

# ExtreMe Matter Institute EMMI

EMMI Rapid Reaction Task Force

Understanding light (anti-)nuclei  
production at RHIC and LHC

April 8 - 12, 2024

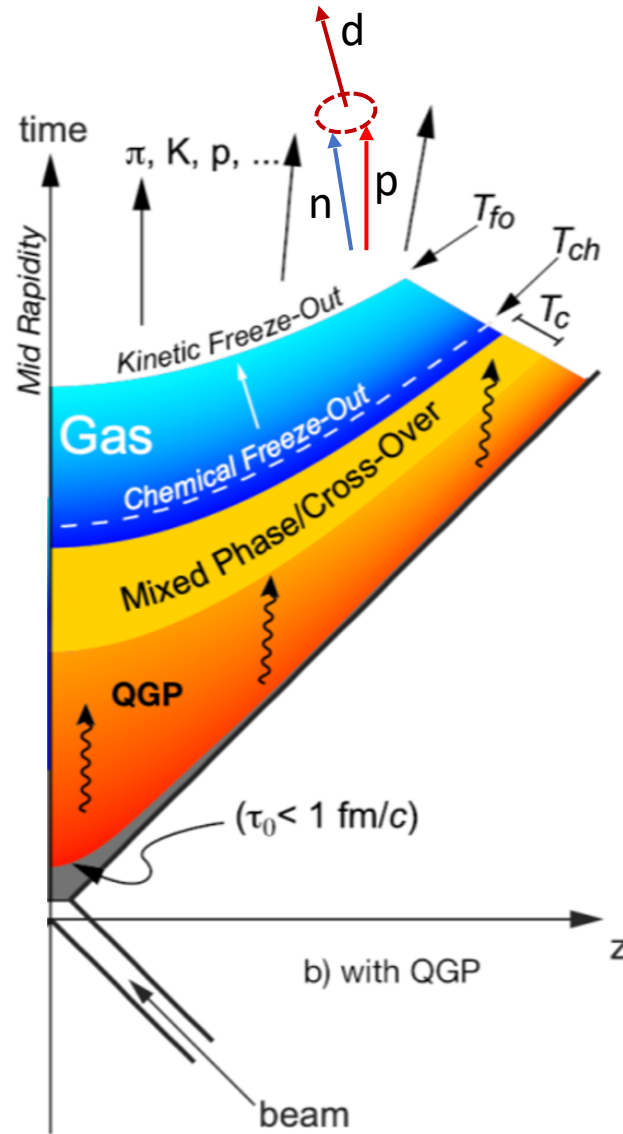
SB1 Lecture Hall, GSI, Darmstadt , Germany



## Experimental overview on light (anti) (hyper)nuclei production at the LHC

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Chiara Pinto  
(CERN)



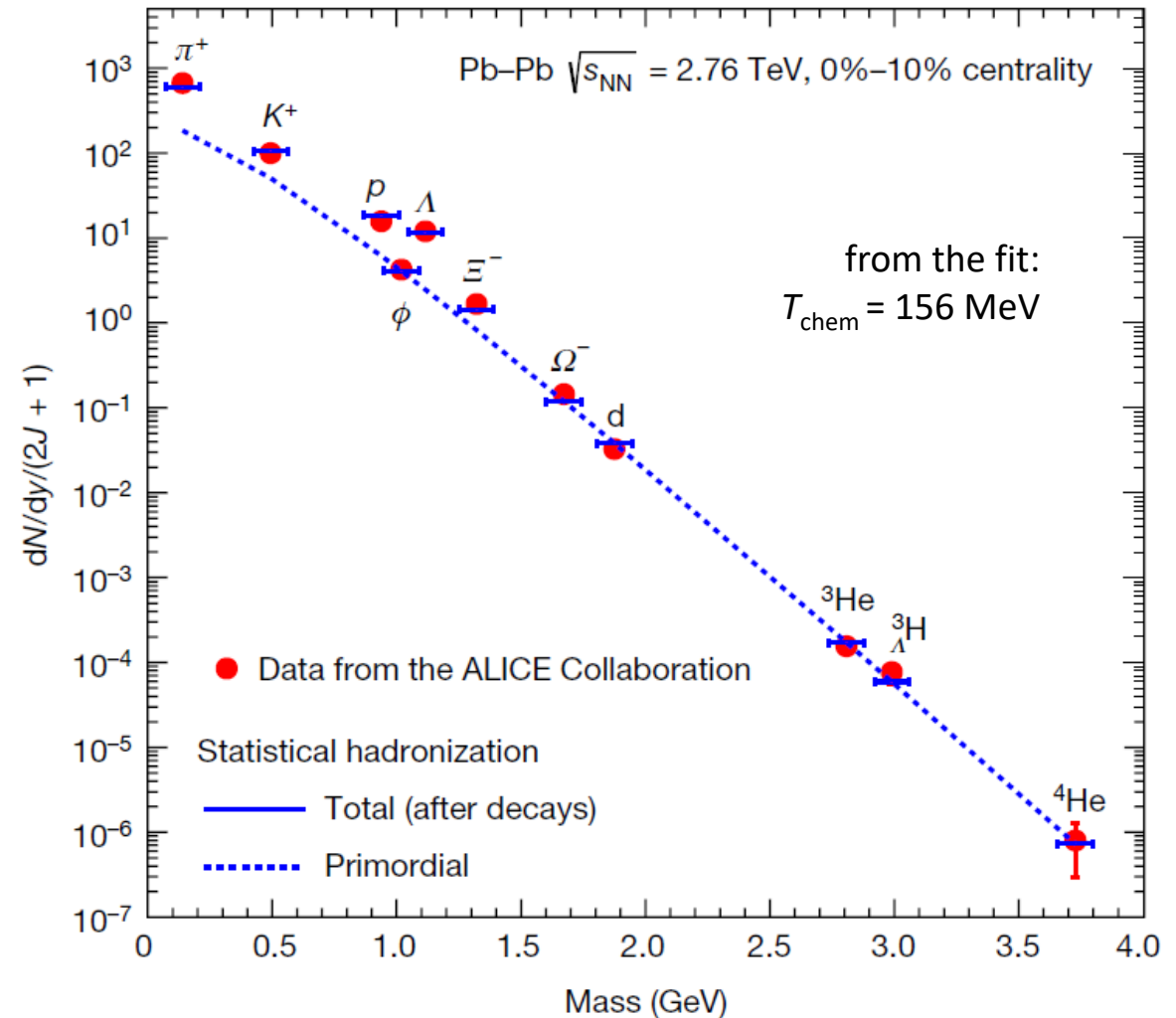
- At LHC energies ( $\sqrt{s} \sim 1\text{--}13 \text{ TeV}$ ) same amount of matter and anti-matter is measured ( $\mu_B \sim 0$ )
- Production mechanism still under debate
- Two classes of phenomenological models available:
  - **statistical hadronization**  $\rightarrow$  works very well for integrated yields (even for nuclei!)
  - **coalescence**  $\rightarrow$  describes fairly well the ratio to protons of integrated yields

## Statistical models (SHMs)

- Hadrons emitted from a system in statistical and chemical equilibrium
- $dN/dy \propto \exp(-m/T_{\text{chem}})$   
 $\Rightarrow$  Nuclei (large  $m$ ): large sensitivity to  $T_{\text{chem}}$
- Light nuclei are produced during phase transition (as other hadrons)
- Typical binding energy of nuclei  $\sim$  few MeV ( $E_B \sim 2$  MeV for  $d$ )

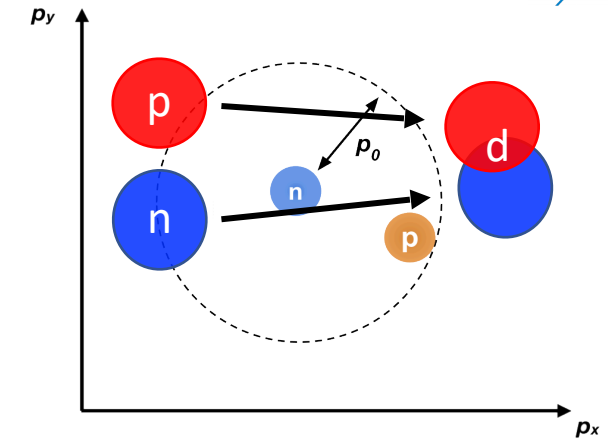
$\Rightarrow$  *how can they survive the hadronic phase*

*environment ( $T_{\text{chem}} \sim 156$  MeV)?*



## Coalescence models

- Simplest implementation: *spherical approximation*  $\rightarrow$  if (anti)nucleons are close in phase space ( $\Delta\mathbf{p} < \mathbf{p}_0$ ) and match the spin state, they can form a (anti)nucleus

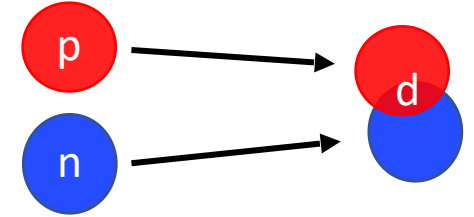



 Butler et al., Phys. Rev. 129 (1963) 836



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- State-of-the-art models use the *Wigner function formalism*  $\rightarrow$  (anti)nuclei arise from the overlap of the (anti)nucleons phase-space distributions with the Wigner density of the bound state

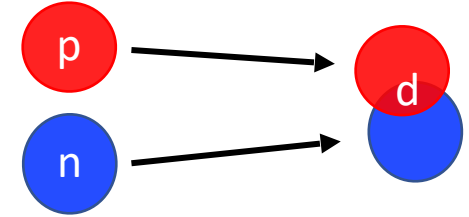


 Butler et al., Phys. Rev. 129 (1963) 836

 Mahlein et al., EPJC 83 (2023) 9, 804

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- Microscopic description
- Key observable is the coalescence parameter  $B_A \rightarrow$  experimental observable tightly connected to the coalescence probability: Larger  $B_A \iff$  Larger coalescence probability



 Butler et al., Phys. Rev. 129 (1963) 836

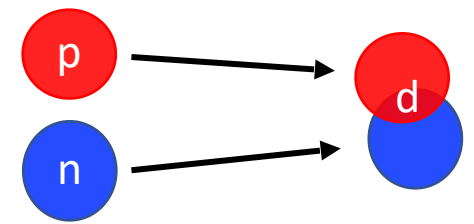
 Mahlein et al., EPJC 83 (2023) 9, 804

$$B_A(p_T^p) = E_A \frac{d^3 N_A}{d p_A^3} / \left( E_p \frac{d^3 N_p}{d p_p^3} \right)^A \Bigg|_{p_T^p = p_T^A / A}$$

# Modelling the production of (anti)nuclei

## Coalescence models

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Butler et al., Phys. Rev. 129 (1963) 836  
 Mahlein et al., EPJC 83 (2023) 9, 804

pp<sup>1</sup>, p—Pb<sup>2</sup>:  $r_0 = 1-1.5$   
fm

$\iff$  large  $B_A$

$$B_A(p_T^p) = E_A \frac{d^3 N_A}{d p_A^3} / \left( E_p \frac{d^3 N_p}{d p_p^3} \right)^A$$

Small distance in space  
 (Only momentum correlations matter)

Pb—Pb<sup>3</sup>:  $r_0 = 3-6$  fm

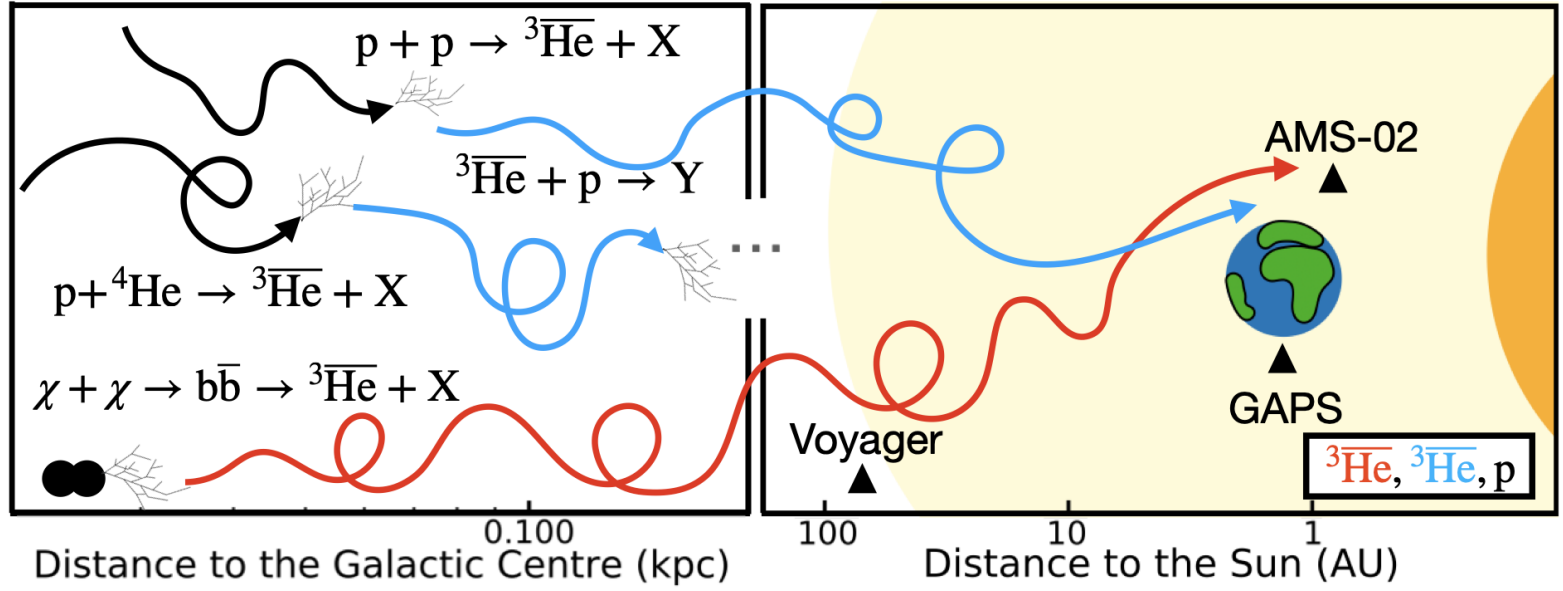
$\iff$  small  $B_A$

$p_T^p = p_T^A / A$

<sup>1</sup> PRC 99 (2019) 024001  
<sup>2</sup> PRL 123 (2019) 112002  
<sup>3</sup> PRC 96 (2017) 064613

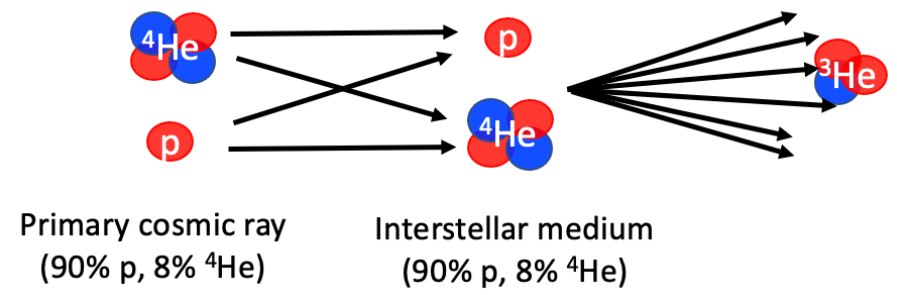
Large distance in space  
 (Both momentum and space correlations matter)

# Astrophysics applications

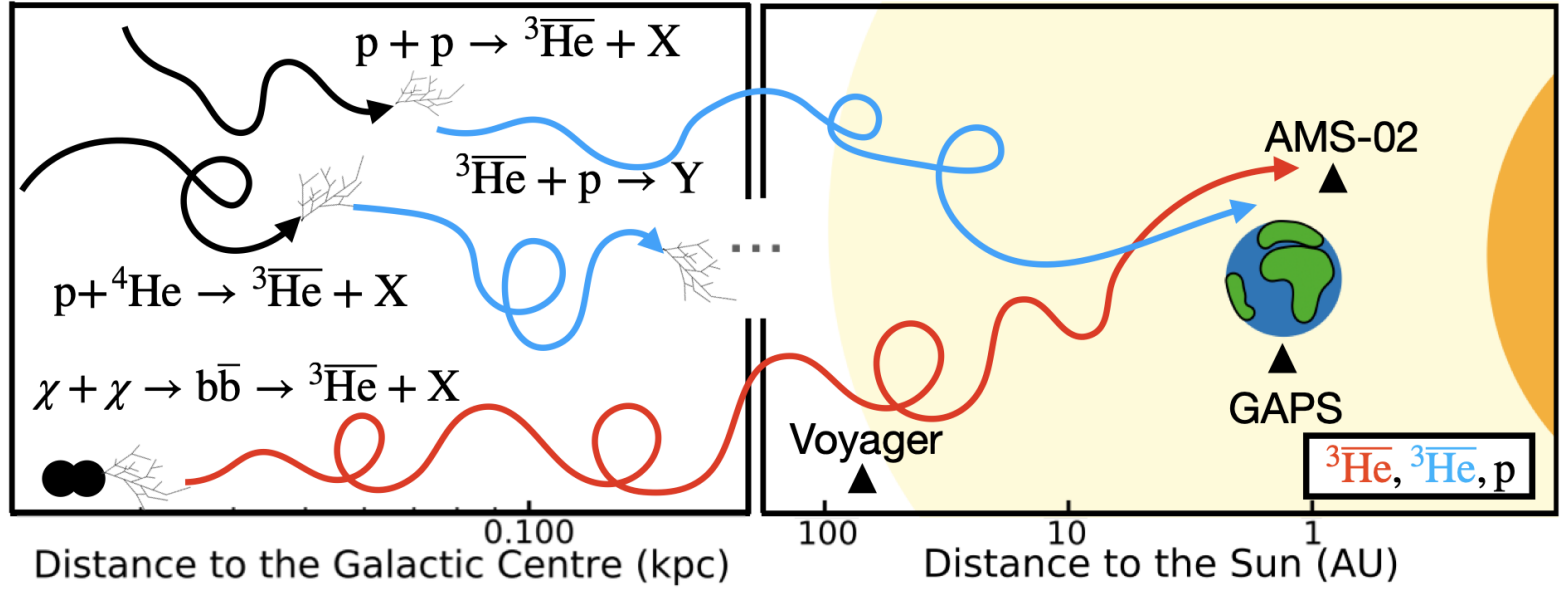


## Antinuclei production:

- pp, pA and (few) AA reactions between primary **cosmic rays** and the interstellar medium
- **dark-matter** annihilation processes

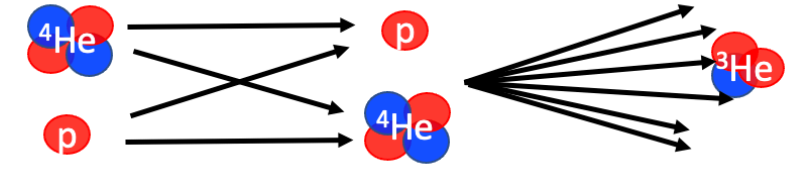


# Astrophysics applications

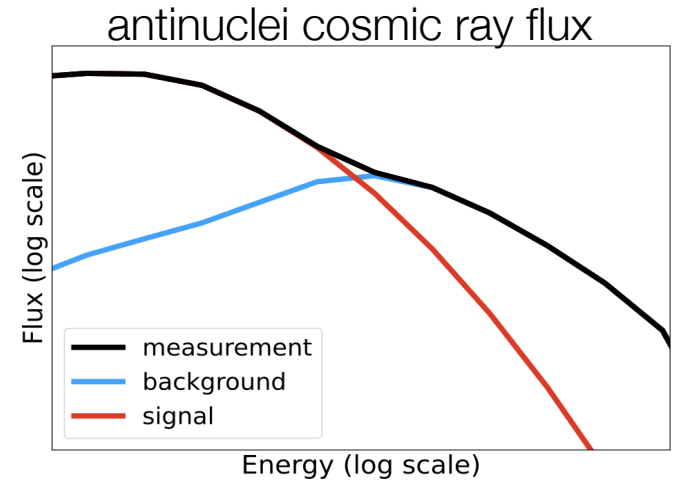


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- pp, pA and (few) AA reactions between primary cosmic rays and the interstellar medium
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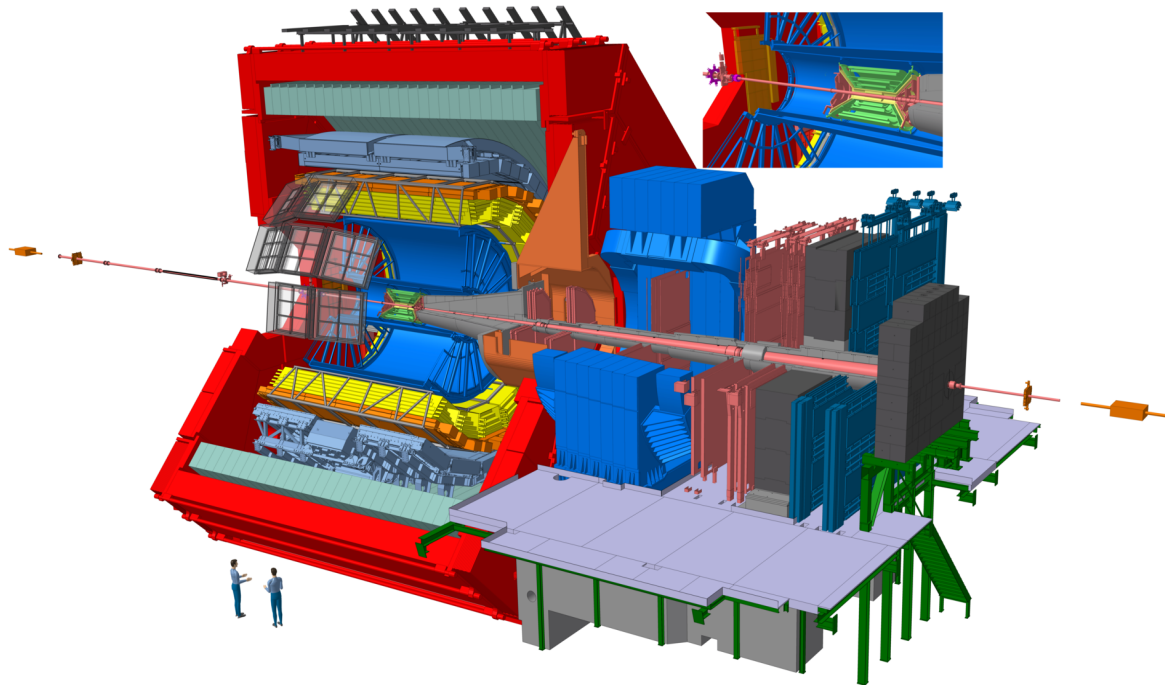


- High Signal/Noise ratio ( $\sim 10^2 - 10^4$ ) at low  $E_{\text{kin}}$  expected by models
- To correctly interpret any future measurement, we need precise knowledge of
  1. production of antinuclei
  2. annihilation



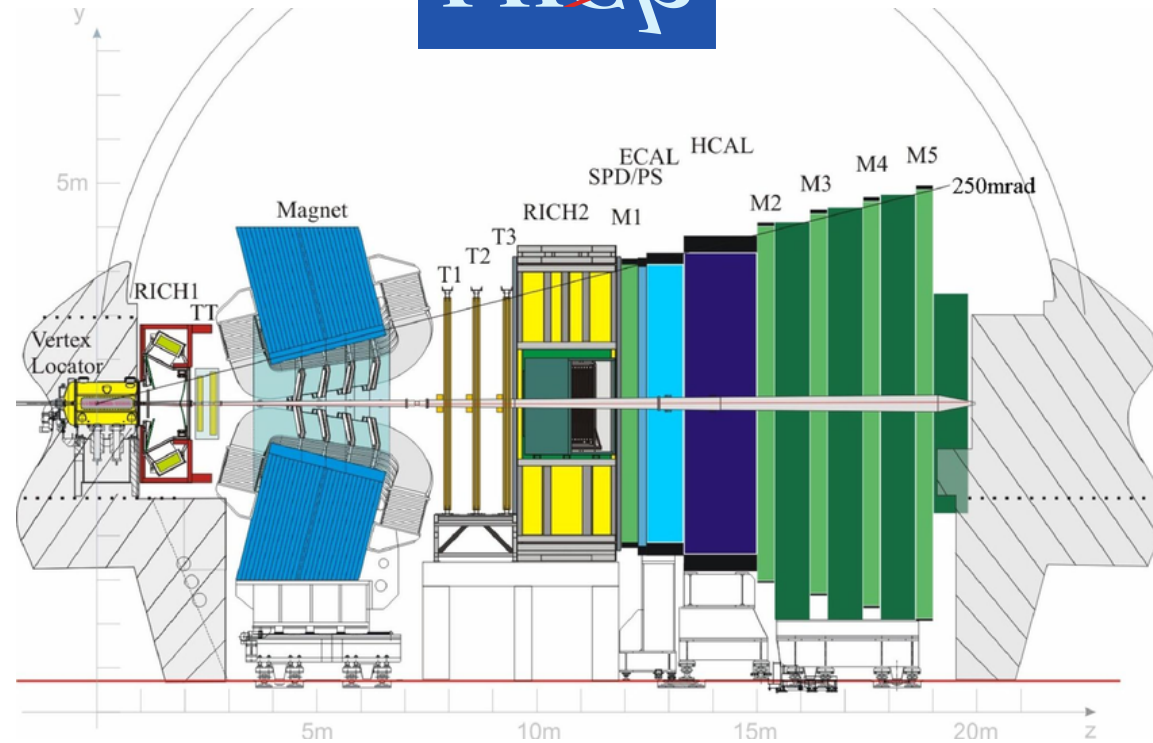


ALICE



ALICE Collaboration, 2008 *JINST* 3 S08002

- excellent tracking & PID capabilities over broad  $p$  range
  - low material budget
- **most suited detector at the LHC for the study of nuclei**



LHCb Collaboration, 2008 *JINST* 3 S08005

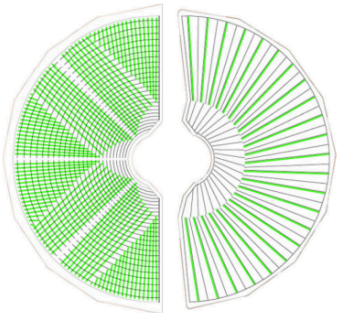
- excellent vertexing ( $\sigma_{IP} = 15+29/p_T$  [GeV]  $\mu\text{m}$ ,  $\sigma_p = 0.5\% - 1.0\%$ )
  - excellent PID separation for  $K$ ,  $\pi$  and  $p$  with  $O(10)$  GeV/c
- **recently joined the nuclei-business!**



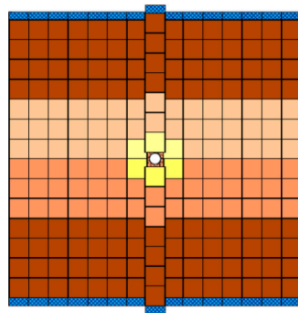
LHCb detector not designed to identify light (anti)nuclei

Use information from the tracking system

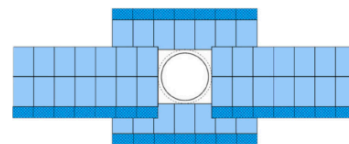
**1** Ionisation losses in silicon sensors:  $Z^2$  dependence in Bethe-Bloch  
 →  $dE/dx$  in VELO, TT, IT to identify He



VELO: 2 x 21 layers

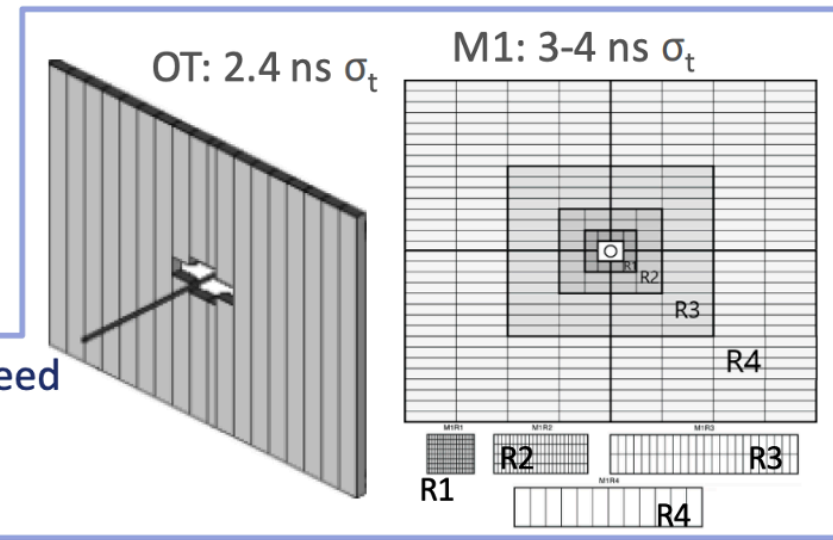
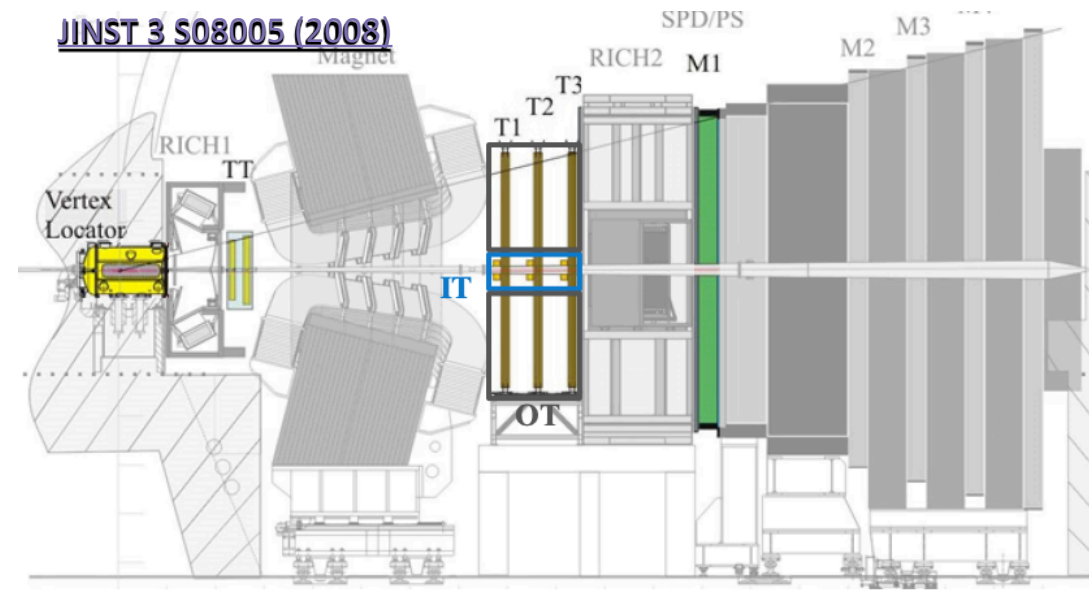


TT: 4 layers

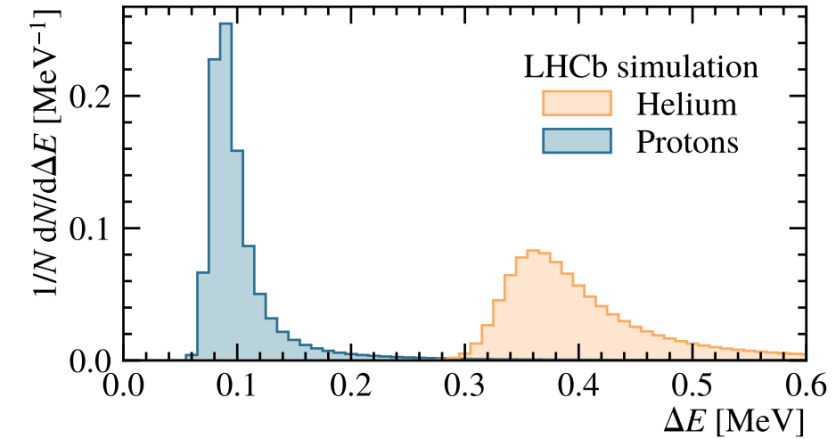


IT: 12 layers

**2** Light nuclei slower than  $c$ :  $M$  dependence of particle speed  
 → Time-of-flight in OT and M1 to identify  $d$ , distinguish  $^3\text{He}$  and  $^4\text{He}$



# Identification of Helium-3 with LHCb



Bethe-Bloch: Z=2 particles deposits  $\sim 4$  times the energy of Z=1 particles  
 $\rightarrow$  He: higher ADC counts and wider cluster size

**First (anti-)Helium candidates observed in  $pp$  in LHCb data!**

Define Likelihood discriminators based on cluster size and ADC counts:

$$\mathcal{L}^X = \left( \prod_{i=1}^n \text{PDD}_i^X \right)^{1/n}, \quad X = \{\text{He, Bkg}\}$$

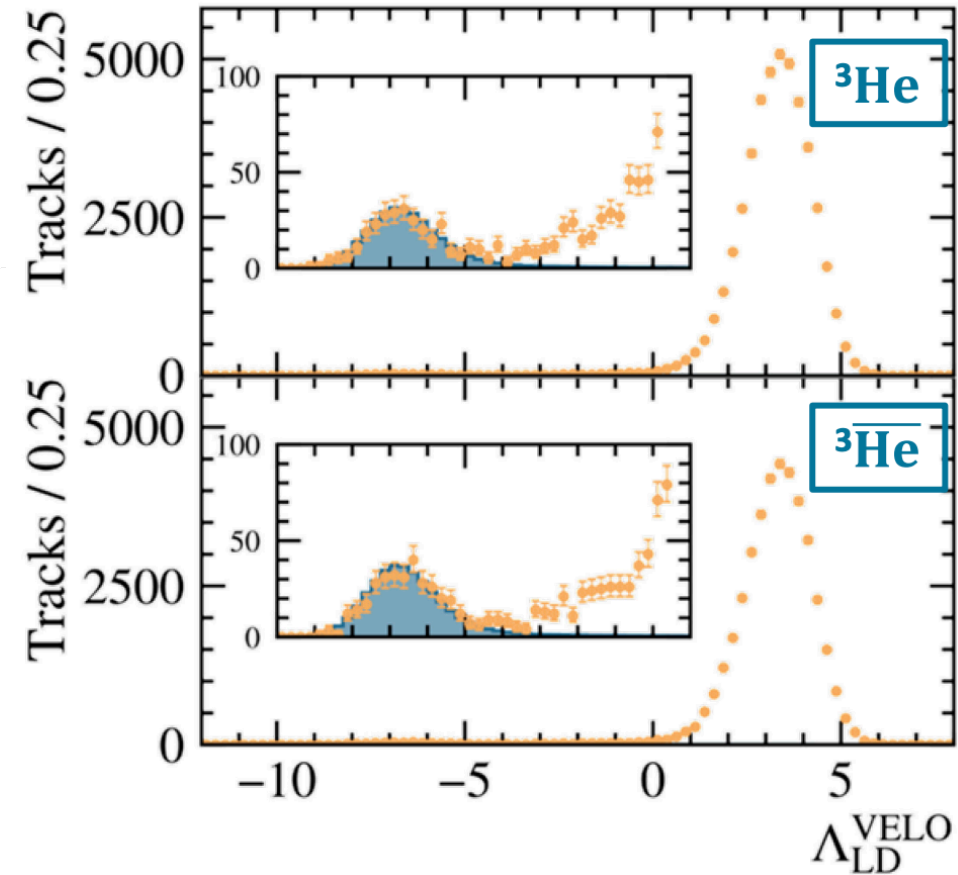
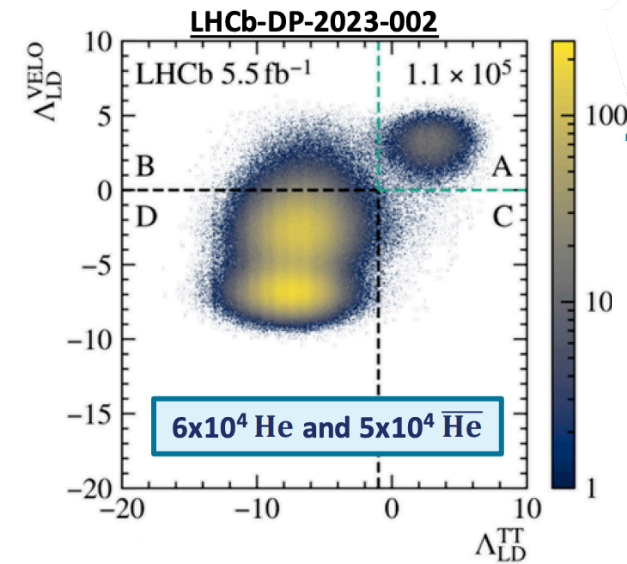
$$\Lambda_{\text{LD}} = \log \mathcal{L}^{\text{He}} - \log \mathcal{L}^{\text{Bkg}}$$

One discriminator for each subdetector:

- $\Lambda_{\text{LD}}^{\text{VELO}}$
- $\Lambda_{\text{LD}}^{\text{TT}}$
- $\Lambda_{\text{LD}}^{\text{IT}}$

**Performance:**

- MisID probability:  $\mathcal{O}(10^{-12})$
- Signal efficiency:  $\sim 50\%$

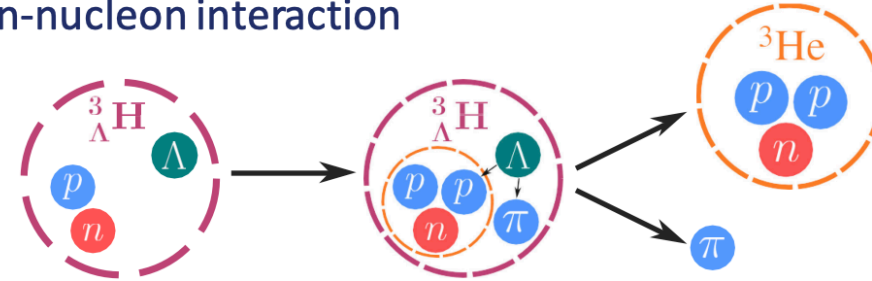




- Hypertriton life-time and binding energy gives access to hyperon-nucleon interaction

→ Constrains on maximum mass of neutron stars

**Search for 2-body decay into He:**



## Results:

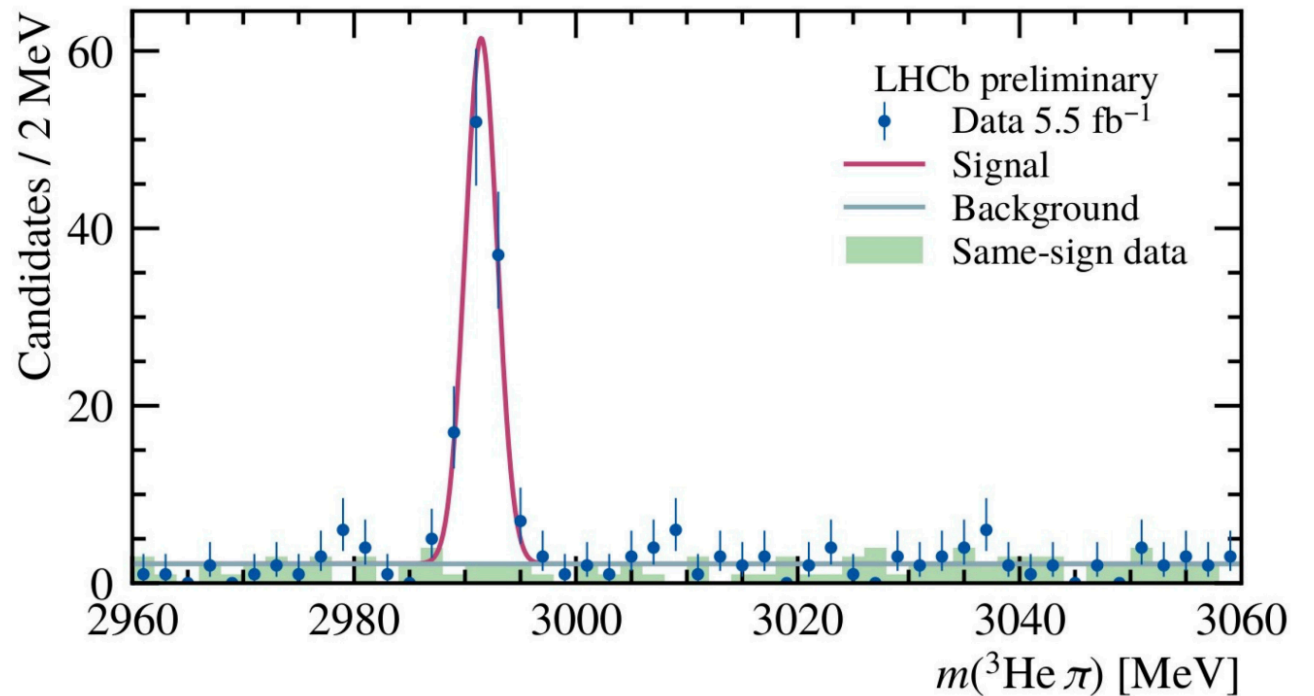
(Run2  $pp$  collisions at  $\sqrt{s} = 13$  TeV)

### • Yields:

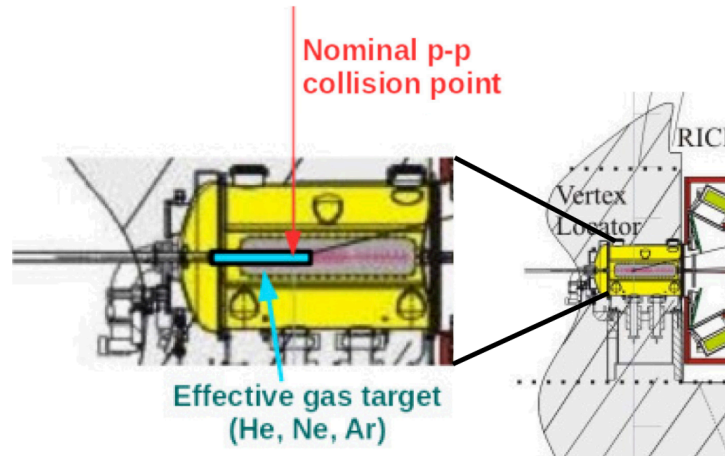
- $61 \pm 8$  Hypertriton
- $46 \pm 7$  anti-Hypertriton

- Statistical mass precision: 0.16 MeV

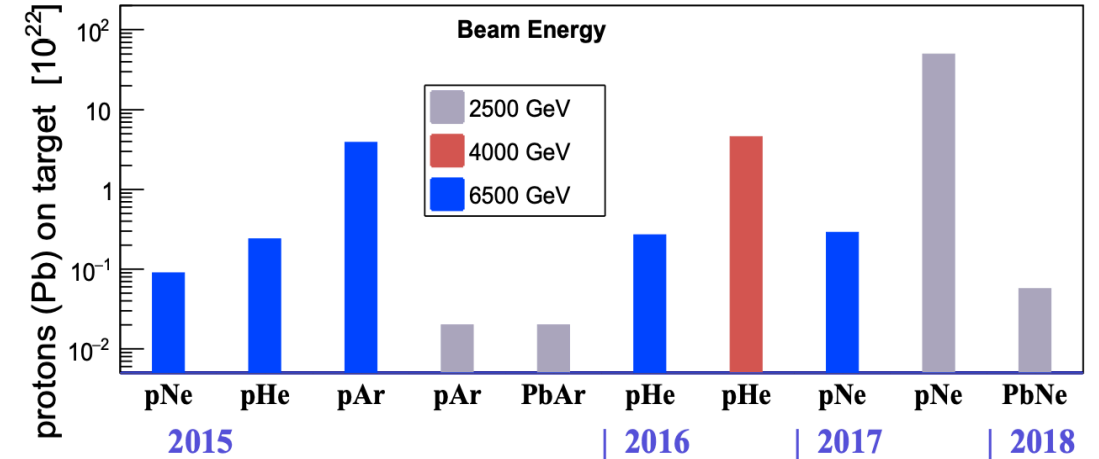
This measurement shows the applicability of  ${}^3\text{He}$  reconstruction and paves the way for future measurements of astrophysical interest



- The *System for Measuring Overlap with Gas (SMOG)* can inject gas in LHC beam pipe around  $\pm 20$  m from the LHCb IP
- SMOG exploited for LHCb **fixed-target physics programme**  
→ Collected physics samples with different **targets** and different **centre of mass energies**



**LHCb contribution is relevant for astrophysics applications!**



Unique opportunities at the LHC:

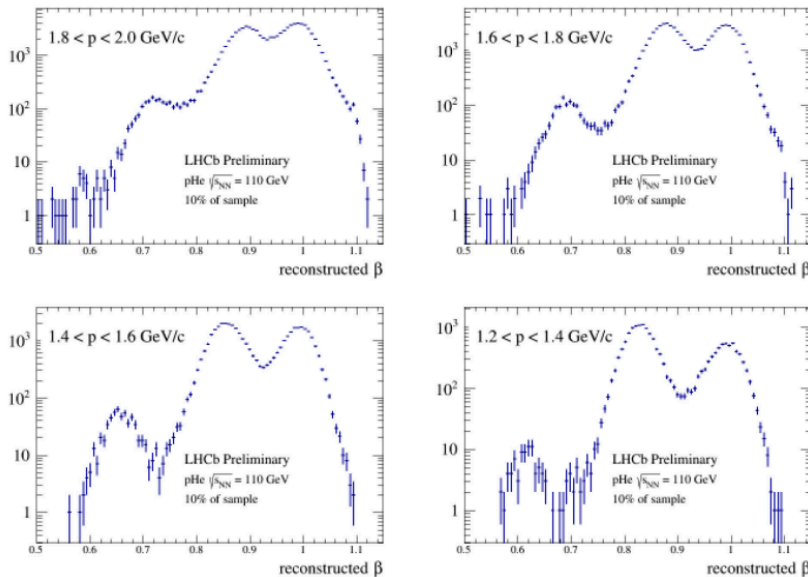
- Collisions with targets of mass number  $A$  intermediate between  $p$  and  $Pb$  → **Reproduce CR interactions (pp, pHe)**
- Energy range  $\sqrt{s_{NN}} \in [30, 115]$  GeV for beam energy in  $[0.45, 7]$  TeV → Unexplored gap between SPS and LHC/RHIC

LHCb is now also capable of measuring (anti)deuterons

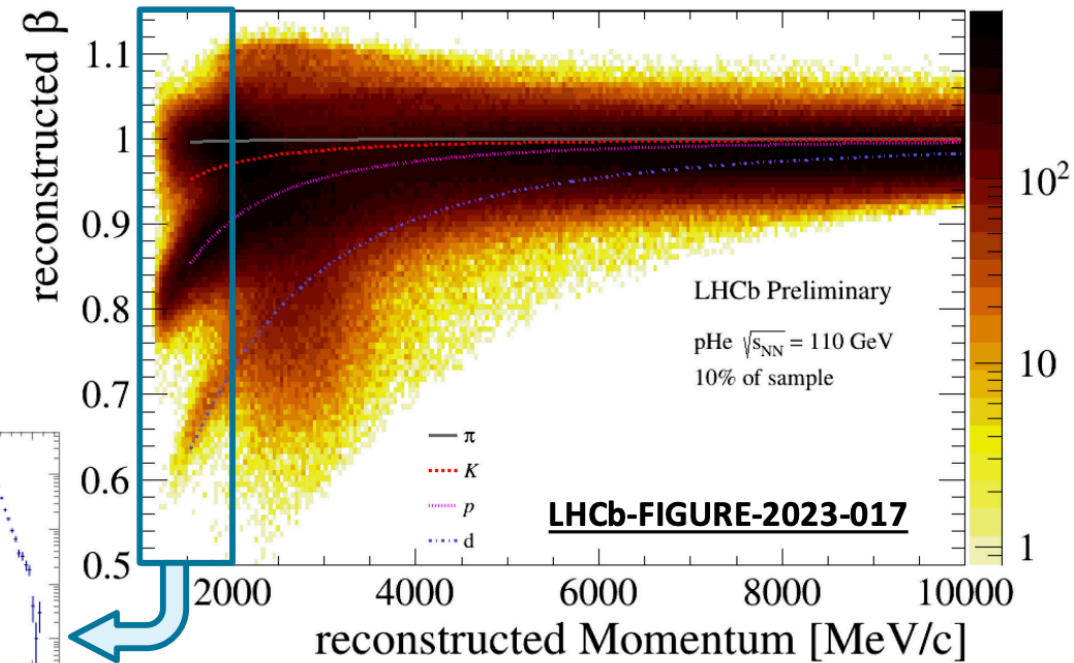
- *Time-of-flight based technique*
- Reconstructed tracks refitted to determine  $\beta$   
→ iterative procedure rerunning Kalman fit with different  $\beta$  hypotheses

- **~10% of SMOG pHe**  
( $\sqrt{s_{NN}} = 110$  GeV) dataset
- **Background suppression:**  
 $\sigma(\beta) < 0.02$ ,  $\chi^2_{\text{OThits}}/\text{ndf} < 2$

**First deuteron candidates  
observed in pHe data!**



LHCb-FIGURE-2023-017

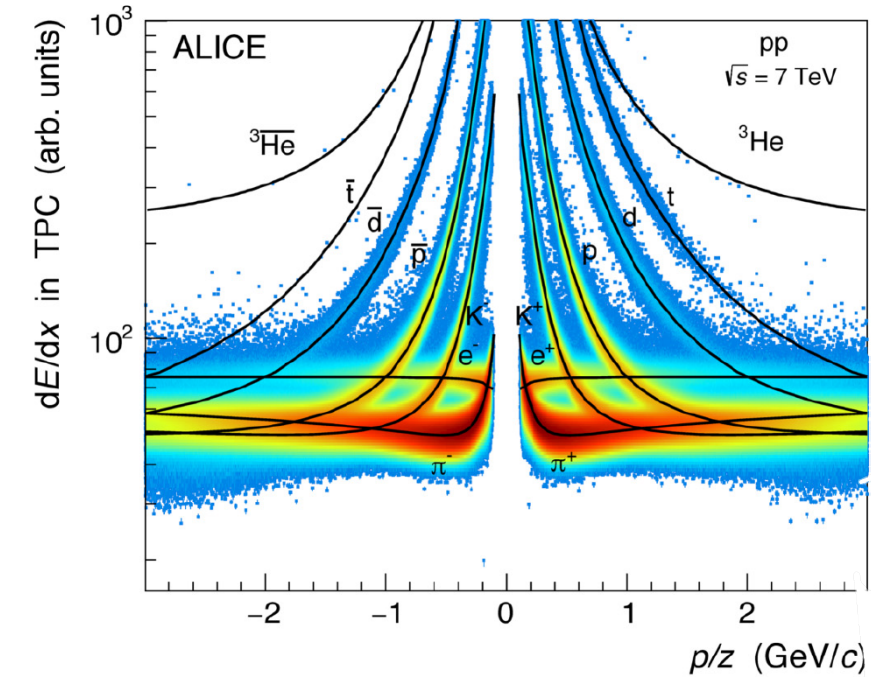


LHCb-FIGURE-2023-017

[https://cds.cern.ch/record/2881940/files/MPI23\\_v1.pdf](https://cds.cern.ch/record/2881940/files/MPI23_v1.pdf)

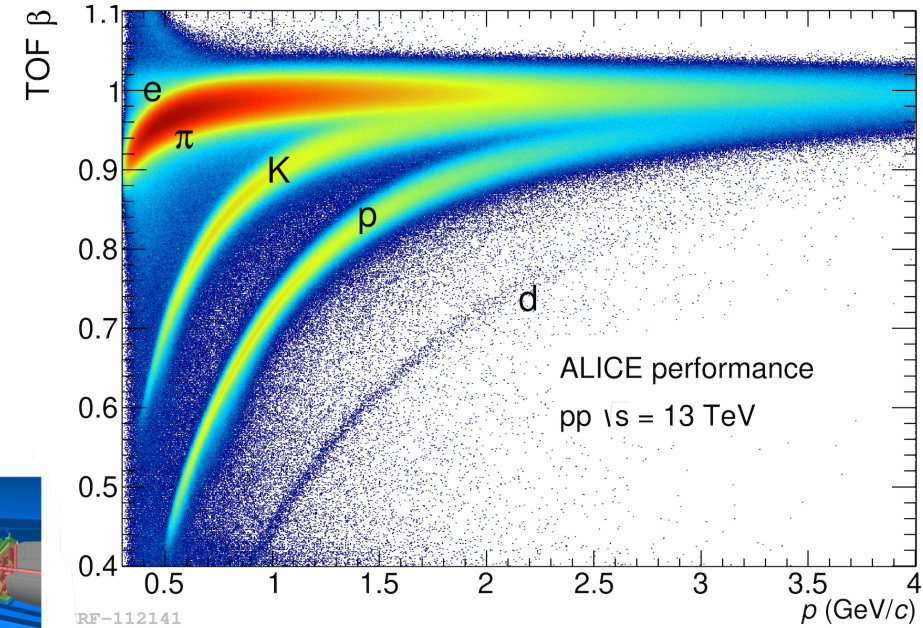
# Identification of nuclei with ALICE

Low  $p$  region (below 1 GeV/c)  $\rightarrow$  PID via  $dE/dx$  measurements in TPC



Time Of Flight

PID via  $\beta$   
 $\sigma_{PID} \sim 70$  ps

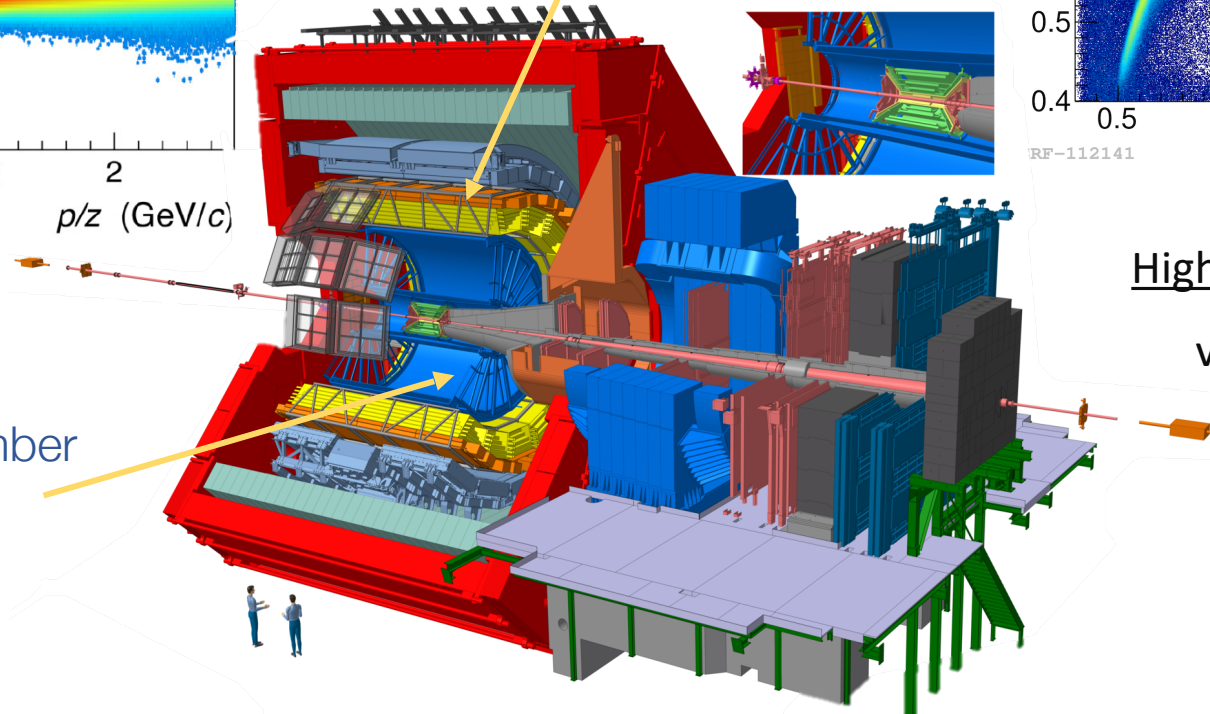


Higher  $p$  region (above 1 GeV/c)  $\rightarrow$  PID via velocity  $\beta$  measurements in TOF

Time Projection Chamber

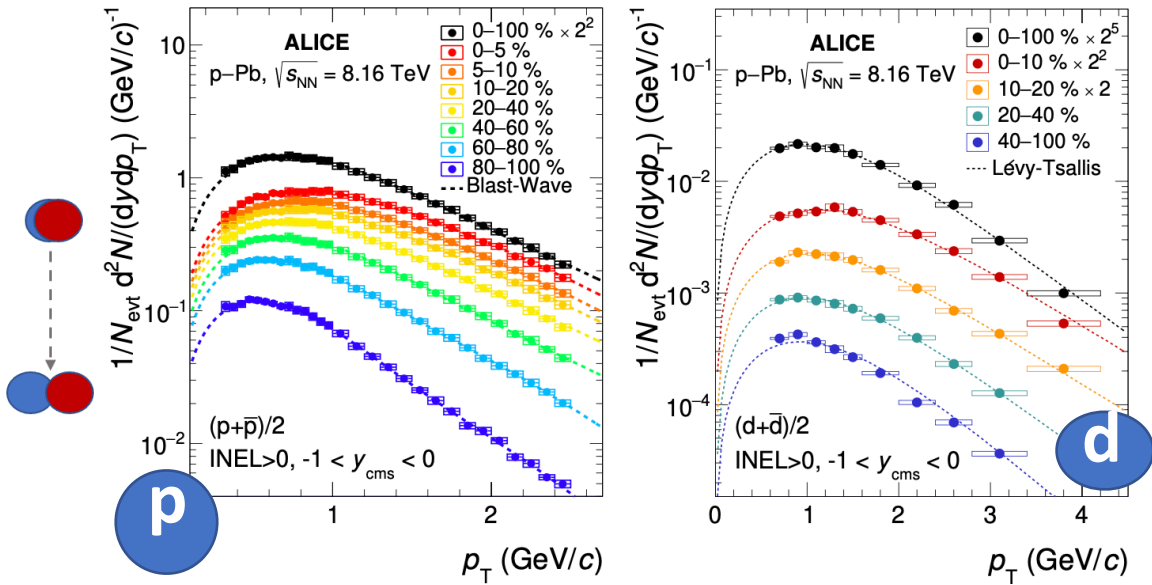
tracking, PID via  $dE/dx$

$\sigma_{dE/dx} \sim 6\%$

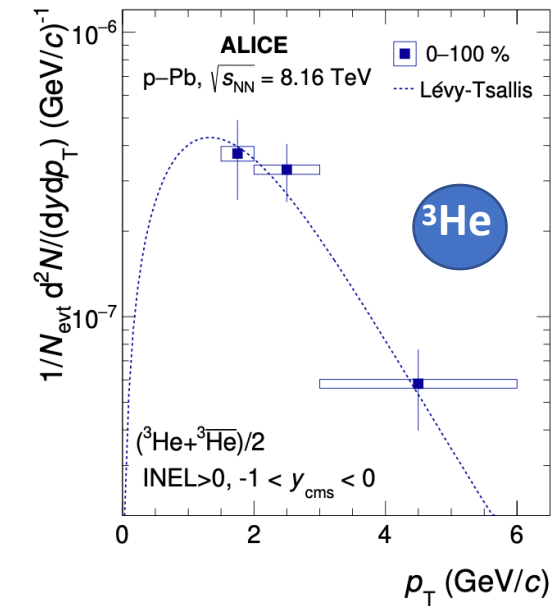




# Measurement of light (anti)nuclei with ALICE

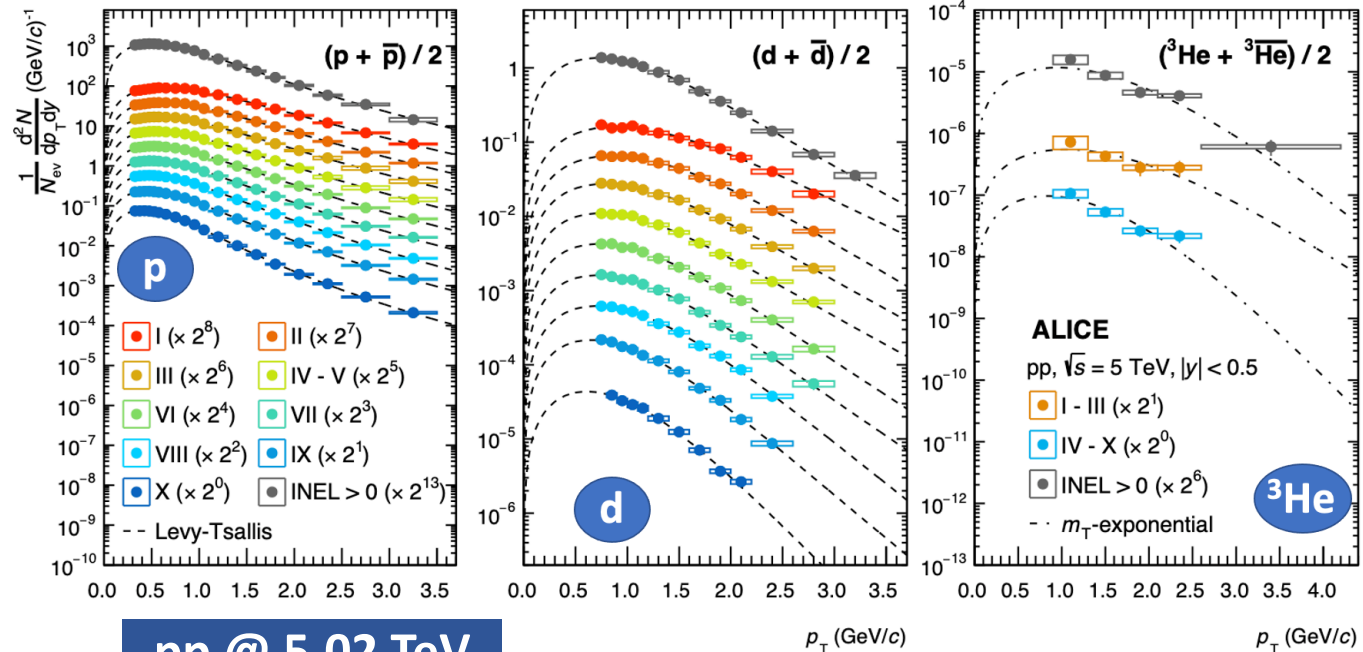


- In small systems such as pp and p-Pb collisions, all nuclei species have been measured, from p to  $^3\text{He}$
- Momentum distributions fitted to extrapolate the yield in the unmeasured regions



**p-Pb @ 8.16 TeV**

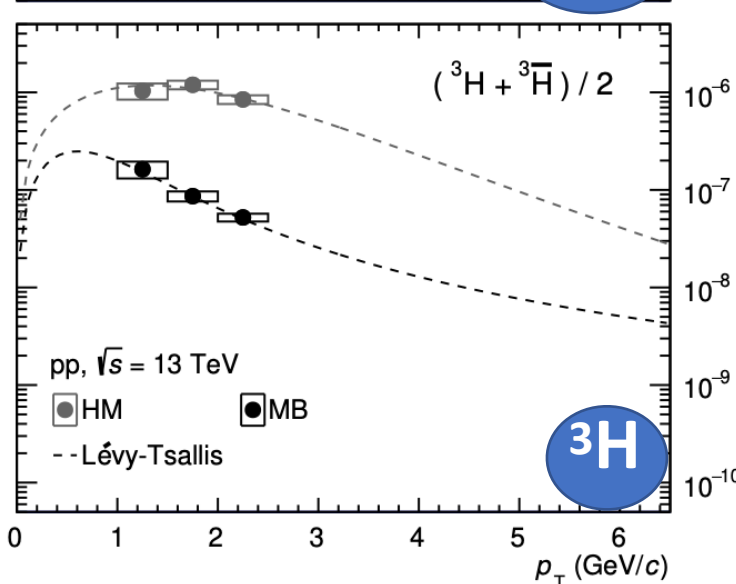
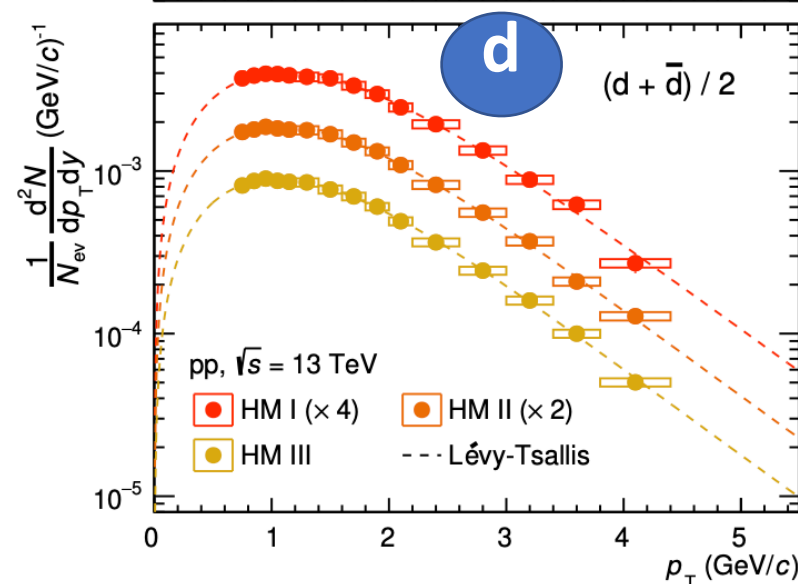
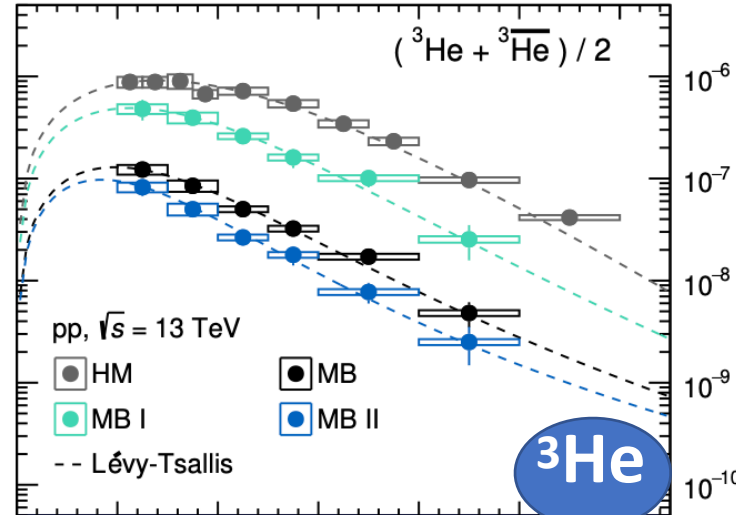
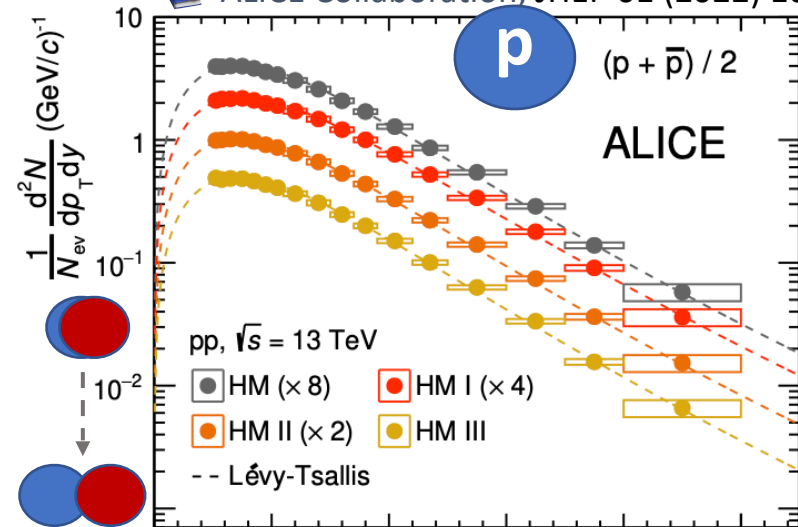
ALICE  
Collaboration, PLB 811  
(2020) 135849



ALICE Collaboration, EPJC (2022) 82:289

# Measurement of light (anti)nuclei with ALICE

ALICE Collaboration, JHEP 01 (2022) 106

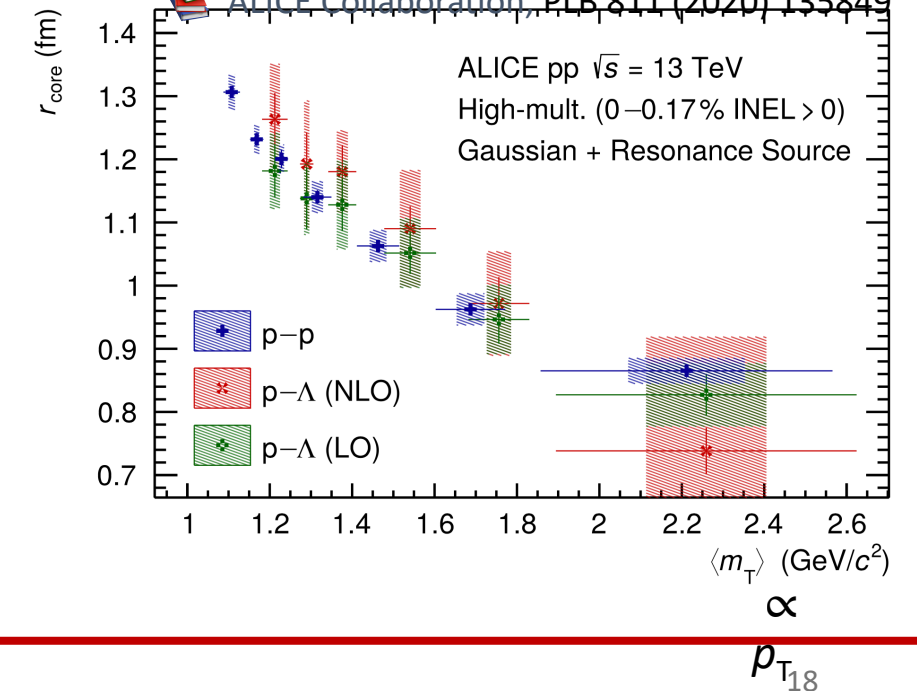


## HM pp @ 13 TeV

- Focus on the **HM data sample**  $\rightarrow$  narrow multiplicity interval covered (0-0.1%)
- Precise measurement of the emission source size  $r_{\text{core}}$  using femtoscopy is available

$\rightarrow$  **crucial to test the coalescence model**

ALICE Collaboration, PLB 811 (2020) 135849



# Coalescence parameter

- Important observable in accelerator measurements: coalescence parameter  $B_A$

$$B_A(p_T^p) = E_A \frac{d^3 N_A}{dp_A^3} / \left( E_p \frac{d^3 N_p}{dp_p^3} \right)^A$$

- Theoretical prediction [1]

$$B_2(\vec{p}) \approx \frac{3}{2m} \int d^3 q D(\vec{q}) e^{-R^2(p_T) q^2}$$

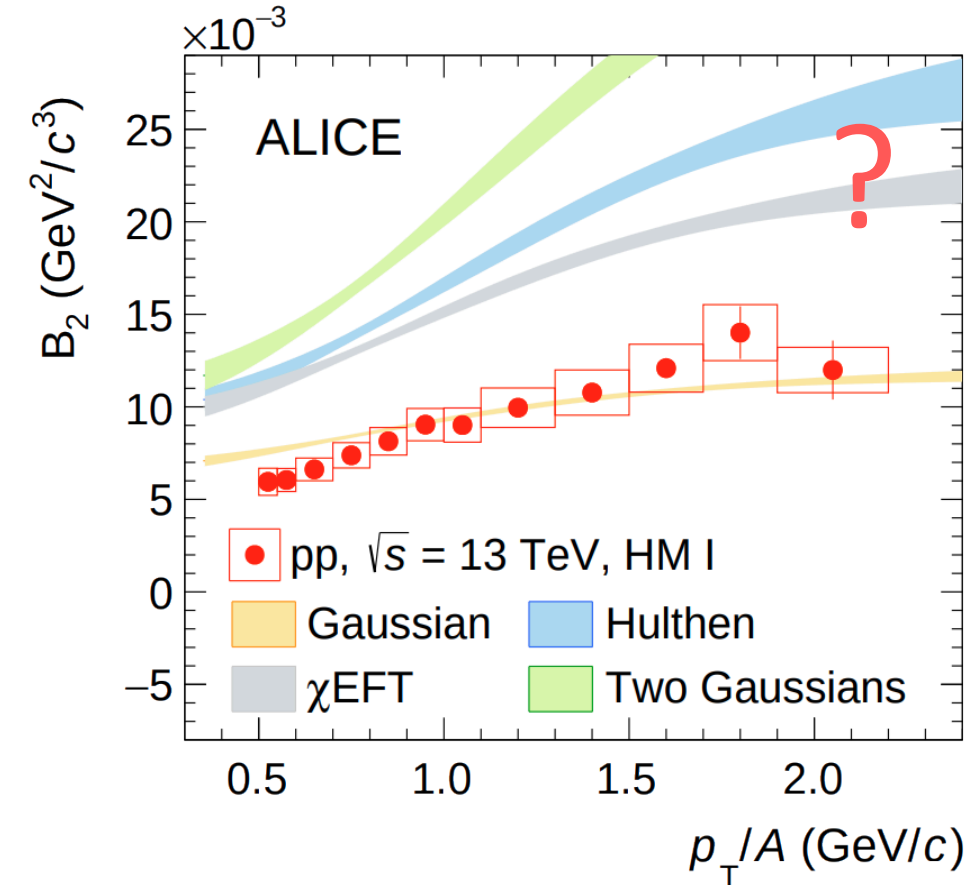
*emission source size*

*deuteron wave function (size  $d = 3.2$  fm)*

$$D(\vec{q}) = \int d^3 r |\phi_d(\vec{r})|^2 e^{-i\vec{q}\cdot\vec{r}}$$

Testing different wave functions:

- **Hulthén:** Favoured by low energy scattering experiments
- **Gaussian:** Best description of currently available ALICE data
- **Two Gaussians:** Approximates Hulthén, easy to use in calculations
- $\chi$ EFT: Favoured by modern nuclear interaction experiments (e.g. Femtoscopy)
- **Argonne v18** (missing here)



ALICE Collaboration, JHEP 01 (2022) 106

<sup>1</sup> Blum, Takimoto, PRC 99 (2019) 044913

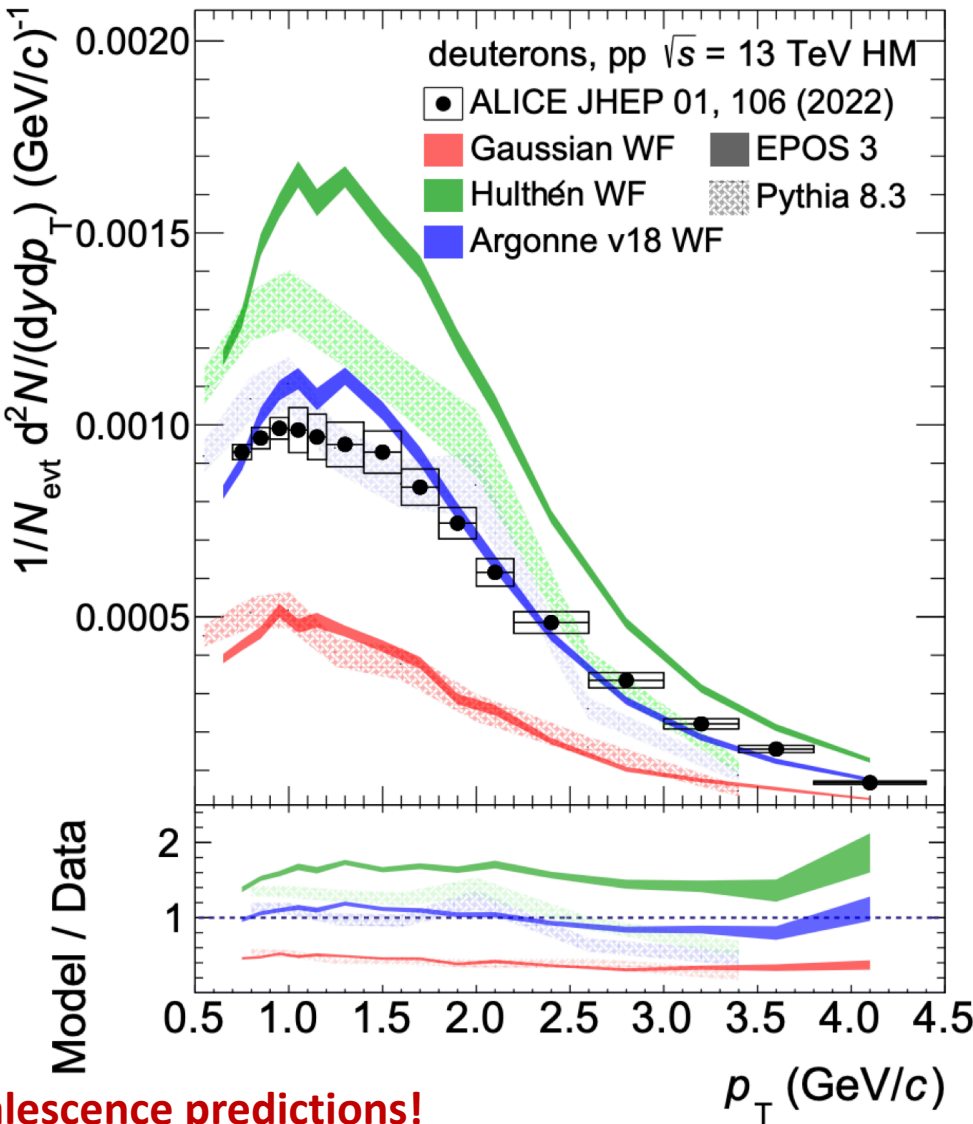
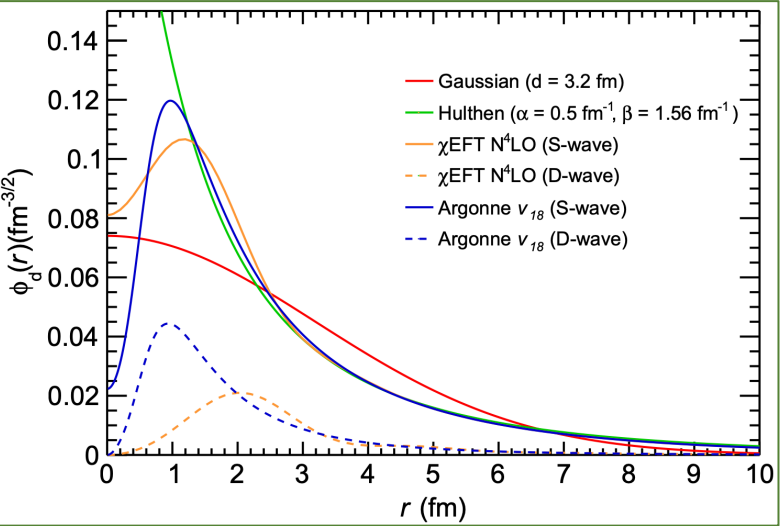
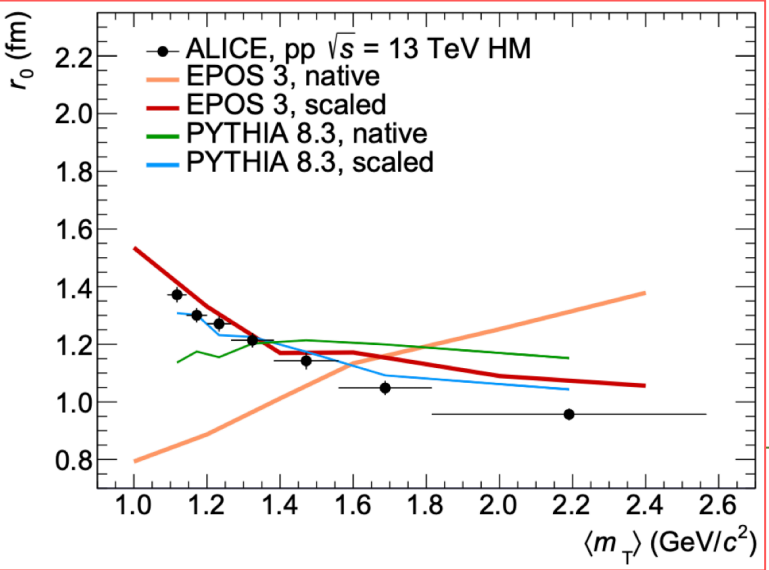
Kachelrieß et al., EPJA 19 (2020) 4

# State-of-the-art coalescence model

Coalescence afterburner based on Wigner function formalism

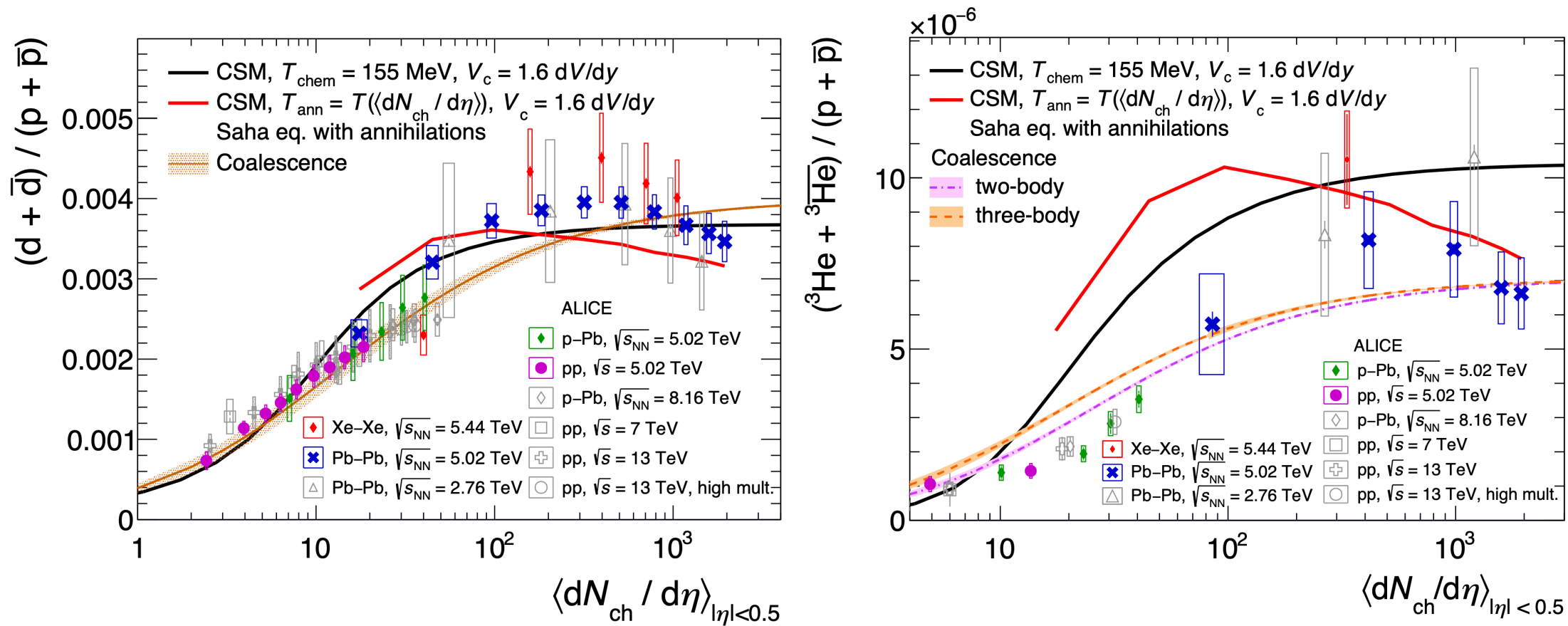
- Use event generators (PYTHIA 8.3 & EPOS 3)
- Emulate experimental multiplicity trigger
- Calibrate (anti)nucleon momentum distribution
- Take resonance cocktail from SHM
- Tune emission source
- Employ realistic wavefunction

→ model predicts deuteron  $p_T$  spectrum



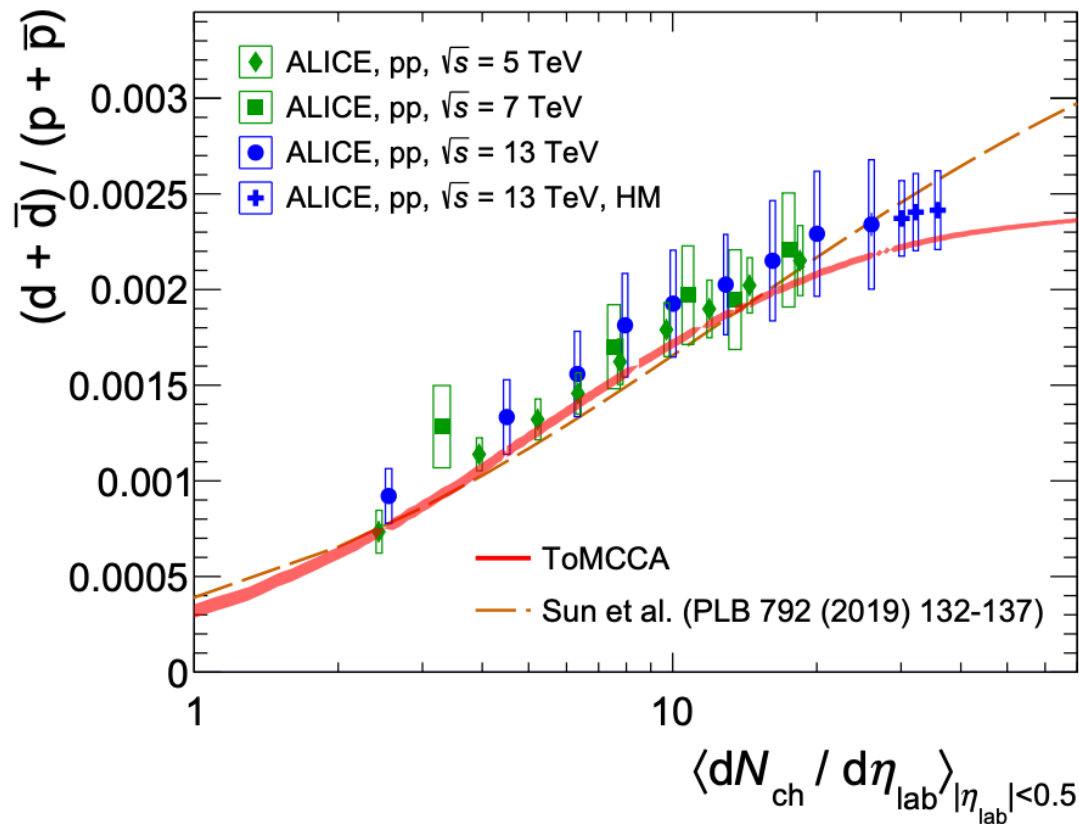
**Realistic wavefunction is key for coalescence predictions!**





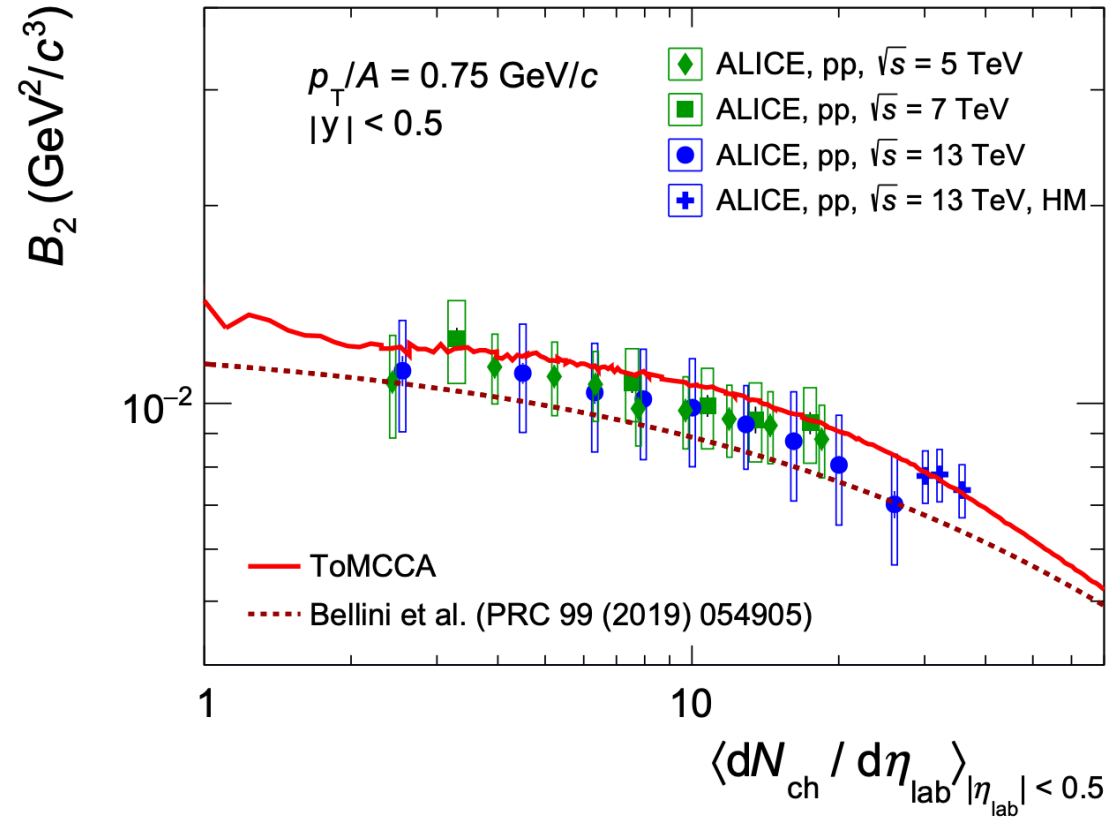
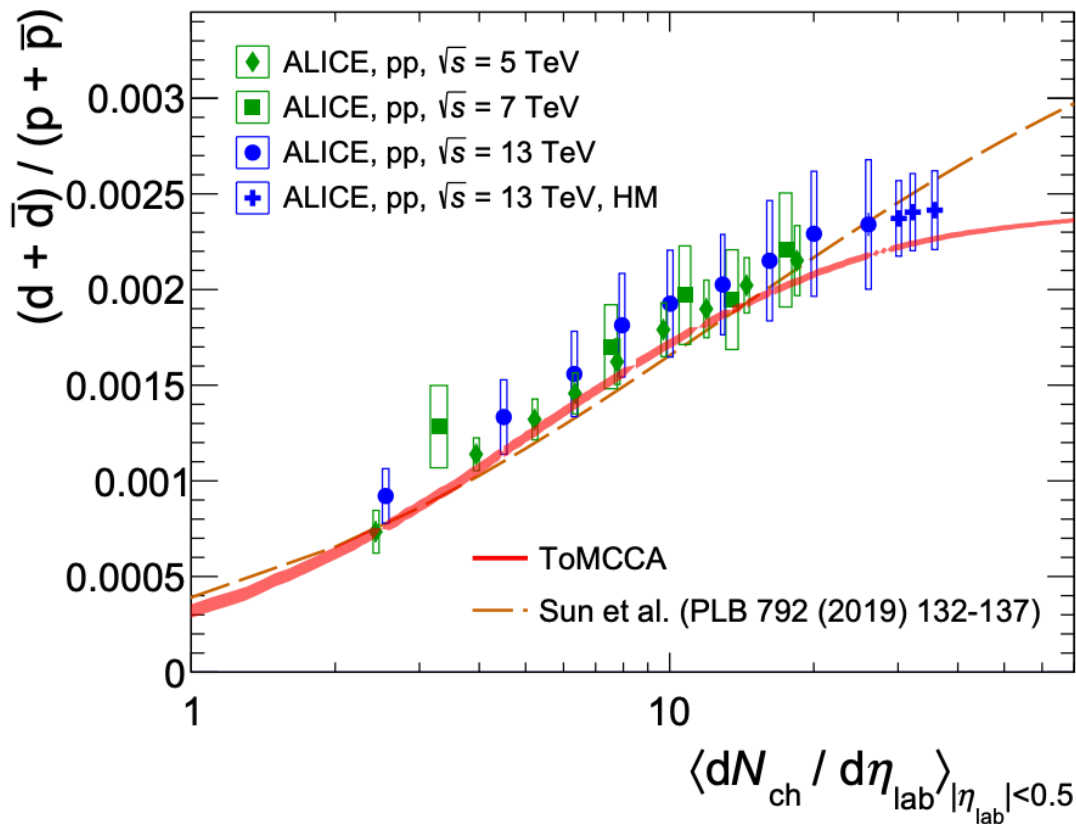
- $V_c=1.6$  dV/dy is the correlation volume needed to describe the net-deuteron number fluctuations in Pb–Pb collisions<sup>1</sup>
- CSM  $\rightarrow$  either with fixed chemical temperature (black) or with **annihilation** temperature depending on multiplicity<sup>2</sup> (red)
- Both CSM and **coalescence**<sup>3</sup> predictions qualitatively reproduce the trend and overall yields, but neither of the models

calculated data points, PRL 131 (2023) 041901 <sup>2</sup> Vovchenko, Koch, PLB 835, 137577 (2022) <sup>3</sup> Sun, Ko, Doenigus, PLB 792 (2019) 132-137



Predictions available only for the pp multiplicity range (1-70)

- **Coalescence** predictions of ToMCCA using Wigner function formalism & multiplicity-dependent input (momentum distributions of nucleons, source size and multiplicity distributions) reproduce all data points within 1sigma
- No  $^3\text{He}$  coalescence predictions yet



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- No  $^3\text{He}$  coalescence predictions yet
- Also coalescence parameter  $B_2$  vs multiplicity is well reproduced by ToMCCA

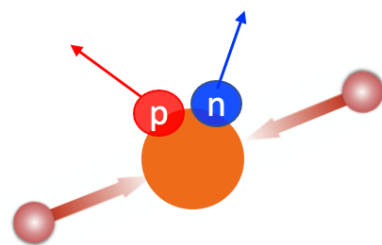
More about ToMCCA during Maxi's talk!

Mahlein, Pinto, Fabbietti, arXiv:2404.03352

# Testing production models with hypertriton

- In small collision systems (as pp) size of system created in the collision is smaller or equal to that of the nucleus under study
- For small systems model predictions are quite different
- Coalescence is sensitive to the **interplay** between the **size of the collision system** and the spatial extension of the **nucleus wave function**

$\Lambda$  ratio provides a powerful tool to investigate nuclear production mechanism



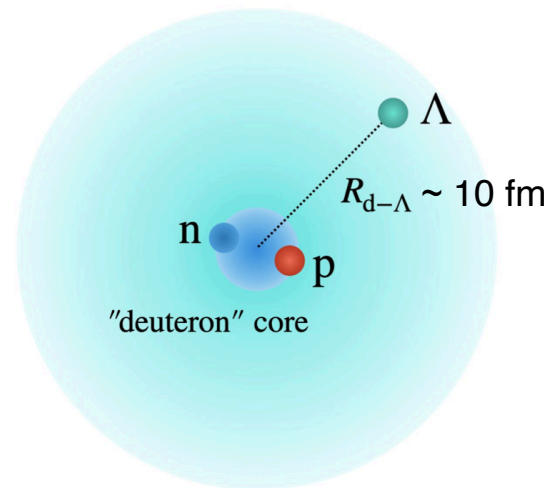
System size (pp, p—Pb): 1–1.5 fm

$r_d$ : 1.96 fm

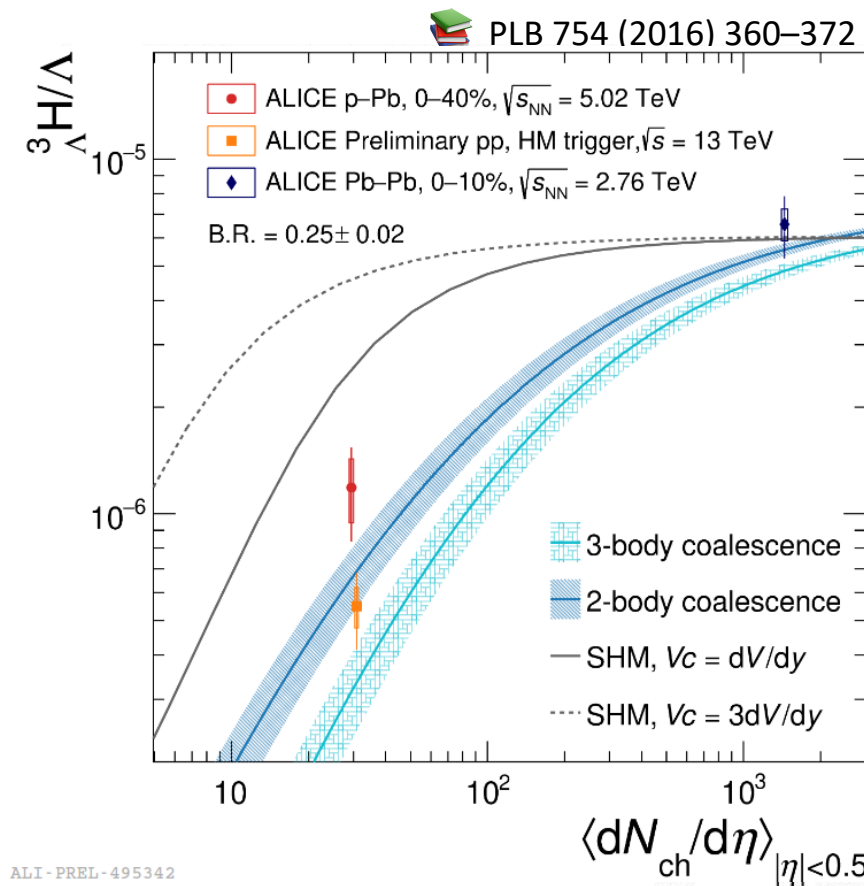
$r_{3\text{He}}$ : 1.76 fm

$r_{(np\Lambda)}$ : 4.9 fm ( $B_\Lambda = 2.35$  MeV)

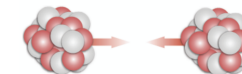
$r_{(d\Lambda)}$ : 10 fm ( $B_\Lambda \sim 0.13$  MeV)



powerful probe for investigating the nucleon –  $\Lambda$  interaction

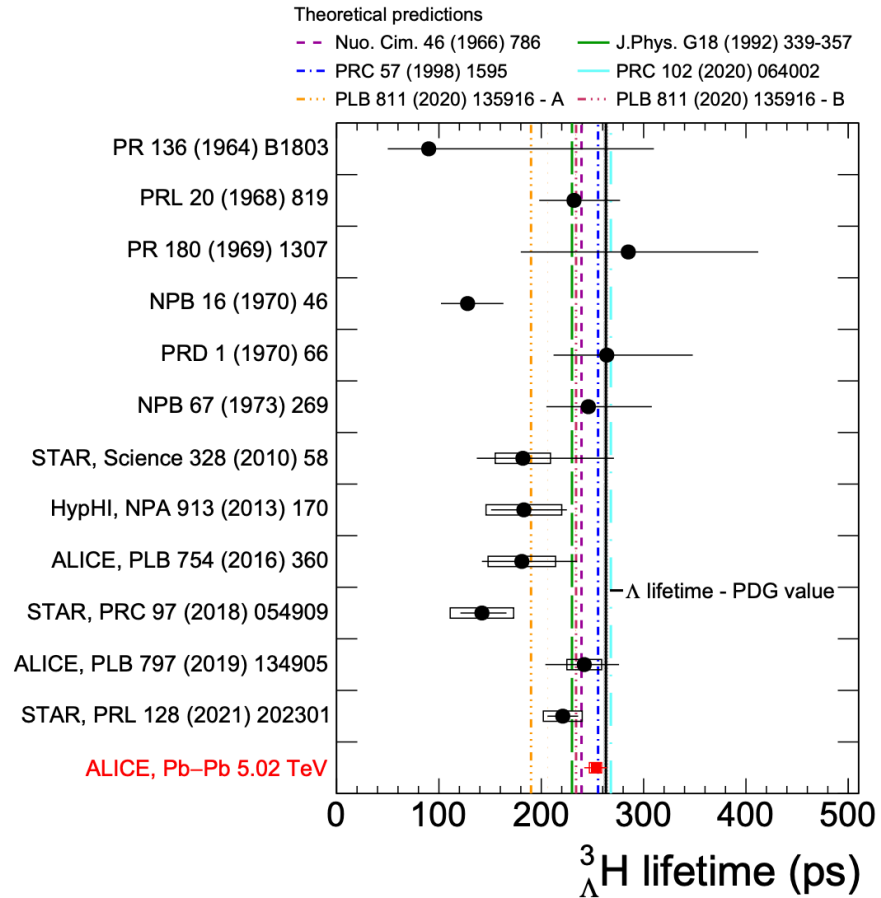


ALI-PREL-495342

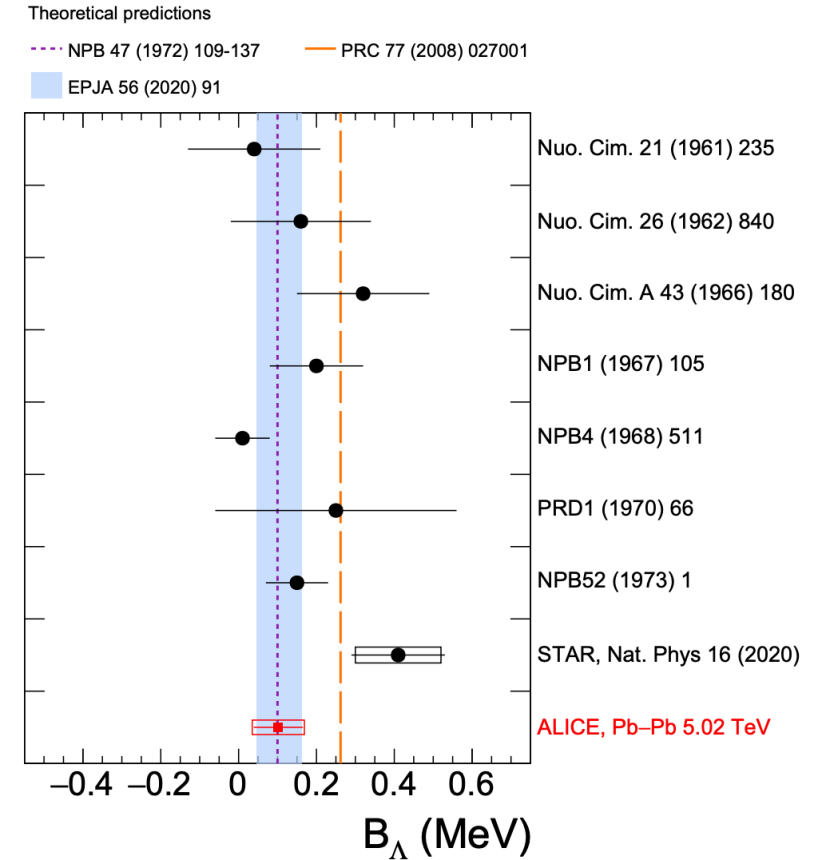


# Hypertriton lifetime & binding energy (Pb–Pb collisions)

$$\tau = [253 \pm 11 \text{ (stat.)} \pm 6 \text{ (syst.)}] \text{ ps}$$



$$B_{\Lambda} = [102 \pm 63 \text{ (stat.)} \pm 67 \text{ (syst.)}] \text{ keV}$$



- Models predicting a lifetime close to the free  $\Lambda$  one are favoured
- Strong hint that hypertriton is weakly bound

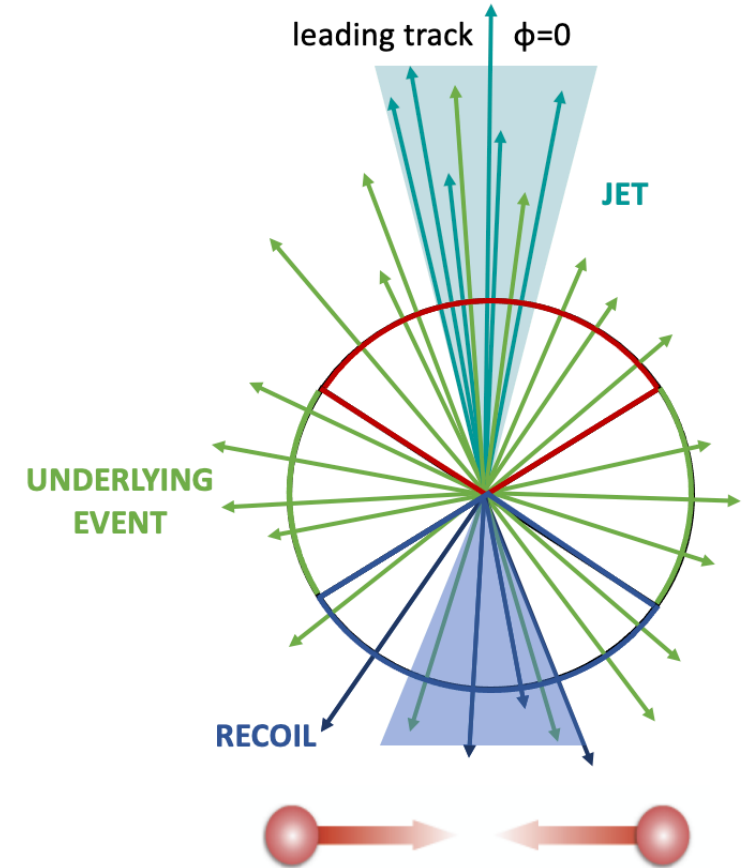
- $B_{\Lambda}$  compatible with zero  $\rightarrow$  Weakly bound nature of  ${}^3_{\Lambda}\text{H}$  is confirmed

Phys. Rev. Lett. 131 (2023) 102302

# Nuclear production in and out of jets

- Powerful tool to investigate coalescence mechanism is the study of nuclear production in and out of jets
- In jets nucleons have strong phase-space constraint

→ **Study  $B_2$  in and out of jets:** jets obtained simply by subtracting the **UE** from the **Toward** region (**Jet** + **UE**)



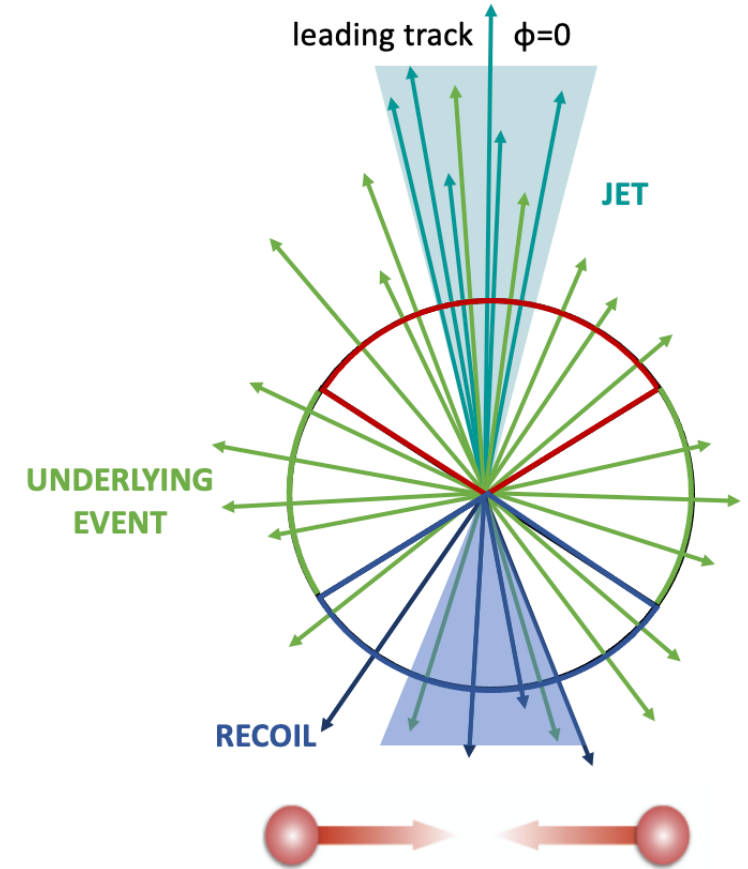
**Toward:**  $|\Delta\phi| < 60^\circ$

**Transverse:**  $60^\circ < |\Delta\phi| < 120^\circ$

**Away:**  $|\Delta\phi| > 120^\circ$

# Nuclear production in and out of jets

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- In jets nucleons have strong phase-space constraint
  - **Study  $B_2$  in and out of jets:** jets obtained simply by subtracting the **UE** from the **Toward** region (**Jet** + **UE**)
- Studying the antideuteron production in jets in small systems (pp, pA) is important to understand and model nuclear production
- Implications for cosmic ray physics
- Antideuteron in the Galaxy is produced in interactions of cosmic rays (p,  $^4\text{He}$ ) with kinetic energies of  $\sim 300$  GeV



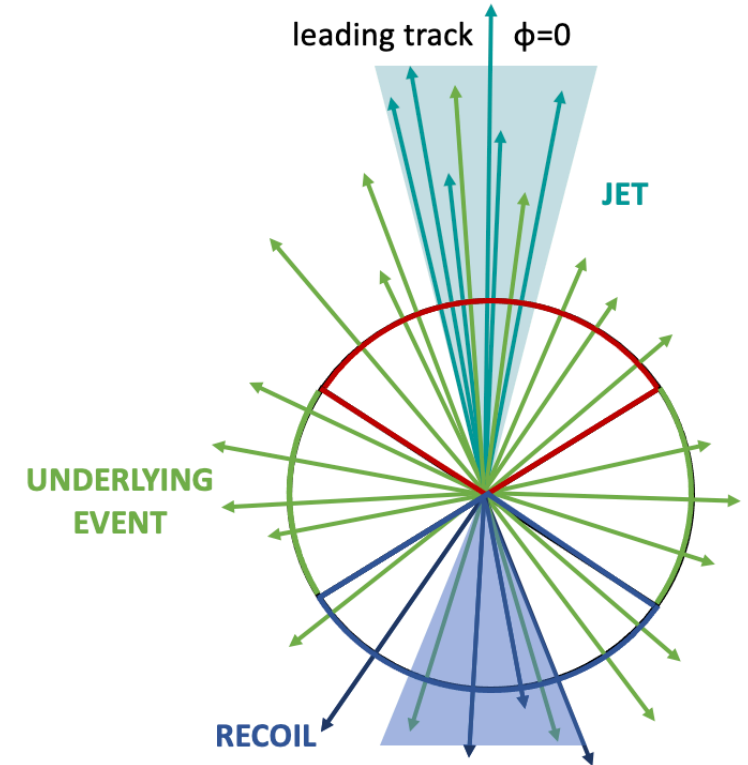
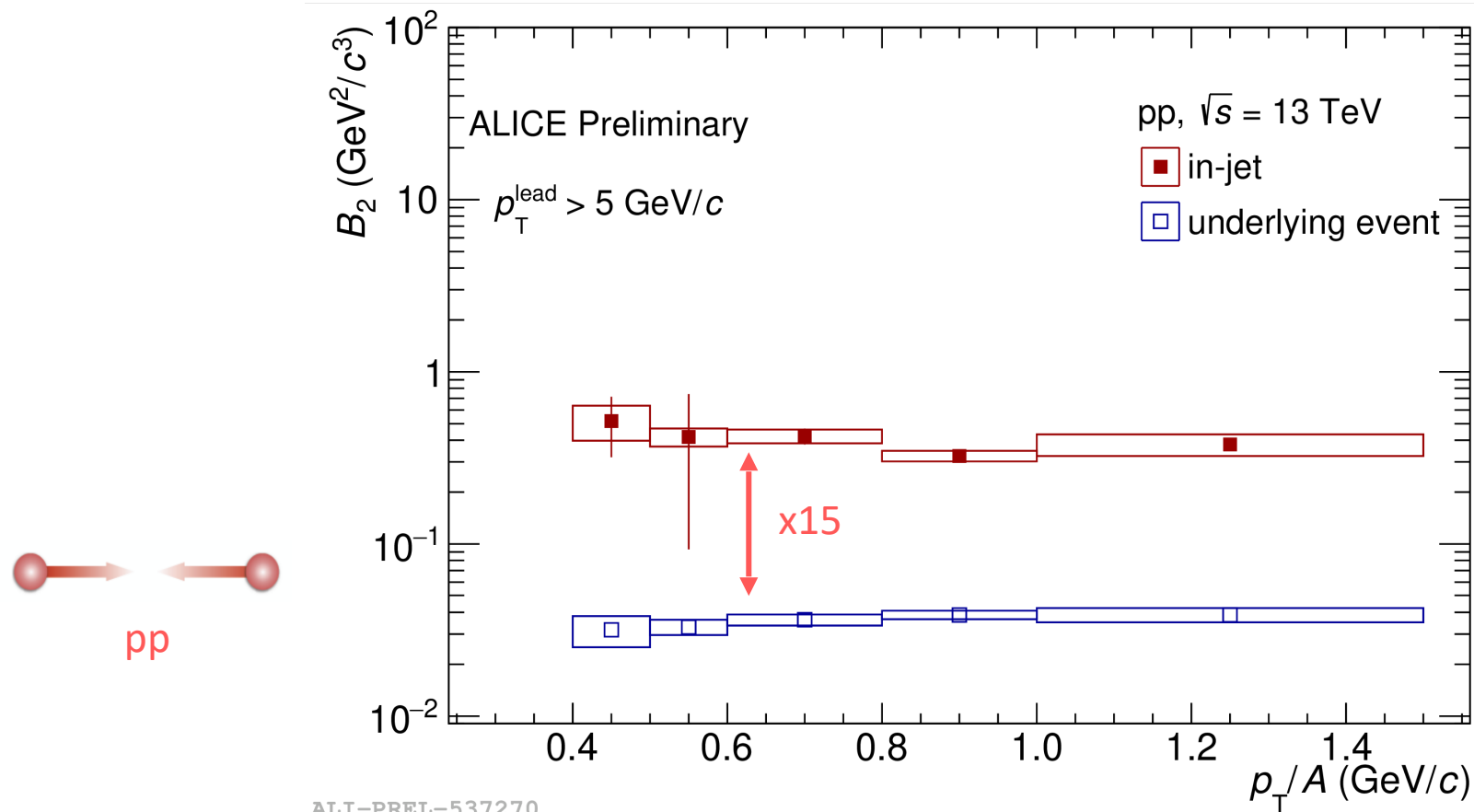
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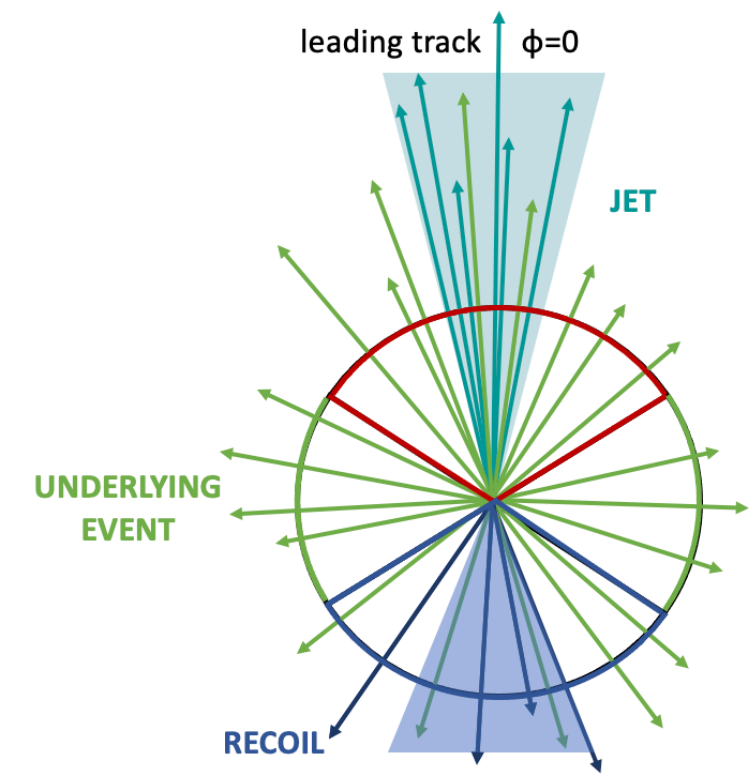
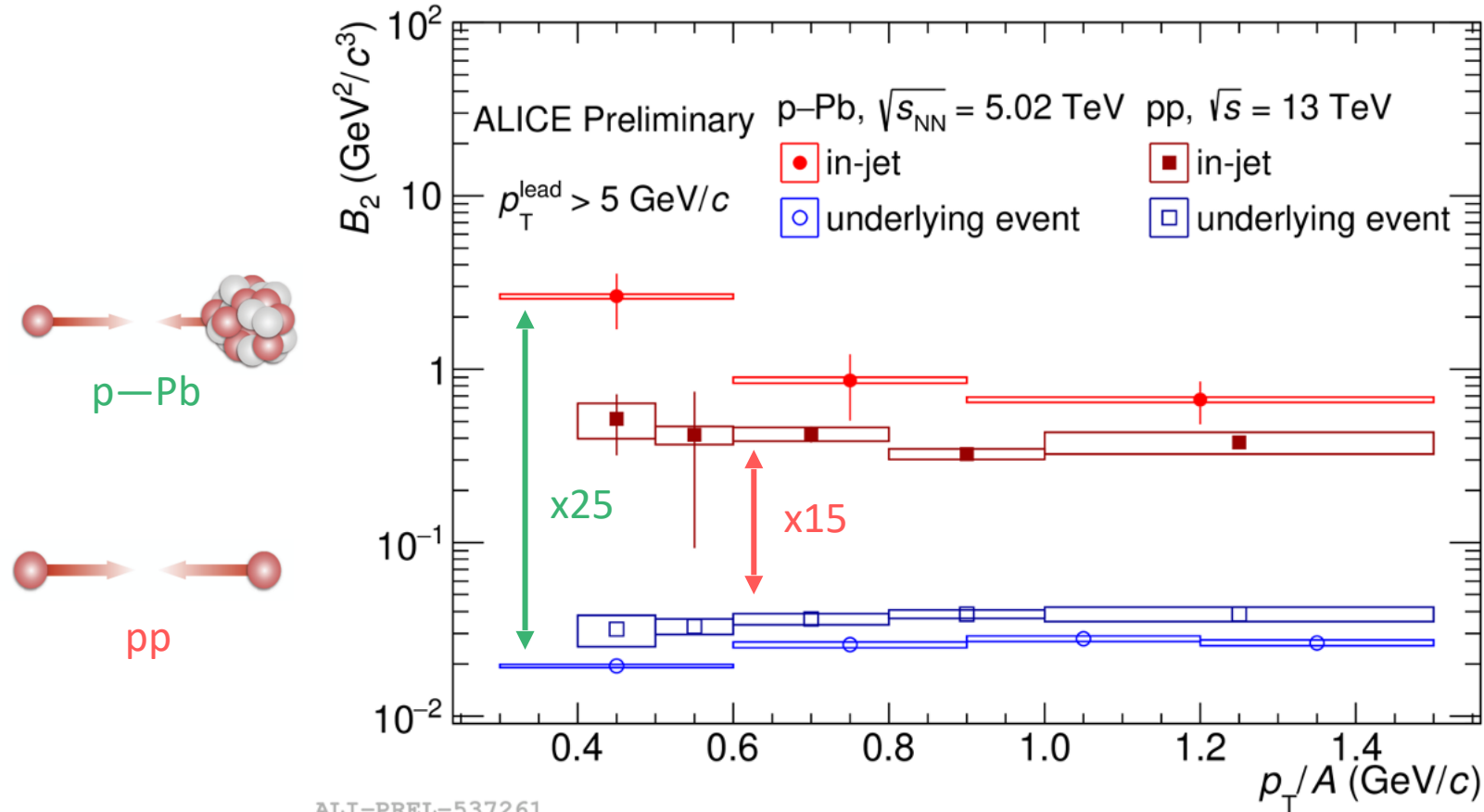
# Coalescence parameters in and out of jets



- *Enhanced deuteron coalescence probability in jets wrt UE is observed for the first time in pp collisions*
- Due to the reduced distance in phase space of hadrons in jets compared to those out of jets → favors coalescence picture

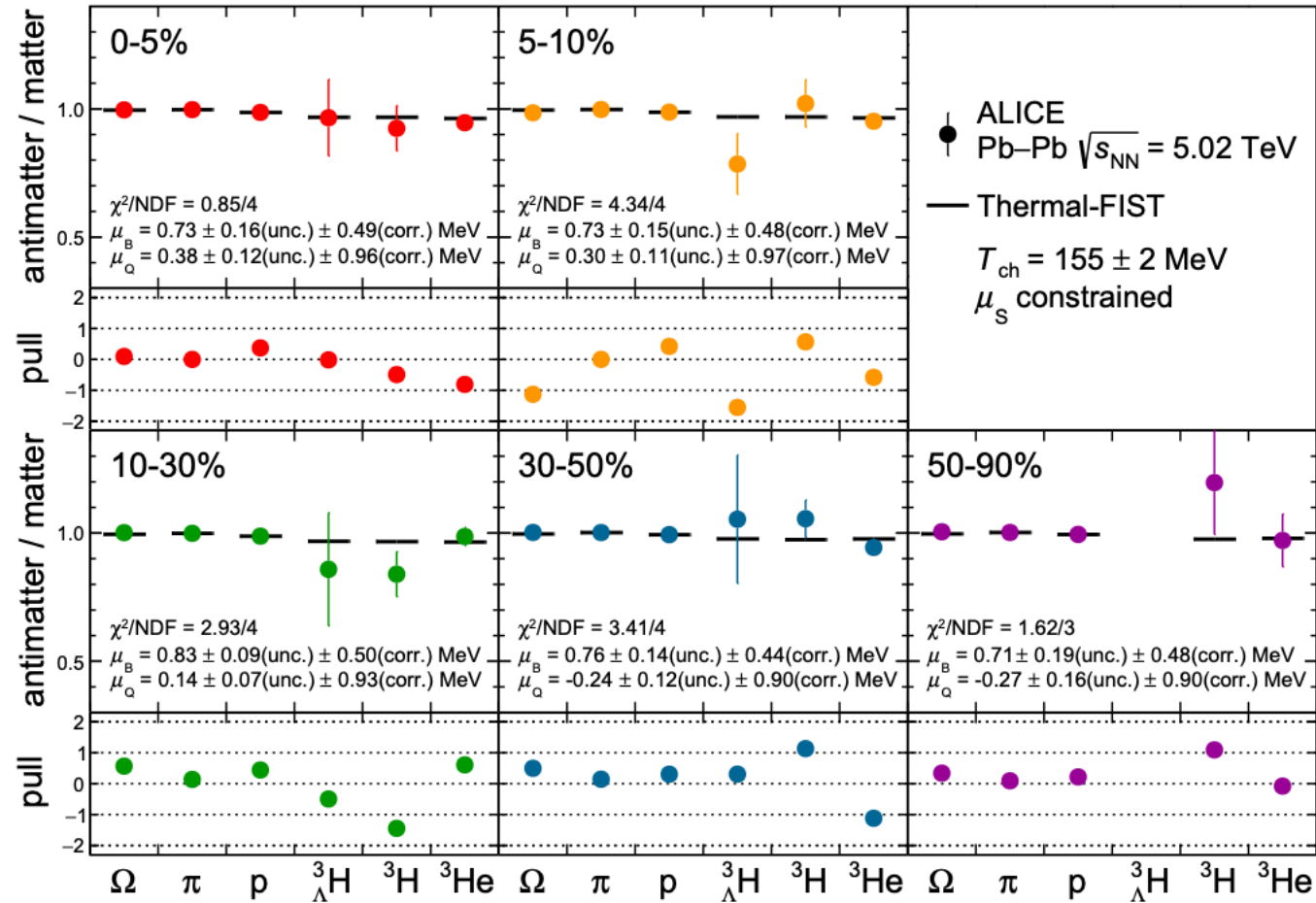


# Coalescence parameters in and out of jets

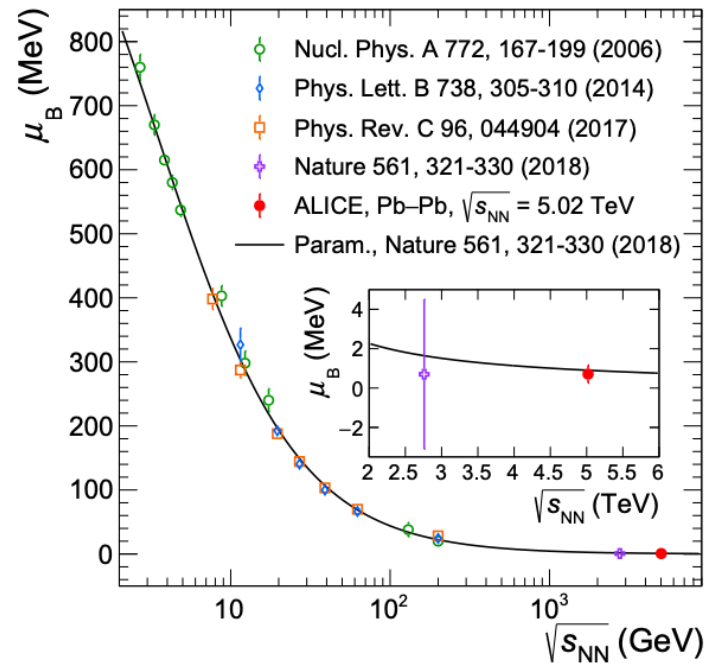
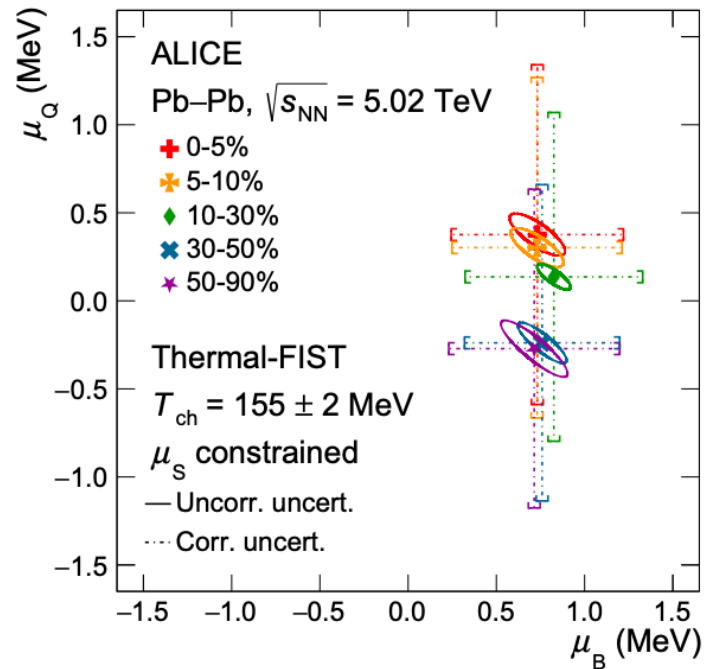
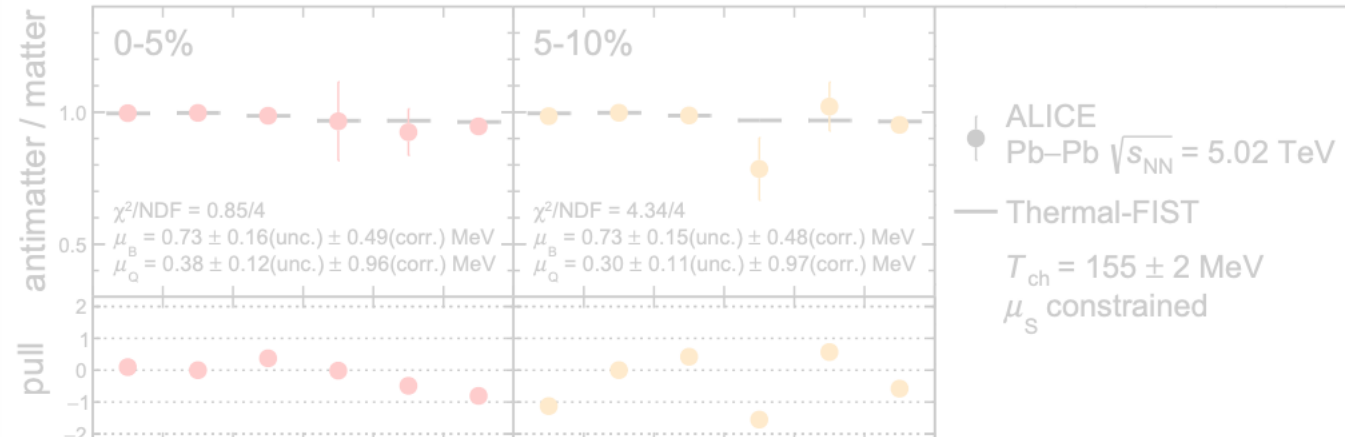


- $B_2$  in-jet in **p—Pb** is larger than  $B_2$  in-jet in **pp**  
 → *could be* related to the different particle composition of jets in pp and p—Pb → *to be further investigated*
- $B_2$  in UE in **p—Pb** is smaller than  $B_2$  in UE in **pp** due to the larger source size in p—Pb

1 Phys.Rev.C 99 (2019) 024001  
 2 Phys.Rev.Lett. 123 (2019) 112002  
 3 Phys.Rev.Lett. 131 (2023) 04, 042301



- $\mu_B$  and  $\mu_Q$  are extracted fitting the antiparticle-to-particle yield ratios with the predictions of the grand-canonical SHM using the Thermal-FIST code



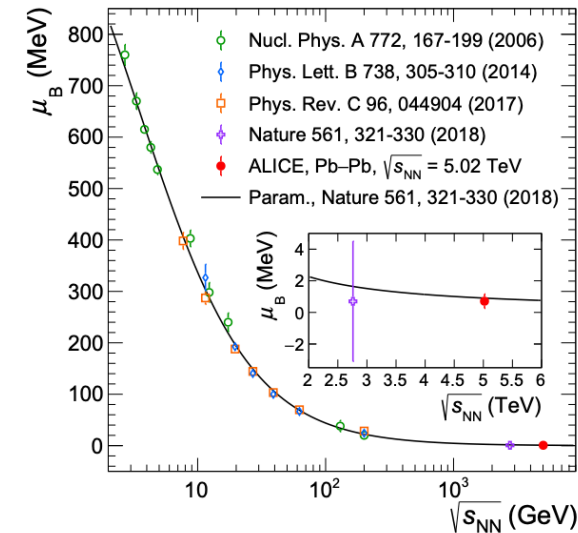
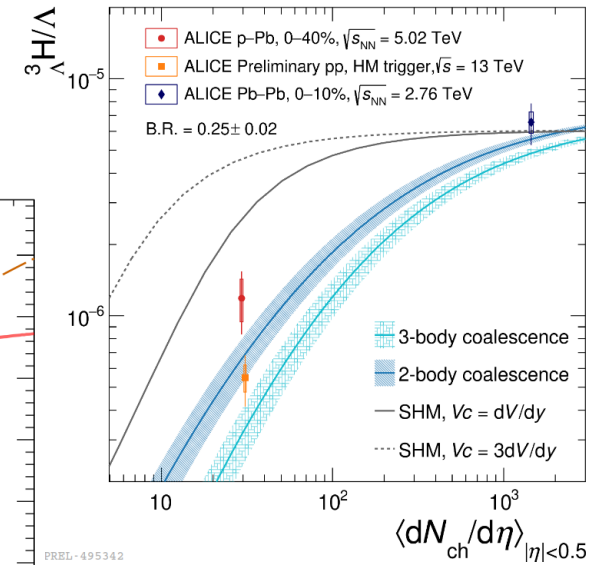
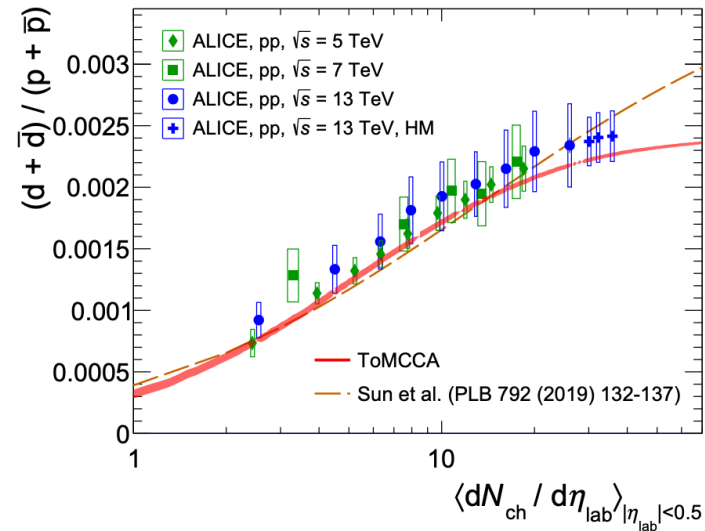
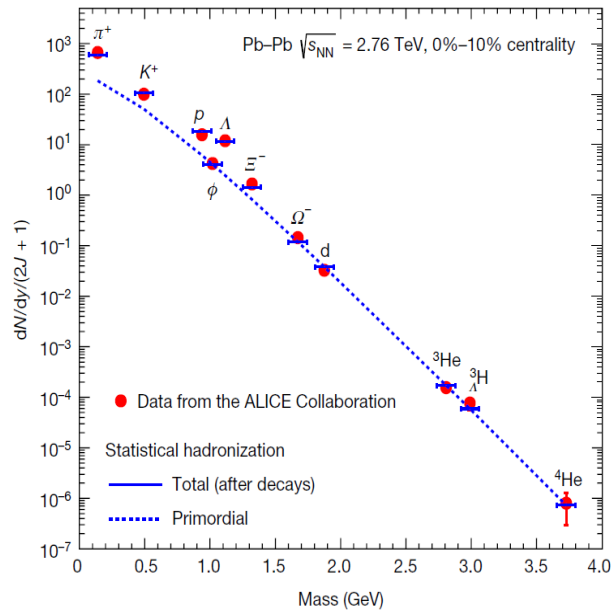
- $\mu_B$  and  $\mu_Q$  are extracted fitting the antiparticle-to-particle yield ratios with the predictions of the grand-canonical SHM using the Thermal-FIST code
- $\mu_Q = -0.18 \pm 0.90$  MeV
- $\mu_B = 0.71 \pm 0.45$  MeV (~8 times more precise than previous measurement)

- **Nuclear transparency regime** is reached ( $\rightarrow$  baryon transport from the colliding ions to the interaction region is negligible)
- **No centrality dependence**  $\rightarrow$  nuclear transparency also in central Pb-Pb (despite  $\mu_B > 0$  could be expected from a more significant baryon number transport at midrapidity)

**The system created in Pb-Pb collisions at the LHC is on average baryon-free and electrically neutral at midrapidity  $\rightarrow$  approaching the early Universe more than any other experimental facility**

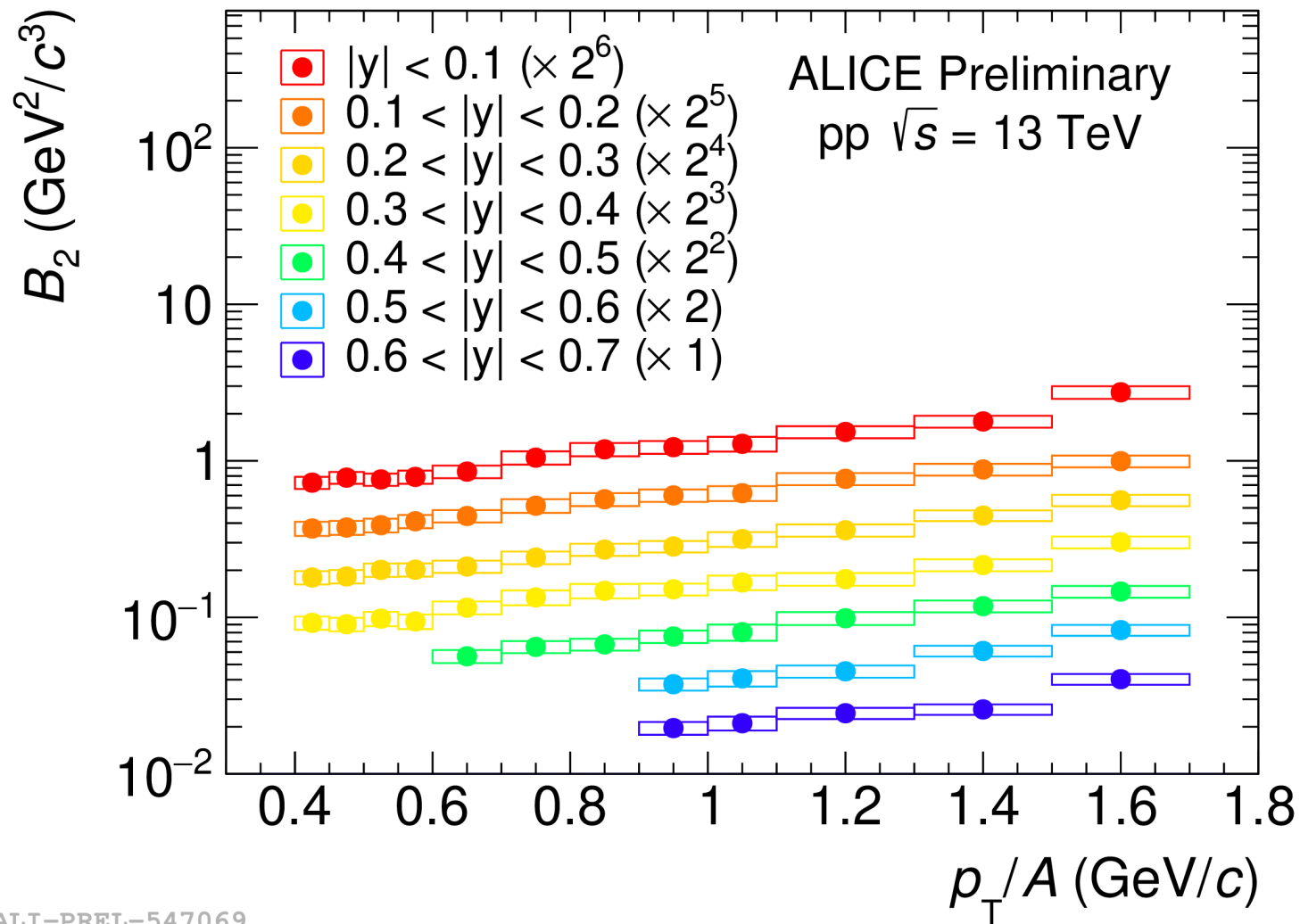
## Production mechanism still not understood

- State-of-the-art coalescence model using Wigner function formalism describes  $d/p$  and  $B_2$  vs multiplicity
- Hypertriton measurements in small systems favor coalescence
- SHM describes the integrated yields of all particles, from pions to hypernuclei
- Using SHM the chemical potentials are calculated showing that nuclear transparency regime is reached at the LHC



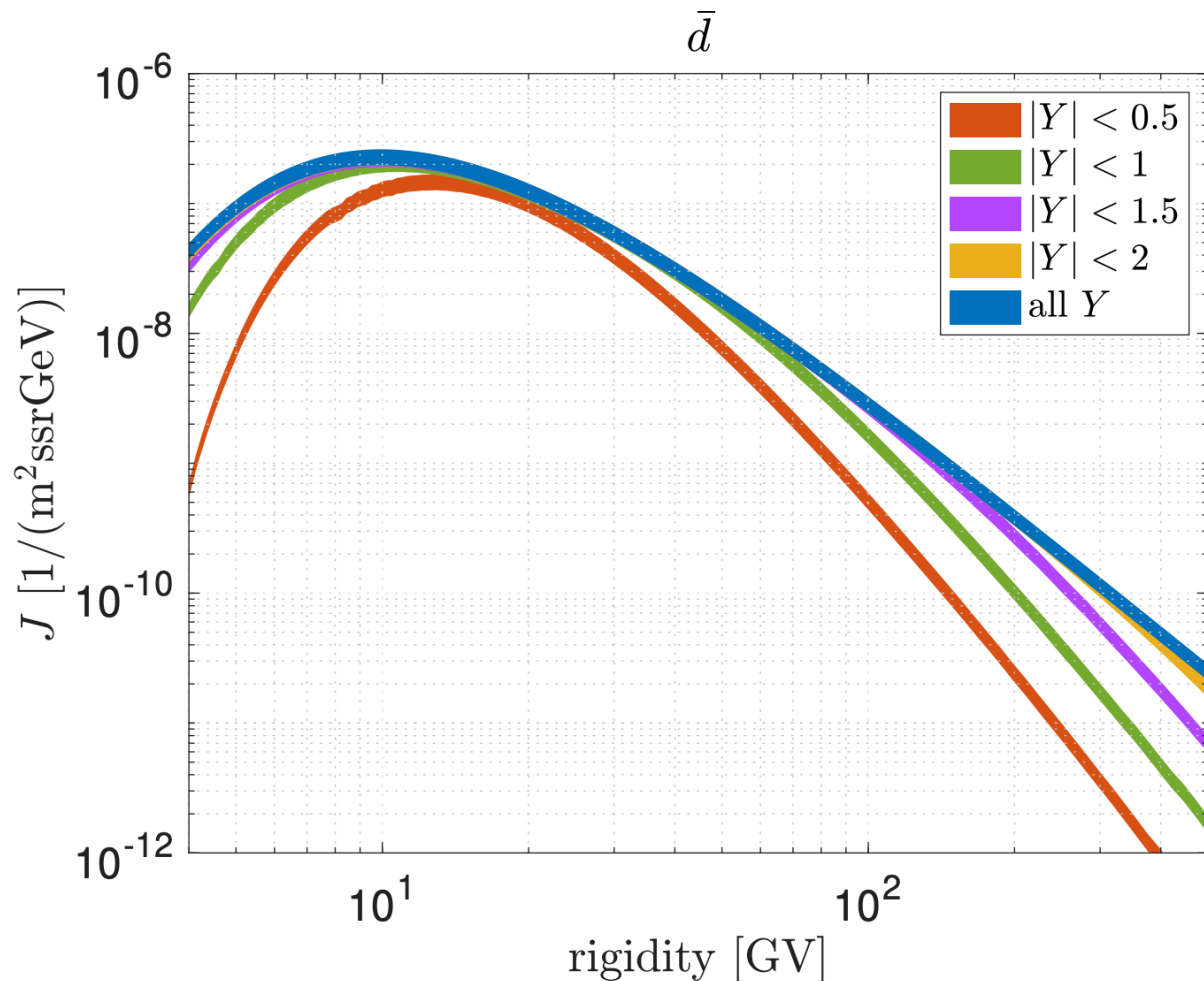


# Coalescence parameter vs. rapidity



- ALICE measurements cover the midrapidity region ( $|y| < 0.5$ ), while astrophysical models extrapolate to forward region
- Current acceptance of ALICE detector allows us to extend the measurement of antinuclei up to  $y = 0.7$
- Rapidity and  $p_T$  dependence of  $B_2$  is extrapolated to forward rapidity using coalescence model + Pythia 8.3 and EPOS as event generators

ALI-PREL-547069



- Model predictions based on ALICE measurements are used as input to calculate antideuteron flux from cosmic rays\*  $\rightarrow$  dominant background in dark matter searches
- Most of the antideuteron yield from  $|y| < 1.5 \rightarrow$  well in reach with future ALICE3<sup>(1)</sup> detector acceptance ( $|y| \lesssim 4$ ) and current LHCb

Production models needed in astrophysics  
 $\rightarrow$  Rapidity coverage is in reach of accelerator experiments  
 $\rightarrow$  Extrapolation to lower energies ( $\sim \text{GeV}$ ) is needed

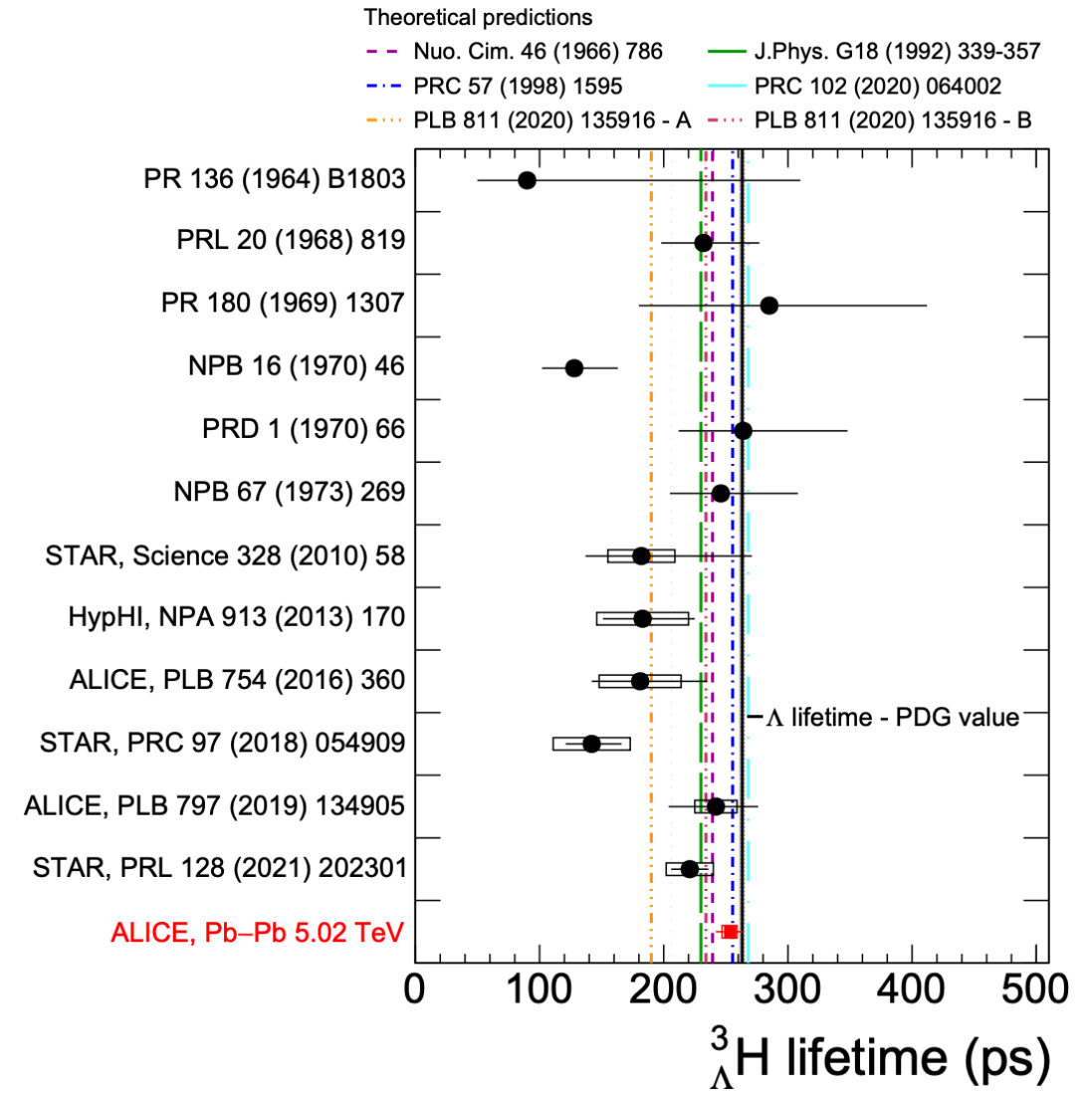
K. Blum, arXiv:2306.13165

\* K. Blum, Phys.Rev.D 96 (2017) 10, 103021

$$\tau = [253 \pm 11 \text{ (stat.)} \pm 6 \text{ (syst.)}] \text{ ps}$$

- **Most precise measurement**
- Compatible with latest **ALICE** and **STAR** measurements
- Models predicting a lifetime close to the free  $\Lambda$  one are favoured
- Strong hint that hypertriton is weakly bound, but  $B_\Lambda$  is still needed to solve the puzzle

**$\geq 2020$  models:** assuming  $B_\Lambda = 70$  keV  
 **$< 2020$  models:** assuming  $B_\Lambda = 130$  keV





- From the mass measurement to  $B_\Lambda$

$$B_\Lambda = M_\Delta + M_d - M_{^3_\Lambda\text{H}}$$

- Weakly bound nature of  $^3_\Lambda\text{H}$  is confirmed by the latest ALICE measurement


- $B_\Lambda$  compatible with zero
- in agreement within  $1\sigma$  with Dalitz and  $\chi\text{EFT}$ -based predictions
- fully consistent with the lifetime measurement according to recent theoretical calculations [1, 2]

$$r_{(np\Lambda)}: 4.9 \text{ fm } (B_\Lambda = 2.35 \text{ MeV})$$

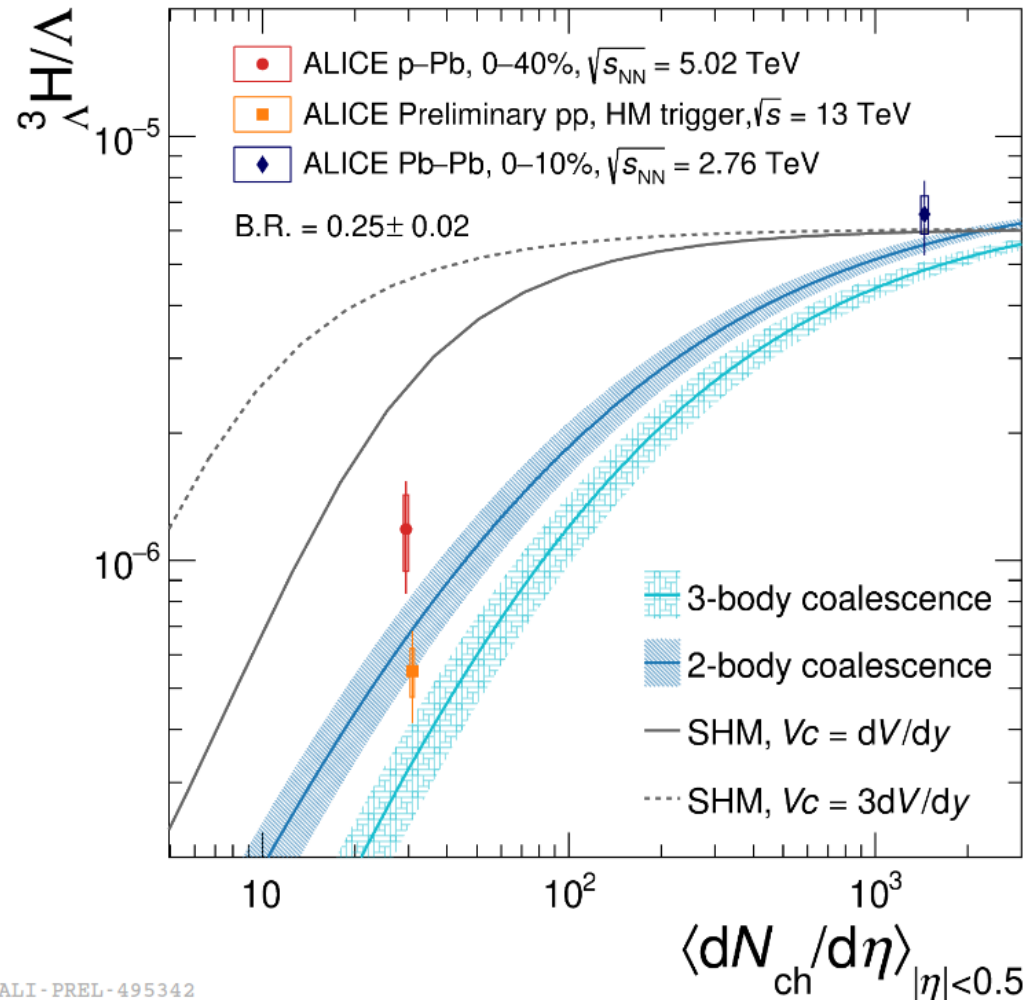
$$r_{(d\Lambda)}: \sim 10 \text{ fm } (B_\Lambda \sim 0.13 \text{ MeV})$$

$$B_\Lambda = [102 \pm 63 \text{ (stat.)} \pm 67 \text{ (syst.)}] \text{ keV}$$

$$r_{^3_\Lambda\text{H}(np\Lambda)}: 4.9 \text{ fm } (B_\Lambda = 2.35 \text{ MeV})$$
$$r_{^3_\Lambda\text{H}(d\Lambda)}: \sim 10 \text{ fm } (B_\Lambda \sim 0.13 \text{ MeV})$$

 <sup>1</sup> Hildenbrand et al., PRC 102 (2020) 6

 <sup>2</sup> Pérez-Obiol et al., PLB (2020) 811



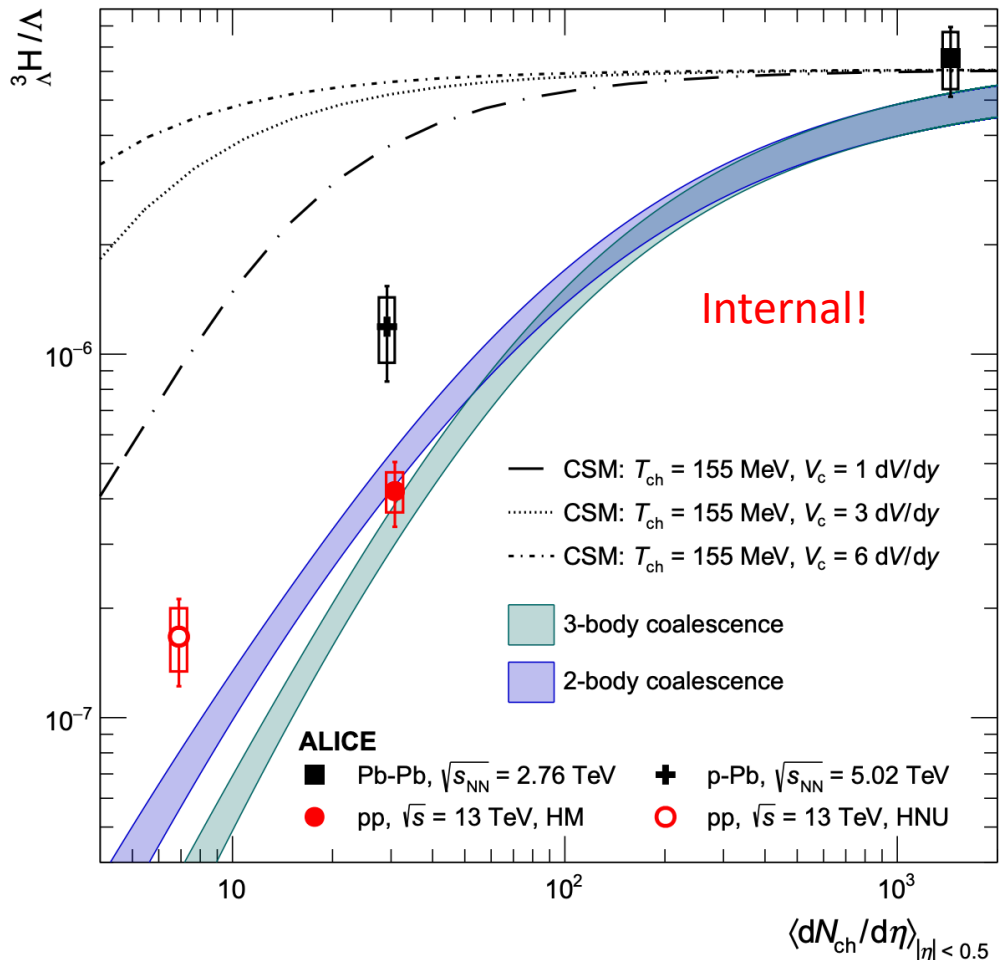
- **Pb—Pb collisions:**
  - small difference between the predictions from SHM and coalescence
- **pp and p—Pb collisions:**
  - large separation between production models
  - **measurements are in good agreement with 2-body coalescence**
  - tension with SHM at low charged-particle multiplicity density
  - configuration with  $V_c = 3dV/dy$  is excluded by more than  $6\sigma$

Coalescence quantitatively describes the suppression in small systems  
 > the nuclear size matters at low charged-particle multiplicity

ALI-PREL-495342

p—Pb: PRL 128 (2022) 25, 252003

Pb—Pb: PLB 754 (2016) 360-372



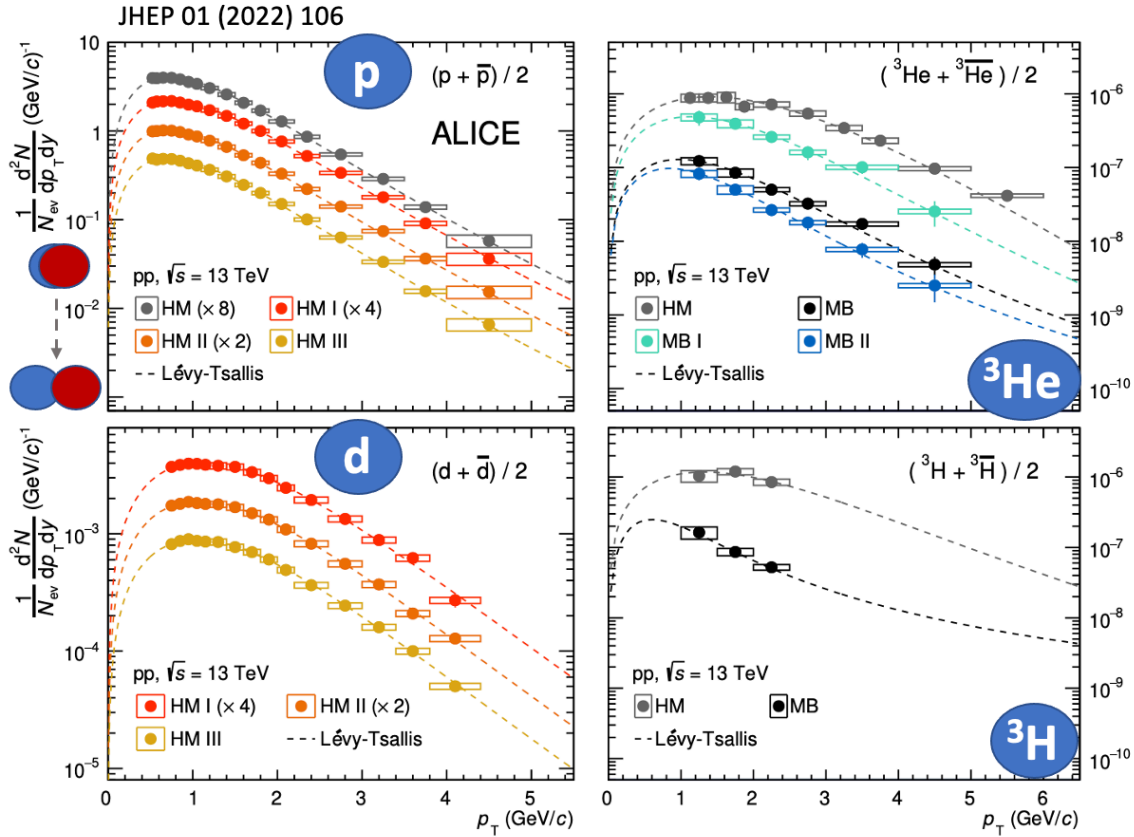
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  - **measurements are in good agreement with 2-body coalescence**
  - tension with SHM at low charged-particle multiplicity density
  - configuration with  $V_c = 3$  dV/dy is excluded by more than  $6\sigma$

Coalescence quantitatively describes the suppression in small systems  
 > the nuclear size matters at low charged-particle multiplicity

p—Pb: PRL 128 (2022) 25, 252003

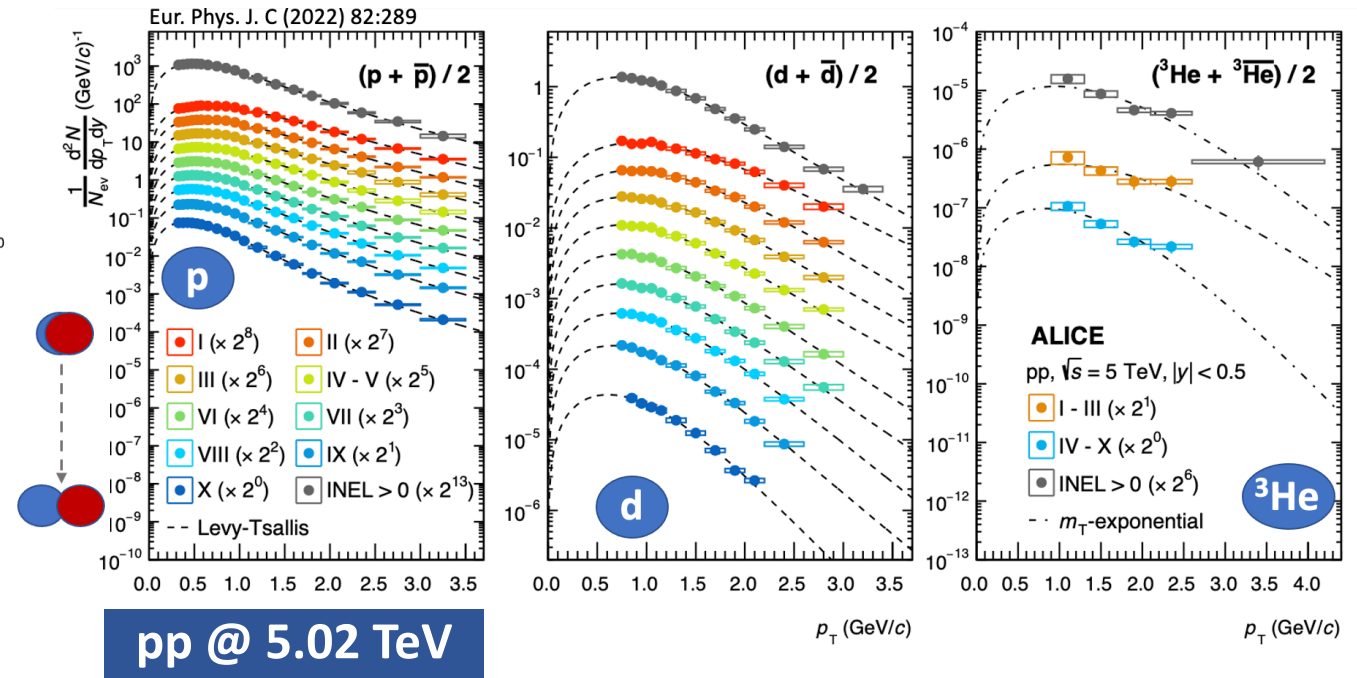
Pb—Pb: PLB 754 (2016) 360-372

# Measurement of light (anti)nuclei with ALICE



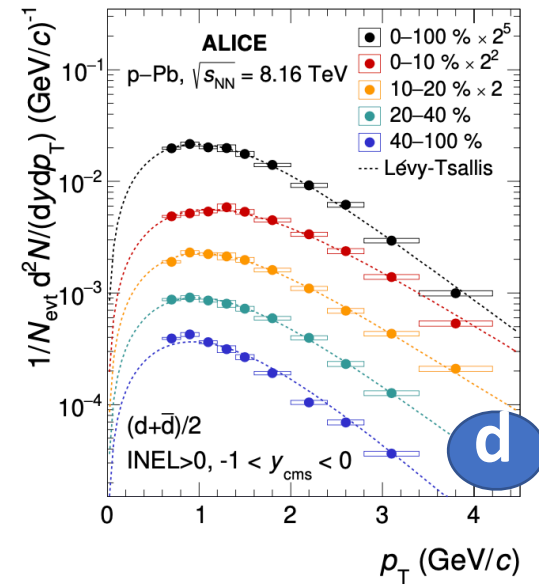
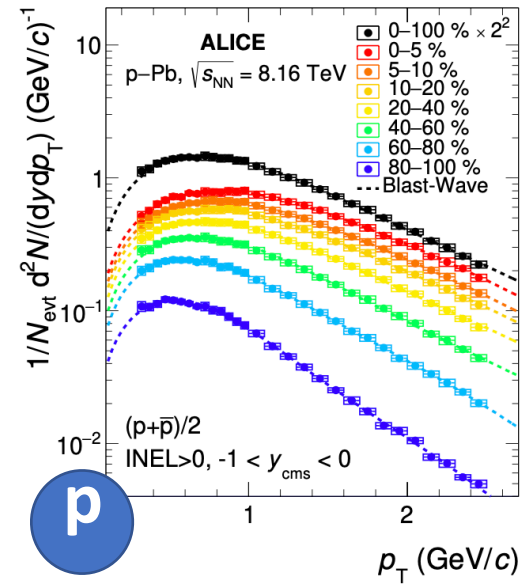
pp @ 13 TeV

- In small systems such as pp and p–Pb collisions, all nuclei species have been measured, from p to  $^3\text{He}$
- Momentum distributions fitted with Lévy-Tsallis function to extrapolate the yield in the unmeasured regions

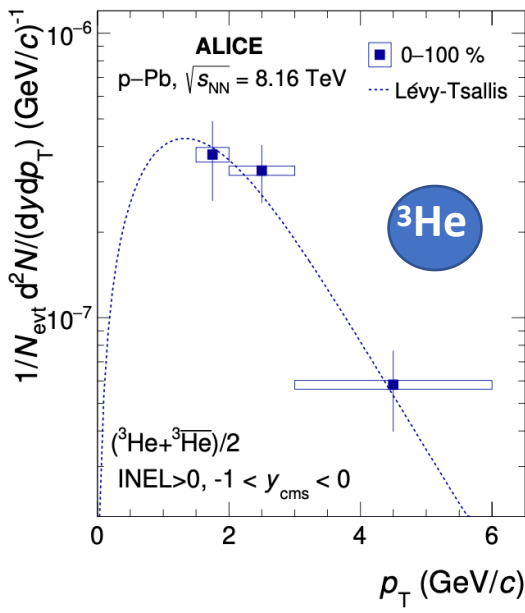


pp @ 5.02 TeV

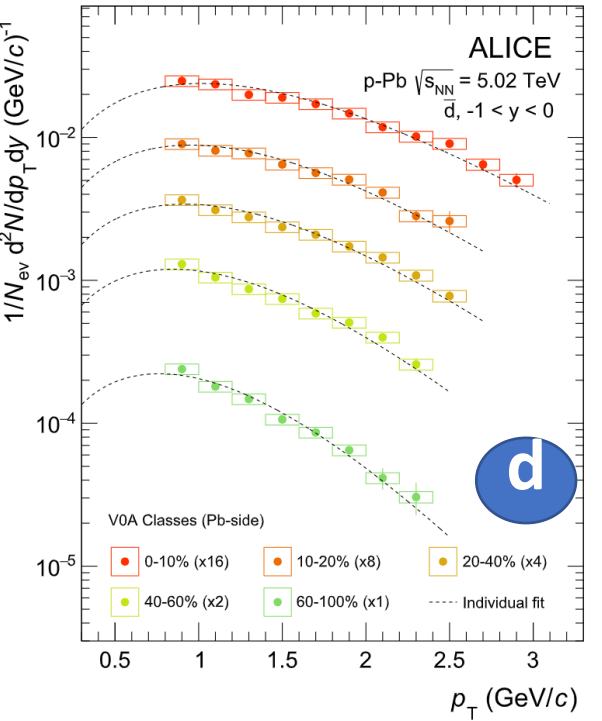
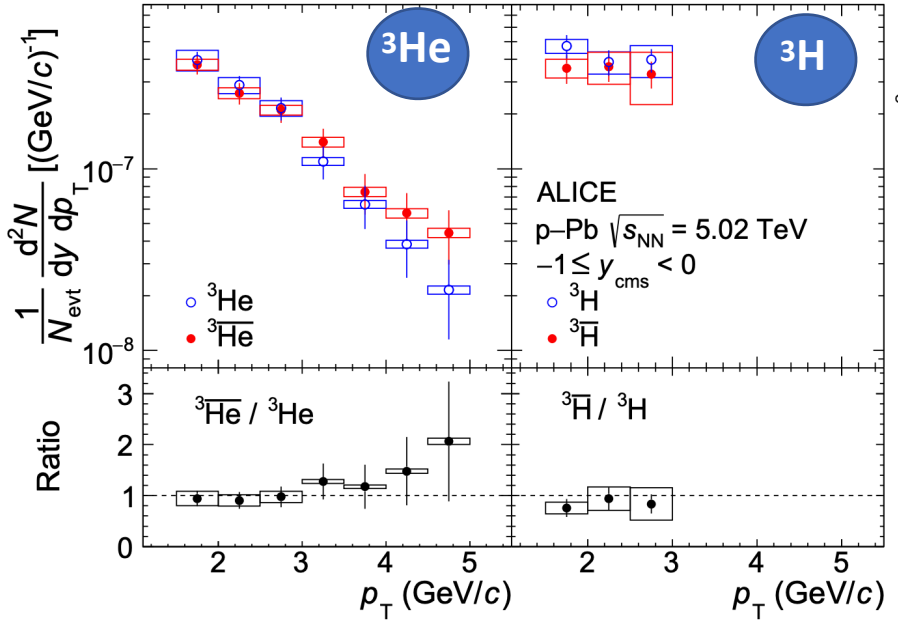
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**p-Pb @ 8.16 TeV**



PLB 800 (2020) 135043

# Inelastic cross section of antinuclei



ALICE measured the **inelastic cross section** for **antinuclei** using the LHC as antimatter factory and the ALICE detector as a target

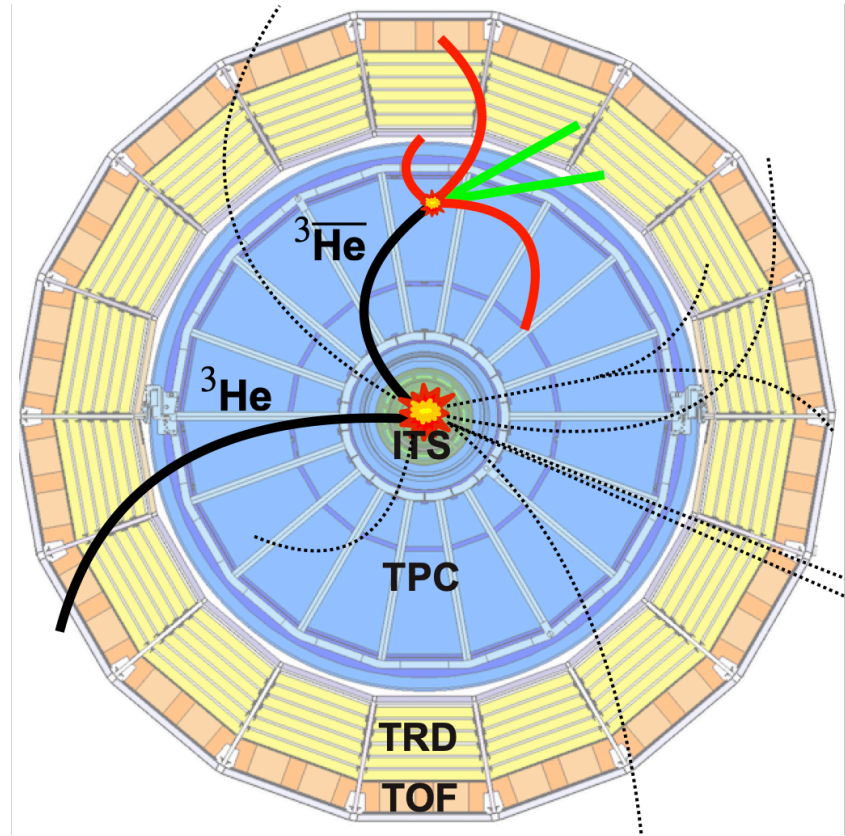
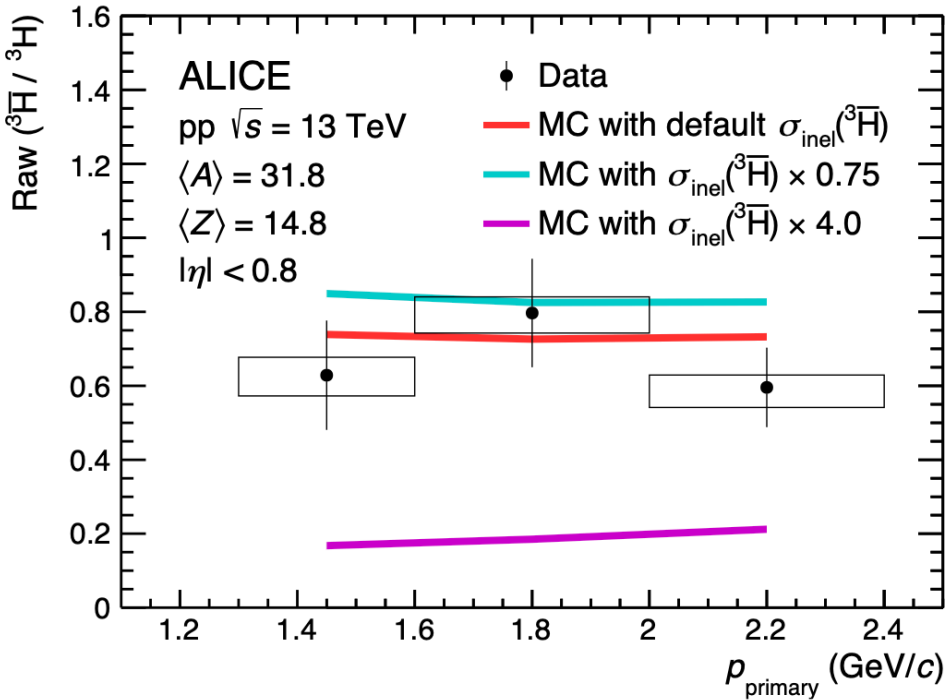


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## Antimatter-to-matter ratio

- Measurement of reconstructed **anti<sup>3</sup>H/<sup>3</sup>H** ratio and compare to MC simulation expectations



**New!**

Phys.Rev.Lett. 125, 162001 (2020)

arXiv:2307.03603 [nucl-ex]

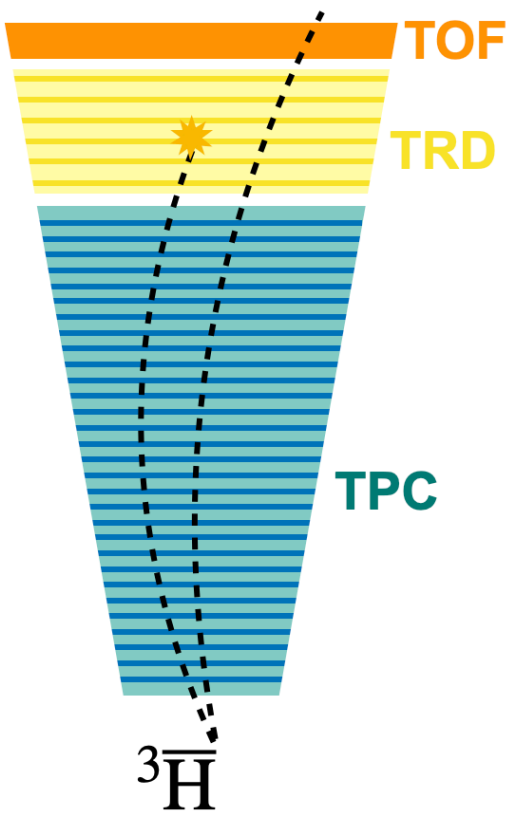
chiara.pinto@cern.ch

Sketch adapted from: Nature Phys. (2023) 19, 61–71



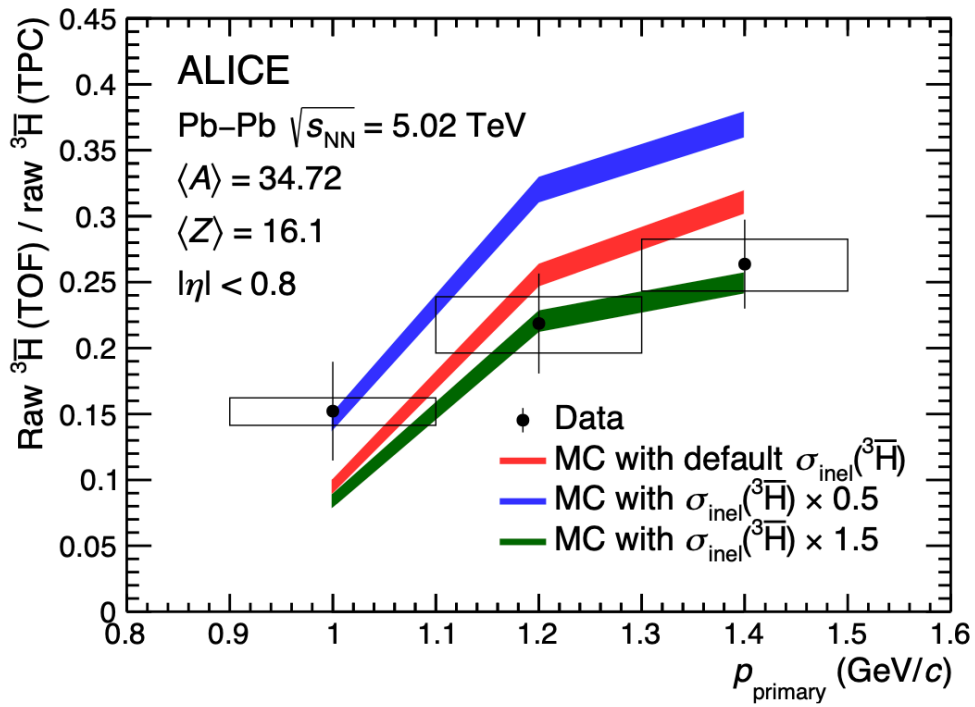
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## TOF/TPC-matching ratio

- Measurement of reconstructed  $\text{anti}^3\text{H}_{\text{TOF}} / \text{anti}^3\text{H}_{\text{TPC}}$  ratio and compare to MC simulation expectations



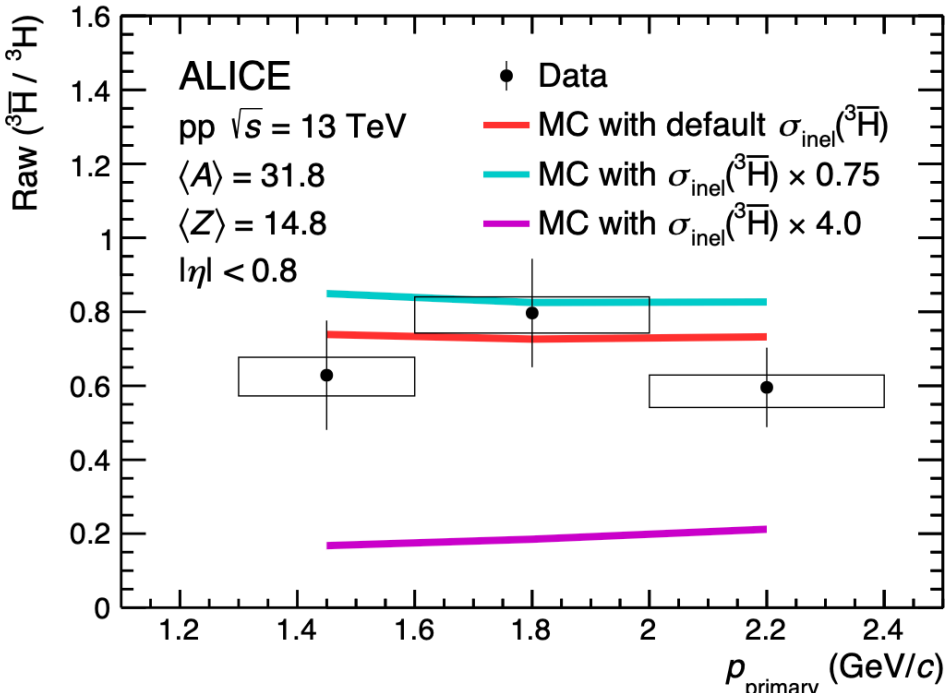
**New!**

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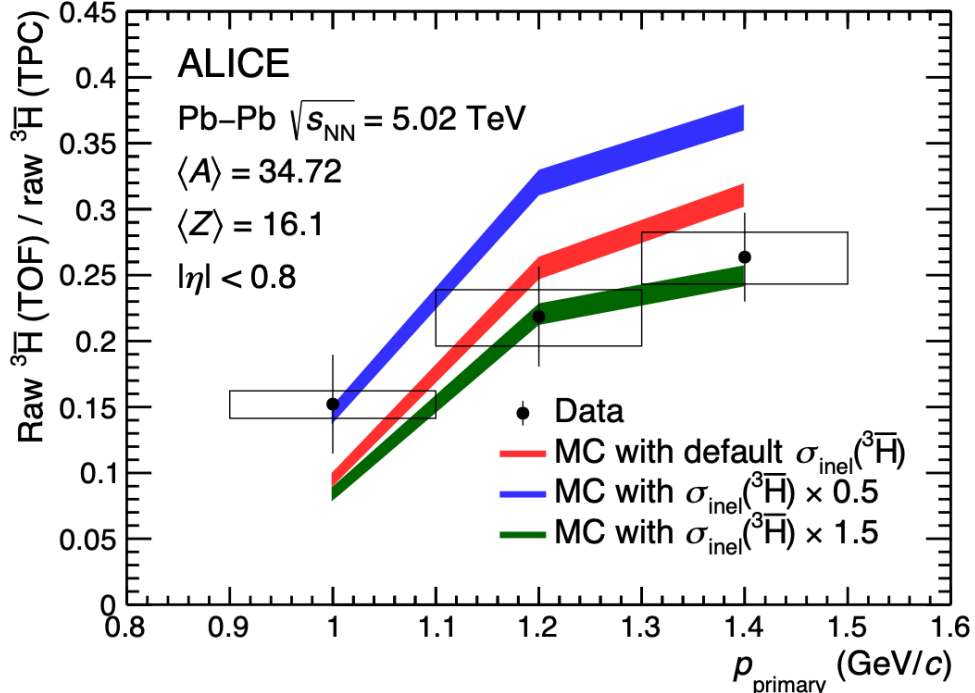
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## TOF/TPC-matching ratio

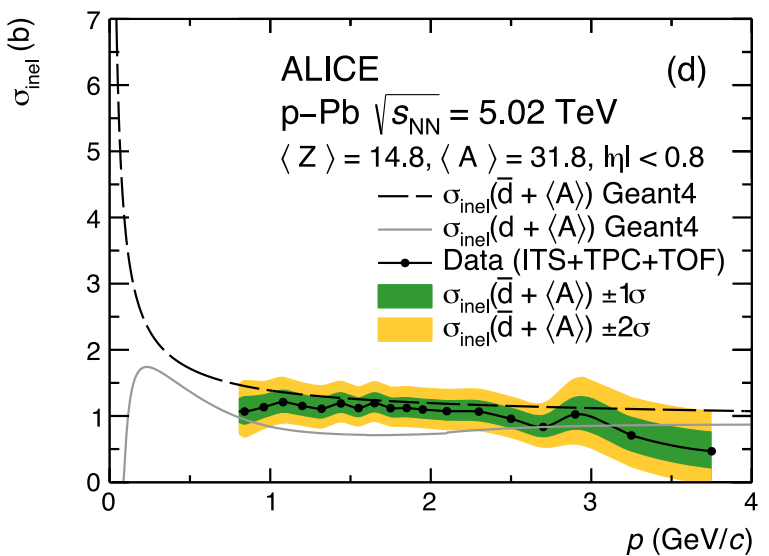
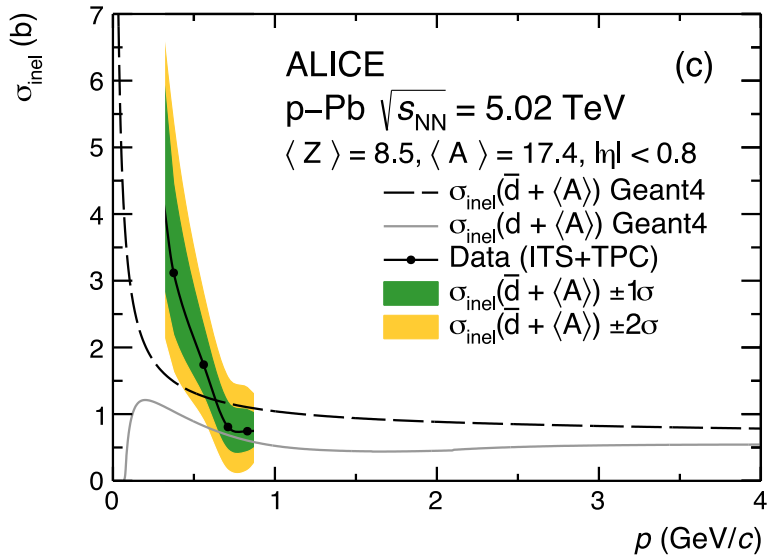
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**New!**

# Inelastic cross section of antinuclei

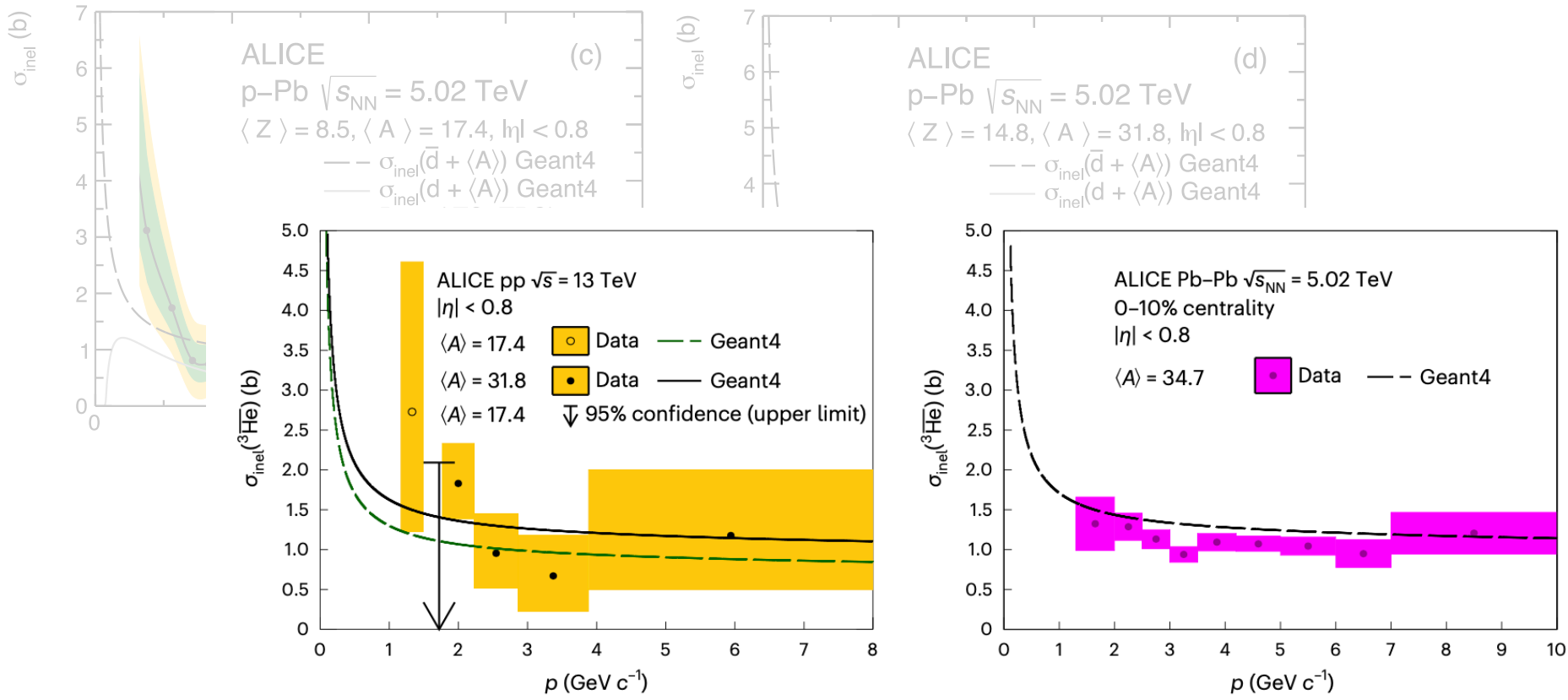
ALICE measured the **inelastic cross section** for **antinuclei** using the LHC as antimatter factory and the ALICE detector as a target



antid: Phys.Rev.Lett. 125, 162001 (2020)

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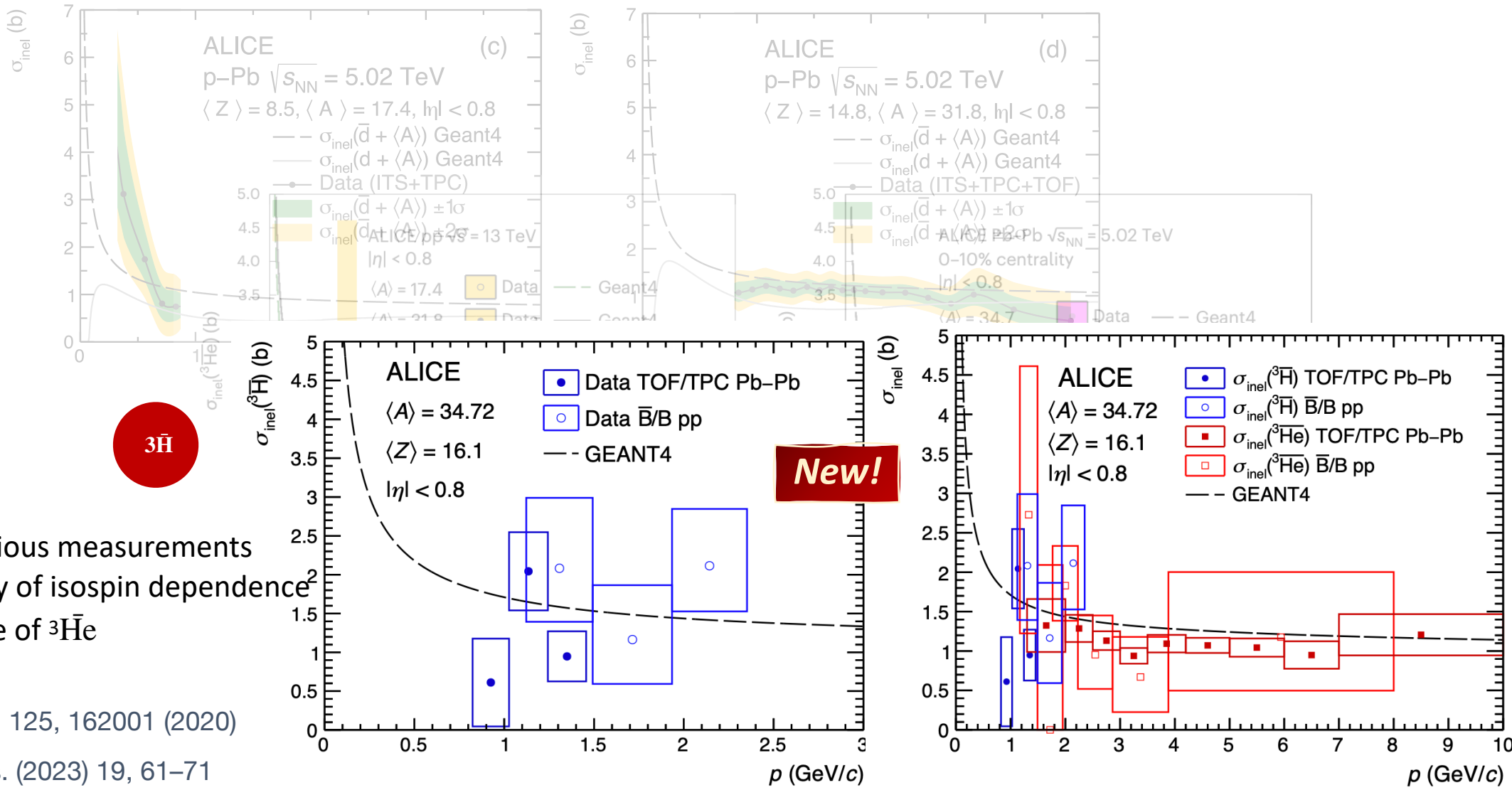


antid: Phys.Rev.Lett. 125, 162001 (2020)

anti<sup>3</sup>He: Nature Phys. (2023) 19, 61-71

# Inelastic cross section of antinuclei

ALICE measured the **inelastic cross section for antinuclei** using the LHC as antimatter factory and the ALICE detector as a target



**$3\bar{H}$**

**New!**

- Complements previous measurements
- Allows for the study of isospin dependence
- Extends the  $p$  range of  ${}^3\bar{H}e$

antid: Phys.Rev.Lett. 125, 162001 (2020)

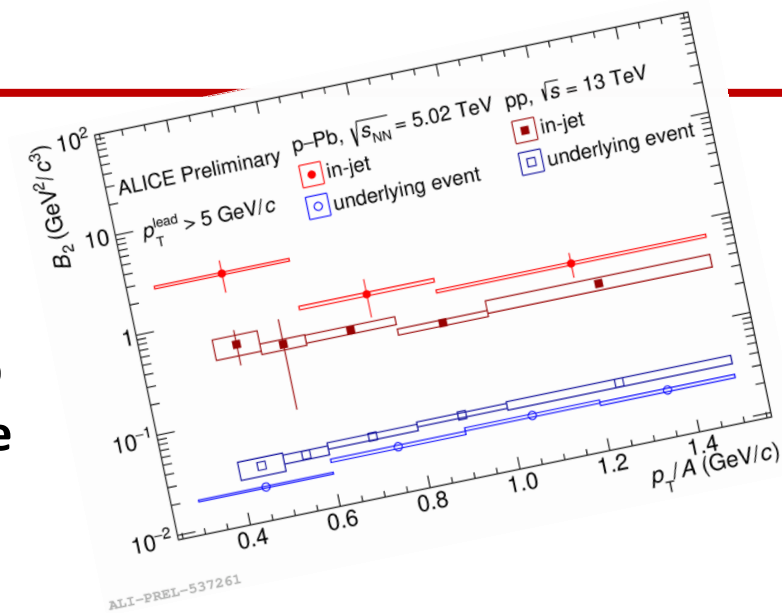
anti ${}^3He$ : Nature Phys. (2023) 19, 61–71

anti ${}^3H$ : arXiv:2307.03603 [nucl-ex]

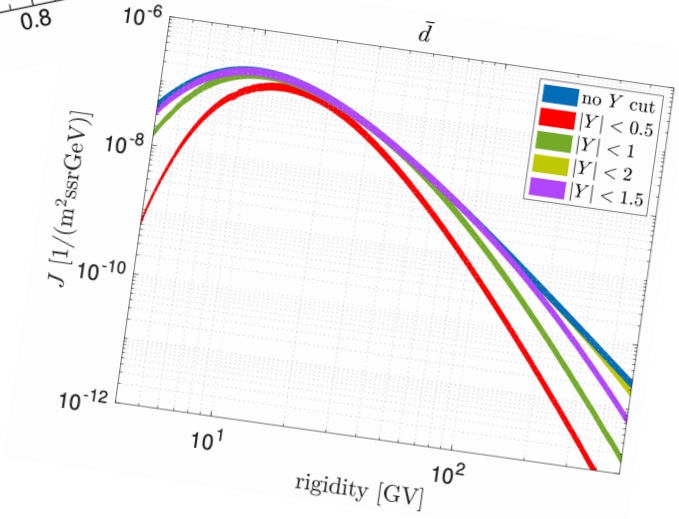
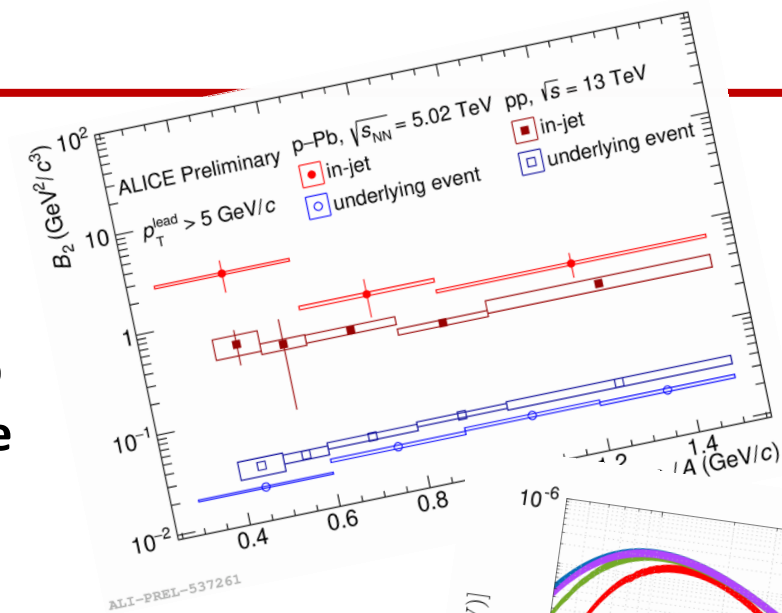
# Summary



- Production of antinuclei measured at accelerators are crucial input in **astrophysical searches** for dark matter
- Antinuclear production **measurements in and out of jets** in pp and p—Pb collisions helps to further constrain the **coalescence** model

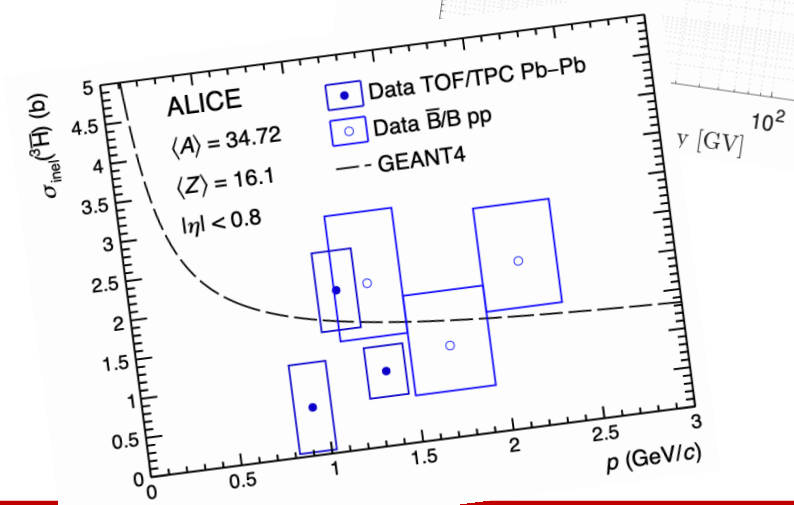
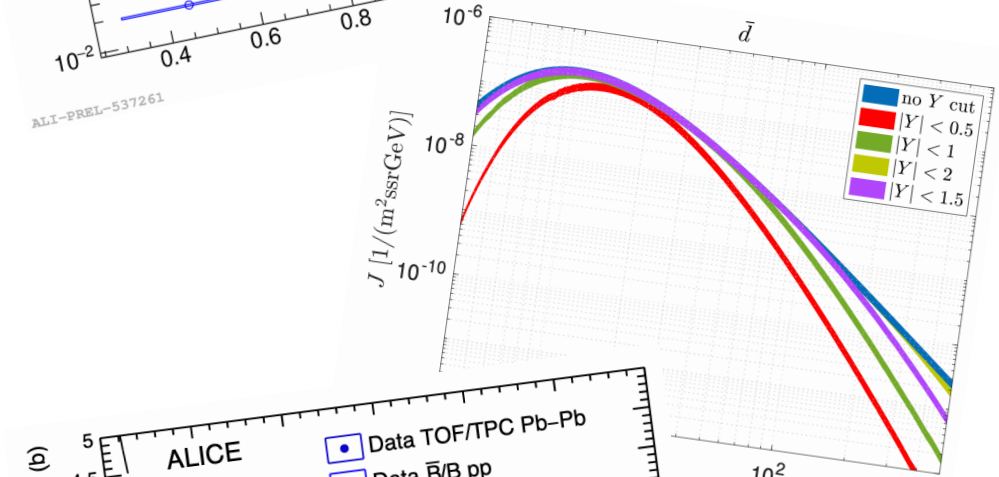
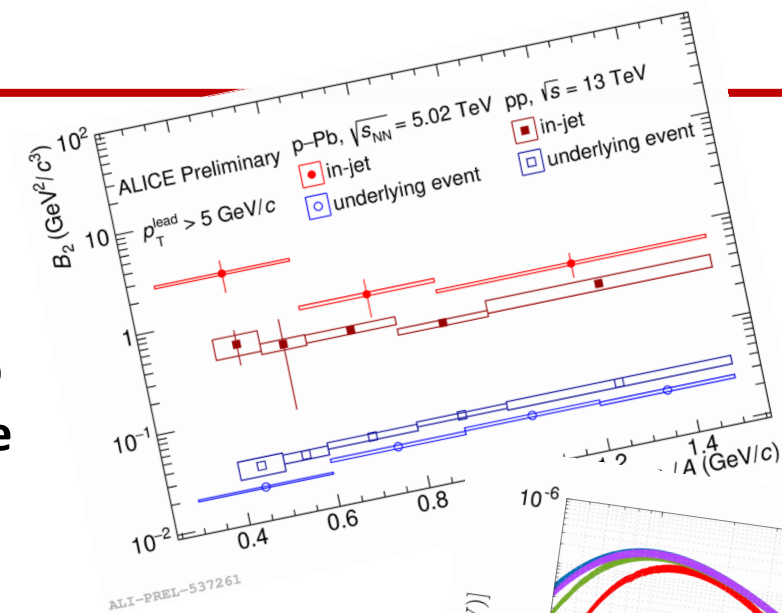


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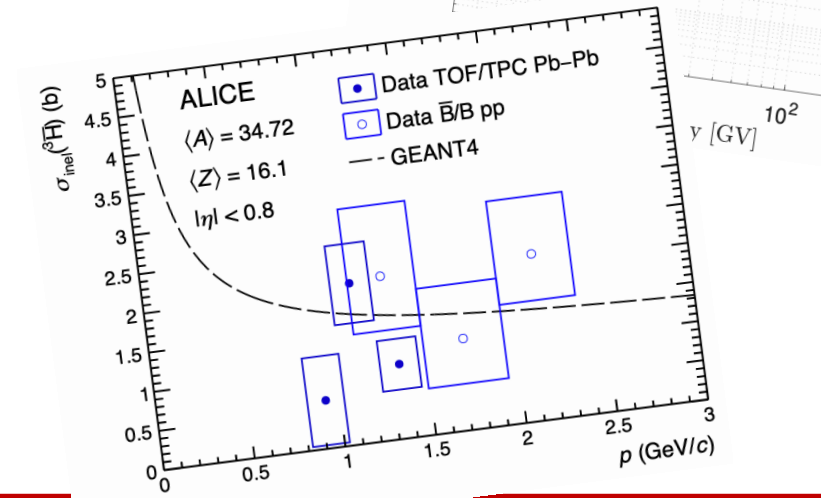
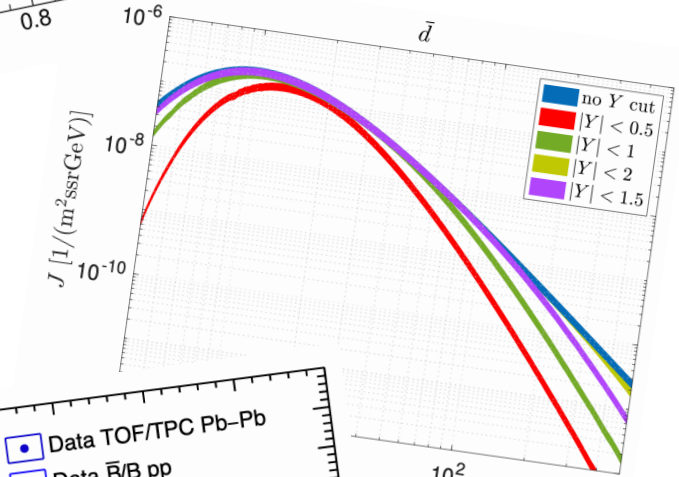
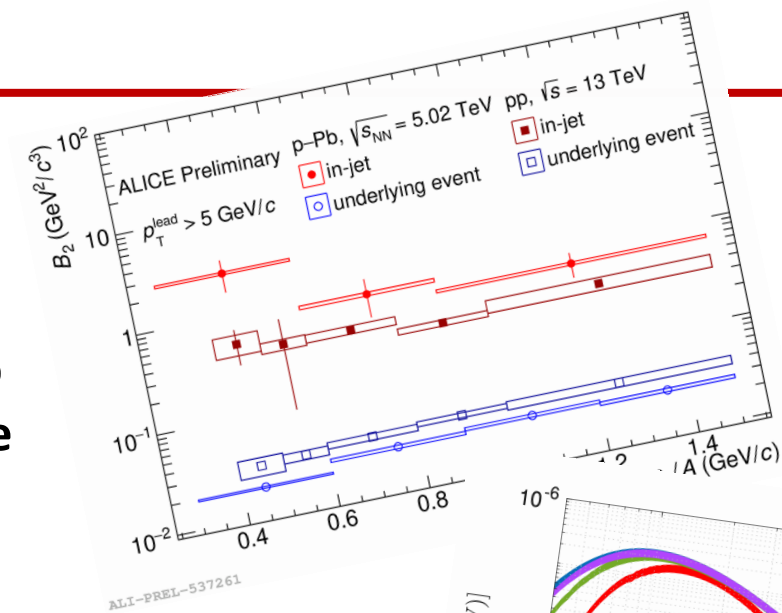




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- **Annihilation processes** have been studied with ALICE, from antiprotons to  ${}^3\bar{\text{H}}\text{e}$  and  ${}^3\bar{\text{H}}$



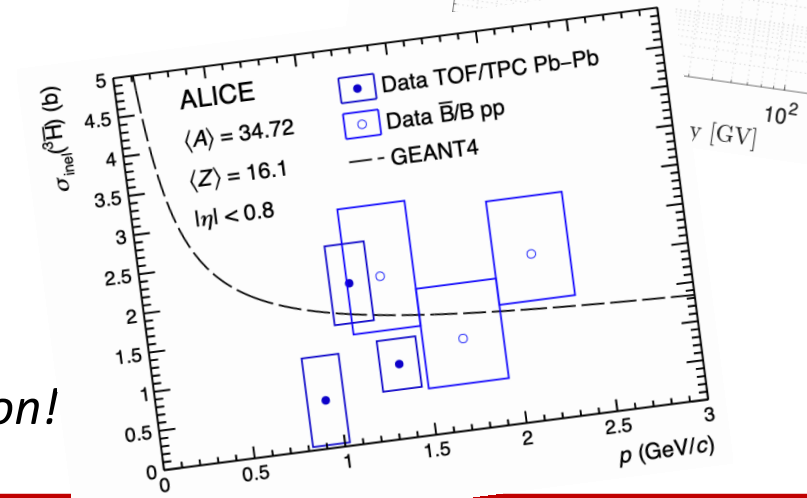
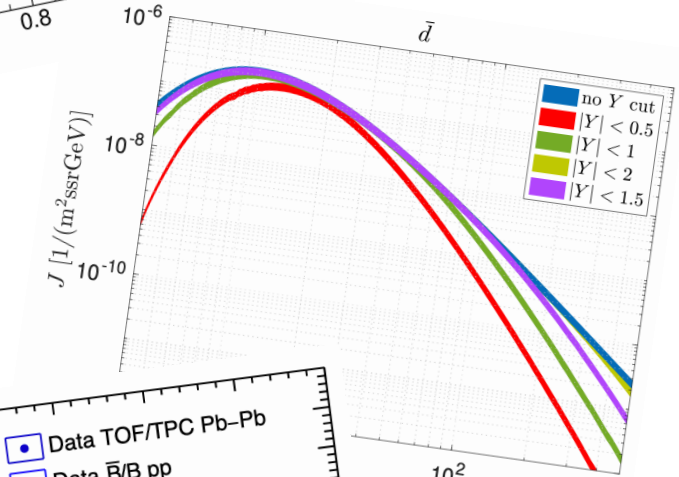
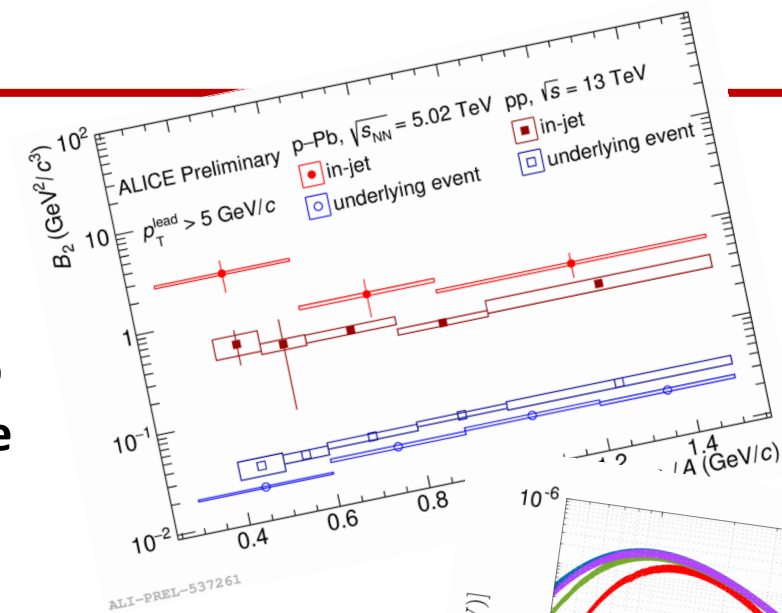
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- More to come with **LHC Run3** increased statistics!



I. Vorobyev's talk  
Wed. 8:50

# Summary

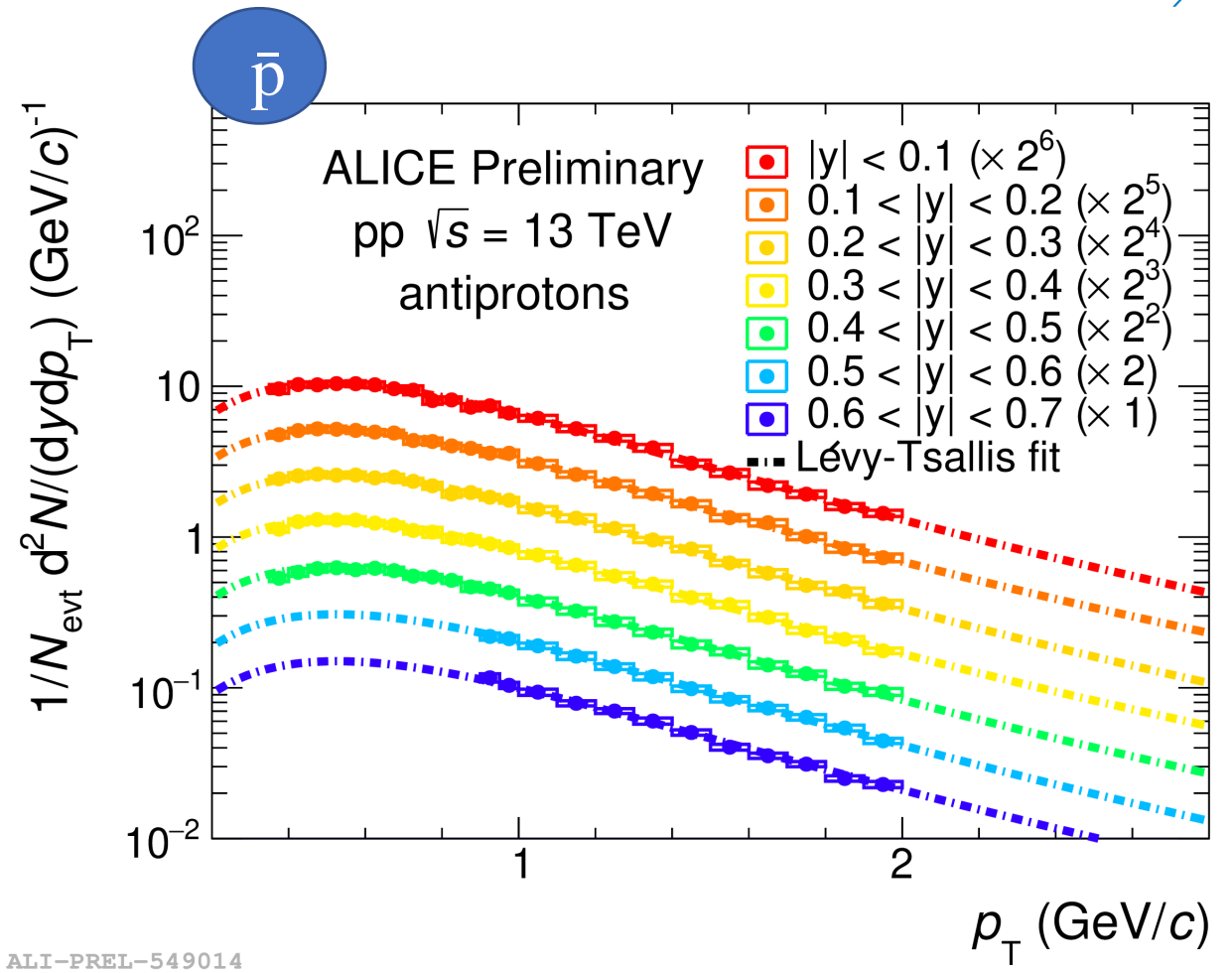
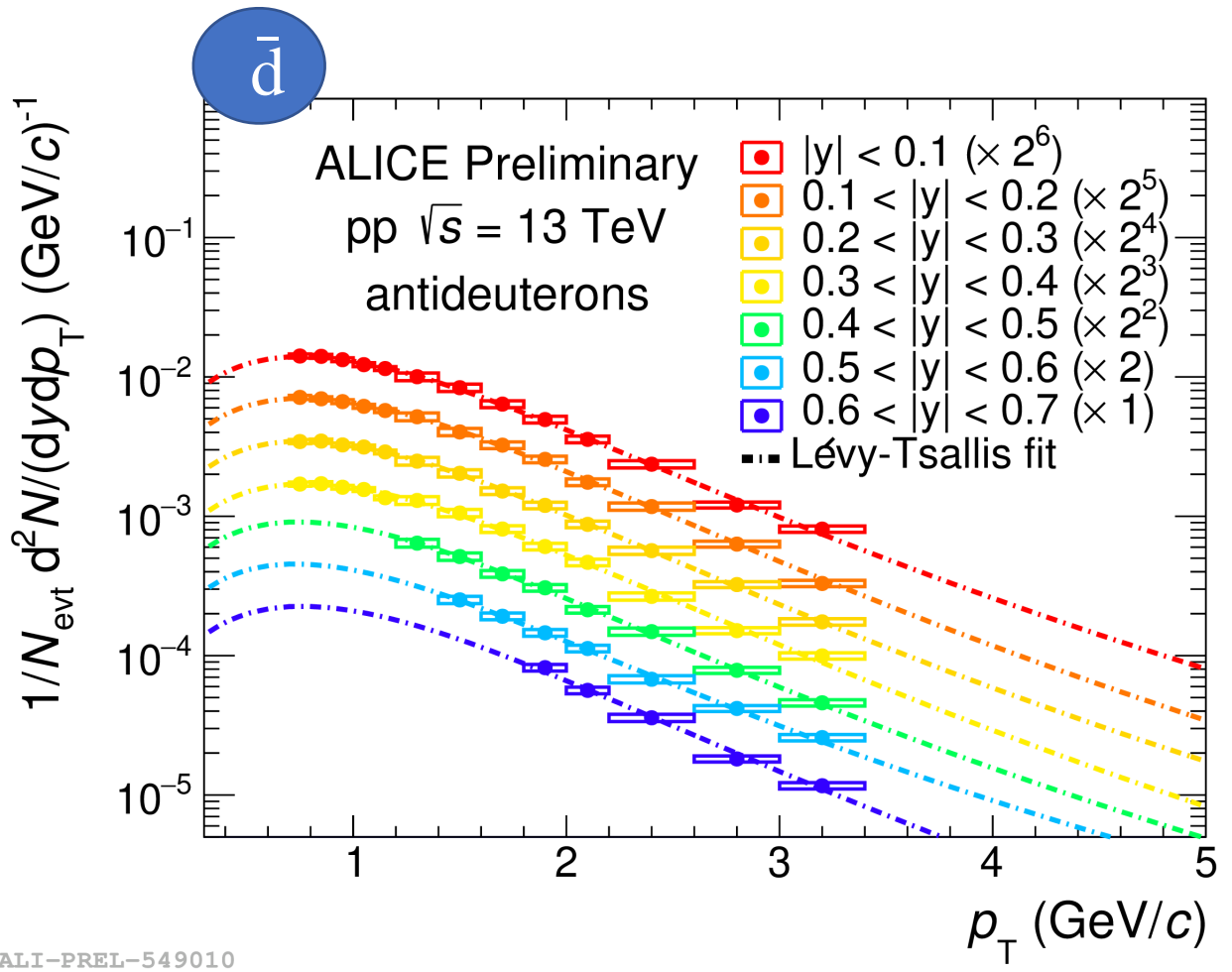
- Production of antinuclei measured at accelerators are crucial input in **astrophysical searches** for dark matter
- Antinuclear production **measurements in and out of jets** in pp and p—Pb collisions helps to further constrain the **coalescence** model
- Measurements of **antinuclear production vs. rapidity** used to extrapolate  $B_2$  at forward rapidity  $\rightarrow$  predict antinuclear **flux** from cosmic rays
- **Annihilation processes** have been studied with ALICE, from antiproton to  $^3\bar{\text{He}}$  and  $^3\bar{\text{H}}$
- More to come with **LHC Run3** increased statistics!



I. Vorobyev's talk  
Wed. 8:50

*Thank you for your attention!*

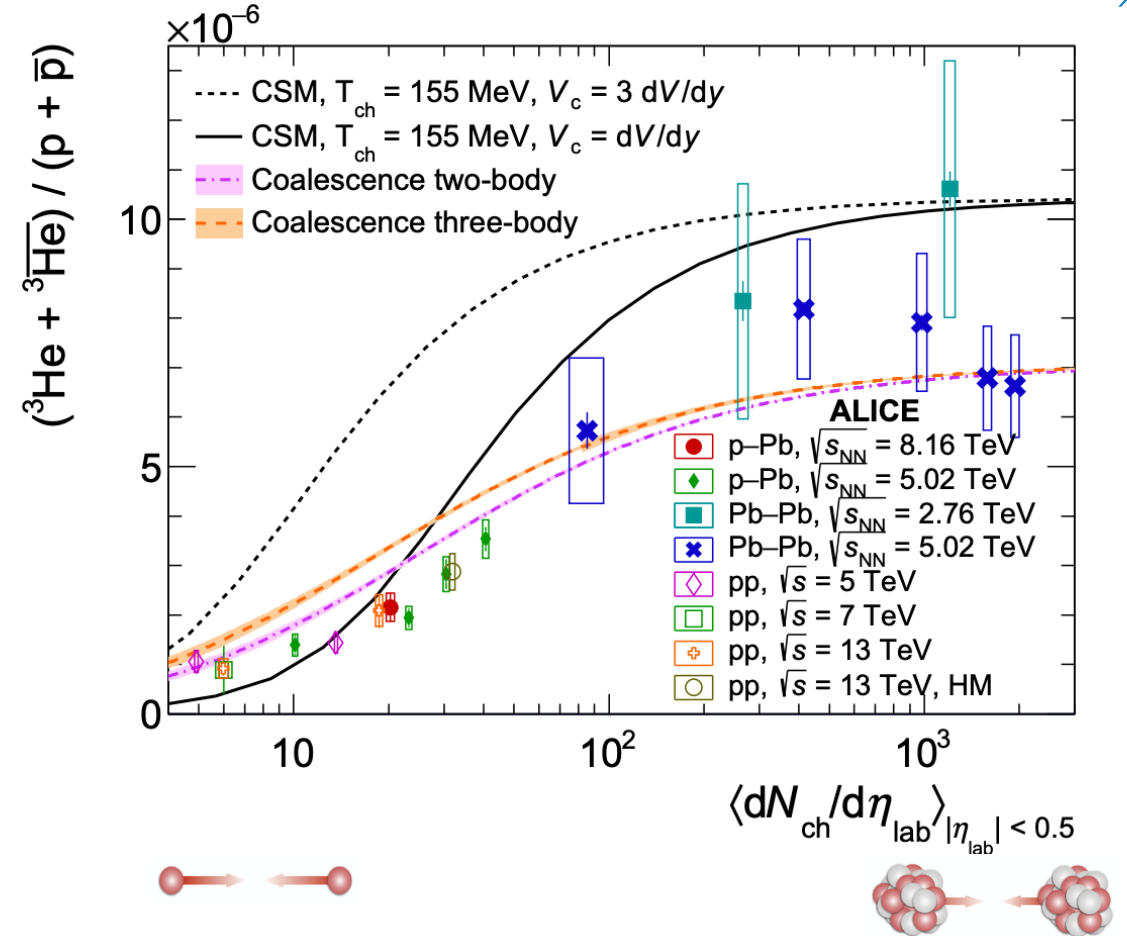
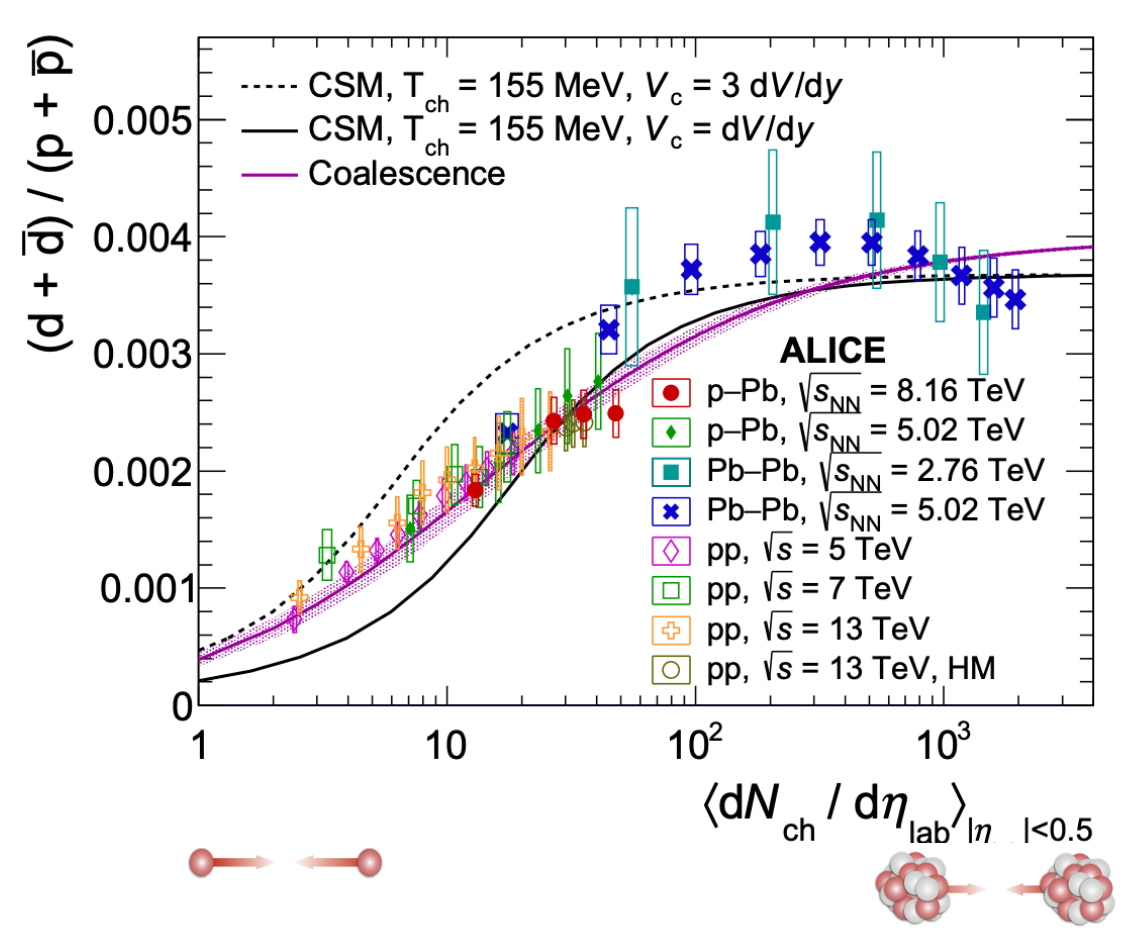
# Spectra as a function of rapidity



- Current acceptance of ALICE detector allows to extend the measurement of antinuclei up to  $y = 0.7$
- All rapidity classes show a common trend with  $y$ , for both species (ratio to  $|y| < 0.1$  is  $\sim 1$ )

**New!**

# Production of (anti)nuclei



- Production of (anti)nuclei has been extensively measured by ALICE
- Coalescence model describes well the data for  $A = 2, 3$
- ALICE measurements cover the midrapidity region ( $|y| < 0.5$ ), while astrophysical models extrapolate to forward region

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## IDEA

- Study of rapidity dependence of antiprotons and antideuterons
- Coalescence parameter  $B_2$  as a function of rapidity
- Comparison with a simple coalescence model

## DATASET

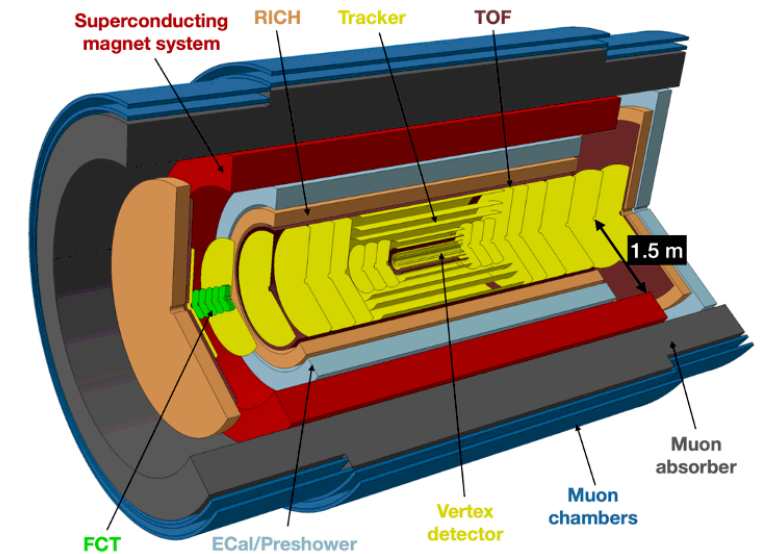
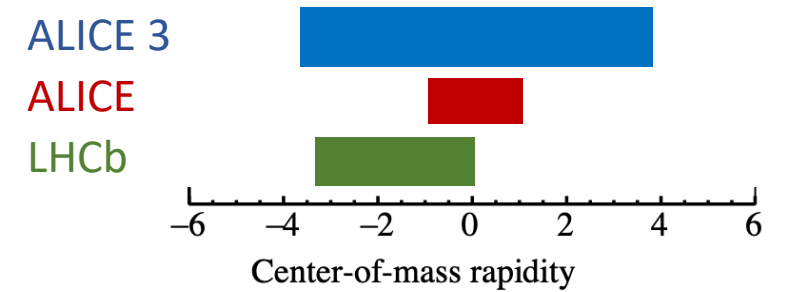
- pp collisions @ 13 TeV, full 2016 + 2017 + 2018 ESD tracks
- $\sim 1.6 \cdot 10^9$  events (after selection cuts)

## MC (JIRA)

- 2016 pp, 13 TeV - Pythia8 Monash2013 + injected (hyper)nuclei – based on G4

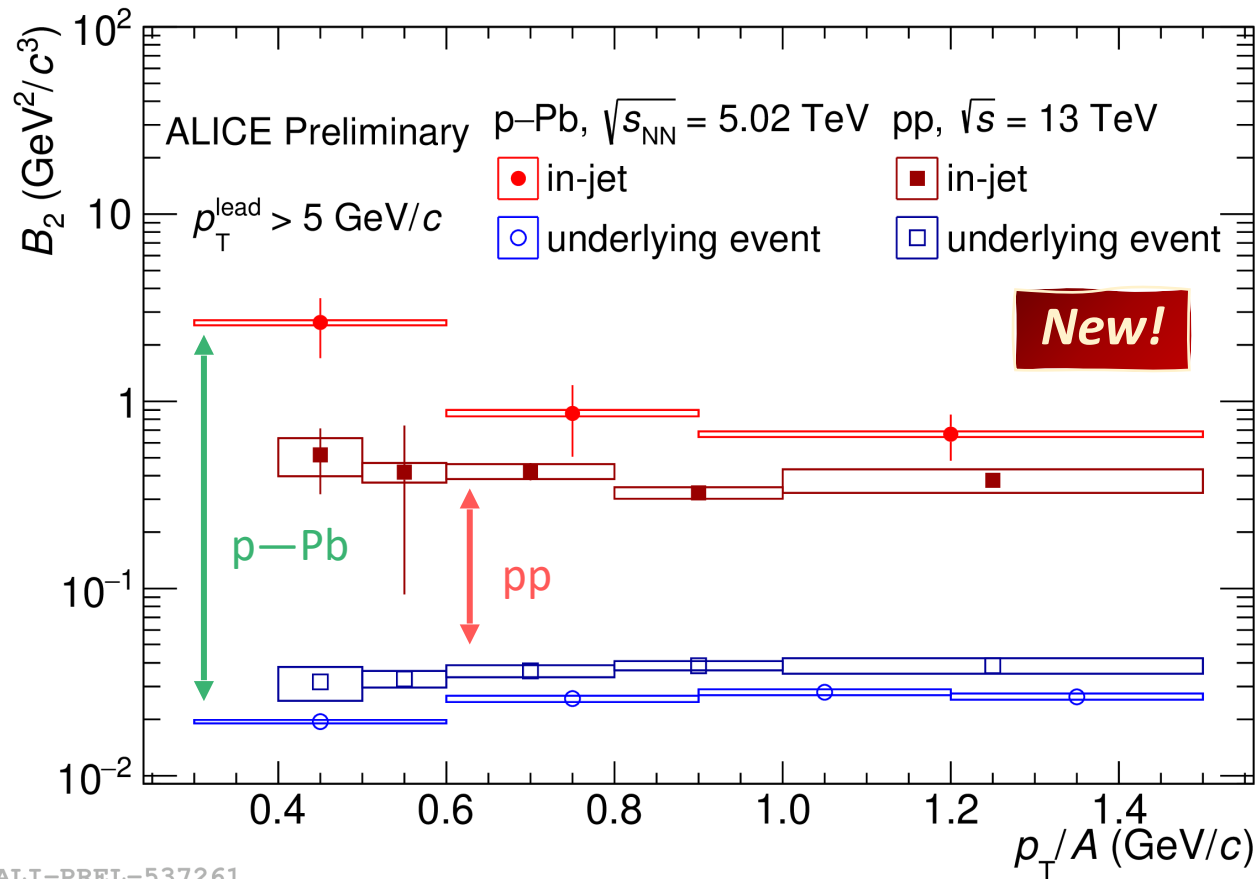
## RESULTS

- Measurements up to  $y=0.7$
- $y$ -differential measurements will be possible with ALICE 3 (rapidity coverage  $\rightarrow |y| \lesssim 4$ ) ([eprint:1902.01211](https://arxiv.org/abs/1902.01211) [[physics.ins-det](https://arxiv.org/abs/1902.01211)])

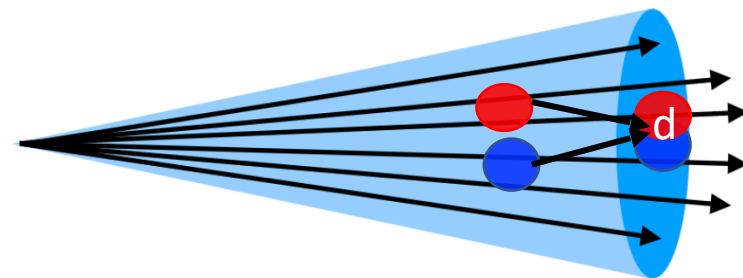




# Coalescence parameters in and out of jets



- $B_2$  in-jet even more enhanced than  $B_2$  in UE in  $p\text{-Pb}$  collisions (factor  $\sim 25$ )
- $B_2$  in-jet in  $p\text{-Pb}$  is larger than  $B_2$  in-jet in  $pp$   $\rightarrow$  could be related to the different particle composition of jets in  $pp$  and  $p\text{-Pb}$
- $B_2$  in UE in  $p\text{-Pb}$  is smaller than  $B_2$  in UE in  $pp$  due to the larger source size in  $p\text{-Pb}$  ( $pp^1: r_0 \sim 1 \text{ fm}$ ,  $p\text{-Pb}^2: r_0 \sim 1.5 \text{ fm}$ )



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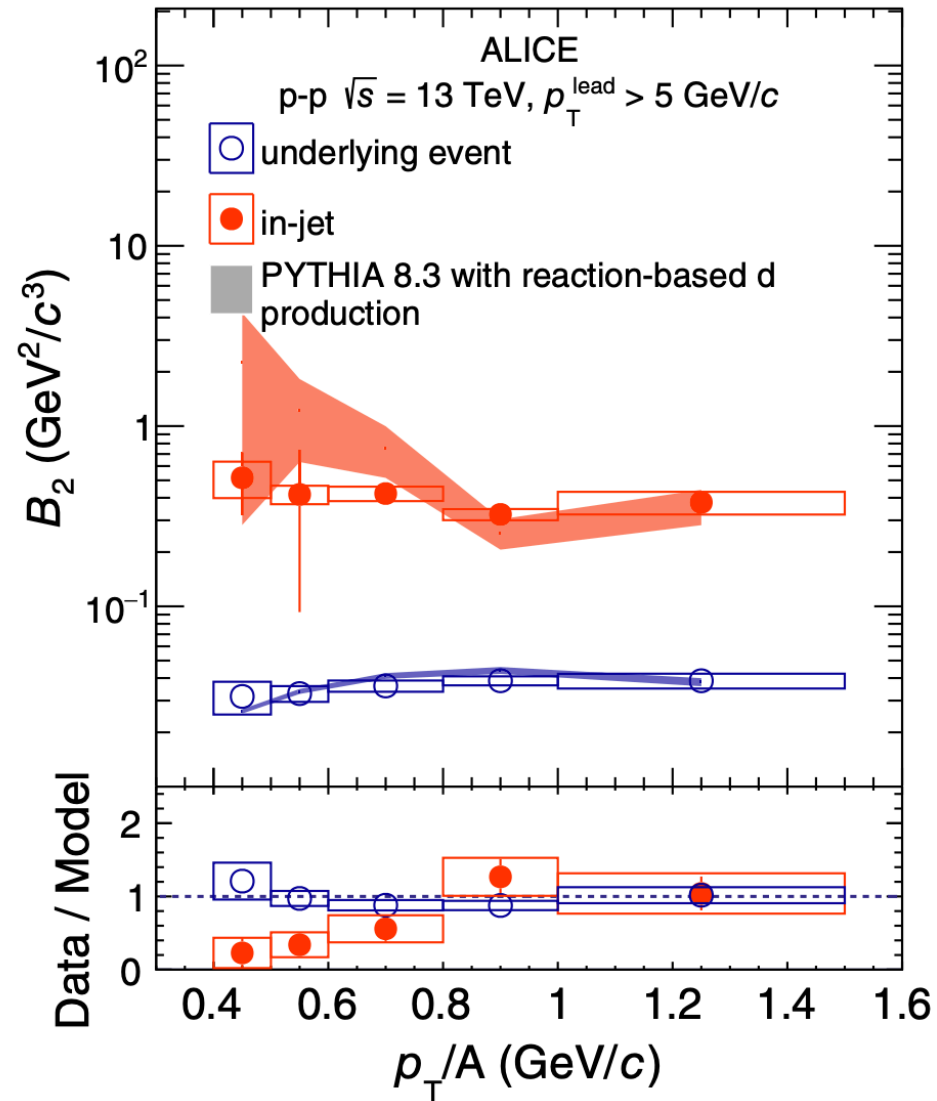
Phys.Rev.Lett. 131 (2023) 4, 042301

1 Phys.Rev.C 99 (2019) 024001

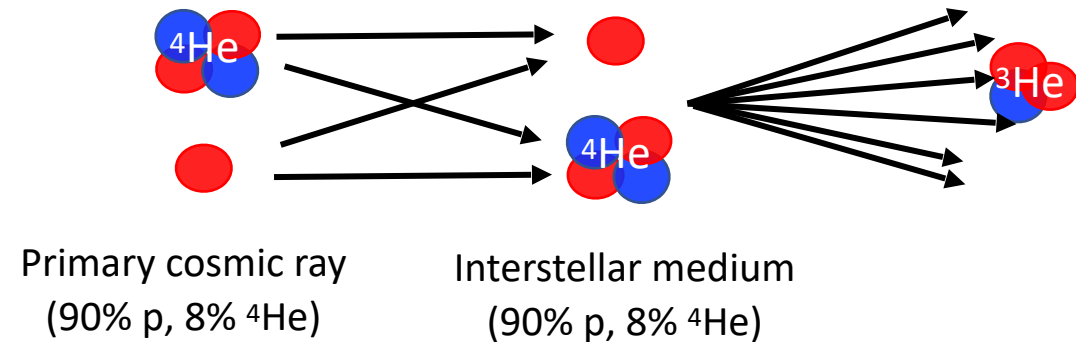
2 Phys.Rev.Lett. 123 (2019) 112002  
 chiara.pinto@cern.ch



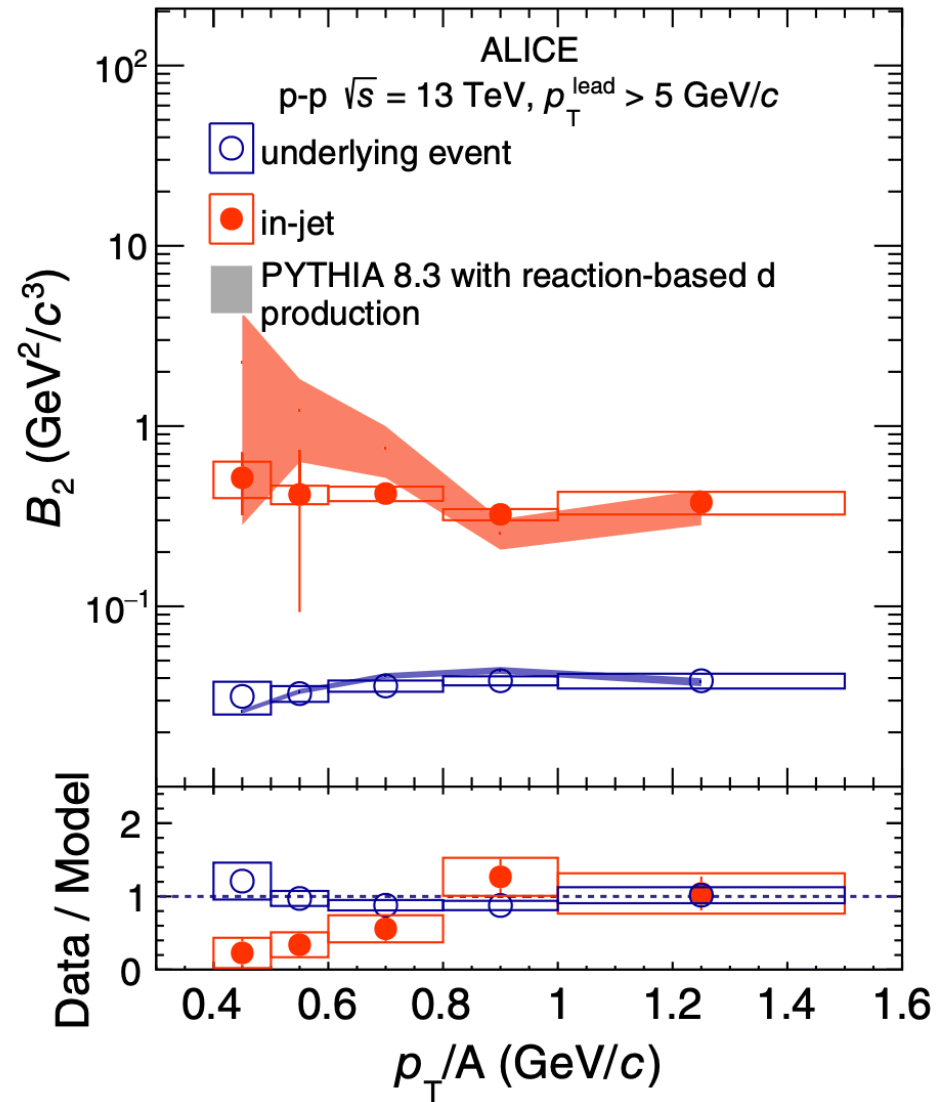
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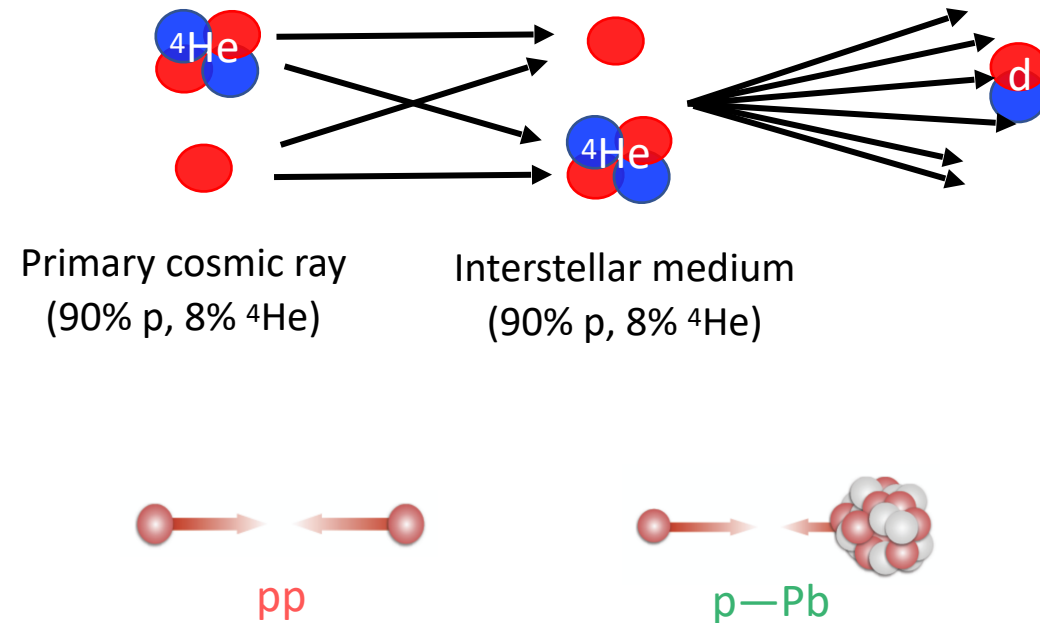
- $B_2$  in-jet  $\sim 15$  times larger than  $B_2$  in UE
- *Enhanced deuteron coalescence probability in jets wrt UE is observed for the first time in pp collisions*
- Due to the reduced distance in phase space of hadrons in jets compared to those out of jets  $\rightarrow$  favors coalescence picture



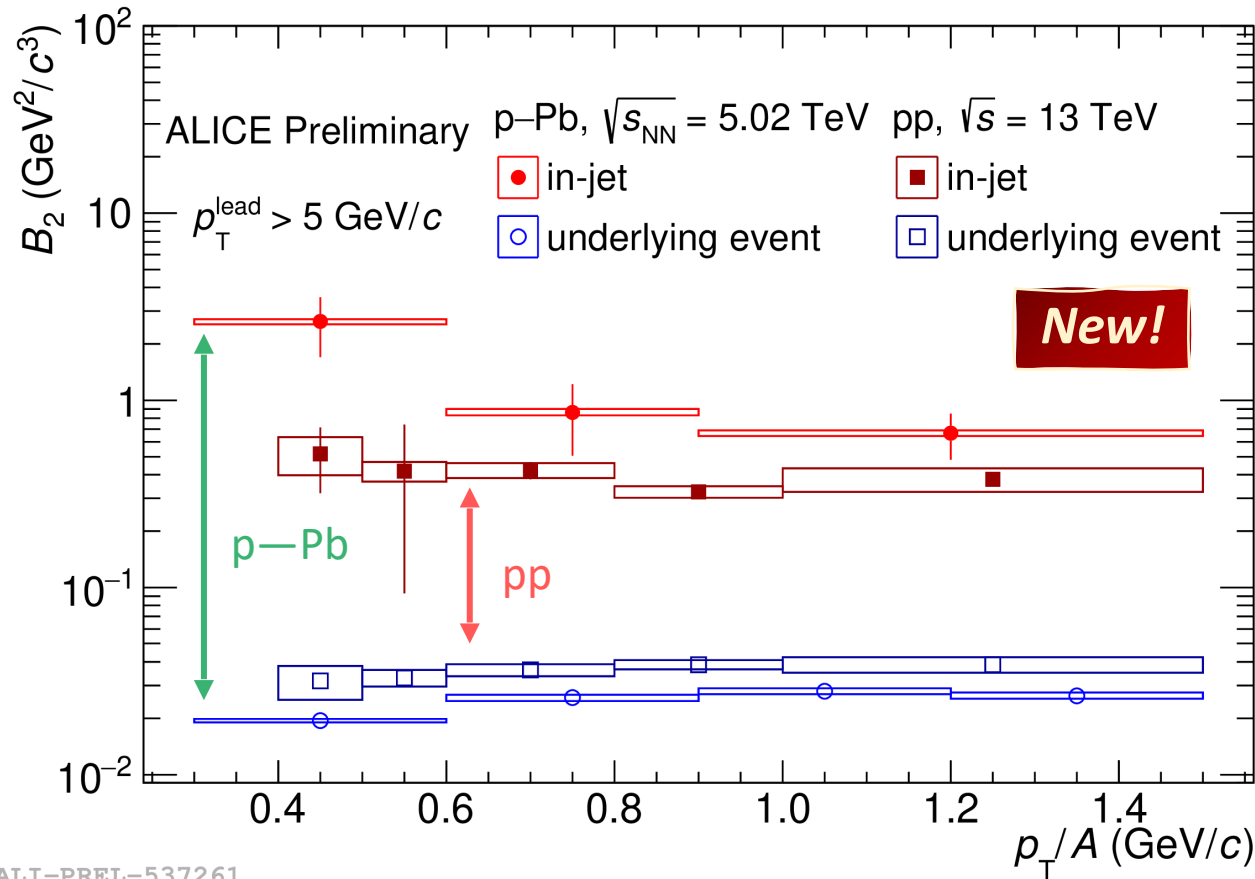
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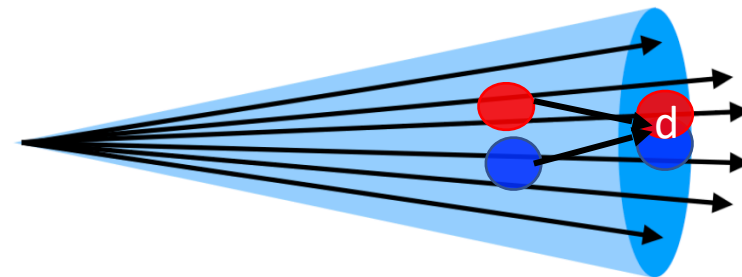
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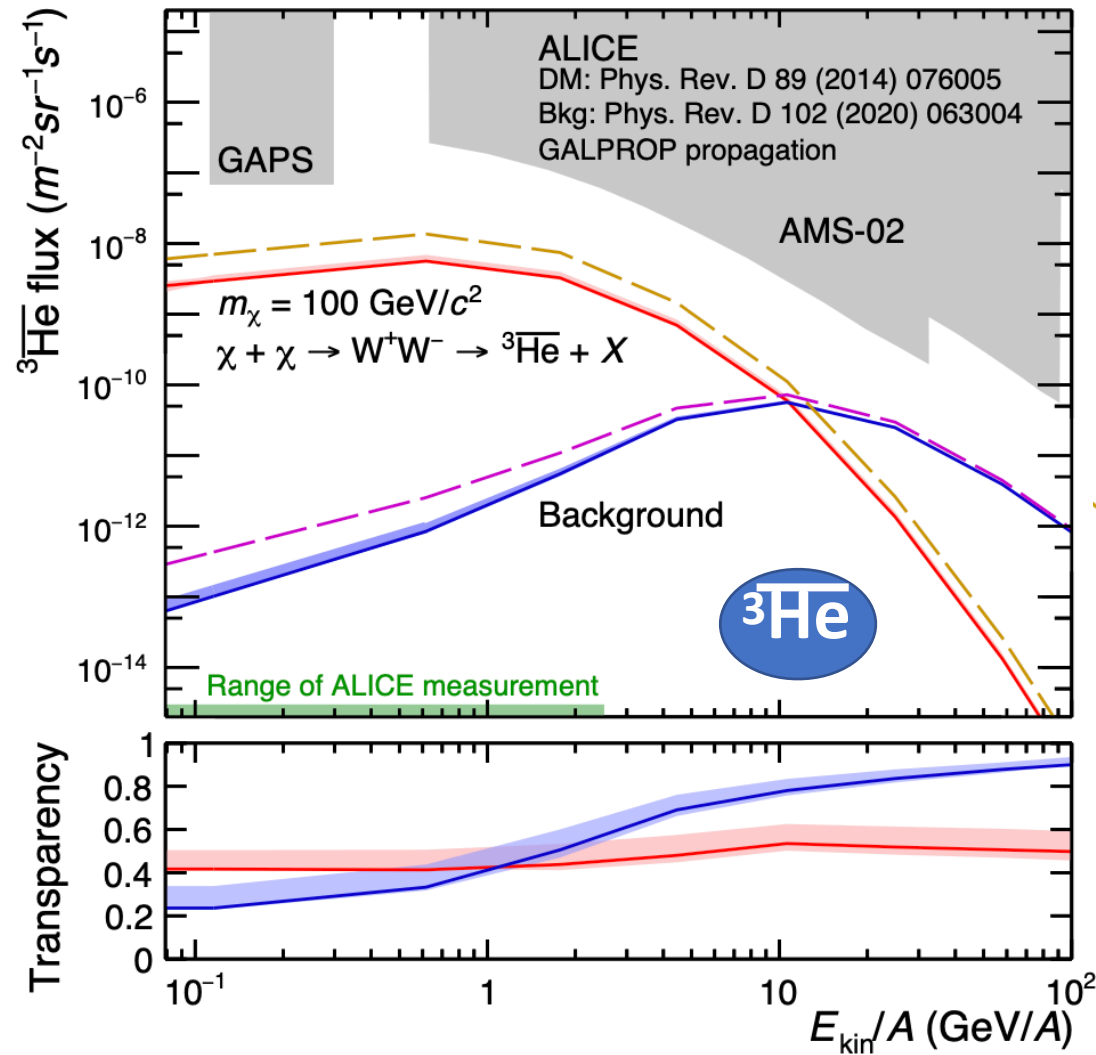
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# Transparency of Galaxy to anti<sup>3</sup>He



Solar modulated flux



$$\text{Transparency} = \frac{\text{flux with annihilation}}{\text{flux without annihilation}} = \frac{\text{---} (\text{---})}{\text{---} (\text{---})} \text{ for bkg}$$

- (DM)
- $\sigma_{\text{inel}}^{\text{GEANT4 DM}}$
  - $\sigma_{\text{inel}}^{\text{GEANT4 bkg}}$
  - $\sigma_{\text{inel}}^{\text{ALICE DM}}$
  - $\sigma_{\text{inel}}^{\text{ALICE bkg}}$
  - $\sigma_{\text{inel}} = 0 \text{ DM}$
  - $\sigma_{\text{inel}} = 0 \text{ bkg}$

- Fluxes are model dependent
- **Our Galaxy is rather constantly transparent to <sup>3</sup>He passage**
- Data are in good agreement with Geant4 predictions
- Uncertainties on Transparency only due to absorption measurements (10-20%)