

Finite Volume QFT & QCD Spectroscopy

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Correlation Functions & Lattice QCD

In previous lectures, we have learned about how to describe hadrons with non-perturbative scattering theory, and also how to compute low-energy hadronic observables from non-perturbative QCD, namely lattice QCD.

In these lectures, we will connect the static energy spectrum of hadrons in a box to scattering systems via a non-perturbative mapping, the so-called Lüscher formalism.

Let us first review some aspects of lattice QCD calculations. Lattice QCD is a numerical technique to stochastically estimate QCD correlation functions. To accomplish this, QCD is formulated in a discrete, Euclidean spacetime which is bounded in a finite volume subject to some boundary conditions. Note that both the Euclidean & finite volume nature of spacetime forbids any notion of computing scattering dynamics directly from lattice QCD, since Euclidean spacetimes do not have real time evolutions & finite volumes do not allow free asymptotic states.

However, we can circumvent these issues by examining the behavior of the scaling of finite volume correlators to corresponding infinite volume correlator functions.

Since the LSZ formalism relates QFT correlation functions to scattering amplitudes, these correlators allow one to relate finite-volume energies to scattering.

Essentially,

$$C_L(P) = C_\infty(P) + \delta C_L(P)$$

Diagram illustrating the decomposition of the finite-volume correlator $C_L(P)$ into its infinite-volume counterpart $C_\infty(P)$ and a correction term $\delta C_L(P)$.

- $C_L(P)$: Finite-volume correlator (LGCD)
- $C_\infty(P)$: Infinite-volume correlator (LSZ \Rightarrow Scattering)
- $\delta C_L(P)$: Correction that connects finite volume (FU) and infinite volume (IV)

Our objectives are as follows:

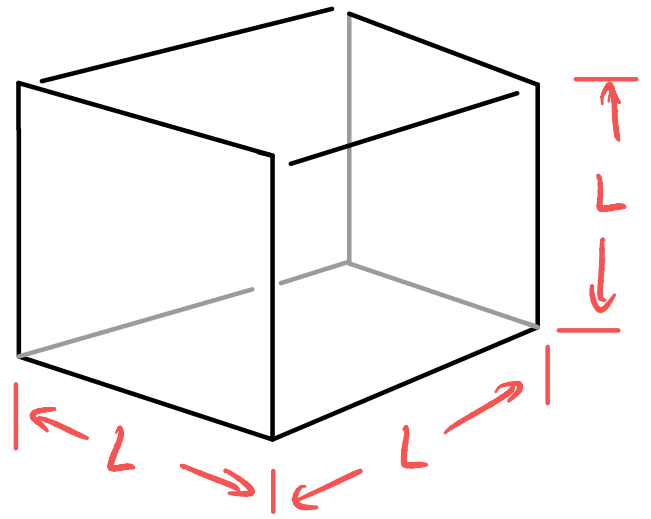
- Examine the spectral rep. of 2-point correlators
- Determine the scaling relation for stable hadrons
- Compute the FU corrections to free two-particle correlators
- Study weakly interacting systems & deriving the Lüscher quantization condition
- Extracting resonances from LGCD

Preliminaries

We work in a continuous 3+1 Minkowski spacetime, $\text{diag } g = (+1, -1, -1, -1)$, which is confined in a cubic volume L^3 & infinite temporal extent, $T \rightarrow \infty$. The spatial volume is subject to periodic boundary conditions (PBCs)

$$x_j = x_j + L \hat{e}_j, \quad j=1,2,3.$$

- The spectrum of a confined system is discrete, E_n , $n \in \mathbb{N}_0$, & since the Hamiltonian is Hermitian, $H^\dagger = H$,



A eigenstate $|E_n, L\rangle$ has real energies, $E_n \in \mathbb{R}$

$$H|E_n, L\rangle = E_n |E_n, L\rangle$$

The energies are independent of the spacetime signature,

$$t \rightarrow -i\tau \quad \Rightarrow \quad e^{-iE_n t} \rightarrow e^{-E_n \tau}$$

- Euclidean time

So, we will work in Minkowski time

- We assume spacetime lattice discretization errors are negligible, thus work with continuous spacetime.
- Most practical spectral calculations use anisotropic lattices, $T \gg L$, so we again assume that finite T effects are negligible, thus $T \rightarrow \infty$.
- Spatial PBCs impose quantization conditions on the momentum. Consider a field operator \mathcal{O} . PBCs enforce

$$\mathcal{O}(t, \vec{x}) = \mathcal{O}(t, \vec{x} + L \hat{e})$$

Introducing the Fourier transform,

$$\mathcal{O}(t, \vec{x}) = \int_{(2\pi)^3} d^3 \vec{p} e^{i\vec{p} \cdot \vec{x}} \tilde{\mathcal{O}}(t, \vec{p}),$$

we find

$$\int_{(2\pi)^3} d^3 \vec{p} e^{i\vec{p} \cdot \vec{x}} \tilde{\mathcal{O}}(t, \vec{p}) = \int_{(2\pi)^3} d^3 \vec{p} e^{i\vec{p} \cdot (\vec{x} + L \hat{e})} \tilde{\mathcal{O}}(t, \vec{p})$$

$$\Rightarrow e^{i\vec{p} \cdot L \hat{e}} = 1 \Rightarrow \vec{p} = \frac{2\pi}{L} \vec{n}, \quad \vec{n} \in \mathbb{Z}^3$$

Note the Fourier transform pair conventions

$$\tilde{f}(p) = \mathcal{F}[f(x)] = \int_{-\infty}^{\infty} dx e^{-ipx} f(x)$$

$$f(x) = \mathcal{F}^{-1}[\tilde{f}(p)] = \frac{1}{L} \sum_p e^{ipx} \tilde{f}(p)$$

Infinite Volume

Finite Volume

$$\int \frac{dp}{2\pi}$$

\longleftrightarrow

$$\frac{1}{L} \sum_p$$

$$2\pi \delta(p'-p)$$

\longleftrightarrow

$$L \delta_{p'p}$$

A key concept we will use in relating IV & FV physics is the Poisson Summation formula, which we write in the form

$$\frac{1}{L^3} \sum_{\vec{p}} f(\vec{p}) = \sum_{\vec{n}} \int_{(2\pi)^3} \frac{d^3 \vec{p}}{(2\pi)^3} e^{iL\vec{n}\cdot\vec{p}} f(\vec{p})$$

The PSF is the key we need in relating
 FV & IV quantities. We define the
 sum-integral operation

$$\frac{1}{L^3} \sum_{\vec{p}} f(\vec{p}) = \left[\frac{1}{L^3} \sum_{\vec{p} \in \frac{2\pi}{L} \mathbb{Z}^3} - \int \frac{d^3 \vec{p}}{(2\pi)^3} \right] f(\vec{p})$$

By the PSF, we have

$$\left[\frac{1}{L^3} \sum_{\vec{p}} - \int \frac{d^3 \vec{p}}{(2\pi)^3} \right] f(\vec{p}) = \sum_{\vec{n} \neq \vec{0}} \int \frac{d^3 \vec{p}}{(2\pi)^3} e^{i \vec{p} \cdot \vec{n} L} f(\vec{p})$$

Suppose $f(\vec{p})$ is a smooth function in \vec{p} (physical)

$$f(\vec{p}) = \sqrt{4\pi} \sum_{l=0}^{\infty} \sum_{m_l=-l}^l Y_{lm_l}(\hat{p}) f_{lm_l}(\rho)$$

↑ spherical harmonics

Recall plane-wave expansion

$$e^{i \vec{p} \cdot \vec{n} L} = 4\pi \sum_{l=0}^{\infty} i^l j_l(n\rho L) \sum_{m_l=-l}^l Y_{lm_l}(\hat{n}) Y_{lm_l}^*(\hat{p})$$

↑ spherical Bessel functions

Inserting expansions into sum-integral, & converting the measure to spherical coordinates,

$$\begin{aligned}
 \frac{1}{L^3} \sum_{\vec{n} \neq \vec{0}} \int_{\vec{p}} f(\vec{p}) &= \sum_{\vec{n} \neq \vec{0}} \int d\hat{p} \int_0^\infty \frac{dp p^2}{(2\pi)^3} e^{i\vec{p} \cdot \vec{n} L} f(\vec{p}) \\
 &= \sum_{\vec{n} \neq \vec{0}} \frac{(4\pi)^{3/2}}{(2\pi)^3} \sum_{l, m_l} \sum_{l', m_{l'}} i^{l'} \int_0^\infty dp p^2 j_{l'}(pnL) f_{l'm_{l'}}(p) \\
 &\quad \times \underbrace{Y_{l'm_{l'}}(\hat{n}) \int d\hat{p} Y_{l'm_{l'}}^*(\hat{p}) Y_{lm}(\hat{p})}_{\delta_{ll'} \delta_{m_l m_{l'}}} \\
 &= \sum_{\vec{n} \neq \vec{0}} \pi^{-3/2} \sum_{l, m} i^l Y_{lm}(\hat{n}) \int_0^\infty dp p^2 j_l(pnL) f_{lm}(p)
 \end{aligned}$$

Consider $l=0$ mode only, $j_0(pnL) = \frac{\sin pnL}{pnL}$

$$\Rightarrow \frac{1}{L^3} \sum_{\vec{p}} f(\vec{p}) = \frac{1}{2\pi^2} \sum_{\vec{n} \neq \vec{0}} \frac{1}{nL} \int_0^\infty dp p \sin(pnL) f_{00}(p)$$

Consider a function like $f_\infty(p) = \frac{1}{p^2 + \Lambda^2}$, $\Lambda > 0$



then,

$$I = \int_0^{\infty} dp \, p \sin(pnL) \frac{1}{p^2 + \Lambda^2}$$

← integrate even under $p \rightarrow -p$

$$= \frac{1}{2} \int_{-\infty}^{\infty} dp \frac{p \sin(pnL)}{p^2 + \Lambda^2} = \frac{1}{2} \text{Im} \int_{-\infty}^{\infty} dp \frac{p}{p^2 + \Lambda^2} e^{ipnL}$$

$$\hookrightarrow \sin x = \text{Im} e^{ix}$$

Consider Cauchy theorem: $\oint_{\gamma} dz f(z) = 0$

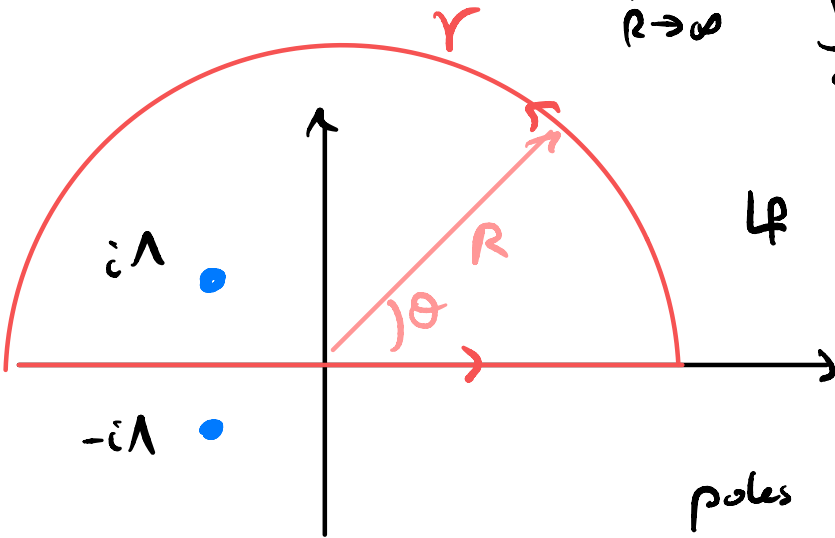
if $f(z)$ analytic in domain bounded by γ

$$\oint_{\gamma} dp \frac{p}{p^2 + \Lambda^2} e^{ipnL} = \int_{-\infty}^{\infty} dp \frac{p}{p^2 + \Lambda^2} e^{ipnL}$$

$$p = R e^{i\theta}$$

$$+ \lim_{R \rightarrow \infty} R i \int_0^{\pi} d\theta e^{i\theta} \frac{R e^{i\theta}}{R^2 e^{i2\theta} + \Lambda^2} e^{ipnL}$$

Vanishes $\rightarrow R \rightarrow \infty$



poles \downarrow

$$p^2 + \Lambda^2 = (p + i\Lambda)(p - i\Lambda) = 0$$

By residue theorem,

$$\begin{aligned} \oint_{\gamma} dp \frac{p}{p^2 + \Lambda^2} e^{ipnL} &= 2\pi i \operatorname{Res}[p = i\Lambda] \\ &= 2\pi i \left(\frac{p}{p + i\Lambda} e^{ipnL} \right) \Big|_{p = i\Lambda} \\ &= 2\pi i \left(\frac{i\Lambda}{2i\Lambda} e^{-n\Lambda L} \right) \\ &= \pi i e^{-n\Lambda L} \end{aligned}$$

So, we find the integral I is

$$\begin{aligned} I &= \frac{1}{2} \operatorname{Im} \int_{-\infty}^{\infty} dp \frac{p}{p^2 + \Lambda^2} e^{ipnL} \\ &= \frac{1}{2} \operatorname{Im} \left(\pi i e^{-n\Lambda L} \right) = \frac{\pi}{2} e^{-n\Lambda L} \end{aligned}$$

Therefore, sum-integral difference (for our example) is

$$\begin{aligned} \frac{1}{L^3} \sum_{\vec{p}} f(\vec{p}) &= \frac{1}{2\pi^2} \sum_{\vec{n} \neq \vec{0}} \frac{1}{nL} I \\ &= \frac{1}{4\pi} \sum_{\vec{n} \neq \vec{0}} \frac{1}{nL} e^{-n\Lambda L} \end{aligned}$$

exponentially suppressed scaling in box volume L

This illustrates an important property, we will exploit.

Let $\frac{1}{L^3} \sum_{\vec{p}} f(\vec{p})$ be the FV quantity of interest.

then,

$$\frac{1}{L^3} \sum_{\vec{p}} f(\vec{p}) = \int_{\left(\frac{2\pi}{L}\right)^3} d^3\vec{p} f(\vec{p}) + \frac{1}{L^3} \int_{\vec{p}=0} f(\vec{p})$$

this is FV correction
to IV term

if $f(\vec{p})$ is smooth, then by PSF

$$\frac{1}{L^3} \int_{\vec{p}} f(\vec{p}) \sim e^{-mL}$$

$m = \text{lightest mass scale in theory}$
i.e., $m = m_\pi$ in QCD.

If $f(\vec{p})$ has some non-analytic behavior, e.g., poles/cuts,
then this need not be so, & it would be correction
scales like $1/L^3$.

A comment about regularization. Recall in QFTs, one often needs to introduce a regularization scheme to tame UV divergences. This is true for FV quantities as well, e.g.,

$$\frac{1}{L^3} \sum_{\vec{p}} f(\vec{p}) \rightarrow \frac{1}{L^3} \sum_{\vec{p}} f^R(\vec{p})$$

← same regularization scheme
e.g., Pauli-Villars

Looking at the correction, i.e., the sum-integral difference

$$\begin{aligned} \frac{1}{L^3} \sum_{\vec{p}} f^R(\vec{p}) &= \frac{1}{L^3} \sum_{\vec{p}} f^R(\vec{p}) - \int \frac{d^3\vec{p}}{(2\pi)^3} f^R(\vec{p}) \\ &= \sum_{\vec{n} \neq \vec{0}} \int \frac{d^3\vec{p}}{(2\pi)^3} e^{i\vec{p} \cdot \vec{n} L} f^R(\vec{p}) \end{aligned}$$

But, notice that as $\vec{p} \rightarrow \infty$

$$\frac{1}{L^3} \sum_{\vec{p}} f^R(\vec{p}) - \int \frac{d^3\vec{p}}{(2\pi)^3} f^R(\vec{p}) \sim \Lambda - \Lambda \rightarrow 0$$

↑ ↑
cutoffs characterizing divergence

⇒ the FV correction is independent of the regularization scheme!

Spectral Decomposition

The primary objects of interest in QCD calculations are correlation functions, or correlators. Here we focus on 2-point correlators,

$$C_L(x) = \langle T O(x) O^\dagger(0) \rangle$$

$x = (t, \vec{x})$

↑ finite volume ↑ time-ordering operator ↑ some scalar operator (simplicity)

Consider Fourier transform,

$$C_L(P) = \int_L d^4x e^{iP \cdot x} C_L(x)$$

$P = (E, \vec{P}) \quad \uparrow$

$$= \int_L d^4x e^{iP \cdot x} \langle T O(x) O^\dagger(0) \rangle$$

Let us first consider $t > 0$ time ordering

$$T O(x) O^\dagger(0) = \underline{O(x) O^\dagger(0) \theta(t)} + O^\dagger(0) O(x) \theta(-t)$$

So,

$$C_L(P) = \int_0^\infty dt \int_L d^3\vec{x} e^{iP \cdot x} \langle O(x) O^\dagger(0) \rangle + (t < 0)$$

Next, we insert a complete set of energy eigenstates

$$\hat{H} |E_n, \vec{P}, L\rangle = E_n(\vec{P}, L) |E_n, \vec{P}, L\rangle$$

so, with normalization $\langle E_n, \vec{P}', L | E_n, \vec{P}, L \rangle = \delta_{nn} \delta_{\vec{P}'\vec{P}}$
the resolution of identity is

$$1 = \sum_n \sum_{\vec{P}} |E_n, \vec{P}, L\rangle \langle E_n, \vec{P}, L|$$

We will suppress all other quantum numbers.

so, find

$$C_L(P) = \sum_n \sum_{\vec{P}'} \int_0^\infty dt \int_V d^3\vec{x} e^{i\vec{P}\cdot\vec{x}} \langle 0 | \mathcal{O}(x) | E_n, \vec{P}', L \rangle \langle E_n, \vec{P}', L | \mathcal{O}^\dagger(0) | 0 \rangle$$

Under spacetime translations, $\mathcal{O}(x) = e^{i\hat{P}\cdot x} \mathcal{O}(0) e^{-i\hat{P}\cdot x}$

$$\text{with } \hat{P} = (\hat{H}, \hat{\vec{P}})$$

To ensure convergence in $t \rightarrow \infty$, we let $\hat{H} \rightarrow \hat{H} - i\epsilon$
with $\epsilon \rightarrow 0^+$ implicit. Then,

$$e^{-i\hat{P}\cdot x} |E_n, \vec{P}', L\rangle = e^{-i(E_n - i\epsilon)t} e^{i\vec{P}'\cdot\vec{x}} |E_n, \vec{P}', L\rangle$$

We find,

$$C_L(P) = \sum_n \sum_{\vec{p}'} \int_0^{\infty} dt e^{i(E-E_n+ie)t} \times \int_L d^3\vec{x} e^{-i(\vec{p}-\vec{p}')\cdot\vec{x}} |z_n(\vec{p}',L)|^2 + (t<0)$$

where $z_n(\vec{p},L) = \langle 0 | \mathcal{O}(0) | E_n, \vec{p}, L \rangle$

Now, $\int_L d^3\vec{x} e^{-i(\vec{p}-\vec{p}')\cdot\vec{x}} = L^3 \delta_{\vec{p}\vec{p}'}$

$$\begin{aligned} \& \int_0^{\infty} dt e^{i(E-E_n+ie)t} &= \frac{e^{i(E-E_n+ie)t}}{i(E-E_n+ie)} \Big|_0^{\infty} \\ &= \frac{i}{E-E_n+ie} \end{aligned}$$

So, we find

$$C_L(P) = L^3 \sum_n \frac{i}{E-E_n+ie} |z_n(\vec{p},L)|^2 + (t<0)$$

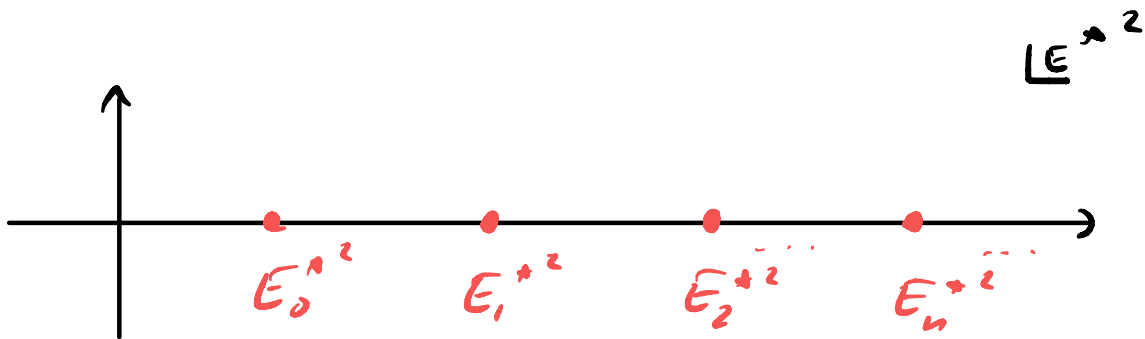
Considering the $t<0$ term gives a pole of the form

$$E + E_n - i\epsilon$$

$$\Rightarrow C_L(P) = \sum_n 2E_n L^3 \frac{i}{E^2 - E_n^2 + i\epsilon} |z_n(\vec{p},L)|^2$$

So, the analytic structure of FV correlators is a sequence of poles in E^2 (or E). Recall the center-of-momentum (CM) frame, $\vec{P}^A = \vec{0}$

$$\Rightarrow E^{*2} = E^2 - \vec{P}^2 = P^2$$



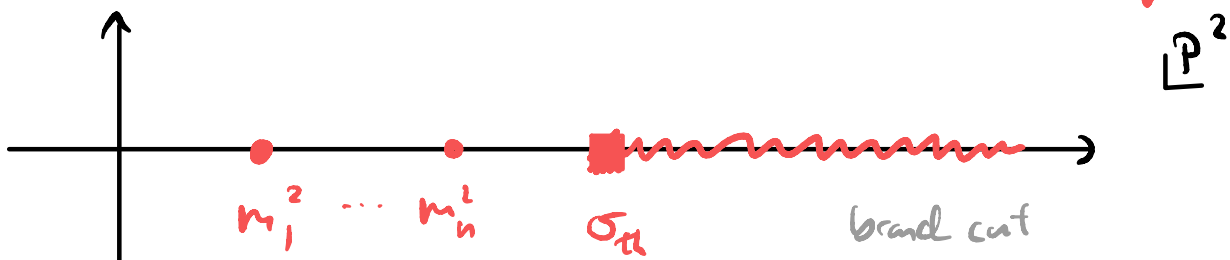
For IV correlators, we can follow the same idea, & derive the Källén-Lehman spectral representation,

$$C_{\infty}(P) = \int d^4x e^{iP \cdot x} \langle T \mathcal{O}(x) \mathcal{O}^\dagger(0) \rangle$$

$$= \sum_{j=0}^n \frac{i z_j}{P^2 - m_j^2} + \int_{\sigma_{th}}^{\infty} \frac{d\sigma}{2i\pi} \frac{\rho(\sigma)}{P^2 - \sigma + i\epsilon}$$

↑ possible stable/bound states

↑ scattering continuum



The analytic structure is different for IV correlators & FV correlators. To FV correlation is

$$C_L(P) = C_{\text{lo}}(P) + \delta C_L(P)$$

↑ ↑ ↑
poles cuts poles & cuts

By LSZ theorem, can get access to amplitudes iM in C_{lo} . From lattice QCD, get access to E_n . Thus, can relate $E_n \leftrightarrow iM$ via FV correlation $\delta C_L(P)$. To do so, we will use an all-orders approach in QFT to construct a diagrammatic representation for these objects. We will show that the results are general, independent on any particular QFT.

Single Particles

We will first examine FV corrections to single particle states. To be concrete, consider real scalar ϕ^4 theory,

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 - \frac{\lambda}{4!} \phi^4$$

↙ mass parameter (not physics)

Our results are generally QFT independent, but we use a particular one, as a generalized EFT, as a catalyst.

Recall the Feynman diagram expansion for the fully dressed propagator. First, ∞ -volume

$$\begin{aligned} C_\infty(P) &= \text{---} \bullet \text{---} \\ &= \text{---} + \text{---} \bigcirc \text{---} + \text{---} \bigcirc \text{---} \bigcirc \text{---} + \dots \\ &= \text{---} + \text{---} \bigcirc \bullet \text{---} \end{aligned}$$

Can show, ↑ self-energy $i\Pi(P^2)$ (1PI)

$$C_\infty(P) = \frac{i}{P^2 - m^2 + \Pi(P^2)}$$

The physical mass is given by

$$C_{\infty}(P^2) = \frac{i}{P^2 - m_{\text{phys}}^2 + i\epsilon} + iS(P^2)$$

↑ regular about $P^2 = m^2$

with

$$m_{\text{phys}}^2 - m^2 + \Pi(m_{\text{phys}}^2) = 0$$

The self energy is of the form,

$$i\Pi(P^2) = \underbrace{\text{loop}}_{\mathcal{O}(\lambda)} + \underbrace{\text{two-loop}}_{\mathcal{O}(\lambda^2)} + \text{three-loop} + \dots$$

At $\mathcal{O}(\lambda)$,

$$i\Pi(P^2) = -i\lambda \int \frac{d^4k}{(2\pi)^4} \frac{i}{k^2 - m^2 + i\epsilon} + \mathcal{O}(\lambda^2)$$

↑ divergent, needs some regulator

... then renormalize, blah blah...

example: Pauli-Villars,


$$\frac{1}{k^2 - m^2 + i\epsilon} \rightarrow \frac{1}{k^2 - m^2 + i\epsilon} + \sum_j \frac{a_j}{k^2 - \Lambda_j^2 + i\epsilon}$$

w/ a_j to be determined & $\Lambda_j \rightarrow \infty$ at end

↳ to keep degree of divergence low

Repeating the same exercise on the FV, we find

$$C_L(P) = \frac{i}{P^2 - m^2 + i\pi_L(P^2)}$$

with $i\pi_L(P^2) =$  + ...

↑ FV loops

$$= -i\lambda \frac{1}{L^3} \sum_{\vec{k}} \int_{-\infty}^{\infty} \frac{dk^0}{2\pi} \frac{i}{k^2 - k^2 + i\epsilon} + \mathcal{O}(\lambda^2)$$

The FV pole position is $P^2 = m_L^2$ (NOT necessarily m_{phys}^2 !)

$$m_L^2 - m^2 + \pi_L(m_L^2) = 0$$

$$m_L = E_{n=0}^+(L)$$

Now, take $m_L^2 - m_{phys}^2$,

$$\delta m_L^2 = m_L^2 - m_{phys}^2 = - \left[\pi_L(m_L^2) - \pi(m_{phys}^2) \right]$$

LQ us assume $\frac{\delta m_L^2}{m_L} \ll 1 \Rightarrow \pi_L(m_L^2) \approx \pi_L(m_{phys}^2)$

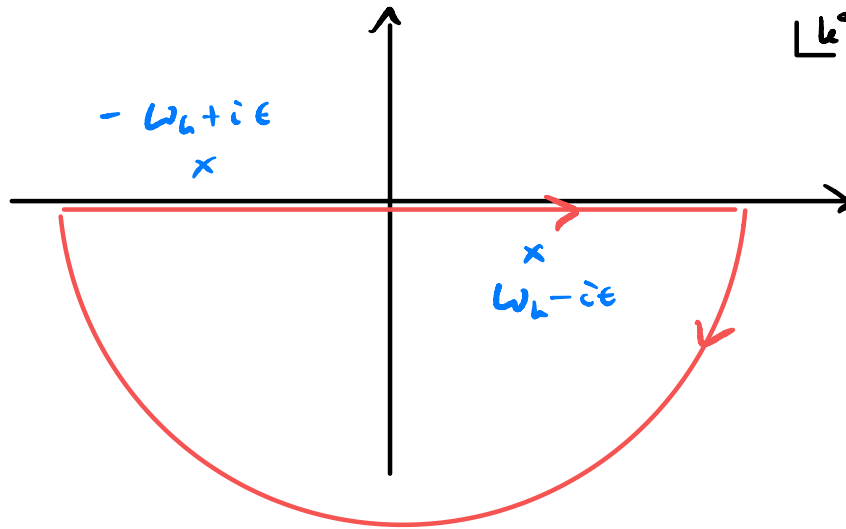
$$\text{So, } \delta m_L^2 = - \left[\Pi_L(m_{pl,1}^2) - \Pi(m_{pl,1}^2) \right]$$

$$= + i\lambda \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 \vec{k}}{(2\pi)^3} \right] \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{i}{\omega^2 - k^2 + i\epsilon}$$

↓
going to implicitly assume regularization
& renormalization program

Let's first do temporal integral

$$\int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{i}{\omega^2 - k^2 + i\epsilon} = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{i}{(\omega - \omega_k + i\epsilon)(\omega + \omega_k - i\epsilon)} = \frac{1}{2\omega_k}$$



$$\omega_k = \sqrt{k^2 + m^2}$$

$$\Rightarrow \delta m_L^2 = i\lambda \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 \vec{k}}{(2\pi)^3} \right] \frac{1}{2\sqrt{k^2 + m^2}}$$

Can show, via PSF, that

↑ smooth function in integration region!

$$\delta m_L^2 = \frac{\lambda}{2(2\pi mL)^{3/2}} e^{-mL} + \mathcal{O}(e^{-\sqrt{2}mL})$$

so, we find that

$$m_L^2 = m_{\text{phys}}^2 + \mathcal{O}(e^{-m_{\text{phys}} L})$$

In QCD, the lightest mass scale is the pion, m_π

so, the FV correction to a hadron mass m_h is

$$E_{n=0}^*(L) = m_h + \mathcal{O}(e^{-m_\pi L})$$

In practical calculations, want to keep L sufficiently large to suppress these effects, e.g., $m_\pi L \sim 4$ gives $\sim 1\%$ correction, which we hope to ignore.

The energy of the lowest state in a generic moving frame is simply

$$\begin{aligned} E_0(\vec{P}, L) &= \sqrt{m_h^2 + \vec{P}^2} + \mathcal{O}(e^{-m_\pi L}) \\ &= \sqrt{m_h^2 + \left(\frac{2\pi\vec{n}}{L}\right)^2} + \mathcal{O}(e^{-m_\pi L}) \end{aligned}$$

Two Particle Systems

In Describing excited states, we must address that most hadrons are resonances of scattering processes. Thus, to rigorously describe excited states, we must access the scattering amplitude with LQCD.

Let us first consider non-interacting two particles

Non-Interacting Two-Particle Spectrum

Let's again consider real scalar ϕ^4 theory. We construct a local operator as

$$O(x) = \xi \int_L d^4y \int_L d^4z A(y,z) \phi(x+y) \phi(x+z)$$

↑
symmetry factor, $\xi = \frac{1}{2!}$

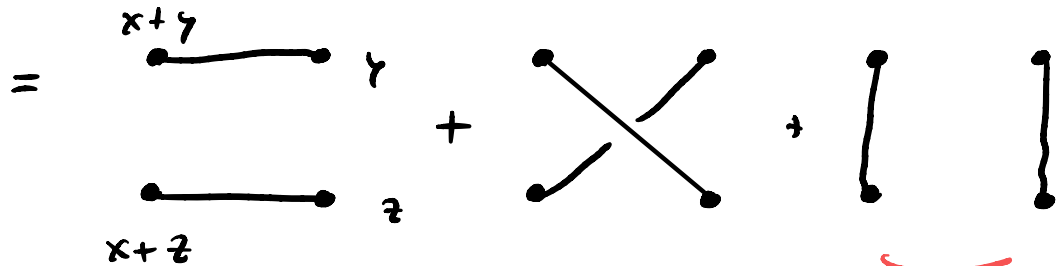
↑
same local function associated by wave function of 2-particle state created by $O(x)$

So,

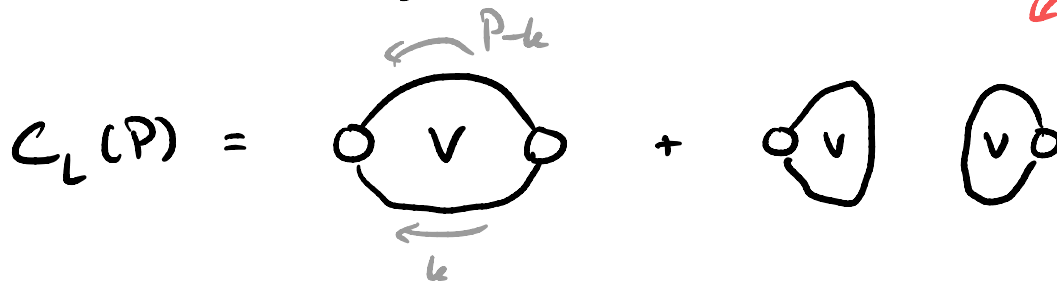
$$C_L(p) = \int_L d^4x e^{i p \cdot x} \langle T O(x) O^\dagger(0) \rangle$$
$$= \xi^2 \int_L d^4x e^{i p \cdot x} \int_L d^4y' \int_L d^4z' \int_L d^4y \int_L d^4z A(y',z') A(y,z)$$
$$\times \langle T \phi(x+y') \phi(x+z') \phi(y) \phi(z) \rangle$$

Diagrammatically,

$$\langle T \varphi(x+y) \varphi(x+z) \varphi(y) \varphi(z) \rangle$$



So, we write "generalized Feynman rules"



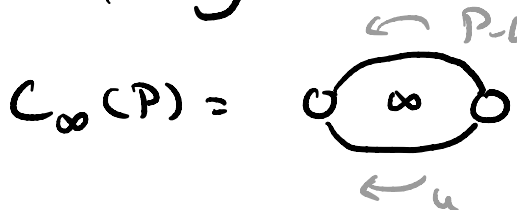
Discarded
Does not contribute (Drop)

$$= \mathbb{Z} \frac{1}{L^3} \sum_{\vec{h}} \int_{-\infty}^{\infty} \frac{d\bar{h}^0}{2\pi} A(\vec{h}, P) \frac{i}{h^2 - m^2 + i\epsilon} \frac{i}{(P-h)^2 - m^2 + i\epsilon} A^*(\vec{h}, P)$$

with

$$A(\vec{h}, P) = \int_L d^4 y \int_L d^4 z e^{i\vec{h} \cdot \vec{y}} e^{i(P-\vec{h}) \cdot \vec{z}} A(\vec{y}, \vec{z})$$

The corresponding IV correlator is



$$= \mathbb{Z} \int \frac{d^4 h}{(2\pi)^4} A(\vec{h}, P) \frac{i}{h^2 - m^2 + i\epsilon} \frac{i}{(P-h)^2 - m^2 + i\epsilon} A^*(\vec{h}, P)$$

Let's examine the spectrum via FV contour

$$\delta C_L(P) = C_L(P) - C_\infty(P)$$

$$= i \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 \vec{k}}{(2\pi)^3} \right] \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} A(\vec{k}, P) \frac{i}{\omega^2 - k^2 + i\epsilon} \frac{i}{(P-\vec{k})^2 - \omega^2 + i\epsilon} A^*(\vec{k}, P)$$

We will again assume some UV regulator & assume $A(\vec{k}, P)$ is sufficiently smooth in k , i.e., non-singular.

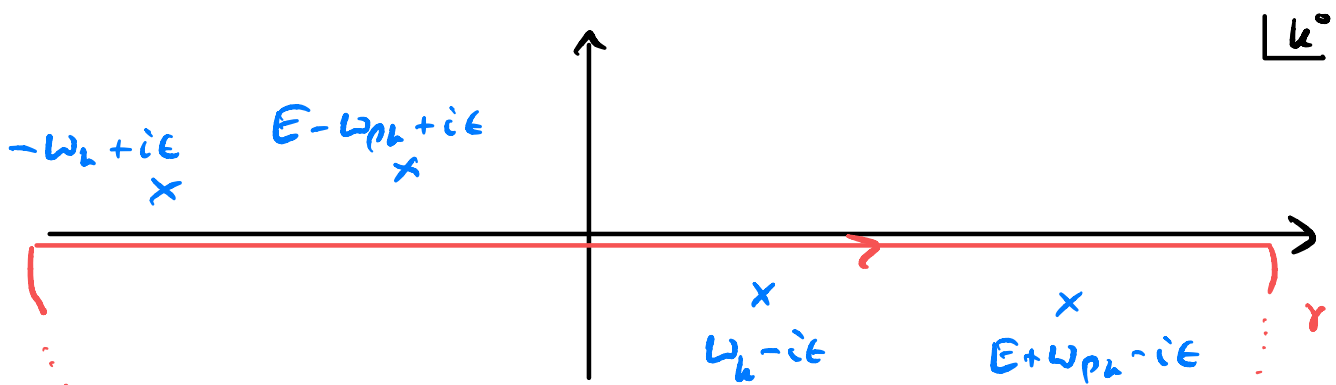
Let's perform ω integral. Poles at

$$(\omega^2 - k^2 + i\epsilon)(\omega^2 - (P-\vec{k})^2 + i\epsilon) = 0$$

$$\Rightarrow (\omega - \omega_k + i\epsilon)(\omega + \omega_k - i\epsilon)(E - \omega - \omega_{P-k} + i\epsilon)(E - \omega + \omega_{P-k} - i\epsilon)$$

$$\omega / \omega_k = \sqrt{k^2 + \vec{k}^2}$$

$$\omega_{P-k} = \sqrt{m^2 + |\vec{P} - \vec{k}|^2}$$



Assuming integrand is well-behaved at $k \rightarrow \infty$,
we have

$$\int_{-\infty}^{\infty} dk^0 f(k^0) = \oint_{\Gamma} dk^0 f(k^0) = \sum_n \text{Res}[f(k^0); k^0 = k_n^0]$$

We pick up two poles, finding

$$\delta C_L(P) = \frac{1}{L^3} \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 \vec{k}}{(2\pi)^3} \right] A(\omega_k, \vec{k}, P) \frac{i}{2\omega_k (E - \omega_k)^2 - \omega_{pk}^2 + i\epsilon} A^*(\omega_k, \vec{k}, P) \\ + \text{term w/ } k^0 = E + \omega_{pk} - i\epsilon$$

For physical energies, $P^2 > 0$, $E^* > 2m$

We find (exercise) that the second term does not
have any singularities in physical E^* region

\Rightarrow FV correction is of $\mathcal{O}(e^{-mL})$

Further,

$$\frac{1}{(E - \omega_k)^2 - \omega_{pk}^2 + i\epsilon} = \frac{1}{2\omega_{pk} (E - \omega_k - \omega_{pk} + i\epsilon)} \quad \leftarrow \text{Singular in } E^*$$

$$+ \frac{1}{2\omega_{pk} (E - \omega_k + \omega_{pk} - i\epsilon)}$$

\uparrow
non-singular in E^*

So, we have

$$\delta C_L(P) = i \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 \vec{k}}{(2\pi)^3} \right] \frac{A(\omega_k, \vec{k}, P) i A^*(\omega_k, \vec{k}, P)}{2\omega_k 2\omega_{pk} (E - \omega_k - \omega_{pk} + i\epsilon)} + \mathcal{O}(e^{-\kappa L})$$

Now, introduce partial waves for $A(\omega_k, \vec{k}, P)$. How?

So far, fixed 1 & 2 particles on mass-shell. Fixing the second gives (in part CM frame)

$$\vec{k}^* = \xi^* \hat{k}^*$$

where

$$\xi^* = \frac{1}{2} \sqrt{E^{*2} - 4m^2}, \quad E^* = \sqrt{P^2}$$

this point is equivalent to where $E = \omega_k + \omega_{pk}$

so, expand about this point

$$\begin{aligned} A(\vec{k}^*, P) &= A(\xi^* \hat{k}^*, P) + [A(\vec{k}^*, P) - A(\xi^* \hat{k}^*, P)] \\ &= A(\xi^* \hat{k}^*, P) + (k^{*2} - \xi^{*2}) \frac{\partial}{\partial k^{*2}} A(k^*, P) \Big|_{k^{*2} = \xi^{*2}} \\ &\quad + \mathcal{O}((k^{*2} - \xi^{*2})^2) \end{aligned}$$

$$\equiv A(\xi^* \hat{k}^*, P) + \delta A(k^*, P)$$

Terms by one or more $\delta A(\hat{k}^+, P)$ we suppressed as e^{-hL} like before since there is no singularity

$$\Rightarrow \delta C_L(P) = \mathcal{Z} \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 \vec{k}}{(2\pi)^3} \right] \frac{A(\hat{q}^+ \hat{k}^+, P) \dot{=} A^*(\hat{q}^+ \hat{k}^+, P)}{2\omega_k 2\omega_{p_k} (E - \omega_k - \omega_{p_k} + i\epsilon)} + \mathcal{O}(e^{-hL})$$

Can now PW expand

$$A(\hat{q}^+ \hat{k}^+, P) = \sqrt{4\pi} \sum_{\ell, m_\ell} A_{\ell, m_\ell}(P) Y_{\ell, m_\ell}^*(\hat{k}^+) \left(\frac{k^+}{q^+} \right)^\ell$$

Basis factors to regulate artificial singularity of $Y_{\ell, m}$ induced by expansion

So,

$$\delta C_L(P) \equiv \sum_{\ell', m'} \sum_{\ell, m_\ell} A_{\ell', m'}(P) \dot{=} F_{\ell', m'; \ell, m_\ell}(P, L) A_{\ell, m_\ell}(P) + \mathcal{O}(e^{-hL})$$

where F is purely geometric & kinematic

$$F_{\ell', m'; \ell, m_\ell}(P, L) \equiv \mathcal{Z} \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 \vec{k}}{(2\pi)^3} \right] \left(\frac{k^+}{q^+} \right)^{\ell'} \frac{Y_{\ell', m'}^*(\hat{k}^+) Y_{\ell, m_\ell}(\hat{k}^+)}{2\omega_k 2\omega_{p_k} (E - \omega_k - \omega_{p_k} + i\epsilon)} \left(\frac{k^+}{q^+} \right)^\ell$$

The summand/integrand has a pole that is physically associated with on-shell 2-particle states. We cannot use PSF as before, so the FV correction scales like $1/L^3$.

The FV function $F(P, L)$ characterizes the distortions induced by the periodic volume. It contains information on both the FV energies & IV continuum. To see this, first consider the imaginary part of $F(P, L)$, this comes only from the integral term,

$$\text{Im} F_{\ell' \ell, \ell \ell'}(P, L) = \int \frac{d\hat{h}}{(2\pi)^3} \frac{Y_{\ell' \ell'}^*(\hat{h}^*) Y_{\ell \ell}(\hat{h}^*)}{2\omega_{p\ell} 2\omega_h} \delta(E - \omega_h - \omega_{p\ell})$$

Going to CM frame, we find

$$\begin{aligned} \text{Im} F_{\ell' \ell, \ell \ell'}(P, L) &= \frac{\int q^4}{8\pi E^4} \delta_{\ell' \ell} \delta_{\ell \ell'} \\ &= \rho \delta_{\ell' \ell} \delta_{\ell \ell'} \quad \text{two-body phase-space} \end{aligned}$$

The real part carries information and the FV part,
 which exist $\partial C_L^{-1}(P) = \delta C_L^{-1}(P) = 0$

$$\Rightarrow \partial \partial [F^{-1}(P, L)] = 0$$

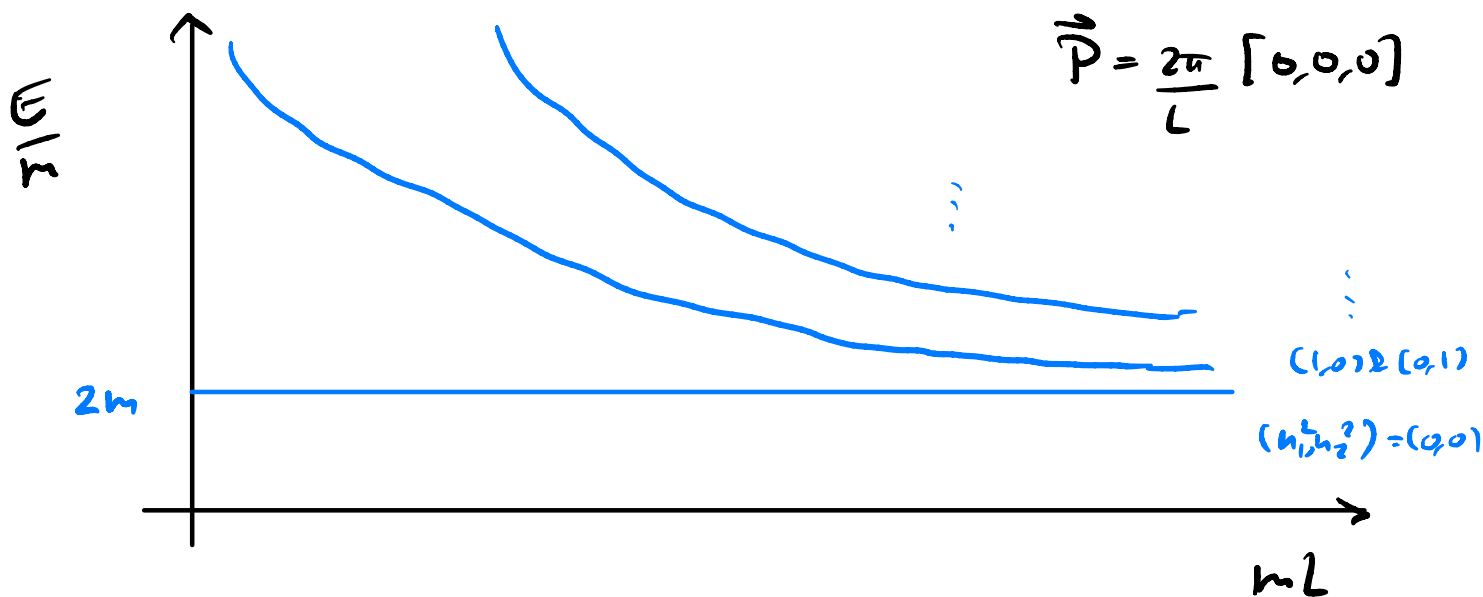
Matrix in $(2n \times 2n)$ -space

the addition of this gives

$$\begin{aligned} E &= \omega_h + \omega_{ph} \\ &= \sqrt{m^2 + \vec{h}^2} + \sqrt{m^2 + (\vec{P} - \vec{h})^2} \\ &= \sqrt{m^2 + \frac{4\pi^2}{L^2} \vec{n}_1^2} + \sqrt{m^2 + \frac{4\pi^2}{L^2} \vec{n}_2^2} \end{aligned}$$

where $\vec{n}_1, \vec{n}_2 \in \mathbb{Z}^3$

This is exactly the FV spectrum of 2 free particles



Diagrammatisch, we define $\delta C_L \leftrightarrow$

$$\delta C_L(P) = \left(\begin{array}{c} \xleftarrow{P-k} \\ \text{---} \text{V} \text{---} \\ \xleftarrow{k} \end{array} \right) - \left(\text{---} \infty \text{---} \right) \leftarrow P$$

$$\equiv \left(\text{---} \text{V} \text{---} \right)$$

↖ FV "cut"

⇒ unterschiedliche states un-shell

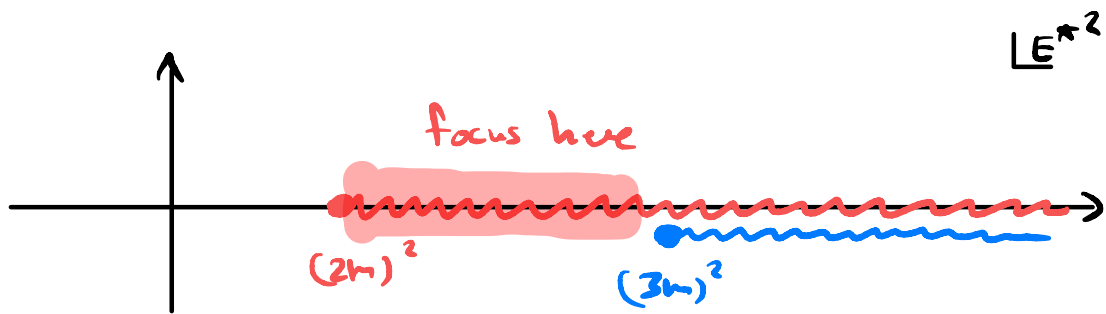
$$= \tilde{A}^*(P) \cdot F(P, L) \cdot \tilde{A}^*(P)$$

↑

Matrices du (core) -space

Interacting Two-Particles

Let us now turn on interactions, focusing on φ^4 theory. Here we want to use our tools to determine scattering amplitudes. Let us focus on the elastic $2 \rightarrow 2$ scattering amplitude iM , & restrict our energy range & interest to the elastic region. In doing so, we systematically have full control over the analytic structure without approximation



From S-matrix unitarity, we know that

$$M_{l_1' l_2' l_1 l_2}(E^*) = \delta_{l_1' l_1} \delta_{l_2' l_2} M_l(E^*) \quad \text{partial wave expansion}$$

and

$$M_l = K_l \frac{1}{1 - ipK_l}$$

\downarrow
 $2 \rightarrow 2$ K-matrix

The K -matrix contains all short-distance interactions not constrained by unitarity. Can relate to phase shifts via

$$K_{\ell}^{-1} = \rho \cot \delta_{\ell}$$

So, also know $M_{\ell} = \frac{1}{\rho} \frac{1}{\cot \delta_{\ell} - i} = \frac{1}{\rho} e^{i\delta_{\ell}} \sin \delta_{\ell}$

Then I use typical relativistic normalization,

$$\langle \vec{p}' | \vec{p} \rangle = (2\pi)^3 2E_p \delta^{(3)}(\vec{p}' - \vec{p})$$

so that $\rho = \frac{2q^*}{8\pi E^*}$

For a weakly interacting system near threshold, effective range parametrization is useful.

Consider S-wave scattering

$$q^* \cot \delta_{\ell=0} = -\frac{1}{a_0} + \mathcal{O}(q^{*2})$$

near threshold, find

$$M_{\ell=0} = -\frac{16\pi}{i} m a_0 + \mathcal{O}(q^*)$$

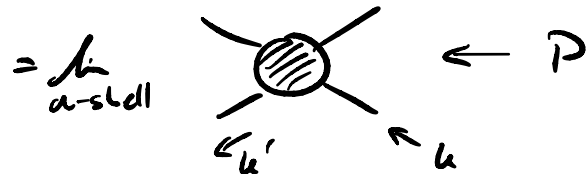
Within QFT, the amplitude is given by applying LSZ to correlators. To all orders, we can write the correlator as

$$C_{\infty}(P) = \text{diagram 1} + \text{diagram 2}$$

fully dressed ↑ propagator

↑ off-shell $z \rightarrow z$ amplitude

The amplitude is $iM = \lim_{\omega \rightarrow \text{shell}} iM(k', k; P)$



The fully dressed propagator

is defined by unit residue, $i\Delta(k) = \frac{i}{k^2 - m^2 + i\epsilon} + iS(k^2)$

↑ non-singular at $k^2 = m^2$

The analogous FV correlator is

$$C_L(P) = \text{diagram 1} + \text{diagram 2}$$

here, recall that $m_L - m_{\infty} = \mathcal{O}(e^{-mL})$

$$\Rightarrow (\text{line})_L = (\text{line})_{\infty} + \mathcal{O}(e^{-mL})$$

↑ drop

We have introduced iM_L as the FV analogue of the off-shell amplitude,

$$iM_L = \left(\text{diagram: a circle with a vertical line through it and a shaded region} \right)_L$$

Let's examine the structure of the FV correction,

$$\delta C_L(P) = C_L(P) - C_\infty(P)$$

$$= \text{diagram: circle with V and two external lines} - \text{diagram: circle with } \infty \text{ and two external lines}$$

$$+ \text{diagram: two circles with V and shaded region} - \text{diagram: two circles with } \infty \text{ and shaded region}$$

= add & subtract useful zeroes

⋮

$$= \text{diagram: circle with V and vertical line}$$

$$+ \text{diagram: circle with } \infty \text{ and vertical line} + \text{diagram: circle with } \infty \text{ and V and vertical line} + \text{diagram: circle with V and vertical line}$$

$$+ \text{diagram: circle with V and shaded region and vertical line}$$

hence, $\text{diagram: circle with V and vertical line} = A(P) \cdot iF(P, L) \cdot A^\dagger(P, L)$ as before,

this why has non-interacting poles

We have introduced the FV correction to the "amplitude",

$$\text{Diagram} = \left(\text{Diagram} \right)_L - \left(\text{Diagram} \right)_\infty$$

$$\Rightarrow i\delta M_L \equiv iM_L - iM$$

Clearly, this is key to getting relationship between FV & IV objects.

We first look at weak coupling expansion,

$$iM = -i\lambda + \mathcal{O}(\lambda^2)$$

but, $iM_L = -i\lambda + \mathcal{O}(\lambda^2)$ too, so

$$\delta M_L = 0 + \mathcal{O}(\lambda^2)$$

\Rightarrow This gives no info on diverging energies.

To see this, consider poles of δC_L

So, δC_L goes like

$$\delta C_L(P) = \text{Diagram 1} + \text{Diagram 2} + \mathcal{O}(\lambda^2)$$

$$= A(P) \cdot iF(P, L) \cdot A^*(P)$$

$$+ A(P) \cdot iF(P, L) \cdot [-i\lambda] \cdot iF(P, L) \cdot A^*(P) + \mathcal{O}(\lambda^2)$$

where,

$$\begin{aligned} M_{\text{triangle}} &= \delta_{ll'} \delta_{r'r'} \frac{1}{2} \int_{-1}^1 d\cos\theta P_2(\cos\theta) [-i\lambda + \mathcal{O}(\lambda^2)] \\ &= -i\lambda \delta_{l0} \delta_{r'0} + \mathcal{O}(\lambda^2) \end{aligned}$$

So, infinite-dimensional matrix is truncated @ S-wave,

$$\delta C_L(P) \approx A(P) iF_S(P, L) \left[1 + \lambda F_S(P, L) \right] A^*(P) + \mathcal{O}(\lambda^2)$$

with

$$F_S(P, L) \equiv F_{000}(P, L) = \frac{1}{L^3} - \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k 2\omega_{p-k} (\epsilon - \omega_k - \omega_{p-k} + i\epsilon)}$$

Poles in C_L are same as poles in δC_L .

Recall that $F(P, L)$ gives free states

But, we run into an issue. Consider systems
 $\partial \mathcal{L}, \vec{P} = \vec{0}$

Poles in δC_L are $\partial \frac{1}{1 + \lambda F(E, L)} = 0$

$$\Rightarrow \left[1 + \lambda \approx \frac{1}{L^3} \frac{1}{4m^2(E-2m)} + \sum_{l \neq 0} \text{terms} \right]^{-1} = 0$$

Thus, again, only has poles of free energies!

A truncated expansion does not yield information
 on decaying energies, due to the fact that $E-2m = \mathcal{O}(\lambda)$,
 & we have neglected terms in the expansion.

Let us consider amplitude of $\mathcal{O}(\lambda^2)$,

$$\partial \mathcal{M} = -i\lambda + (-i\lambda)^2 \int \frac{d^4 k}{(2\pi)^4} \frac{i}{k^2 - m^2 + i\epsilon} \frac{i}{(P-k)^2 - m^2 + i\epsilon} + (t, u) + \mathcal{O}(\lambda^3)$$

$$= \text{X} + \text{loop with } \infty \text{ on } k \text{ line} + \text{loop with } \infty \text{ on } P-k \text{ line} + \text{loop with } \infty \text{ on } k \text{ line} + \mathcal{O}(\lambda^3)$$

FV "amplitude"

$$i\mathcal{M}_L = -i\lambda - \lambda^2 \int \frac{d^4 k^0}{L^3} \frac{1}{L^2} \int \frac{d^4 k}{2\pi} \frac{i}{k^2 - m^2 + i\epsilon} \frac{i}{(P-k)^2 - m^2 + i\epsilon} + (t, u) + \mathcal{O}(\lambda^3)$$

Examine δM_L ,

$$\delta M_L = -\lambda^2 \left[\frac{1}{L^3} \sum_{\vec{k}} - \int \frac{d^3 k}{(2\pi)^3} \right] \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{i}{k^2 - \omega^2 + i\epsilon} \frac{i}{(p-k)^2 - \omega^2 + i\epsilon} \\ + (t, u) + \mathcal{O}(\lambda^3)$$

The t - & u -channel terms have no singularities
in $2m \leq E^* < 3m$. Conclude

$$\delta M_L \Big|_{t,u} \sim \mathcal{O}(e^{-nL})$$

Further, the s -channel term is exactly like what we
considered for the F -function,

$$\Rightarrow \delta M_L = -\lambda^2 i F_S(p, L) + \mathcal{O}(e^{-nL}, \lambda^3)$$

As we have seen, it is enough to look at poles of
 δM_L . But, again, find only non-advancing poles.
What is going on? Pole structures emerge only
when summing the infinite series of interactions.

Consider the Dyson-Schwinger eq. for iM ,

$$\begin{aligned}
 \text{shaded circle} &= \text{circle} + \text{circle with shaded loop} \\
 &= \text{circle} + \text{circle with 2 shaded loops} + \text{circle with 3 shaded loops} + \dots
 \end{aligned}$$

\hookrightarrow ZPE Bethe-Salpeter kernel

$$\text{circle} = \text{circle} + \text{circle with loop} + \text{circle with 2 loops} + \dots$$

This is an integral eqn. for iM ,

$$iM(p, p) = iB(p, p) + \int \frac{d^4 k}{(2\pi)^4} iB(p, k) i\Delta(k) i\Delta(p-k) iM(k, p)$$

Similar expressions exists for iM_L .

Suppose we truncate $iB = -i\lambda + \mathcal{O}(\lambda^2)$,

but sum the infinite s-channel series. We find

for δM_L ,

$$\begin{aligned}
 \delta M_L &= -i\lambda \sum_{n=0}^{\infty} \left[F_S(p, L) \lambda \right]^n \\
 &= -i\lambda \frac{1}{1 - \lambda F_S(p, L)}
 \end{aligned}$$

Now, poles of δM_L are at

$$1 - \lambda F_s(p, L) = 0$$

Let $\vec{p} = \vec{0}$, & consider near threshold ∂E ,

$$1 - \lambda \left\{ \frac{1}{4m^2 L^3} \frac{1}{(E_0 - 2m)} + (\vec{0} \neq \vec{0}) \right\} = 0$$

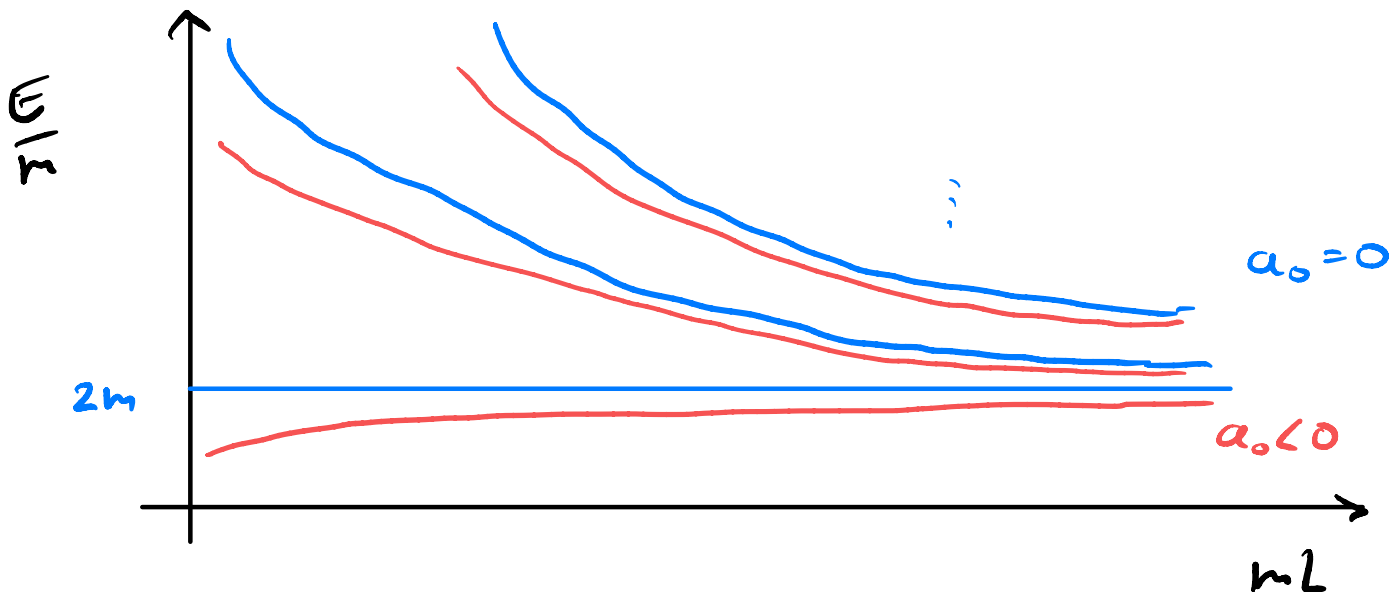
$$\Rightarrow E_0 - 2m = \frac{\lambda \{ \dots \}}{4m^2 L^3} + \mathcal{O}(\lambda^2)$$

But, $M = -\lambda = -\frac{16\pi m a_0}{\hbar}$ at threshold

So,

$$E_0 - 2m = \frac{4\pi a_0}{m L^3} + \mathcal{O}(a_0^2)$$

The presence of interactions shifts the energy levels!



We can expand this argument to all orders in λ , & arrive at a non-perturbative relation between the scattering amplitude & FV spectrum.

The result is

$$\begin{aligned} \delta M_L &= iM(P) \cdot \sum_{n=0}^{\infty} [iF(P,L) \cdot iM(P)]^n \\ &= iM(P) \cdot \frac{1}{1 + F(P,L) \cdot M(P)} \end{aligned}$$

↑
matrix in (L, L) space

So, poles exist when

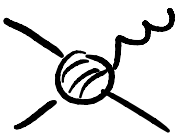
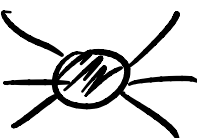

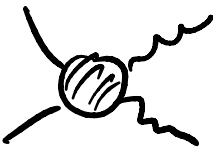
$$\det [1 + F(P,L) \cdot M(P)] = 0$$

↑
in (L, L) space

This is the Lüscher quantization condition. It links the IV scattering amplitude & FV spectrum. It has been used to access scattering amplitudes & resonance physics from LQCD.

Outlook

We have only just touched on the basics of accessing amplitudes via LQCD. Current state of the art includes:

-  radiative transitions
-  $3 \rightarrow 3$ processes
-  resonance form-factors
-  Two-photon transitions