

“An experimentalist’s guide to the outer crust of accreting neutron stars”

or:

“Nuclear masses and nuclear processes in the crust of neutron stars”

Alfredo Estrade
The University of Edinburgh



530th WE Heraeus Seminar, Bad Honnef, April 2012



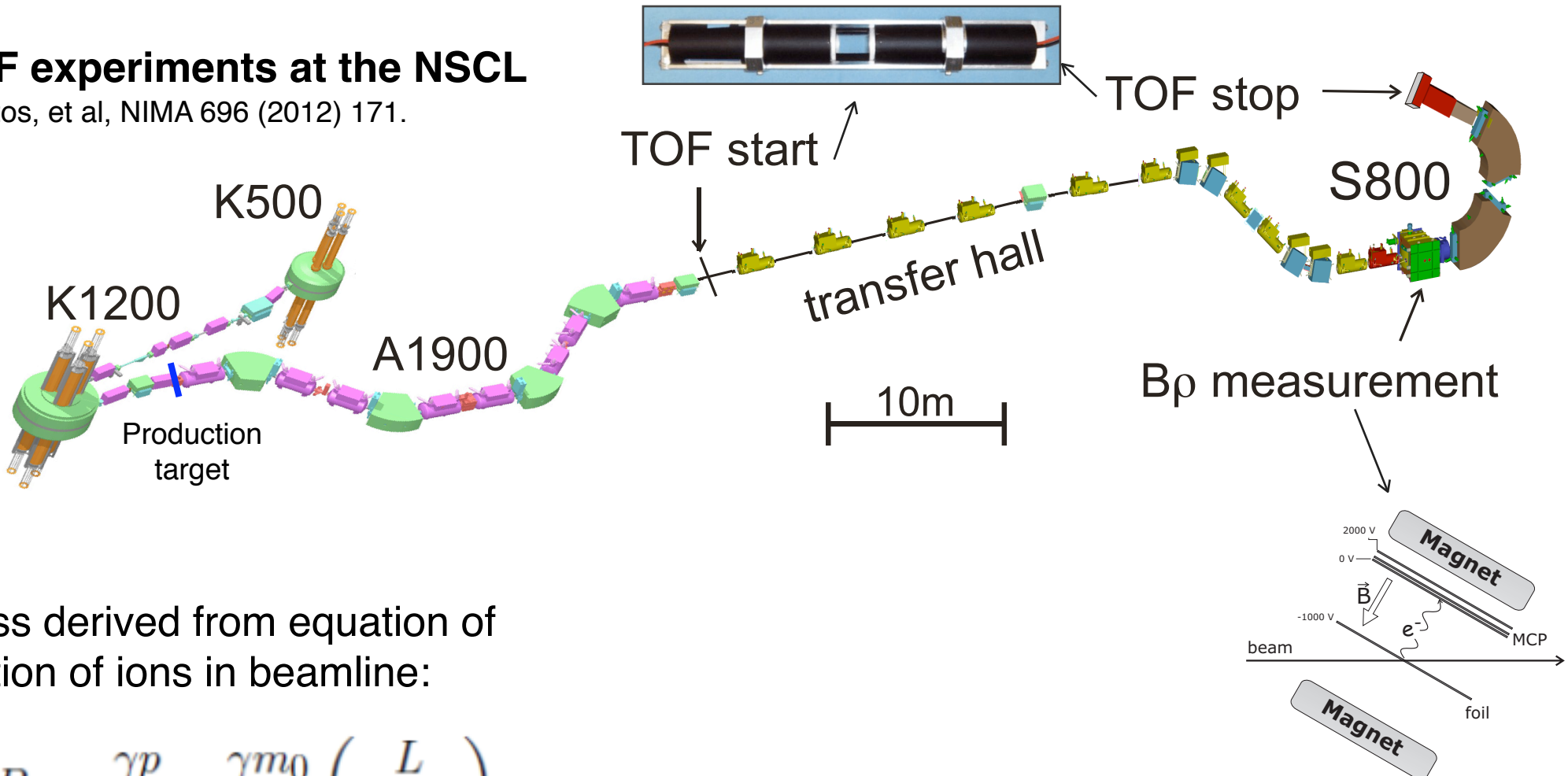
Outline

- TOF mass measurements of very neutron-rich nuclei: the case of ^{66}Mn .
- Electron Capture processes in accreting neutron stars.
- Nuclear mass models for EC calculations.
- Regions of interest for future experiments?

TOF mass measurement technique

TOF experiments at the NSCL

Matos, et al, NIMA 696 (2012) 171.

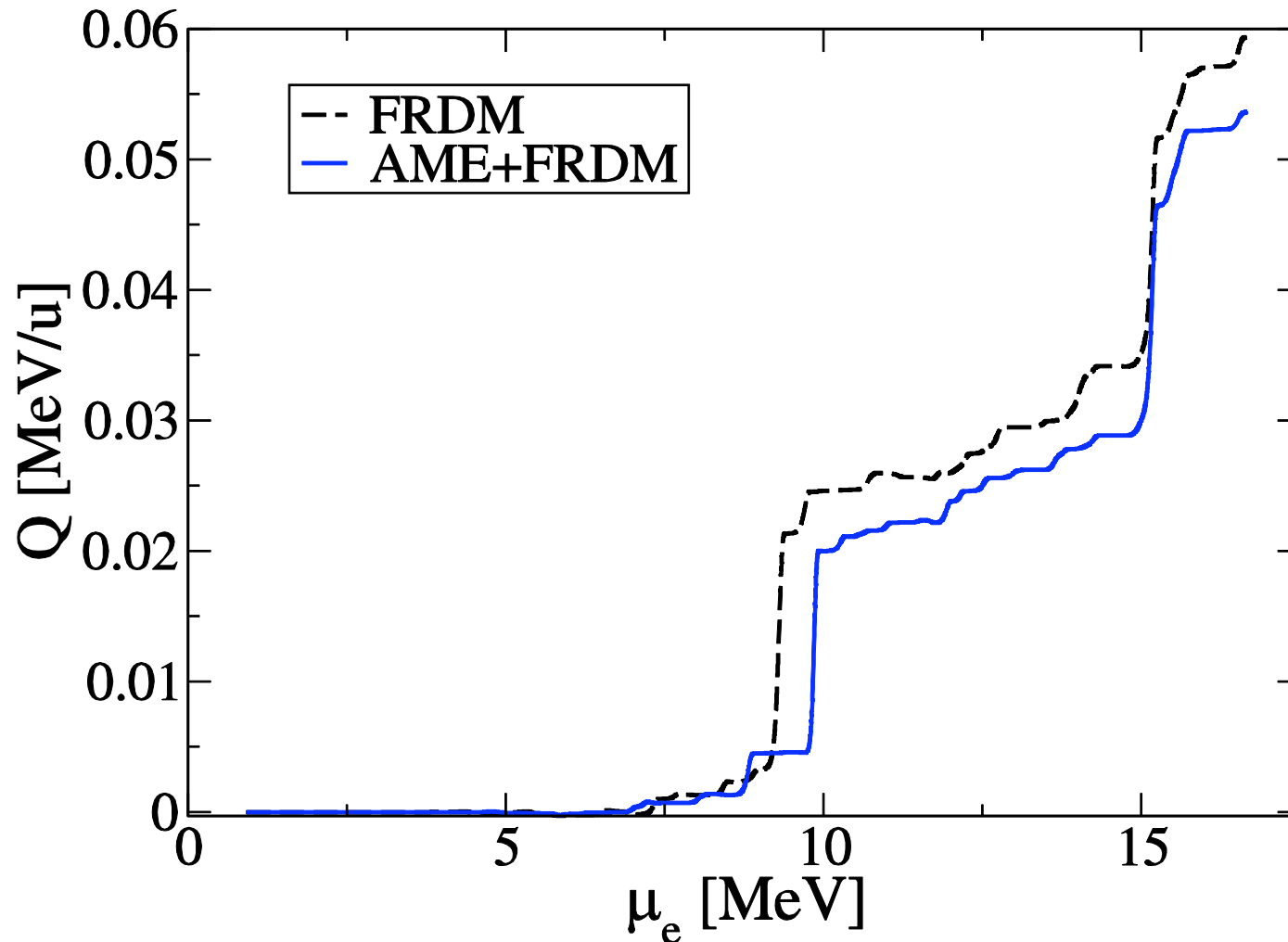


Mass derived from equation of motion of ions in beamline:

$$B\rho = \frac{\gamma p}{q} = \frac{\gamma m_0}{q} \left(\frac{L}{TOF} \right)$$

See talk by Z. Meisel on Thursday

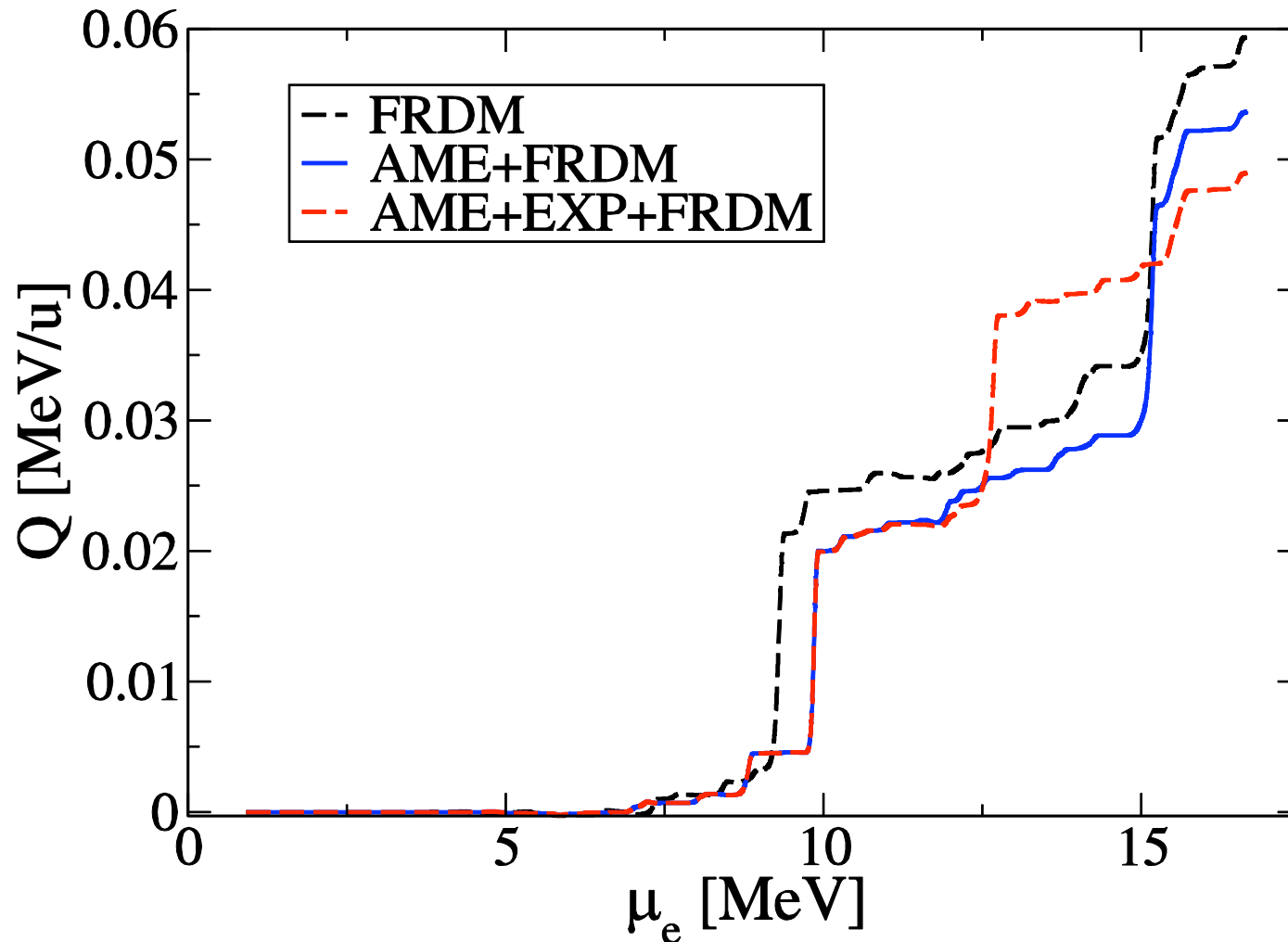
Experimental data and crust calculations



FRDM: $Q_{\text{EC}}(^{66}\text{Ni}) = -9.35$ MeV

AME: $Q_{\text{EC}}(^{66}\text{Ni}) = -9.9$ MeV

Experimental data and crust calculations



FRDM: $Q_{\text{EC}}(^{66}\text{Ni}) = -9.35$ MeV

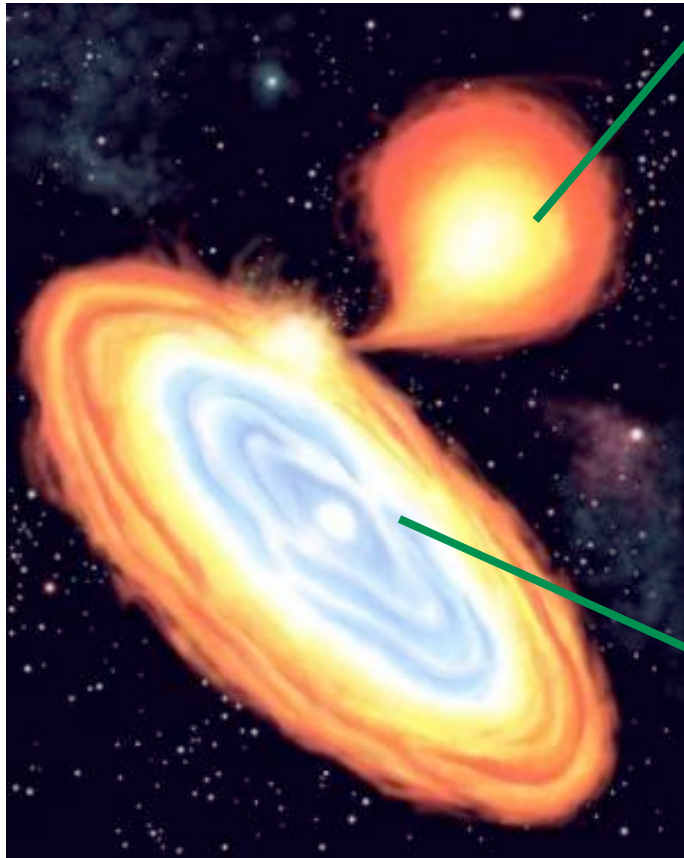
AME: $Q_{\text{EC}}(^{66}\text{Ni}) = -9.9$ MeV

FRDM: $Q_{\text{EC}}(^{66}\text{Fe}) = -15.3$ MeV

EXP: $Q_{\text{EC}}(^{66}\text{Fe}) = -12.7$ MeV

Accreting Neutron Stars

An artist's rendition...



low mass companion

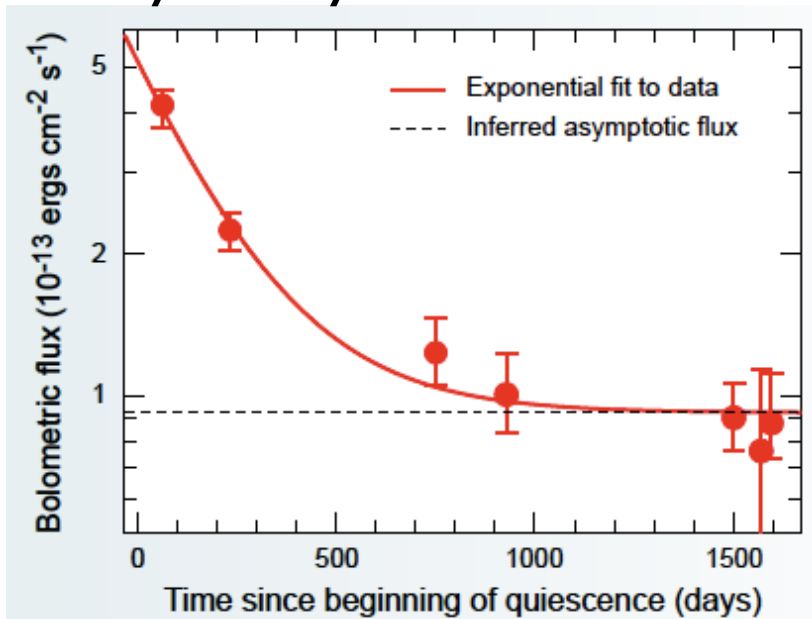
neutron star

The Chandra X-ray observatory.



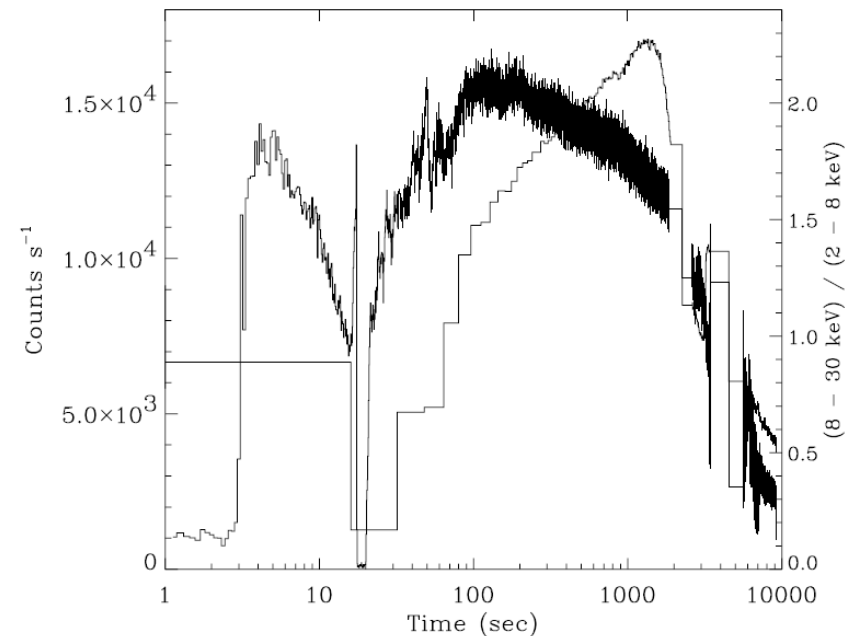
Observables affected by thermal properties of the neutron star

Cooling curve from transient X-ray binary



E. Cackett 2006

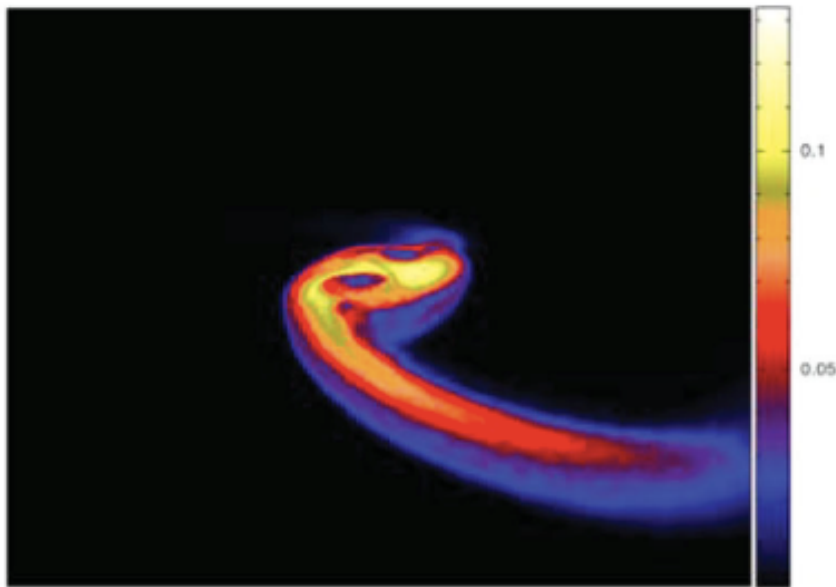
Superburst event from 4U 1820-30



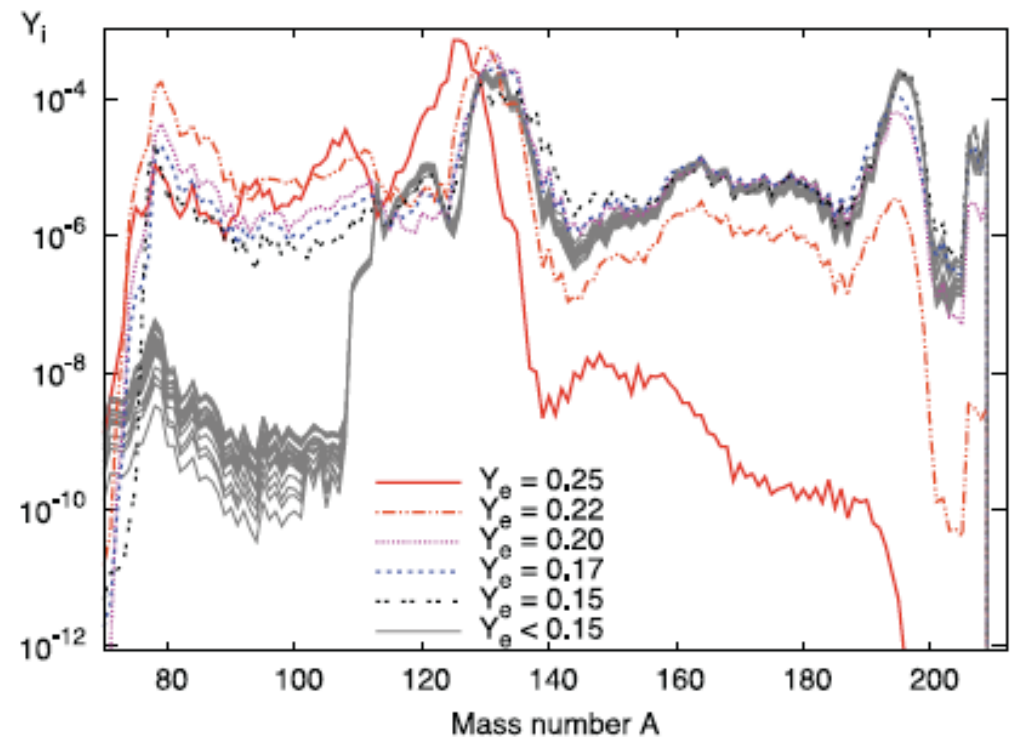
T. Strohmayer and E. Brown 2003

Crust chemical composition and nucleosynthesis in NS-NS mergers

Y_e distribution in merger event



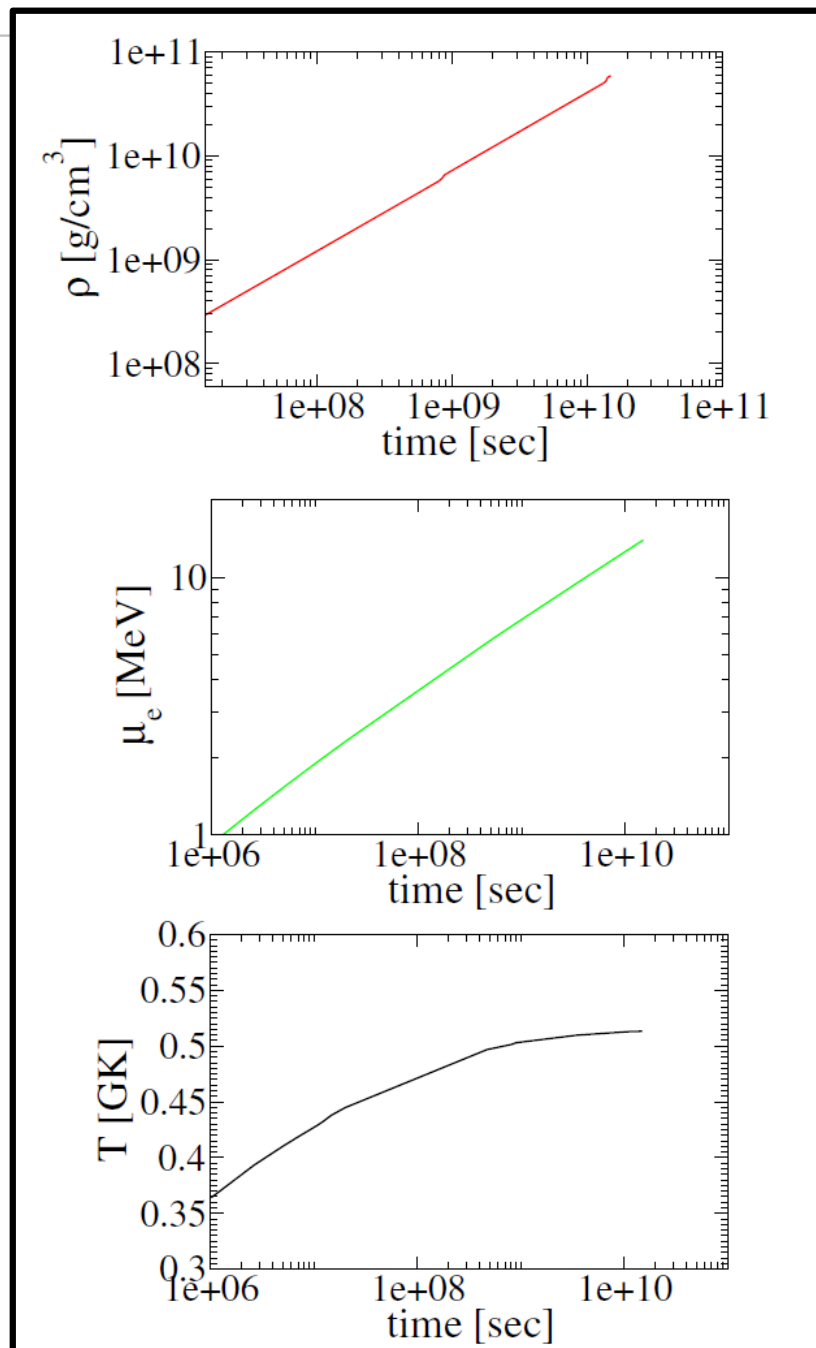
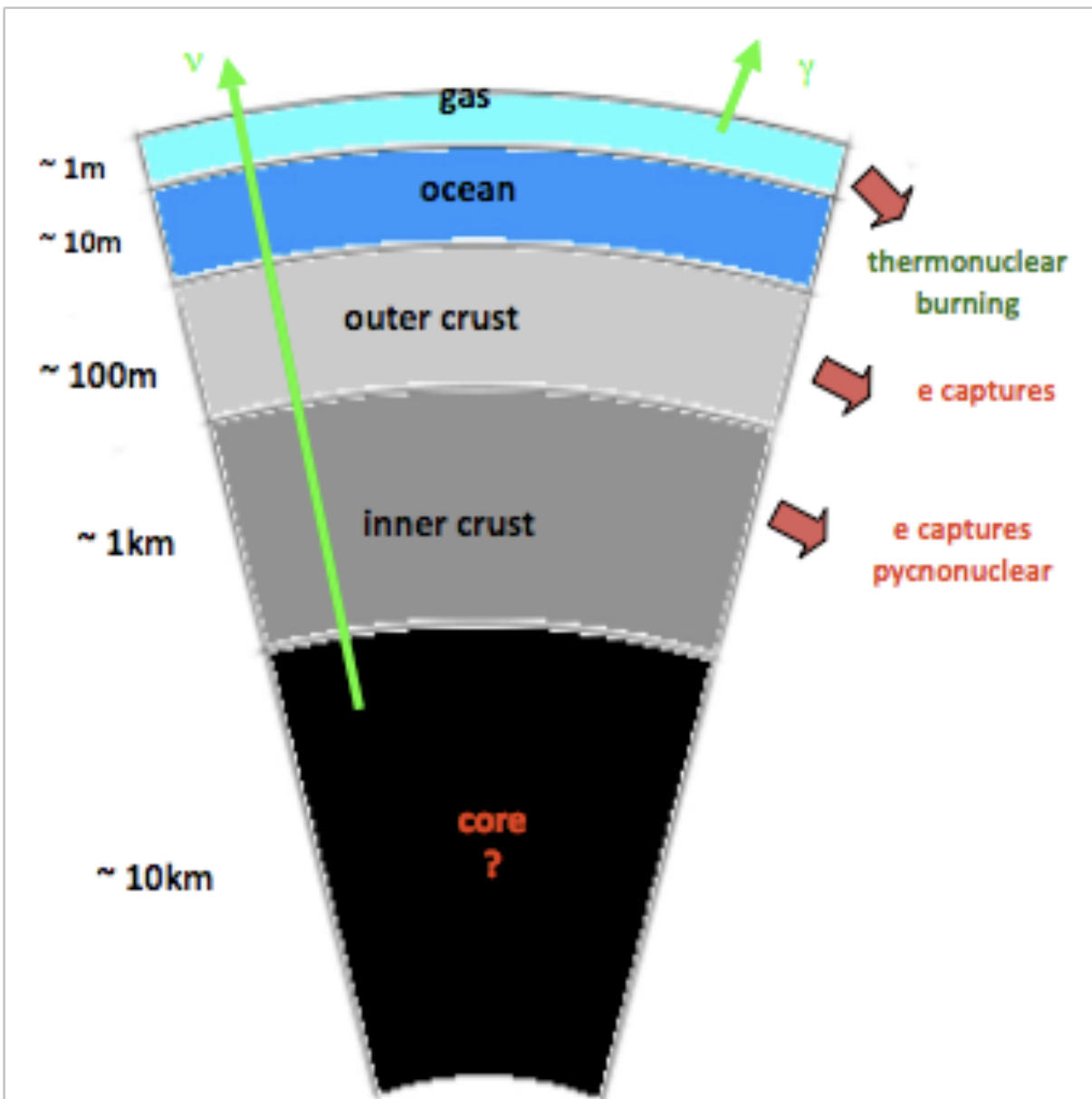
Abundance distributions resulting of trajectories with different Y_e



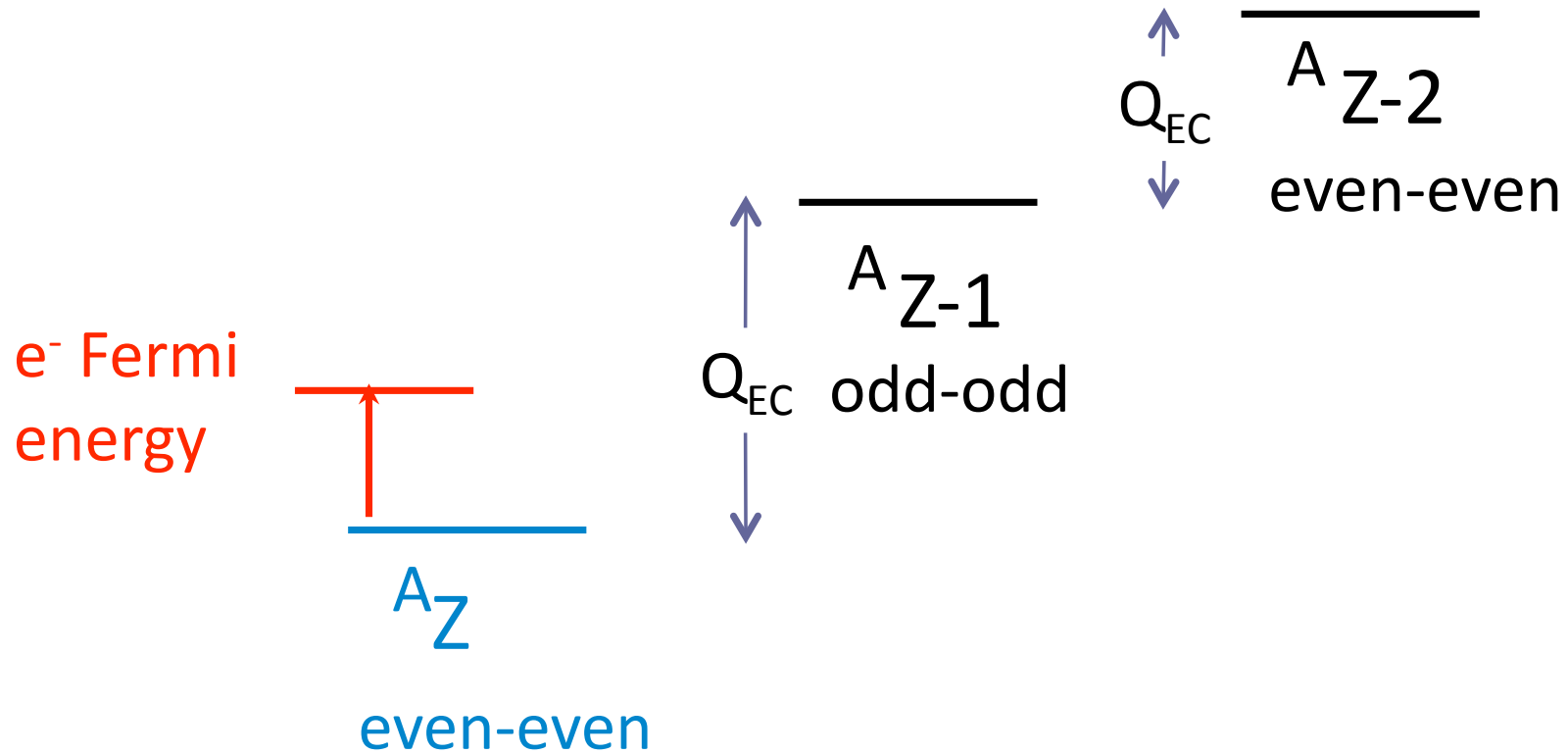
O. Korobkin et al 2012

talk by S. Rosswog on Friday

Density-driven nuclear processes in accreting neutron stars



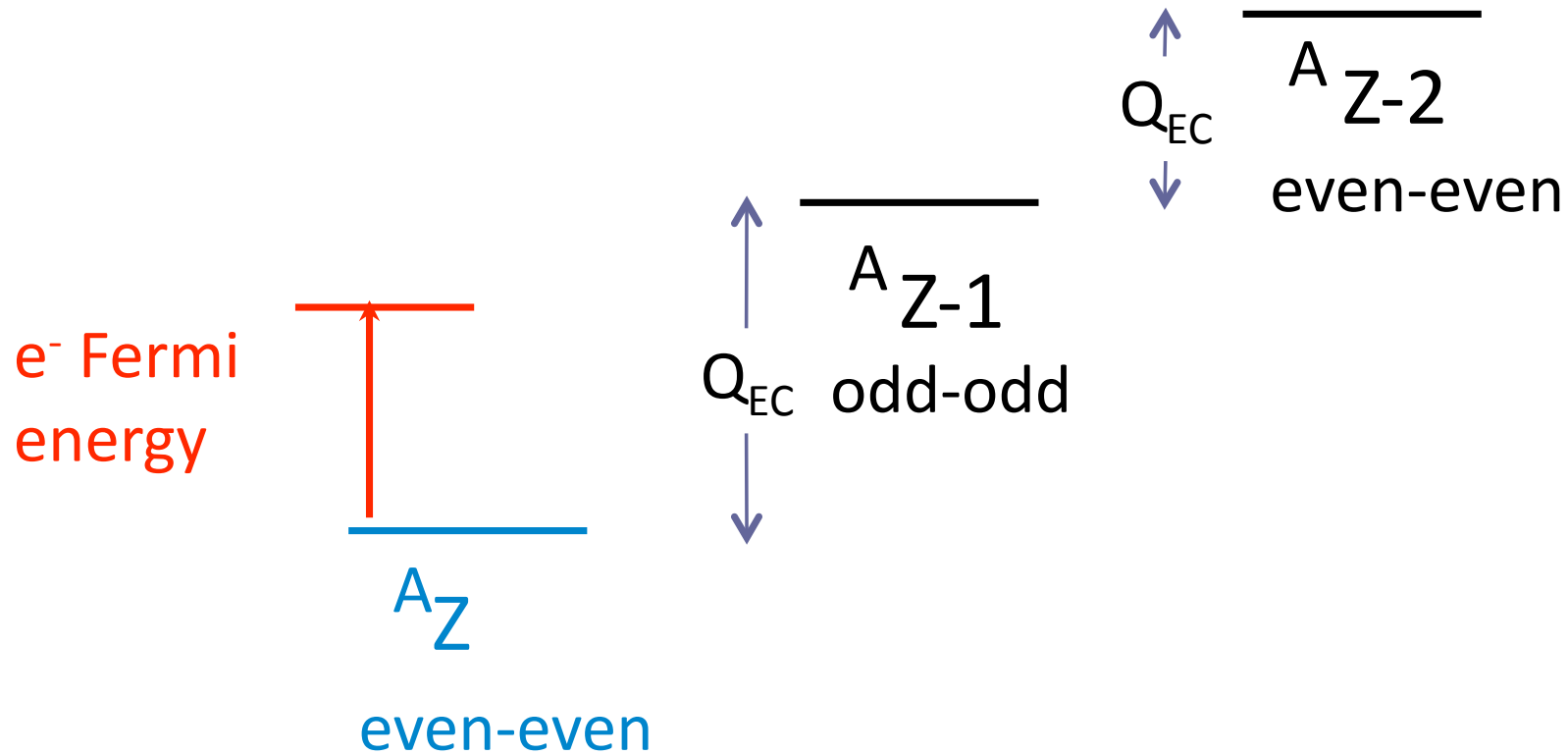
Electron capture processes in the outer crust



**For even A
isobaric chain**

$$Q_{EC} = M(A, Z) - M(A, Z - 1)$$

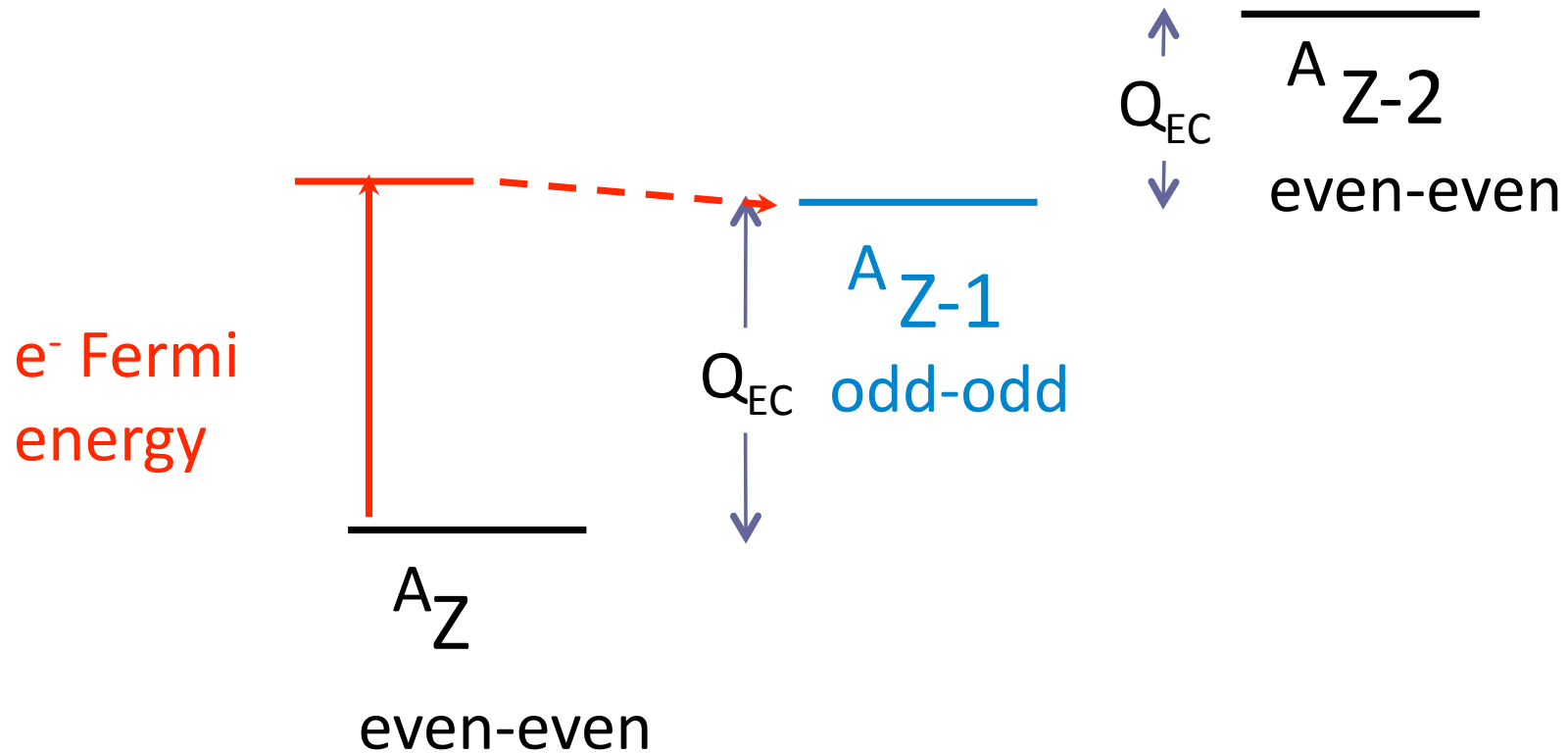
Electron capture processes in the outer crust



**For even A
isobaric chain**

$$Q_{EC} = M(A, Z) - M(A, Z - 1)$$

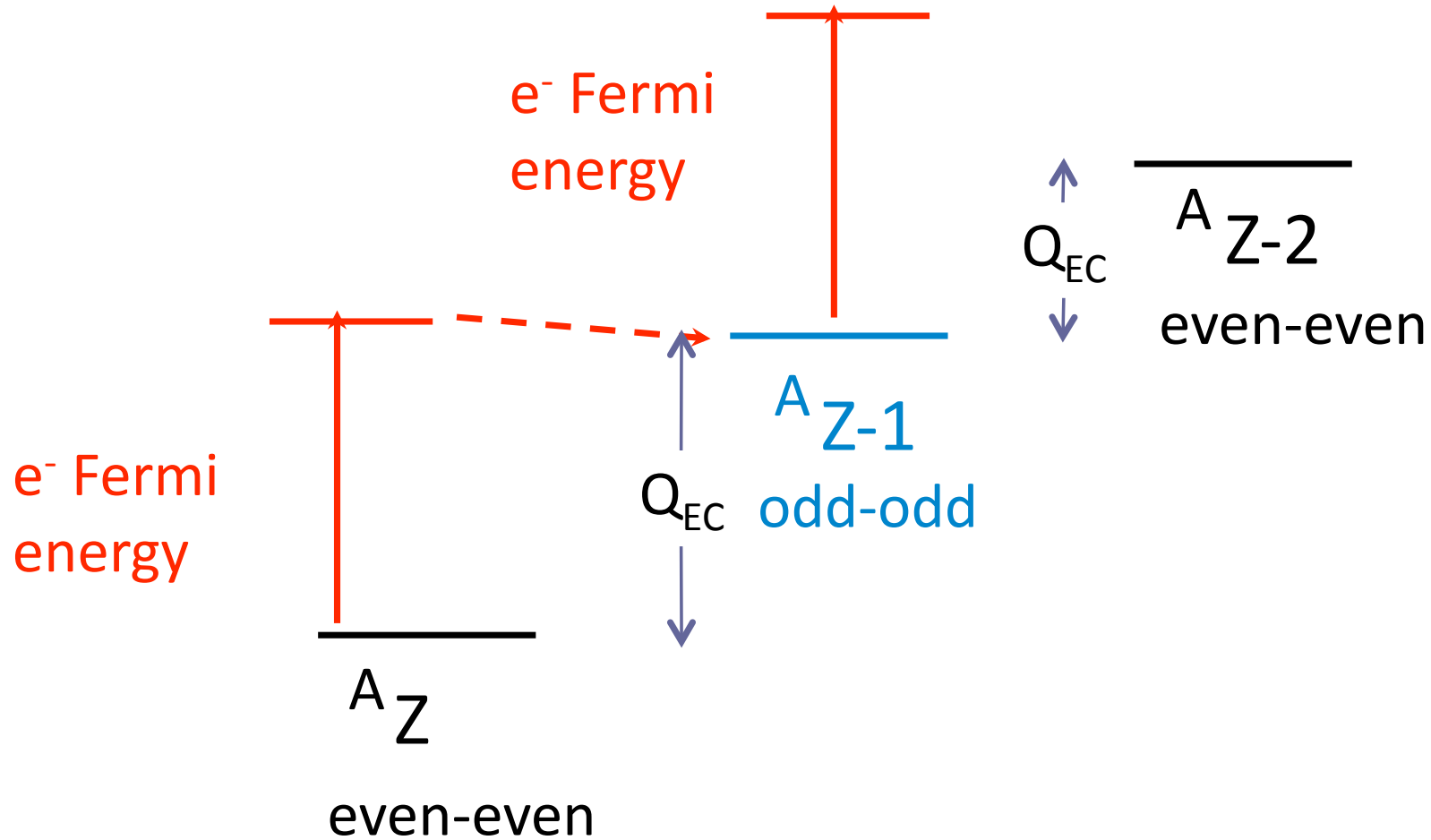
Electron capture processes in the outer crust



**For even A
isobaric chain**

$$Q_{EC} = M(A, Z) - M(A, Z - 1)$$

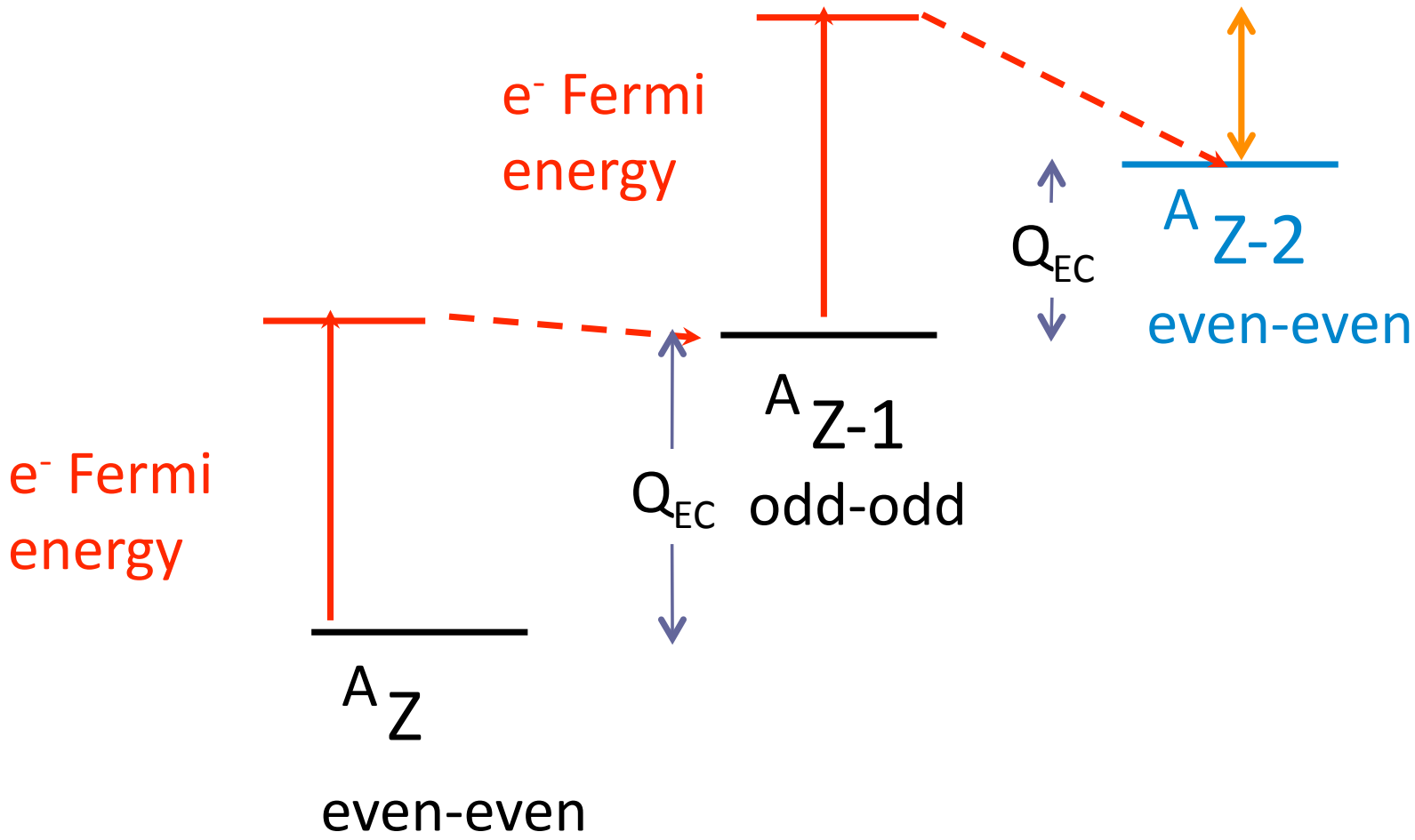
Electron capture processes in the outer crust



**For even A
isobaric chain**

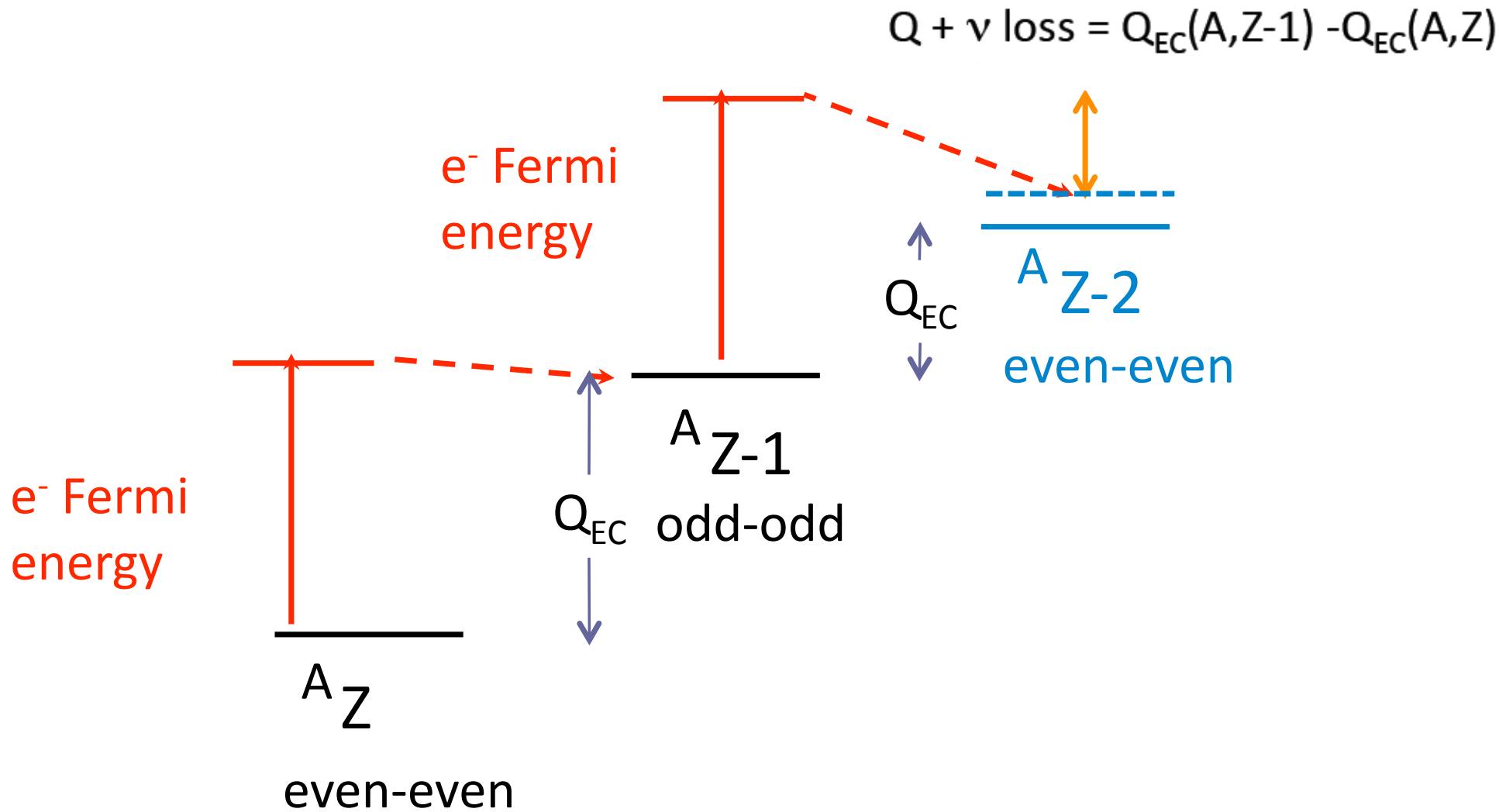
$$Q_{EC} = M(A, Z) - M(A, Z - 1)$$

$$Q + \nu \text{ loss} = Q_{\text{EC}}(A, Z-1) - Q_{\text{EC}}(A, Z)$$



**For even A
isobaric chain**

$$Q_{\text{EC}} = M(A, Z) - M(A, Z - 1)$$



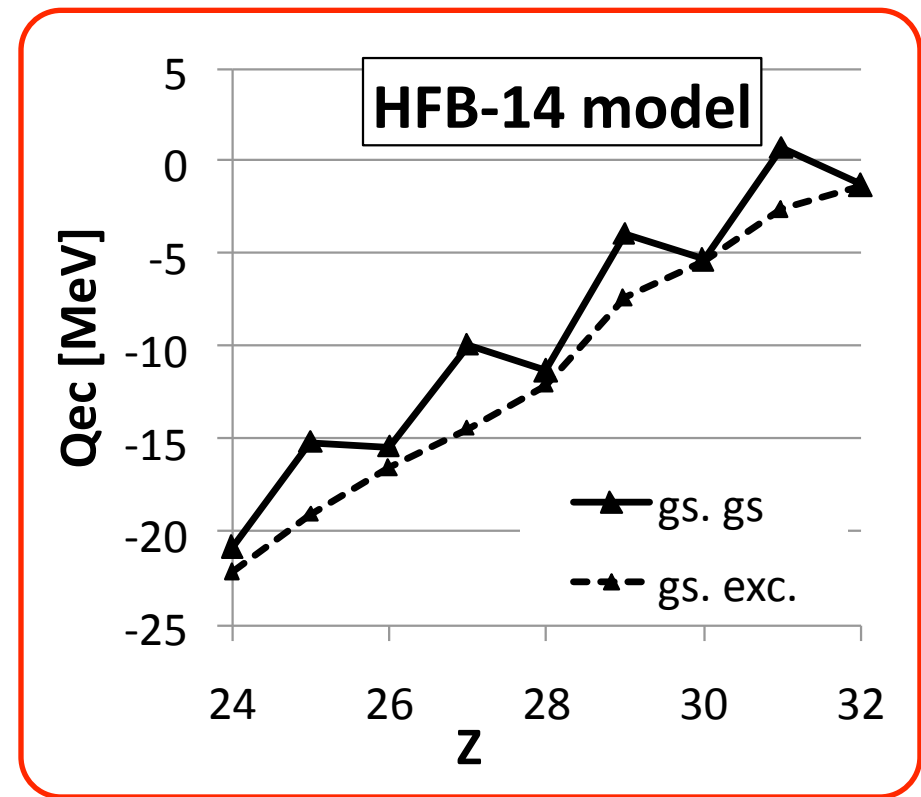
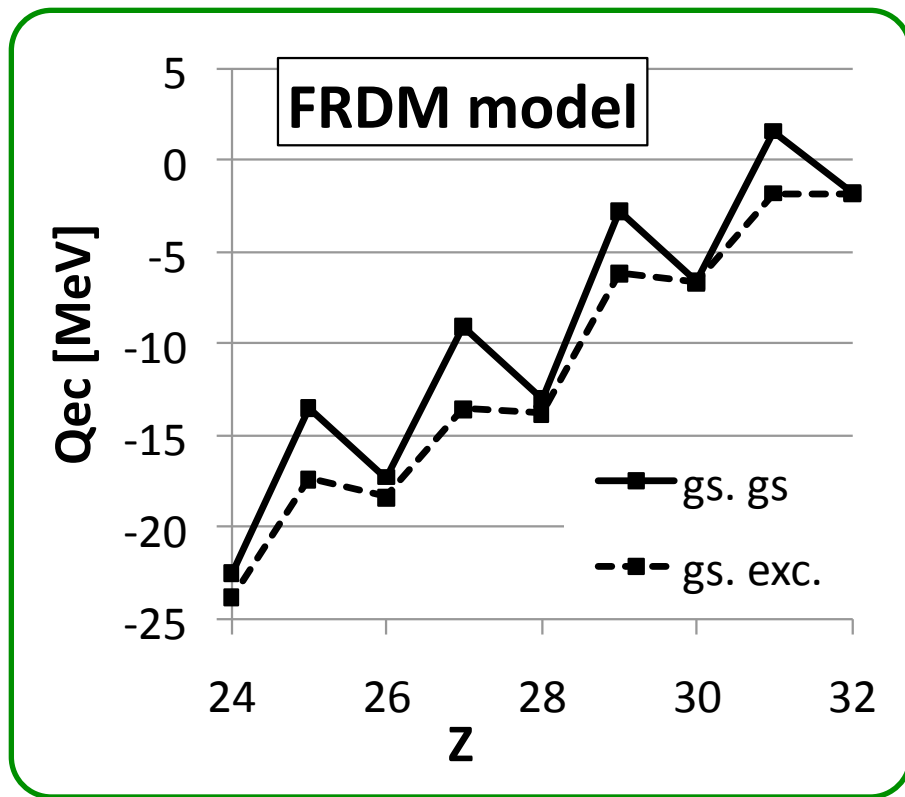
Masses (and excitation energies) set **location** of and **strength** of nuclear heating sources.

$$Q_{\text{EC}} = M(A, Z) - M(A, Z - 1)$$

Electron captures for $A=70$ mass chain

MOVIES!

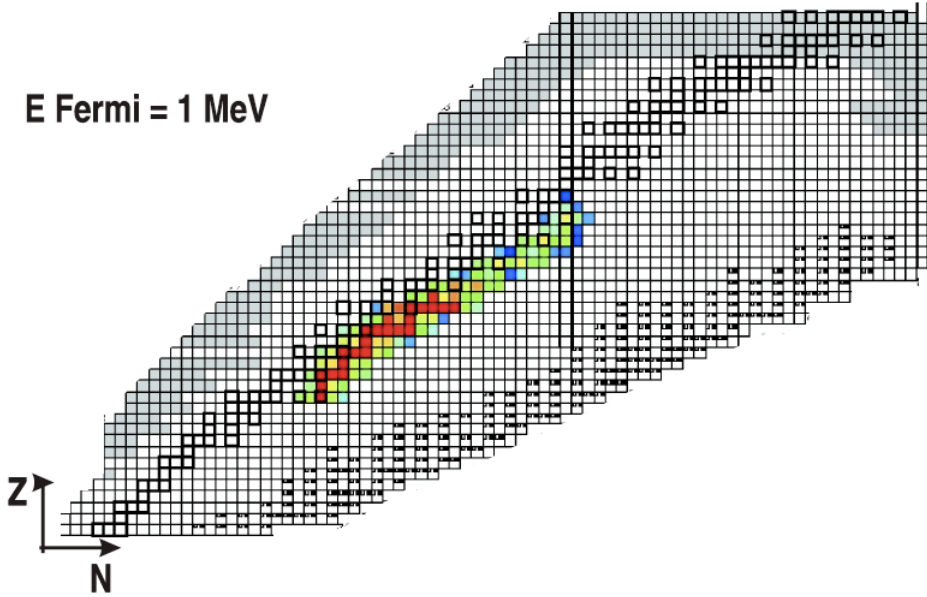
Electron captures for A=70 mass chain



Different staggering of EC Q-value results in different behavior of electron captures along this isobaric chain.

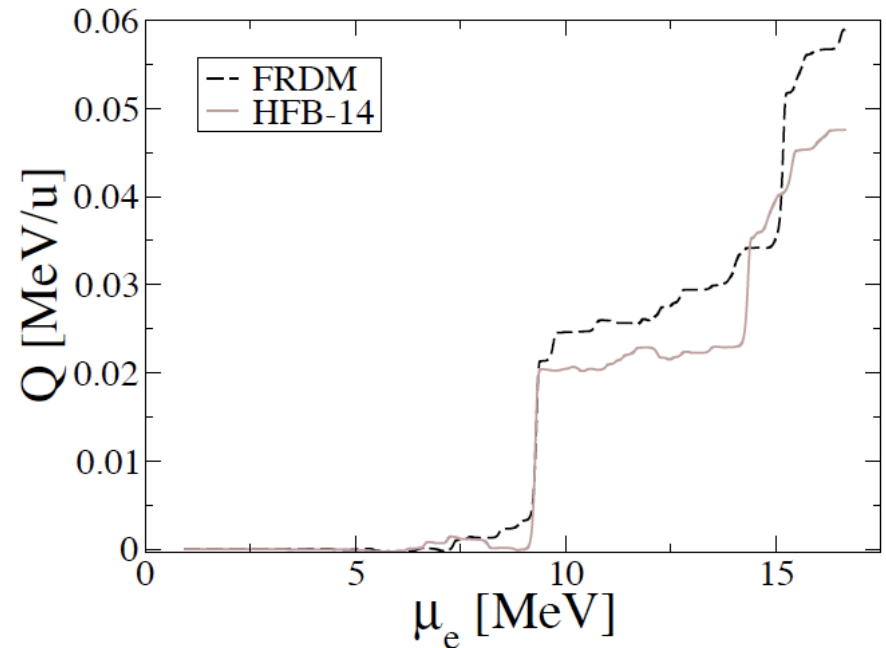
Results for Carbon superburst ashes

$E_{\text{Fermi}} = 1 \text{ MeV}$



Carbon superburst ashes in
 $48 < A < 70$ range

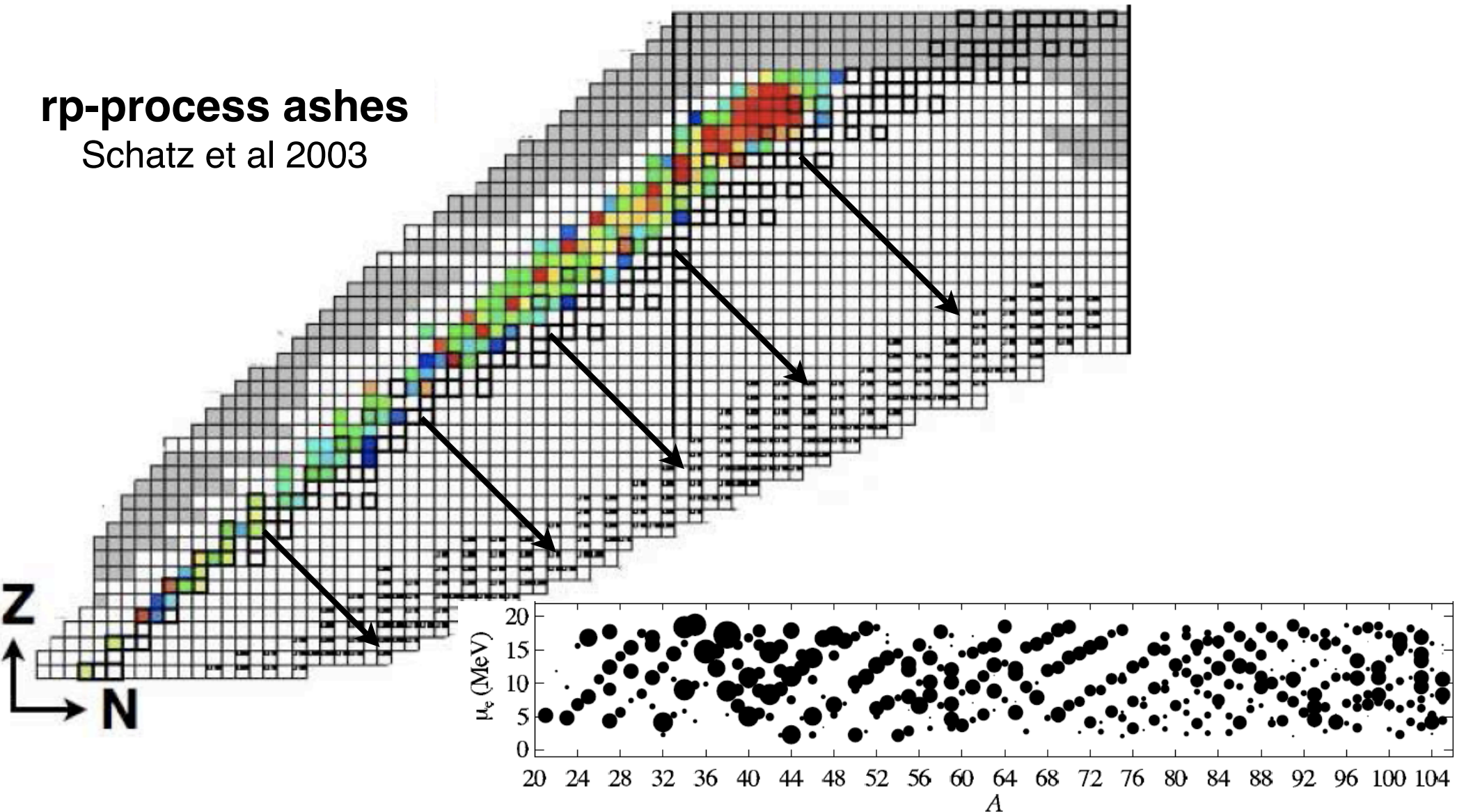
Heat deposition in outer crust



Nuclear masses relevant to nuclear processes in neutron star crusts

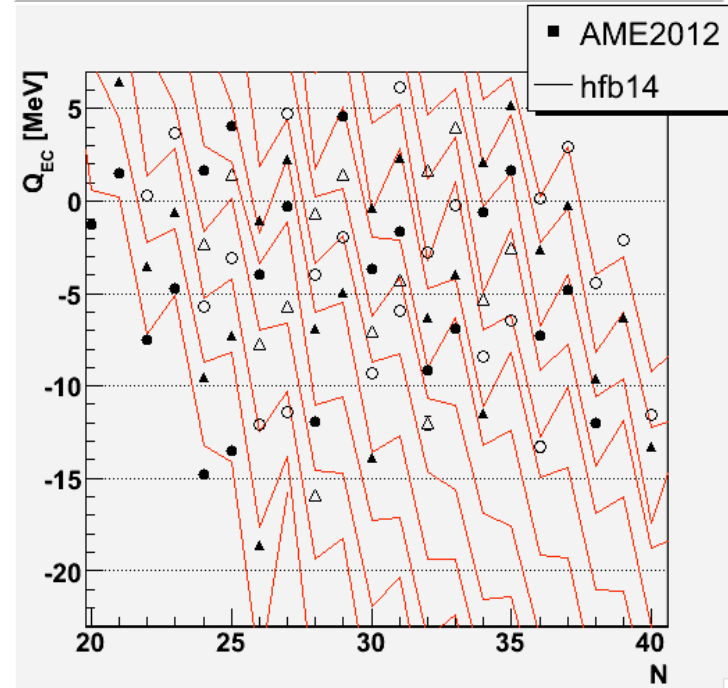
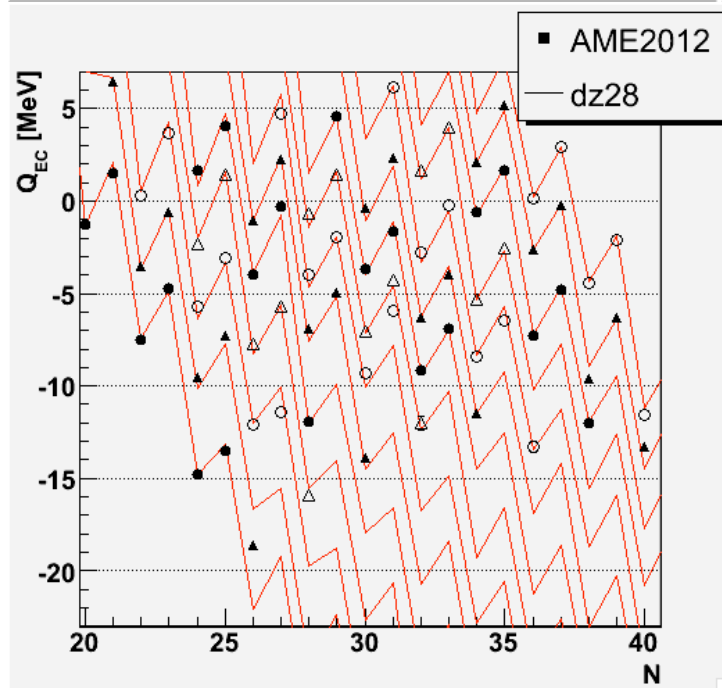
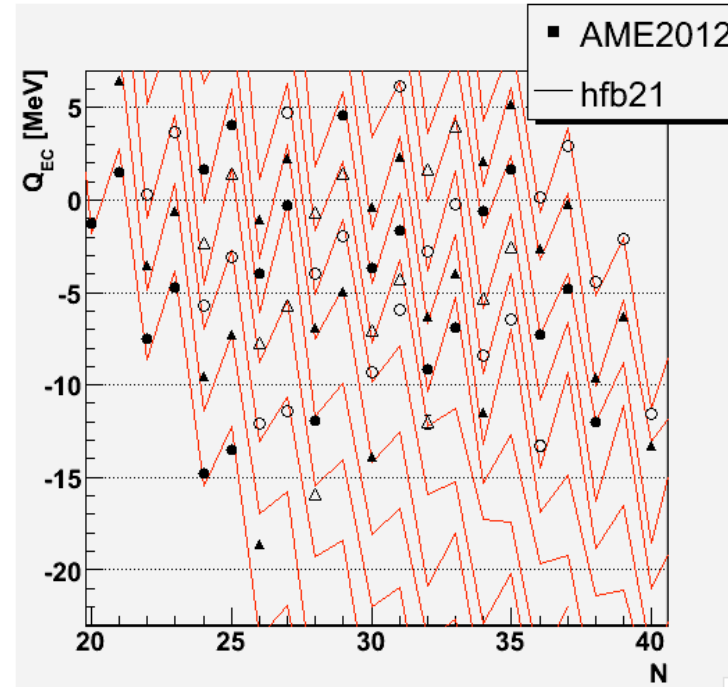
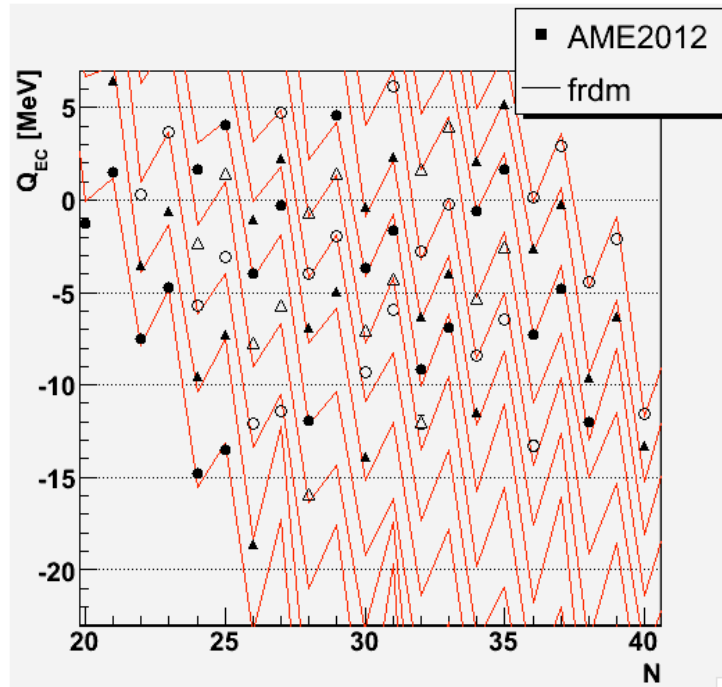
rp-process ashes

Schatz et al 2003

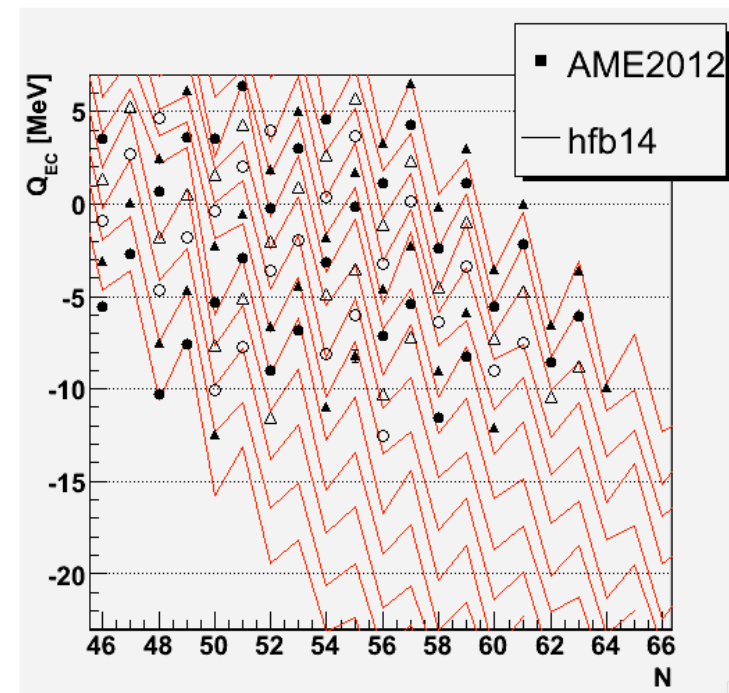
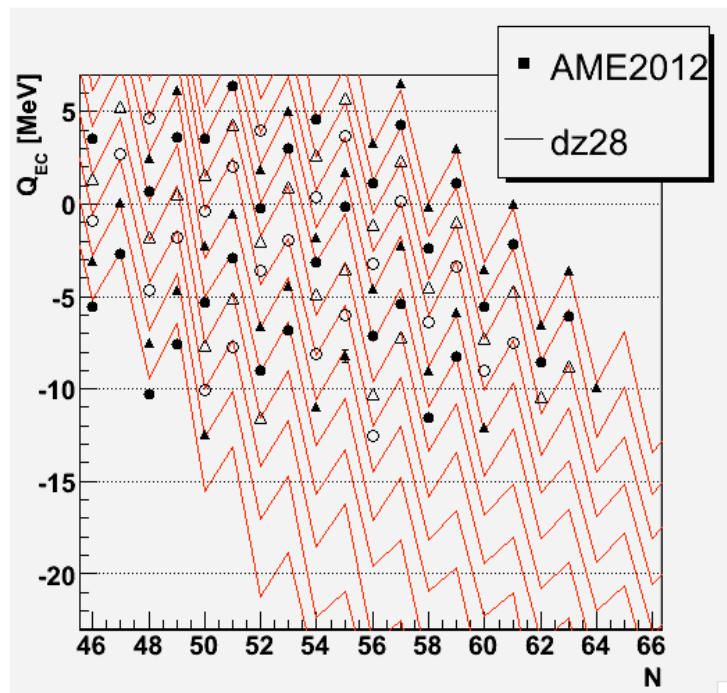
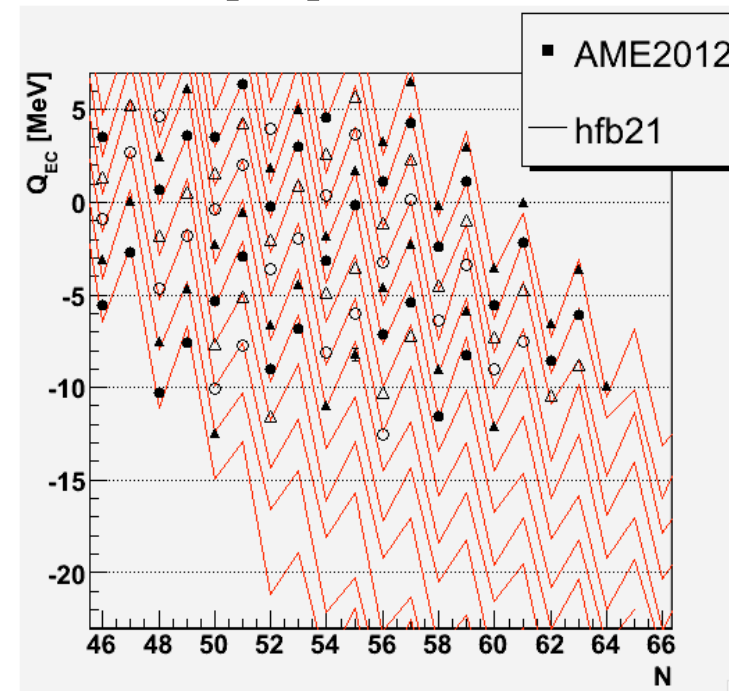
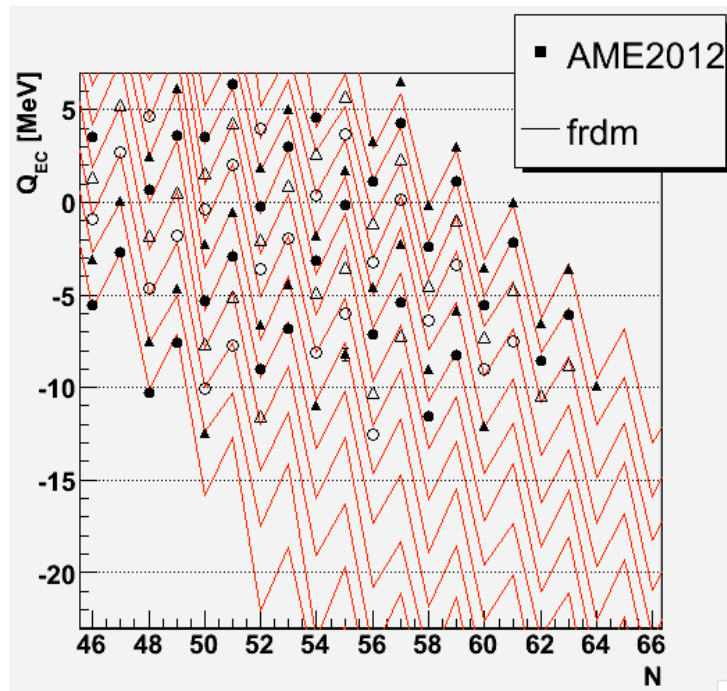


Gupta et al 2007

Models vs AME2012: Q_{EC} for C burst ashes

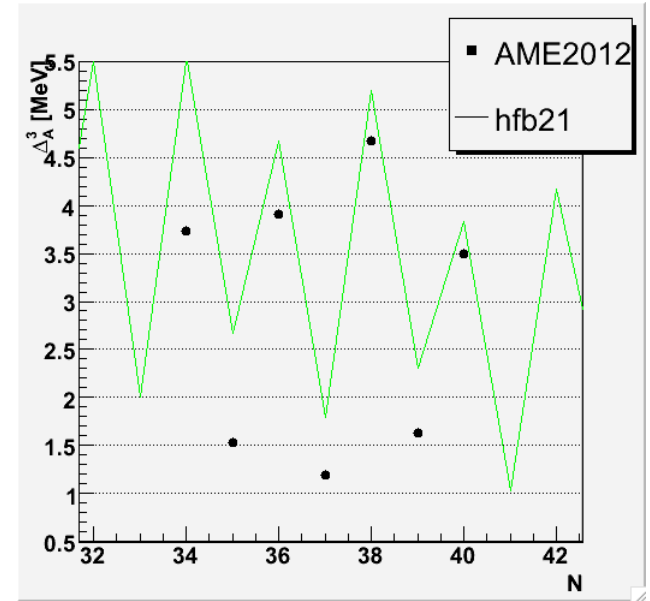
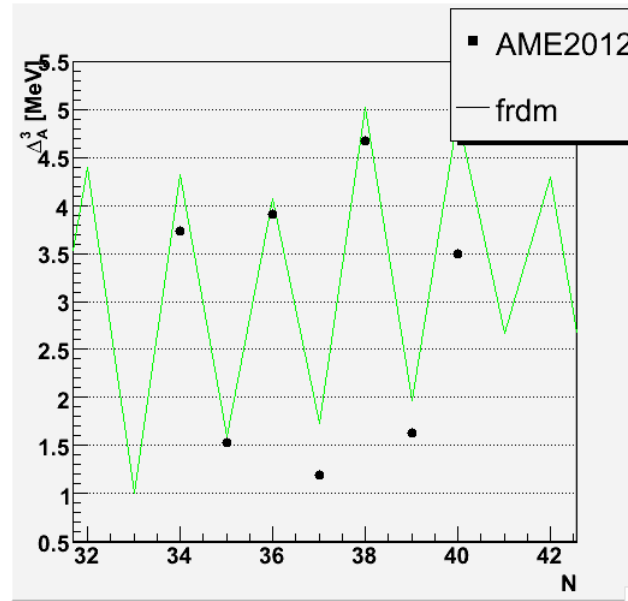
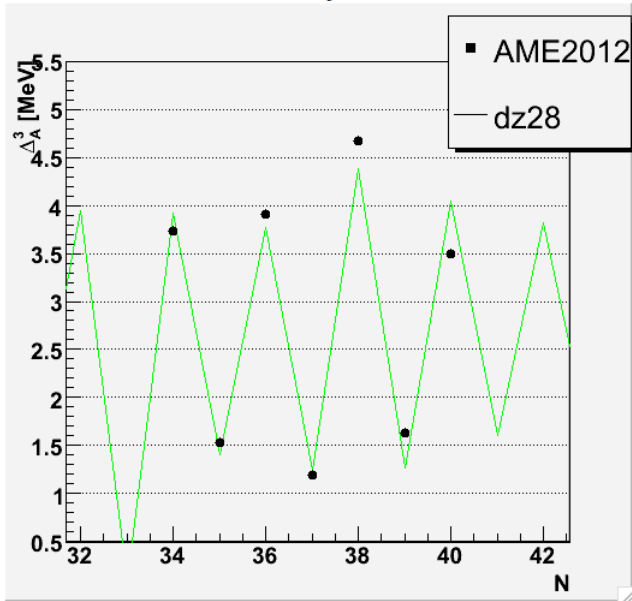


Models vs AME2012: Q_{EC} for rp-process ashes

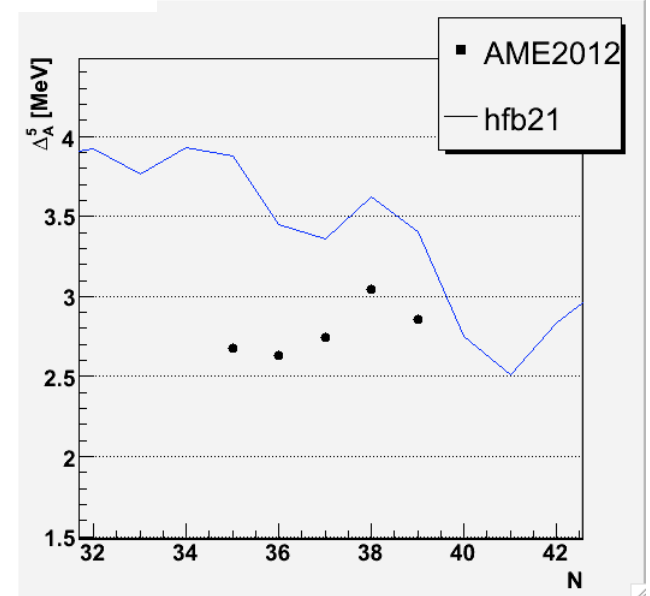
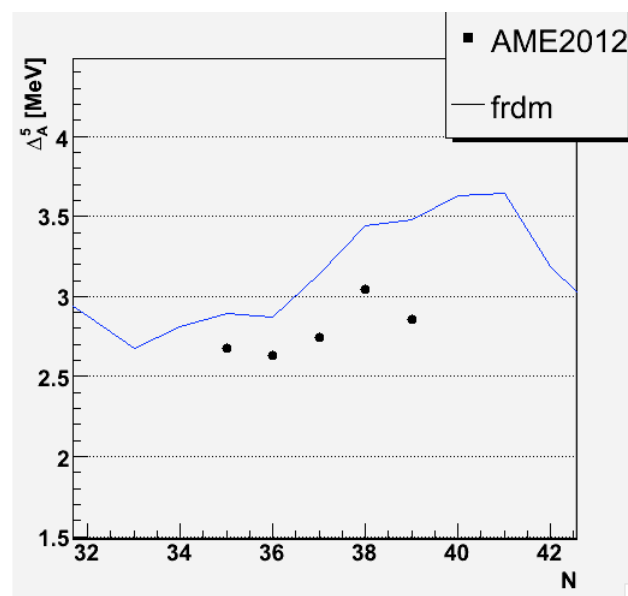
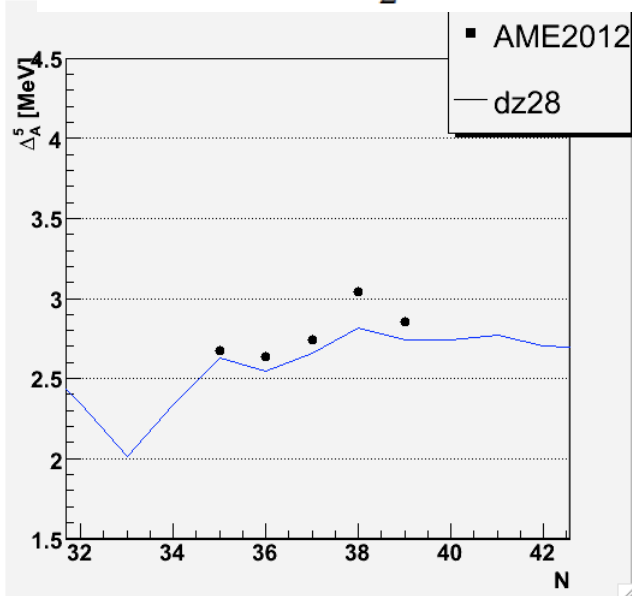


A=66 isobaric chain: 3- and 5-point mass differences

$$\Delta_A^{(3)}(N, Z) = \frac{-1^N}{2} [E(Z+1, N-1) - 2E(Z, N) + E(Z-1, N+1)]$$

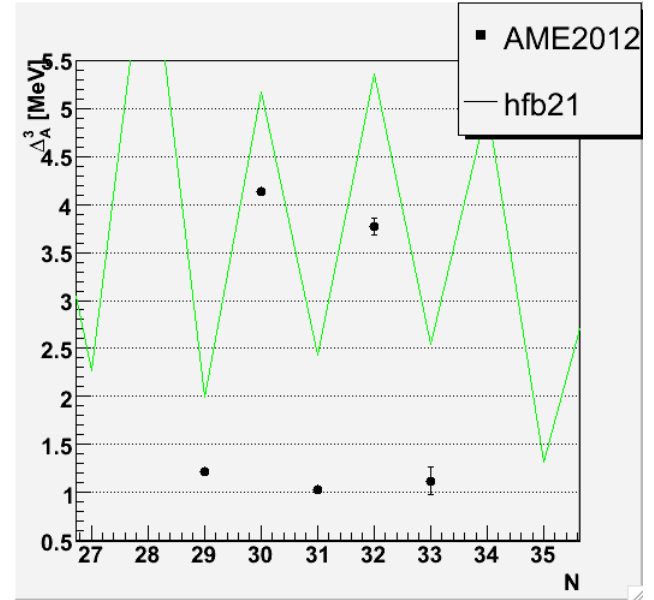
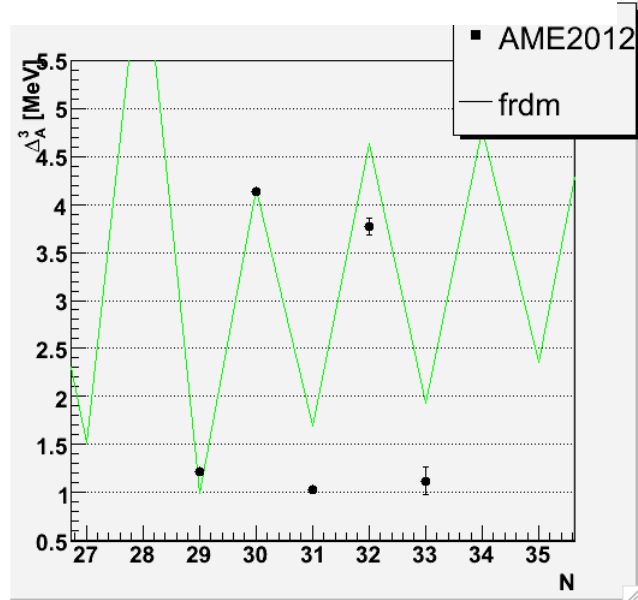
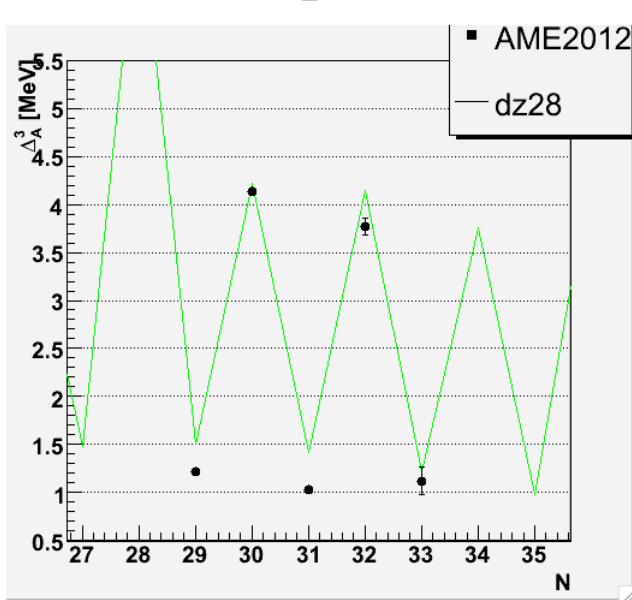


$$\Delta_A^{(5)}(N, Z) = -\frac{(-1)^N}{2} [E(Z+2) - 4E(Z+1) + 6E(Z) - 4E(Z-1) + E(Z-2)]$$

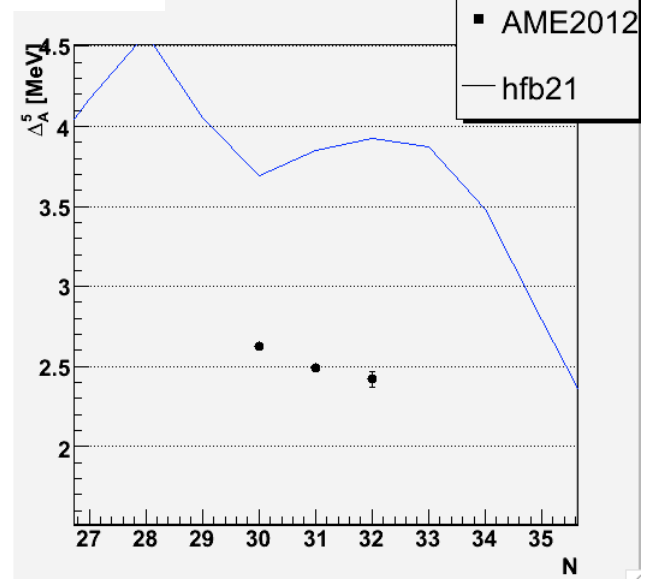
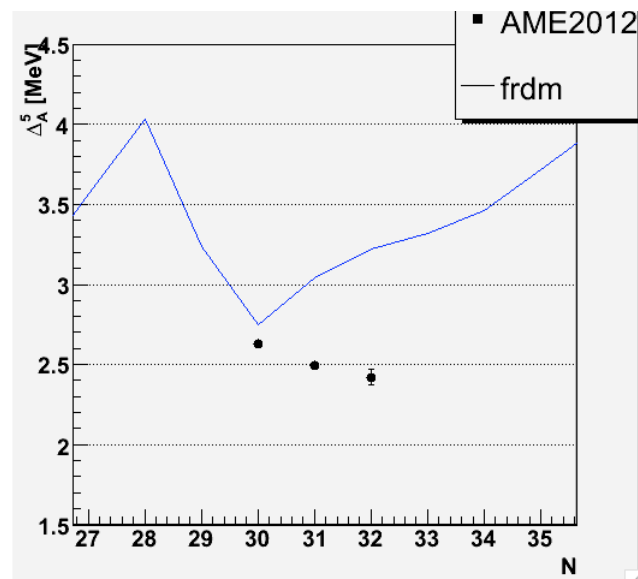
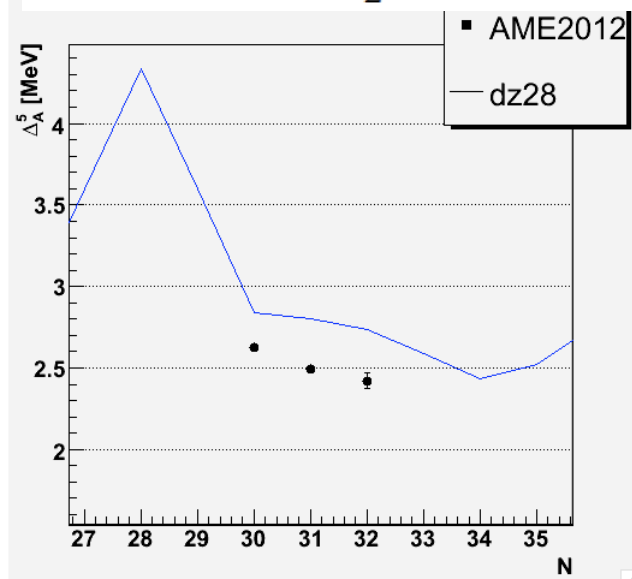


A=56 isobaric chain: 3- and 5-point mass differences

$$\Delta_A^{(3)}(N, Z) = \frac{-1^N}{2} [E(Z+1, N-1) - 2E(Z, N) + E(Z-1, N+1)]$$

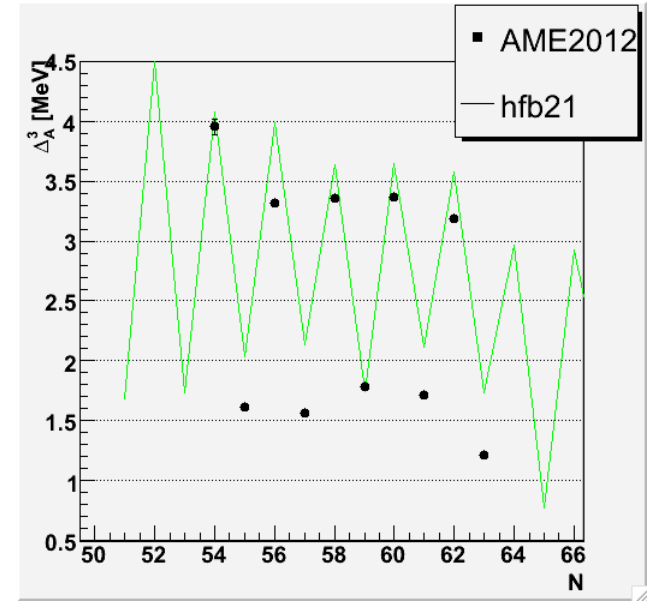
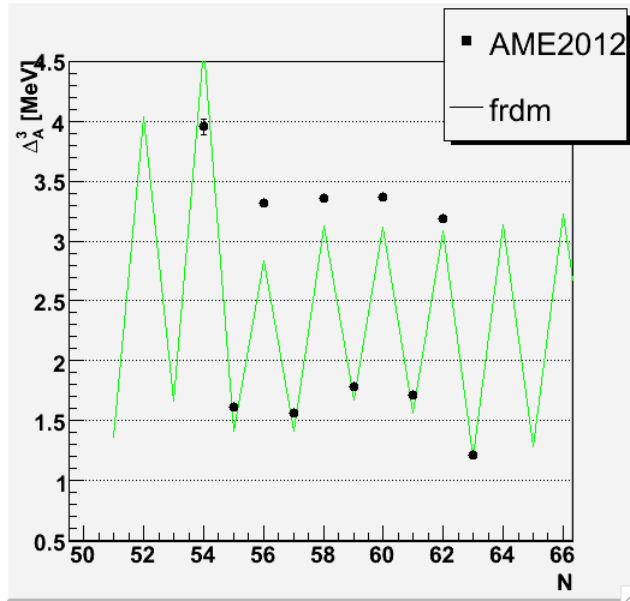
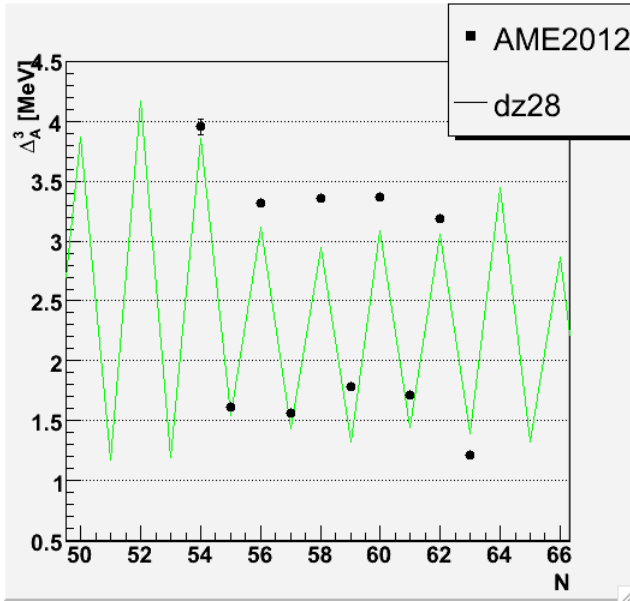


$$\Delta_A^{(5)}(N, Z) = -\frac{(-1)^N}{2} [E(Z+2) - 4E(Z+1) + 6E(Z) - 4E(Z-1) + E(Z-2)]$$

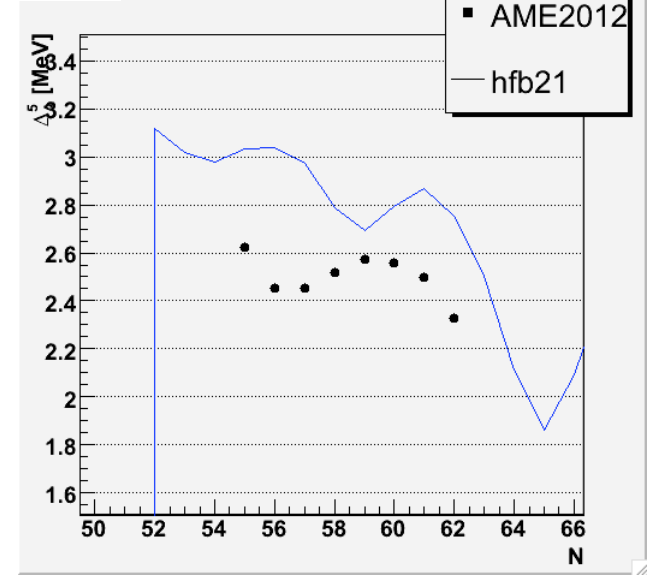
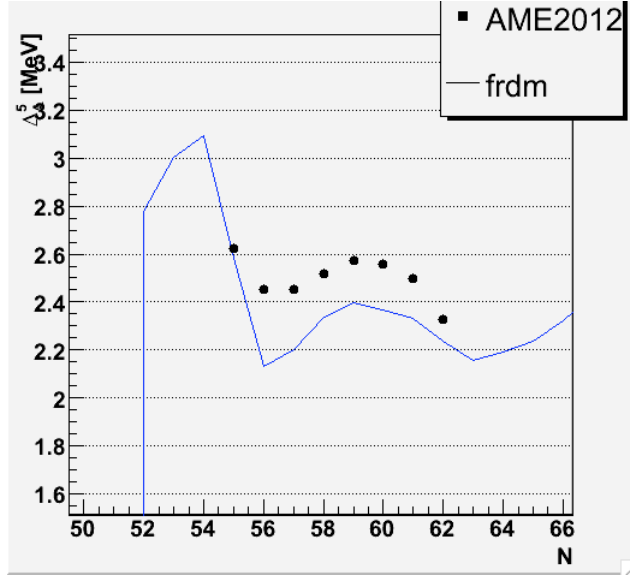
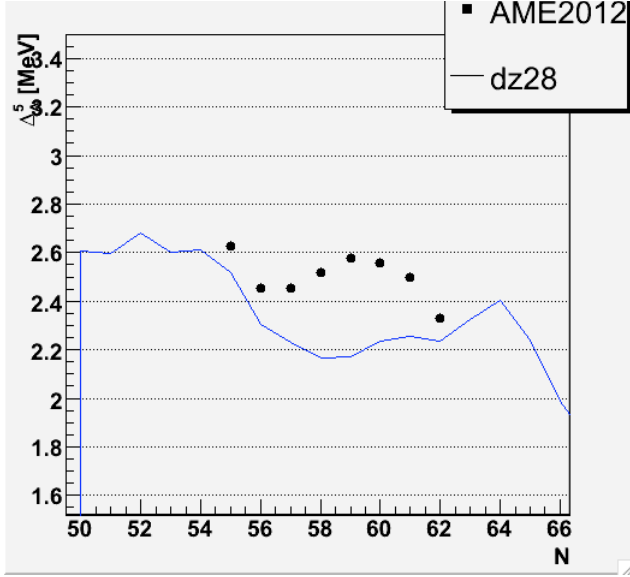


A=104 isobaric chain: 3- and 5-point mass differences

$$\Delta_A^{(3)}(N, Z) = \frac{-1^N}{2} [E(Z+1, N-1) - 2E(Z, N) + E(Z-1, N+1)]$$



$$\Delta_A^{(5)}(N, Z) = -\frac{(-1)^N}{2} [E(Z+2) - 4E(Z+1) + 6E(Z) - 4E(Z-1) + E(Z-2)]$$



RMS versus AME2012

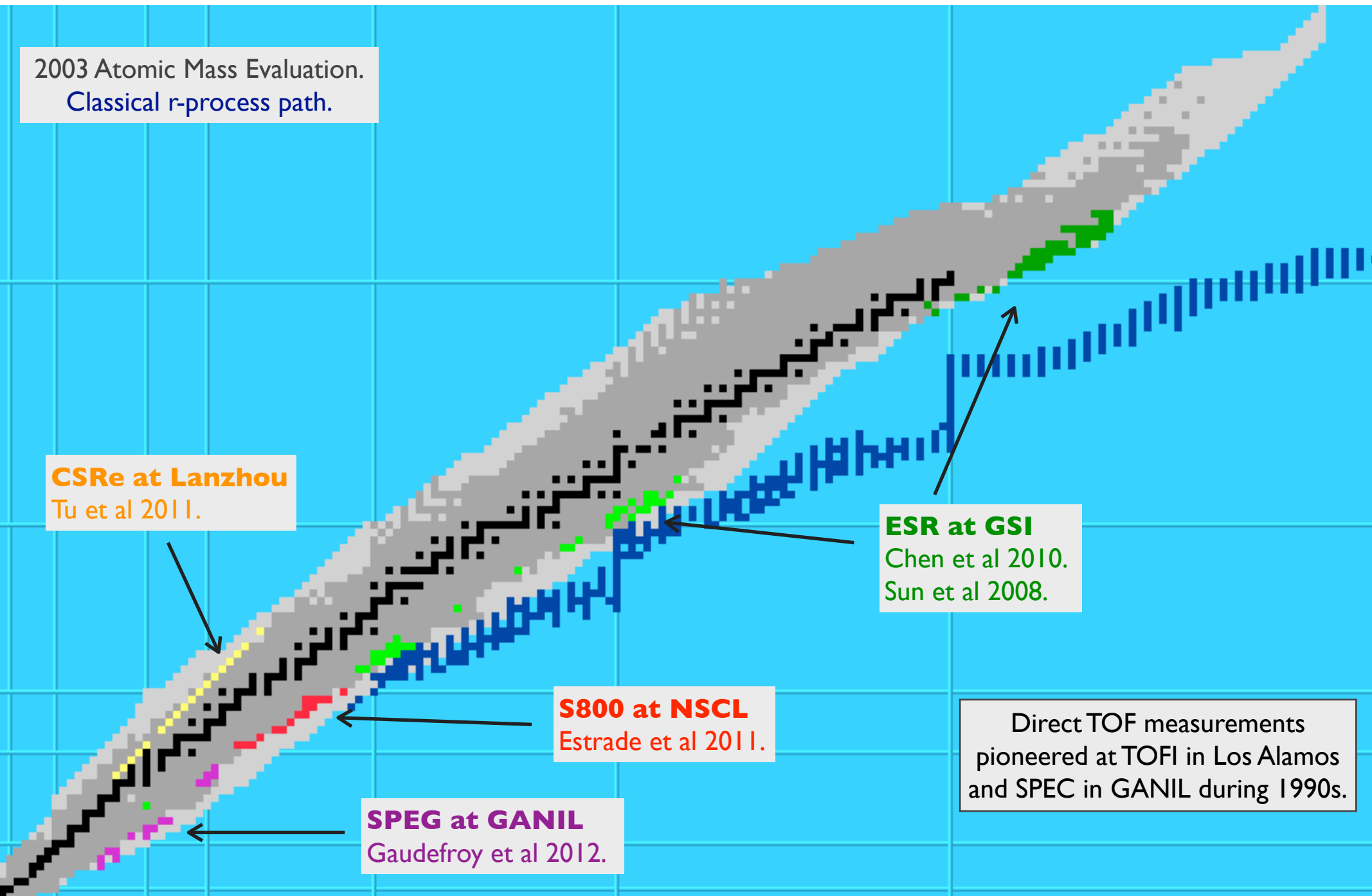
RMS(δ^3) [keV]

| model | 20<A<106 | 20<A(even)<106 | 40<A(even)<68 |
|--------------|----------|----------------|---------------|
| DuZu | 369 | 360 | 331 |
| FRDM | 456 | 445 | 500 |
| HFB21 | 541 | 538 | 609 |
| HFB14 | 692 | 595 | 508 |

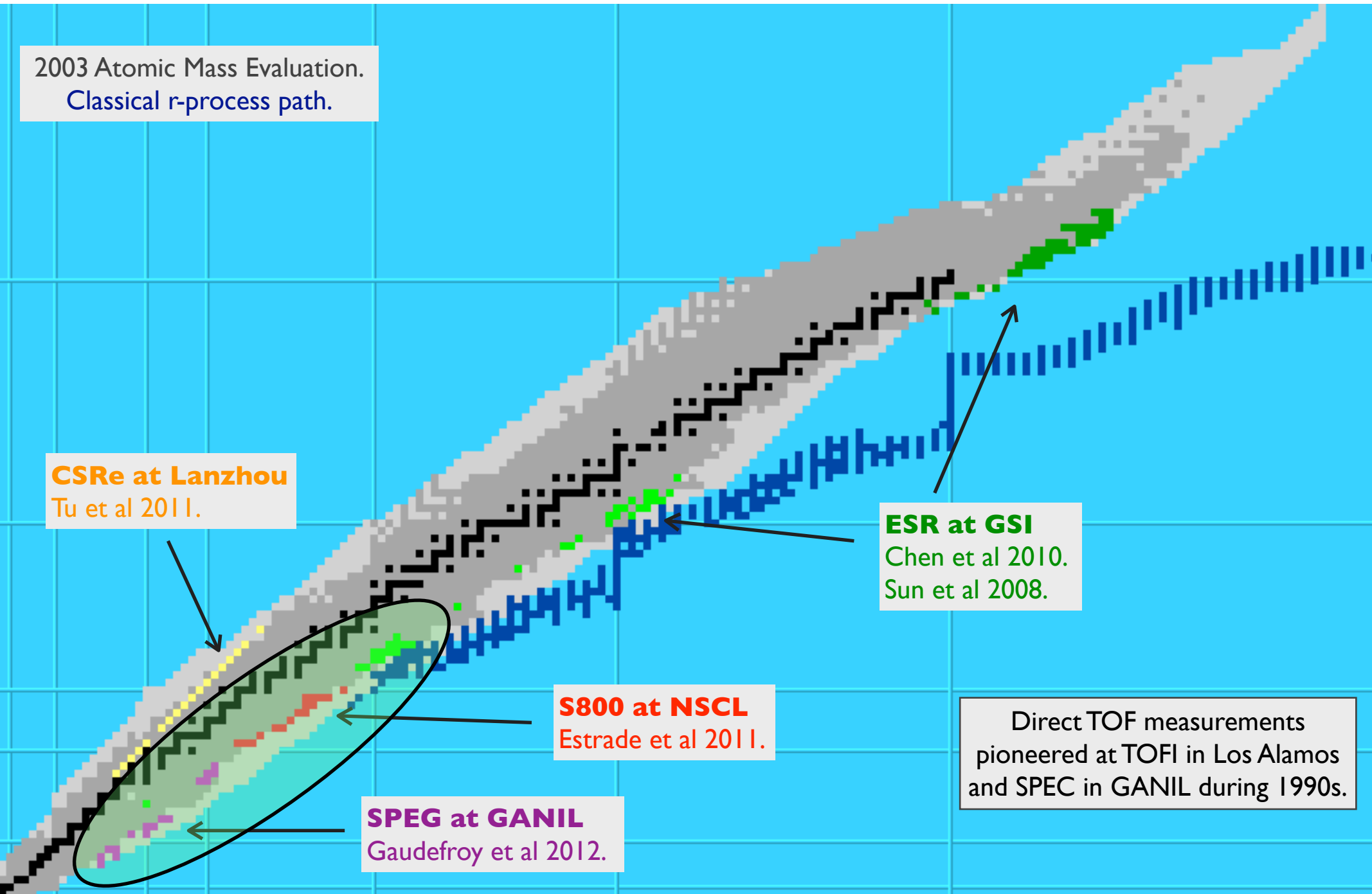
RMS(Q_{ec}) [keV]

| model | 20<A<106 | 20<A(even)<106 | 40<A(even)<68 |
|--------------|----------|----------------|---------------|
| DuZu | 539 | 558 | 601 |
| FRDM | 748 | 780 | 942 |
| HFB21 | 826 | 914 | 1104 |
| HFB14 | 867 | 956 | 933 |

Recent TOF mass measurements

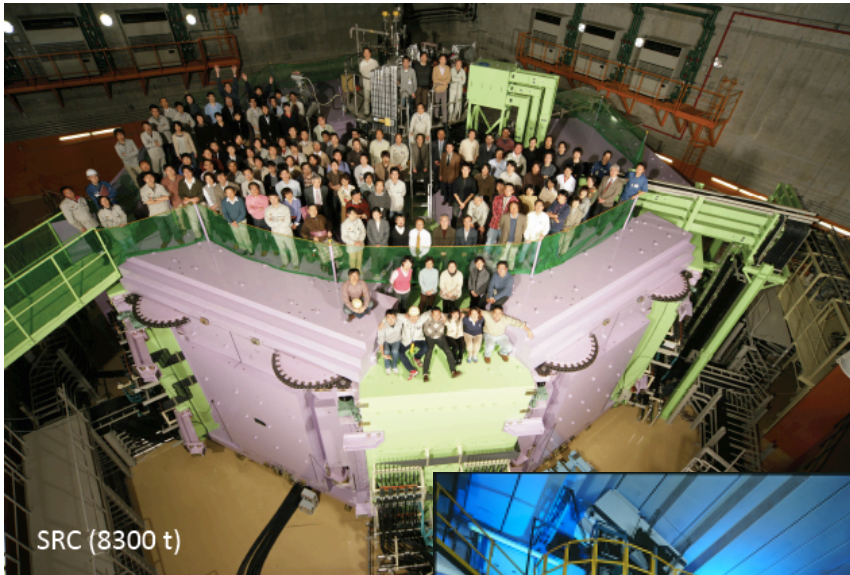


Recent TOF mass measurements



The future is bright for TOF experiments but...

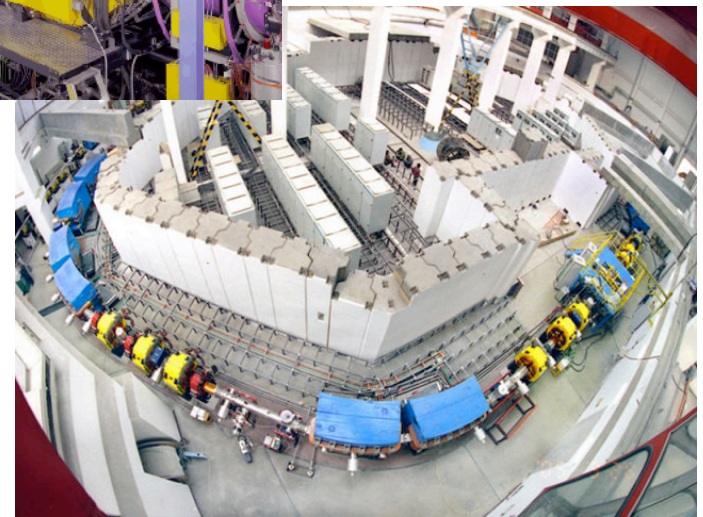
RIBF at RIKEN



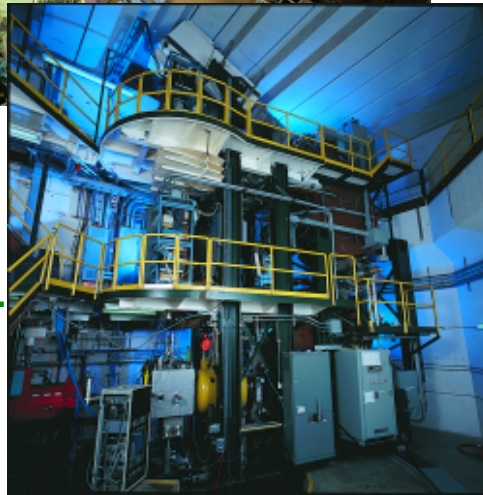
ESR at GSI (FAIR)



CSRe at IMP



S800 at NSCL
(FRIB)



... do we measure relevant along isobaric chains (large abundance/highly excited transitions), or can we constrain models better?

Collaborators

TOF experiment: Hendrik Schatz, Milan Matos, Sebastian George, Zach Meissel, Ana Becerril, Matt Amthor, Daniel Bazin, Thom Elliot, Alexandra Gade, Daniel Galaviz, Rita Lau, Giuseppe Lorusso, Jorge Pereira, Andrew Rogers, Andreas Stolz, John Yurkon (NSCL), Dan Shapira (ORNL), Mark Wallace (LANL), Ed Smith (OSU), Mike Famiano (WMU).

EC simulations: Hendrik Schatz, Ed Brown, Rita Lau (NSCL/MSU), Mary Beard, Michael Wiescher (Notre Dame), Sanjib Gupta, Peter Moller (LANL).