

# Improving nuclear structure calculations through new measurements of β-delayed neutron emitters

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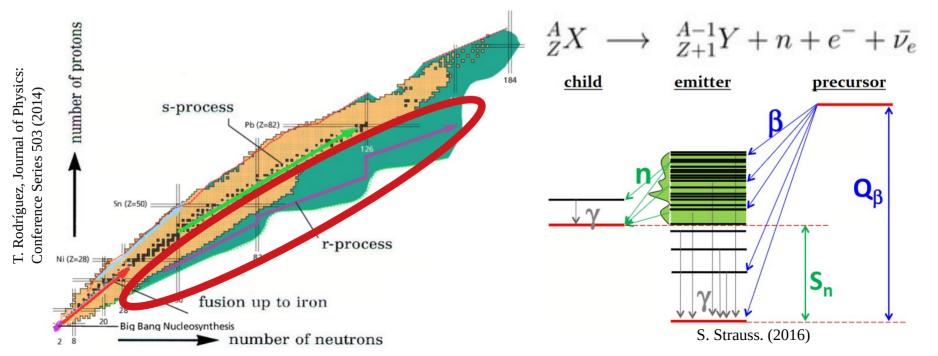




# β-delayed neutron emission



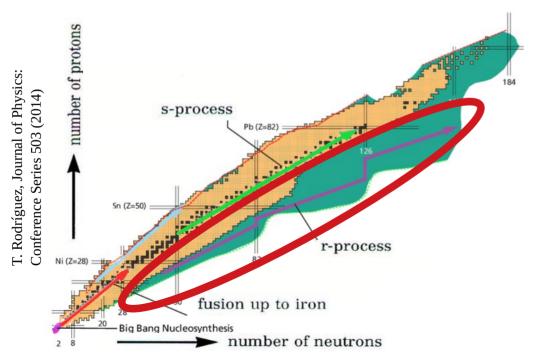
Neutron-rich nuclei typically decay via  $\beta^-$  decay. The energy available for this process is given by the  $\mathbf{Q}_{\beta}$  value. If  $\mathbf{Q}_{\beta}$  is greater than the neutron separation energy  $(S_{xn})$ , neutrons may be emitted following the  $\beta^-$  decay. This process is known as  $\beta$ -delayed neutron emission.



# β-delayed neutron emission



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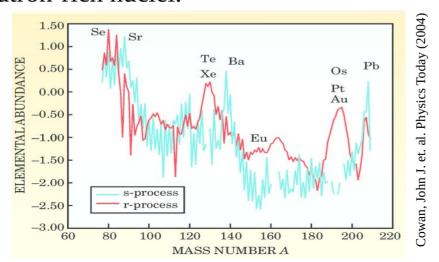
Physical magnitudes of interest for r-process simulations:

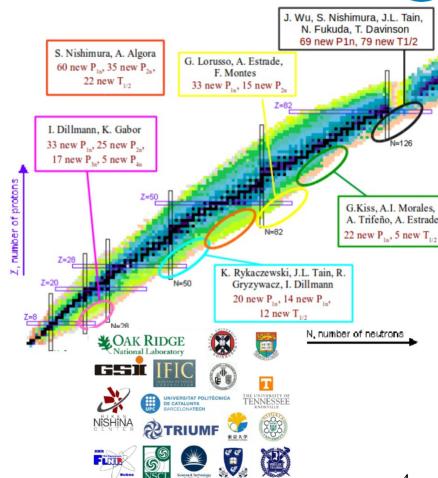
- Masses
- Neutron capture rates
- Half-life  $(T_{1/2})$
- $\beta$ -delayed neutron emission probabilities ( $P_{xn}$ )

#### BRIKEN project

UPC

- Built the largest β-delayed neutron detector in the world to conduct experiments at the RIKEN Nishina Center (Japan).
- Over 80 participants from 18 international institutions.
- Studied the most experimentally accessible neutron-rich nuclei.





#### BRIKEN project



J. Wu, S. Nishimura, J.L. Tain,

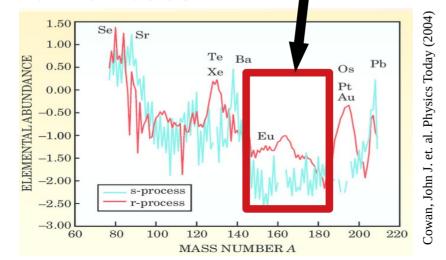
N. Fukuda, T. Davinson 69 new P1n. 79 new T1/2

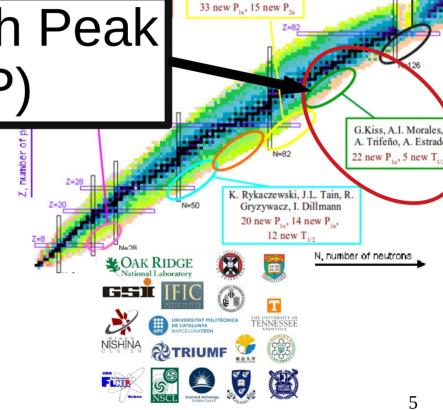
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Nishina Center (Japan). Rare-Earth Peak

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G. Lorusso, A. Estrade,

F. Montes

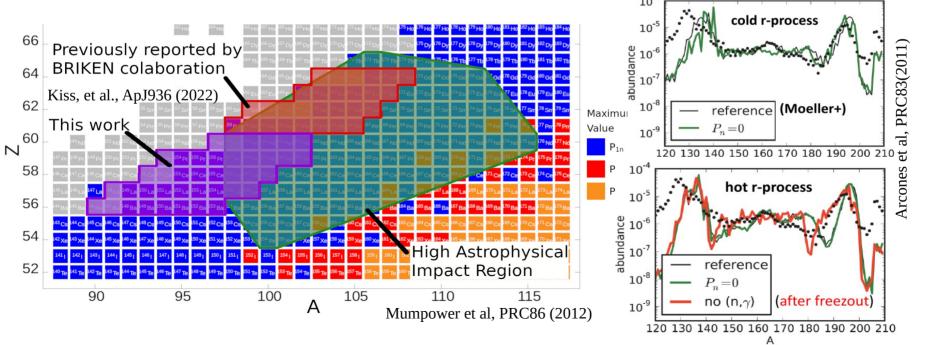
S. Nishimura, A. Algora

60 new P. . 35 new P.

# β-delayed neutron emission on REP



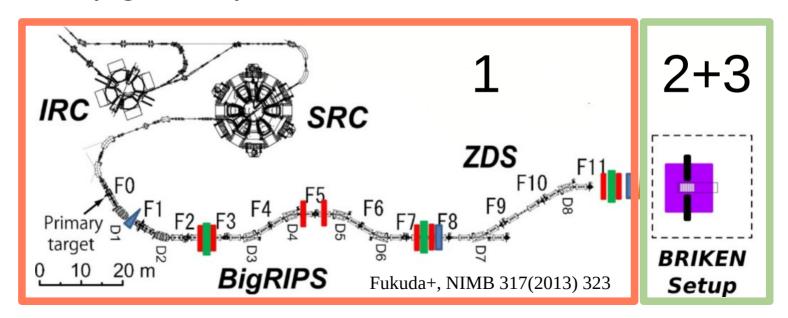
According to theoretical models and sensitivity studies,  $T_{1/2}$  and  $P_{xn}$  of very neutron-rich nuclei for  $55 \le Z \le 64$  are the most influential ones on the formation of the REP.



# Experimental setup



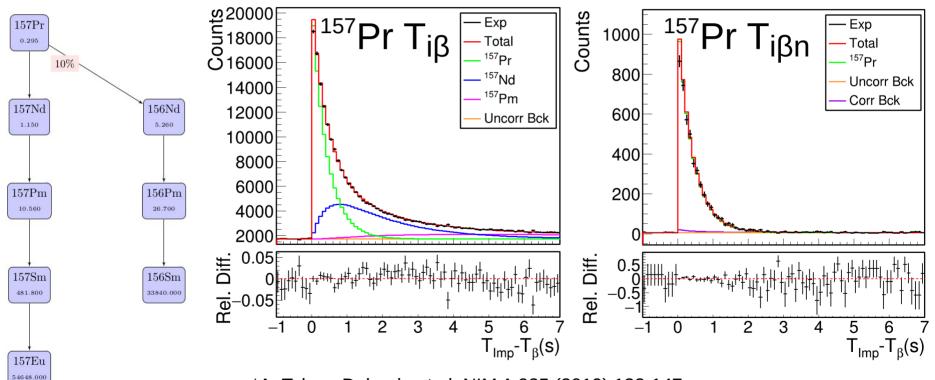
- 1. Production + Separation + Identification
- 2. Implantation +  $\beta$ -decay detection
- 3. Neutron (+gamma) detection



# Fitting procedure



For the analysis, we perform a self-consistent fitting procedure\* to obtain  $T_{1/2}$  and  $P_{1n}$  values using Bateman equations.

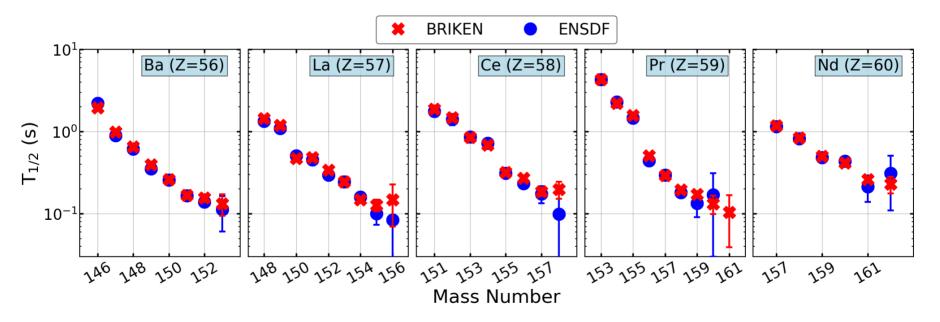


\*A. Tolosa-Delgado et al. NIM A 925 (2019) 133-147

# $T_{1/2}$ results



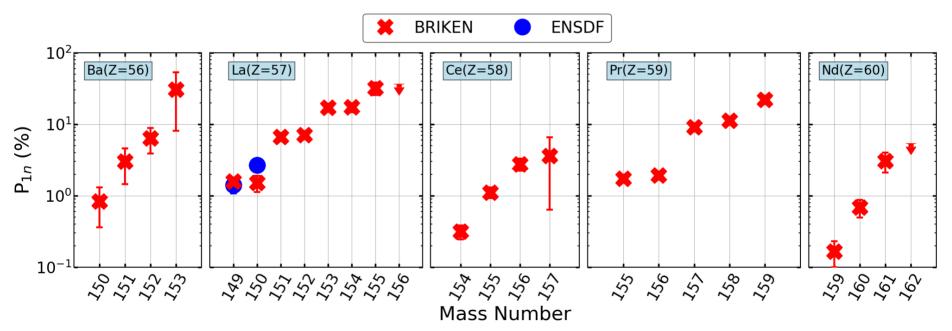
- 39 T<sub>1/2</sub> values have been remeasured.
- 1 new T<sub>1/2</sub>



# P<sub>1n</sub> results

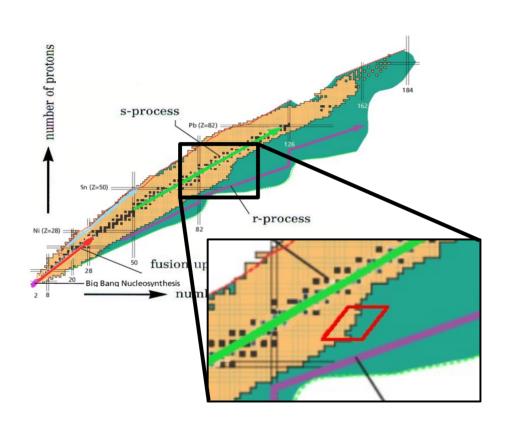


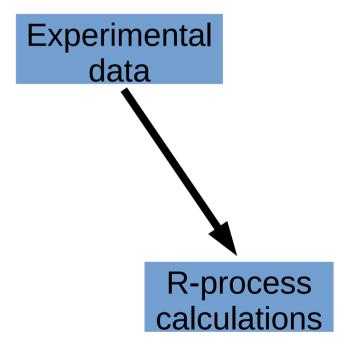
- 21 new P<sub>1n</sub> values.
- 2 P<sub>1n</sub> values have been remeasured.



# UPC

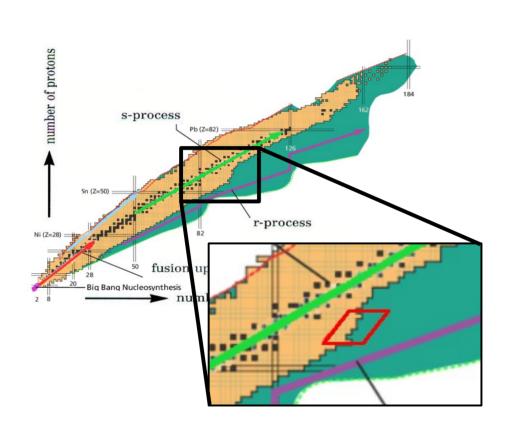
### What can we do with this data?

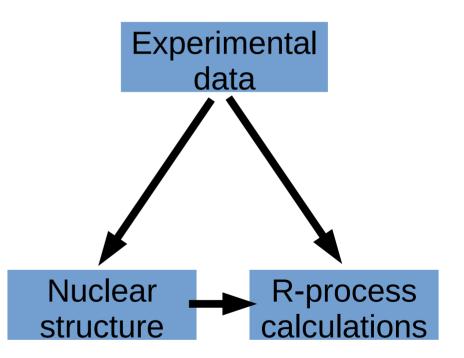






#### What can we do with this data?





#### Nuclear structure calculations

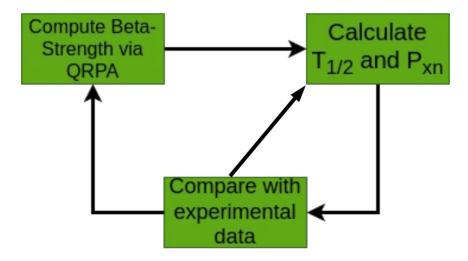


In collaboration with GSI Theoretical Nuclear Astrophysics Group.

- 1. Compute the  $\beta$ -Strength.
- 2. Calculate  $T_{1/2}$  and  $P_{xn}$ .
- 3. Compared with the experimental data from REP-BRIKEN.

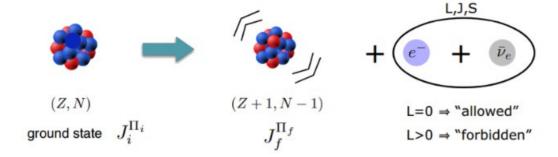


Diana Alvear Terrero



# β-decay rates of r-process nuclei

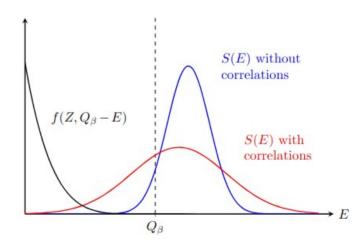




• Allowed decays (GT):

$$\lambda = \frac{\ln 2}{T_{1/2}} \propto \int^{Q_{\beta}} f(Z, Q_{\beta} - E) S(E) dE$$
$$S(E) = \sum_{f} |\langle f | \hat{F} | i \rangle|^{2} \delta(E - E_{f} + E_{i})$$

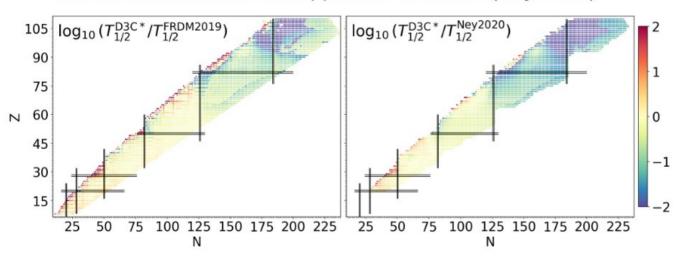
- Gamow-Teller (GT):  $\Delta S = 1$
- First Forbidden: L=1,  $\Pi_i \neq \Pi_f$



# Global β-decay calculations



- FRDM + gross theory for FF<sup>1</sup>
- relativistic spherical approach with D3C\*<sup>2</sup> and pairing correlations with D1S
- non-relativistic deformed approach with SKO' (Ney 2020)<sup>3</sup>



<sup>&</sup>lt;sup>1</sup>P. Möller et al., Phys. Rev. C **67**, 055802 (2003), P. Möller et al., Atomic Data and Nuclear Data Tables **125**, 1−192 (2019).

<sup>&</sup>lt;sup>2</sup>T. Marketin et al., Phys. Rev. C **93**, 025805 (2016), M. Eichler et al., The Astrophysical Journal **808**, 30 (2015).

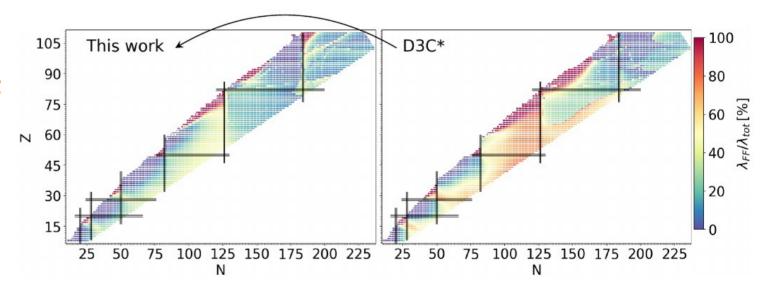
<sup>&</sup>lt;sup>3</sup>E. M. Ney et al., Phys. Rev. C **102**, 034326 (2020).

# Global β-decay calculations



This work model uses the D3C\* model but adds corrections¹ to the RHB+RQRPA code.

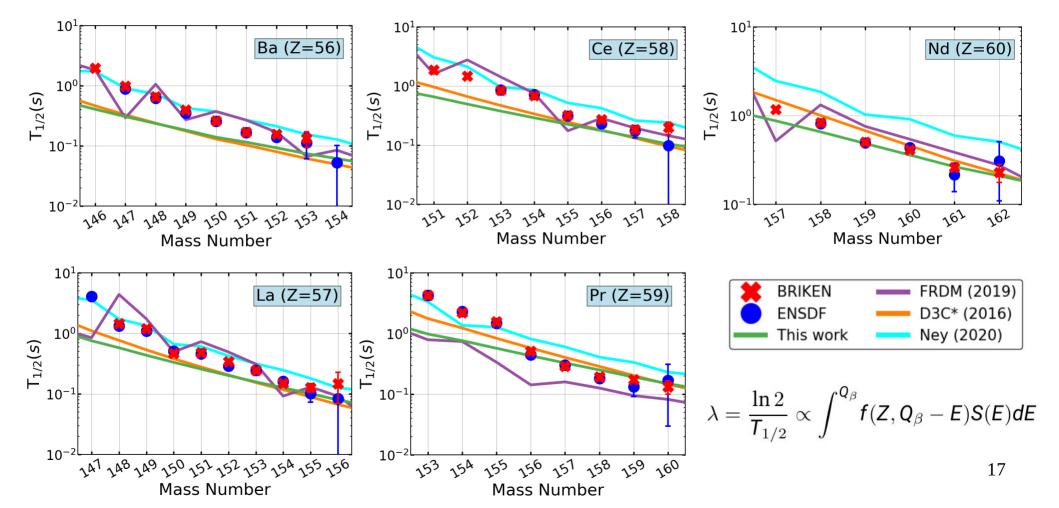
This varies FF contributions significantly



<sup>&</sup>lt;sup>1</sup> C.E.P Robin and G. Martinez-Pinedo, arXiv:20403.17115 (2024)

# $T_{1/2}$ for the REP





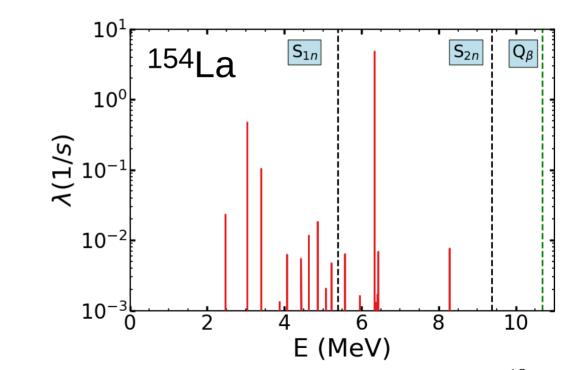


<u>Cut off method:</u> If the neutron emission is energetically possible, it will take place.

$$P_{xn} = rac{\displaystyle\sum_{i,E_i = S_{xn}}^{\min(Q_{eta},S_{(x+1)n})} \lambda_i}{\displaystyle\sum_{i}^{\sum} \lambda_i}.$$

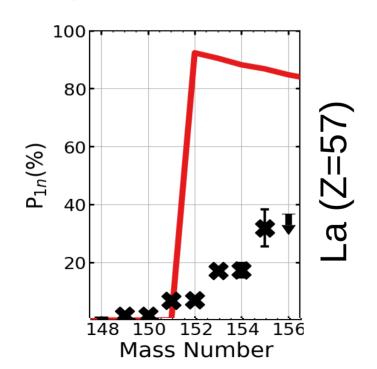
The  $P_{xn}$  calculations depends on:

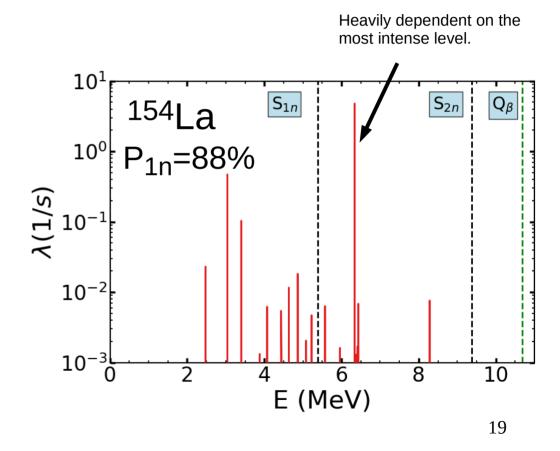
- 1. The Rate distribution.
- 2. Nuclear data ( $Q_{beta}$  and  $S_{xn}$ )





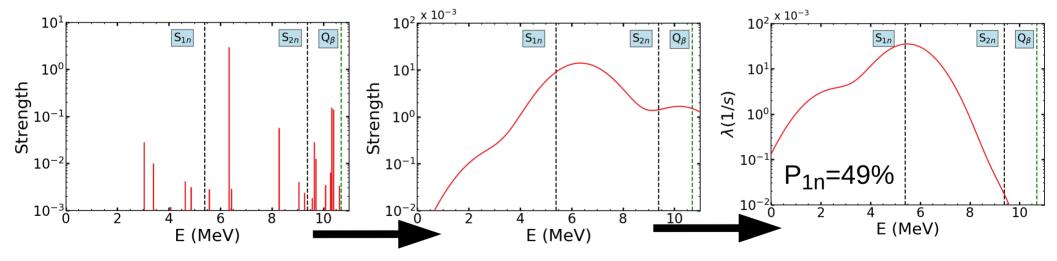
<u>Cut off method:</u> If the neutron emission is energetically possible, it will take place.







The QRPA calculations offers single discrete levels. As in reality the strength is a continous distribution, we must apply gausian distribution of certain width ( $\sigma$ ) on each energy level of the Strength.

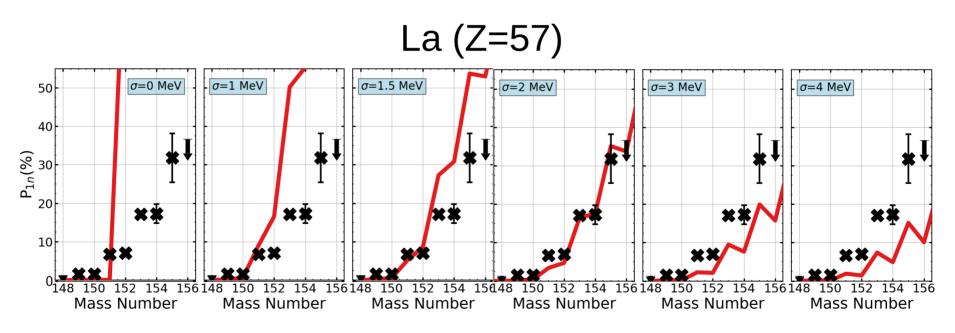


Gaussian distribution  $C(W) = k + kaW + kb/W + kcW^2$ .

$$\lambda = \frac{\ln 2}{K} \int_0^{p_0} p_e^2 (W_0 - W)^2 F(Z, W) C(W) dp_e,$$



The QRPA calculations offers single discrete levels. As in reality the strength is a continous distribution, we must apply gausian distribution of certain width ( $\sigma$ ) on each energy level of the Strength.



# Other methods to compute P<sub>xn</sub>



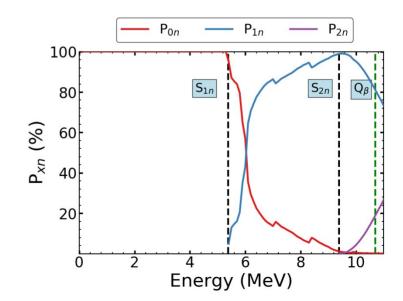
#### **Cut off method:**

- 1. If the neutron emission is energetically possible, it will take place.
- 2. By definition offers an upper limit.

$$P_{xn} = rac{\displaystyle\sum_{i,E_i=S_{xn}}^{\min(Q_{eta},S_{(x+1)n})}\lambda_i}{\displaystyle\sum_{i}^{\sum_{j}\lambda_i}\lambda_i}.$$

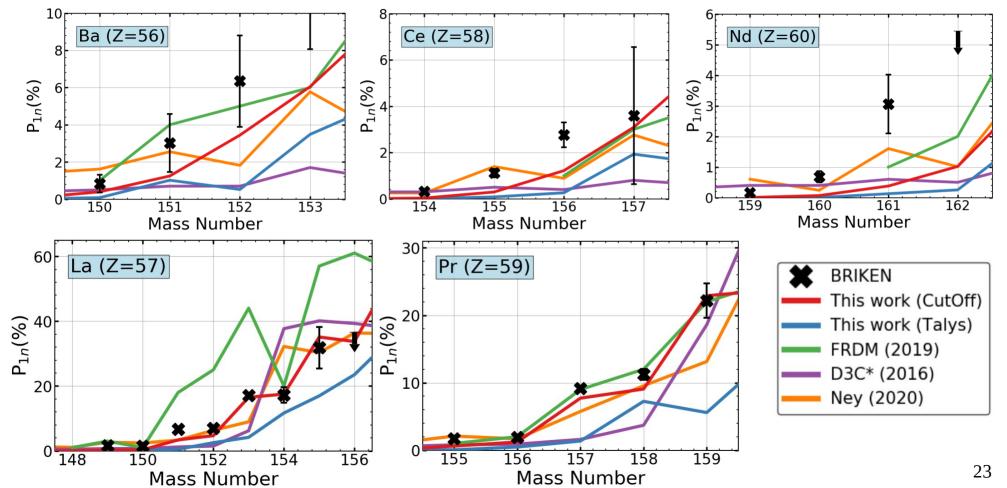
#### **Statistical HF method:**

- 1. If neutron emission is energetically possible, there will be a probability of it occurring.
- 2. We employ TALYS.



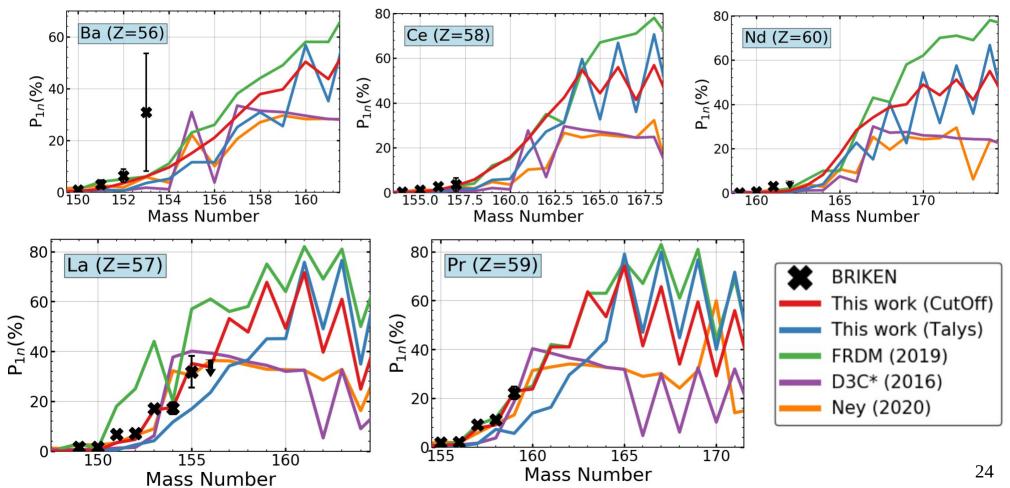
# P<sub>1n</sub> for REP-BRIKEN





# P<sub>1n</sub> for REP





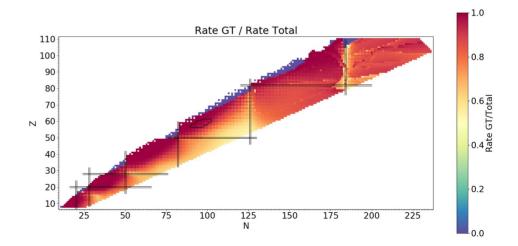
# Talys Spin-Parity



We do not have a detailed description of these excited states but we can assume that all are populated by GT.

The other required information is the initial state from which the parent nucleus decayed. We can assume it was the ground state.

Most of our nuclei present a λ distribution with 90% GT

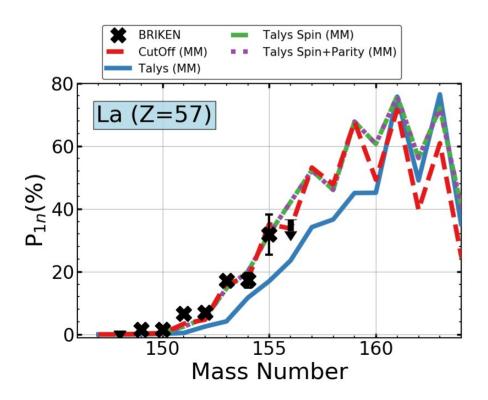


# **Example: Talys Spin-Parity**



Ground state spin taken from JINA.

Assuming all decays are GT.



Слреппена	140La 141La 142La 143La 144La 145La 146La 147La 148La	3- (7/2+) 2- (7/2)+ 3- (5/2+) (2-) (5/2+)
	148La 149La	(2-) (3/2-)
	150La	(3+)

Explore: Level Densities y-Strength function Optical model



#### Conclusions and outlook

- We performed high-precision measurements of beta-decay properties for dozens of very neutron-rich REP isotopes.
- Nuclear structure calculations using a global QRPA model offered consistent results.
- Future work:
  - Explore the impact of different parameters on the TALYS calculations for the P<sub>xn</sub> values.
  - Combine the new experimental data from BRIKEN with updated nuclear structure calculations to perform r-process nuclear reaction network calculations.



# Acknowledgments

This work has been supported by the Spanish Ministerio de Economía y Competitividad under Grants nos. FPA2014-52823-C2-1-P, FPA2014-52823-C2-2-P, FPA2017-83946-C2-1-P, FPA2017-83946-C2-2-P and grants from Ministerio de Ciencia e Innovacion nos PID2019-104714GB-C21, PID2019-104714GB-C22 and PID2022-138297NB-C22. Also supported by SANDA project funded under H2020-EURATOM-1.1 Grant No. 847552.