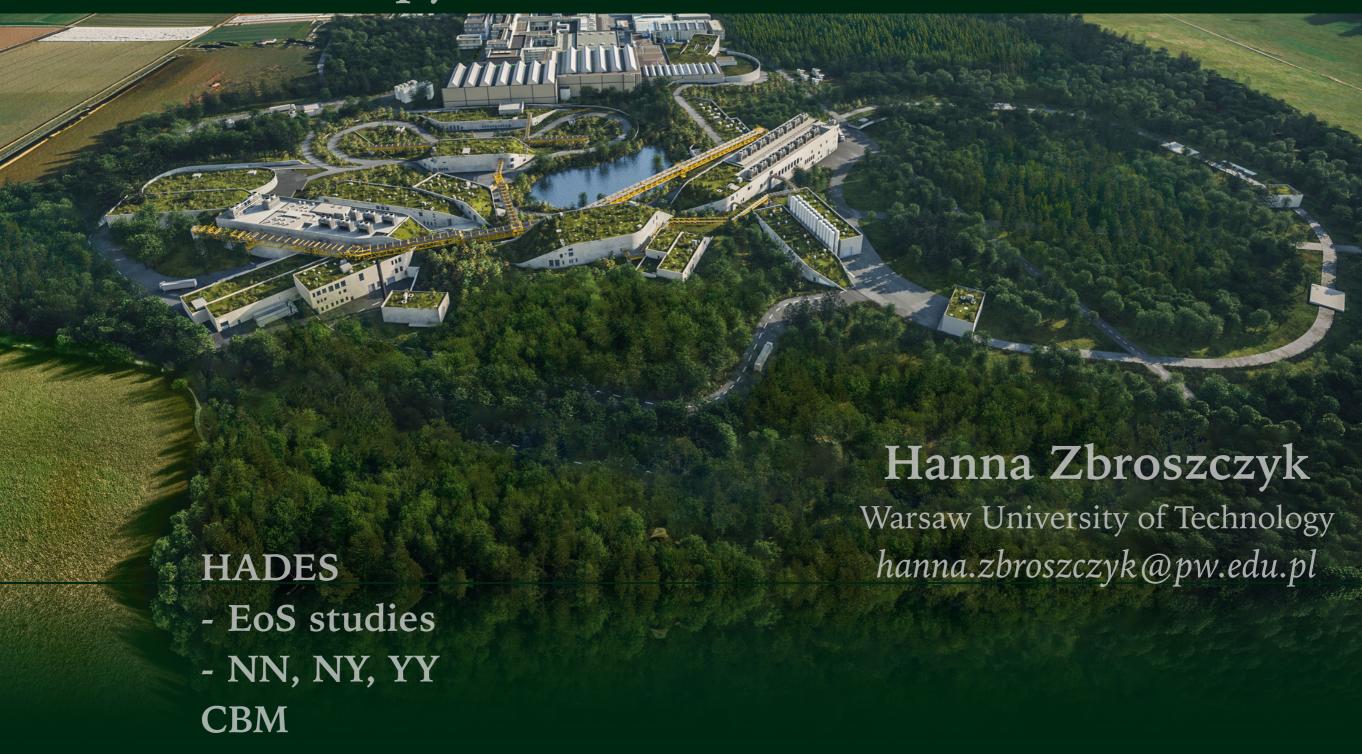


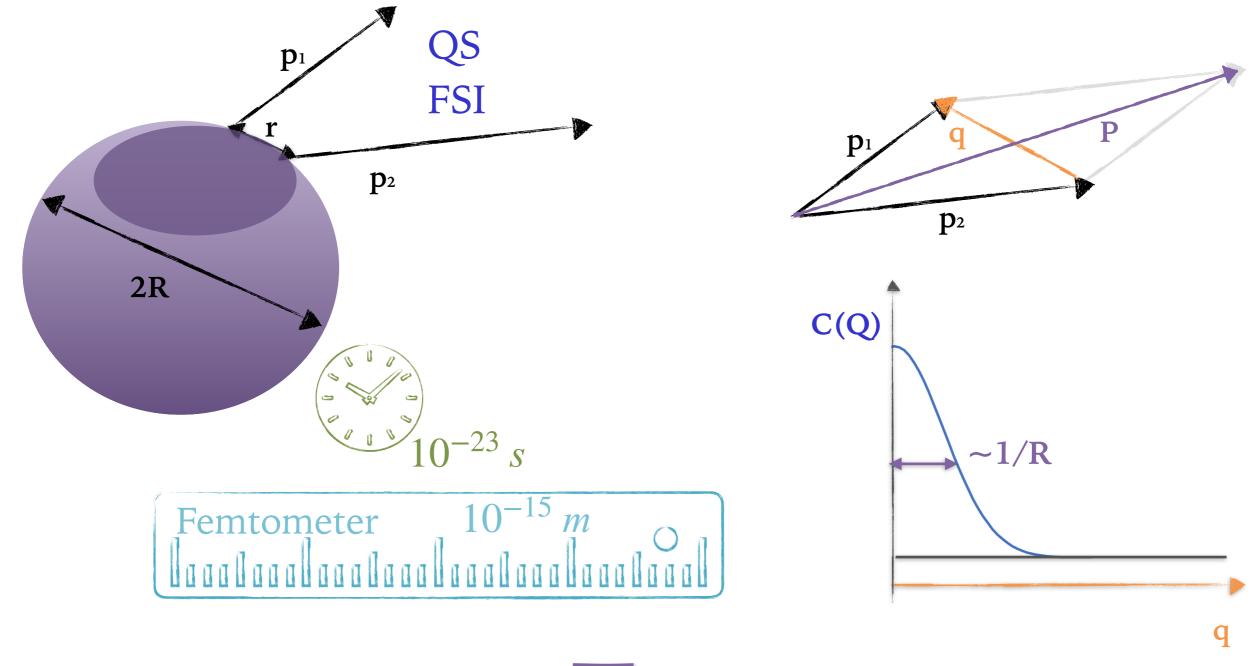




Femtoscopy and EOS at HADES and CBM



EMMI Workshop: Collective phenomena and the EOS of dense baryonic matter, 10-13th November 2025, GSI



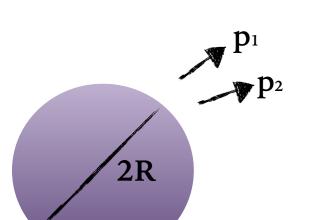
Femtoscopy

... the method to probe **geometric** and **dynamic** properties of the source (emission region, range of correlations-interactions, phase-space cloud, ...)

Femtoscopy does not measure the whole source, but homogeneity length.

Classic femtoscopy

Femtoscopy (originating from HBT): the method to probe **geometric** and **dynamic** properties of the source



Space-time properties $(10^{-15}m, 10^{-23}s)$ determined thanks to two-particle correlations:

Quantum Statistics (Fermi-Dirac, Bose-Einstein);

Final State Interactions (Coulomb, strong)

determined assumed measured
$$C(k^*, r^*) = \int S(r^*) |\Psi(k^*, r^*)|^2 d^3 r^* = \frac{Sgnl(k^*)}{Bckg(k^*)}$$

$$S(r^*)$$
 - source function

k* - momentum of the first particle inthe Pair Rest Frame reference

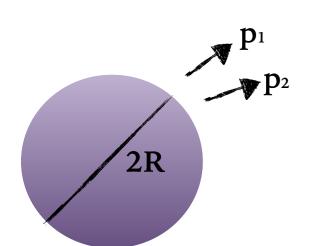


 $\Psi(k^*, r^*)$ - two-particle wave function (includes e.g. FSI interactions)

$$\frac{Sgnl(k^*)}{Bckg(k^*)}$$
 - correlation function

Gateway to study interactions

If we assume we know the **source function**, measured **correlations** are used to determine **interactions** in the final state.



Space-time properties $(10^{-15}m, 10^{-23}s)$ determined thanks to two-particle correlations:

Quantum Statistics (Fermi-Dirac, Bose-Einstein); Final State Interactions (Coulomb, strong)

assumed determined measured
$$C(k^*, r^*) = \int S(r^*) |\Psi(k^*, r^*)|^2 d^3r = \frac{Sgnl(k^*)}{Bckg(k^*)}$$

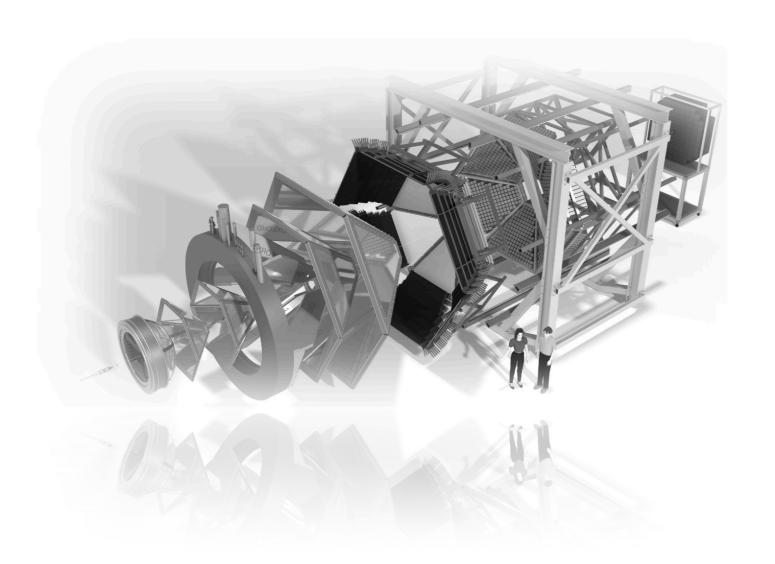
$$S(r^*) - \text{source function}$$

*k** - momentum of the first particle in the Pair Rest Frame reference



 $\Psi(k^*, r^*)$ - two-particle wave function (includes e.g. FSI interactions)

$$\frac{Sgnl(k^*)}{Bckg(k^*)}$$
 - correlation function

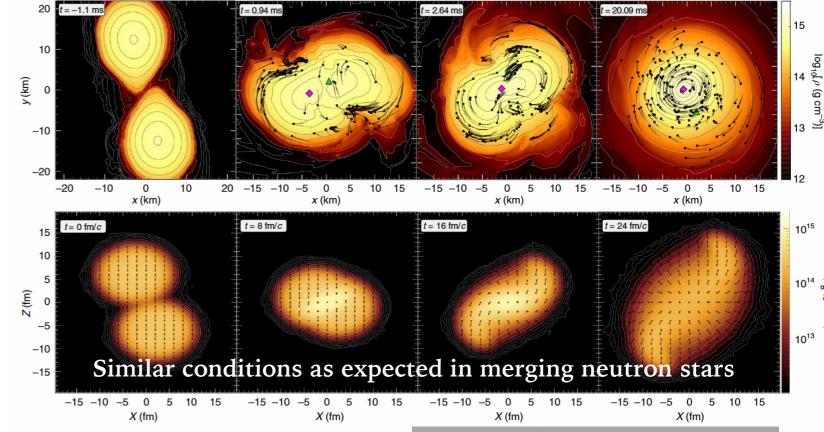


HADES

HADES physics goals

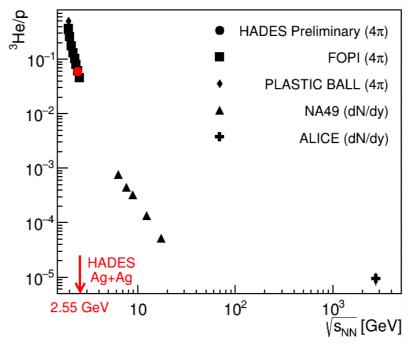
Heavy-ion collisions at $\sqrt{s_{NN}}$ up to 2.7 GeV

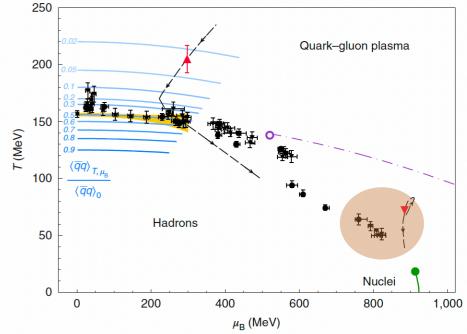
- Microscopic properties of baryon dominated matter
- EoS observables



 π^- (\sqrt{s} up to 2.35 GeV) and $\frac{2}{9}$ nucleon (\sqrt{s} up to 3.46 GeV) beams:

- Reference measurements (vacuum, cold QCD matter)
- Electromagnetic structure of baryons and hyperons





HADES, Nature Phys. 15, 1040–1045 (2019)

High Acceptance Di-Electron Spectrometer

Fixed target experiment at SIS-18 accelerator (GSI, Germany)

Magnet spectrometer

Low mass Mini-Drift-Chambers (MDCs)

Time of flight walls: RPC and TOF

RICH and ECAL for e^+/e^- and photon identification

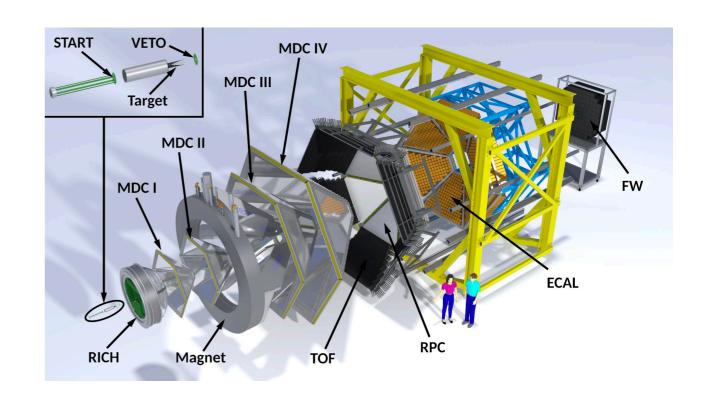
Full azimuthal angle and polar angles between 18° and 85° covered

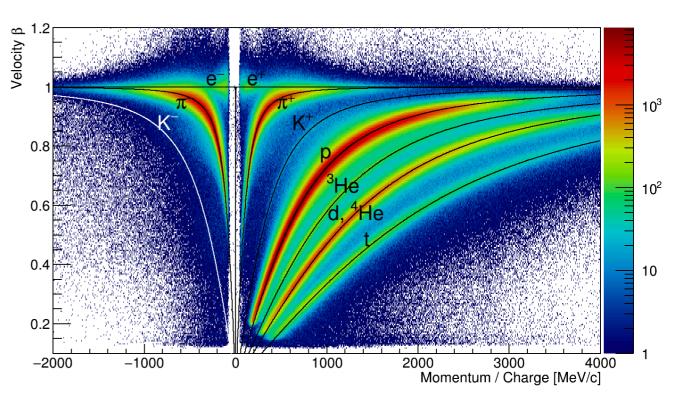
2012: Au+Au,
$$\sqrt{s_{NN}} = 2.42 \text{ GeV}$$
 (7 billion)

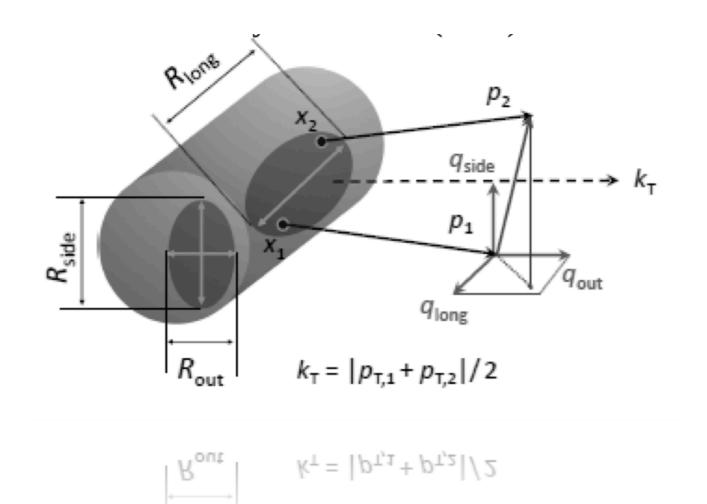
2019: Ag+Ag,
$$\sqrt{s_{NN}} = 2.55 \text{ GeV}$$

and 2.42 GeV (14 billion)

2024: Au+Au,
$$\sqrt{s_{NN}} = 2.24 \text{ GeV}$$
 (1.8 billion)







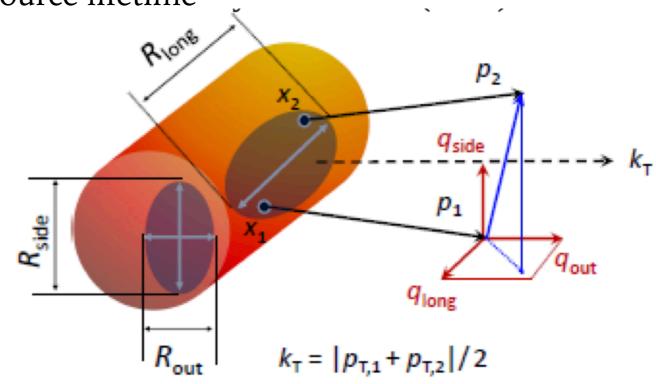
Sizes and dynamics

Bertsch-Pratt parametrization, 3D- and 1D-dimensional cases

- \rightarrow R_{side} spatial source evolution in the transverse direction
- \rightarrow R_{out} related to spatial and time components
- $\rightarrow R_{out}/R_{side}$ signature of phase transition
- \rightarrow R_{out}²- R_{side}² = $\Delta \tau^2 \beta_t^2$; $\Delta \tau$ emission time
- \rightarrow R_{long} temperature of kinetic freeze-out and source lifetime

long - determined by the beam directionout - determined by the pair transverse momentumside - perpendicular to long and side

3D case is considered if statistics is enough and two-particle correlations are easy to describe (Quantum Statistics and Coulomb FSI).

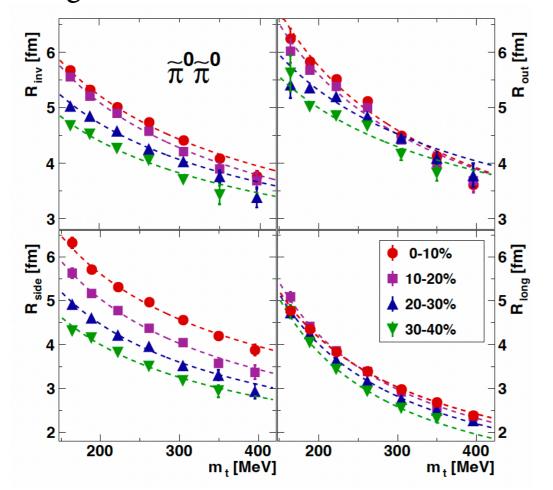


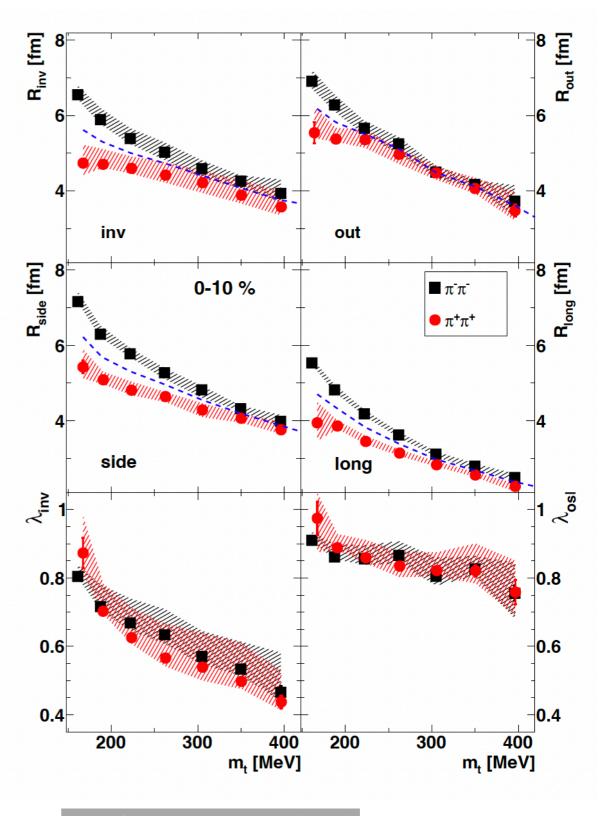
$$C(\vec{q}) = (1 - \lambda) + K_{\text{Coul}}(q_{\text{inv}})\lambda$$

$$\times \exp\left(-q_o^2 R_o^2 - q_s^2 R_s^2 - q_l^2 R_l^2 - 2q_o q_s R_{os}^2 - 2q_o q_l R_{ol}^2\right)$$

Identical pion correlations

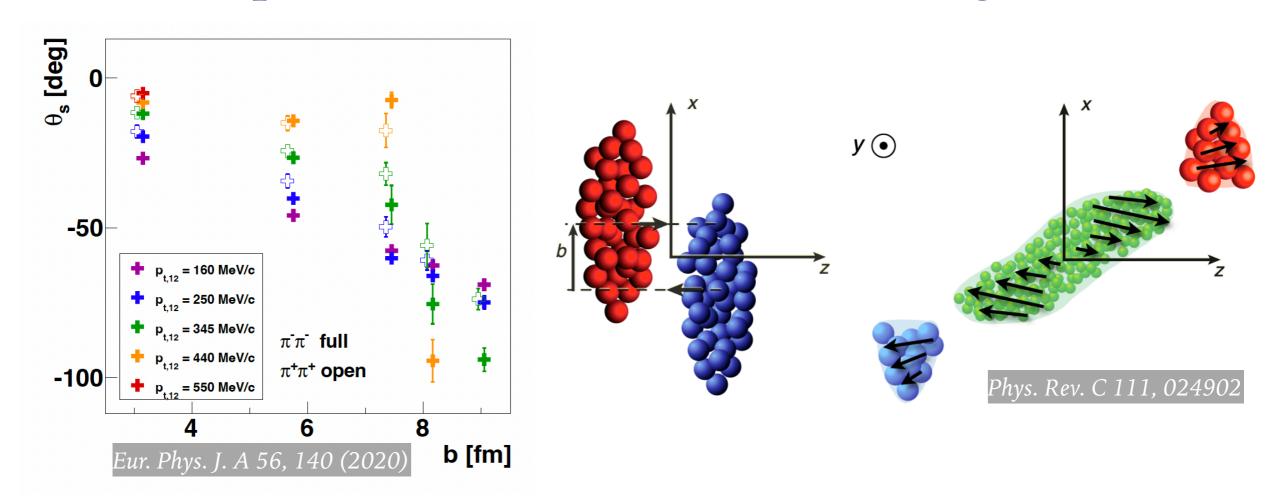
- 10% of most central collisions.
- 7 k_T intervals.
- Differences in HBT parameters for π^- and π^+ (especially low p_T) due to Coulomb effect.
- Neutral π deduced from interpolations of the charged π data.





Eur. Phys. J. A 56, 140 (2020)

Identical pion correlations - the tilt angle

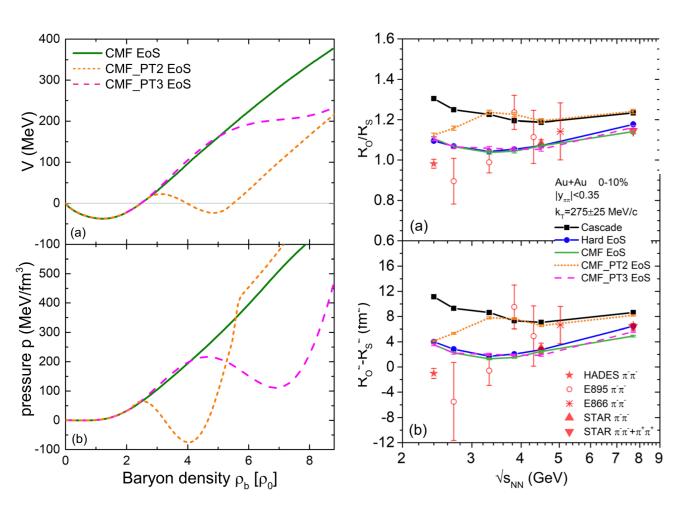


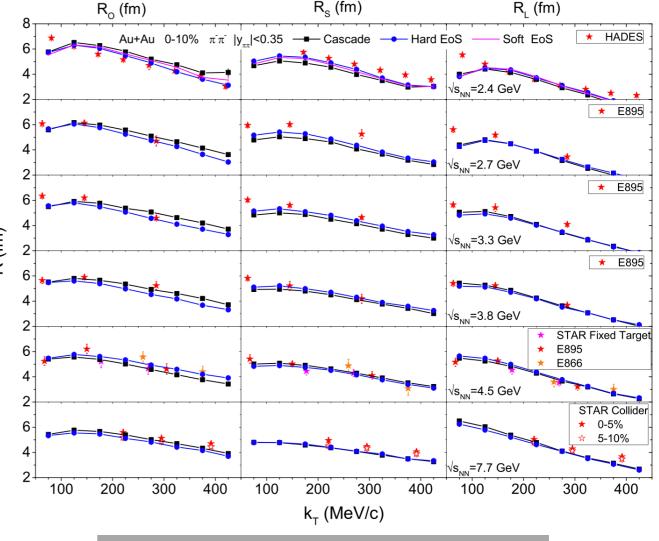
- Directed flow is attributed to the tilted source (from forward-backward asymmetry of the initial geometry).
- Tilt angle of the source describes the azimuthal distribution of the particles.

$$\theta_{out-long} = \frac{1}{2} tan^{-1} \left(\frac{4R_{out-long,1}^2}{R_{long,0}^2 - R_{side,0}^2 + 2R_{side,2}^2} \right) \qquad \theta_{side-long} = \frac{1}{2} tan^{-1} \left(\frac{4R_{side-long,1}^2}{R_{long,0}^2 - R_{side,0}^2 + 2R_{side,2}^2} \right)$$

Sensitivity of the HBT to the EoS with UrQMD

- HBT parameters sensitive to the stiffness of the EoS, can be used to constrain and understand the QCD EoS.
- PT with a significant softening of the EoS below 4 times nuclear saturation density can be excluded.

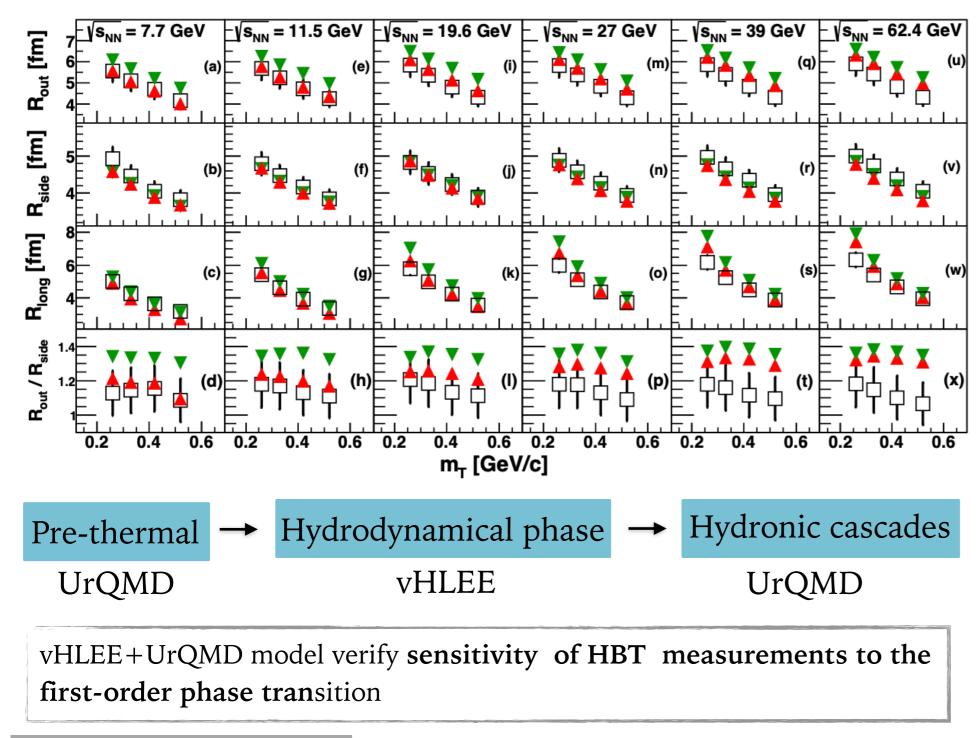




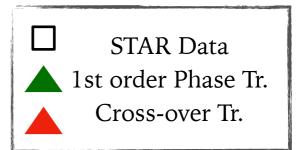
Sci. China-Phys. Mech. Astron. 66, 232011 (2023)

EMMI Workshop: Collective phenomena and the EOS of dense baryonic matter, 10-13th November 2025, GSI

How to measure phase transition?



Phys. Rev. C 96 (2017) no.2, 024911



vHLLE (3+1)-D viscous hydrodynamics: Iu. Karpenko, P. Huovinen, H. Petersen, M. Bleicher; Phys.Rev. C 91, 064901 (2015), arXiv:1502.01978, 1509.3751

HadronGas + Bag Model → 1^{St} order PT; P.F. Kolb, et al, PR C 62, 054909 (2000)

Chiral EoS → crossover PT (XPT); J. Steinheimer, et al, J. Phys. G 38, 035001 (2011)

3D / asHBT to be studied for p-p correlations



NN, NY interactions

Neutron star puzzle

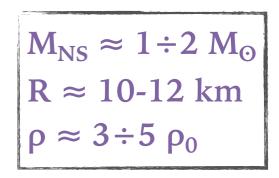
- Hyperons: expected in the core of neutron stars; conversion of N into Y energetically favorable.
- Appearance of Y: The relieve of Fermi pressure → softer
 EoS → mass reduction (incompatible with observation).

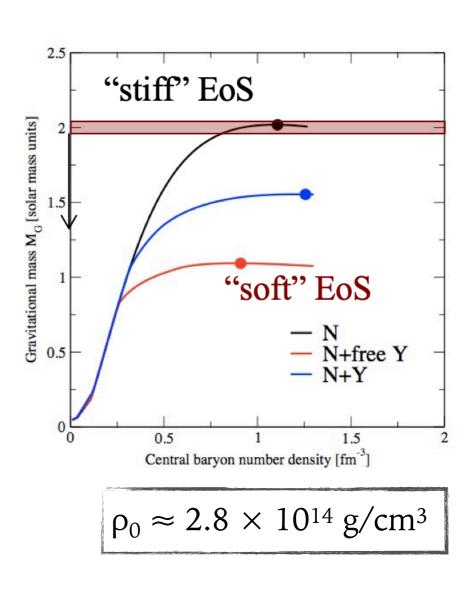
The solution requires a mechanism that could provide the **additional pressure** at high densities needed to make the EoS stiffer.

A few possible mechanisms, one of them:

Two- (and more) body YN, YY, NNY, NYY, ... interactions.

The existence of **hypernuclei** (confirmed by attractive YN interaction) → indicates the possibility to bind Y to N. The measurement of the YN and YY interactions leads to important implications for the possible formation of **YN, YY, NNY, NYY, ... bound states**.





Lednicky-Lyuboshitz model

The correlation function can be calculated analytically by averaging Ψ over the total spin S and the distribution of the relative distances $S(r^*)$

Ref: Lednicky, Richard & Lyuboshits, V.L.. (1982). Sov. J. Nucl. Phys. (Engl. Transl.); (United States). 35:5.

$$C(k^*) = \int S(r^*) |\Psi(r^*, k^*)|^2 d^3r$$

The normalized pair separation distribution (source function) $S(r^*)$ is assumed to be Gaussian,

$$S(r^*) = (2\sqrt{\pi}r_0)^{-3}e^{-rac{r^{*2}}{4r_0^2}},$$

$$\Psi^{S}(r^*,k^*) = e^{-ik^*r^*} + f^{S}(k^*) \frac{e^{ik^*r^*}}{r^*}$$

$$f^{S}(k^{*}) = (\frac{1}{f_{0}^{S}} + \frac{1}{2}d_{0}^{S}k^{*2} - ik^{*})^{-1}$$
 Strong

$$|\Psi^{C}(r*,k*)| = \sqrt{A_{C}}e^{-ik^{*}r^{*}}F(-i\eta,1,i\zeta)$$

$$A_C(\eta) = \frac{2\pi}{k * a_c} (exp(\pm \frac{2\pi}{k * a_c}) - 1)^{-1}$$
 Coulomb

F- confluent hypergeometric function

 f_0 and d_0 - parameters of strong interaction.

Theoretical correlation function (k*) depends on: R, f₀ and d₀.

f₀ - the scattering length, determines low-energy scattering.

The elastic cross section, σ_e , (at low energies) determined by

the scattering length, $\lim_{k\to 0} \sigma_e = 4\pi f_0^2$

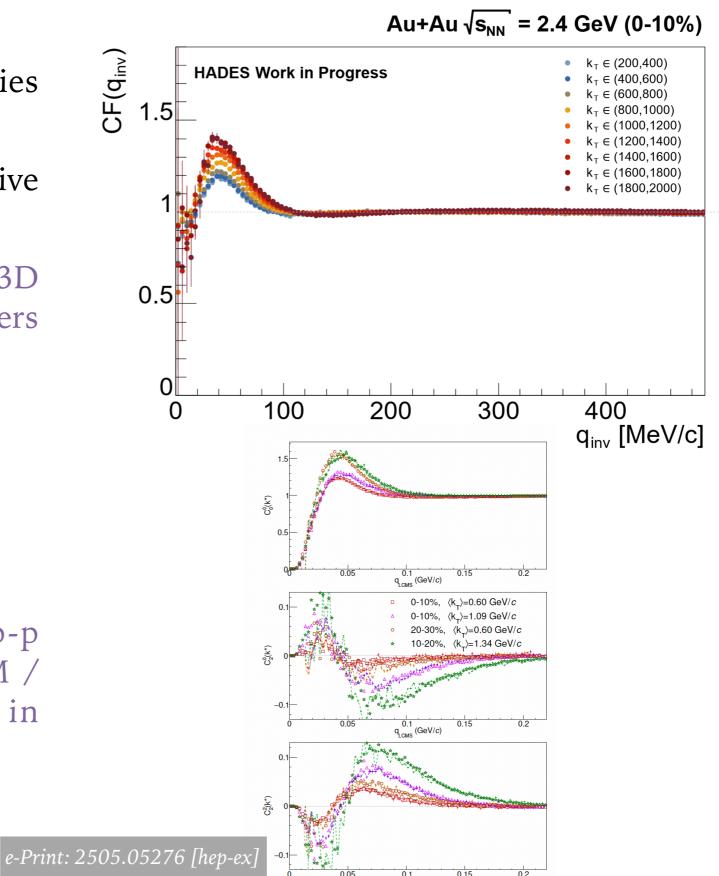
 d_0 - the effective range, corresponds to the range of the potential (simplified scenario - the square well potential.

For identical systems one has to include QS (Fermi-Dirac / Bose-Einstein) as well.

Proton - proton correlations

- High stopping in HADES implies large abundance of protons
- Protons expected to be more sensitive to the EOS
- So far only 1D analysis done, 3D analysis coming soon, BP parameters from protons to compare to pions

- Plan to use SH decompositions $C_l^m(k^*) = \int C(\overrightarrow{k^*}) Y_l^m(\theta_k, \phi_k) d\cos(\theta) d\phi_k$
- SMASH predictions for 3D p-p correlations for HADES (and CBM / STAR FXT) for various EoS in progress

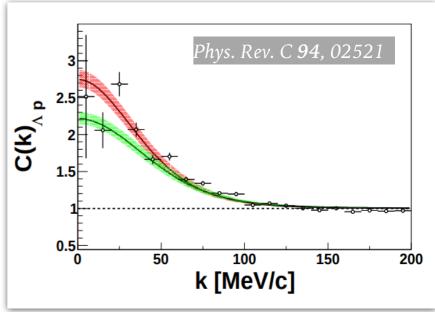


Proton - lambda correlations

Model		$f_0^{S=0}$ (fm)	$f_0^{S=1}$ (fm)	$d_0^{S=0}$ (fm)	$d_0^{S=1}$ (fm)	n_{σ}
ND [77]		1.77	2.06	3.78	3.18	1.1
NF [78]		2.18	1.93	3.19	3.358	1.1
NSC89 [79]		2.73	1.48	2.87	3.04	0.9
NSC97 [80]	a	0.71	2.18	5.86	2.76	1.0
	b	0.9	2.13	4.92	2.84	1.0
	С	1.2	2.08	4.11	2.92	1.0
	d	1.71	1.95	3.46	3.08	1.0
	e	2.1	1.86	3.19	3.19	1.1
	f	2.51	1.75	3.03	3.32	1.0
ESC08 [81]		2.7	1.65	2.97	3.63	0.9
χEFT	LO [25]	1.91	1.23	1.4	2.13	1.8
	NLO [26]	2.91	1.54	2.78	2.72	1.5
Jülich	A [82]	1.56	1.59	1.43	3.16	1.0
	J04 [83]	2.56	1.66	2.75	2.93	1.4
	J04c [83]	2.66	1.57	2.67	3.08	1.1

S. Acharya *et al.* Phys. Rev. C 99, 024001 – Published 13 Feb 2019 https://doi.org/10.1103/PhysRevC.99.024001

parameter scan boundaries : f_0 [0.01, 5.0], d_{0s} [0.01, 2.0] and d_{0t} [0.01, 5.0]

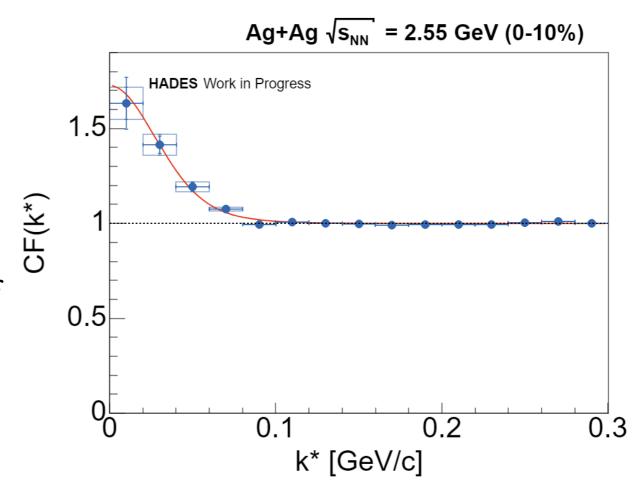


Parame- ters	p-Nb (LO)	p-Nb (NLO)
f_{0s}	1.91 fm	2.91 fm
d_{0s}	1.40 fm	2.78 fm
f_{0t}	1.23 fm	1.54fm
d _{0t}	2.13 fm	2.72fm
\mathbf{r}_{0}	1.71土 0.10	1.62土 0.02

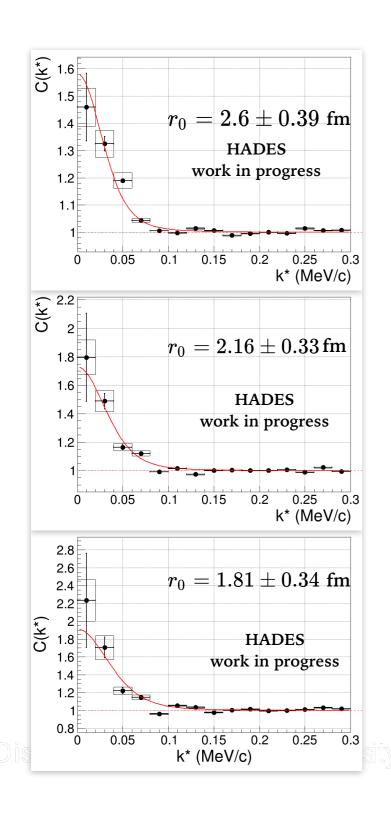
Proton - lambda correlations

- $p \Lambda$ correlations first step to study N-Y interactions in HIC
- Ag+Ag collisions at $\sqrt{s_{NN}} = 2.55$ GeV fitted using L-L parametrisation
- First spin separated measurements of strong interaction parameters of the $p-\Lambda$ interaction obtained from HIC:

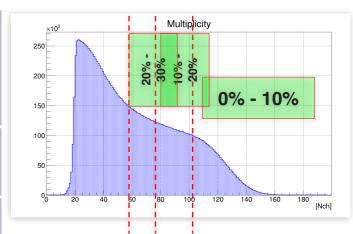
$$f_0 = 1.93^{+0.48}_{-0.42}$$
 $d_0 = 0.01$ for S=0
 $f_0 = 1.76^{0.14}_{-0.49}$ $d_0 = 3.44^{+0.36}_{-0.49}$ for S=1

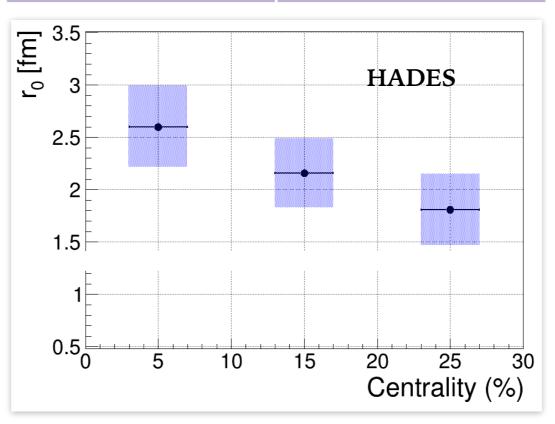


Proton - lambda correlations: centrality dependence

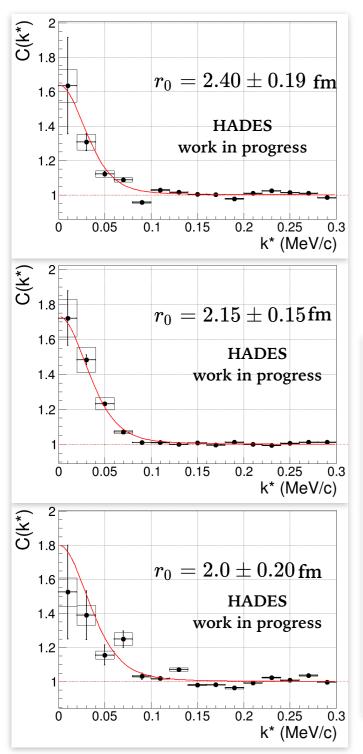


Centrality	Systematic Uncertainty
0 - 10 %	15.30 %
10 - 20 %	15.49 %
20 - 30 %	19.00 %

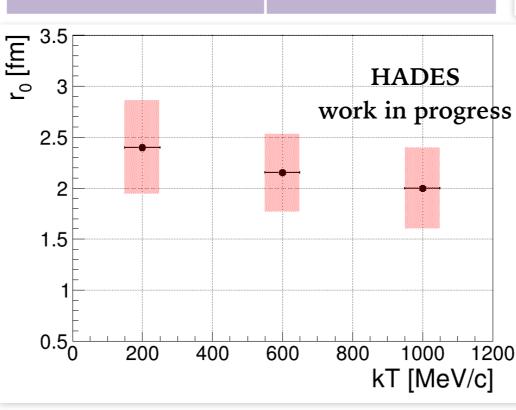


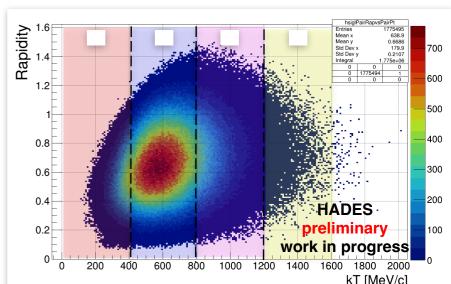


Proton - lambda correlations: k_T dependence



kт [GeV/c]	Systematic Uncertainty
0 - 400	19 %
400 - 800	15 %
800 - 1200	22 %





$$k_T = \frac{p_{T_1} + p_{T_2}}{2}$$

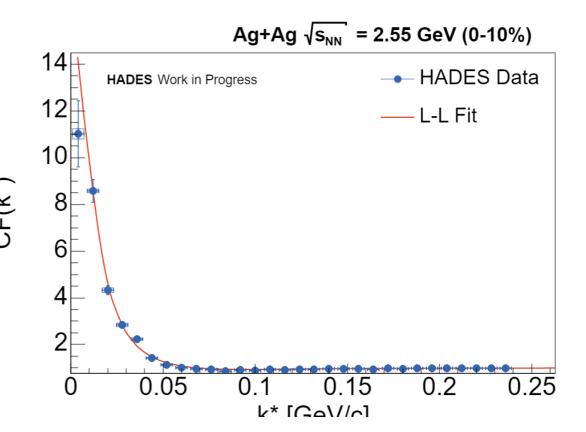
Deuteron - lambda correlations

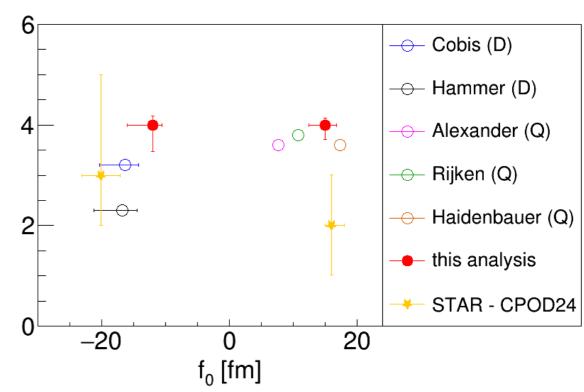
- $d \Lambda$ correlations next step to study N-Y interactions in HIC
- Insight to study production of hypertriton
- Ag+Ag collisions at $\sqrt{s_{NN}} = 2.55$ GeV fitted using L-L parametrisation
- First look at spin separated measurements of strong interaction parameters of the $d-\Lambda$ interaction obtained from HIC:

$$f_0 = -12^{+1.44}_{-3.92}$$
 $d_0 = 4^{+0.18}_{-0.53}$
 $f_0 = 15^{+1.73}_{-2.58}$ $d_0 = 4^{+0.13}_{-0.28}$

Quartet

Doublet ∰





Deuteron - lambda correlations

Expected centrality dependence:

$$R_{0-10\%} > R_{10-20\%} > R_{20-30\%}$$

• First look at spin separated measurements of strong interaction parameters of the $d-\Lambda$ interaction obtained from HIC:

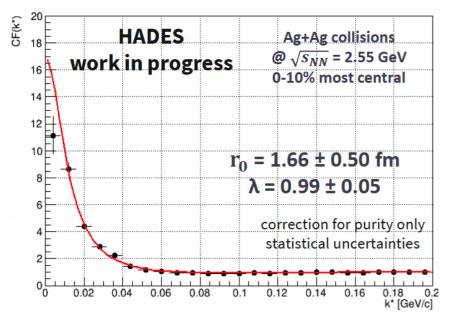
$$f_0 = -12^{+1.44}_{-3.92}$$
 $d_0 = 4^{+0.18}_{-0.53}$

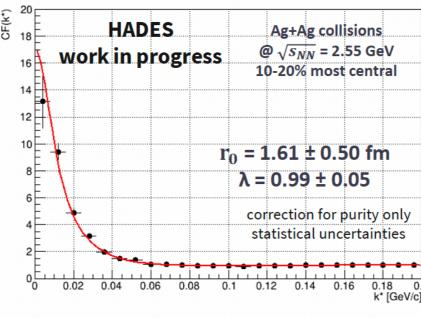
Doublet

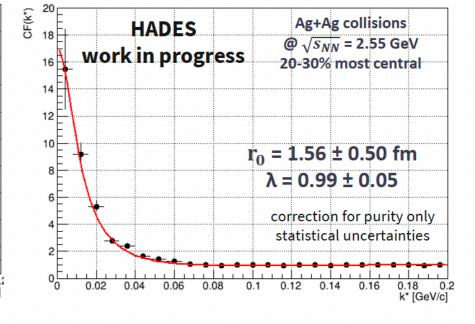
$$f_0 = 15^{+1.73}_{-2.58}$$
 $d_0 = 4^{+0.13}_{-0.28}$

$$d_0 = 4^{+0.13}_{-0.28}$$

Quartet







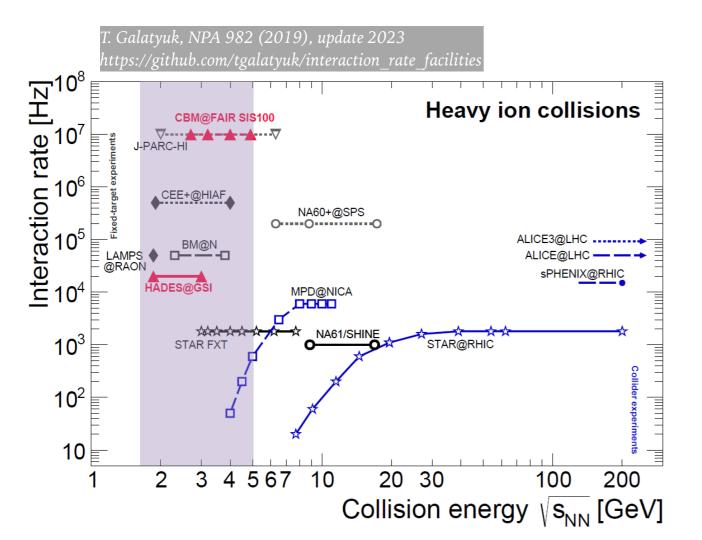


CBM

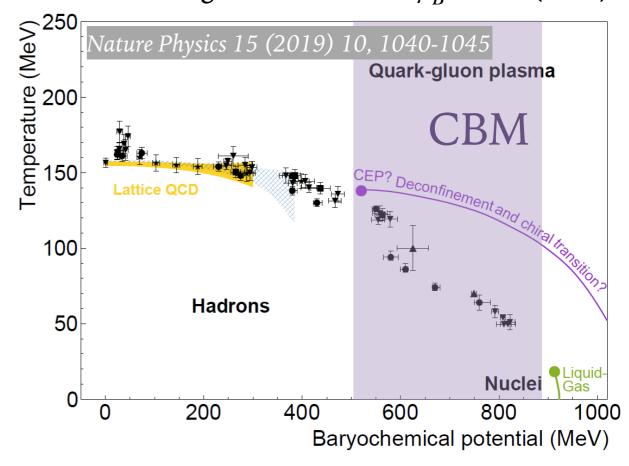
Compressed Baryonic Matter experiment

Key observables – systematic measurements!

- Fluctuations: System transition via first order phase transition line, Critical Point;
- Hadrons, Strangeness, Charm: System in equilibrium, Hypernuclei, Vorticity, Flow, EOS;
- **Dileptons**: Emissivity of dense baryonic matter: lifetime, temperature, density, in-medium properties;
- Correlations: Flow, Vorticity, YN, YNN, ... interactions and 3D studies possible (femtoscopy).



Experimental investigation of the region of the $500 < \mu_B < 850$ (MeV)



Summary

Wrap -up

Femtoscopy is sensitive to different EoS

Depending on the available statistics one can focus on 1D- or 3D studies

It is possible to use traditional HBT technique to deduce 3D HBT parameters

One can determine parameters of interactions, search for possible bound states and look for an evidence to make the EoS softer

HADES has already measured many systems

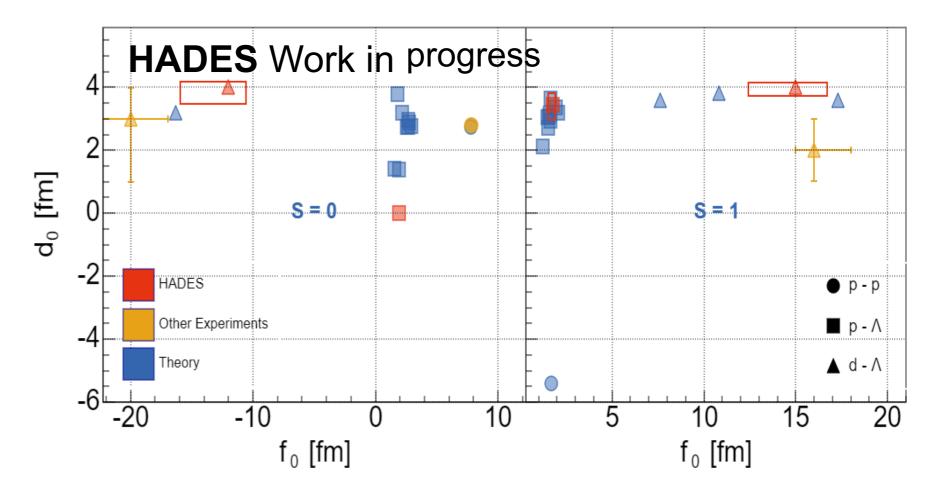
Future CBM (and HADES) will use higher statistics

Thank you!

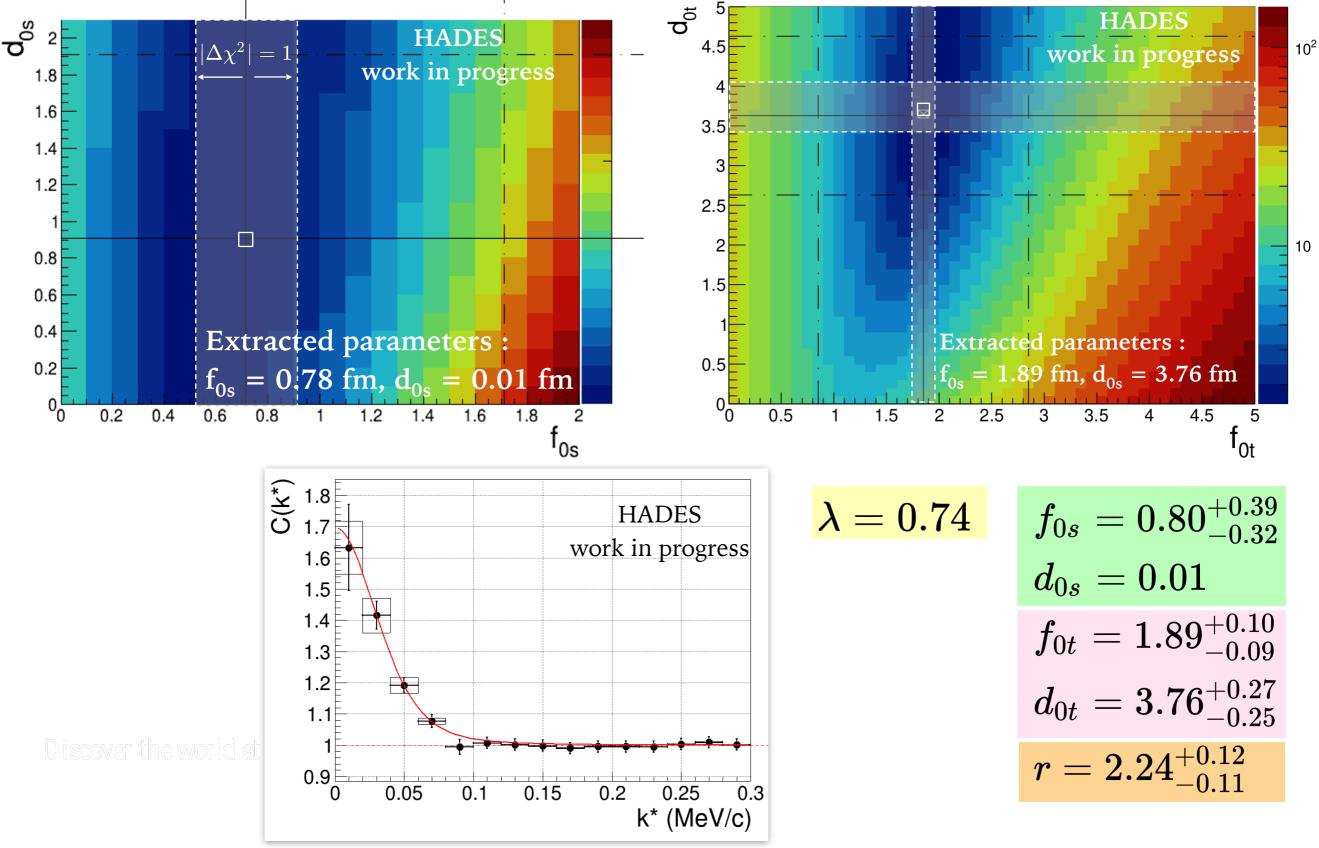
Supporting slides

FSI parameters

Clear separation of FSI

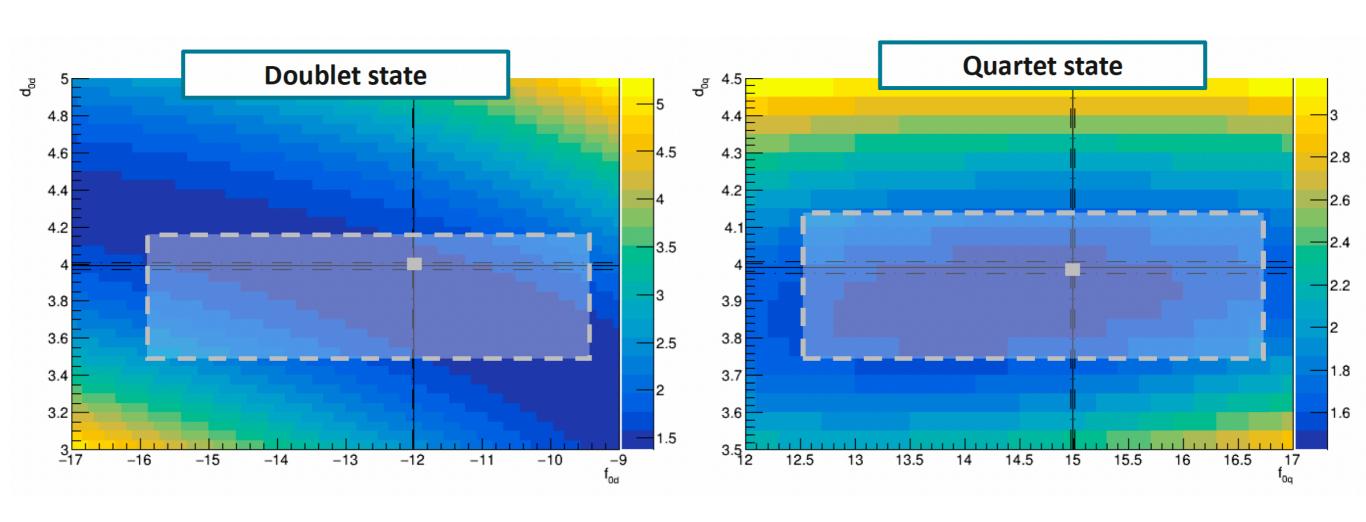


YN $(p\Lambda)$ correlations at HADES



YN $(d\Lambda)$ correlations at HADES

 $|\Delta\chi^2|=1$



Strange hadronic matter in the inner core

The inner core of the neutron star is totally unknown. One of the most probable scenarios is that hyperons (baryons with strange quarks) appear at a density larger than (2–3) ρ_0

 Λ hyperons, being free from Pauli exclusion principle by neutrons, are allowed to stay at the bottom of the attractive nuclear potential made by neutrons. When the kinetic energy of a neutron on the Fermi surface of the degenerate neutron matter exceeds the Λ -n mass difference of 176 MeV, it converts into a Λ hyperon via weak interaction.

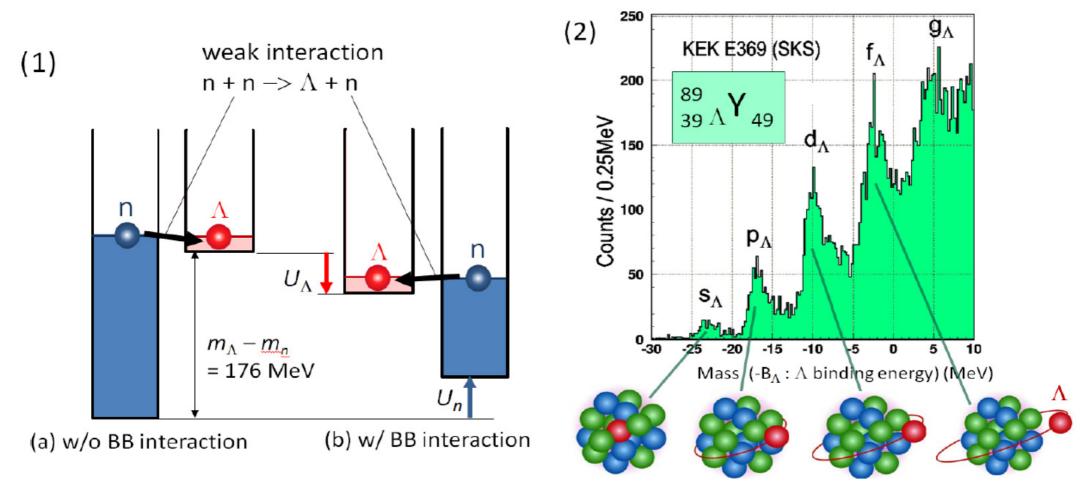
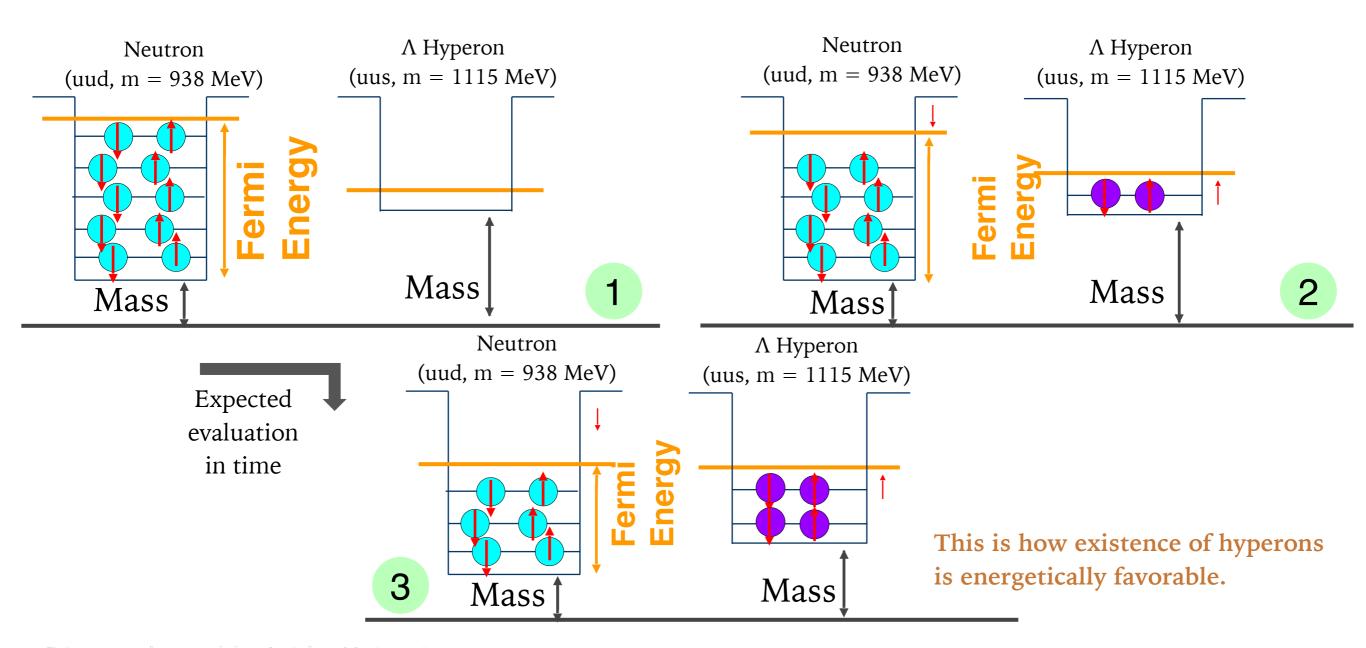
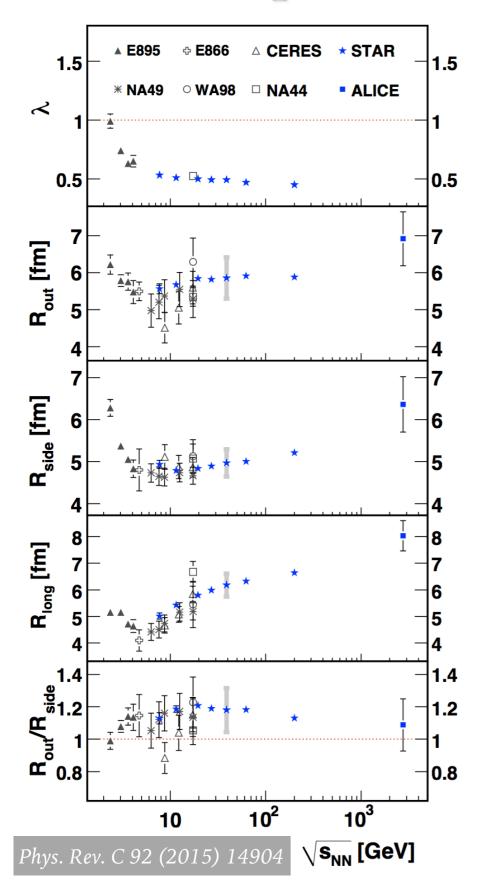


Fig. 3. (1) Energies of neutrons and Λ hyperons in high density neutron matter confined in the potential made by gravity. See text for details. (2) Excitation spectrum of a Λ hypernucleus $^{89}_{\Lambda}{\rm Y}$ via the (π^+, K^+) reaction on $^{89}{\rm Y}$ target [6].

Hyperon production



Identical pion femtoscopy at STAR



- \rightarrow R_{side} spatial source evolution in the transverse direction
- \rightarrow R_{out} related to spatial and time components
- \rightarrow R_{out}/R_{side} signature of phase transition
- \rightarrow R_{out}²- R_{side}² = $\Delta \tau^2 \beta_t^2$; $\Delta \tau$ emission time
- \rightarrow R_{long} temperature of kinetic freeze-out and source lifetime

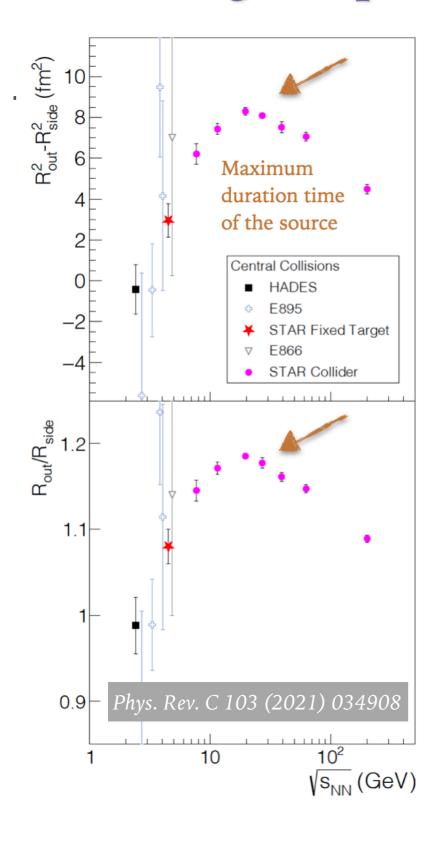
$$C(\vec{q}) = (1 - \lambda) + K_{\text{Coul}}(q_{\text{inv}})\lambda$$

$$\times \exp\left(-q_o^2 R_o^2 - q_s^2 R_s^2 - q_l^2 R_l^2 - 2q_o q_s R_{os}^2 - 2q_o q_l R_{ol}^2\right)$$

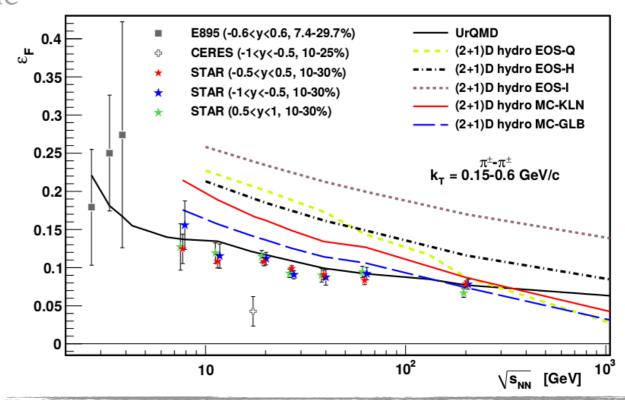
HBT source sizes determined for wide range of collision energy;

Non-monotonic behavior seen in three directions

Looking for phase transition at STAR



- \rightarrow R_{side} spatial source evolution in the transverse direction
- → R_{out} related to spatial and time components
- $\rightarrow R_{out}/R_{side}$ signature of phase transition
- ightarrow R_{out}^2 $R_{side}^2 = \Delta \tau^2 \ \beta_t^2$; $\Delta \tau$ emission time
- \rightarrow R_{long} temperature of kinetic freeze-out and source lifetime



System evolves faster in the reaction plane

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