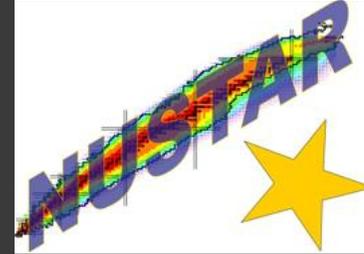


# S122@GSI: constraining symmetry energy at supra-saturation densities by means of HIC



*Fabio Risitano*

*On behalf of the ASY-EOS collaboration*

- *Centro Siciliano di Fisica Nucleare e Struttura della Materia (CSFNSM), Catania, Italy*
- *INFN Section of Catania, Catania, Italy*
- *Dipartimento MIFT, University of Messina, Messina, Italy*

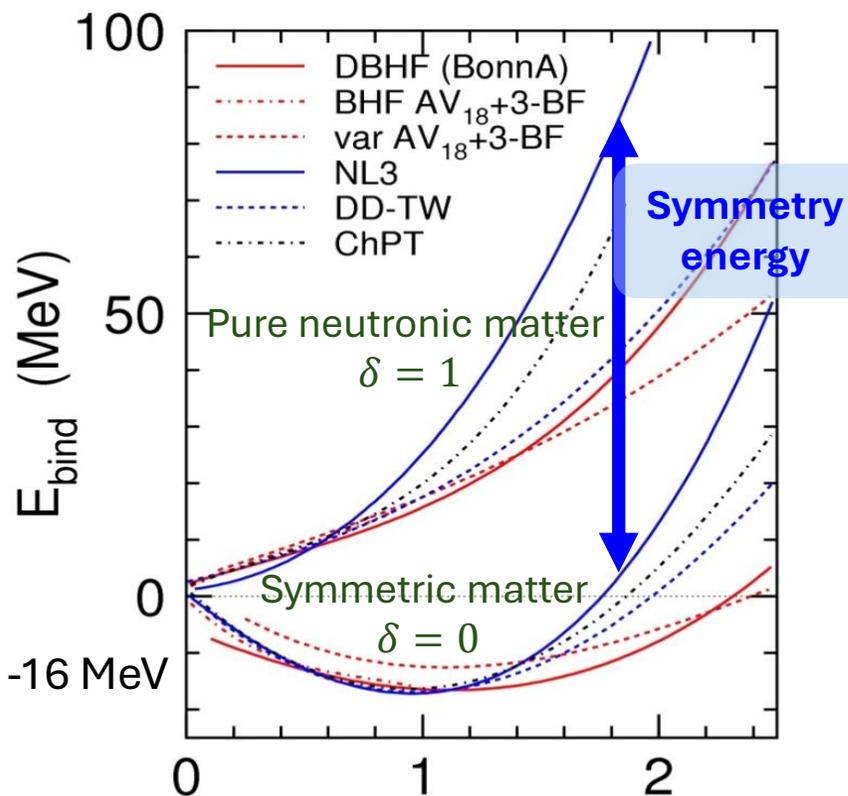


# EOS of asymmetric nuclear matter and neutron matter

Negligible

$$E(\rho, \delta) = E(\rho, \delta = 0) + E_{sym}(\rho)\delta^2 + o(\delta^4) + \dots$$

$$\delta = \frac{(\rho_n - \rho_p)}{\rho} = \frac{N - Z}{A}$$



$$E_{sym}(\rho) = E(\rho, \delta = 1) - E(\rho, \delta = 0)$$

Curvature Parameter

$$K_{sym} \equiv 9\rho_0^2 \left. \frac{\partial^2 E_{sym}}{\partial \rho^2} \right|_{\rho=\rho_0}$$

$$E_{sym}(\rho) = E_{sym,0} + \frac{L}{3} \left( \frac{\rho - \rho_0}{\rho_0} \right) + \frac{K_{sym}}{18} \left( \frac{\rho - \rho_0}{\rho_0} \right)^2 + \dots$$

Strength Parameter  
 $S_0 \equiv E_{sym}(\rho = \rho_0)$

Slope Parameter

$$L \equiv 3\rho_0 \left. \frac{\partial E_{sym}}{\partial \rho} \right|_{\rho=\rho_0}$$

$$P_{sym}(\rho) = \rho_0^2 \frac{dE_{sym}}{d\rho} = \frac{\rho_0}{3} L$$

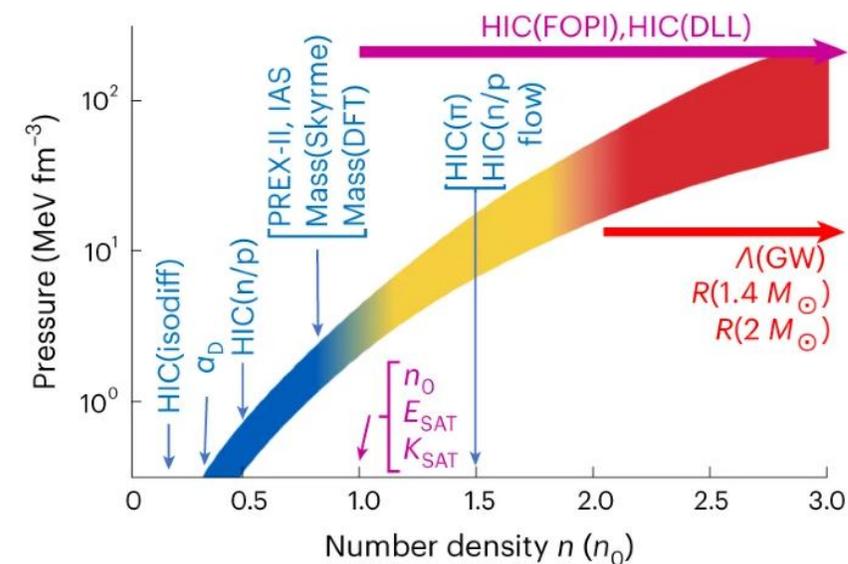
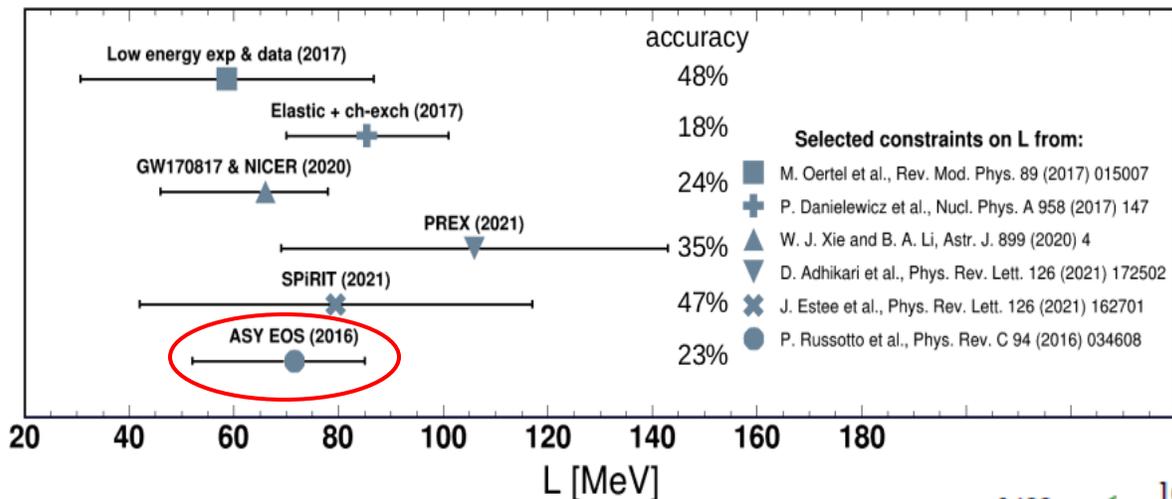
$$\rho_0 = 0.16 \text{ fm}^{-3} \quad \rho/\rho_0$$

Fuchs and Wolter, EPJA 30 (2006)

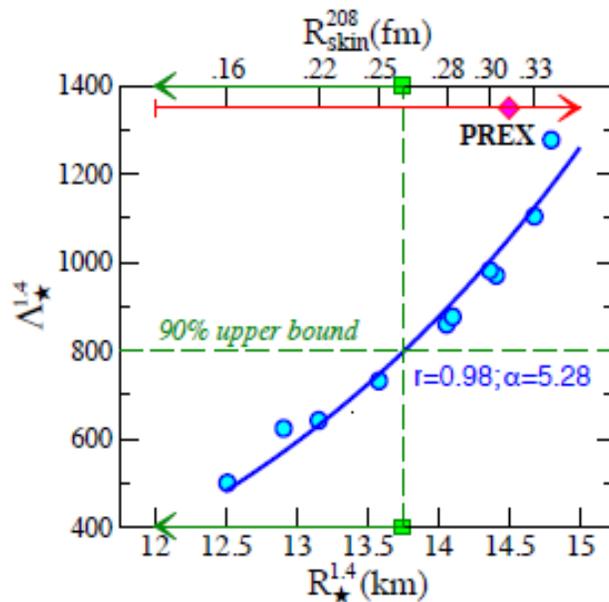
Russotto P. et al., La Rivista del Nuovo Cimento 46, (2023) 1-70

# Investigate symmetry energy in experiments

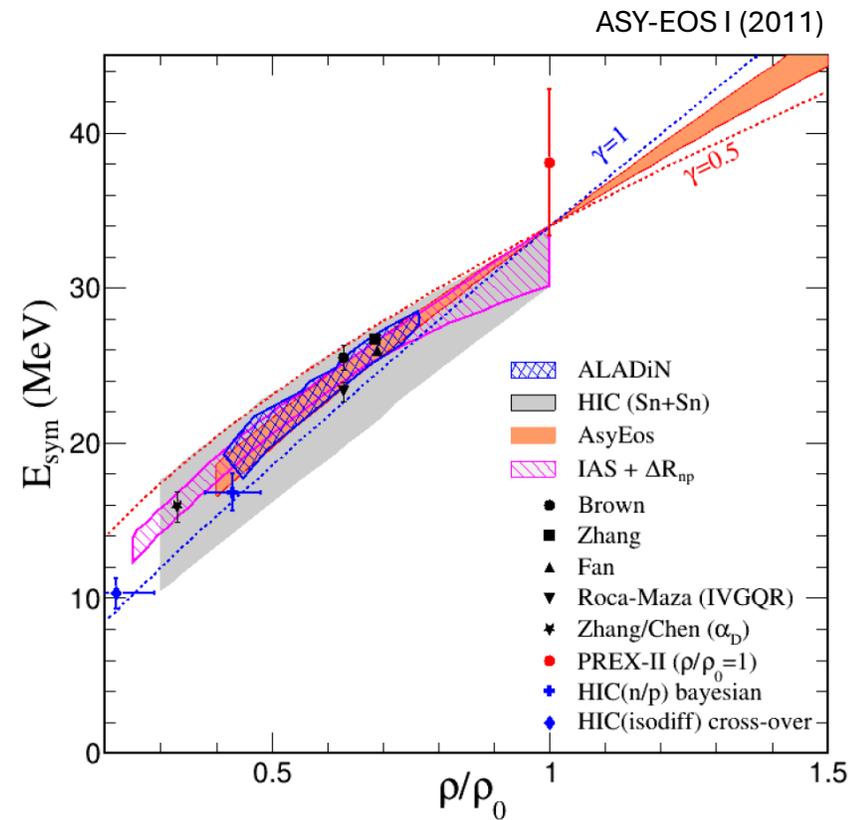
Courtesy of J. Lukasik, NUSYM2024, Caen



Chun Yuen Tsang et al., Nature Astronomy 8 328 (2024)



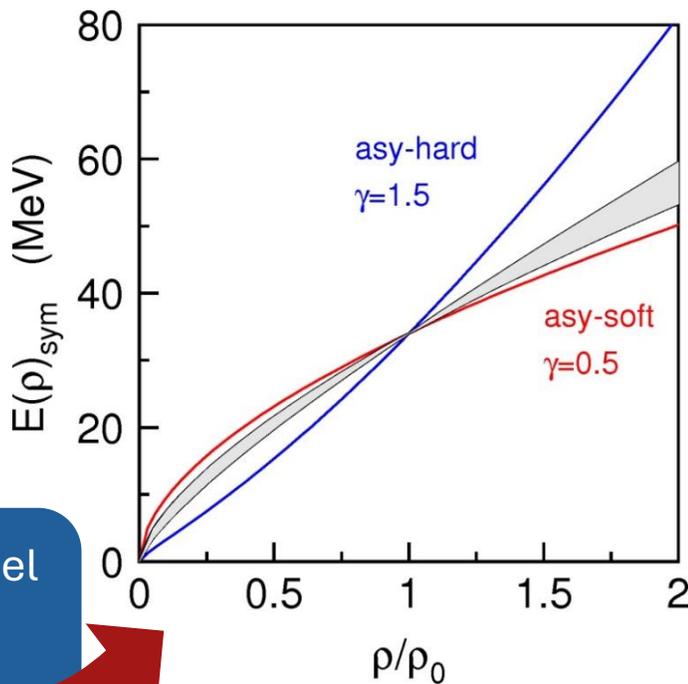
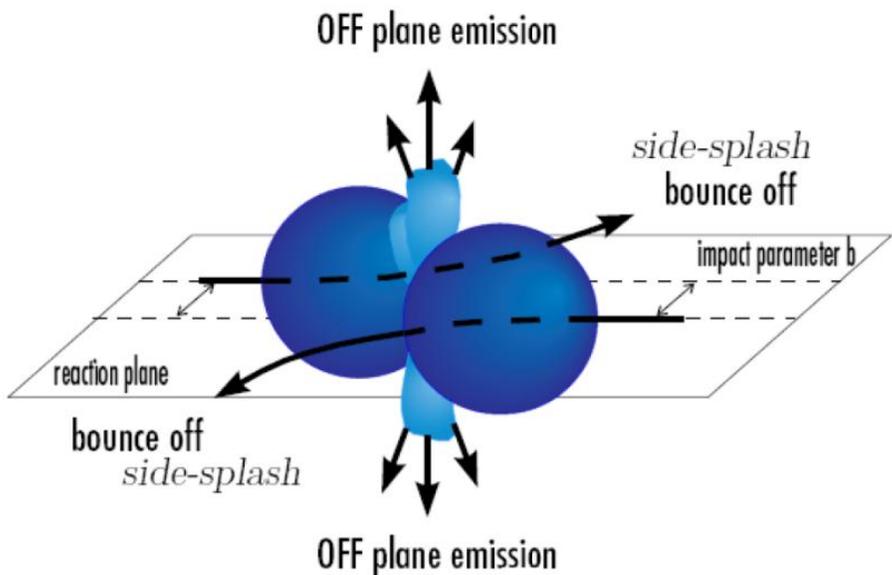
Fattoyev F.J., Piekarewicz J., Horowitz C.J., PRL 120, 172702 (2018)



Russotto P. et al., La Rivista del Nuovo Cimento 46, (2023) 1-70

# Symmetry energy at high densities: neutron/proton elliptic flow

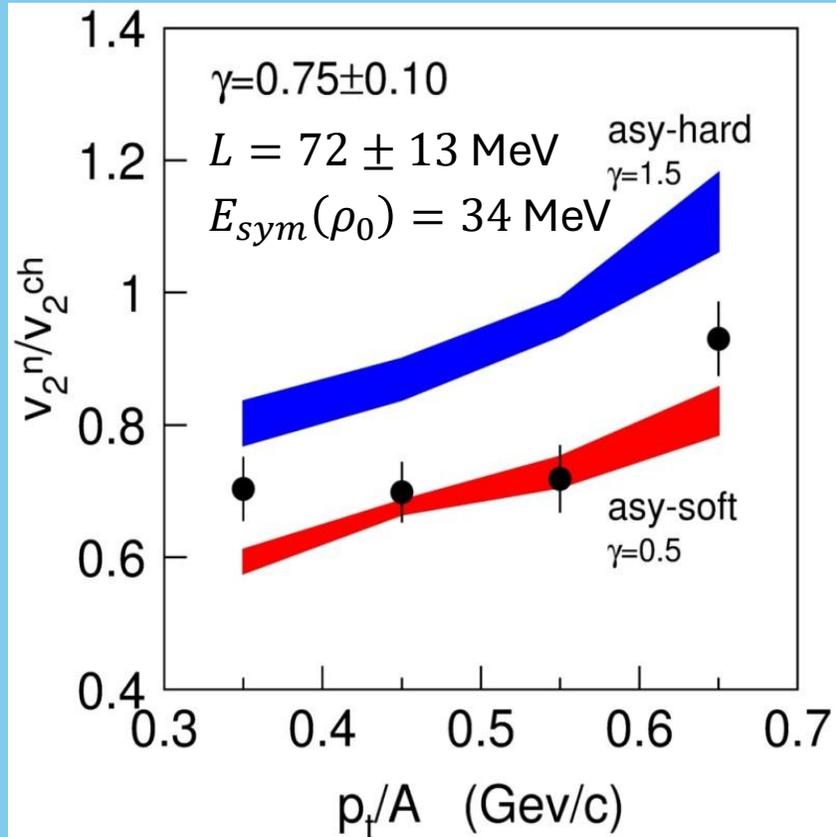
$$\frac{dN}{d(\phi - \phi_R)}(y, p_t) = \frac{N_0}{2\pi} \left( 1 + 2 \sum_{n \geq 1} v_n \cos n(\phi - \phi_R) \right)$$



Parametrization for  $E_{sym}$  used in UrQMD model

$$E_{sym} = E_{sym}^{kin} + E_{sym}^{pot} = 12 \text{ MeV} (\rho/\rho_0)^{2/3} + 22 \text{ MeV} (\rho/\rho_0)^\gamma$$

Elliptic flow from ASY-EOS I experiment (Au+Au@400 A MeV)

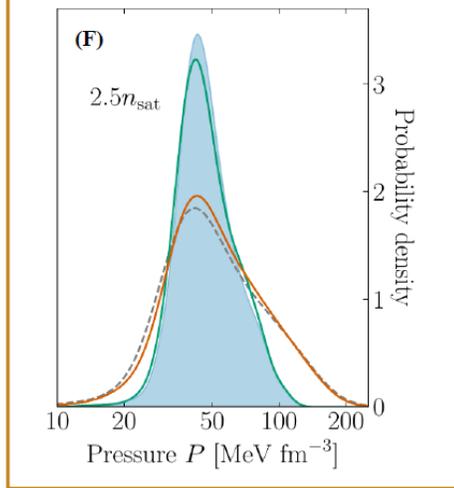
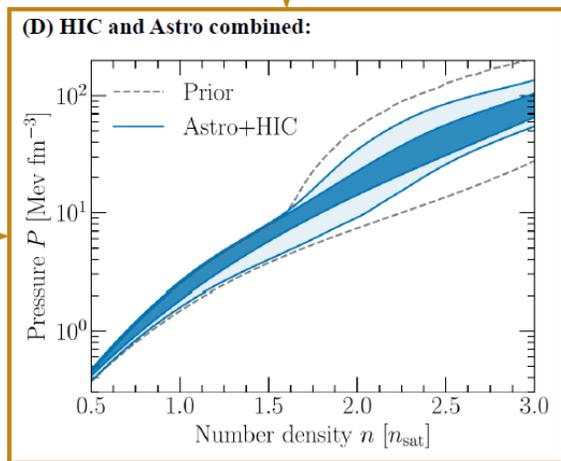
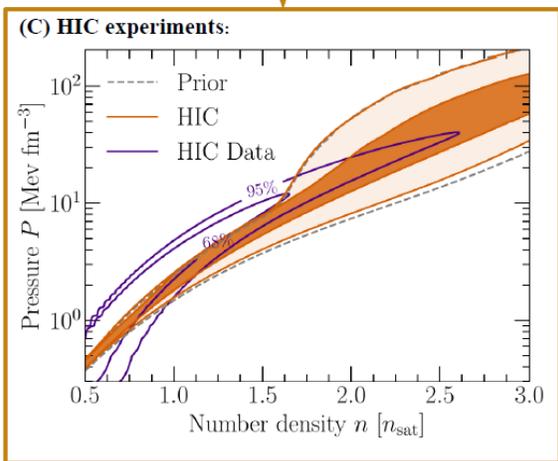
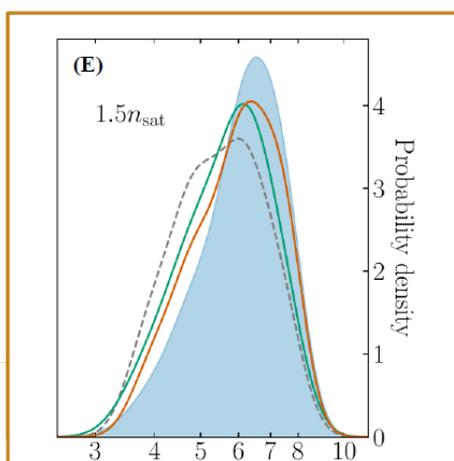
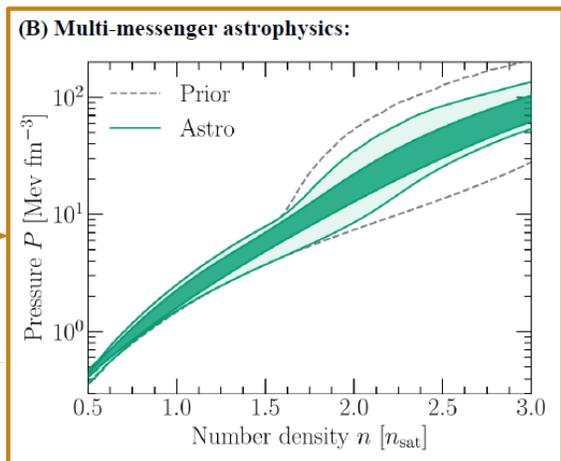
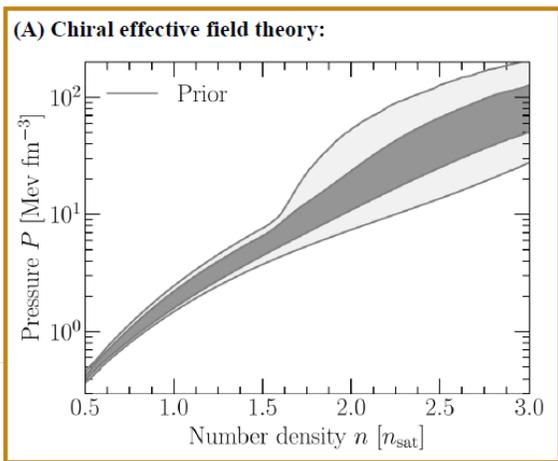


P. Russotto et al., PRC 94, 034608 (2016)

# Constraining neutron matter EOS combining HIC and astrophysics

«HIC» = FOPI+ASYEOS+AGS

«Astro» = GW, NICER (pulsar X-ray hot spots)



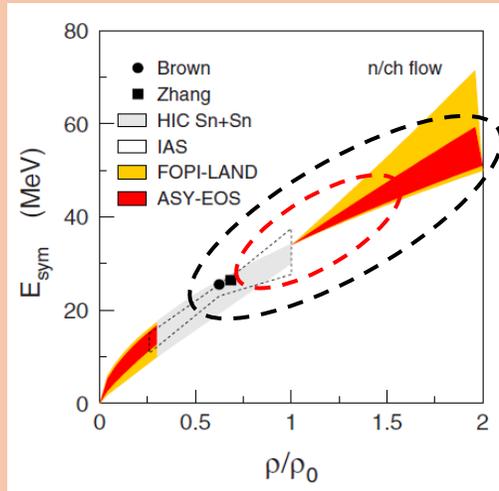
- HIC experiments probe EOS around 1–1.5  $\rho_0$  with high experimental sensitivity
- NICER measurement constrain independently giving similar observations
- Astrophysical observations dominate the current limit of EOS above  $2\rho_0$



Extending HIC experiments toward higher densities is essential to bridge **laboratory** and **astrophysical** constraints

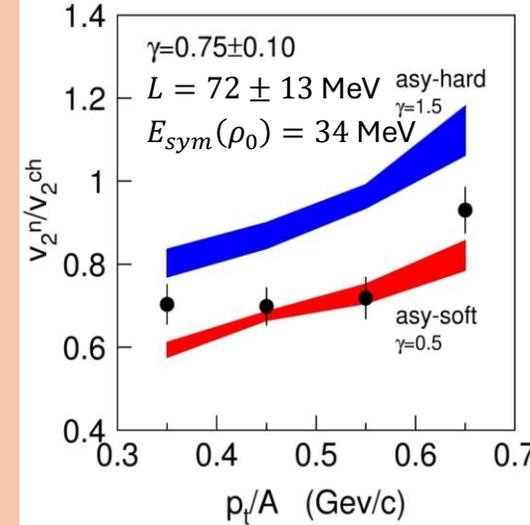
# S122 (ASY-EOS II): advancing Symmetry Energy studies towards higher densities

P. Rusotto et al., PRC 94, 034608 (2016)

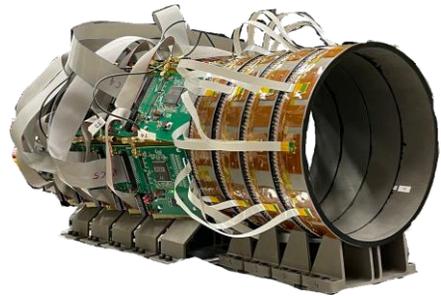


Density probed by the ASY-EOS I exp

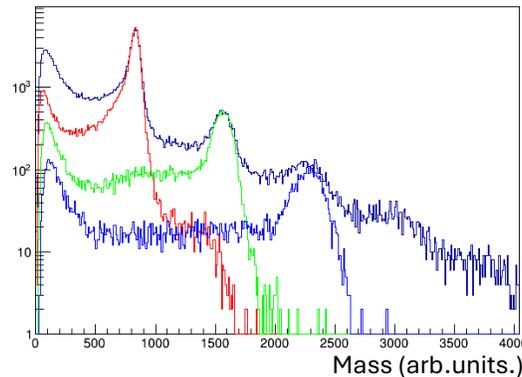
Density to be probed in the ASY-EOS II exp, most relevant for neutron star physics



Improved resolution



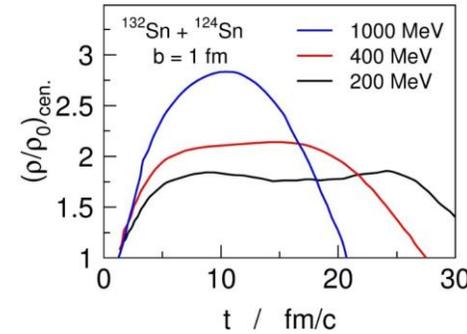
**KRAB:** new detector for reaction plane determination and on-beam centrality selection



**NeuLAND:** capability to resolve **p, d, t**

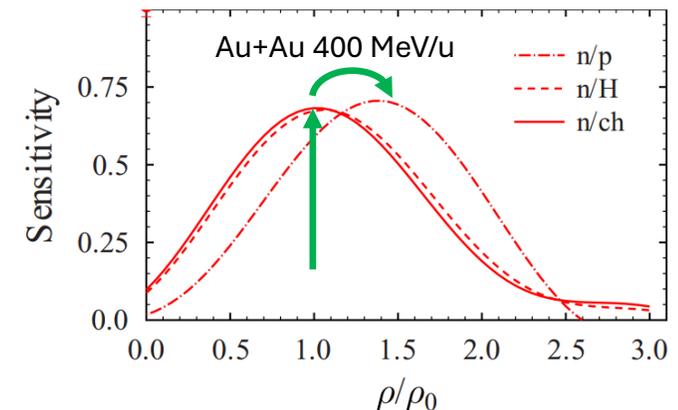
Higher densities

Li, Bao-An, NPA 708, 365 (2002)



Higher incident energies → higher densities reached (probed?)

P. Rusotto et al., PRC 94, 034608 (2016)



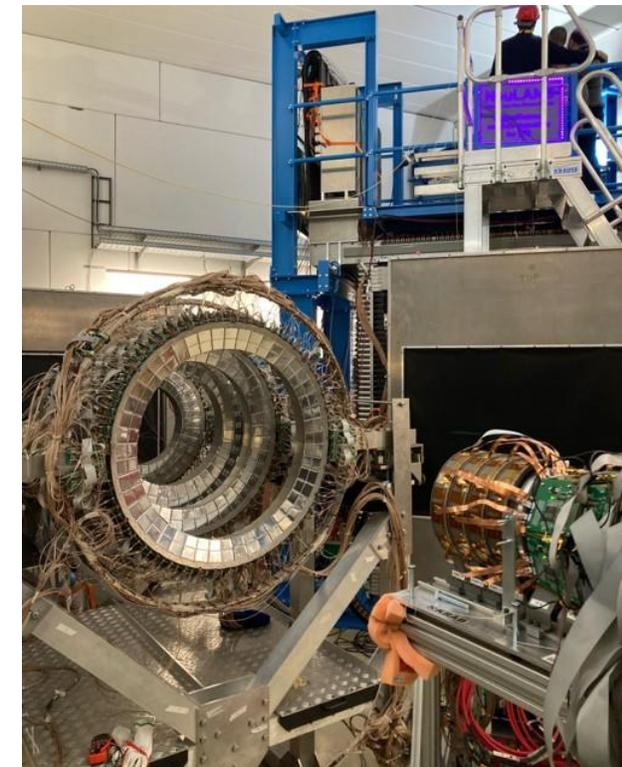
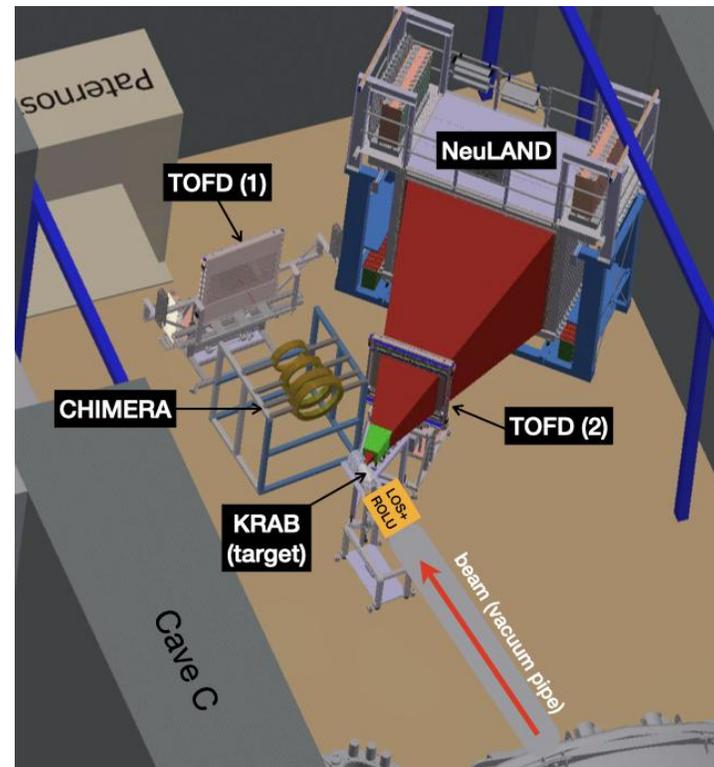
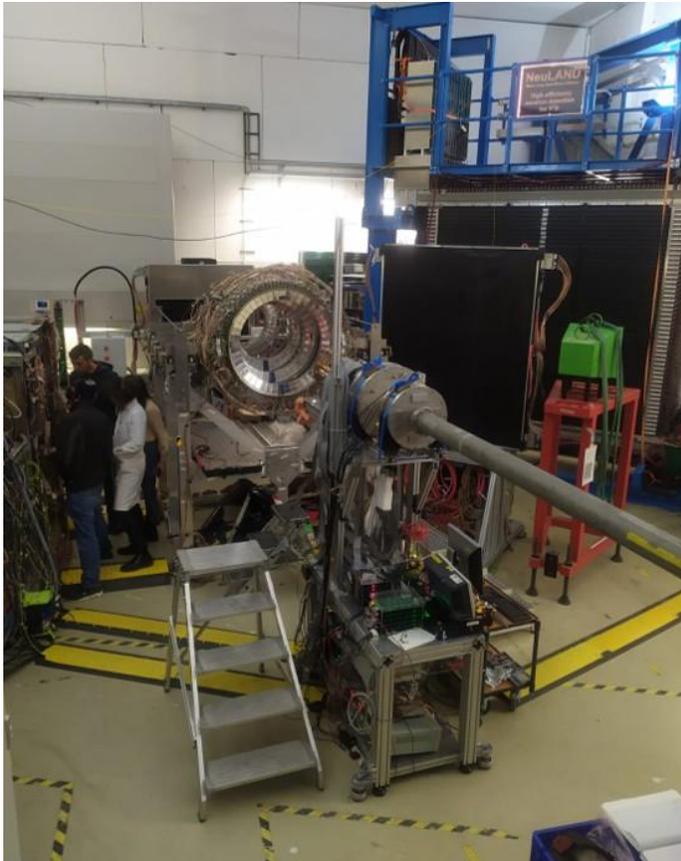
n/p observable → higher densities probed

# S122 (ASY-EOS II): advancing Symmetry Energy studies towards higher densities

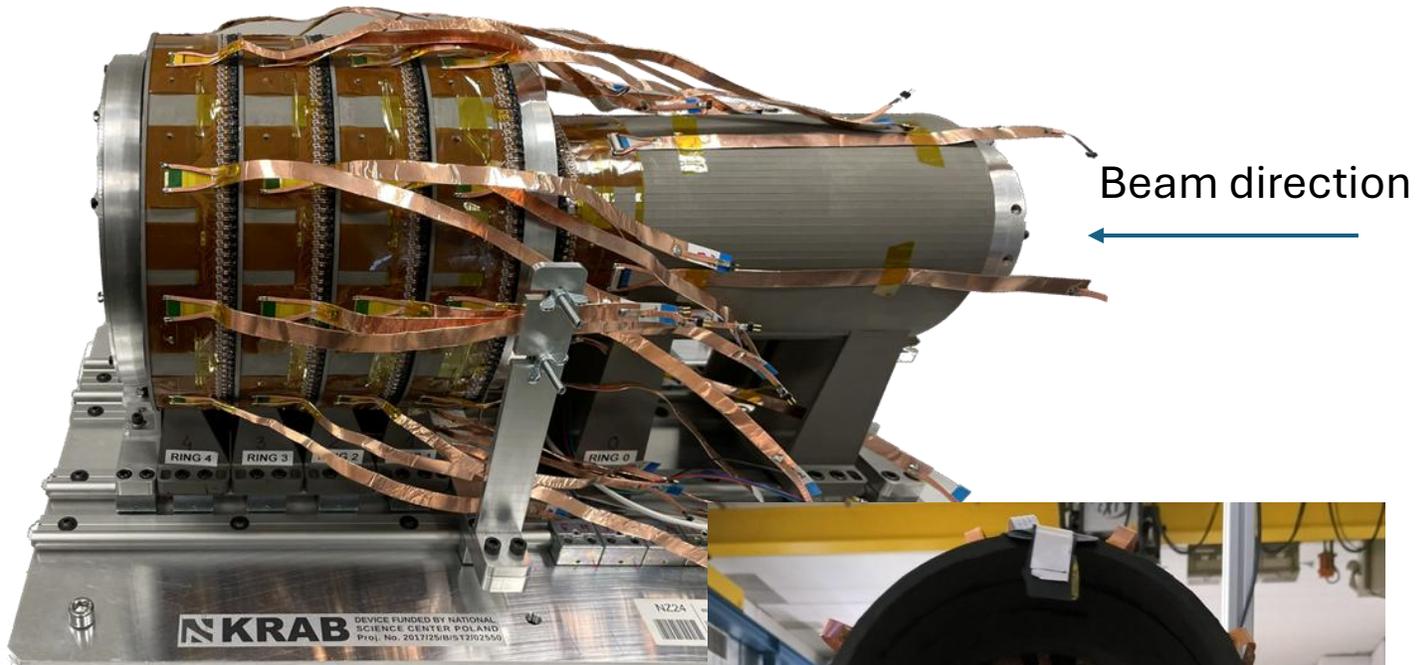
March 2025 measurements at GSI Darmstadt:

- Au+Au @ 280, 400, 600, 1000 AMeV.
- $10^5$  pps, 2 weeks

Energy (AMeV)	Main motivation
280	highest sensitivity on $K_{\text{sym}}$
400	reference to ASY-EOS I, n/p tests higher densities than before
600	highest sensitivity on L, high-density tested
1000	Maximum energy ( $\approx$ density) where models agree on $v_2$ sensitivity



# KRAB: multiplicity trigger, centrality and reaction plane detector

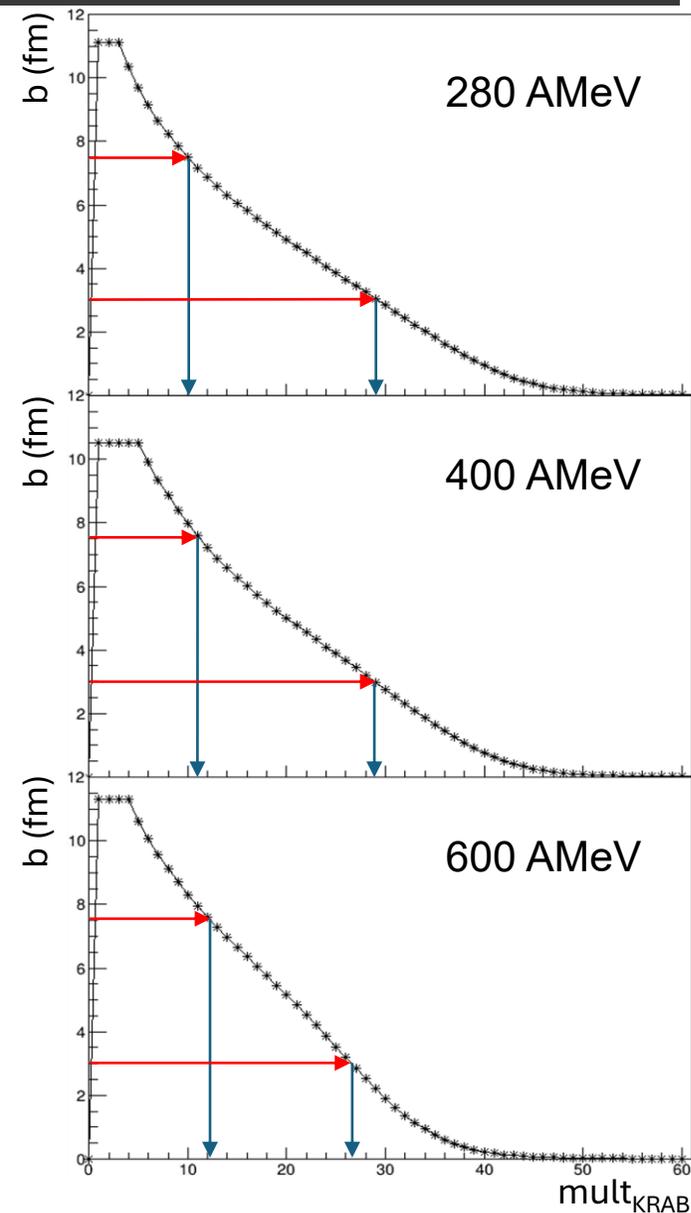
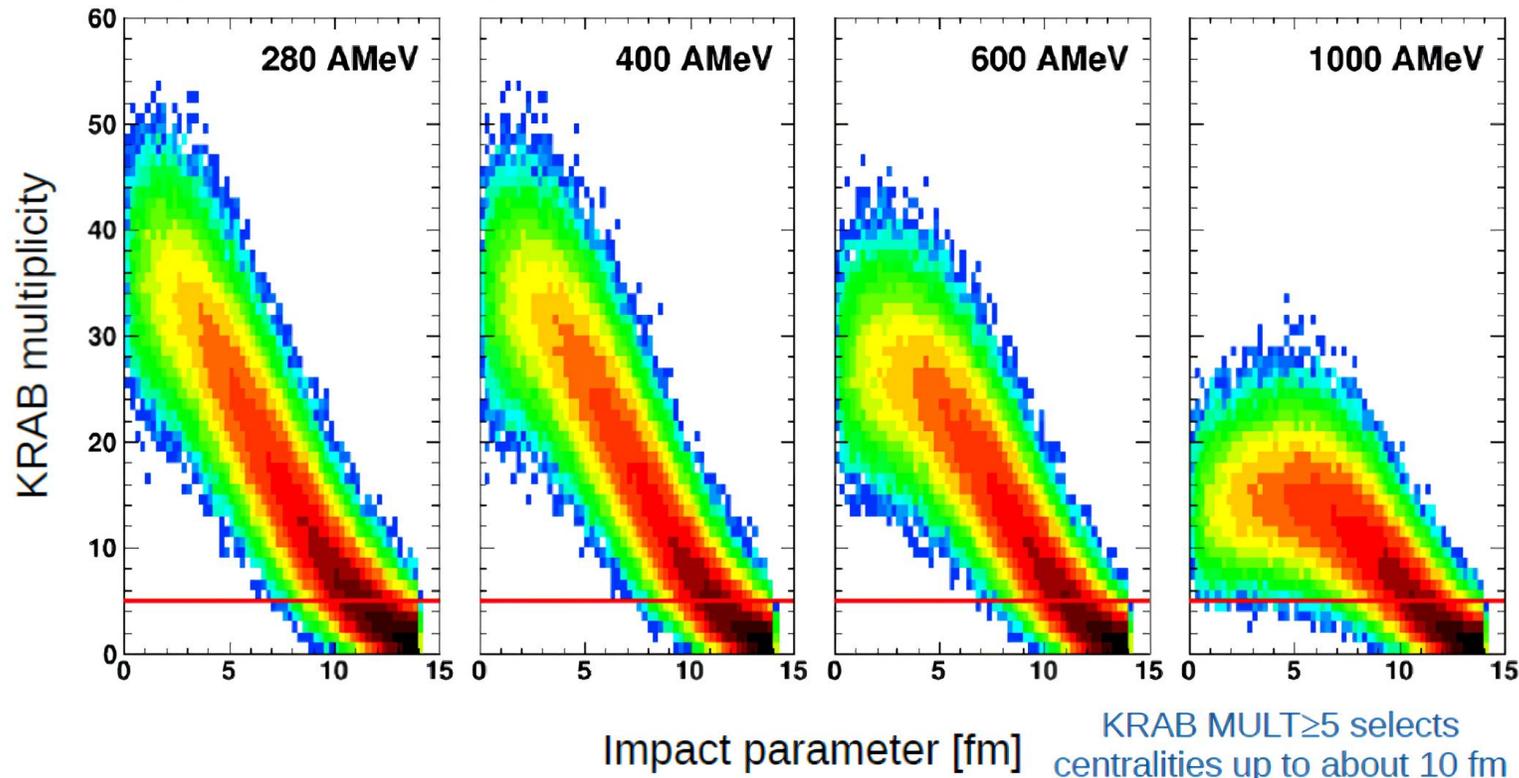


- 5 rings of  $4 \times 4 \text{ mm}^2$  segments of fast scintillating fibers, read out by  $3 \times 3 \text{ mm}^2$  SiPMs
- $4 \times 160$  segments in forward rings, 96 segments in backward ring
  - 736 total channels
- broad coverage from  $30^\circ$  to  $165^\circ$  ( $\sim 87\%$  geometric efficiency)
- $\sim 5\%$  multihit probability (for 1 AGeV Au+Au reaction)
- max radius 12 cm
- length  $\sim 43 \text{ cm}$ , 3D printed

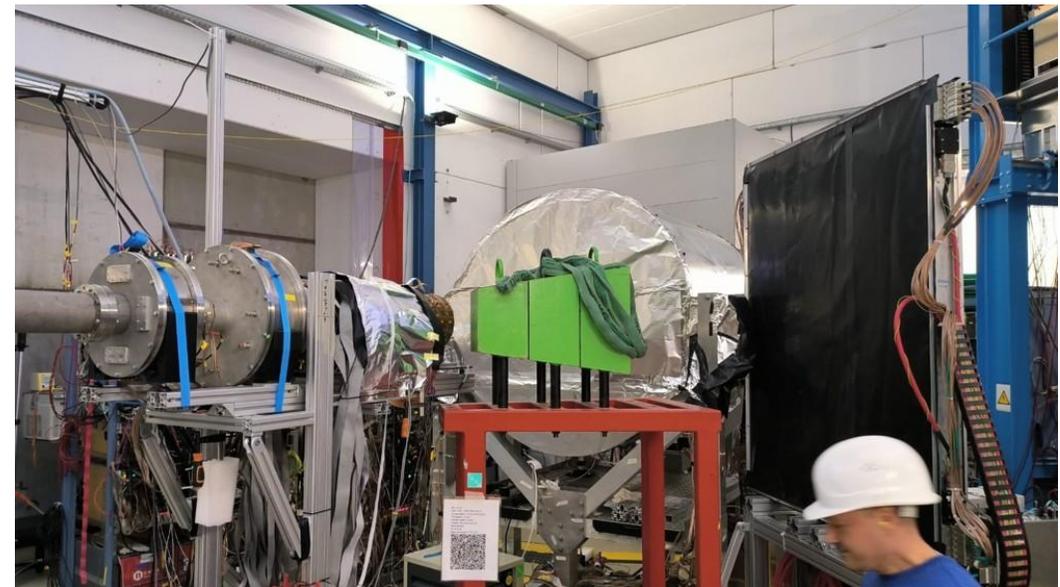
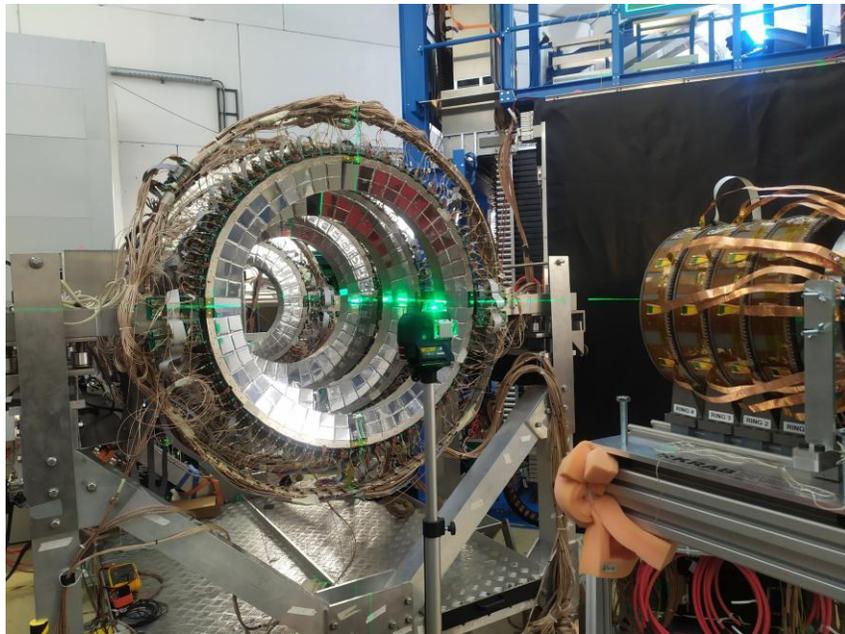
# KRAB: impact parameter through multiplicity selection

## KRAB multiplicity of fired segments vs $b$ for Au+Au collisions

UrQMD 3.4 + Clustering + GEANT4 + KRAB + photon number thresholds applied

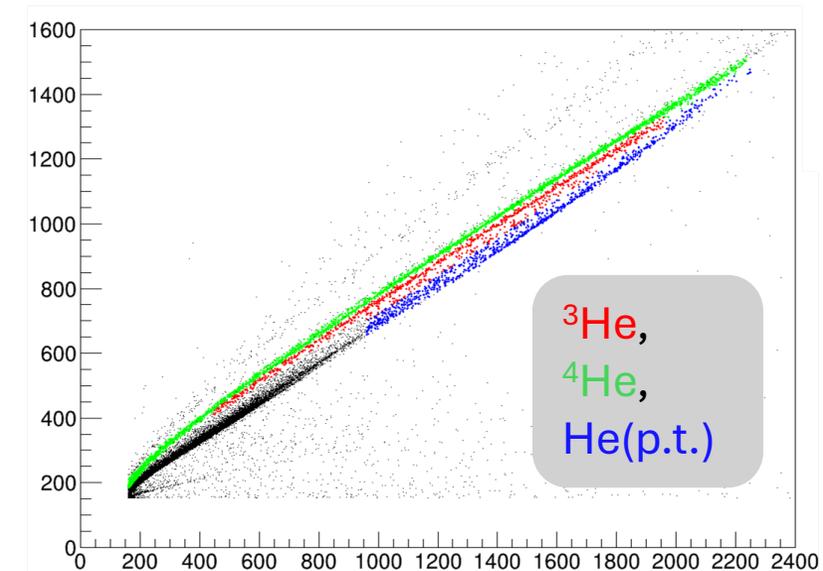
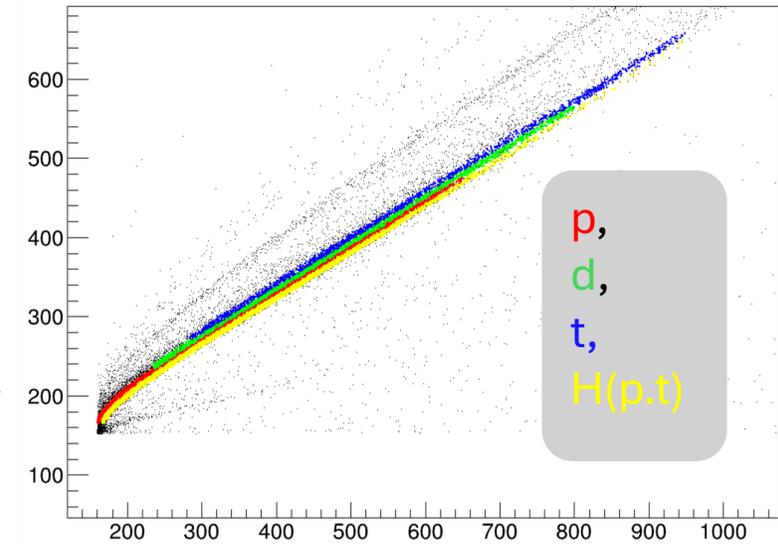
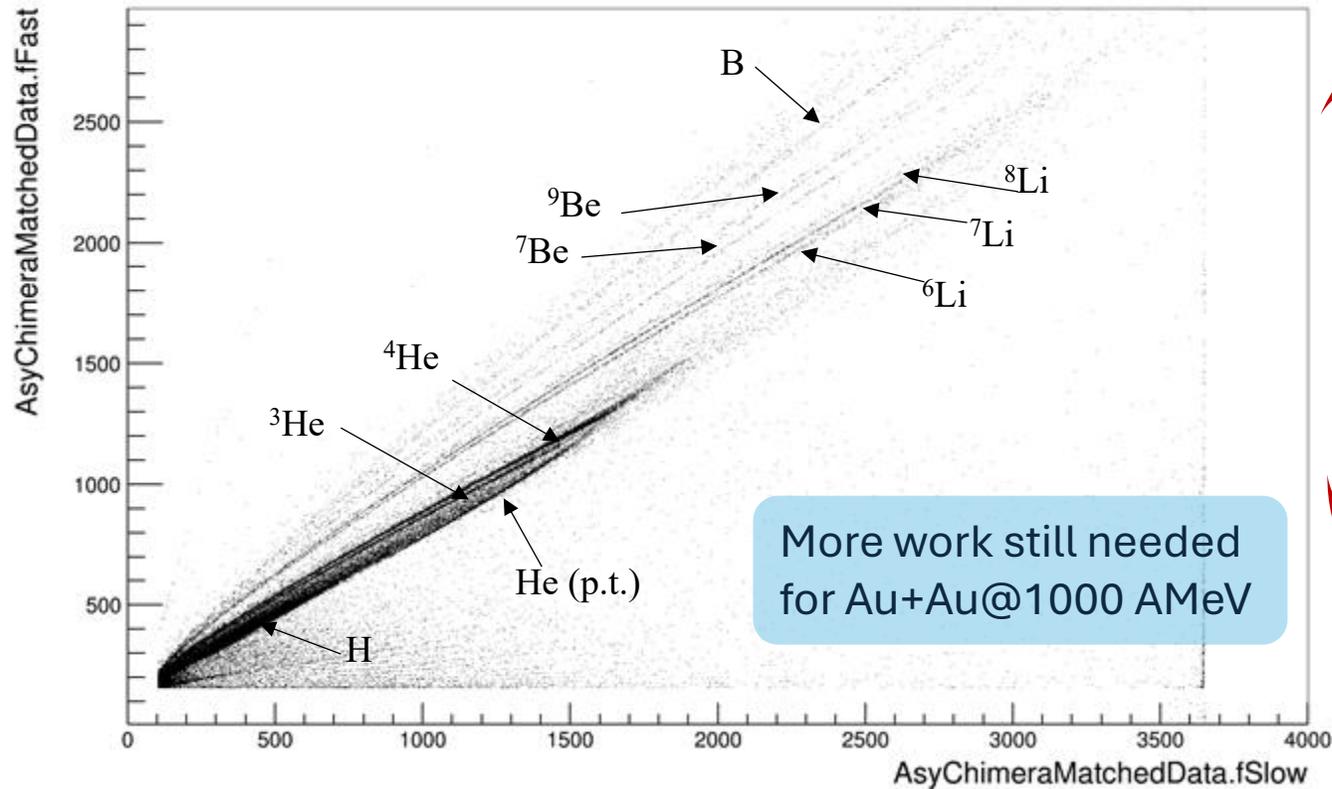


made by J. Łukasik

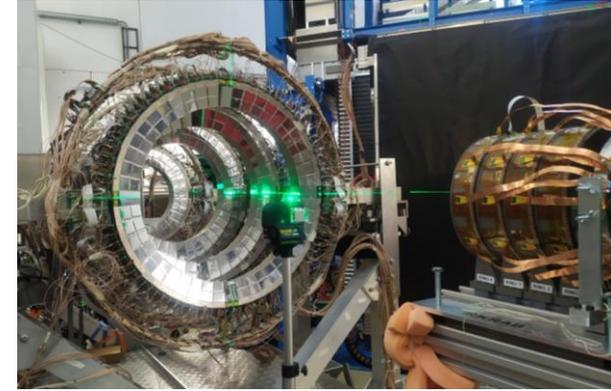
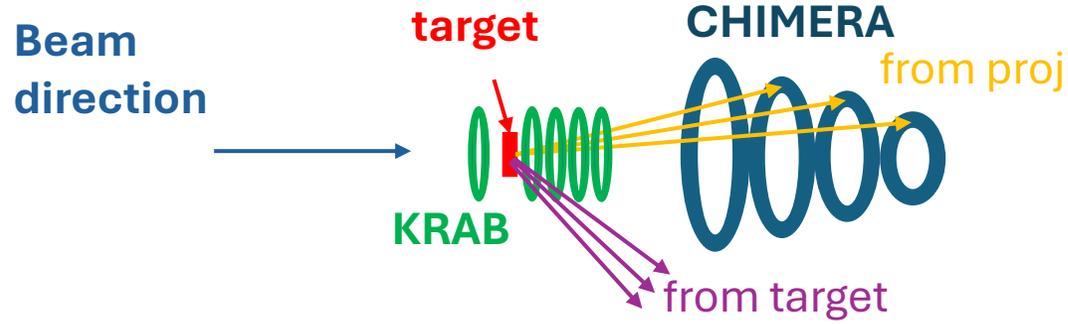


4 wheels (8 rings)  
 320 CsI(Tl) scintillators 12 cm thick  
 Polar angles  $4.6^\circ \leq \theta \leq 16^\circ$

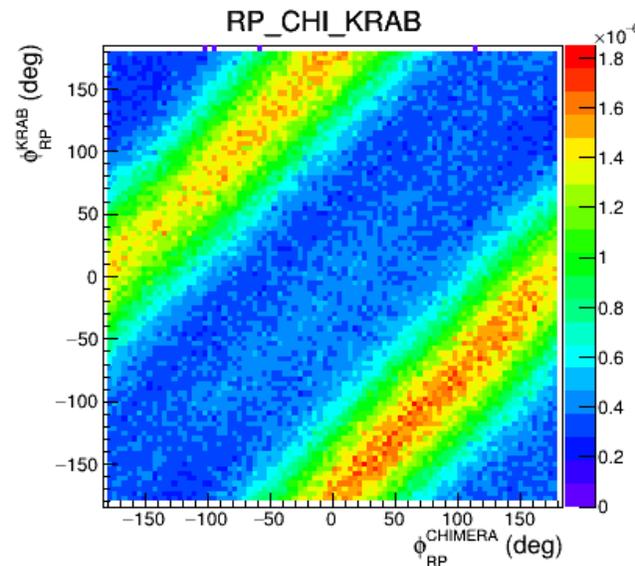
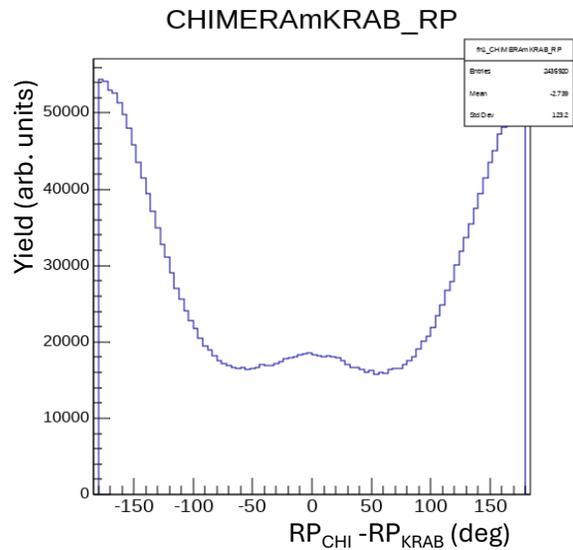
Identification plot of CsI(Tl) fast-slow signals at 400 AMeV



# CHIMERA-KRAB correlations

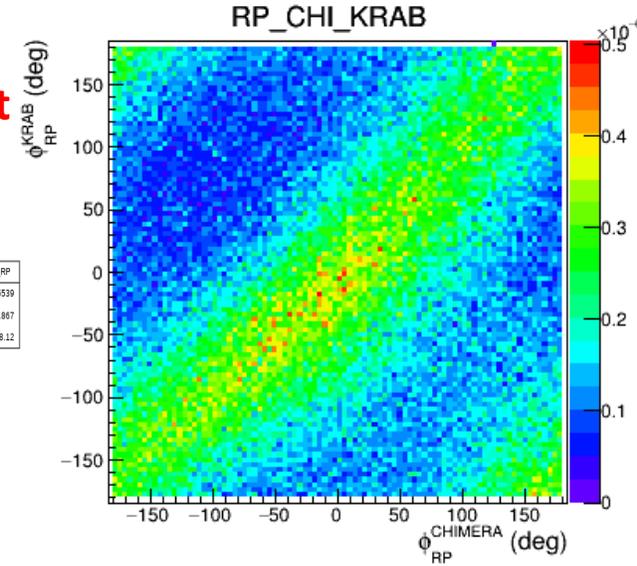
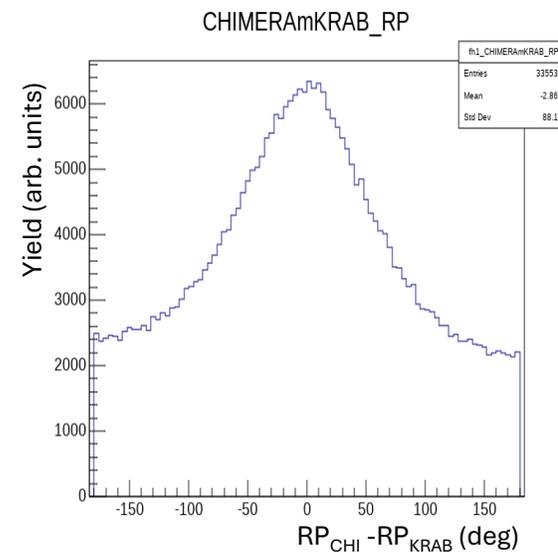


**Au 400 AMeV + Au 2%**



$\Delta\phi \approx 180^\circ$

**Au 400 AMeV + EMPTY target**



$\Delta\phi \approx 0^\circ$

# CHIMERA-KRAB correlations: reaction plane determination

ASY-EOS I  $\chi = 1.58$

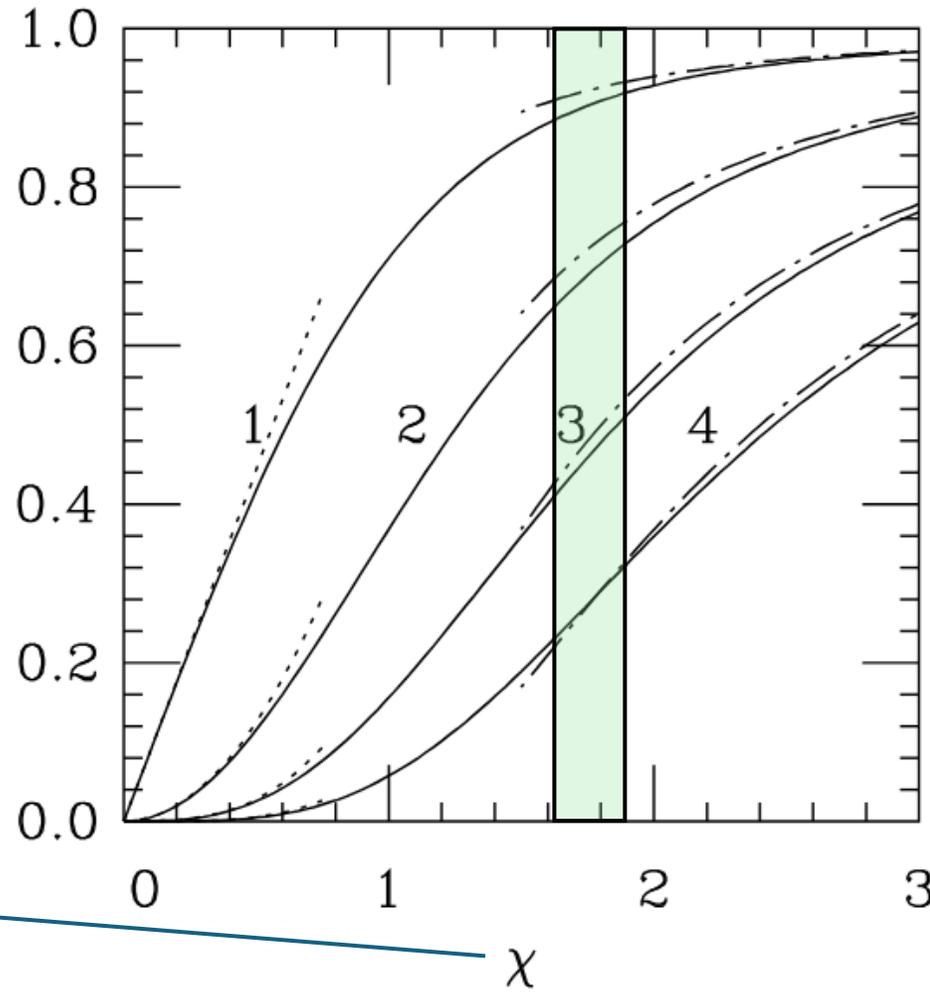
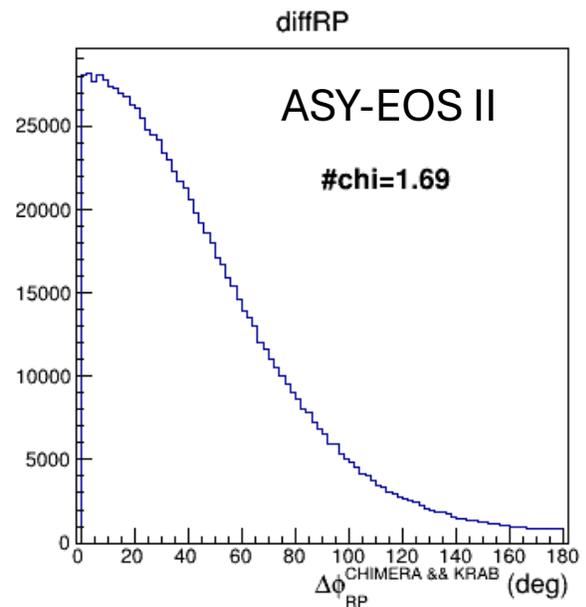
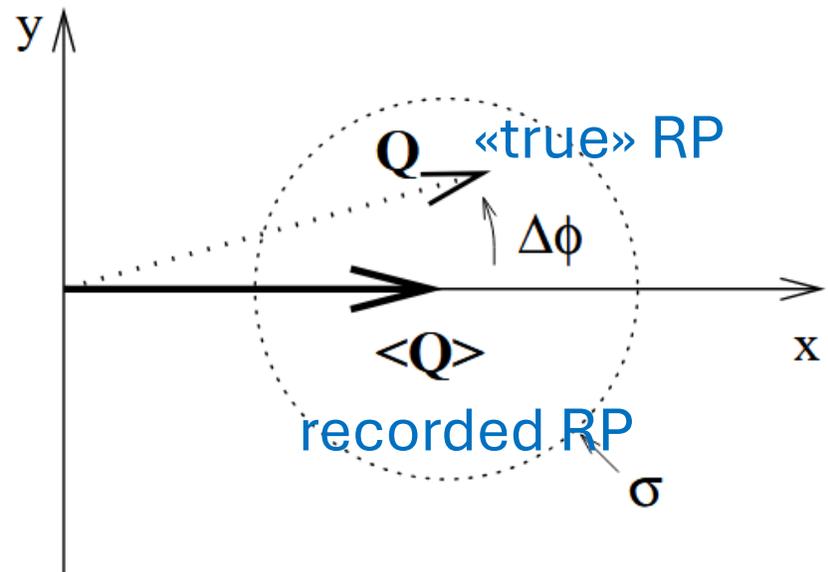
$$\frac{dN}{d\phi} = \frac{1}{2\pi} \left( 1 + 2 \sum_{n \geq 1} c_n \cos n\phi \right)$$

Measured flow

$$\langle \cos n\psi \rangle = \langle \cos n\phi \rangle \langle \cos n\Delta\phi \rangle$$

Attenuation coefficient

$\langle \cos n \Delta\phi \rangle$



Jean-Yves Ollitrault, arXiv:nucl-ex/9711003v2

# Forward TOFD: impact parameter ( $Z_{bound}$ )

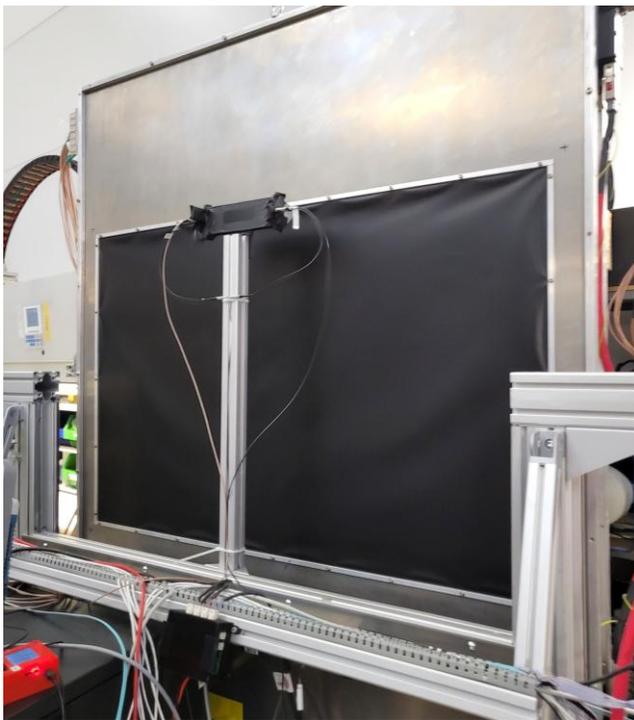
- 44 vertical plastic scintillating bars
  - Each readout by 2 photomultipliers
- Precise time-of-flight measurements
- Charge identification
- Position calibration

- Forward planes: charge, impact parameter
- 50° double plane: veto for charged particles for NeuLAND

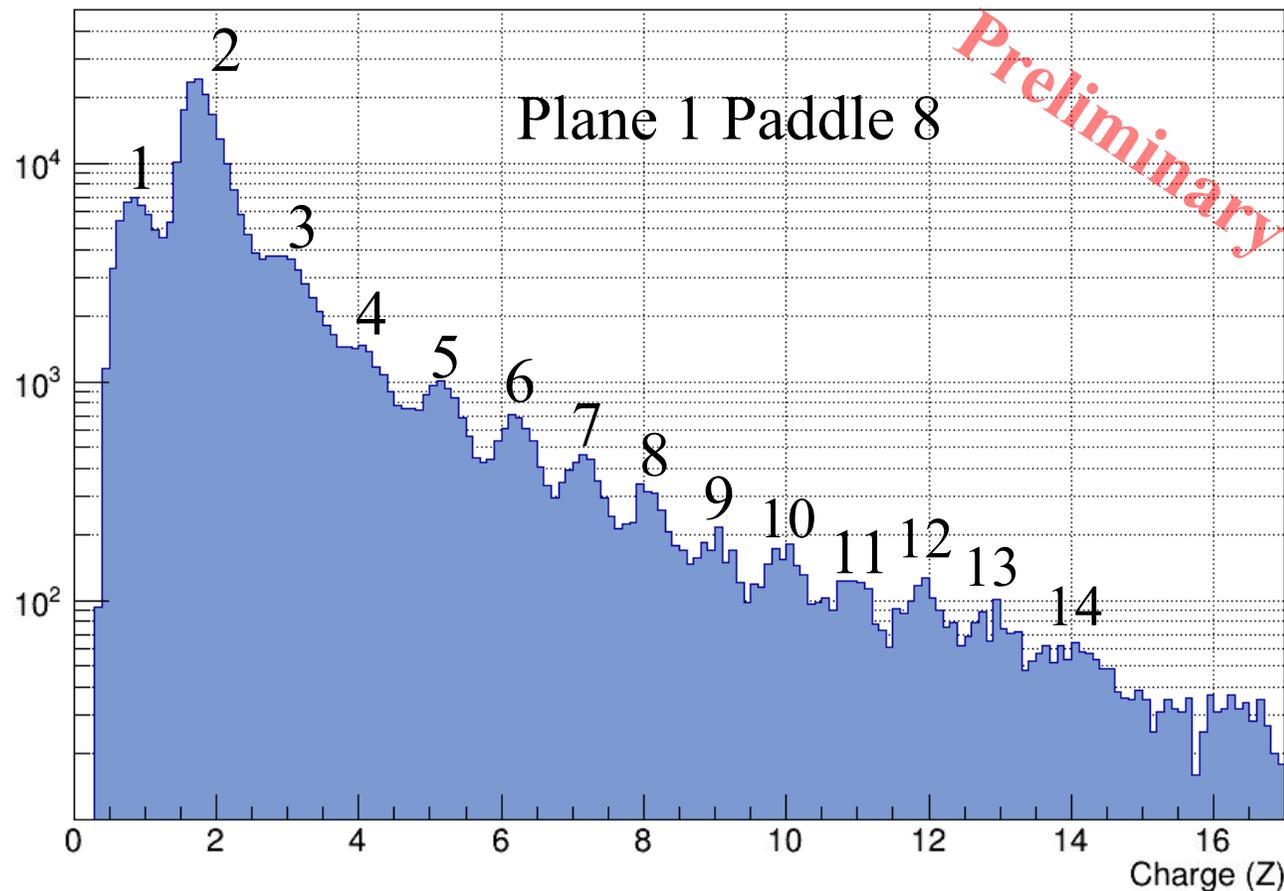
$$Z_{Bound} = \sum_{i=1}^M Z_i$$

(with  $Z_i \geq 2$ )

made by F.R.

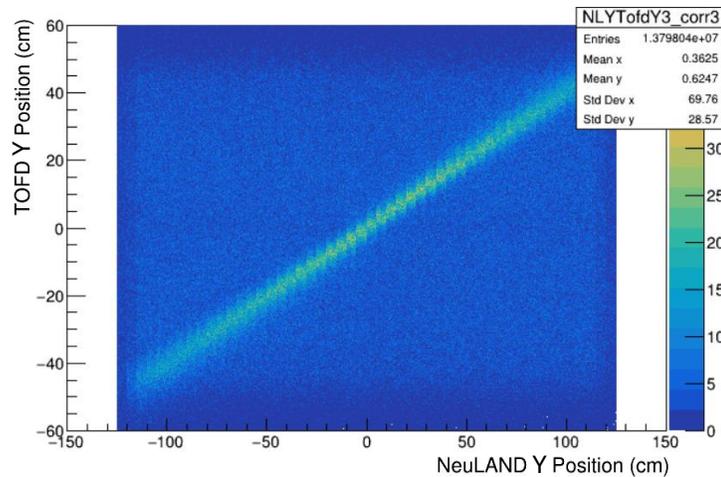
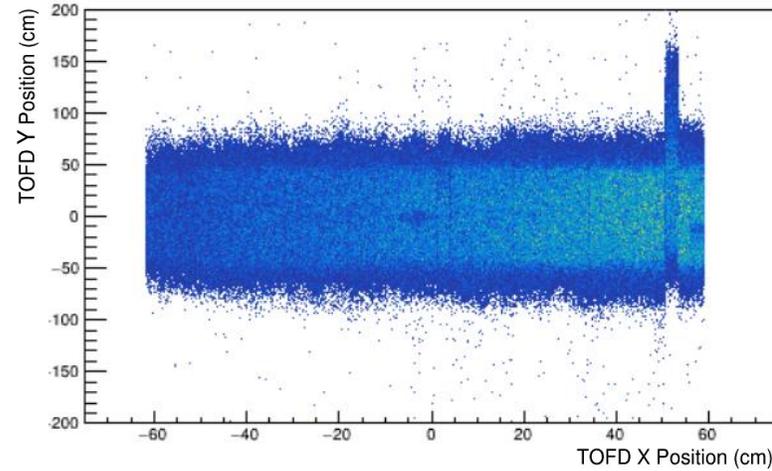


Will be further improved with coincidence of plane 1-2 adjacent paddles



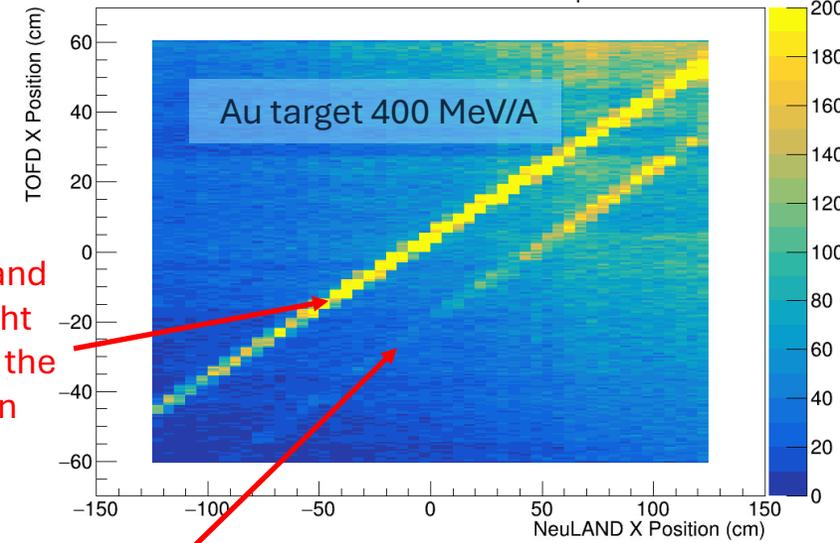
# Veto TOFD for NeuLAND

## Y position calibration and correlation

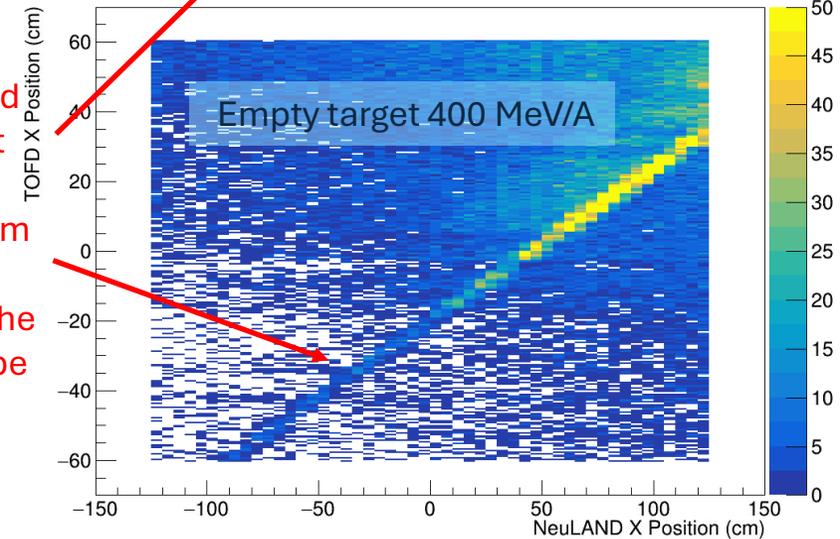


## X position correlation

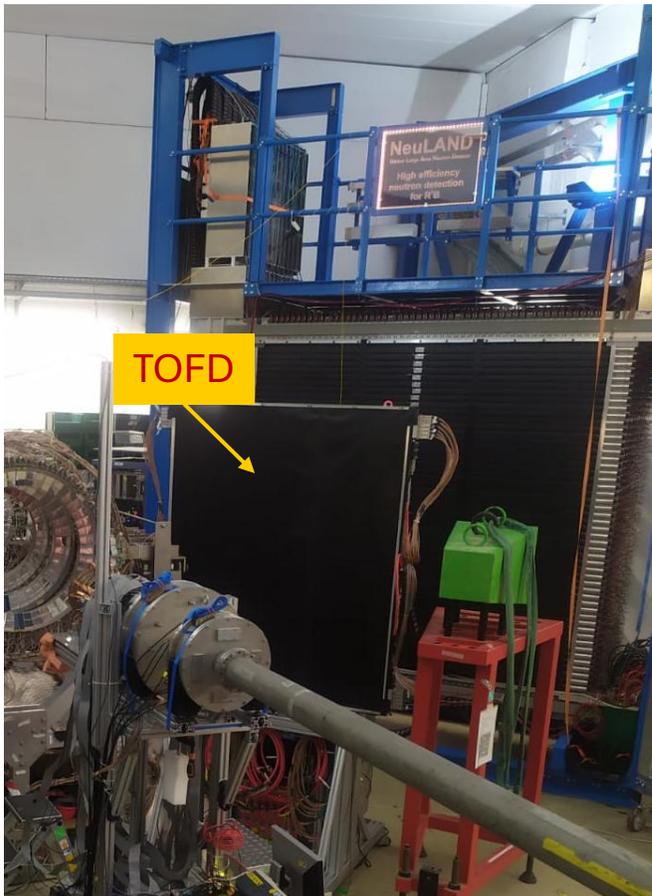
Neuland Plane 2 vs ToFD plane 4



Protons and other light ions from the reaction target

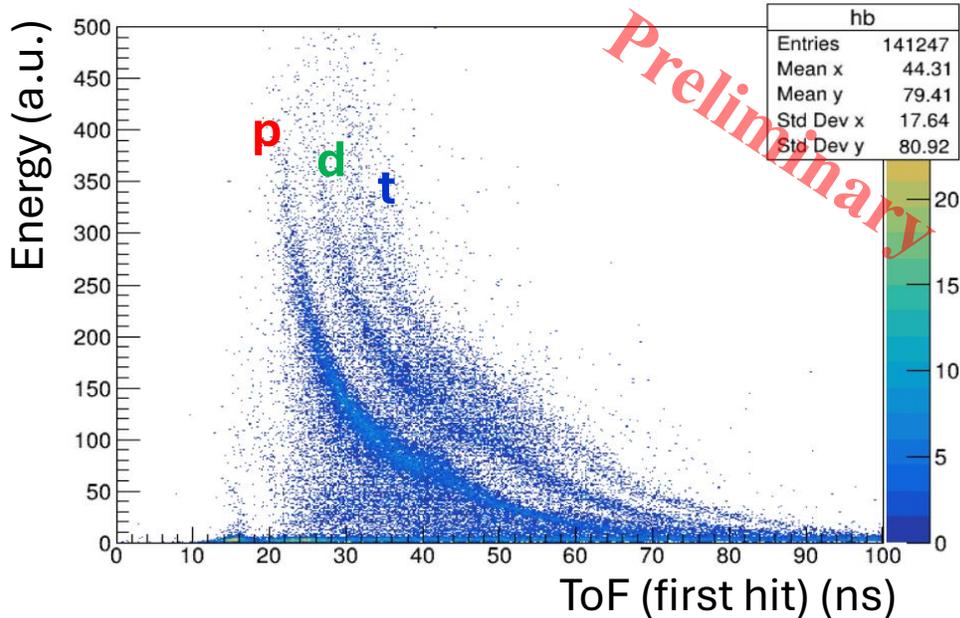


Protons and other light coming (mainly) from the exit window of the vacuum pipe

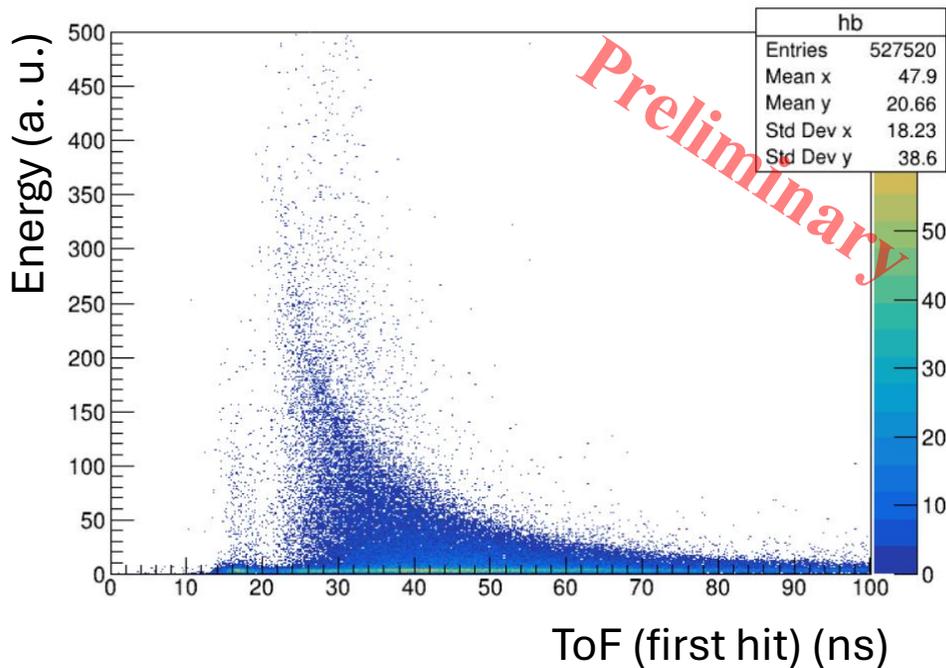


made by E. Gambera and I. Gasparic

# NeuLAND: n/p elliptic flows

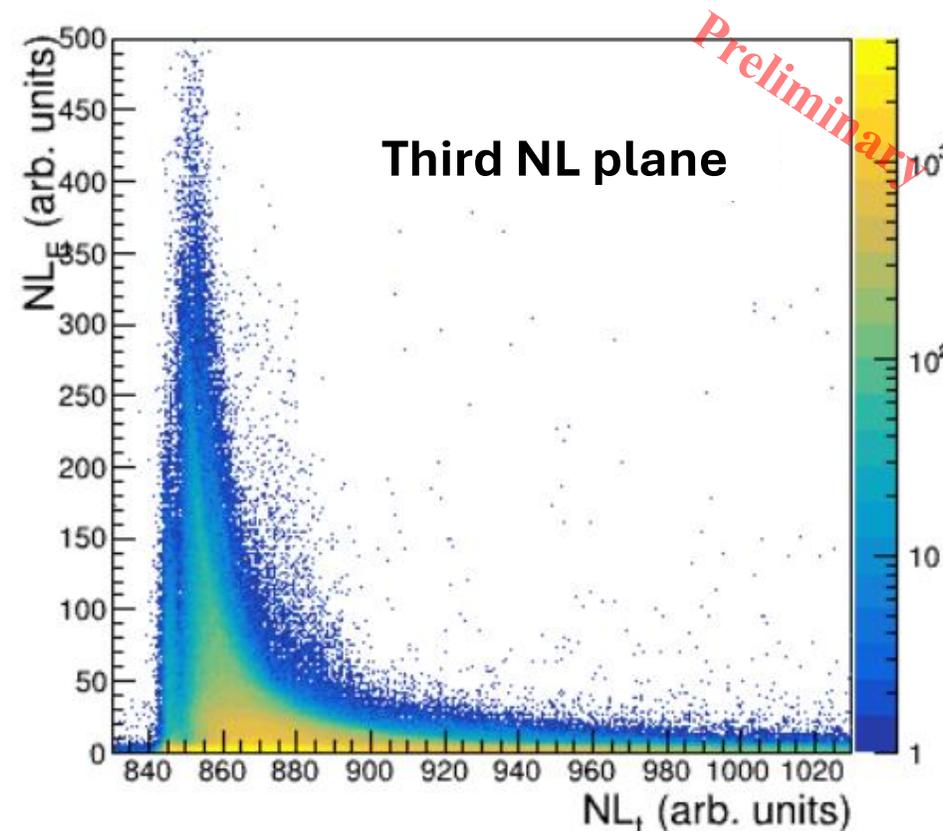
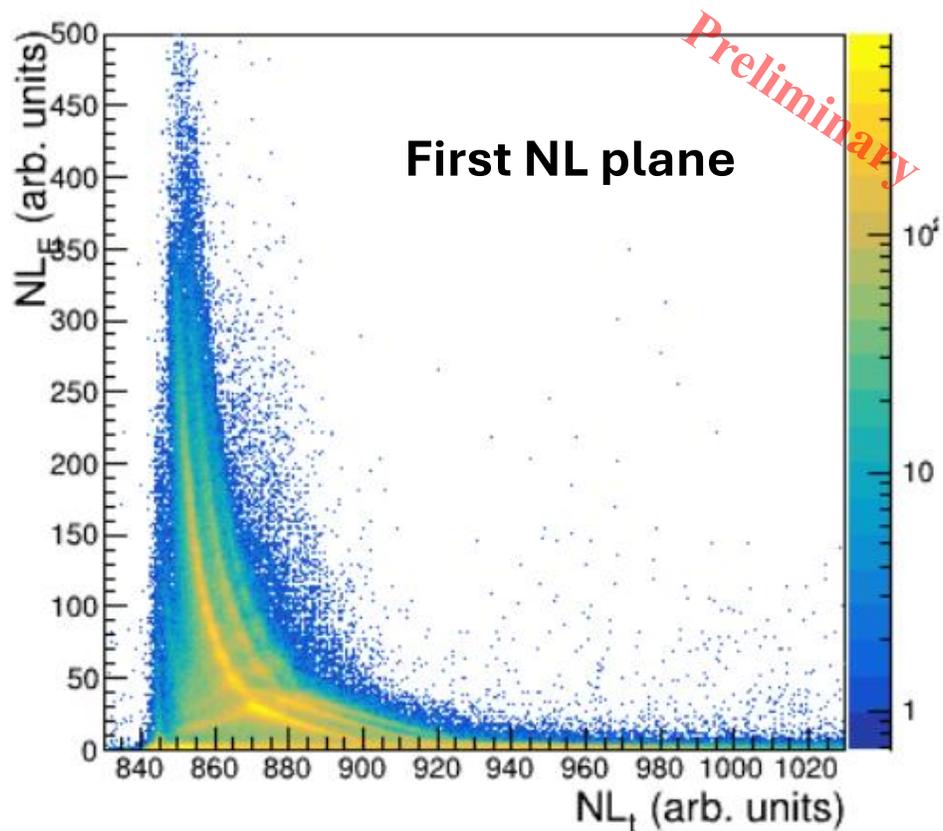


Raw **neutron** identification obtained using the complementary condition (no first plane fired)



- Preliminary hit clusterizations and raw isotopic identification in Au+Au (280 MeV/A)
  - obtained using the NeuLAND first plane as «veto» detector.
- Charged particles (**p**, **d**, **t**) are well resolved.

# NeuLAND: preliminary cluster analysis



Time-of-Flight vs cluster energy deposited on NeuLAND, in which the first hit of the “cluster” is in the first plane; p, d, t are clearly visible

Time-of-Flight vs cluster energy deposited on NeuLAND, with the first hit of the “cluster” occurring in the third plane: n and  $\gamma$  rays

# NeuLAND-Veto [ $\Delta x$ , $\Delta y$ ] position correlation

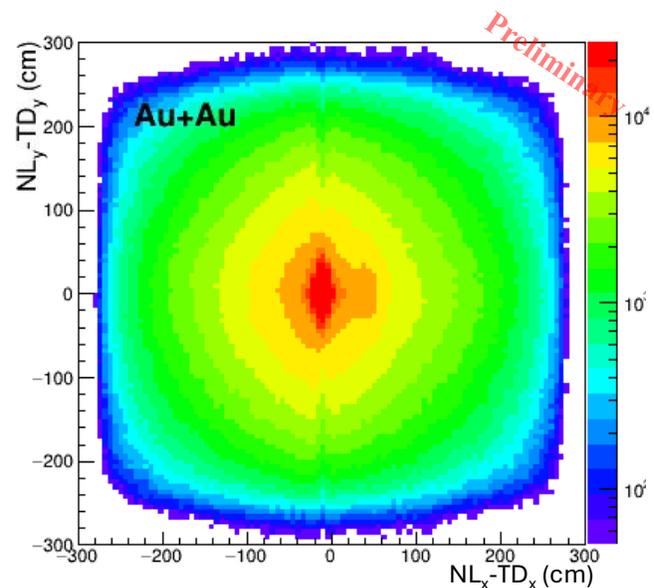
Horizontal  $\Delta x$  vs Vertical  $\Delta y$  position difference between NeuLAND and TOFD (veto) (normalized)



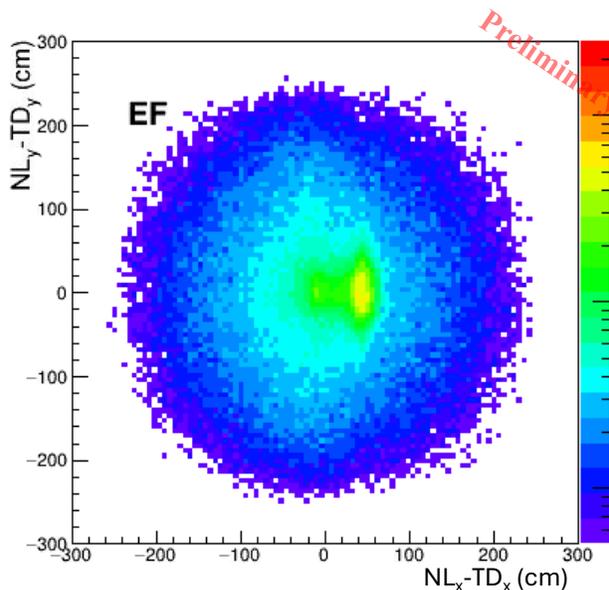
Quantify various sources of background for corrections of reaction plane and flow

Au+Au@400 AMeV  
(physics run)

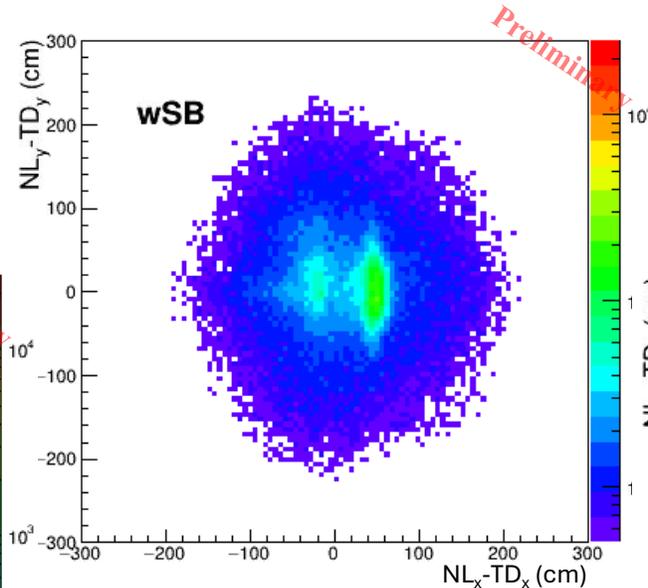
Au+Au@400 AMeV  
(w/ Shadow Bars)



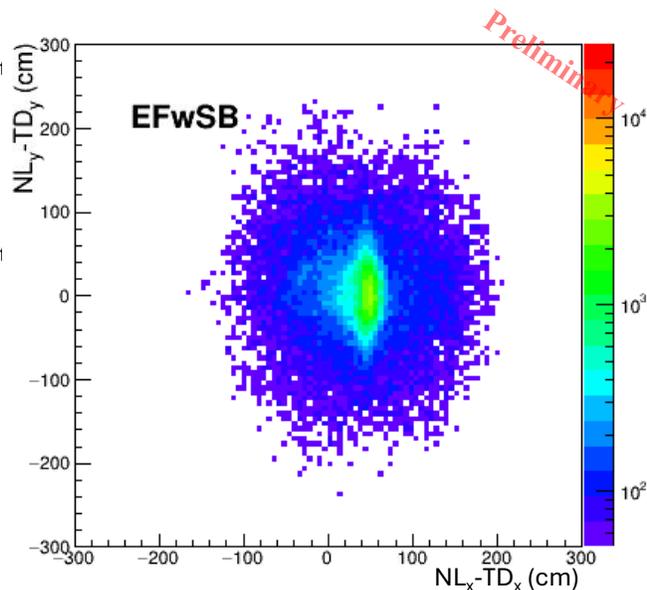
Au+EmptyFrame@  
400 AMeV



wSB



Au+EF@400 AMeV  
(w/ Shadow Bars)



# Conclusions

## Experiment

- S122 completed (March 2025, ~16 days)
- Au+Au @ 280, 400, 600, 1000 AMeV
- Stable beam, high statistics

## Analysis status

- CHIMERA & KRAB: advanced data reduction
- Reaction plane determination seems robust
- Preliminary impact parameter determination: (comparison with simulations needed)

## Detector work

- ToFD calibration ongoing
- NeuLAND + veto refinement in progress
- Joint n/p flow extraction requires careful tuning

## Next steps

- Optimizing data analysis ongoing
- Background evaluation and subtraction
- Setting up a theoretical taskforce (various transport models, TMEP)
- Consistent experiment–theory comparison

- This project has received funding from the European Union's Horizon Europe Research and Innovation programme under Grant Agreement No 101057511
- This project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 824093 (STRONG2020)

# ***Thank you for your attention!***



**This was made possible by a large collaboration from many countries:**

**R<sup>3</sup>B – GSI (Germany)**

**CHIMERA (Italy)**

**IFJ-PAN Krakow (Poland)**

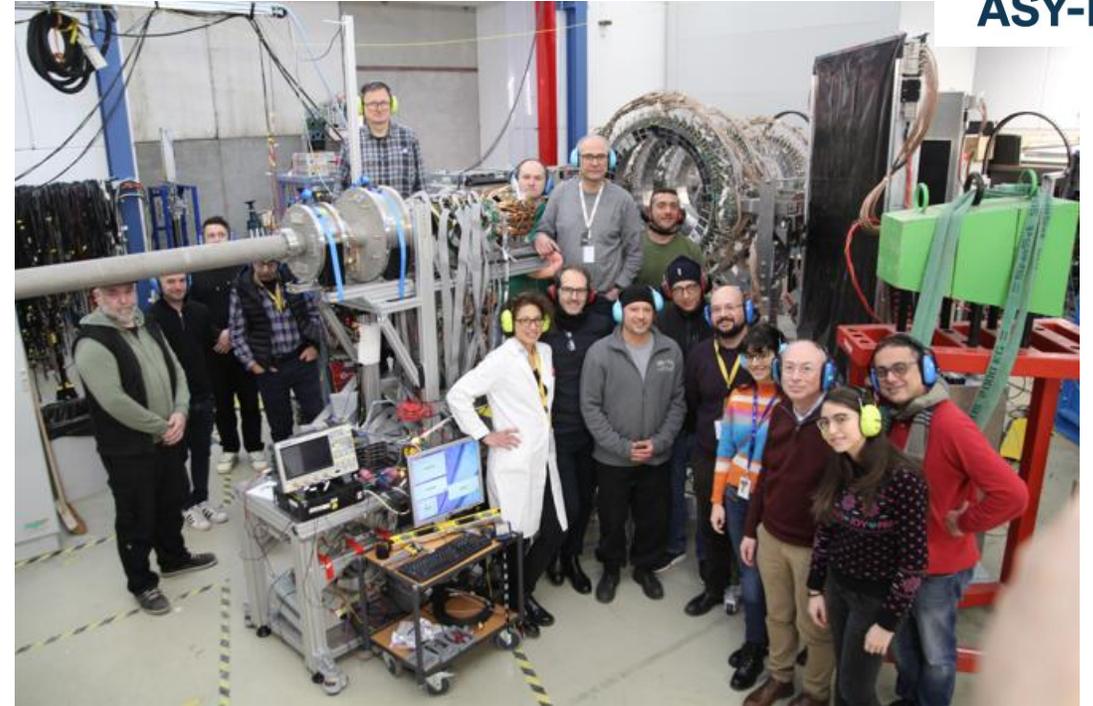
**INDRA-FAZIA (France-Italy)**

**MSU-FRIB (USA)**

**Texas A&M (USA)**

**GANIL (France)**

**RIKEN (Japan)**



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**J. Łukasik (IFJ-PAN Krakow, Poland)**

**A. Le Fèvre (GSI Darmstadt, Germany)**