

Following ALICE's Footsteps: A Modern Tale of Neutrino Interactions in the Few-GeV Realm

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University of Warwick

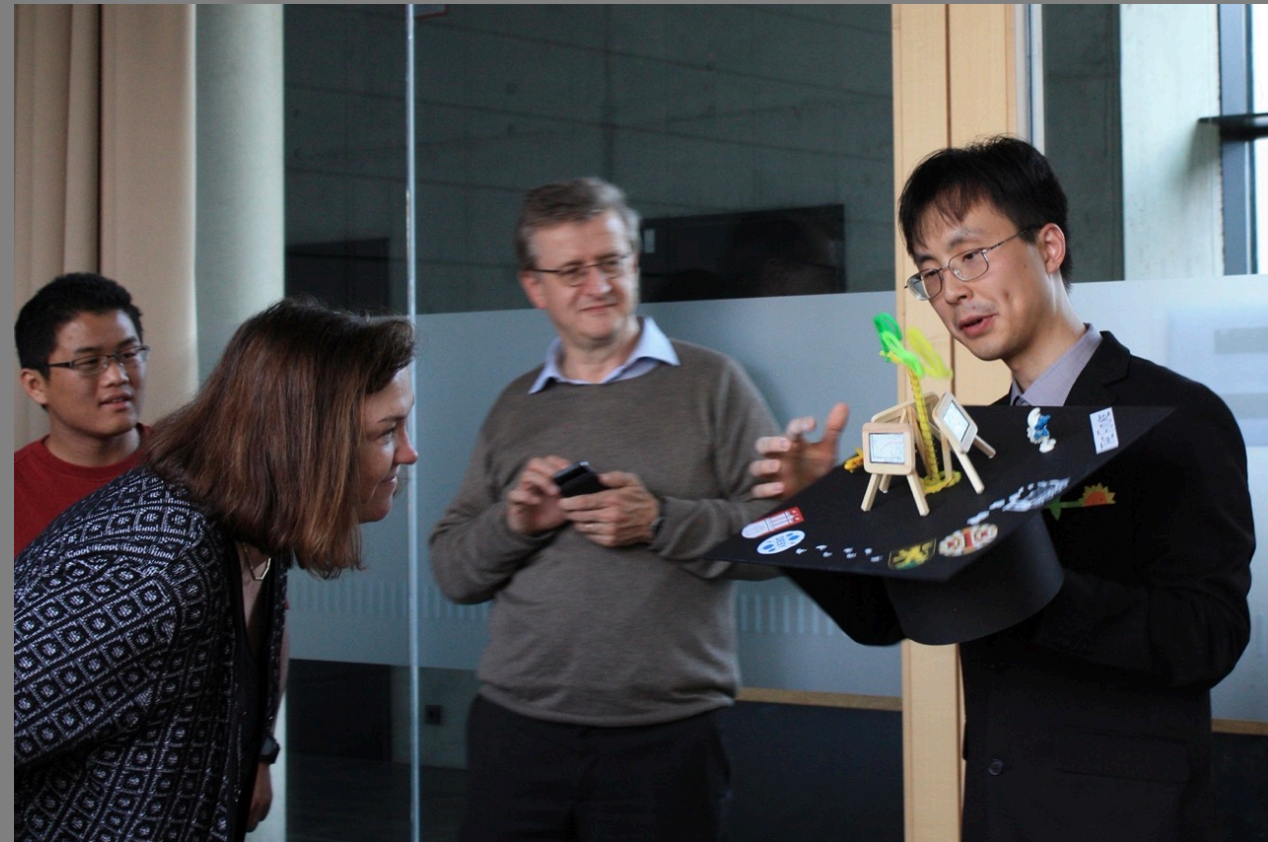
Never at Rest: A Lifetime Inquiry of QGP
Celebrating the 70th birthday of Prof. Johanna Stachel

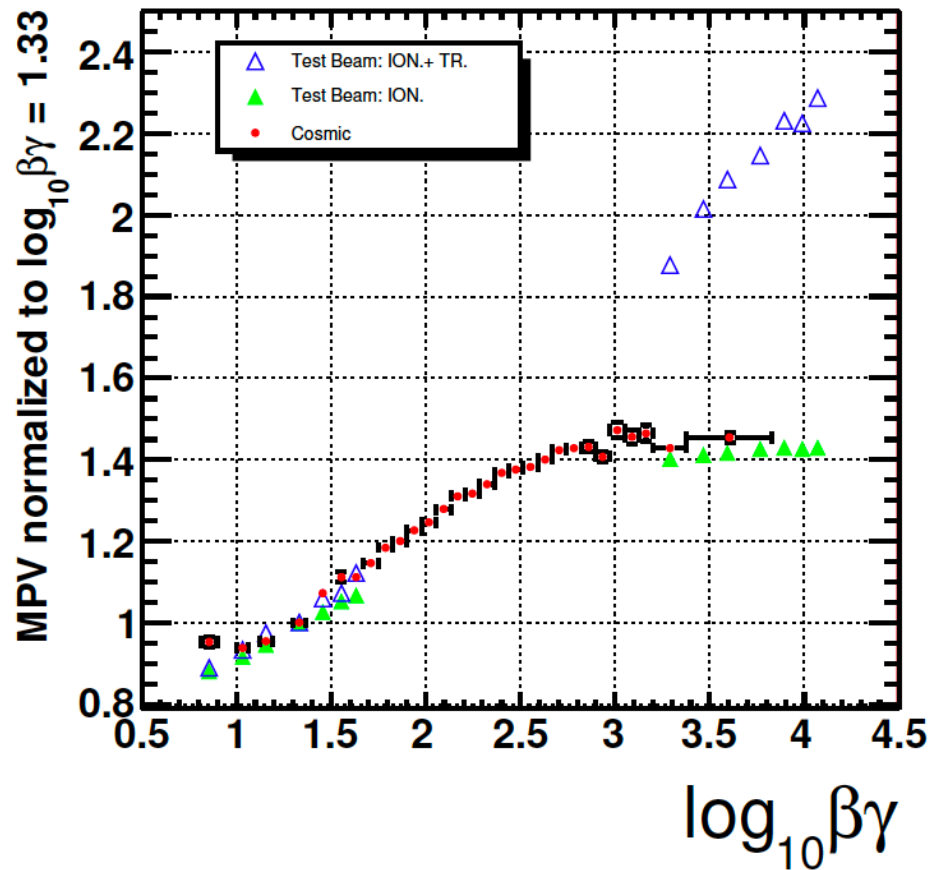
2025 February 10, Bad Honnef

- ❑ Started PhD under Johanna in September 2009
- ❑ Thesis defence in October 2013
- ❑ First PostDoc until August 2014, then moved to neutrino

Member of neutrino experiments w/ GeV neutrinos, in particular:

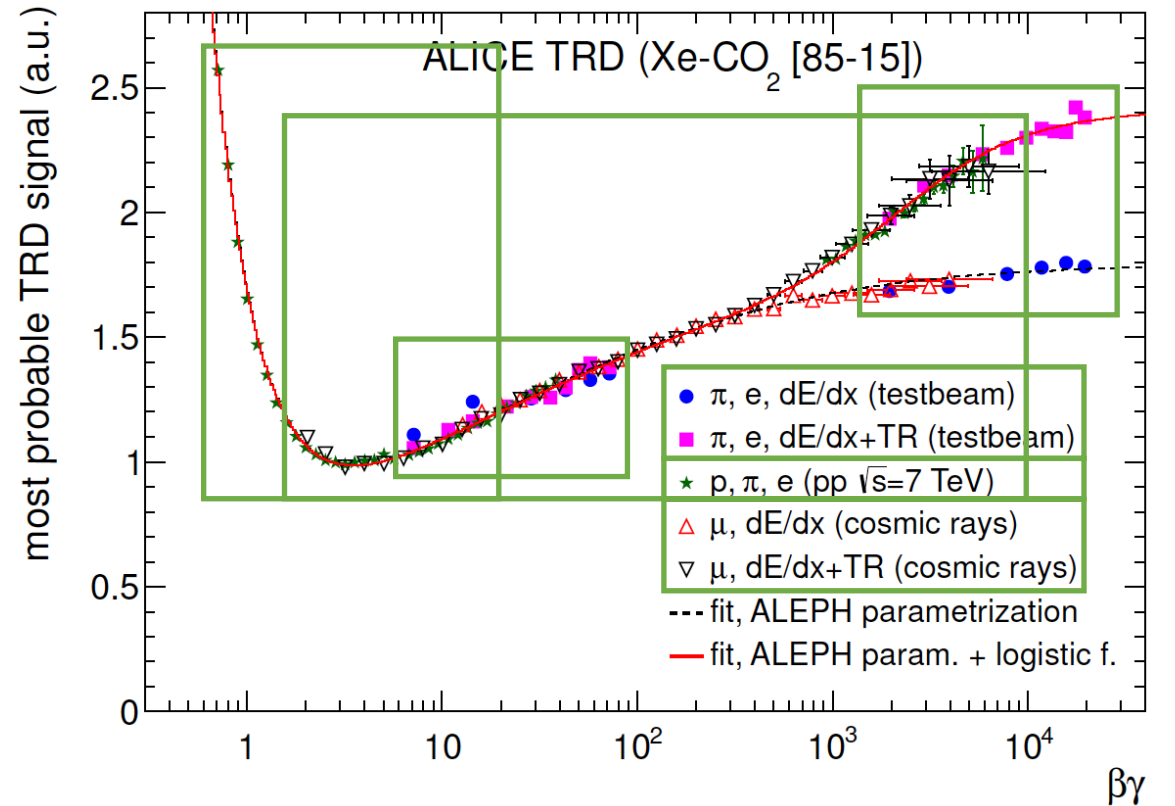
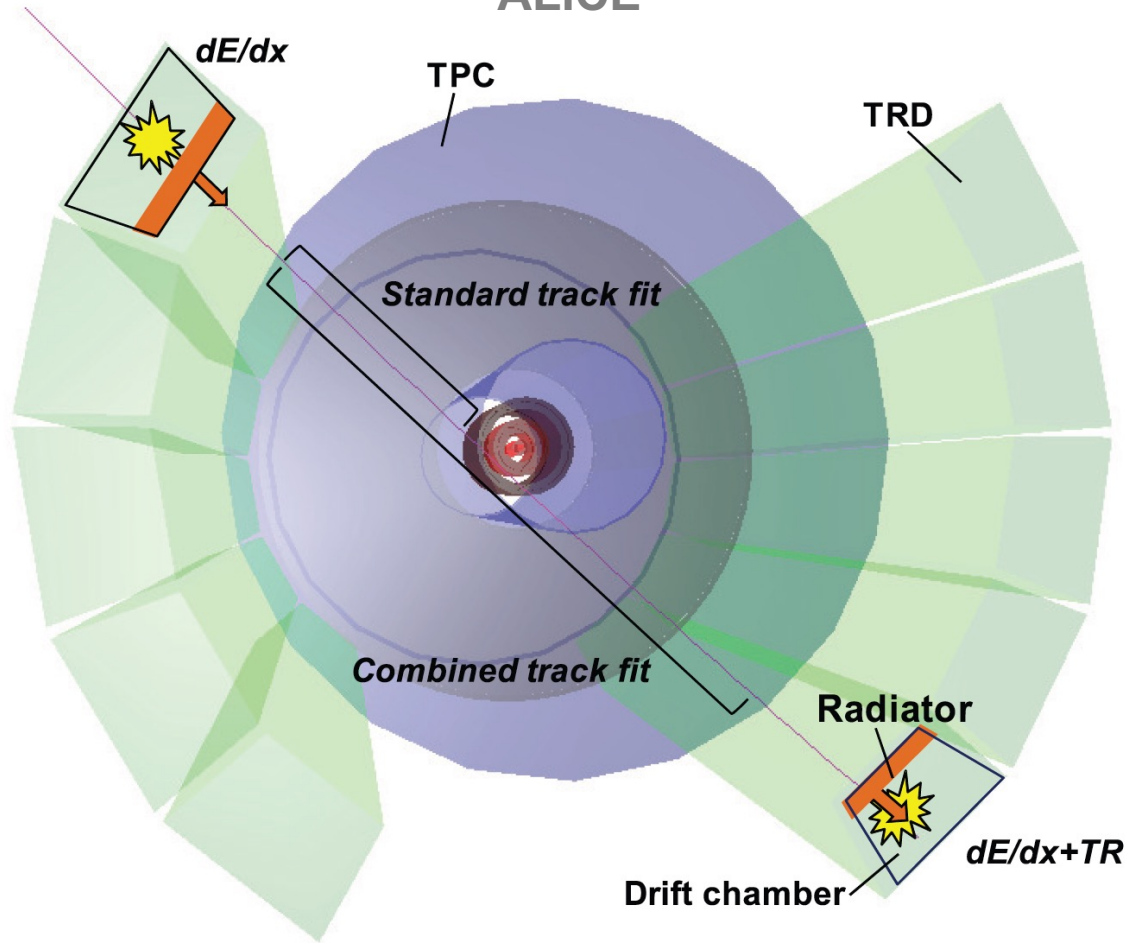
- ❖ nuSTORM (neutrinos from STORed Muons)
- ❖ JUNO (Jiangmen Underground Neutrino Observatory)





- A plot from ~ 1 June 2010
- MinJung and I were trying to understand transition radiation (TR) in full ALICE set up with cosmic muons
- We were surprised that the TRD signal was consistent with *pure* dE/dx from test beam...
- We tried very hard to understand it and in the end, I sent an email
 “Anton, do you think the *theory* of TR could be wrong?”

ALICE



One of my thesis topics: ALICE TRD performance in full ALICE set up

- ❑ Turned out could measure onset of TR → in vs. out cosmics
- ❑ Need TPC tracking up to TeV → combined track fit
- ❑ Need MIP for energy scale → dedicated cosmic data with $B = 0.1$ T

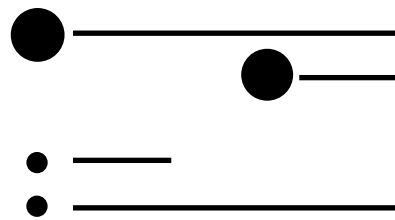
Outline

1. Problems and opportunities with neutrino masses
 - Call for a GeV ν_e and $\bar{\nu}_e$ machine: nuSTORM
2. Using lepton-hadron correlations to study GeV neutrino interactions
 - Transverse Kinematic Imbalance (TKI)
3. JUNO news!

Neutrino Mass and Mixing



Mass Ordering



Normal Inverted

Δm^2 leads to neutrino oscillations

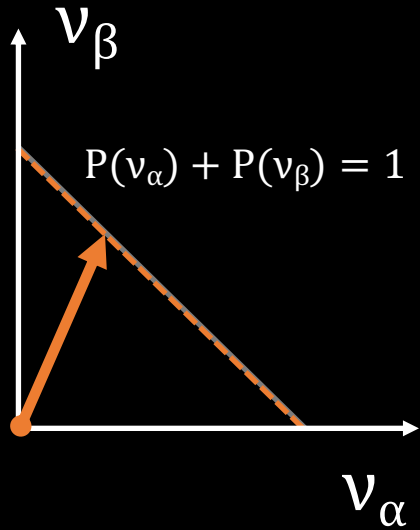
Problems

- What are the neutrino masses?
 - ❖ Mass gaps (Δm_{21}^2 , $|\Delta m_{32}^2|$) and ordering ($\text{sgn } \Delta m_{32}^2$)?
- What are the mixing parameters?
 - ❖ Mixing angles (θ_{12} , θ_{23} , θ_{13}) and CP-phase (δ_{CP})?

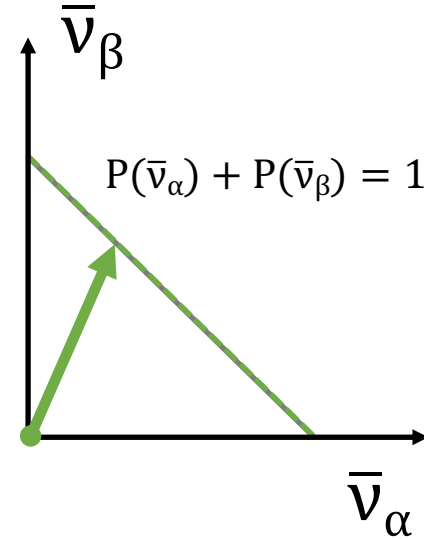
$$\theta_{13} = 0$$

Neutrinos

2-flavor oscillation



Antineutrinos

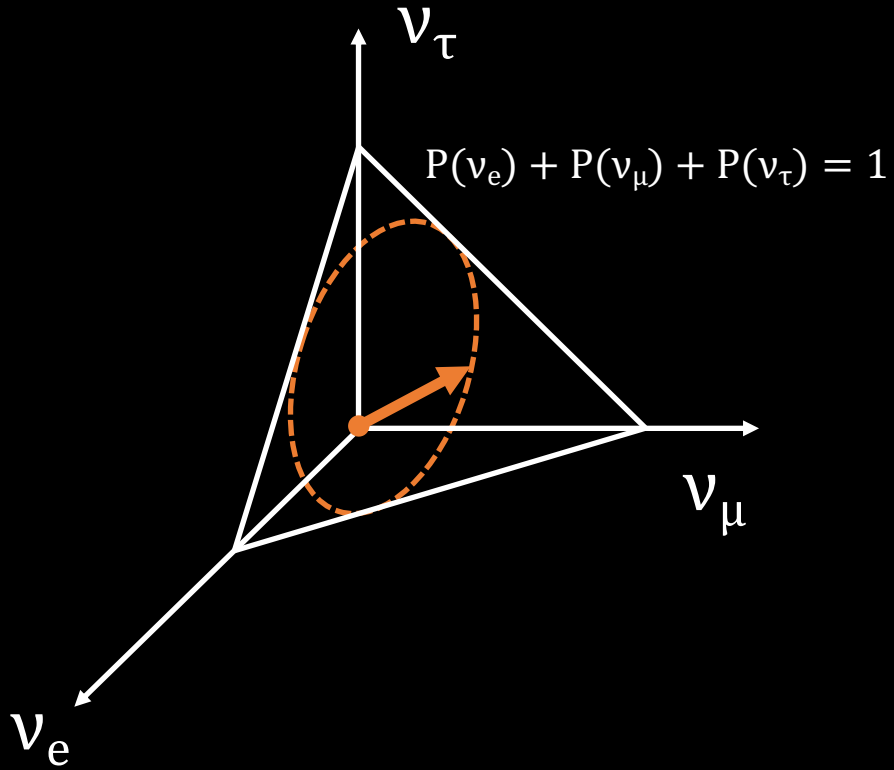


□ 2-flavor oscillation: CP not observable

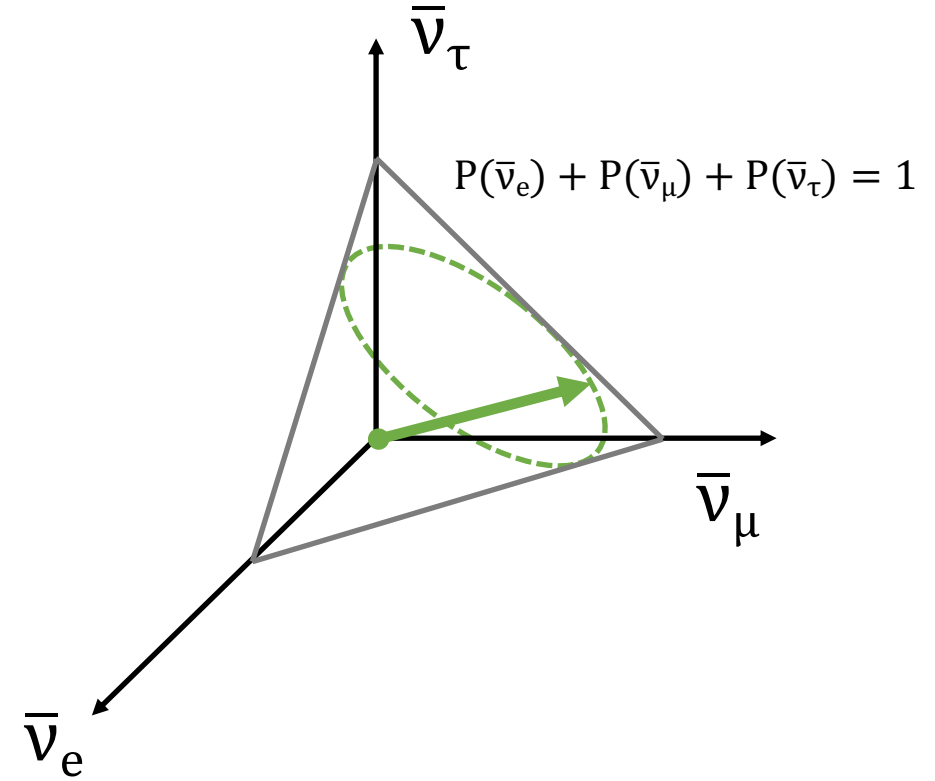
$$\theta_{13} \neq 0$$

3-flavor oscillation

Neutrinos



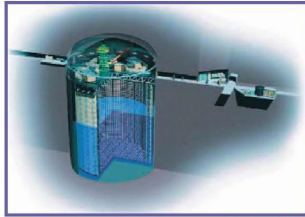
Antineutrinos



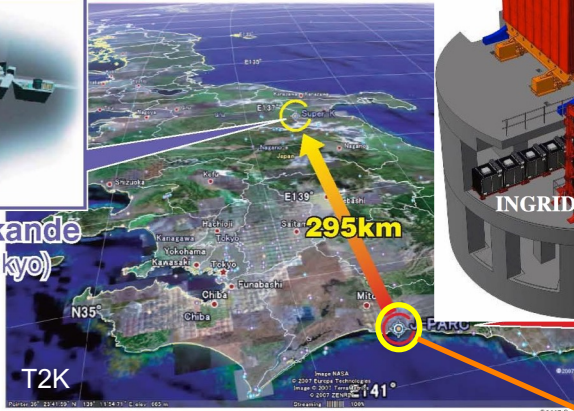
Opportunities

- 3-flavor oscillation: CP-violation possible

Accelerator Neutrino Experiments

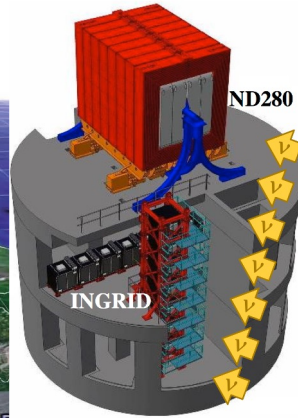


Super-Kamiokande
(ICRR, Univ. Tokyo)



295km

T2K



ND280

INGRID

T2K / Hyper-K



Far Det.

Minnesota

Wisconsin

Michigan

Iowa

Illinois

Indiana

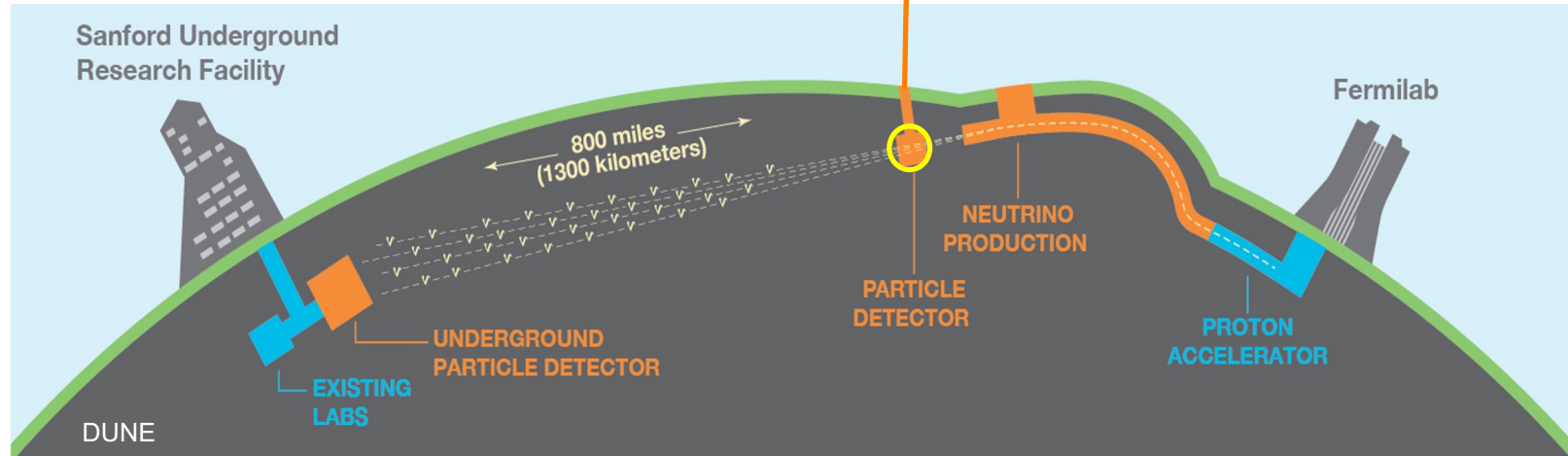
Missouri

NOvA

NOvA

ν beams via hadron decay

DUNE



Sanford Underground
Research Facility

800 miles
(1300 kilometers)

Fermilab

NEUTRINO
PRODUCTION

PARTICLE
DETECTOR

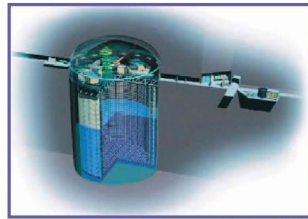
PROTON
ACCELERATOR

UNDERGROUND
PARTICLE DETECTOR

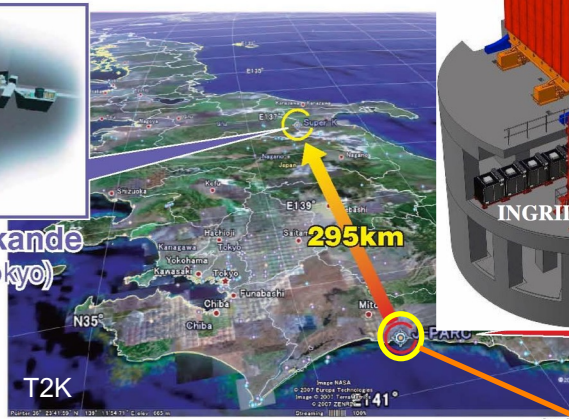
EXISTING
LABS

DUNE

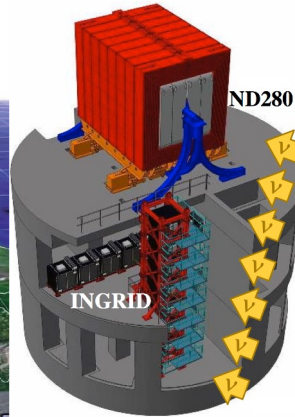
Accelerator Neutrino Experiments



Super-Kamiokande
(ICRR, Univ. Tokyo)



T2K / Hyper-K



ND280

INGRID

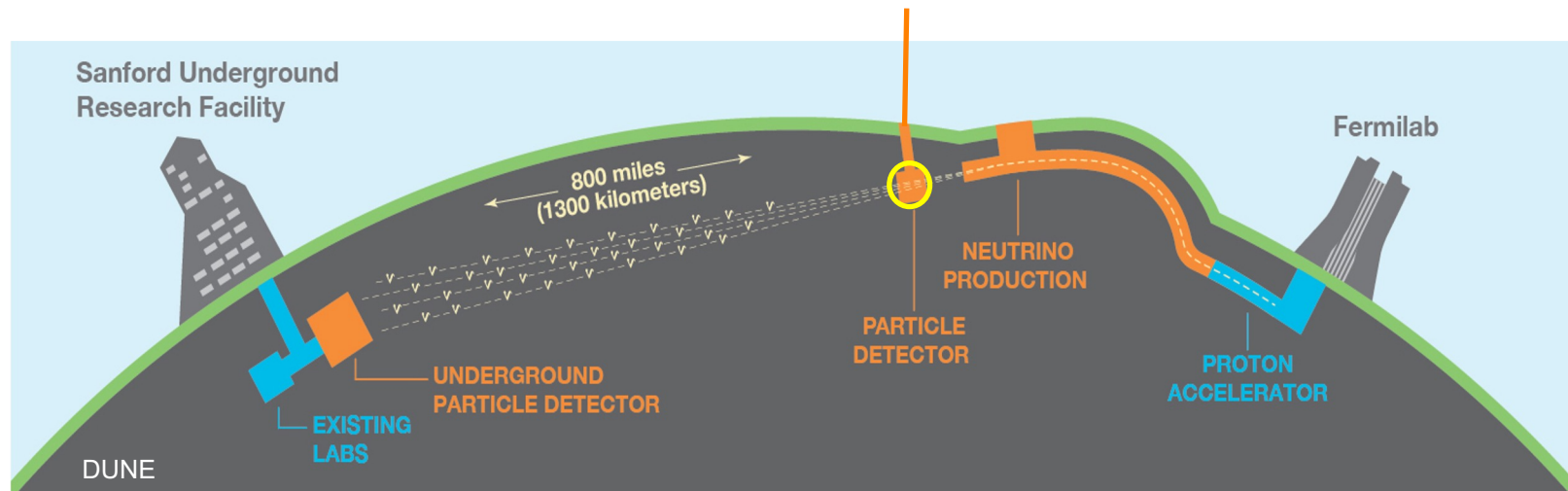


NOvA

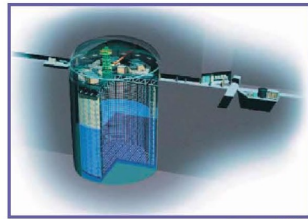
NOvA

Near Detectors to constrain ν flux *and* cross section *together* (no oscillation)

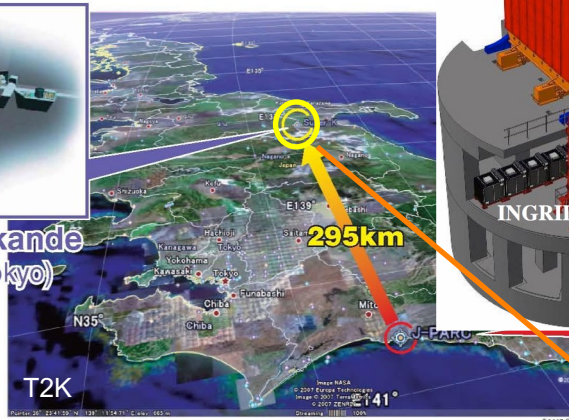
DUNE



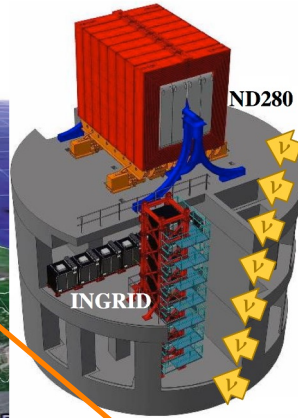
Accelerator Neutrino Experiments



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T2K / Hyper-K



ND280

INGRID

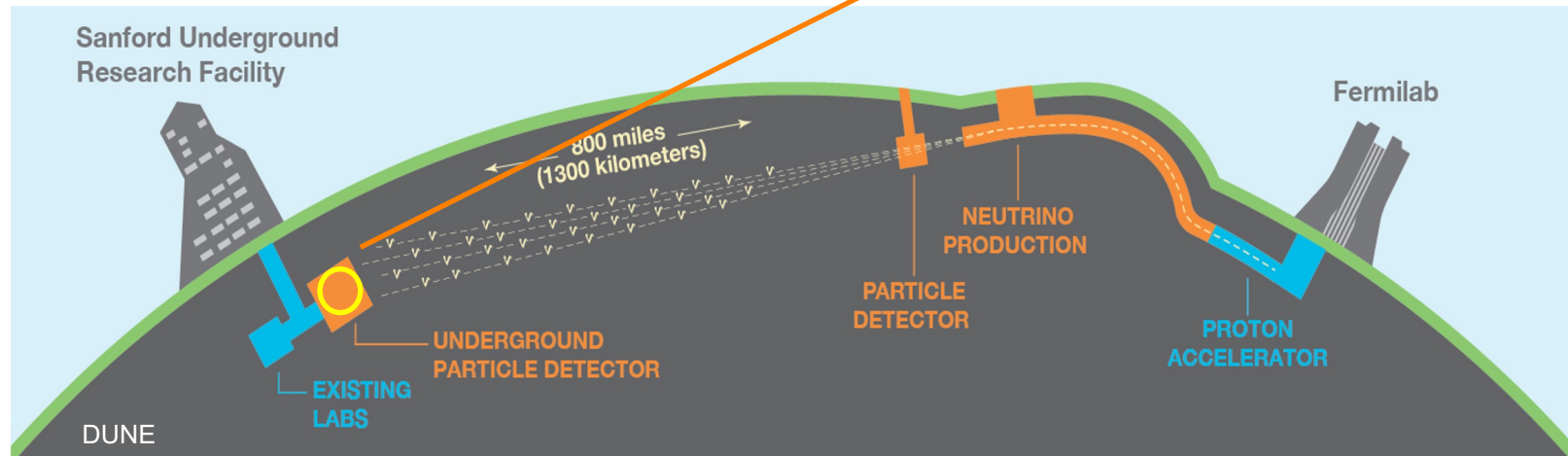
Far Detectors to measure ν oscillation



NOvA

NOvA

DUNE



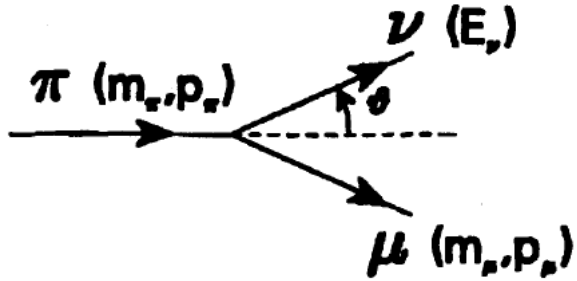
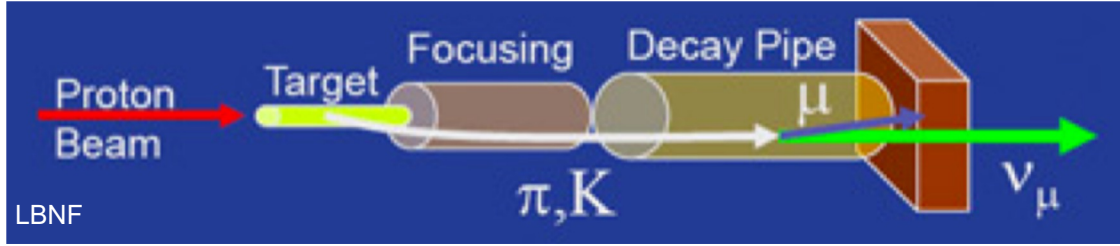
Accelerator Neutrino Experiments

Accelerator ν Experiment	E_ν/GeV @ Flux Peak	Far Detector Technology	Target Nuclei
T2K / Hyper-K	0.6	Water Cherenkov	H ₂ O
NOvA	2	Liquid Scintillator	CH
DUNE	2.4	LAr TPC	Ar

Signal = (**Beam flux** · **Oscillation probability** · Cross section) \oplus Detector effects

- Beam: ν_μ and $\bar{\nu}_\mu$
- Oscillation
 - ❖ ν_μ and $\bar{\nu}_\mu$ disappearance (most oscillated to ν_τ and $\bar{\nu}_\tau$)
 - ❖ **ν_e and $\bar{\nu}_e$ appearance, then CP violation**

Accelerator Neutrino Experiments



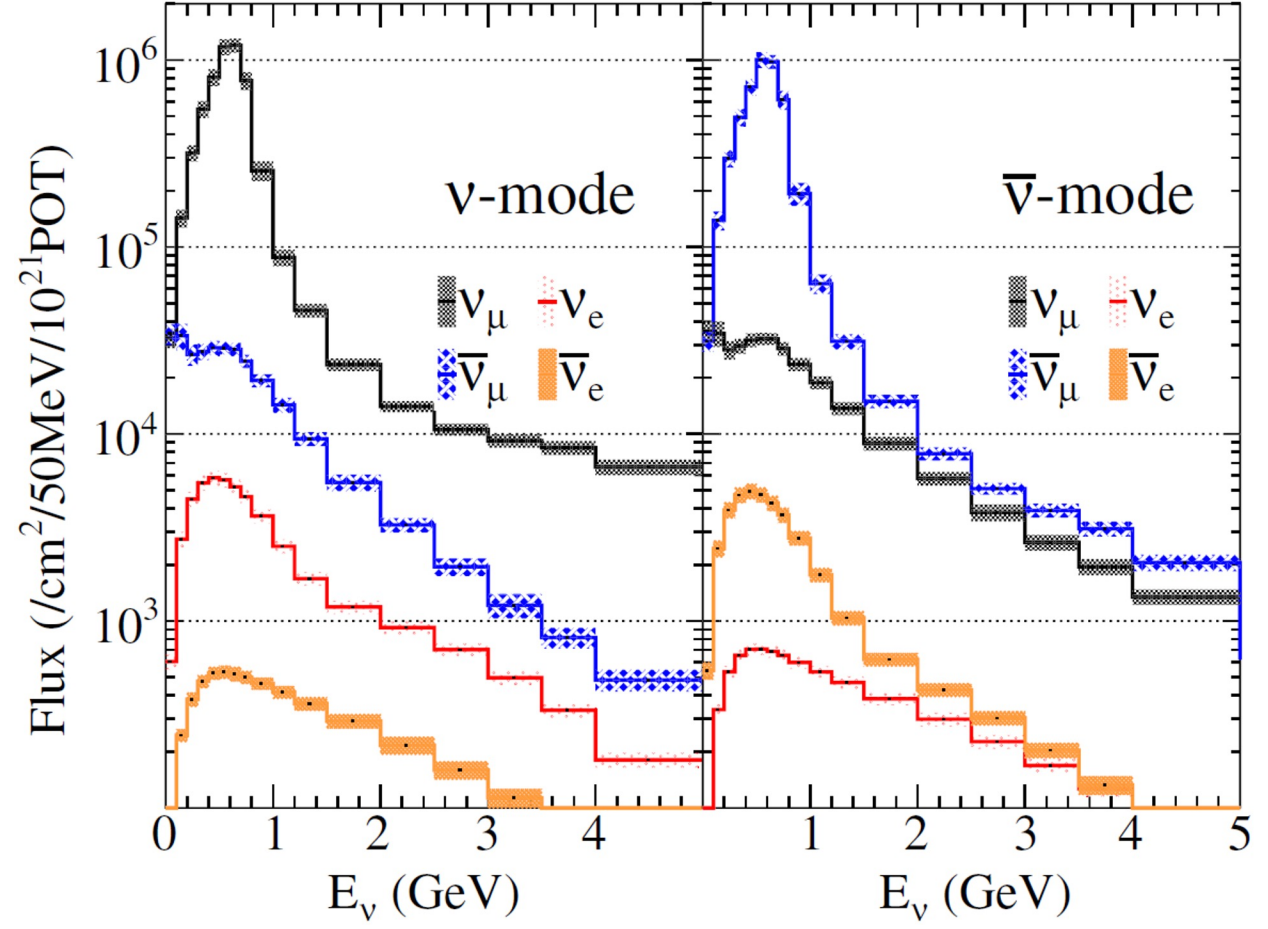
“β decay” of collision products
(ν_μ from π , ν_e from K)

Neutrino beams from accelerators
→ Directional

Charge selection on π
→ High purity ν or $\bar{\nu}$ beams

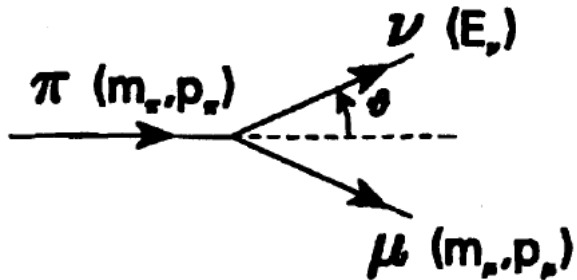
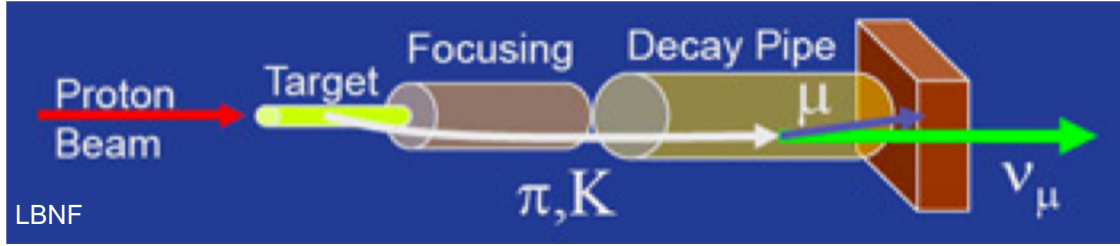
T2K flux

T2K, Phys.Rev.Lett. 116, 181801 (2016)



□ Maximise ν_μ and $\bar{\nu}_\mu$ flux

Accelerator Neutrino Experiments



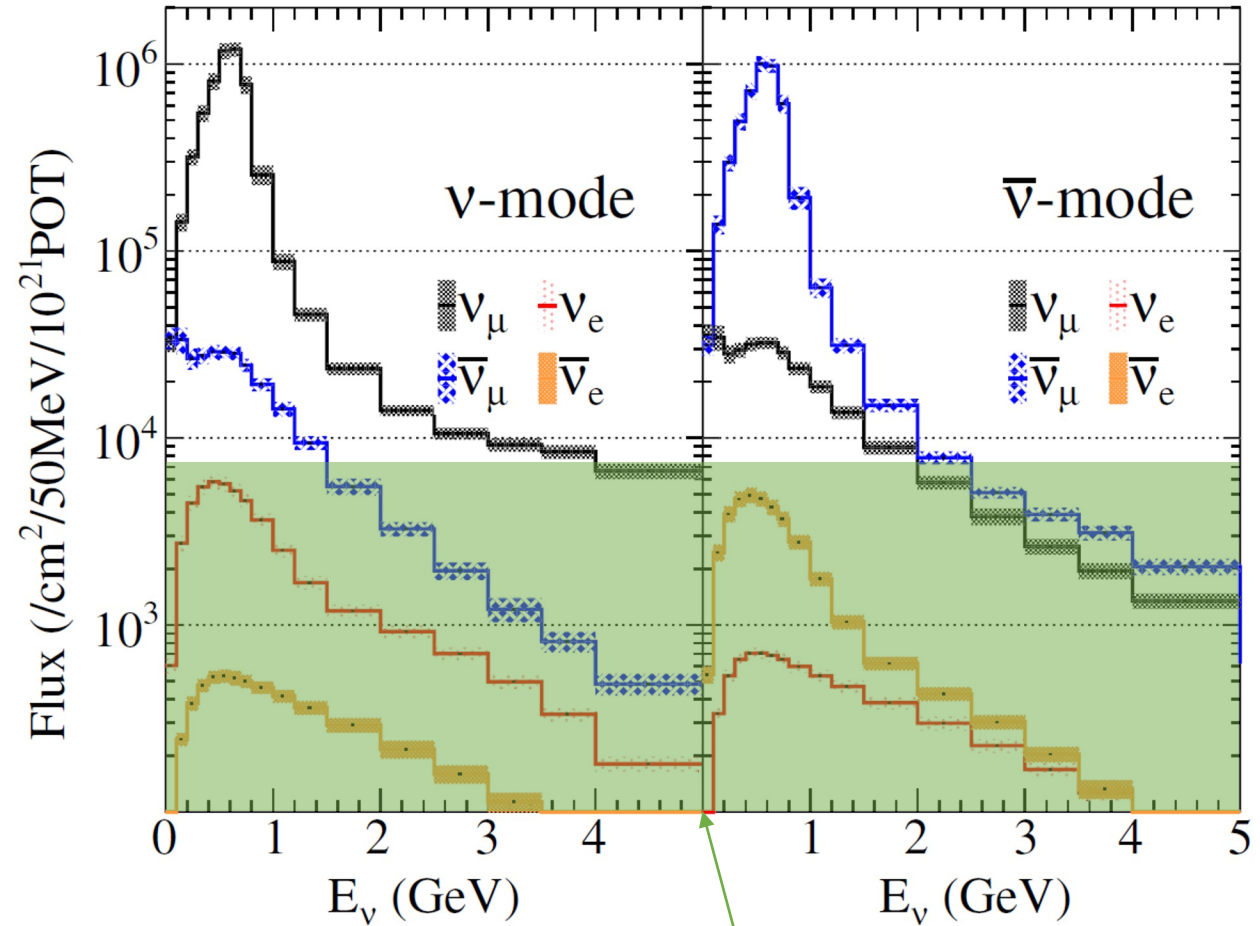
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T2K flux

T2K, Phys.Rev.Lett. 116, 181801 (2016)



- Maximise ν_μ and $\bar{\nu}_\mu$ flux
- Minimise ν_e and $\bar{\nu}_e$ to observe appearance**
- ν_e and $\bar{\nu}_e$ flux are background**

Call for a GeV ν_e and $\bar{\nu}_e$ Machine

Oscillation Signal = (Beam flux · Oscillation probability · **Cross section**) ⊕ **Detector effects**

ν_e ($\bar{\nu}_e$) cross sections: major δ_{CP} systematics

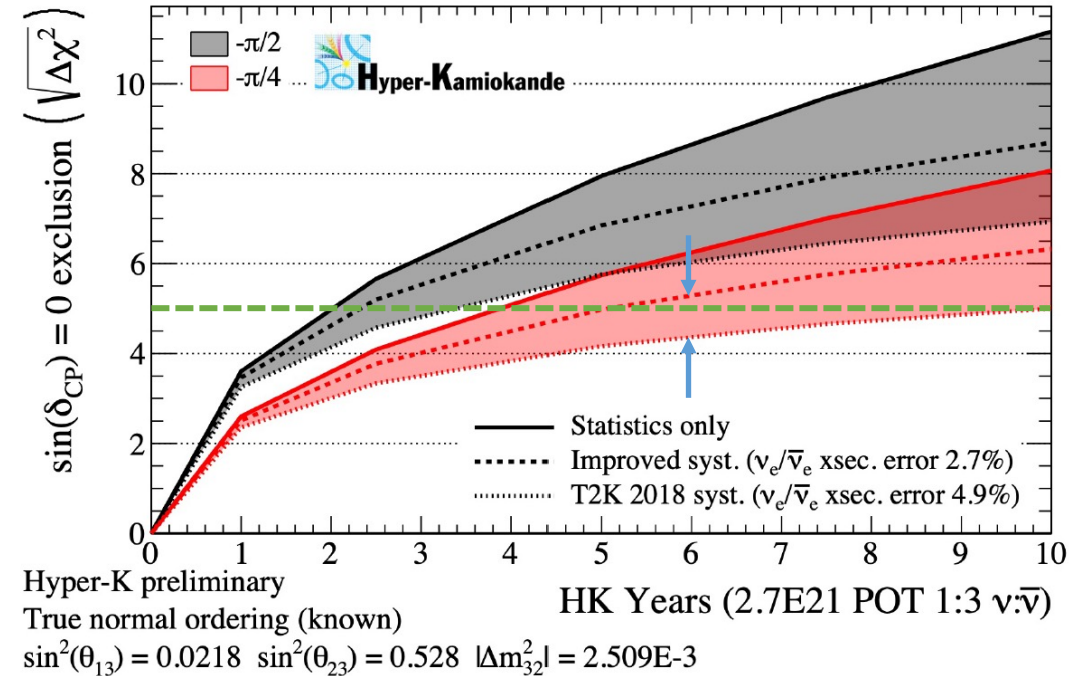
- Very few ν_e scattering data
- $\delta_{CP} \sim \nu_e$ appearance \sim no ν_e in beams
 - ❖ No *in situ* ν_e measurements
- ν_μ for ν_e via lepton universality, but higher precision needed for δ_{CP}

Hyper-K example

Improving error of $\nu_e/\bar{\nu}_e$ xsec ratio 4.9% \rightarrow 2.7%

- Improve δ_{CP} sensitivity by $\sim 1 \sigma$ for 6 year
- Significantly shorten running time to reach 5σ**

Jeanne Wilson, Neutrino2022



Host

Cost

Cost

Time

Call for a GeV ν_e and $\bar{\nu}_e$ Machine

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ν_e ($\bar{\nu}_e$) cross sections: major δ_{CP} systematics

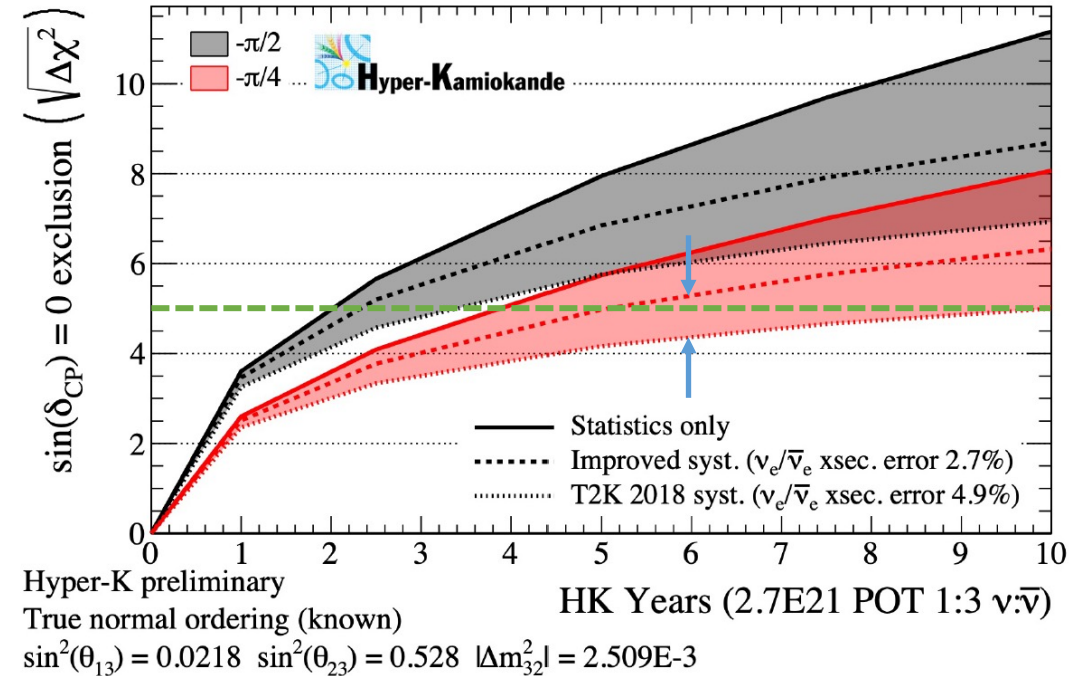
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- ❑ **Significantly shorten running time to reach 5σ**
- ❑ **Shorten running time = ~~reduce share cost~~**

Jeanne Wilson, Neutrino2022



Host

Cost

Co-Host

Cost

Time

Call for a GeV ν_e and $\bar{\nu}_e$ Machine

Oscillation Signal = (Beam flux · Oscillation probability · **Cross section**) \oplus **Detector effects**

ν_e ($\bar{\nu}_e$) cross sections: major δ_{CP} systematics

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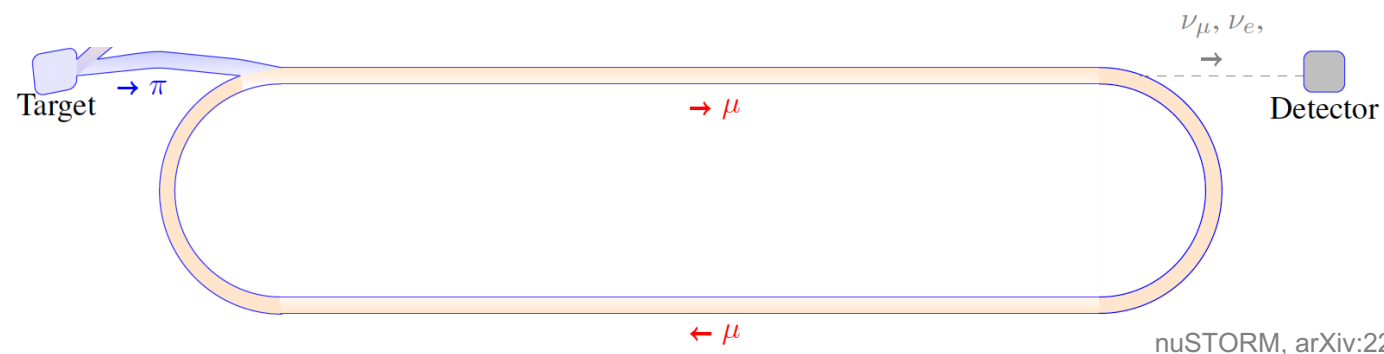
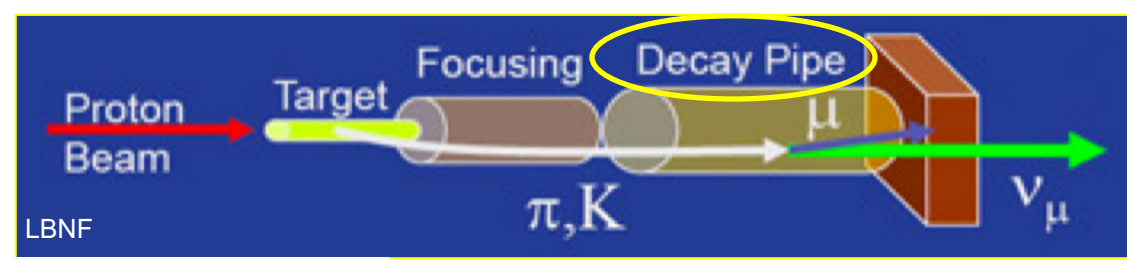
✓ Wishlist

1. ν_e and $\bar{\nu}_e$ beams with the relevant energy for appearance
2. Well-understood fluxes
3. High statistics
4. Low ν_μ background

- Shorten running time = ~~reduce share cost~~

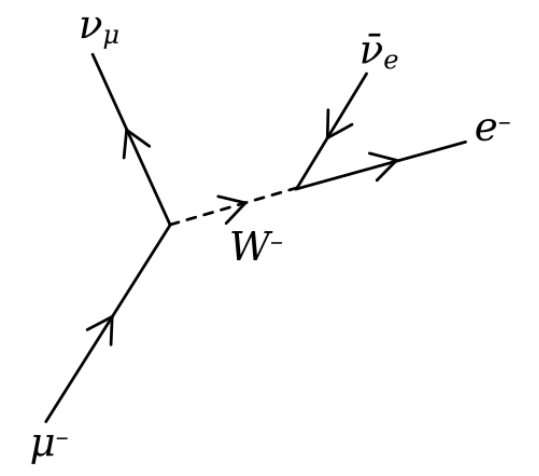


ν from *STORed* Muons (*nuSTORM*)

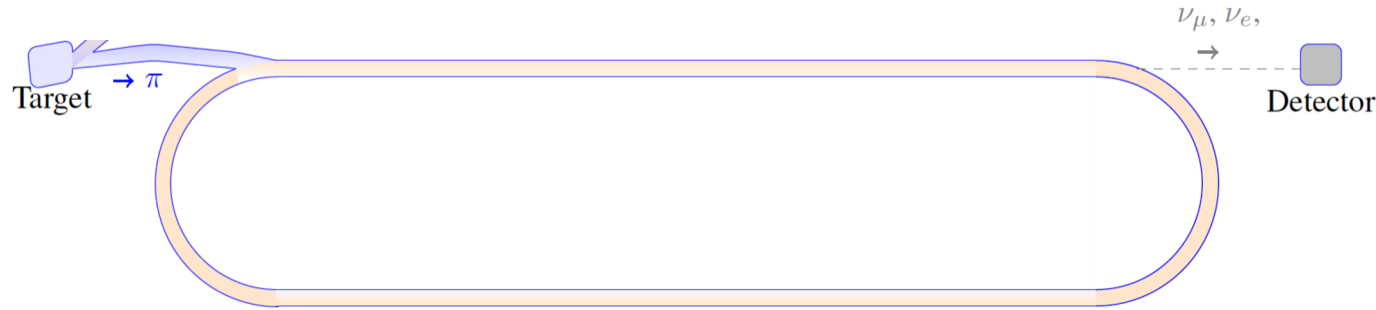


nuSTORM, arXiv:2203.07545

- ❑ π captured by storage ring $\rightarrow \mu$, instead of decay pipe $\rightarrow \nu$
- ❑ $\bar{\nu}_\mu + \nu_e$ and $\nu_\mu + \bar{\nu}_e$ fluxes from μ^\pm decays
 - ❖ **Storage ring: tunable fluxes**
 - ❖ **μ decay: perfect understanding of flux shape and normalisation**
- ❑ Scientific objectives
 - ❖ %-level ν cross sections
 - ❖ BSM searches, e.g. steriles beyond Short Baseline Neutrino program at FNAL
 - ❖ Test bed for muon collider technology

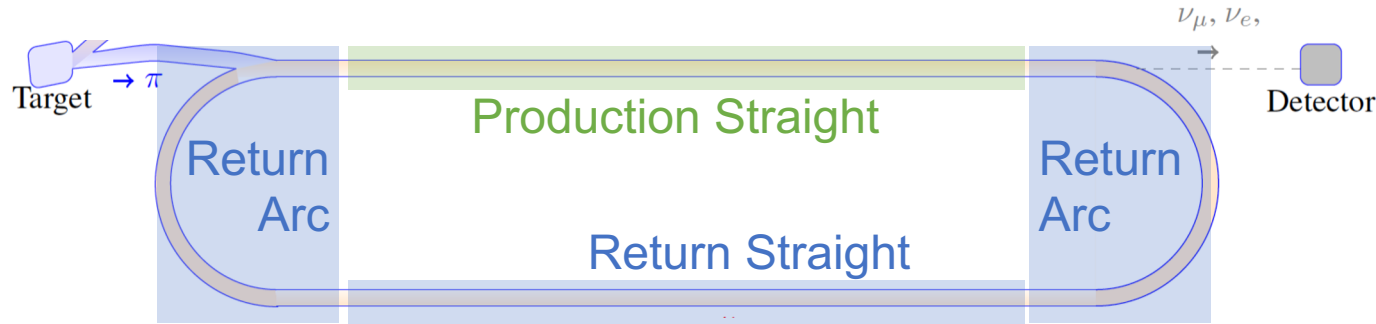


ν from STORed Muons (nuSTORM)



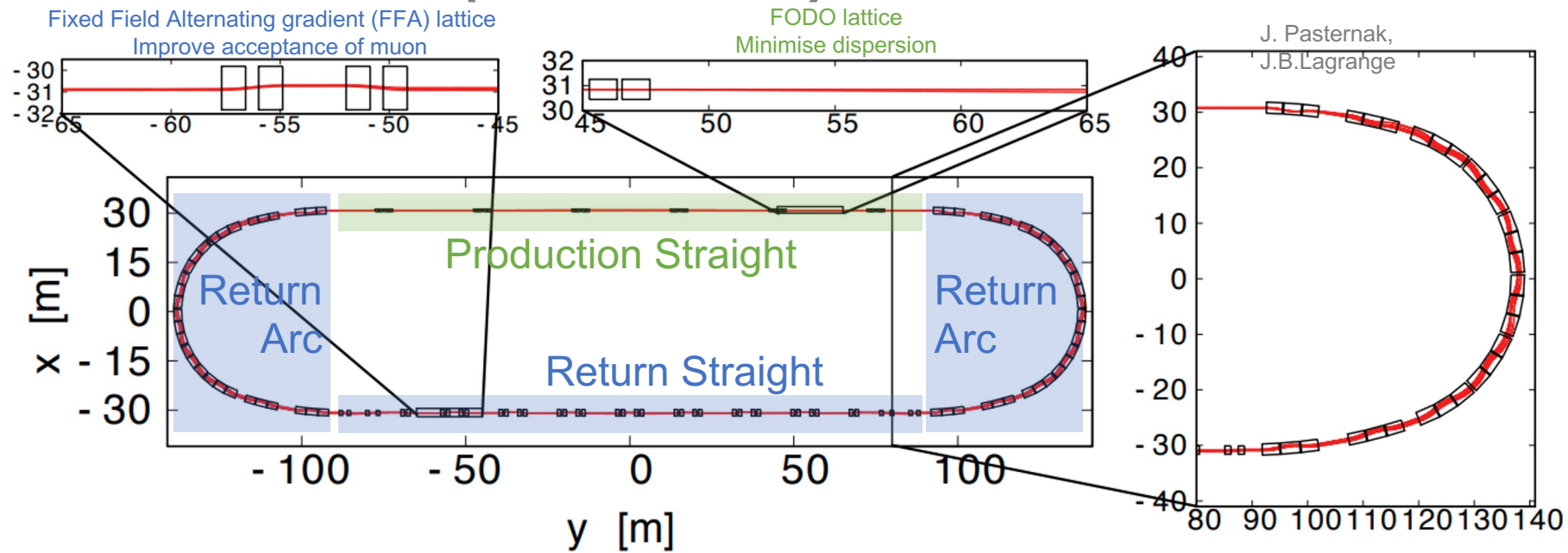
- ❑ **1st ν beam facility & highest ever beam power based on stored muons**
- ❑ Fine tune neutrino fluxes via μ acceptance
 - ❖ p_μ spread: flux shape
 - ❖ p_μ mean: neutrino beam energy scan (ν BES)

ν from *STORed* Muons (*nuSTORM*)



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ν from *STORed* Muons (*nuSTORM*)



❑ **1st ν beam facility & highest ever beam power based on stored muons**

❑ Fine tune neutrino fluxes via μ acceptance

❖ p_μ spread: flux shape

❖ p_μ mean: neutrino beam energy scan (ν BES)

❑ Production Straight (example w/ π^+ injection)

❖ ν_μ flux from $\pi^+ \rightarrow \mu^+ \nu_\mu$ ("pion flash")

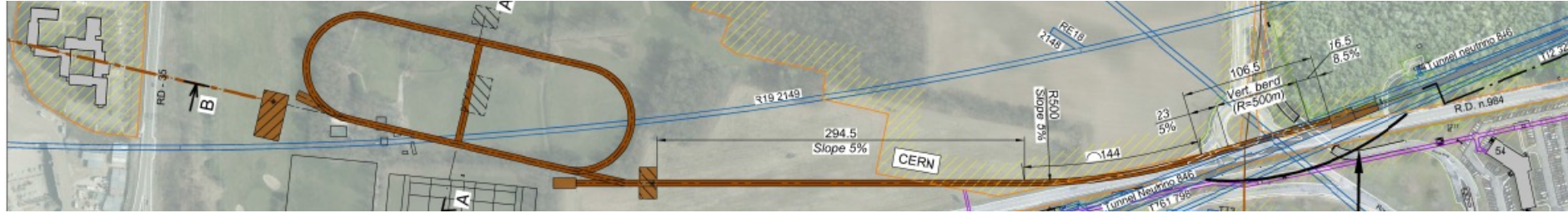
❖ $\nu_e + \bar{\nu}_\mu$ flux from $\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$

❖ Maximise μ capture efficiency

❑ Return Arcs and Straight

❖ $\langle p_\mu \rangle$ tunable between 1 and 6 GeV/c, spread $\pm 16\%$

A Brief History of nuSTORM



2012-13 Lol and Proposal to FNAL PAC [arXiv:1206.0294, arXiv:1308.6822], EoI to CERN [arXiv:1305.1419]

- Sterile neutrinos
- Neutrino-nucleus scattering
- Technology test bed for muon accelerators

2014 Steriles sensitivity [Phys.Rev.D 89, 071301 (2014)]

- nuSTORM at FNAL

2019 Feasibility of nuSTORM at CERN [CERN-PBC-REPORT-2019-003]

- SPS 100 GeV proton beam
- Optimised for neutrino-nucleus scattering, maintaining sensitivity to BSM (steriles + non-unitarity, NSI, Lorentz-invariance/CPT violation)

2022 Snowmass 2021 [arXiv:2203.07545]

- Advocating synergy with ENUBET and Muon Collider Demonstrator
 - ✓ Muon Collider demonstrator
 - ❖ 6-D cooling
 - ✓ nuSTORM as test bed for muon storage ring
 - ❖ Complete implementation for large acceptance
 - ❖ R&D for very precise determination of stored-muon energy and spread

❑ Why do we need nuSTORM?

✓ Wishlist

1. ν_e and $\bar{\nu}_e$ beams with the relevant energy for appearance
2. Well-understood fluxes
3. High statistics
4. Low ν_μ background

❑ What other physics can we do with it?

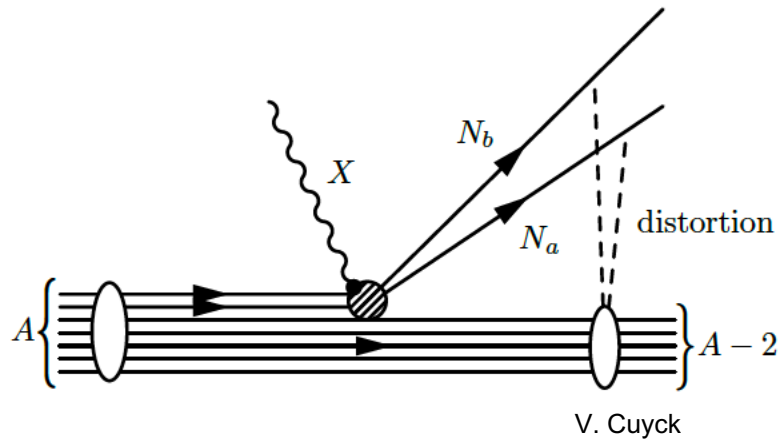
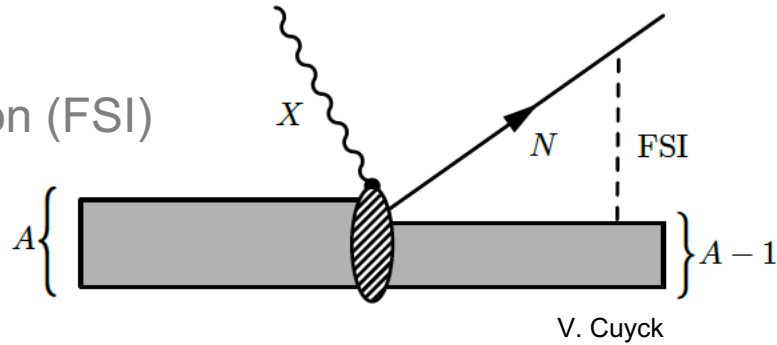
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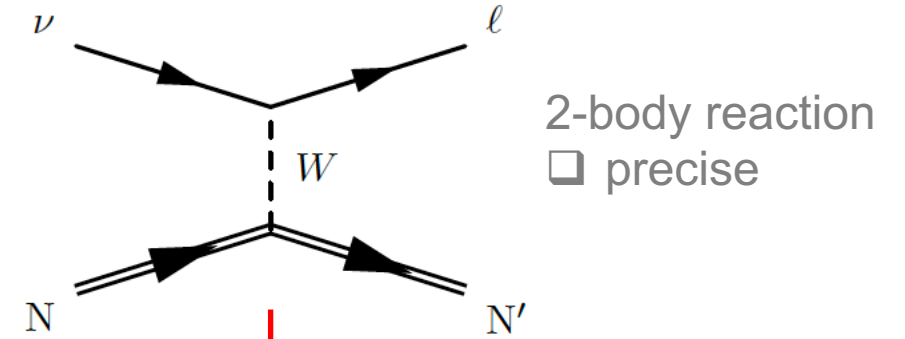
Why Neutrino Interactions?

How well can we measure neutrino energy?
 (reminder: oscillation very sensitive to baseline and energy, L/E)

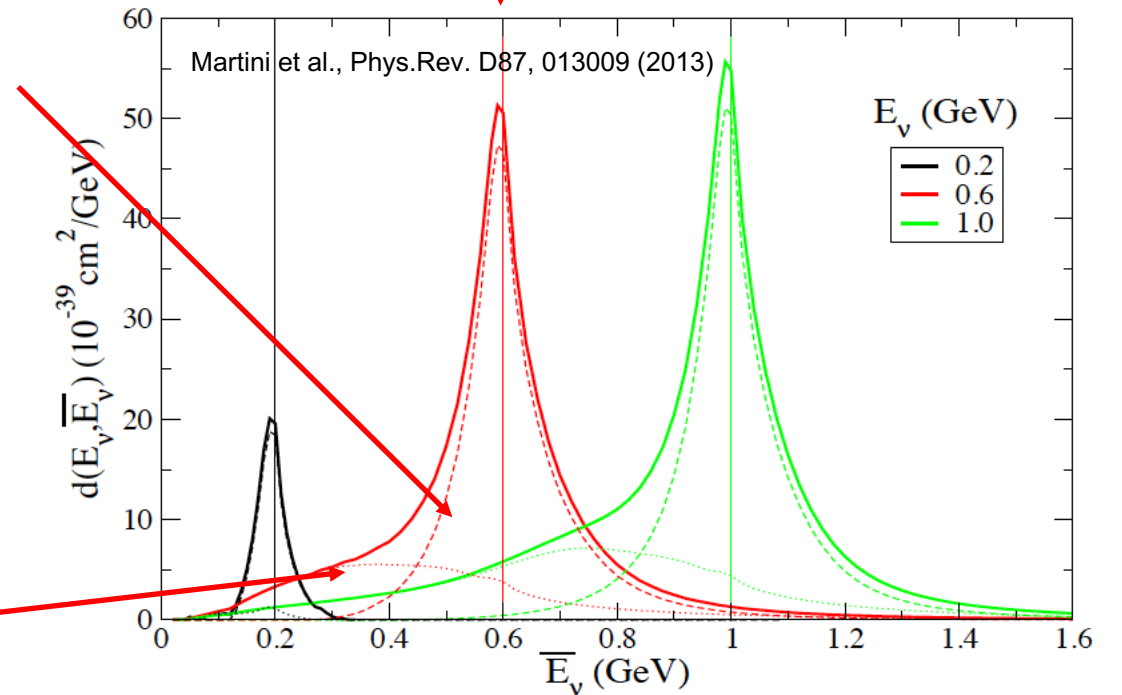
Fermi motion,
 Final State Interaction (FSI)
 spread



2-particle-2-hole (2p2h),
 pion absorption (missing particles)
 Large fraction of large bias and spread



2-body reaction
 precise

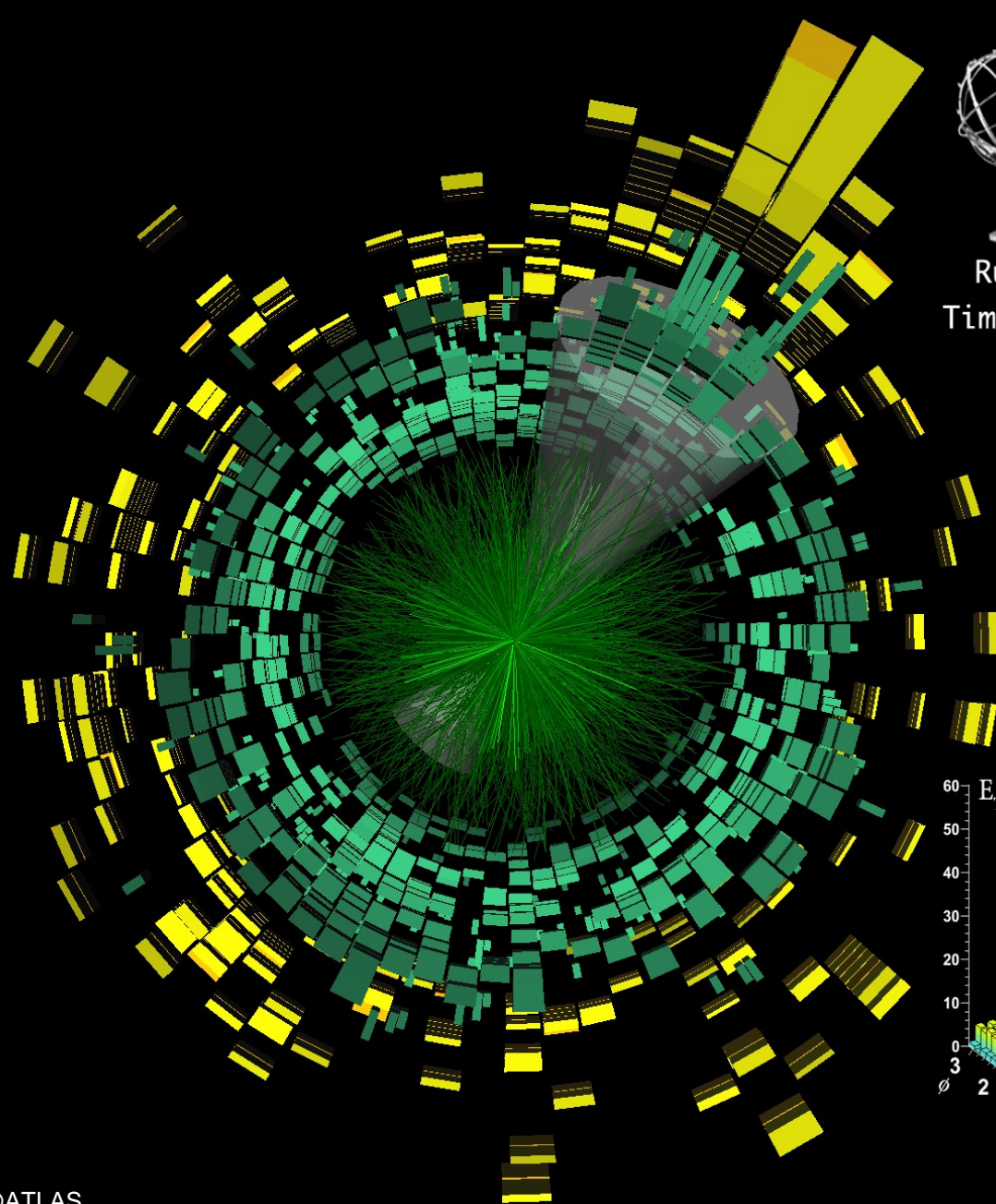


2010: Watched a live ATLAS seminar in the Tower—didn't see what the big deal was...



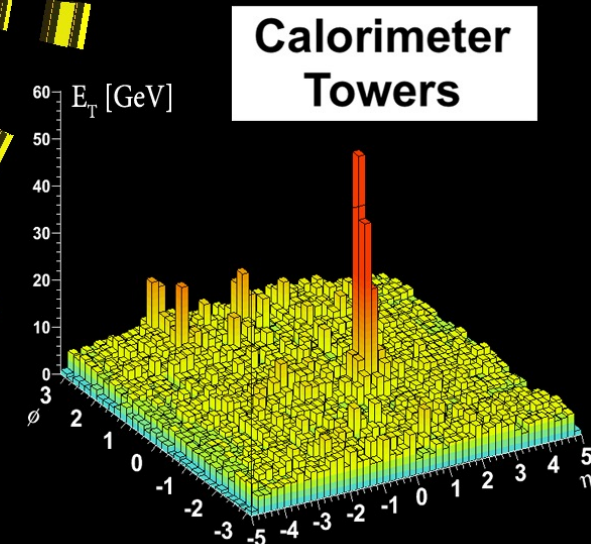
- Over time in ALICE, learnt that *imbalance & correlations* are powerful tools for studying medium effects.
- 2014/2015: Moved into neutrinos and *started to* study neutrino interactions with hadron-lepton correlations.
- Transverse Kinematic Imbalance (TKI) now active topic in T2K, MINERvA, MicroBooNE, SBND, ICARUS with real data, and also in future projects like DUNE and nuSTORM.

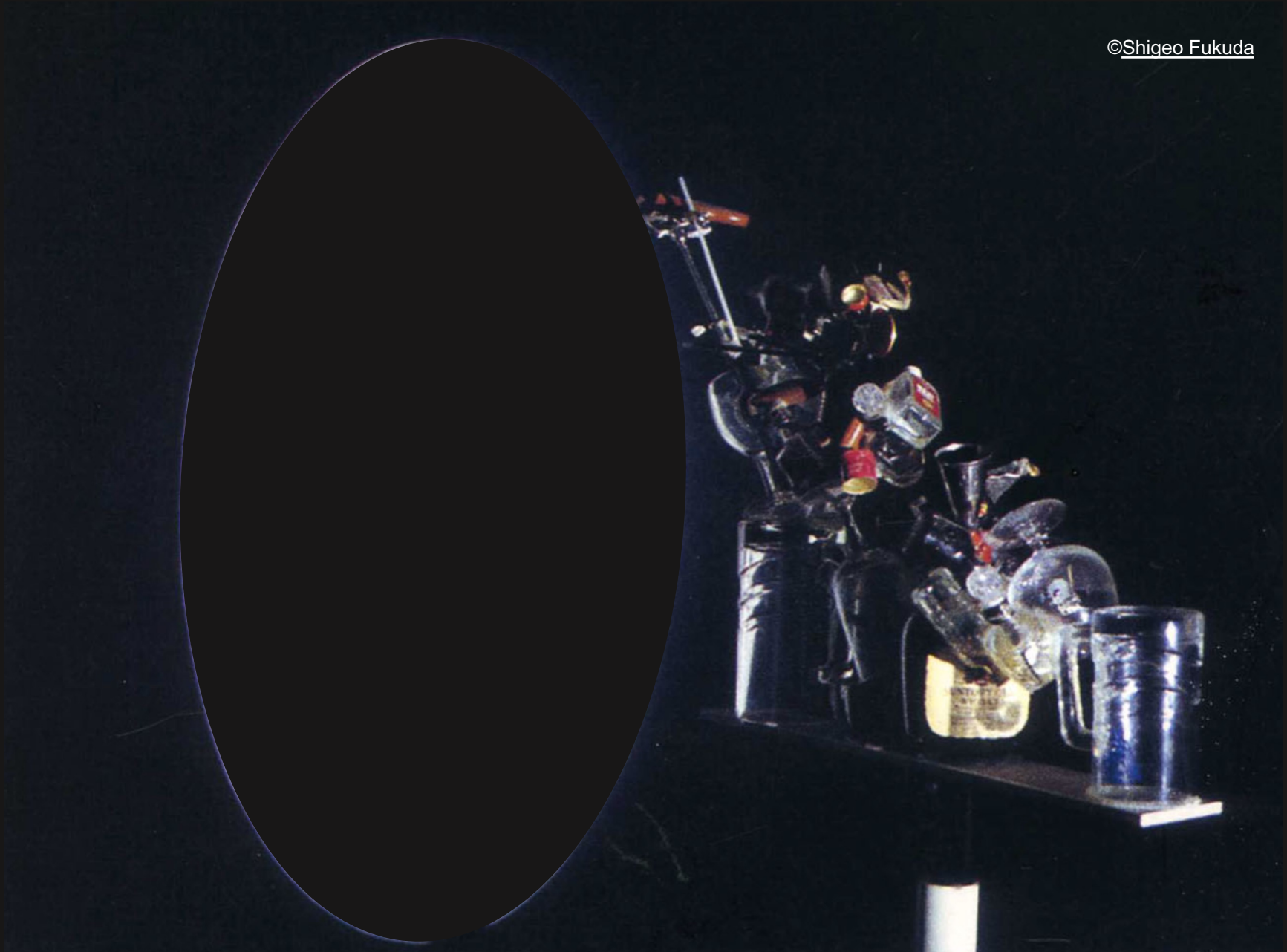
@ATLAS



 **ATLAS**
EXPERIMENT

Run 168795, Event 7578342
Time 2010-11-09 08:55:48 CET

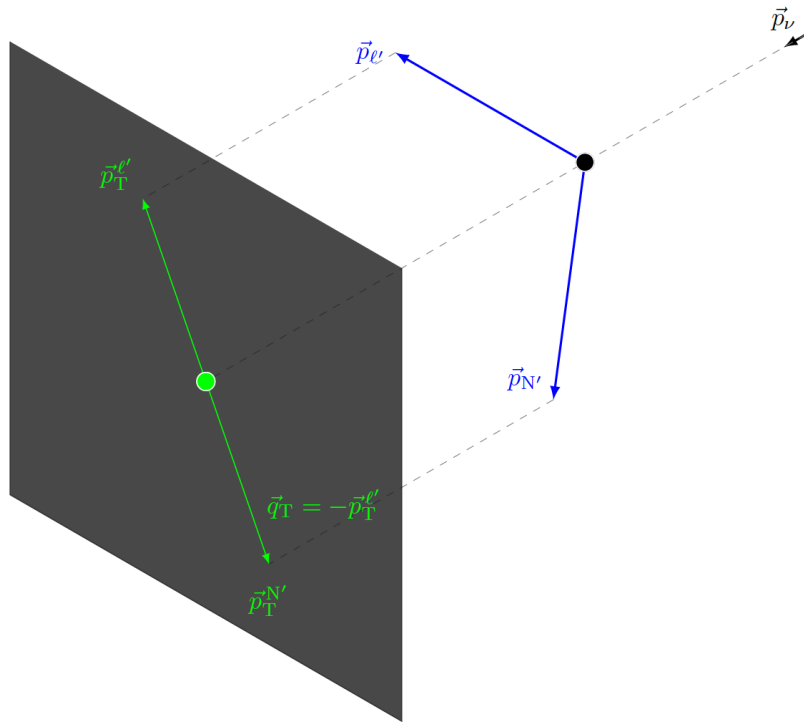




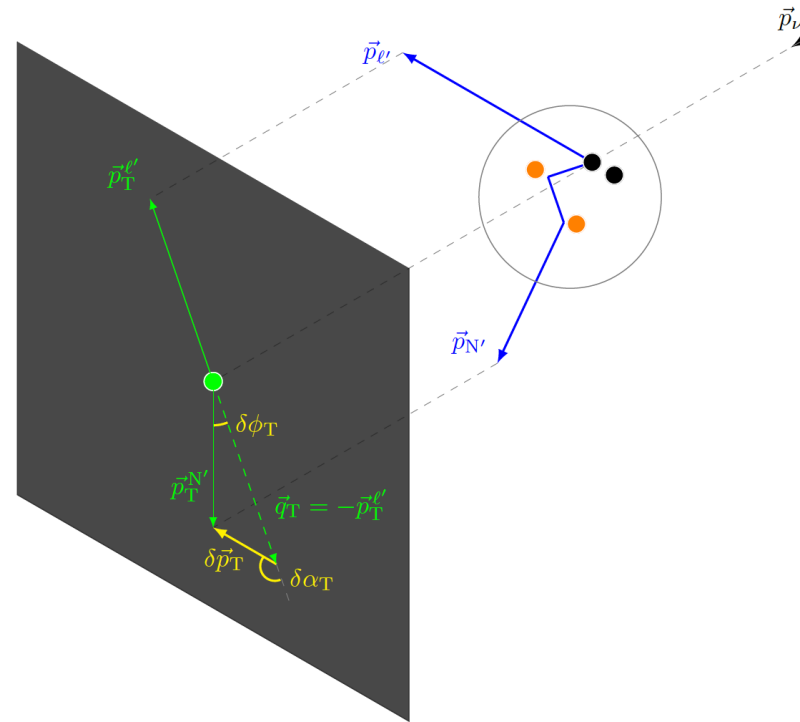


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Transverse Kinematic Imbalance (TKI)



Stationary free nucleon target



Nuclear target ($A > 1$)

- Fermi motion
- Final-state interactions (FSI)
- 2-particle-2-hole (2p2h)

Missing energy



From Wikipedia, the free encyclopedia

[...]
[neutrinos](#).^[1] In general, missing energy is used to infer the presence of non-detectable particles and is expected to be a signature of many theories of [physics beyond the Standard Model](#).^{[2][3][4]}
 [...]
[hadron colliders](#).^[5] The initial momentum of the colliding [partons](#) along the beam axis is not known —

TKI

Multi-dimensional observation

- Momentum (magnitude)
- Angle
- Asymmetry

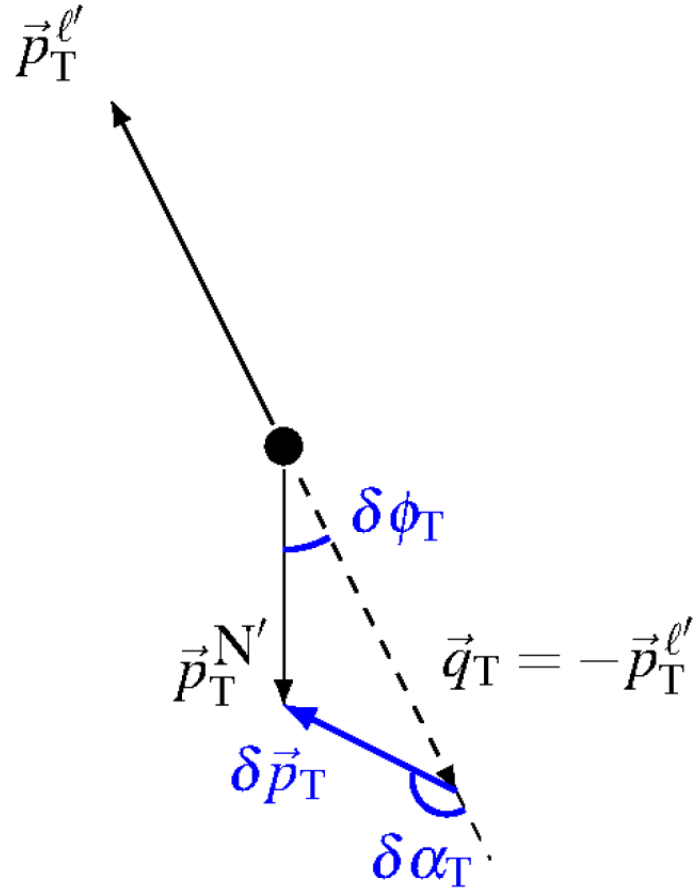
[Lu, et al., Phys.Rev.D 92, 051302 \(2015\)](#)

[Lu, et al., Phys.Rev.C 94, 015503 \(2016\)](#)

[Lu & Sobczyk, Phys.Rev.C 99, 055504 \(2019\)](#)

[Cai, Lu, Ruterbories, Phys.Rev.D 100, 073010 \(2019\)](#)

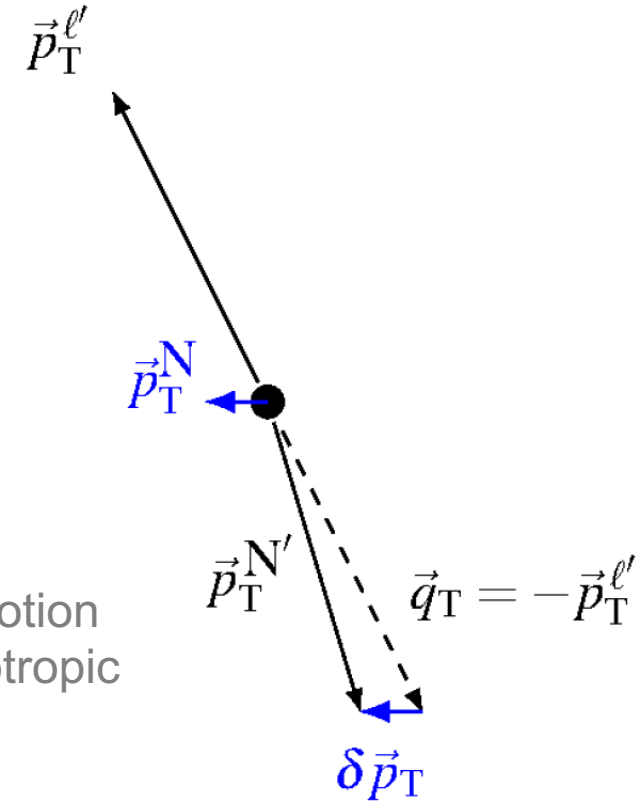
Transverse Boosting Angle $\delta\alpha_T$



if Fermi motion only



$\delta\vec{p}_T = \vec{p}_T^N$
 $\delta\alpha_T$ is Fermi motion
 direction \rightarrow isotropic



$\delta\vec{p}_T$

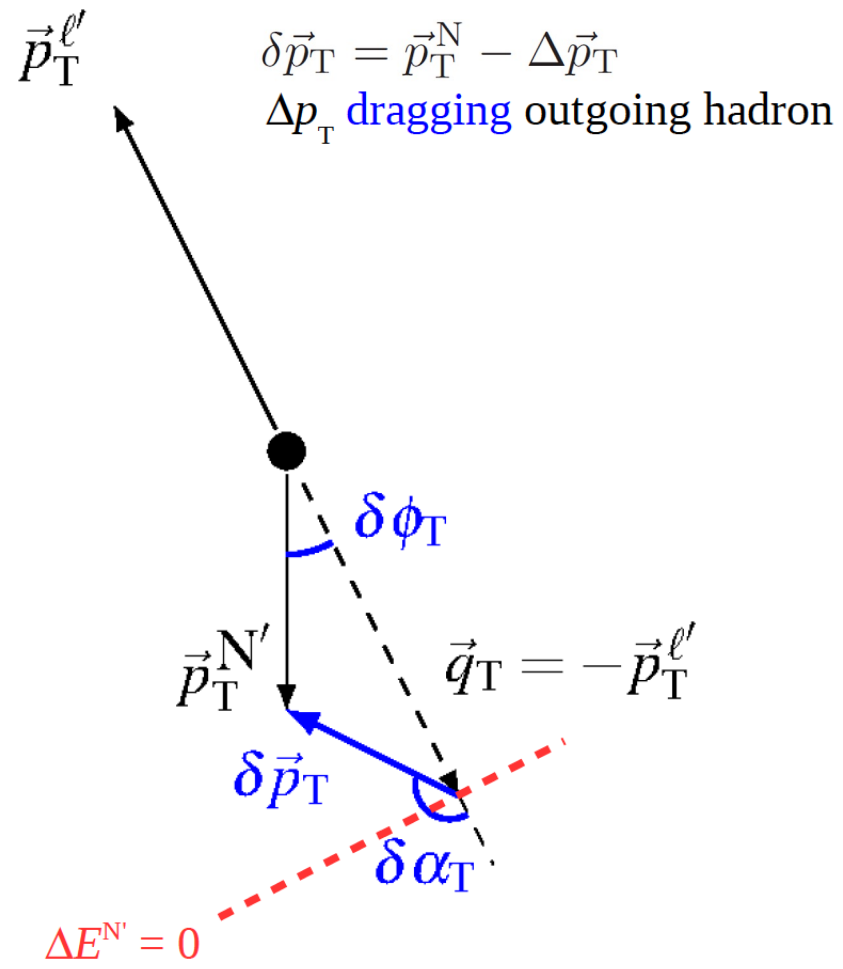
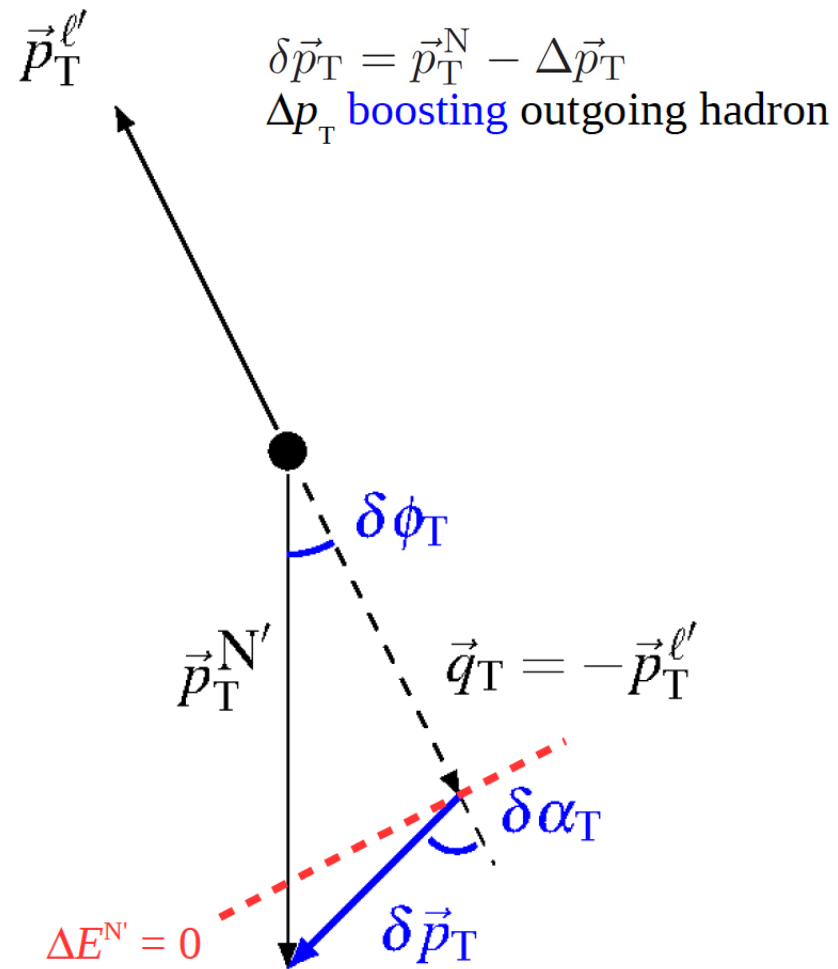
- total transverse momentum
- transverse momentum imbalance
- missing p_T
- ...

In full

$$\delta\vec{p}_T = \vec{p}_T^N - \Delta\vec{p}_T$$

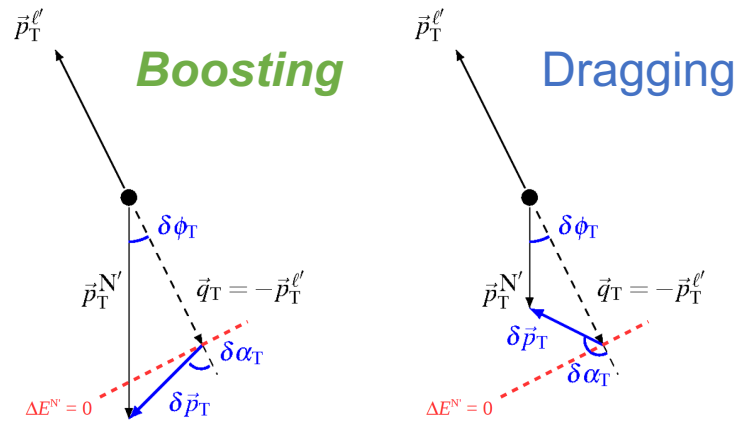
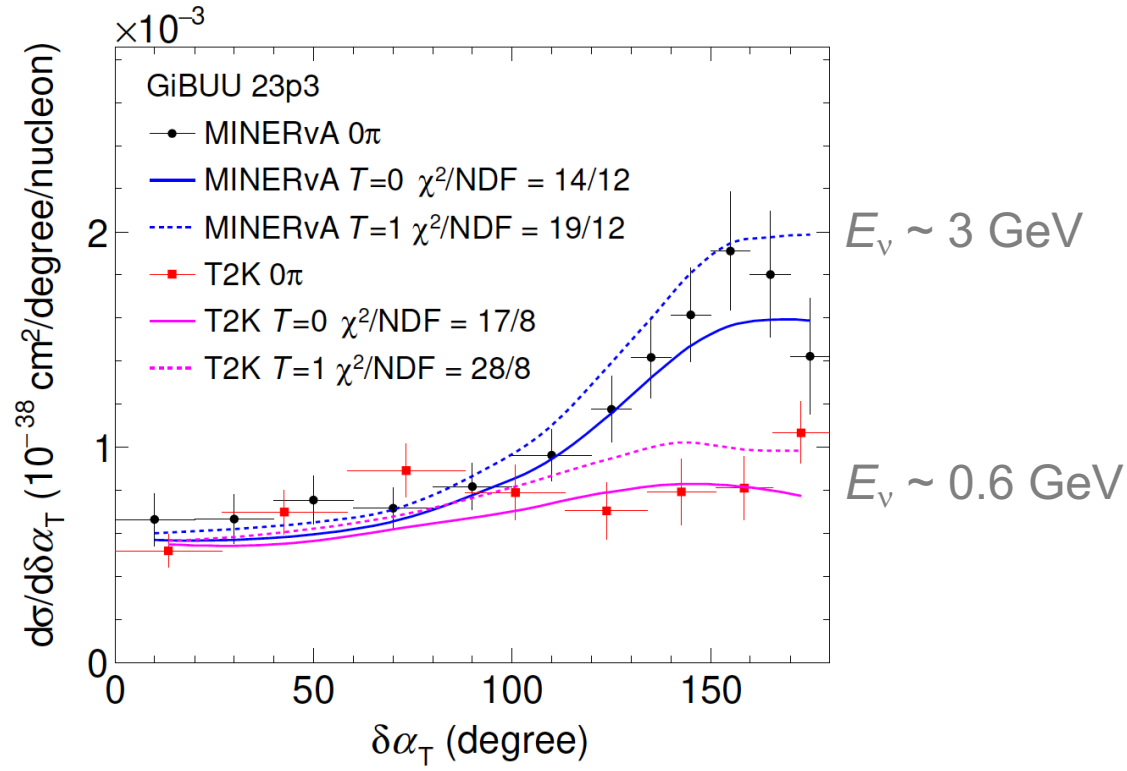
— FSI and missing particles

Transverse Boosting Angle $\delta\alpha_T$

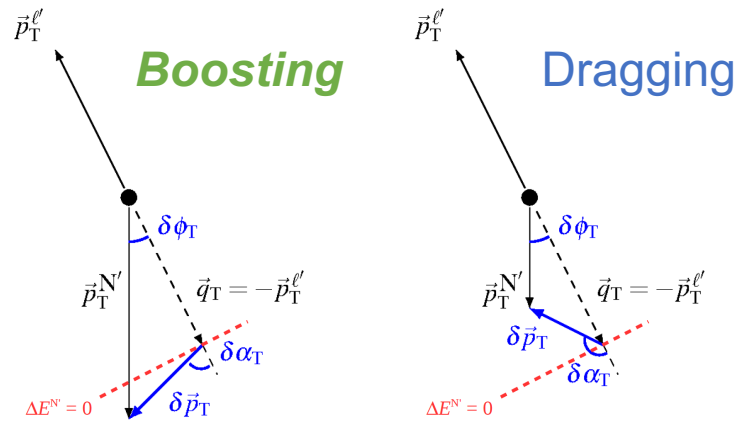
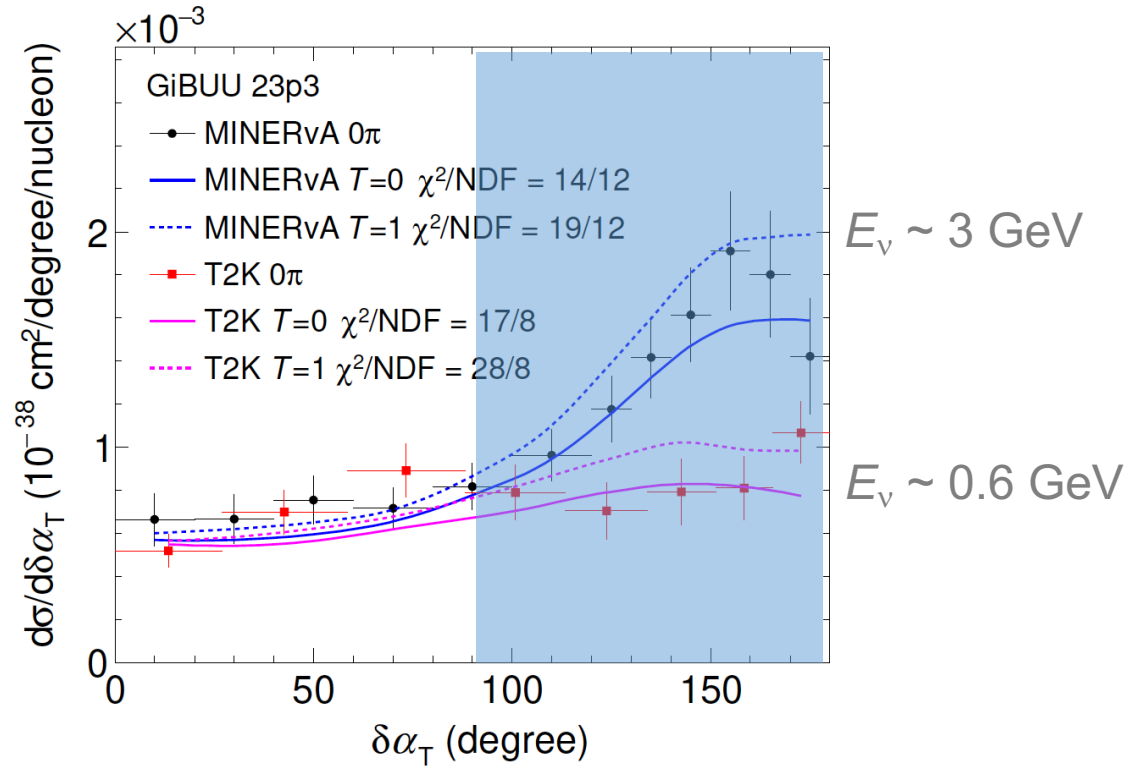


- FSI and momentum sharing with extra particles
- pion absorption
 - 2p2h

TKI Measurements at T2K, MINERvA, and nuSTORM

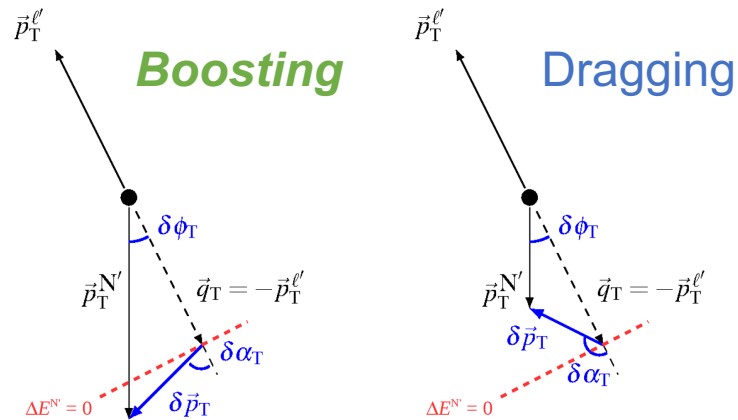
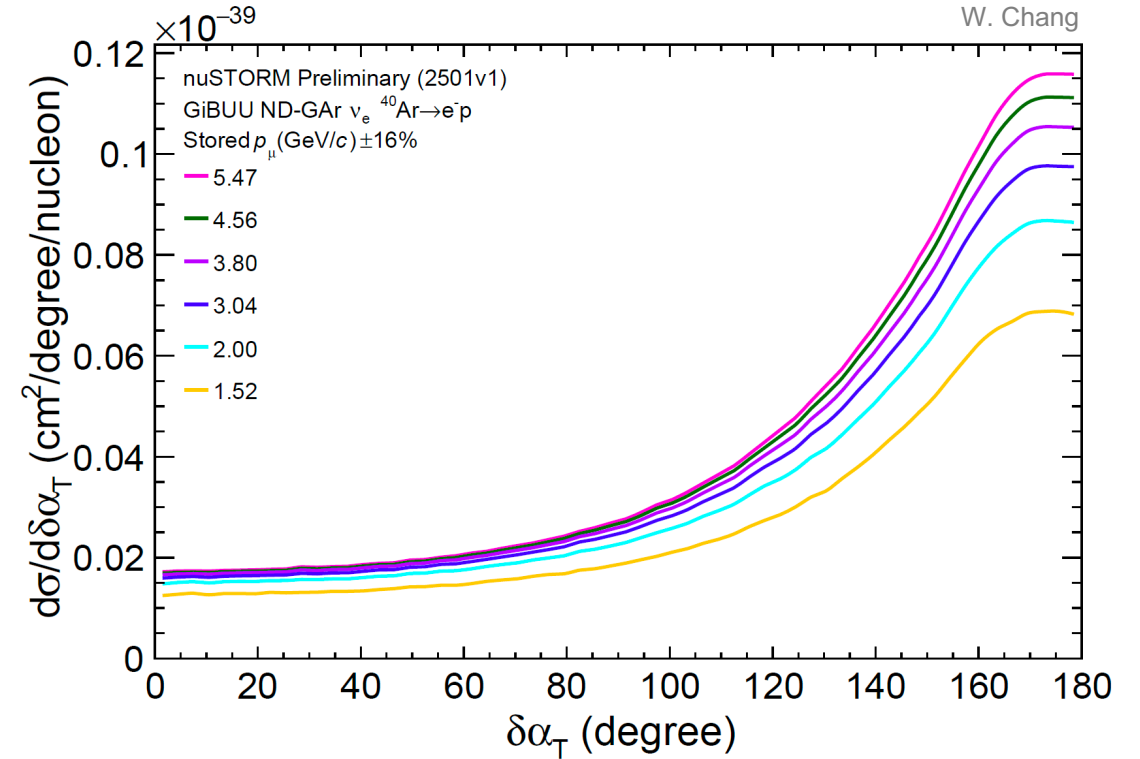
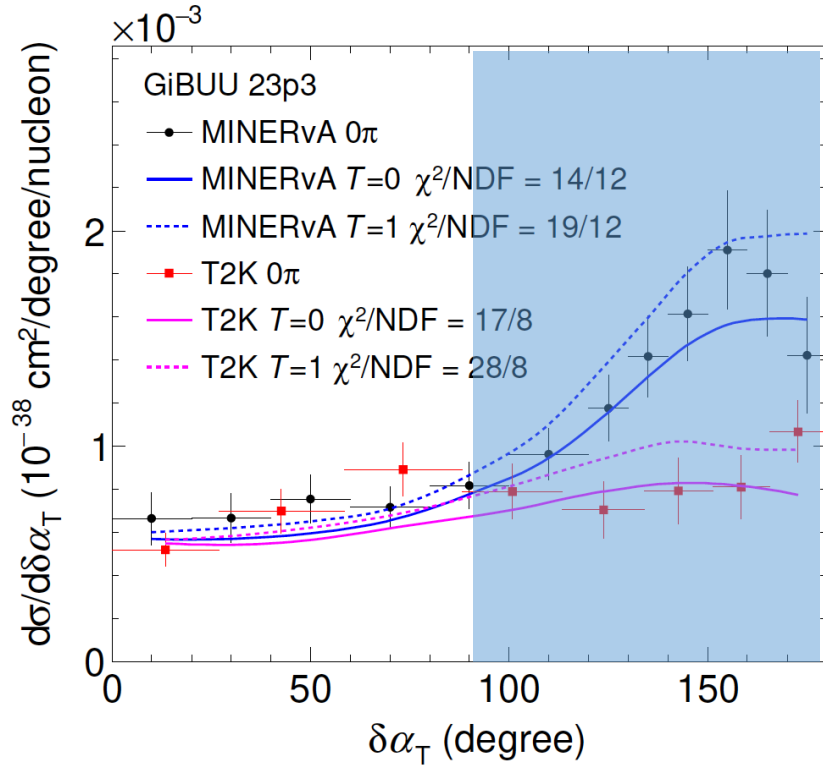


TKI Measurements at T2K, MINERvA, and nuSTORM



□ Dissipative processes: 2p2h, FSI

TKI Measurements at T2K, MINERvA, and nuSTORM

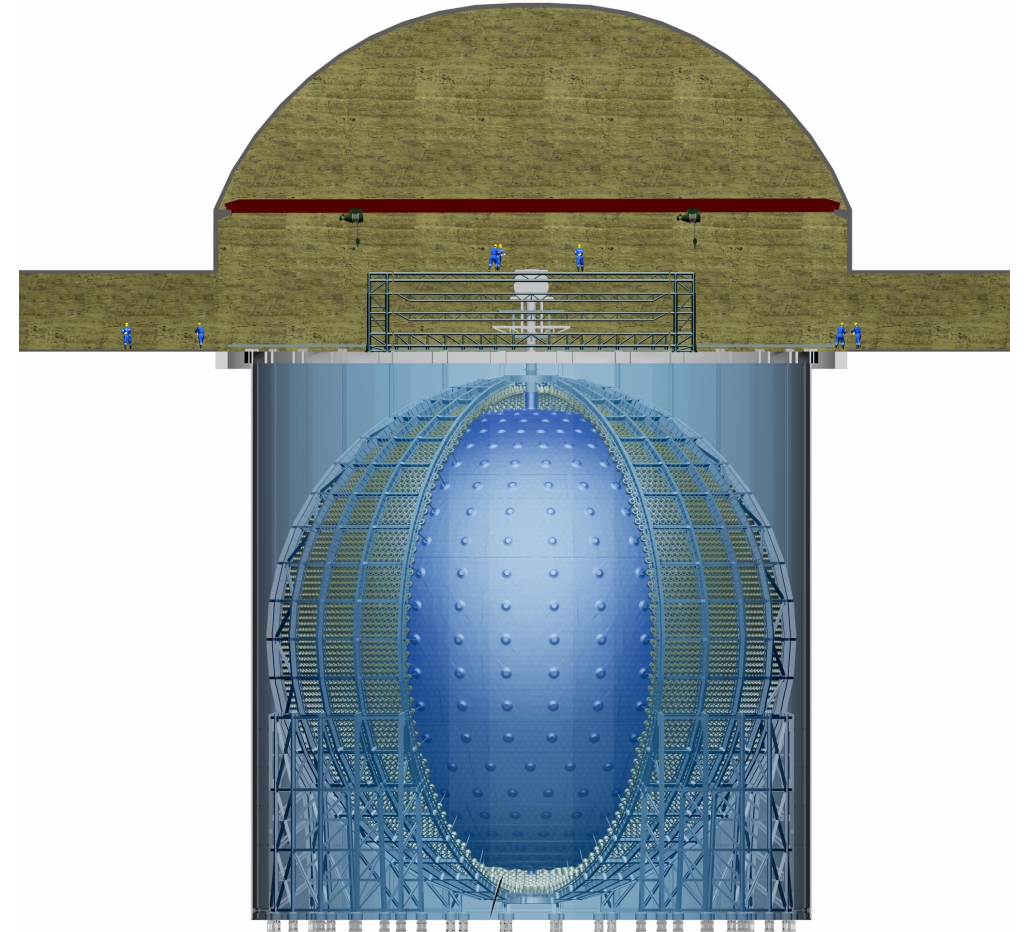
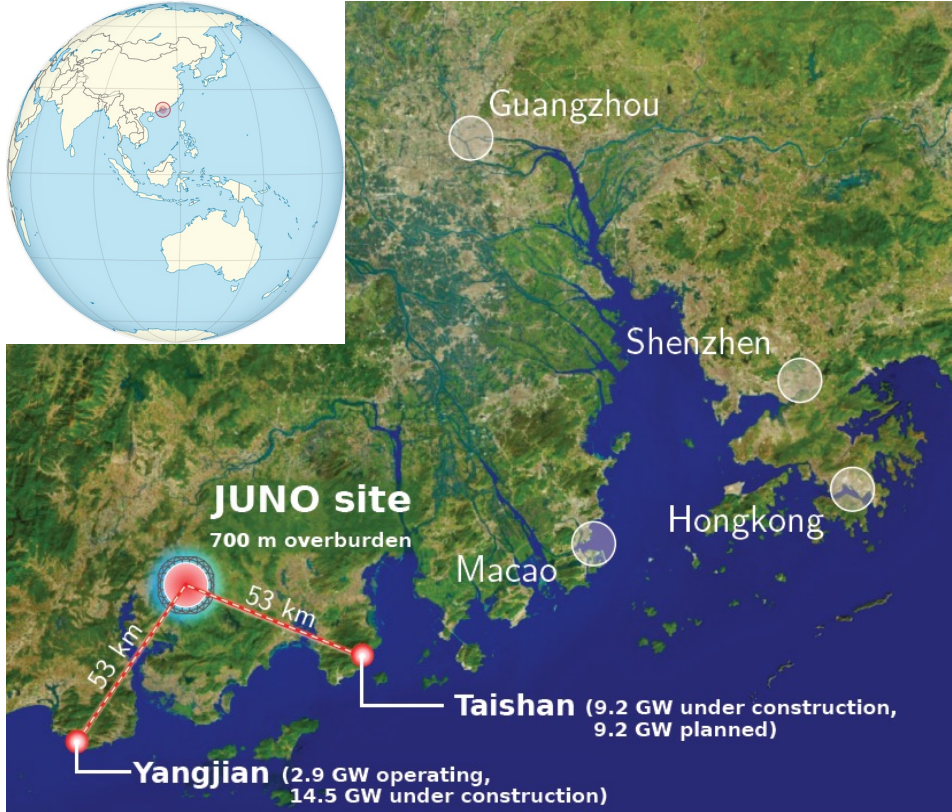


- Dissipative processes: 2p2h, FSI
- Dynamical evolution of nuclear medium effects can be mapped out via ν BES at nuSTORM

Outline

1. Problems and opportunities with neutrino masses
 - Call for a GeV ν_e and $\bar{\nu}_e$ machine: nuSTORM
2. Using lepton-hadron correlations to study GeV neutrino interactions
 - Transverse Kinematic Imbalance (TKI)
3. JUNO news!

Jiangmen Underground Neutrino Observatory (JUNO)



❑ JUNO: *primarily* a reactor neutrino experiment

- ❖ 20 kt liquid scintillator
- ❖ 17,612 20-inch PMTs + 25,600 3-inch PMTs
 - 77.9% photocathode coverage

❖ **Expected mass ordering sensitivity with reactor neutrinos: 2.9 sigma in 6 years**

❑ Turns out big enough to contain atmospheric ν up to 10 GeV

- 3-10 GeV sensitive to ν mass ordering

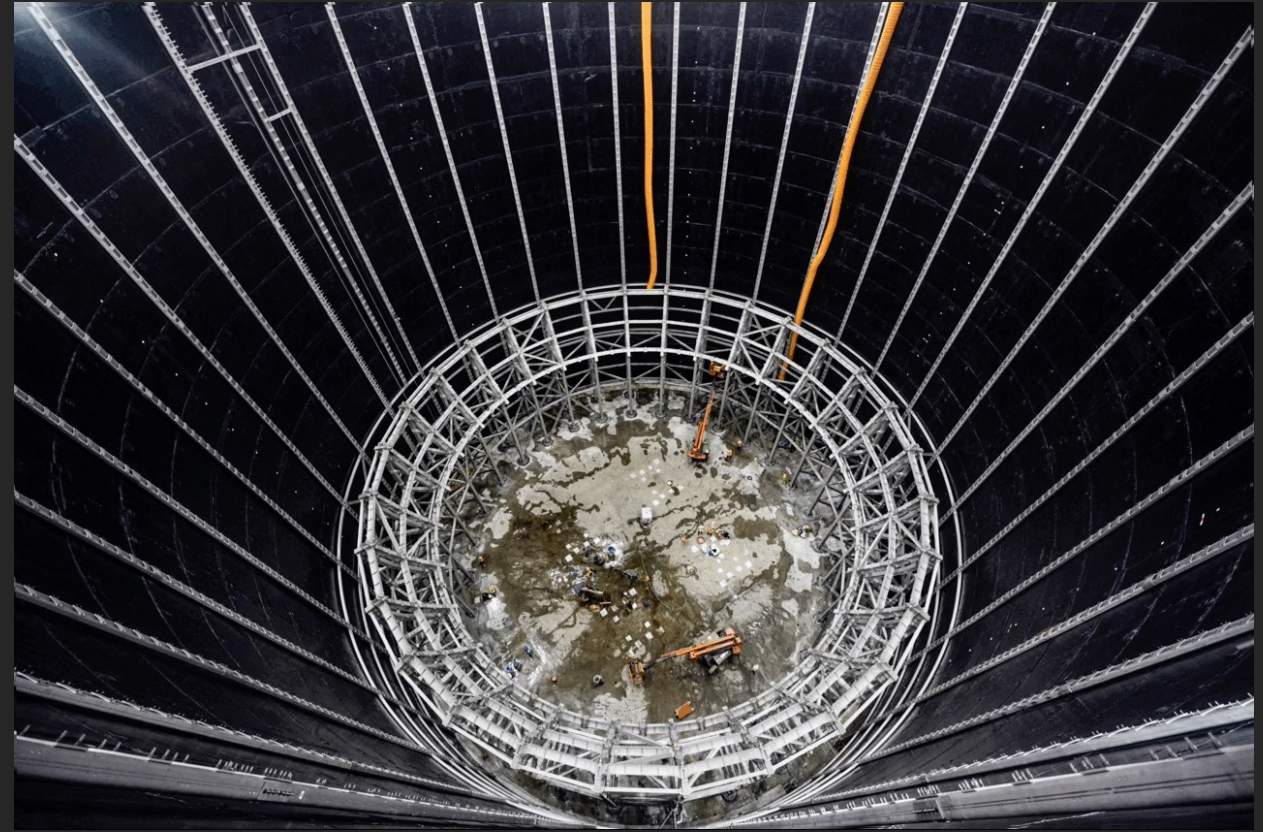
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2022/03—2024/12

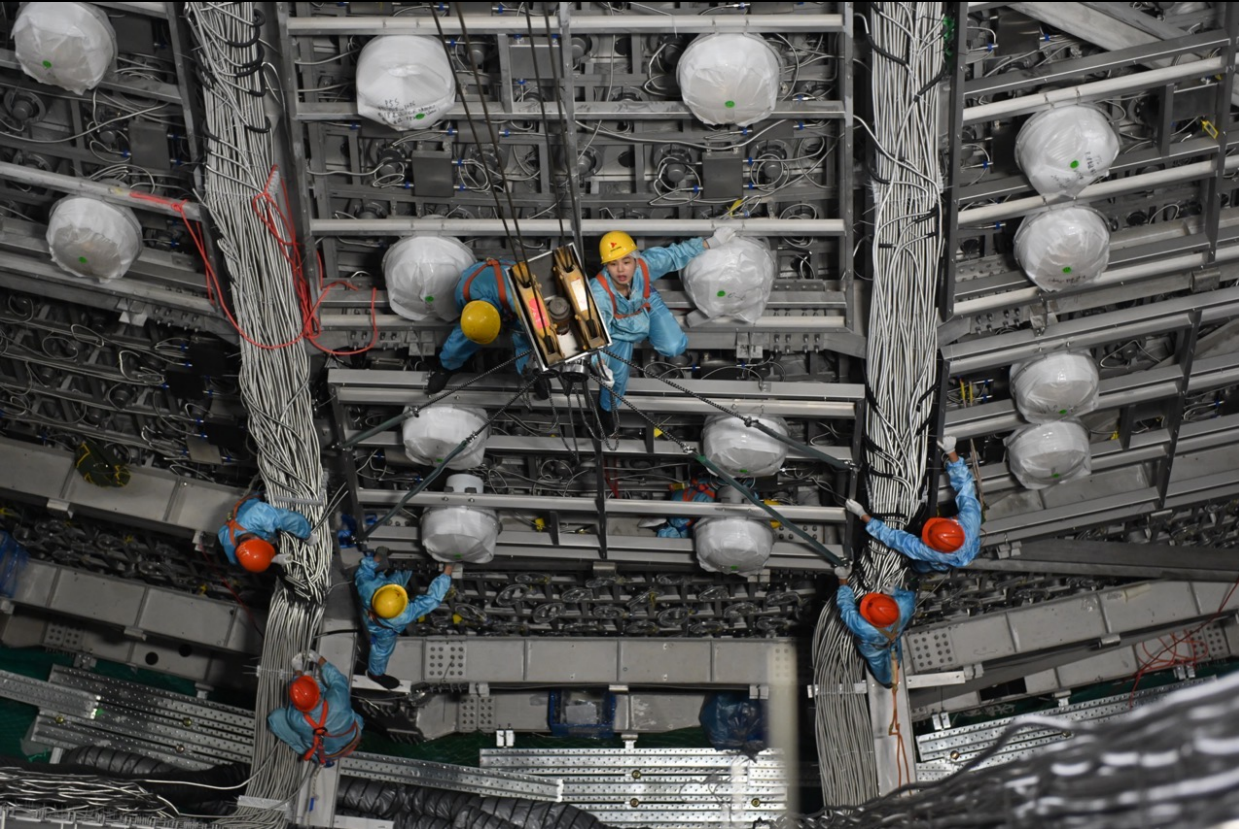
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<https://www.nature.com/articles/d41586-024-00694-5>



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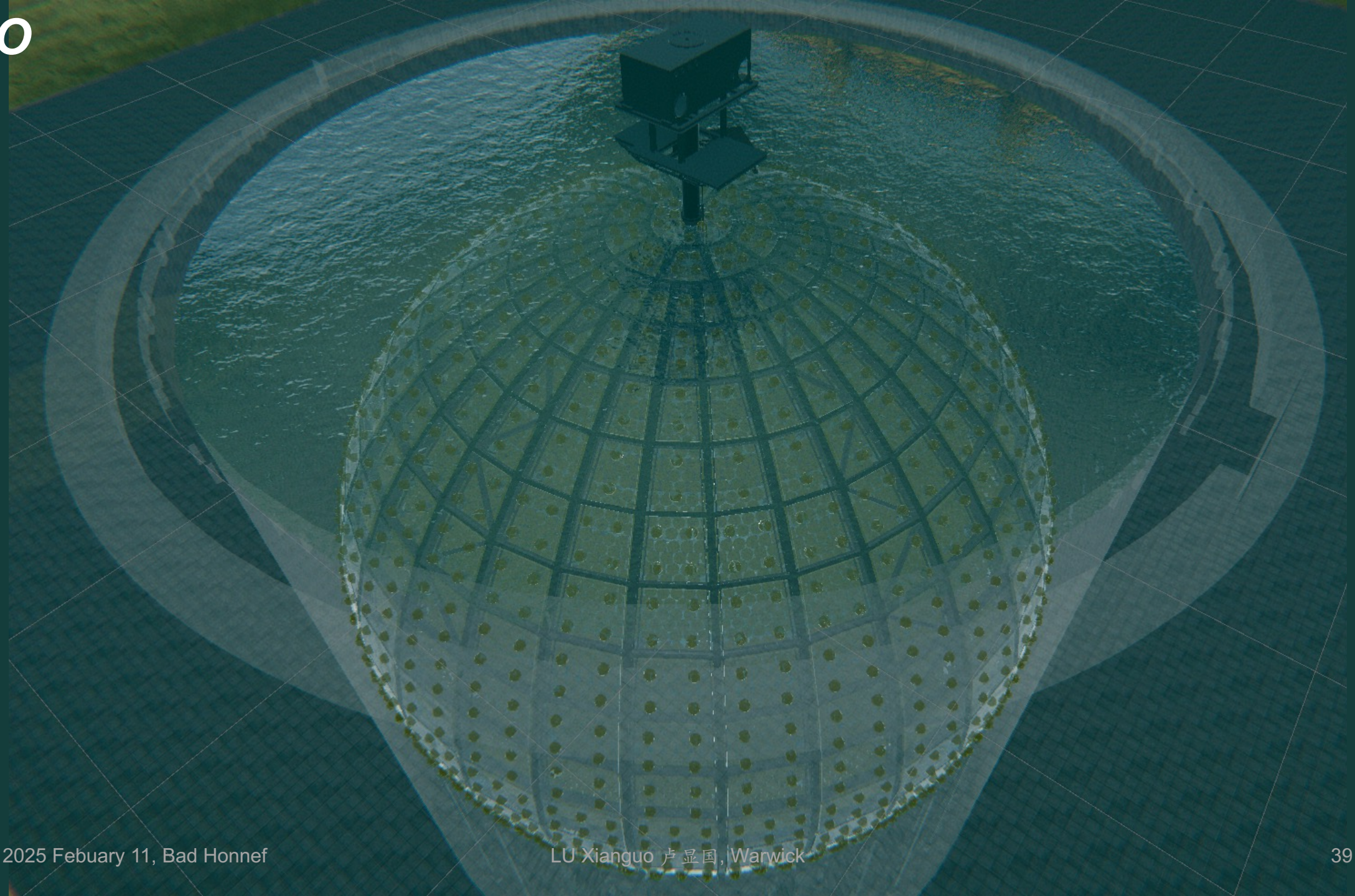


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When I left Heidelberg and the field, I wasn't sure what to expect. But looking back, I realise those years had prepared me well for the journey ahead. Thank you, Johanna!

Now, let's celebrate Johanna's special day and this wonderful gathering!

2025 February 11, Bad Honnef

松齡萬古春

壬子年
石年



松齡鶴壽
壬子年
石年

鶴僊千辛壽