



Facultad de
Ciencias
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Explicit Regularization and medium effects in the NJL model

International School of Nuclear Physics, 46th course

Bernardo Andrade in collaboration with A. Sánchez - September 2025



Hadronic Matter in extreme conditions



- Temperature

[Dominguez PRD 86, 034030](#)

- Magnetic Field

[Skokov IJMPA 2009 24:31, 5925-5932](#)

- Baryonic Density

[Askawa Nucl. Phys. A 504 \(1989\) 668](#)

- Angular Momentum

[Becattini PRC.77.024906](#)

- Isospin

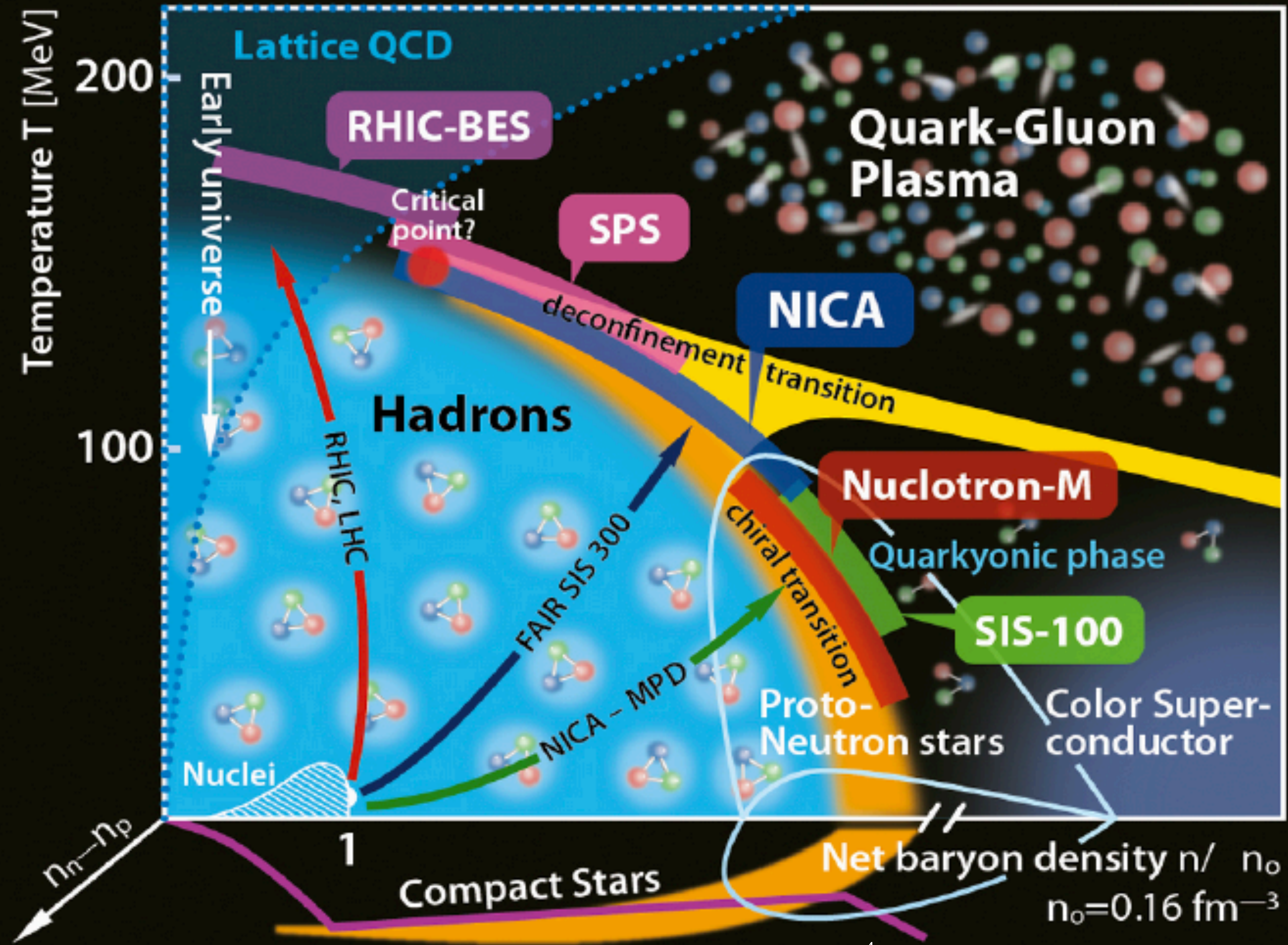
[Son PRL .86.592](#)

- Big Bang

- Heavy Ion Collisions

- Neutron Stars

QCD Phase Diagram



- It mostly shows our ignorance.
- Has theoretical predictions
- Can have many other axes
 - eB
 - I_3
 - L

Taken from Compact Stars in the QCD Phase Diagram VI

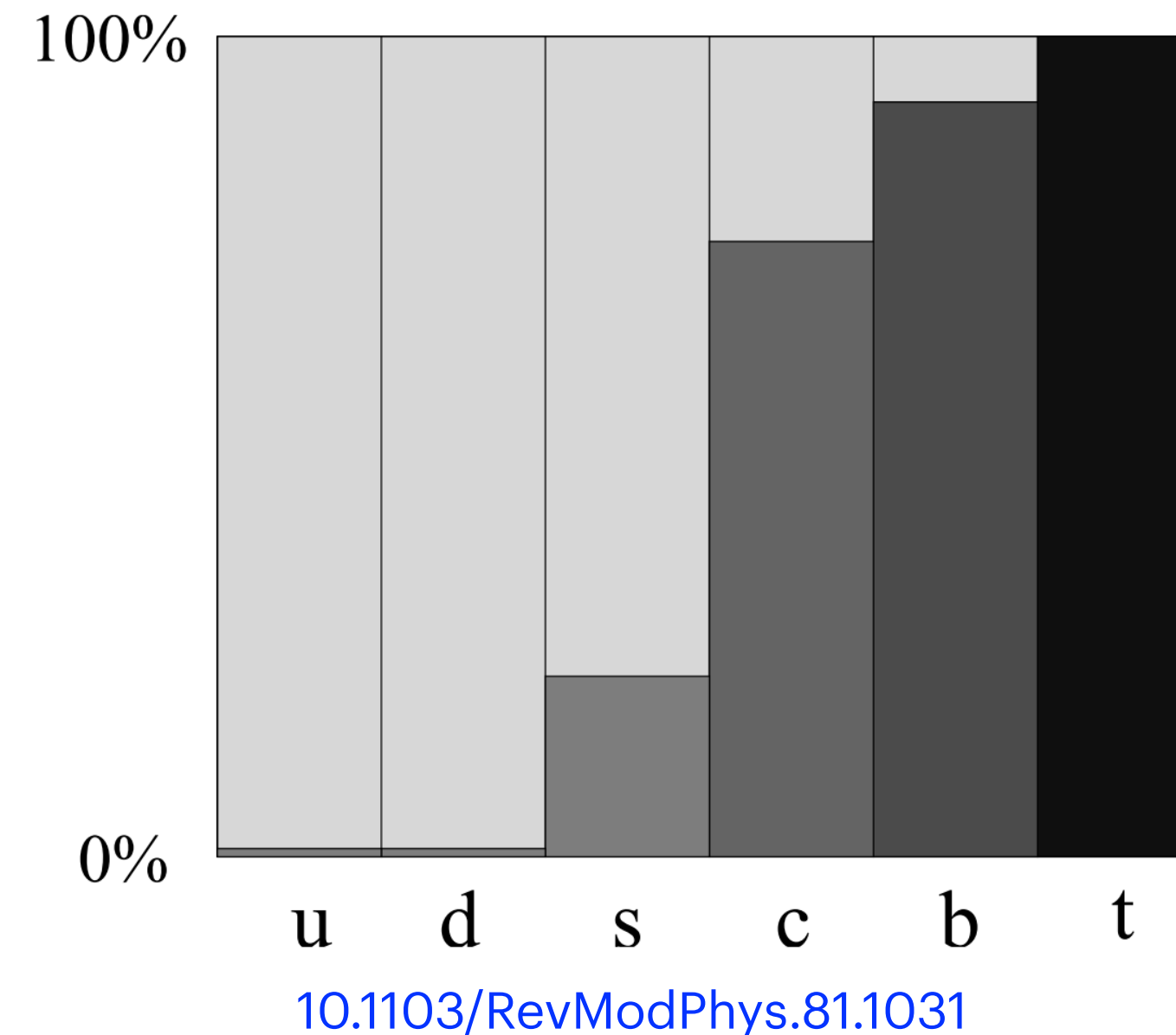
Two chiral phases

Restored - Broken

- Dynamically generated mass (Not from the Higgs Mechanism)
 - But from QCD vacuum phenomena — condensates, instantons, gluons
- Order parameter
 - Quark condensate

[Eur. Phys. J. A \(2023\) 59:252](#)

[Int.J.Mod.Phys.A24:5925–5932,2009](#)



Ways to study

- Lattice QCD

[Aoki, Endödi 10.1038/nature05120](#)

- Dyson-Schwinger Equations

[Fischer 10.1016j.ppnp.2019.01.002](#)

- Effective Field Theories

- Nambu - Jona-Lasinio

[Buballa2005 10.1016/j.physrep.2004.11.004](#)

[Klevansky 10.1103/RevModPhys.64.649](#)

[Fischer 10.1016j.ppnp.2019.01.002](#)

Schwinger-Dyson equations (QED)

First three of an infinite tower of **SDEs** in **QED**:

[Bashir Hugs 2023](#)

Going into NJL

Lagrangian Density

- Contact Interaction

$$\mathcal{L}_{NJL} = \bar{q}(i\gamma^\mu \partial_\mu - m)q + G \left((\bar{q}q)^2 + (\bar{q}i\gamma_5\tau^a q)^2 \right)$$

- BCS - like ground state
 - Dynamically breaks symmetry with quark condensation
- $G = [Energy]^{-2} \rightarrow$ Non renormalizable

Going into NJL

Properties

- Mean Field Approximation \rightarrow Linearize the interaction

$$\mathcal{L}_{NJL} = \bar{\psi}(i\gamma_{\mu}\partial^{\mu} - m)\psi - \bar{\psi}\sigma\psi - \frac{1}{4G}\sigma^2$$

- $M = m + \sigma$, $\sigma = -2G \langle \bar{\psi}\psi \rangle$
- When divergences occur, associated quantities become dependent on a cutoff scale Λ .
- Parameters must be fixed to reproduce physical quantities
 - Set, G , Λ , m to fit m_{π} , f_{π} , M

Divergences

A first example: Effective Potential

- Effective Potential

$$\begin{aligned} V_{\text{eff}}(\sigma) &= V_{\text{tree}}(\sigma) + V^{(1)}(\sigma) \\ &= \frac{\sigma^2}{4G} + i\Omega^{-1} \ln \det \left[-\frac{\delta^2 \mathcal{L}}{\delta \bar{\psi} \delta \psi} \right] \end{aligned}$$

- In a convenient base, the second term is diagonal

$$V^{(1)} = tr \int \frac{d^4 p_E}{(2\pi)^4} \ln(p_E^2 + M^2)$$

- Power counting is sufficient to identify the divergence

Divergences

A first example

$$V^{(1)} = tr \int \frac{d^4 p_E}{(2\pi)^4} \ln(p_E^2 + M^2) = - 2N_c N_f \int_{REG} \frac{d^4 p_E}{(2\pi)^4} \ln(p_E^2 + M^2)$$

- A regularization scheme must be implemented
 - Sharp Cutoff
 - Pauli-Villars
 - Dimensional Regularization
 - Form Factor

How to regulate; schemes

Cutoff

- Cutoff - restrict the region of integration

- 4D $V^{(1)} = -2N_c N_f \int d\Omega_3 \int_0^{\Lambda_p} \frac{dp_E}{(2\pi)^4} p_E^3 \ln(p_E^2 + M^2)$

Integration of p_4

- 3D $V^{(1)} = -2N_c N_f \int d\Omega_2 \int_0^{\Lambda_p} \frac{dp_E}{(2\pi)^4} p_E^2 \sqrt{p_E^2 + M^2}$

- Proper Time $V^{(1)} = \frac{2N_c N_f}{(2\pi)^4} \int_{1/\Lambda_s^2}^{\infty} \frac{ds}{s} \left(\sqrt{\frac{\pi}{s}} \right)^4 e^{-sM^2}$

Use of identity

$$\ln(x) = - \int \frac{ds}{s} e^{-sx}$$

Switching infinities

$$p \rightarrow s$$

- When using $\ln(x) = - \int \frac{ds}{s} e^{-sx}$, a new divergence appears when $s \rightarrow 0$
 - Cutoff on the nature of the logarithm
- The procedure of this scheme involves $\int_{-\infty}^{\infty} dp f(p)$
- The original cutoff Λ_p is no more, it gets switched for Λ_s
- What is the meaning of Λ_s ?

Switching infinities

Is this valid?

- Λ_p can be interpreted as the energy for which the model predictions are accurate.
 - Λ_s comes from a mathematical identity.
 - We totally discarded the divergence on p in favor of the divergence on s .
 - Let's see if there is a way to preserve the divergence on momentum.

Another Representation

Can we still capture the essence of proper time?

- A quick revision in the literature shows that

$$\ln x = \int_0^\infty \frac{ds}{s} (e^{-s} - e^{-sx})$$

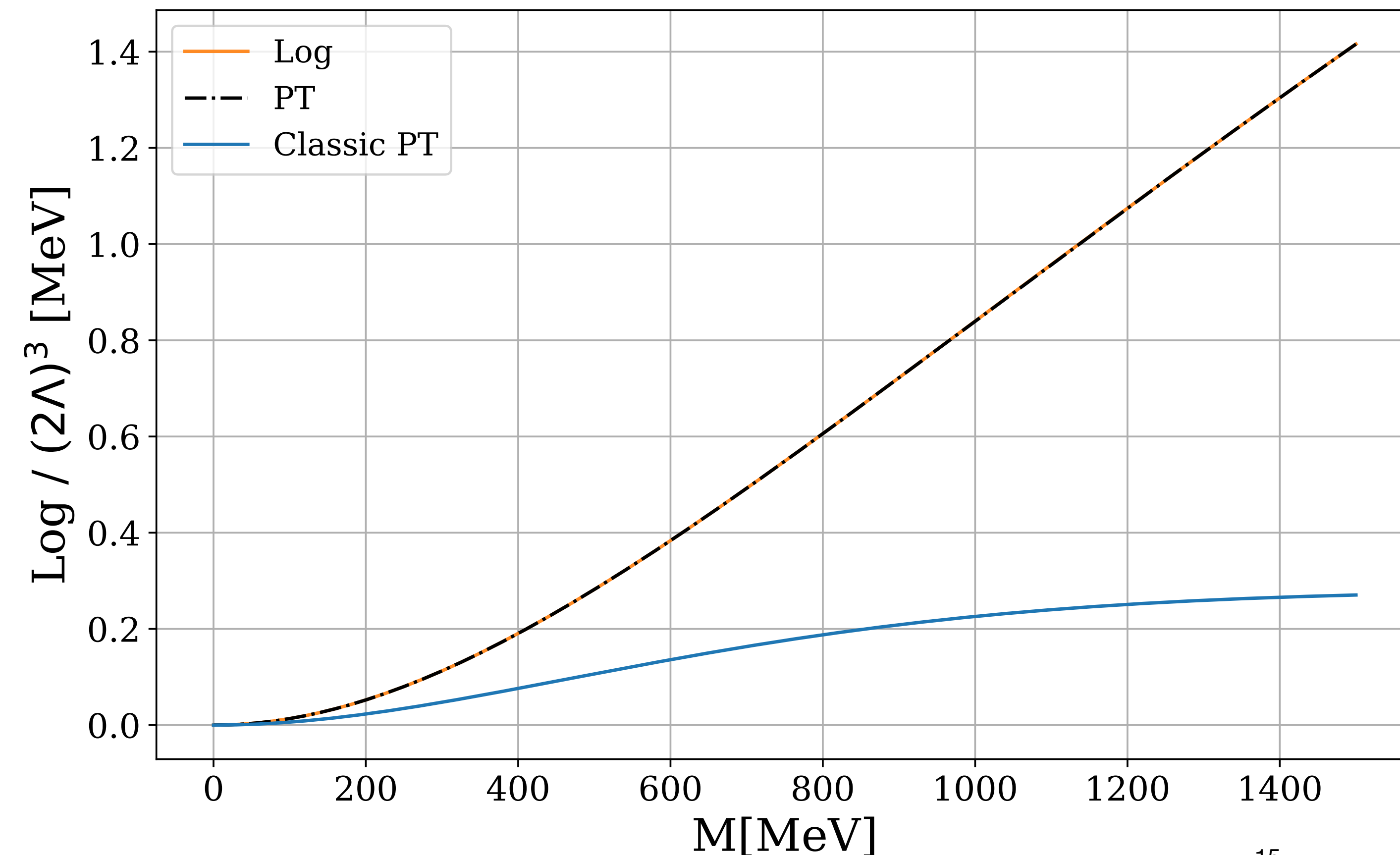
- Plugging into the effective potential, we can compute the momentum integrals

$$V_{eff} = \frac{1}{4G} \sigma^2 - \frac{N_c N_f}{8\pi^2} \int_0^\infty \frac{ds}{s} \left((2\Lambda)^4 e^{-s} - \frac{e^{-sM^2} \text{Erf}(\sqrt{s}\Lambda)^4}{s^2} \right)$$

Another Representation

How does this hold up?

- Lets compare the three representations



- $\int_{-\Lambda_p}^{\Lambda_p} \frac{d^4 p}{(2\pi)^4} \ln(p_4^2 + \omega^2 + M^2) \rightarrow \text{Log}$

- $-\int_{-\Lambda_p}^{\Lambda_p} \frac{dp_1 dp_2 dp_3 dp_4}{(2\pi)^4} \int_{1/\Lambda_s^2}^{\infty} \frac{ds}{s} e^{-s(p_4^2 + \omega^2 + M^2)} \rightarrow \text{Classic PT}$

- $\int_{-\Lambda_p}^{\Lambda_p} \frac{dp_1 dp_2 dp_3 dp_4}{(2\pi)^4} \int_0^{\infty} \frac{ds}{s} \left(e^{-s} - e^{-s(p_4^2 + \omega^2 + M^2)} \right) \rightarrow \text{PT}$

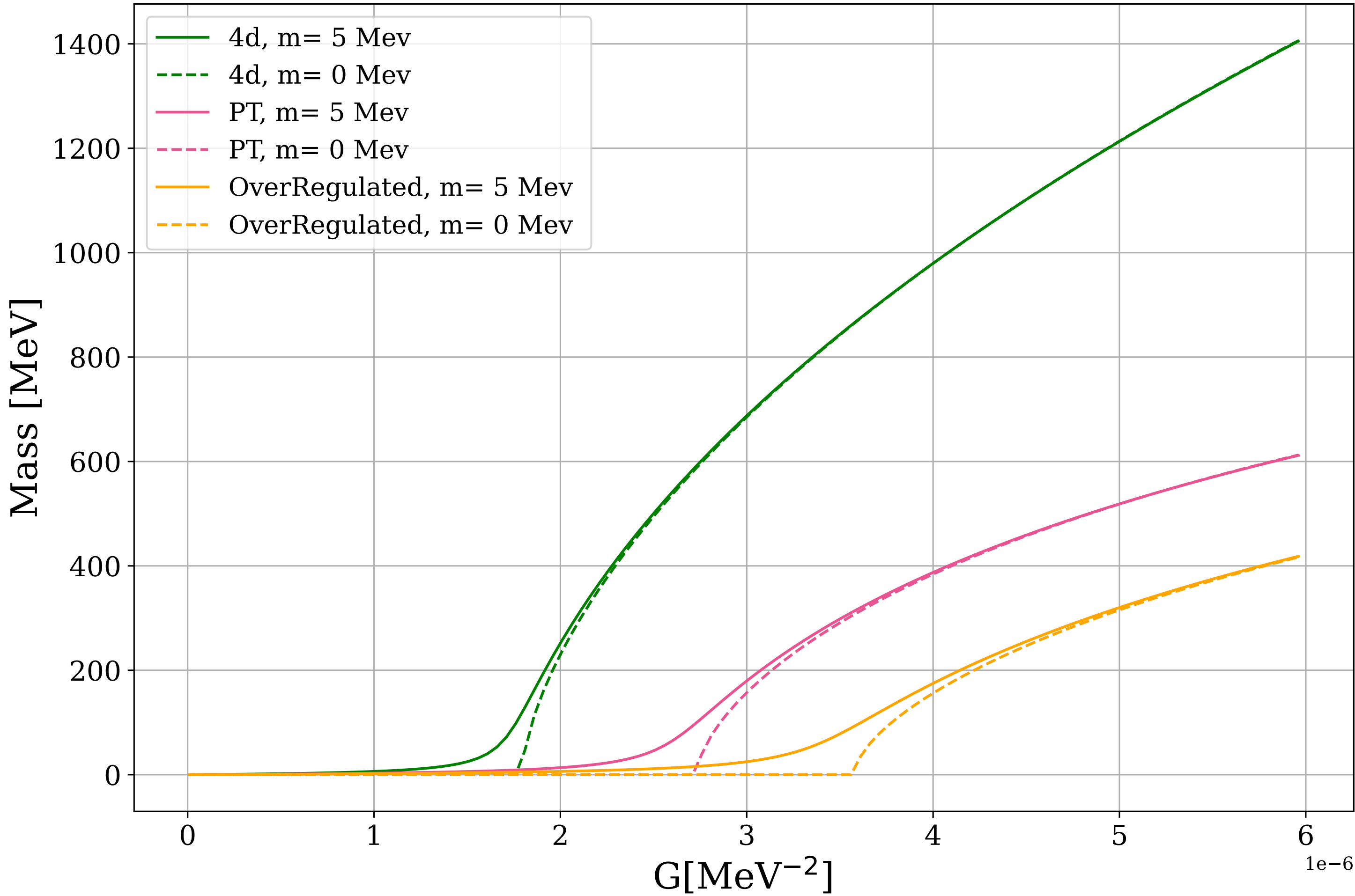
Effects of cutoffs

Gap equations

$$\frac{\partial V_{eff}}{\partial \sigma} = 0$$

	Λ_4	Λ	Λ_s
4D	1027	1027	∞
3D	∞	665	∞
PT	∞	∞	1097
Over	1027	1027	1097

Data from Nuclear Physics B, 896:682–715



Choosing how to integrate

Let's rethink our approach

- In the proper time representation, cartesian coordinates are the easiest to compute!
 - (But they are not the only ones)
 - Using cartesian coordinates, the integration is gaussian
- What would happen if other geometries in p are used?

Geometry matters

Physical Scenario

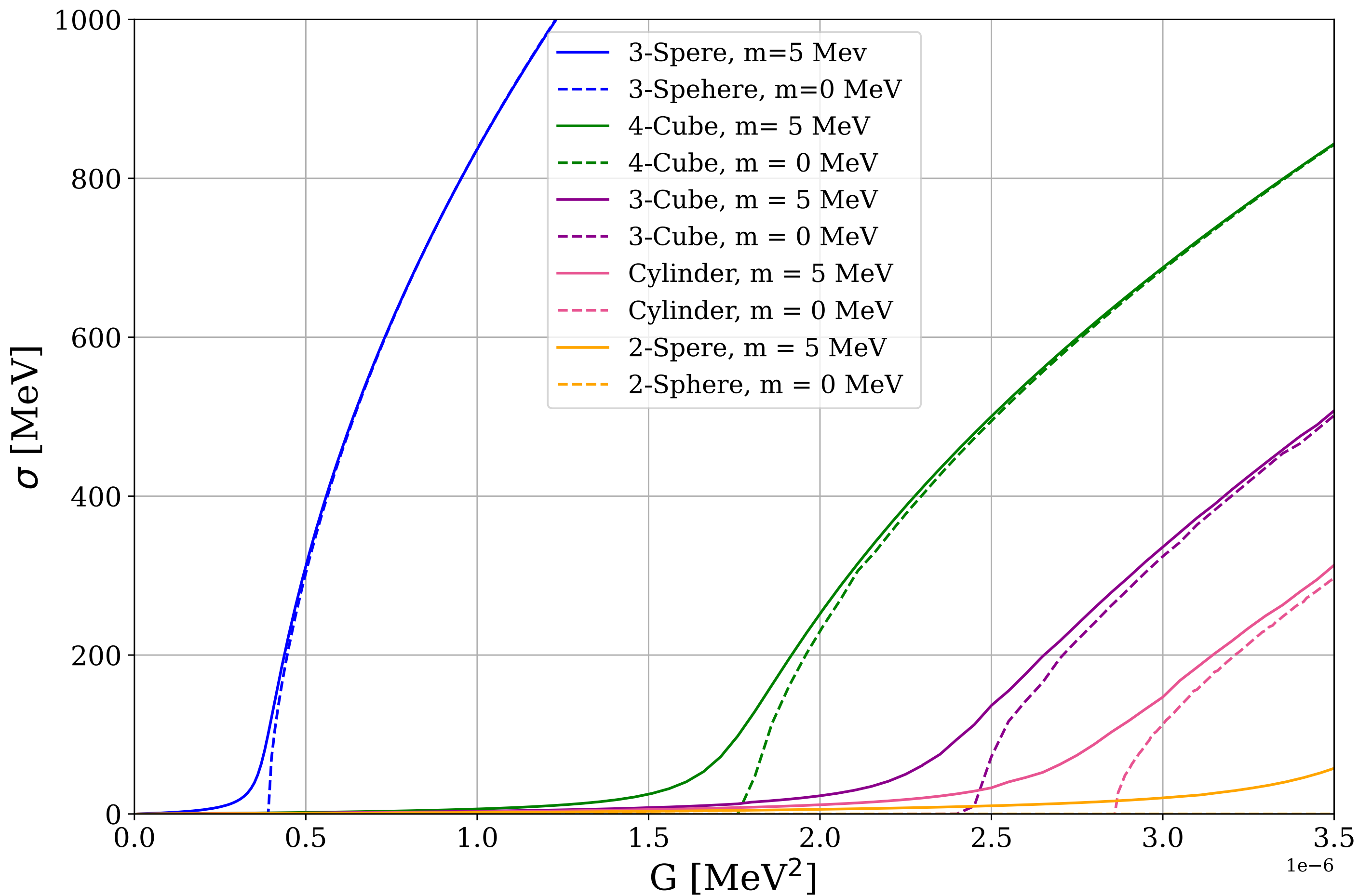
- Different geometries involve different scenarios that every scenario (Landau Level, Matsubara Frequency and normal mode) is taken into account

Volume Element	p^μ Domain
$p^3 dp d\Omega_3$	$p \in [0, \Lambda_{4sphere}]$
$dp_4 dp_1 dp_2 dp_3$	$p_i \in [-\Lambda_{4cube}, \Lambda_{4cube}]$
$dp_4 dp_1 dp_2 dp_3$	$p_4 \in [-\Lambda_4, \Lambda_4], p \in [-\Lambda_{3cube}, \Lambda_{3cube}]$
$dp_4 p^2 dp d\Omega_2$	$p_4 \in [-\Lambda_4, \Lambda_4], p \in [0, \Lambda_{3sphere}]$
$dp_4 dp_3 p dp d\Omega_1$	$p_4 \in [-\Lambda_4, \Lambda_4], p \in [0, \Lambda_{pol}], p_3 \in [-\Lambda, \Lambda]$

Geometry matters

They behave like cutoffs

	Λ_4	Λ	Λ_z
4 Cube	—	1027	—
3-Sphere	—	1027	—
3 Cube	∞	665	—
2 sphere	∞	665	—
Cylinder	∞	665	665



Geometry matters

Symmetry breaking is different

- The symmetry breaking depends on the geometry, not only on the volume
 - If one renormalizes, the divergence is an aspect of little matter
 - We do not have access to counterterms, we have to get the most of the non-divergent quantity .

Temperature

What if we have enough time...

- We introduced a new representation
- To include Temperature the Matsubara formalism must hold

$$p_0 \rightarrow \omega_n = \pi T(2n + 1) \quad \int dp_0 f(p_0) \rightarrow T \sum_{n=-\infty}^{\infty} f(\omega_n)$$

- This assumes thermalization
- And the difference must hold term to term

$$\Omega = \frac{1}{4G} \sigma^2 - \frac{2N_c N_f}{(2\pi)^3} \sum_n \int_0^\infty \frac{ds}{s} \left((2\Lambda)^3 e^{-s} - \pi^{3/2} \frac{e^{-sM^2} \text{Erf}(\sqrt{s}\Lambda)^3}{s^{3/2}} e^{-s\omega_n^2} \right)$$

Temperature

First Approach

- The sum diverges because of the first term
 - We can take the first N terms
- We lose the vacuum-medium separation
 - This is not good
 - If there is not a clear separation, the medium contribution induces unphysical phenomena

[Allen PRD.92.074041](#)

[Tavares PRD.109.016011](#)

[Lopes arXiv:2507.14343](#)

[Duarte PRD.99.016005](#)

Temperature

Separating the medium

- A way we can do this is using the Euler-MacLaurin Formula

$$f(x) = \int_0^\infty \frac{ds}{s} \left((2\Lambda)^3 e^{-s} - \left(\frac{\pi}{s} \right)^{3/2} \text{Erf}(\sqrt{s}\Lambda)^3 e^{-sM^2} e^{-s(\pi^2 T^2)(2x+1)^2} \right)$$

$$\sum_{n=a}^b f(n) = \boxed{\int_a^b f(x) dx} + \boxed{\frac{1}{2} (f(a) + f(b)) + \sum_{k=1}^m \frac{B_{2k}}{(2k)!} (f^{(2k-1)}(b) - f^{(2k-1)}(a)) + R_m}$$

Vacuum
Divergent!

Medium
Well Behaved

$$\Omega = \Omega_{tree} + \Omega_{vacuum}^{(1)}(\Lambda) + \Omega_{medium}^{(1)}(\Lambda)$$

Temperature

Separating the medium

- The medium still depends on the cutoff, although we can still reproduce some of the physics.
- This cutoff dependance disappears when taking the gap equation or other derivatives so the limit $\Lambda \rightarrow \infty$ can be taken.

Temperature Comparison

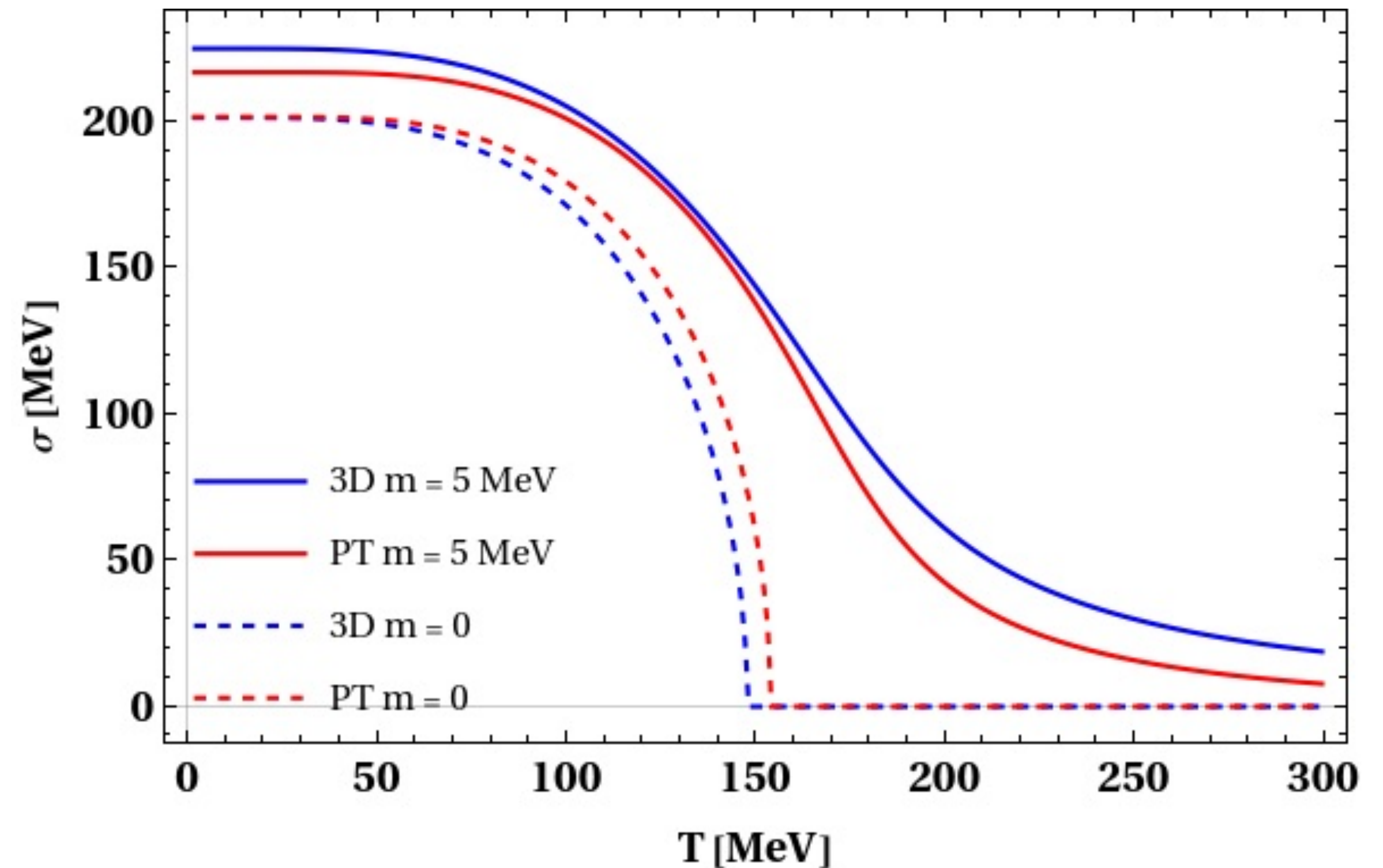
$$\Omega = \frac{1}{4G}\sigma^2 - \frac{2N_c N_f}{(2\pi)^3} \sum_n \int_0^\infty \frac{ds}{s} \left((2\Lambda)^3 e^{-s} - \pi^{3/2} \frac{e^{-sM^2} \text{Erf}(\sqrt{s}\Lambda)^3}{s^{3/2}} e^{-s\omega_n^2} \right)$$

$$\Omega = \frac{\sigma^2}{4G} + \frac{N_c N_f T}{4\pi^{3/2}} \int_{1/\Lambda_s^2}^\infty \frac{ds}{s^{5/2}} e^{-sM^2} \theta_2 \left(0, e^{-4\pi^2 T^2 s} \right).$$

Euler-Maclaurin
At sufficiently high N

The behavior is the same, and T_c varies a little

$$\begin{aligned} G_{PT} &= 3.14 \times 10^{-6} \text{MeV}^{-2} & \Lambda_{PT} &= 1097 \text{MeV} \\ G_{3D} &= 2.72 \times 10^{-6} \text{MeV}^{-2} & \Lambda_{3D} &= 665 \text{MeV} \end{aligned}$$



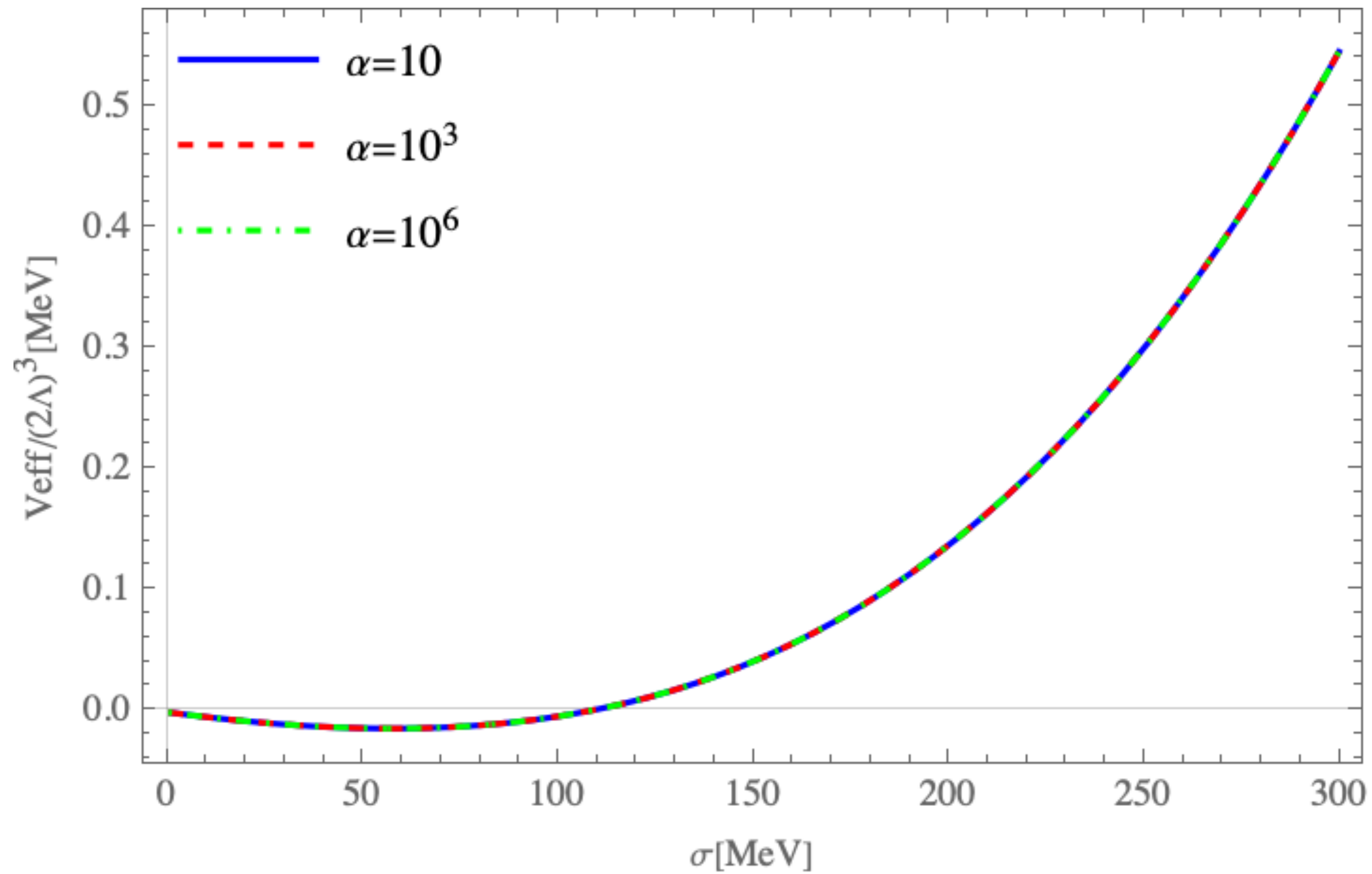
Summary and Conclusions

- We showed the effects of interchanging divergences and when working on a nonrenormalizable model.
 - We also showed that the regularization scheme takes a toll on the parameters.
- There is also a strong dependence on the way we integrate, so the geometry and the physical scenario gain more importance.

Thank you!
¡Gracias!
Grazie!

Backup

Dependence of a dimension full parameter



$$\int_{-\Lambda_p}^{\Lambda_p} \frac{dp_1 dp_2 dp_3 dp_4}{(2\pi)^4} \int_0^\infty \frac{ds}{s} \left(\boxed{e^{-s\alpha^2}} - e^{-s(p_4^2 + \omega^2 + M^2)} \right)$$