

First Results of 2023 Juelich B-field measurements

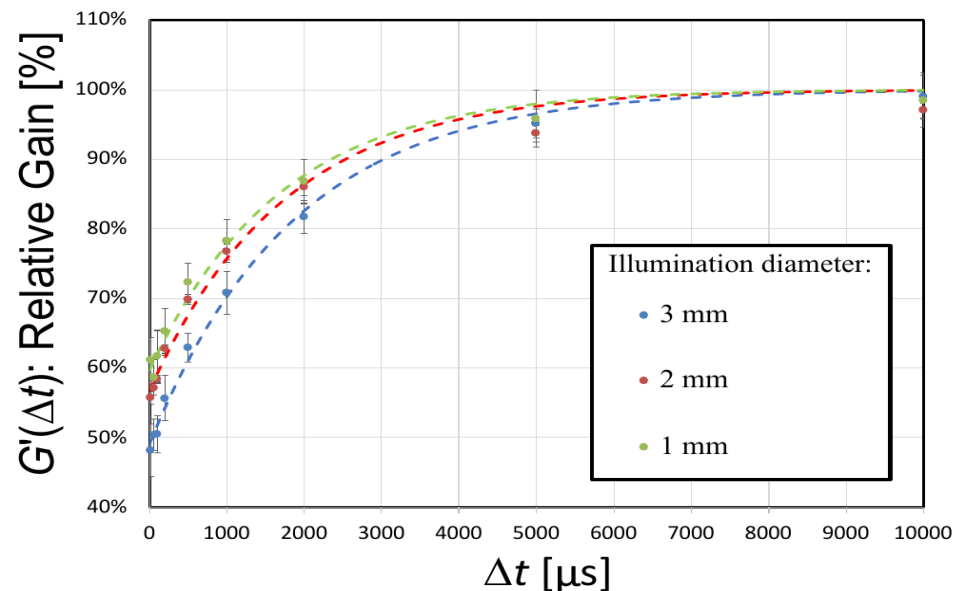
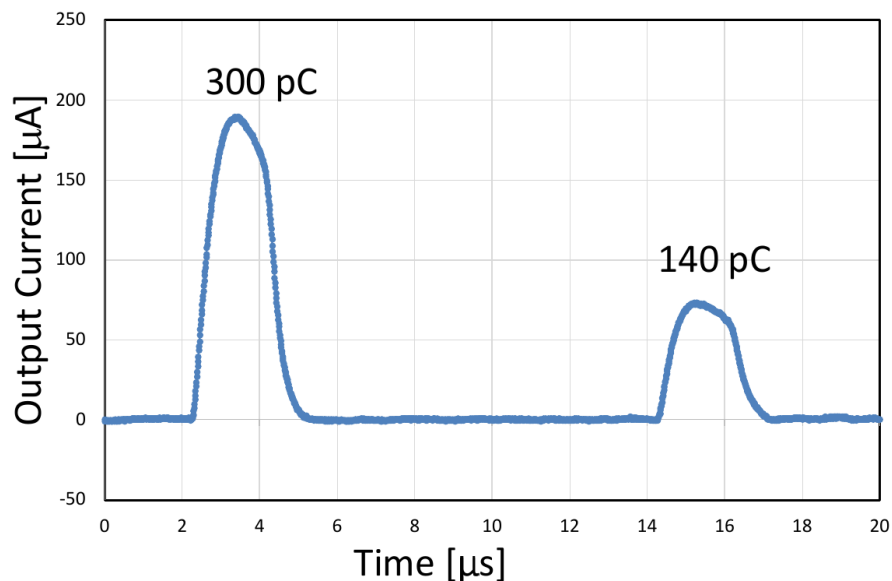
ERLANGEN CENTRE
FOR ASTROPARTICLE
PHYSICS

D. Miehling, M. Götz, K. Gumbert,
S. Krauss, A. Lehmann

PANDA-Meeting 23/2, Jun 13, 2023



Gain-recovery time by the double-pulse method

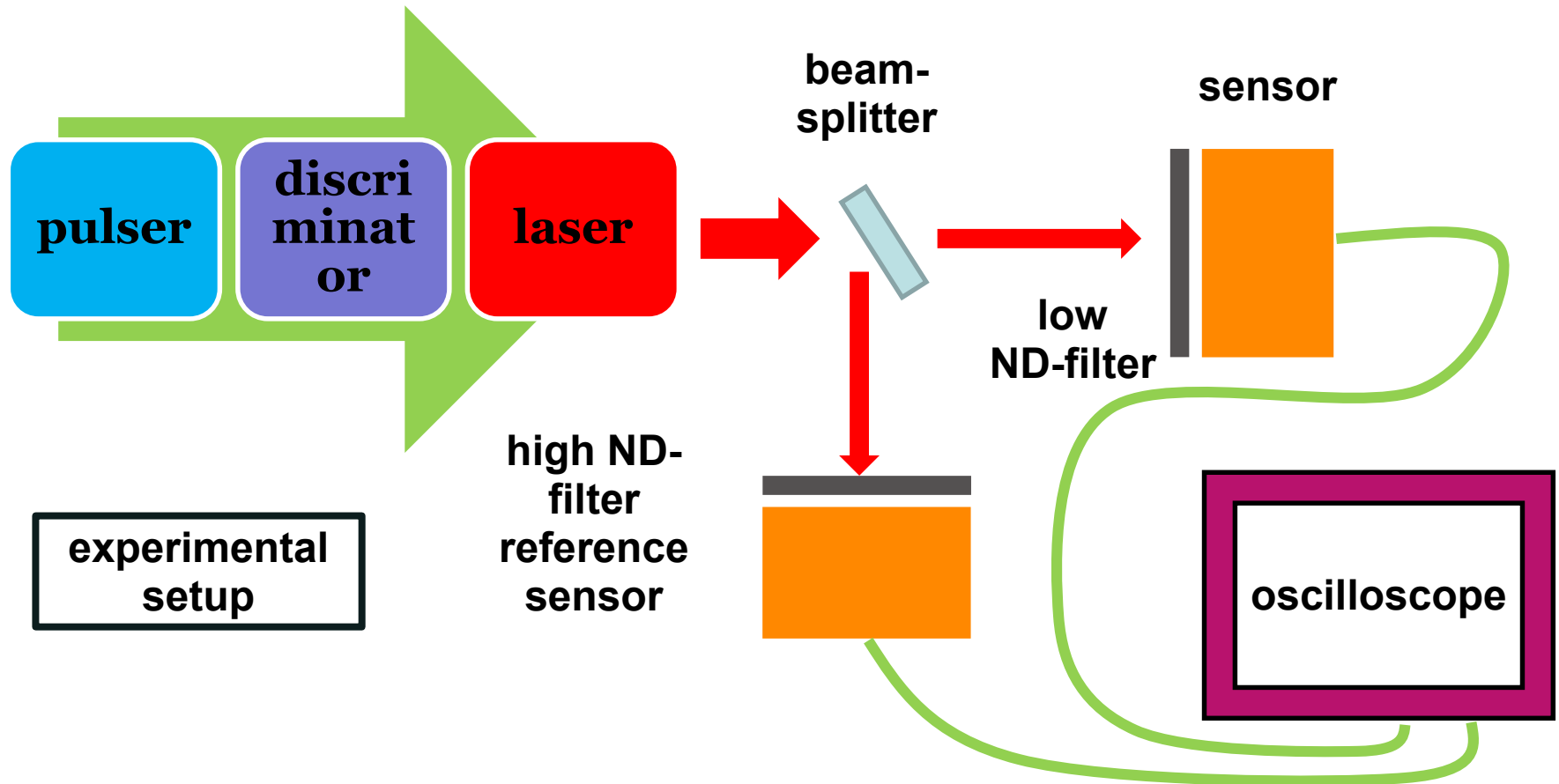


Result of the research group in Japan:

Left: waveform acquired by averaging 100 waveforms at $\Delta t = 10$ ns, illumination area 3 mm in diameter.

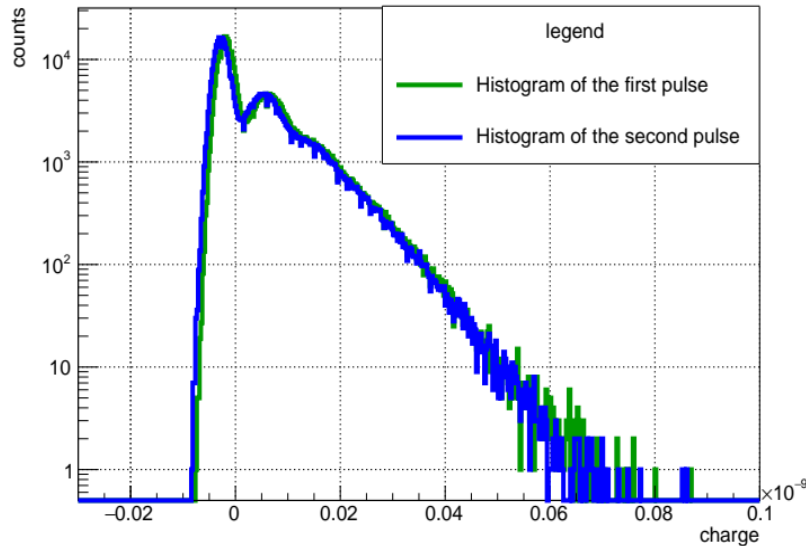
Right: relative gain over the time interval Δt between both pulses → Our goal: comparable curve

Gain-recovery time by the double-pulse method



Gain-recovery time by the double-pulse method

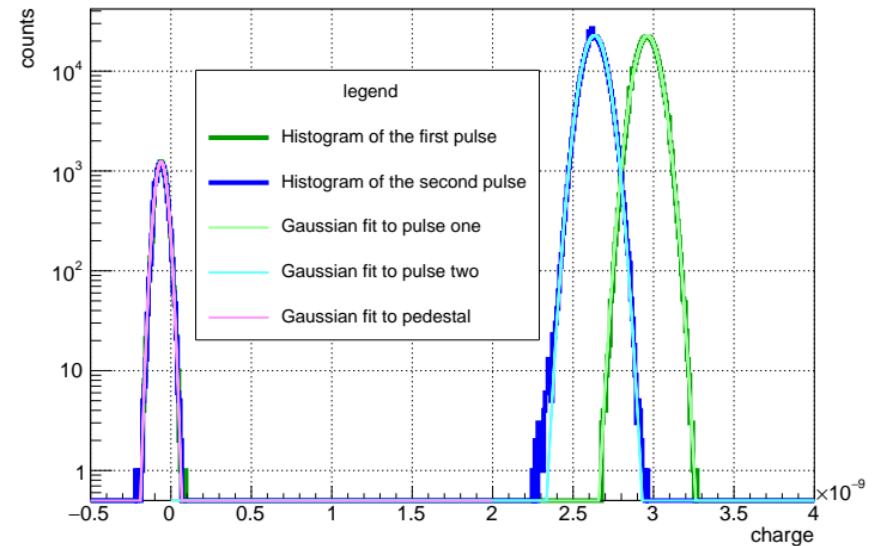
Histogram of reference sensor charge



Left: Histogram of the reference sensor at a delay of 360 ns and ND filter of 4

→ laser pulsed correctly

Histogram of sensor charge



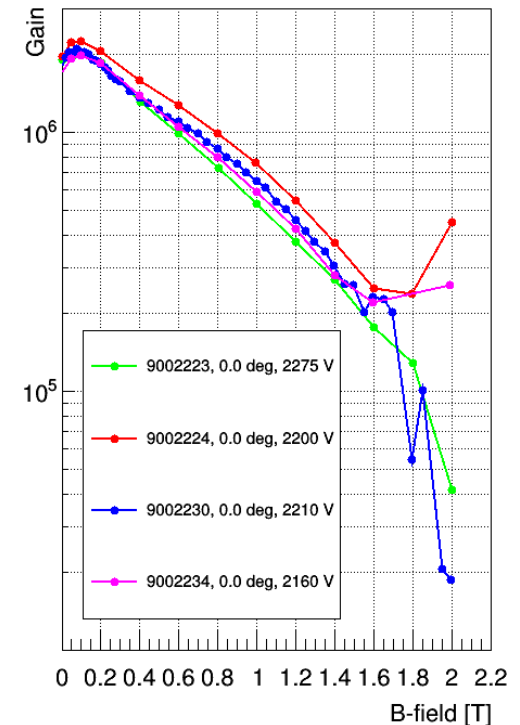
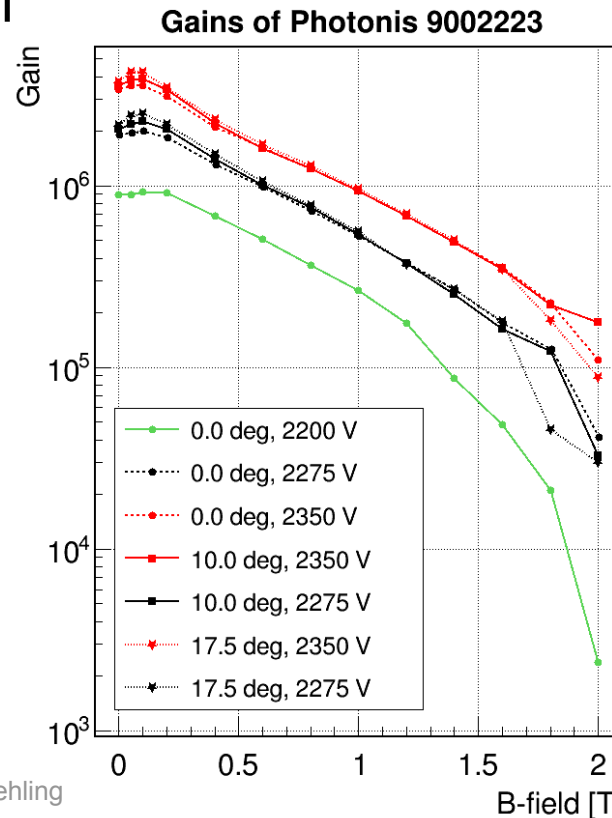
Right: Histogram of the sensor at a delay of 360 ns and ND filter of 0.6

→ Relative Gain = 87%

Experiment still in early stage, more data will be taken in the next weeks

gain curves in magnetic field

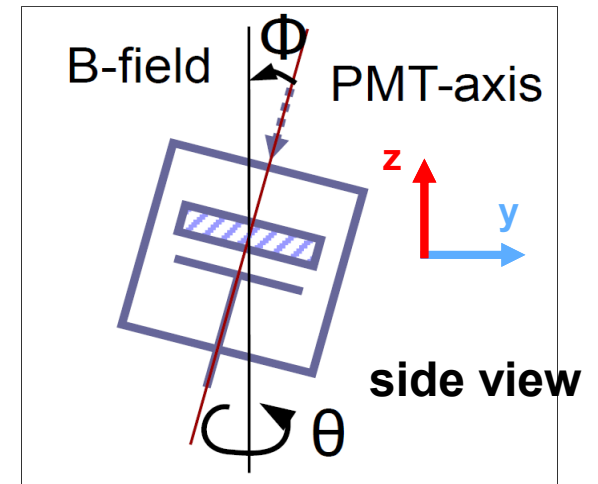
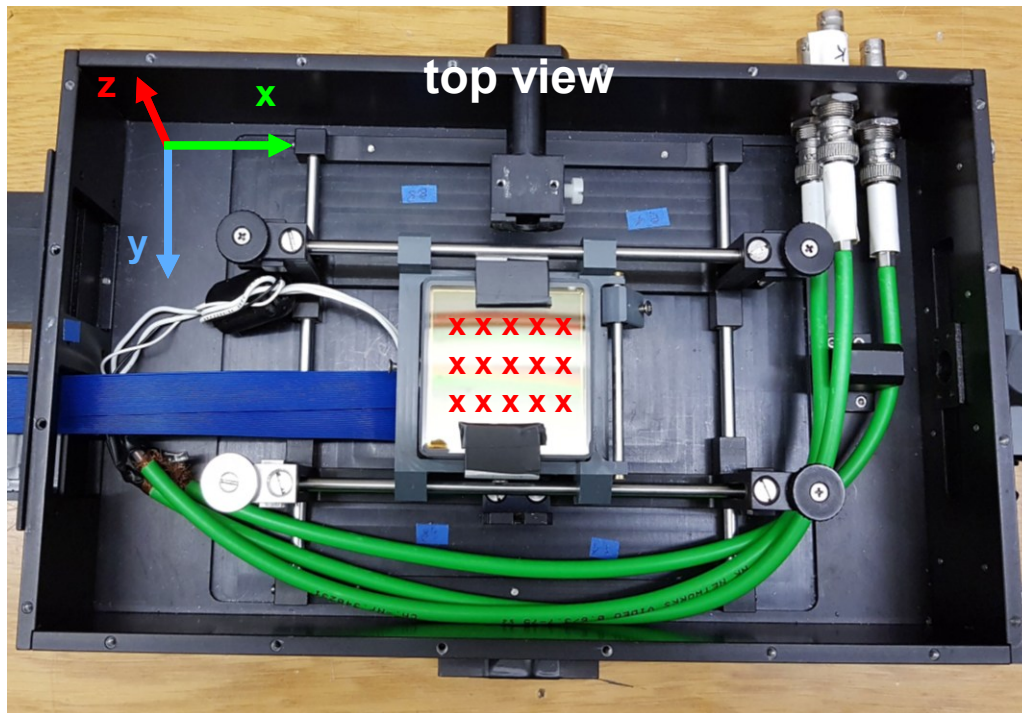
- measurement of gain for different magnetic fields, voltages and tilting angles
- at least 1 sensor per batch measured, only 4 shown
- last points: gain too low for fit
- factor ~ 3 gain loss at 1T
- all tilting angles look similar
- all tubes look similar



shifts of charge cloud centroid in B-field

● Measurement:

- tilting of the tube in **yz** plane around **x** axis, scanning in **y** direction, B-field in **z** direction, E-Field along **yz** plane



Simulation of charge cloud centroid in B-Field



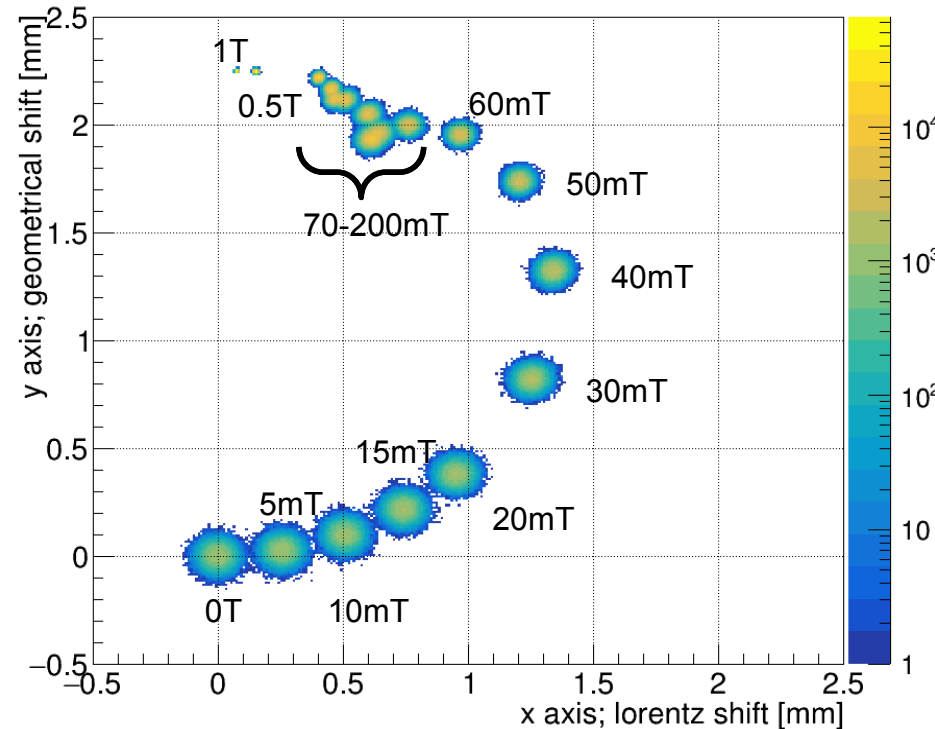
ERLANGEN CENTRE
FOR ASTROPARTICLE
PHYSICS

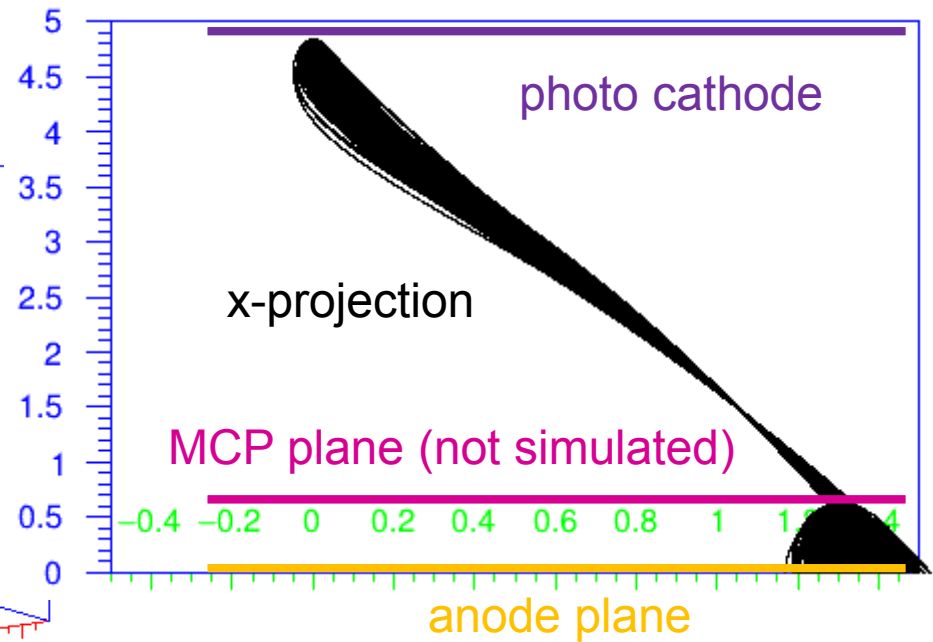
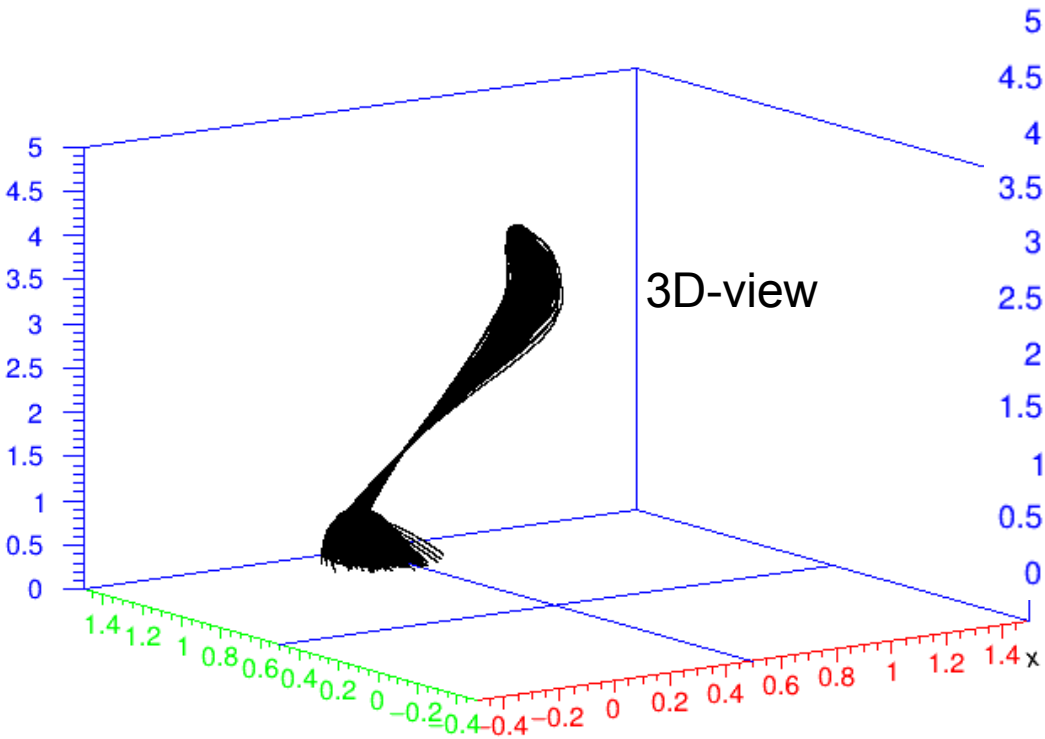


FRIEDRICH-ALEXANDER
UNIVERSITÄT
ERLANGEN-NÜRNBERG
NATURWISSENSCHAFTLICHE
FAKULTÄT

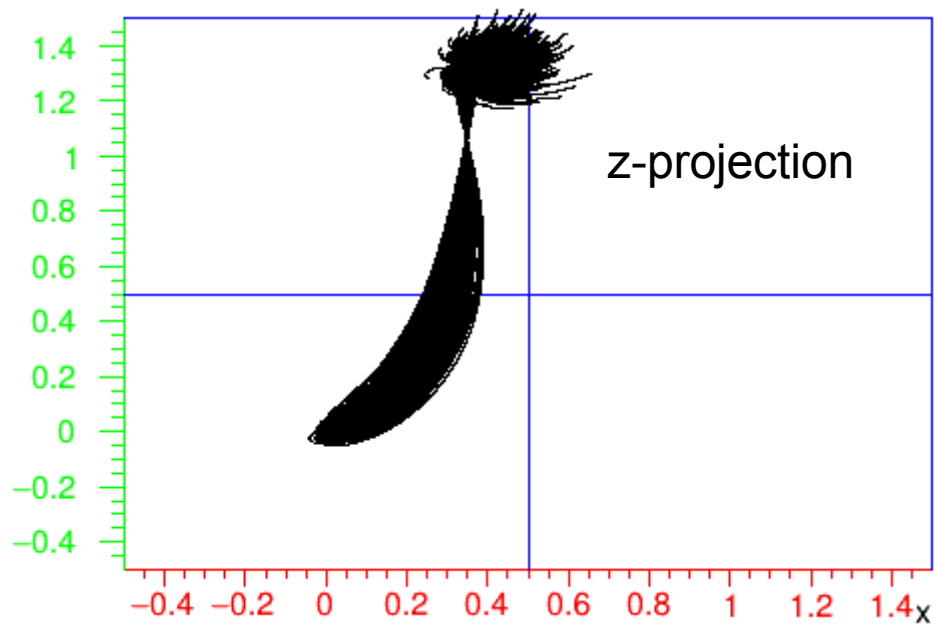
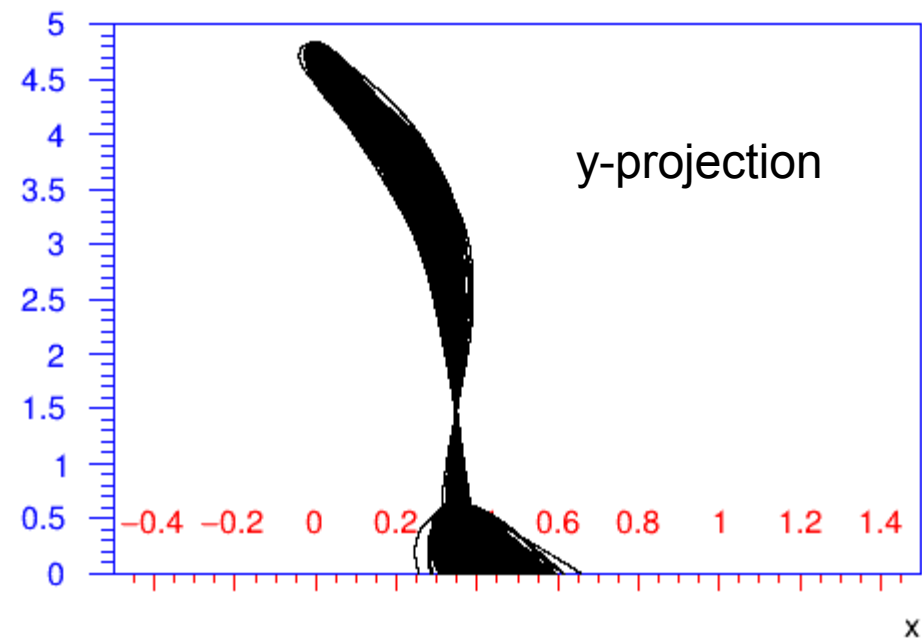
- **geometrical shift** of charge cloud in **y** direction
 - electrons will follow B-field direction, depends on tilt angle, saturates at certain B-field strength
- **lorentz shift** of charge cloud in **x** direction
 - electrons experience Lorentz force perpendicular to E- and B-field depending on **B-field strength & tilt angle** (higher Φ increases E_{\perp} component)

simulated hit positions @ 25deg tilt angle





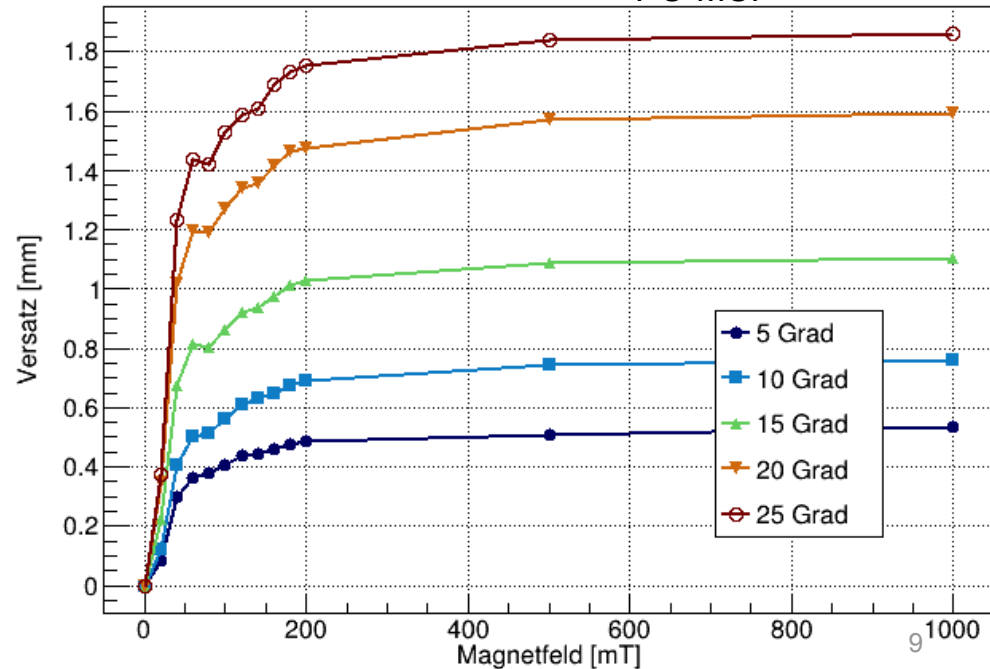
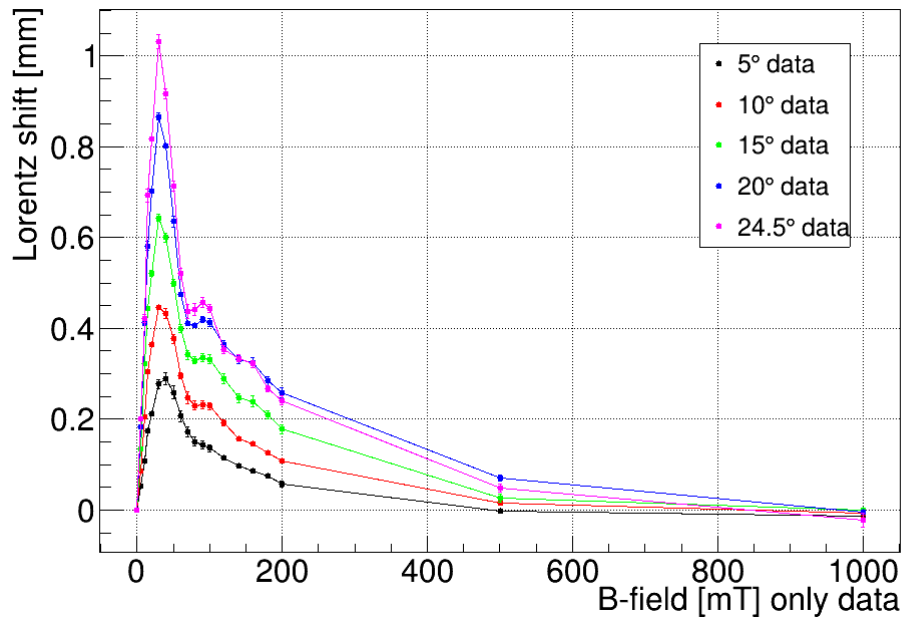
simulation for 17.5 deg and 80mT



shifts of charge cloud centroid in B-field

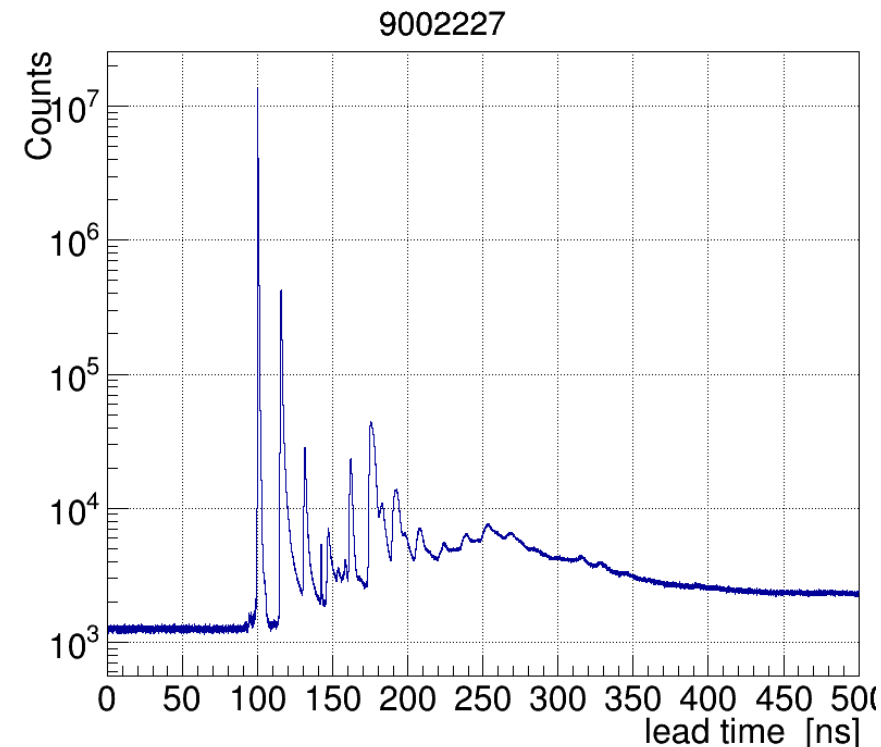
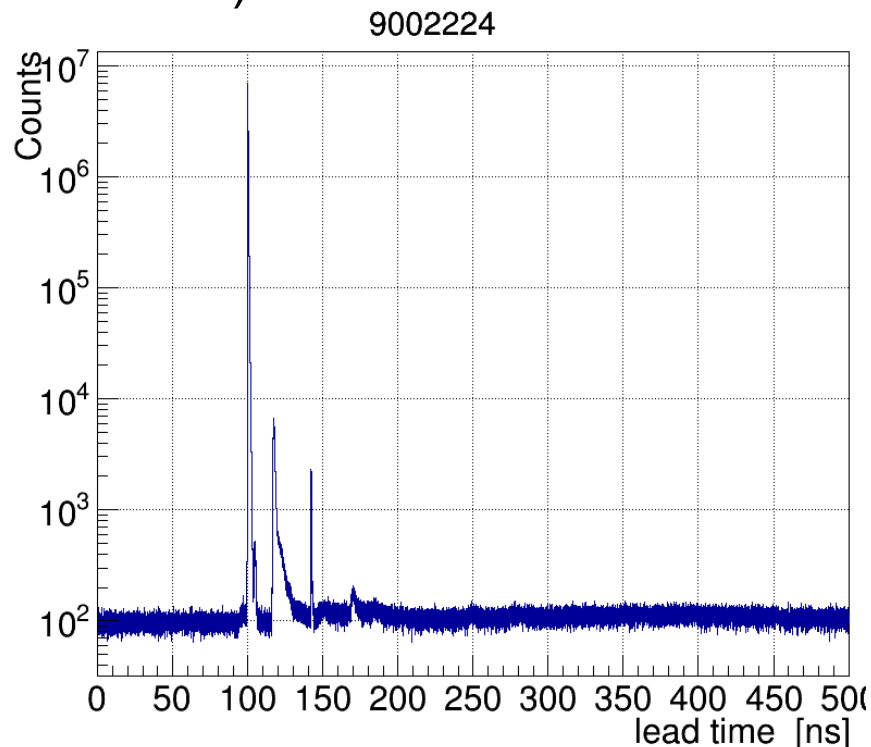


- measurement of Photonis 946P541 (3x100 pixels)
- **lorentz shift** and **geometrical shift** shown for different tilting angles below
- **lorentz shift** peaks at ~ 40 mT and goes back to almost zero
- **geometrical shift** saturation starts at a few hundred mT, as expected
 - at 25° tilt angle, saturation at ~ 1.8 mm shift \rightarrow 3 – 4 pixels shift for EDD
 - also at 15° tilt angle & 1 T B-field \sim mm shift \rightarrow geometrical shift not negligible for Barrel DIRC
- both depend on the inner dimensions of the tube (especially d_{PC-MCP} !)



Afterpulsing of 9002224 and 9002227

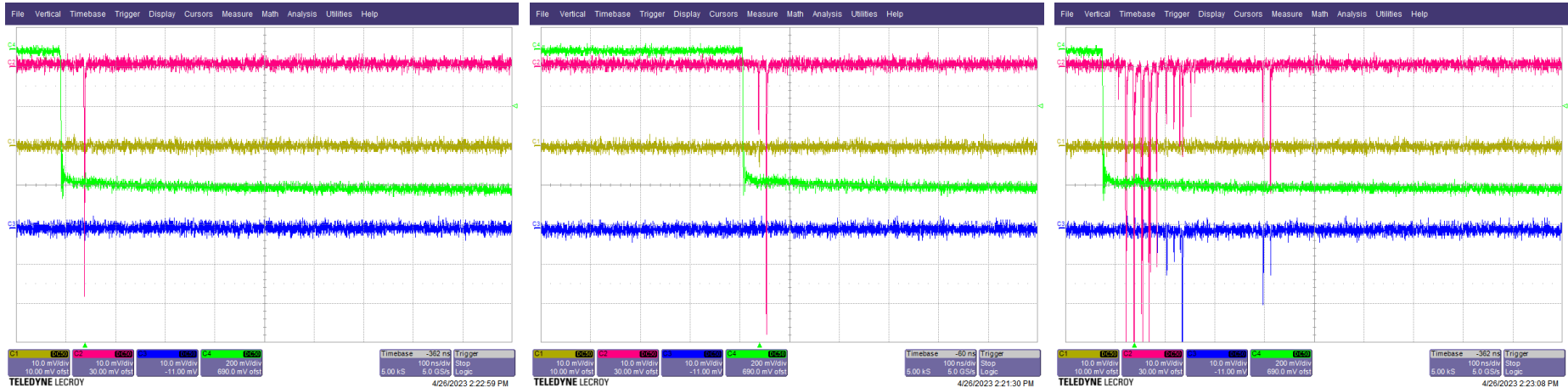
- afterpulse ratio = probability of creating a feedback ion which produces an additional delayed signal
- probably bad impact on the lifetime → should be low
- looks very different for different tubes (see TRB-measurements below)



Comparison of 9002224 and 9002227 afterpulsing

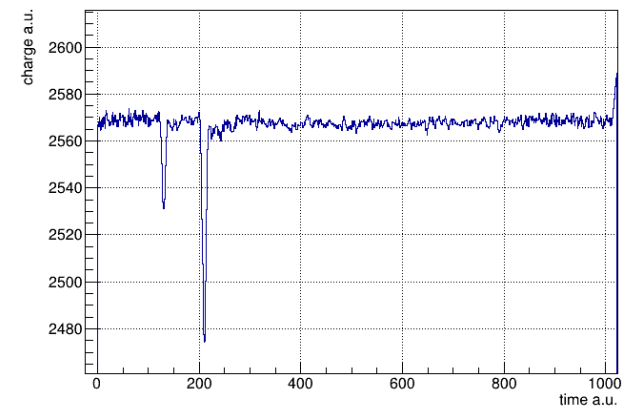
- further investigation on very high afterpulsing of 9002227

scope screenshots of a normal (a little high) signal, a normal + afterpulse and multiple afterpulses



- non-bipolar pulses → indicates real afterpulses and no (electronic) crosstalk
- measuring waveforms with CAEN DT5742-digitizer
- analysis of pulse arrival time and charge

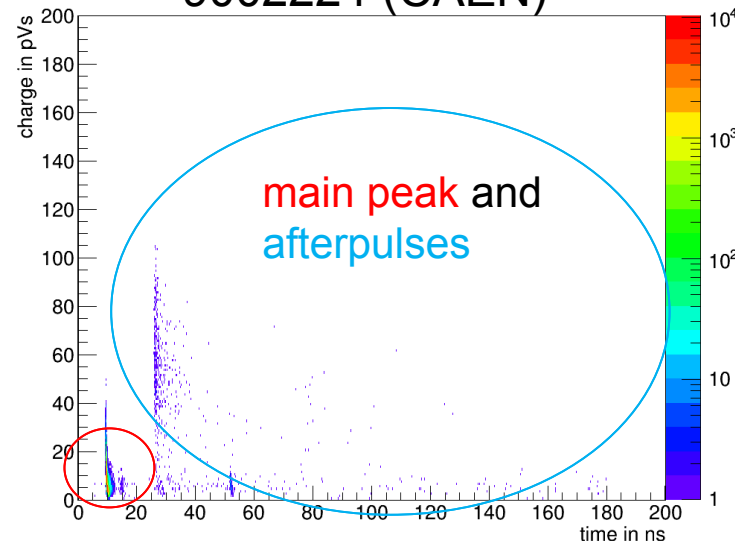
hData: pulse



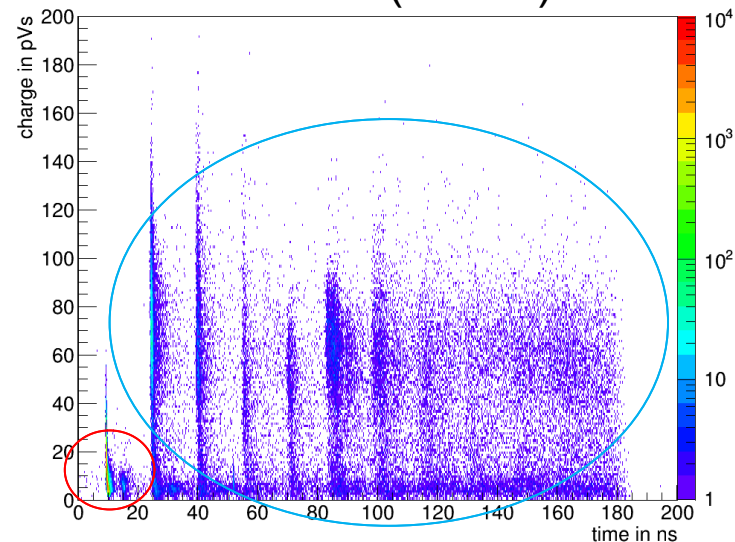
Comparison of 9002224 and 9002227 afterpulsing

- illumination at one single position, readout of one complete row (CAEN) or complete sensor (TRB)
- charge vs arrival time of illuminated pixel for $\sim 1e6$ gain and 0T for 9002224 and 9002227 (CAEN)
- a lot more entries for 9002227
- afterpulses: higher charge than main peak
- afterpulses have a chance of creating another feedback ion (and again and again) \rightarrow equidistant stripes (~ 15 ns in this case)

9002224 (CAEN)



9002227 (CAEN)



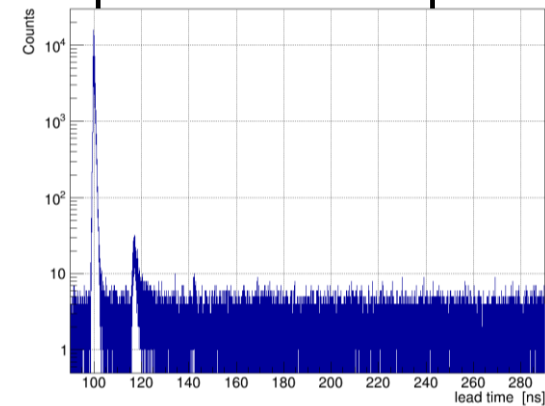
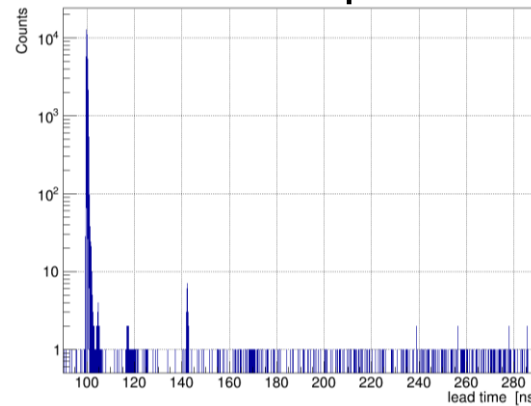
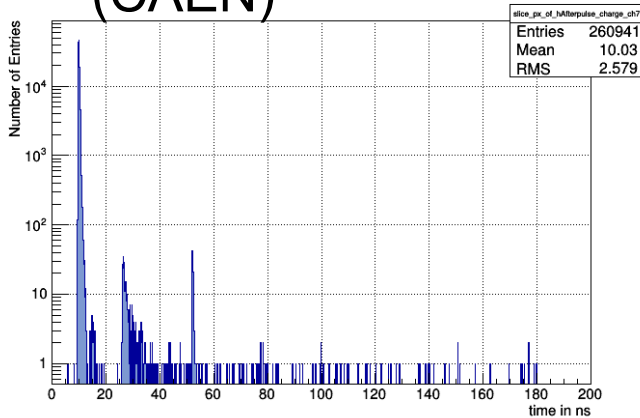
Comparison of 9002224 and 9002227 afterpulsing

projection to x-axis
(CAEN)

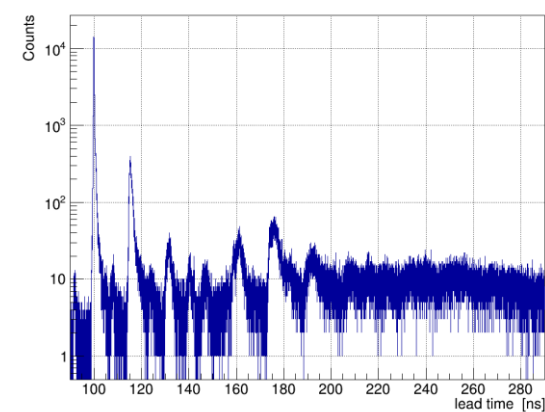
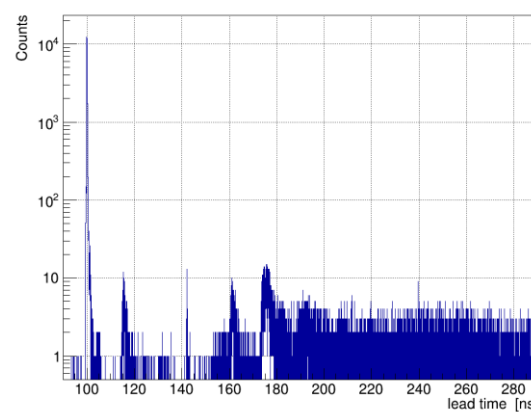
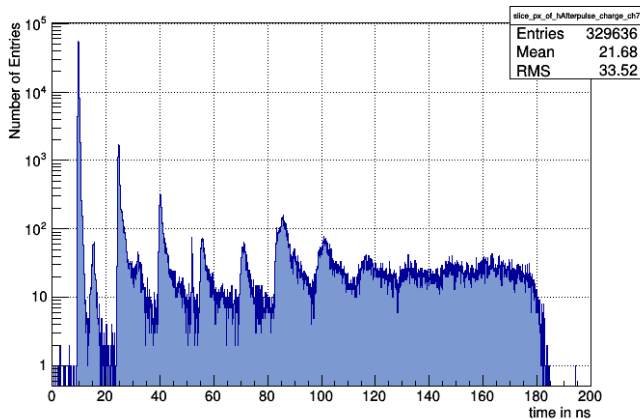
TRB afterpulse spectrum
of illuminated pixel

TRB afterpulse
spectrum of all pixel

900
2224



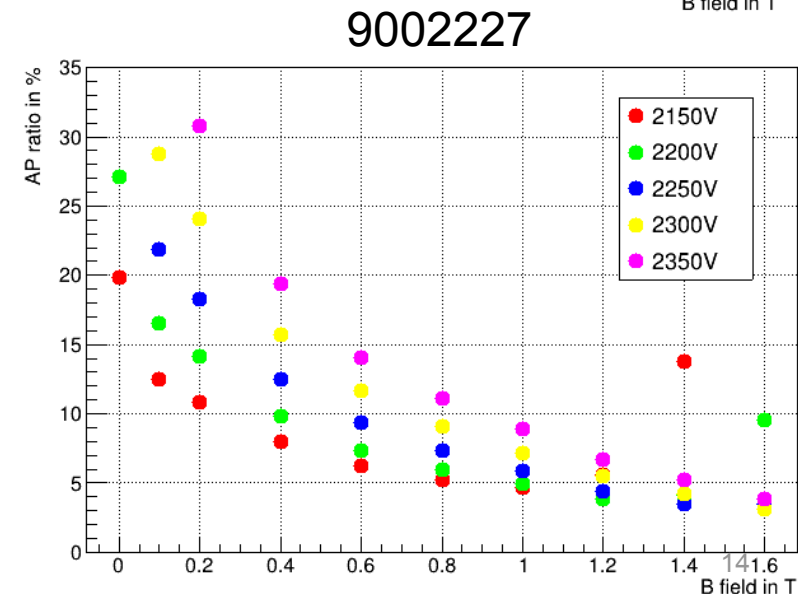
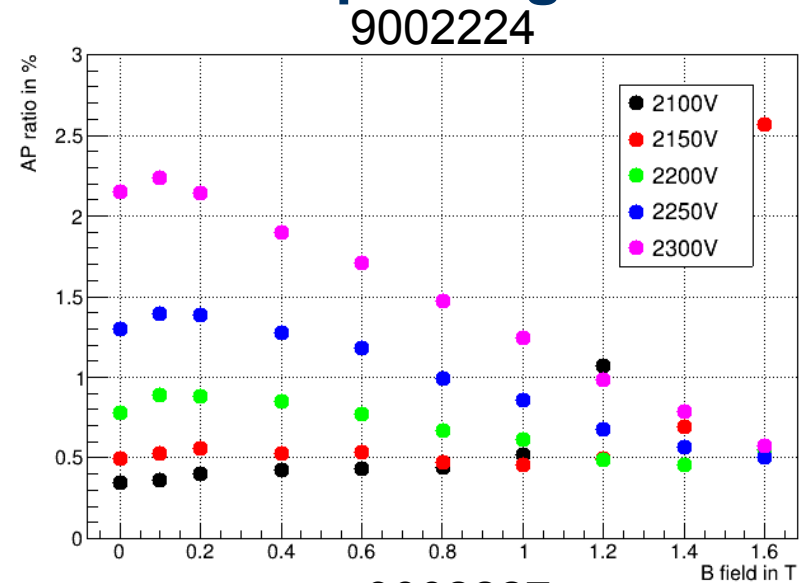
900
2227



- same setup, but: TRB has 30 ns dead time after hit → 1 pixel spectrum not useful, better: use of all pixel but possibility of counting one afterpulse ion multiple times

Comparison of 9002224 and 9002227 afterpulsing

- afterpulse ratio (CAEN) for different magnetic fields and voltages
- missing first points for 9002227 are due to escalation
- „rising“ afterpulsing for high magnetic fields is due to threshold effect (main peak is below threshold, while afterpulses are still above)
- higher afterpulse ratio for higher voltage and smaller B-field
 → afterpulse ratio depends mainly on gain (not voltage!)



Summary

- first try of measuring the time constant of MCPs with the double pulse method
- gain in B-field looks **as expected**
- shifts in x- and y-direction in magnetic fields (Lorentz and geometrical) are now **understood**
- high afterpulse ratio of 9002227 caused by real afterpulses, not (electronic) crosstalk or backplane effects etc.
- there is still a lot of data from Jülich to be analyzed
 - voltage and magnetic field dependent AP-ratio with the TRB-system
 - convert voltage and magnetic field information to a gain information to get the gain dependent AP-ratio
 - cuts on neighboring pixels to look for crosstalk waveforms with the CAEN system
 - ...

Thank you for your attention!

ERLANGEN CENTRE
FOR ASTROPARTICLE
PHYSICS



GEFORDERT VOM



Bundesministerium
für Bildung
und Forschung



ERLANGEN CENTRE
FOR ASTROPARTICLE
PHYSICS



FRIEDRICH-ALEXANDER
UNIVERSITÄT
ERLANGEN-NÜRNBERG

NATURWISSENSCHAFTLICHE
FAKULTÄT