



Swansea University
Prifysgol Abertawe

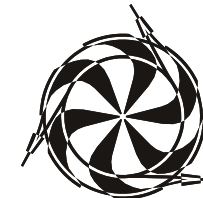
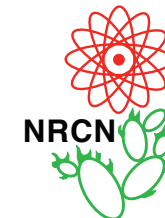


ALPHA: Antihydrogen

and fundamental physics

Prof. Niels Madsen
Swansea University

and the ALPHA collaboration



Low Energy Antiproton Physics conference, Uppsala, Sweden, 10 June 2013

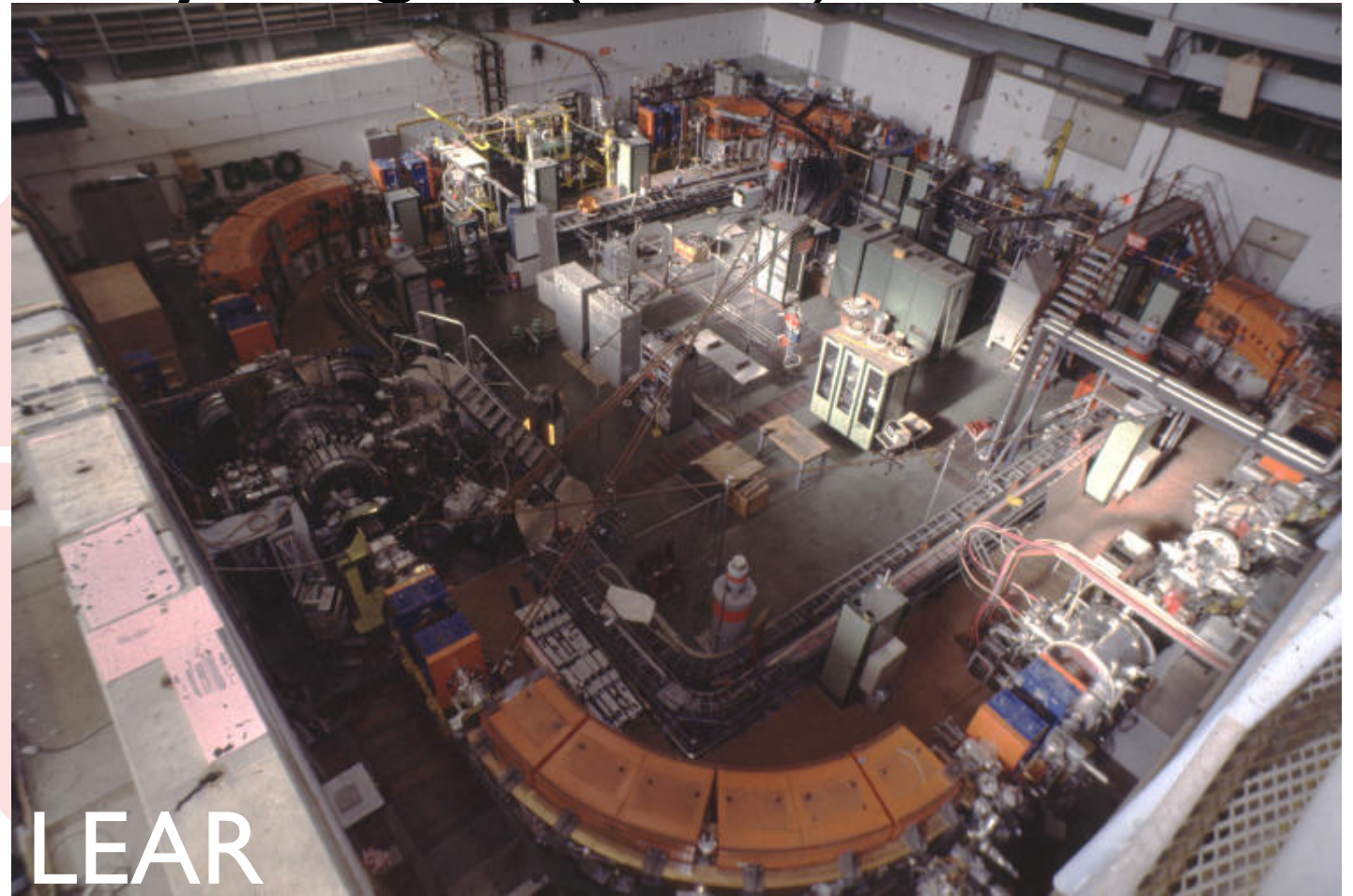
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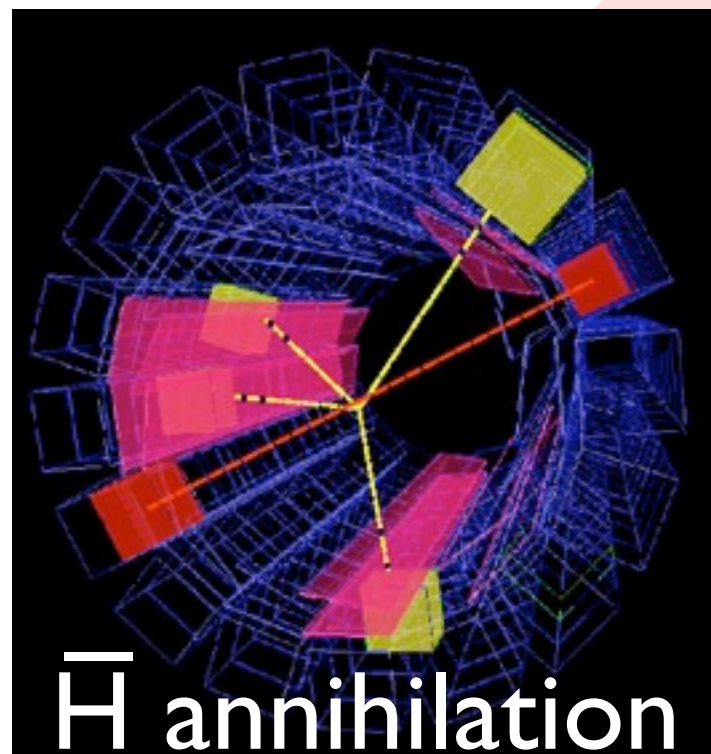
ALF
LEAR



Brief History of Antihydrogen

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ALPHA



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Stories by subject

• [Physics](#)

Stories by keywords

• [Antimatter](#)

Published online 17 November 2010 | *Nature* **468**, 355 (2010) | doi:10.1038/468355a

News

Antimatter held for questioning

Magnetically trapped atoms could test fundamental physics.

Eugenie Samuel Reich

For physicists, a bit of antimatter is a precious gift indeed. By

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2011: ALPHA: $\bar{\text{H}}$ held 1000s \Rightarrow in ground state!



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2013: ALPHA: New method for gravitational measurements.



Why make Antihydrogen



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ALPHA

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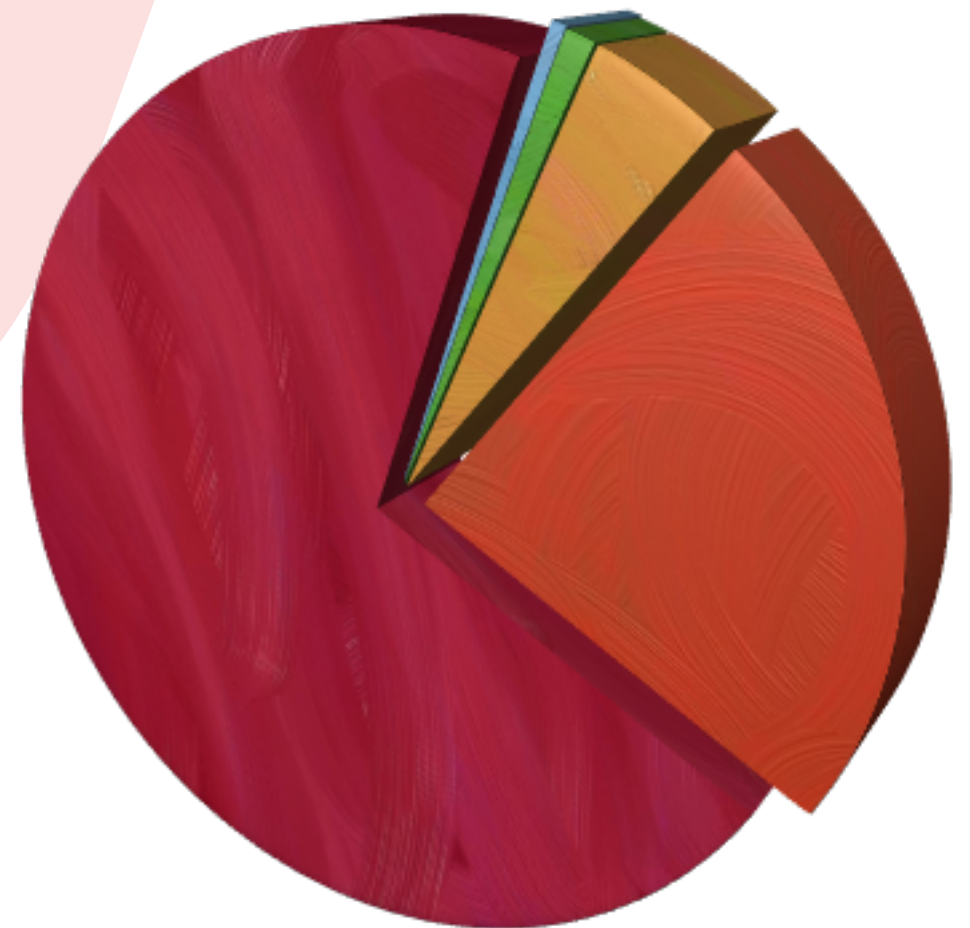
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- Baryon Asymmetry: The universe seems made almost entirely of matter! Really? Why?
- Gravity: How does antimatter respond to gravity? Weak equivalence principle! Insights for quantum gravity?
- Note: **ANY** difference between \bar{H} and H will imply new physics!

Energy Budget of Universe

- Stars and galaxies ~ 0.5%
- Neutrinos ~0.1-1.5%
- Rest of ordinary matter (electrons, protons & neutrons) ~ 4.4%
- Dark Matter ~ 23%
- Dark Energy ~ 73%
- Antimatter 0% ?



Why/how does \bar{H} help?



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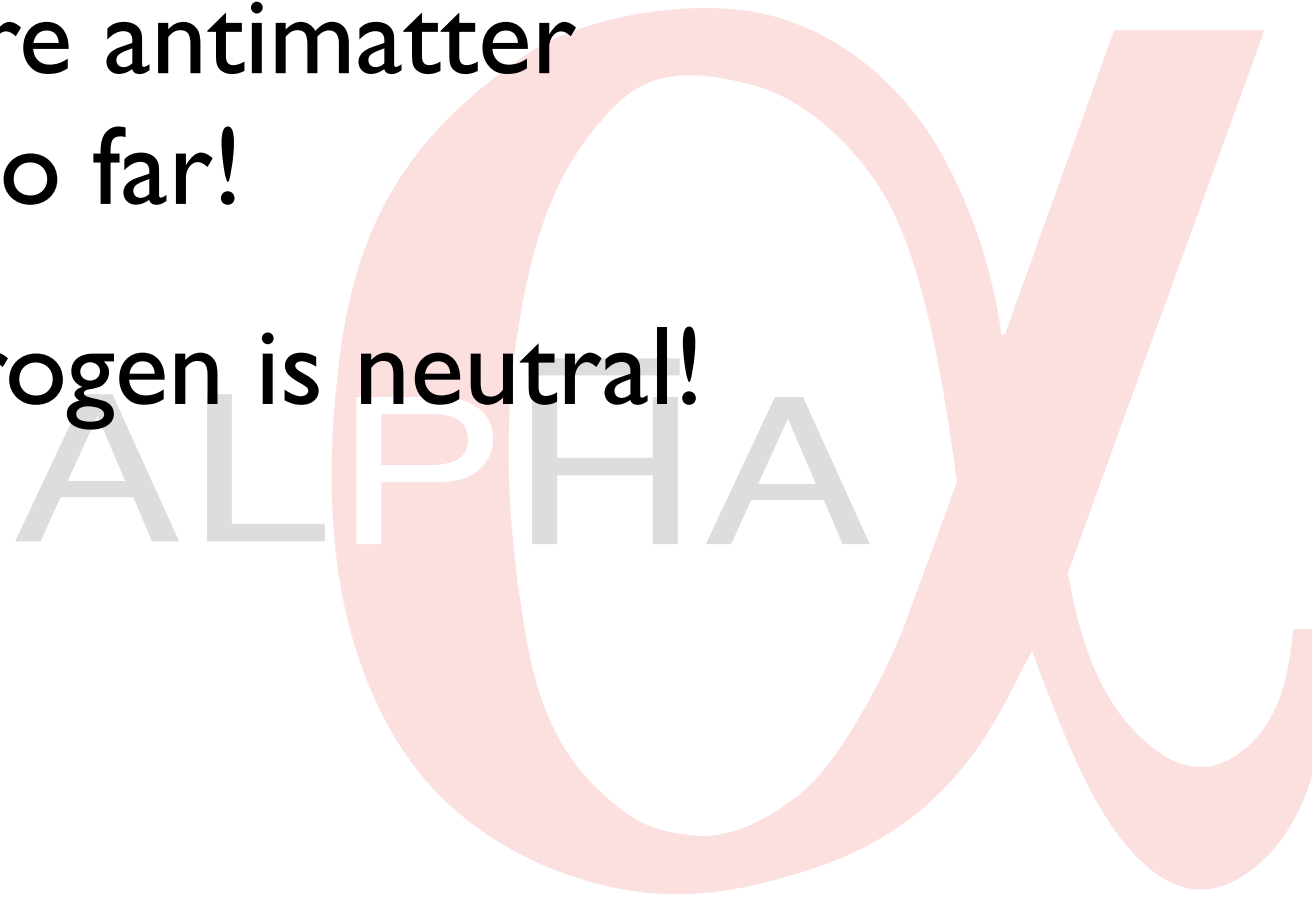
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ALPHA



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- Antihydrogen is neutral!

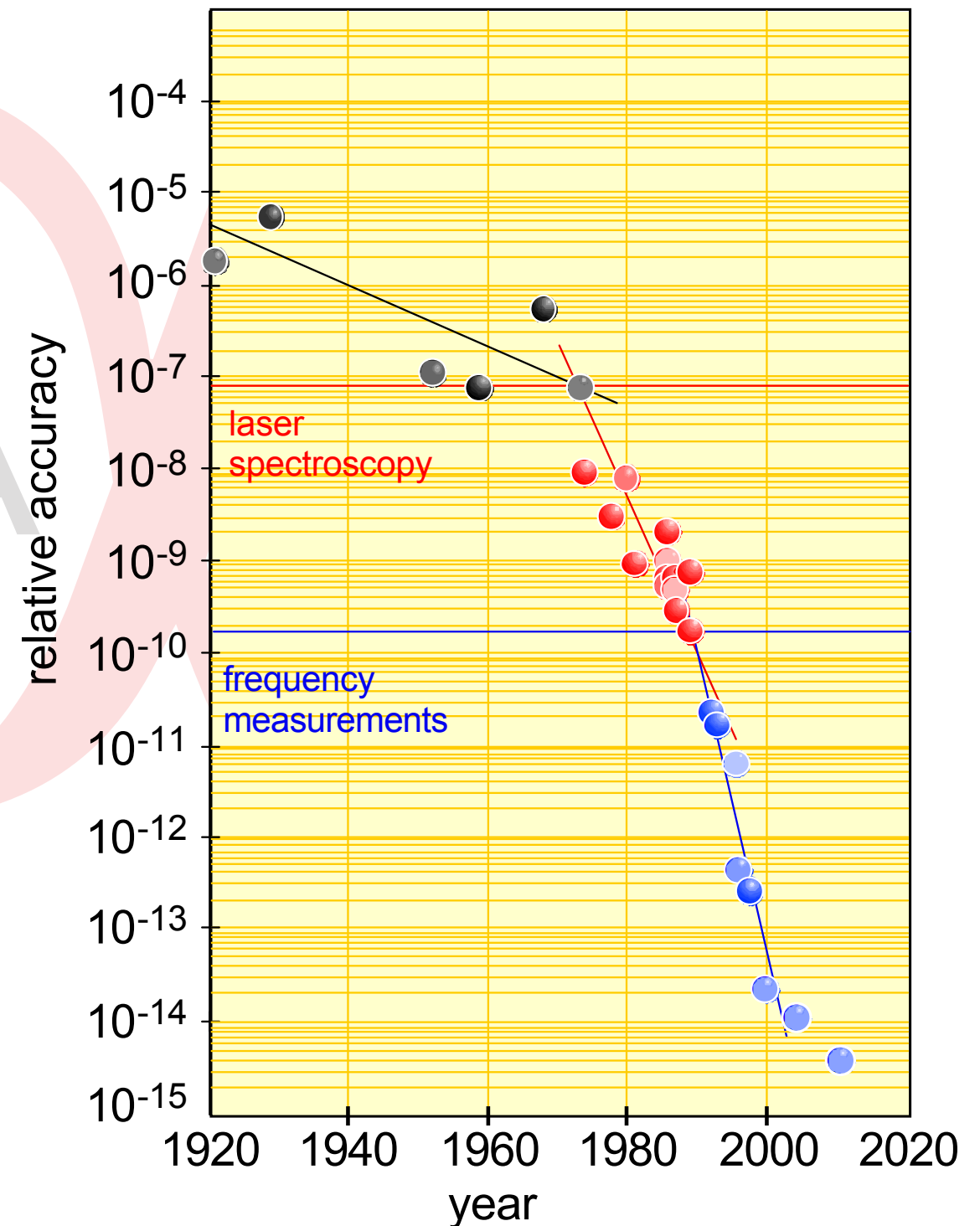


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- Spectroscopic techniques can be brought to bear.
- Ex: H - \bar{H} comparison by $1s$ - $2s$ two photon spectroscopy.



Immediate Concerns



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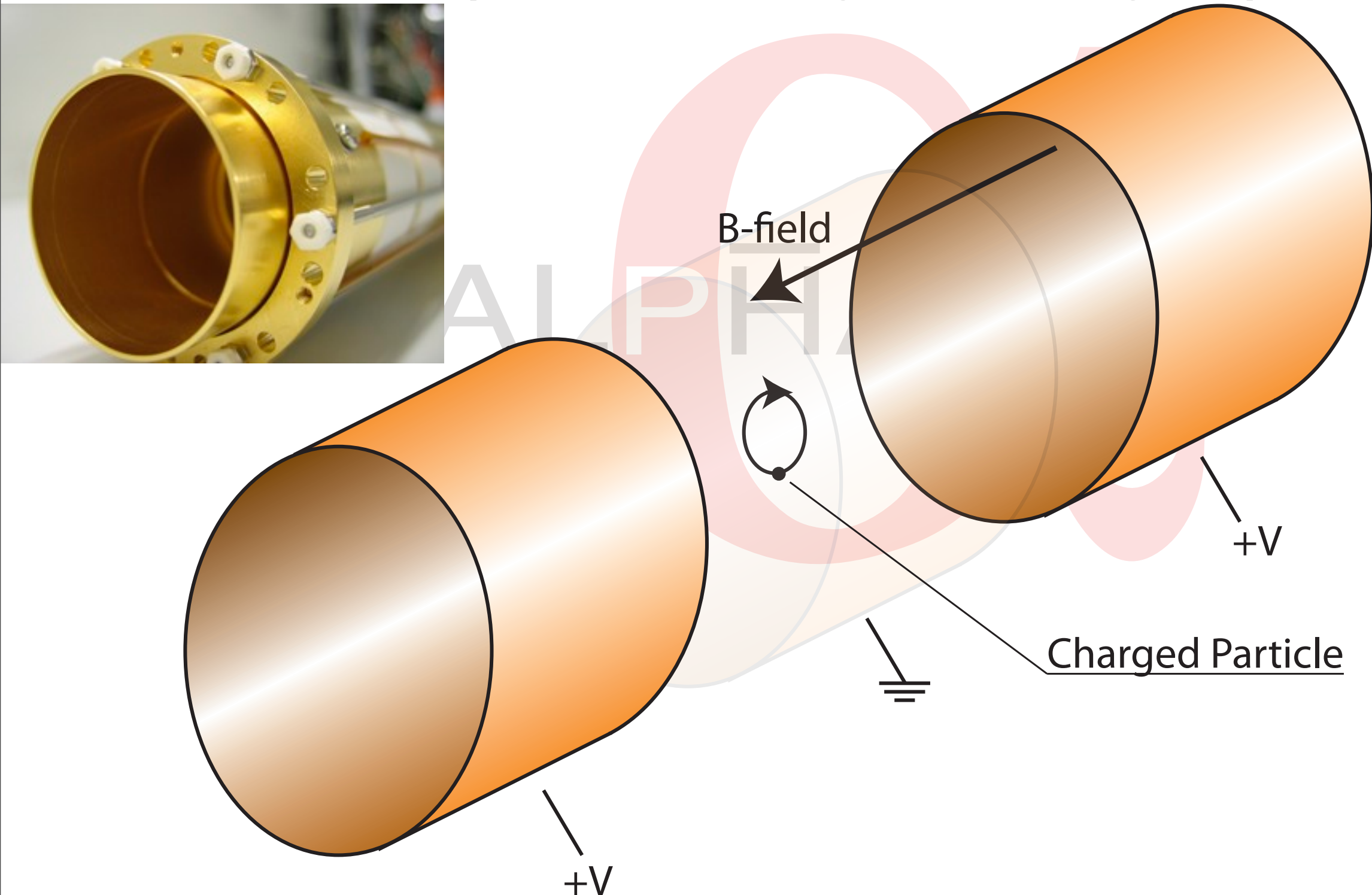
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 - We need cold bare \bar{p} and e^+ !
 - We must form \bar{H} cold and trapped.

Charged Particle Traps

- All our traps are Penning-Malmberg traps



Where do Positrons come from?



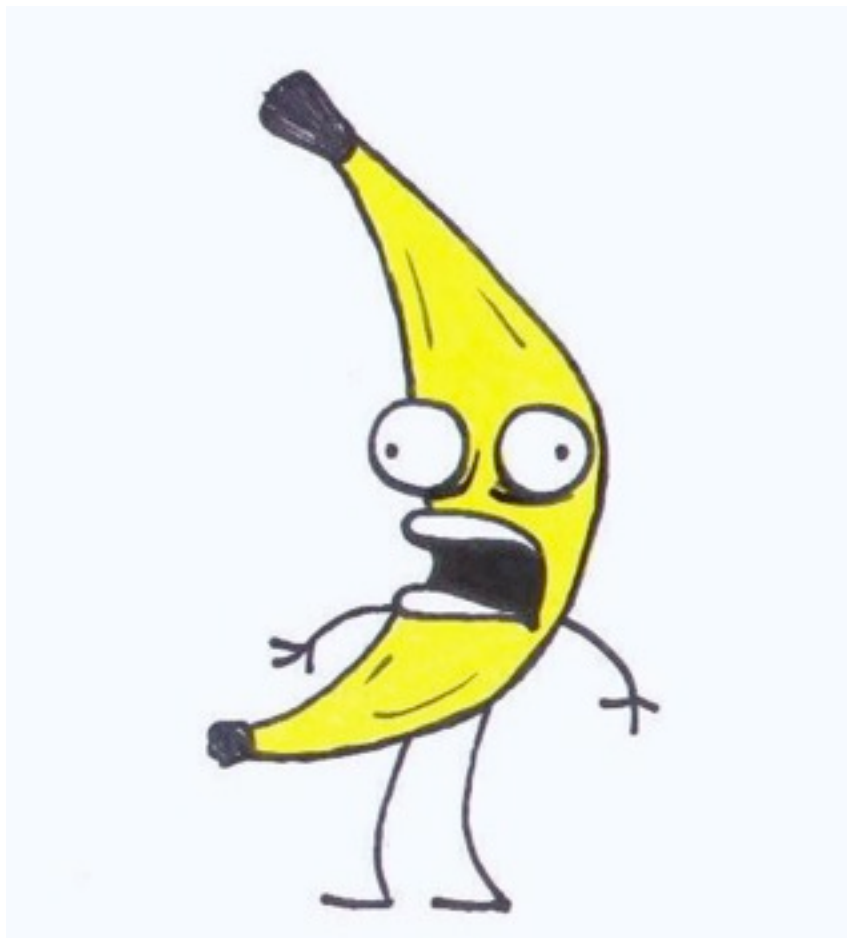
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ALPHA α

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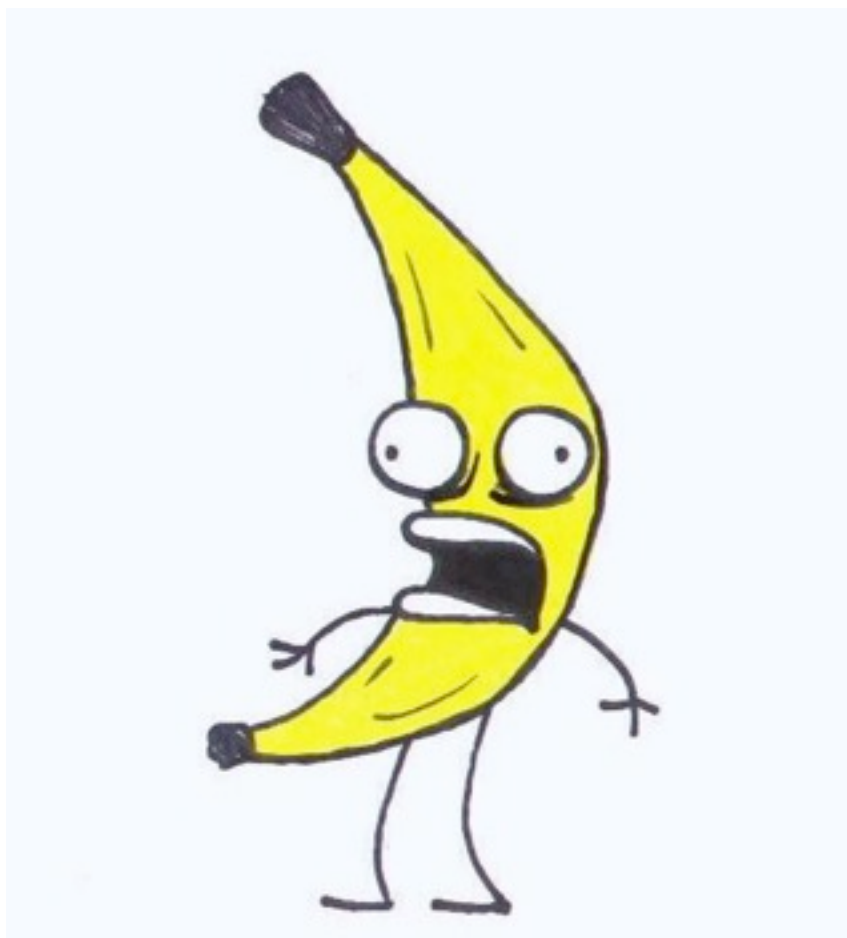
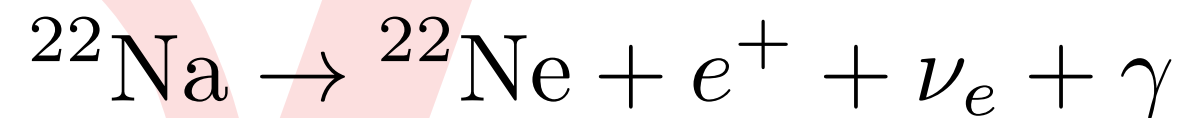


"I am a banana!" Don Hertzfeld

α

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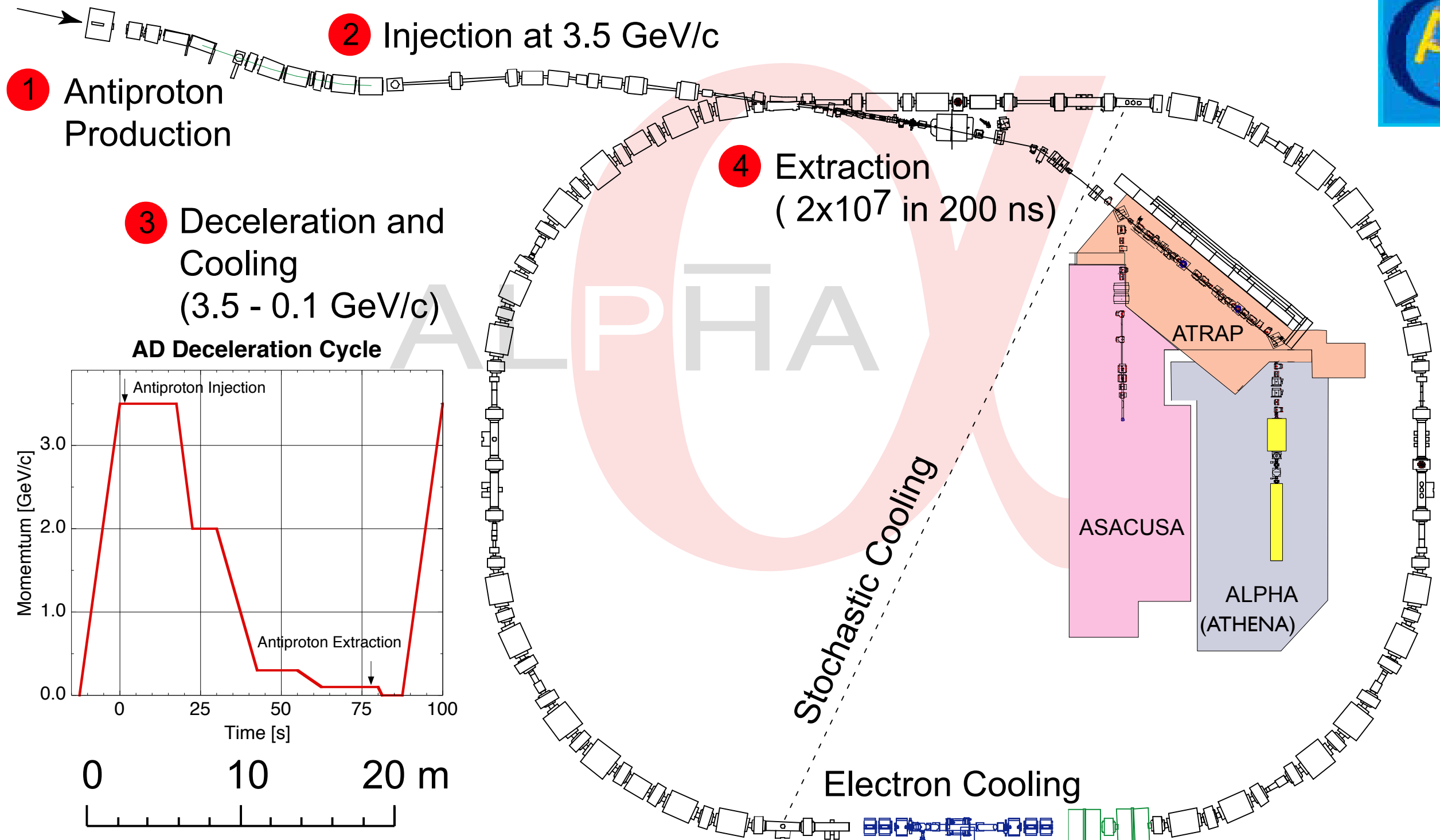
- Fairly Easy: Positive β^+ decay in radioactive isotopes
 - Potassium-40 in Bananas: ~ 15 Positrons / sec
 - We use Sodium-22 source: ~ 10 M / sec



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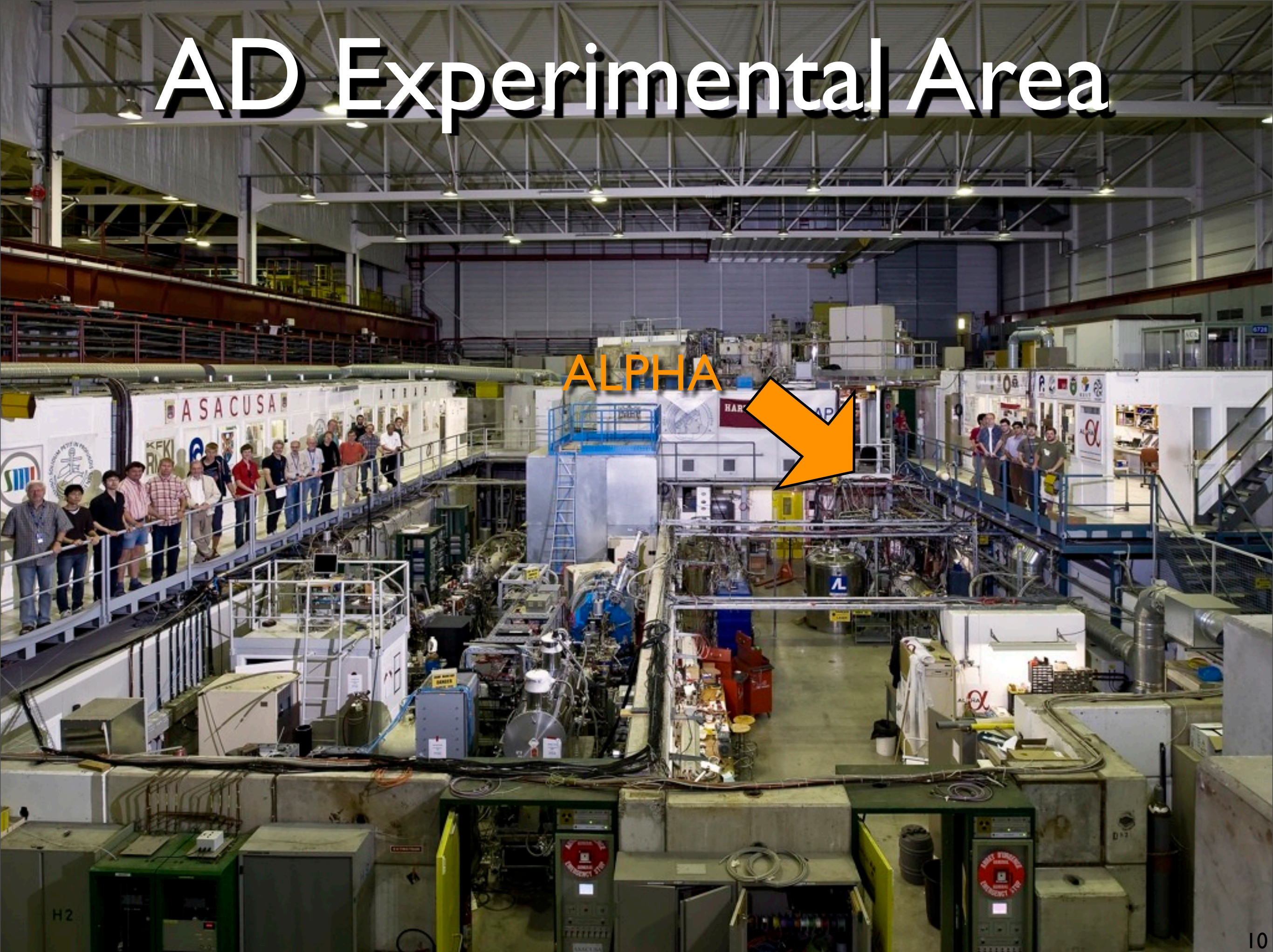
Antiproton Decelerator



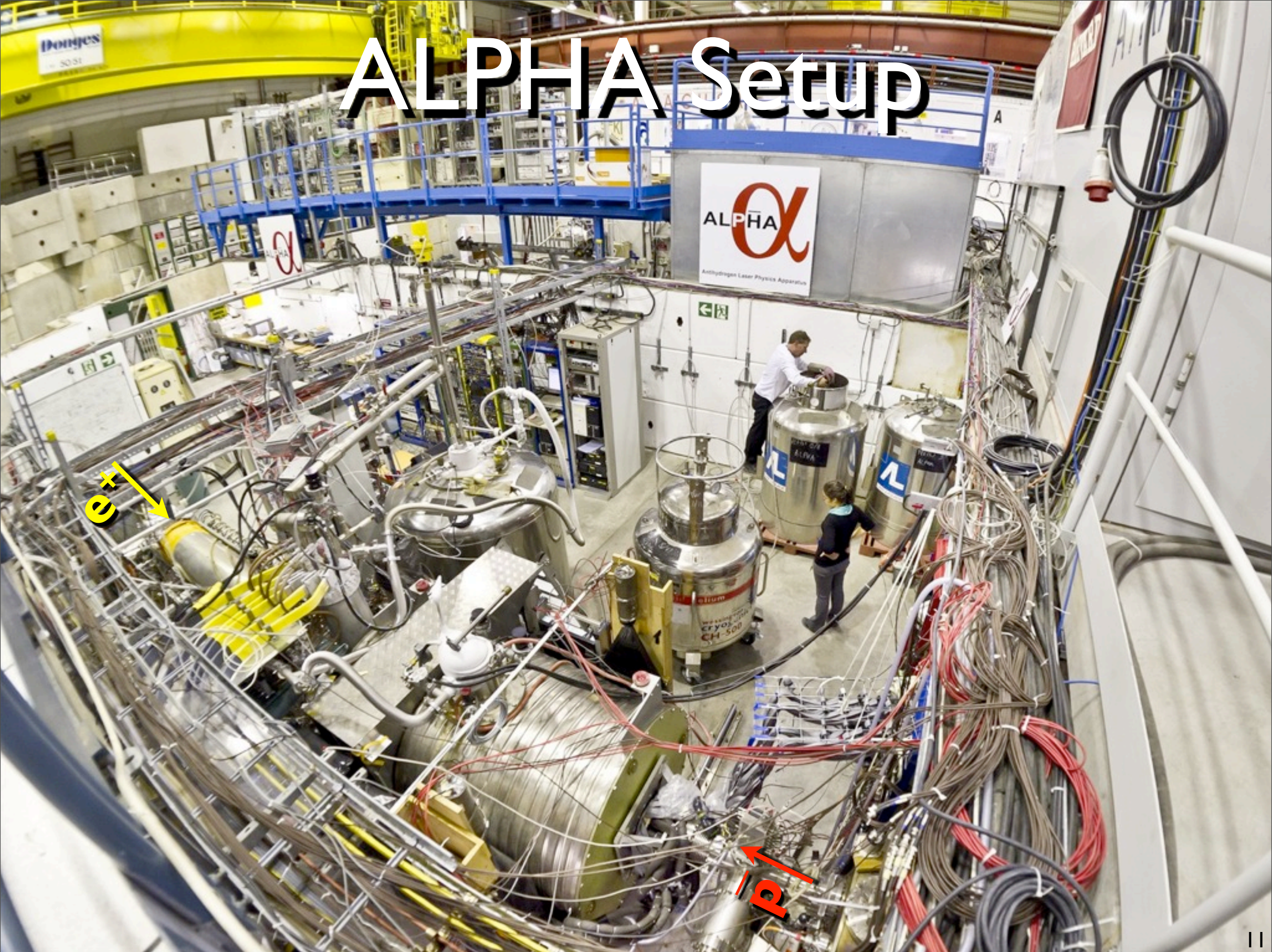
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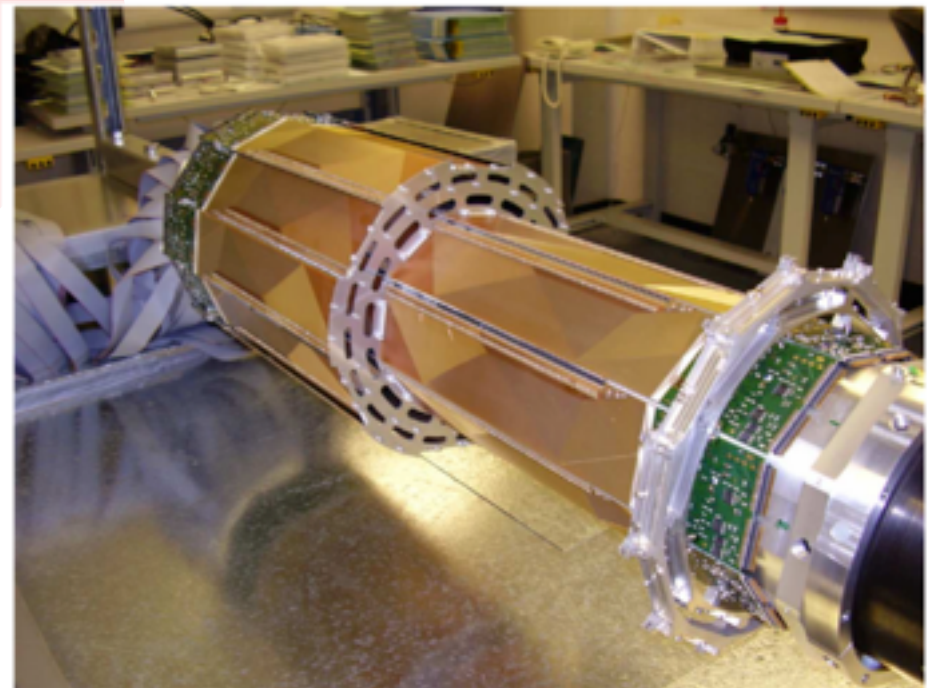
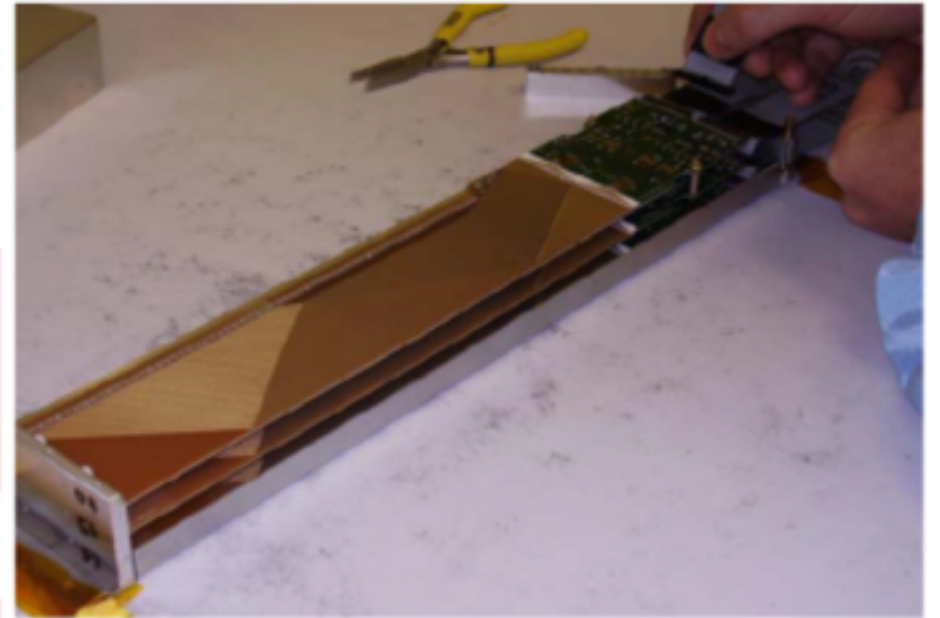
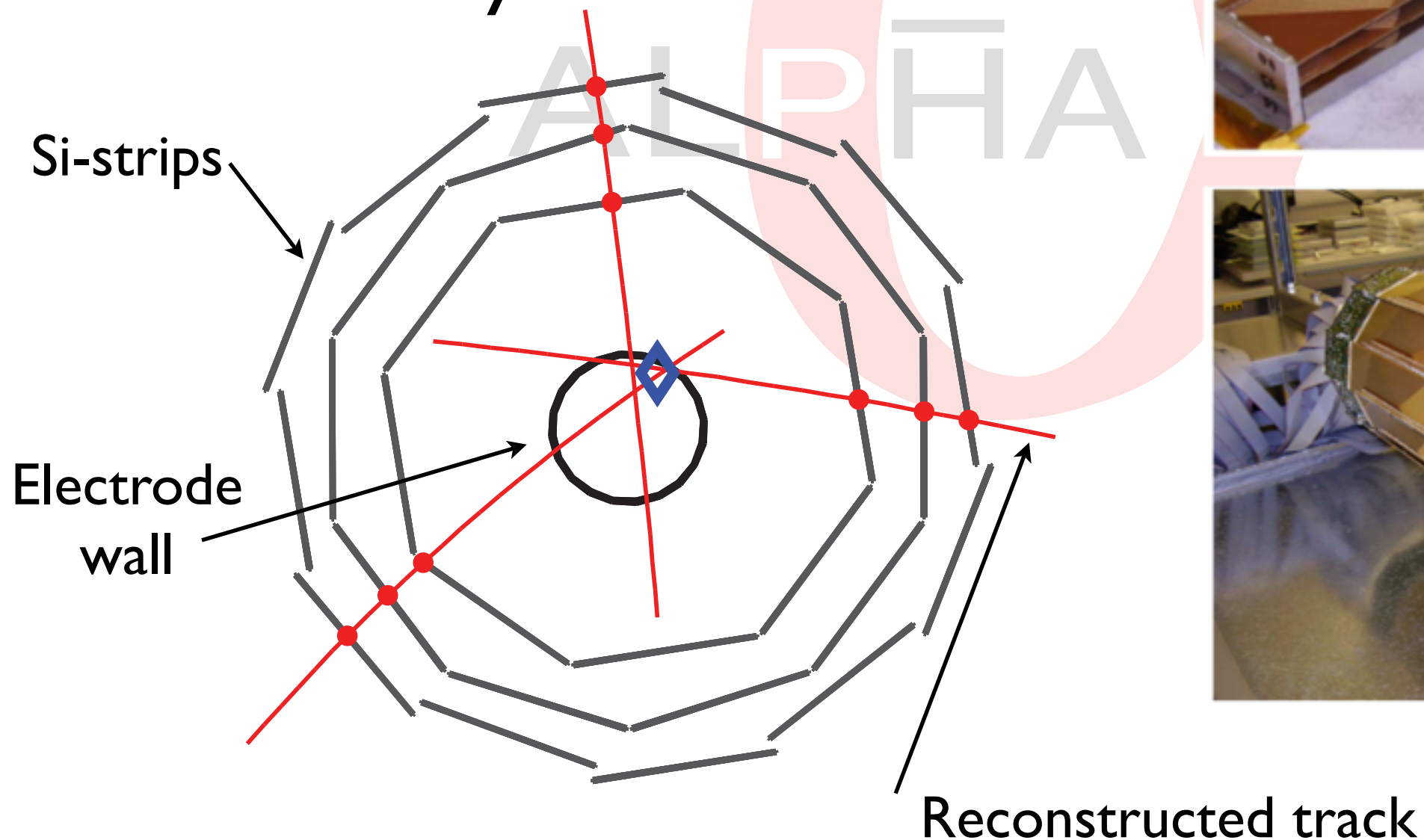


ALPHA Setup



Annihilation Detection

- Si-strip detection
- Vertex resolution $\sim 1\text{mm}$
- Efficiency $\sim 50\%$



(Anti)Atom Trap



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- Atoms can be trapped on their magnetic dipole-moment. $U = -\bar{\mu} \cdot \bar{B}$

ALPHA

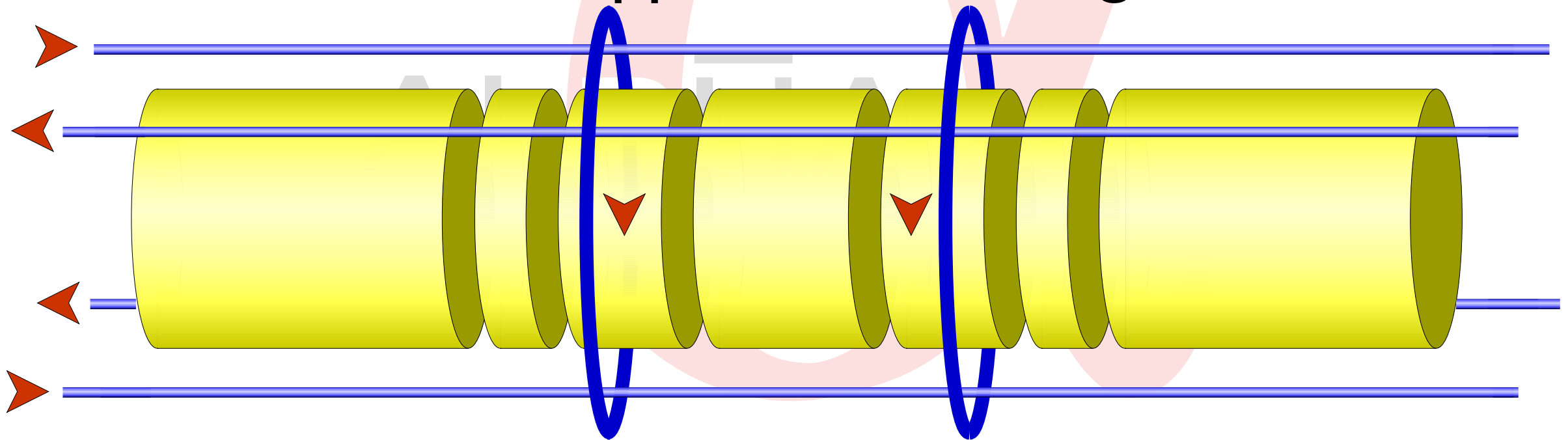
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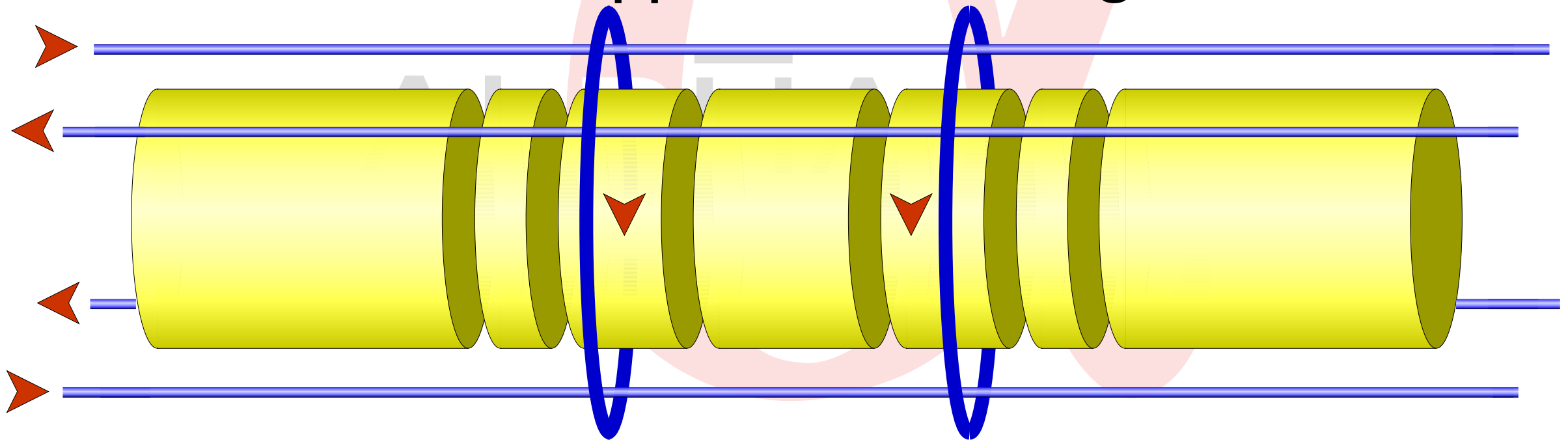


Ioffe-Pritchard Geometry

$$\Delta B = \sqrt{B_{sol}^2 + B_{wall}^2} - B_{sol}$$

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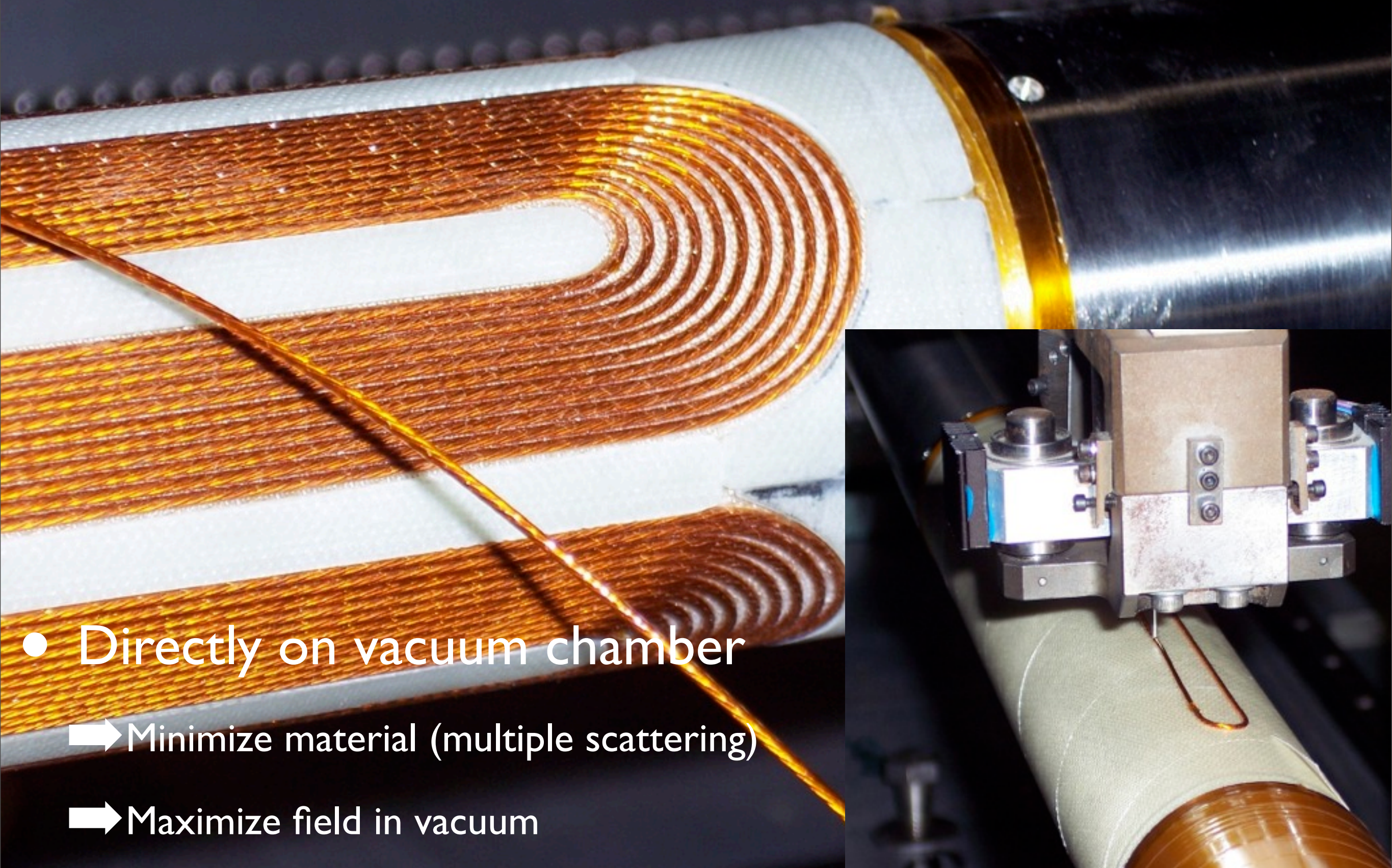


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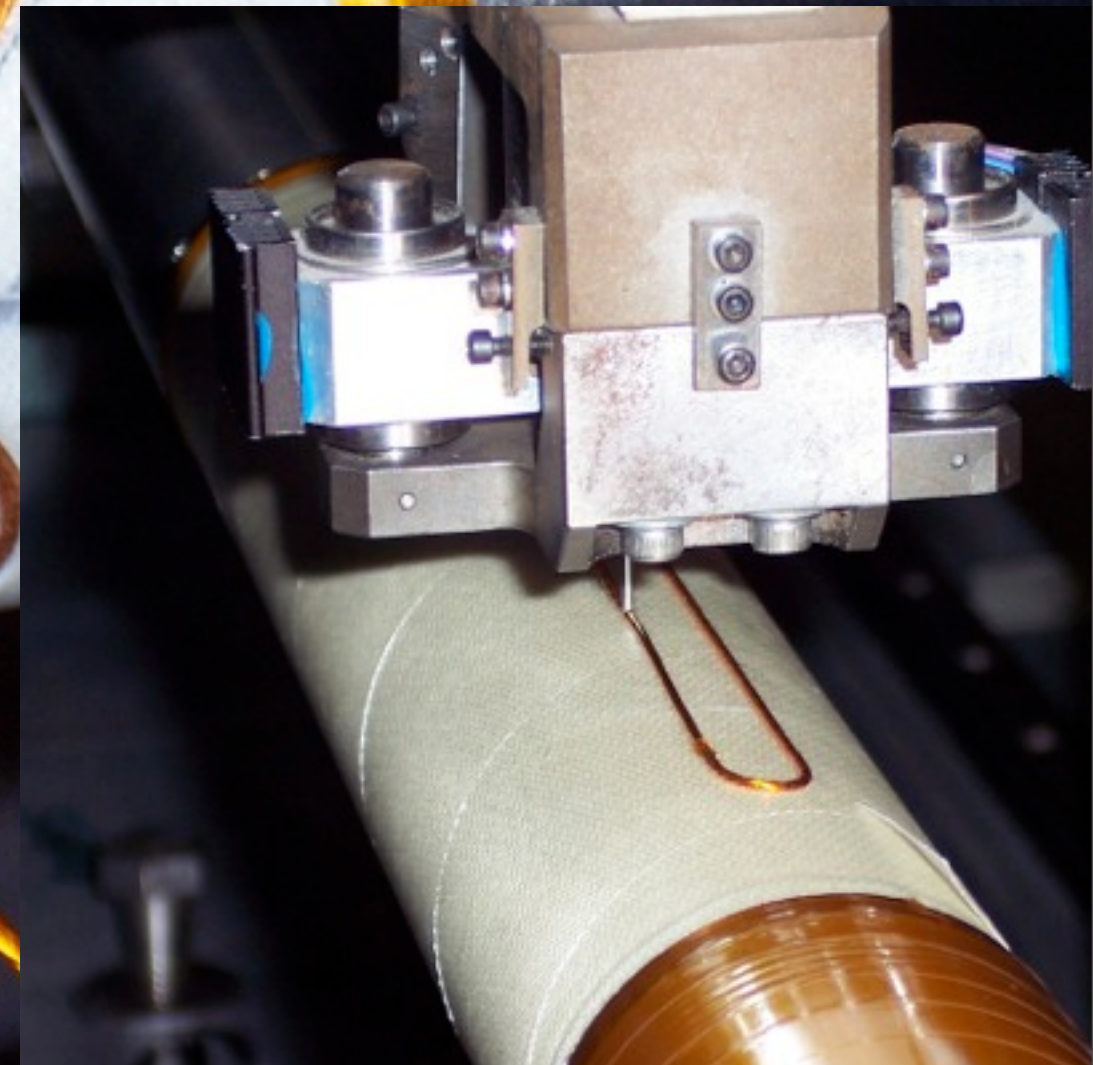
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Shallow : ~ 0.7 K/T for H ground state

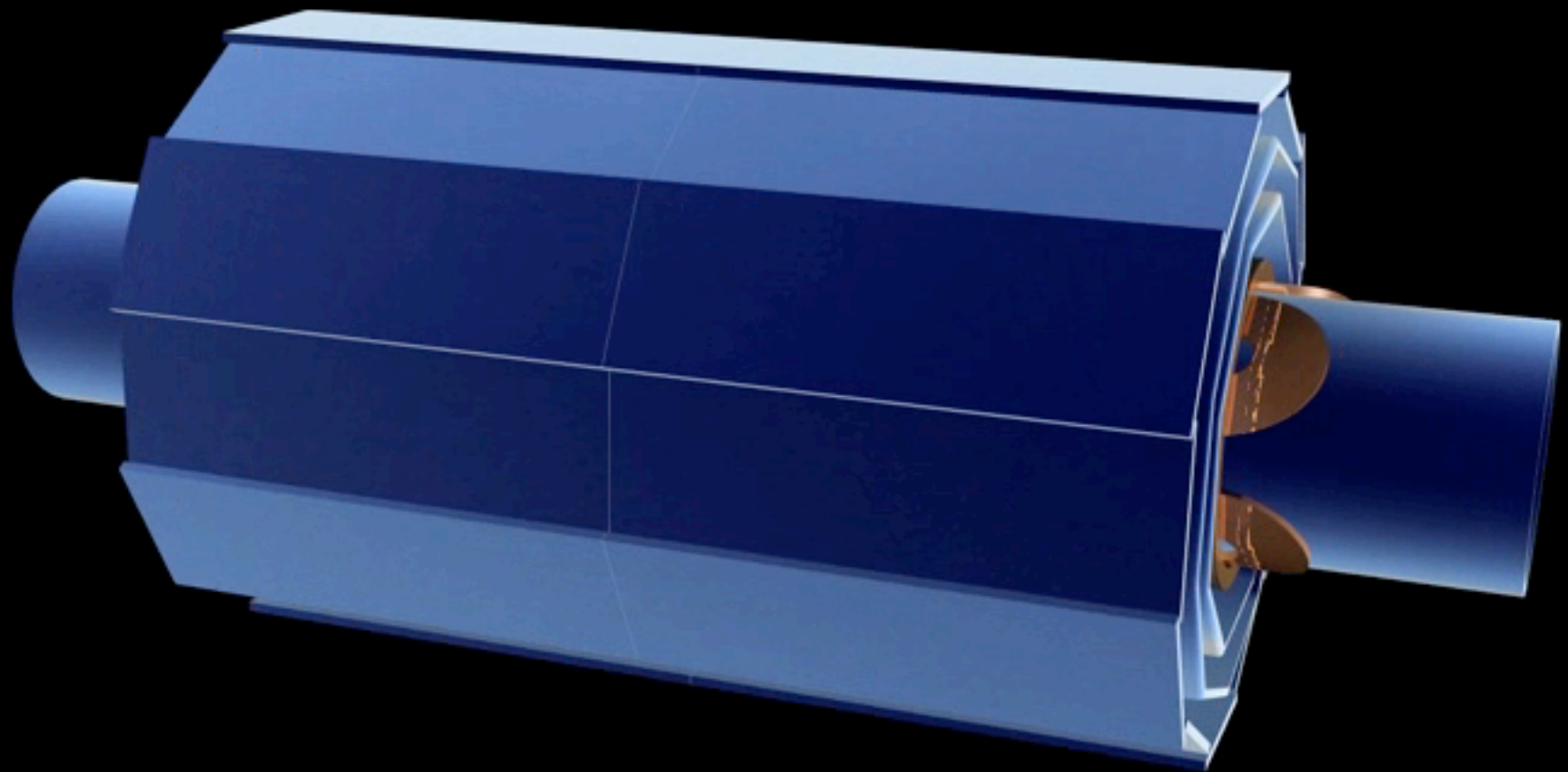
ALPHA Octupole



- Directly on vacuum chamber
 - ➡ Minimize material (multiple scattering)
 - ➡ Maximize field in vacuum

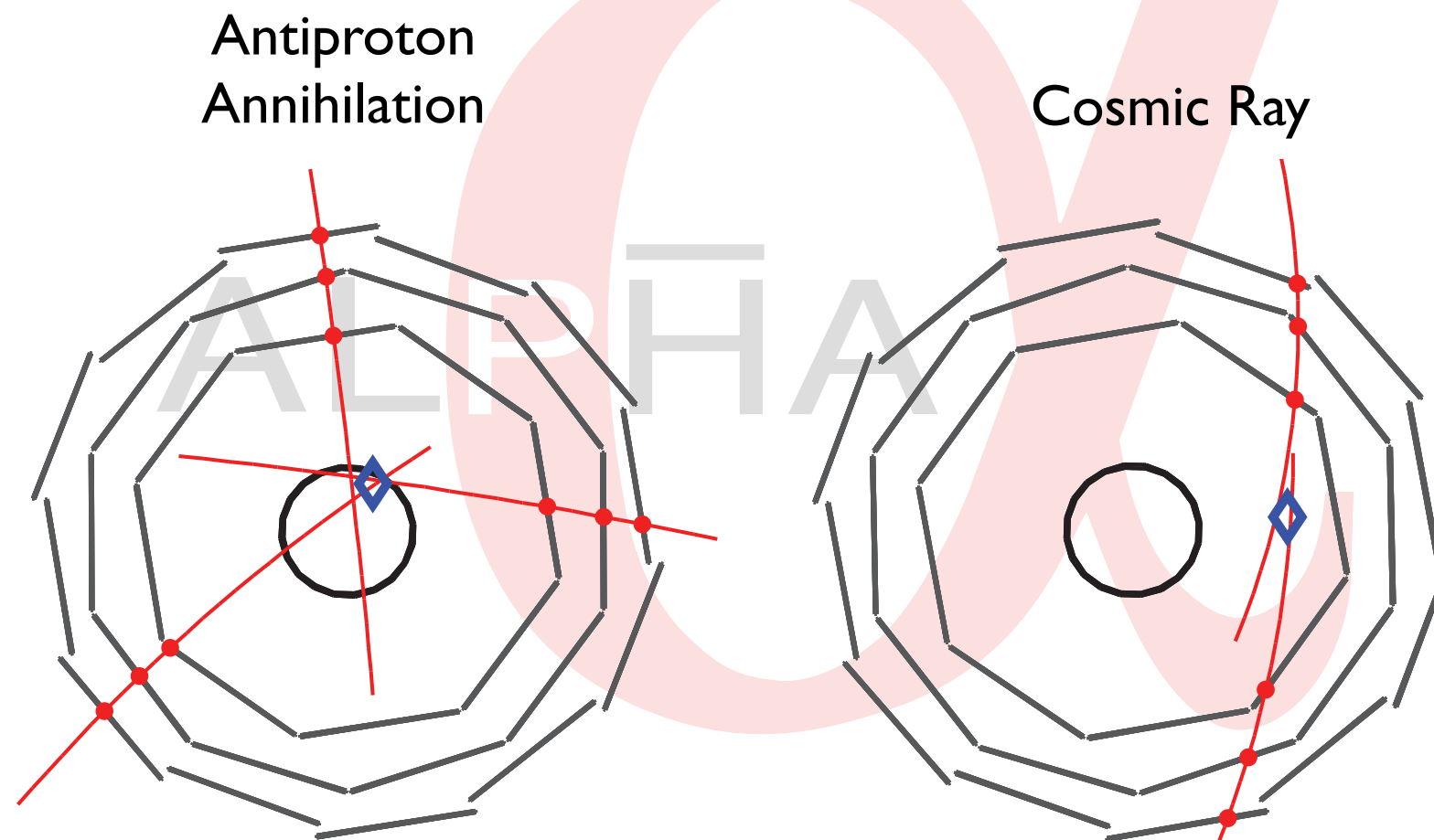


Antihydrogen trapping



Cosmic Background

- One fake signal to worry about : cosmic rays!



- Standard (2010/11) : 1.4/1000 cosmic “fakes” / experiments

Mirror Trapping



Mirror Trapping

- But \bar{p} are not (necessarily) \bar{H} !



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- To avoid these we apply clearing fields before the trap is turned off.

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ALPHA

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- Clean-out not guaranteed ($>20\text{eV}$)



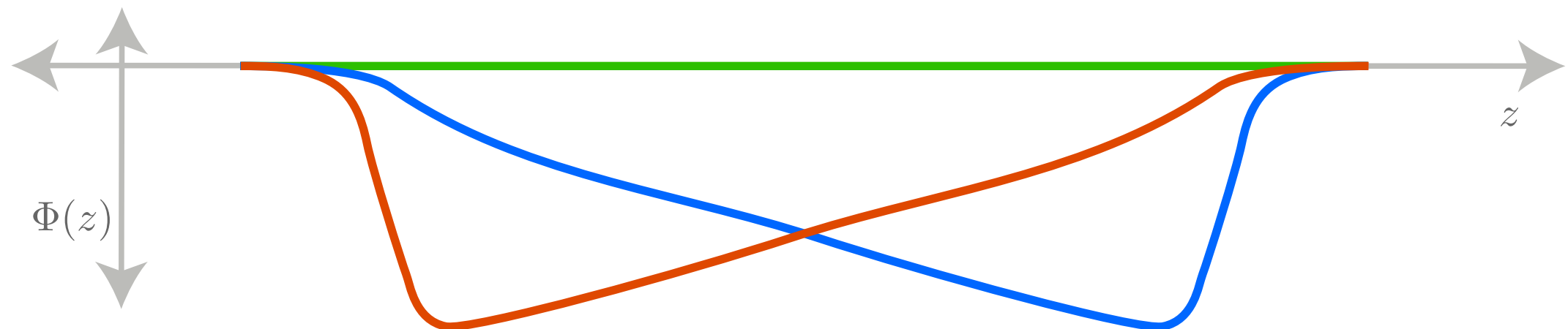
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ALP \bar{H} A

Procedure to check \bar{p} is \bar{H}

- Clean-out not guaranteed ($>20\text{eV}$)
- Heat the positrons and turn off antihydrogen production.
- We distinguished charged particles from neutral using a bias-field (during quench) which does not influence the neutrals!



Trapping Results

- 335 cold e^+ cycles: Each: 30000 \bar{p} , 2×10^6 e^+ , 5000 \bar{H}
- 0.46 ± 0.01 mis-identified cosmons expected

Experiment	Number of attempts	Annihilation events
No bias	137	15
Left bias	101	11
Right bias	97	12
No bias, heated e^+	132	1
Left bias, heated e^+	60	0
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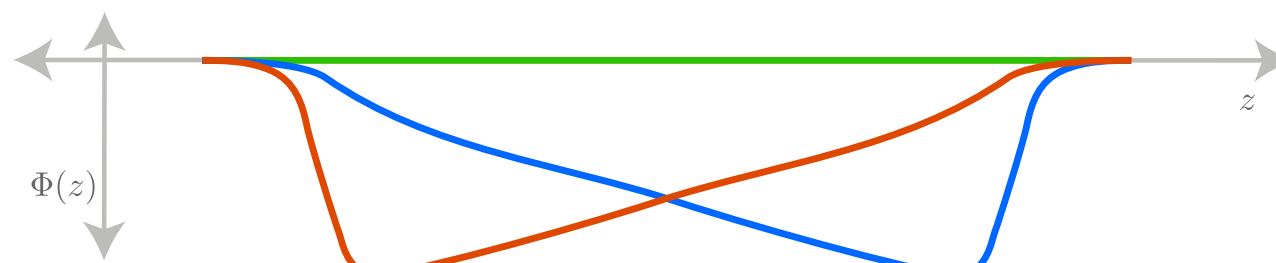
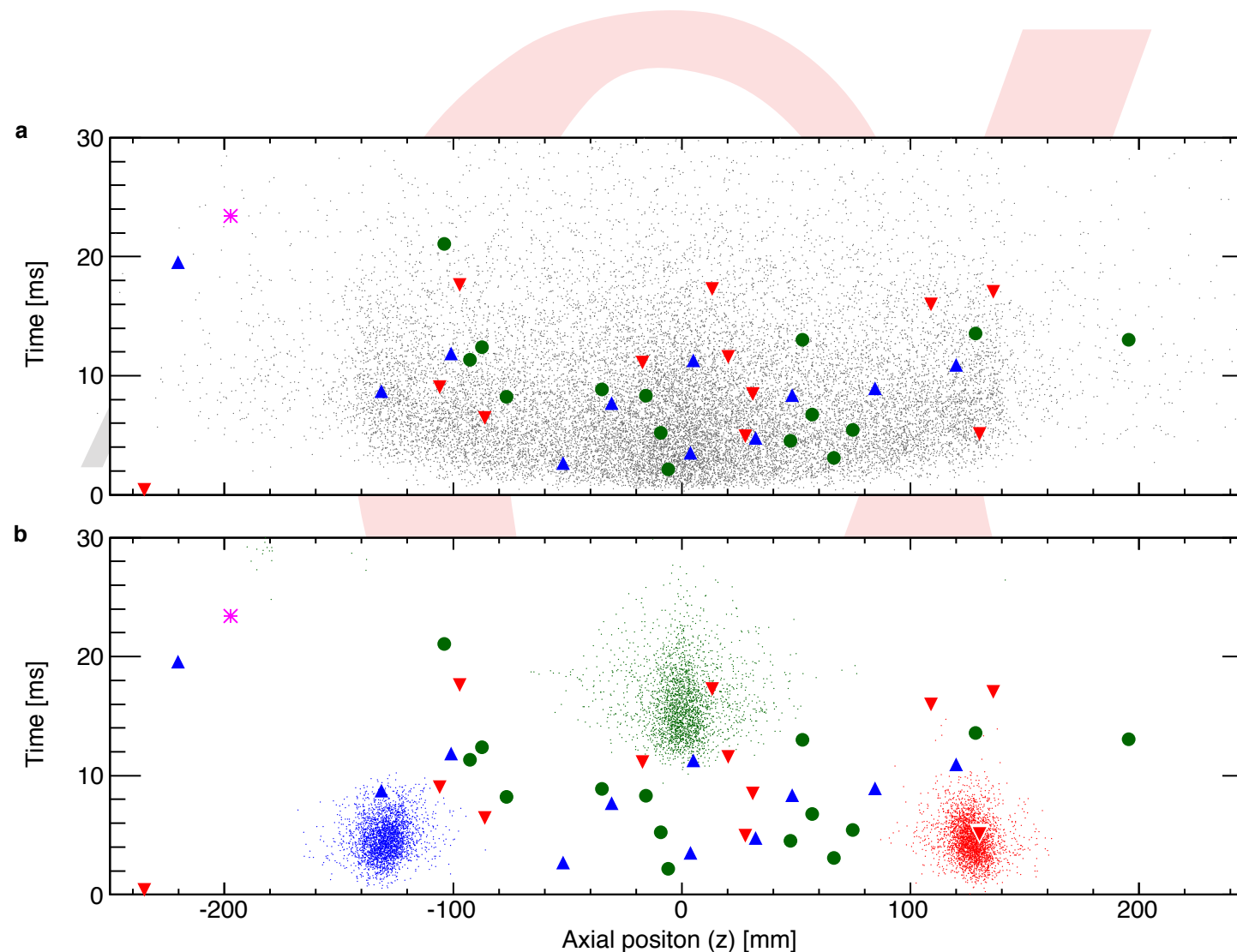
1 “hot” event

Trapping Results

- No spatial bias in signal; Heating 'turns off' signal

Simulation:
Antihydrogen

Simulation:
Antiprotons



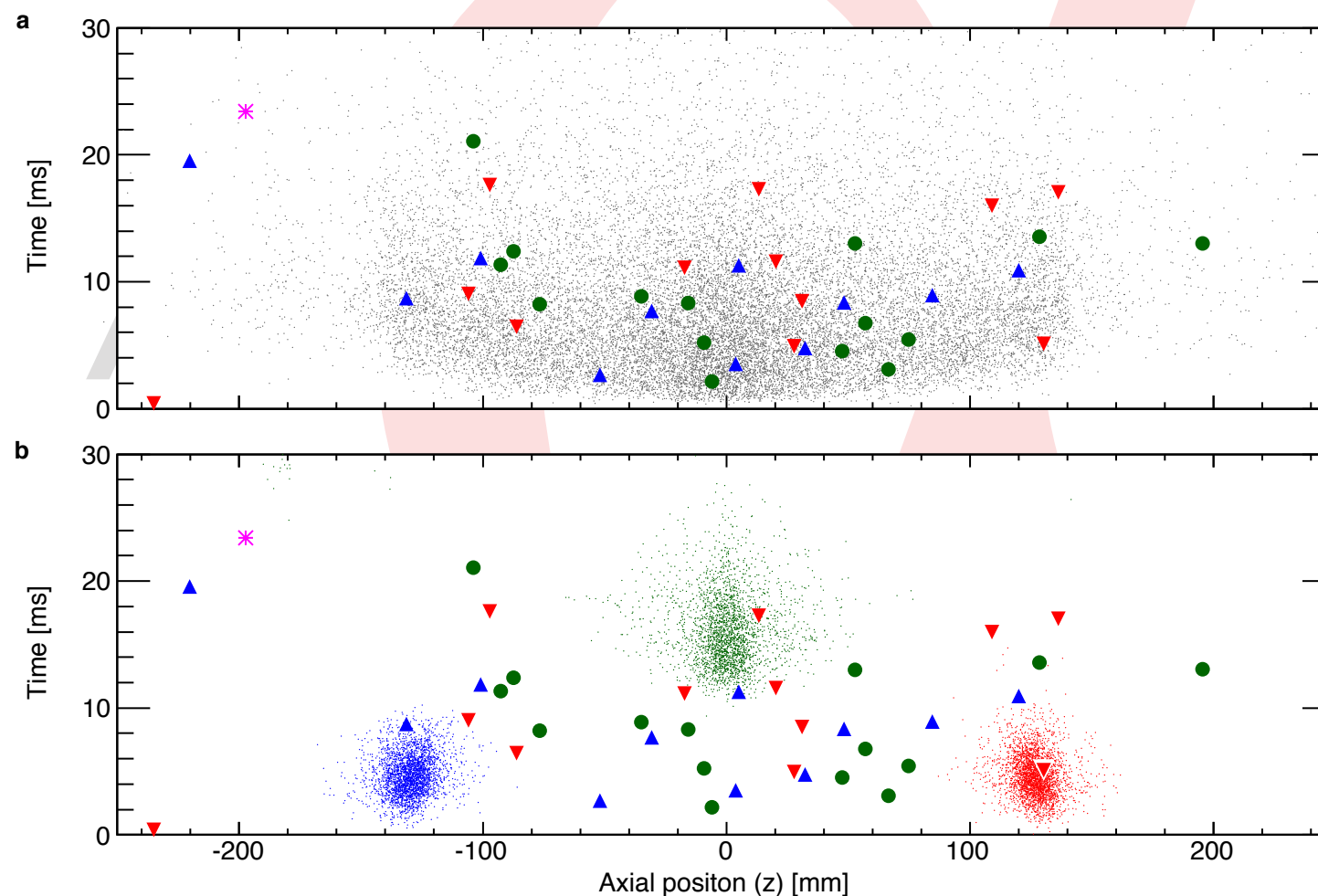
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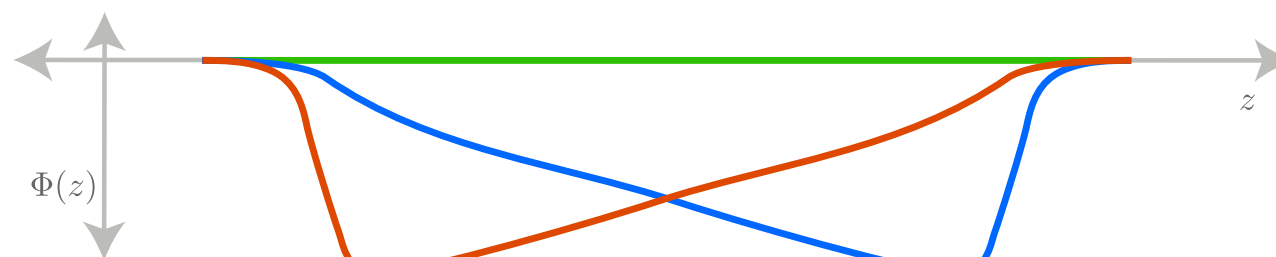
38 Antihydrogen atoms trapped! Background 1.4 ± 1.4

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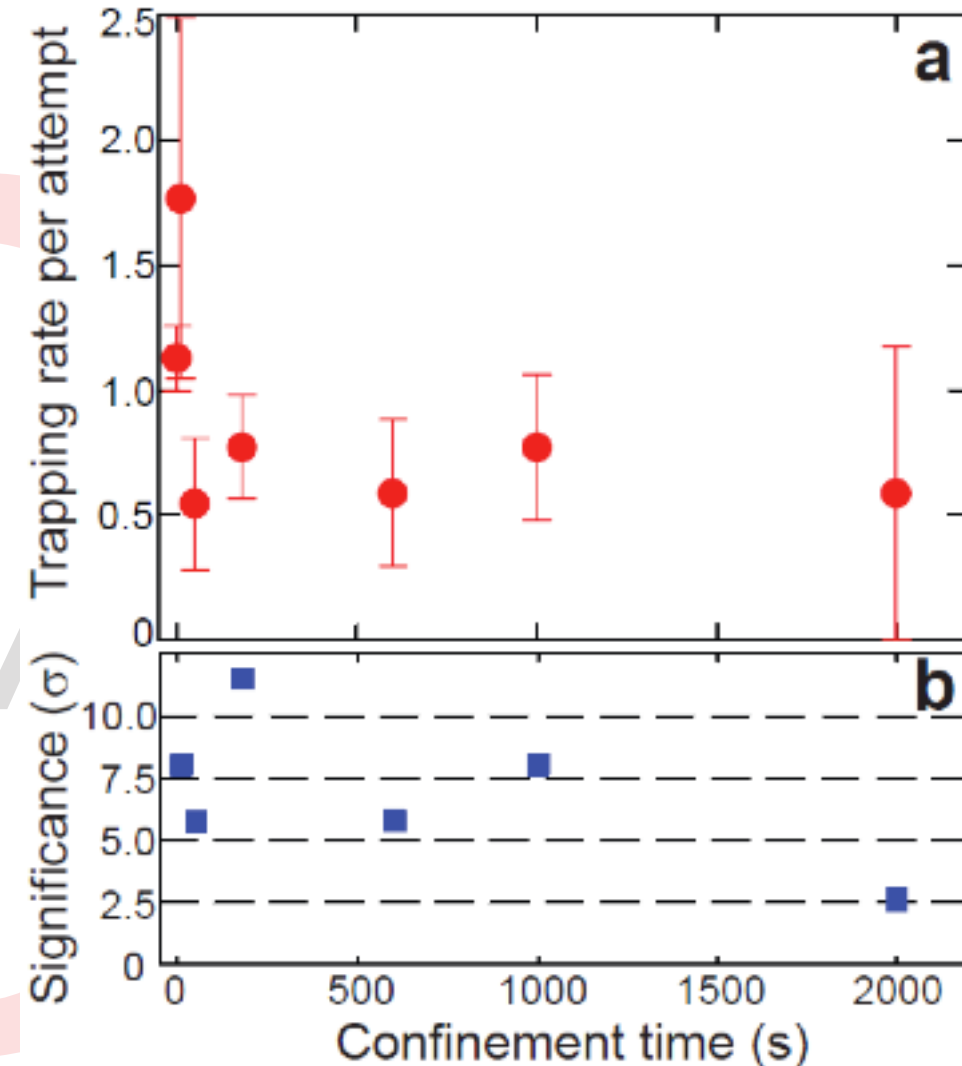
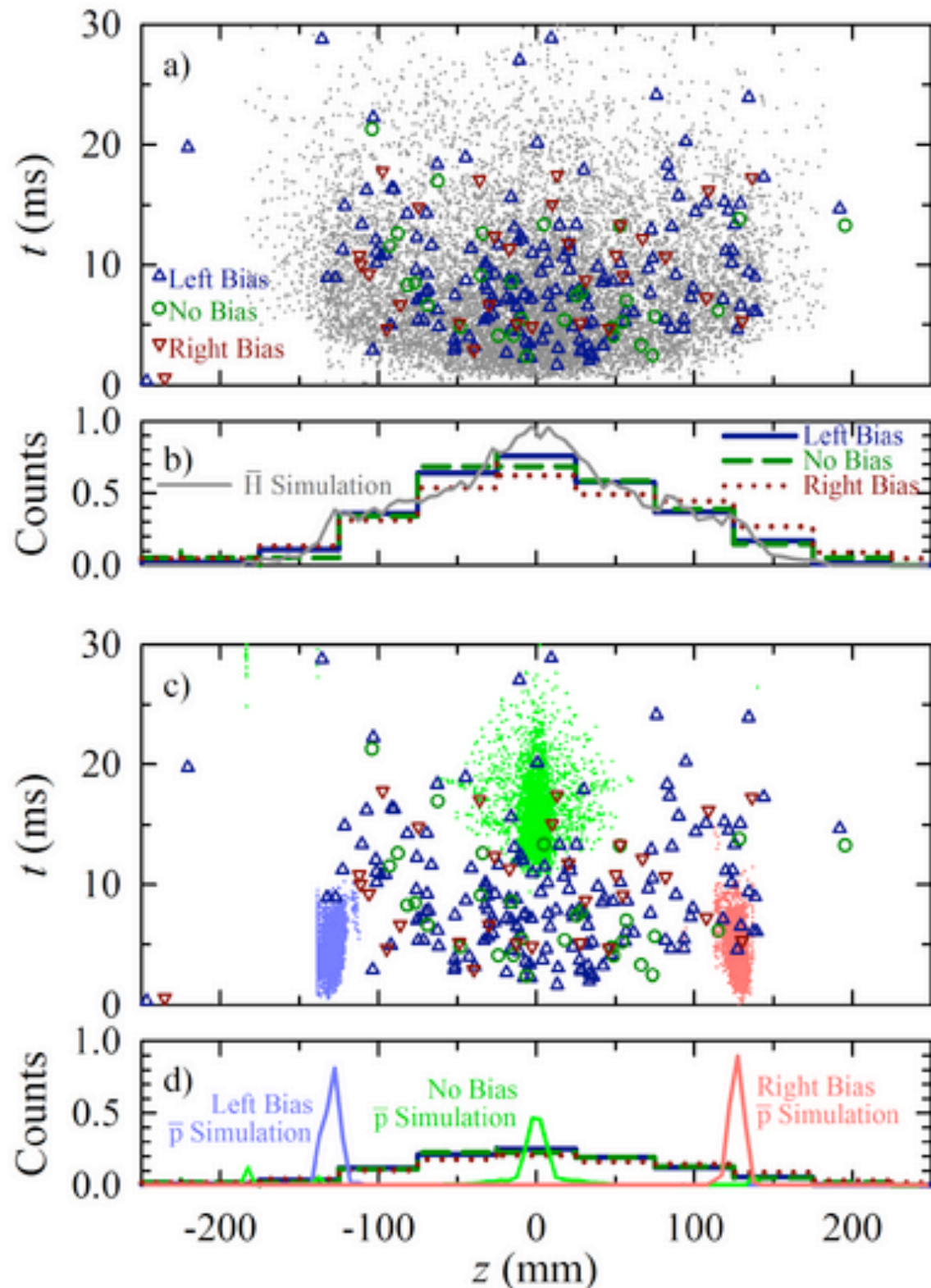


No Bias
Left Bias
Right Bias
(* Heating)



Nature **468**, 673 (2010)

Long time confinement



>300 events!
stored for a 1000s!
I trapped/experiment!

Nature Physics **7**, 558 (2011)

Why “only” ~ 1 per exp. ?

ALP \bar{H} A



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- \bar{H} must be cold to be trapped!
- Many techniques developed to reduce energy.
- BUT : Even 1 atom can be interrogated!

Quantum Transitions



Quantum Transitions

- Trapped atom(s) in the ground state - even if there's only one it is a platform for starting to compare antihydrogen and hydrogen.

ALPHA

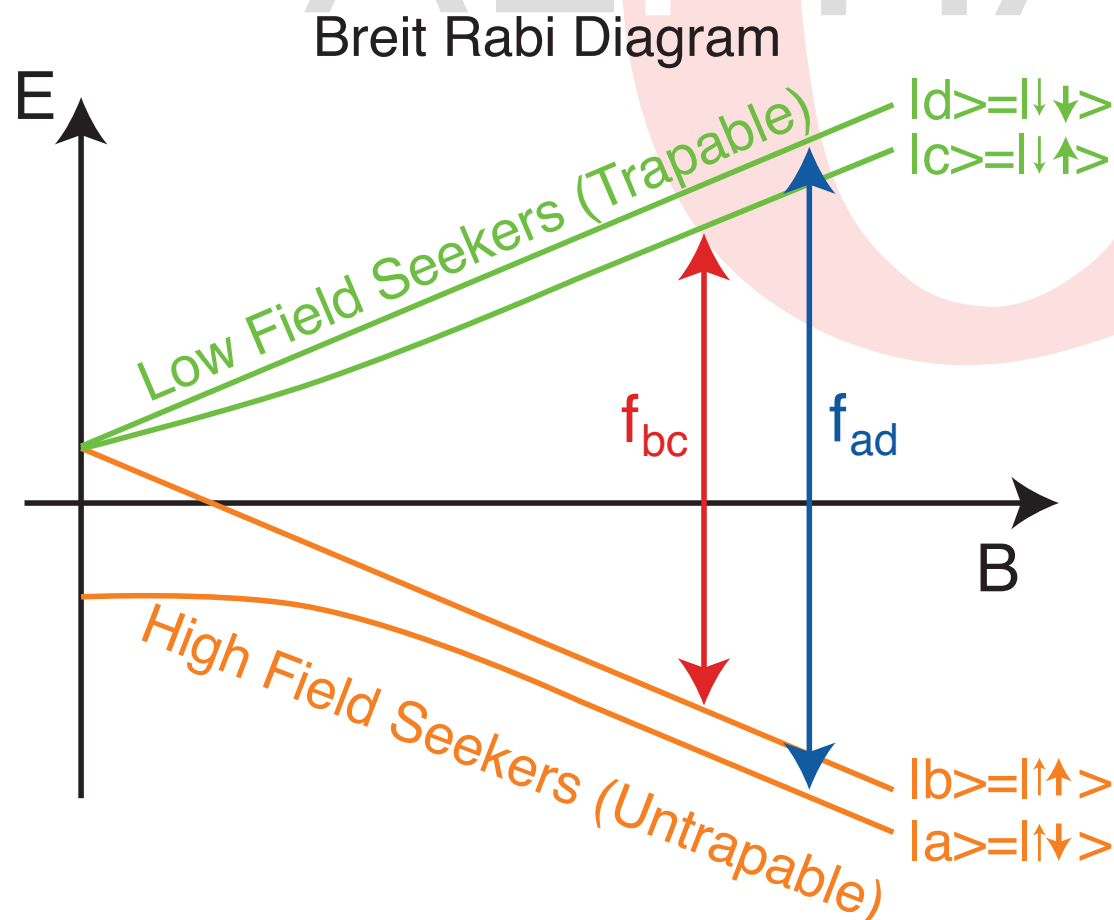
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- Diagnostic of one $\bar{\text{H}}$: Annihilation detection
- Method : Lose $\bar{\text{H}}$ resonantly from trap - spin flip.



Details : Prof. M. Hayden,
Monday 11h00

Nature, March 7th (2012)

Microwave Spectroscopy



microwave
spectroscopy

Disappearance Summary

- Clear decrease observed in on-resonance relative to off-resonance

	Attempts	Antihydrogen left	Rate
ON resonance	103	2	0.02 ± 0.01
OFF resonance	110	23	0.21 ± 0.04
NO μ -waves	100	40	0.40 ± 0.06

$$p = 1.0 \times 10^{-5}$$

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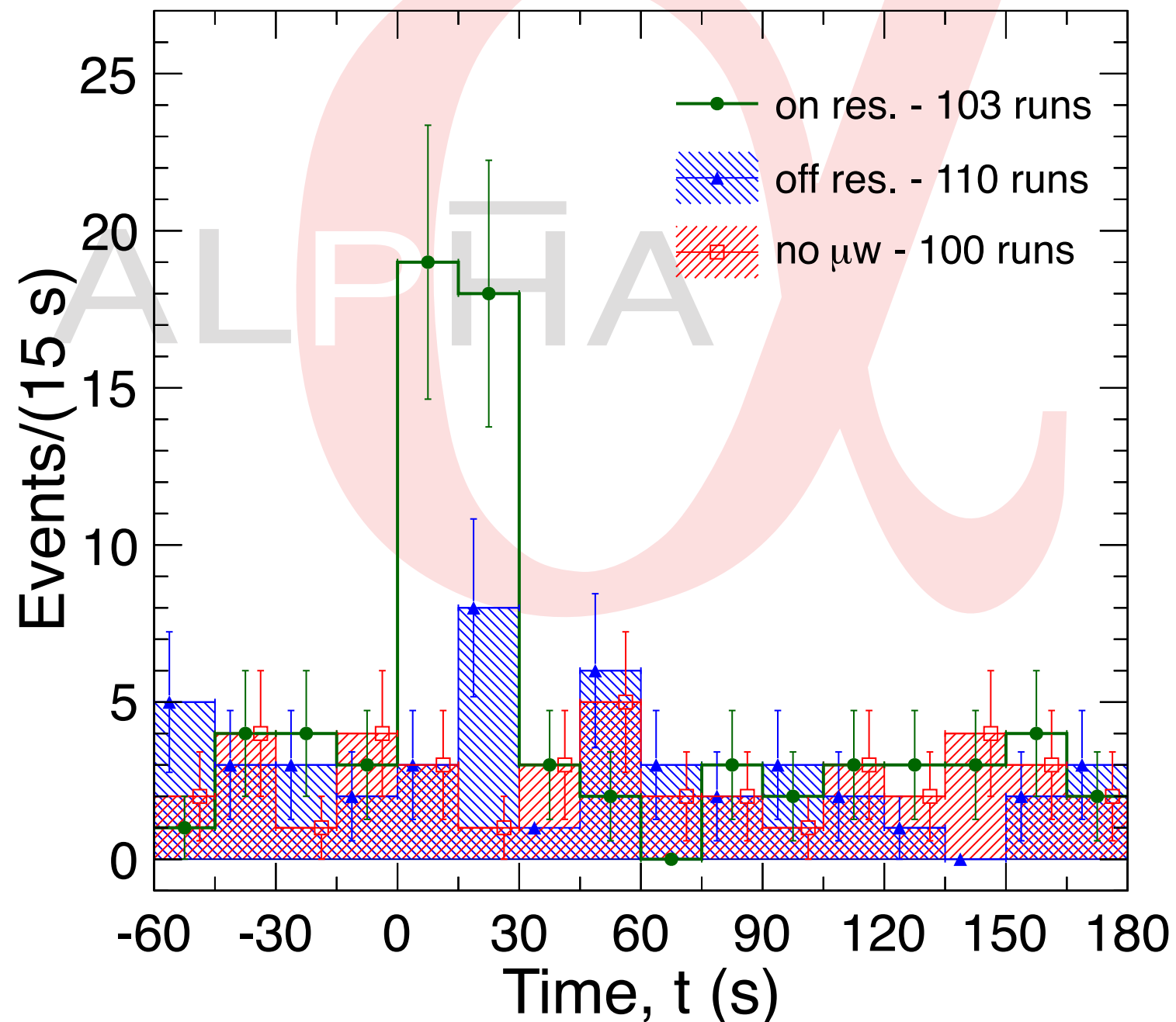
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313 attempts 20min each - 104h - 2w (no hickups)

Nature, March 7th (2012)

Appearance Measurement

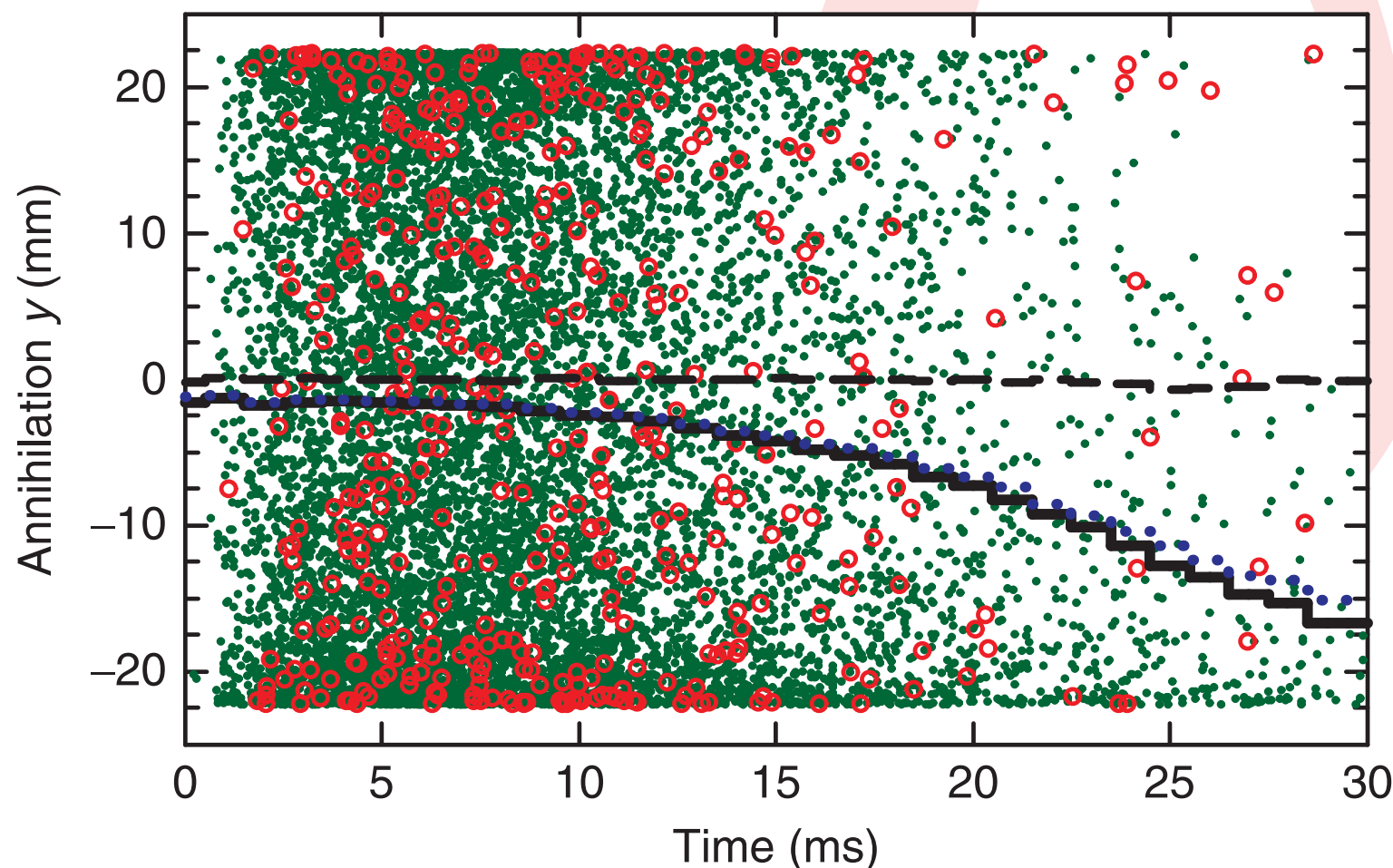
- Significant excess in first 2 sweeps in the 0-15s + 15-30s windows



Nature, March 7th (2012)

First “stab” at Gravity

- Using the position sensitivity of our detector and the “long” shutdown time we’re sensitive to some deviations from standard gravity...



Simulation for
 $M_{\text{gravity}}/M_{\text{inertial}} = 100$

- Details: Friday morning 9h30, Prof. J. Fajans

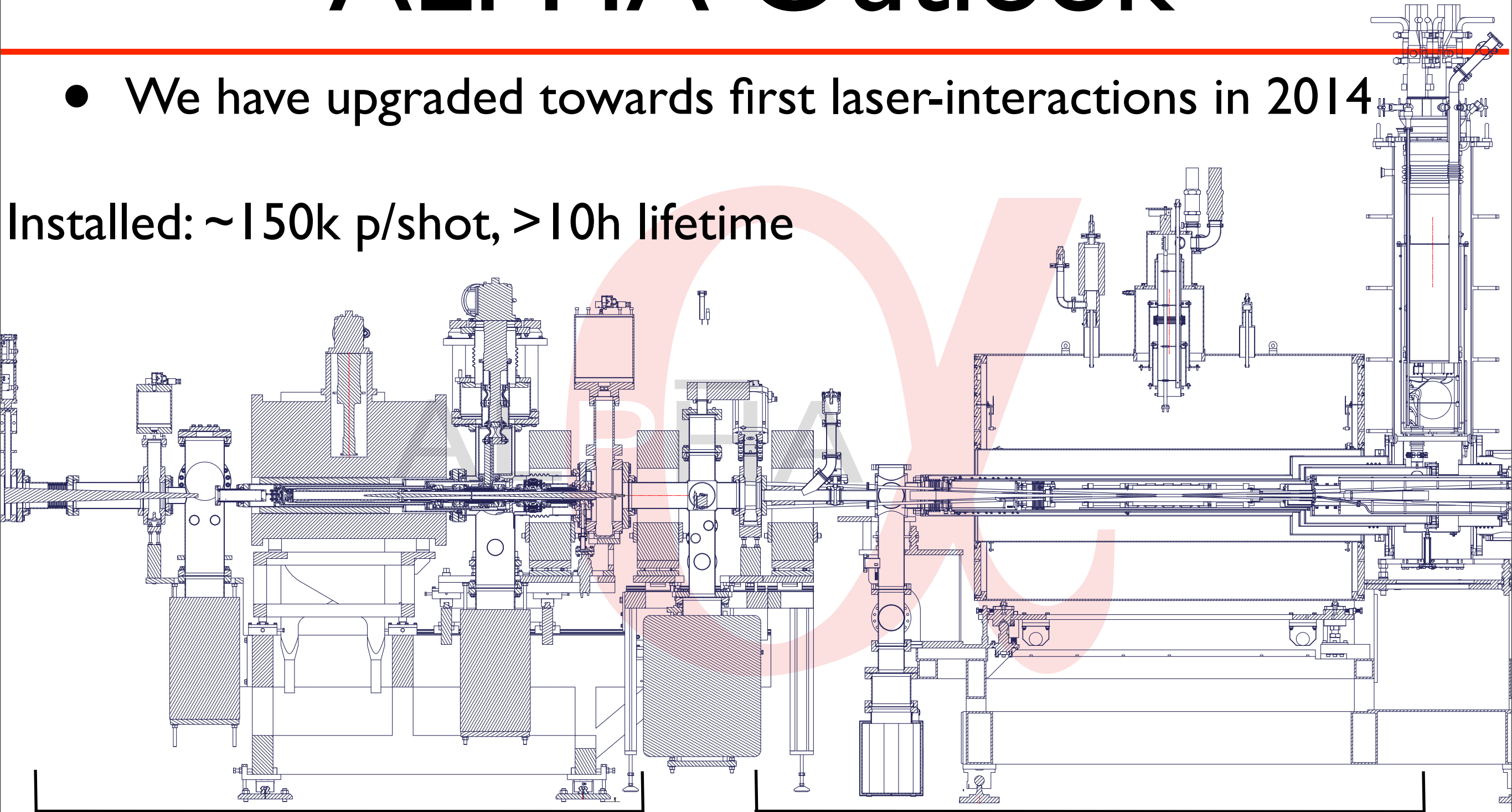
Where are we now



ALPHA Outlook

- We have upgraded towards first laser-interactions in 2014

Installed: $\sim 150\text{k p/shot}$, $> 10\text{h}$ lifetime



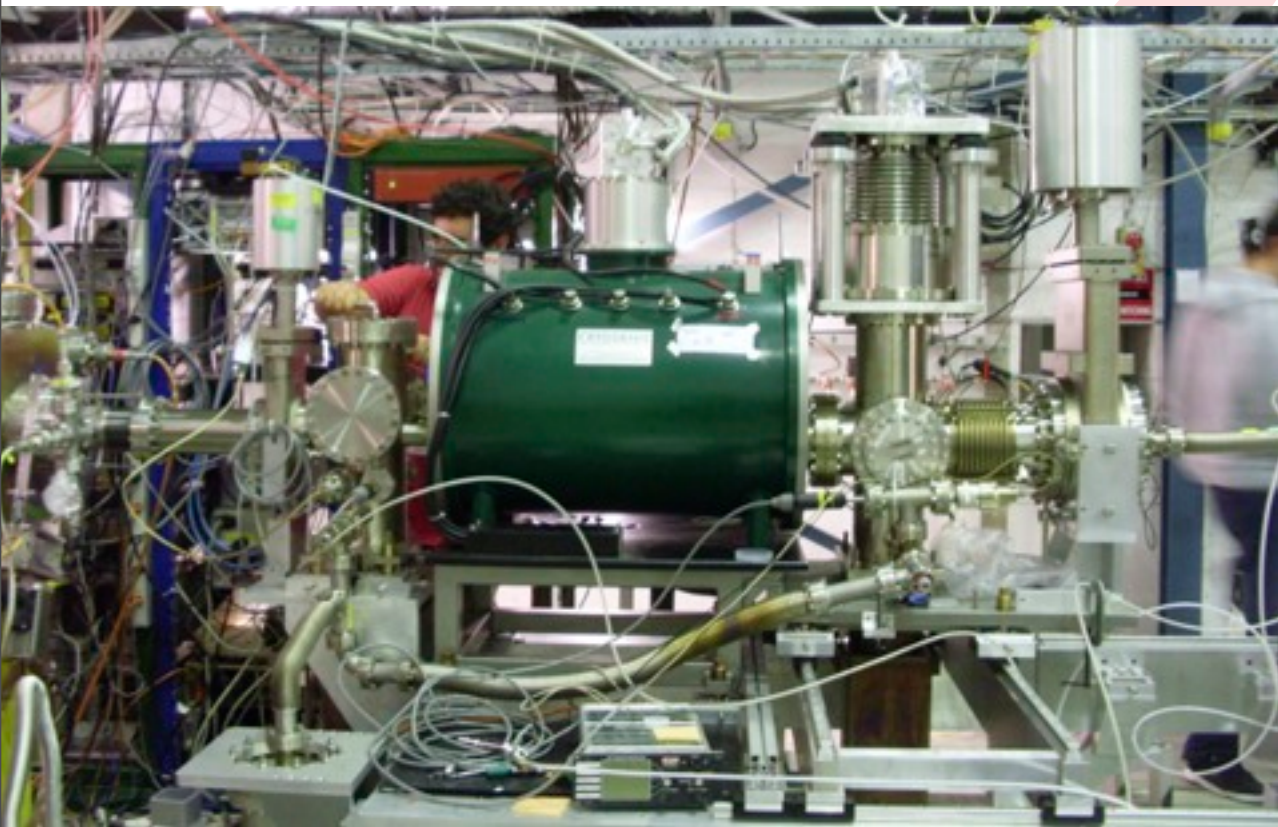
\bar{p} catch and accumulation

\bar{H} formation, trap and spectroscopy

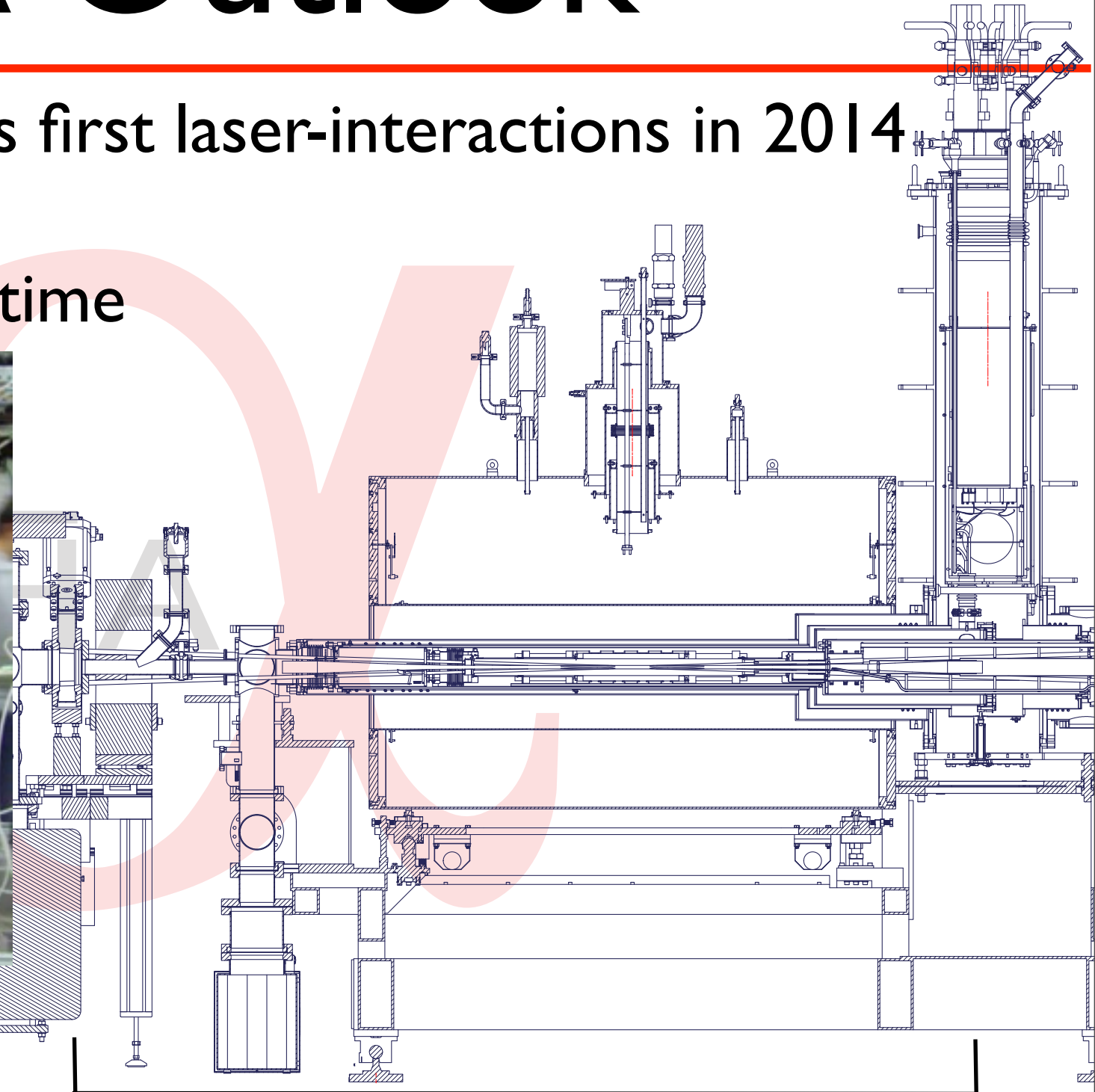
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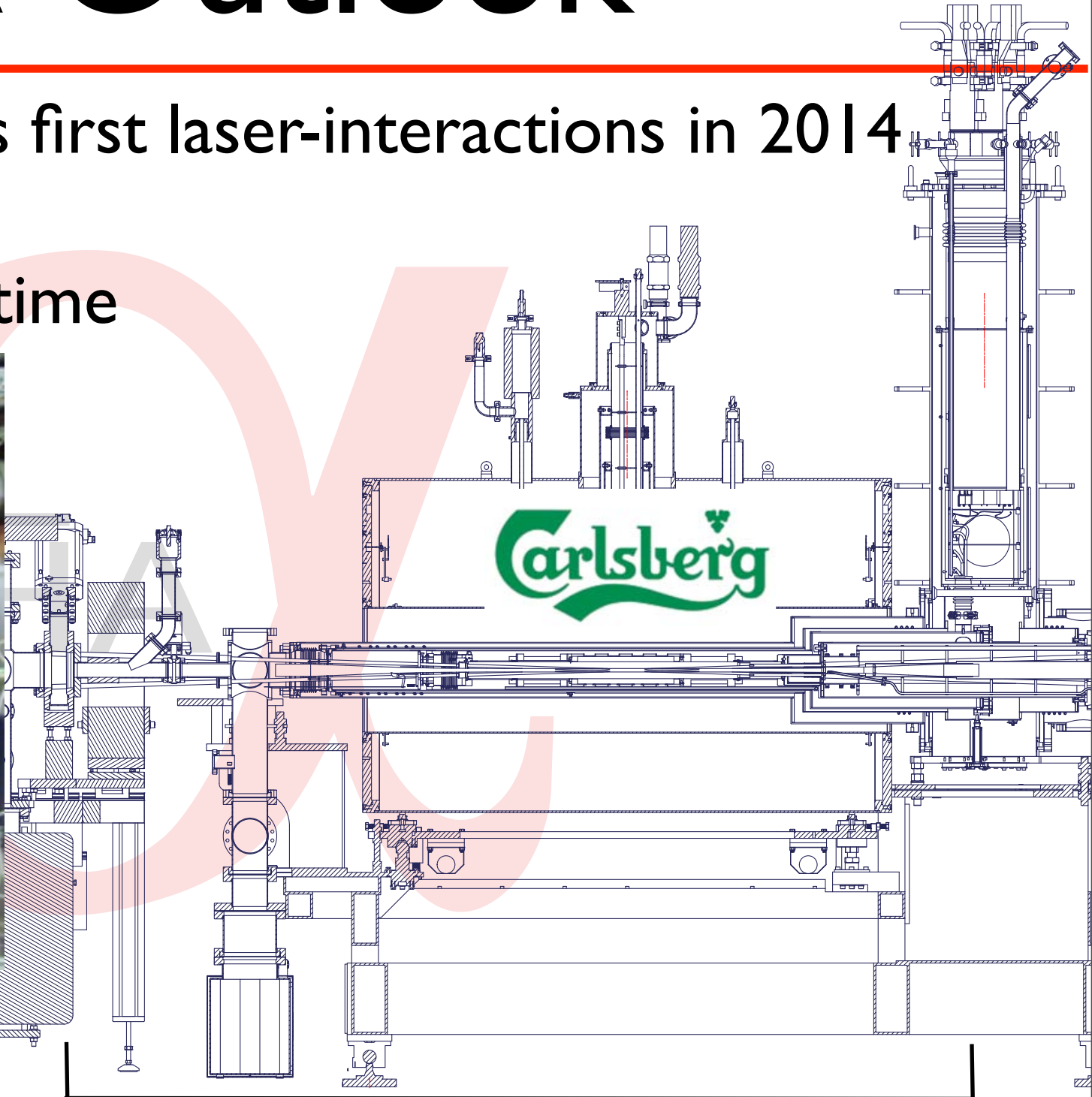
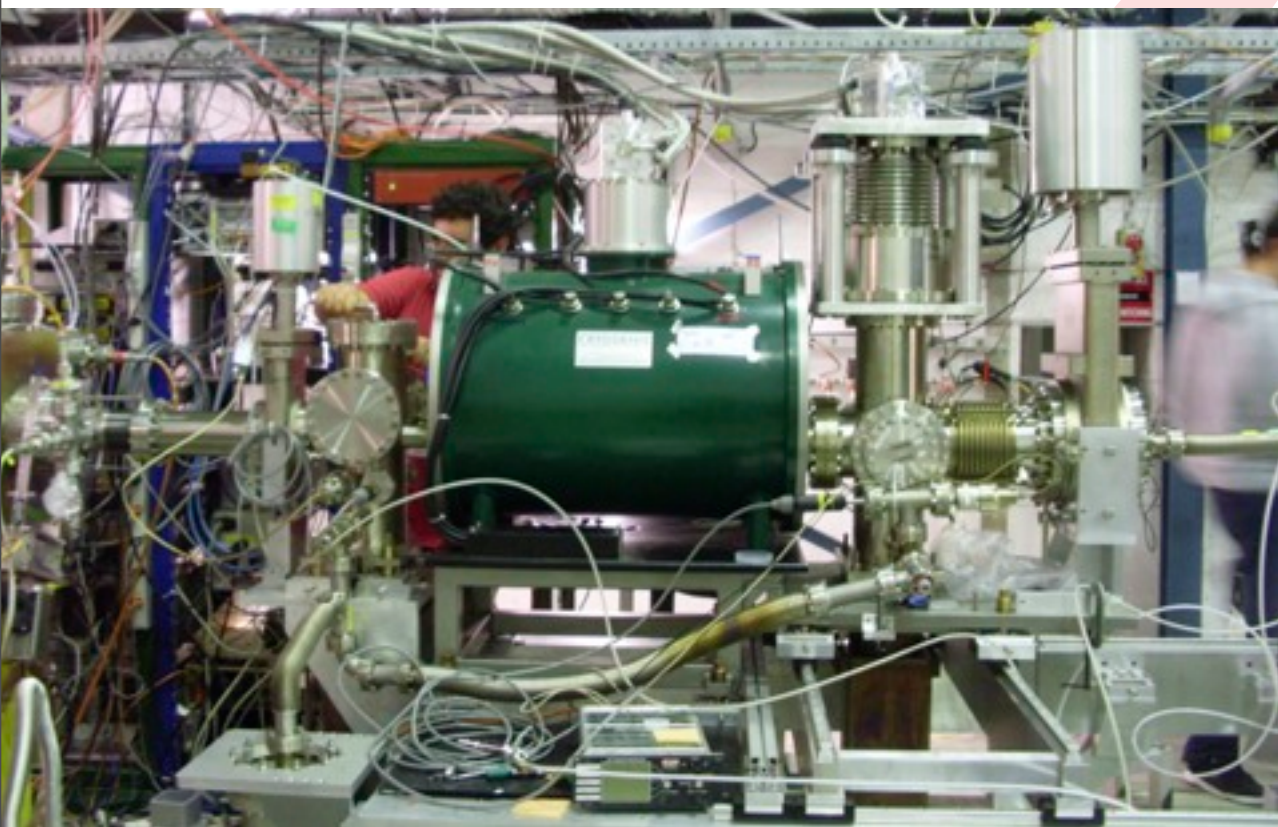


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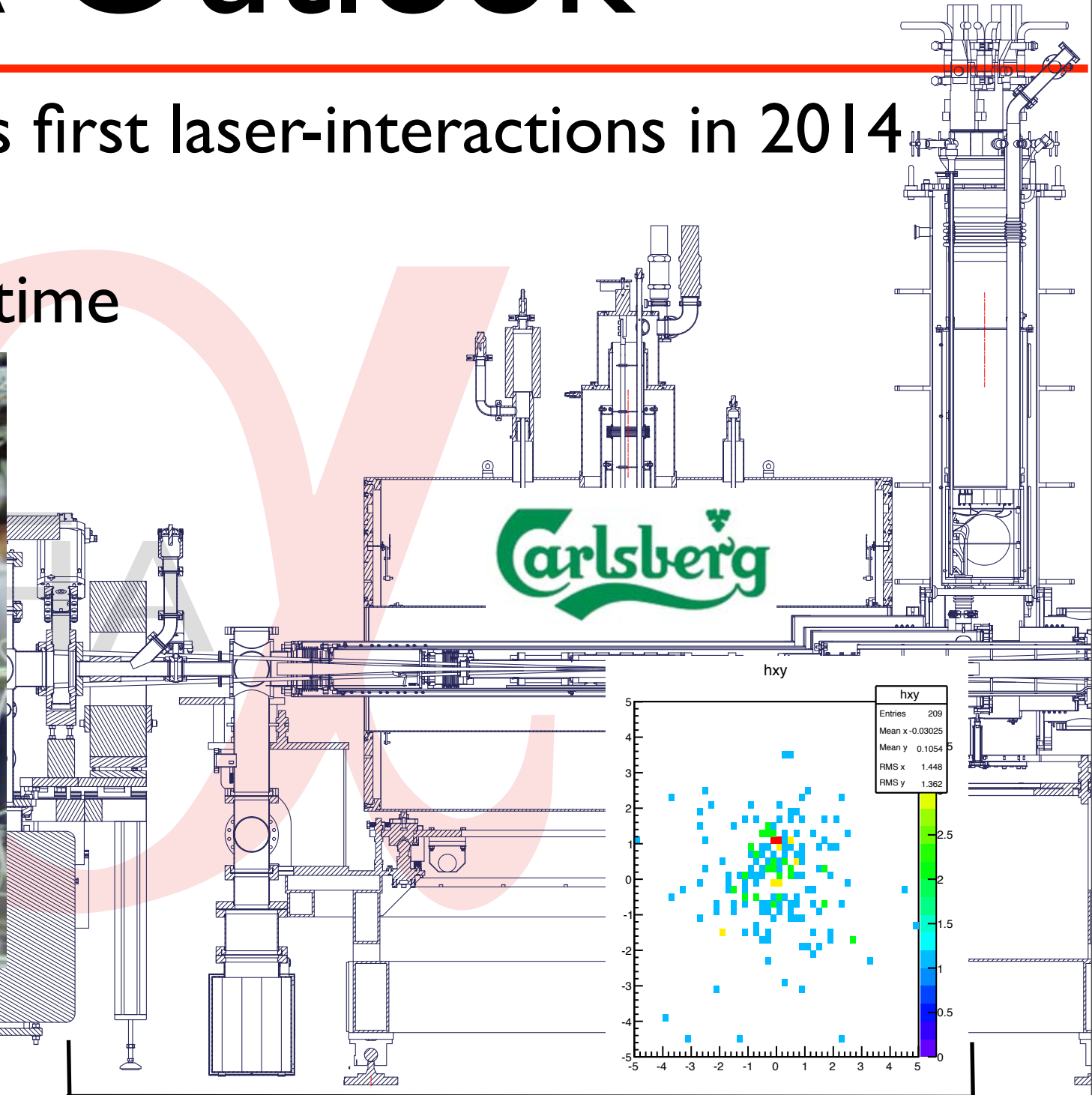
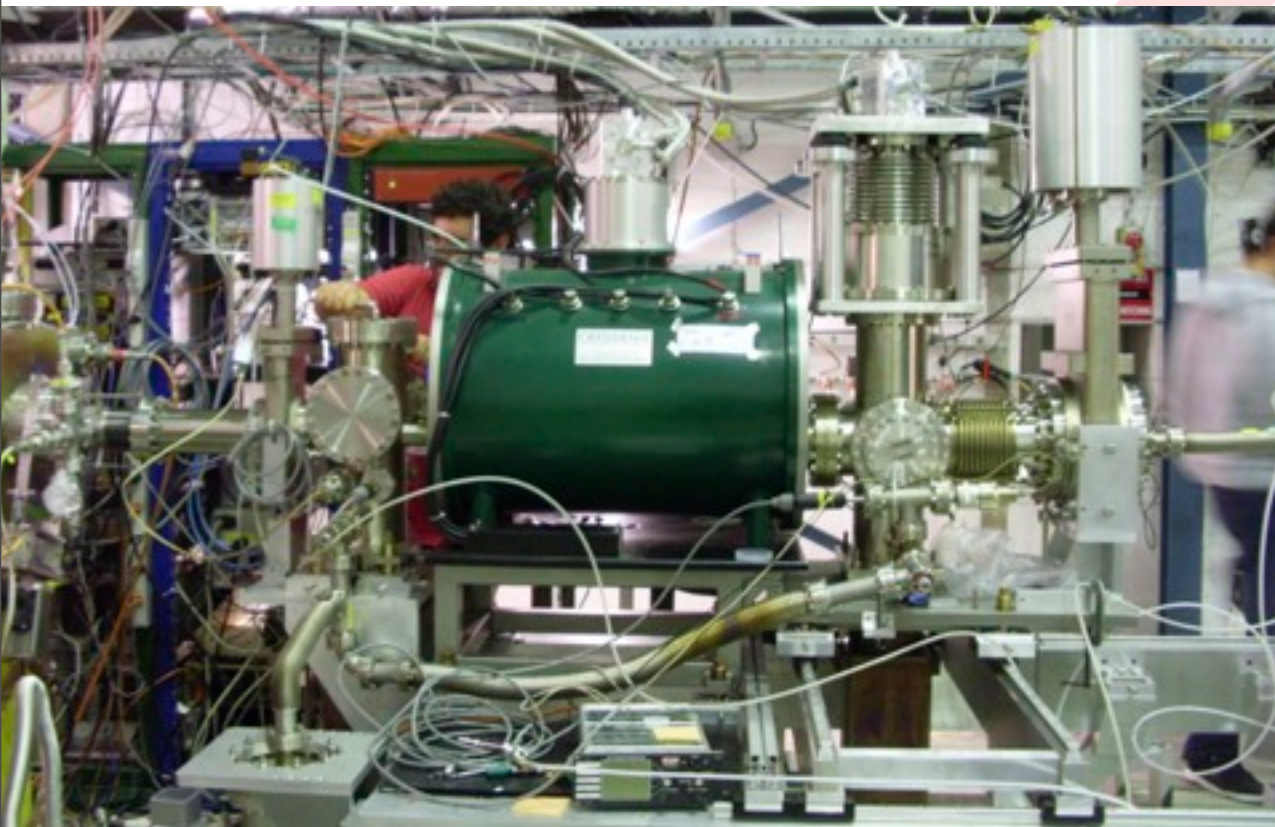
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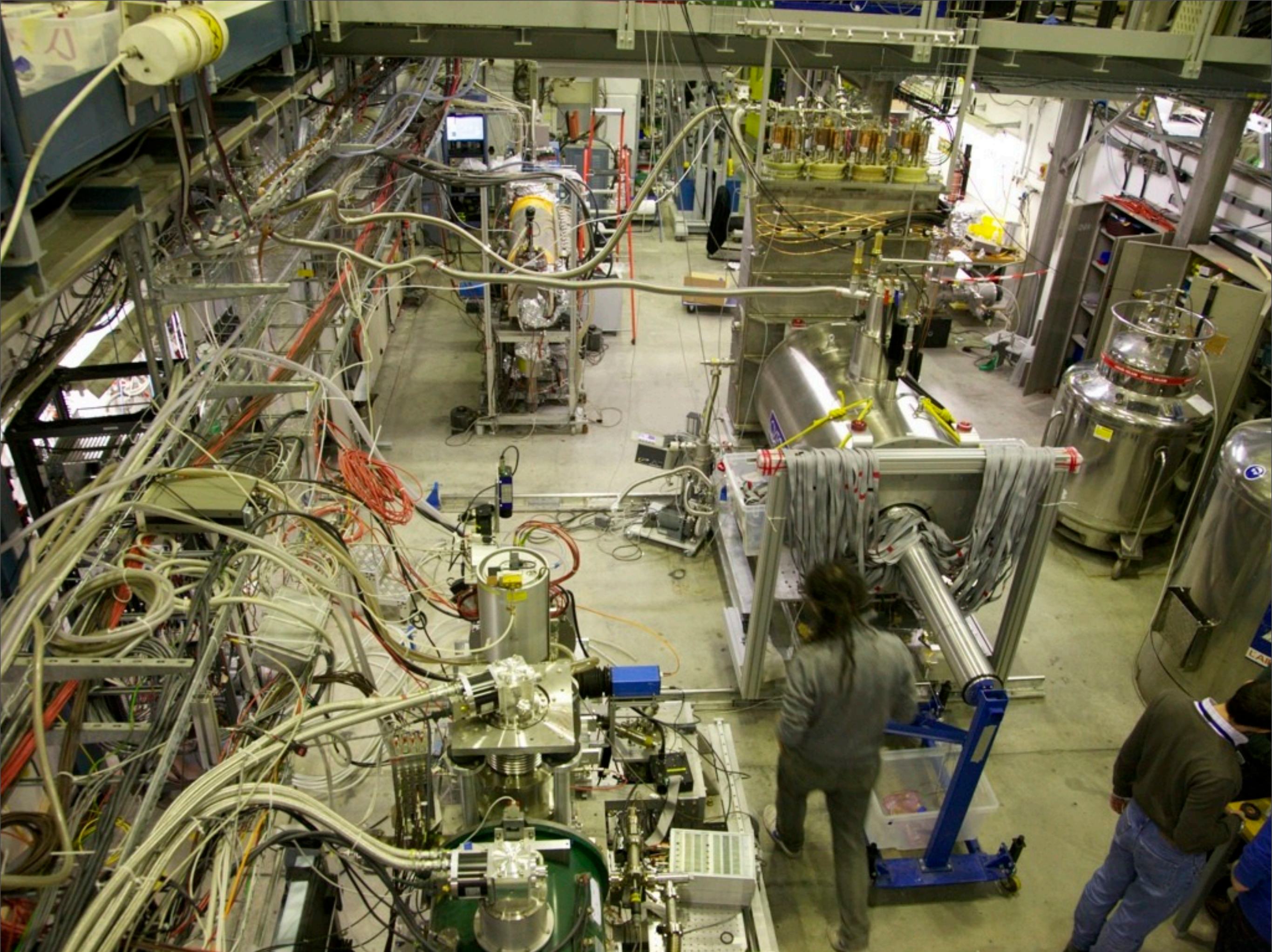
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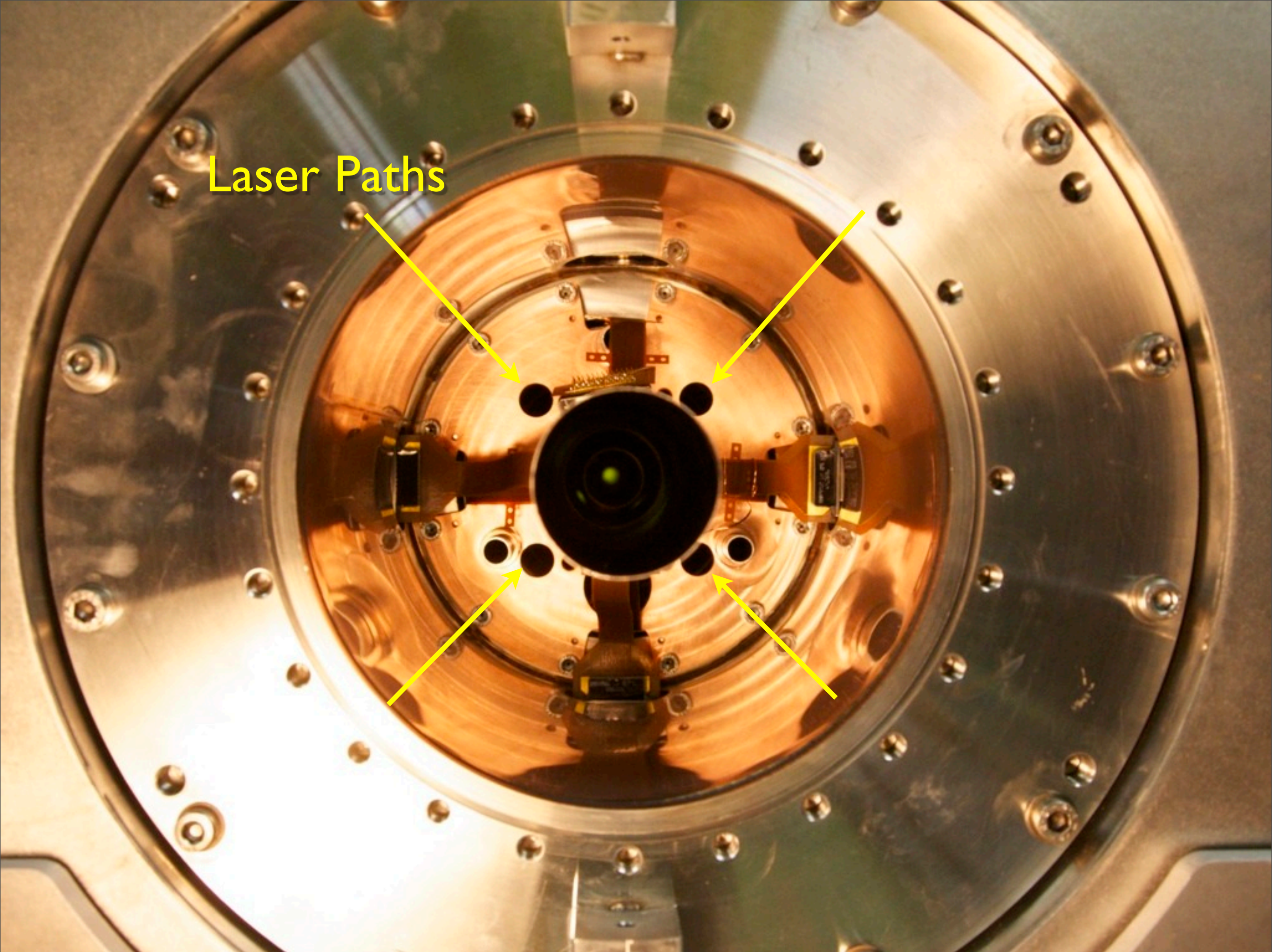
\bar{H} formation, trap and spectroscopy

Modular approach - ELENA & “ \bar{g} ” ready

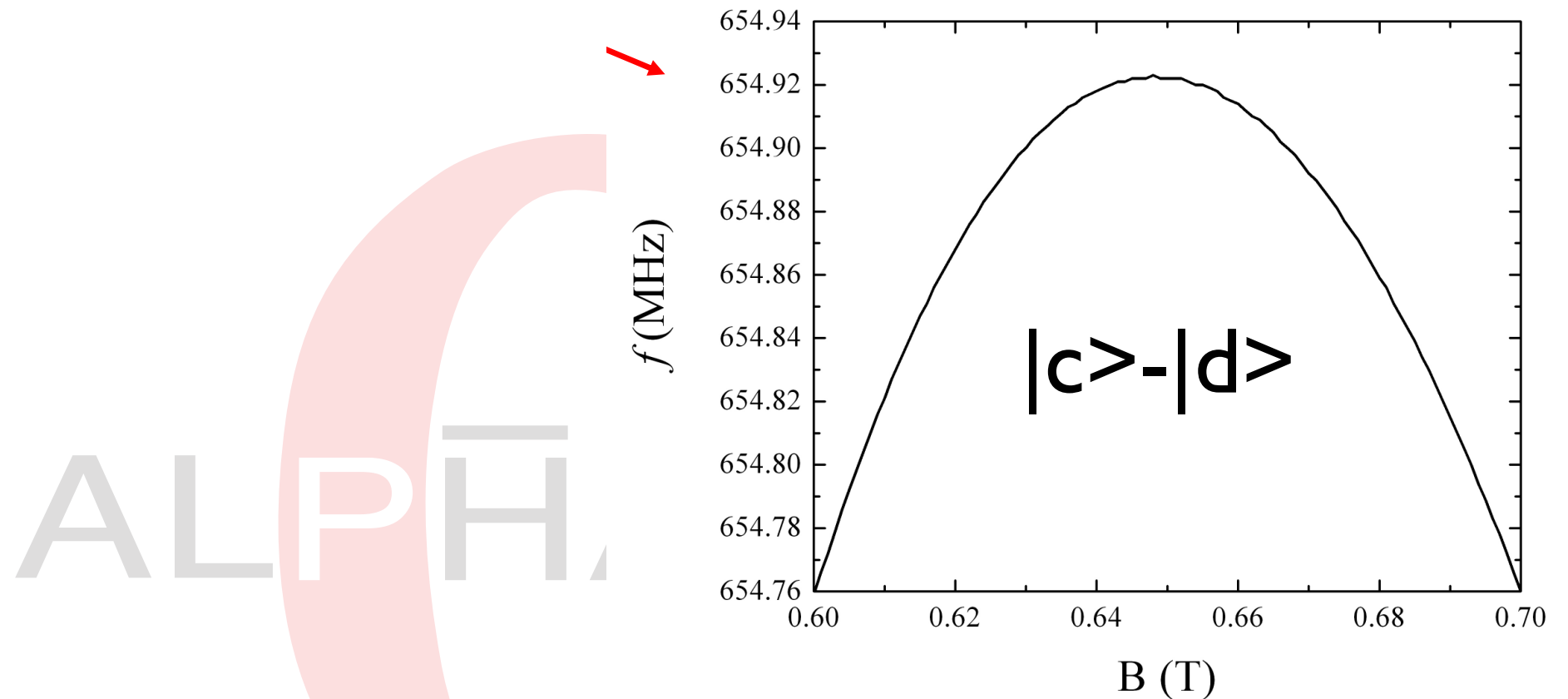


Tuesday, June 11, 13

Laser Paths

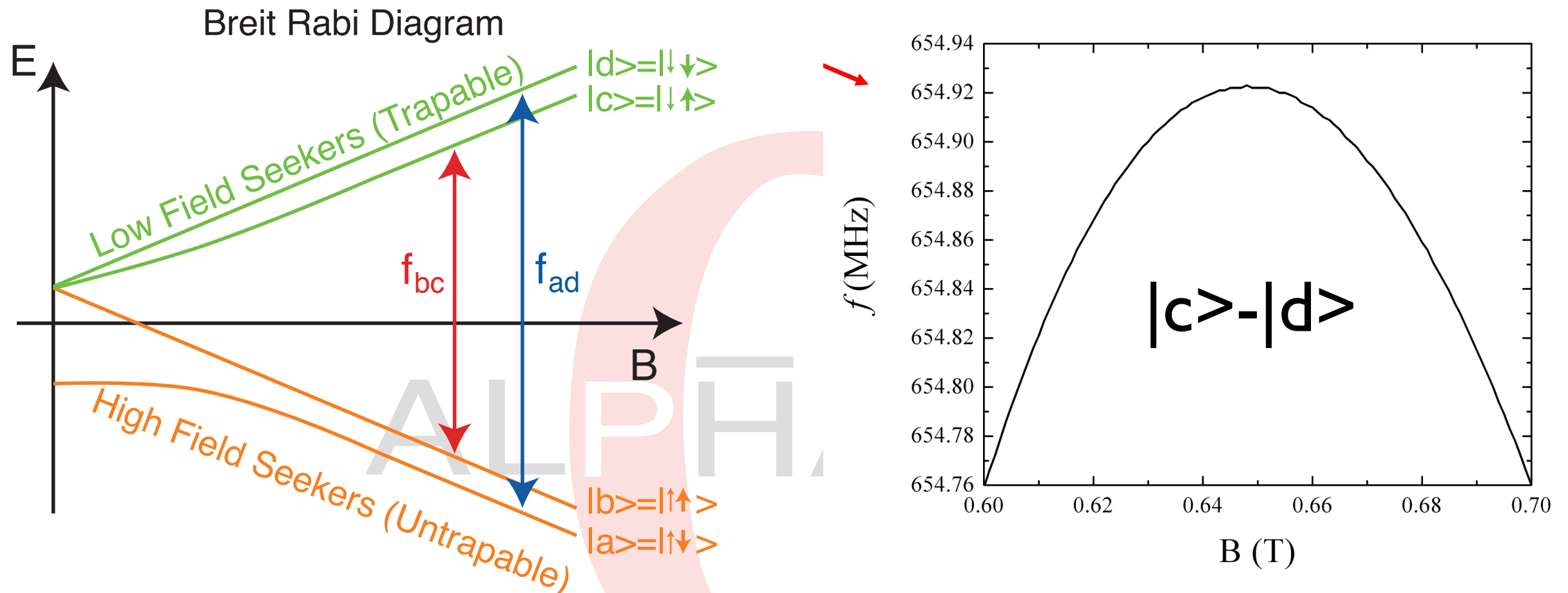


μ -waves continued



- NB: Antiproton μ - known only to $\sim 10^{-3}$
- We can also measure the Hyperfine splitting ($f_{ad} - f_{bc}$) which is independent of B !
- Upgrade includes μ -waves! $10^{-6} \rightarrow 10^{-8}$ potential!

μ -waves continued



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Pre-requisite... and preps...

- While micro-wave and gravity measurements have been demonstrated with ~ 1 atom trapped/experiment (every ~ 15 min) and laser experiments are pending return of \bar{p} s - it's not ideal and will pose a limit to the ultimate precision.
- We seem limited by the plasma (in particular e^+) temperatures...
- Upgrade and other initiatives are aimed at making the e^+ colder for increased trapping rates.
- Successful commissioning of the ALPHA-upgrade cryostat + ~ 1 y to tune all our “guns” = steady progress towards first laser-interaction with \bar{H} in the fall of 2014...



Thank you for listening.