Probing Quark-Gluon Plasma at the Large Hadron Collider in the Charmonia sector using Machine Learning



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CBM Juniors' Meeting



Experimental High-Energy Physics Group

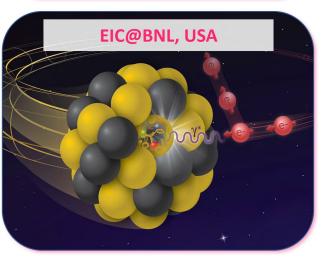






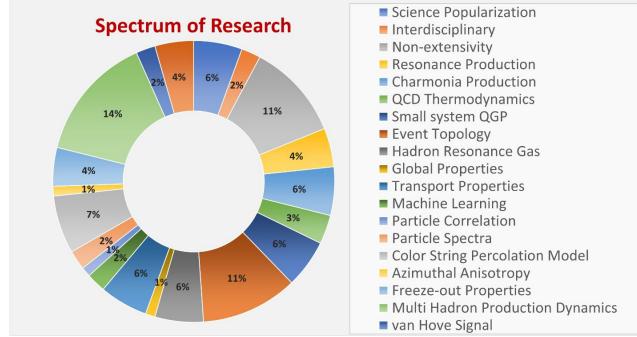
Contact: Prof. Raghunath Sahoo, FinstP e-mail: <u>raghunath@iiti.ac.in</u> Web: www.iiti.ac.in/~raghunath



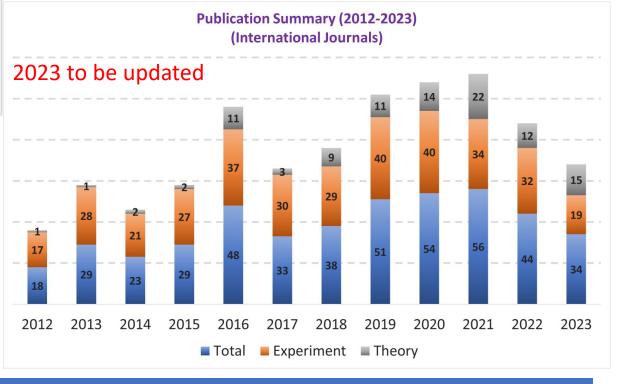




Research Spectrum



Details: www.iiti.ac.in/~raghunath





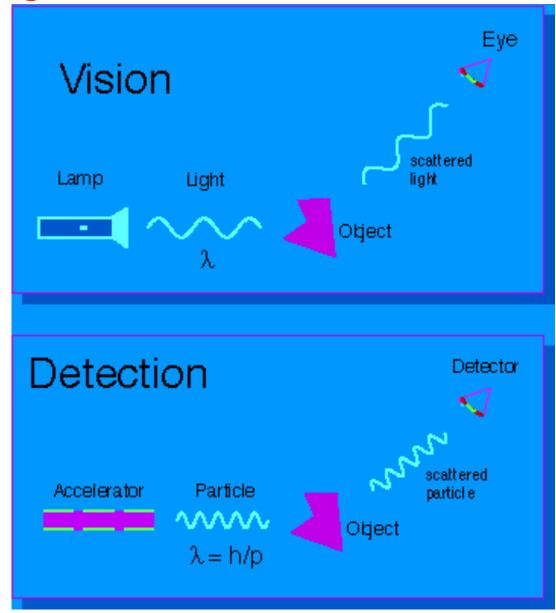
The Right Light to Look Inside

Vision works by scattering of 'visible' light

$$\lambda = 400-700 \text{ nm}$$

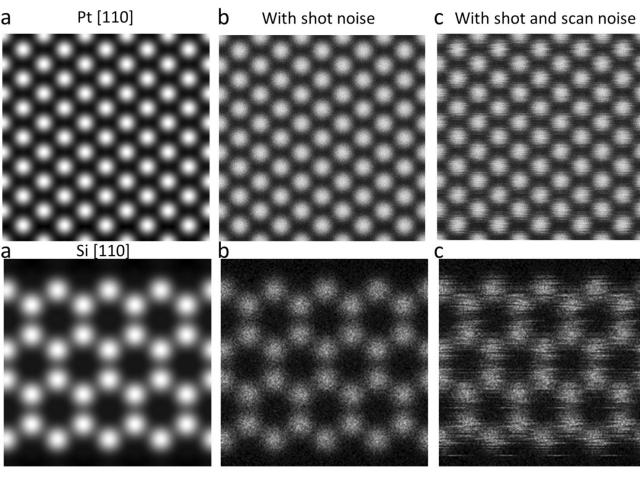
"Vision" of even smaller structures via scattering of particles

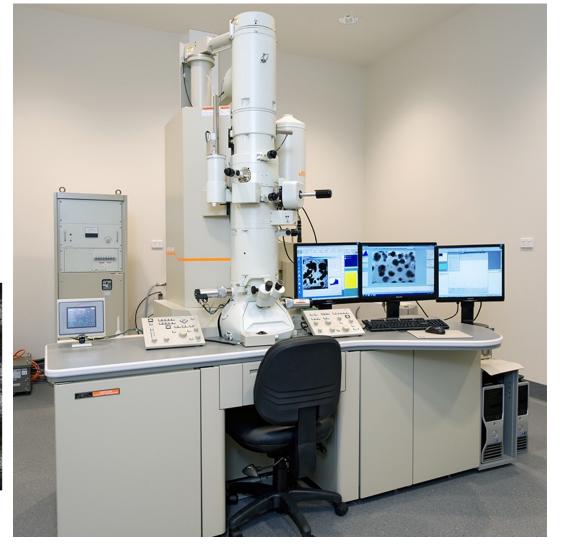
$$\lambda = h/p$$





Seeing the Atoms





(for demonstration purpose)

Transmission Electron Microscope



A Journey from Electrons to Quarks

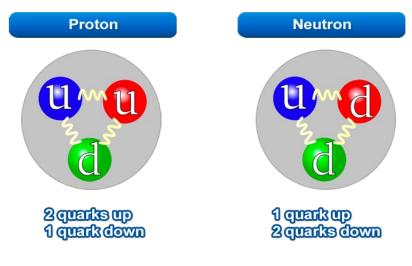


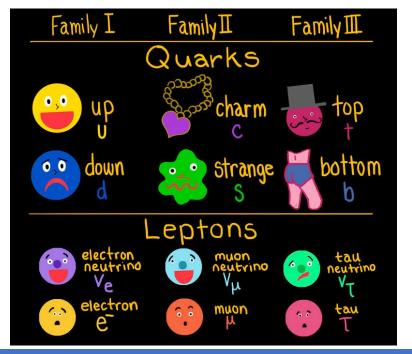
SLAC: Stanford Linear Accelerator Center (3.2 Kms) Electron-positron energy ~ 50 GeV

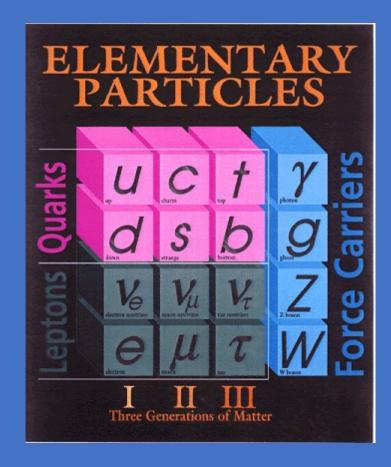
 $\lambda = h/p: \alpha 1/E$

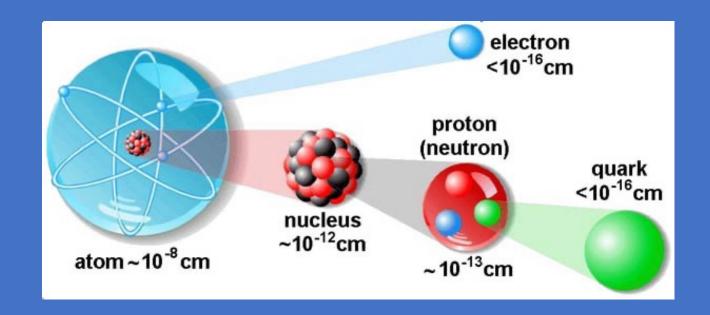
This universal formula of de Broglie helps in deciding the energy of the probe.

Proton charge radius \sim 0.843 fermi. (1 fermi = 10^{-15} meter)



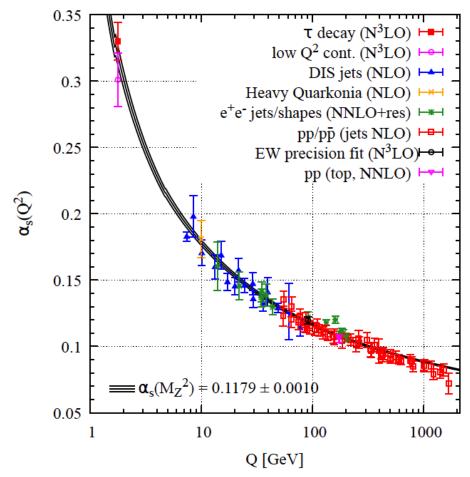








Can we find free quarks?



To the leading order:

$$\alpha_s(Q^2) = \frac{4\pi}{(33-2n_f)ln(Q^2/\Lambda^2)}$$

 n_f is the no. of quark flavors, with a mass less than Q/2, Λ is the QCD scale parameter, obtained experimentally \sim 200 MeV

Coupling becomes weaker as the momentum transfer increases,

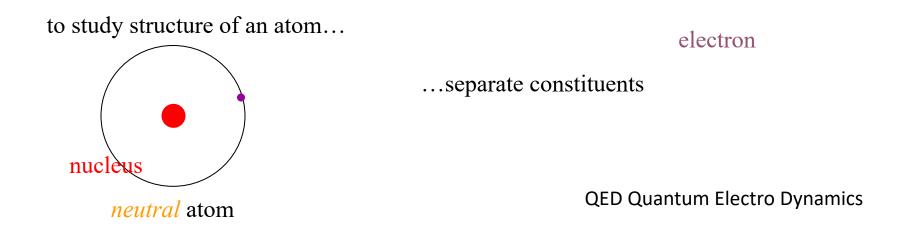
Or as we go smaller in the length-scale \rightarrow small distance, partons move freely (asymptotic freedom of QCD)

←As the inter-quark distance increases, strong coupling grows faster making the quarks confined inside the cage of the hadrons → Quark confinement

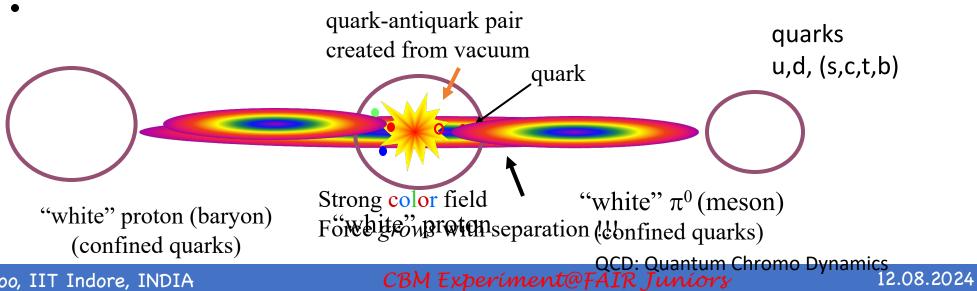
Asymptotic freedom and infrared slavery is an inbuilt property of QCD (Strong interaction) > No free quarks in nature!



Analogies and differences between QED and QCD



Confinement: fundamental & crucial (but *not* understood!) feature of strong force - colored objects (quarks) have ∞ energy in normal vacuum





Can we find free quarks?

1973: Asymptotic freedom

D.J. Gross, F. Wilczek, H.D. Politzer

1975: Asymptotic QCD & deconfinement:

N. Cabibbo and G. Parisi;

J. Collins and M. Perry

2004: Nobel Prize







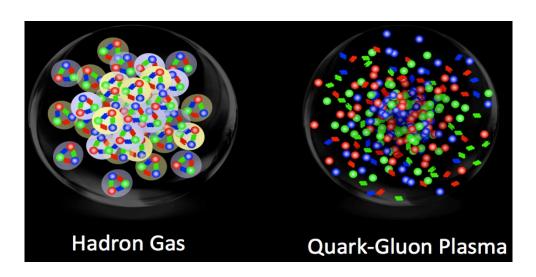
H. David Politzer



Frank Wilczek

QCD predicts that normal nuclear matter undergoes a phase transition to quark-gluon plasma (QGP) under extreme temperatures and energy densities.

Quark Gluon Plasma

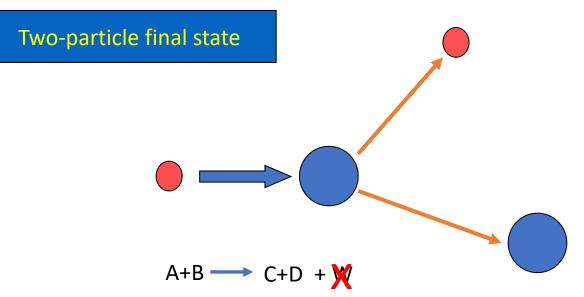


Quark Gluon Plasma (QGP): (locally) thermally equilibrated state of matter in which quarks and gluons are deconfined from hadrons, so that color degrees of freedom become manifest over nuclear, rather than merely nucleonic, volumes.

A journey to the beginning of the universe.....

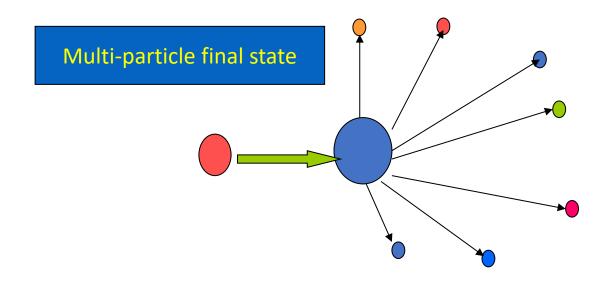


Collisions: Classical vs Relativistic



Classical collisions:

- a) Mass is conserved: $m_A + m_B = m_C + m_D$
- b) Momentum is conserved: $p_A + p_B = p_C + p_D$
- c) Kinetic energy may(not) be conserved
 - i. Sticky- K.E. decreases: $T_A+T_B > T_C+T_D$
 - ii. Explosive- K.E. increases: $T_A+T_B < T_C+T_D$
 - iii. Elastic- K.E. is conserved: $T_A + T_B = T_C + T_D$



Relativistic collisions:

- a) Energy is conserved: $E_A + E_B = E_C + E_D$
- b) Momentum is conserved: $p_A + p_B = p_C + p_D$
- c) Kinetic energy may(not) be conserved

Mass-energy conversions in "explosive collisions"!

If mass is conserved: collision is elastic.







Energy available for particle production:

In laboratory frame (fixed target):

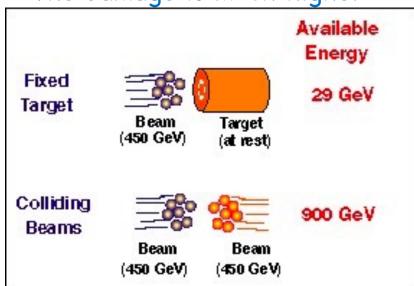
 $E_{cm} = \sqrt{(2m_t E_{beam})}$

In CM frame (collider):

 $E_{cm} = (2E_{beam})$

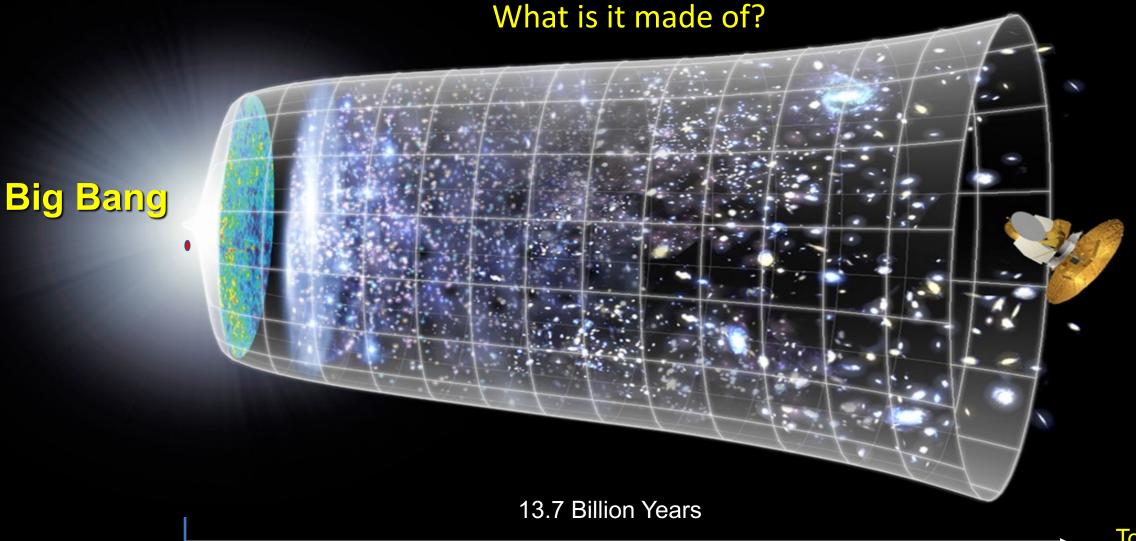


The damage is much higher!





Our Universe How did it start?

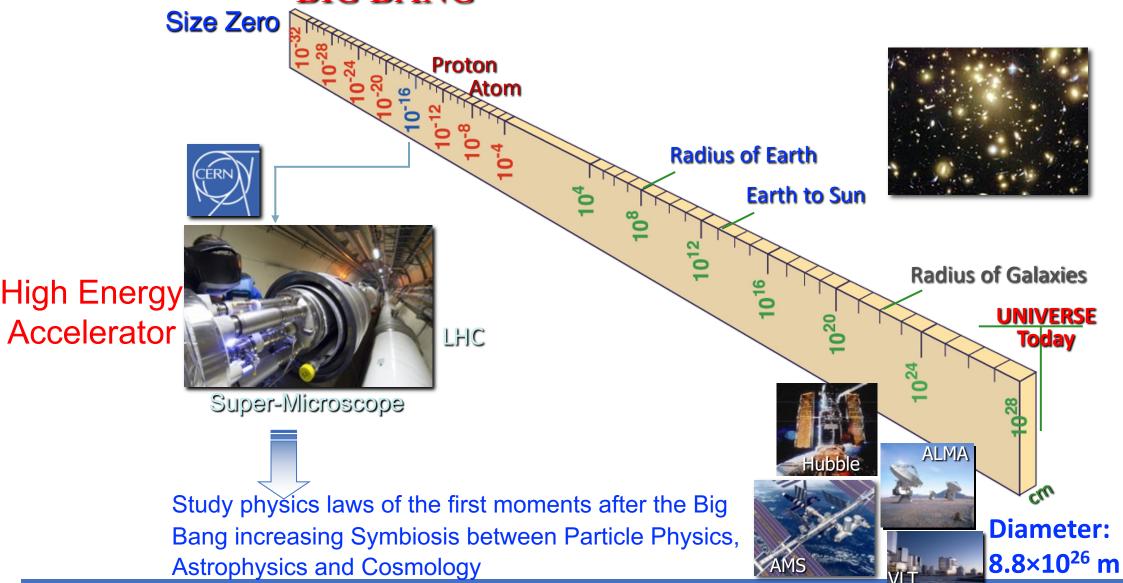




Accessing Big Bang conditions





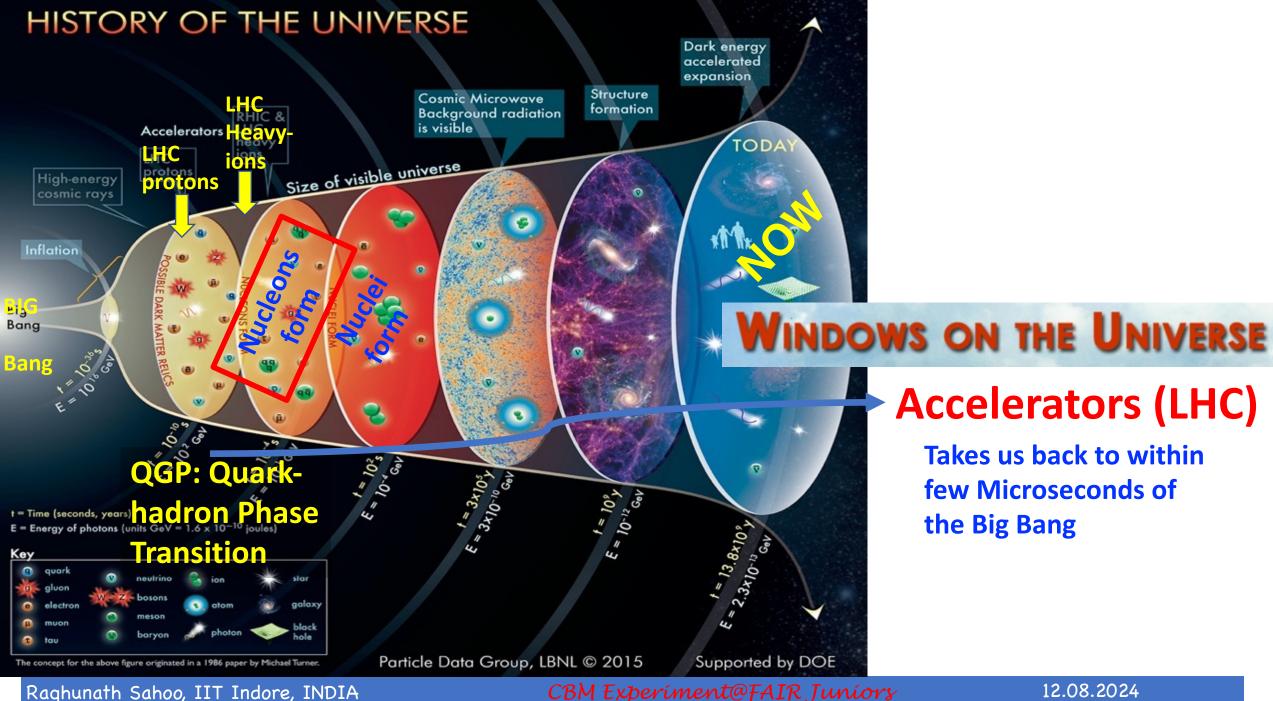


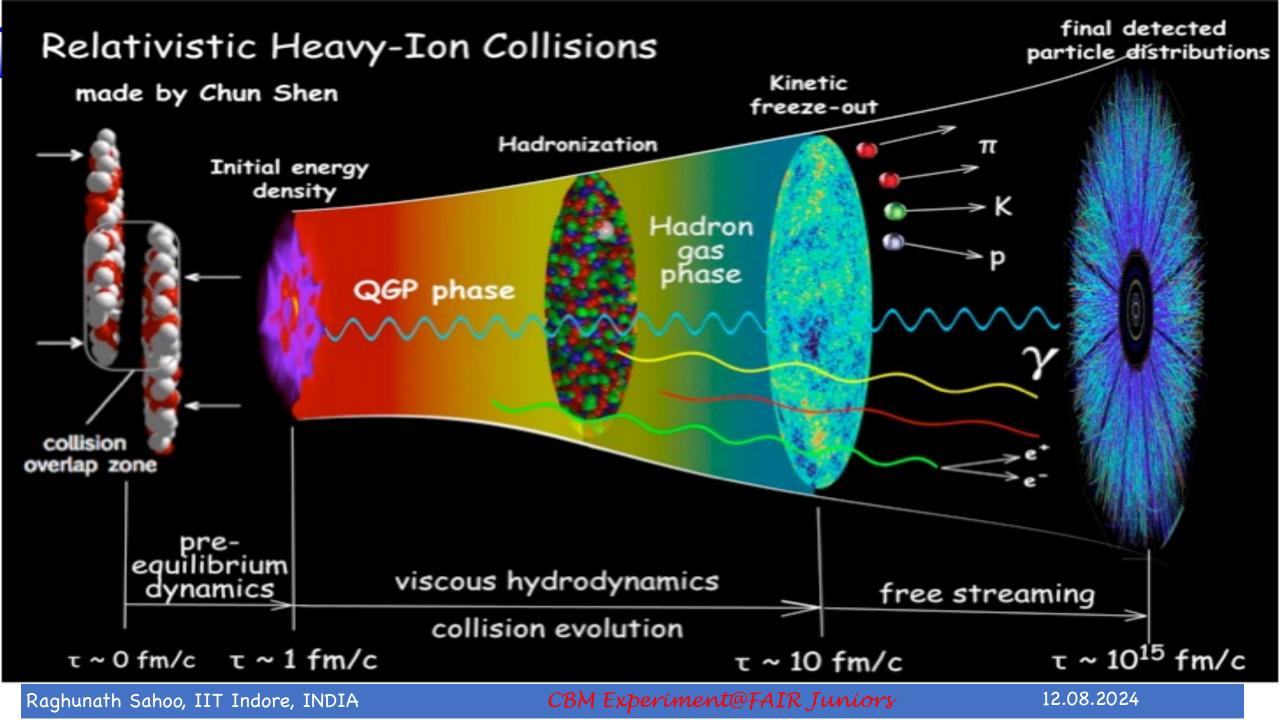
HISTORY OF THE UNIVERSE ... energy accelerated expansion Structure Cosmic Microwave formation Background radiation RHIC & is visible Accelerators TODAY LHC Size of visible universe protons High-energy cosmic rays Inflation <u> Pia</u> Bang 1 = 3x105 E = 10-12 Gay t = Time (seconds, years) $I \approx 13.8 \times 10^{\circ} \text{y}$ E = Energy of photons (units GeV = 1.6×10^{-10} joules) Key Particle Data Group, LBNL © 2015 Supported by DOE CBM Experiment@FAIR Juniors Raghunath Sahoo, IIT Indore, INDIA

Astrophysical Probes

Takes us back to 380,000 years after the Big Bang

WINDOWS ON THE UNIVERSE







How much high is high-energy?

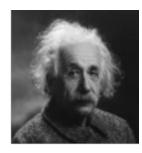
High energies allow us:

To look deeper into Nature (Ε α 1/size), ("powerful microscopes")



de Broglie

To discover new particles with high(er) mass $(E = mc^2)$



Einstein



To study the early universe (E = kT)

Boltzmann

Lesson: To probe the subatomic universe, we need very high energies, which could produce many particles in the final state with very high temperature system.

How do we do that?



Which particles do we detect?

Baryon Summary Table

 $\Lambda_c(2595)^+ 1/2^- ***$

1/2+ ****

* $\Xi(1690)$ *** $A_c(2880)^+$ 5/2+ *** $A_c(2940)^+$ ***

Meson Summary Table

This short table gives the name, the quantum numbers (where known), and the status of baryons in the Review. Only the baryons with 3- or 4-star status are included in the Baryon Summary Table. Due to insufficient data or uncertain interpretation, the other entries in the table are not established baryons. The names with masses are of baryons that decay strongly. The spin-parity J^P (when known) is given with each particle. For the strongly decaying particles, the J^P values are considered to be part of the names.

 $N(1680) \quad 5/2^{+} \quad **** \quad \Delta(1910) \quad 1/2^{+} \quad **** \quad \Sigma(1620) \quad 1/2^{-} \quad * \quad \Xi(2030) \quad \geq \frac{5}{2}? \quad *** \quad \Sigma_{c}(2520) \quad 3/2^{+} \quad ****$

1/2⁺ **** =0

1/2+ **** Ξ-

1/2+ **** Ξ(1530) 3/2+ ****

See also the table of suggested $q\overline{q}$ quark-model assignments in the Quark Model section. . Indicates particles that appear in the preceding Meson Summary Table. We do not regard the other entries as being established.

	LIGHT UNFLAVORED (S = C = B = 0)				STRANGE $(S = \pm 1, C = B = 0)$		CHARMED, STRANGE $(C = S = \pm 1)$		$c\overline{c}$ $f^{G}(J^{PC})$	
	$I^G(J^{PC})$		$I^G(J^{PC})$		$I(J^{\circ})$		$I(J^P)$	• η _c (1S)	0+(0-+)	
• π [±]	1-(0-)	 π₂(1670) 	1-(2-+)	• K±	1/2(0-)	• D _s [±]	0(0-)	 J/ψ(1S) 	0-(1)	
 π⁰ 	1-(0-+)	 φ(1680) 	0-(1)	• K ⁰	1/2(0-)	• D _s *±	$0(?^{?})$	 X_{c0}(1P) 	0+(0++)	
• η	0+(0-+)	 ρ₃ (1690) 	1+(3)	• K ⁰ _S	1/2(0-)	 D_{s0} (2317)[±] 	$0(0^{+})$	 χ_{c1}(1P) 	0+(1++)	
• f ₀ (500)	0+(0++)	 ρ(1700) 	1+(1)	• K ⁰ _L	1/2(0-)	 D_{s1} (2460) ± 	$0(1^{+})$	 h_c(1P) 	?(1 + -)	
 ρ(770) 	1+(1)	$a_2(1700)$	1-(2++)	$K_0^*(800)$	$1/2(0^{+})$	 D_{s1} (2536) ± 	$0(1^{+})$	 Xc2(1P) 	0+(2++)	
• ω(782)	0-(1)	 f₀(1710) 	0+(0++)	 K*(892) 	1/2(1-)	 D_{s2} (2573) 	0(??)	 η_c(2S) 	0+(0-+)	
 η'(958) 	0+(0-+)	$\eta(1760)$	0+(0-+)	 K₁(1270) 	$1/2(1^+)$	 D_{s1} (2700)[±] 	0(1-)	 ψ(2S) 	0-(1)	
• f ₀ (980)	0+(0++)	 π(1800) 	1-(0-+)	 K₁(1400) 	$1/2(1^+)$	$D_{sJ}^{*}(2860)^{\pm}$	0(??)	• ψ(3770)	0-(1)	
 a₀(980) 	1-(0++)	$f_2(1810)$	0+(2++)	 K*(1410) 	1/2(1-)	$D_{sJ}(3040)^{\pm}$	0(??)	• X(3872)	0+(1++) 0+(0++)	
 φ(1020) 	0-(1)	X(1835)	??(?-+)	 K₀*(1430) 	1/2(0+)			• χ _{c0} (2P)	0.(0,.)	

There are hundreds of particles ...

7/2- ****

5/2+ *** 3/2 * 9/2+ ***

A(2100)

 $\Lambda(2350)$

* $\Delta(1920)$ 3/2+ *** $\Sigma(1660)$ 1/2+ *** $\Xi(2120)$

N(1520) 3/2 **** $\Delta(1700)$ 3/2 **** $\Sigma(1385)$ 3/2 **** $\Xi(1620)$

N(1675) 5/2- **** $\Delta(1905)$ 5/2+ **** $\Sigma(1580)$ 3/2- * $\Xi(1950)$

 $\Sigma(1480)$

 $1/2^{+}$ **** $\Delta(1232)$ $3/2^{+}$ **** Σ^{+}

 $1/2^{+}$ **** $\Delta(1600)$ $3/2^{+}$ *** Σ^{0}

N(1440) 1/2+ **** $\Delta(1620)$ 1/2- **** Σ^-

N(1650) 1/2 **** $\Delta(1900)$ 1/2 ** $\Sigma(1560)$

N(1535) $1/2^-$ **** $\Delta(1750)$ $1/2^+$ *

however most of them are so short-lived that we'll never see them directly in our detectors.

*** \(\Sum_c(2455) \) 1/2+ ****

* \(\Sum_c(2800)\)

Track length: $I_{\text{track}} = v\tau = c\beta\gamma\tau_0$ with τ_0 being the lifetime at rest.

Only if l_{track} (at GeV scale) ≥ 1 mm, we have a chance to measure them.

81/ 2										
N(2190)	1/2	71	4/4		Z (2230)					
N(2220)	9/2+ ****	$\Lambda(1405)$	1/2-	****	$\Sigma(2455)$	**		=+ ec		*
N(2250)	9/2 ****	$\Lambda(1520)$	3/2-	****	$\Sigma(2620)$	**		-ce		
N(2300)	1/2+ **	$\Lambda(1600)$	1/2+	***	$\Sigma(3000)$	*		A0,	1/2+	***
N(2570)	5/2 **	$\Lambda(1670)$	1/2	****	$\Sigma(3170)$	*		$A_b(5912)^0$	1/2-	***
N(2600)	11/2 ***	A(1690)	3/2-	****				$A_h(5920)^0$	3/2-	***
N(2700)	13/2+ **	$\Lambda(1800)$	1/2-	***				Σ_h		***
		$\Lambda(1810)$	1/2+	***				Σ_b^*	3/2+	***
		$\Lambda(1820)$	5/2 ⁺	****				\equiv_b^0 , \equiv_b^-	1/2+	***
		A(1830)	5/2-	****				$\Xi_b(5945)^0$	3/2+	***
		$\Lambda(1890)$	$3/2^{+}$	****				Ω_b^-	1/2+	***
		A(2000)		*				32 b	1/2	
1		4/00000	- 10+		I		I	I		

• ρ(1450) 1+(1) • η(1475) 0+(0 - +) • δ(1500) 0+(0 + +) • δ(1510) 0+(1 + +) • δ(1510) 0+(2 + +) • δ(1565) 0+(2 + +)	$f_0(2200)$ $0^+(0^{++})$ $f_J(2220)$ $0^+(2^{++})$ $0^+(2^{++})$ $0^+(0^{-+})$ $0^+(0^{-+})$ $0^+(0^{-+})$ $0^+(0^{-+})$ $0^+(0^{++})$	$K_4(2500)$ 1/2(4 ⁻) K(3100) ? ² (? [?] ?) CHARMED (C = ±1)	• B_s^* 0(1 ⁻) • $B_{s1}(5830)^0$ 0(1 ⁺) • $B_{s2}^*(5840)^0$ 0(2 ⁺) • $B_{sJ}^*(5850)$?(? ²)	$\eta_b(1S)$ 0+(0-+) • $\Upsilon(1S)$ 0-(1) • $\chi_{b0}(1P)$ 0+(0++) • $\chi_{b1}(1P)$ 0+(1++) • $h_b(1P)$? [?] (1+-) • $\chi_{b2}(1P)$ 0+(2++)
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	D± 1/2(0 [¬]) D ⁰ 1/2(0 [¬]) D*(2007) ⁰ 1/2(1 [¬]) D*(2010)± 1/2(1 [¬]) D* ₀ (2400) ⁰ 1/2(0 ⁺) D* ₀ (2400)± 1/2(0 ⁺) D* ₁ (2420) ⁰ 1/2(1 ⁺)	$(B = C = \pm 1)$ $\bullet B_c^{\pm} \qquad 0(0^-)$	$\eta_b(2S)$ 0+(0-+) • $\Upsilon(2S)$ 0-(1) • $\Upsilon(1D)$ 0-(2) • $\Upsilon(1D)$ 0-(2) • $\Upsilon(1D)$ 0-(10-+) • $\Upsilon(1D)$ 0-(10-+)
• ω_3 (1670) 0 - (3)	OTHER LIGHT Further States	$\begin{array}{cccccccccccccccccccccccccccccccccccc$		• 7(3S) 0 - (1) • \(\lambda_0(3P)\) ? (??+) • 7(4S) 0 - (1) • \(\lambda(010)^\pi ?^+(1^+)\) • \(\cap(10860)\) 0 - (1) • \(\cap(11020)\) 0 - (1)

Leptons

ν_e

ν_τ

Gauge bosons

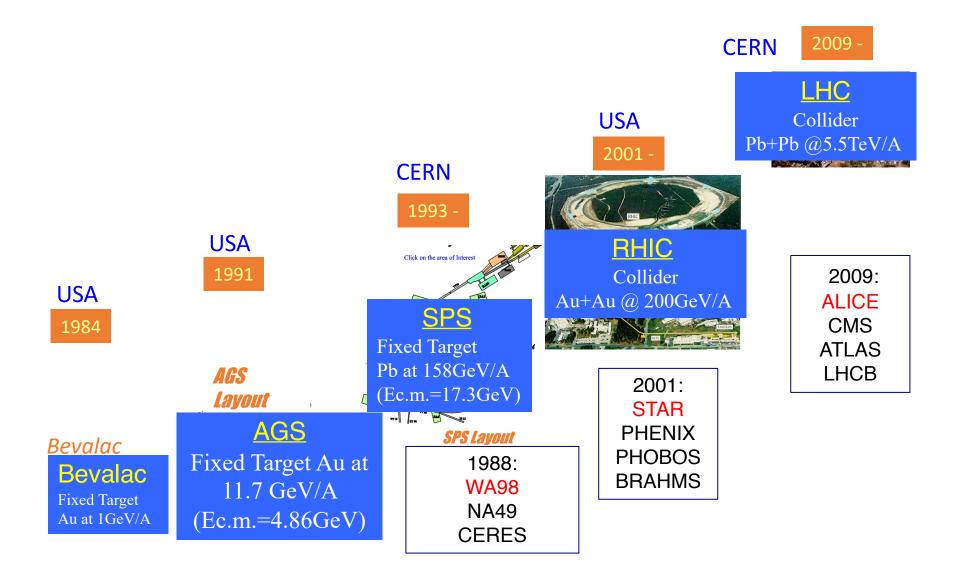
- W^{+/-}
- Z
- H

Which are left then? These 8 particles (and their antiparticles).

	γ	р	n	e [±]	μ [±]	π^{\pm}	K [±]	$\mathbf{K_0} (K_S/K_L)$
τ_0	∞	∞	∞	8	2.2μs	26 ns	12 ns	89 ps / 51 ns
I _{track (p=1GeV)}	∞	8	8	8	6.1 km	5.5 m	6.4 m	5 cm / 27.5 m



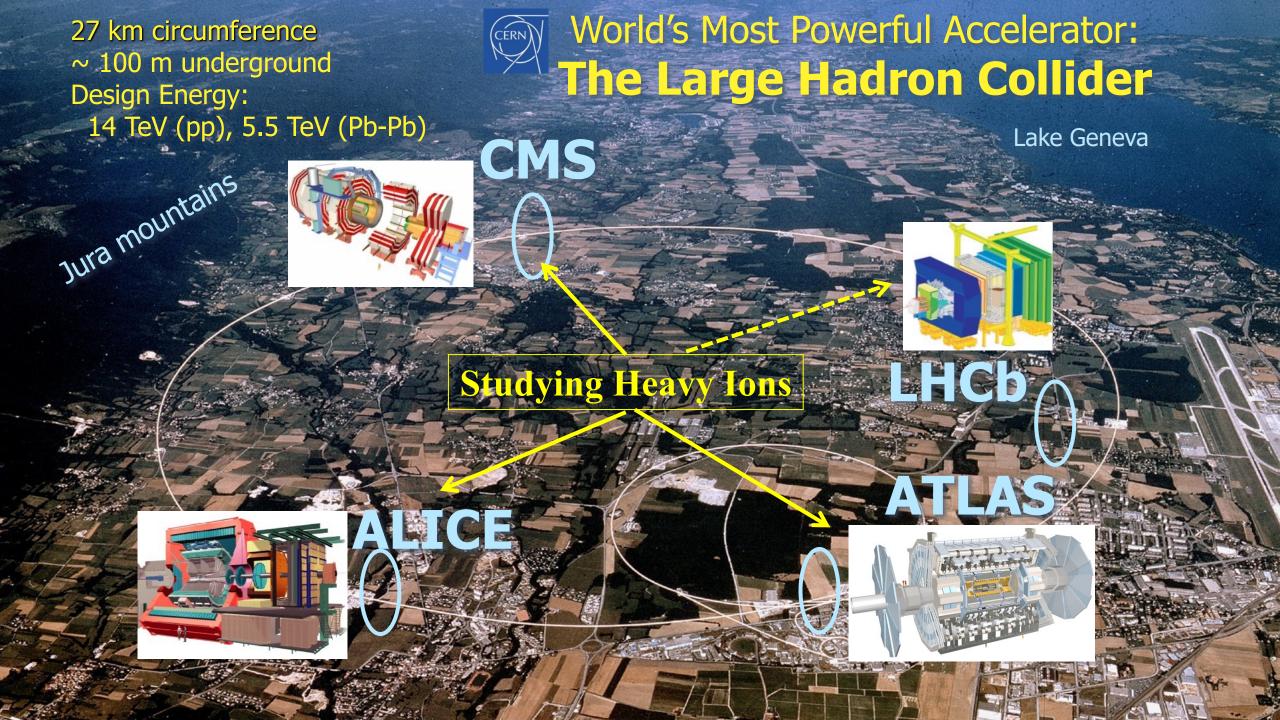
Accelerator facilities for heavy-ions











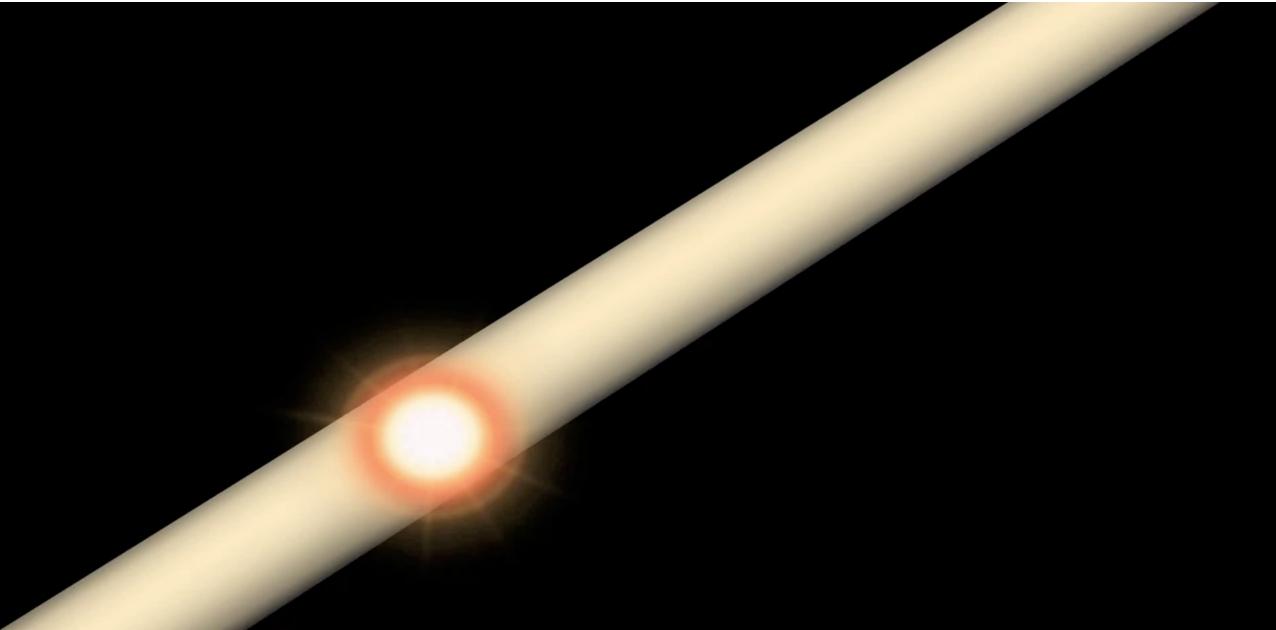


27km tunnel:

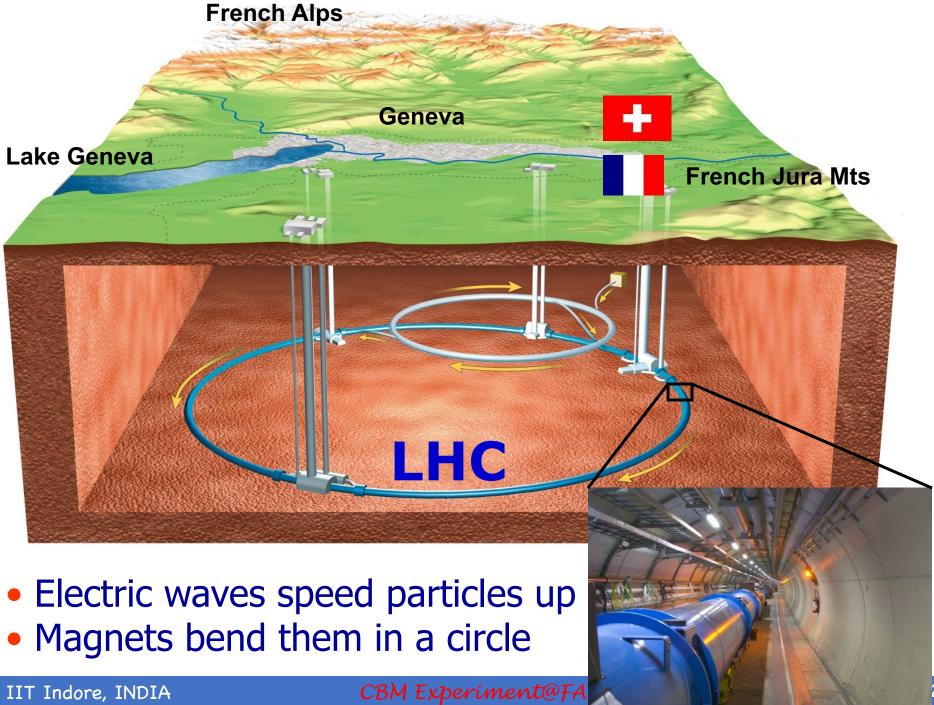
- 50-150m below ground
- Two beams circulating in opposite directions
- Total of 9300 magnets: beams controlled by 1800 superconducting magnets (up to 8T)

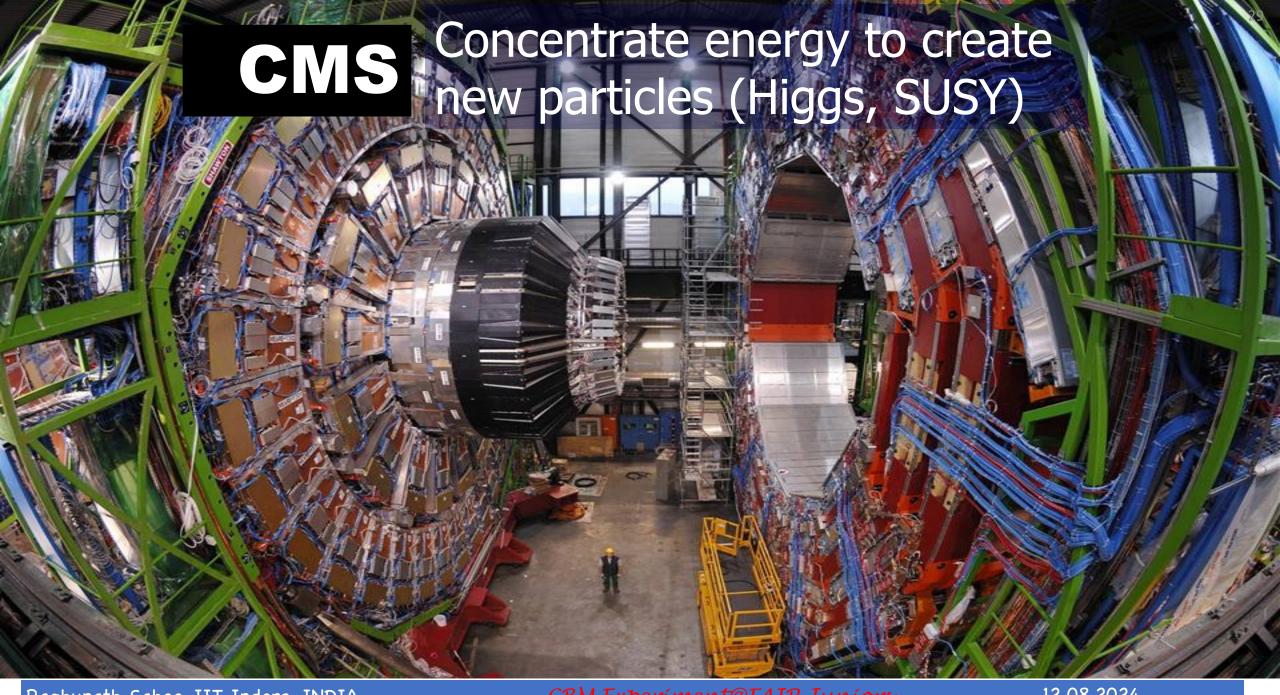


The Large Hadron Collider

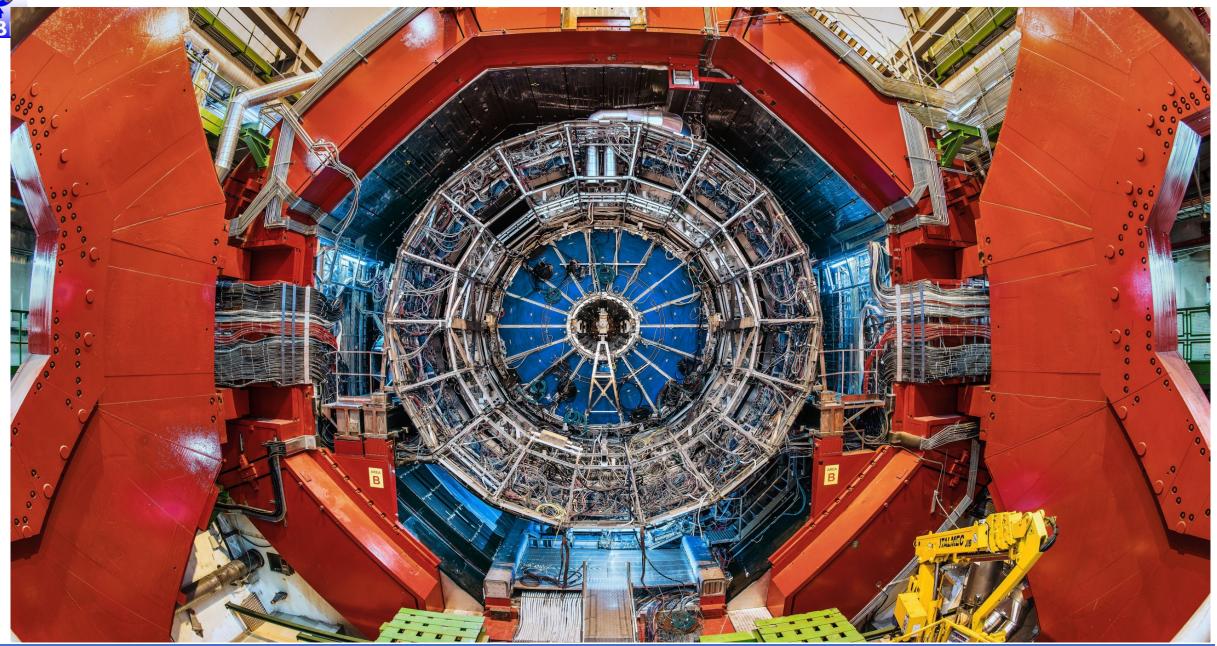






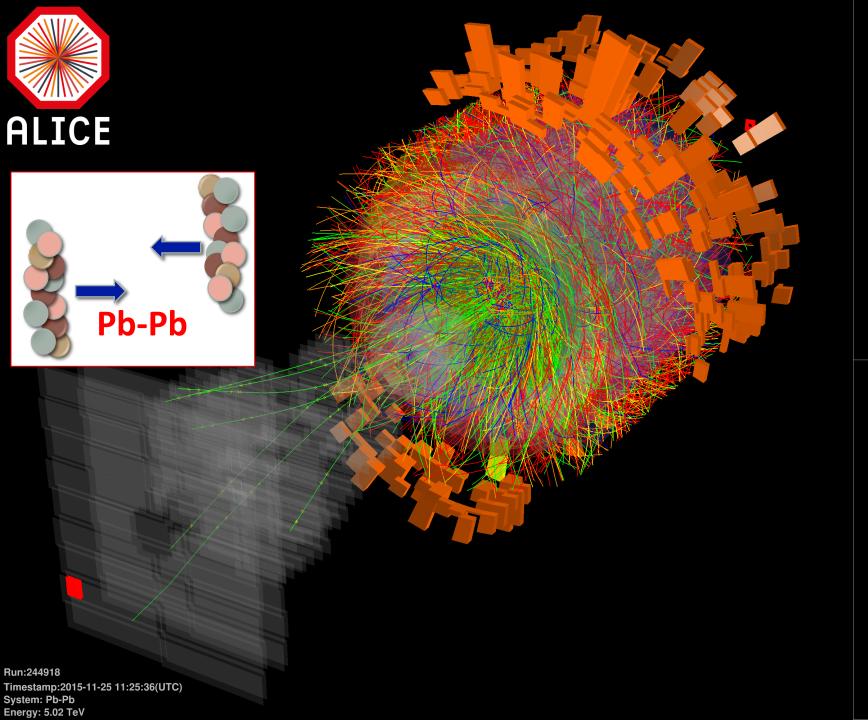


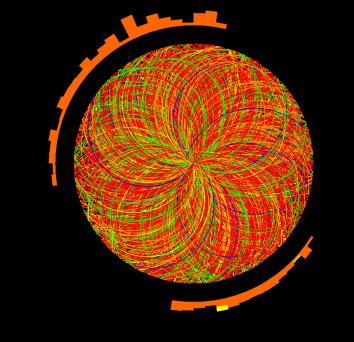
ALICE at Point-2 of the LHC

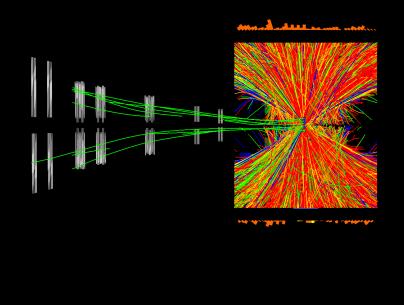








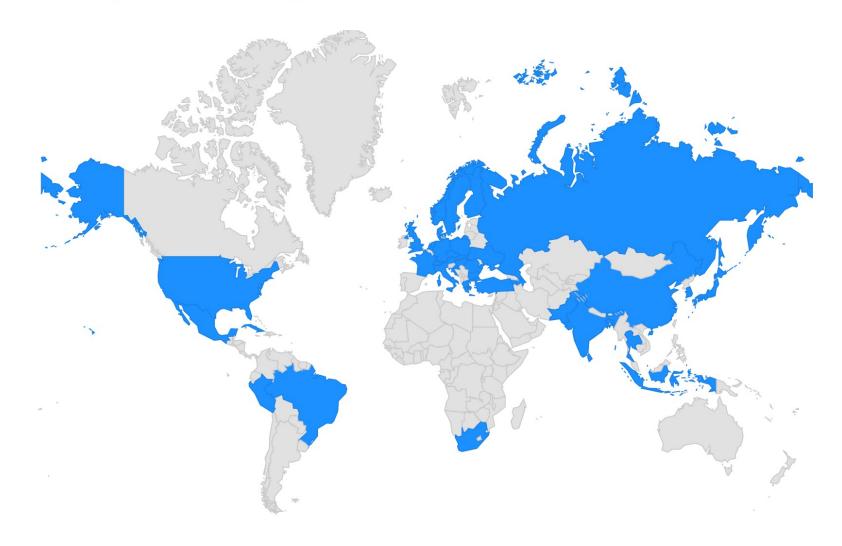




ALICE Collaboration

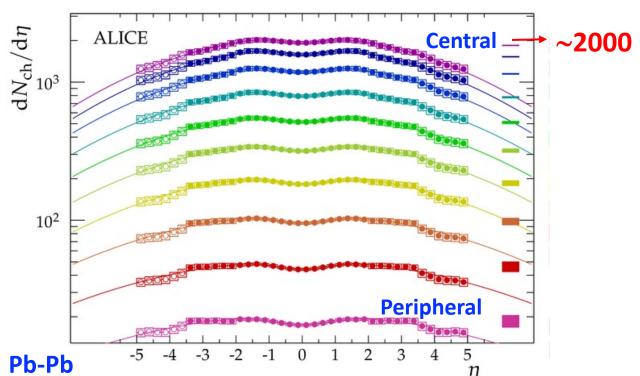


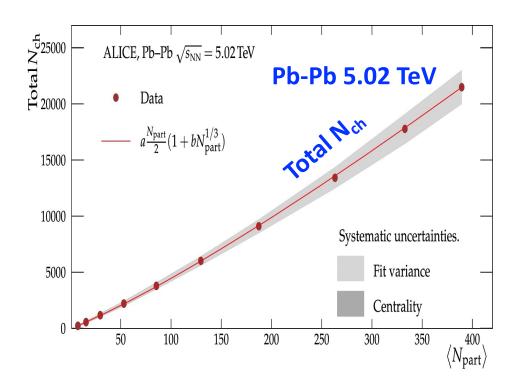
40 countries, 171 institutes, 1996 members





Charged particle multiplicity





Number of charged particles in one collision:

Central collisions: 21400 ± 1300

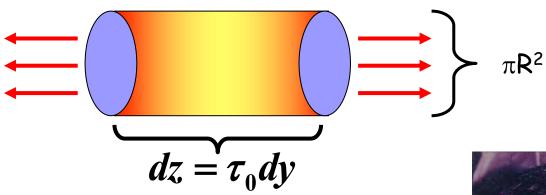
• Peripheral collisions: 230 ± 38

Phys.Lett. B 772 (2017) 567577 Phys. Rev. Lett. 116 (2016) 222302

VERY LARGE NUMBER OF PRODUCED PARTICLES

Particle density & Energy density

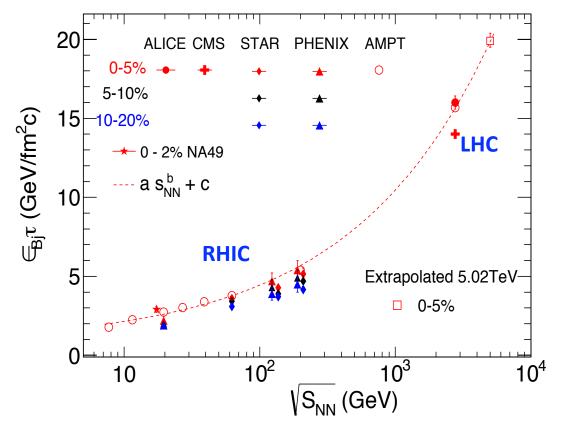
J. D. Bjorken, Phys. Rev. D 27, 140 (1983).



$$\varepsilon_{Bj}(\tau) = \frac{1}{\pi R^2 \tau} \frac{dE_T}{dy}$$

$$\approx \frac{1}{\pi R^2 \tau} < m_T > \frac{3}{2} \frac{dN_{ch}}{d\eta}$$





 $\varepsilon.\tau \sim 16 \text{ GeV/fm}^2\text{c}$

S. Basu et al. PRC 93 (2016) 064902

R. Sahoo et al. Adv. in HEP, Vol. 2015

LARGEST ENERGY DENSITIES EVER ACHIEVED

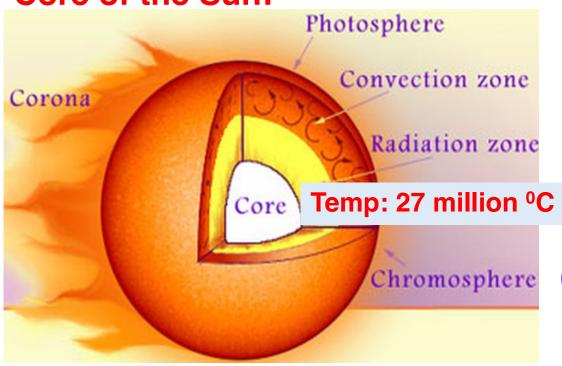


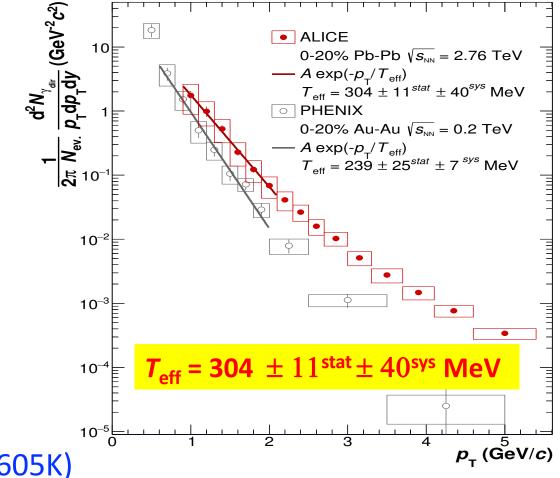
Photon Spectra and QGP temperature

Phys. Lett. B 754 (2016) 235-248

- Photons do not interact via the nuclear force → transparent to the medium
- Photons are emitted in all stages and are unaffected by the medium.

Core of the Sun:





(1eV=11605K)

 $T_{\rm eff}$ = 3,527,920 million deg

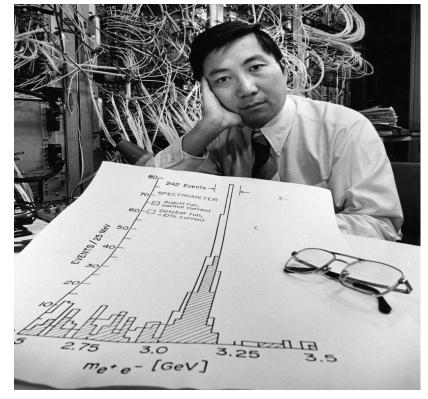
LARGEST EVER TEMPERATURE REACHED IN THE LAB

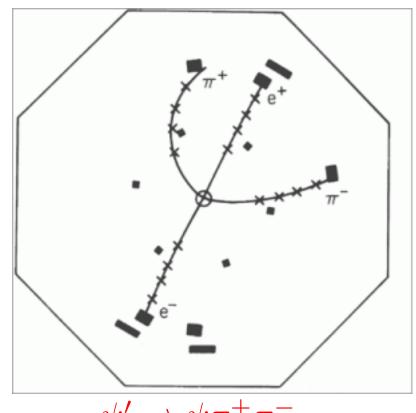


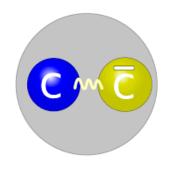
Charmonia, Machine Learning and QGP



The November Revolution: 1974







Crosses: spark chamber hits Dark recatangles: ToF counters





1.14. Using four quarks (u,d, s, and c), construct a table of all the possible baryon species. How many combinations carry a charm of +1? How many carry charm +2, and +3?



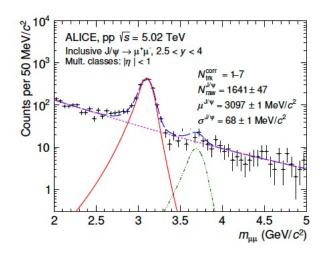
Burton Richter@SLAC: e+e-Sam Ting@BNL: pp

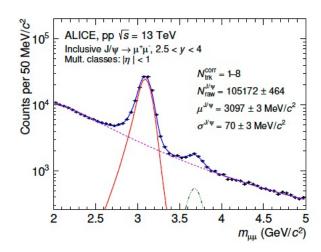


The Nobel Prize in Physics 1976



Our discovery (measurement) of J/ψ





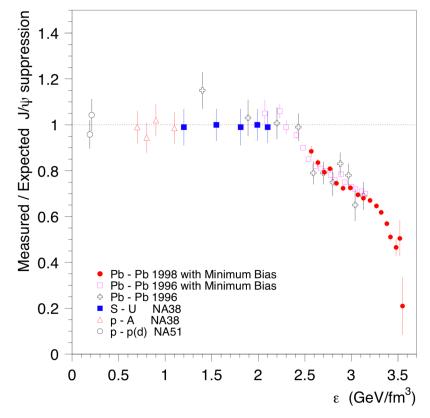


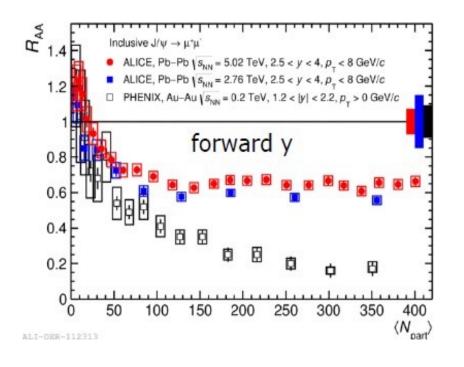
D. Thakur, R. Sahoo et al. ALICE Collaboration JHEP06(2022)015

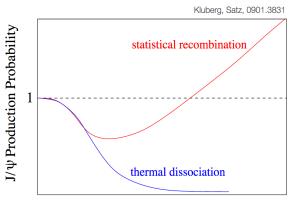
https://alice-notes.web.cern.ch/node/734, ANA-734, D. Thakur and R. Sahoo https://alice-publications.web.cern.ch/node/5122, D. Thakur, R. Sahoo et al.



Experimental observation of J/ψ Suppression







Energy Density $R_{\text{AA}}(p_{\text{T}}) = \frac{d N_{\text{AA}}/dp_{\text{T}}}{\langle T_{\text{AA}} \rangle \times d\sigma_{\text{pp}}/dp_{\text{T}}}$

RAA < 1 : Suppression of yield due to presence of medium

RAA > 1: No suppression and hence no medium

(Physics Letters B 766 (2017) 212-224)

http://alice.web.cern.ch/content/mystery-jpsi

CERN SPS (NA50): observed J/ψ suppression as a function of energy density for various collision species. Note that the critical energy density for a partonic medium is 1 GeV/fm³.

- Big question: use of pp as reference at LHC
- Need : pp should be investigated properly

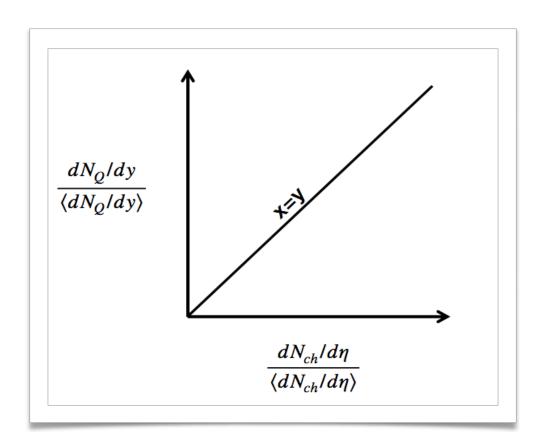


Study of J/ψ production in pp collisions

□ Investigate charmonium production in pp as a function of multiplicity

Observables:

- X axis: relative chargedparticle density
- Y axis: relative quarkonium yield (w.r.t. minimum bias)

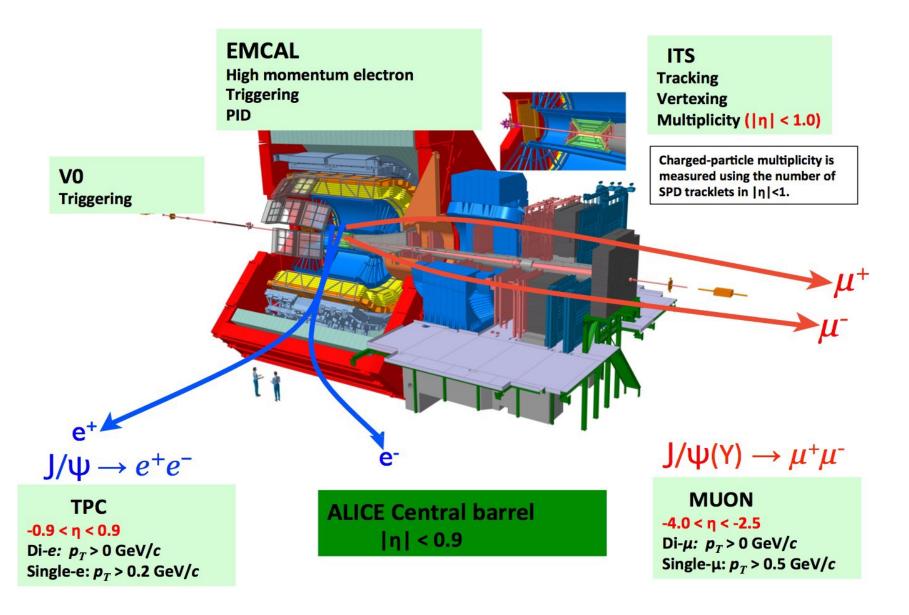


Advantages:

- Several correction factors cancel out in relative yields
- Easy to compare the results across collision systems and energies

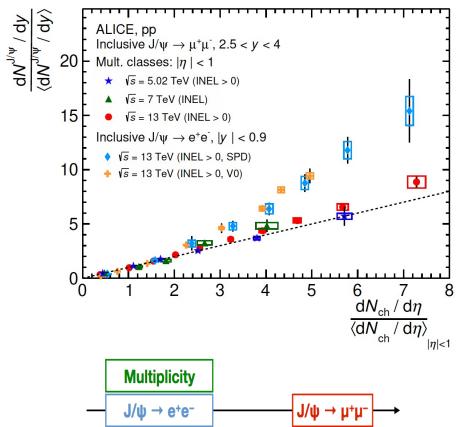


Study of J/ψ production in ALICE





Study of J/ψ production in ALICE pp collisions



ALICE: JHEP06 (2022) 015

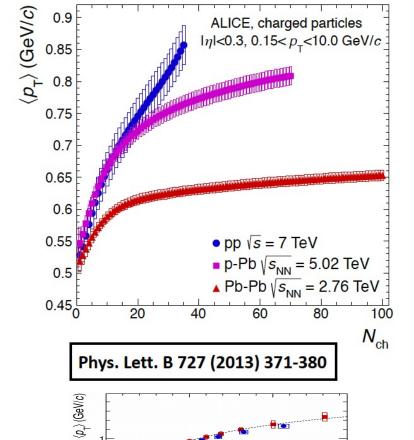
https://alice-notes.web.cern.ch/node/734, ANA-734, D. Thakur and R. Sahoo https://alice-publications.web.cern.ch/node/5122, D. Thakur, R. Sahoo et al.

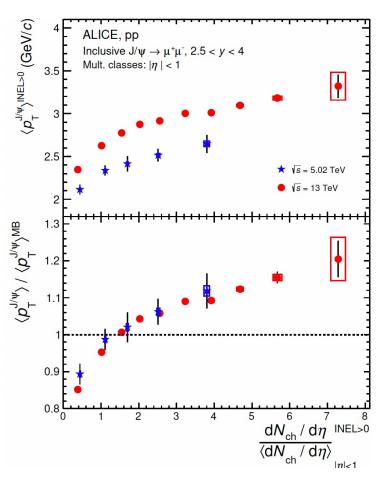
- \blacktriangleright Relative J/ Ψ yield at midrapidity is compared to the forward rapidity yield as a function of midrapidity relative charged-particle multiplicity
- Midrapidity yields exhibit faster than linear increase
- The results using midrapidity multiplicity selection based on the SPD detector ($|\eta| < 1$) and forward-rapidity multiplicity selection based on the VO detector ($-3.7 < \eta < -1.7$ and $2.8 < \eta < 5.1$) are found to be compatible within the uncertainties
- Therefore, the different trends in the multiplicity dependence of the J/Ψ production observed at midrapidity and forward rapidity are not due to a possible auto-correlation bias



Data/Fit

$\langle p_T \rangle$ of J/ ψ vs. multiplicity in pp collisions



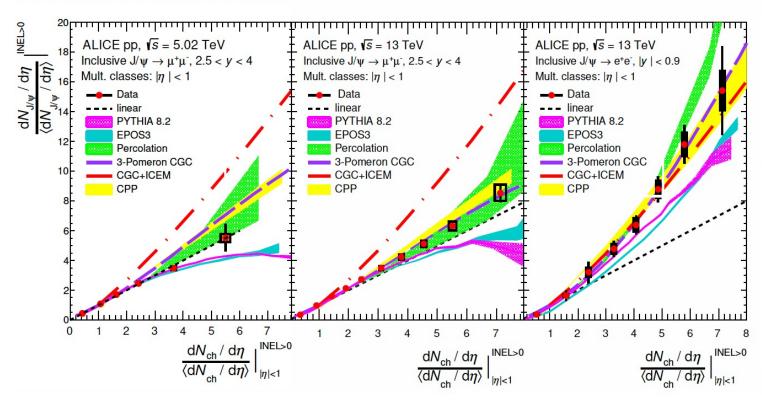


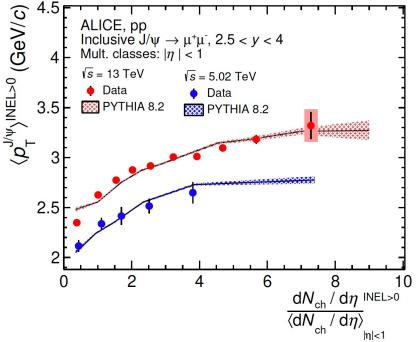
- $\langle p_T \rangle$ of J/ ψ increases with increasing multiplicity with a little saturation towards higher multiplicities
 - collectivity? (Phys. Lett. B 727 (2013) 371-380)
- ❖ Possible explanation of MPI prospect: the high multiplicity events are produced by MPIs and in the absence of CR, incoherent superposition of MPIs would lead to constant $\langle p_T \rangle$ at high multiplicity.
- \Leftrightarrow $\langle p_T \rangle$ of J/ψ vs multiplicity trend is energy-independent
- The multiplicity dependent trend is the same for light-flavor and heavy-flavor particles

https://alice-notes.web.cern.ch/node/734, ANA-734, D. Thakur and R. Sahoo https://alice-publications.web.cern.ch/node/5122, D. Tahkur, R. Sahoo et al.

 $\langle dN_{ch}/d\eta \rangle$

Theoretical model comparison: J/ψ production





None of the theoretical models give a complete picture of the J/ψ prodution dynamics in pp collisions

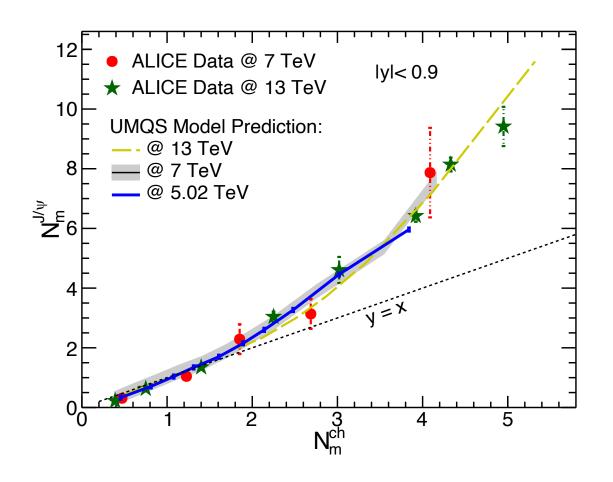
PYTHIA8 with CR gives a reasonable description of the multiplicity dependence of J/ψ <p_T> in pp collisions

https://alice-notes.web.cern.ch/node/734, ANA-734, D. Thakur and R. Sahoo https://alice-publications.web.cern.ch/node/5122, D. Tahkur, R. Sahoo et al.

A lot of room for theory/phenomenology!



UMQS model explains ALICE J/ψ production

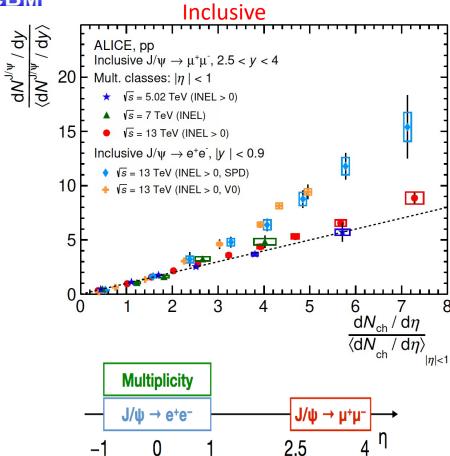


- J/ψ self-normalized yield as a function of selfnormalized multiplicity follows a scaling across collision energies.
- ✓ Unified Model of Quarkonia Suppression (UMQS) model which incorporates the suppression of J/Ψ through color screening, gluonic dissociation, and collision damping and regeneration of charmonium due to correlated c cbar pairs.

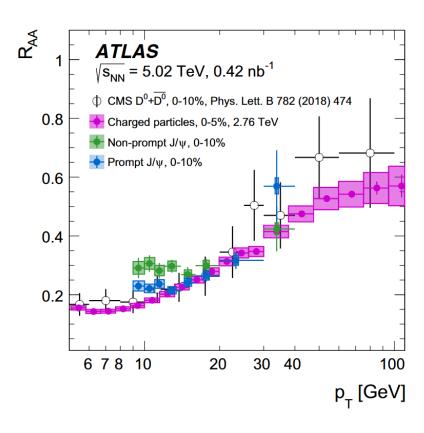
C.R. Singh, S. Deb, R. Sahoo, J. Alam, Eur. Phys. J. C, 82, 542 (2022)



Need of ML: Separating prompt from nonprompt





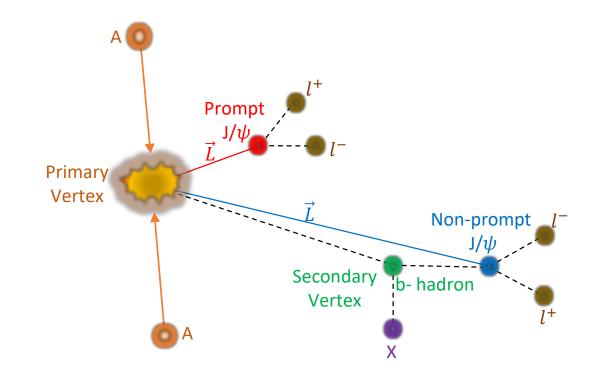


Nuclear modification factor for prompt and non-prompt J/ψ are different.



ML: Separating prompt and non-prompt J/ ψ @LHC

- J/ ψ (3.096 GeV/ c^2)
- In experiments, $J/\psi \rightarrow \mu^+ + \mu^-$ or $J/\psi \rightarrow e^+ + e^-$
- Prompt Production: Direct production/ decay of heavier charmonia states
- Non-prompt Production: Products of beauty hadron weak decays
- Prompt and non-prompt J/ ψ are topologically different



S. Prasad, N. Mallick and R. Sahoo, *Phys. Rev. D 109, 014005 (2024)*



ML to separate prompt and non-prompt J/ ψ @LHC

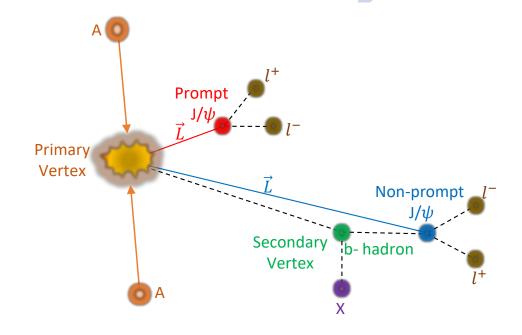
Simulating pp collisions at $\sqrt{s} = 13$ TeV using PYTHIA8

Training machine-learning models

Using the machine-learning models to predict prompt and non-prompt yields at different energies

Results

- PYTHIA8 (4C-tune) 20 billion minimum bias events for pp $\sqrt{s} = 13 \, \text{TeV}$
- The coordinates of the primary vertex are randomised following a Gaussian distribution: experiment-like scenario
- J/ $\psi \to \mu^+ + \mu^-$ channel is used to reconstruct invariant mass $(m_{\mu\mu})$, transverse momentum $(p_{T,\mu\mu})$, pseudorapidity $(\eta_{\mu\mu})$ and rapidity $(y_{\mu\mu})$ of the dimuons
- Pseudoproper decay length (c au) of the reconstructed dimuon pairs along with $m_{\mu\mu}$, $p_{T,\mu\mu}$, and $\eta_{\mu\mu}$ are taken as inputs



$$c au = rac{c \; m_{\mathrm{J/\psi}} \; \vec{L} \; . \; \vec{p_{\mathrm{T}}}}{|\vec{p_{\mathrm{T}}}|^2}$$

Training Validation



ML parameters to separate prompt and non-prompt J/ ψ @LHC

- Background : Prompt : Non-prompt = 20 : 10 : 1
- Classification models required to be trained on similar number of training instances → oversampling of data is done
- Dataset for Training: Testing: Validation = 81:10:9
- Parameters are chosen through a grid search method (Making an array of all possible parameters and training to find the parameter values for minimum loss)
- Grid Search Method to find best parameters

	XGB	LGBM
Learning rate	0.3	0.1
Sub-sample	1.0	1.0
No. of trees	60	60
Maximum depth	3	3
Objective	softmax	softmax
Metric	mlogloss	$\boxed{multilogloss}$

S. Prasad, N. Mallick and R. Sahoo, *Phys. Rev. D* 109, 014005 (2024)

 0.6°

0.5

0.4 SSOJ

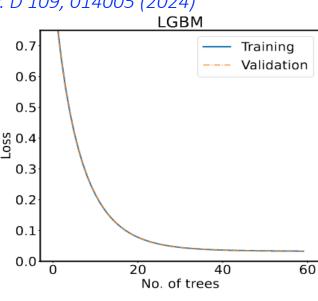
0.3

0.2

0.1

- Loss saturates around 25 and 45
 trees for XGB and LGBM
- Training and validation curves are on top of each other → No overfitting/underfitting

LGBM: Light Gradient Boosting Machine XGB: Extreme Gradient Boosting



20

No. of trees

40

XGB



Prompt

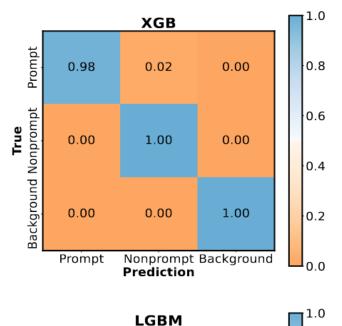
True Background Nonprompt 0.98

0.00

0.00

Prompt

ML Model performance to separate prompt and non-prompt J/ ψ @LHC



0.02

1.00

0.00

Prediction

0.00

0.00

1.00

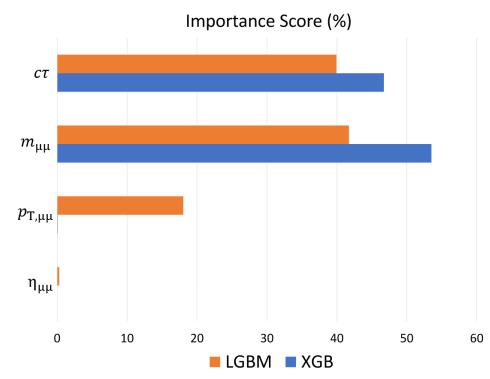
0.8

0.6

0.4

0.2

- Confusion Matrix talks about the mispredictions given by the model for each class
- ullet Both XGB and LGBM perfectly separate the inclusive J/ ψ from the uncorrelated background pairs
- Both models mispredict 2% of prompt J/ ψ as the non-prompt \rightarrow Raises non-prompt yield

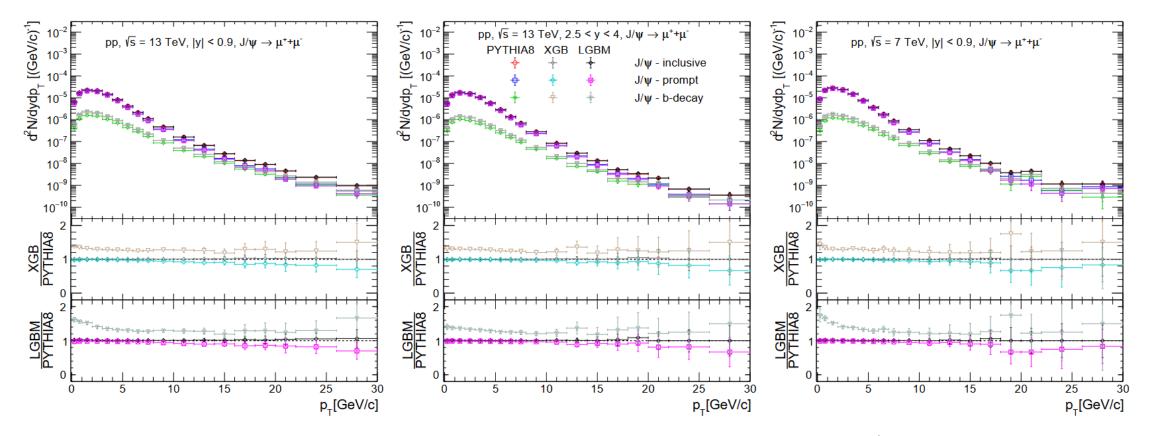


- The importance score tells how important a feature is for the decision-making of the models
- The importance score of the invariant mass of dimuons is highest for both models
- $c\tau$ contributes to decision making of the models significantly

Nonprompt Background



ML Results: Transverse momentum spectra

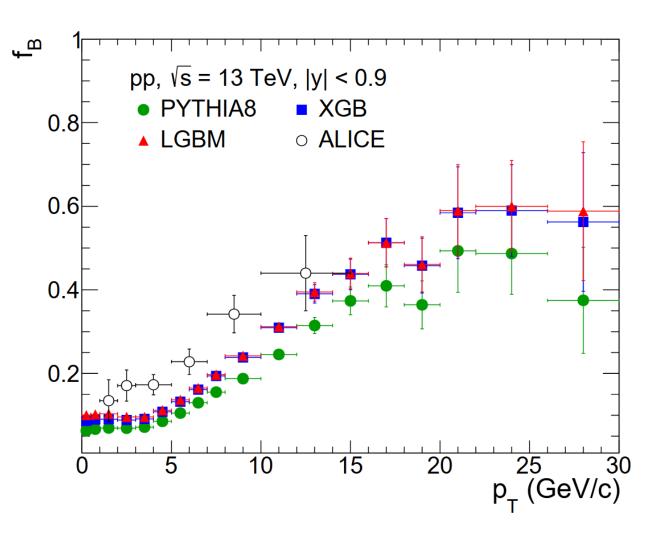


- Both XGB and LGBM give accurate predictions for $p_{\rm T}$ -spectra for inclusive and prompt-J/ ψ both in mid and forward rapidity in pp collisions at $\sqrt{s}=13$ TeV and 7 TeV
- ullet The ML models overpredict the non-prompt J/ ψ throughout the $p_{
 m T}$ spectra for both the collision energy and rapidity
 - → Expected from the confusion matrix

S. Prasad, N. Mallick and R. Sahoo, *Phys. Rev. D* 109, 014005 (2024)



ML Results: Fraction of non-prompt J/ ψ yield



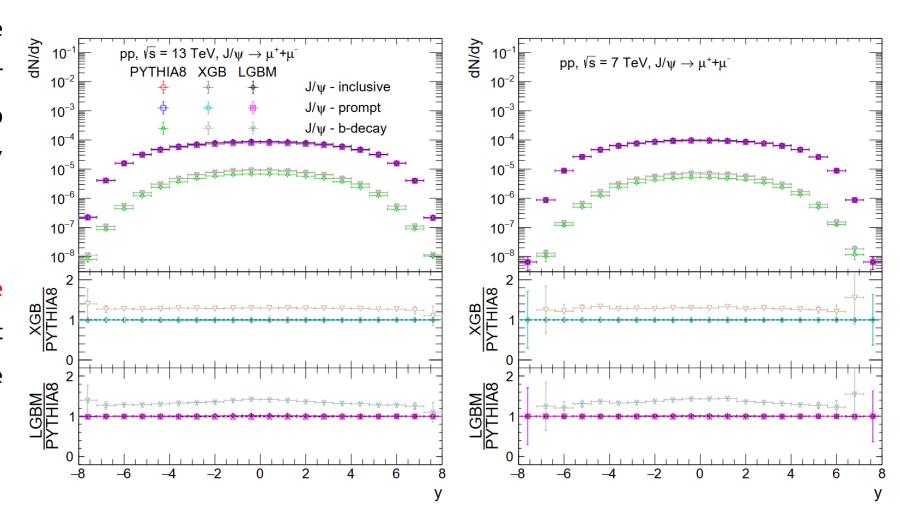
$$egin{aligned} \sigma_{
m nonprompt\ J/\psi} &= f_{
m B} \cdot \sigma_{
m J/\psi}, \ \sigma_{
m prompt\ J/\psi} &= (1-f_{
m B}) \cdot \sigma_{
m J/\psi}. \end{aligned}$$

- $f_{\rm B}$ is the fraction of the non-prompt production (b-hadron decays)
- $f_{\rm B}$ increases with increase in $p_{\rm T}$ ightharpoonup The b-hadron production is favoured towards higher $p_{\rm T}$
- PYTHIA8 underestimates the experimental data following a similar trend
- Both XGB and LGBM overestimate PYTHIA8
- As this method does not require fitting, it can be used in both low and high statistics without affecting its efficiency

S. Prasad, N. Mallick and R. Sahoo, *Phys. Rev. D* 109, 014005 (2024)



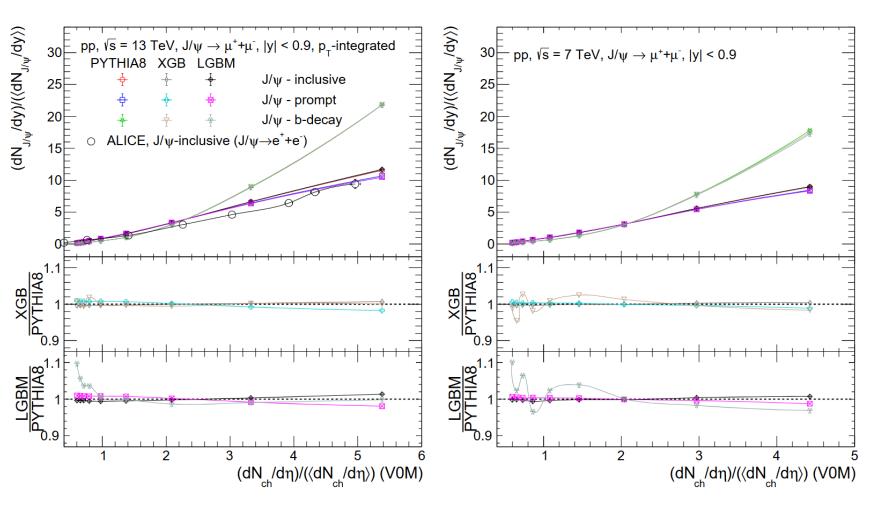
- Both XGB and LGBM give accurate predictions for rapidity spectra for inclusive and prompt-J/ ψ in pp collisions at $\sqrt{s}=13$ TeV and 7 TeV
- The ML models overpredict the non-prompt J/ψ throughout rapidity region for both the collision energies



S. Prasad, N. Mallick and R. Sahoo, *Phys. Rev. D* 109, 014005 (2024)



ML Results: Normalized J/ ψ Yields



- The normalised yield for inclusive J/ψ from PYTHIA8 matches qualitatively with the ALICE results
- Both XGB and LGBM reproduce the PYTHIA8 results very precisely for inclusive and prompt ${\rm J}/\psi$
- The predictions for non-prompt J/ ψ from both XGB and LGBM matches PYTHIA8 findings within 10% uncertainty

12.08.2024

S. Prasad, N. Mallick and R. Sahoo, *Phys. Rev. D* 109, 014005 (2024)



ML Results J/ ψ : Summary

- We have used BDT based ML models such as XGBoost and LGBM to segregate the prompt, non-prompt and inclusive J/ ψ production in pp collisions at $\sqrt{s}=13$ TeV
- The models the parameters such as, pseudo-proper decay length $(c\tau)$, invariant mass $(m_{\mu\mu})$, transverse momentum $(p_{T,\mu\mu})$, pseudorapidity $(\eta_{\mu\mu})$ of the dimuons as the input, which are accessible in the experiments
- The model almost achieves 99% overall accuracy
- The estimations for the prompt and inclusive J/ ψ from the ML models match with the PYTHIA8 for the inclusive and non-prompt J/ ψ
- Using these models, track label identification is possible, and it avoids the necessity of fit for the identification



Open Charms: D⁰

PHYSICAL REVIEW D VOL..XX, 000000 (XXXX)

8

22 23 24

Machine learning-based study of open-charm hadrons in proton-proton collisions at the Large Hadron Collider

Kangkan Goswami, Suraj Prasad, Neelkamal Mallick, and Raghunath Sahoo, Department of Physics, Indian Institute of Technology Indore, Simrol, Indore 453552, India

Gagan B. Mohanty

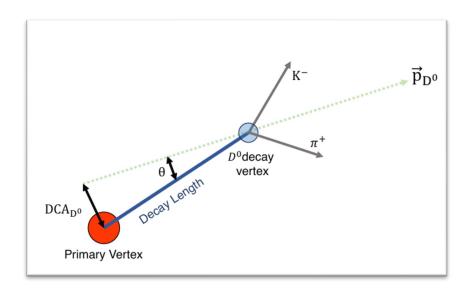
Tata Institute of Fundamental Research, Homi Bhabha Road, Mumbai 400005, India

(Received 16 April 2024; accepted 2 July 2024)

In proton-proton and heavy-ion collisions, the study of charm hadrons plays a pivotal role in understanding the QCD medium and provides an undisputed testing ground for the theory of strong interaction, as they are mostly produced in the early stages of collisions via hard partonic interactions. The lightest open charm, D^0 meson $(c\bar{u})$, can originate from two separate sources. The prompt D^0 originates from either direct charm production or the decay of excited open charm states, while the nonprompt stems from the decay of beauty hadrons. In this paper, using different machine learning (ML) algorithms such as XGBoost, CatBoost, and Random Forest, an attempt has been made to segregate the prompt and nonprompt production modes of the D^0 meson signal from its background. The ML models are trained using the invariant mass through its hadronic decay channel, i.e., $D^0 \to \pi^+ K^-$, pseudoproper time, pseudoproper decay length, and distance of closest approach of D^0 meson, using PYTHIA8 simulated pp collisions at $\sqrt{s}=13$ TeV. The ML models used in this analysis are found to retain the pseudorapidity, transverse momentum, and collision energy dependence. In addition, we report the ratio of nonprompt to prompt D^0 yield, the self-normalized yield of prompt and nonprompt D^0 , and explore the charmonium, J/ψ to open charm, D^0 yield ratio as a function of transverse momenta and normalized multiplicity. The observables studied in this paper are well predicted by all the ML models compared to the simulation.

Phys. Rev. D (2024: In Press)

arXiv: 2404.09839

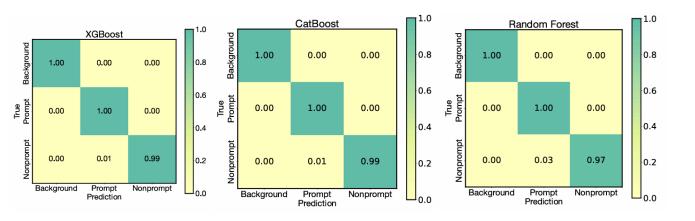


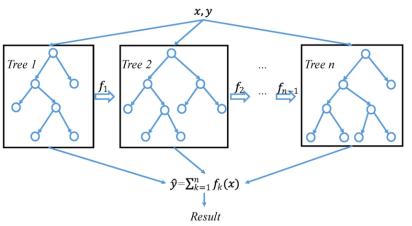
DOI:



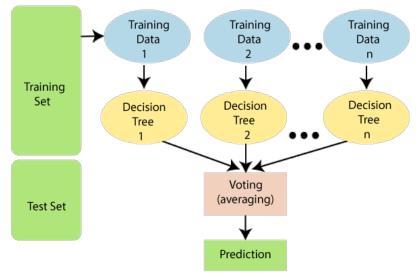
ML Techniques

- Extreme Gradient Boost (XGBoost): Combines the predictions of multiple weak models to produce a stronger model.
- Categorical Boosting (CatBoost): Similar working principle as XGBoost but faster and more efficient when working with categorical data.
- Random Forest: In a Random Forest classifier, multiple decision trees are created, each on a different subset of the data. Each tree gets a vote on the class label for a new instance. The class that gets the most votes is chosen as the final prediction.





XGBoost and CatBoost Architecture



Random Forest Classifier Architecture

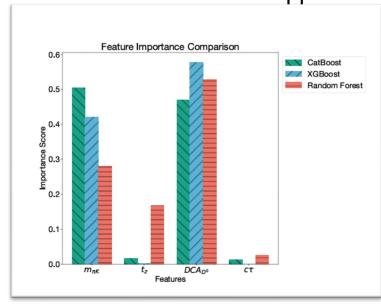


Training ML Models

Input Variables

- 1. Invariant Mass
- 2. The pseudo-proper time:

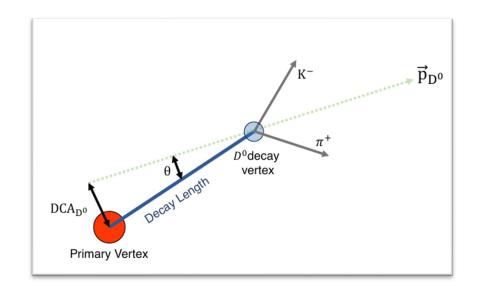
4. Distance of closest approach:



$$t_z = \frac{(z_{D^0} - z_{PV}) \times m_{D^0}}{p_z}$$

$$c\tau = \frac{cm_{D^0} \vec{L}.\vec{p_T}}{p_T^2}$$

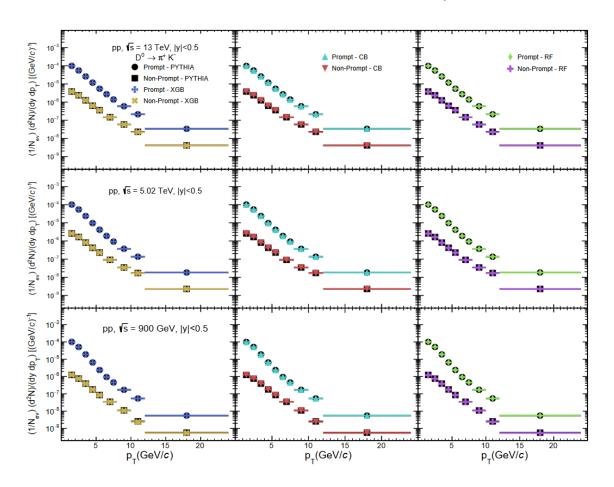
$$DCA_{D^0} = L \times \sin \theta$$

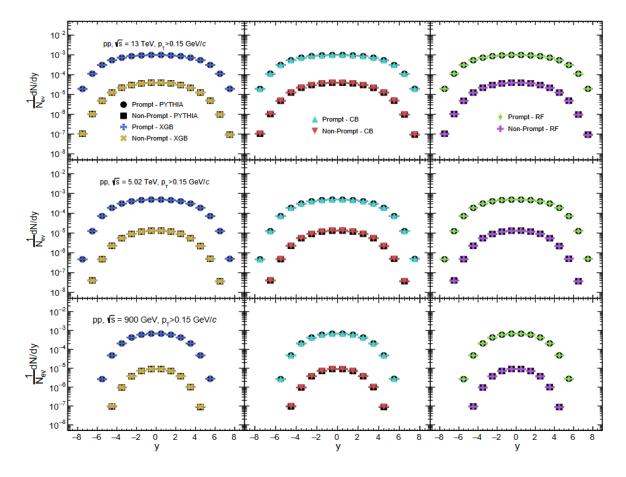


 \vec{L} is the vector pointing from the primary vertex towards D^0 decay vertex, i.e. $\vec{L} = \vec{V} - \vec{S}$. \vec{V} is the position of the primary vertex and \vec{S} is the position of the D^0 decay vertex given by,

$$S_i = \frac{(t_1 + d_{i,1}m_1/p_{i,1}) - (t_2 + d_{i,2}m_2/p_{i,2})}{m_1/p_{i,1} - m_2/p_{i,2}}$$

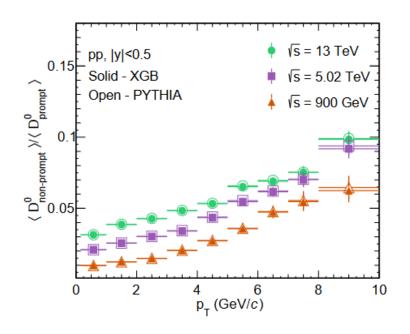
- CBM Transverse momenta and rapidity spectra
 - \circ Trained the ML models only at one centre-of-mass energy: $\sqrt{s}=13~{\rm TeV}$
 - \circ Predicted normalized D^0 meson yield at $\sqrt{s}=13$ TeV, 5.02 TeV and 900 GeV

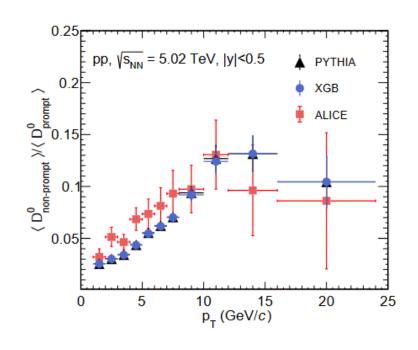






- ❖ Predicted the non-prompt to prompt ratio at 5.02 TeV and compared with ALICE data





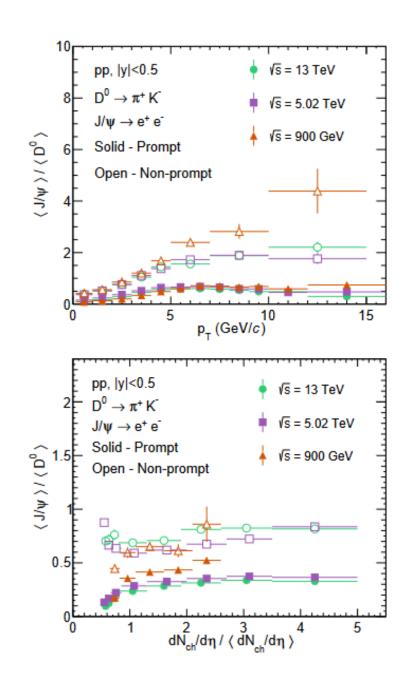
- \blacktriangleright One can clearly notice the increase in the ratio with increasing p_T across all the collision energies.
- > We observe an energy-dependent hierarchy in the ratio.



This ratio tells us about the production dynamics of charmonium state wrt open-charmed hadrons.

- \blacktriangleright The ratio shows a similar rising trend as a function of p_T , upto 5 GeV.
- The prompt J/ψ to D^0 ratio is always less than 1. However, the non-prompt ratio rises above 1, indicating a higher contribution of beauty hadrons towards J/ψ states.

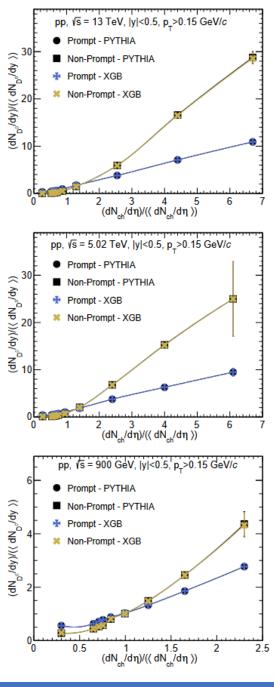
- For the prompt ratio, there is a slight increase and then it follows a flat trend.
- > However, the non-prompt ratio is independent of the charged particle multiplicity.





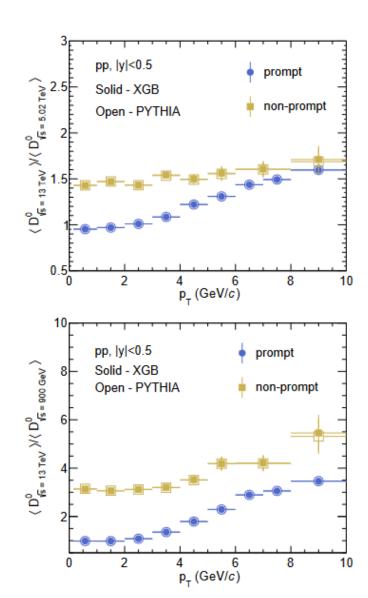
- \gt Self-normalized yield of prompt and non-prompt D^0 meson as a function of charged particle multiplicity.
- \triangleright A linear rise can be seen for the prompt D^0 meson.

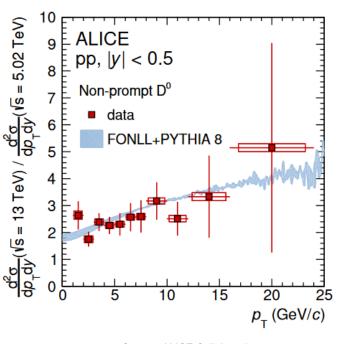
 \triangleright A non-linear rise is observed for non-prompt D^0 meson.





- ightharpoonup Role of centre-of-mass energy in D^0 meson production.
- \clubsuit In both plots, a clear increase in the ratio for prompt D^0 meson can be observed.
- In the top panel, a flat trend can be seen for the non-prompt case. Similar results have been observed in ALICE.
- In the bottom panel, the same ratio tends to increase slightly due to the significant difference in energies.
- \circ A higher value of non-prompt ratio indicates more abundant production of beauty hadrons at higher \sqrt{s} .

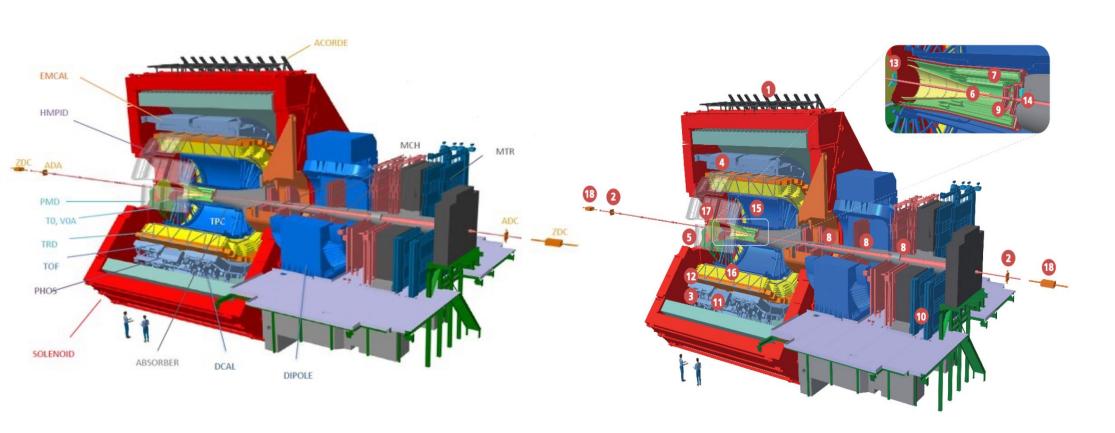




Source: ALICE Collaboration, arXiv:2402.16417



ALICE in RUN 3 with Muon Forward Tracker (MFT) will enhance the charmonia measurement Capability while distinguishing prompt vs non-prompt J/ψ .



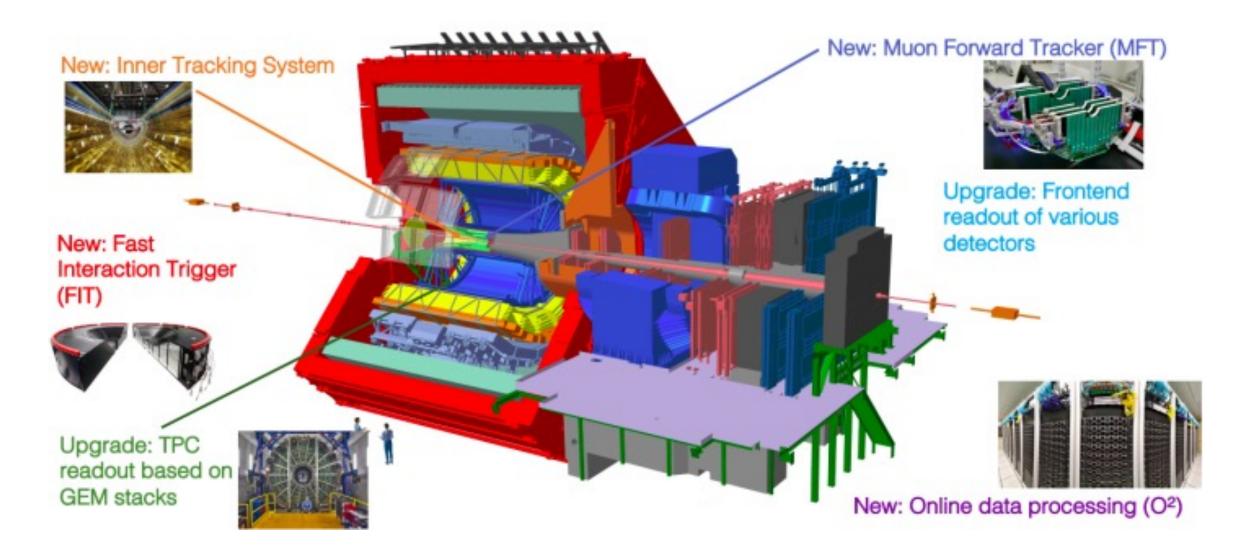
- ACORDE | ALICE Cosmic Rays Detector
- AD | ALICE Diffractive Detector
- DCal Di-jet Calorimeter
- 4 EMCal | Electromagnetic Calorimeter
- HMPID | High Momentum Particle Identification Detector
- 6 ITS-IB | Inner Tracking System Inner Barrel
- 7 ITS-OB | Inner Tracking System Outer Barrel
- 8 MCH | Muon Tracking Chambers
- MFT | Muon Forward Tracker
- MID | Muon Identifier
- PHOS / CPV | Photon Spectrometer
- 12 TOF | Time Of Flight
- 13 T0+A | Tzero + A
- 10+C | Tzero + C
- 15 TPC | Time Projection Chamber
- 16 TRD | Transition Radiation Detector
- 17 V0+ | Vzero + Detector
- 18 ZDC | Zero Degree Calorimeter

ALICE RUN 2 Setup

ALICE RUN 3 Setup

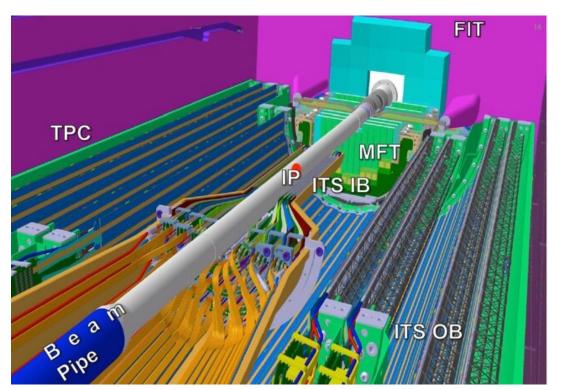


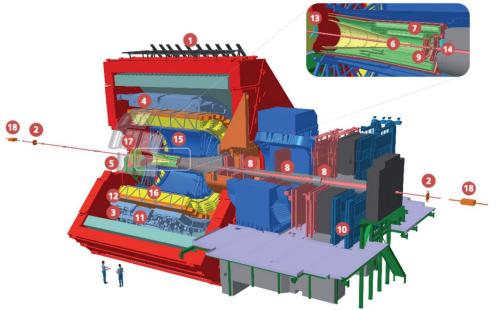
ALICE RUN 3 Upgrades





ALICE in RUN 3 with Muon Forward Tracker (MFT) will enhance the charmonia measurement Capability while distinguishing prompt vs non-prompt J/ψ .





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ALICE RUN 3 Setup showing MFT

ALICE RUN 3 and beyond stays interesting given HF measurements at the LHC!

