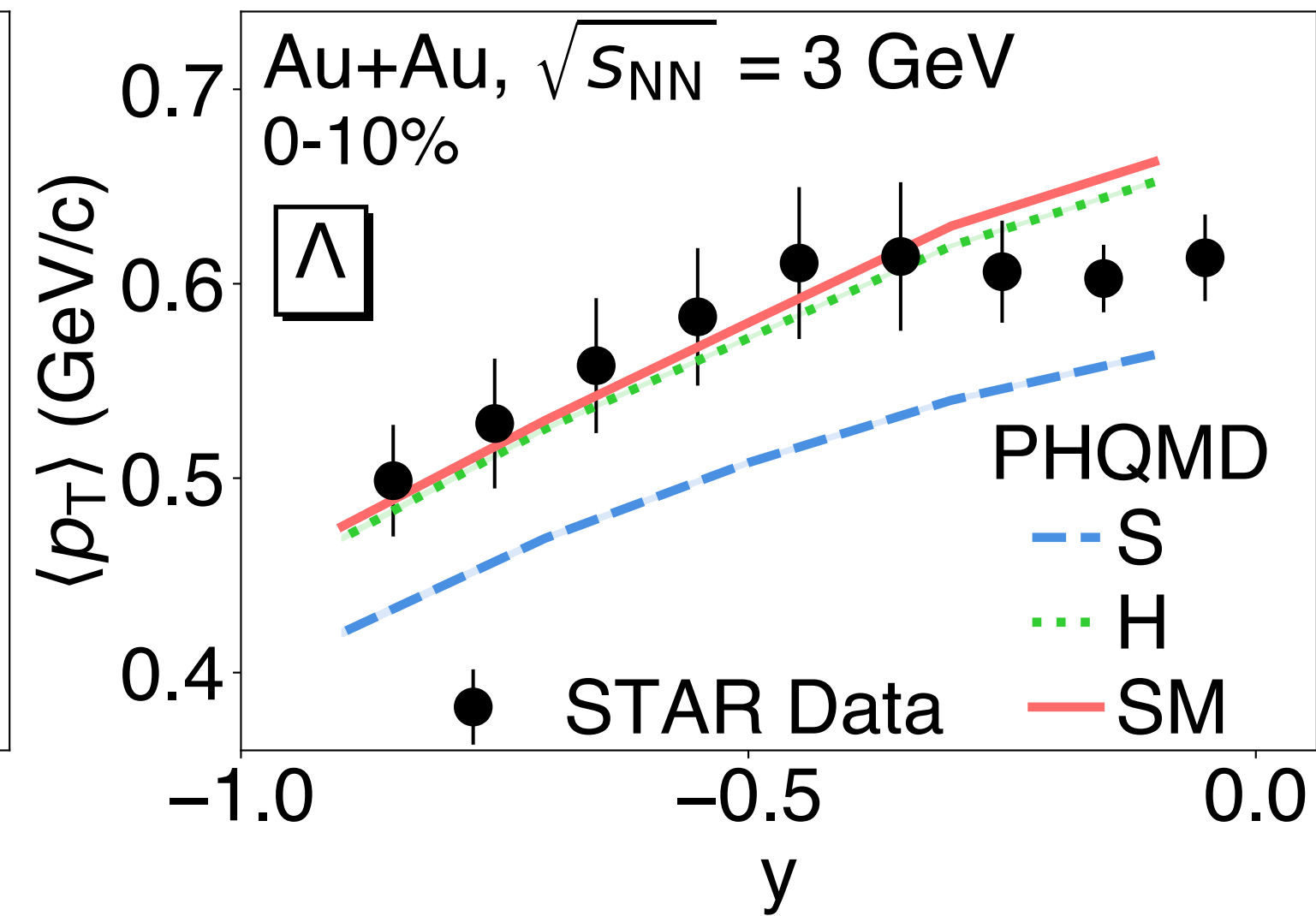
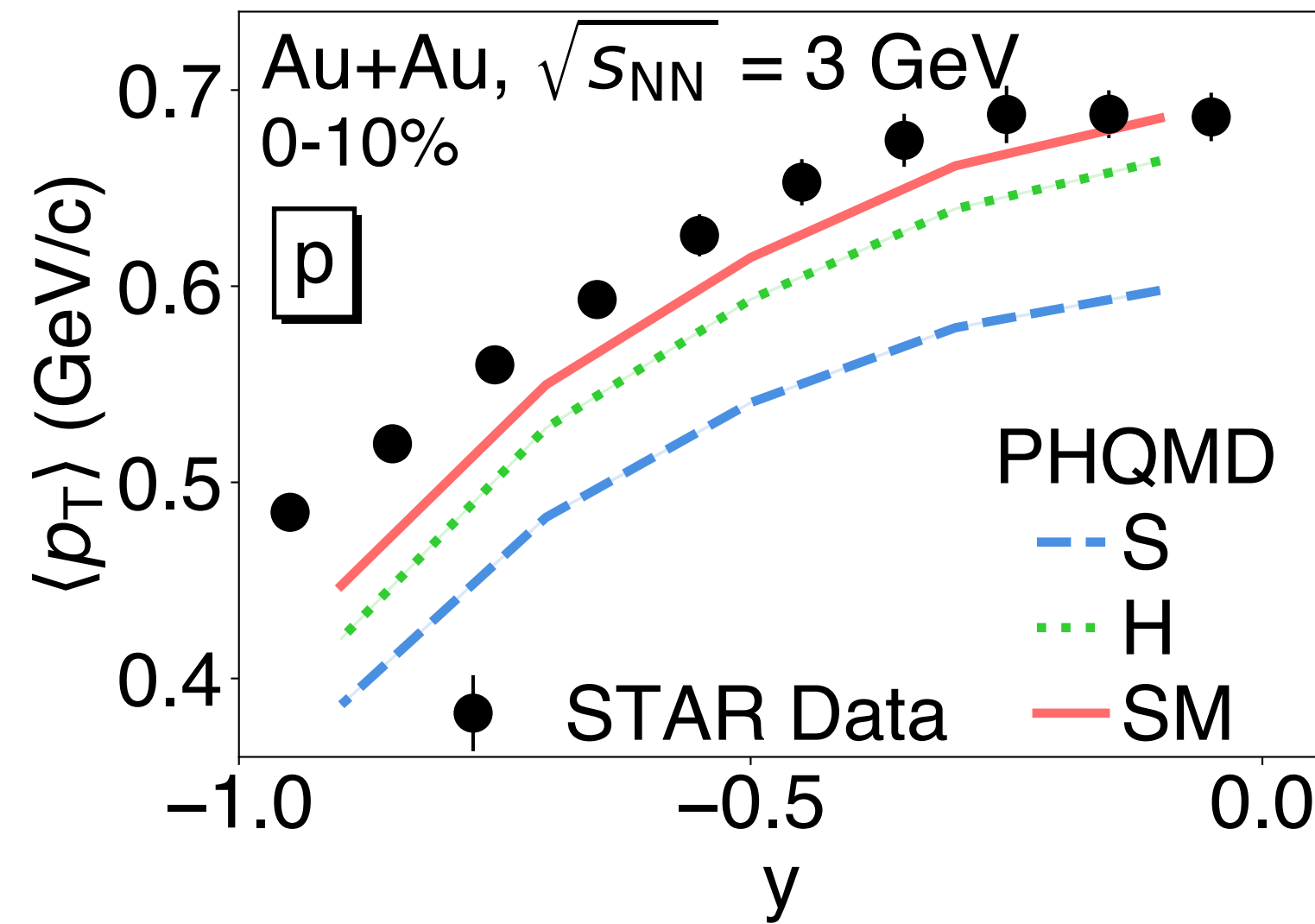




# Mean Transverse Momentum $\langle p_T \rangle$

$$\langle p_T \rangle = \frac{\int p_T \frac{dN}{dp_T} dp_T}{\int \frac{dN}{dp_T} dp_T}$$

- **S** EOS underestimate both  $p$  and  $\Lambda$  data
- **H** and **SM** EOS come much closer to the data
  - Deviations at  $y=0$  may warrant further study
- $\langle p_T \rangle$  of  $p > \langle p_T \rangle$  of  $\Lambda$ 
  - Fraction of protons may be spectator protons at such low energies



# Kinetic Freeze-Out Parameters

- Blast-wave fits to separate influence from the temperature and the collective expansion on the transverse momentum spectra

$$\frac{1}{2\pi p_T} \frac{d^2 N}{dp_T dy} \propto \int_0^R r dr m_T I_0\left(\frac{p_T \sinh \rho(r)}{T_{kin}}\right) K_1\left(\frac{m_T \cosh \rho(r)}{T_{kin}}\right)$$

$$\rho(r) = \tanh^{-1}(\beta_T)$$

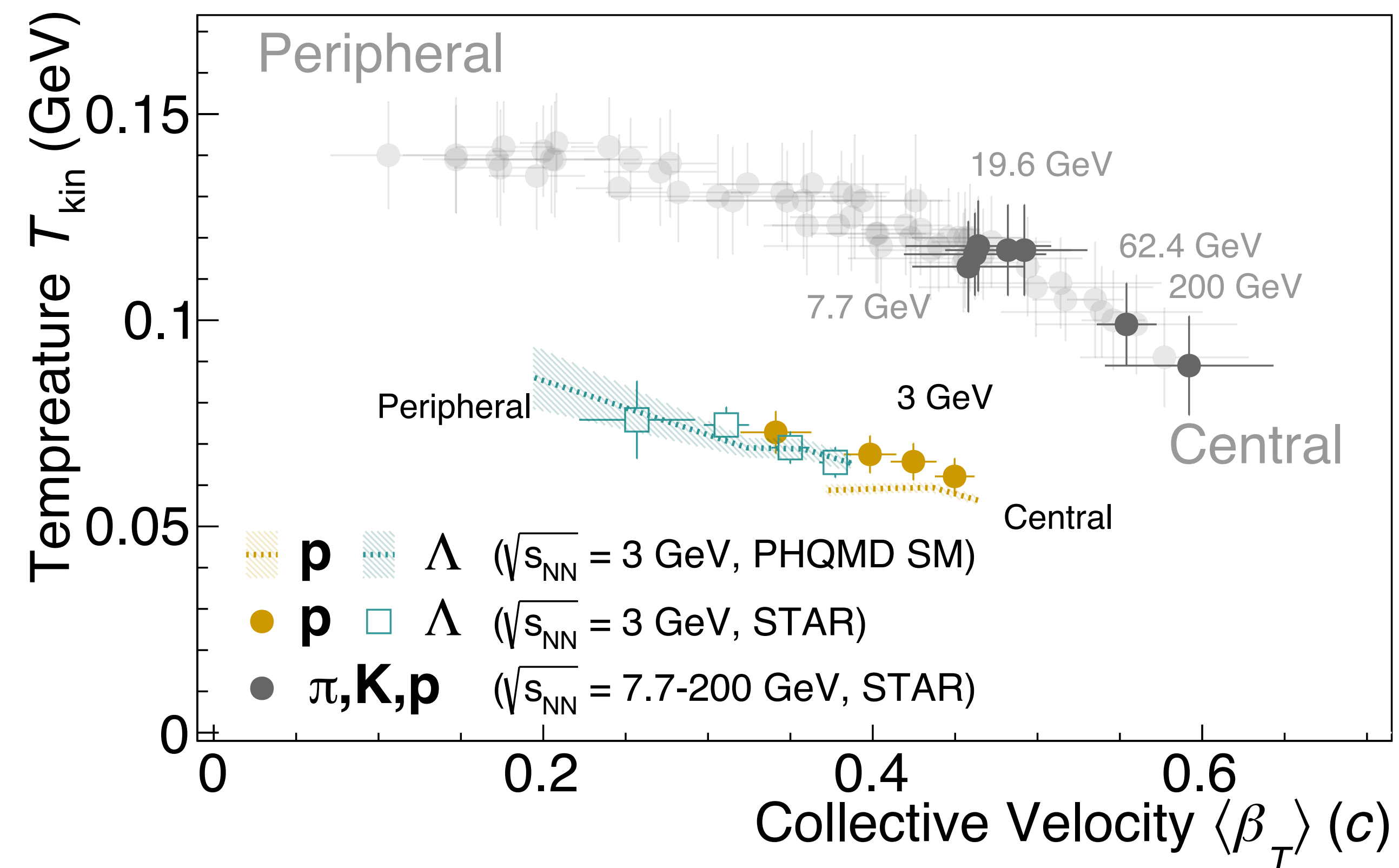
- Kinetic freeze-out temperature:  $T_{kin}$
- Average transverse velocity:  $\langle \beta_T \rangle$

- Similar  $T_{kin}$ , but larger  $\langle \beta_T \rangle$  for protons
- At fireball center, nucleon density is higher, collective velocity is smaller

**Production probability of  $\Lambda$ s higher close to the center of the fireball**

- PHQMD describes the trends observed in the data well

Au+Au Collisions at RHIC





# Azimuthal Anisotropy

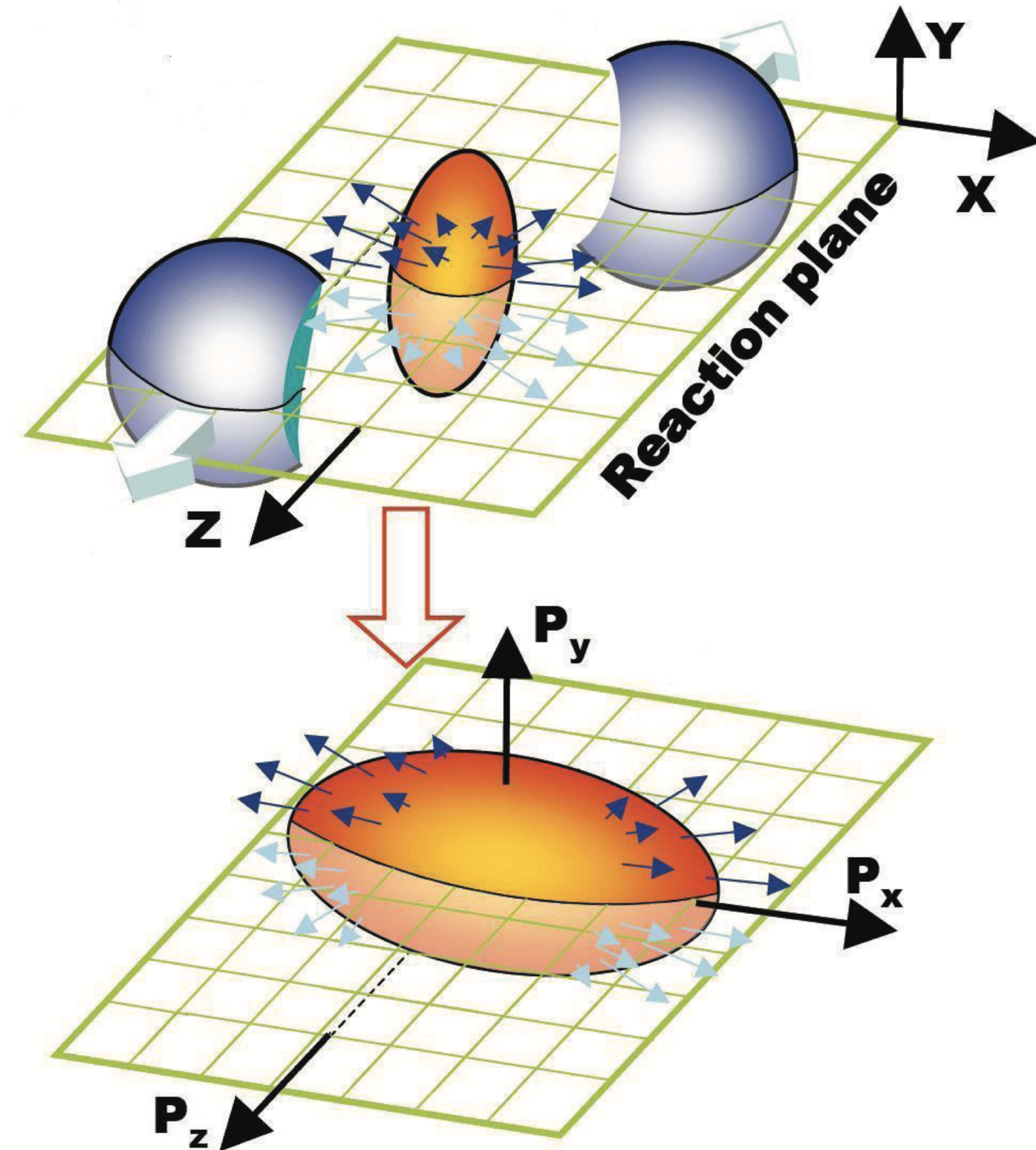
A. M. Poskanzer and S. A. Voloshin, PRC 58 (1998) 1671

- In non-central collisions, initial spatial anisotropy  $\rightarrow$  pressure gradients  $\rightarrow$  momentum space anisotropy

$$E \frac{d^3N}{dp^3} = \frac{1}{2\pi} \frac{d^2N}{p_T dp_T dy} \left( 1 + \sum_1^{\infty} 2v_n \cos[n(\phi - \Psi_{rp})] \right)$$

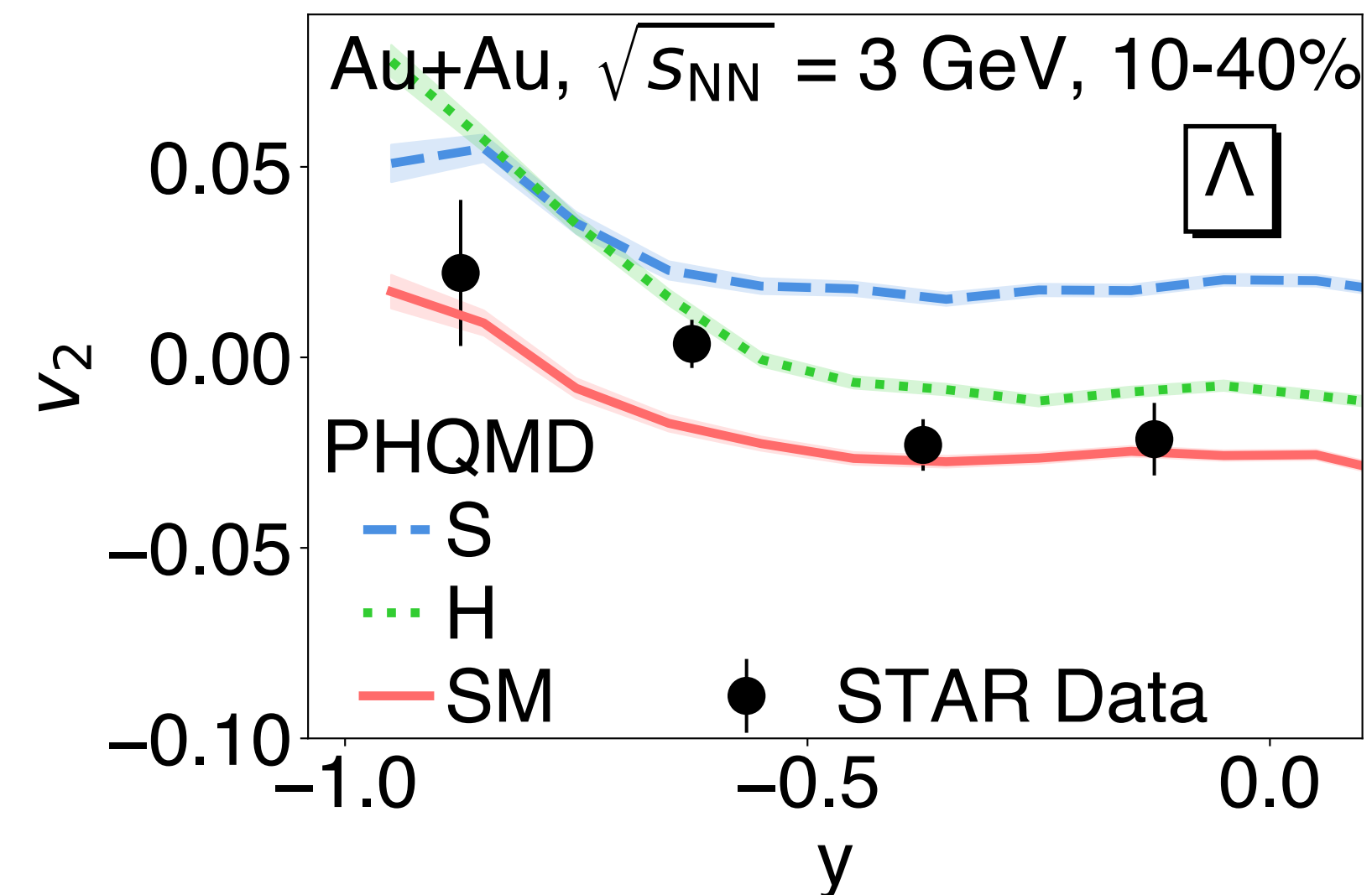
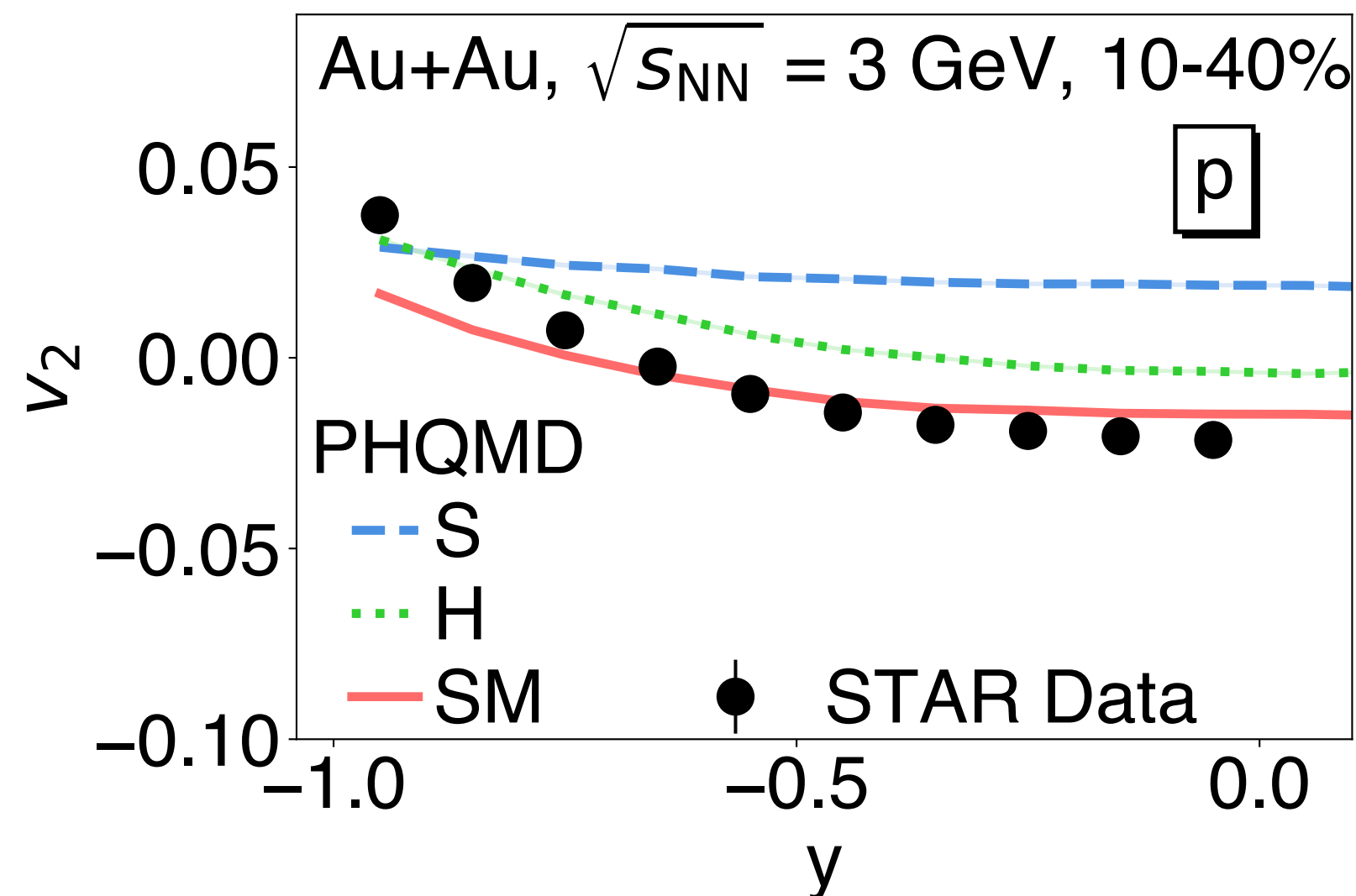
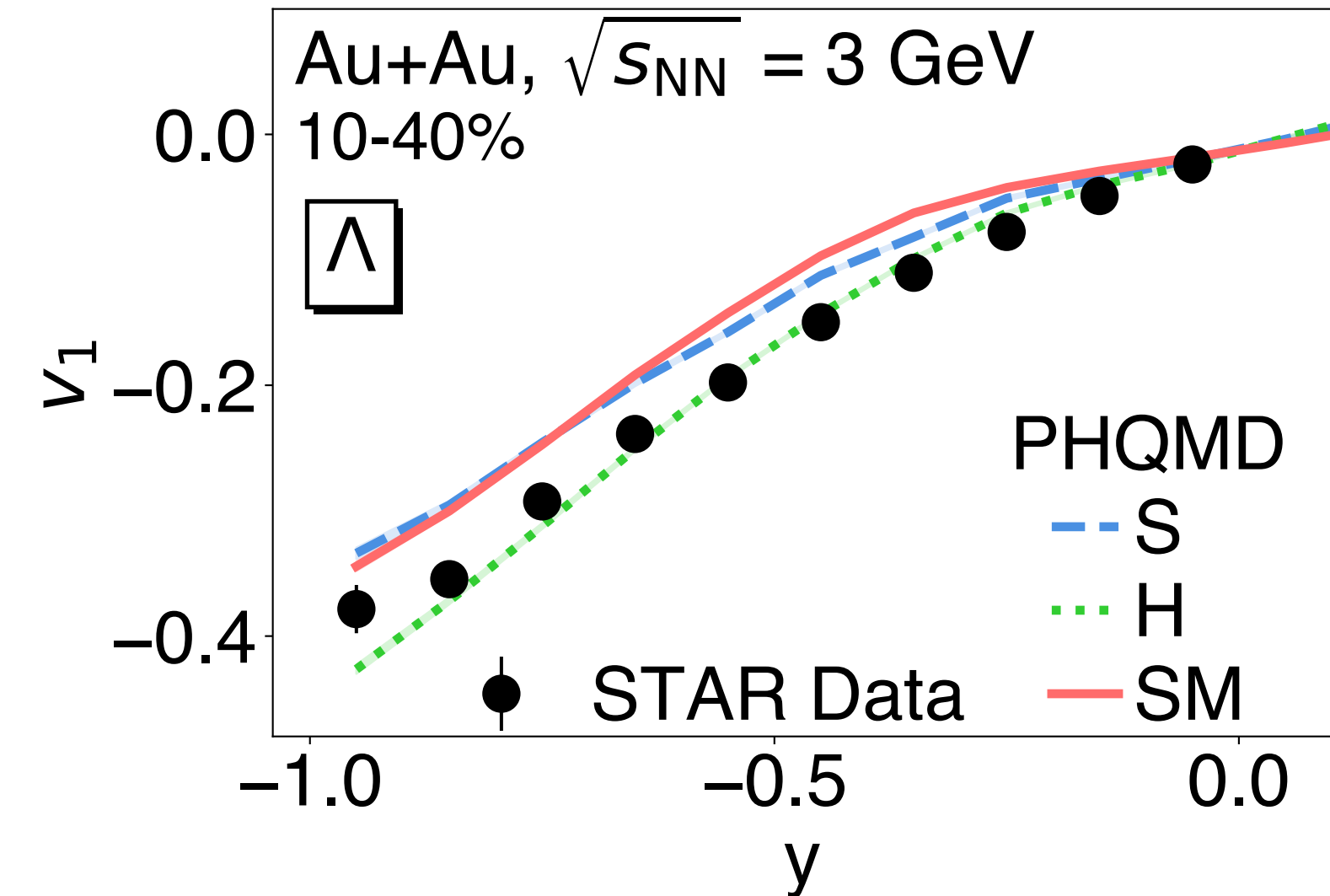
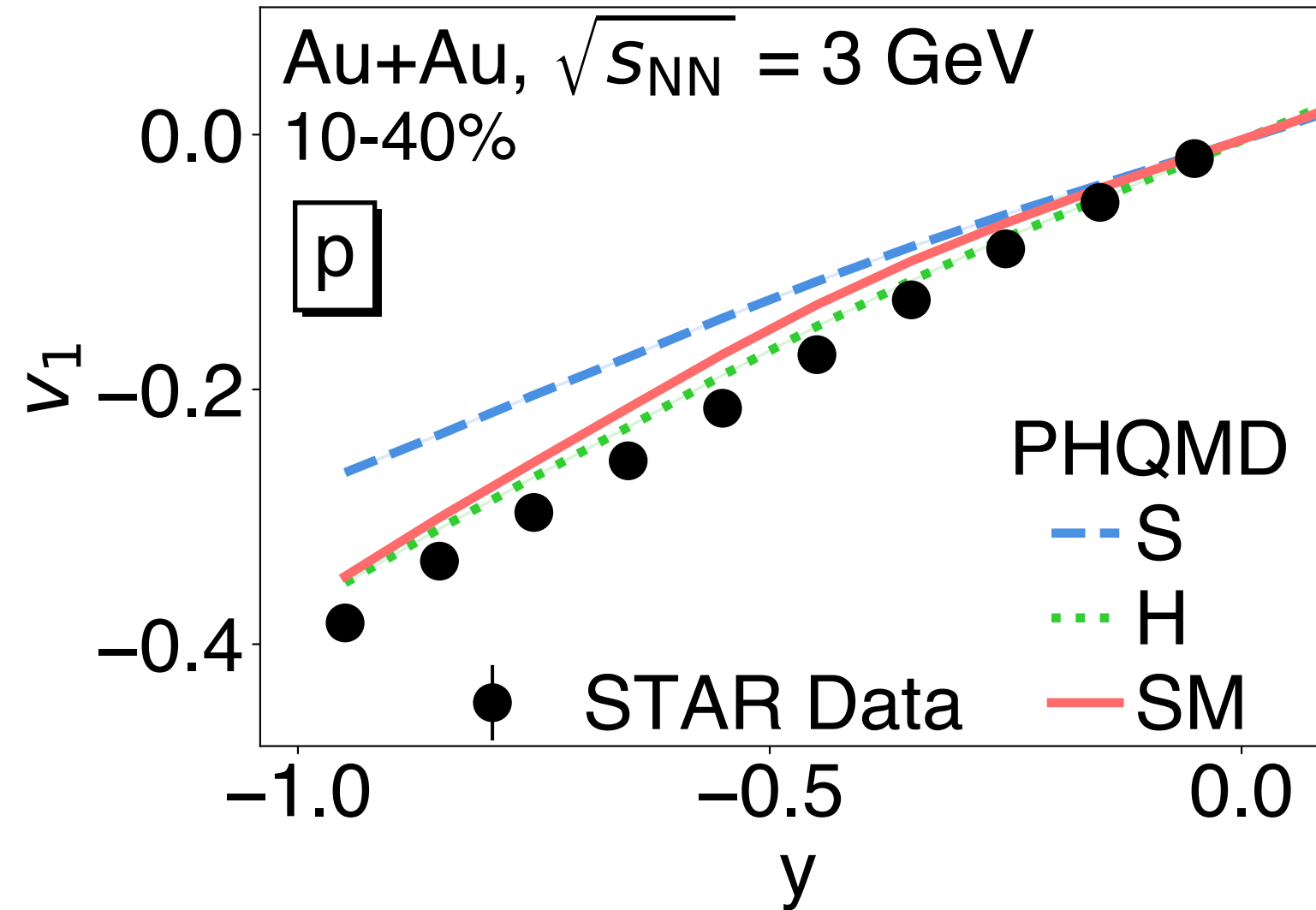
- Directed flow:  $v_1 = \langle \cos(\phi - \Psi_{rp}) \rangle$
- Elliptic flow:  $v_2 = \langle \cos[2(\phi - \Psi_{rp})] \rangle$

**Sensitive probes of the EoS of nuclear matter at high baryon density**



# Directed Flow and Elliptic Flow

- Directed flow  $v_1$ :
  - Similar to  $\langle p_T \rangle$ , **S** EOS do not provide good description, **H** EOS describe data well
  - SM** EOS underestimate  $\Lambda$ 
    - Choice of  $\Lambda$  potential?
- Elliptic flow  $v_2$ :
  - SM** EOS describe data well, **S** predicts wrong sign at mid-rapidity



Most of the 3 GeV Au+Au data best described by **SM** EOS



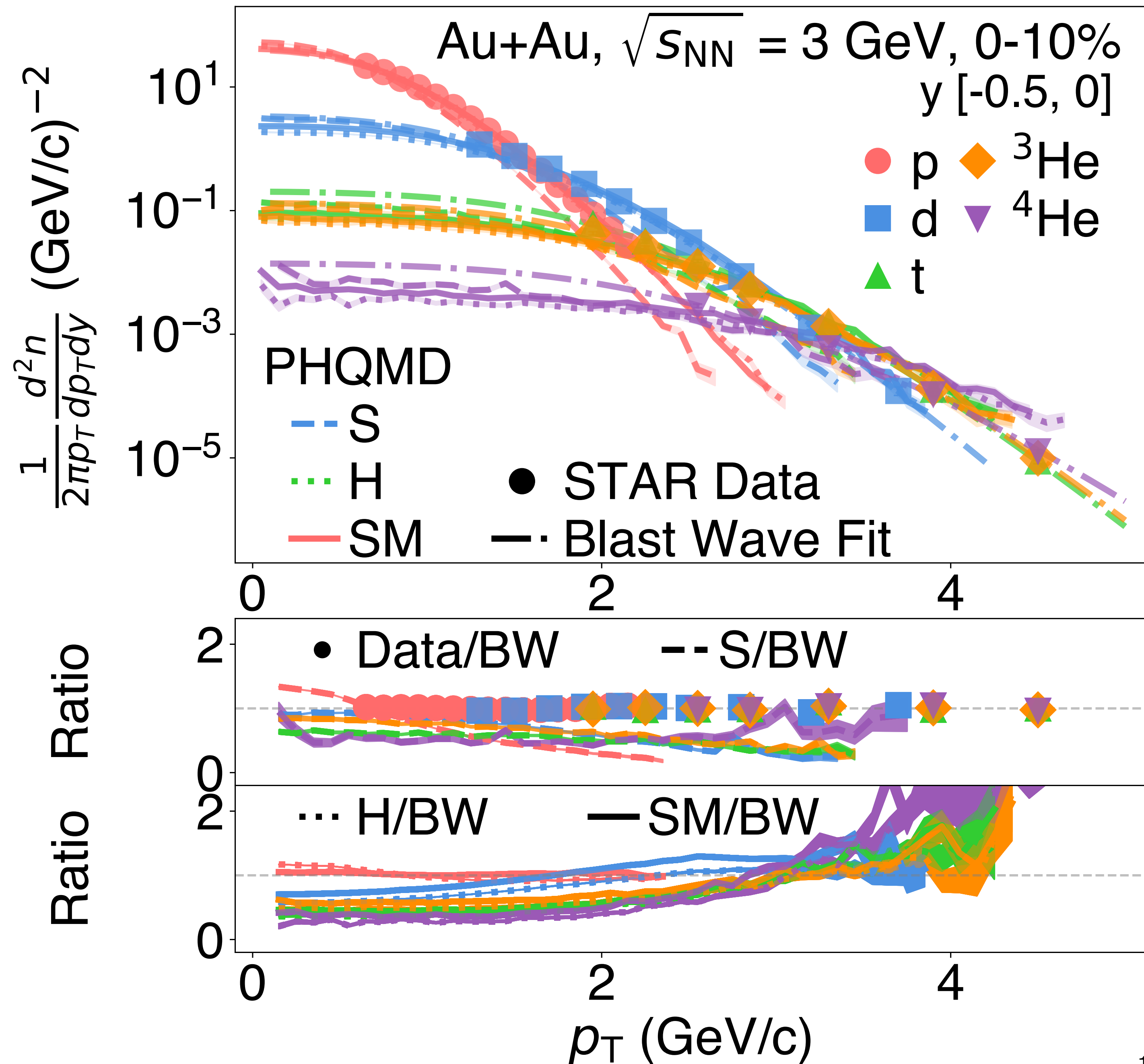
# Nuclei and Hypernuclei

## STAR data

- $d, t, {}^3\text{He}, {}^4\text{He}$   $p_T$  spectra: Phys. Rev. C 110 (2024) 54911
- ${}^3_\Lambda\text{H}, {}^4_\Lambda\text{H}$   $p_T$  spectra: Phys. Rev. Lett. 128 (2022) 202301
- $d, t, {}^3\text{He}, {}^4\text{He}$   $v_1, v_2$ : Phys. Lett. B 827 (2022) 136941
- ${}^3_\Lambda\text{H}, {}^4_\Lambda\text{H}$   $v_1$  : Phys. Rev. Lett. 130 (2023) 212301

# Nuclei $p_T$ spectra

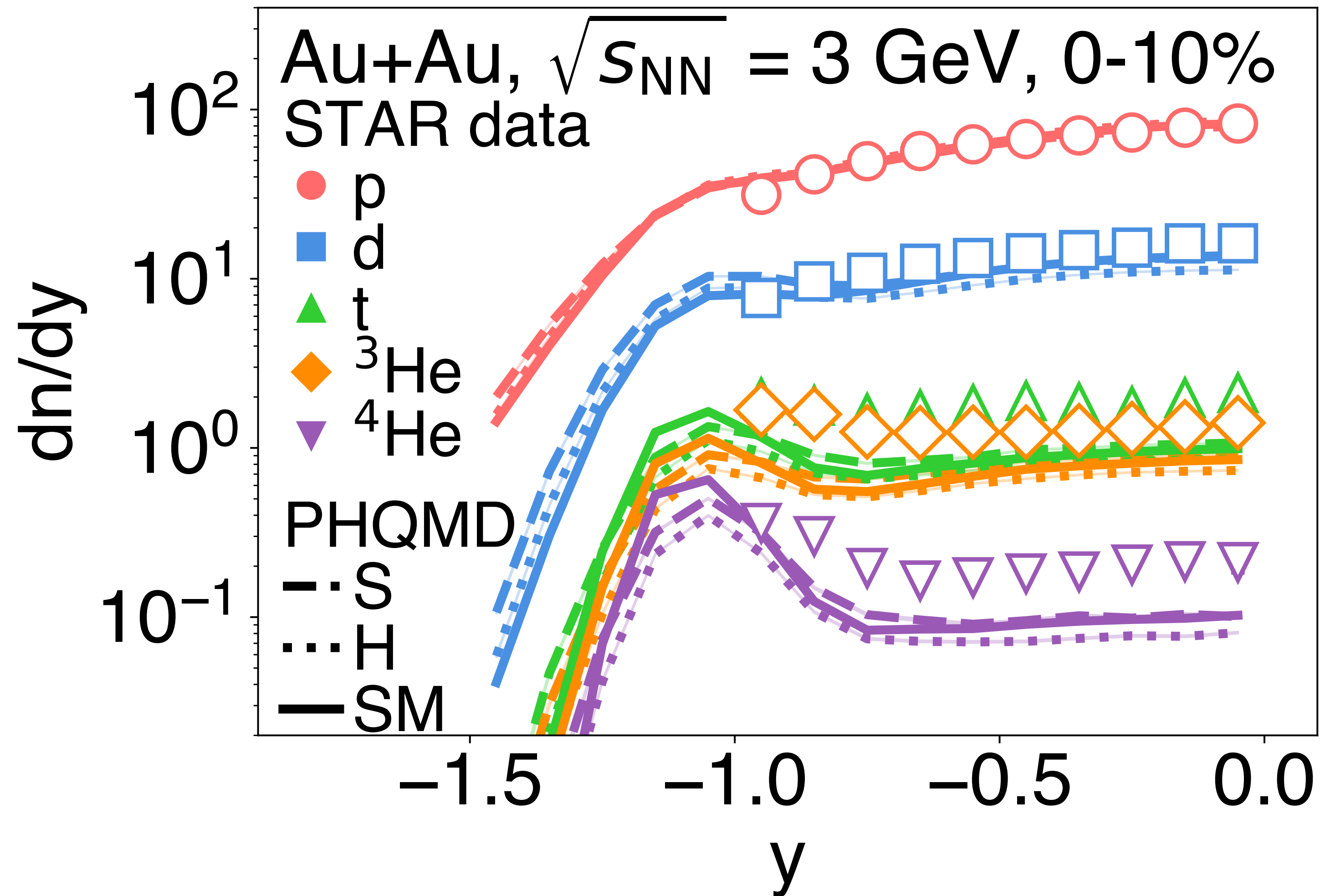
- **S** EOS is systematically lower than data for all clusters
- For  $d$ , **H** and **SM** EOS are mostly consistent with data, however there is a **tension in the slope**
- For  $t$ ,  ${}^3\text{He}$ , and  ${}^4\text{He}$ , the deviations are larger, with PHQMD underestimating the data





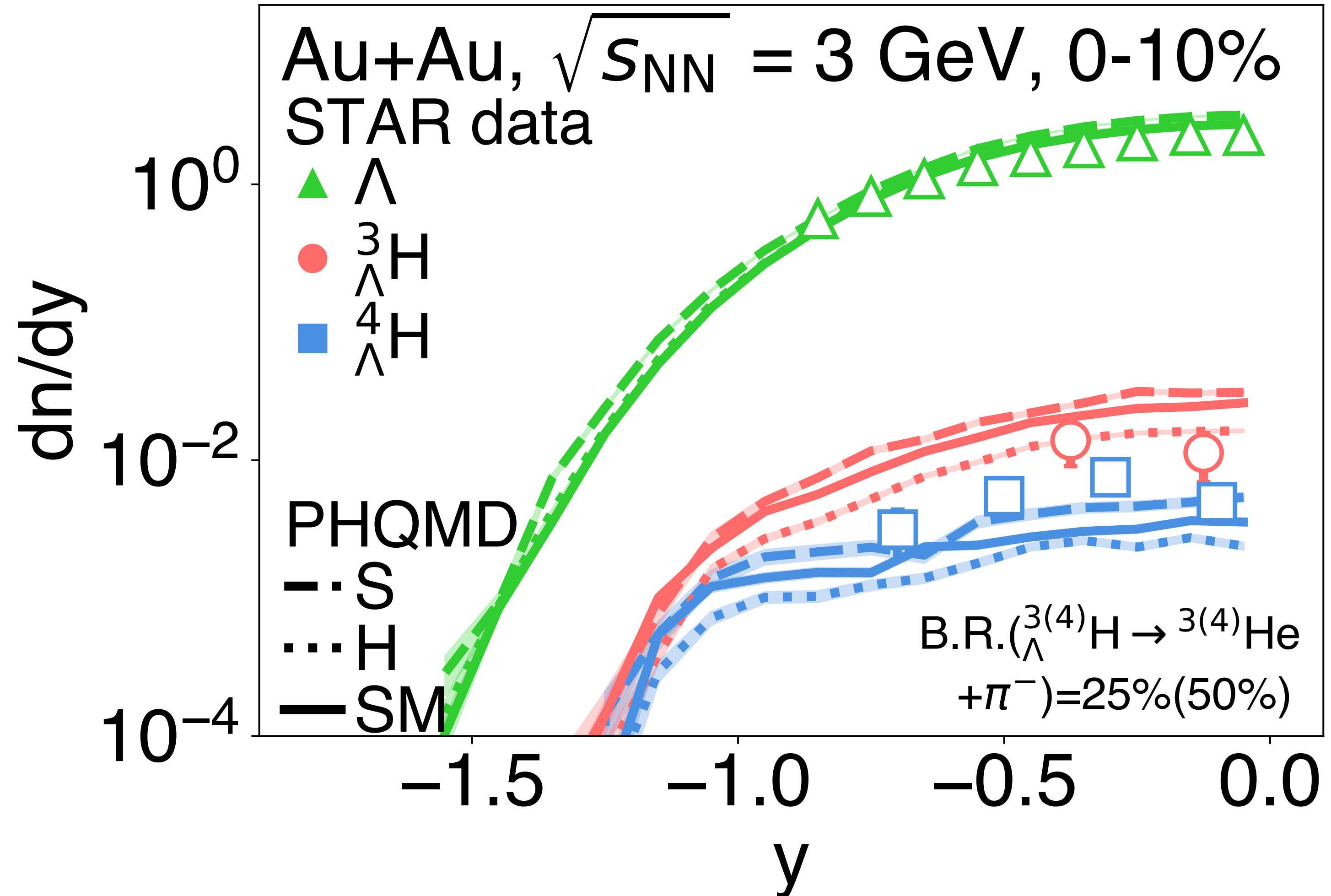
# Nuclei dN/dy

- $p$  yields well reproduced
- H EoS predicts a lower  $d$  yield compared to S and SM EoS
  - $d$  yields reproduced within 20% with S and SM
- **Larger deviations for heavier clusters**



# Hypernuclei dN/dy

- $\Lambda$  yields yield has a mild sensitivity to the EoS
- ${}^3_{\Lambda}\text{H}$ ,  ${}^4_{\Lambda}\text{H}$ : S EoS predicts the largest yield, followed by SM and H EoS
  - Approx. factor of 2 difference b/w S and H EoS

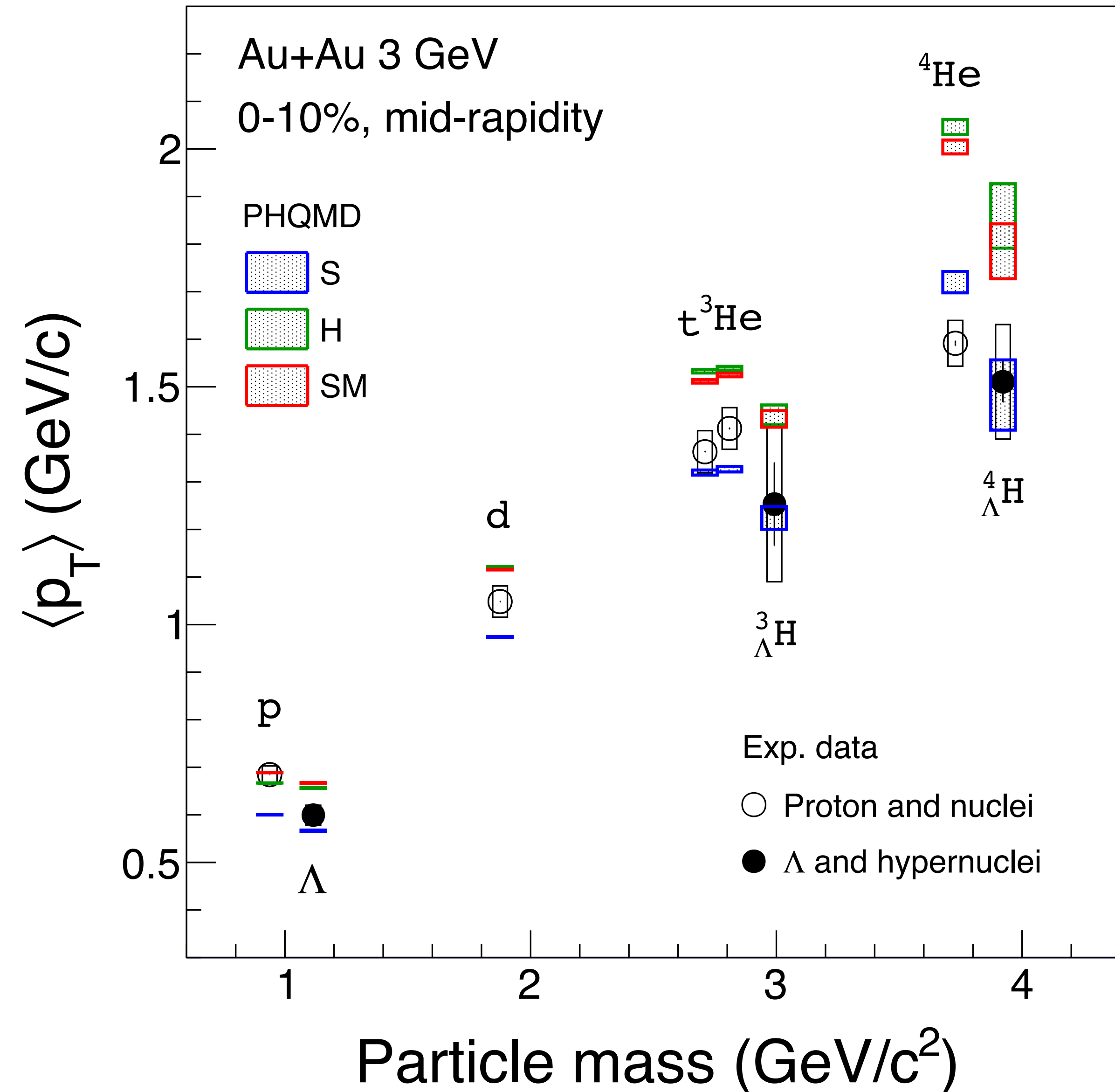


**Higher sensitivity of the hypernuclei yields to EoS compared to other particles**



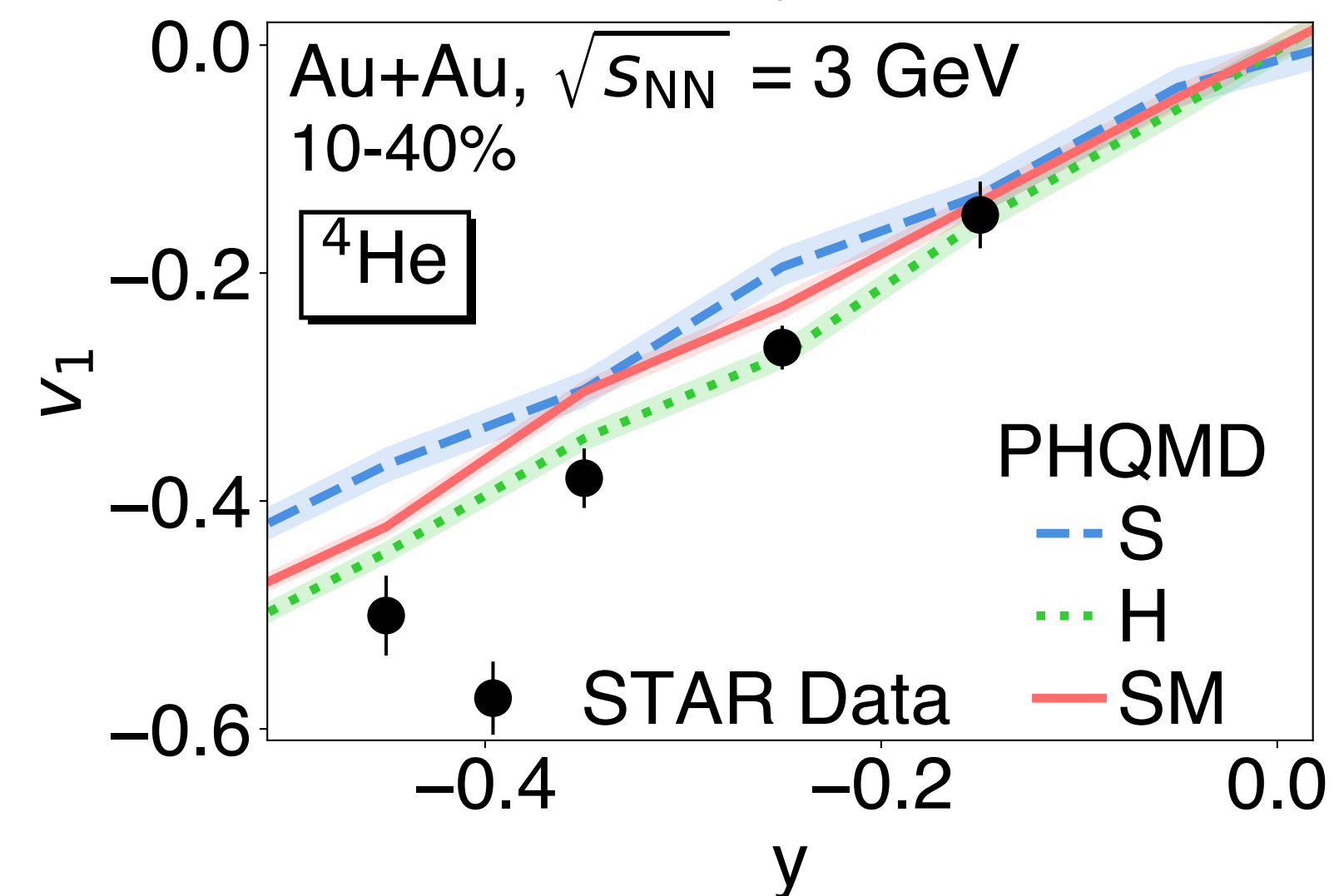
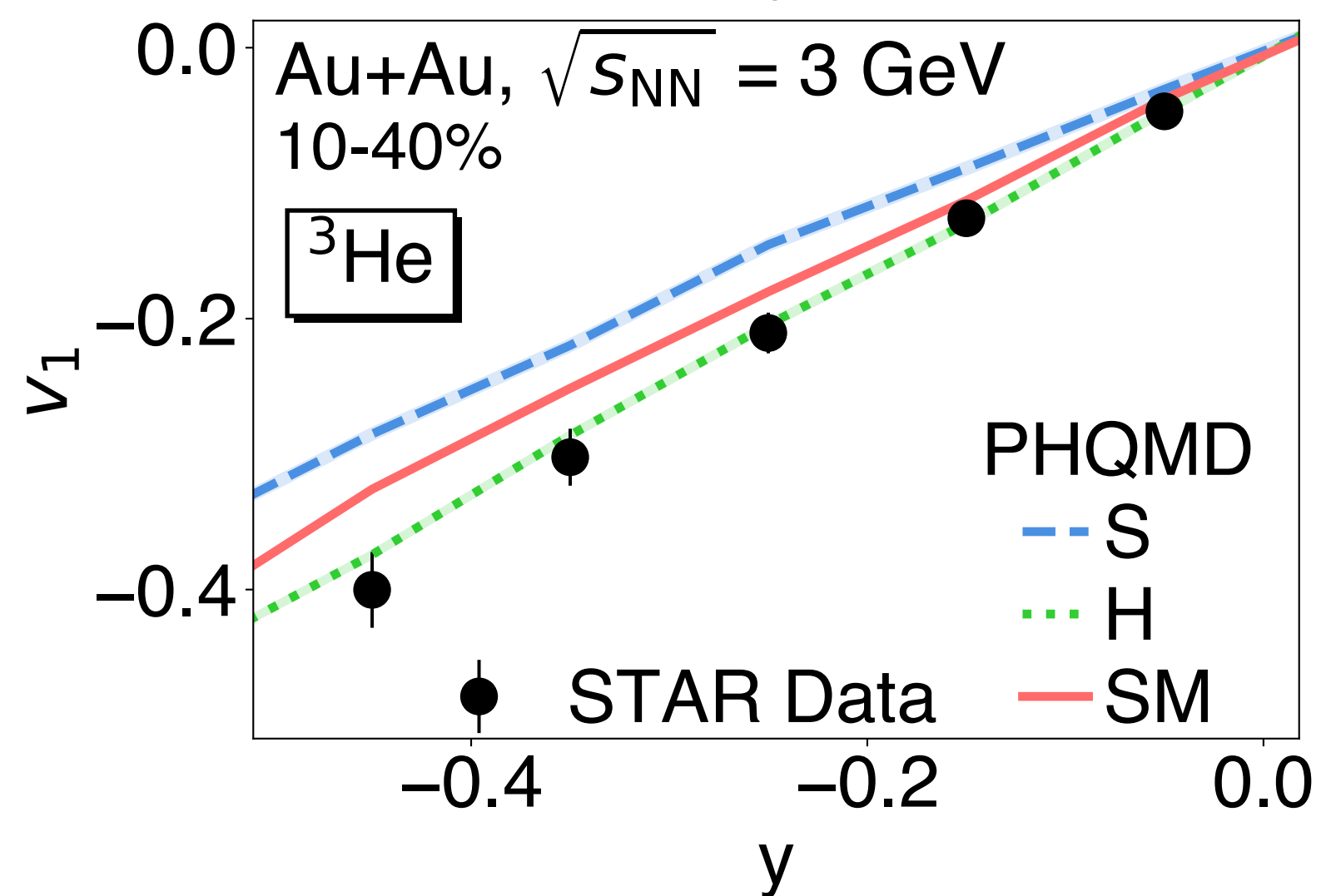
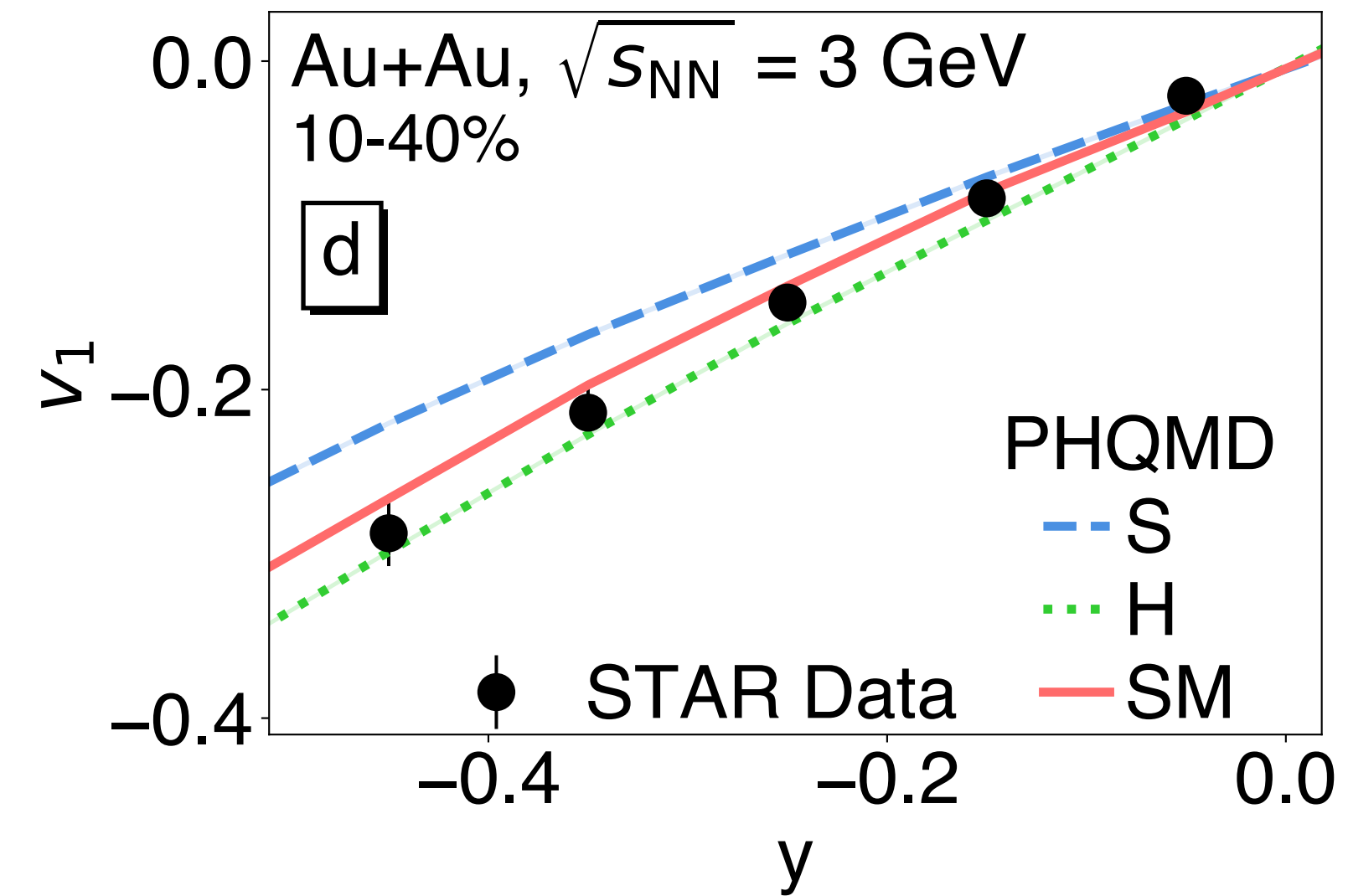
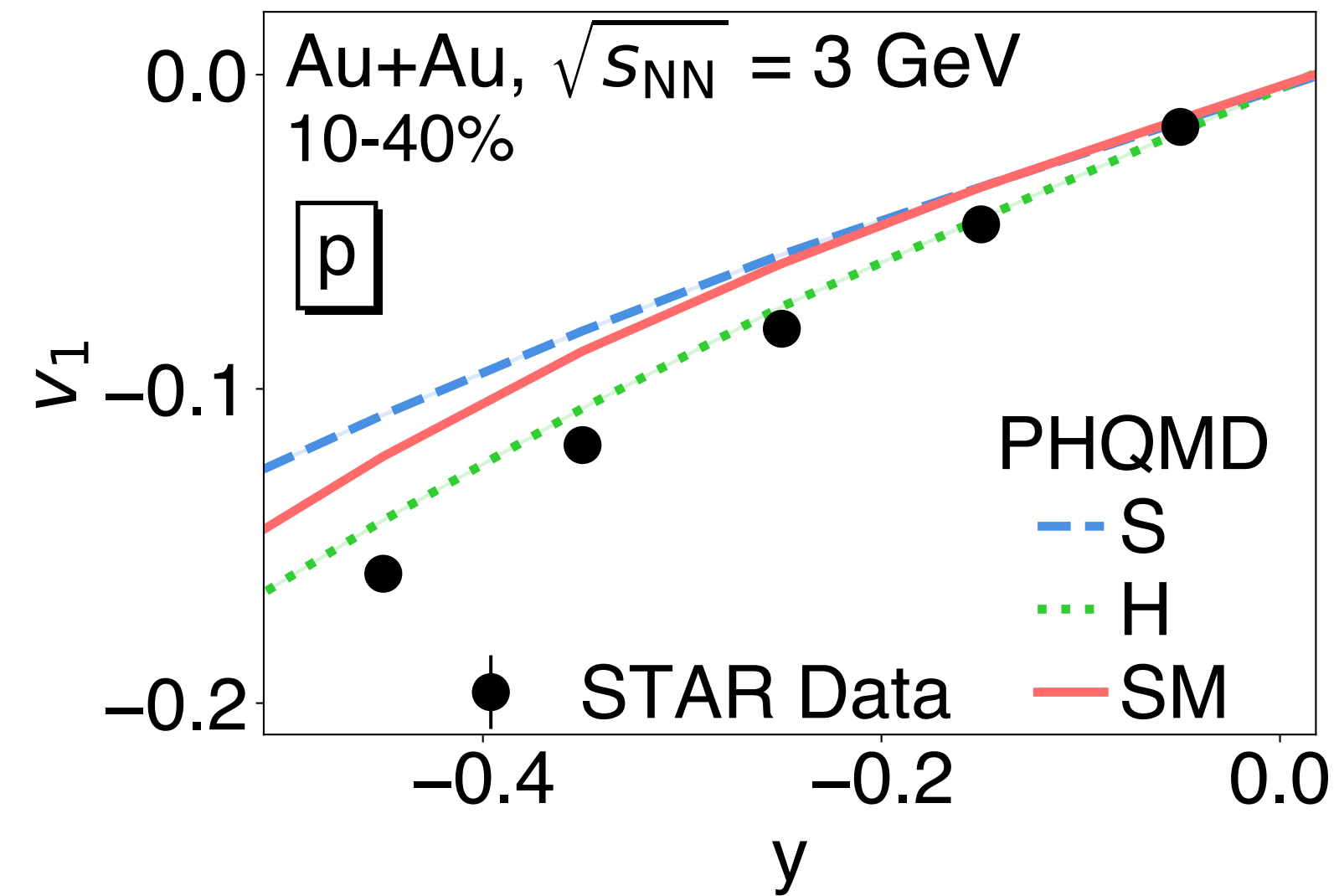
# Nuclei and Hypernuclei $\langle p_T \rangle$

- Although **H** and **SM** EOS describe the  $\langle p_T \rangle$  of protons, they **overestimate** the  $\langle p_T \rangle$  of all the nuclei
  - Underprediction of the yield of low  $p_T$  clusters and disagreement of the extrapolation of the experimental data to  $p_T \rightarrow 0$  w.r.t. PHQMD
- H** and **SM** EOS also tend to overestimate the  $\langle p_T \rangle$  of hypernuclei



# Rapidity Dependence of Nuclei Directed Flow

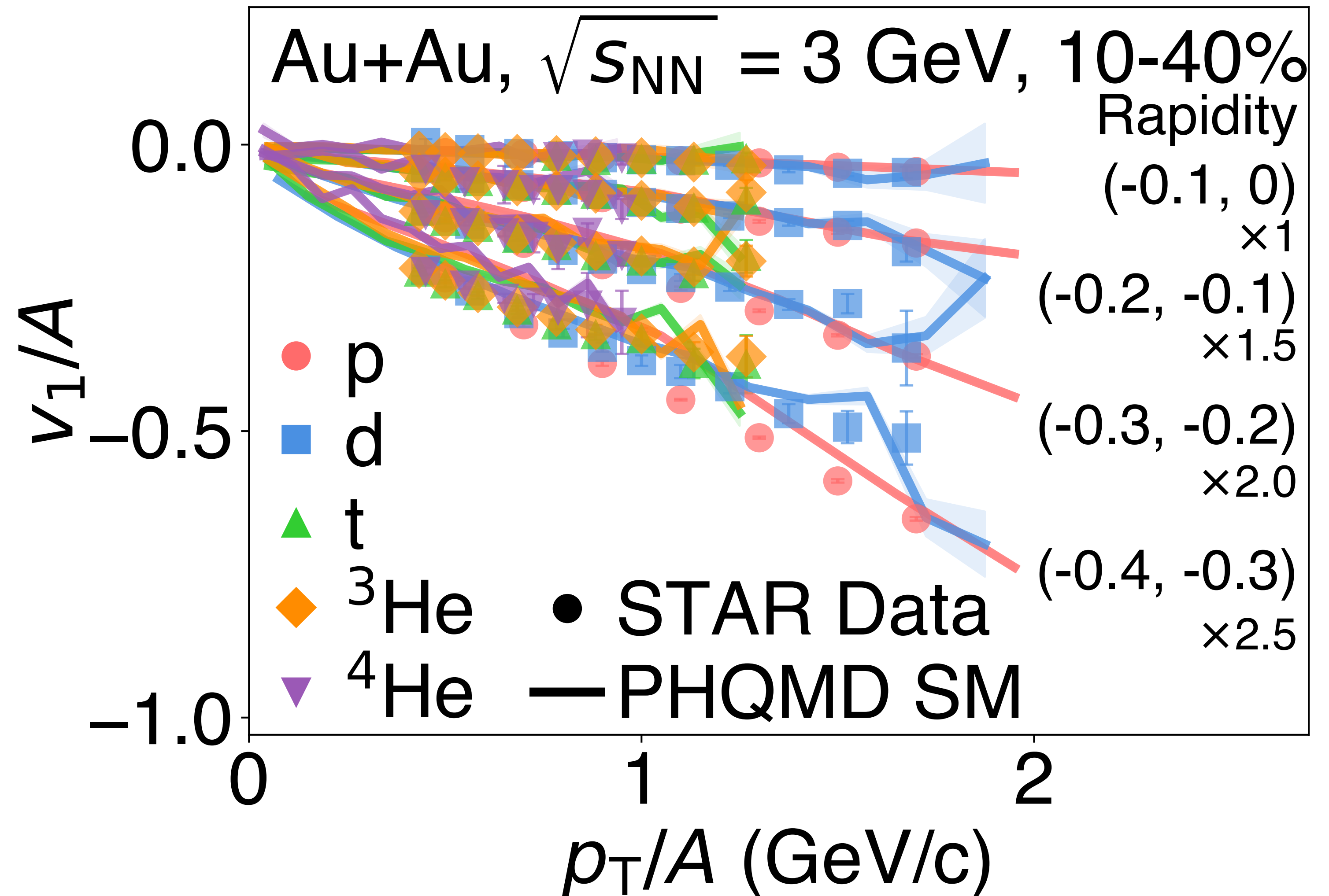
- **S** EOS underestimates proton and all clusters
- **H** EOS provides a better description compared to **SM**





# $p_T$ Dependence of Nuclei Directed Flow ( $v_1$ )

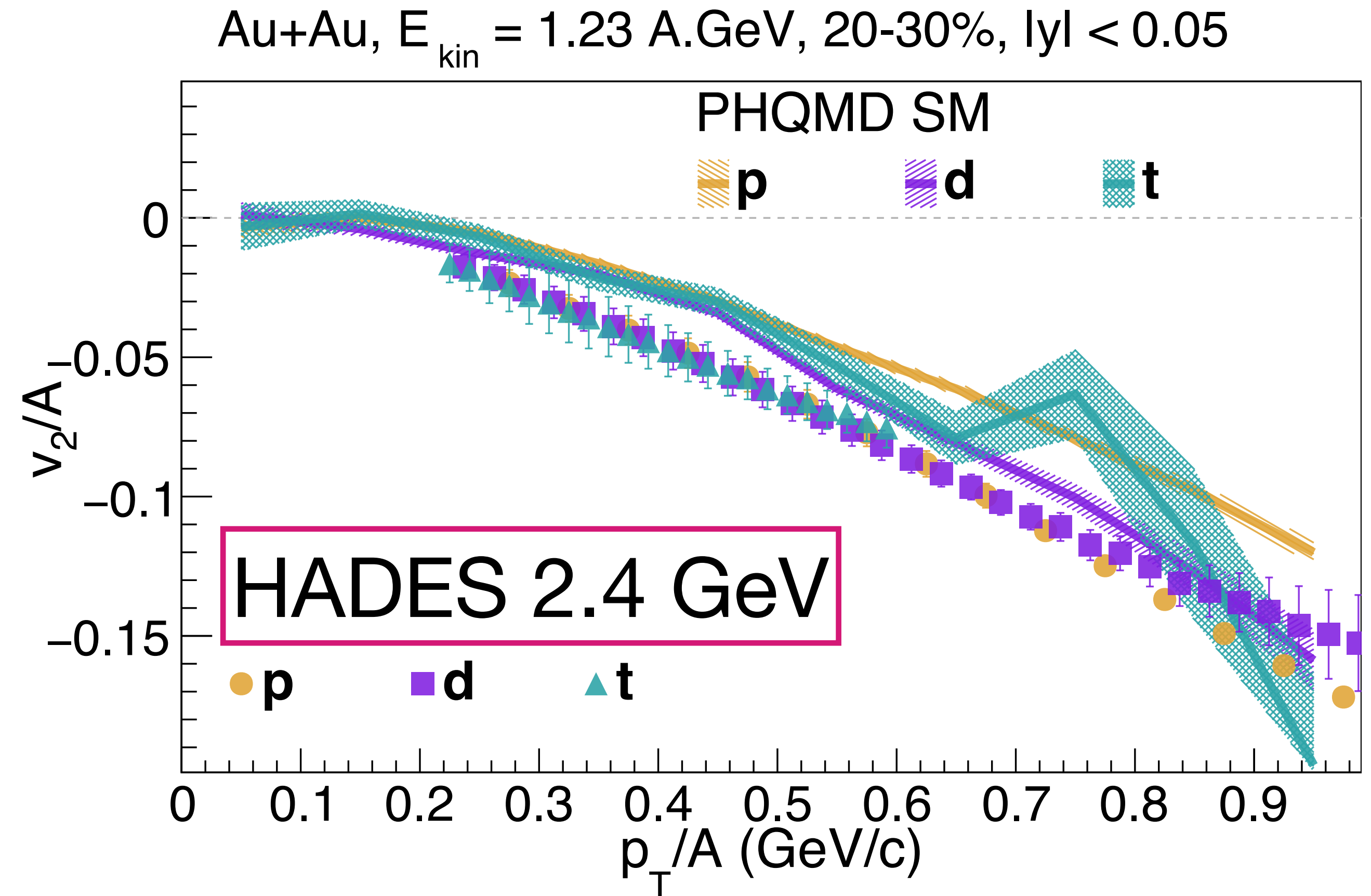
- $v_1$  vs  $p_T$  scales with mass number for  $y = (-0.3, 0)$ , reproduced by PHQMD
- Deviations from mass number scaling for  $y = (-0.4, -0.3)$
- Measurements at backward rapidity are important



# Elliptic Flow ( $v_2$ ) of Nuclei

V. Kireyeu et al., arXiv:2411.04969

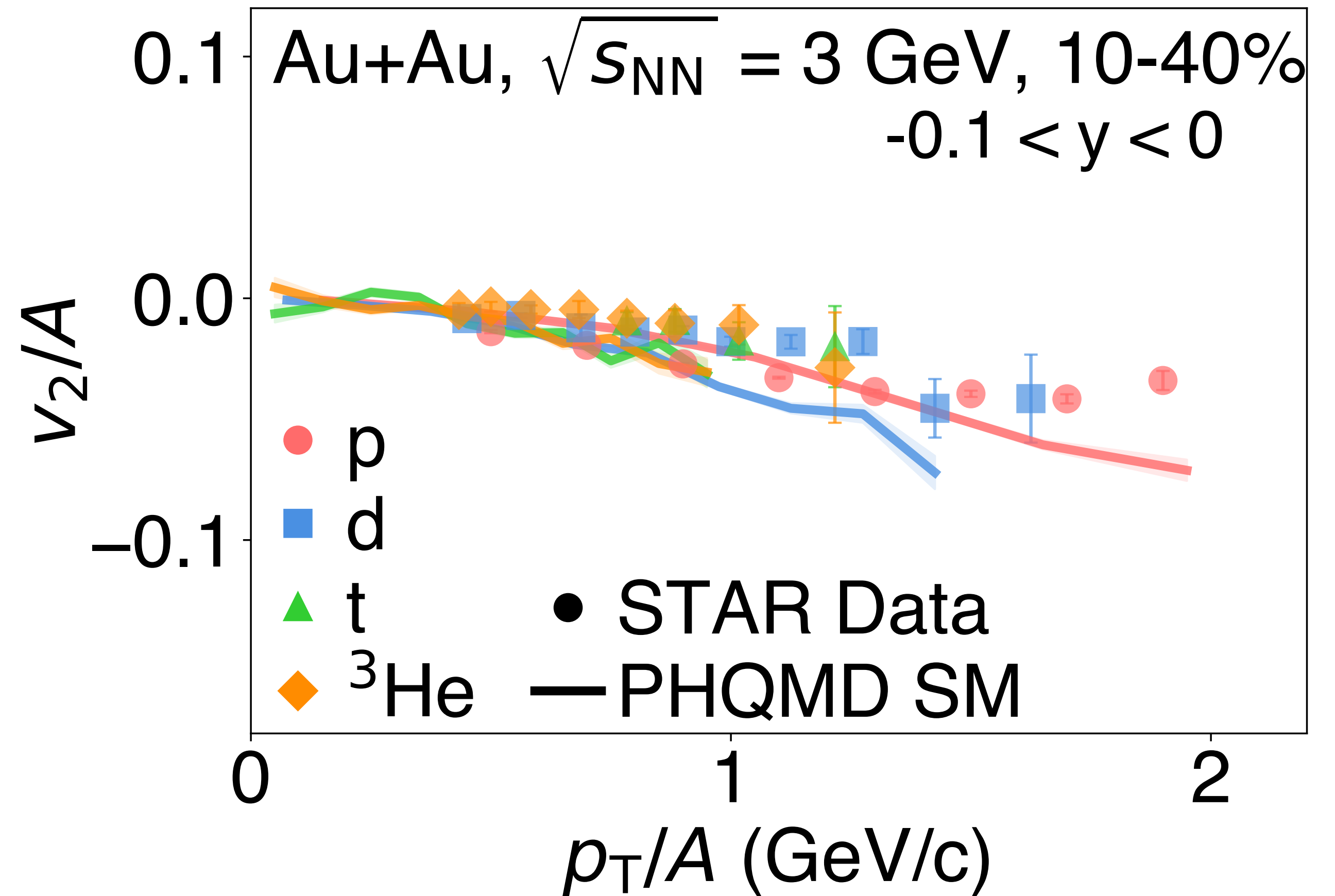
- HADES 2.4 GeV data show mass number scaling for elliptic flow  $v_2$  up to  $p_T/A = 1$  GeV





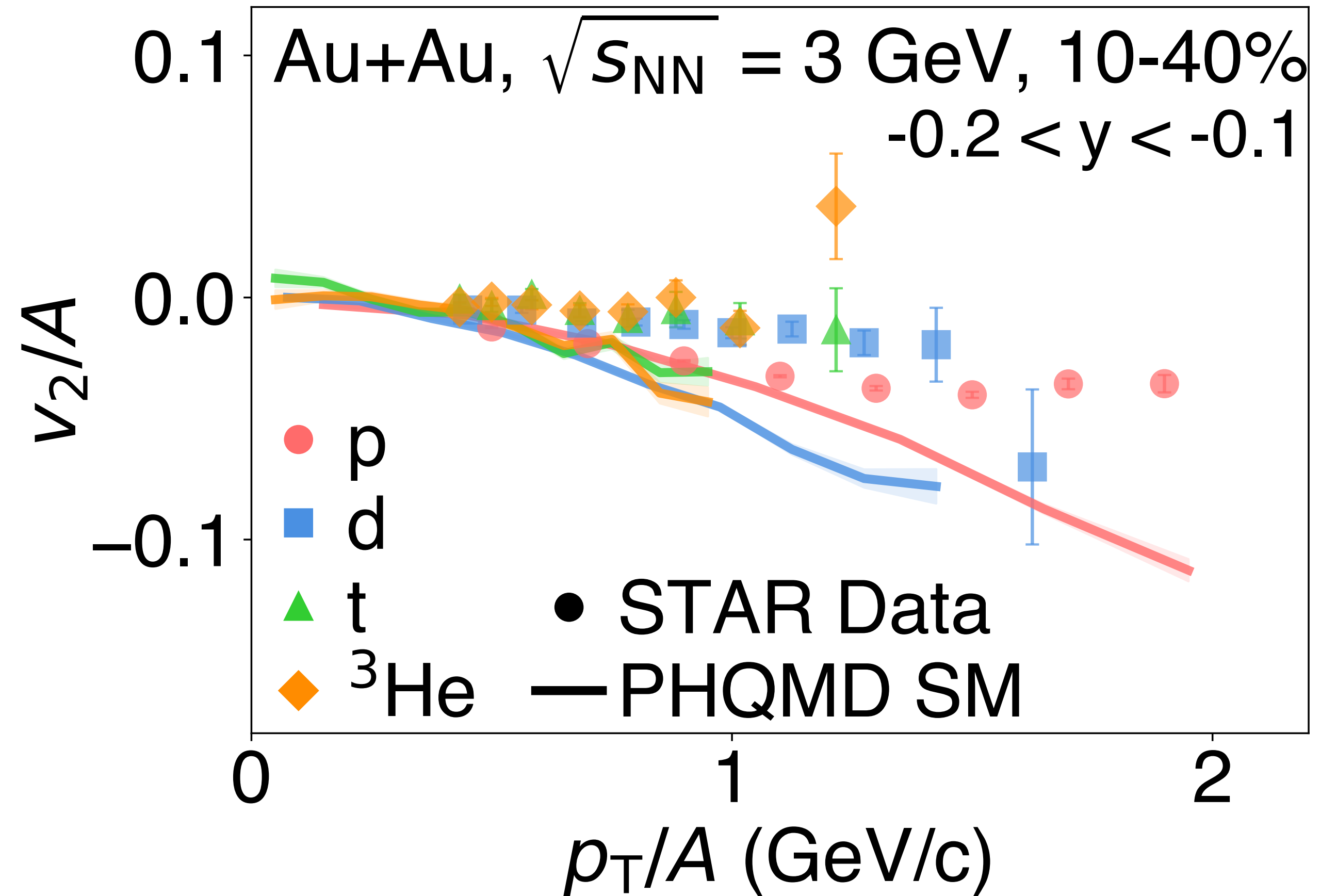
# Elliptic Flow ( $v_2$ ) of Nuclei

- HADES 2.4 GeV GeV data show mass number scaling for elliptic flow  $v_2$  up to  $p_T/A = 1$  GeV
- Elliptic flow for the data at 3 GeV **do not show mass number scaling**



# Elliptic Flow ( $v_2$ ) of Nuclei

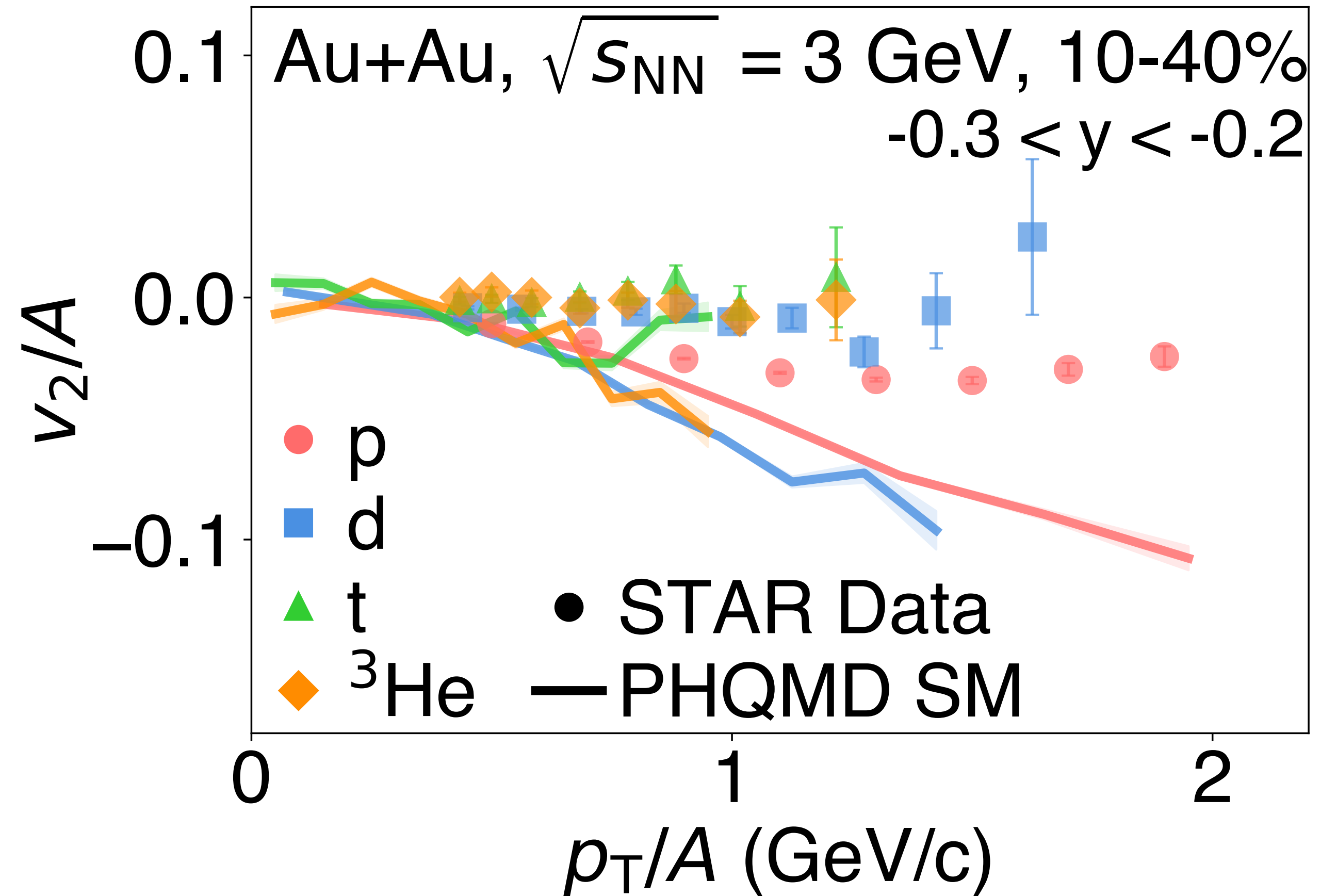
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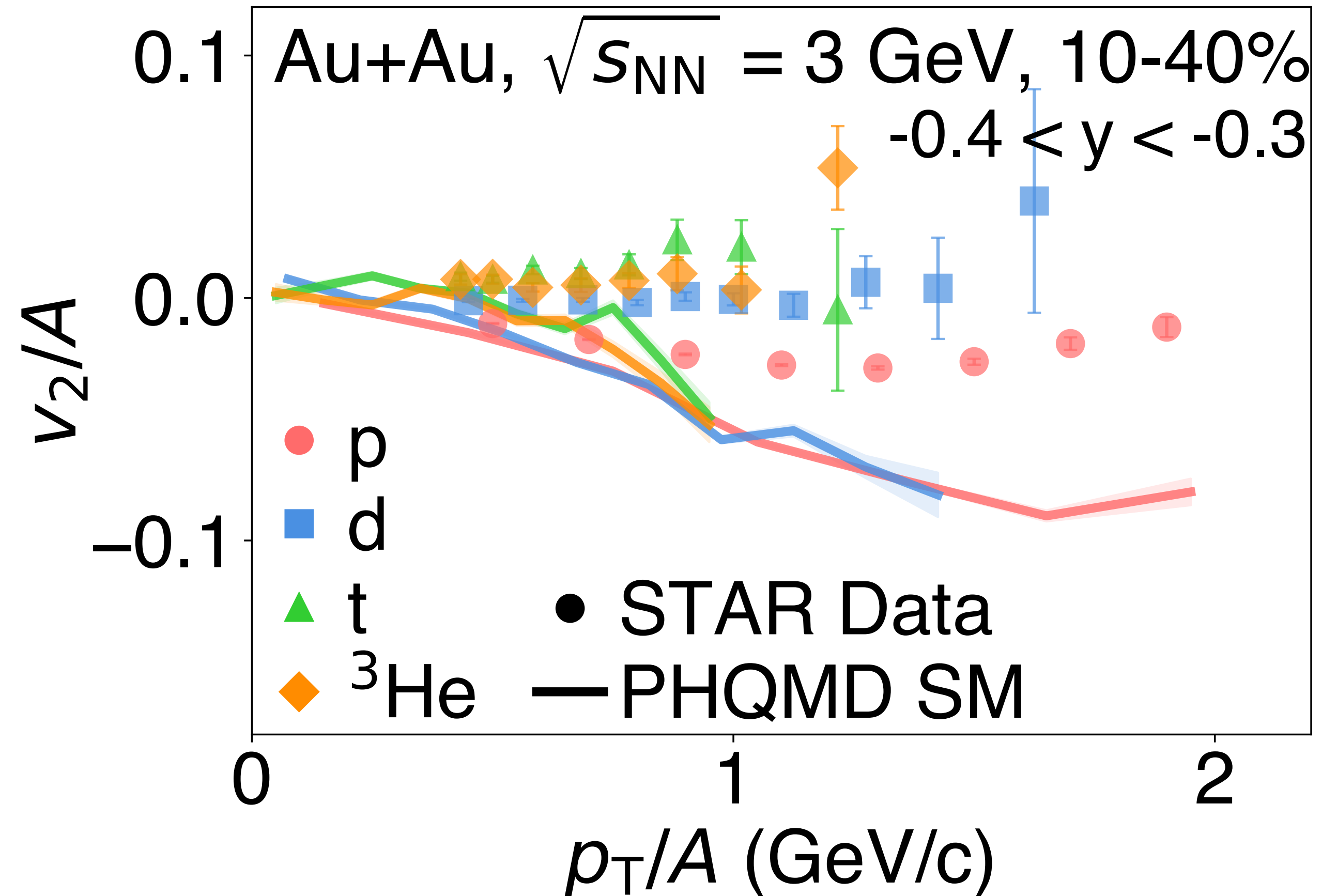
# Elliptic Flow ( $v_2$ ) of Nuclei

- HADES 2.4 GeV GeV data show mass number scaling for elliptic flow  $v_2$  up to  $p_T/A = 1$  GeV
- Elliptic flow for the data at 3 GeV **do not show mass number scaling**



# Elliptic Flow ( $v_2$ ) of Nuclei

- HADES 2.4 GeV GeV data show mass number scaling for elliptic flow  $v_2$  up to  $p_T/A = 1$  GeV
- Elliptic flow for the data at 3 GeV **do not show mass number scaling**
- Deviations are more obvious away from mid-rapidity



**Suggests that at 3 GeV clusters are not a random selection of nucleons but a selective process**



# Summary

- Compared the results of PHQMD calculations using **3 different EoS, soft (S), hard (H), and soft with momentum dependence (SM)** with STAR 3 GeV Au+Au data

## Probes of EOS

- **Different EoS gives quite different results for  $\langle p_T \rangle$ ,  $v_1$  and  $v_2$  of baryons and clusters**
- **Hypernuclei yields show high sensitivity to EoS**

## Proton and $\Lambda$

- **Proton and  $\Lambda$  data mostly described by SM EoS**
- $\Lambda$  has similar  $T_{kin}$  with proton while having smaller  $\beta$ , suggests higher production probability of  $\Lambda$ s close to center of fireball

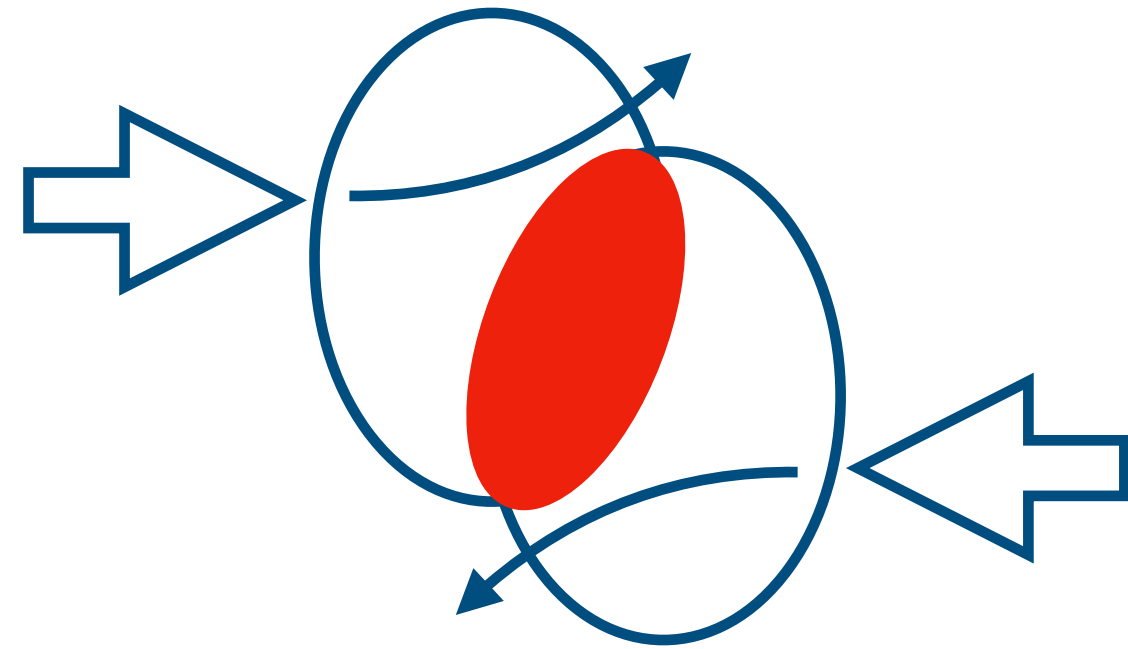
## Clusters and hypernuclei

- Stronger discrepancy b/w PHQMD results and data for heavier clusters
- **$v_2$  for data do not show A scaling, suggests that clusters are not a random selection of nucleons but a selective process**

**Thank you for listening!!**

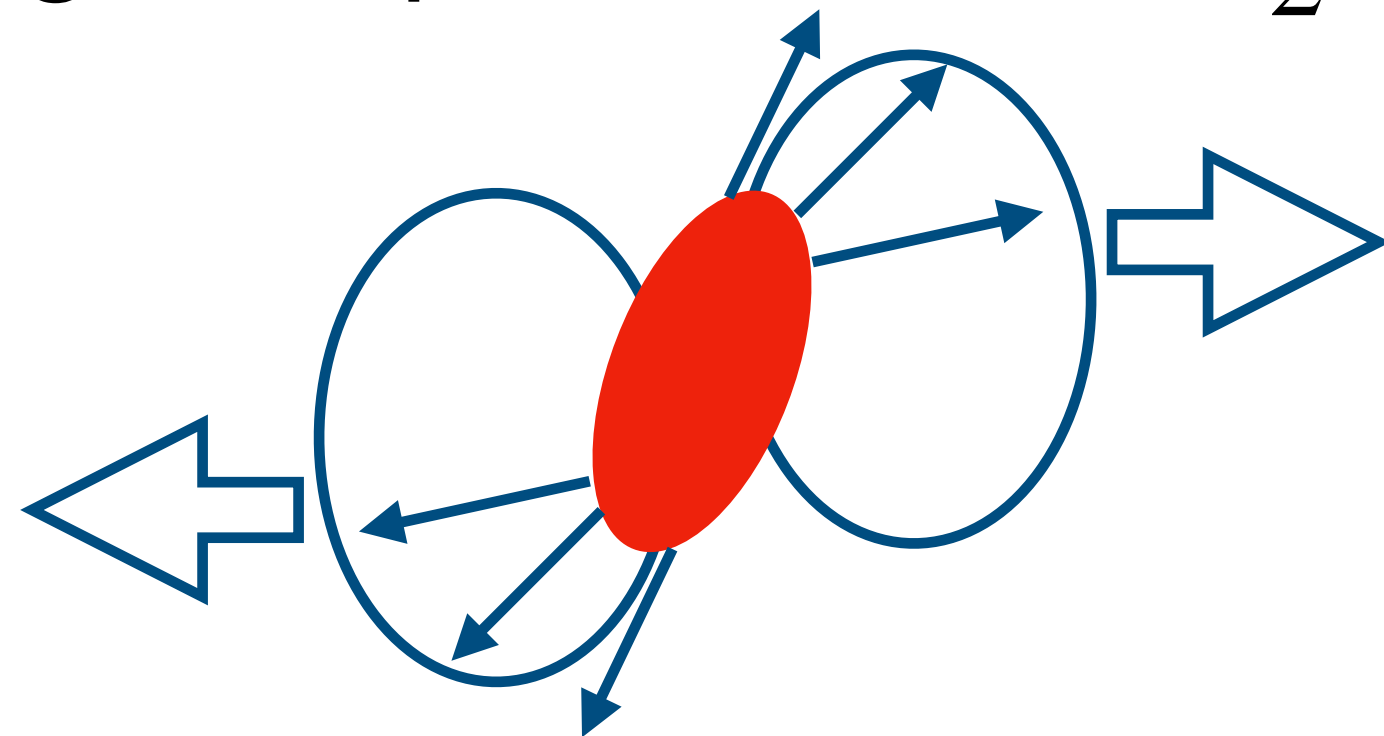
# Energy Dependence of Directed and Elliptic Flow

- Larger proton and  $\Lambda$   $v_1$  at 3 GeV

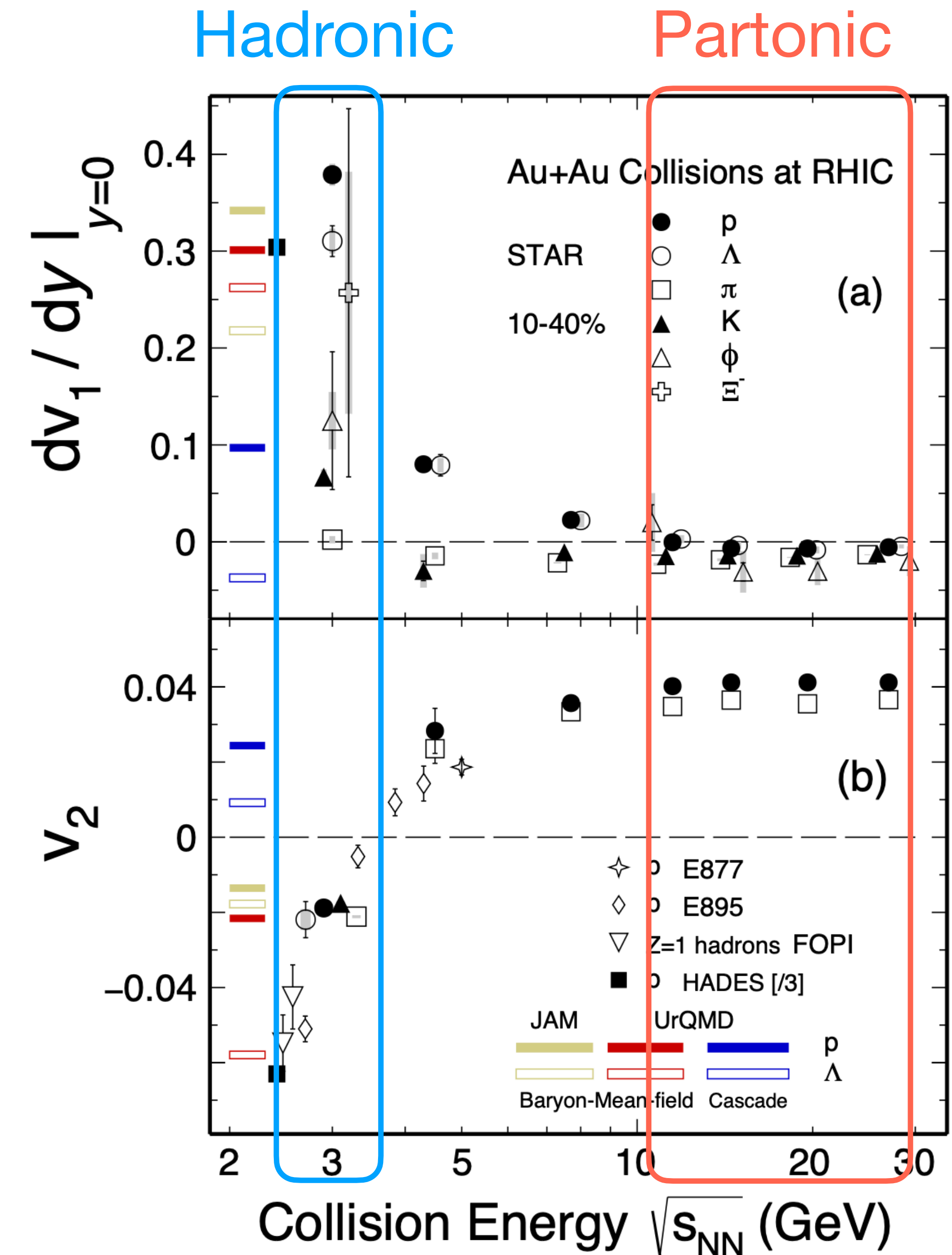


- “*Side splash*” due to repulsive baryon-baryon interactions

- Negative proton and  $\Lambda$   $v_2$  at 3 GeV



- Nuclear shadowing*: spectators inhibit in-plane flow





# Parton-Hadron-Quantum-Molecular Dynamics

= n-body microscopic transport approach for the description of heavy-ion collisions with dynamical cluster formation

PHSD



+ QMD

+ MST

Relativistic considerations

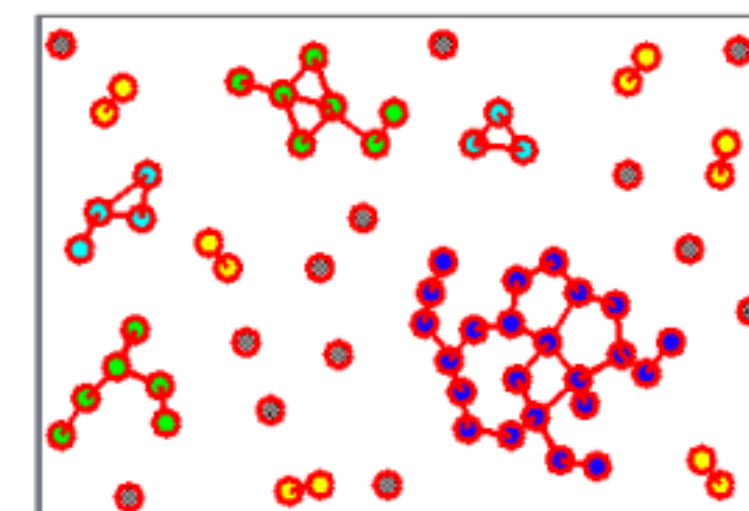
but: mean-field potentials  
=> correlations are smeared out

Correlations between baryons

n-body transport approach  
=> formation of clusters due to potential interactions

Cluster recognition

search for accumulations of particles in coordinate space



Initial A+A collisions

Formation of QGP

Partonic phase

Hadronization

Hadronic phase

QMD

Initialization nuclei

propagation of baryons

local  $\epsilon > \epsilon_c$  : dissolution of pre-hadrons

propagation of partons

scattering of partons

scattering of hadrons

propagation of mesons

PHSD

Primary collisions  
pre-hadronic states

MST or SACA



[See S. Gläsel talk on 11/11 \(Tue\)!](#)