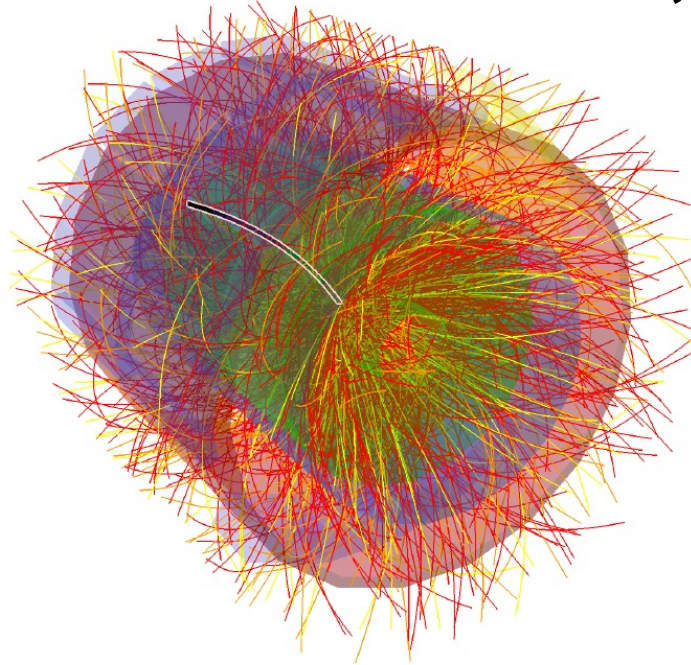


# Production of loosely-bound objects in heavy-ion collisions at HADES, RHIC and LHC



**April 14, 2025**

**GSI**

**Physics of High net-baryon Densities - PHD 2025**

**Benjamin Dönigus**

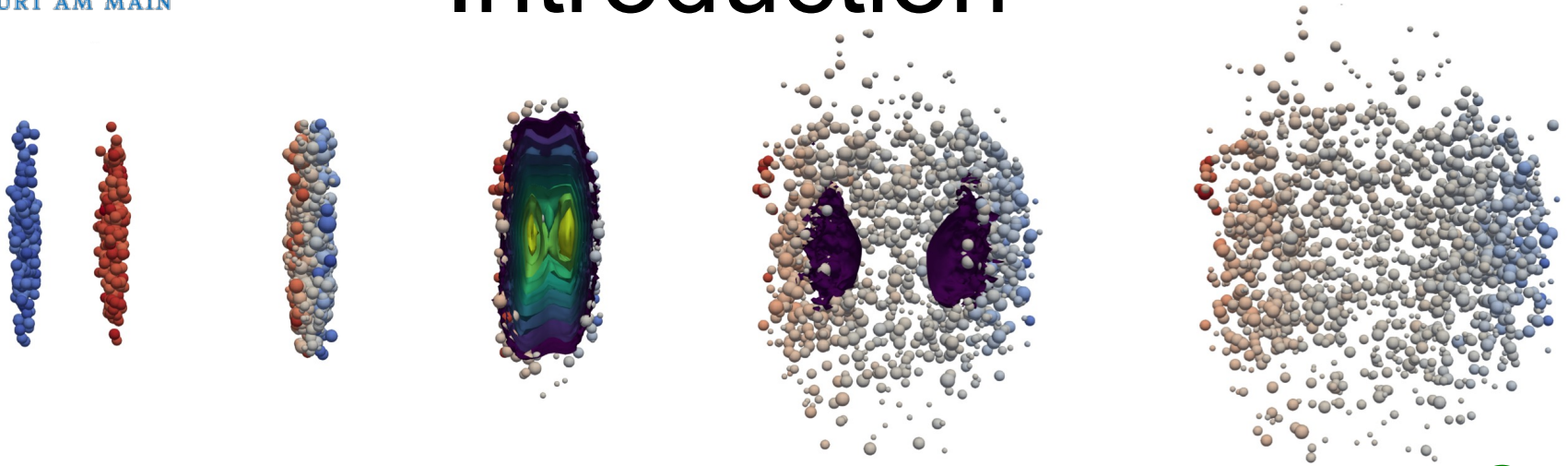
**Institut für Kernphysik**

**Goethe Universität Frankfurt**

# Content

- Introduction
- Nuclei and Exotica
  - (Anti-)nuclei
  - (Anti-)hypernuclei
- Summary & Outlook

# Introduction

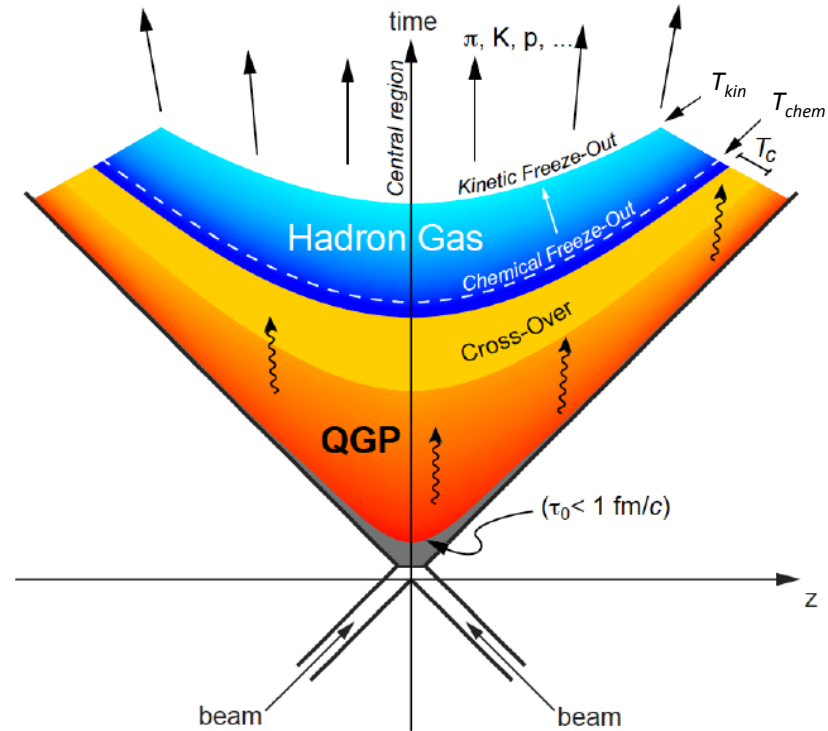


Cartoon of a Ultra-relativistic heavy-ion collision

Left to right:

- the two Lorentz contracted nuclei approach,
- collide,
- form a Quark-Gluon Plasma (QGP),
- the QGP expands and hadronizes,
- finally hadrons rescatter and freeze

# Introduction

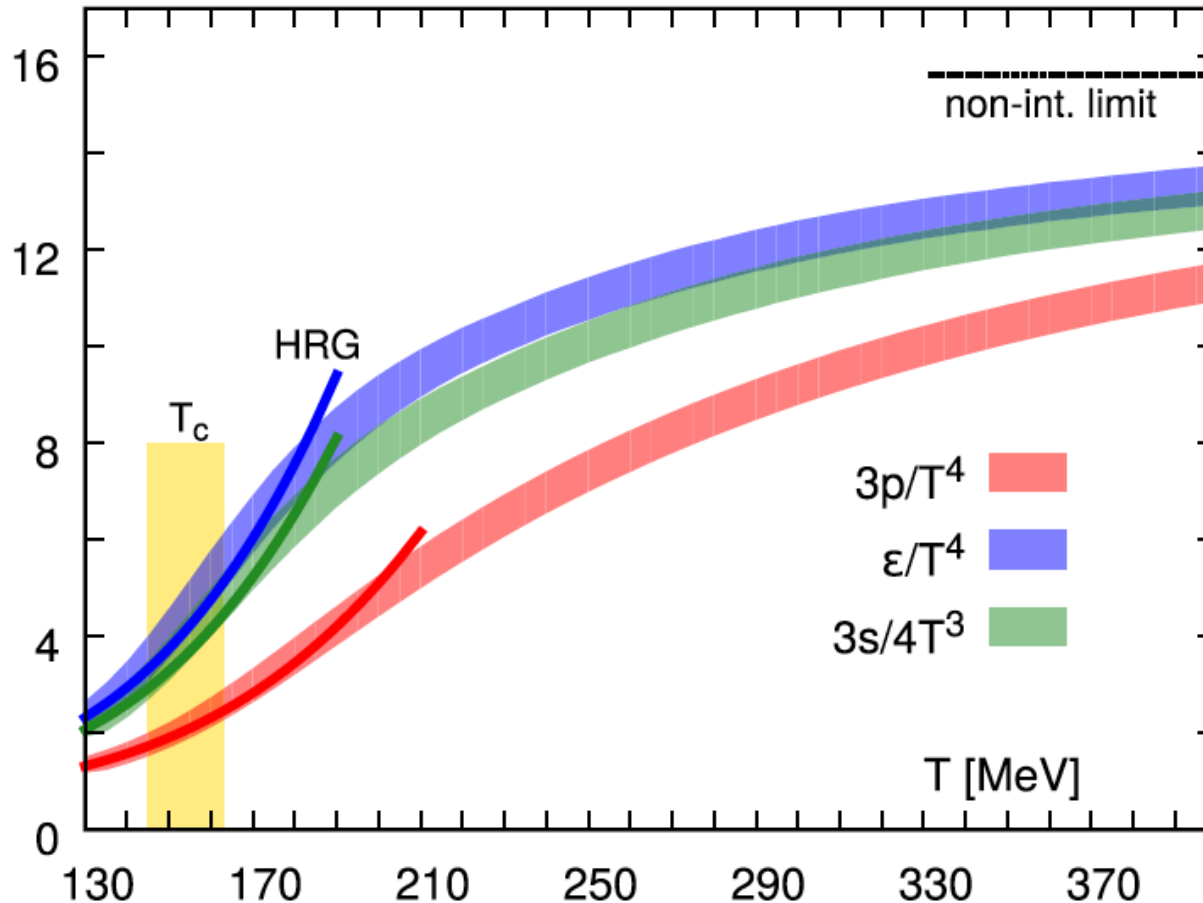


The fireball evolution:

- Starts with a “pre-equilibrium state”
- Forms a Quark-Gluon Plasma phase (if  $T$  is larger than  $T_c$ )
- At *chemical freeze-out*,  $T_{\text{ch}}$ , *hadrons stop being produced*
- At *kinetic freeze-out*,  $T_{\text{fo}}$ , *hadrons stop scattering*



# Lattice QCD results



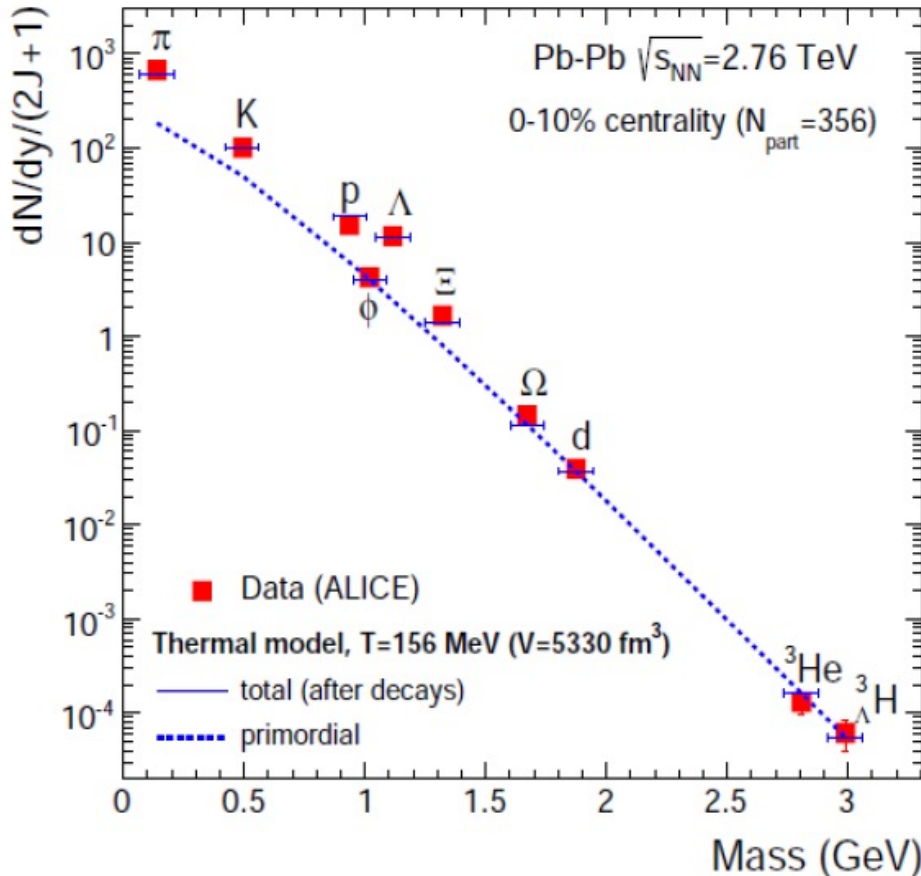
Lattice QCD  
tells us where  
to expect the  
phase  
transition

Critical energy density:  
 $\epsilon_c = 0.34 \pm 0.16 \text{ GeV/fm}^3$

Critical temperature  
 $T_c = (157 \pm 2) \text{ MeV}$

*A. Bazavov et al. (hotQCD) Phys. Rev. D90 (2014) 094503 & PLB 795 (2019) 15  
Similar results from Budapest-Wuppertal group: S. Borsányi et al. JHEP 09 (2010) 073  
& PRL 125 (2020) 052001*

# Temperature of the source



Analogy:

Light source → particle source

- Multiplicity described best with  
 $T = 1\,900\,000\,000\,000$  °C  
(1,9 trillion degree centigrade)

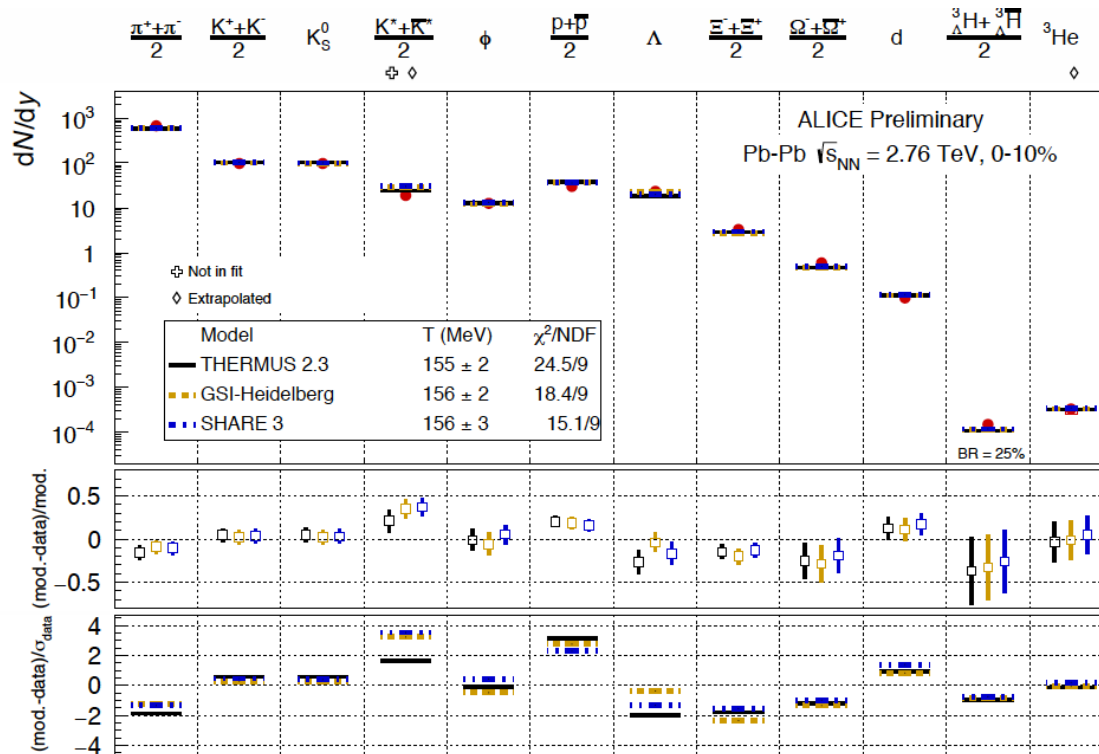
→ 100 000 times hotter than in the  
interior of the sun!

1/40 eV = 20 °C

Plot by A. Andronic, GSI-Heidelberg group  
[arXiv:1407.5003 \[nucl-ex\]](https://arxiv.org/abs/1407.5003)

# Thermal model

- Statistical (thermal) model with only three parameters able to describe particle yields (grand canonical ensemble)

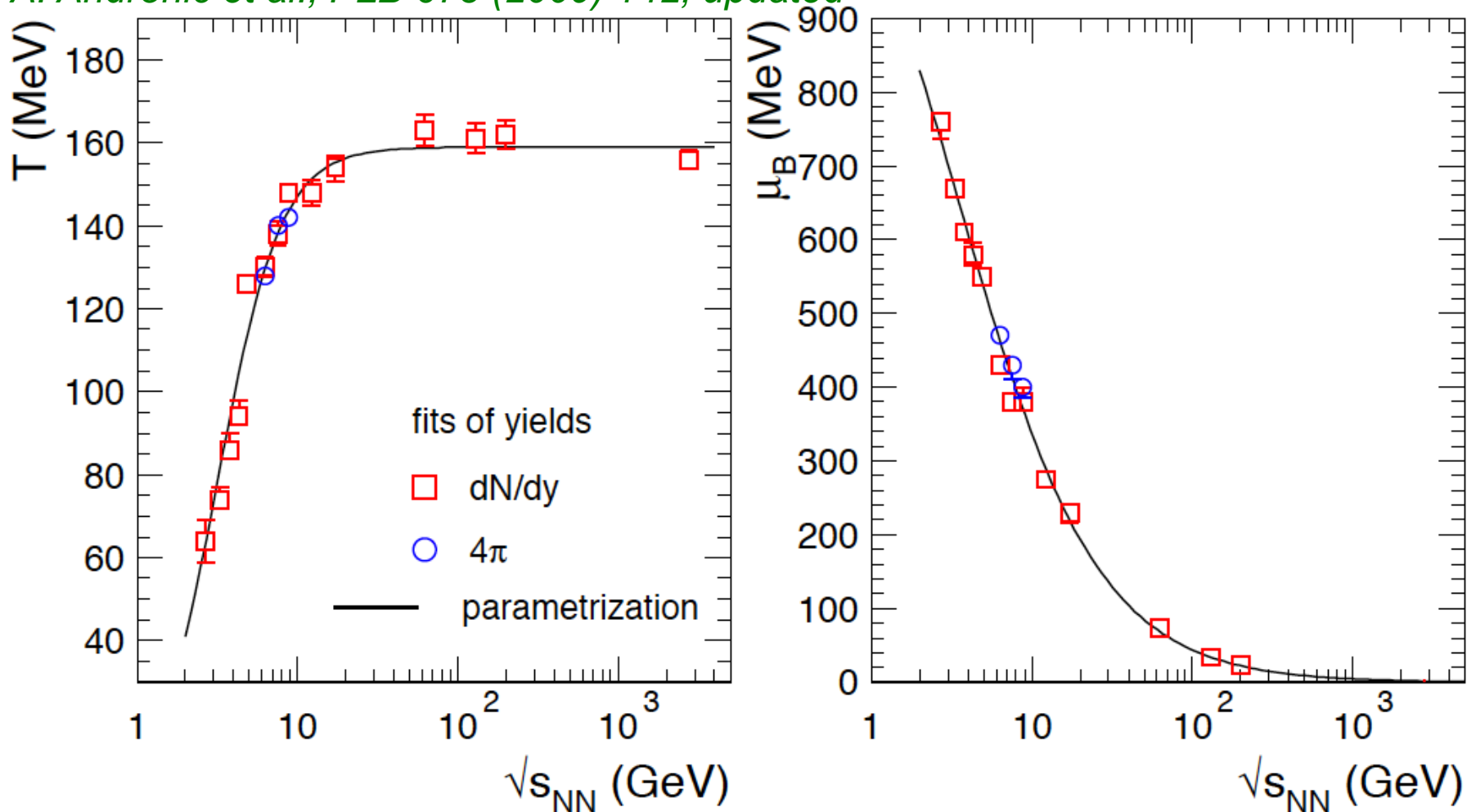


- chemical freeze-out temperature  $T_{ch}$
- baryo-chemical potential  $\mu_B$
- Volume  $V$

→ Using particle yields as input to extract parameters

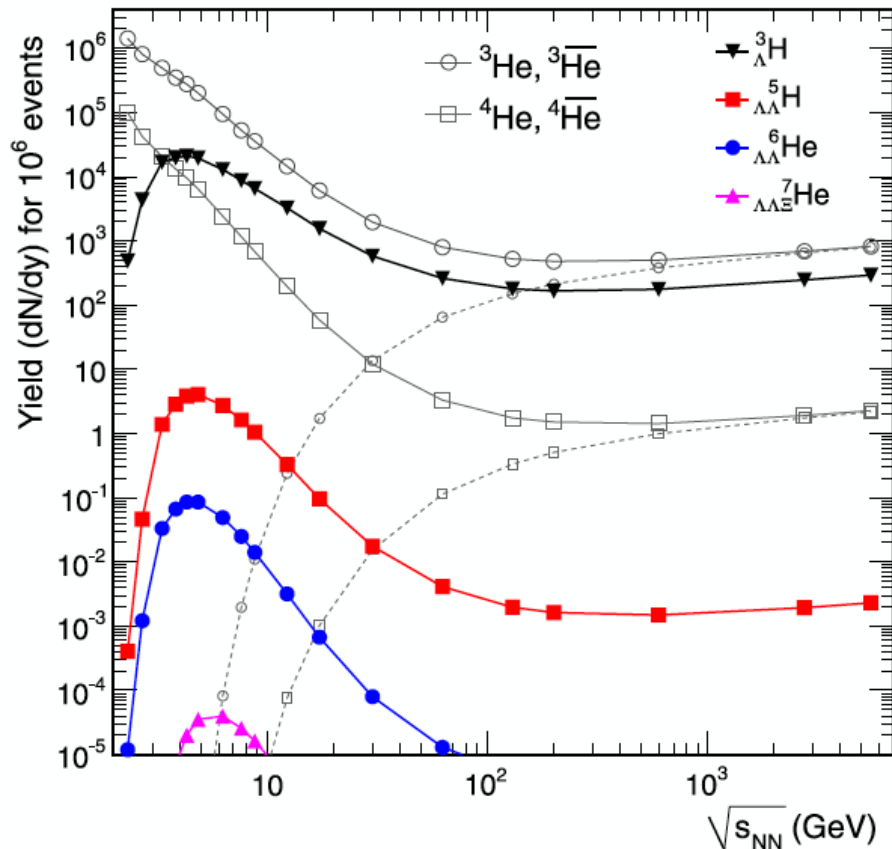
# Energy dependence

*A. Andronic et al., PLB 673 (2009) 142, updated*



Thermal model fits show limiting temperature:  $T_{lim} = (159 \pm 2) \text{ MeV}$

# Predicting yields of bound states



*A. Andronic et al., PLB 697 (2011) 203*

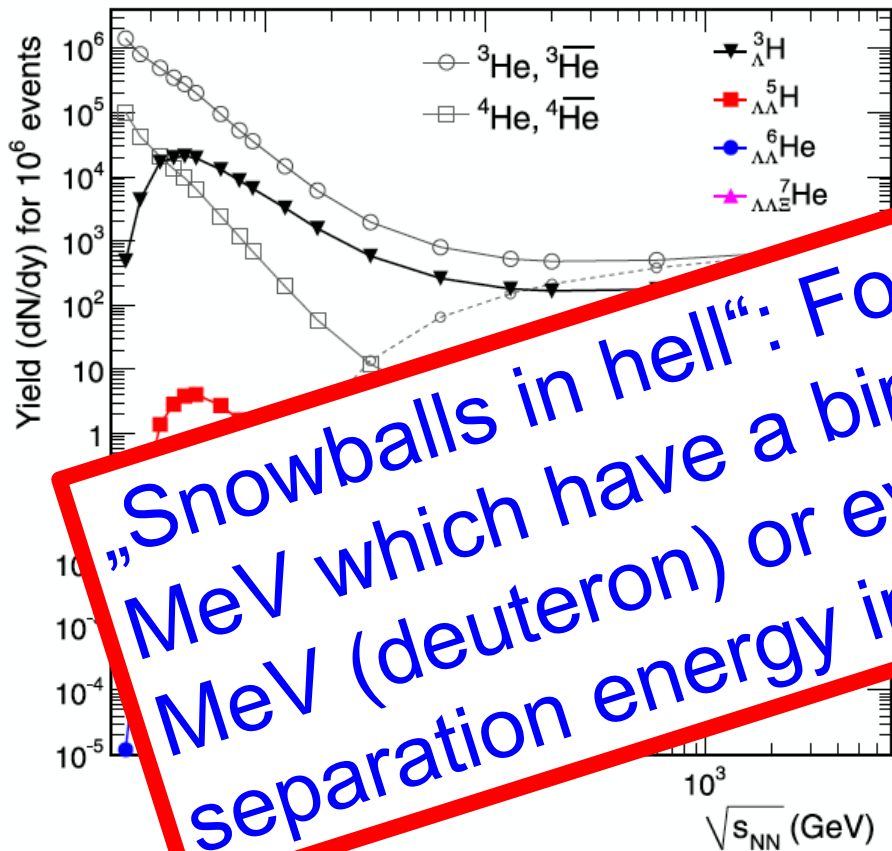
Key parameter at LHC energies:

chemical freeze-out temperature  $T_{\text{ch}}$

Strong sensitivity of abundance of nuclei to choice of  $T_{\text{ch}}$  due to:

1. large mass  $m$
  2. exponential dependence of the yield  $\sim \exp(-m/T_{\text{ch}})$
- Binding energies small compared to  $T_{\text{ch}}$

# Predicting yields of bound states



Key parameter at high energies

"Snowballs in hell": Forming objects at 150 MeV which have a binding energy of only 2.2 MeV (deuteron) or even only 130 keV ( $\Lambda$  separation energy in the hypertriton)

A. Andronic et al., PLB 697 (2011) 203

of  $T_{ch}$  due to:

1. large mass  $m$
  2. exponential dependence of the yield  $\sim \exp(-m/T_{ch})$
- Binding energies small compared to  $T_{ch}$

# Snowballs in hell

*Recipe for a Hot Universe* 93

- Already mentioned in S. Weinberg's book: The first three minutes

quantities from the conditions for thermal equilibrium - for instance, we can make the density of water molecules plus hydrogen ions a little greater or less than  $3.3 \times 10^{22}$  molecules per cubic centimetre by raising or lowering the pressure - so we need to specify them in order to know what is in our glass.

This example also helps us to understand the shifting meaning of what we call 'conserved' quantities. For instance, if our water is at a temperature of millions of degrees, as inside a star, then it is very easy for molecules or ions to dissociate, and for the constituent atoms to lose their electrons. The conserved quantities are then the numbers of electrons and of oxygen and hydrogen nuclei. The density of water molecules plus hydroxyl atoms under these conditions has to be *calculated* from the rules of statistical mechanics rather than *specified in advance*; of course it turns out to be quite small. (Snowballs are rare in hell.) Actually, nuclear reactions do occur under these conditions, so even the numbers of nuclei of each species are not absolutely fixed, but these numbers change so slowly that a star can be regarded as evolving gradually from one equilibrium state to another.



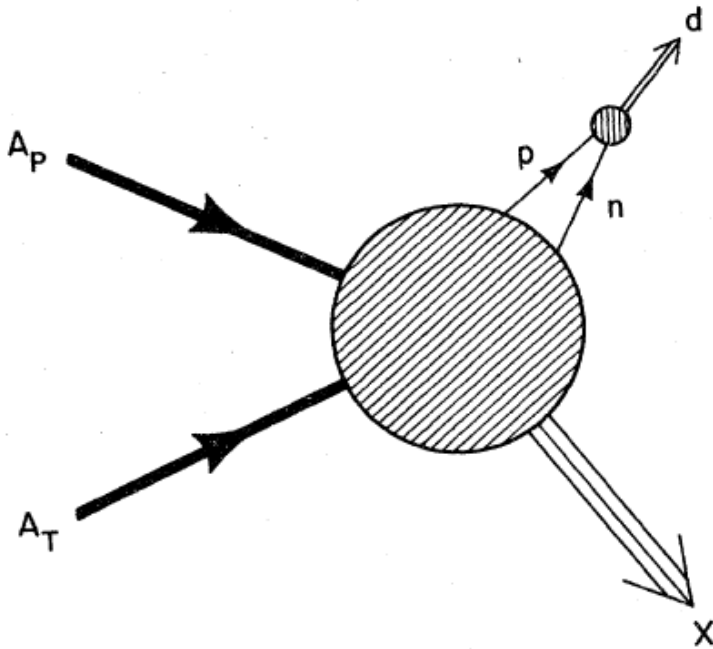
# AGS predictions

*P. Braun-Munzinger, J. Stachel, J. Phys. G: Nucl. Part. Phys. 21 (1995) L17*

- Already at AGS the predictions from thermal model and coalescence were very close
- Looking at production yields alone might not be enough to disentangle models

Particles	Thermal model		Coalescence model
	$T = 0.120 \text{ GeV}$	$T = 0.140 \text{ GeV}$	
d	15	19	11.7
t+ <sup>3</sup> He	1.5	3.0	0.8
$\alpha$	0.02	0.067	0.018
H <sub>01</sub>	0.09	0.15	0.07
${}^5_{\Lambda\Lambda}\text{H}$	$3.5 \times 10^{-5}$	$2.3 \times 10^{-4}$	$4 \times 10^{-4}$
${}^6_{\Lambda\Lambda}\text{He}$	$7.2 \times 10^{-7}$	$7.6 \times 10^{-6}$	$1.6 \times 10^{-5}$
${}^7_{\Xi^0\Lambda\Lambda}\text{He}$	$4.0 \times 10^{-10}$	$9.6 \times 10^{-9}$	$4 \times 10^{-8}$
${}^{10}\text{Si}^{-8}$	$1.6 \times 10^{-14}$	$7.3 \times 10^{-13}$	
${}^{12}\text{Si}^{-9}$	$1.6 \times 10^{-17}$	$1.7 \times 10^{-15}$	
${}^{14}\text{Si}^{-11}$	$6.2 \times 10^{-21}$	$1.4 \times 10^{-18}$	
${}^{16}_2\text{Si}^{-13}$	$2.4 \times 10^{-24}$	$1.2 \times 10^{-21}$	
${}^{20}_2\text{Si}^{-16}$	$9.6 \times 10^{-31}$	$2.3 \times 10^{-27}$	

# Coalescence



*J. I. Kapusta, PRC 21, 1301 (1980)*

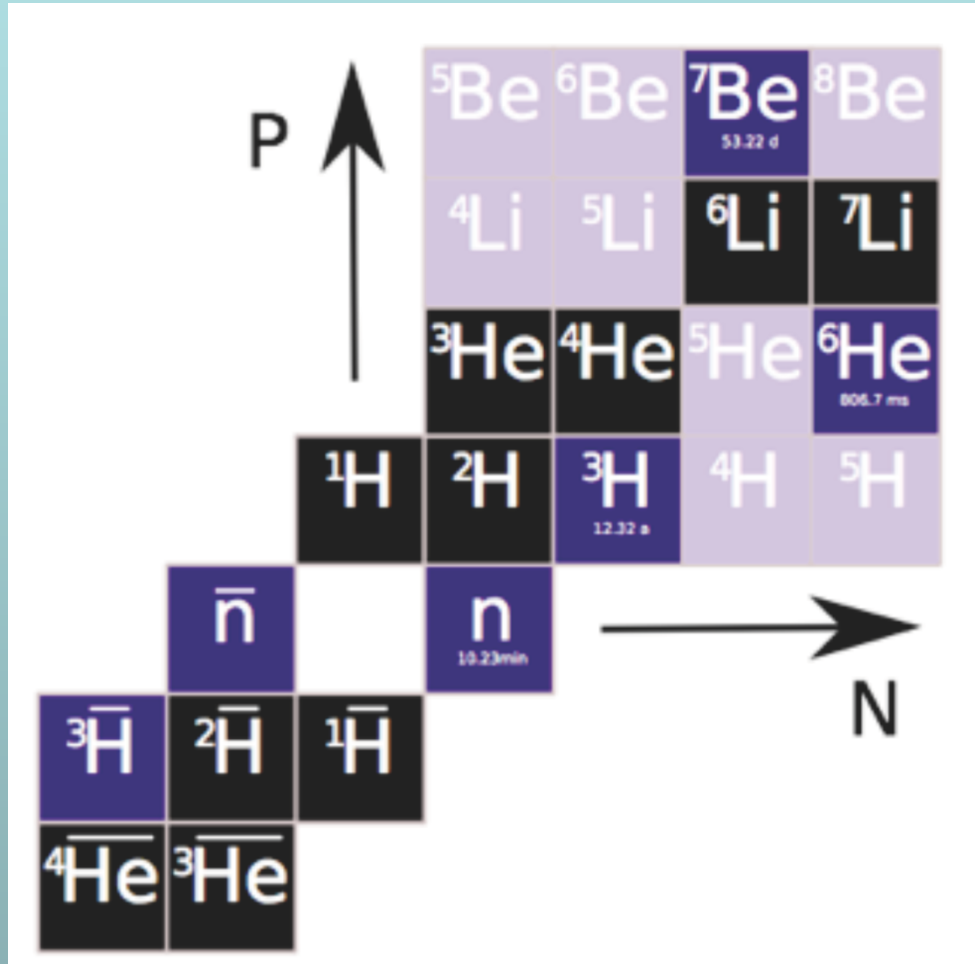
Nuclei are formed by protons and neutrons which are nearby and have similar velocities (after kinetic freeze-out)

Produced nuclei

→ can break apart

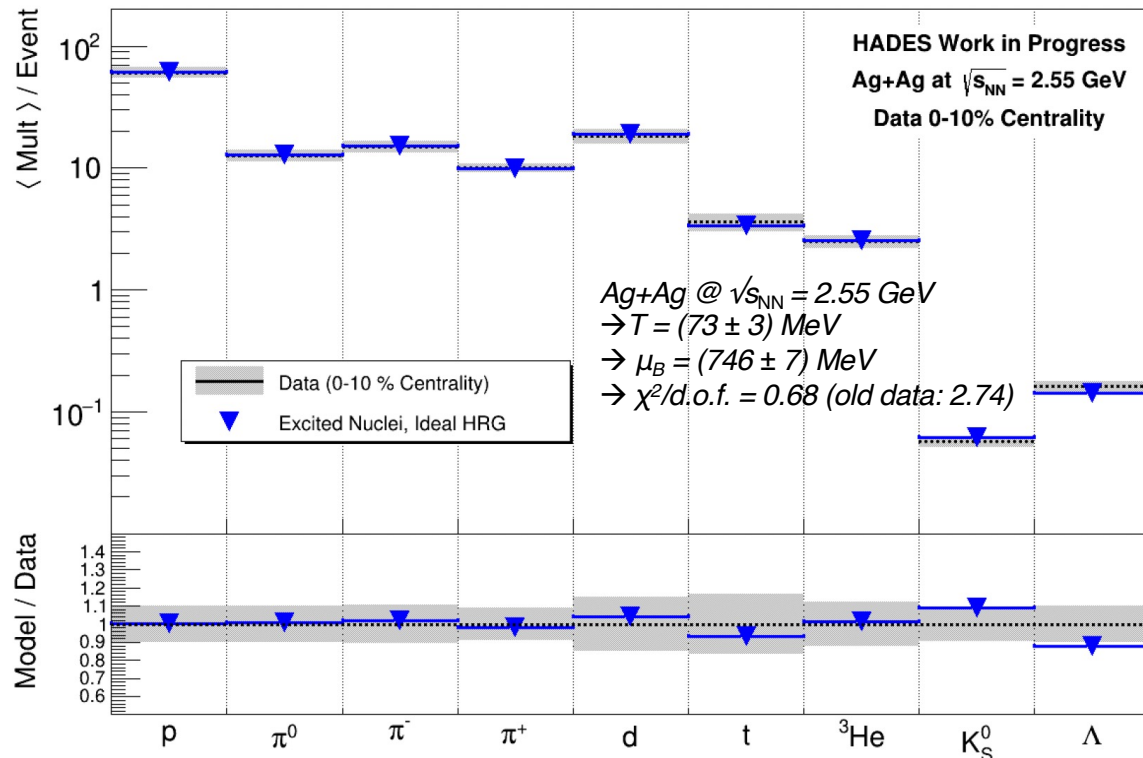
→ created again by final-state coalescence

# (Anti-)Nuclei



# Thermal model: HADES

- For HADES Ag+Ag data the thermal model fit describes the production of all investigated yields, including nuclei
- Rather good description

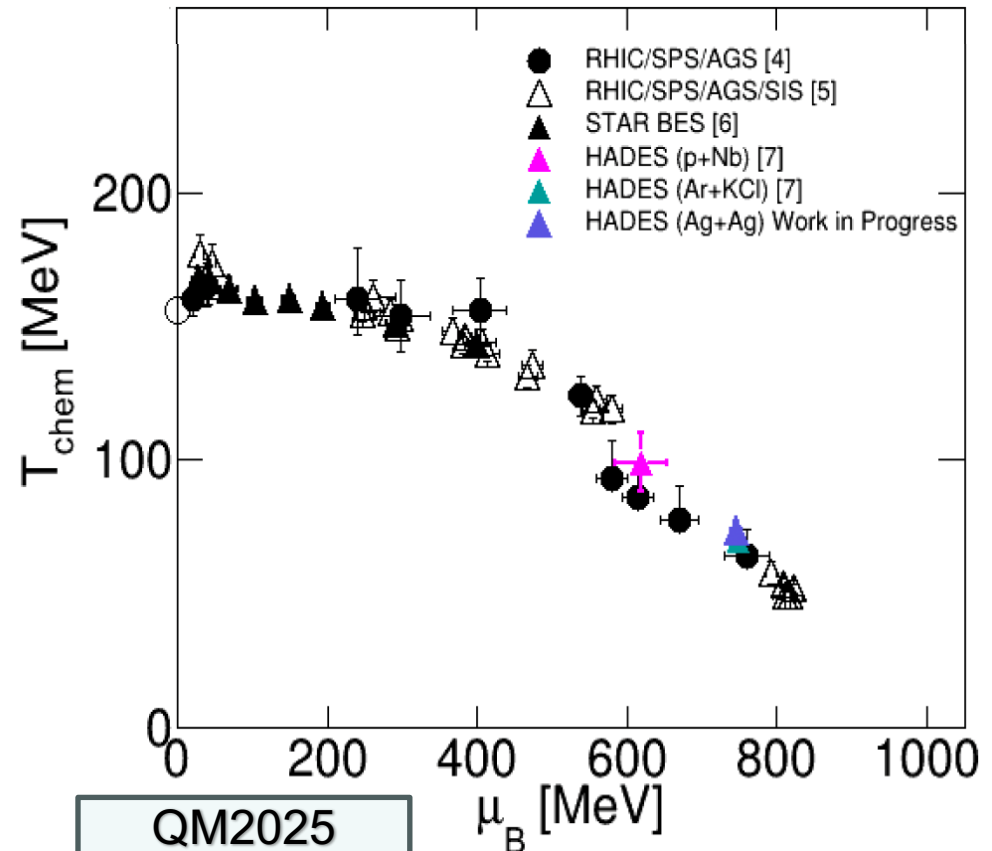


QM2025  
Talk #577  
S. Spies

QM2025  
Poster #576  
M. Kohls

# Phase diagram

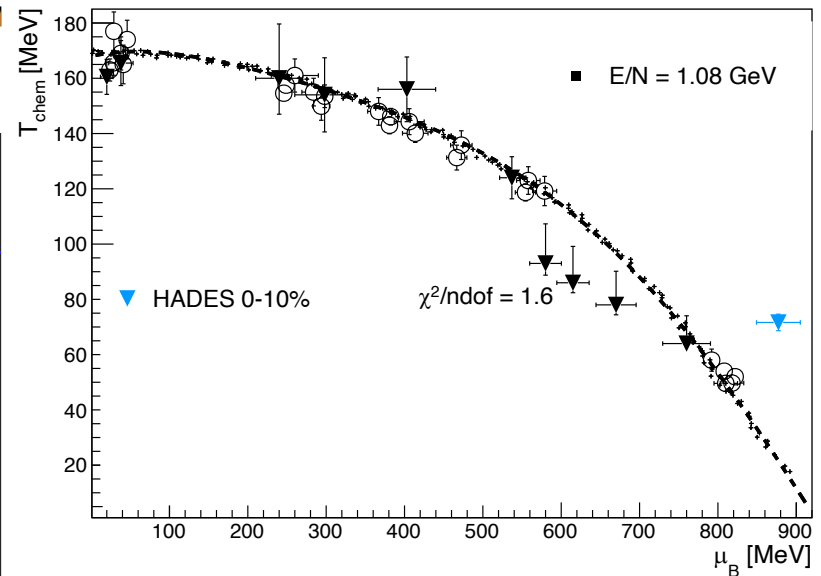
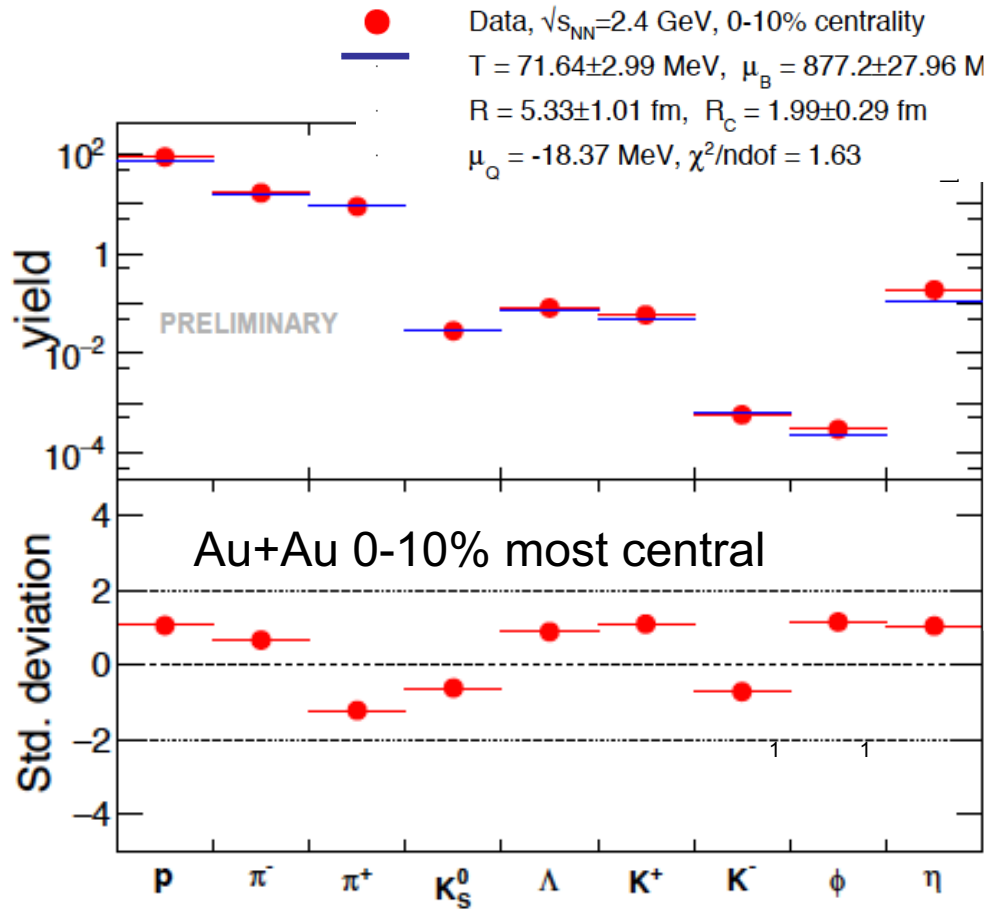
- For HADES Ag+Ag data the thermal model fit describes the production of all investigated yields, including nuclei
- Rather good description
- Including nuclei leads to nice agreements with the trend of the world data



QM2025  
Poster #576  
M. Kohls

# Phase diagram: HADES

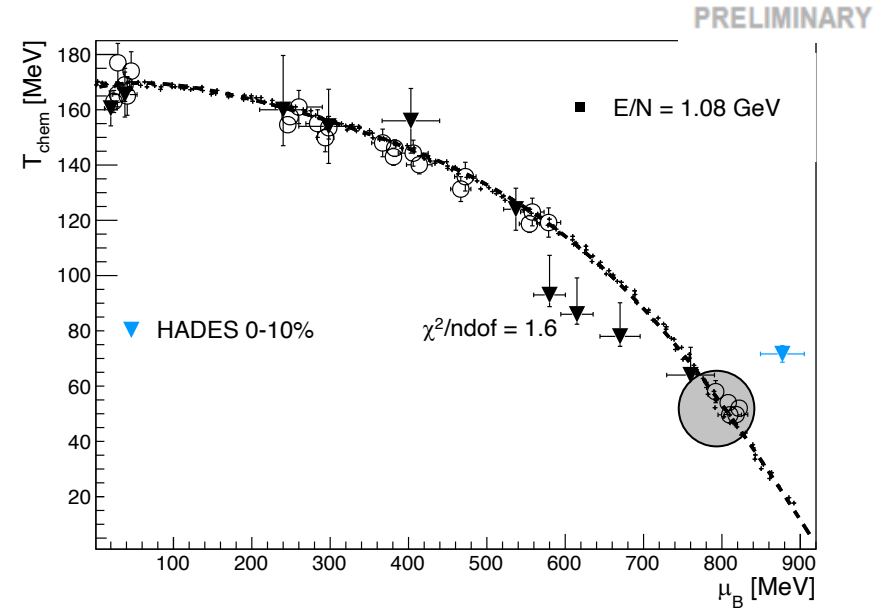
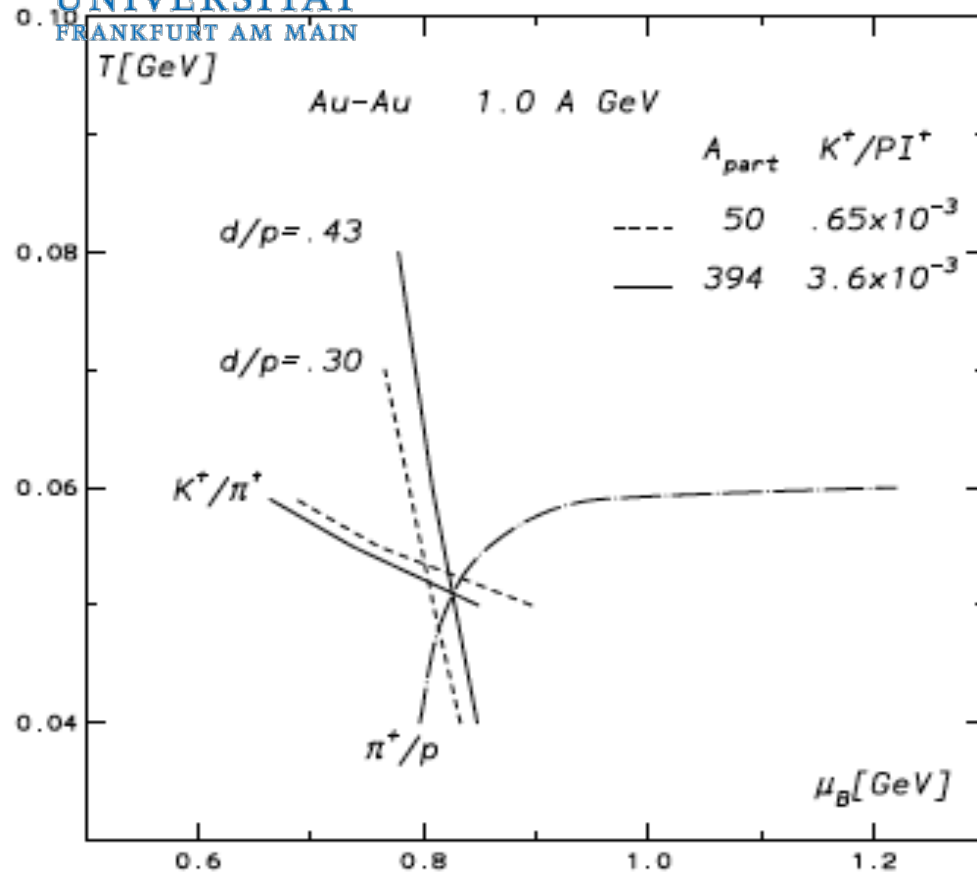
PRELIMINARY



- Freeze-out point stays at higher  $T$  and  $\mu_B$  also for 0-10% most central events

*M. Lorenz 2019*

# Phase diagram: HADES



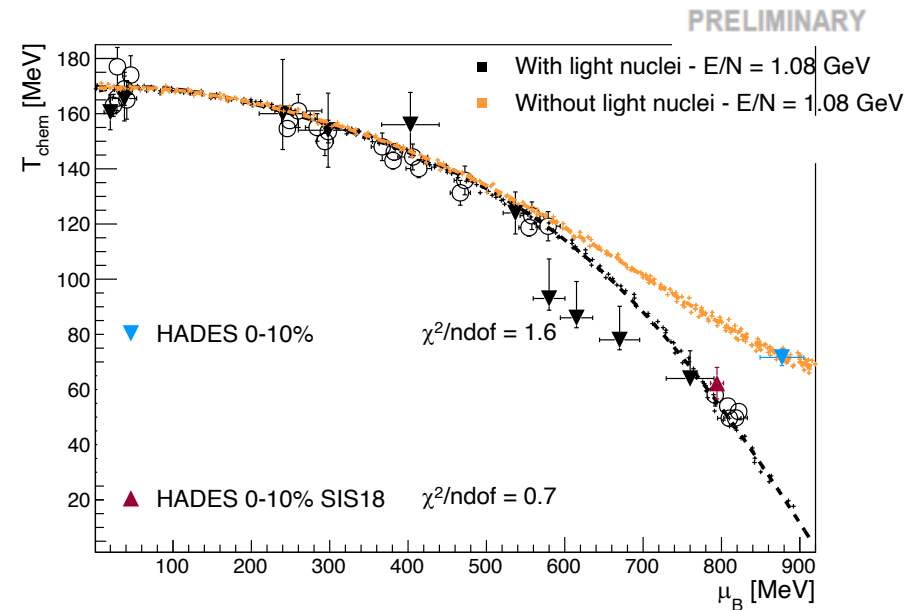
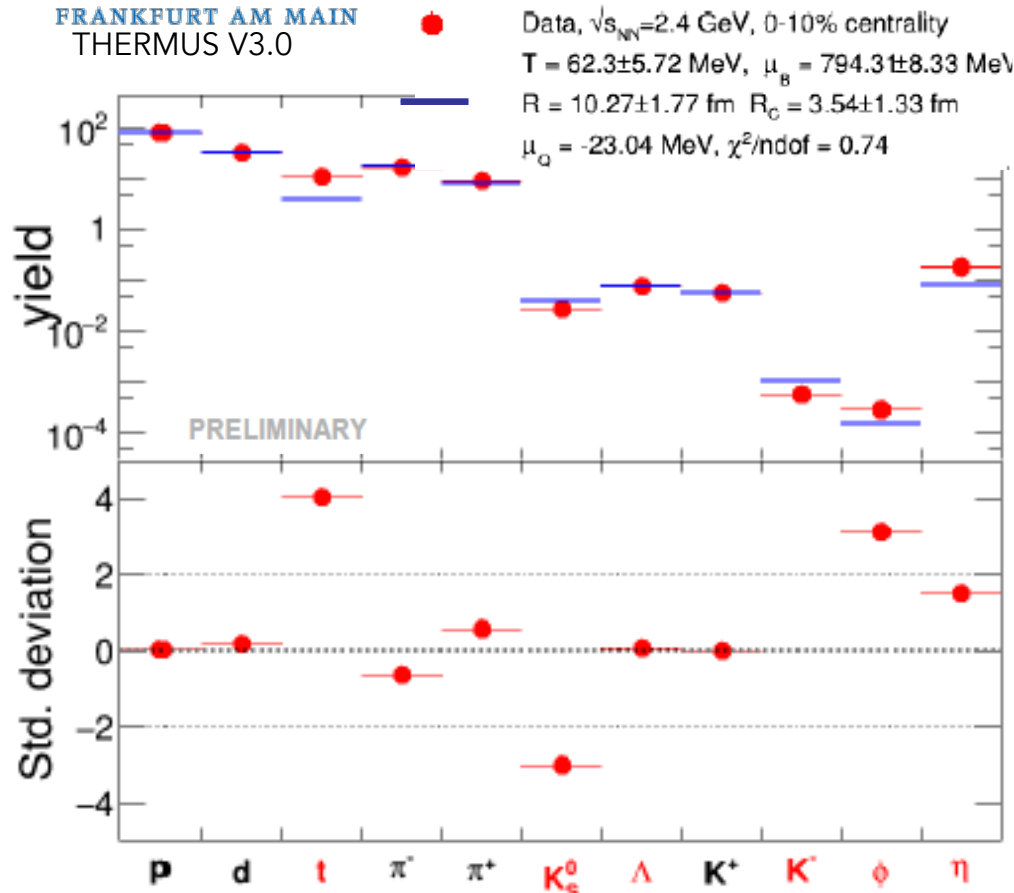
J. Cleymans, H. Oeschler, K. Redlich, Phys.Rev. C59 (1999))

- Freeze-out points previously estimated based on ratios of  $p$ ,  $d$ ,  $K^+$ ,  $\pi^+$
- Light nuclei are not included in Thermus V2.3
- Switch to Thermus V3.0 or Thermal-Fist

*M. Lorenz 2019*



# Phase diagram: HADES



- Fit to HADES data consistent with previous works when same hadron yields are used
- $E/N=1.08$  GeV with or without light nuclei?
- Light nuclei are important to define chemical freeze-out line at high  $\mu_B$ .

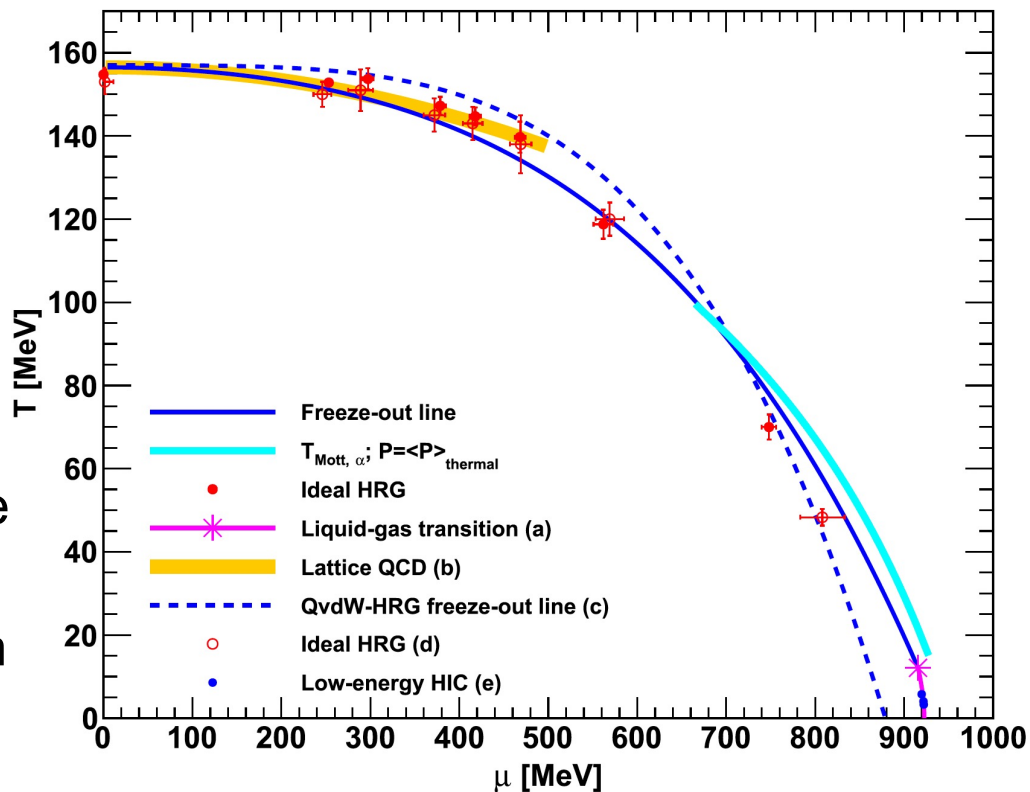
J. Cleymans, H. Oeschler, K. Redlich, Phys.Rev. C59 (1999))

- Freeze-out points previously estimated based on ratios of p, d,  $K^+$ ,  $\pi^+$
- Light nuclei are not included in Thermus V2.3
- Switch to Thermus V3.0 or Thermal-Fist

*M. Lorenz 2019*

# Phase diagram: Mott $\alpha$

- If one maps the energy dependence onto a  $\mu_B$ - $T$  plane one gets experimental access to the QCD phase diagram
- Using parametrizations of the thermal model fits one can get an approximation of the freeze-out curves
- Particular models can even describe their shape, e.g. Mott curves

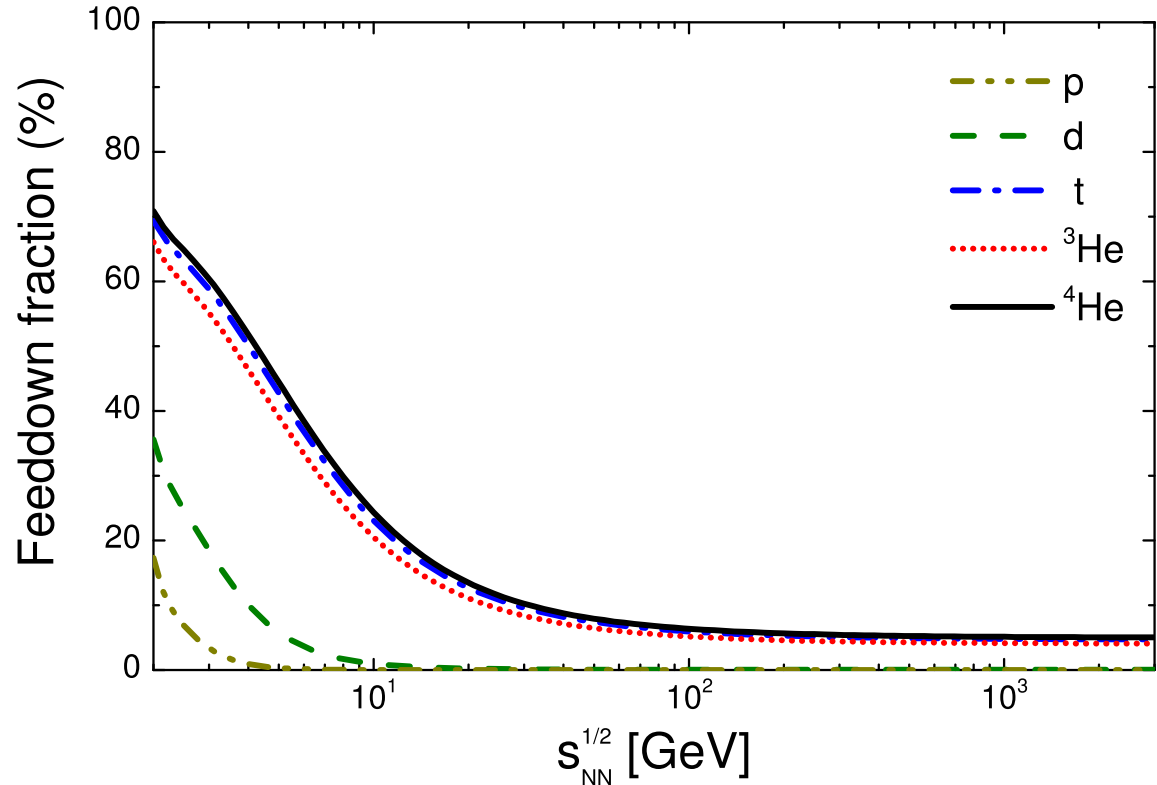


*D. Blaschke, S. Liebing, G. Röpke, BD,  
Phys. Lett. B 860 (2025) 139206*

Poster #1096  
D. Blaschke et al.

# Thermal model

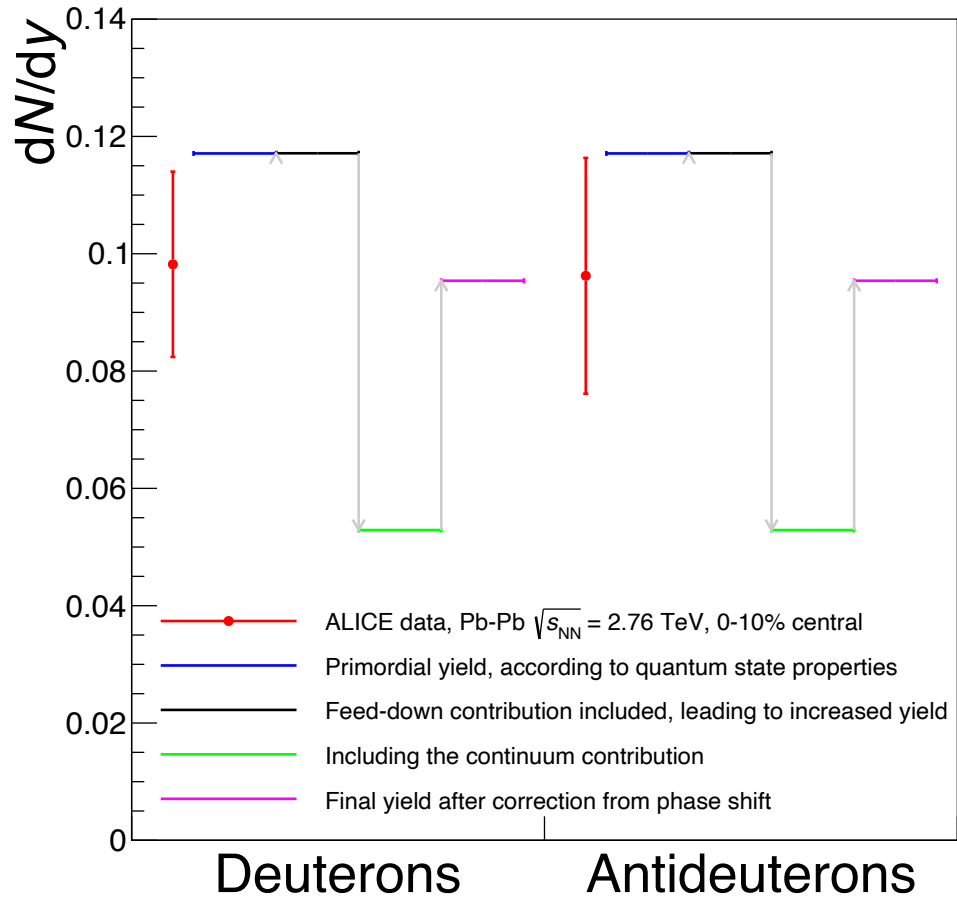
- For the thermal model description of production yields, feed-down is an important ingredient
- All light hadron production yields are populated strongly by resonances
- Seems to not be the case for (hyper-) nuclei at the LHC



*V. Vovchenko, BD, B. Kardan, M. Lorenz,  
H. Stoecker, Phys.Lett.B 809 (2020) 135746*

# Thermal model

- For the thermal model description of production yields, feed-down is an important ingredient
- All light hadron production yields are populated strongly by resonances
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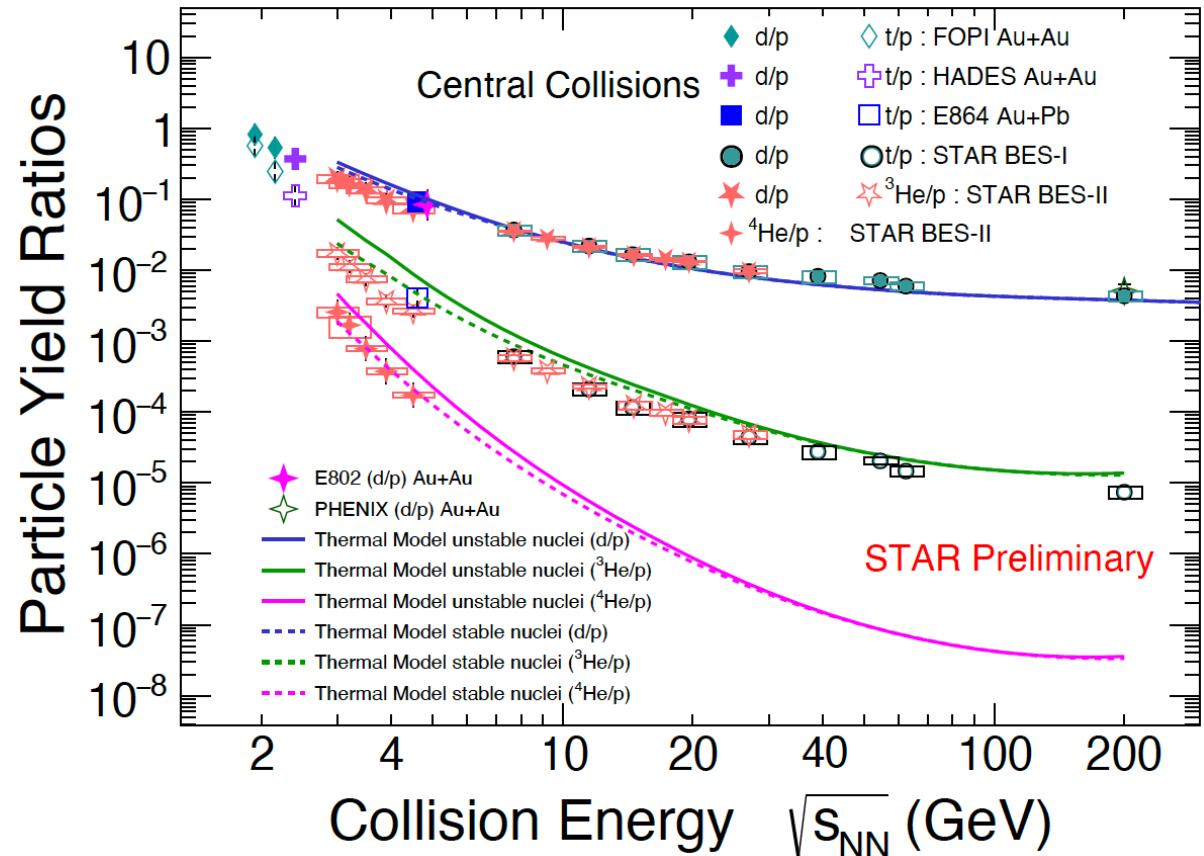


BD, G. Röpke, D. Blaschke,  
*Phys. Rev. C* 106 (2022) 044908

A. Andronic et al., *Phys.Lett.B* 797 (2019) 134836;  
*Nature* 561 (2018) 7723, 321; *Phys.Lett.B* 697  
(2011) 203; *Phys.Lett.B* 792 (2019) 304

# Light nuclei: yield ratios

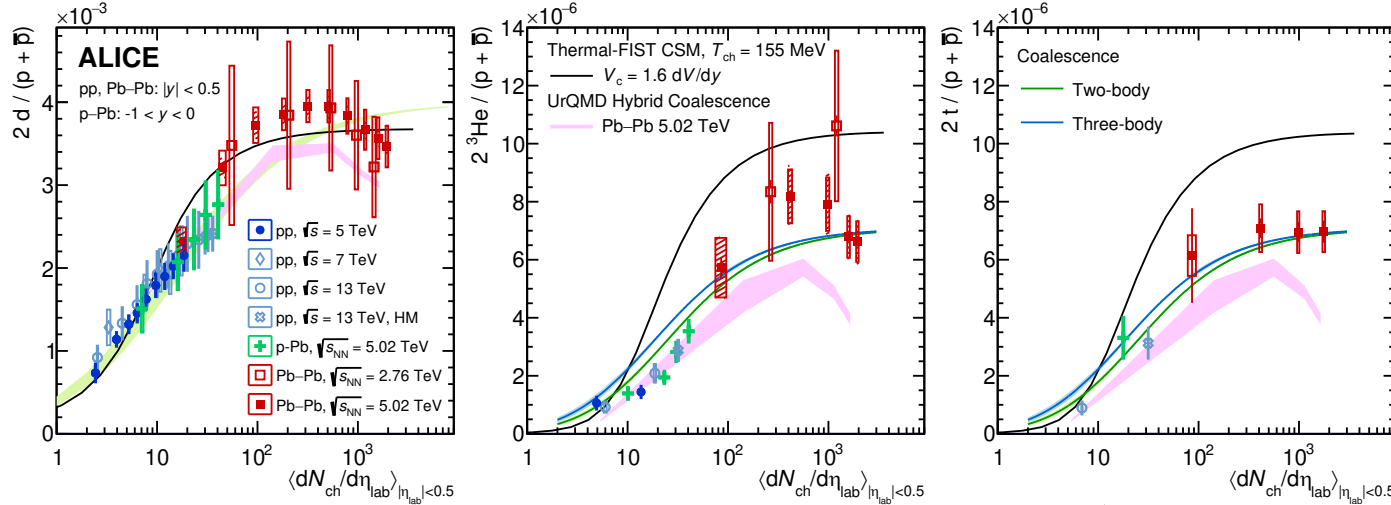
- A/p ratio:  
better  
described by  
thermal model  
with only  
stable nuclei



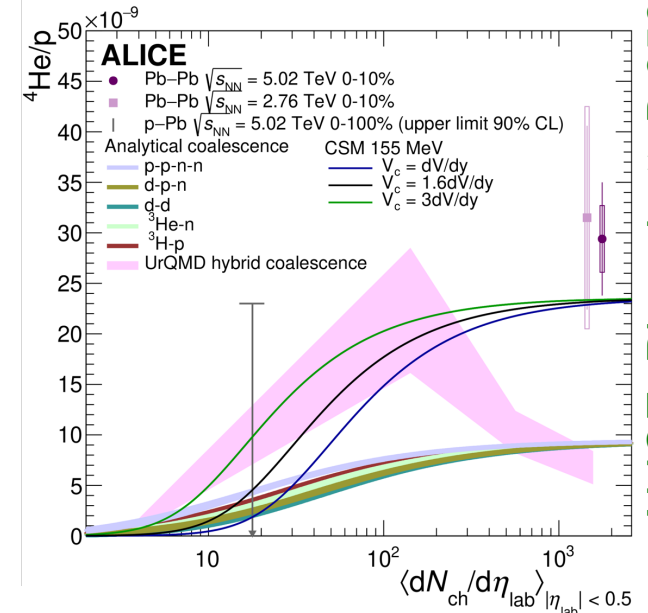
Talk #554  
Y. Zhou

Poster #711  
L. Chen

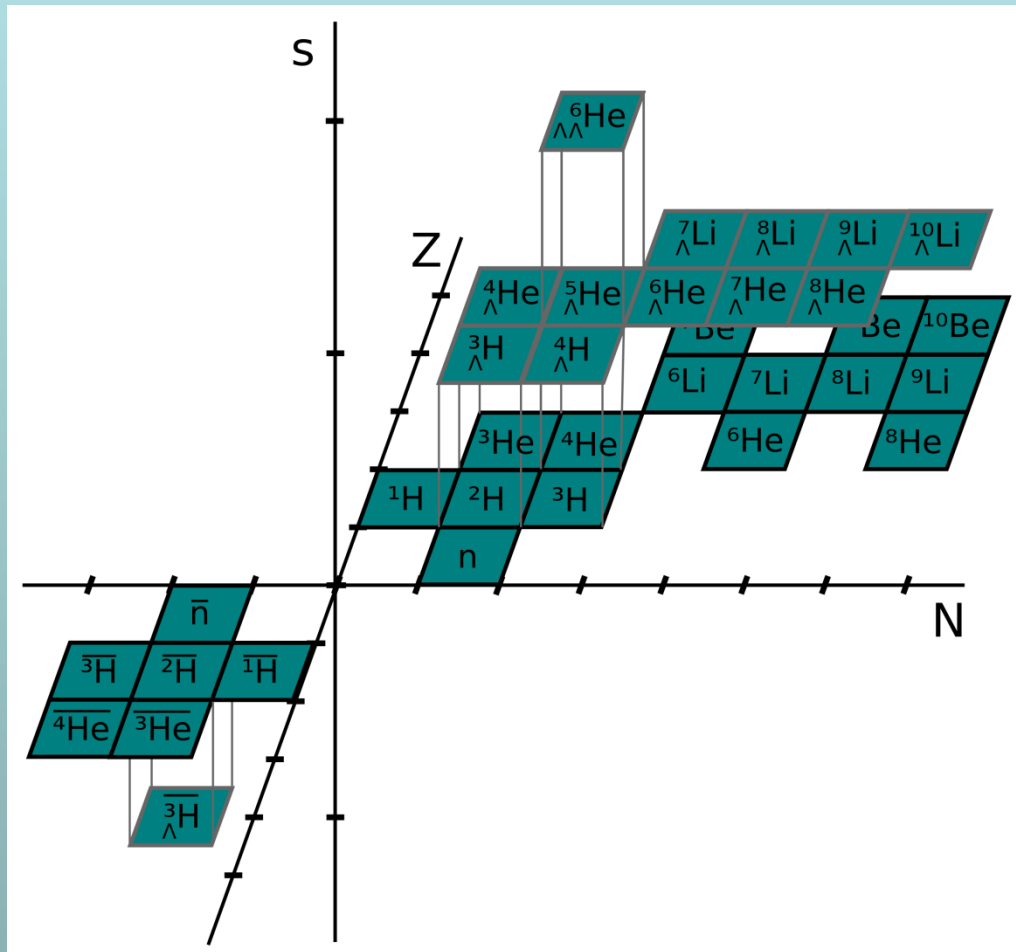
# Light nuclei: model description



- Model comparisons in ratios of A/p
- d/p possible to describe by coalescence and thermal models
- $^3\text{He}$  seems to prefer coalescence,  $^3\text{H}$  prefers coalescence
- $^4\text{He}$  in contrary clearly prefers thermal production



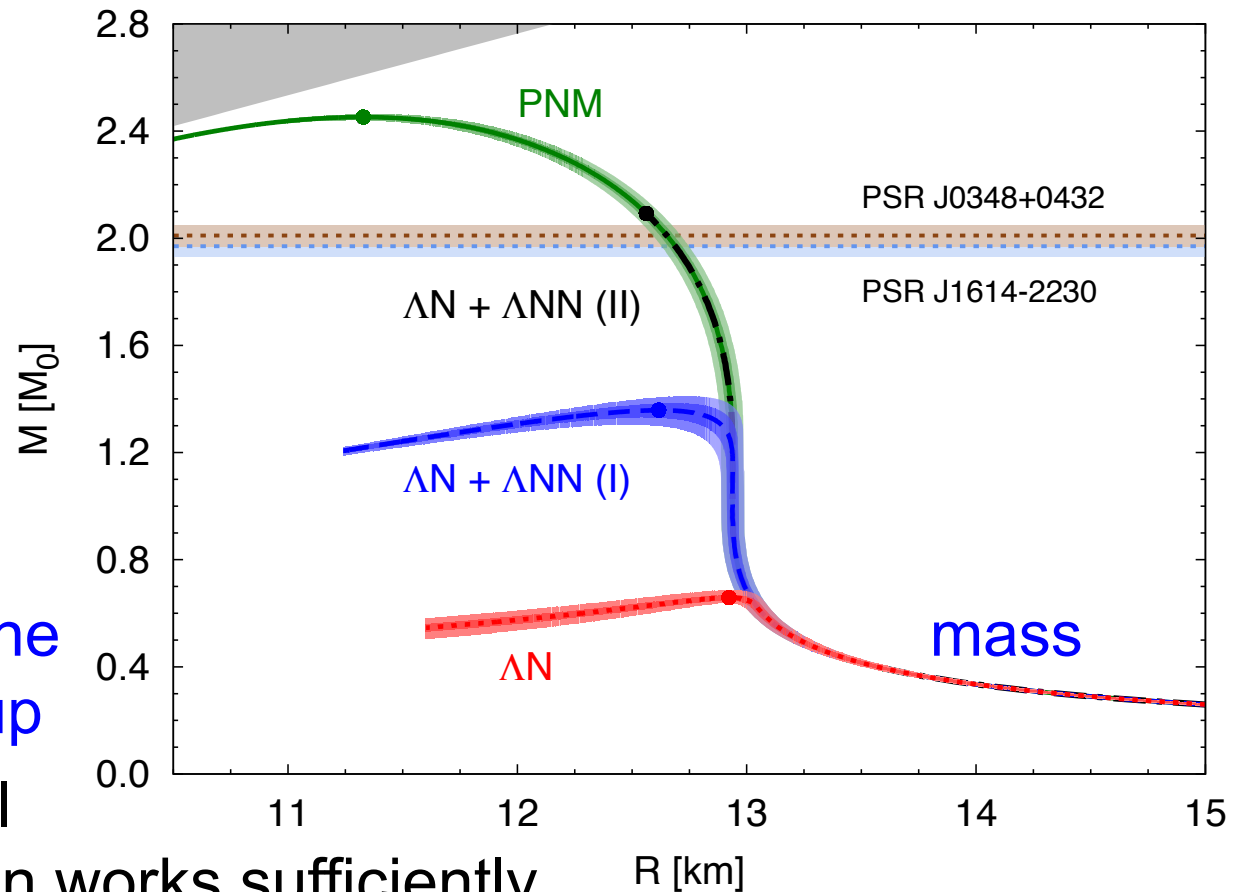
# Hypernuclei





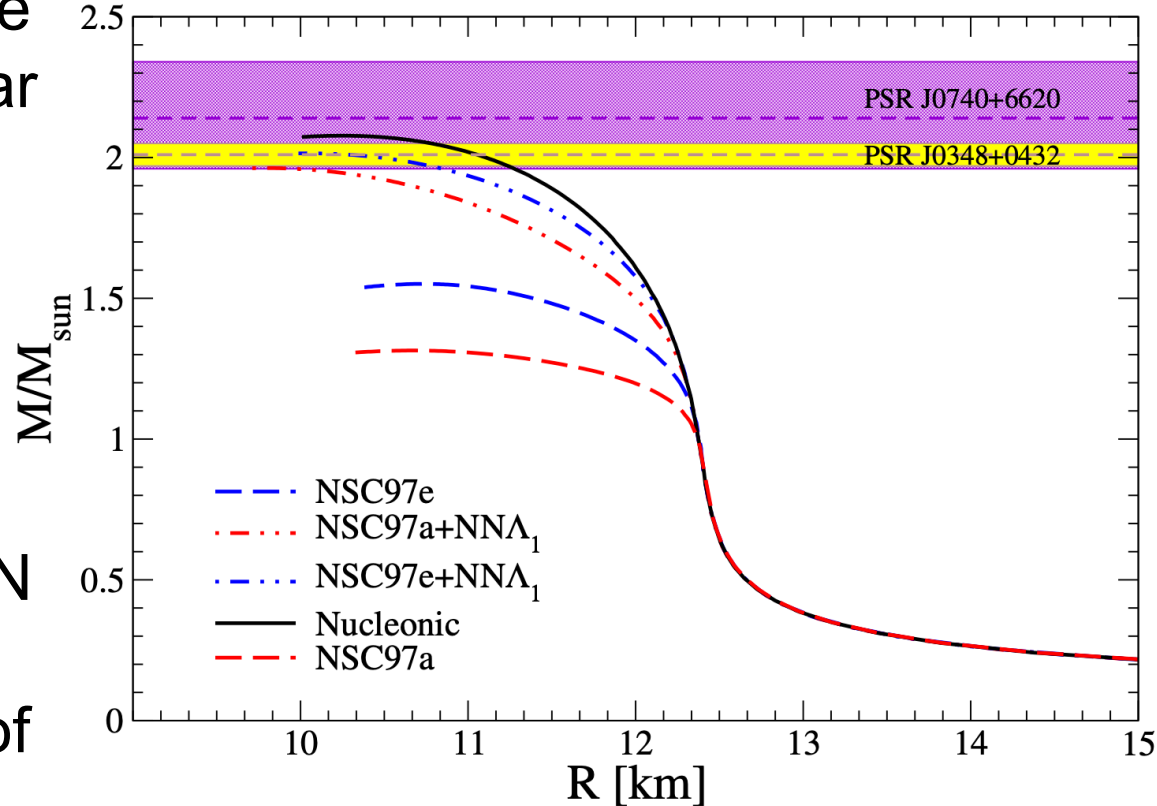
# Neutron stars and interactions

- Hyperon puzzle in neutron stars  $\rightarrow$  hyperons make the EOS softer:
- Pure neutron matter (PNM) works well
- Known  $\Lambda N$  interaction  $\rightarrow$  way to soft
- Including  $\Lambda NN$  forces brings the mass slightly up
- Only additional  $\Lambda NN$  interaction works sufficiently



# Hypernuclei

- Hypernuclei are unique probes to study nuclear structure
- Single  $\Lambda$ -hypernuclei are major source of extracting  $\Lambda$ -N interaction
- Correct  $\Lambda$ -N and  $\Lambda$ -N-N interaction needed to understand structure of neutron stars



*D. Logoteta et al., Astron. Astrophys. 646 (2021) A55*

# Hypertriton

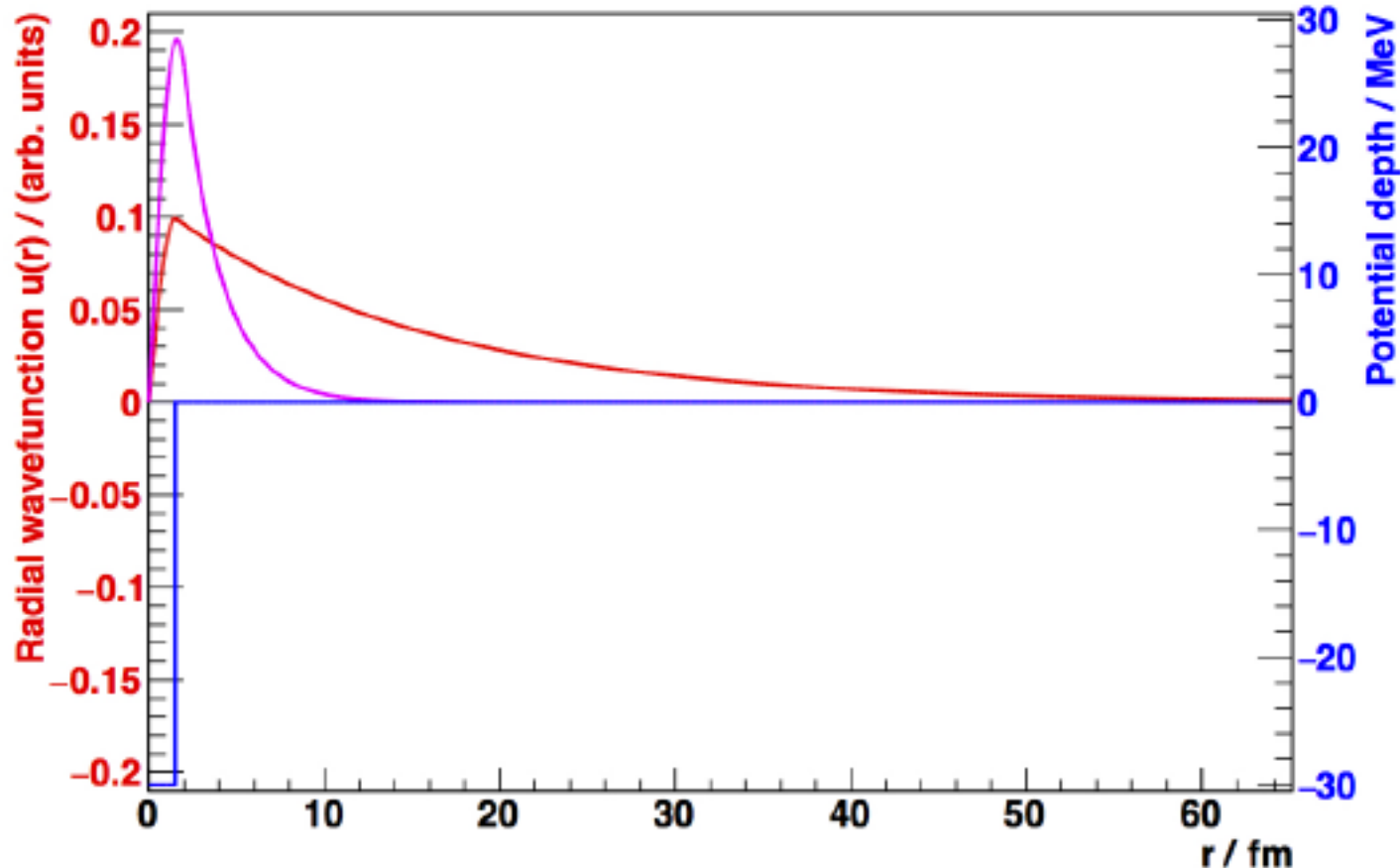
Bound state of  $\Lambda$ , p, n

$$m = 2.991 \text{ GeV}/c^2 \text{ (} B_{\Lambda} = 130 \text{ keV)}$$

# Hypertriton

Bound state of  $\Lambda$ , p, n

$$m = 2.991 \text{ GeV}/c^2 \quad (B_\Lambda = 130 \text{ keV})$$

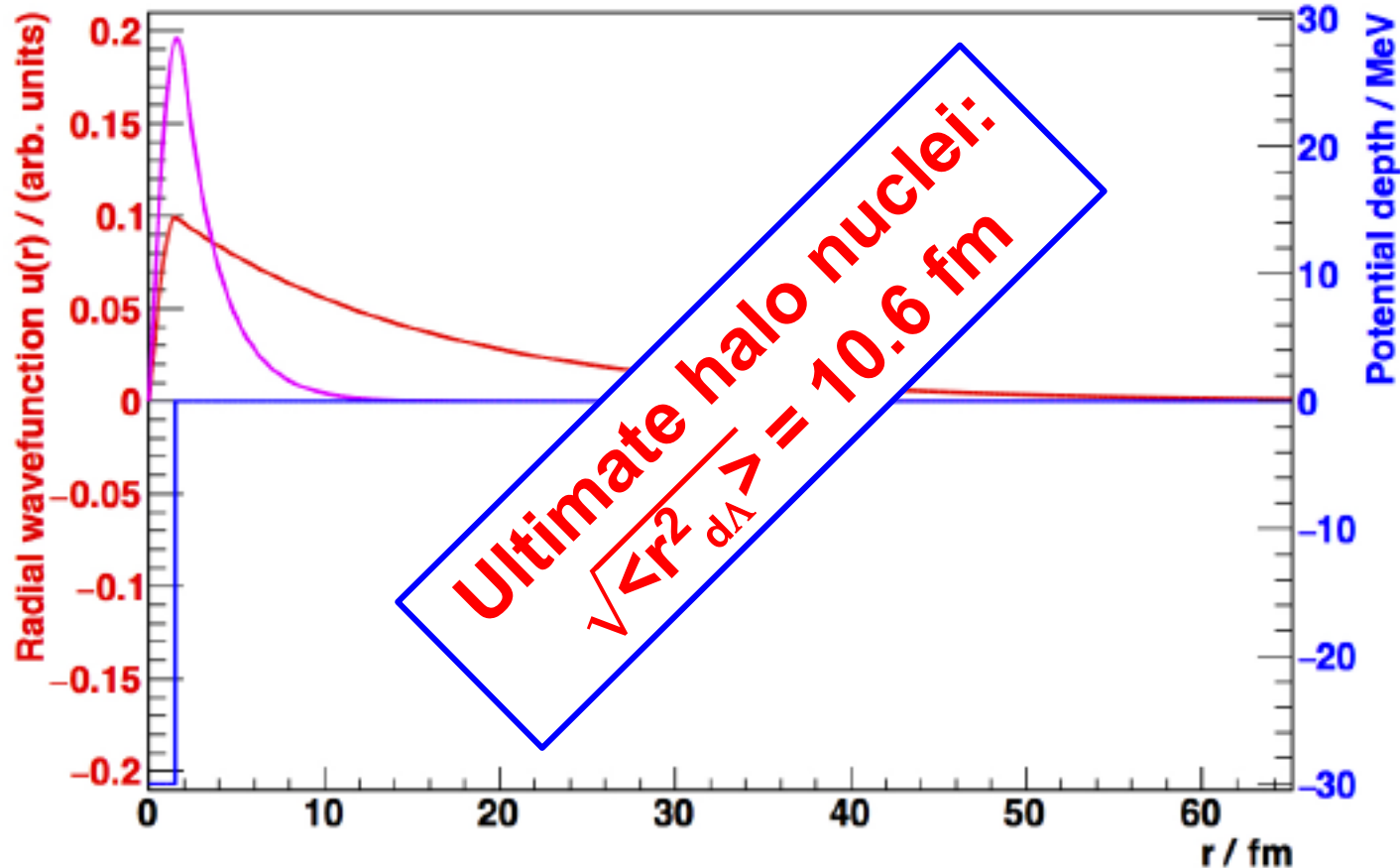


*P. Braun-Munzinger, BD, Nucl. Phys. A 987 (2019) 144*

# Hypertriton

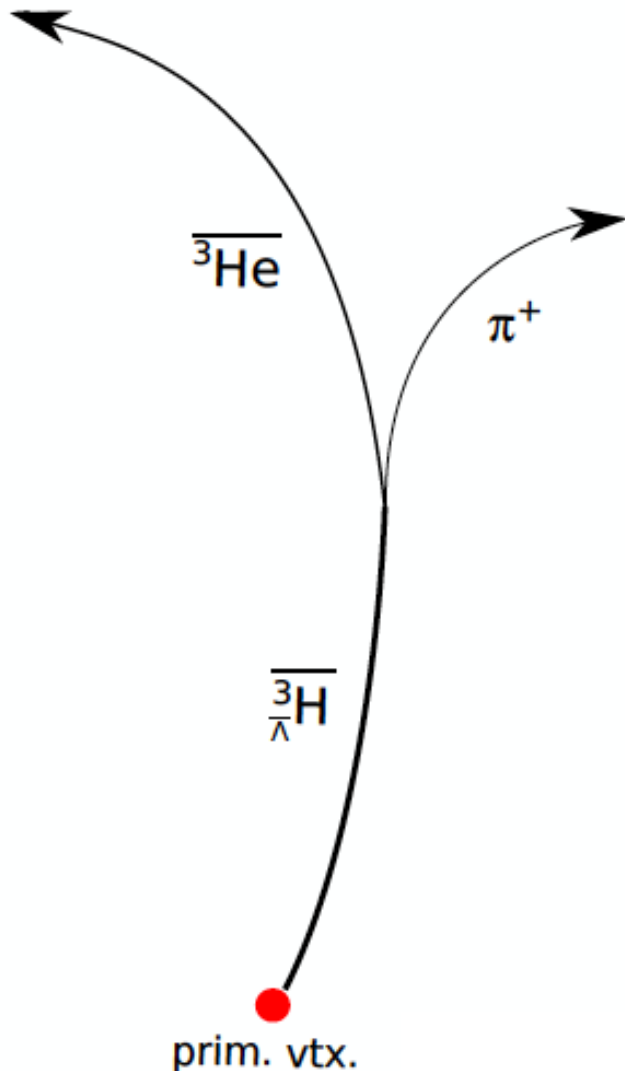
Bound state of  $\Lambda$ , p, n

$$m = 2.991 \text{ GeV}/c^2 \quad (B_\Lambda = 130 \text{ keV})$$



*P. Braun-Munzinger, BD, Nucl. Phys. A 987 (2019) 144*

# Hypertriton Identification

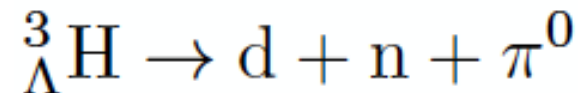
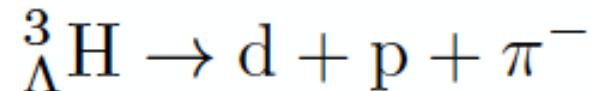
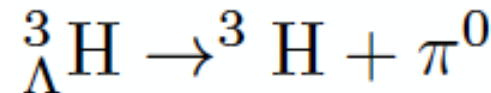
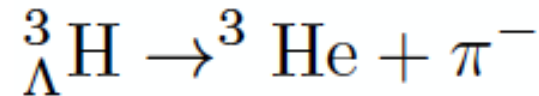


Bound state of  $\Lambda$ , p, n

$m = 2.991 \text{ GeV}/c^2$  ( $B_\Lambda = 130 \text{ keV}$ )

→ Radius of about 10.6 fm

Decay modes:

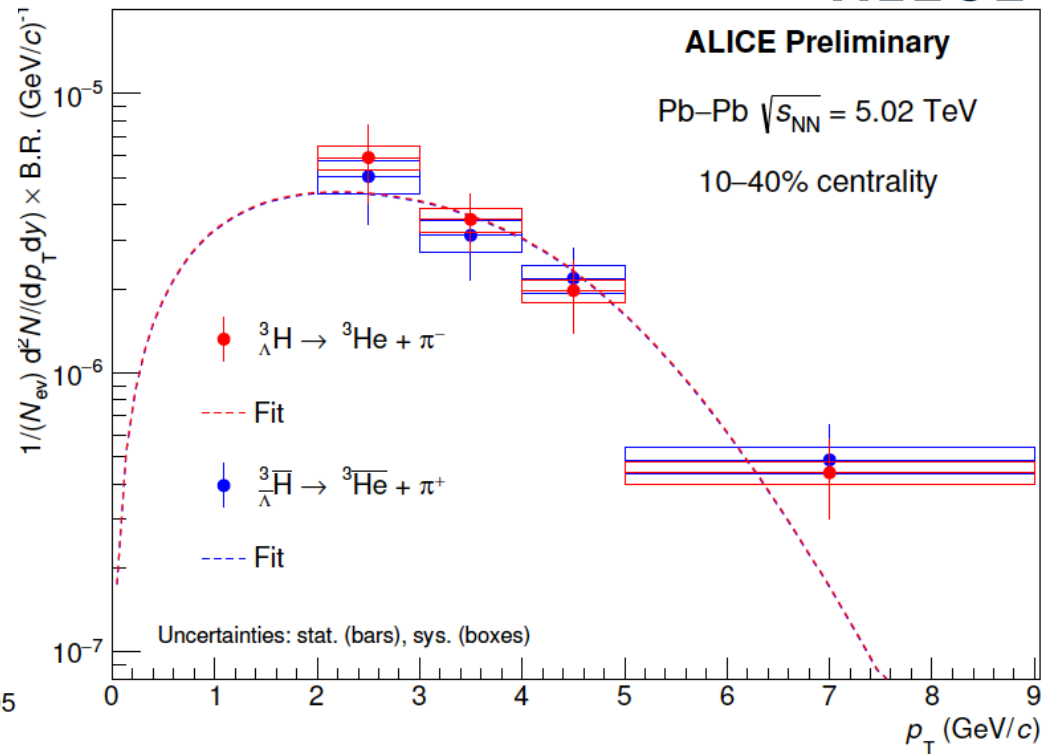
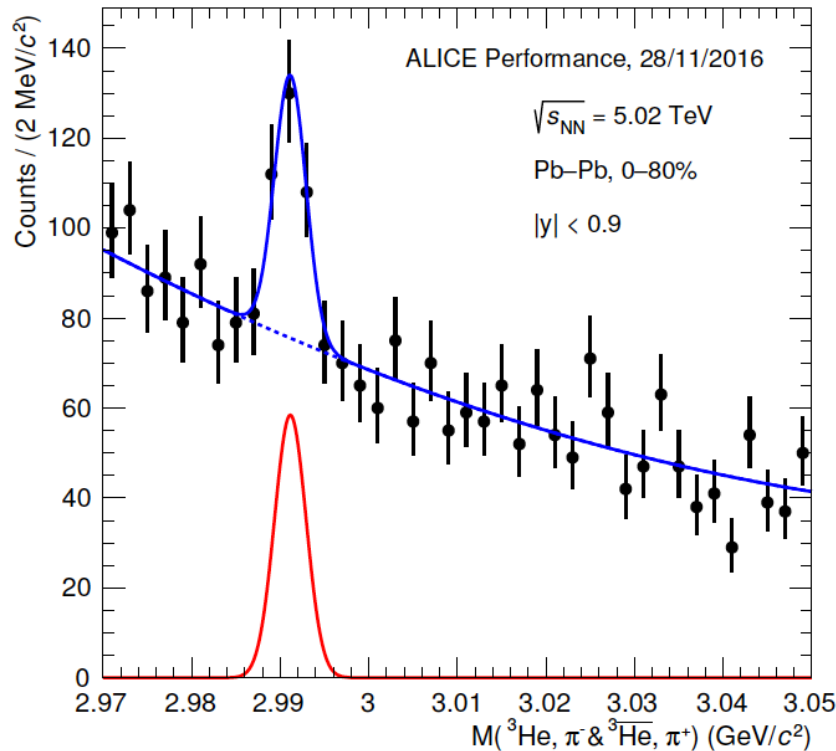


+ anti-particles

→ Anti-Hypertriton first observed by  
STAR Collaboration:

*Science 328,58 (2010)*

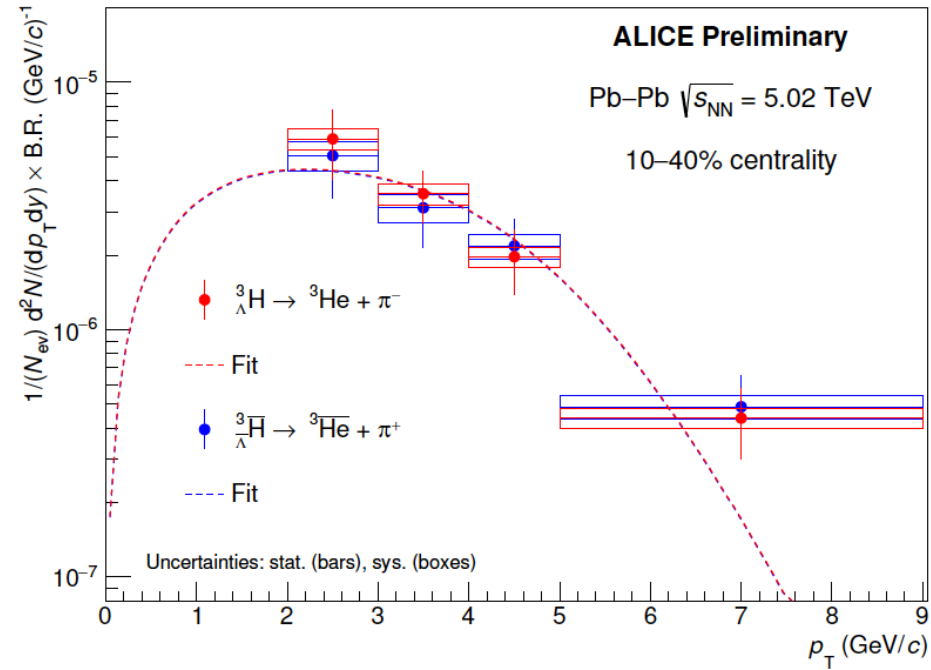
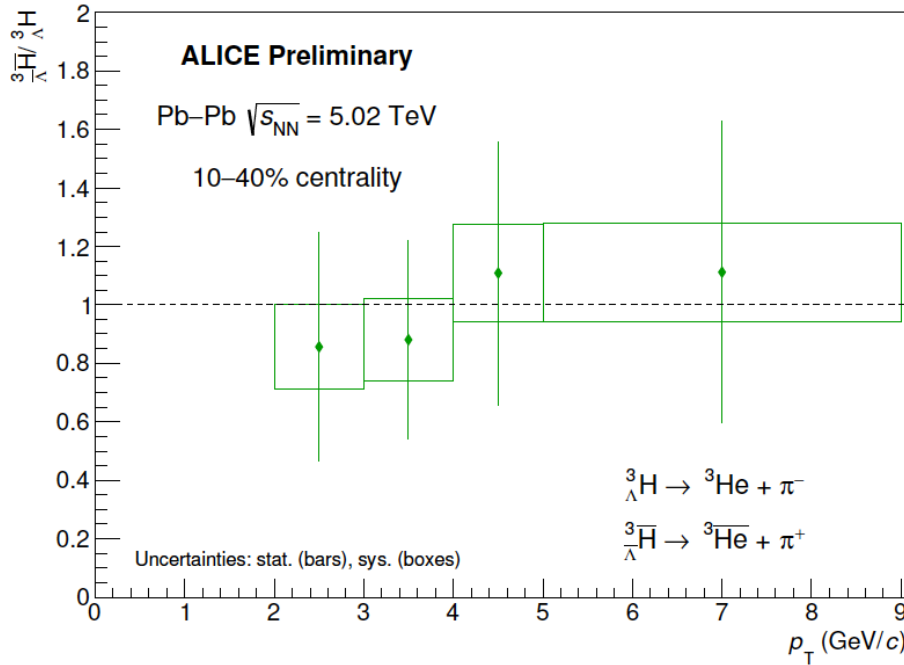
# Hypertriton signal



- Clear signal reconstructed by decay products
- Spectra can also be described by Blast-Wave model  
→ Hypertriton flows as all other particles



# Hypertriton spectra

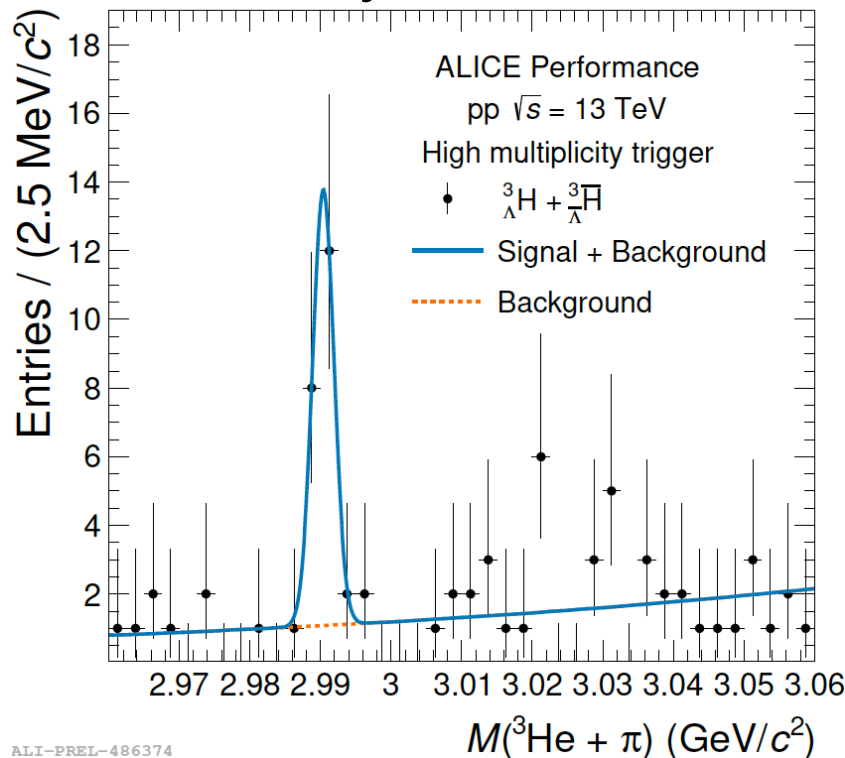


- Anti-hypertriton/Hypertriton ratio consistent with unity vs.  $p_T$

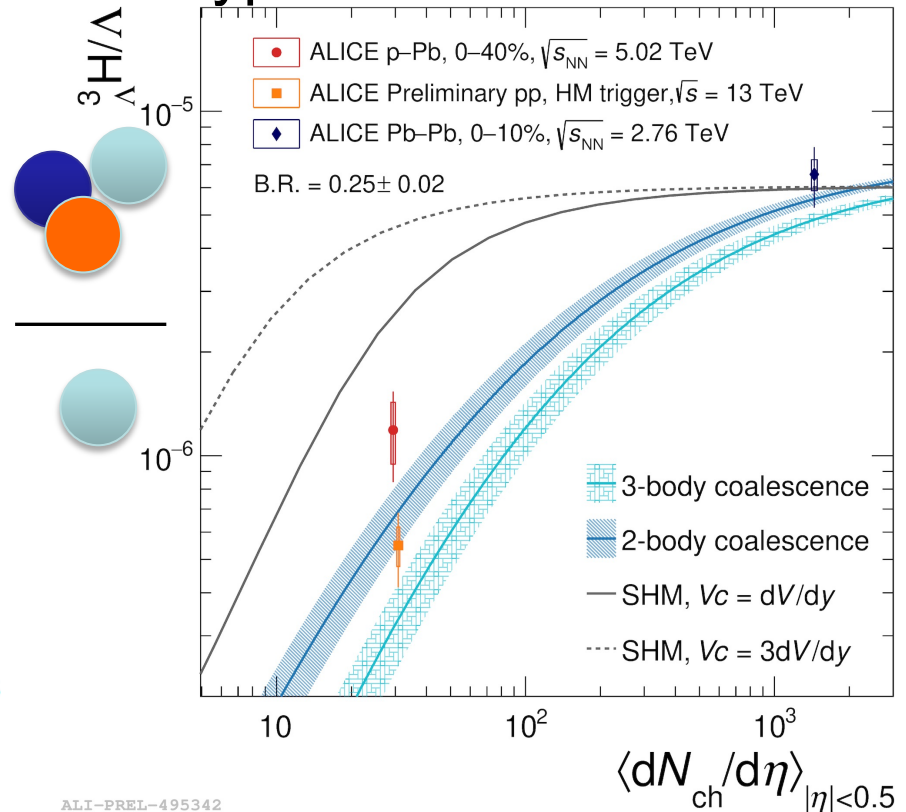


# Hypertriton in pp & p-Pb

- Hypertriton signal recently also extracted in pp and p-Pb collisions
- Stronger separation between models as for other particle ratios, mainly due to the size of the hypertriton



ALI-PREL-486374

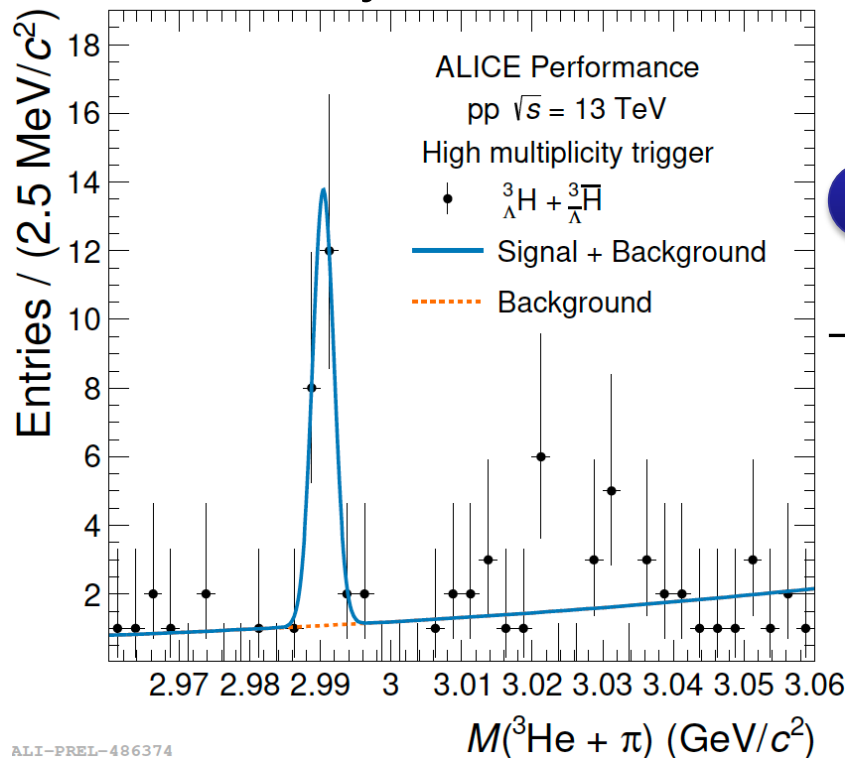


ALI-PREL-495342

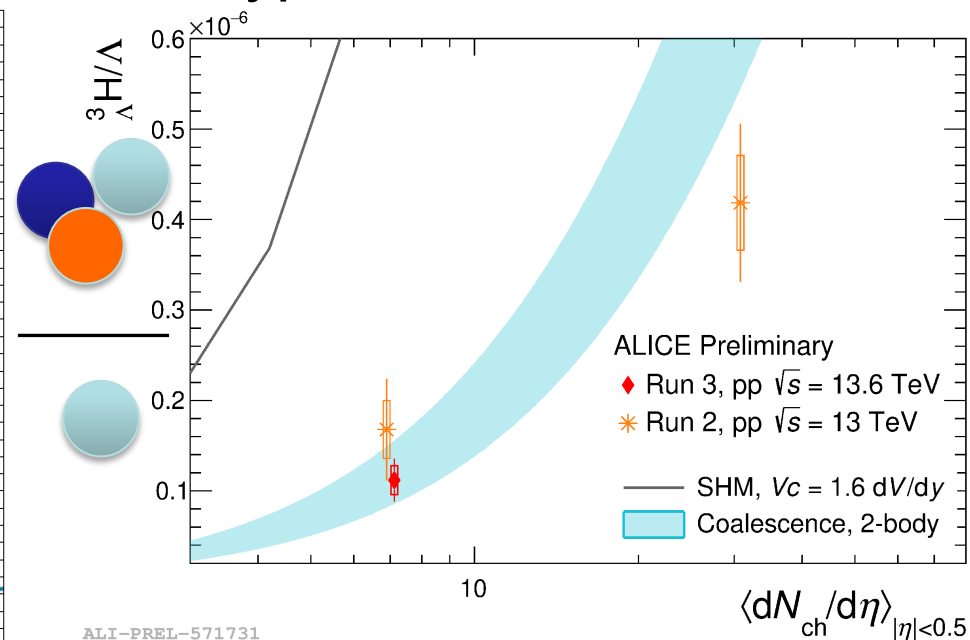


# Hypertriton in pp & p-Pb

- Hypertriton signal recently also extracted in pp and p-Pb collisions
- Stronger separation between models as for other particle ratios, mainly due to the size of the hypertriton



ALI-PREL-486374

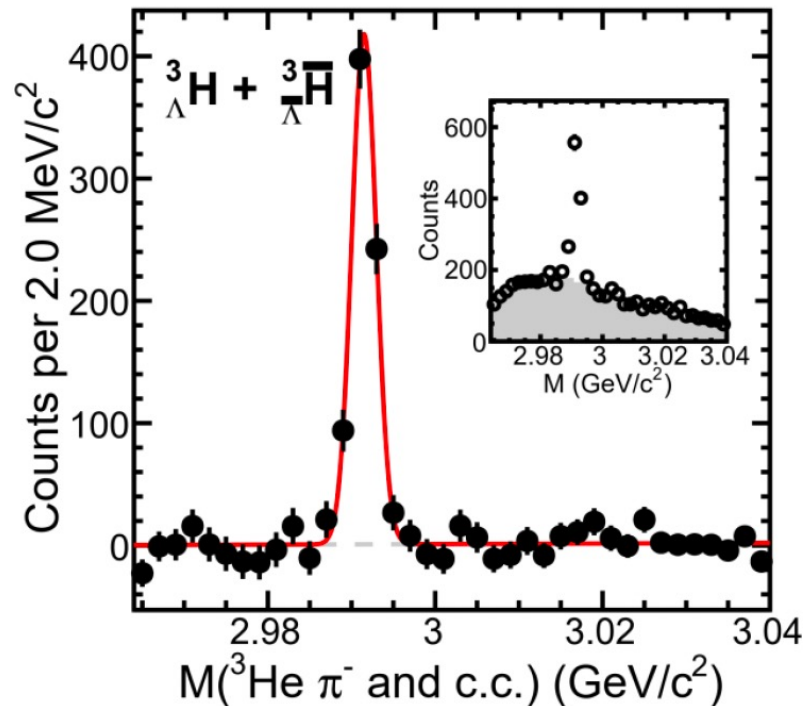


ALI-PREL-571731

# Hypertriton at RHIC



- Hypertriton signal recently also extracted in isobar collisions



*As presented by Dongsheng Li @ SQM2024*

0-80% Zr+Zr & Ru+Ru @  $\sqrt{s_{NN}} = 200$  GeV

$p_T \leq 5$  (GeV/c)

○ Same Event (SE)

■ Mixed Event (ME)

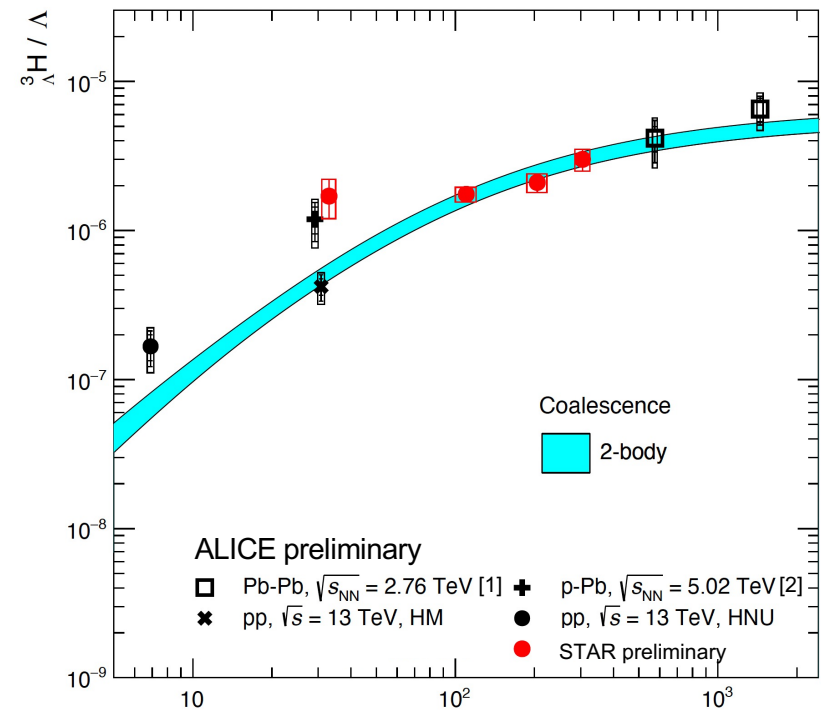
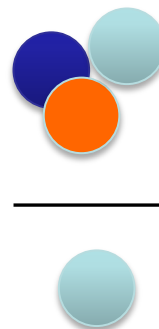
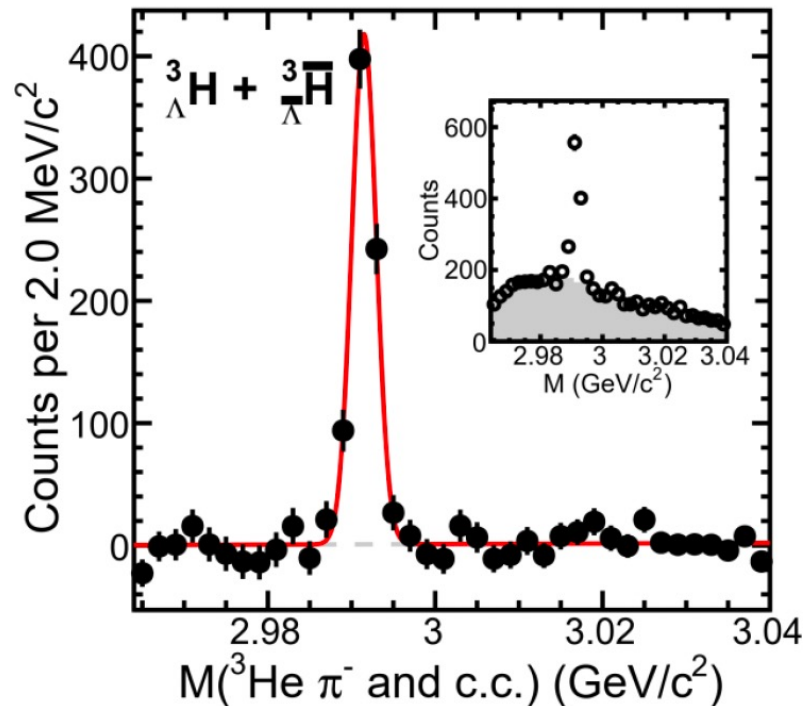
● SE - ME



# Hypertriton/ $\Lambda$ ratio



- Hypertriton signal recently also extracted in isobar collisions
- Stronger separation between models as for other particle ratios, mainly due to the size of the hypertriton

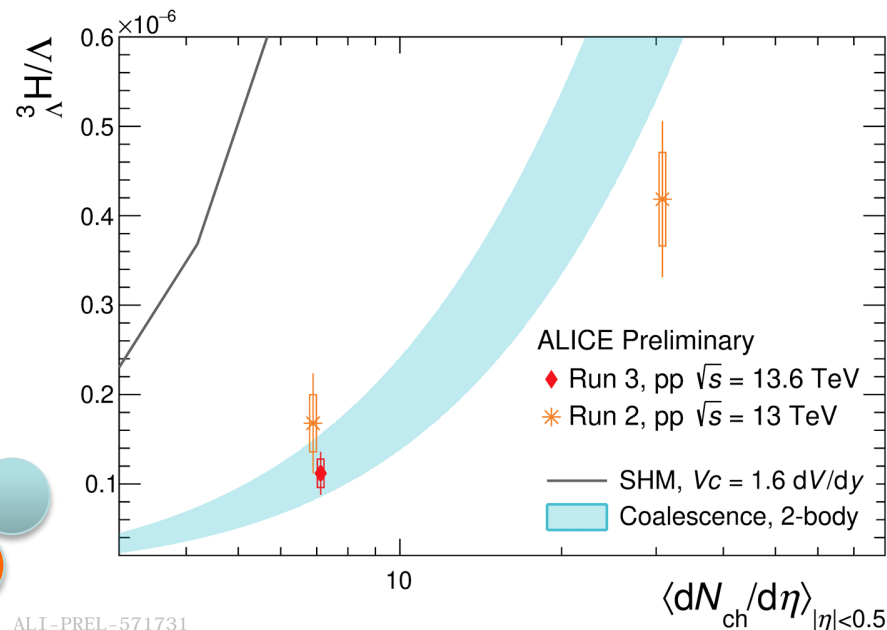
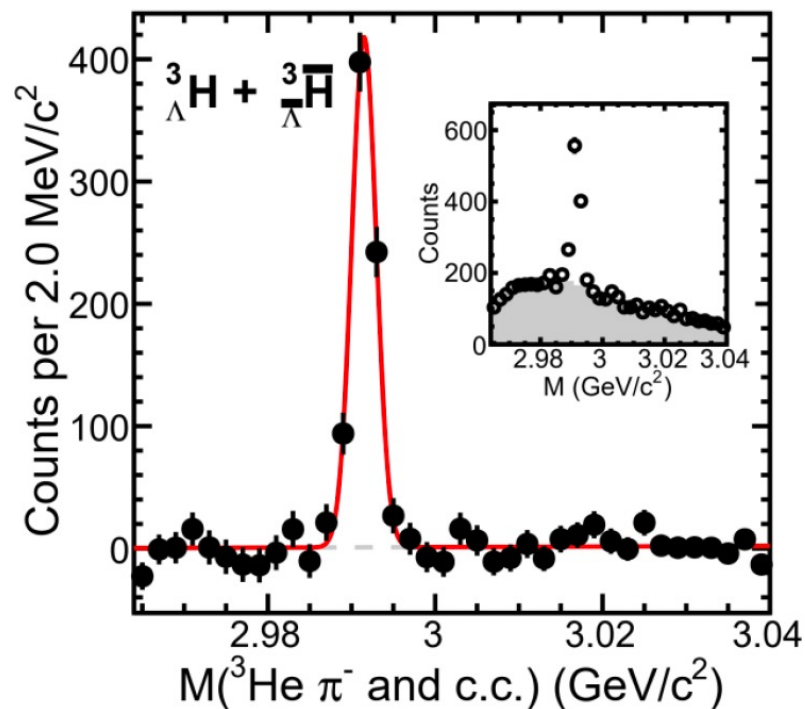


[1] *PLB* 754 (2016) 360  
[2] *PRL* 128 (2022) 252003

$\langle dN_{\text{ch}}/d\eta \rangle_{|\eta| < 0.5}$

# Hypertriton/ $\Lambda$ ratio

- Hypertriton signal recently also extracted in isobar collisions
- Stronger separation between models as for other particle ratios, mainly due to the size of the hypertriton

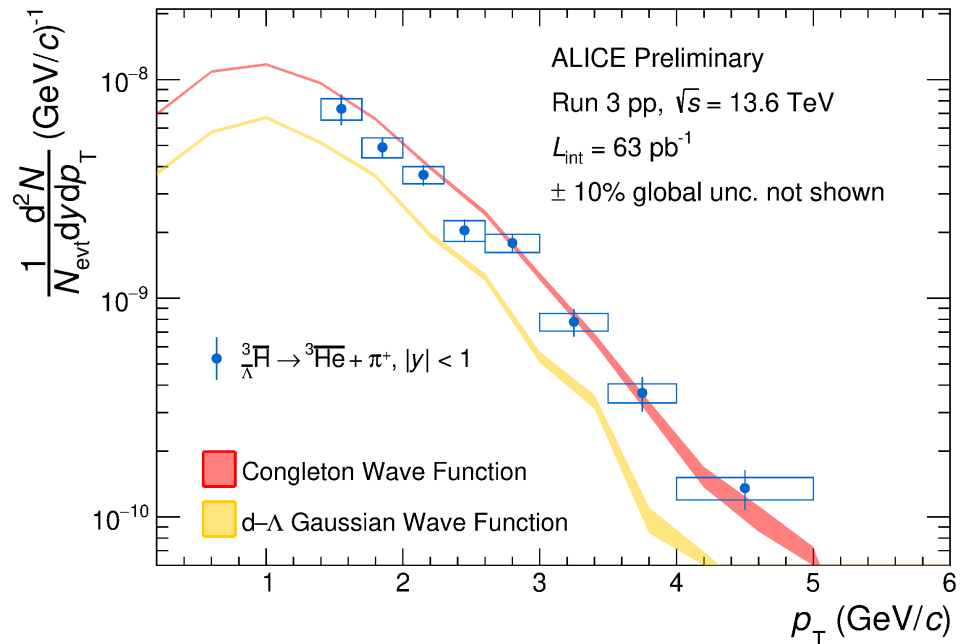


ALI-PREL-571731

# Hypertriton: small systems

- Strong suppression in small systems due to large size of the object compared to the system size
- Coalescence model favoured over canonical statistical model

*M. Mahlein et al. arxiv:2504.02491*



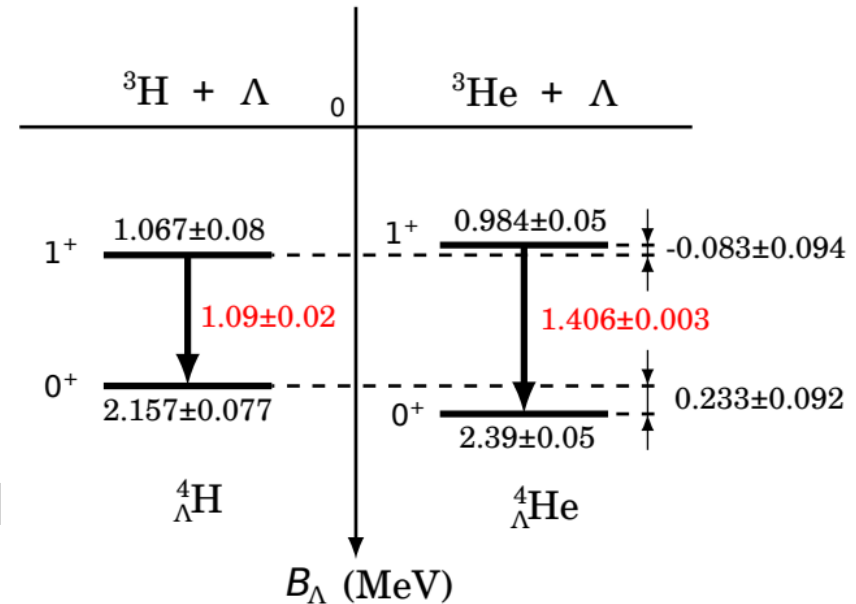
QM2025  
Talk #1118  
J. Ditzel

QM2025  
Poster #493  
C. Reetz



# A = 4 hypernuclei

- Large suppression expected for A = 4 hypernuclei by the SHM wrt A = 3
- A = 4 hypernuclei are more bound and each has an excited state  
[Phys. Rev. Lett. 115, 222501 \(2015\)](#)
- The yields of these hypernuclei are enhanced with respect to the ground state due to the feed-down from excited states



[M. Schäfer, N. Barnea, A. Gal, Phys.Rev.C 106, L031001 \(2022\)](#)

- Also the yields of the SHM scale with the **spin-degeneracy**
- Resulting in a total enhancement of a factor 4 for both hypernuclei

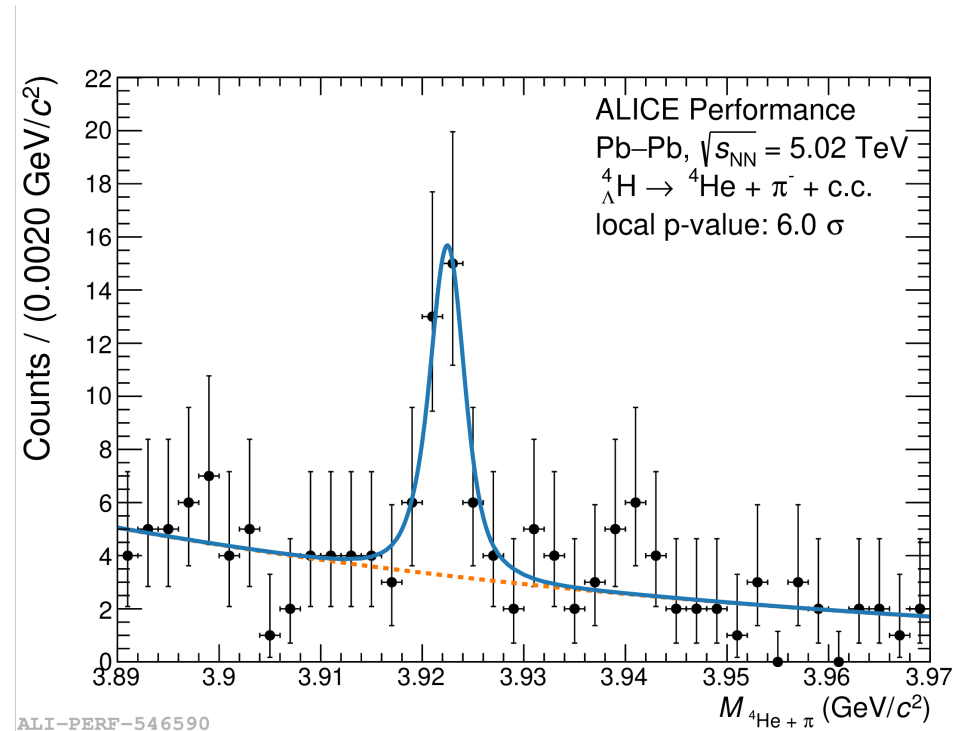
[BD, EPJ Web Conf. 276 \(2023\) 04002](#)



# A = 4 hypernuclei

- For the first time, we are able to reconstruct A = 4 (anti)hypernuclei at the LHC and determine their production yield
- **(Anti)hyperhydrogen-4** invariant-mass spectrum in Run 2  
Pb-Pb collisions at 5.02 TeV
- Examined in the two-body decay:  

$${}^4_{\Lambda}\text{H} \rightarrow {}^4\text{He} + \pi^- + \text{c.c.}$$
- Reaching a local p-value of **6σ**

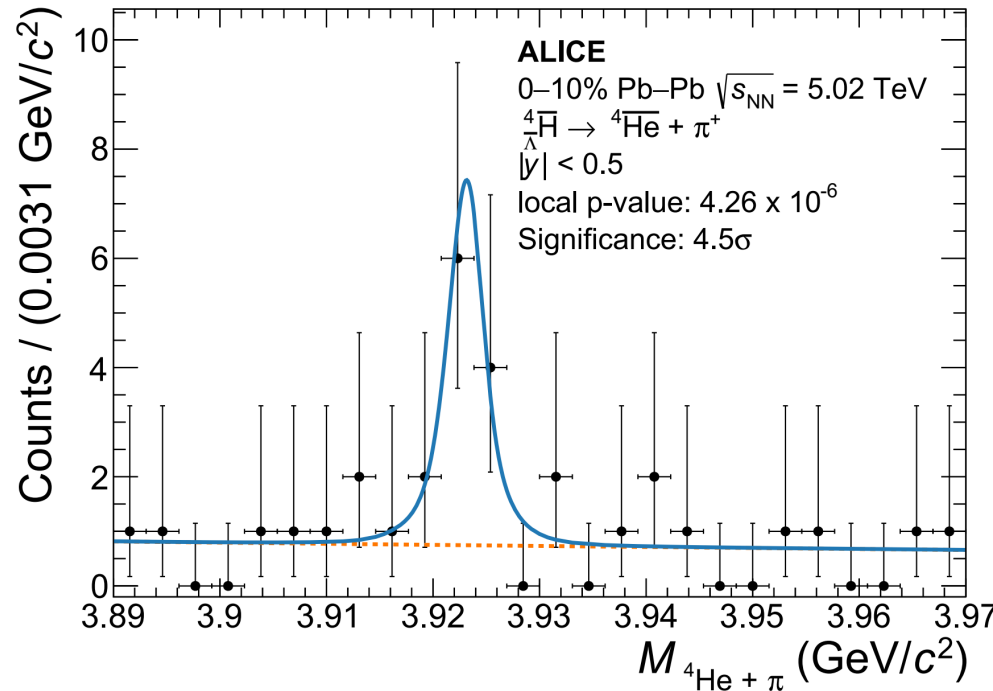


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- Reaching a significance of **4.5 $\sigma$**

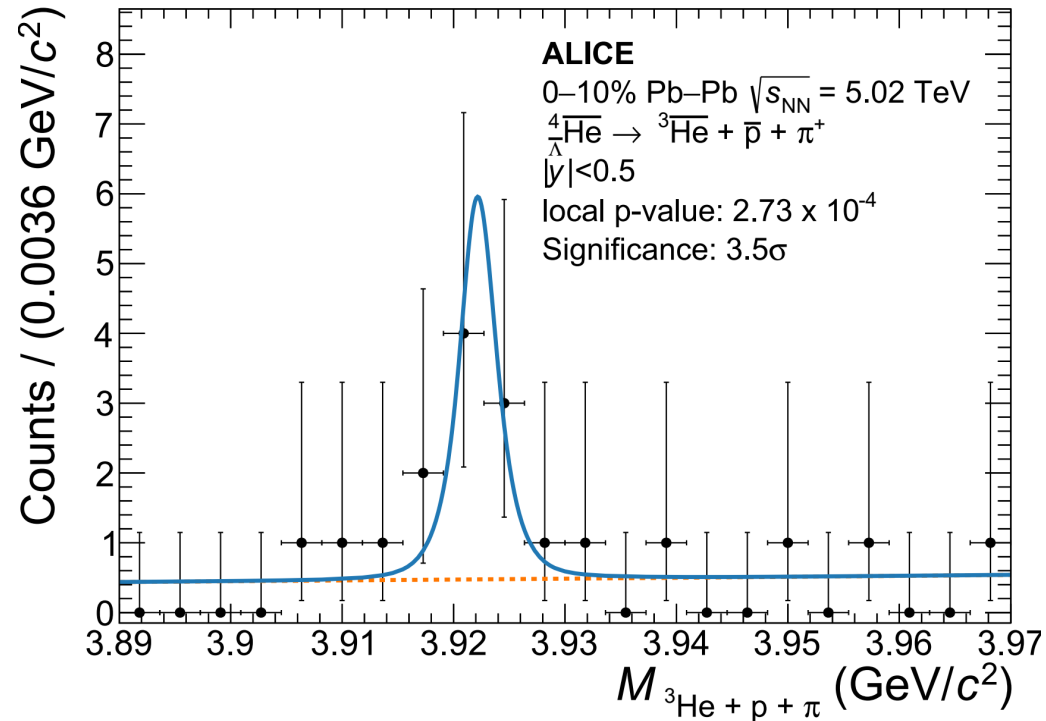
ALICE Collaboration, [arXiv:2410.17769](https://arxiv.org/abs/2410.17769)



# A = 4 hypernuclei

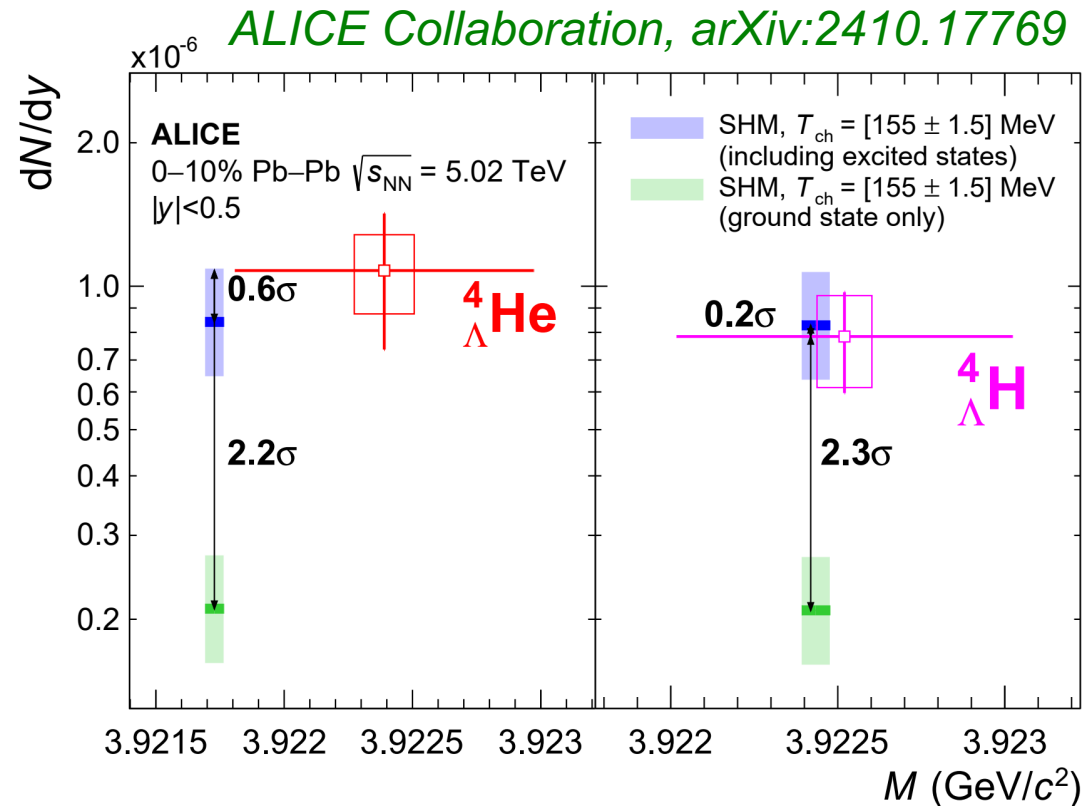
- For the first time, we are able to reconstruct A = 4 (anti)hypernuclei at the LHC and determine their production yield
- First observation of the antihyperhelium-4** in Run 2 Pb-Pb collisions at 5.02 TeV
- Reaching a significance of **3.5 $\sigma$**

ALICE Collaboration, [arXiv:2410.17769](https://arxiv.org/abs/2410.17769)



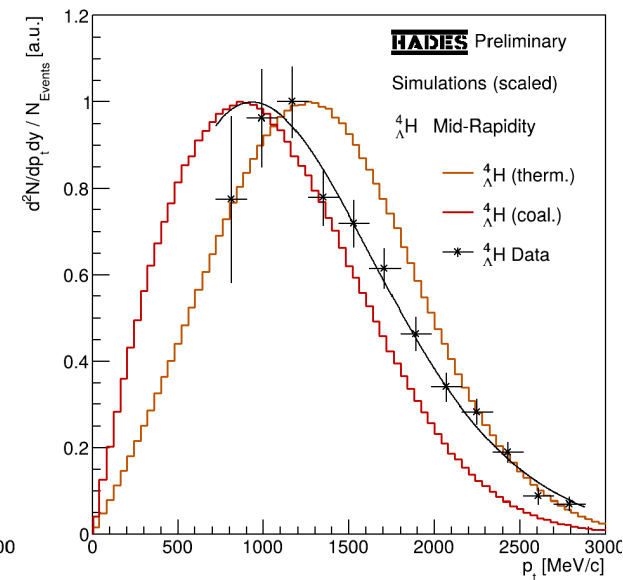
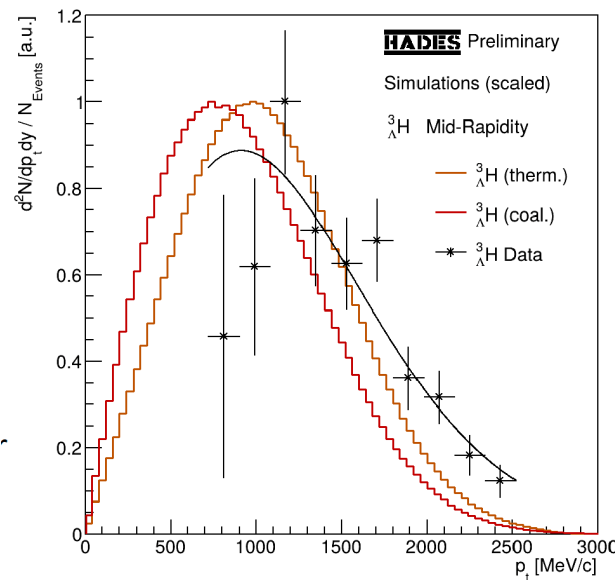
# A = 4 hypernuclei

- First measurement of the (anti)hyperhelium-4 production yield
- Testing the dependence of the yields of the SHM with the spin-degeneracy
- Our yields confirm the hypothesis of excited states for both (anti)hypernuclei within  $2\sigma$
- currently dominated by statistical uncertainties
- with more data, a high precision measurement will be feasible (like for the  $\Lambda$  hyperon)



# Hypernuclei: HADES

- Lifetimes of investigated hypernuclei consistent with world data
- $p_T$  and  $y$  distributions hint to (thermal) production

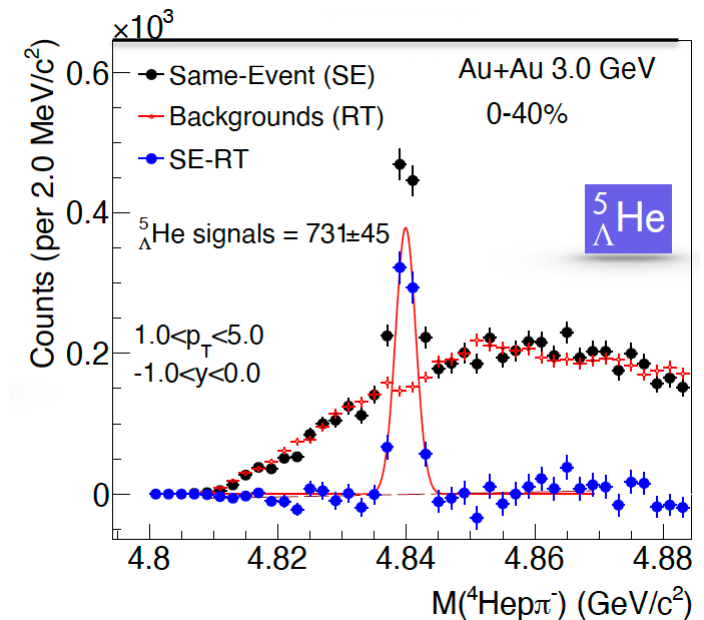


QM2025  
Talk #577  
S. Spies

# Hypernucleus: ${}^5_{\Lambda}\text{He}$



- Highlight: clear signal of hyperhelium-5 in Au+Au at 3.0 GeV

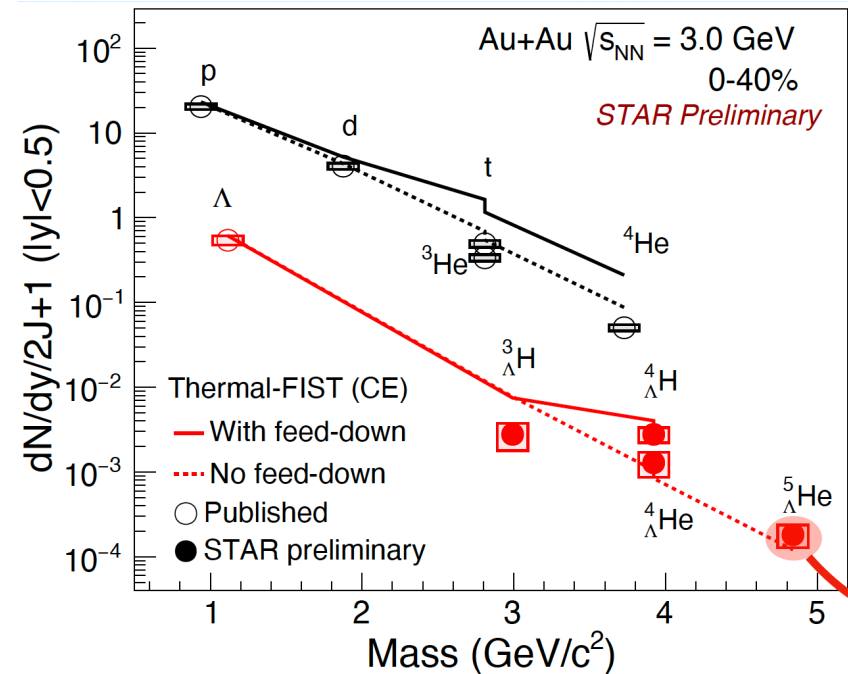


QM2025  
Talk #554  
Y. Zhou

# Hypernucleus: ${}^5_{\Lambda}\text{He}$



- Highlight: clear signal of hyperhelium-5 in Au+Au at 3.0 GeV
- Comparison to thermal model seems to work only up to  $A=2$ , even when excited states are considered

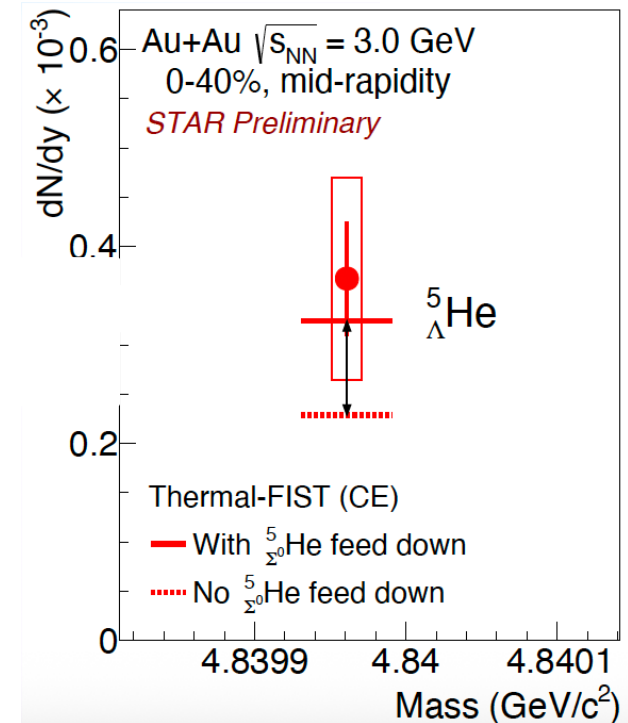


QM2025  
Talk #554  
Y. Zhou

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- Highlight: clear signal of hyperhelium-5 in Au+Au at 3.0 GeV
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- Hypothesis of  $A=5$   $\Sigma$ -hypernucleus would help explaining the data



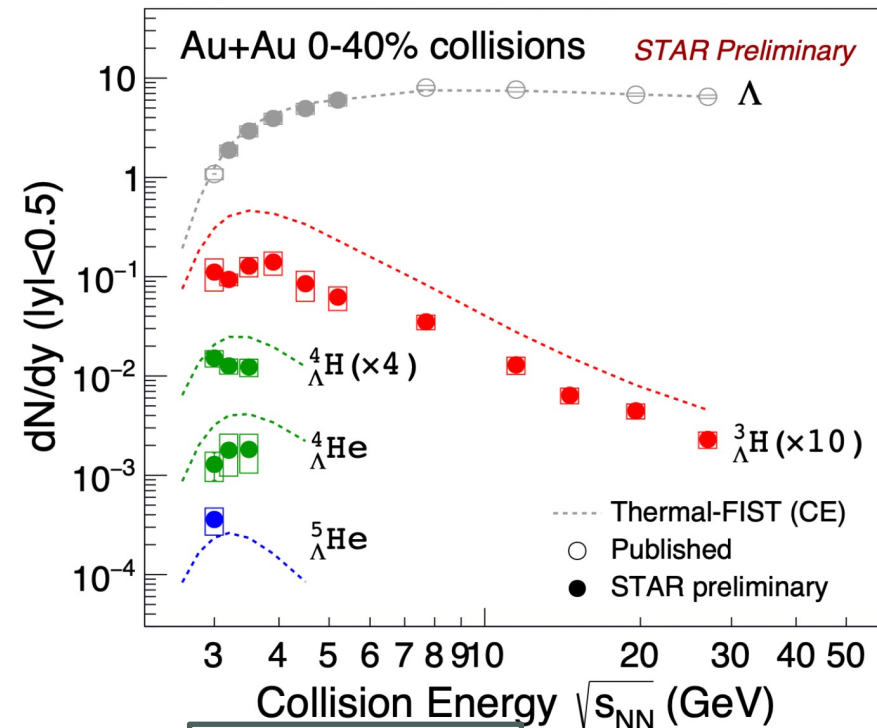
QM2025  
Talk #554  
Y. Zhou



# Hypernucleus: ${}^5_{\Lambda}\text{He}$

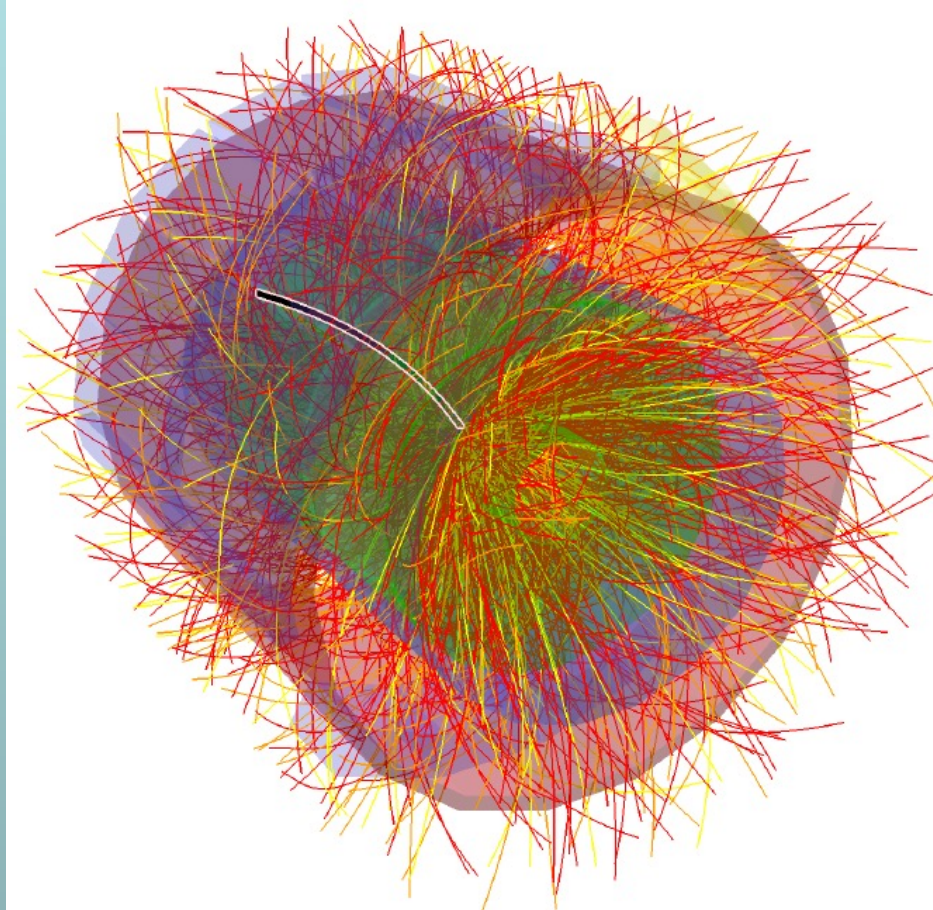


- Highlight: clear signal of hyperhelium-5 in Au+Au at 3.0 GeV
- Comparison to thermal model seems to work only up to  $A=2$ , even when excited states are considered
- Hypothesis of  $A=5$   $\Sigma$ -hypernucleus would help explaining the data
- Energy dep. shows slight offset for all hypernuclei

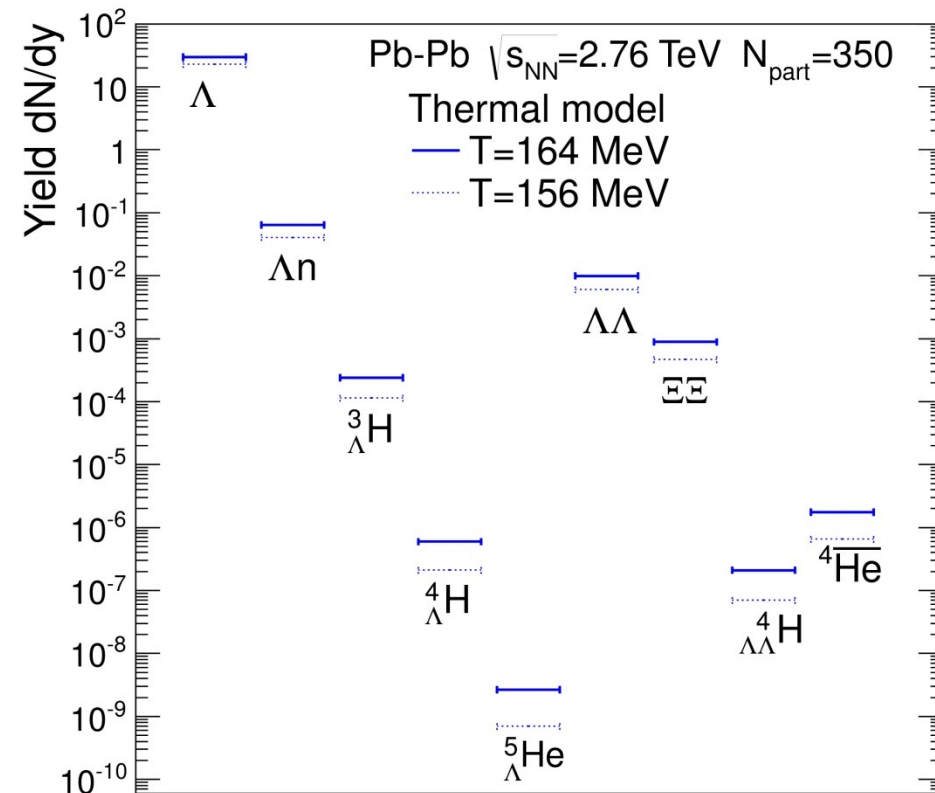


QM2025  
Talk #554  
Y. Zhou

# Outlook & Summary



# Outlook

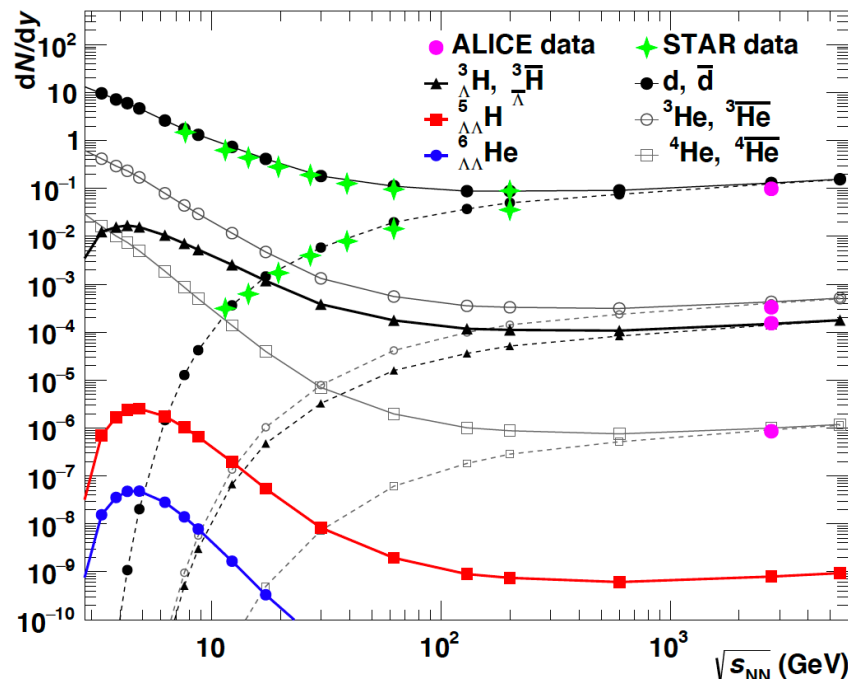


- Explore QCD and QCD inspired model predictions for (unusual) multi-baryon states
- Search for rarely produced anti- and hyper-matter
- Test model predictions, e.g. thermal and coalescence

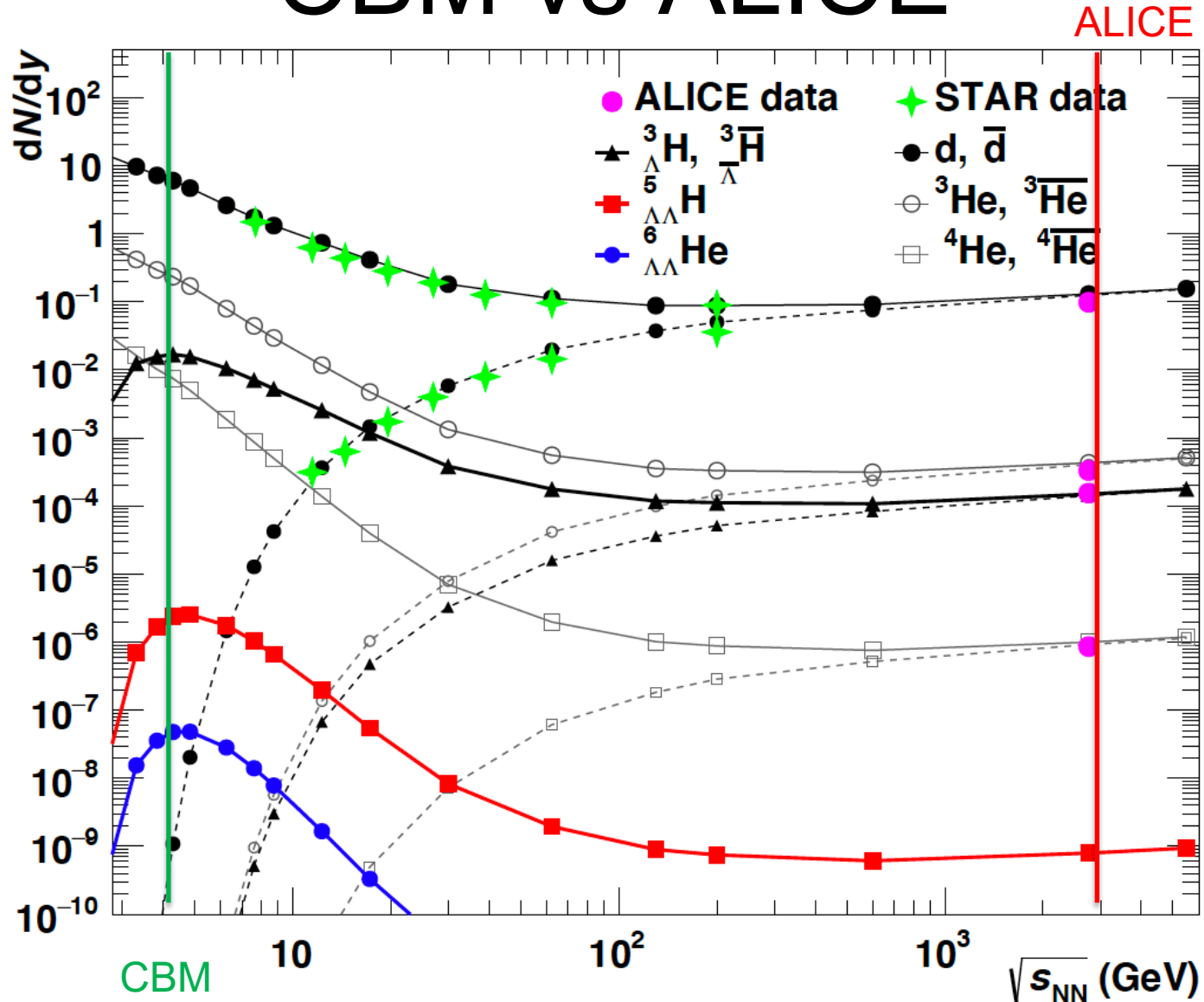
*A. Andronic, private communication, model described in A. Andronic et al., PLB 697, 203 (2011) and references therein*

# Conclusion

- ALICE@LHC, HADES@GSI and STAR@RHIC are well suited to study light (anti-)(hyper-) nuclei and perform searches for exotic bound states ( $A < 5$ )
- Copious production of loosely bound objects measured by ALICE, HADES and STAR as predicted by the thermal model
- Models describe the data rather well
- Ratios vs. multiplicity trend described by both models  
- only tension: Alpha vs.  ${}^3_\Lambda\text{H}$
- New and more precise data can be expected in the next years



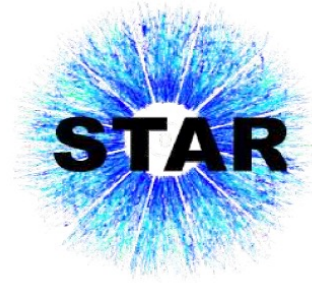
# CBM vs ALICE



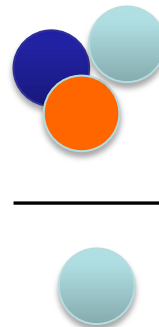
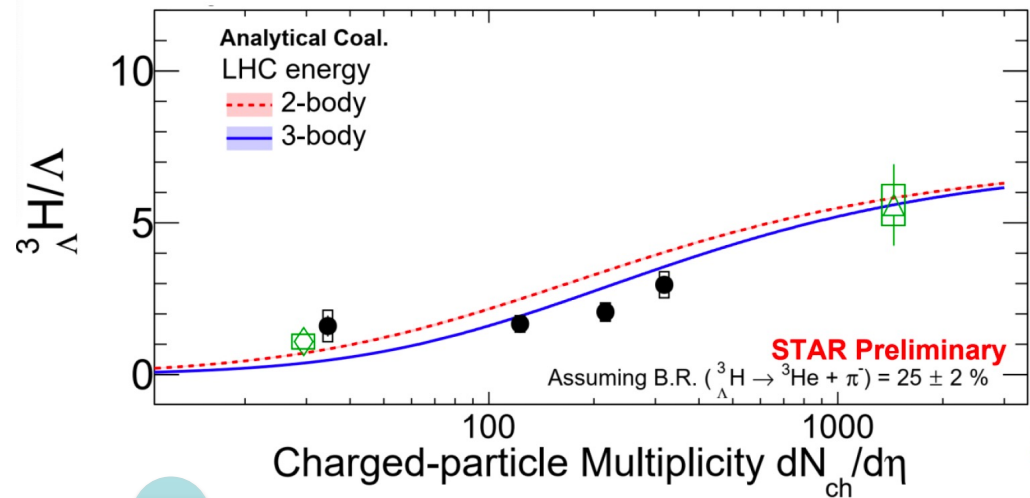
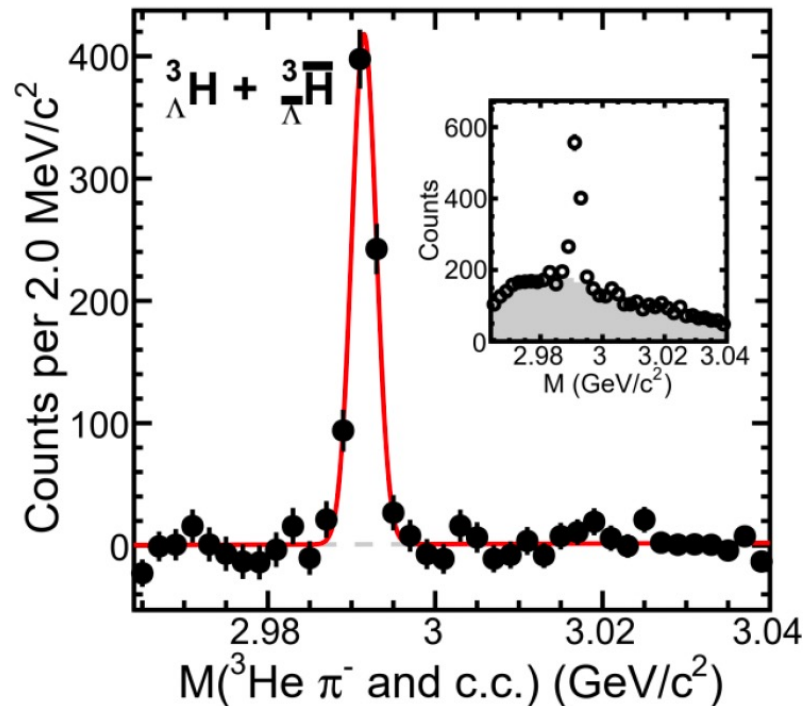
*BD, Eur. Phys. J 56 (2020) 258*

# Backup

# Hypertriton/ $\Lambda$ ratio

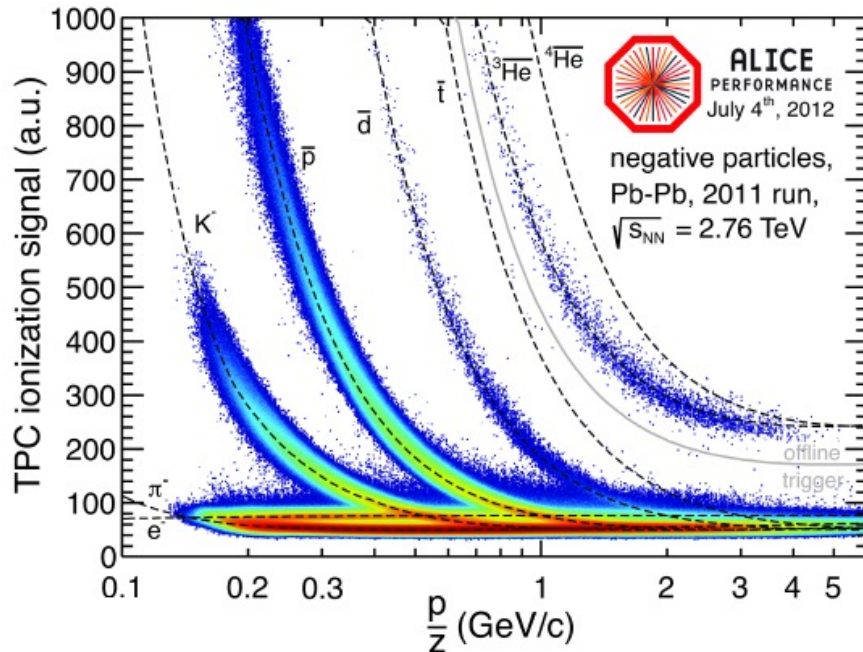


- Hypertriton signal recently also extracted in isobar collisions
- Stronger separation between models as for other particle ratios, mainly due to the size of the hypertriton



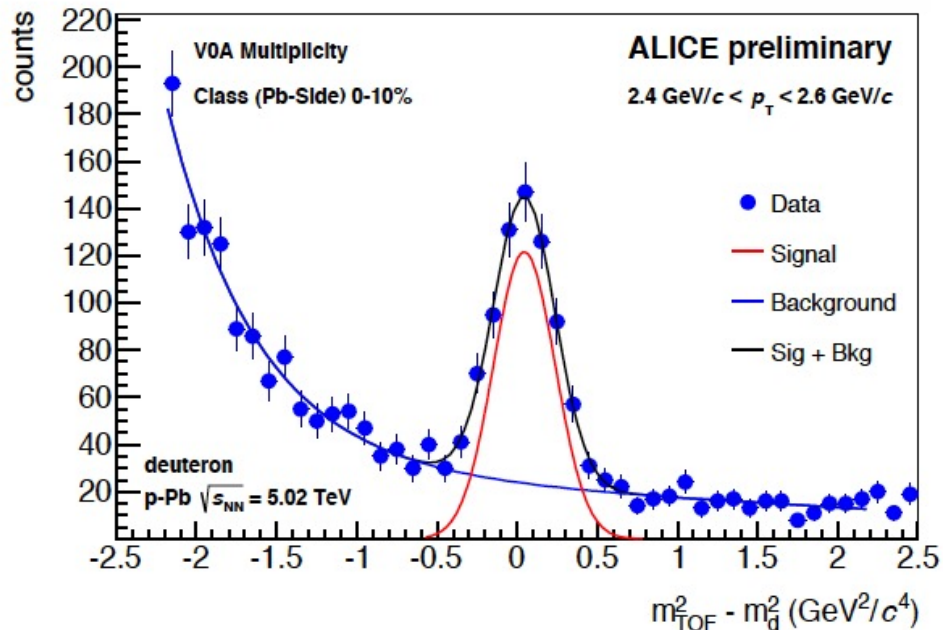


# Particle Identification



## Low momenta:

Nuclei are identified using the  $dE/dx$  measurement in the Time Projection Chamber (TPC)

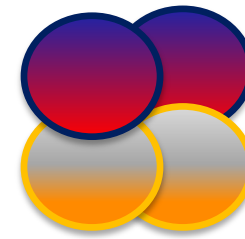


## Higher momenta:

Velocity measurement with the Time-of-Flight (TOF) detector is used to calculate the  $m^2$  distribution



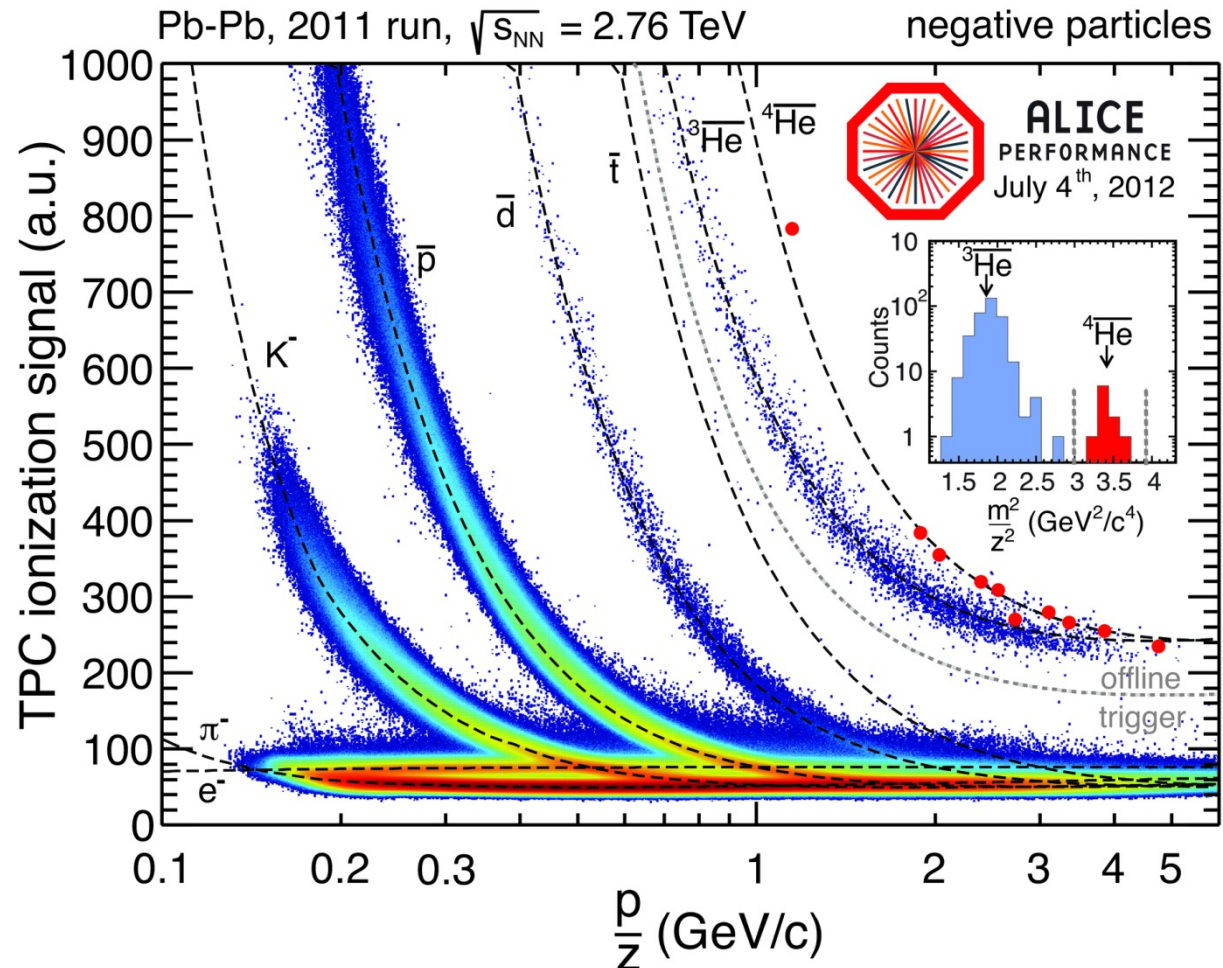
# Anti-Alpha



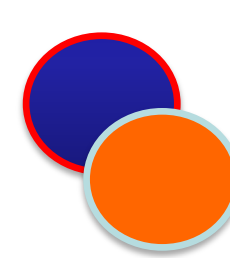
For the full statistics  
of 2011 ALICE  
identified 10 Anti-  
Alphas using  
TPC and TOF

STAR observed the  
Anti-Alpha in 2010:

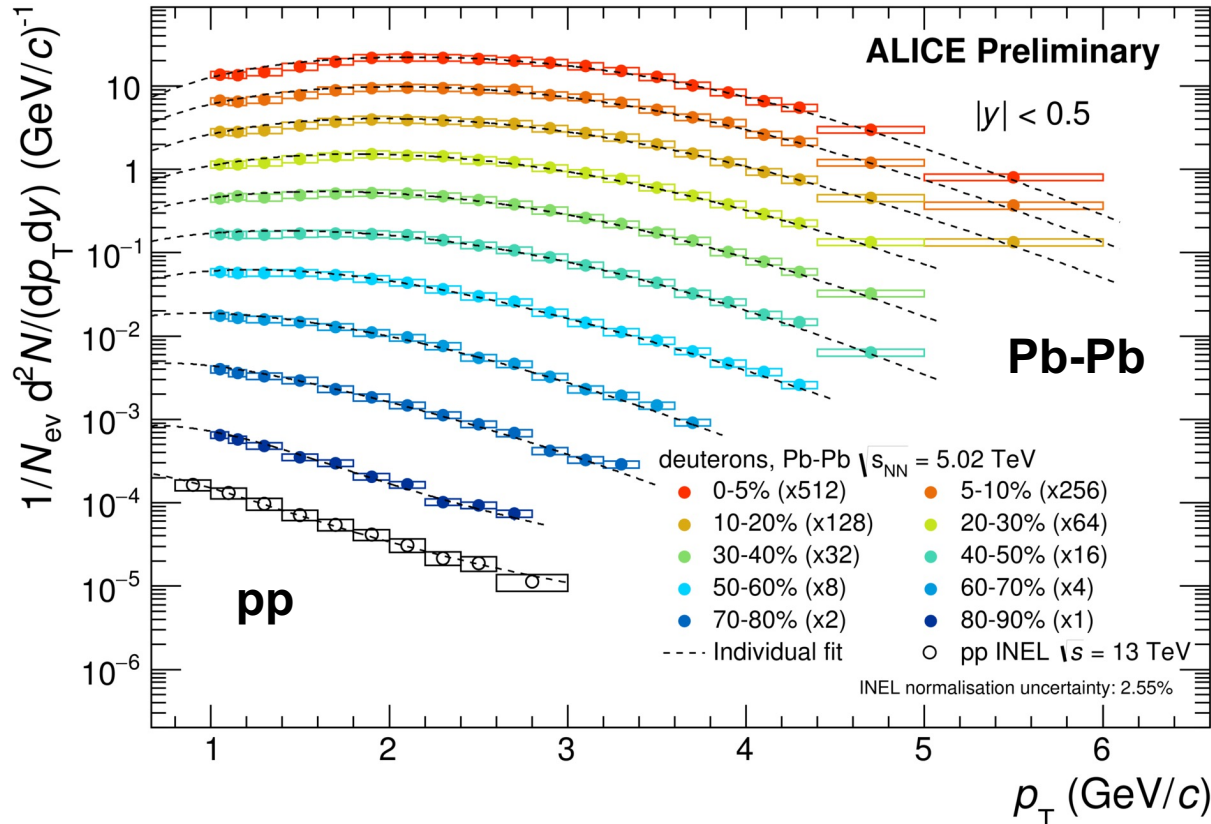
*Nature 473, 353 (2011)*



# Deuterons

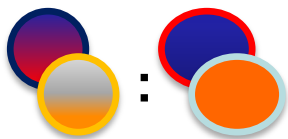
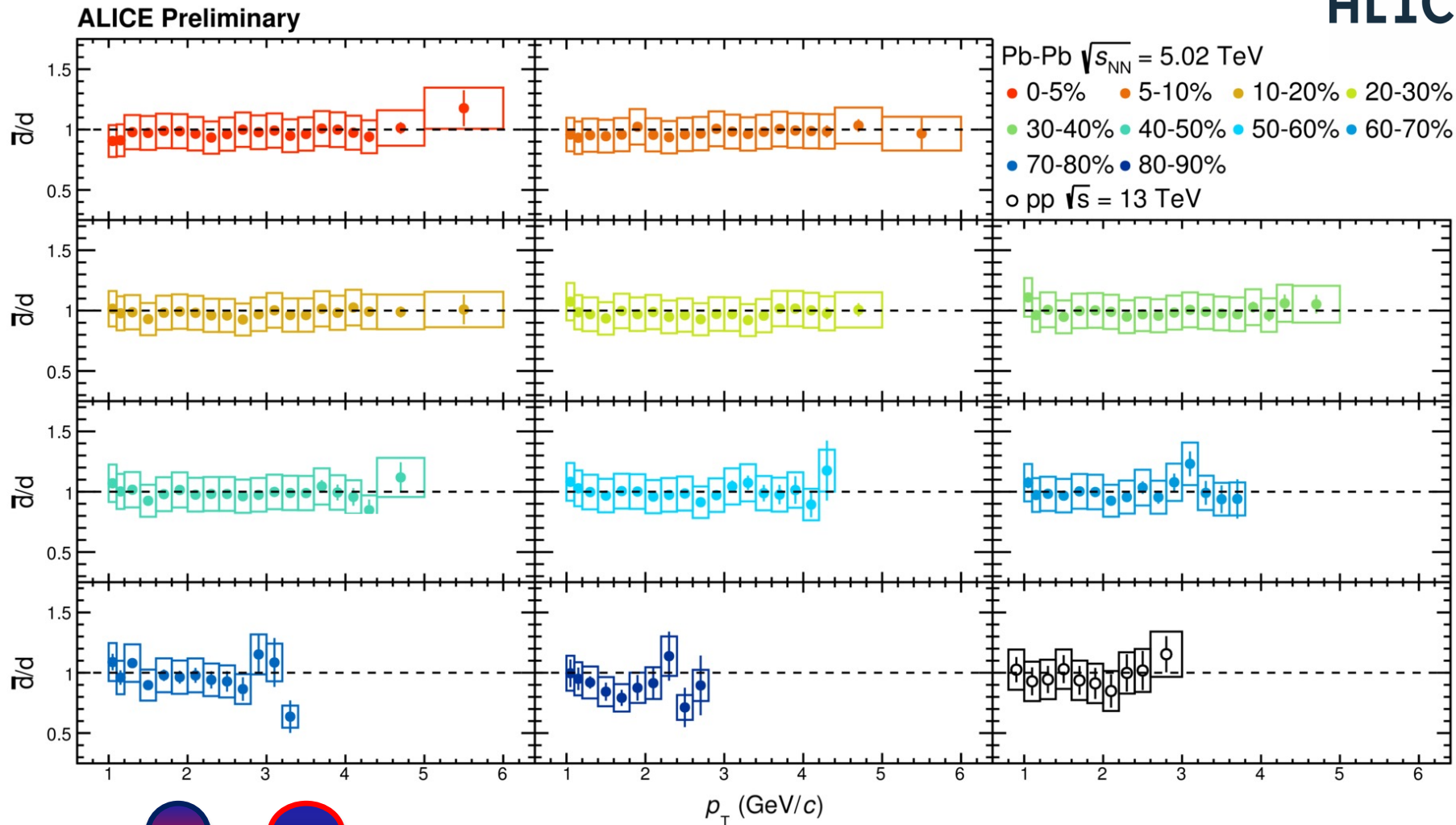


ALICE-PUBLIC-2017-006



- $p_T$  spectra getting harder for more central collisions (from pp to Pb-Pb)  $\rightarrow$  showing clear radial flow
- Blast-Wave fits describe the data in Pb-Pb very well
- No hint for radial flow in pp

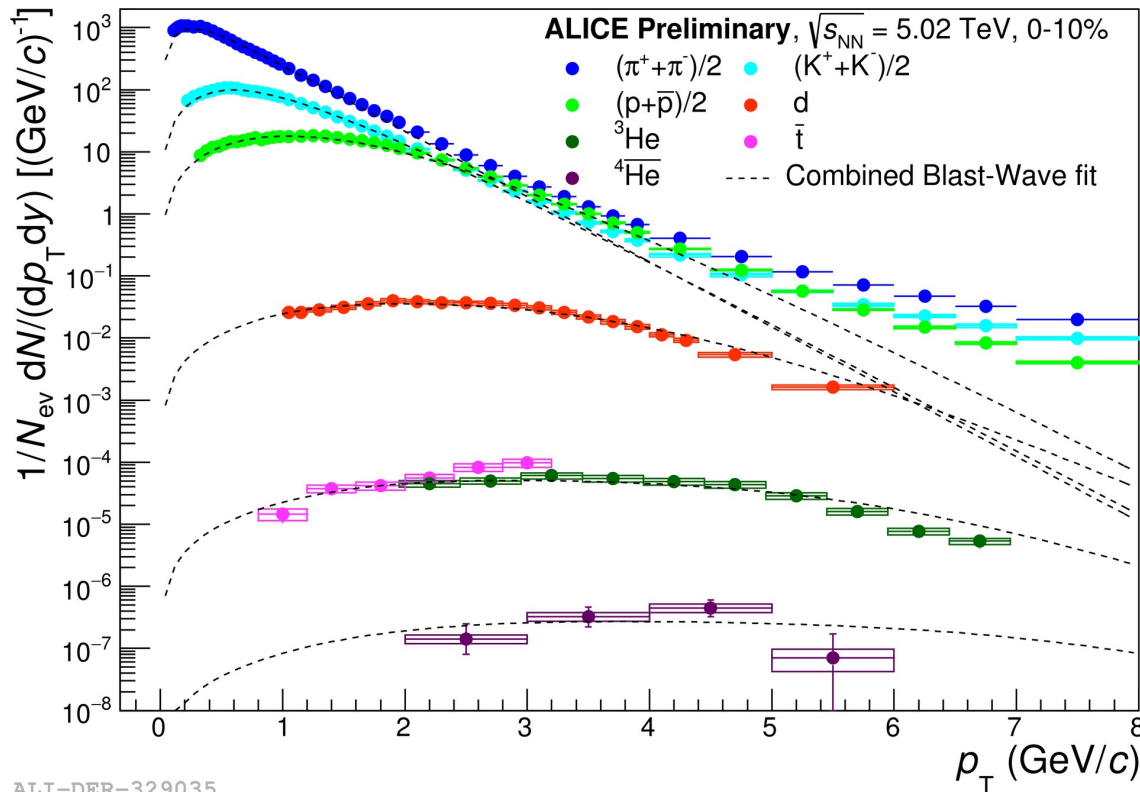
# (Anti-)Deuteron ratio



-ratios consistent with unity, as expected

# Combined Blast-Wave fit

ALICE Collaboration, [arXiv:1910.07678](https://arxiv.org/abs/1910.07678), see also [arXiv:2311.11758](https://arxiv.org/abs/2311.11758)



ALI-DER-329035

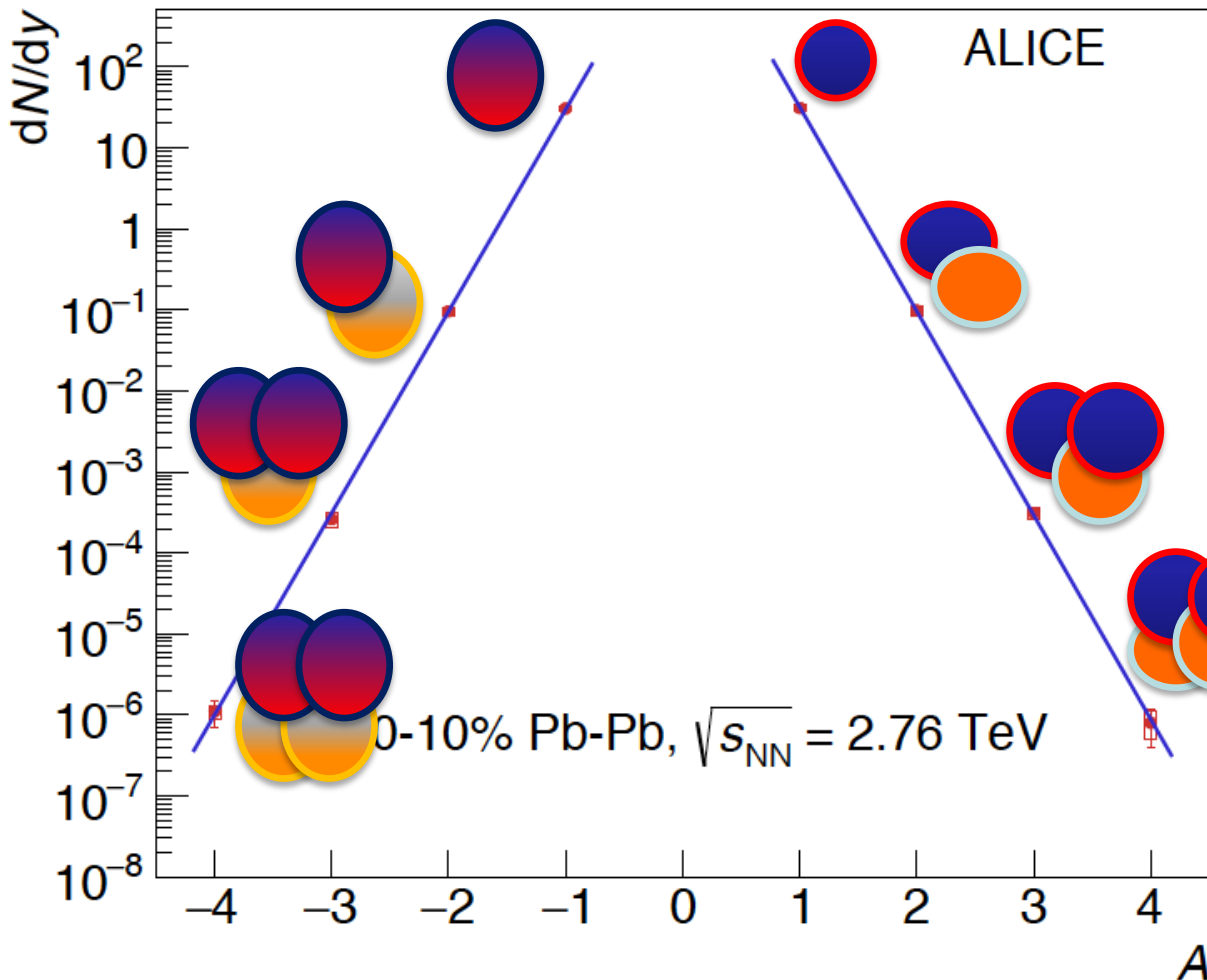
- Simultaneous Blast-Wave fit of  $\pi^+$ ,  $K^+$ ,  $p$ ,  $d$ ,  $t$ ,  ${}^3\text{He}$  and  ${}^4\text{He}$  spectra for central Pb-Pb collisions leads to values for  $\langle\beta\rangle$  and  $T_{kin}$  close to those obtained when only  $\pi, K, p$  are used

- All particles are described rather well with this simultaneous fit

# Mass dependence



ALICE



- Production of (anti-) nuclei is following an **exponential**, and decreases with mass as expected from thermal model
- In Pb-Pb the „penalty factor“ for each additional baryon  $\sim 300$  (for particles and anti-particles)

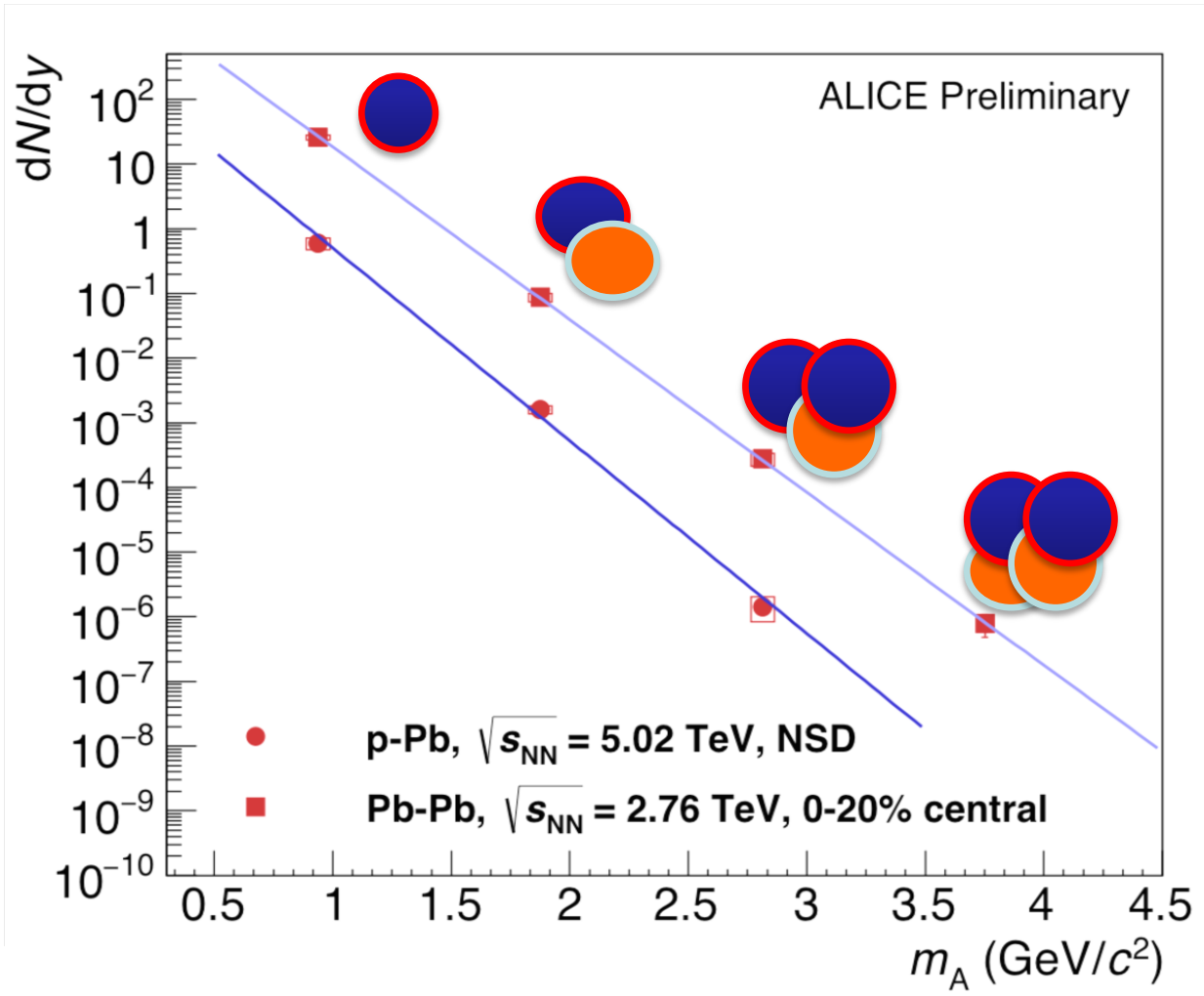
ALICE Collaboration, *arXiv:1710.07531*, NPA 971, 1 (2018)



# Mass dependence

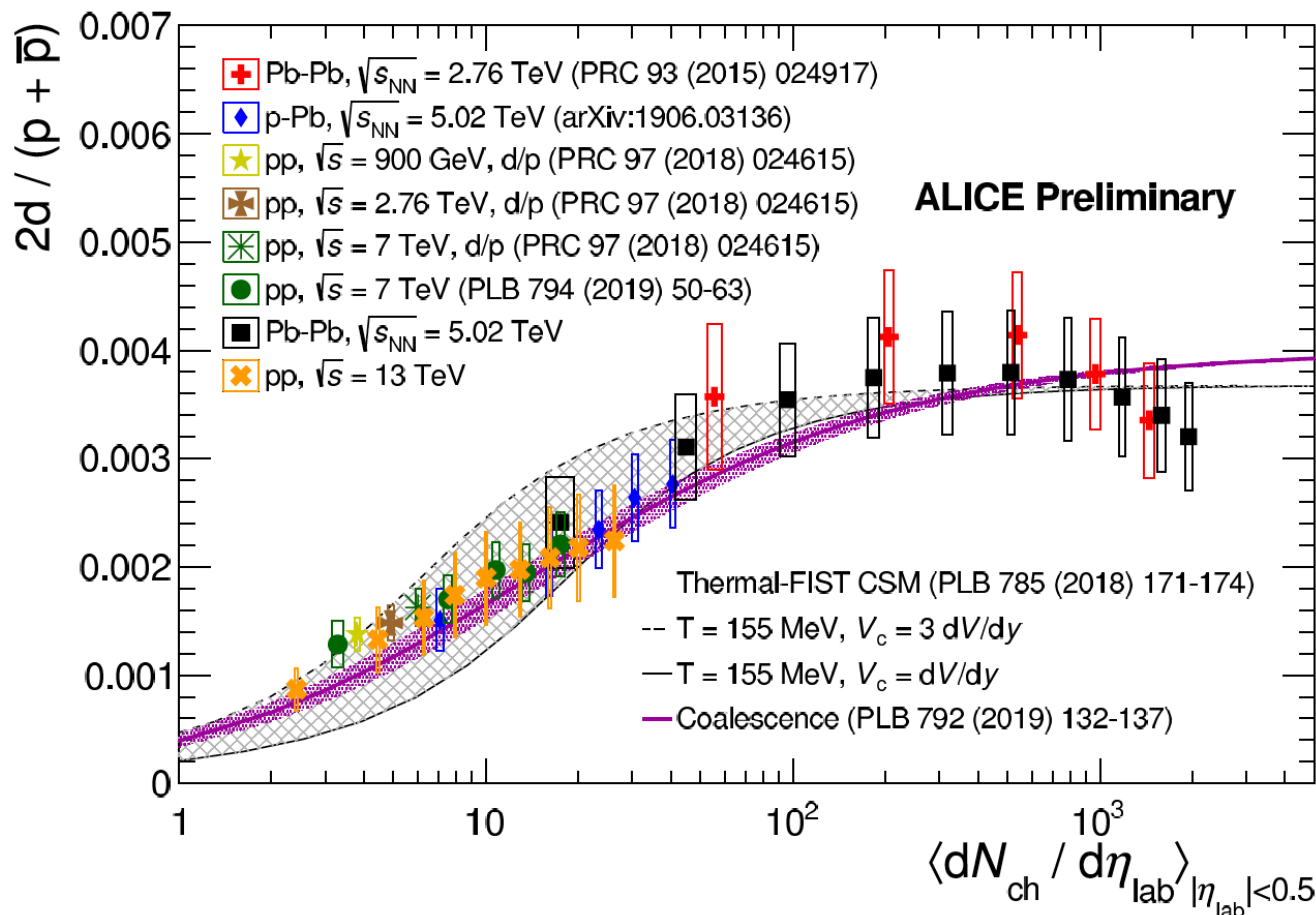


ALICE



- Production of (anti-) nuclei is following an **exponential**, and decreases with mass as expected from thermal model
- In Pb-Pb the „penalty factor“ for each additional baryon  $\sim 300$ , in p-Pb  $\sim 600$  and in pp  $\sim 1000$

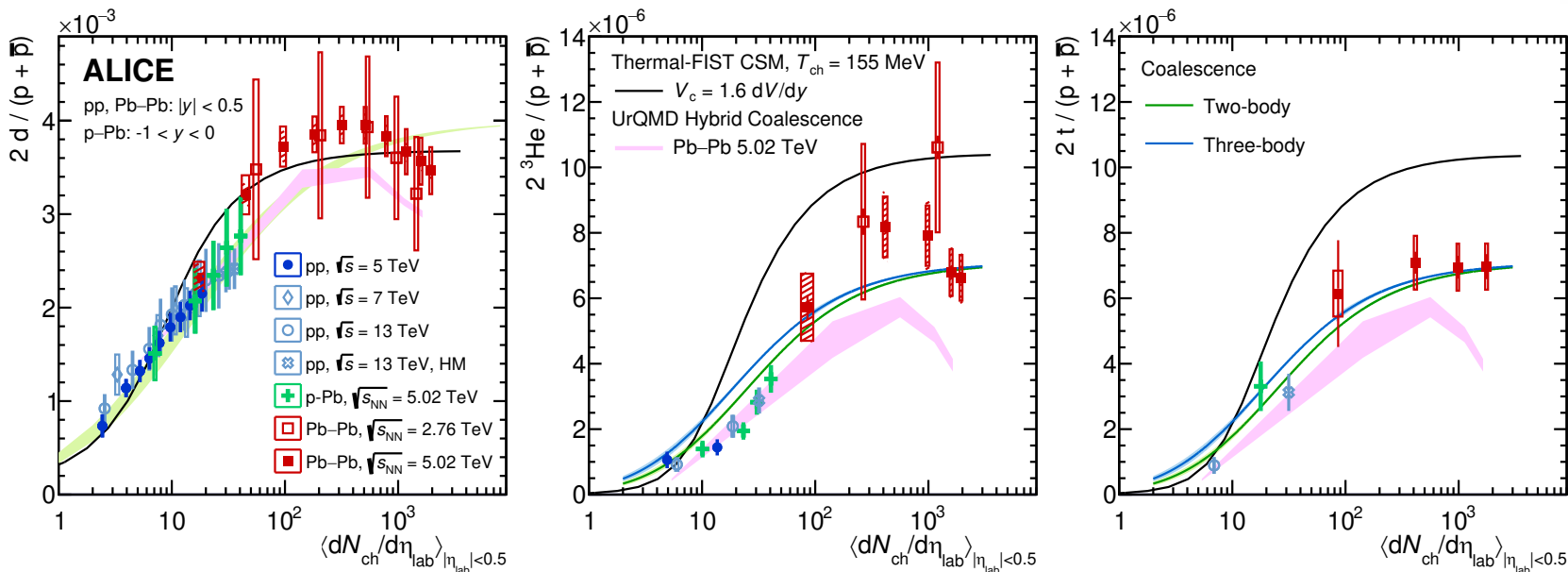
# d/p vs. multiplicity



d/p ratio rather well described by coalescence and (canonical) thermal model

# Ratios vs. multiplicity

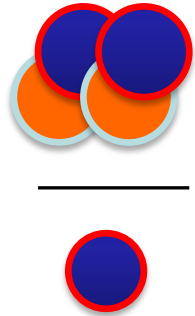
ALICE Collaboration, arXiv:2211.14015, Phys.Rev.C 107 (2023) 064904



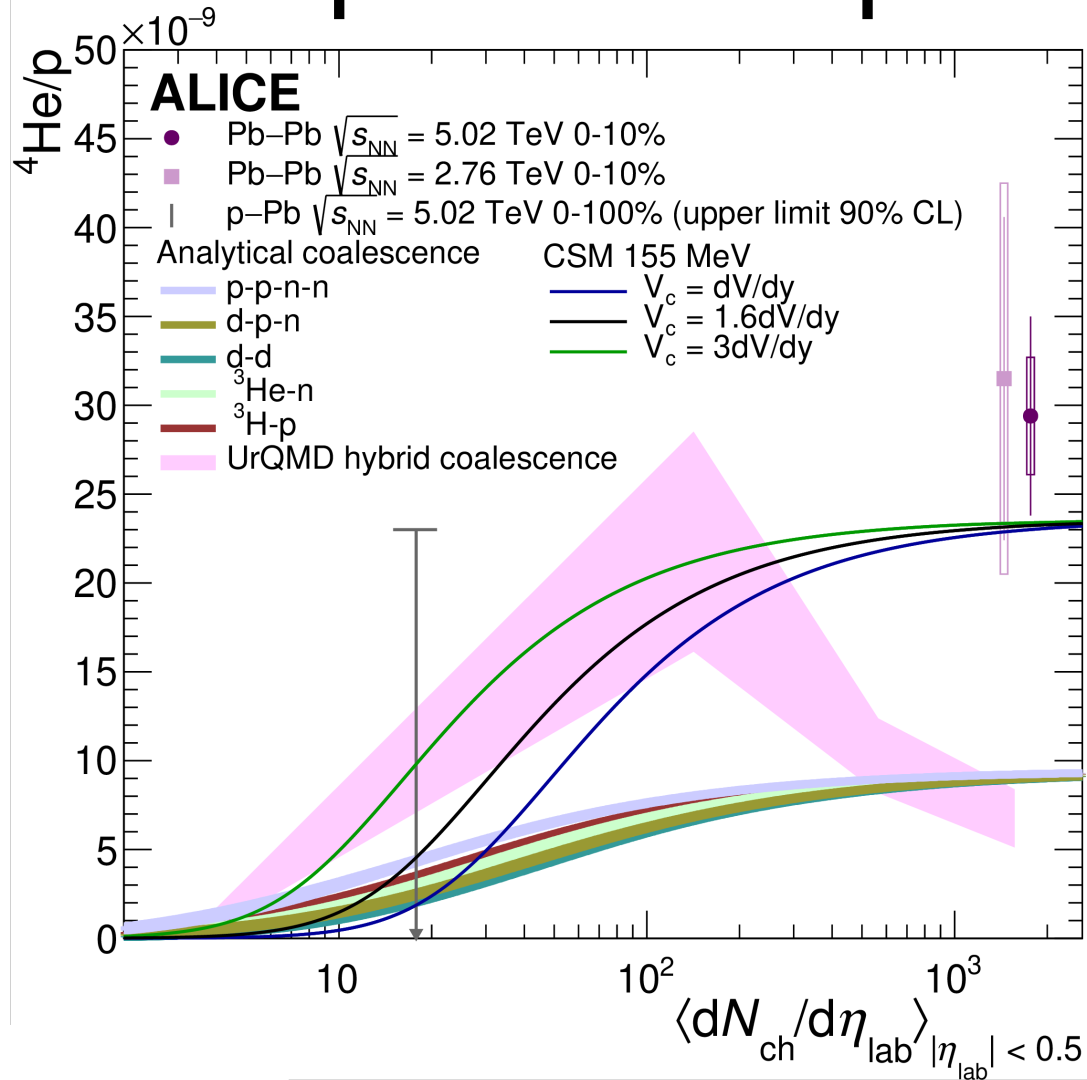
- d/p ratio rather well described by coalescence and (canonical) thermal model
- Some tension for  $^3\text{He}/p$  and  $^3\text{H}/p$  over  $p_T$

Models:  
Coalescence: K.J. Sun, C.M. Ko, BD, PLB 792 (2019) 132  
CSM: V.Vovchenko, BD, H. Stöcker, PLB 785 (2018) 171  
UrQMD Hybrid: T. Reichert, J. Steinheimer, V.Vovchenko, BD, M. Bleicher, Phys. Rev. C 107 (2023) 1





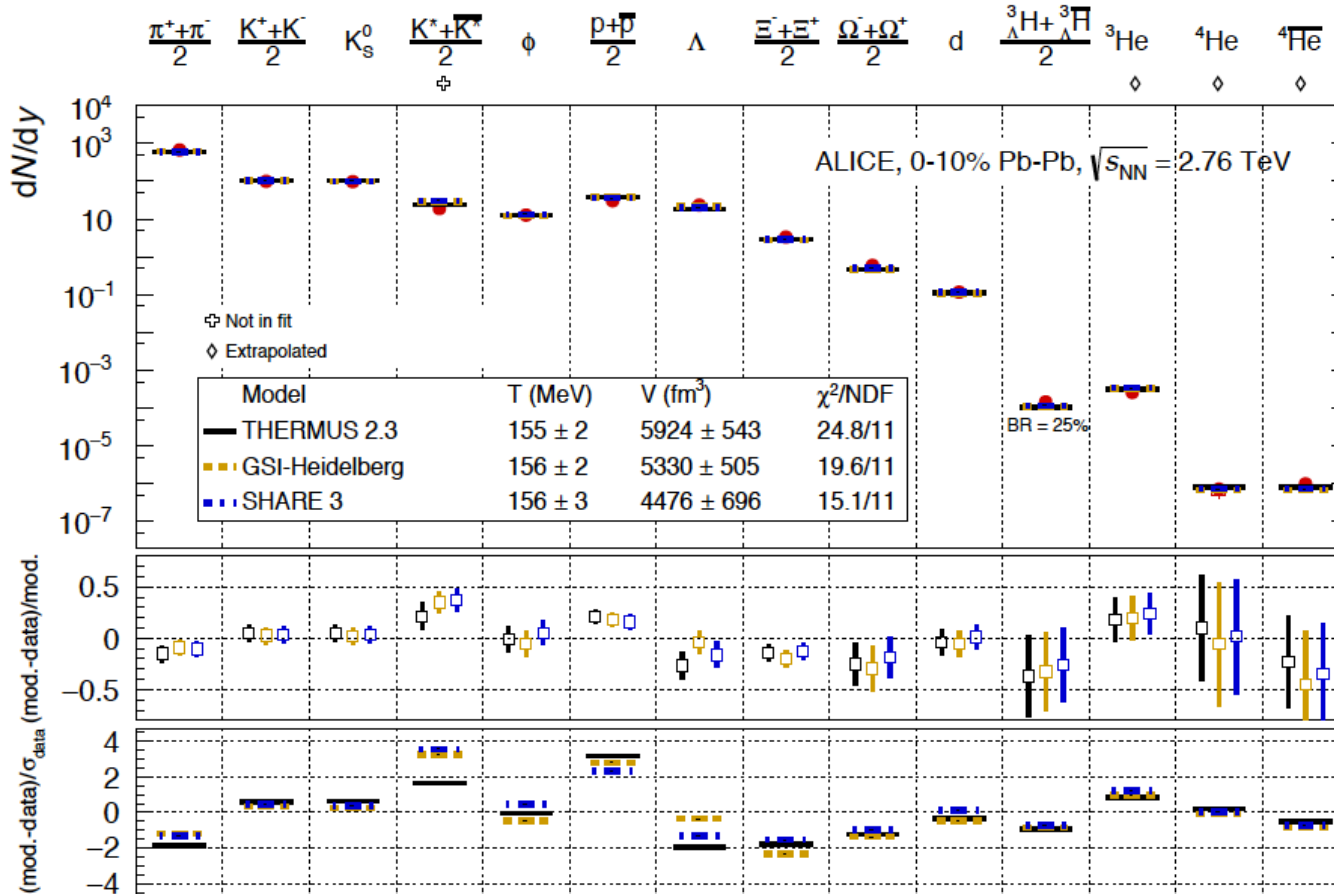
# $^4\text{He}/p$ vs. multiplicity



ALICE Collaboration, arXiv:2311.11758  
Phys. Lett. B 858 (2024) 138943

$^4\text{He}/p$  ratio significantly better described by the thermal model

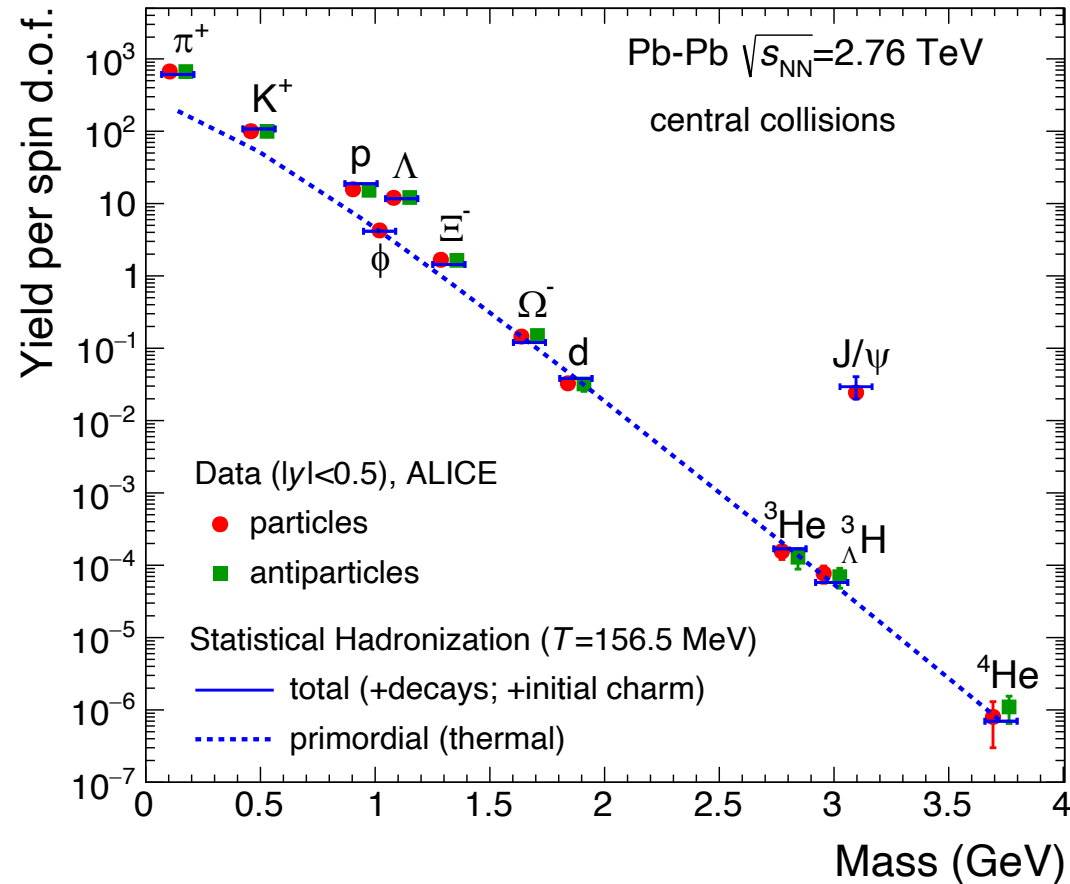
# Thermal model



- Different model implementations describe the production probability, including light nuclei and hyper-nuclei, rather well at a temperature of about  $T_{ch} = 156$  MeV

# Thermal model

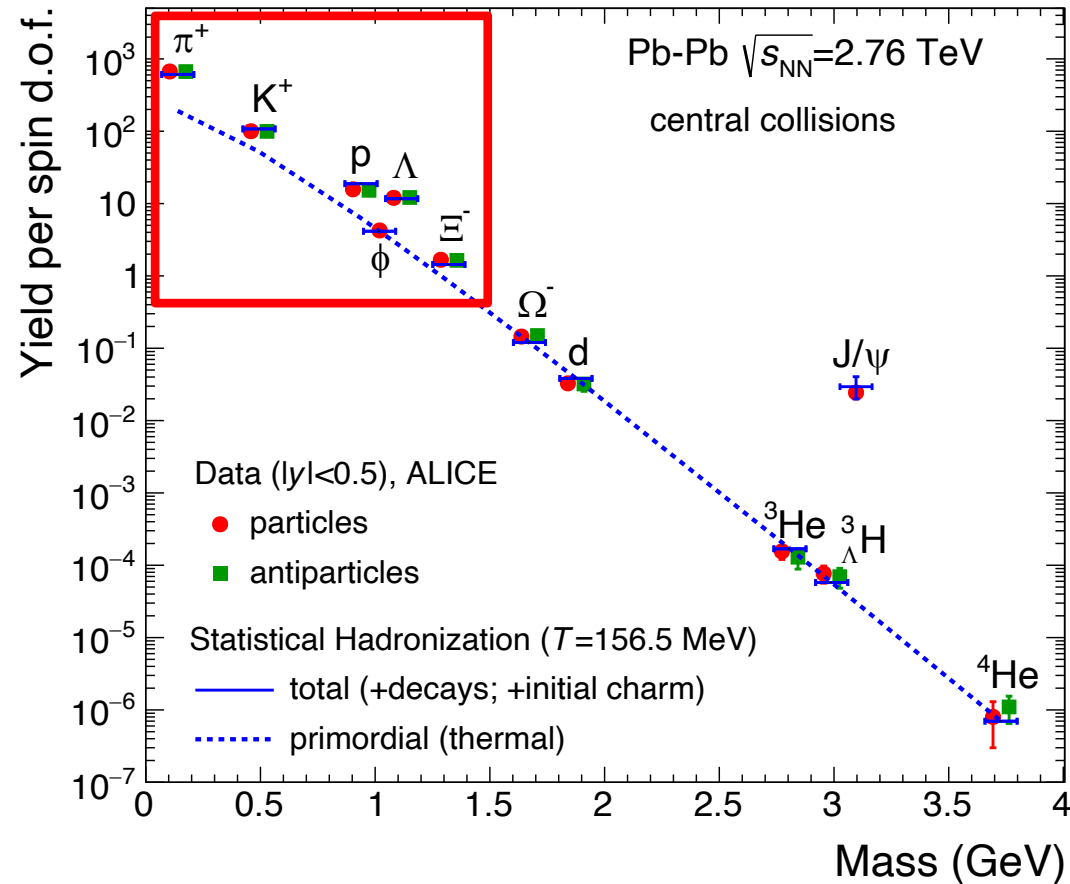
- For the thermal model description of production yields, feed-down is an important ingredient
- All light hadron production yields are populated strongly by resonances
- Seems to not be the case for (hyper-)nuclei



*A. Andronic et al., Phys.Lett.B 797 (2019) 134836*

# Thermal model

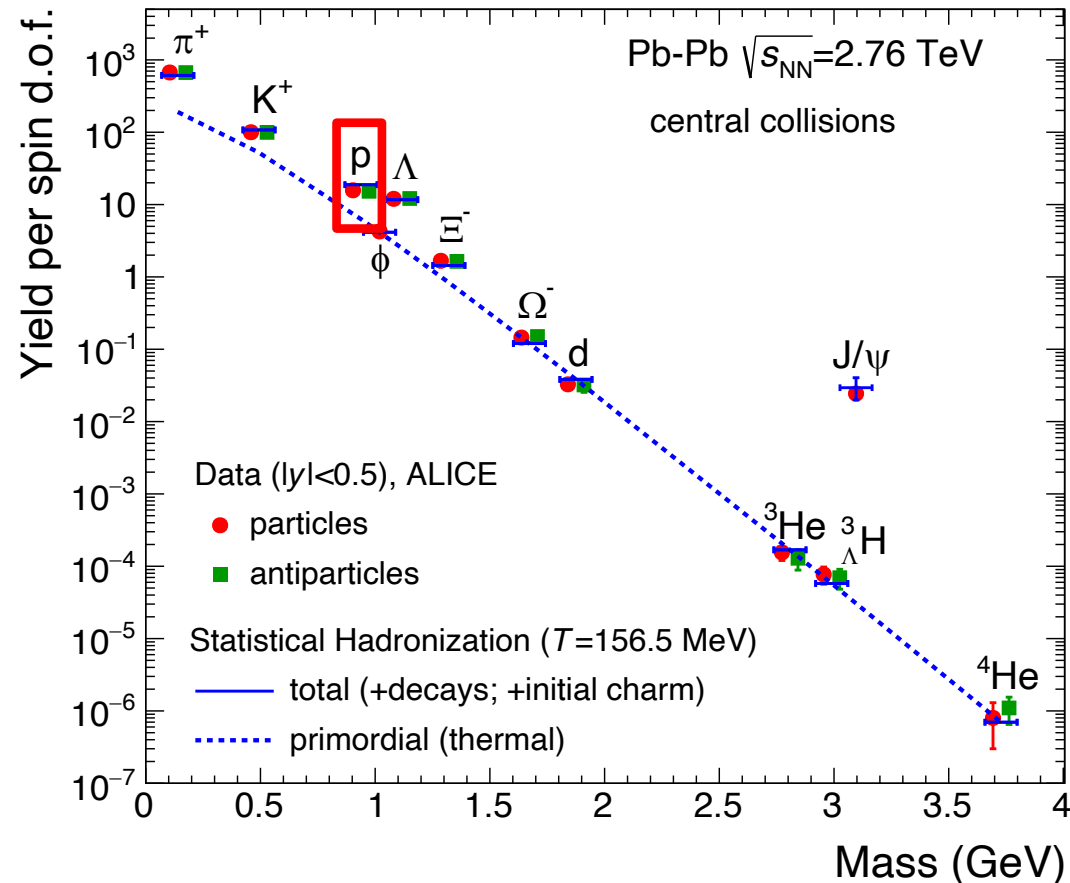
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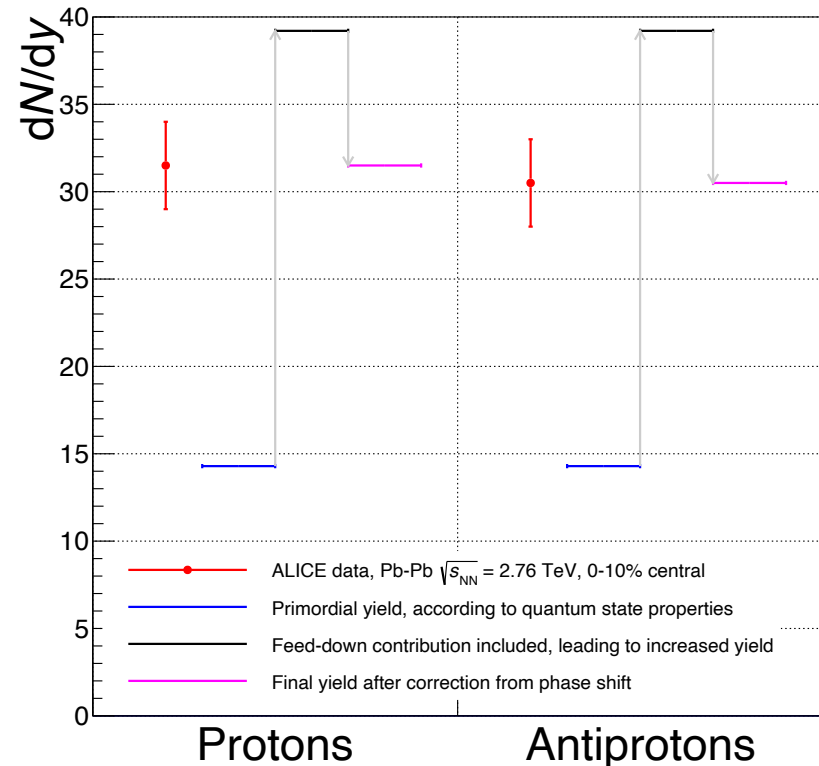


*A. Andronic et al., Phys.Lett.B 797 (2019) 134836*

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*BD, G. Röpke, D. Blaschke,  
Phys. Rev. C 106 (2022) 044908*

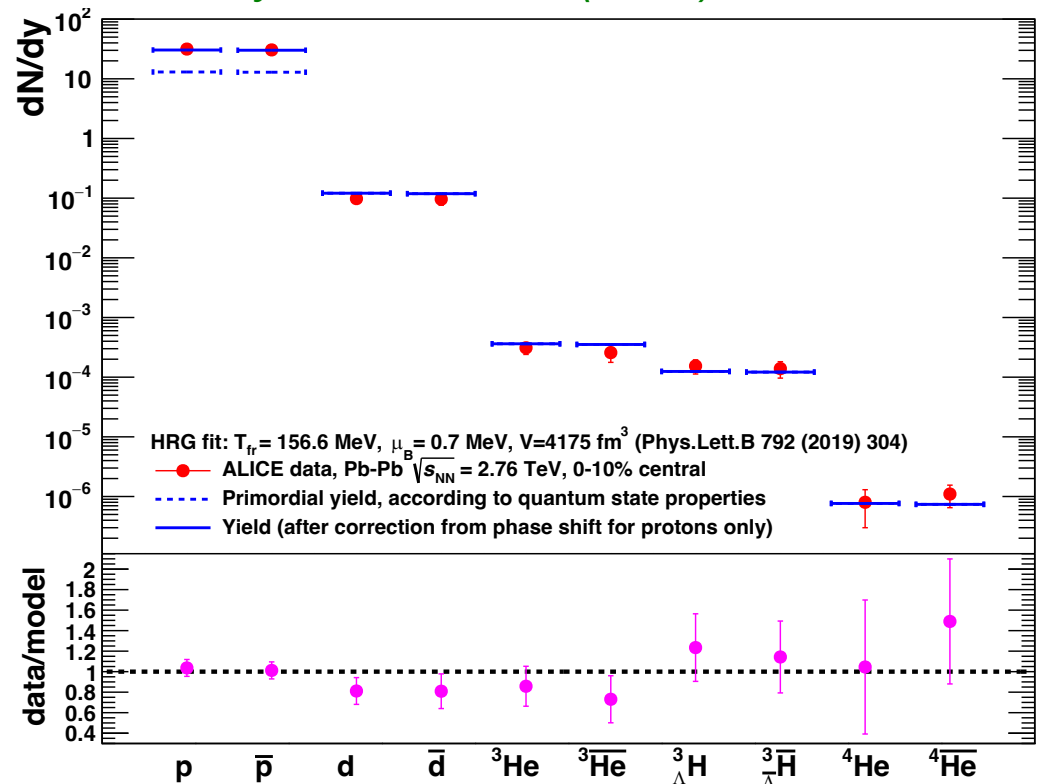


*A. Andronic et al., Phys.Lett.B 797 (2019) 134836;  
Nature 561 (2018) 7723, 321; Phys.Lett.B 697 (2011) 203;  
Phys.Lett.B 792 (2019) 304*

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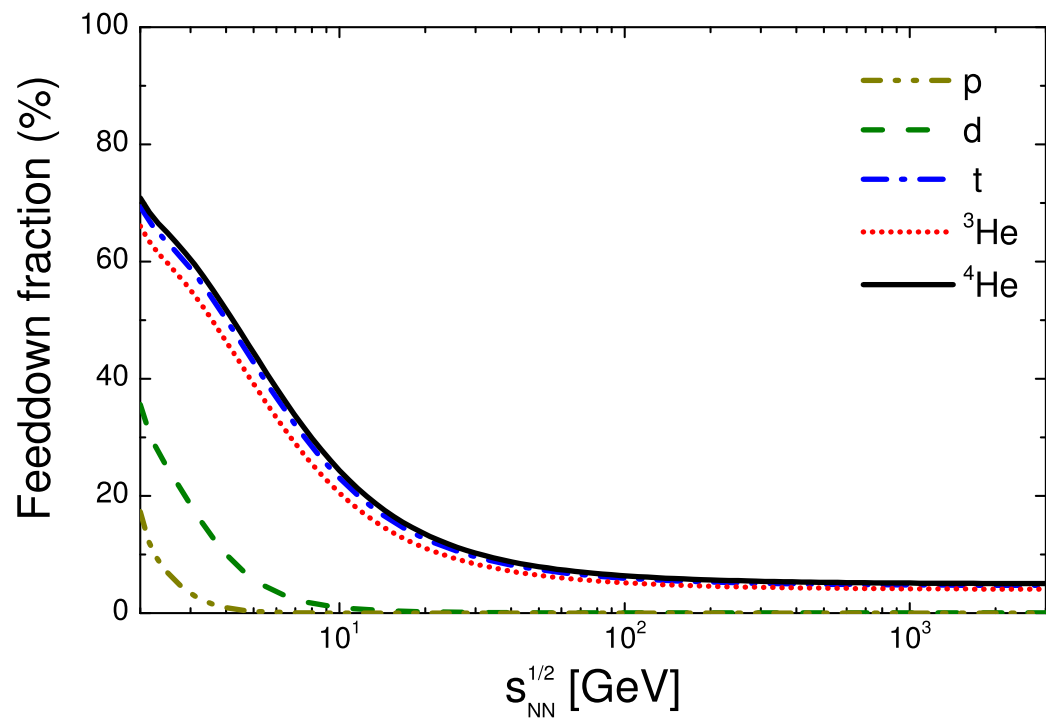


*A. Andronic et al., Phys. Lett. B 797 (2019) 134836;  
Nature 561 (2018) 7723, 321; Phys. Lett. B 697 (2011) 203;  
Phys. Lett. B 792 (2019) 304*

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- For the thermal model description of production yields, feed-down is an important ingredient
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- Seems to not be the case for (hyper-)nuclei at LHC

*V. Vovchenko, BD, B. Kardan, M. Lorenz, H. Stoecker, Phys.Lett.B 809 (2020) 135746*

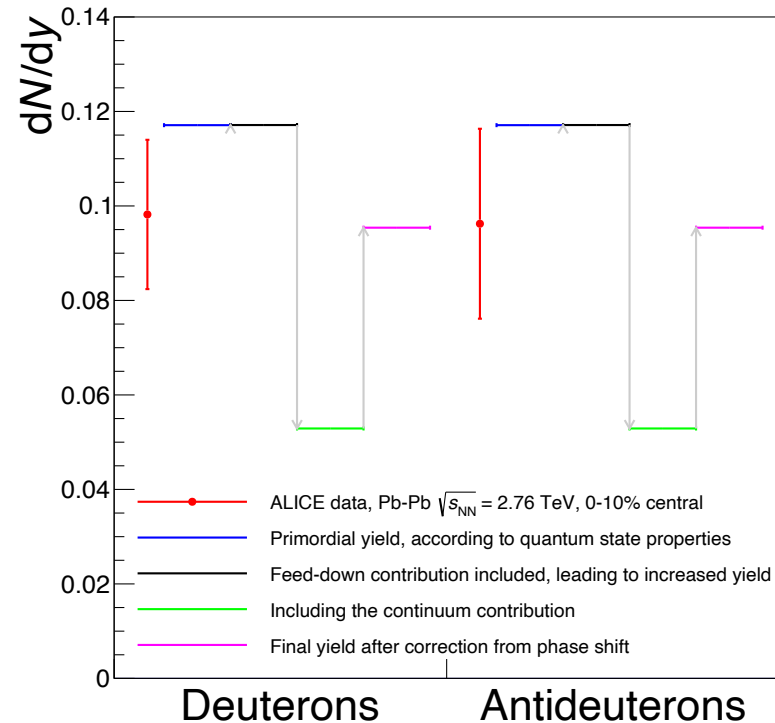




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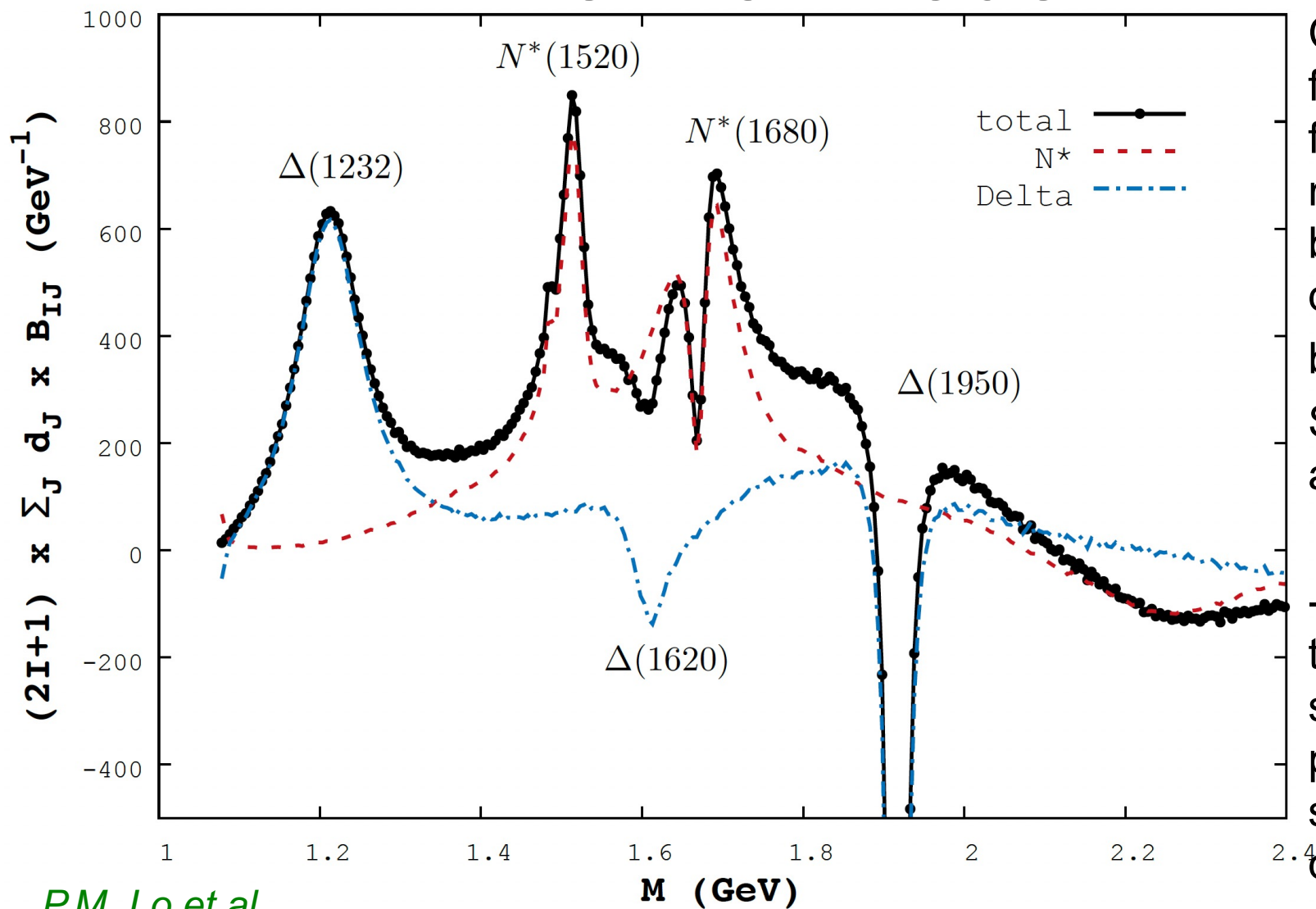
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Nature 561 (2018) 7723, 321; Phys.Lett.B 697 (2011) 203;  
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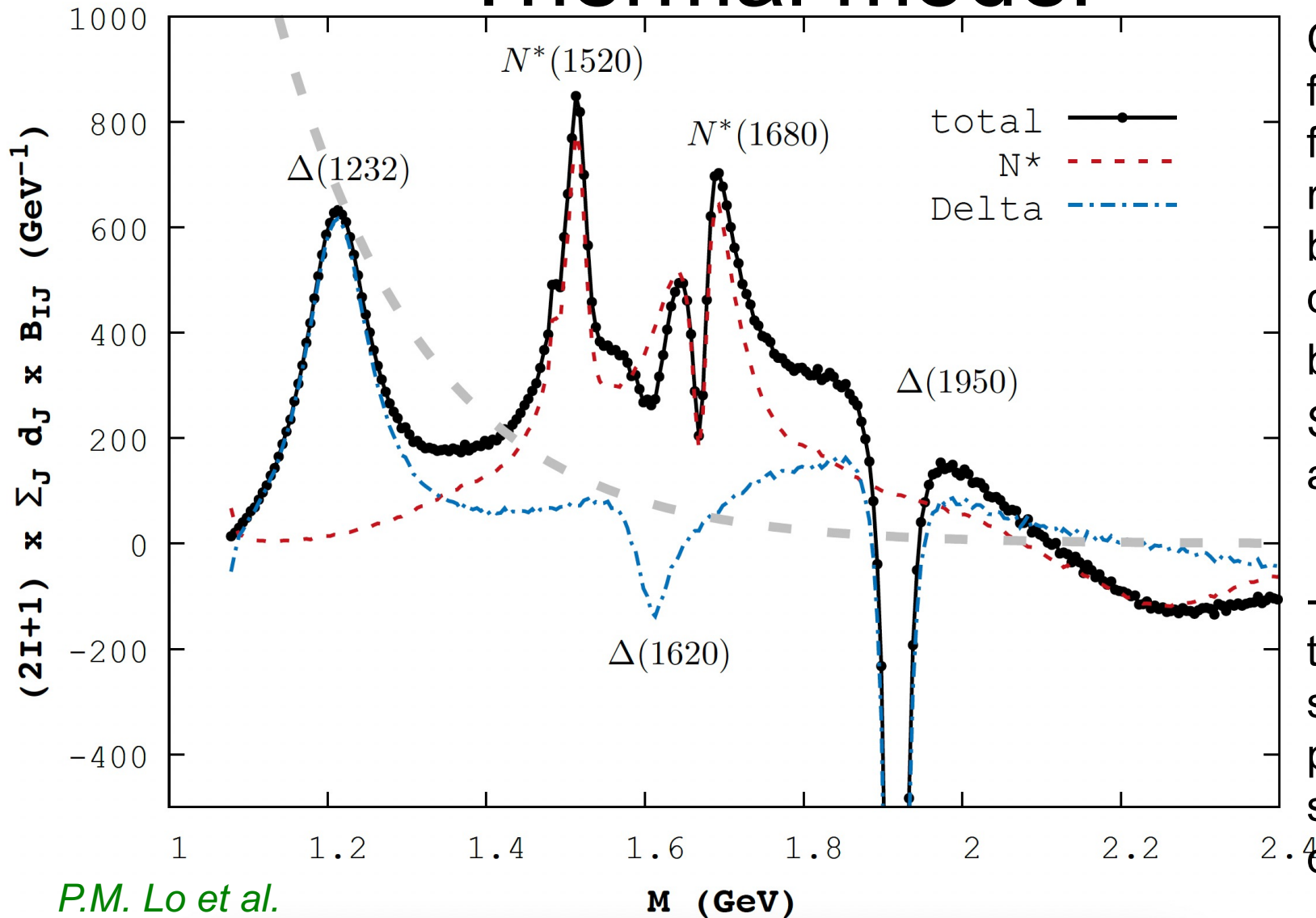


Contribution  
for protons  
from  
resonances  
best  
described  
by the  
S-Matrix  
approach

-connecting  
to phase  
shifts from  
 $p-\pi$   
scattering  
data

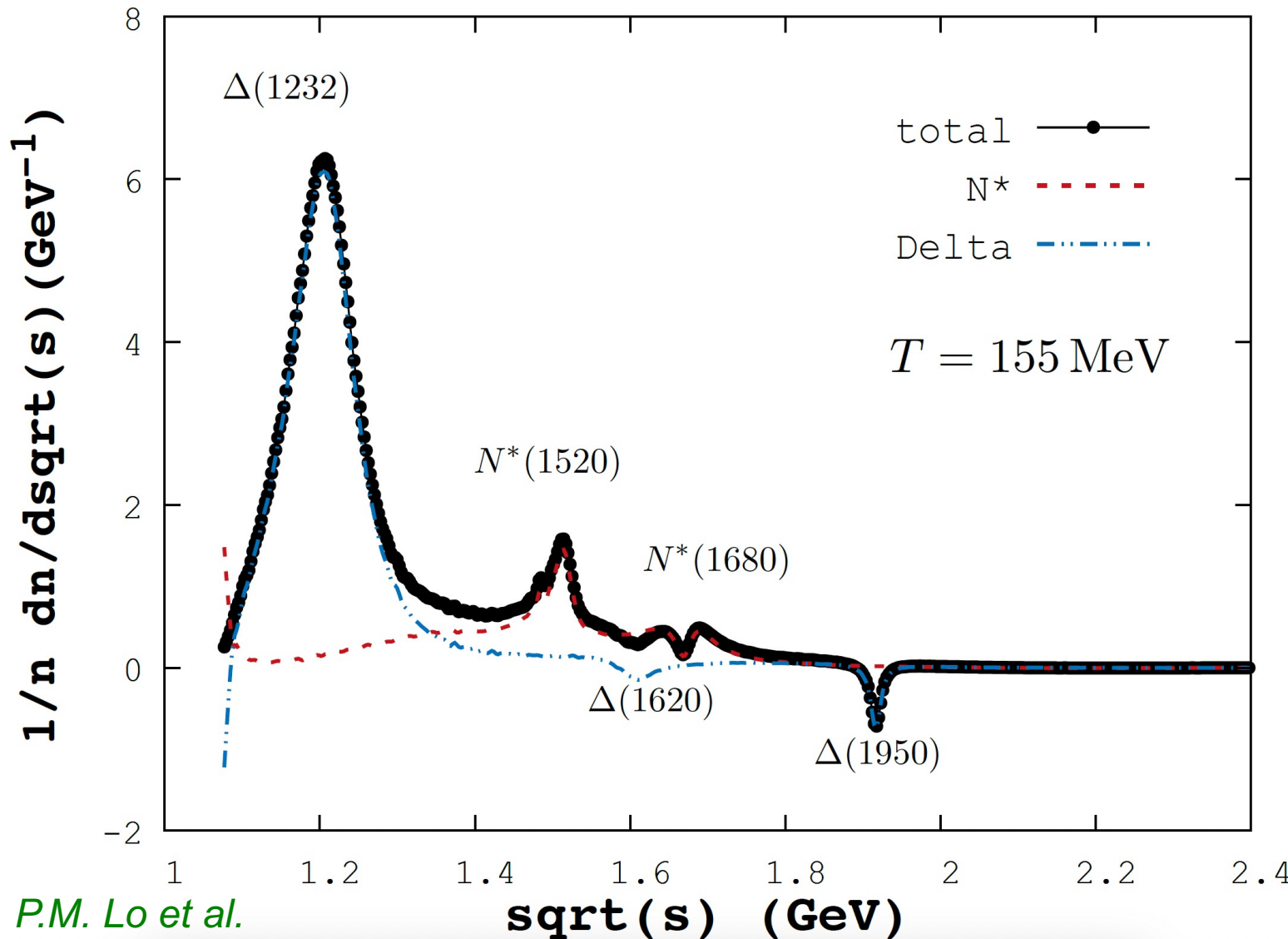
*P.M. Lo et al.*

# Thermal model



P.M. Lo et al.

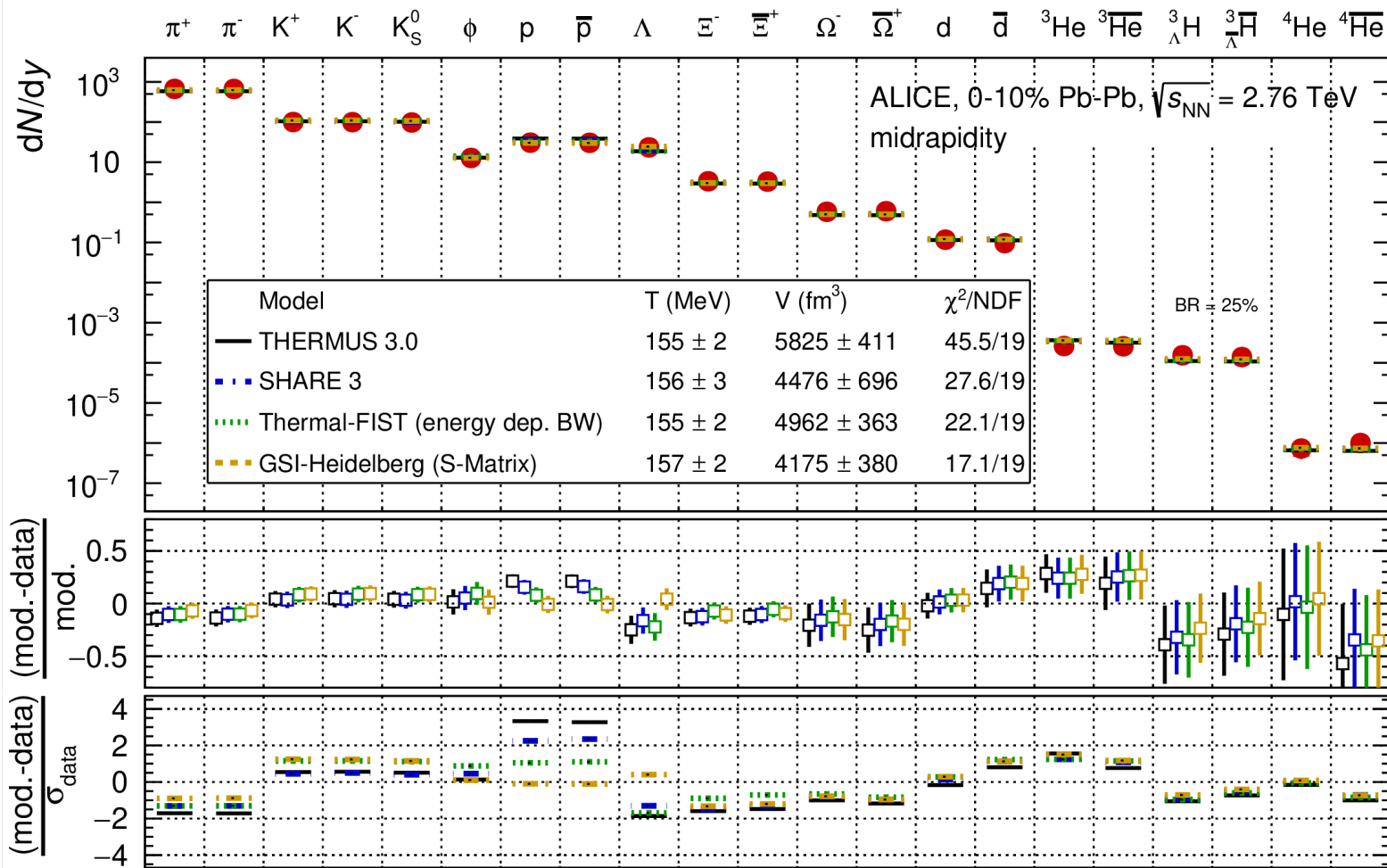
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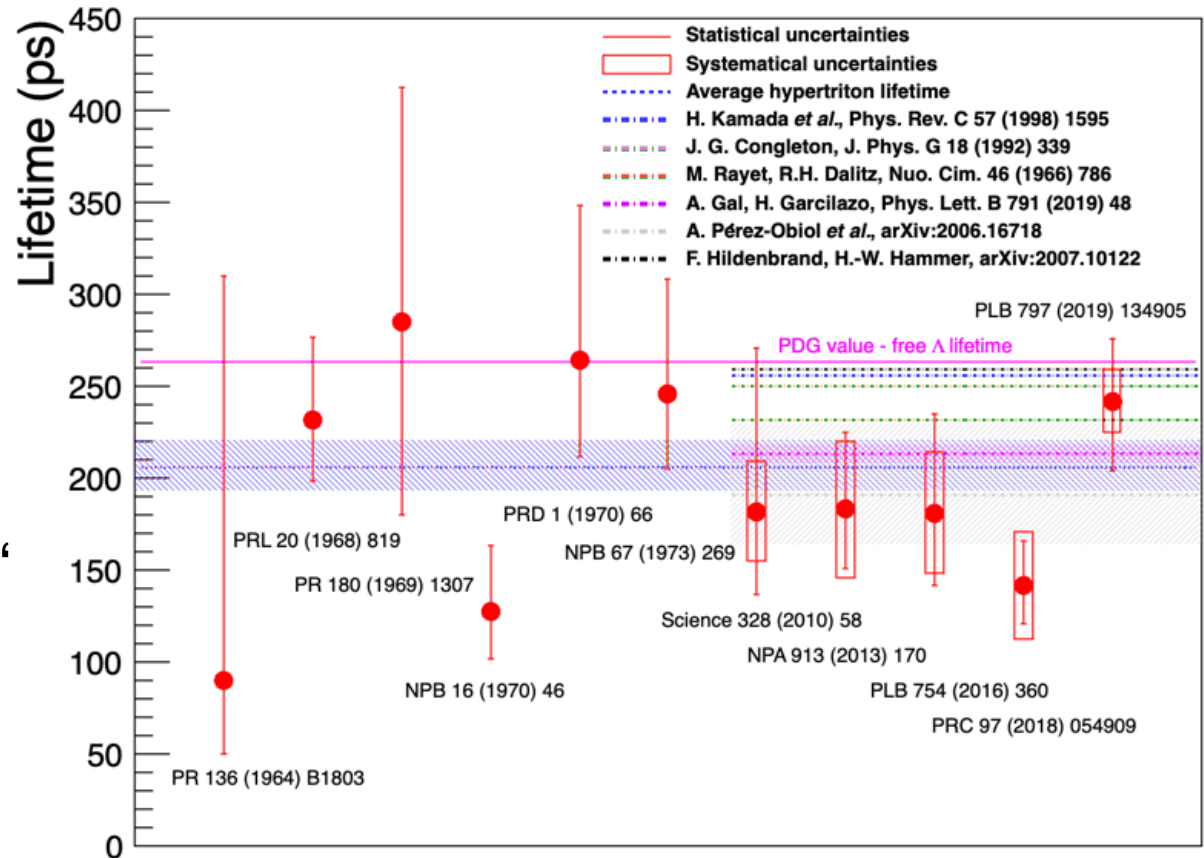
-connecting  
to phase  
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 $\rho$ - $\pi$   
scattering  
data

Energy-dependent Breit-Wigner works similar well

ALICE Collaboration, [arXiv:2211.04384 \[nucl-ex\]](https://arxiv.org/abs/2211.04384)

# Hypertriton „Puzzle“

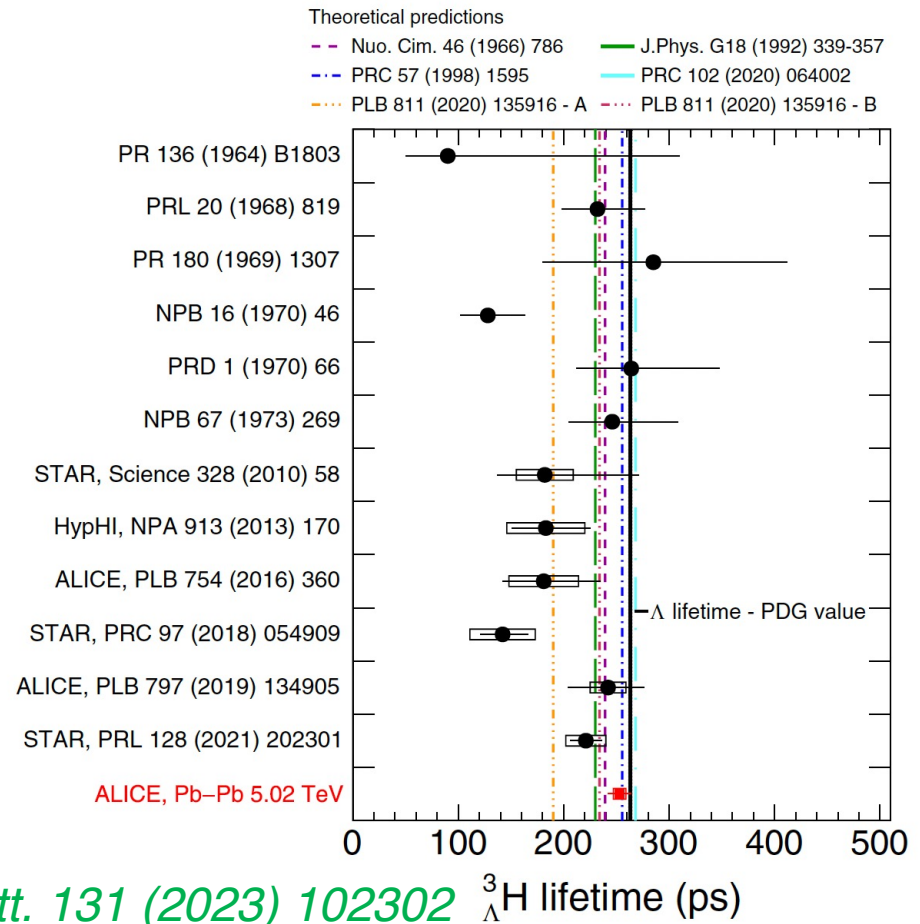
- Recently measured lifetimes are significantly below the lifetime of the free  $\Lambda \rightarrow$  new ALICE results agree with the world average of all known measurements and with the free  $\Lambda$  lifetime
- Most recent calculations include „final-state“ interaction and agree well with the data



*BD, Eur. Phys. J 56 (2020) 258*

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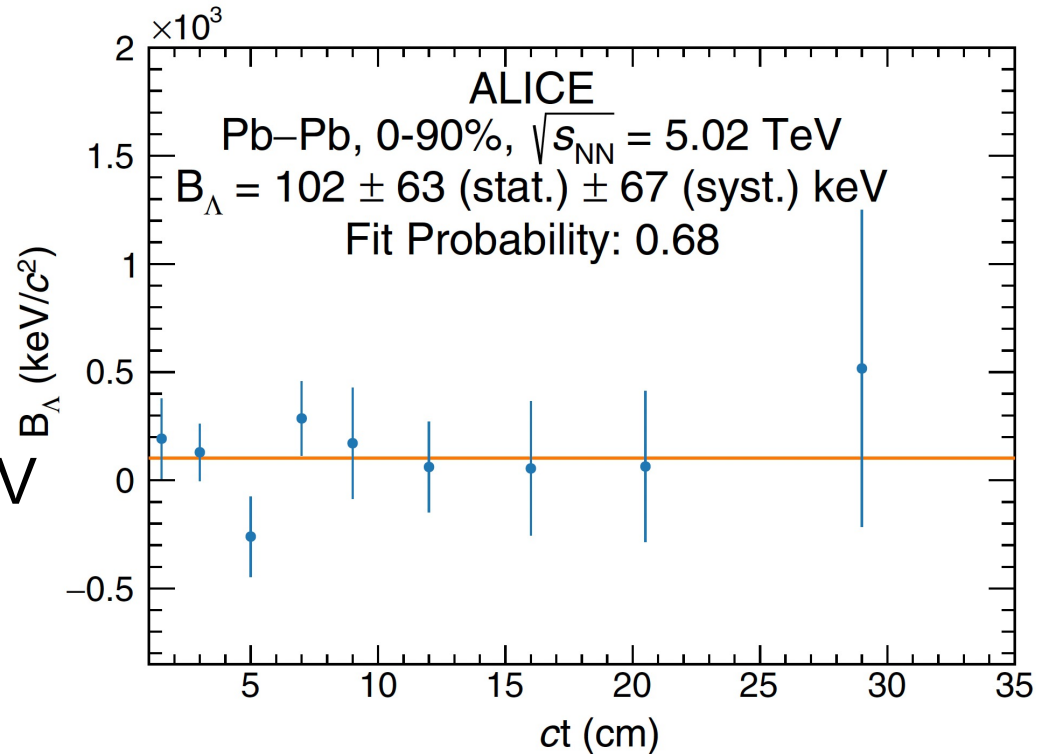


ALICE Collaboration, Phys. Rev. Lett. 131 (2023) 102302



# Binding Energy

- Current studies show a better constraint and smaller statistical uncertainties
- The value obtained by this fit is  
 $B_{\Lambda} = (102 \pm 63 \pm 67) \text{ keV}$
- Is compatible with the theoretical predictions

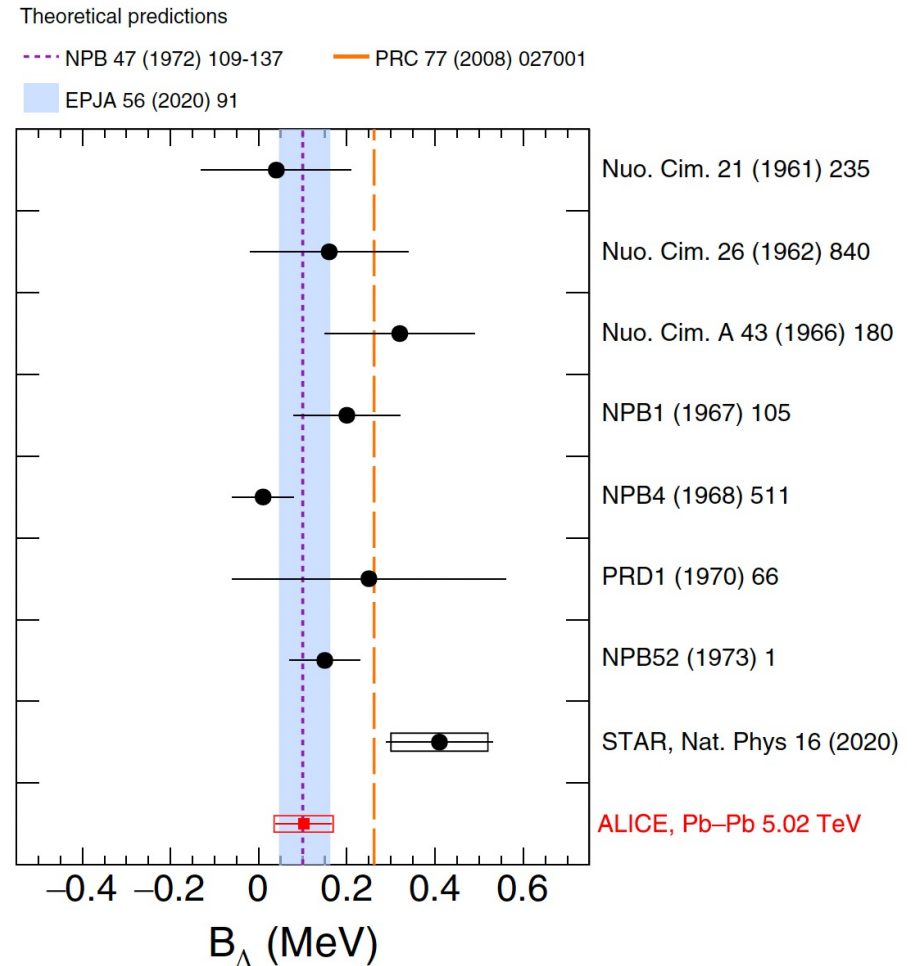


*ALICE Collaboration, Phys. Rev. Lett. 131 (2023) 102302*



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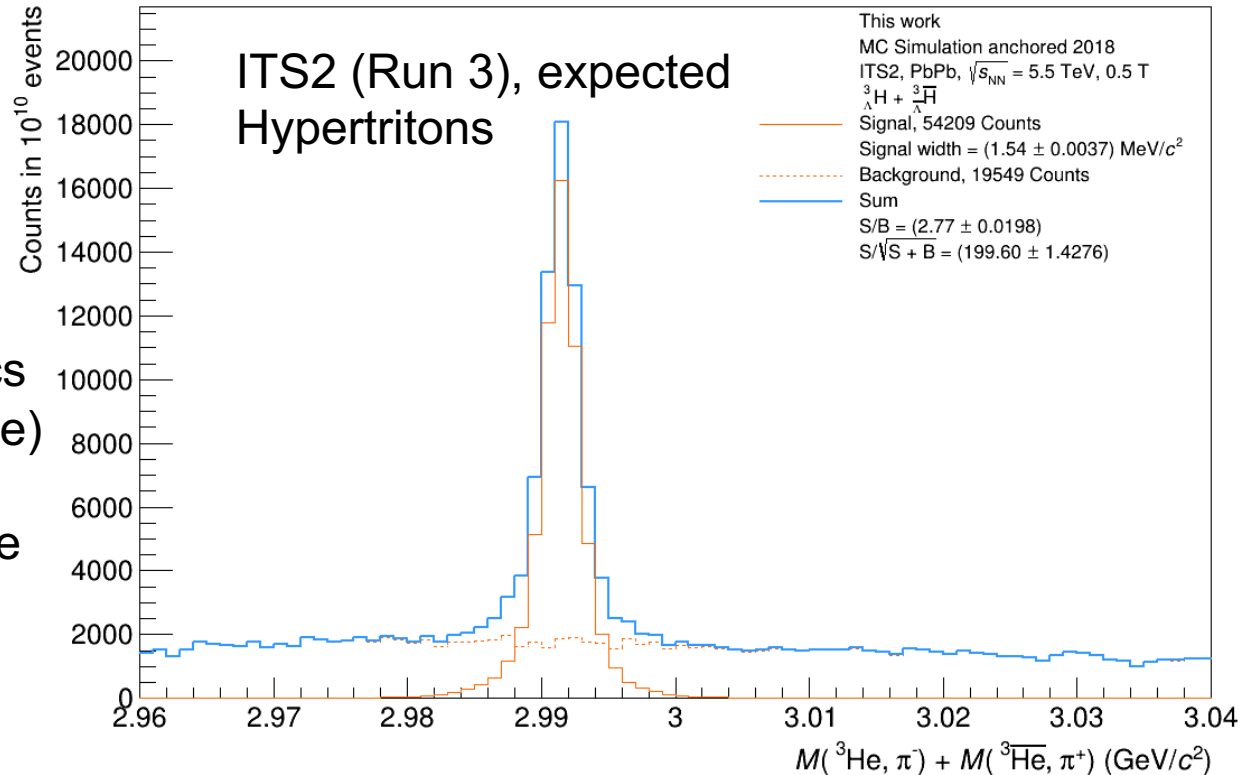
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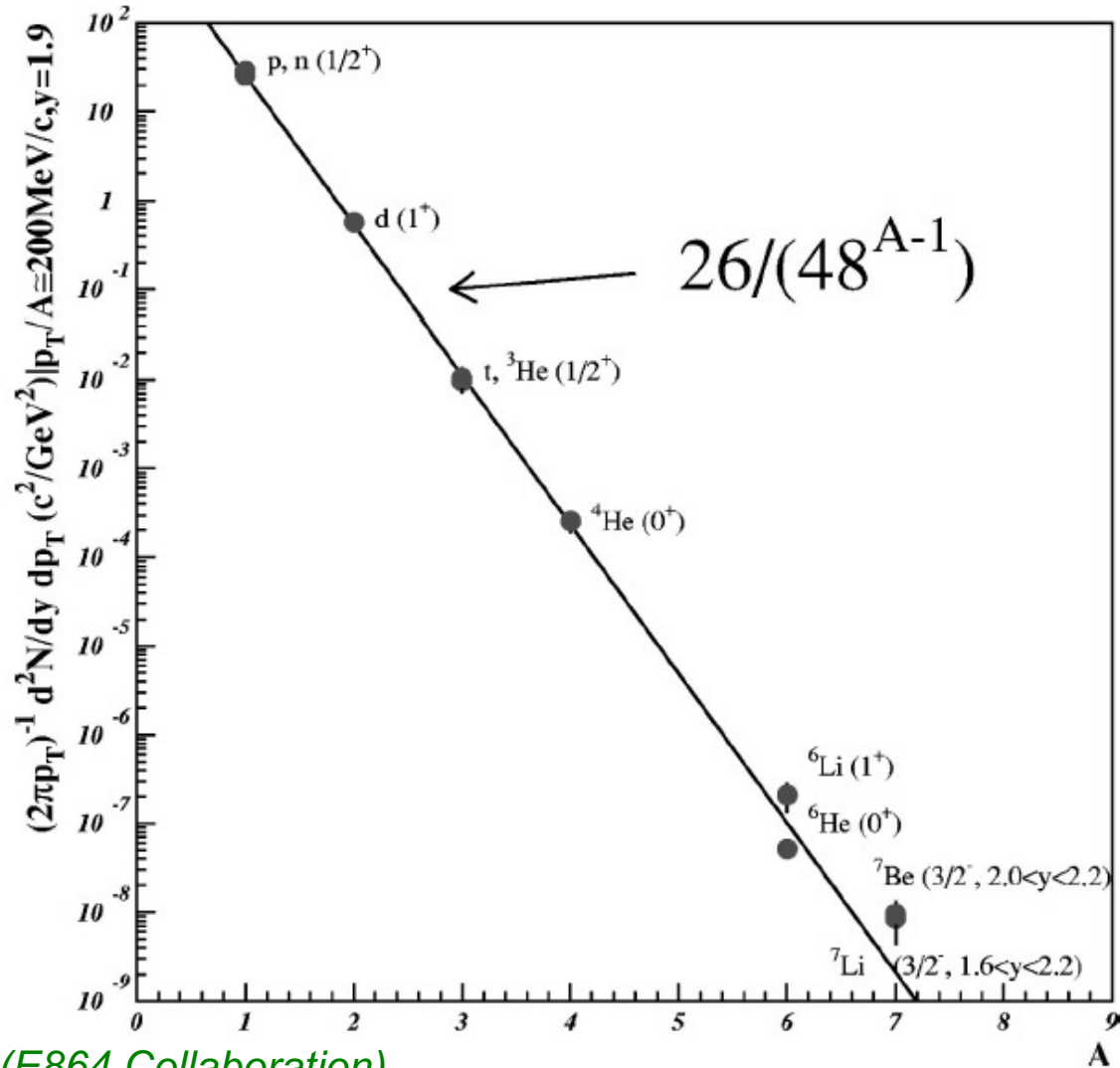
*ALICE Collaboration, Phys. Rev. Lett. 131 (2023) 102302*

# Expectations

- Run 2 of the LHC has started in 2015 and for Pb-Pb collisions ~ factor 10 increase expected in statistics
- Run 3 & Run 4 of LHC will deliver much more statistics (50 kHz Pb-Pb collision rate)
- Upgraded ALICE detector will be able to cope with the high luminosity
- TPC Upgrade: GEMs for continuous readout
- ITS Upgrade: less material budget and more precise tracking for the identification of hyper-nuclei
- Physics which is now done for  $A = 2$  and  $A = 3$  (hyper-)nuclei will be done for  $A = 4$

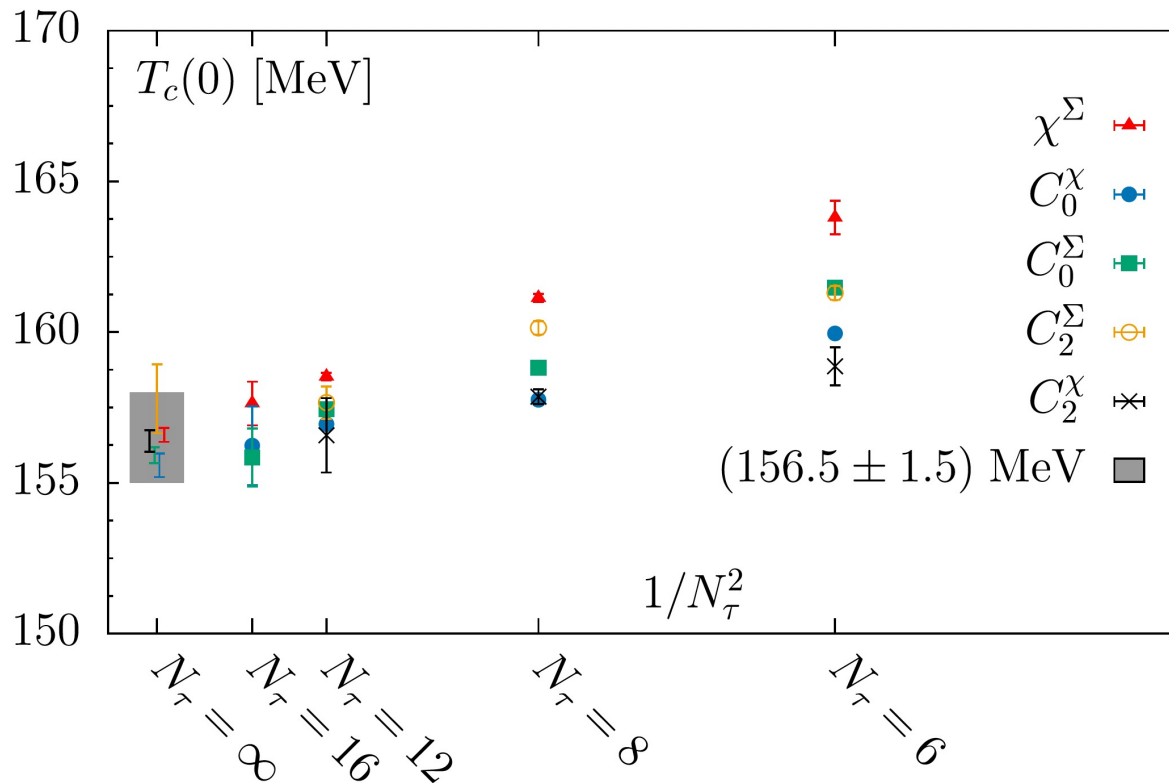


# E864 nuclei result



*T.A. Armstrong et al. (E864 Collaboration),  
Phys. Rev. C 61 (2000) 064908*

# Lattice QCD results



Lattice QCD  
tells us where  
to expect the  
phase  
transition

**Critical energy density:**  
 $\epsilon_c = 0.34 \pm 0.16 \text{ GeV/fm}^3$

**Critical temperature**  
 $T_c = (156.5 \pm 1.5) \text{ MeV}$

*A. Bazavov et al. (hotQCD) PLB 795 (2019) 15*

*Similar results: S. Borsányi et al. (Budapest-Wuppertal group) PRL 125 (2020) 052001*

# Hypertriton/ $\Lambda$ ratio



- Hypertriton signal recently also extracted in isobar collisions
- Stronger separation between models as for other particle ratios, mainly due to the size of the hypertriton

