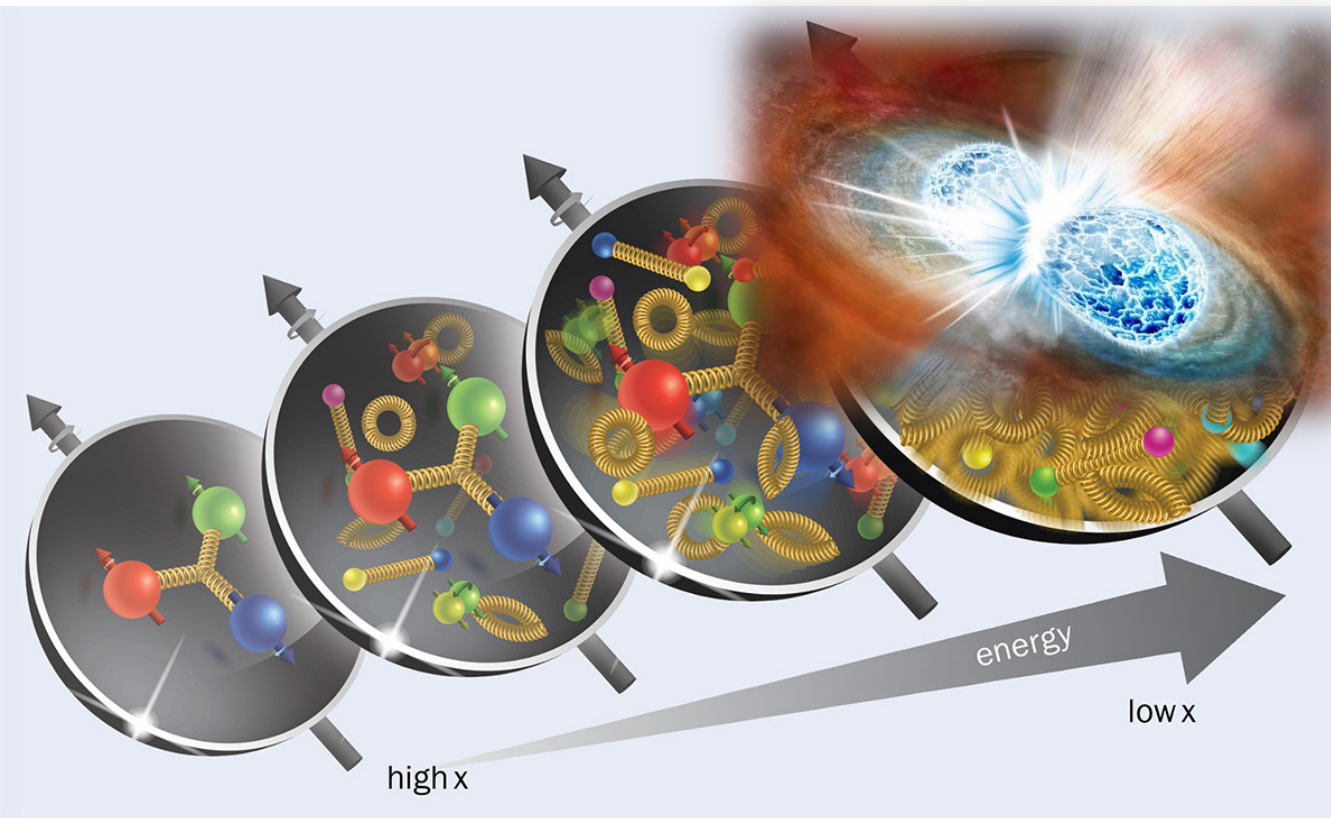


Experiment Overview of nuclear cluster production at RHIC

Zhangbu Xu

(Kent State University)

- $B=1$: simplest gluon topology in QCD
- $B=3,2$ nuclear yield ratio as probe of quantum wavefunction overlaps and density fluctuation
- $B=4,3$ Hypernuclear properties
- Discovery of $B=-4$ hypernucleus
- Future:
charmed hypernuclei and Hoyle states



EMMI Open Symposium:
Understanding light (anti-)nuclei production at RHIC and LHC



Where is Kent State University?

Path of Totality



Three Minutes of Twilight

On April 8, 2024, Kent, Ohio will be in the path of totality for a solar eclipse where the moon will completely blackout the sun for approximately three minutes at 3:14 p.m. There will not be another total solar eclipse in the U.S. until August 2044; and in Northeast Ohio until September 2099.

Partial Eclipse: 1:59 p.m. to 4:29 p.m.

Total Eclipse: 3:14 p.m. to 3:17 p.m.



Tenskwatawa's prediction **June 16, 1806**

It has been called Tecumseh's Eclipse after the [Shawnee](#) chief, [Tecumseh](#). He realized that the only hope for the various tribes in east and central [North America](#) was to join. He was assisted by his brother, [Tenskwatawa](#), called The Prophet, who called for a rejection of European influence and a return to traditional values. This tribal unity threatened [William Henry Harrison](#), the Territorial Governor of Indiana and future 9th President of the United States. Harrison tried to discredit the Shawnee leader by challenging Tenskwatawa to prove his powers. He wrote: "If he (Tenskwatawa) is really a prophet, ask him to cause the Sun to stand still or the Moon to alter its course, the rivers to cease to flow or the dead to rise from their graves."

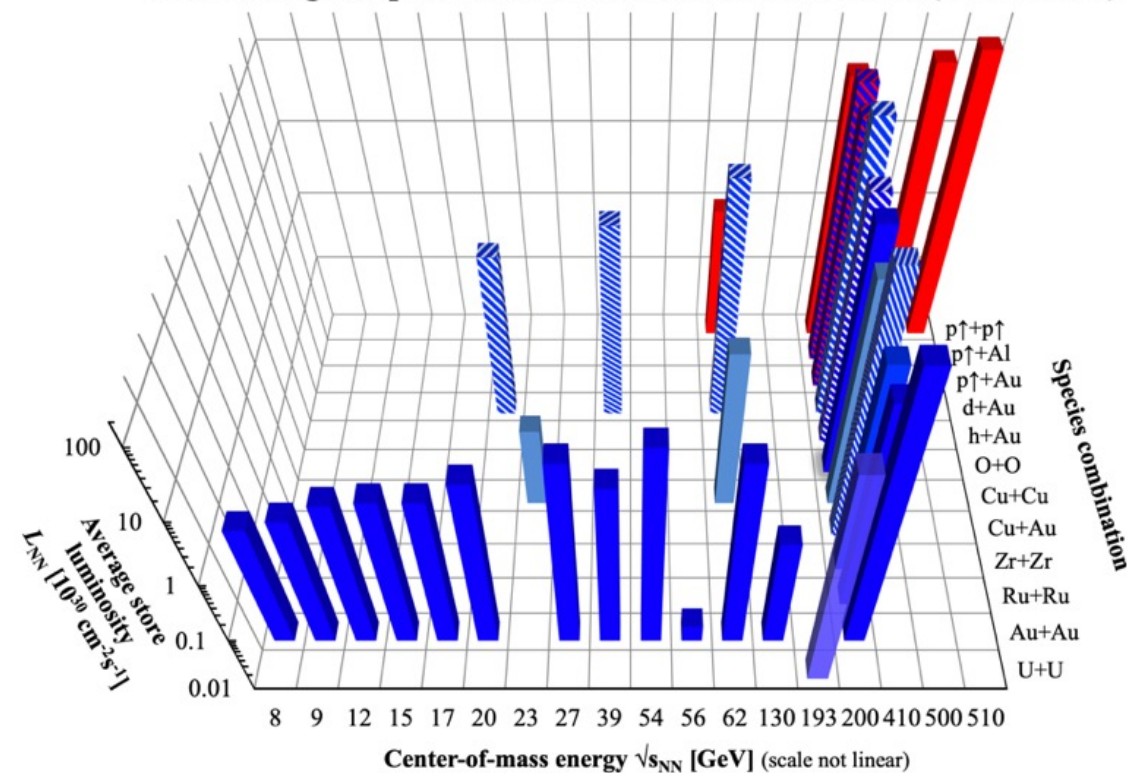
Tenskwatawa declared that the [Great Spirit](#) was angry at Harrison and would give a sign. "Fifty days from this day there will be no cloud in the sky. Yet, when the Sun has reached its highest point, at that moment will the Great Spirit take it into her hand and hide it from us. The darkness of night will thereupon cover us and the stars will shine round about us. The birds will roost and the night creatures will awaken and stir." On that day, there was an eclipse, and Harrison's attempt to divide the Shawnee people backfired spectacularly. Then Tecumseh ordered the Great Spirit to release the sun.^[1]

Example of versatile colliders and detectors

STAR major upgrades over the last twenty years to improve particle identification and vertex reconstruction and is still evolving with an extension to forward rapidity as of today. **pioneered in using new technologies**: MRPC, MAPS, GEM and siPM. Estimate 35M(initial) +75M(upgrades)\$.

PHENIX→sPHENIX

RHIC energies, species combinations and luminosities (Run-1 to 22)

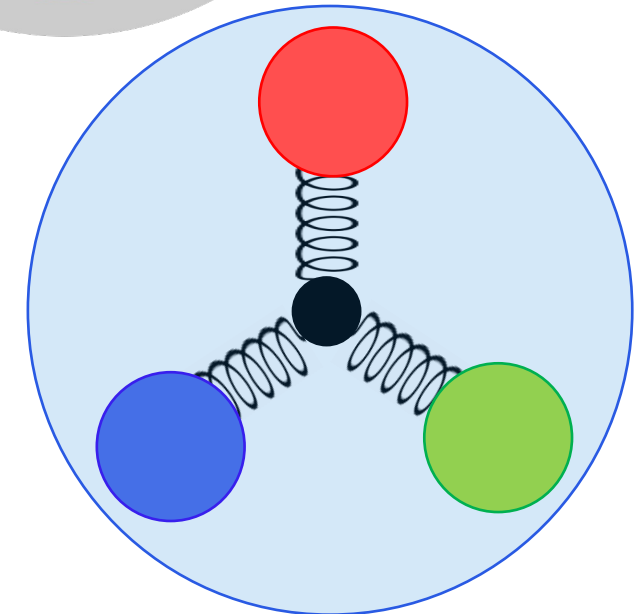
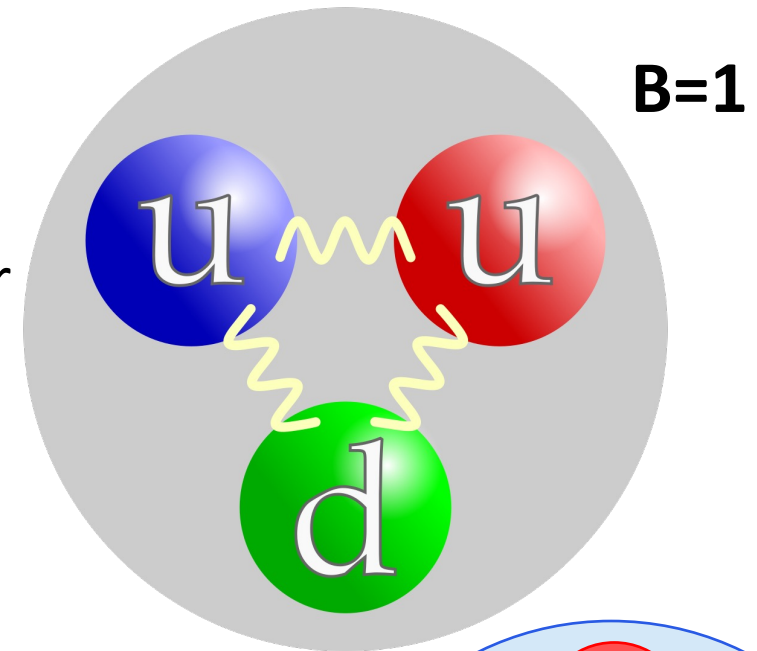


Detector	primary functions	DOE+(in-kind)	year
TPC+Trigger	$ \eta < 1$ Tracking		1999-
Barrel EMC	$ \eta < 1$ jets/ $\gamma/\pi^0/e$		2004-
FTPC	forward tracking	(Germany)	2002-2012
L3	Online Display	(Germany)	2000-2012
SVT/SSD	V0/charm	(France)	2004-2007
PMD	forward photons	(India)	2003-2011
EEMC	$1 < \eta < 2$ jets/ π^0/e	(NSF)	2005-
Roman Pots	diffractive		2009-
TOF	PID	(China)	2009-
FMS/Preshower	$2.5 < \eta < 4.2$	(Russia)	2008-2017
DAQ1000	x10 DAQ rate		2008-
HLT	Online Tracking	(China/Germany)	2012-
FGT	$1 < \eta < 2$ W^\pm		2012-2013
GMT	TPC calibration		2012-
HFT/SSD	open charm	(France/UIC)	2014-2016
MTD	muon ID	(China/India)	2014-
EPD	event plane	(China)	2018-
RHICf	$\eta > 5$ π^0	(Japan)	2017
iTPC	$ \eta < 1.5$ Tracking	(China)	2019-
eTOF	$-2 < \eta < -1$ PID	(Germany/China)	2019-
FCS	$2.5 < \eta < 4$ calorimeter	(NSF)	2021-
FTS	$2.5 < \eta < 4$ Tracking	(NCKU/SDU)	2021-

8 new detectors added to STAR since 2014

Baryon Number (B) Carrier

- Textbook picture of a proton
 - Lightest baryon with strictly conserved baryon number
 - Each valence quark carries 1/3 of baryon number
 - Proton lifetime $>10^{34}$ years
 - Quarks are connected by gluons
- Alternative picture of a proton
 - Proposed at the Dawn of QCD in 1970s
 - A Y-shaped gluon junction topology carries baryon number (B=1)
 - The topology number is the strictly conserved number
 - Quarks do not carry baryon number
 - Valence quarks are connected to the end of the junction always

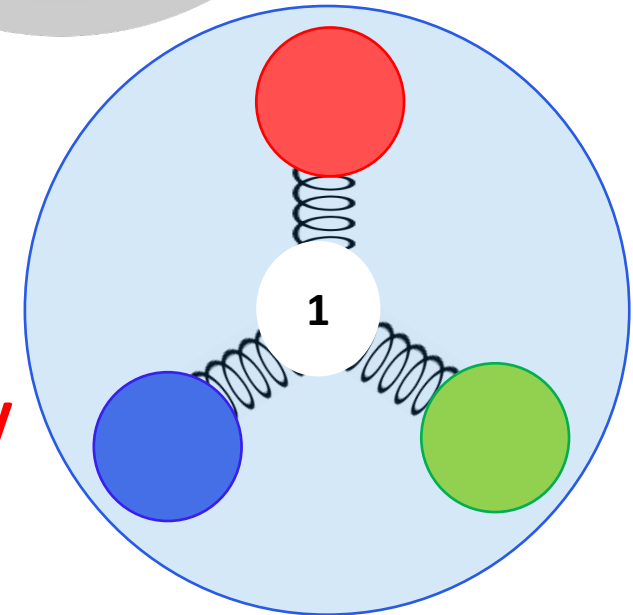
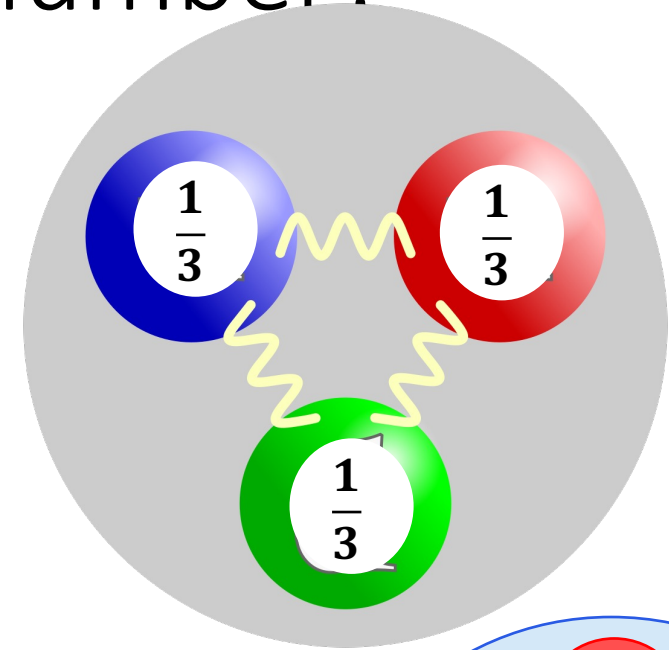


[1]: Artru, X.; String Model with Baryons: Topology, Classical Motion. Nucl. Phys. B 85, 442–460 (1975).

[2]: Rossi, G. C. & Veneziano, G. A; Possible Description of Baryon Dynamics in Dual and Gauge Theories. Nucl. Phys. B 123, 507–545 (1977)

Measurements of quark baryon number?

- Textbook picture of a proton
 - Lightest baryon with strictly conserved baryon number
 - Each valence quark carries $\frac{1}{3}$ of baryon number
 - Proton lifetime $>10^{34}$ years
 - Quarks are connected by gluons
- Alternative picture of a proton
 - Proposed at the Dawn of QCD in 1970s
 - A Y-shaped gluon junction topology carries baryon number ($B=1$)
 - The topology number is the strictly conserved number
 - Quarks do not carry baryon number
 - Valence quarks are connected to the end of the junction always
- **Neither of these postulations has been verified experimentally**



[1]: Artru, X.; String Model with Baryons: Topology, Classical Motion. Nucl. Phys. B 85, 442–460 (1975).

[2]: Rossi, G. C. & Veneziano, G. A; Possible Description of Baryon Dynamics in Dual and Gauge Theories. Nucl. Phys. B 123, 507–545 (1977)

Model implementations of baryons at RHIC

- Many of the models used for heavy-ion collisions at RHIC (HIJING, AMPT, UrQMD) have implemented a nonperturbative baryon stopping mechanism

V. Topor Pop, *et al*, Phys. Rev. C **70**, 064906 (2004)

Zi-Wei Lin, *et al*, Phys. Rev. C **72**, 064901 (2005)

M. Bleicher, *et al*, J.Phys.G **25**, 1859-1896 (1999)

- **Baryon Stopping**

- Theorized to be an effective mechanism of stopping baryons in pp and AA

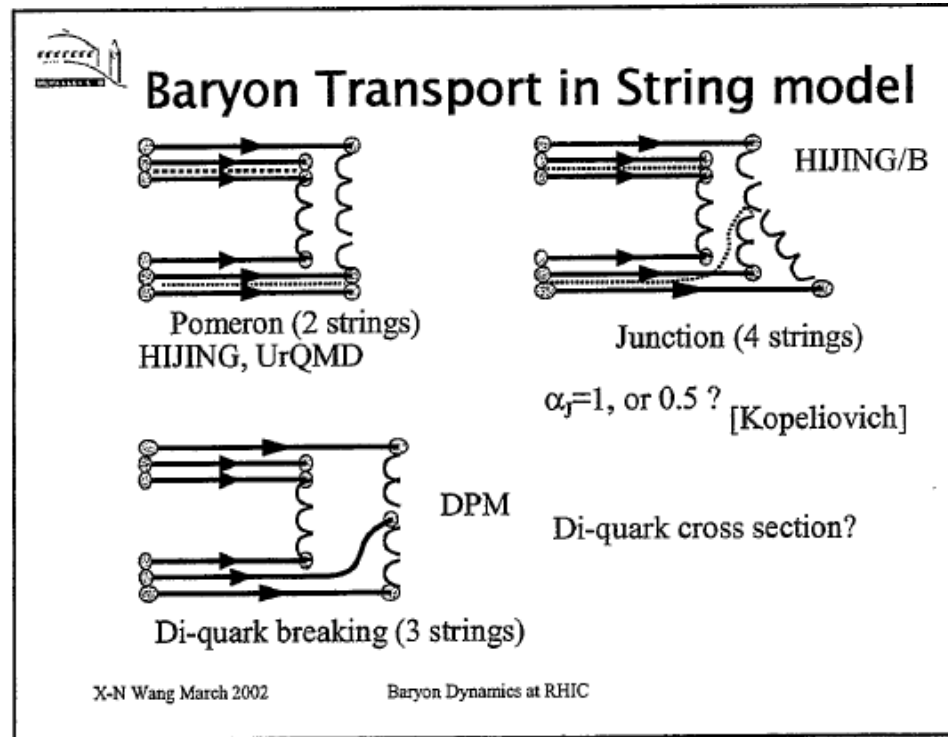
D. Kharzeev, Physics Letters B **378**, 238-246 (1996)

- Specific rapidity dependence is predicted:

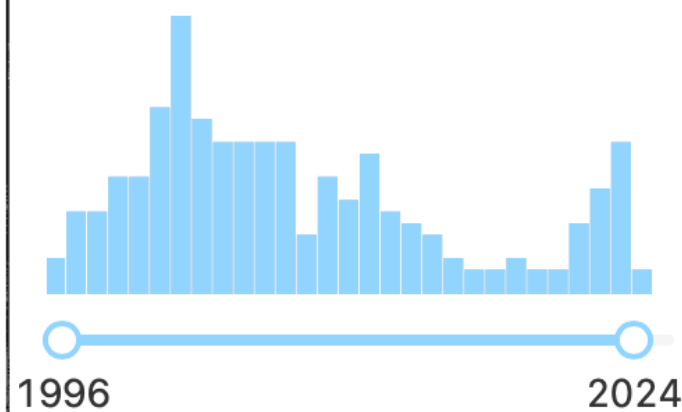
$$p = \sim e^{-\alpha_B y}$$

$$\alpha_B \sim 0.5$$

2003 RBRC Workshop on “Baryon Dynamics at RHIC”



D. Kharzeev, Physics Letters B **378**, 238-246 (1996)
“Can gluons trace baryon number?”



“Science, however, is never conducted as a popularity contest...” --- Michio Kaku

BUT citations ARE

Three approaches toward tracking the origin of the baryon number

1. STAR Method:

Charge (Q) stopping vs baryon (B) stopping:
if valence quarks carry Q and B,
 $Q=B$ at middle rapidity

2. Kharzeev-STAR Method:

If gluon topology (J) carries B as one unit, it should show scaling according to Regge theory

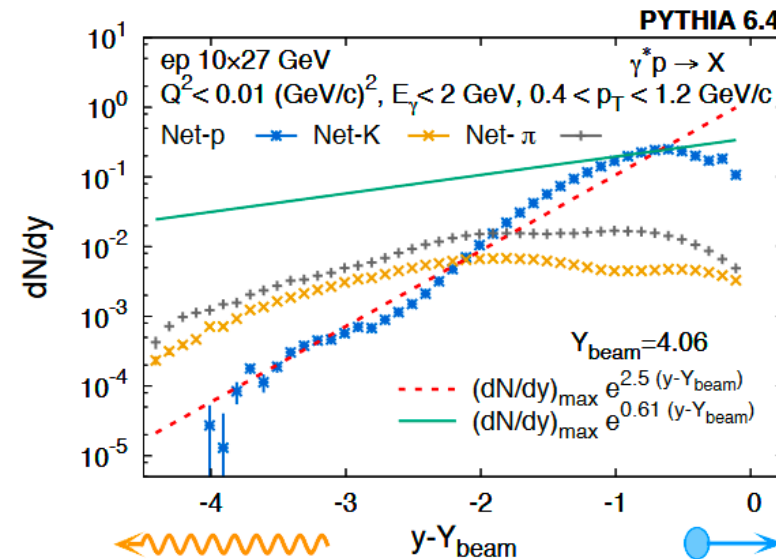
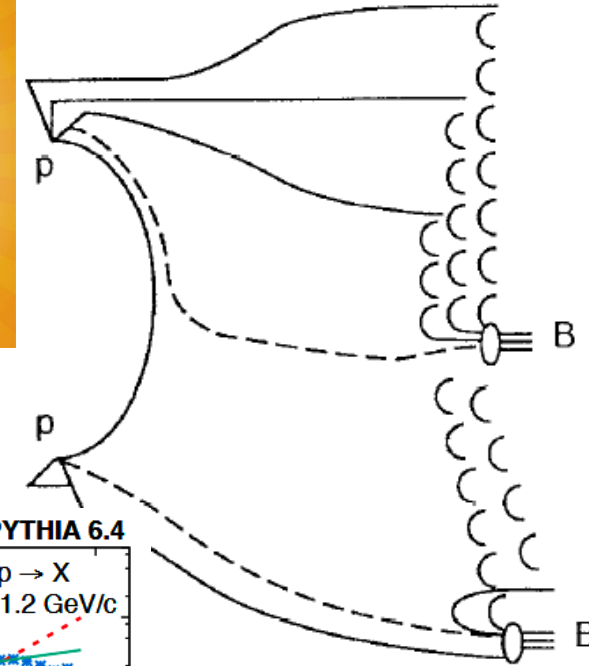
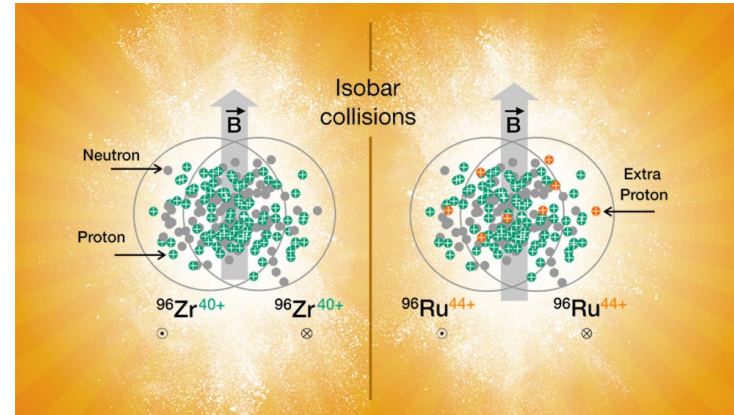
$$p = \sim e^{-\alpha_B Y}$$

$$\alpha_B \sim 0.5$$

3. Artru Method:

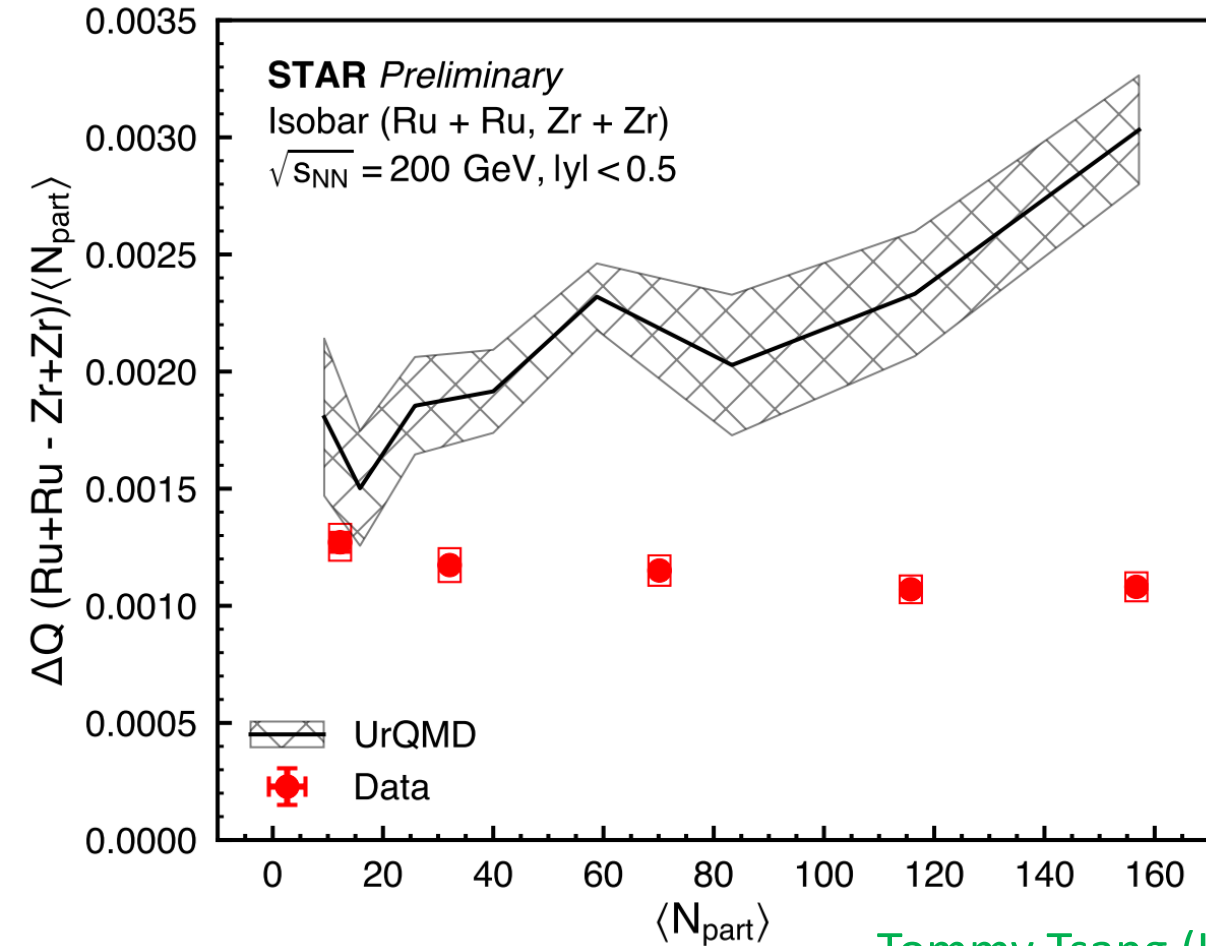
In γ +Au collision, rapidity asymmetry can reveal the origin

arXiv:2205.05685

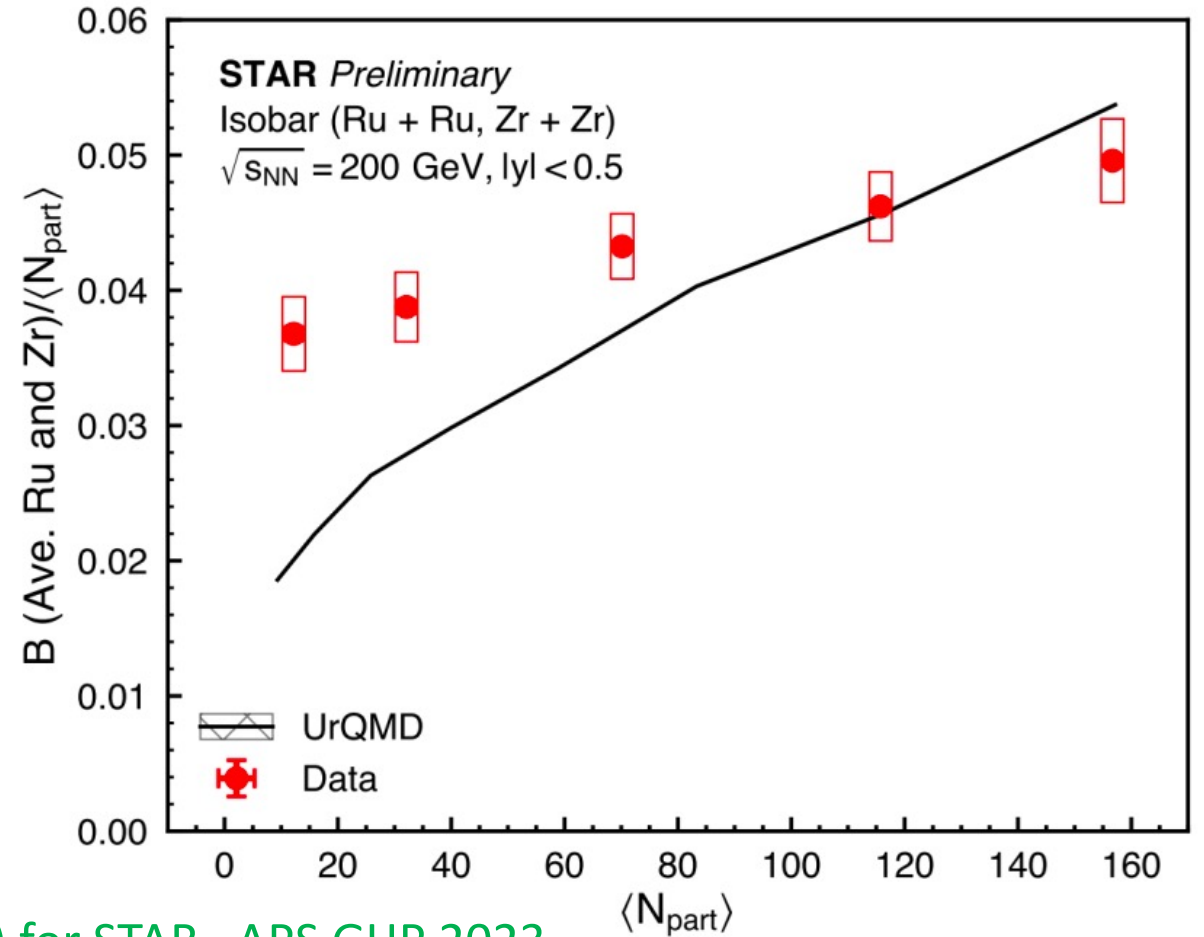


Separate charge and baryon transports

Charge number transport



Baryon number transport



Tommy Tsang (KSU) for STAR, APS GHP 2023

UrQMD matches data on charge stopping better in peripheral; better on baryon stopping in central
overpredicts charge stopping in central; underpredicts baryon stopping in peripheral

Three approaches toward tracking the origin of the baryon number

1. STAR Method:

Charge (Q) stopping vs baryon (B) stopping:

if valence quarks carry Q and B,
Q=B at middle rapidity

$$B/Q=2$$

2. Kharzeev-STAR Method:

If gluon topology (J) carries B as one unit,
it should show scaling according to

Regge theory

$$\alpha_B=0.61$$

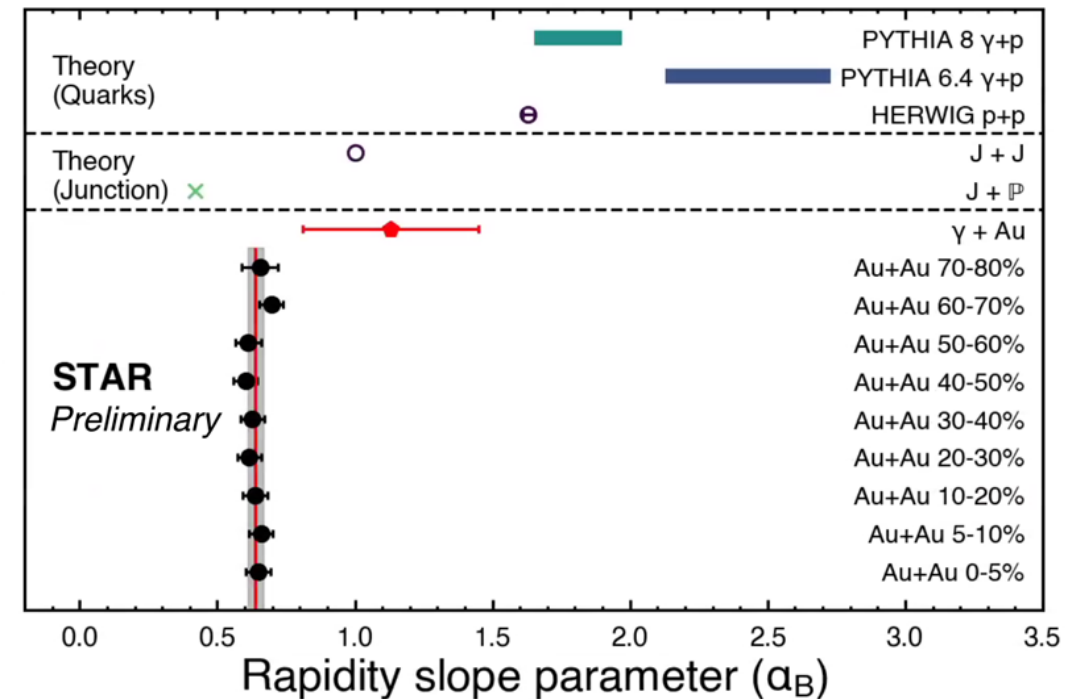
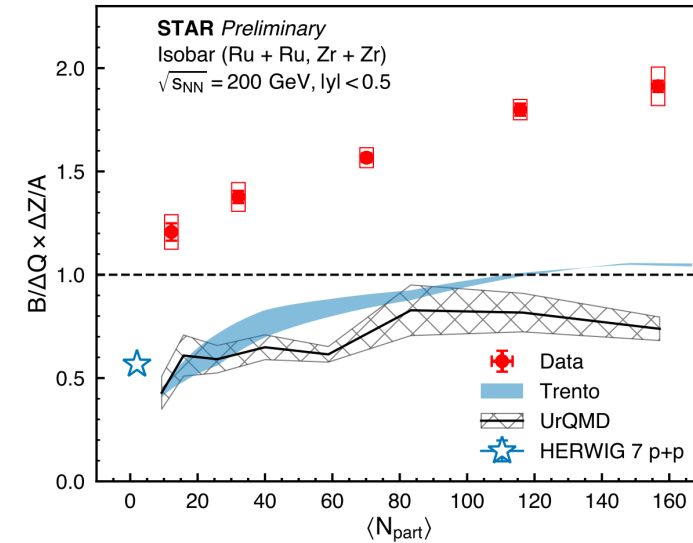
$$p = \sim e^{-\alpha_B y}$$

$$\alpha_B \sim 0.5$$

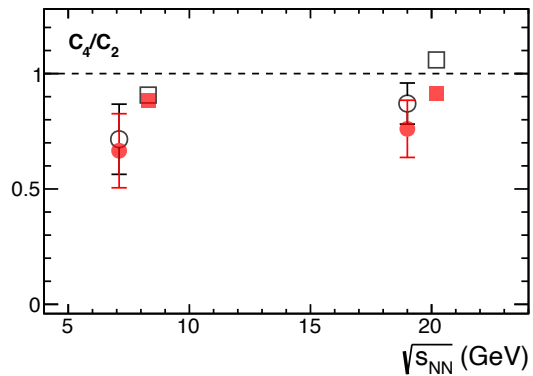
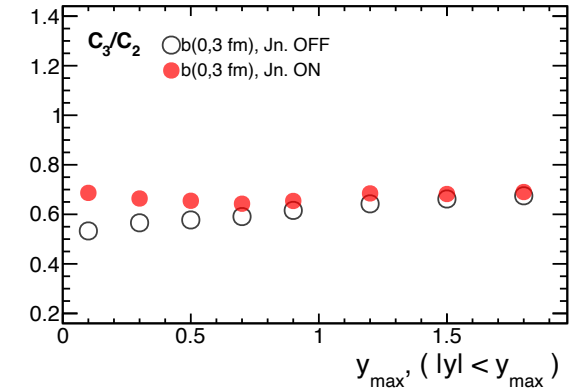
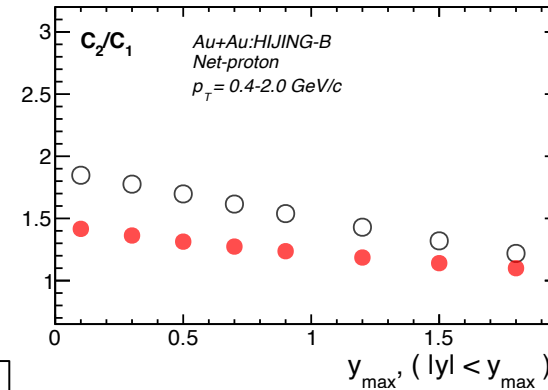
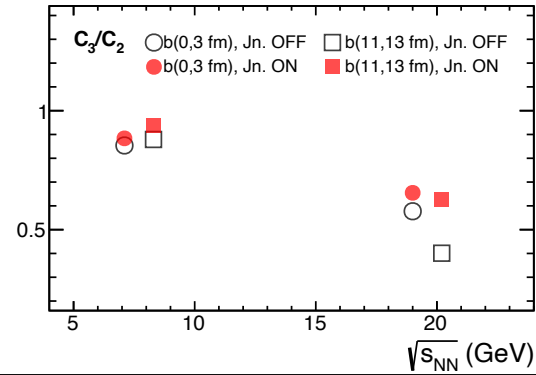
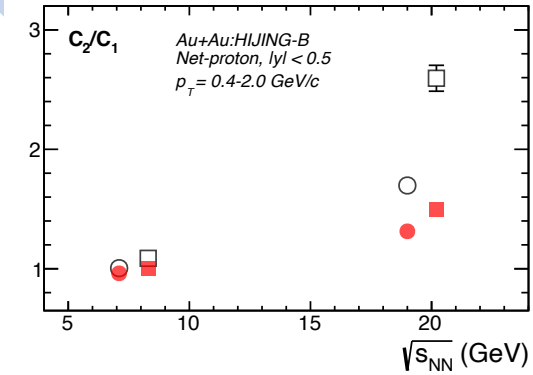
3. Artru Method:

In γ +Au collision, rapidity asymmetry can
reveal the origin

$$\alpha_B(A+A)=0.61 < \alpha_B(\gamma+A)=1.1 < \alpha_B(\text{PYTHIA})$$



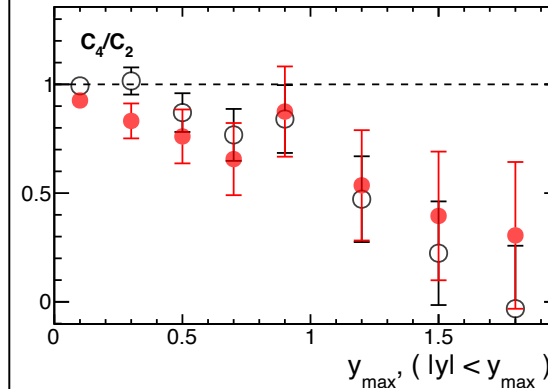
CUMULANTS OF NET-PROTON DISTRIBUTIONS W/W.O BARYON JUNCTION



Baryon Junction:

- Effect at higher collision energy.
- 2nd order normalized fluctuations are suppressed.
- Distribution gets more skewed.

Bedanga Mohanty @
[CFNS Baryon Dynamics](#)



HIJING simulation
w/o junction

19.6 GeV

Baryon Junction:

Normalized 2nd order fluctuations more suppressed towards mid-rapidity

B=1,2,3 nuclear yield ratios

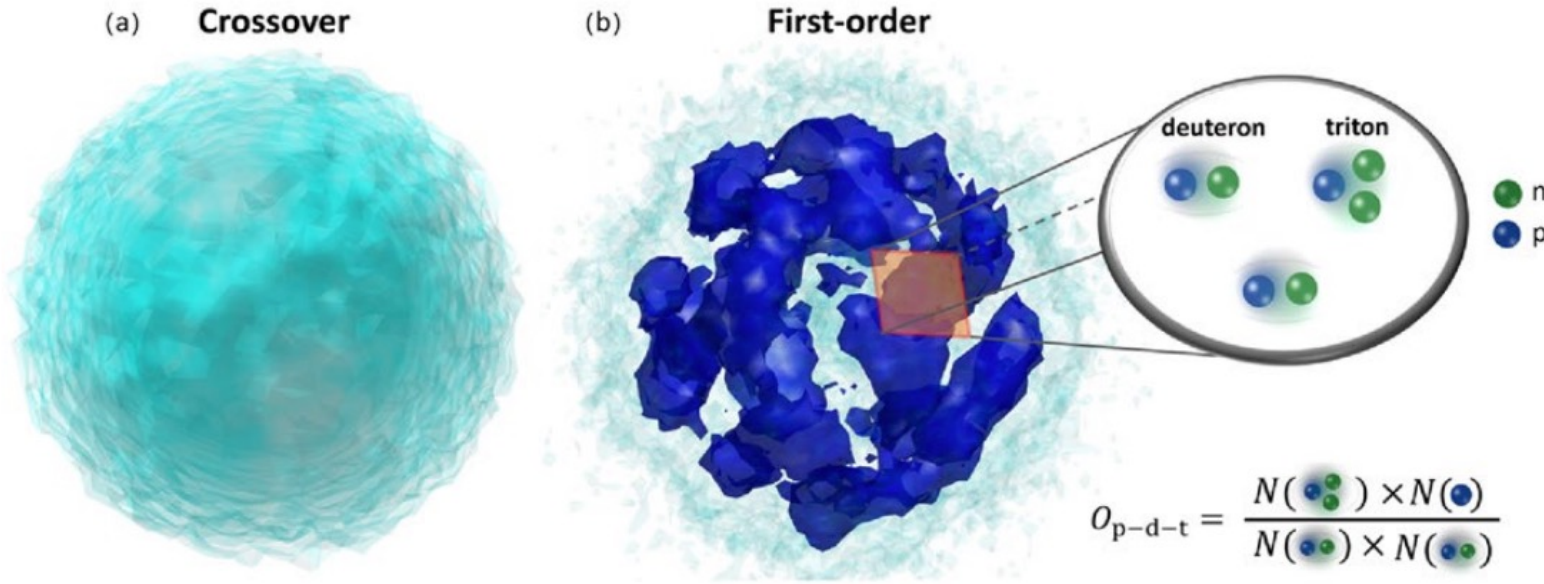


Fig. 1 (Color online) Density distribution of strongly interacting matter in a heavy ion collision after its expansion for the cases of crossover transition (panel **a**) and first-order chiral phase transition (panel **b**). Also shown for illustration of the latter case are deuterons and tritons produced from the density fluctuating hadronic matter and their yield ratio $\mathcal{O}_{p-d-t} = N_t N_p / N_d^2$, which depends on the magnitude of neutron density distribution as discussed in the text

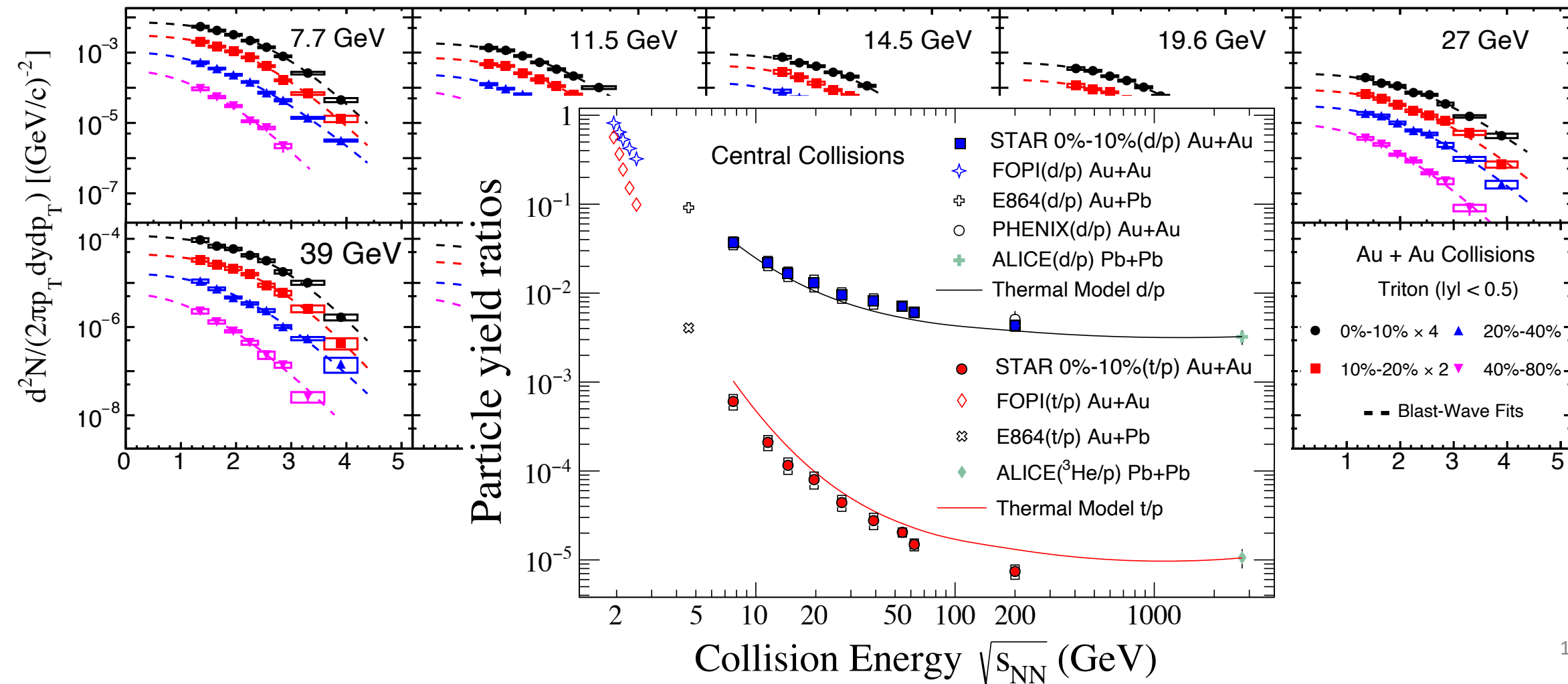
- Light nuclei production as a probe of the QCD phase diagram
K.J. Sun, et al., PLB 781 (2018) 499
- Probing QCD critical fluctuations from light nuclei production in relativistic heavy-ion collisions
K.J. Sun, et al., PLB 774 (2017) 103

C.M. Ko, NST 34 (2023) 80

$$\mathcal{O}_{p-d-t} = \frac{N_{3H} N_p}{N_d^2} = g \frac{1 + (1 + 2\alpha)\Delta n}{(1 + \alpha\Delta n)^2},$$

Spectra and two-particle ratios

STAR, Phys.Rev.Lett. 130 (2023) 202301

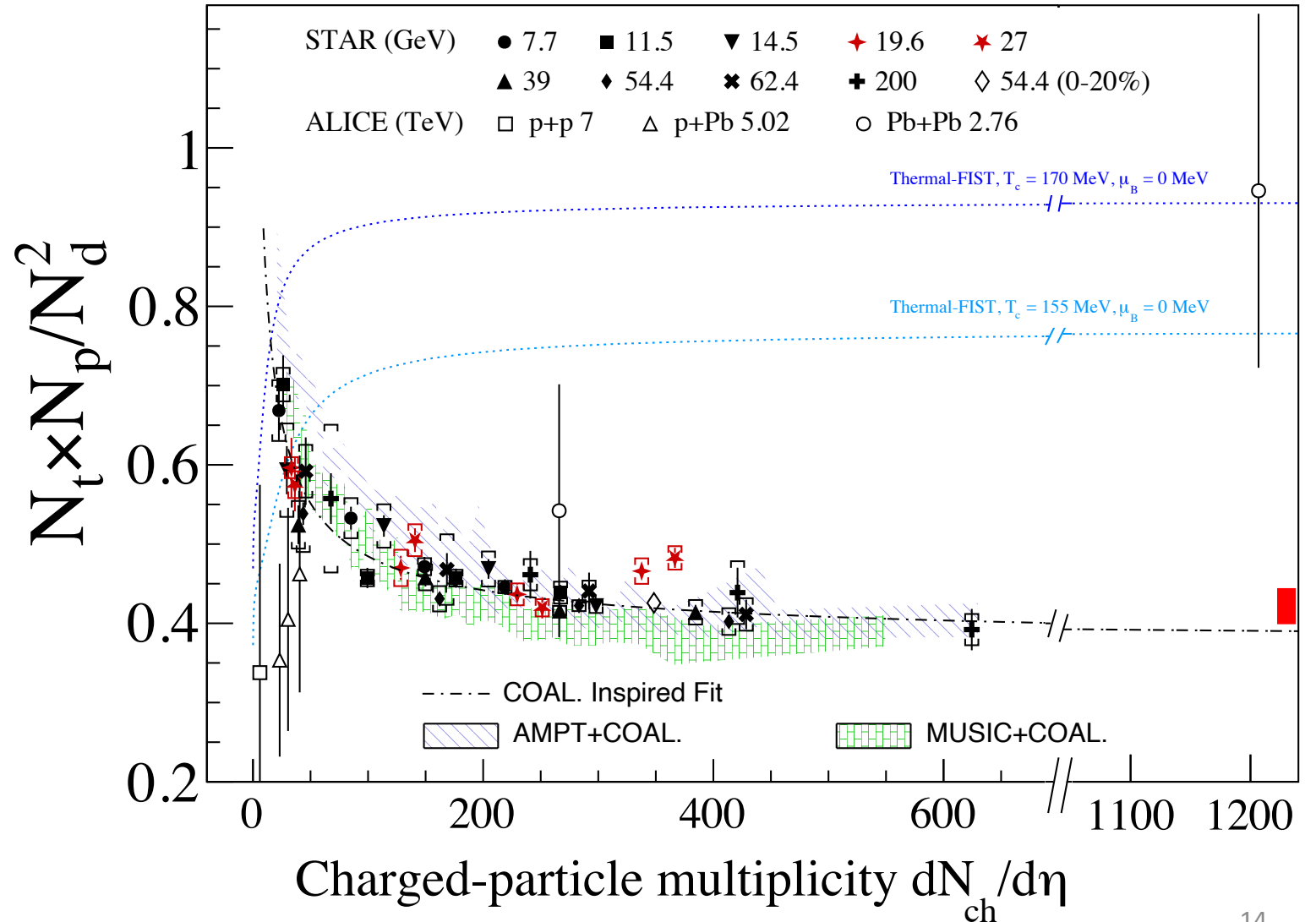


Quantum Wavefunction overlap efficiency

STAR, Phys.Rev.Lett. 130 (2023) 202301

$$\frac{N_t \times N_p}{N_d^2} = p_0 \times \left(\frac{R^2 + \frac{2}{3}r_d^2}{R^2 + \frac{1}{2}r_t^2} \right)^3$$

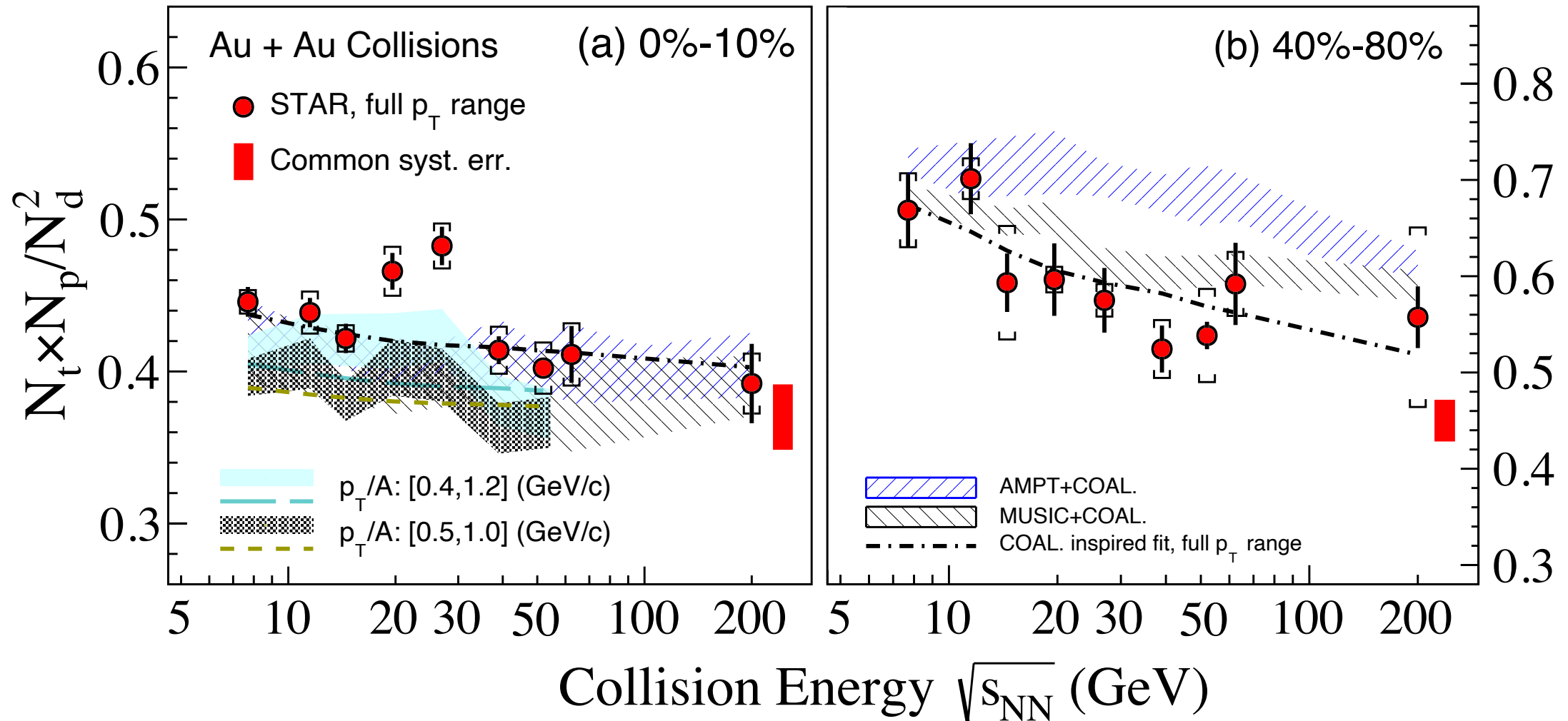
Coalescence wavefunction overlap between nucleus and nucleons



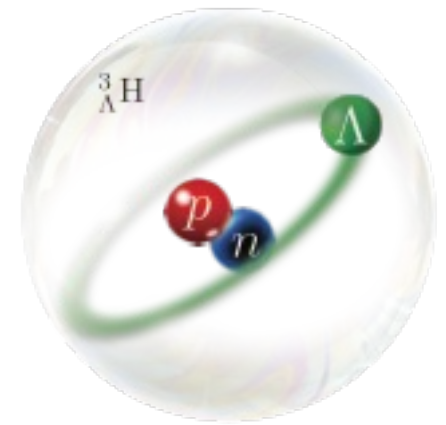
Possible sign of Density Fluctuation

4 σ effect, BES-II data x10 statistics

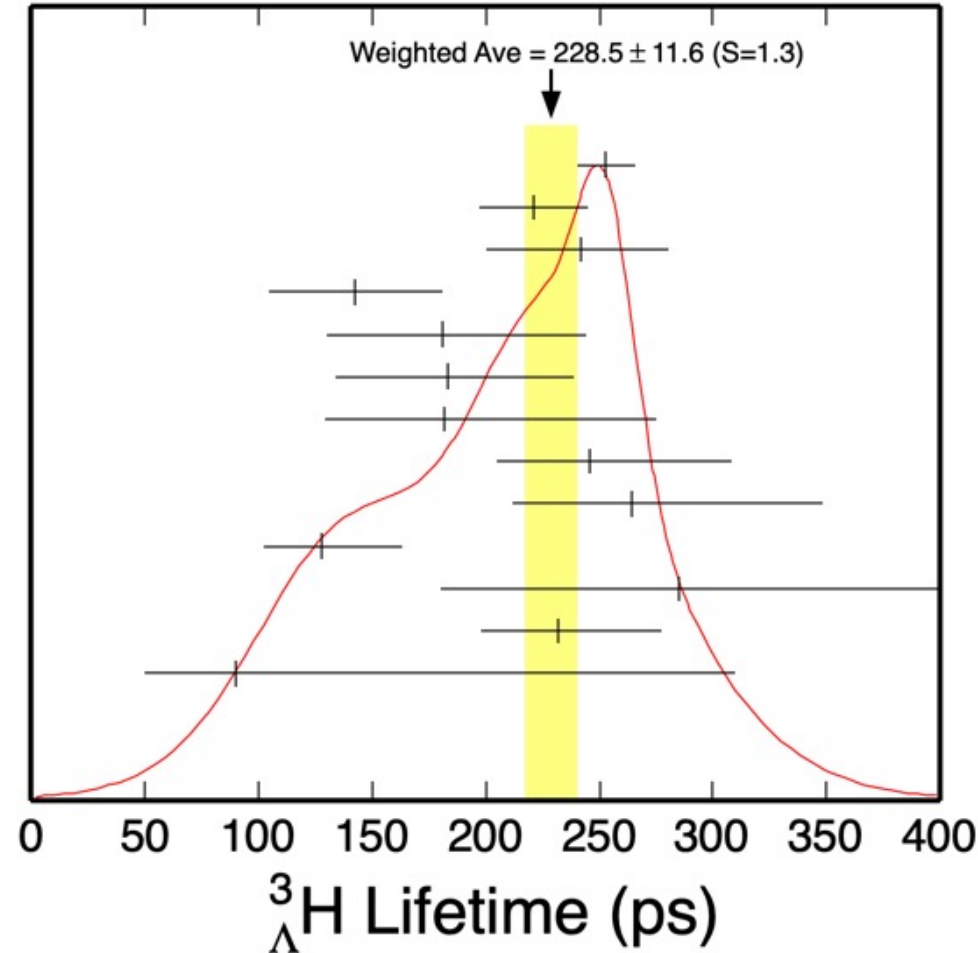
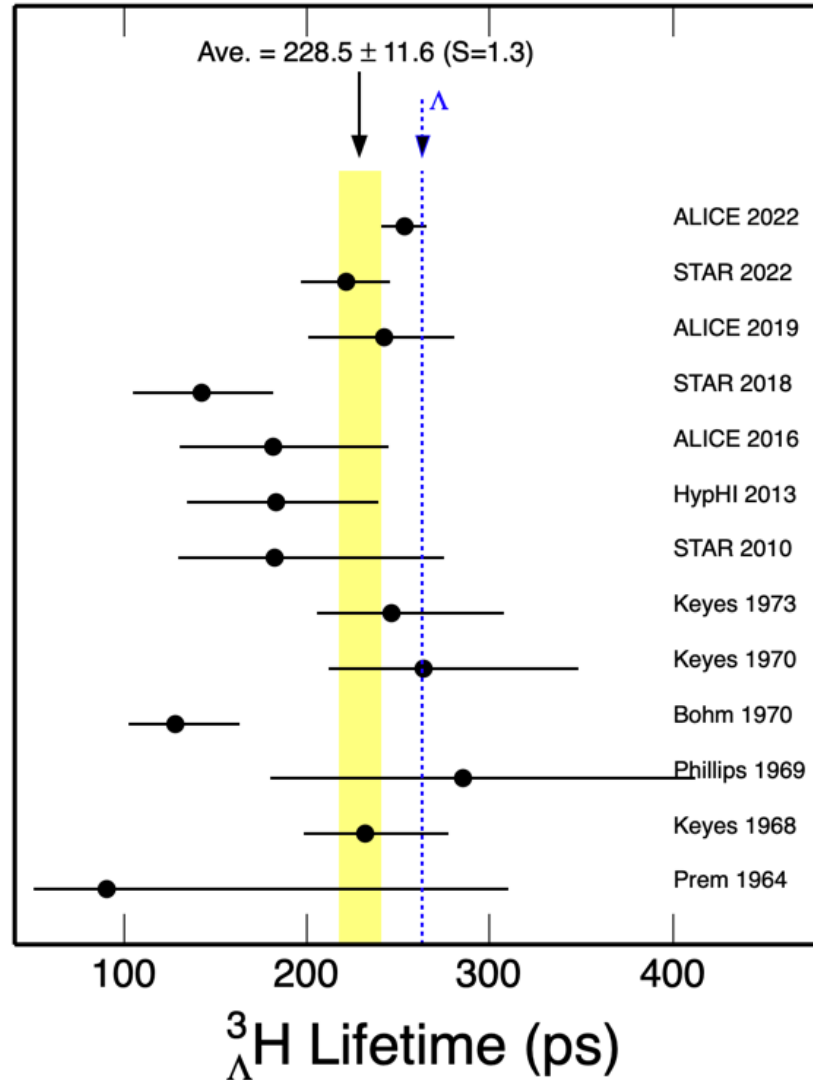
STAR, *Phys.Rev.Lett.* 130 (2023) 202301



$|B|=3$ hypertriton lifetime



arXiv:2311.09877



	χ^2
ALICE 2022	3.8
STAR 2022	0.1
ALICE 2019	0.1
STAR 2018	4.9
ALICE 2016	0.6
HypHI 2013	0.7
STAR 2010	0.3
Keyes 1973	0.1
Keyes 1970	0.4
Bohm 1970	8.2
Phillips 1969	
Keyes 1968	0.0
Prem 1964	
<hr/>	
	19.3
Confidence Level = 0.036	

Potential discrepancy?

arXiv:2311.09877

Simultaneous fit to all heavy-ion data

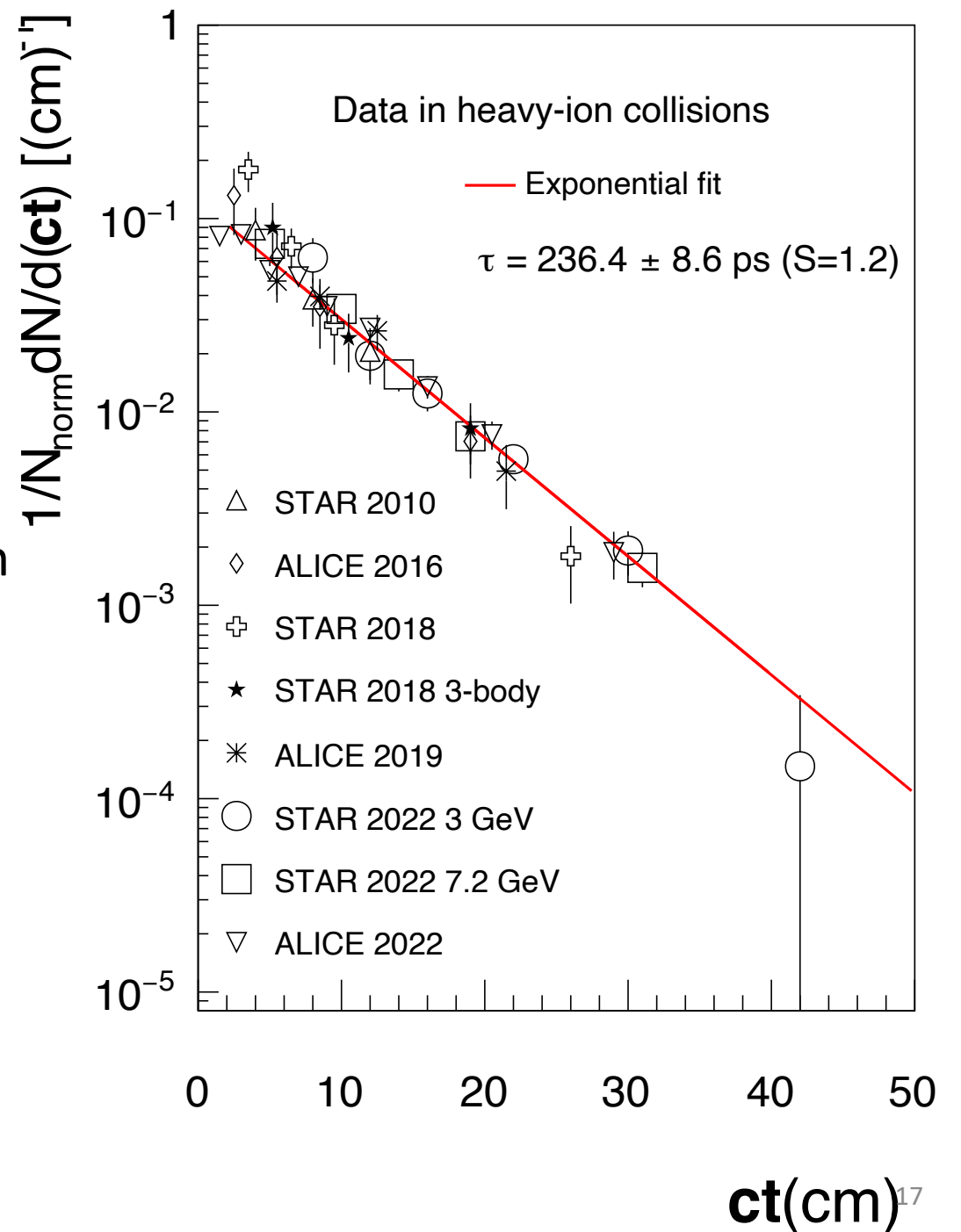
Scale yields to one common exponential function

Result consistent with other (average) methods

About 3σ smaller than Lambda lifetime

STAR 2018 first $c\tau$ point appears high

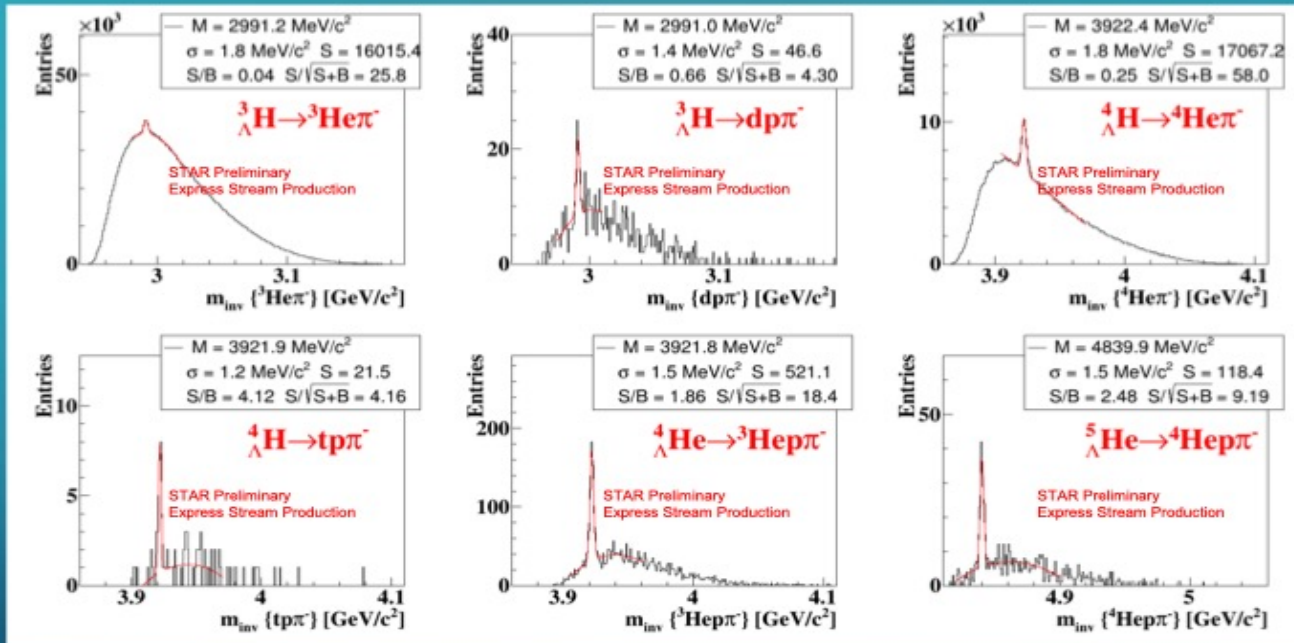
ALICE 2022 first $c\tau$ point appears low



A zoo of hypernucleus measurements

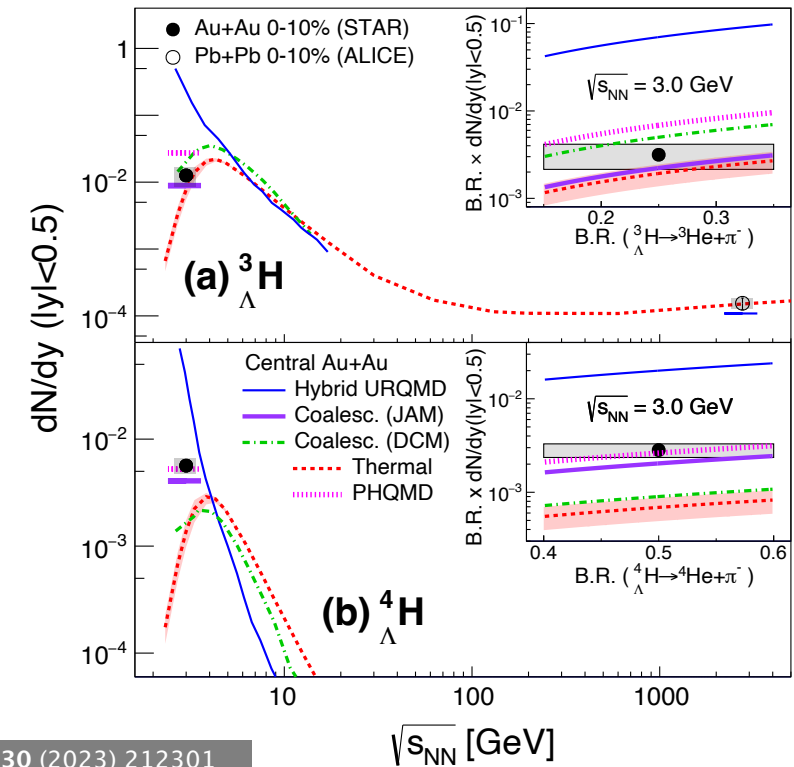
“Fluctuation and correlation measurements”, Hanna Zbroszczyk

437M HLT TRIGGERED EVENTS AT 3 GEV

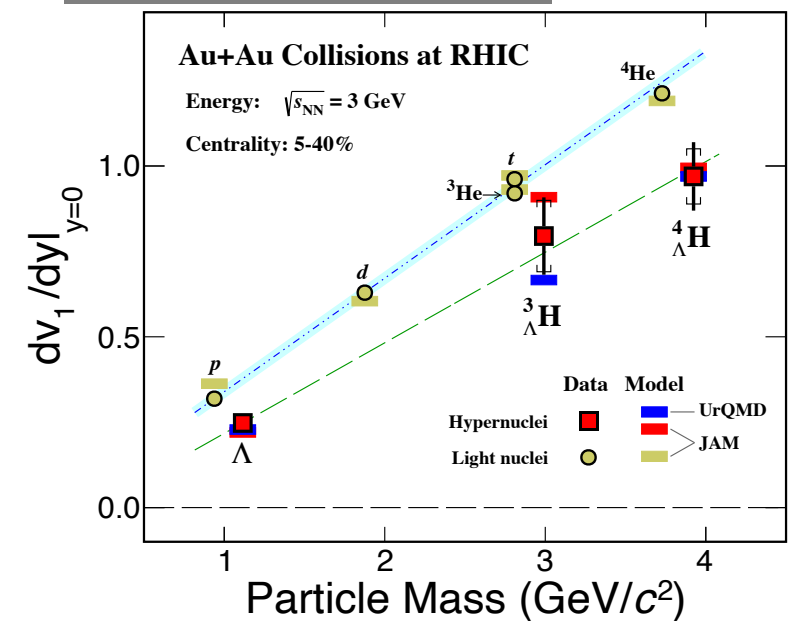


- With increased beam collision intensity in the Fixed Target mode HLT farm had not enough capacities to process all collected data online.
- Therefore a trigger on He has been introduced to enhance hypernuclei.

The collected statistics is enough to measure yields, lifetimes and spectra of these hypernuclei



Phys. Rev. Lett. 130 (2023) 212301



Search for heavy antimatter and baryon objects

Rapid Communication

Strangelet search

B. I. Abelev *et al.* (STAR Collaboration)
Phys. Rev. C **76**, 011901(R) – Published online 24 April 2011

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Published: 24 April 2011

Observation of the antimatter helium-4 nucleus

[The STAR Collaboration](#)

Nature **473**, 353–356 (2011) | [Cite this article](#)

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Letter | Published: 09 March 2020

Measurement of the mass difference and the binding energy of the hypertriton and antihypertriton

[The STAR Collaboration](#)

Nature Physics **16**, 409–412 (2020) | [Cite this article](#)

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HOME > SCIENCE > VOL. 328, NO. 5974 > OBSERVATION OF AN ANTIMATTER HYPERNUCLEUS

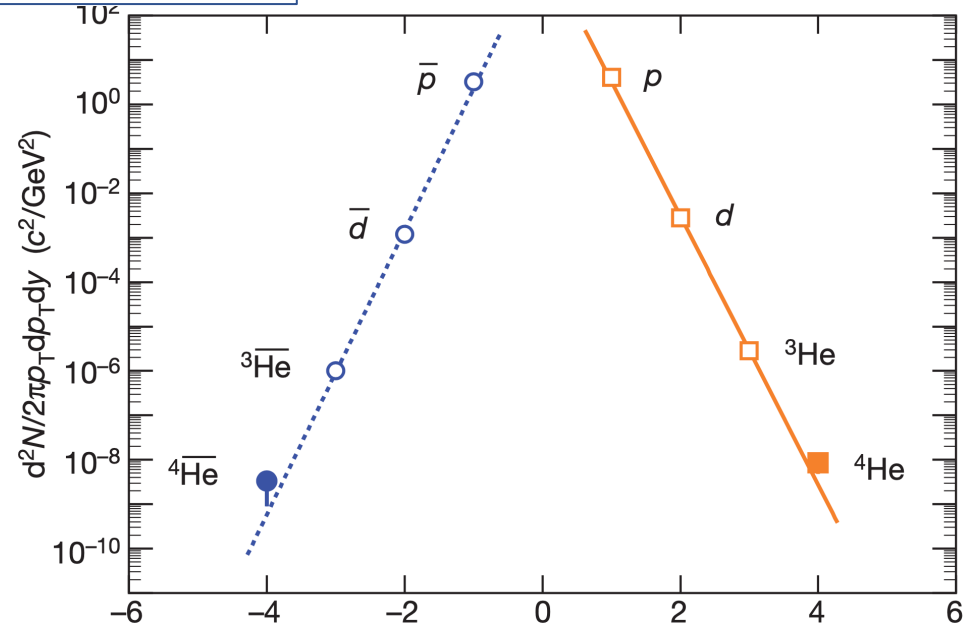
RESEARCH ARTICLE

Observation of an Antimatter Hypernucleus

THE STAR COLLABORATION, B. I. ABELEV, M. M. AGGARWAL, Z. AHAMMED, A. V. ALAKHVERDYANTS, I. ALEKSEEV, B. D. ANDERSON, D. ARK

Y. ZOULKARNEEVA **+382 authors** [Authors Info & Affiliations](#)

SCIENCE • 4 Mar 2010 • Vol 328, Issue 5974 • pp. 58-62 • DOI: 10.1126/science.1183980



Observation of antimatter H4Lambda

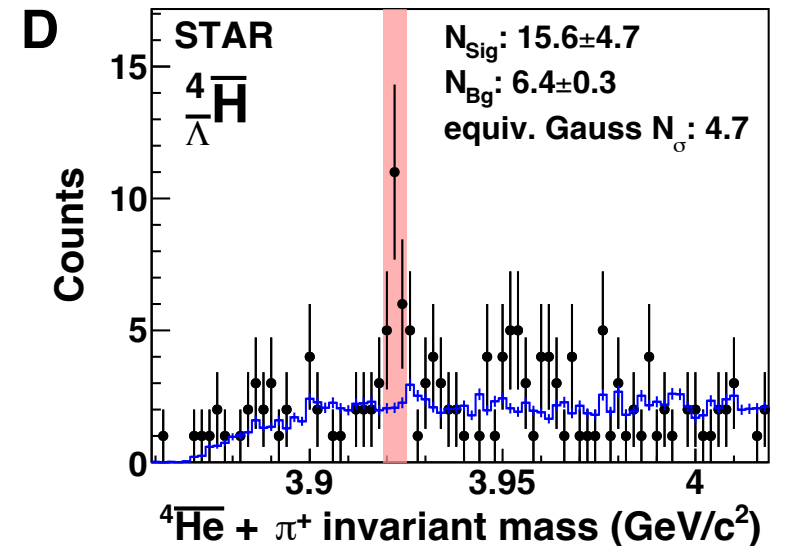
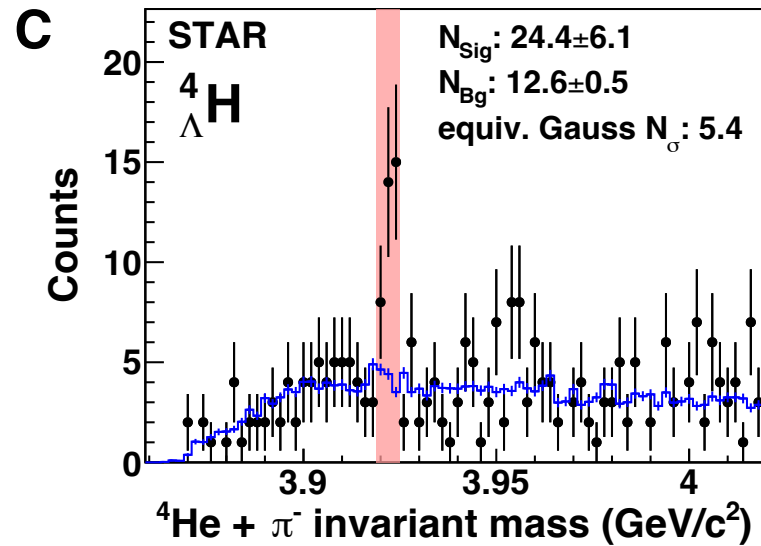
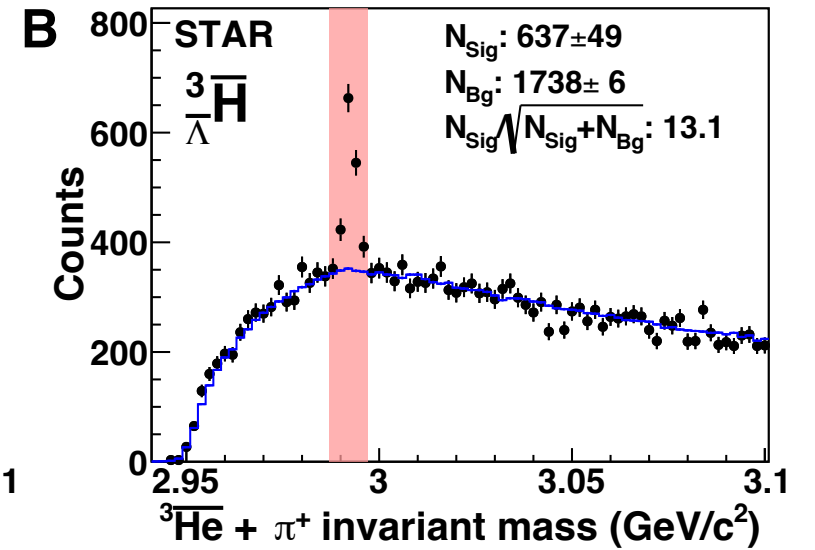
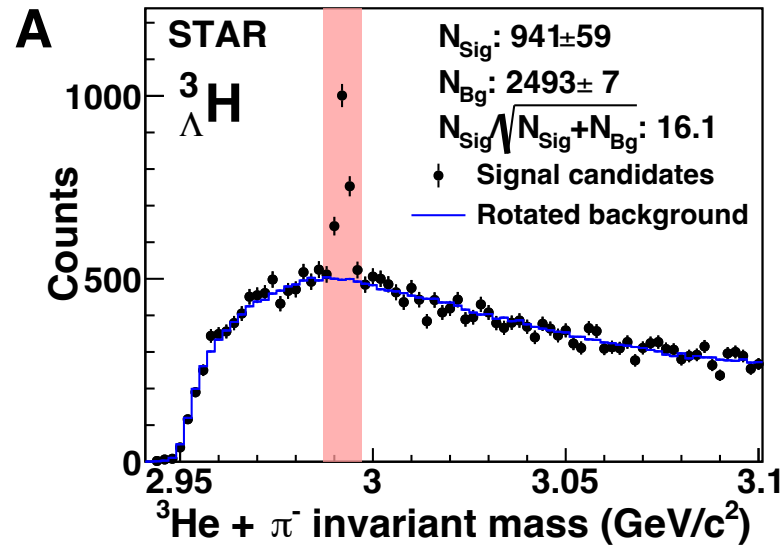
STAR, arXiv: 2310.12674

J.L. Wu (STAR), SQM2022

Au+Au: 2010+2011

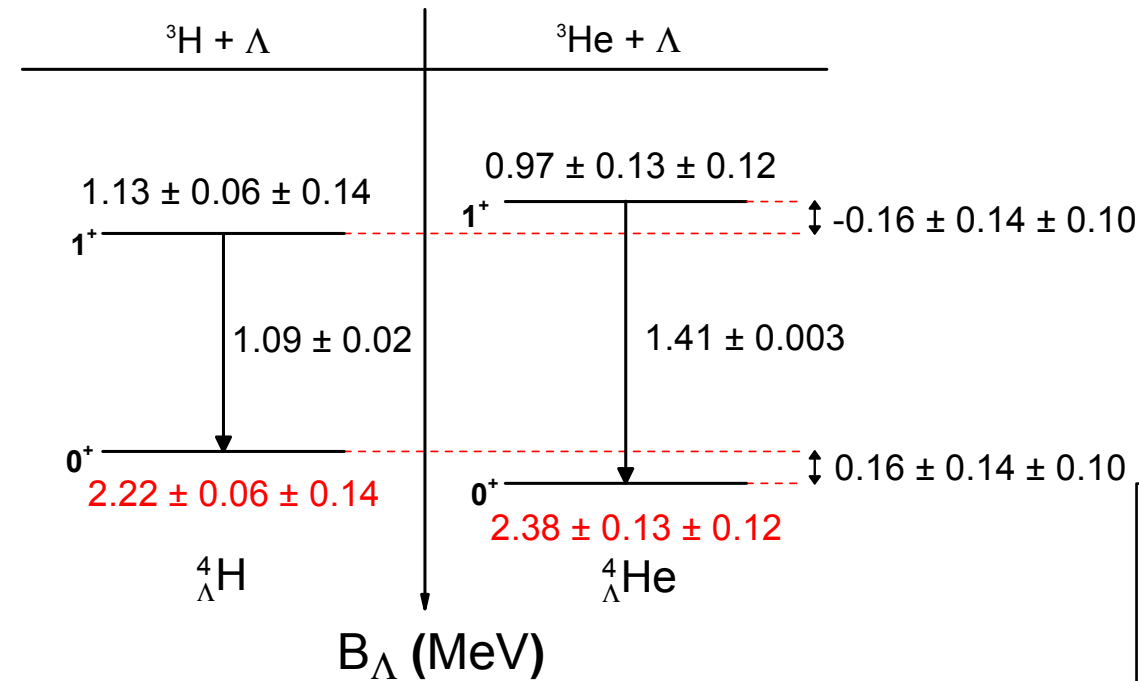
U+U: 2012

Zr+Zr/Ru+Ru: 2018



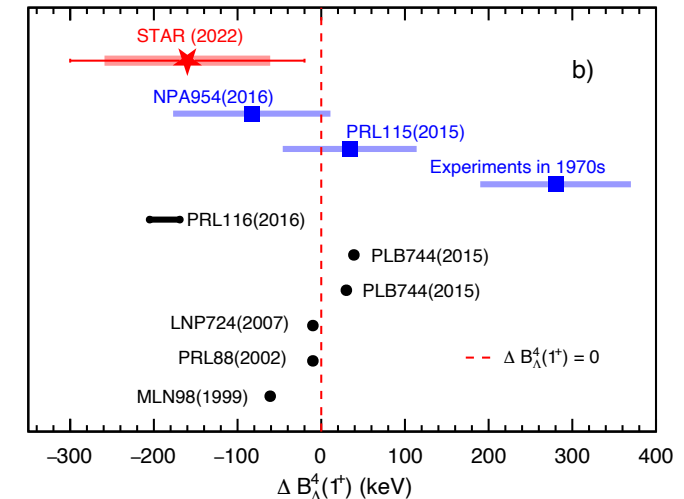
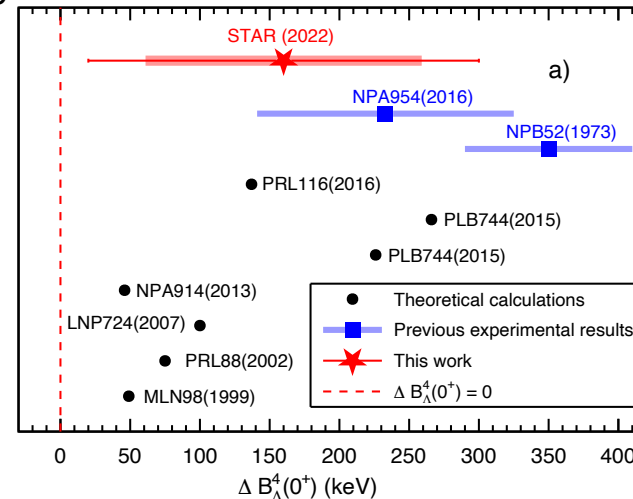
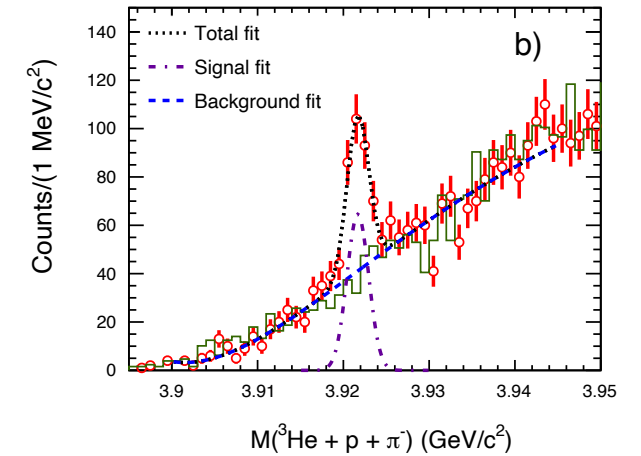
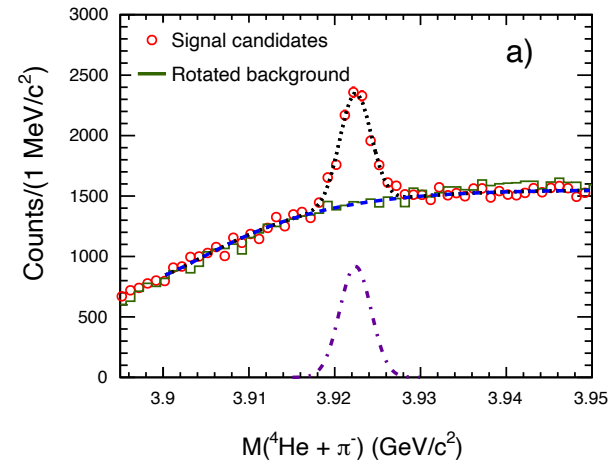
Charge Symmetry Breaking in B=4 hypernuclei

STAR, Phys. Lett. B **834** (2022) 137449



A puzzling CSB (70s):
both 0^+ and 1^+ large and positive ΔB

New measurements:
small and symmetric ΔB





Discovery potential at EIC

Heavy-flavor states

http://belle.kek.jp/belle/talks/moriondQCD10/pakhlov.ppt

Many (>10) states poorly consistent with quark model

State	M (MeV)	Γ (MeV)	J^{PC}	Decay Modes	Production Modes
$Y_s(2175)$	2175 ± 8	58 ± 26	1^{--}	$\phi f_0(980)$	e^+e^- (ISR) $J/\psi \rightarrow \eta Y_s(2175)$
$X(3872)$	3871.4 ± 0.6	< 2.3	1^{++}	$\pi^+\pi^- J/\psi,$ $\gamma J/\psi, D\bar{D}^*$	$B \rightarrow KX(3872), p\bar{p}$
$X(3915)$	3914 ± 4	23 ± 9	$0/2^{++}$	$\omega J/\psi$	$\gamma\gamma \rightarrow X(3915)$
$Z(3930)$	3929 ± 5	29 ± 10	2^{++}	$D\bar{D}$	$\gamma\gamma \rightarrow Z(3940)$
$X(3940)$	3942 ± 9	37 ± 17	$0^{?+}$	$D\bar{D}^*$ (not $D\bar{D}$ or $\omega J/\psi$)	$e^+e^- \rightarrow J/\psi X(3940)$
$Y(3940)$	3943 ± 17	87 ± 34	$?^{?+}$	$\omega J/\psi$ (not $D\bar{D}^*$)	$B \rightarrow KY(3940)$
$Y(4008)$	4008^{+82}_{-49}	226^{+97}_{-80}	1^{--}	$\pi^+\pi^- J/\psi$	e^+e^- (ISR)
$X(4160)$	4156 ± 29	139^{+113}_{-65}	$0^{?+}$	$D^*\bar{D}^*$ (not $D\bar{D}$)	$e^+e^- \rightarrow J/\psi X(4160)$
$Y(4260)$	4264 ± 12	83 ± 22	1^{--}	$\pi^+\pi^- J/\psi$	e^+e^- (ISR)
$Y(4350)$	4361 ± 13	74 ± 18	1^{--}	$\pi^+\pi^- \psi'$	e^+e^- (ISR)
$X(4630)$	4634^{+9}_{-11}	92^{+41}_{-32}	1^{--}	$\Lambda_c^+ \Lambda_c^-$	e^+e^- (ISR)
$Y(4660)$	4664 ± 12	48 ± 15	1^{--}	$\pi^+\pi^- \psi'$	e^+e^- (ISR)
$Z(4050)$	4051^{+24}_{-23}	82^{+51}_{-29}	$?$	$\pi^\pm \chi_{c1}$	$B \rightarrow KZ^\pm(4050)$
$Z(4250)$	4248^{+185}_{-45}	177^{+320}_{-72}	$?$	$\pi^\pm \chi_{c1}$	$B \rightarrow KZ^\pm(4250)$
$Z(4430)$	4433 ± 5	45^{+35}_{-18}	$?$	$\pi^\pm \psi'$	$B \rightarrow KZ^\pm(4430)$
$Y_b(10890)$	$10,890 \pm 3$	55 ± 9	1^{--}	$\pi^+\pi^- \Upsilon(1, 2, 3S)$	$e^+e^- \rightarrow Y_b$

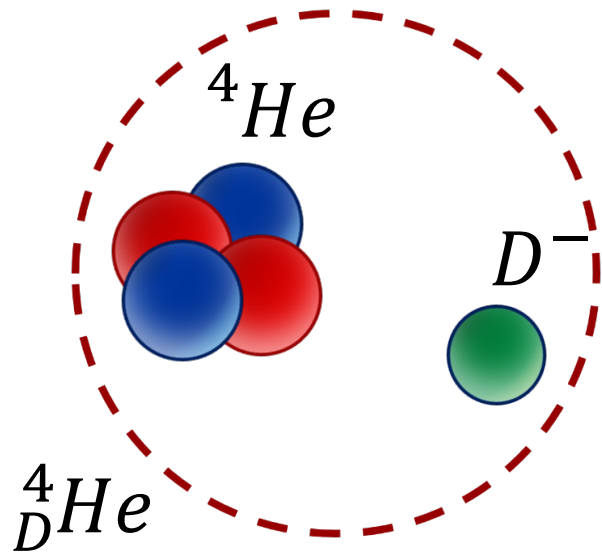
observed last 6 years by B-factories

How about baryon states?

Heavy-flavor hypernuclei

- Predicted to exist (70's)
- Cannot be produced in pp, ep collisions
- Cannot be detected in fixed target experiment
- Only solution: **EIC**
- EIC enough energy for charm and bottom hypernuclei
- Vertex detector at Fragmentation region
Displace vertex: 3cm

Search for Stable Charmed Mesic Nucleus $D^-^4\text{He}$ in Heavy-Ion and EIC



Stable and existence due to Coulomb force

PYTHIA: $D^-/n \sim 5 \times 10^{-4}$ p+p collisions at AGS and RHIC forward kinematics

$D^-^4\text{He}$ yield 10^{-8} per collision

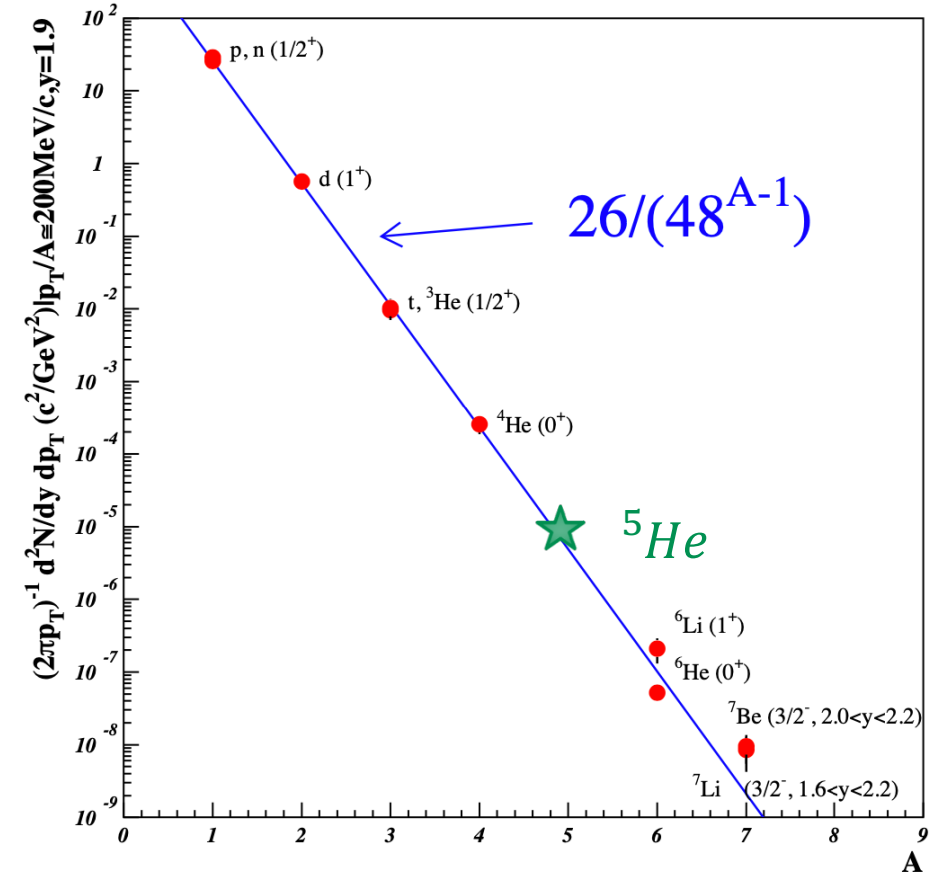
STAR@RHIC:

Estimate 1×10^5 /year in forward acceptance
But without vertex detector

CBM@FAIR high baryon, good vertex
LHCb@LHC forward with good vertex

EIC ion forward direction:
clean environment with good vertex
Nuclear cluster

Zhangbu Xu (BNL)
Cheng-Wei Lin, Yi Yang (NCKU)
DNP (2022), EMMI (2023)

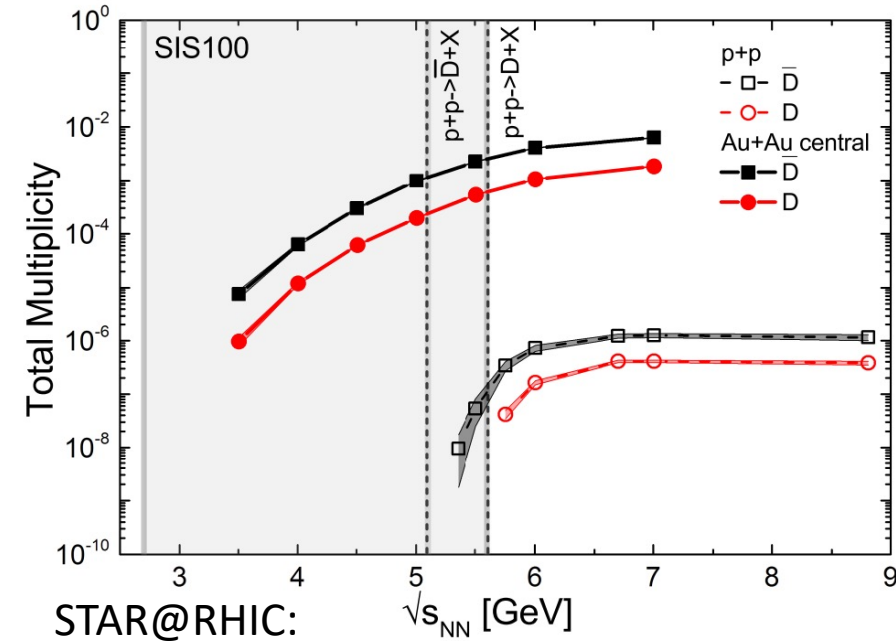
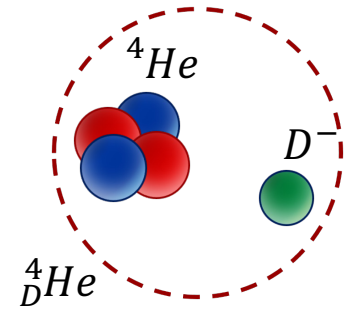


Phys. Rev. C **61**, 064908

Search for Stable Charmed Mesic Nucleus $D^- \text{ } ^4\text{He}$ at CBM

J. Steinheimer, A. Botvina, M. Bleicher, PRC 95 (2017) 014911

Zhangbu Xu (BNL)
Cheng-Wei Lin, Yi Yang (NCKU)
DNP (2022), EMMI (2023)



STAR@RHIC:
Estimate 1×10^5 /year in forward acceptance
But without vertex detector

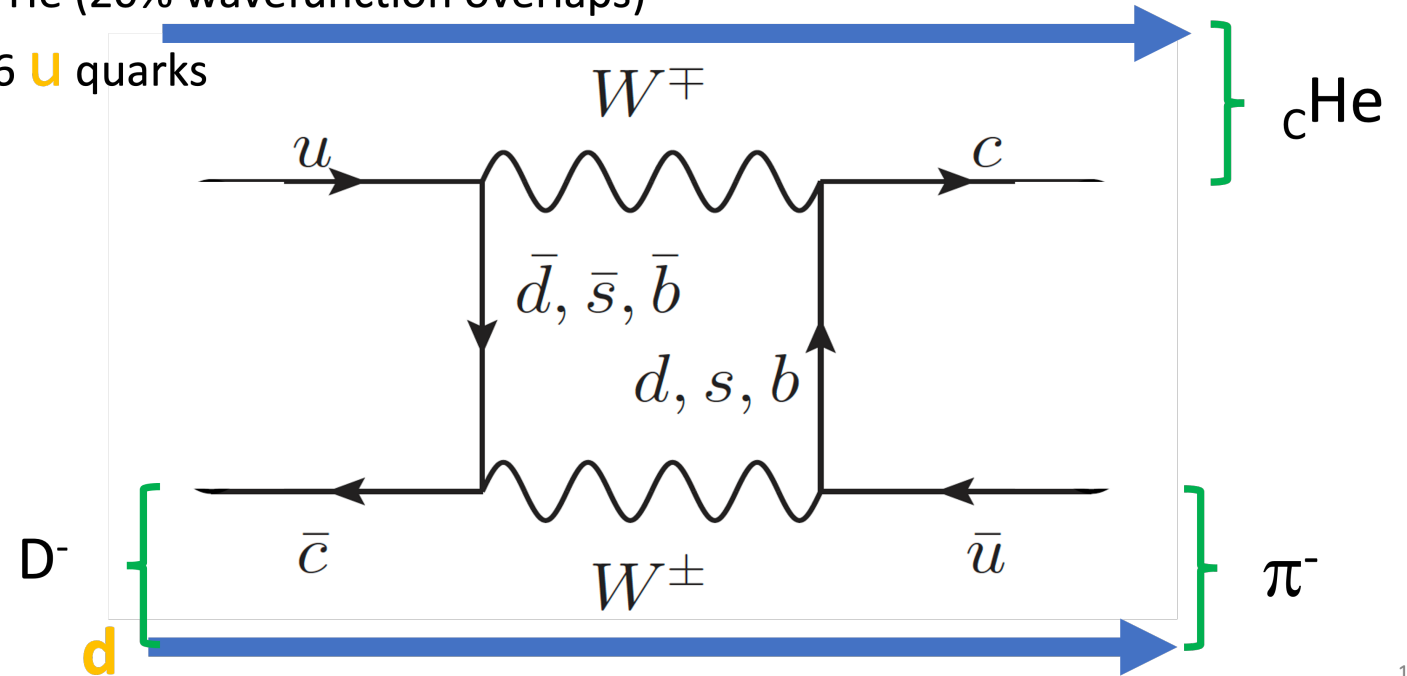
CBM@FAIR high baryon, good vertex
LHCb@LHC forward with good vertex

EIC ion forward direction:
clean environment with good vertex

Charm Quark Oscillation with large mass difference

^4He (20% wavefunction overlaps)

6 u quarks



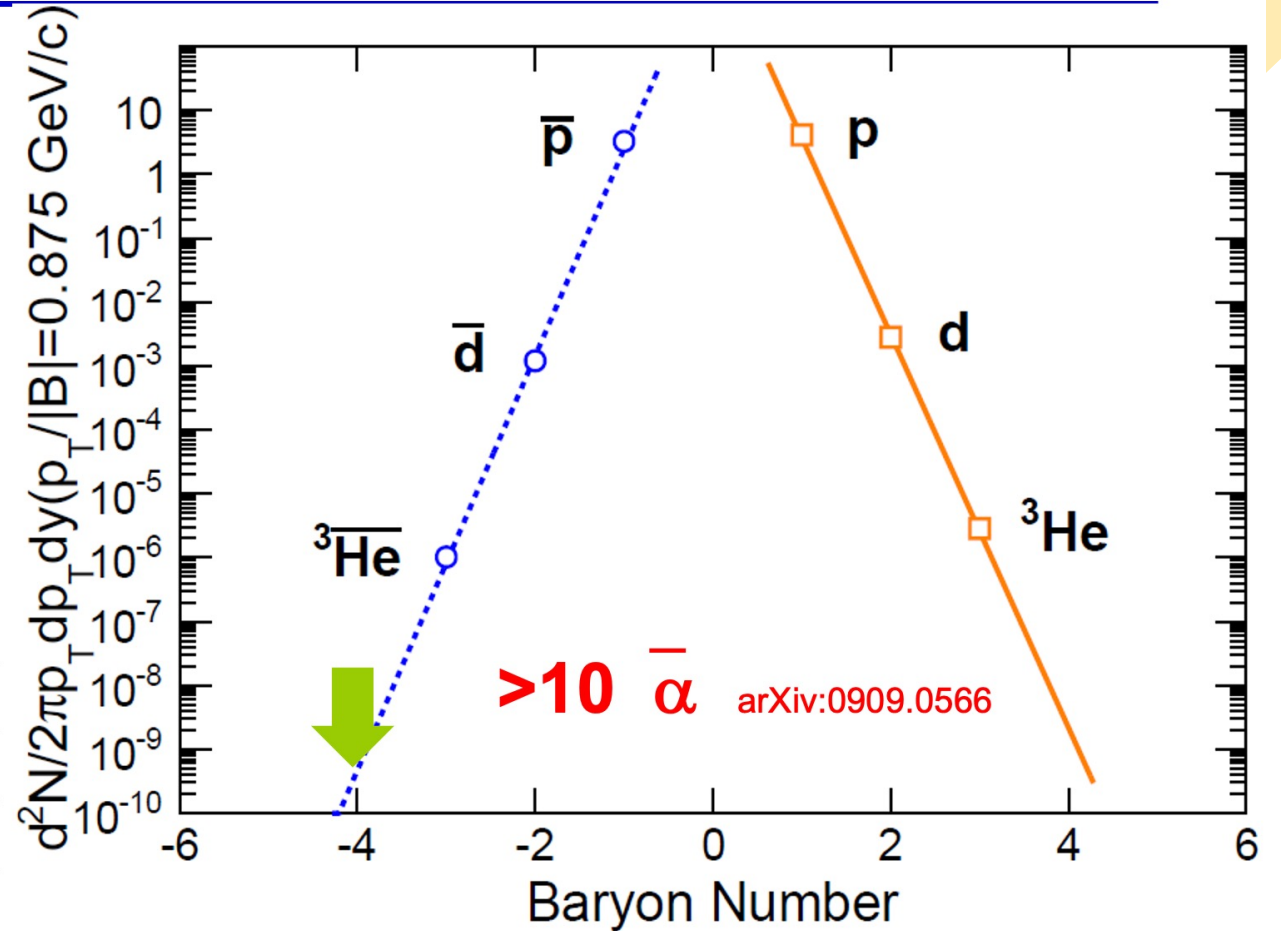
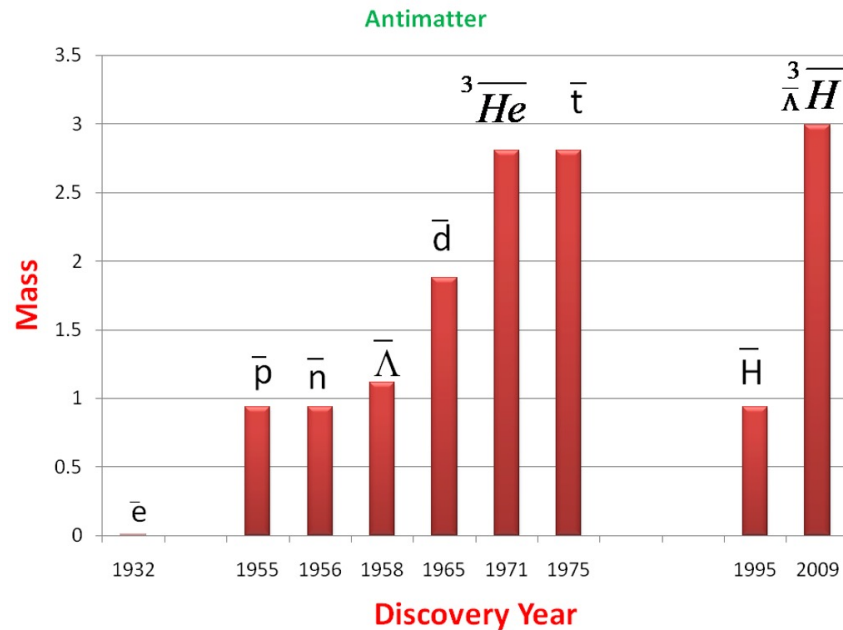


Projected Discovery of antimatter Helium-4 nucleus from STAR

How many possible antimatter nuclei can we discover?

Anti-hypertriton, anti-alpha;
Anti-hyperH4?

Can we get to antimatter ${}^6\text{He}$?
Unless technology and Physics change dramatically, **NO!**



xzb: AAAS annual meeting 01/2011

Heaviest Antimatter Found; Made in U.S. Atom Smasher

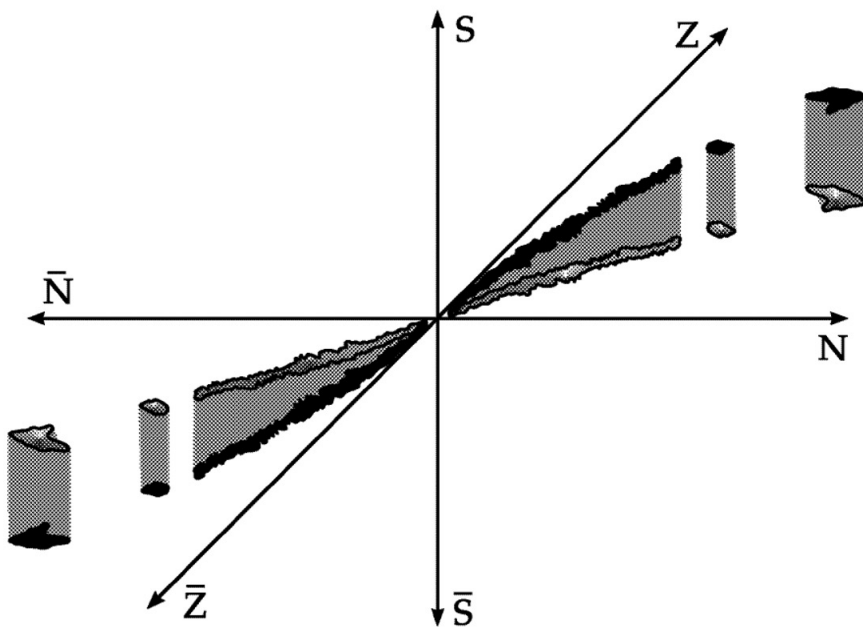


hypernuclei and antimatter from correlations in the Vacuum

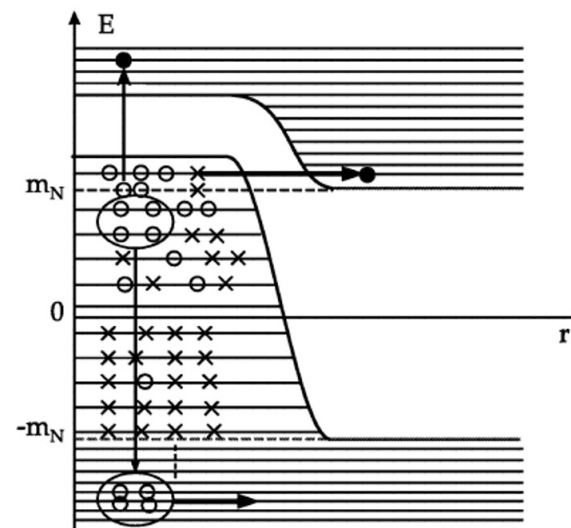
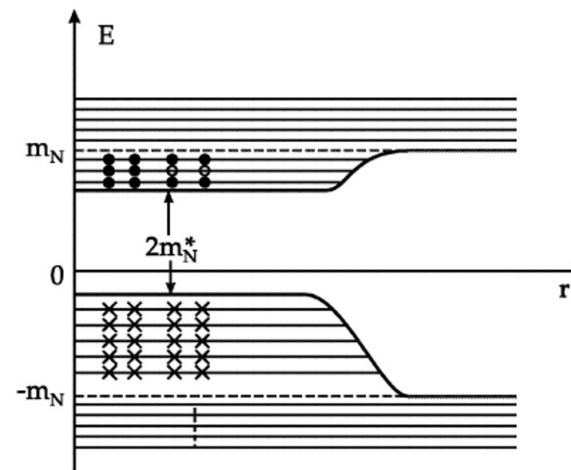
Fundamental Issues in the Physics of Elementary Matter:

Cold Valleys and Fusion of Superheavy Nuclei -
Hypernuclei – Antinuclei – Correlations in the Vacuum

Walter Greiner

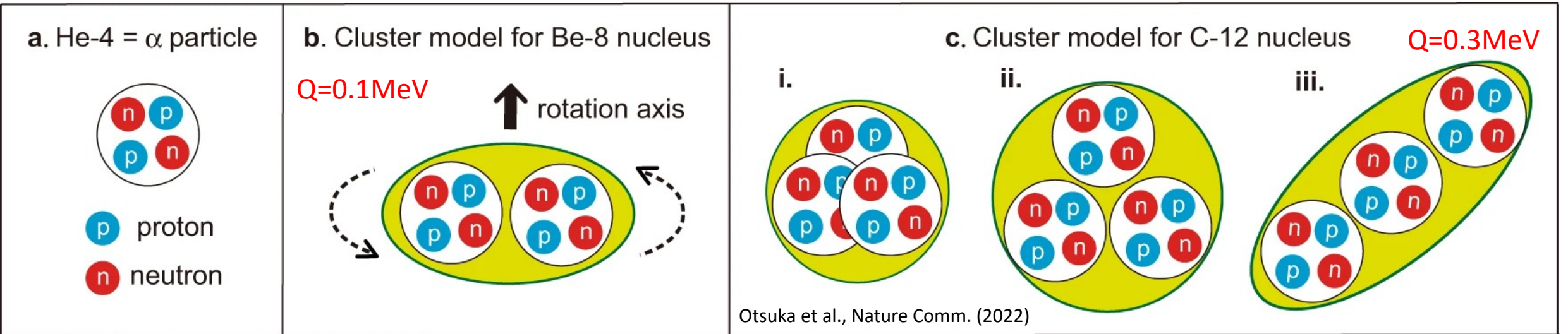


Real 3-D periodic table



Pull α from vacuum (Dirac Sea)

Hoyle Mechanism of creating heavier elements



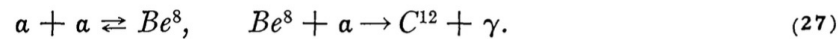
ON NUCLEAR REACTIONS OCCURRING IN VERY HOT STARS. I. THE SYNTHESIS OF ELEMENTS FROM CARBON TO NICKEL

F. HOYLE*

MOUNT WILSON AND PALOMAR OBSERVATORIES
 CARNEGIE INSTITUTION OF WASHINGTON
 CALIFORNIA INSTITUTE OF TECHNOLOGY

Received December 22, 1953

It is convenient to replace reaction (24) by



This is a permissible step, since the lifetime of the unstable Be^8 is appreciably longer than the time required for a "nuclear" collision of two a -particles; that is, longer than the a -particle radius divided by the relative velocity. The merit of reaction (27) is that the number density n_4^8 of Be^8 nuclei is given in terms of the number density n_2^4 of a -particles by the equation of statistical equilibrium (Hoyle 1946),

$$\log n_4^8 = 2 \log n_2^4 - 34.53 - \frac{3}{2} \log T - \frac{0.464}{T}, \quad (28)$$

where the disintegration energy of Be^8 is taken as 0.092 mev, and T is in units of 10^9 K.

PHYSICAL REVIEW

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The 7.68-Mev State in C^{12}

D. N. F. DUNBAR,* R. E. PIXLEY, W. A. WENZEL, AND W. WHALING
 Kellogg Radiation Laboratory, California Institute of Technology, Pasadena, California
 (Received July 21, 1953)

Magnetic analysis of the alpha-particle spectrum from $N^{14}(d,\alpha)C^{12}$ covering the excitation energy range from 4.4 to 9.2 Mev in C^{12} shows a level at 7.68 ± 0.03 Mev. At $E_d = 620$ kev, $\theta_{lab} = 90^\circ$, transitions to this state are only 6 percent of those to the level at 4.43 Mev.

SALPETER¹ and Öpic² have pointed out the importance of the $Be^8(\alpha,\gamma)C^{12}$ reaction in hot stars which have largely exhausted their central hydrogen. Hoyle³ explains the original formation of elements heavier than helium by this process and concludes from the observed cosmic abundance ratios of $O^{16}:C^{12}:He^4$

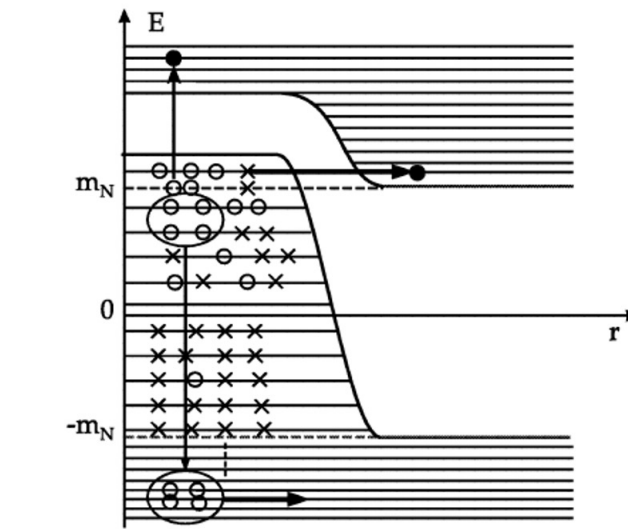
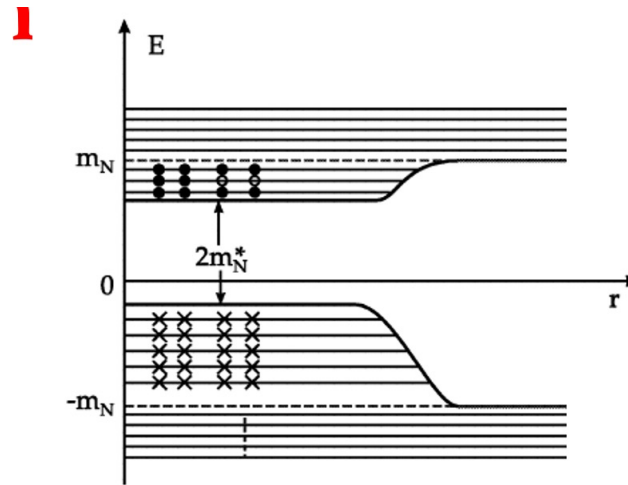
that this reaction should have a resonance at 0.31 Mev or at 7.68 Mev in C^{12} .

An early measurement of the range of the alpha particles from $N^{14}(d,\alpha)C^{12}$ indicated a level in C^{12} at 7.62 Mev.⁴ However, a recent magnetic analysis of this reaction failed to detect a transition to any level

How do Hoyle States appear at RHIC?

In coalescence picture

- Prepare the alpha (anti-alpha) at the early stage (hot and dense)
- Those alpha clusters scatter around, and form heavier ${}^8\text{Be}$ (0.1MeV) and ${}^{12}\text{C}^*$ (0.3MeV) because those excited levels are really close to the free alpha energy level $\delta R \delta p$ (very small δp and very large δR)



Pull $\bar{\alpha}$ from vacuum (Dirac Sea)

In thermodynamics picture

- Prepare the alpha (anti-alpha) at the early stage (hot and dense)
- The alpha clusters re-thermalized at later stage with a large fugacity (similar to charm thermalization) and therefore form Prepare the alpha (anti-alpha) at the early stage (hot and dense) at much later time and lower density and low temperature

Another way of picturing this

No penalty of wavefunction overlap due to late-stage alpha coalescence

$$\frac{N_t \times N_p}{N_d^2} = p_0 \times \left(\frac{R^2 + \frac{2}{3}r_d^2}{R^2 + \frac{1}{2}r_t^2} \right)^3$$

Coalescence wavefunction overlap between nucleus and nucleons

A=8 formation to 4+4 formation (A~5)

Benefit of alpha cluster formation at early stage (two-stage formation)

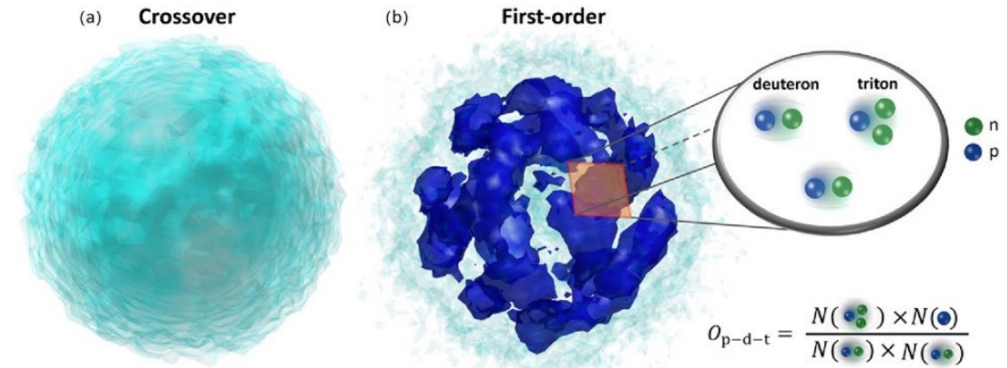
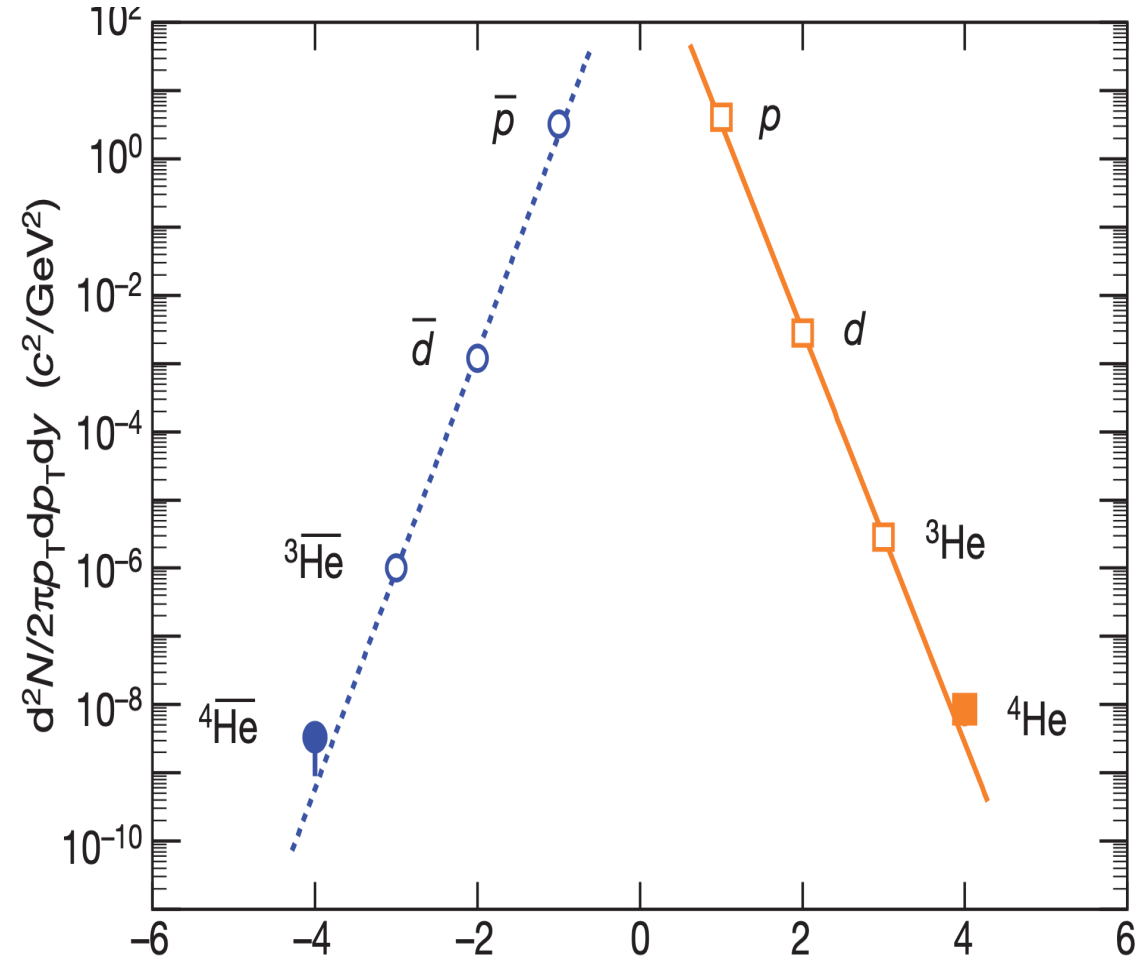
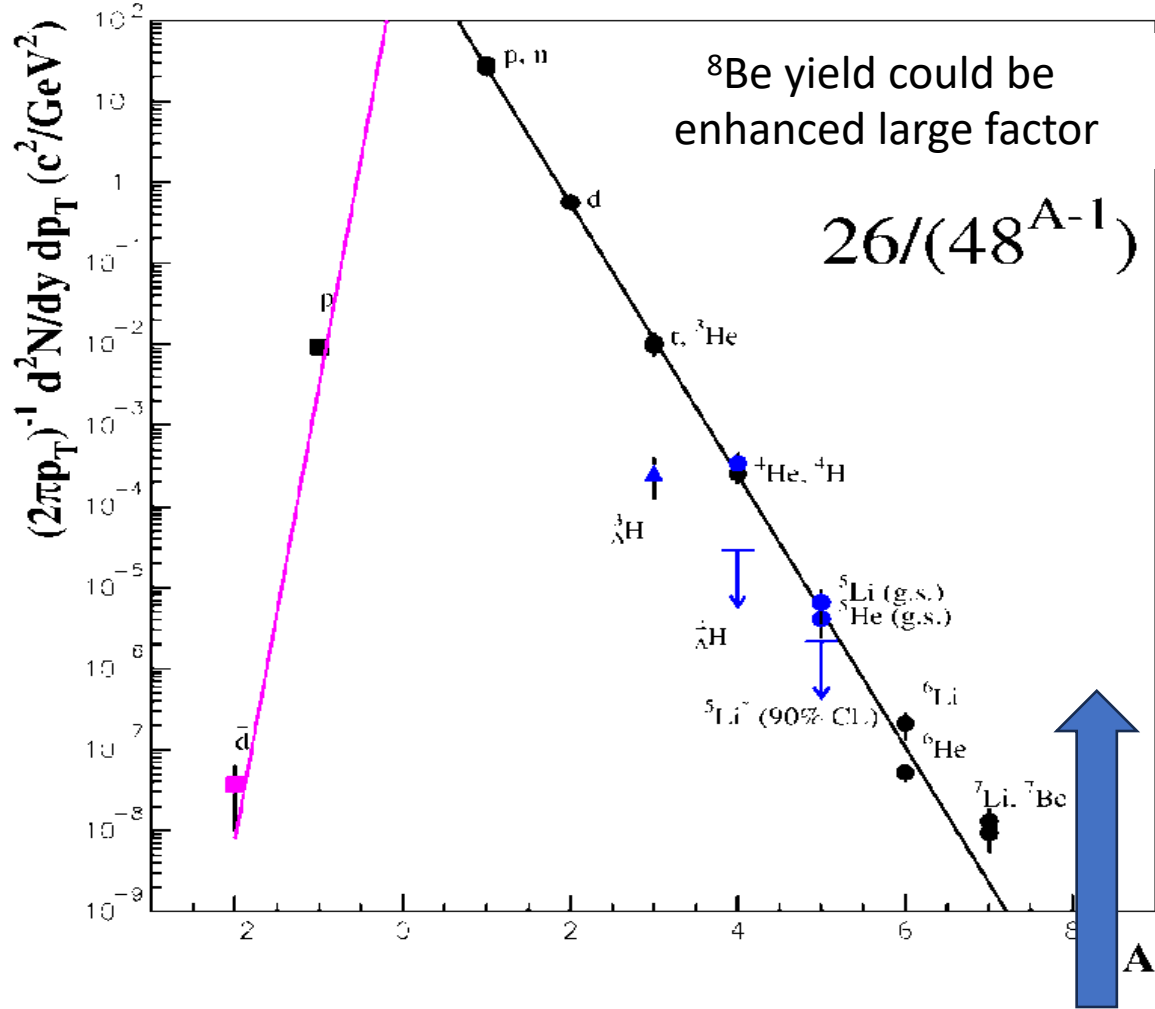


Fig. 1 (Color online) Density distribution of strongly interacting matter in a heavy ion collision after its expansion for the cases of crossover transition (panel **a**) and first-order chiral phase transition (panel **b**). Also shown for illustration of the latter case are deuterons and tritons produced from the density fluctuating hadronic matter and their yield ratio $O_{p-d-t} = N_t N_p / N_d^2$, which depends on the magnitude of neutron density distribution as discussed in the text

^8Be and ^{12}C enhanced yields



Could be between $|A|=5$ and $|A|=8$ extrapolation

Even better at CBM due to slow expansion and redistribution of alpha clusters

Conclusions

- Discovery of the heaviest antimatter nuclear cluster (hyperhydrogen 4)
- Continue to improve our measurements on hypernuclear lifetime and binding energy (CSB)
- Use nuclear yields to study production mechanism, quantum wavefunction overlap: thermal vs coalescence model
- Use nuclear yield ratios as a sensitive probe of nucleon density fluctuation

- Baryon number is a strictly conserved quantum number, keeps the Universe as is
- We did not know what its carrier is; It had not been experimentally verified one way or the other until now;
- Explore other signatures
- Charmed hypernuclei (EIC, LHC, CBM)
- Hoyle States (STAR FXT, **CBM**) and antimatter ${}^8\text{Be}$ (LHC?)