



INSTITUTO DE FÍSICA  
Universidade Federal Fluminense



# Hydrodynamic model of heavy ion collisions

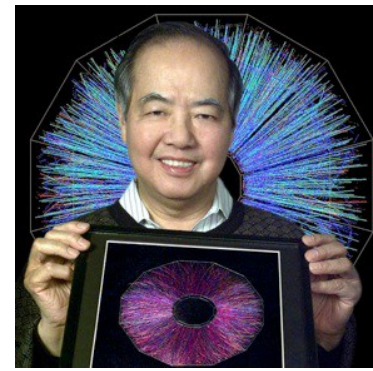
Gabriel S. Denicol  
Universidade Federal Fluminense

Erice: INTERNATIONAL SCHOOL OF NUCLEAR  
PHYSICS, 46th COURSE

# What you will see

- ✓ Motivation: why do we use hydrodynamic models?
- ✓ Why do we believe in fluid-dynamical models?
- ✓ Relativistic hydrodynamics: fundamentals and ingredients
- ✓ Magnetohydrodynamic simulations for heavy ion collisions
- ✓ Final remarks

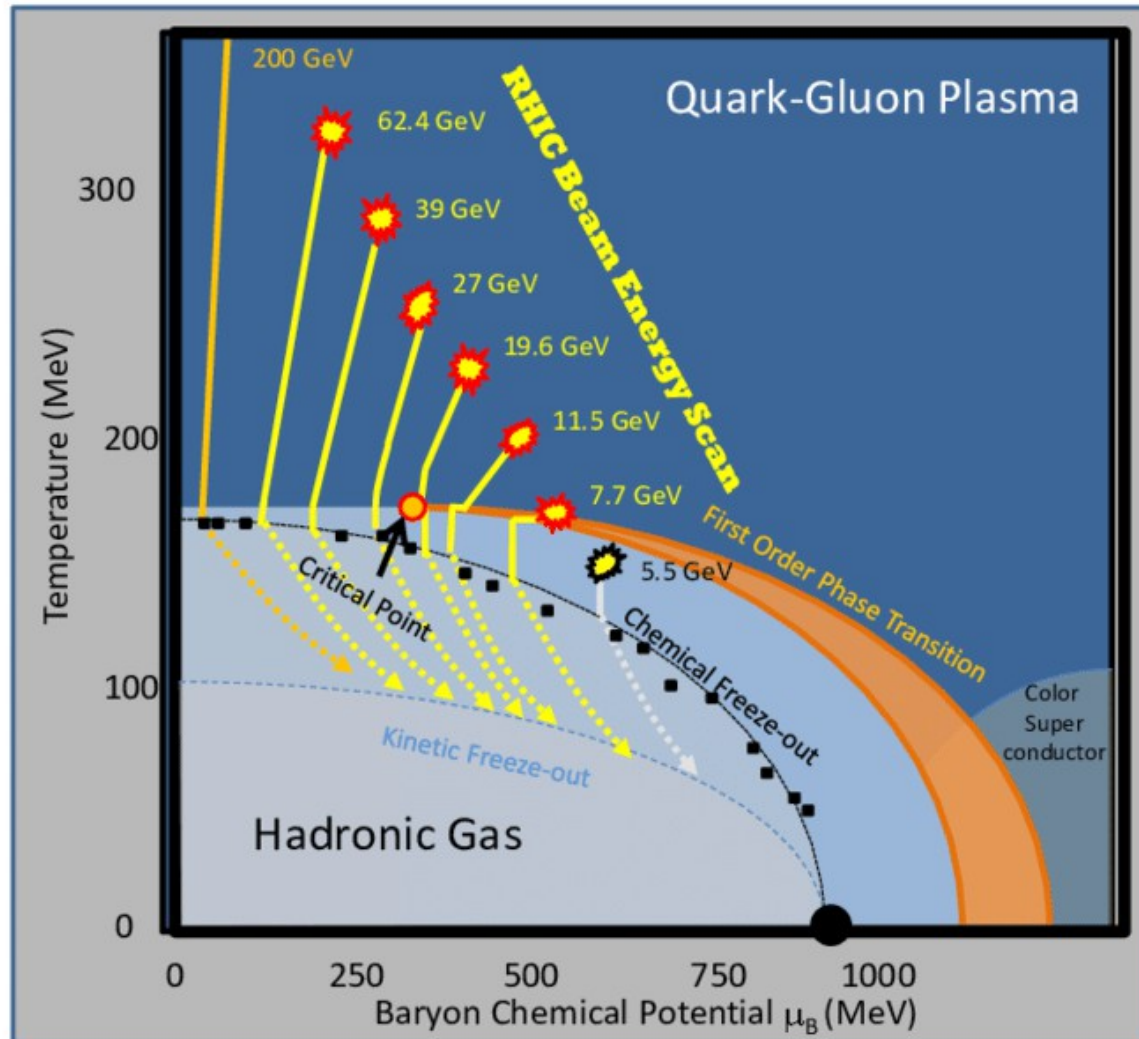
# Goal of Heavy-Ion Collisions: *Produce and study* QCD matter near (local) equilibrium



T.D. Lee

## Challenge

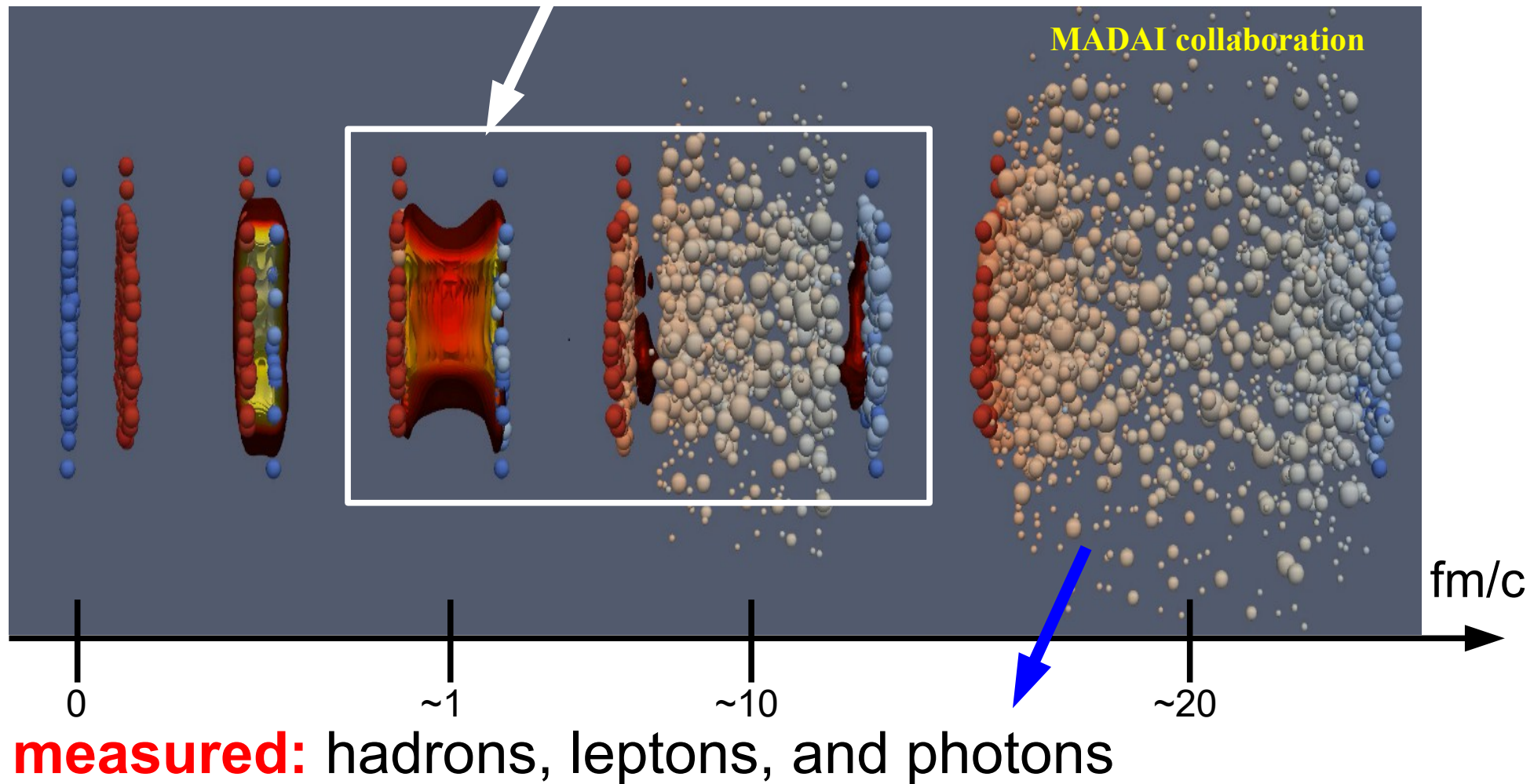
- Reach thermodynamic limit experimentally!
- Extract thermodynamic and transport properties of QCD matter
- Experiment is not to test QCD, but to understand it



# Heavy Ion Collisions

QCD matter is only created *transiently*  $\sim 10$  fm/c

**expectation:** *fluid-dynamical expansion*



*Properties of matter must be reverse-engineered!*

# Current theoretical description

## 1) Initial state and “pre-equilibrium” dynamics:

- description of early-time dynamics and “thermalization”
- initial condition for hydrodynamic evolution



(approach) thermalization (?)

## 2) Fluid-dynamical expansion of QGP and Hadron Gas

- Phase transition
- Matter described by EoS and transport coefficients  
shear and bulk viscosity, charge diffusion ...



fluid elements converted to particles

## 3) Transport description of Hadron Gas

- Late stage description using the *hadron resonance gas* model – using cross sections and decay probabilities



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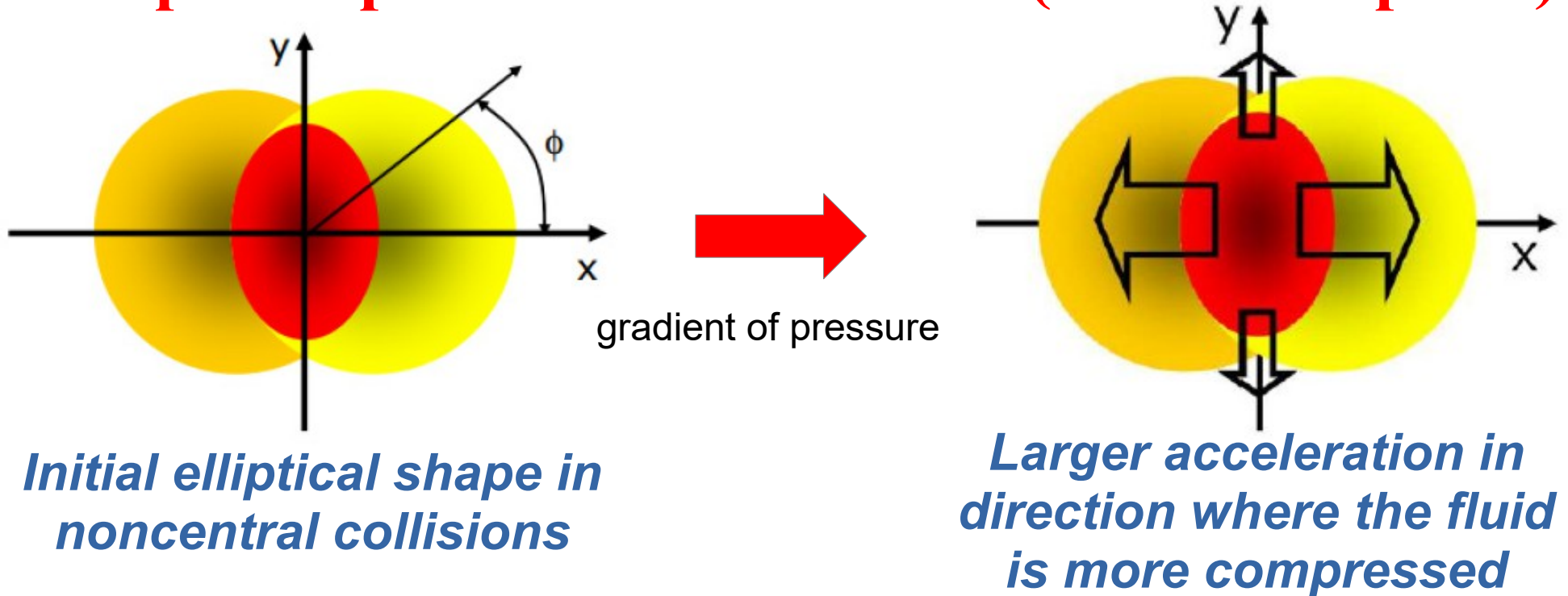
# Why do we think fluid-dynamical models work?





# Anisotropic flow

- Simplified picture of the collision (transverse plane)**

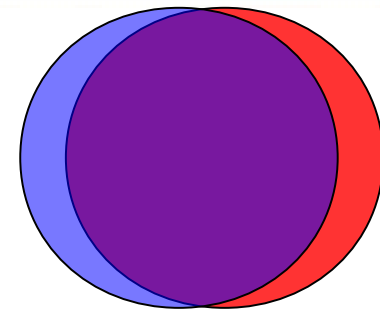
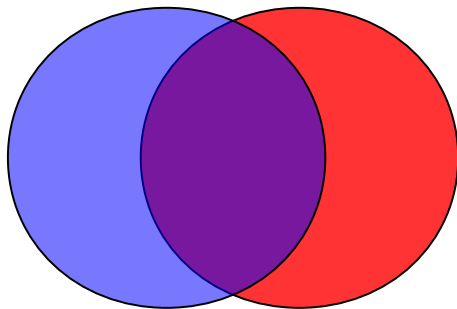


$$\frac{2\pi}{N} \frac{dN}{d\phi} = 1 + 2 \sum_{n=1}^{\infty} v_n \cos n(\phi - \Psi_n) \quad v_2 \text{ is large}$$

particles are not emitted isotropically!

**hydrodynamic response to the initial geometry**

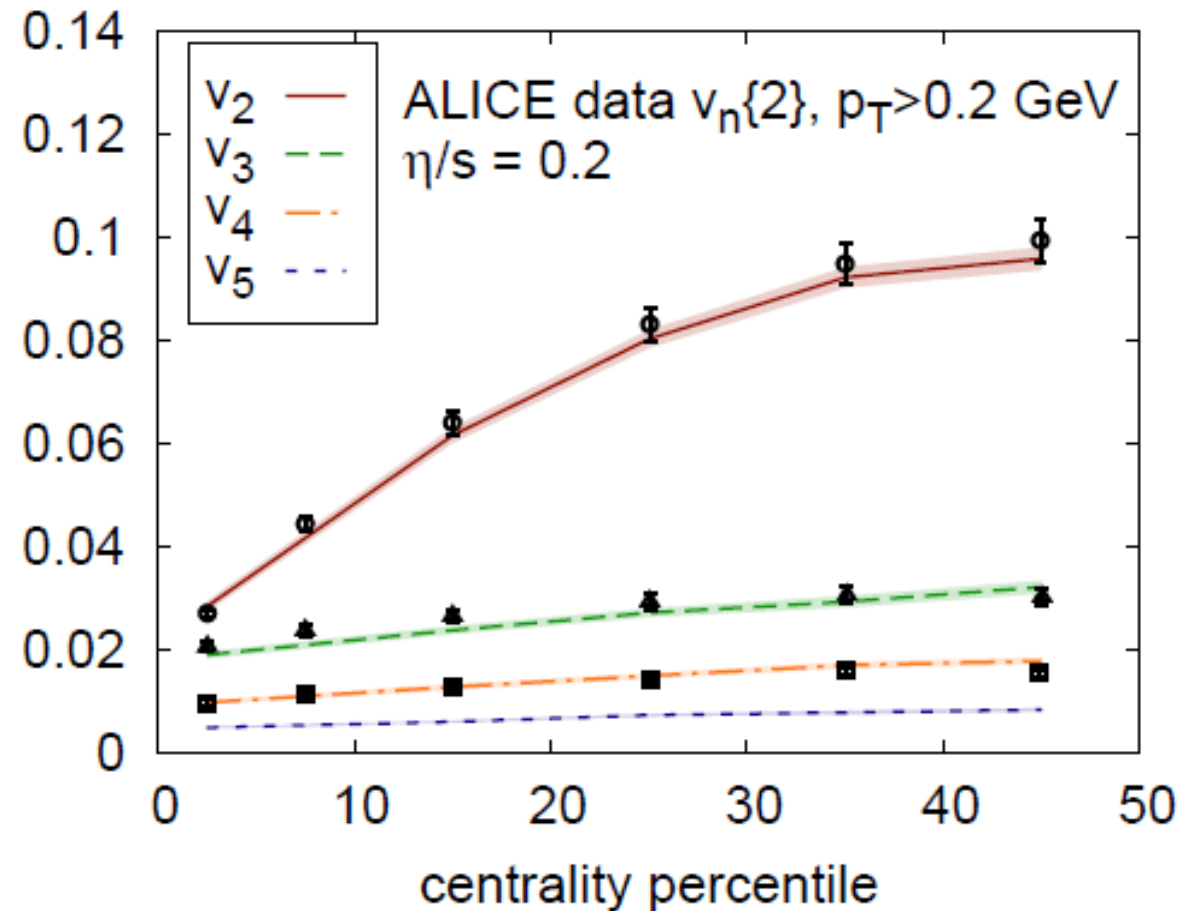
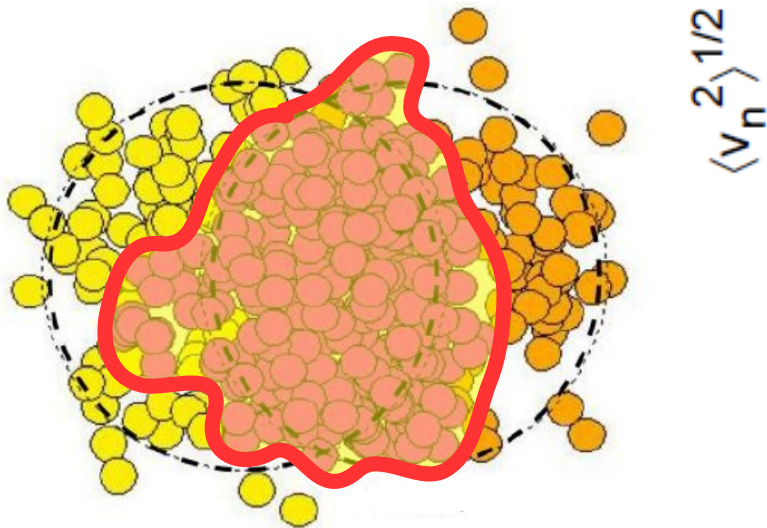
It is possible to “estimate” the centrality



# Hydrodynamic response to initial state can be observed experimentally

Gale *et al*, PRL 110 (2013) no.1, 012302

Quantum effects render  
the initial geometry more  
complicated

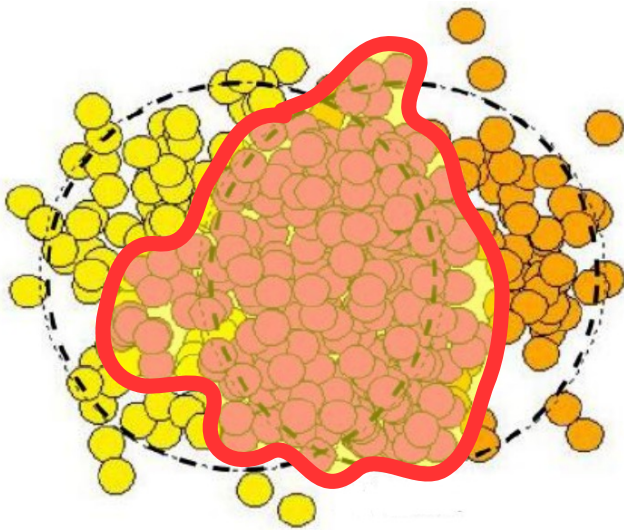


*non-trivial test of fluid-dynamical models*

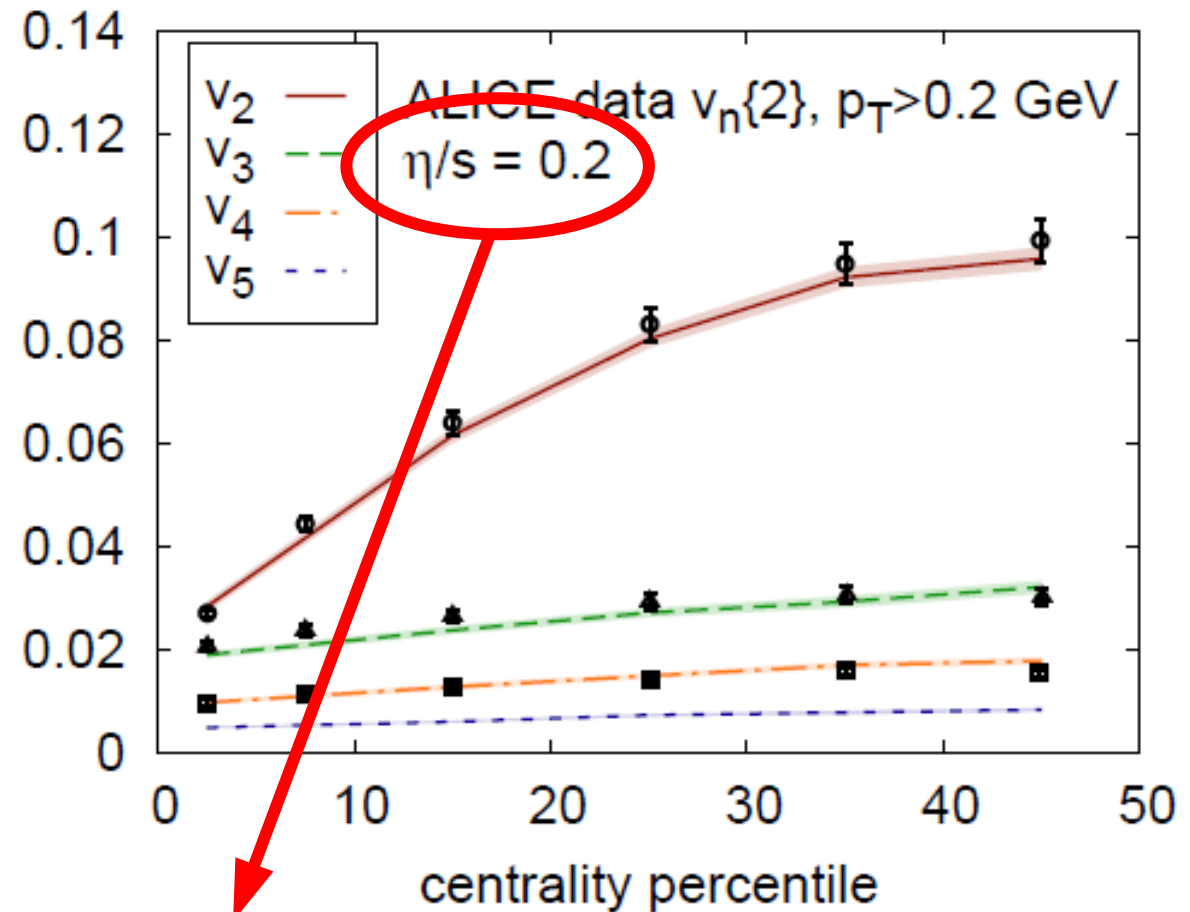
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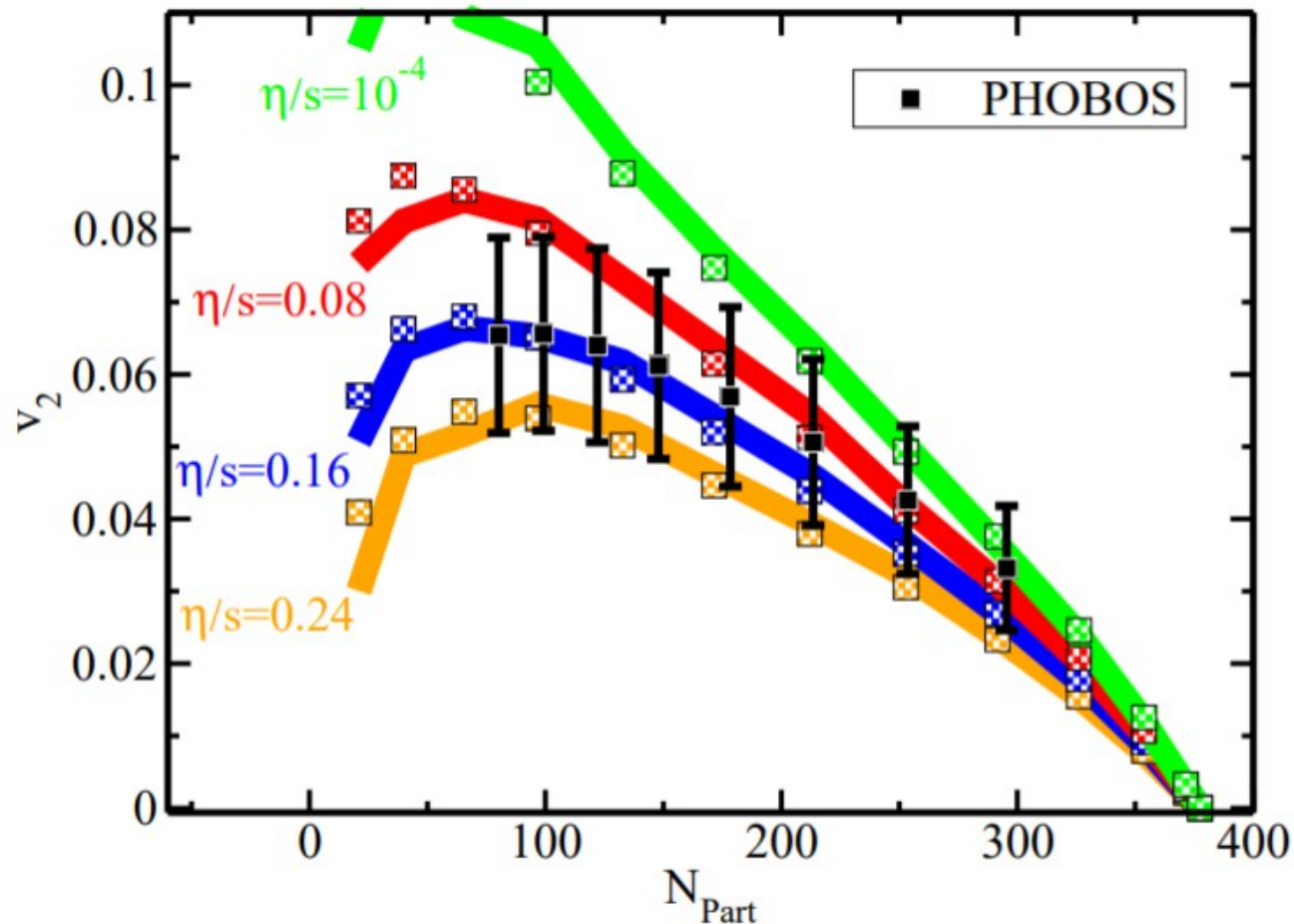
$\langle v_n^2 \rangle^{1/2}$



**Value of shear viscosity extracted**

# Extraction of shear viscosity: an example

M. Luzum and P. Romatschke, PRC78 (2008) 034915



Momentum anisotropy *decreases* with *increasing* viscosity

**Nowadays:** systematic model-to-data comparison with Global fits

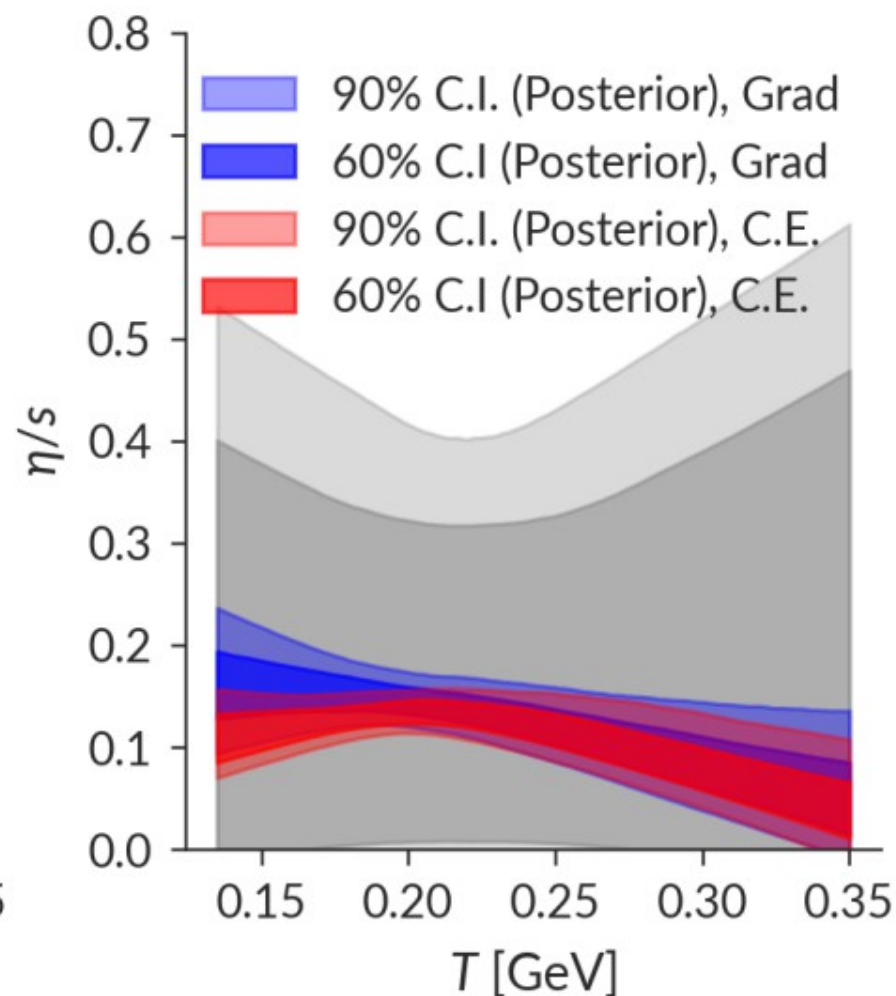
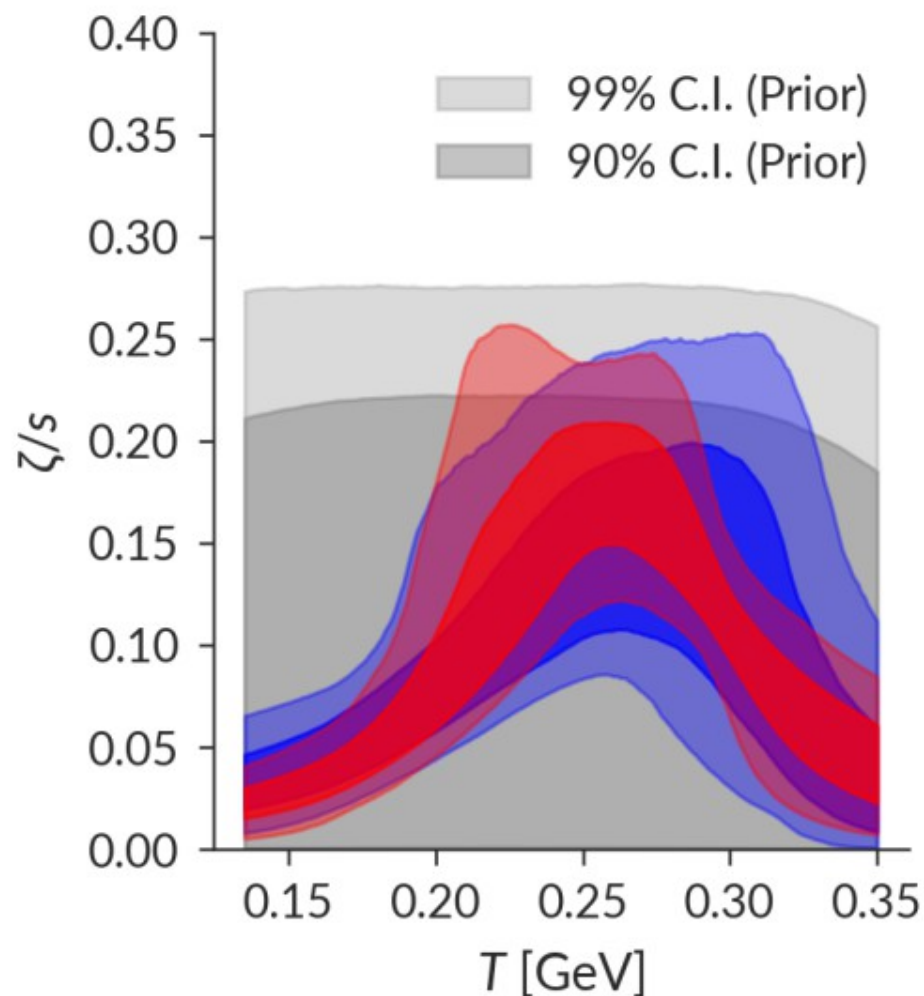
J. Bernhard, J. Moreland & Steffen A. Bass, Nature Physics , 1–5 (2019)



# Systematic model-to-data comparison – Bayesian Analysis

IPGlasma+MUSIC+UrQMD

Heffernan *et al*, Phys.Rev.C 109 (2024) 6, 065207

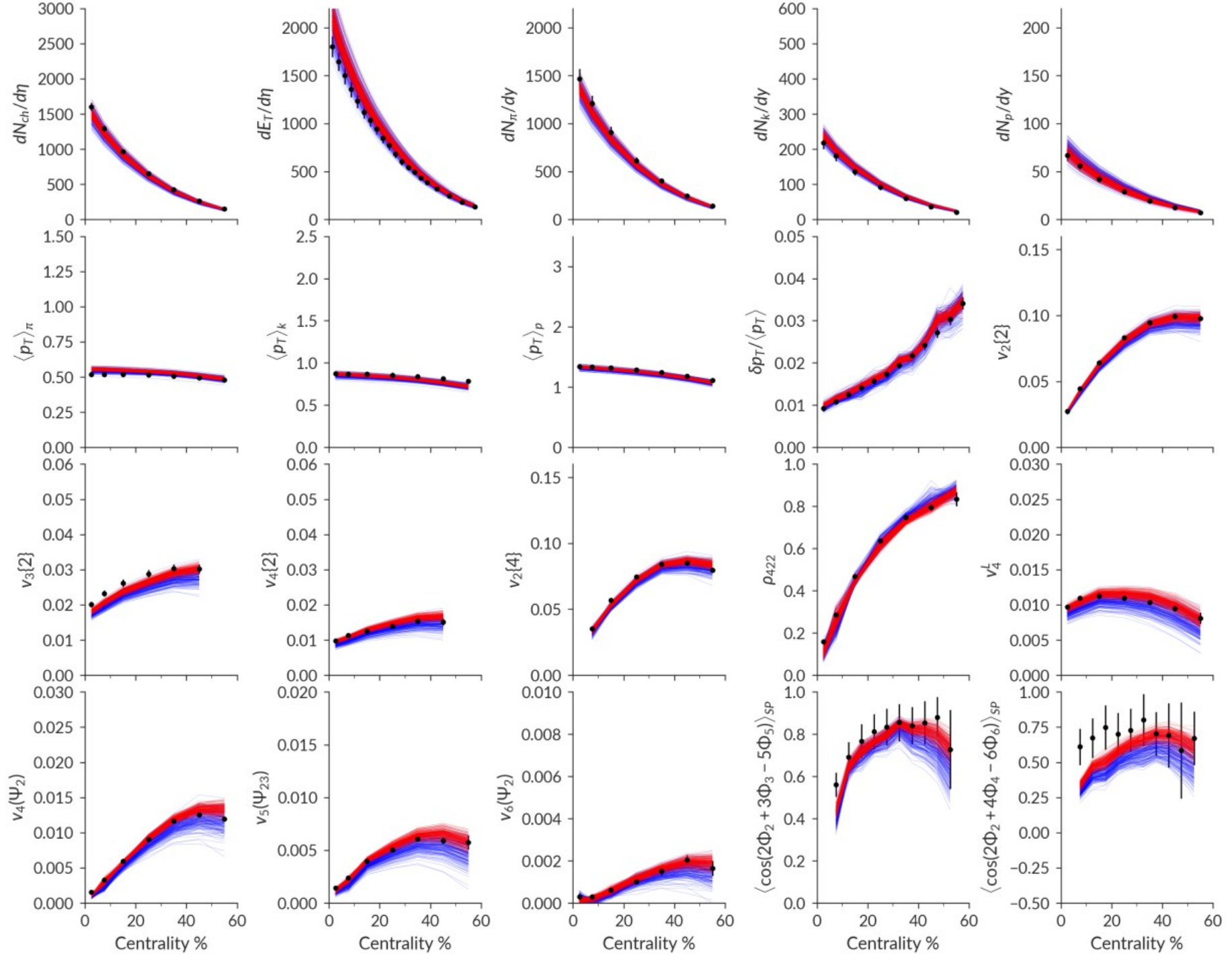


large bulk viscosity around “ $T_c$ ”!

**Almost perfect fluid!**

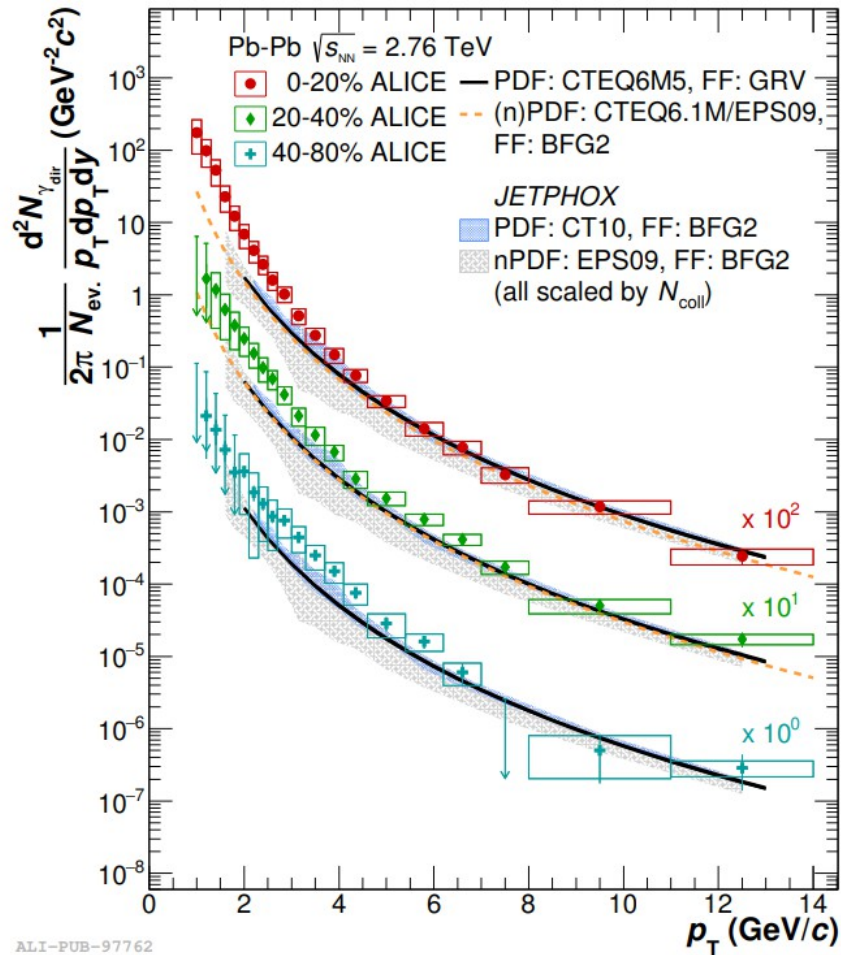
$\eta/s \sim 0.1$

# Systematic model-to-data comparison – Bayesian Analysis



# Significant evidence of a medium

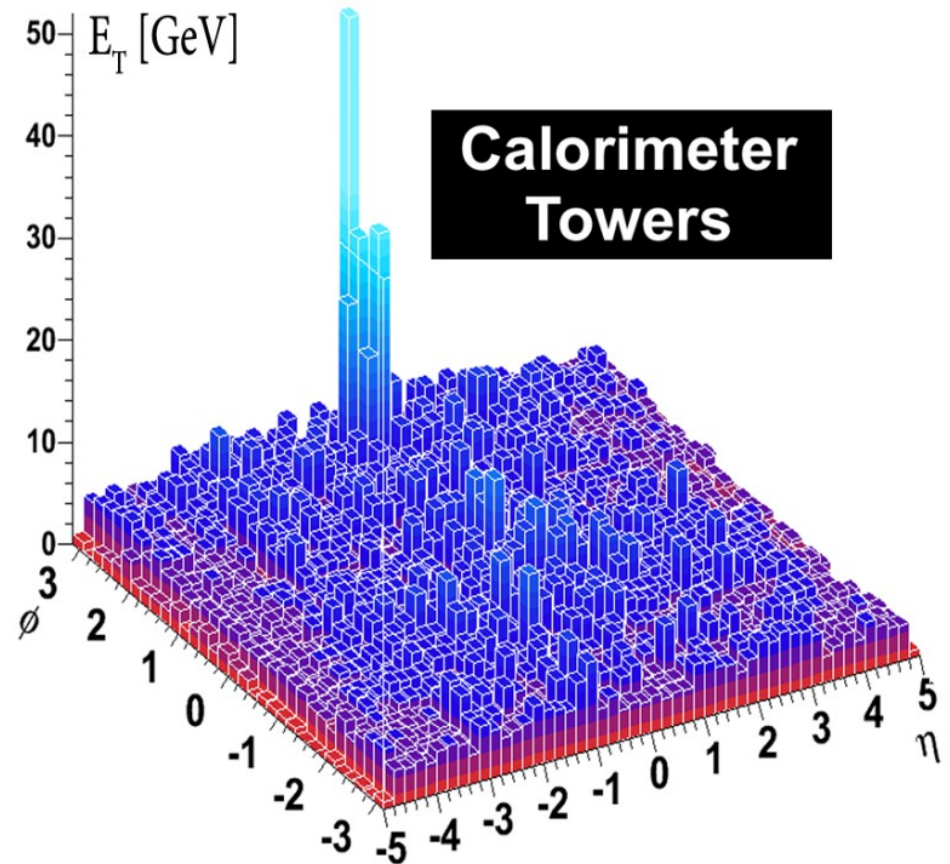
## thermal photons



Excess of direct photons at low  $p_T$  — thermal radiation

ALICE, Phys.Lett.B 754 (2016) 235-248

## Jet quenching



Jets are stopped by the medium

ATLAS, Phys.Rev.Lett. 105 (2010) 252303

# Relativistic fluid dynamics

Effective theory describing the dynamics of a system over long-times and long-distances

**Conservation laws**

+

**Equation of state**

+

**simple constitutive relations**



# Basics of fluid dynamics

## Conservation laws

**energy-momentum  
conservation**

$$\partial_\mu T^{\mu\nu} = 0$$

## Net charge conservation

$$\partial_\mu N_s^\mu = 0$$

strangeness

$$\partial_\mu N_e^\mu = 0$$

electric charge

$$\partial_\mu N_b^\mu = 0$$

Baryon number

## Tensor decomposition

$$N_q^\mu = n_q u^\mu + n_q^\mu$$

$$T^{\mu\nu} = \varepsilon u^\mu u^\nu - (P_0 + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu}$$

**net-charge diffusion  
4-current**

**Bulk viscous  
pressure**

**Shear stress  
tensor**

Projection operator:  $\Delta^{\mu\nu} = g^{\mu\nu} - u^\mu u^\nu$



# Basics of fluid dynamics

## Conservation laws

**energy-momentum  
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$$\partial_\mu T^{\mu\nu} = 0$$

**Net charge conservation**

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**strangeness**

$$\partial_\mu N_e^\mu = 0$$

**electric charge**

$$\partial_\mu N_b^\mu = 0$$

**Baryon number**

## Tensor decomposition

$$N_q^\mu = n_q u^\mu + n_q^\mu$$

**???**

$$T^{\mu\nu} = \varepsilon u^\mu u^\nu - (P_0 + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu}$$

**net-charge diffusion  
4-current**

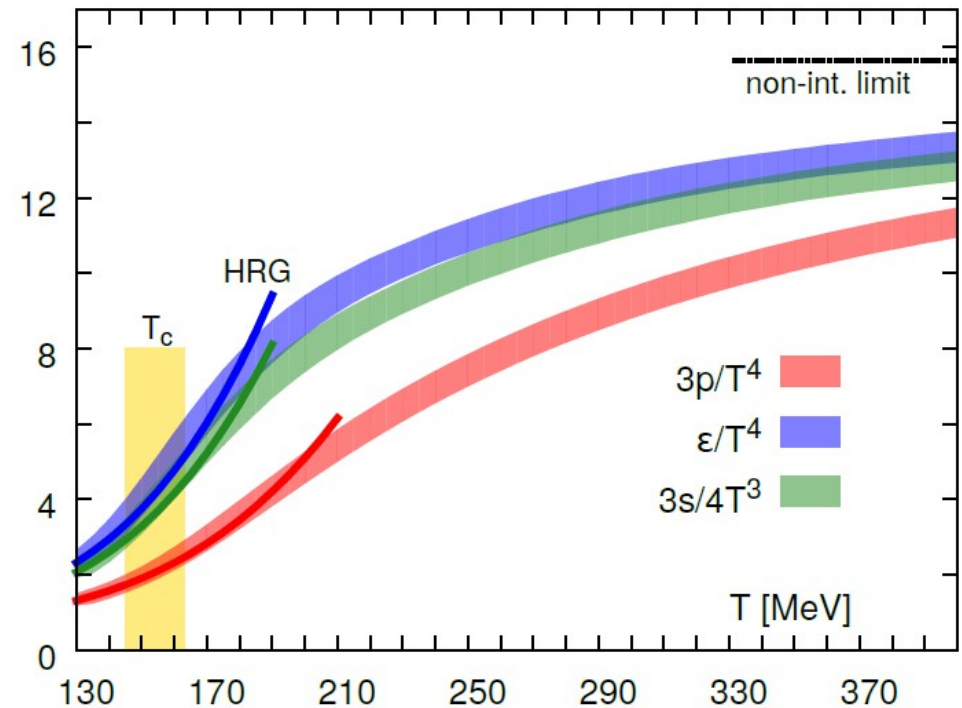
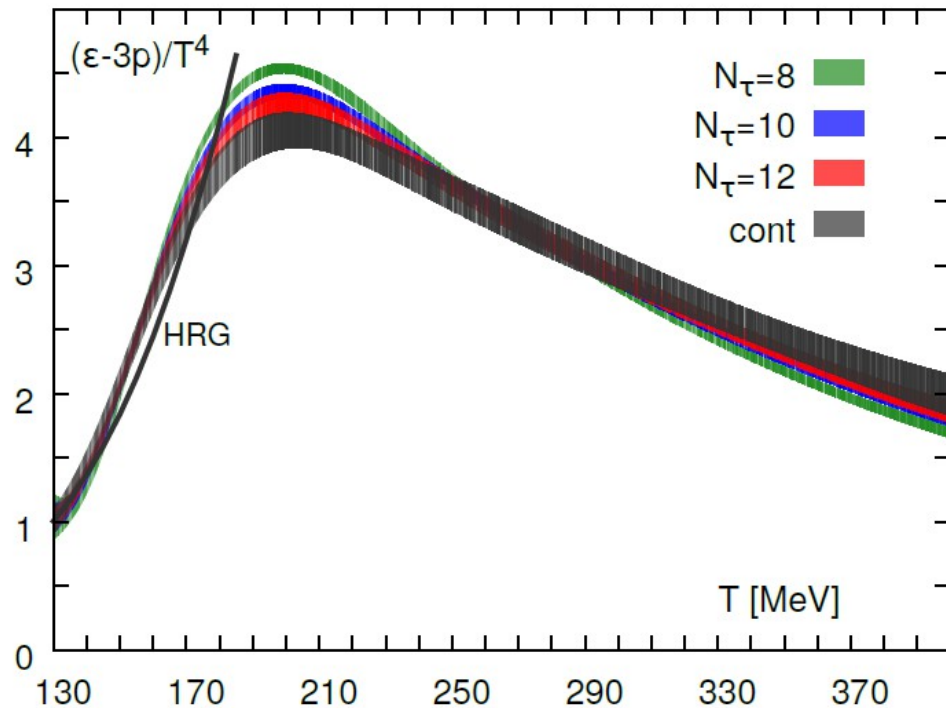
**Bulk viscous  
pressure**

**Shear stress  
tensor**

Projection operator:  $\Delta^{\mu\nu} = g^{\mu\nu} - u^\mu u^\nu$

# EoS for QCD matter – $\mu_B = 0$

**Lattice QCD:** entropy density increases near  $T = 180$  MeV  
**quark and gluon degrees of freedom become manifest**



PRD 90, 094503 (2014)

- QCD predicts a cross-over

**Below  $\sim 150$  MeV:** matches a hadron resonance gas

# EoS at finite $\mu$

Thermodynamic pressure:  $P_0 = P_0(T, \mu_b, \mu_e, \mu_s)$

Taylor expansion  
up to 4th order:  $\frac{P}{T^4} = \frac{P_0}{T^4} + \sum_{l,m,n} \underbrace{\frac{\chi_{l,m,n}^{B,Q,S}}{l!m!n!}}_{\text{1QCD}} \left(\frac{\mu_B}{T}\right)^l \left(\frac{\mu_Q}{T}\right)^m \left(\frac{\mu_S}{T}\right)^n$

- matched to *hadron resonance gas* model at small T
  - matched to Stefan-Boltzmann limit at large T
- 
- Prescription employed by:
    - Monnai, Schenke, Shen, PRC 100, 024907 (2019)
    - Noronha-Hostler, Parotto, Ratti, Stafford, PRC 100, 064910 (2019)

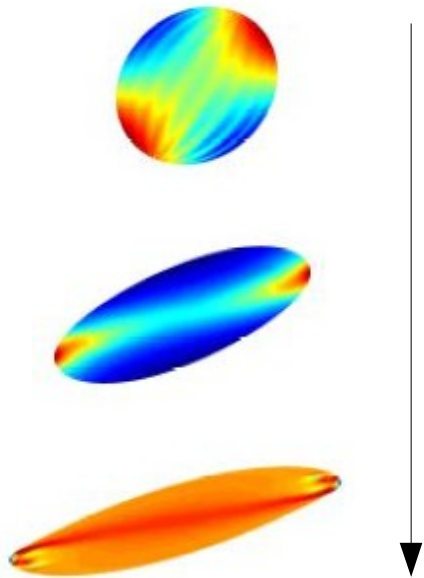
**Critical point?**

# Relativistic Navier-Stokes theory

## Shear Viscosity

(Resistance to deformation)

$$\pi^{\mu\nu} = 2\eta \nabla^{\langle\mu} u^{\nu\rangle}$$



$$\eta(T, \mu_q)$$

## Bulk Viscosity

(Resistance to expansion)

$$\Pi = -\zeta \nabla_\mu u^\mu$$



$$\zeta(T, \mu_q)$$

## Net-Charge Diffusion

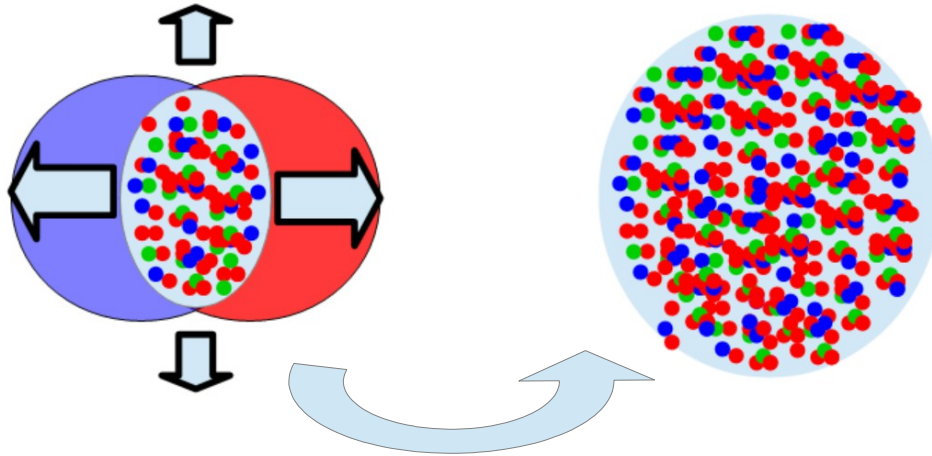
$$n_q^\mu = \kappa_q \nabla^\mu \frac{\mu_q}{T}$$



$$\kappa_q(T, \mu_q)$$

# Shear Viscosity (Resistance to deformation)

*Important for heavy ion collisions*



Change from elliptical to spherical shape

- *Shear viscosity will resist this process*
- *Shear viscosity reduces the flow coefficients*

*Intuition from kinetic theory (gas)*

$$\eta = \frac{1}{3} n p \ell_{\text{mfp}}$$

Annotations for the equation:

- density** (points to  $n$ )
- mean free path** (points to  $\ell_{\text{mfp}}$ )
- average momentum** (points to  $p$ )

- *Proportional to scattering rate*

small  $\eta/s$

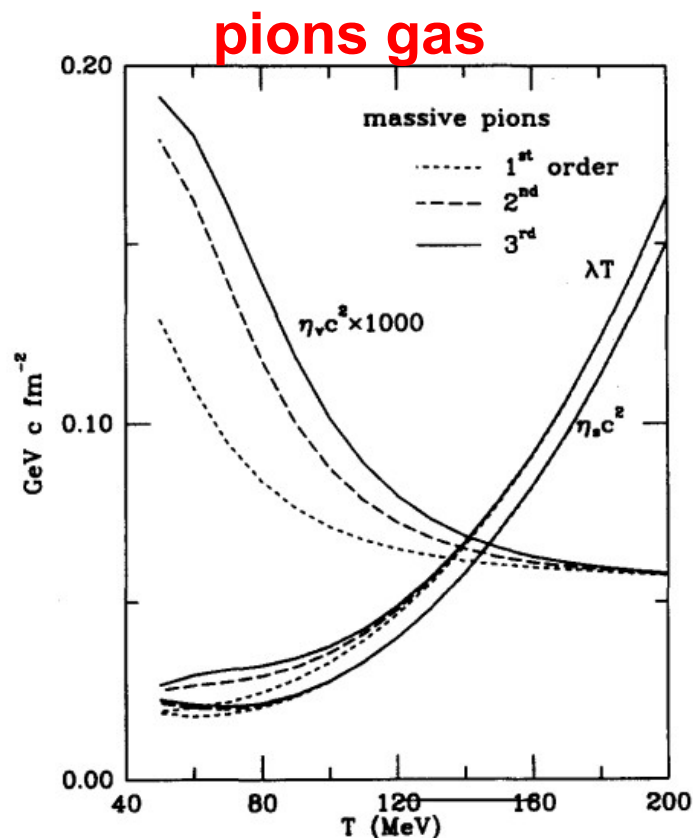
“strong” interaction

(this does not work for liquids)

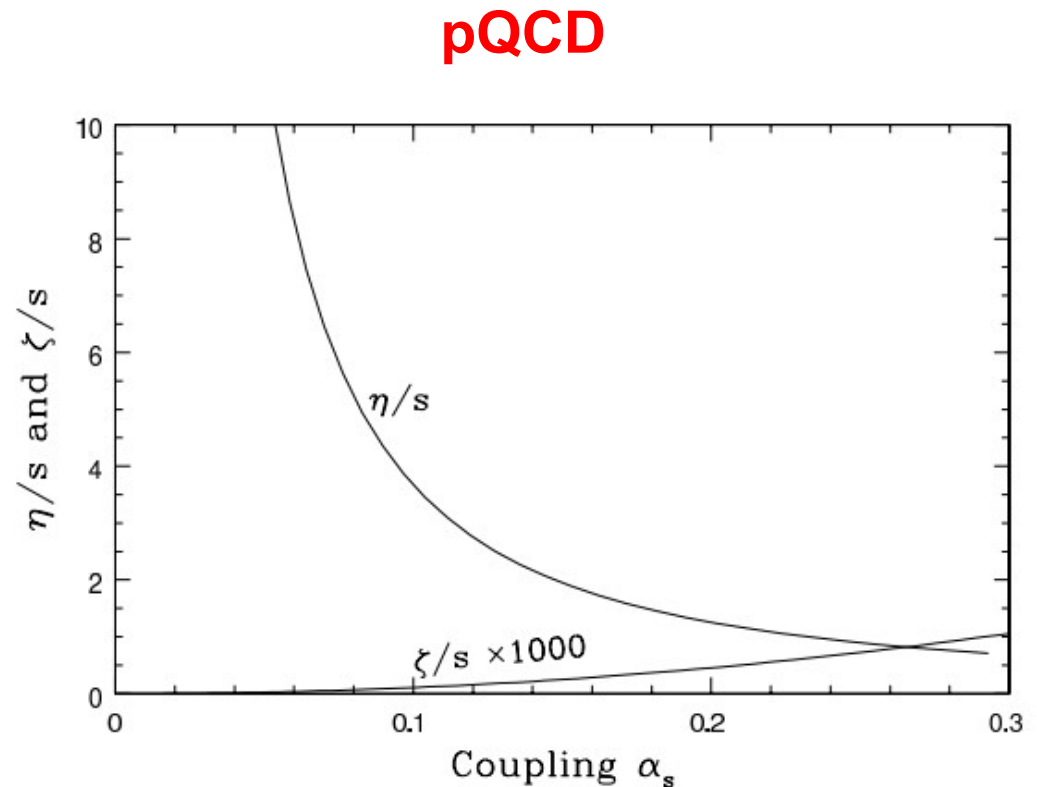


# What about bulk viscosity?

- should be important for highly explosive systems



Prakash *et al*, Phys. Rept. 227, 321 (1993)

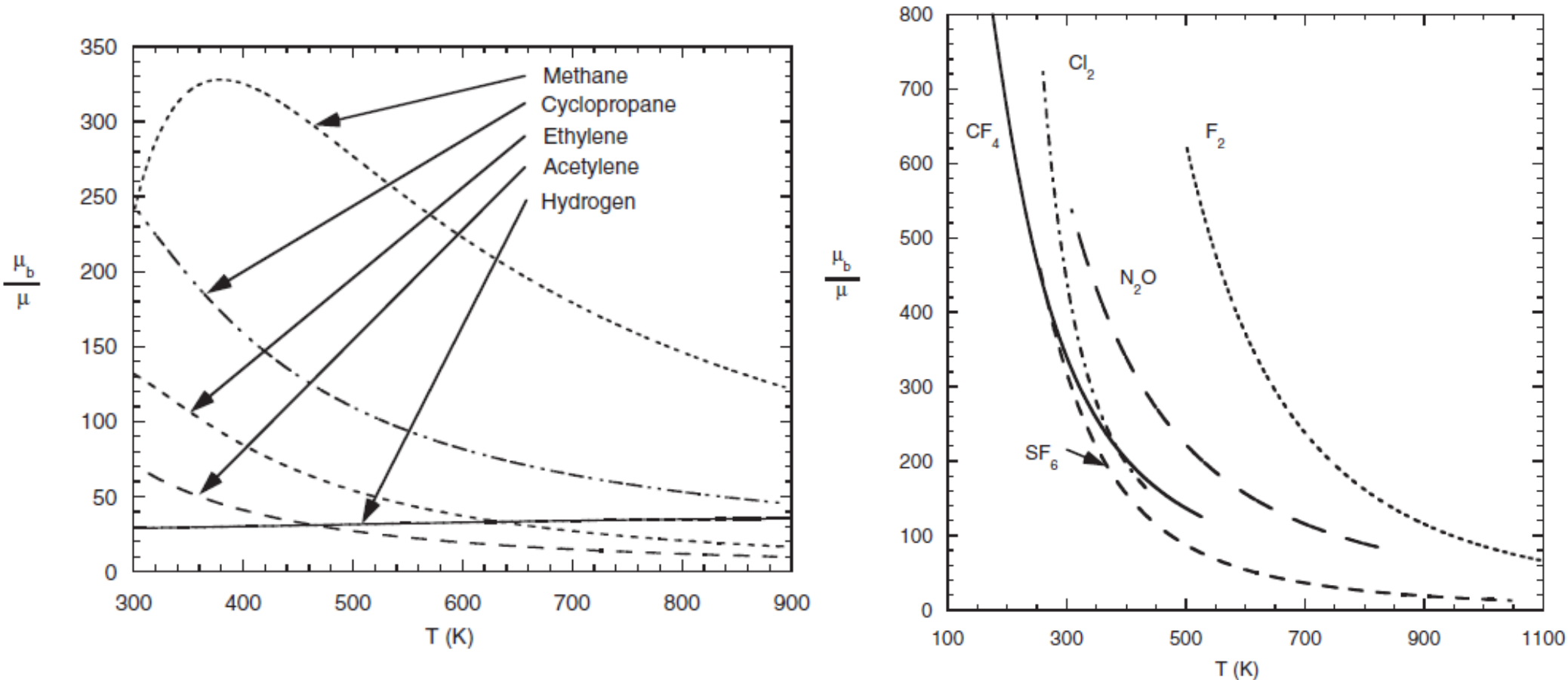


Arnold *et al*, PRD 74, 085021 (2006)

**Bulk viscosity ~1000 times smaller than shear**  
**What about near  $T_c$ ?**

# What about bulk viscosity?

## Bulk to shear viscosity ratio for several gases



M. Cramer, Phys. of fluids 24, 066102 (2012)

Bulk viscosity is large when molecules have excited states  
(vibration and rotation):  $\zeta \sim \tau_{\text{excitation}}$

~~$\zeta \sim \tau_{\text{mean free-time}}$~~

# What about bulk viscosity?

Bulk viscosity is large when molecules have internal degrees of freedom: can vibrate or rotate

i.e., bulk viscosity is large when kinetic energy is transferred to internal energy: production of excited states

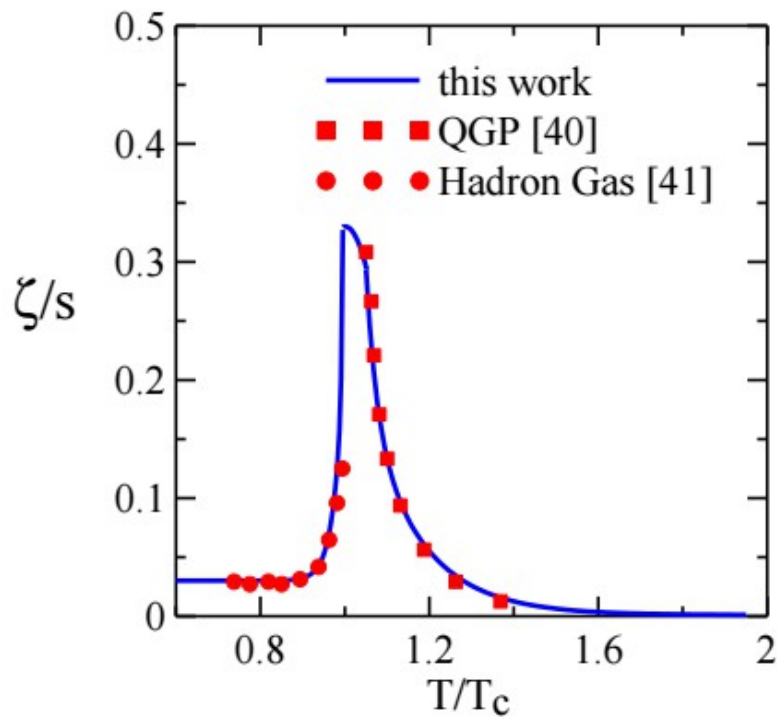
In QCD, we expect many resonances to be formed at the cross-over region: magnitude of bulk viscosity is related to the lifetime of these excited states

# Importance of bulk viscosity

**IP-Glasma** + **MUSIC** + **UrQMD**

Ryu *et al*, PRL 115, 132301 (2015)

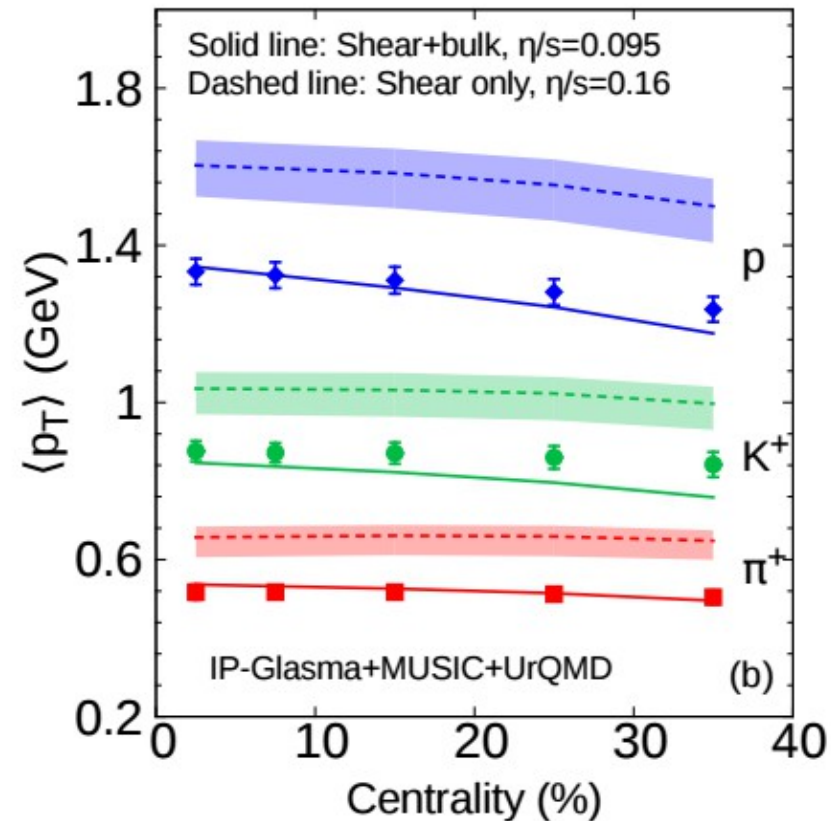
LHC  $\eta/s=0.095$



IP-Glasma initial conditions lead to high mean  $p_T$



Bulk viscosity reduces mean  $p_T$



Value of shear viscosity extracted changes significantly  
 $\eta/s=0.16 \longrightarrow 0.095$

# Relativistic Navier-Stokes theory

## Shear Viscosity

(Resistance to deformation)

$$\pi^{\mu\nu} = 2\eta \nabla^{\langle\mu} u^{\nu\rangle}$$

## Bulk Viscosity

(Resistance to expansion)

$$\Pi = -\zeta \nabla_\mu u^\mu$$

## Net-Charge Diffusion

$$n_q^\mu = \kappa_q \nabla^\mu \frac{\mu_q}{T}$$

- Equations violate causality and display unphysical instabilities

***Global equilibrium state is linearly unstable***

Hiscock and Lindblom, Phys.Rev.D 31 (1985) 725-733

- Relativistic Navier-Stokes theory cannot be used to model relativistic fluids
- This is the reason why it took the field so long to add dissipation into the models



# Why acausal and unstable?

- At large wave numbers, dispersion relations of Navier-Stokes have the typical form associated to the diffusion equation

$$\omega = iDk^2$$

$$\partial_t T = D \nabla^2 T ,$$

**acausal**

- If we perform a (1D) Lorentz boost

$$\begin{pmatrix} \omega \\ k \end{pmatrix} = \begin{pmatrix} \gamma & -\gamma V \\ -\gamma V & \gamma \end{pmatrix} \begin{pmatrix} \omega' \\ k' \end{pmatrix} = \begin{pmatrix} \gamma\omega' - \gamma V k' \\ -\gamma V \omega' + \gamma k' \end{pmatrix}$$

$$\omega = iDk^2 \longrightarrow \omega' - V k' = i\gamma D (k' - V \omega')^2$$

- An unstable mode appears:  $\omega'(k' = 0) = -\frac{i}{\gamma D V^2} ,$

# What we solve is not “traditional” fluid dynamics

Israel and Stewart, Annals Phys. 118 (1979) 341-372

$$\tau_{\Pi} D\Pi + \Pi = -\zeta\theta + \dots ,$$

$$\tau_n \Delta_{\alpha}^{\mu} Dn^{\alpha} + n^{\mu} = \kappa_n \nabla^{\mu} \alpha + \dots ,$$

$$\tau_{\pi} \Delta_{\alpha\beta}^{\mu\nu} D\pi^{\alpha\beta} + \pi^{\mu\nu} = 2\eta\sigma^{\mu\nu} + \dots .$$

**expansion rate**

$$\theta = \nabla_{\mu} u^{\mu}$$

**shear tensor**

$$\sigma^{\mu\nu} = \nabla^{\langle\mu} u^{\nu\rangle}$$

- relaxes to Navier-Stokes theory (asymptotic solution)
- **causality:** relaxation times cannot be arbitrarily small

$$\frac{\zeta}{(\varepsilon_0 + p_0)\tau_{\Pi}} + \frac{4}{3} \frac{\eta}{(\varepsilon_0 + p_0)\tau_{\pi}} \leq 1 - c_s^2 ,$$

# Complete equations: second order hydrodynamics

Denicol *et al* PRD 85 (2012) 114047

$$\begin{aligned} \dot{\Pi} = & -\frac{\Pi}{\tau_{\Pi}} - \beta_{\Pi}\theta - \ell_{\Pi n}\partial \cdot n - \tau_{\Pi n}n \cdot \dot{u} - \delta_{\Pi\Pi}\Pi\theta \\ & -\lambda_{\Pi n}n \cdot \nabla\alpha_0 + \lambda_{\Pi\pi}\pi^{\mu\nu}\sigma_{\mu\nu} , \end{aligned} \quad (20)$$

**bulk**

$$\begin{aligned} \dot{n}^{\langle\mu\rangle} = & -\frac{n^{\mu}}{\tau_n} + \beta_n\nabla^{\mu}\alpha_0 - n_{\nu}\omega^{\nu\mu} - \delta_{nn}n^{\mu}\theta - \ell_{n\Pi}\nabla^{\mu}\Pi \\ & +\ell_{n\pi}\Delta^{\mu\nu}\partial_{\lambda}\pi_{\nu}^{\lambda} + \tau_{n\Pi}\Pi\dot{u}^{\mu} - \tau_{n\pi}\pi_{\nu}^{\mu}\dot{u}^{\nu} \\ & -\lambda_{nn}n^{\nu}\sigma_{\nu}^{\mu} + \lambda_{n\Pi}\Pi\nabla^{\mu}\alpha_0 - \lambda_{n\pi}\pi^{\mu\nu}\nabla_{\nu}\alpha_0 , \end{aligned} \quad (21)$$

**diffusion**

$$\begin{aligned} \dot{\pi}^{\langle\mu\nu\rangle} = & -\frac{\pi^{\mu\nu}}{\tau_{\pi}} + 2\beta_{\pi}\sigma^{\mu\nu} + 2\pi_{\alpha}^{\langle\mu}\omega^{\nu\rangle\alpha} - \tau_{\pi n}n^{\langle\mu}\dot{u}^{\nu\rangle} \\ & +\ell_{\pi n}\nabla^{\langle\mu}n^{\nu\rangle} - \delta_{\pi\pi}\pi^{\mu\nu}\theta - \tau_{\pi\pi}\pi_{\alpha}^{\langle\mu}\sigma^{\nu\rangle\alpha} \\ & +\lambda_{\pi n}n^{\langle\mu}\nabla^{\nu\rangle}\alpha_0 + \lambda_{\pi\Pi}\Pi\sigma^{\mu\nu} . \end{aligned} \quad (22)$$

**shear**

- derived from kinetic theory
- often called DNMR equations

# Equations of motion employed - DNMR

Denicol *et al*, Phys.Rev.D 85 (2012) 114047

$$\begin{aligned}\tau_{\Pi}\dot{\Pi} + \Pi &= -\zeta\theta - \delta_{\Pi\Pi}\Pi\theta + \varphi_1\Pi^2 + \lambda_{\Pi\pi}\pi^{\mu\nu}\sigma_{\mu\nu} + \varphi_3\pi^{\mu\nu}\pi_{\mu\nu} , \\ \tau_{\pi}\dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} &= 2\eta\sigma^{\mu\nu} + 2\pi_{\alpha}^{\langle\mu}\omega^{\nu\rangle\alpha} - \delta_{\pi\pi}\pi^{\mu\nu}\theta + \varphi_7\pi_{\alpha}^{\langle\mu}\pi^{\nu\rangle\alpha} - \tau_{\pi\pi}\pi_{\alpha}^{\langle\mu}\sigma^{\nu\rangle\alpha} \\ &\quad + \lambda_{\pi\Pi}\Pi\sigma^{\mu\nu} + \varphi_6\Pi\pi^{\mu\nu} .\end{aligned}$$

## Transport coefficients – ultrarelativistic gas

$$z = m/T$$

Phys.Rev.C 90 (2014) 2, 024912

$$\begin{aligned}\frac{\eta}{\tau_{\pi}} &= \frac{\varepsilon_0 + P_0}{5} + \mathcal{O}(z^2) , & \frac{\zeta}{\tau_{\Pi}} &= 15 \left( \frac{1}{3} - c_s^2 \right)^2 (\varepsilon_0 + P_0) + \mathcal{O}(z^5) \\ \frac{\delta_{\pi\pi}}{\tau_{\pi}} &= \frac{4}{3} + \mathcal{O}(z^2) , & \frac{\delta_{\Pi\Pi}}{\tau_{\Pi}} &= \frac{2}{3} + \mathcal{O}(z^2 \ln z) , \\ \frac{\tau_{\pi\pi}}{\tau_{\pi}} &= \frac{10}{7} + \mathcal{O}(z^2) , & \frac{\lambda_{\Pi\pi}}{\tau_{\Pi}} &= \frac{8}{5} \left( \frac{1}{3} - c_s^2 \right) + \mathcal{O}(z^4) , \\ \frac{\lambda_{\pi\Pi}}{\tau_{\pi}} &= \frac{6}{5} + \mathcal{O}(z^2 \ln z) , & &\end{aligned}$$

Applied to heavy ion collisions ...  
(MUSIC)

# Particle emission

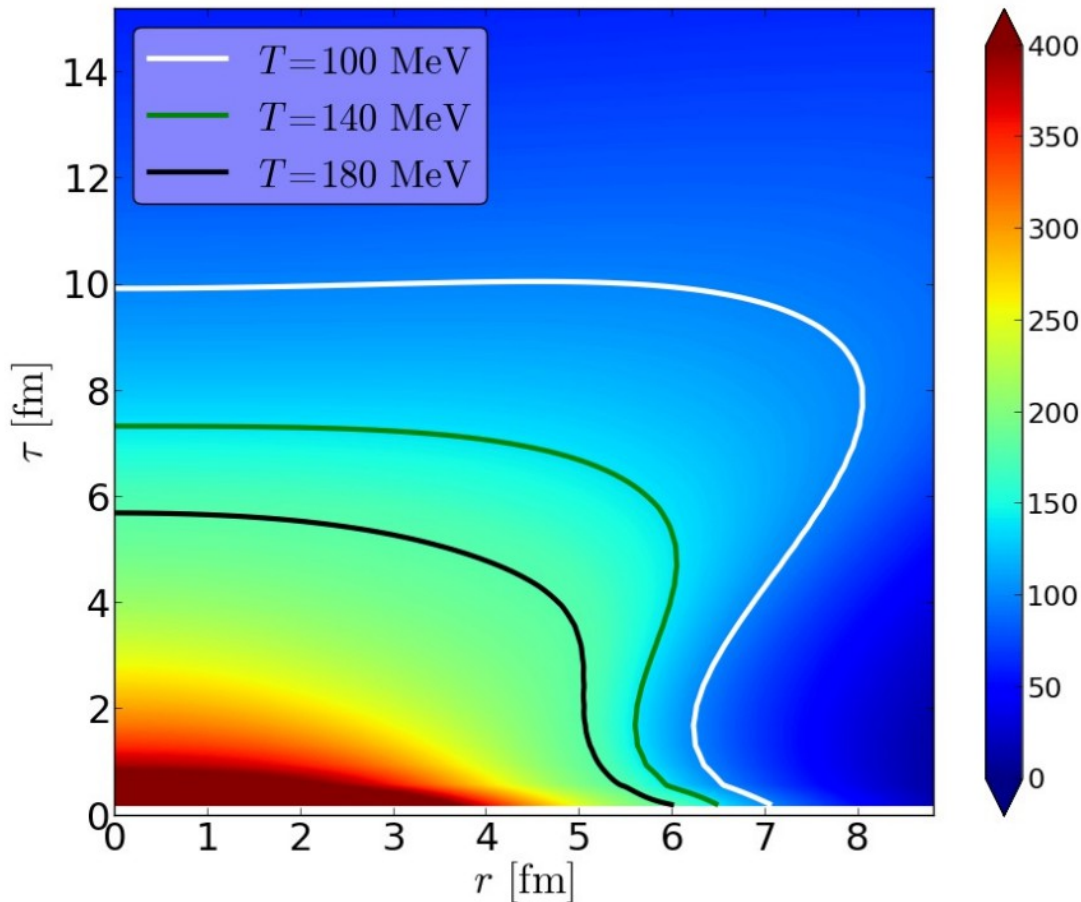
**Fluid elements**



**Particles**

Assumption: hadrons are emitted from a hypersurface

(smooth) central collision



constant temperature



# Particle emission

**Fluid elements**



**Particles**

Assumption: hadrons are emitted from a hypersurface

■ Cooper-Frye procedure

$$\frac{d^2 N_i}{d^2 p_T dy} = \frac{g_i}{(2\pi)^3} \int_{\Sigma} p_i^{\mu} d\sigma_{\mu} (f_0^i + \delta f^i)$$

constant temperature

Freeze-out  
Hypersurface

Non-equilibrium  
single-particle  
momentum distribution

**Common choice:** “Democratic ansatz”

$$\delta f_i = f_{0i} \frac{p_i^{\mu} p_i^{\nu} \pi_{\mu\nu}}{T^2 (e + P)}.$$

# Particle emission

**Fluid elements**



**Particles**

**Remark:** we have to solve a very complicated problem!

Energy-momentum tensor  
hadron resonance gas

$$T^{\mu\nu}(x^\lambda)$$

difficult



Momentum distribution  
of hadrons

$$f(x^\lambda, p^\lambda)$$



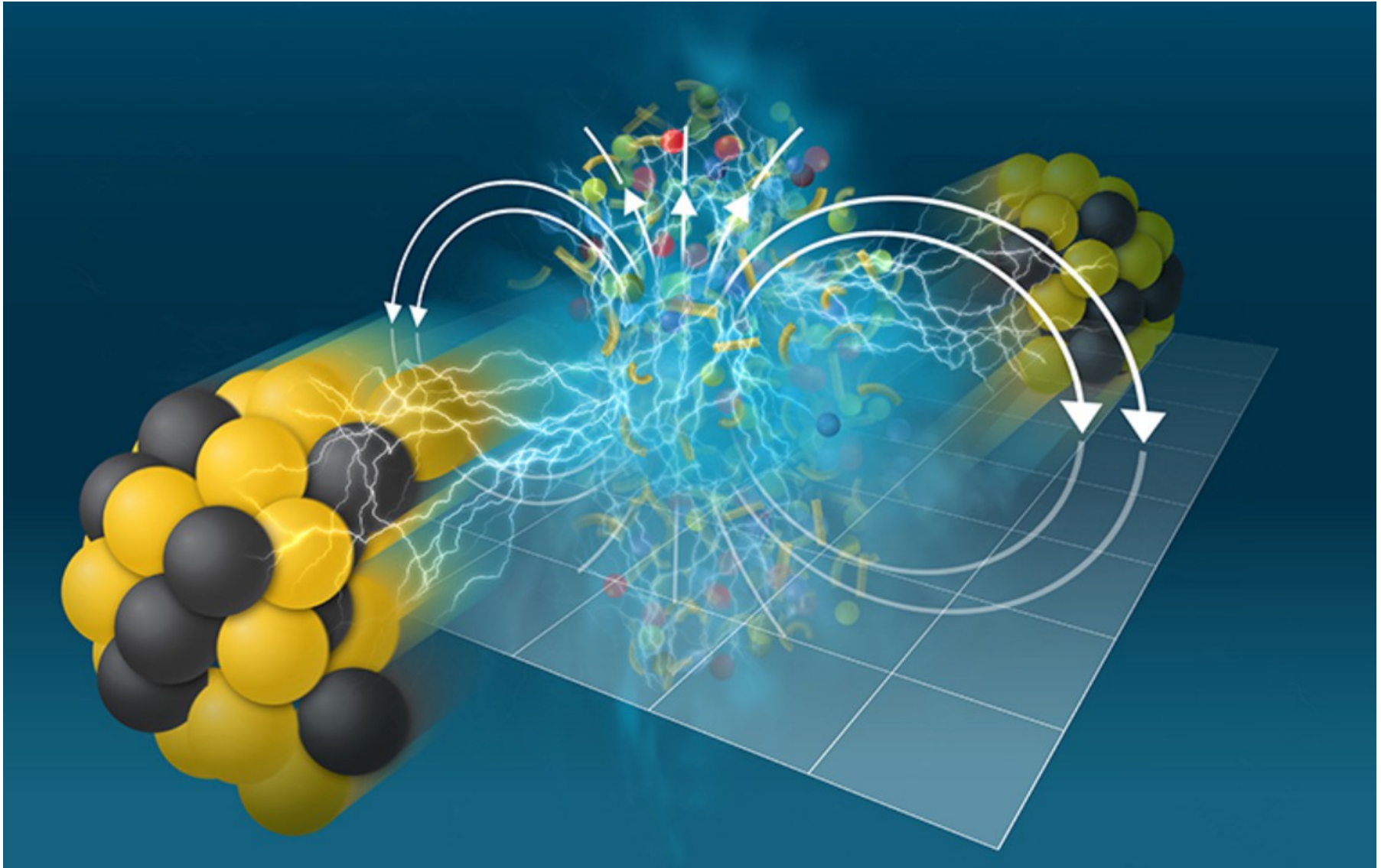
easy

**Assumption:** system is very close to equilibrium ...

**Assumption:** hydrodynamics and transport match ...

**Assumption:** Relaxation time approximation ...

# Large Electromagnetic fields



<https://www.bnl.gov/newsroom/>

# Large Electromagnetic fields

**Spectators** (protons) produce *intense* and *shortlived* electromagnetic fields

electromagnetic fields induced by the medium can increase the life time of this effect

Gursoy&Kharzeev&Rajagopal  
Phys.Rev.C 89 (2014) 5, 054905

This can only be addressed within a framework of relativistic magnetohydrodynamics

# Magnetohydrodynamics

## Conservation laws

**energy-momentum  
conservation**

$$\partial_\mu T^{\mu\nu} = -F^{\nu\lambda} N_\lambda$$

**energy-momentum exchanged with EM fields**

**Net electric-charge conservation**

$$\partial_\mu N^\mu = 0$$

Faraday tensor:  $F^{\mu\nu} = E^\mu u^\nu - E^\nu u^\mu + \epsilon^{\mu\nu\alpha\beta} u_\alpha B_\beta$

## Maxwell's equations

**Electric field**

**Magnetic field**

$$\partial_\mu F^{\mu\nu} = N^\nu,$$

$$\partial_\mu \tilde{F}^{\mu\nu} = 0,$$

$$\tilde{F}^{\mu\nu} = B^\mu u^\nu - B^\nu u^\mu - \epsilon^{\mu\nu\alpha\beta} u_\alpha E_\beta$$

Fields due to spectators are added as an external source



# Magnetohydrodynamics

## Dissipative equations (simplified)

### Viscous contributions

$$\tau_\pi \dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = 2\eta \nabla^{\langle\mu} u^{\nu\rangle} - \frac{4}{3} \tau_\pi \pi^{\mu\nu} \nabla_\lambda u^\lambda$$

$$\tau_\Pi \dot{\Pi} + \Pi = -\zeta \nabla_\lambda u^\lambda - \tau_\Pi \Pi \nabla_\lambda u^\lambda$$

- no effects due to EM fields included

### charge 4-diffusion current

$$\tau_e \dot{n}^\mu + n^\mu = \underbrace{\kappa_e \nabla^\mu \frac{\mu_e}{T}}_{\text{negligible at high energies}} - \tau_e n_e^\mu \nabla_\lambda u^\lambda + \underbrace{\sigma_e E^\mu}_{\text{Ohms Law}}$$

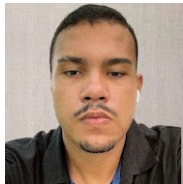
negligible at high energies

Ohms Law

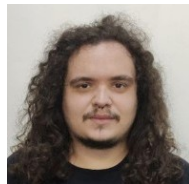
These equations were implemented in a new simulation code: PRAIA



Kevin Pala



Vinicius  
França



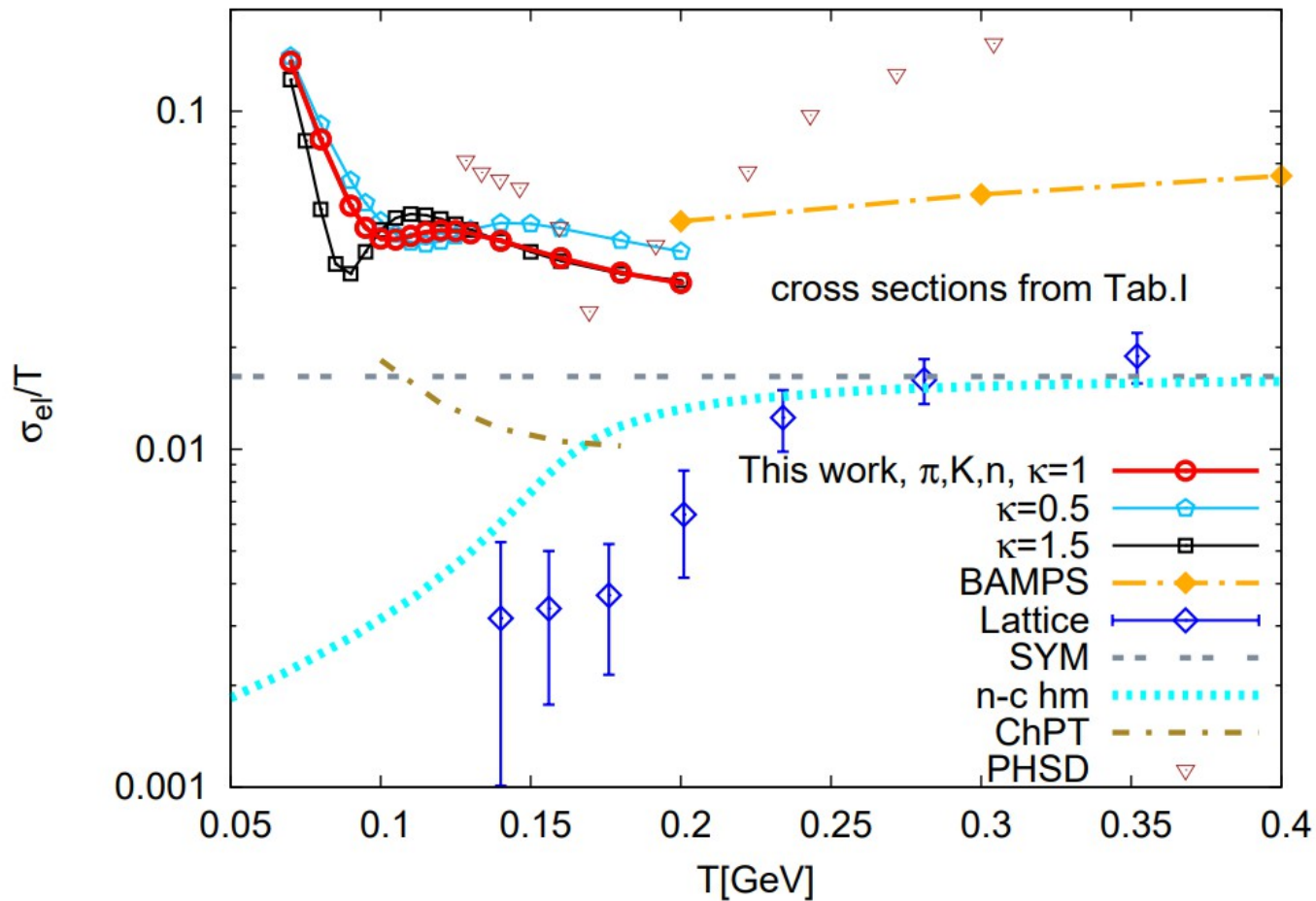
Jonatan Sola

**Smooth Particle Hydrodynamics  
algorithm**

# Magnetohydrodynamics

## Electric conductivity

Greif&Greiner&Denicol Phys.Rev.D 93 (2016) 9, 096012



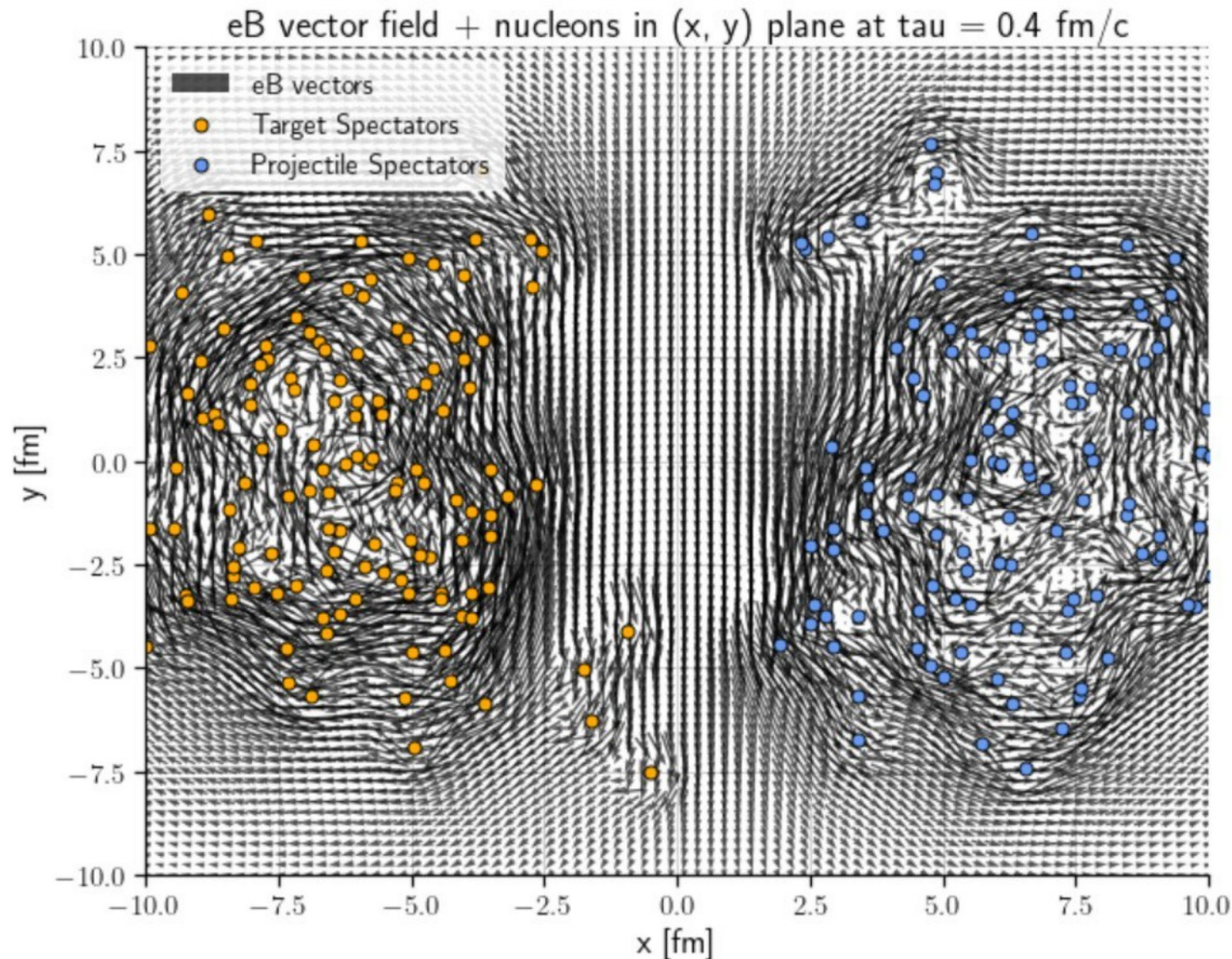
$$\sigma_e \sim 0.001-0.1 \text{ T}$$

# Spectator's magnetic field

*AMPT+PRAIA*,  $\tau_0 = 0.4 \text{ fm}$

30-40%

**PRELIMINARY**



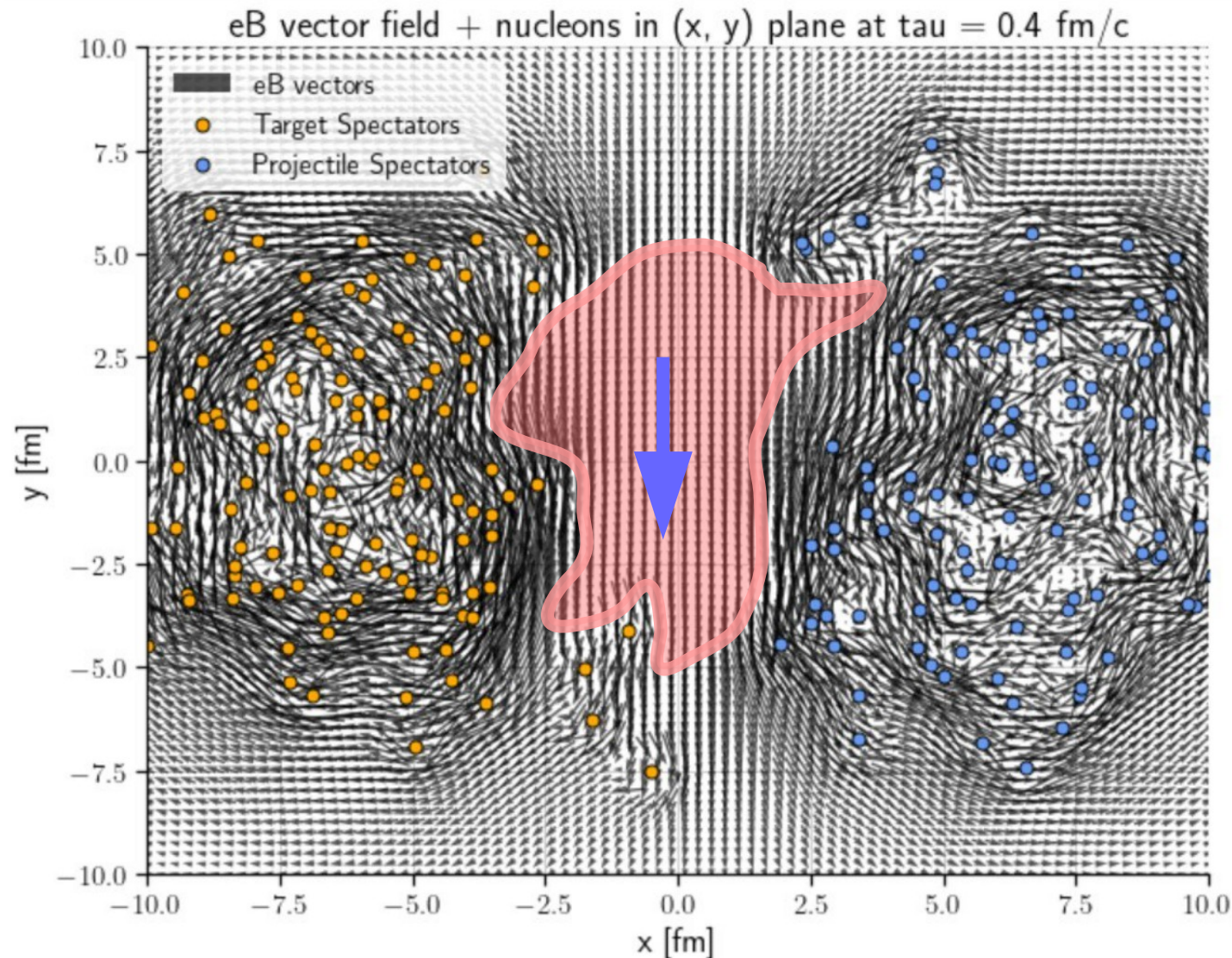


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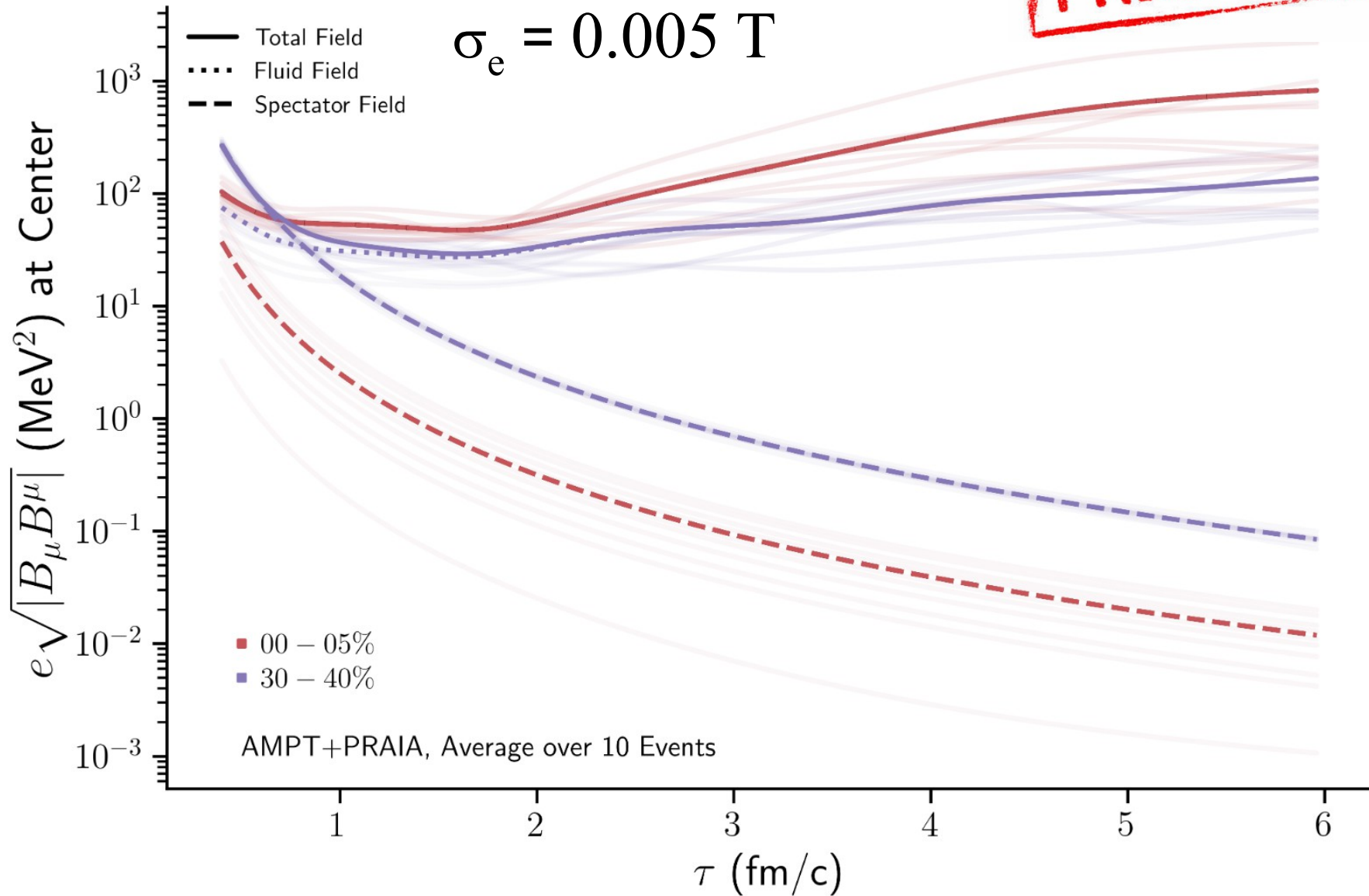


# 3D+1 Event-by-event Simulations

*AMPT+PRAIA*,  $\tau_0 = 0.4 \text{ fm}$

**PRELIMINARY**

$\sigma_e = 0.005 \text{ T}$

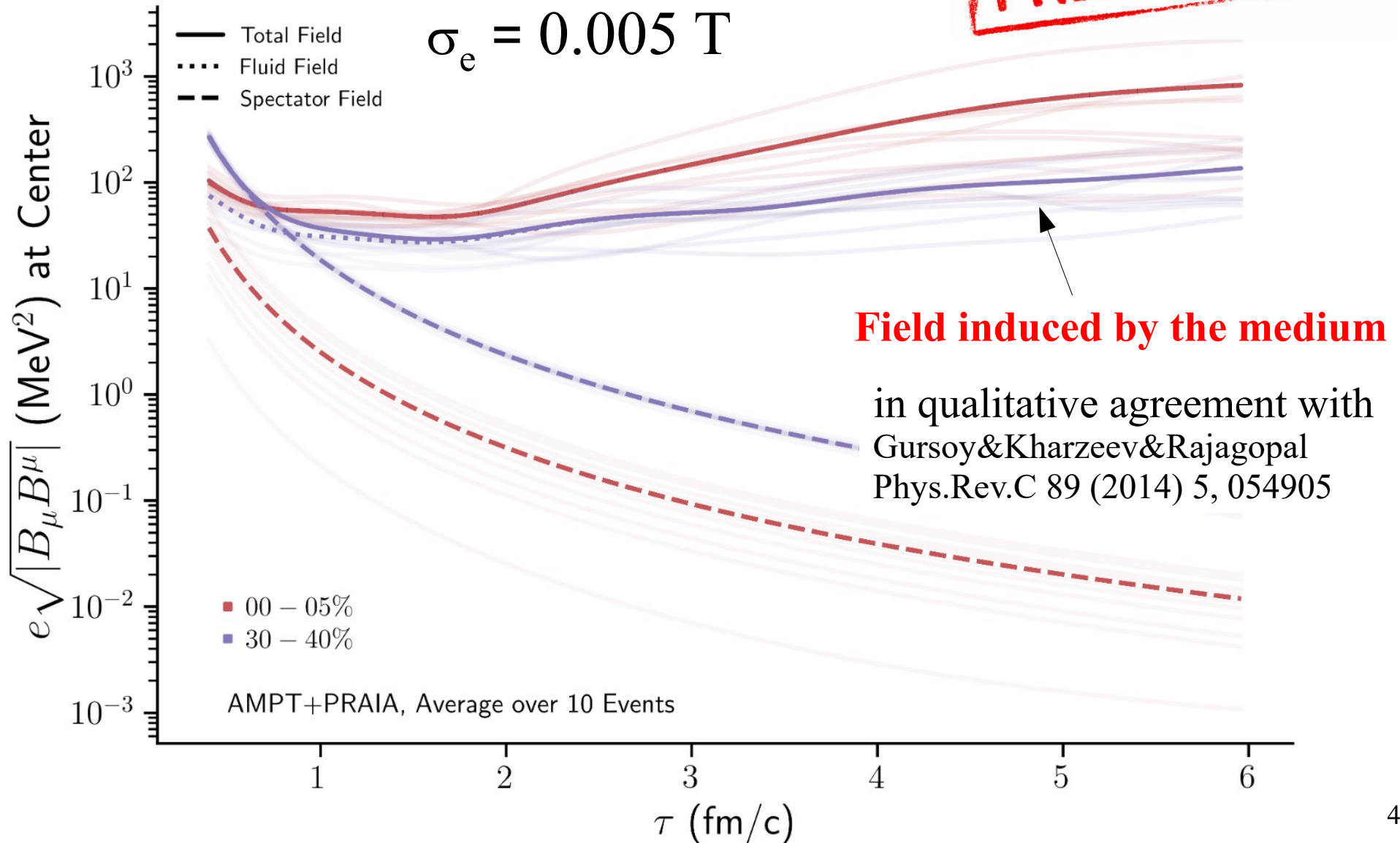


# 3D+1 Event-by-event Simulations

*AMPT+PRAIA*,  $\tau_0 = 0.4 \text{ fm}$

**PRELIMINARY**

$\sigma_e = 0.005 \text{ T}$



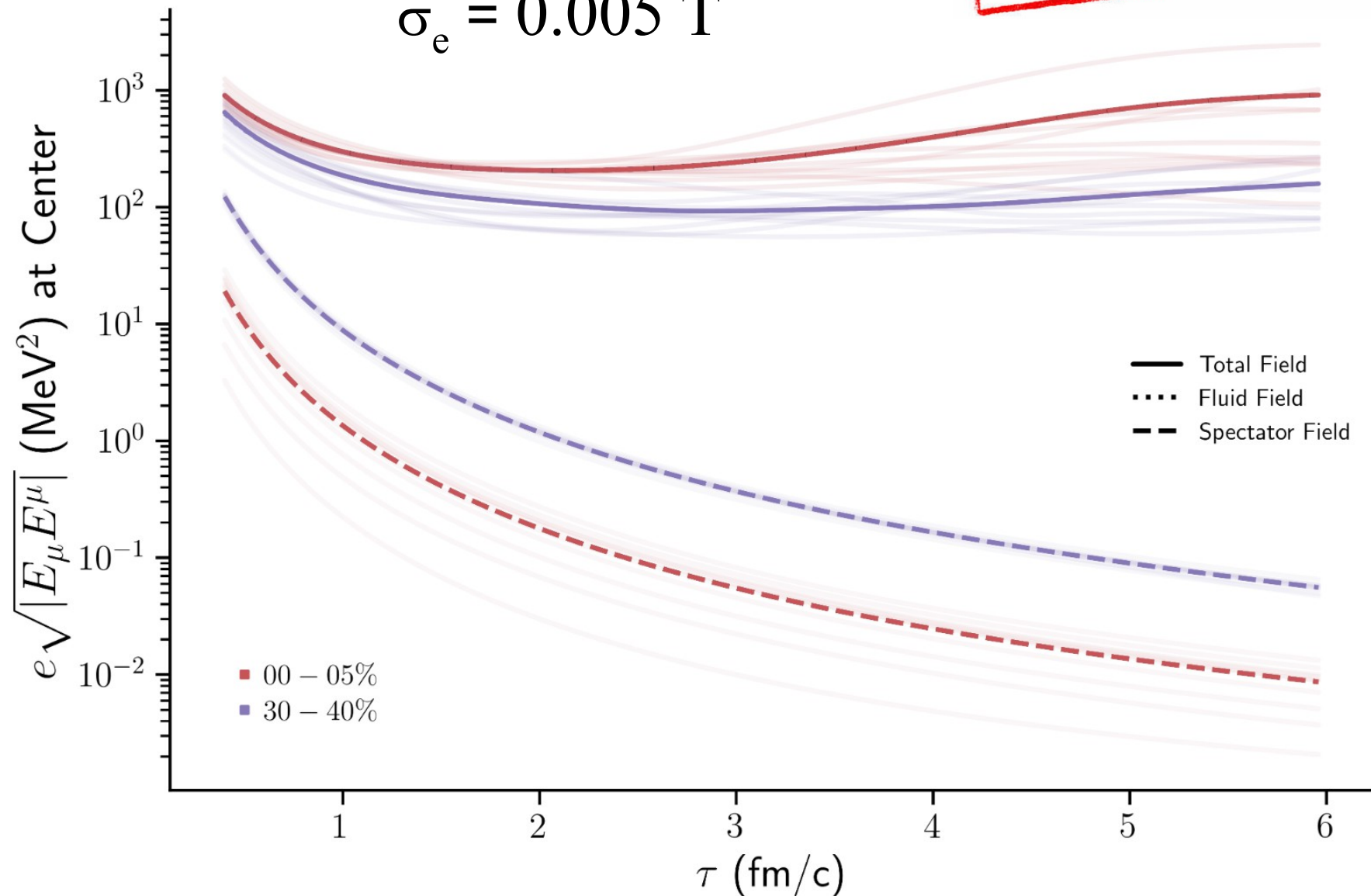


# 3D+1 Event-by-event Simulations

*AMPT+PRAIA*,  $\tau_0 = 0.4 \text{ fm}$

$\sigma_e = 0.005 \text{ T}$

**PRELIMINARY**



# Future directions

- *Field generated by participants? Effect on observables?*
- *Chiral magnetic effect*

**Chiral current:**

$$\partial_\mu N_5^\mu = -\frac{C}{2\pi^2} E_\mu B^\mu$$
$$N_5^\mu = n_5 u^\mu + n_5^\mu$$

**Unstable?**

**Dissipative currents:**

$$\tau_e \dot{n}^\mu + n^\mu = \kappa_e \nabla^\mu \frac{\mu_e}{T} - \tau_e n_e^\mu \nabla_\lambda u^\lambda + \sigma_e E^\mu + C \mu_5 B^\mu$$
$$\tau_5 \dot{n}_5^\mu + n_5^\mu = \kappa_5 \nabla^\mu \frac{\mu_5}{T} - \tau_5 n_5^\mu \nabla_\lambda u^\lambda + C \mu_e B^\mu$$

# Final remarks

Discussed *some* basic aspects of fluid-dynamical models

- Simulations are essential to interpret the data and extract the properties of QCD matter

*Properties of matter must be reverse-engineered!*

- Models are very complicated and include a lot of physics

*Nowadays, we study heavy ion collisions themselves, as a proxy to develop our understanding of QCD*

# Final remarks – a few challenges

## Hydrodynamics with spin degrees of freedom?

- New theories being constructed.
- May be crucial to describe data.

## Hydrodynamics near critical point?

- New theories are required ...
- Inclusion of fluctuations.

## Relativistic Magnetohydrodynamics?

- Causal theories of magnetohydrodynamics must be constructed