

Andy Sproles, ORNL

The **Skyrme** EDF: the view from Brussels

May 31th 2022



wryssens.com

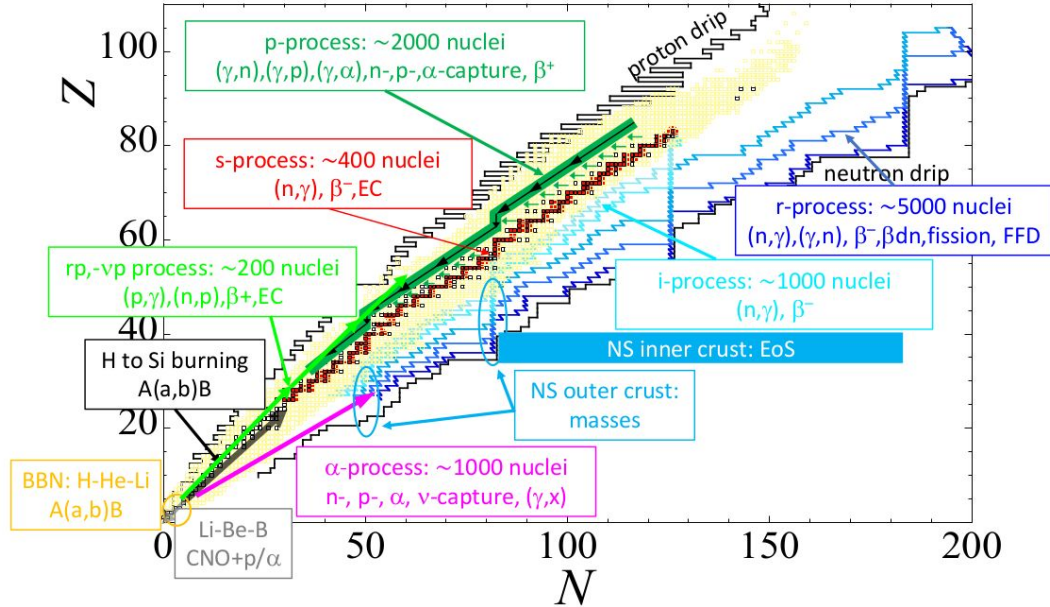
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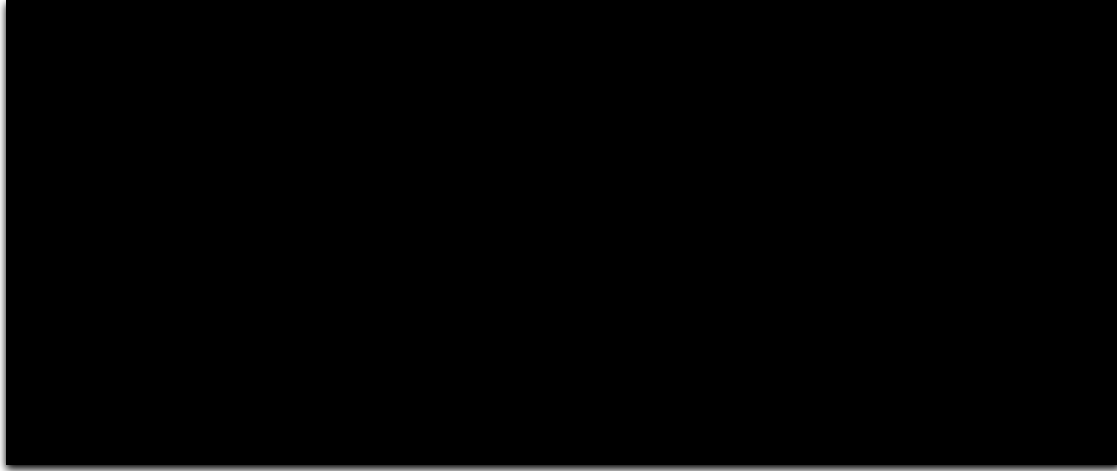
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I. The call of astrophysics



I. The call of astrophysics: **a video!**

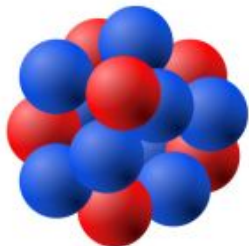


Animation by J.-F. Lemaître (CEA)

Modelling challenge

- predictions for thousands of nuclei
- ... for many **different observables**
- ... in a **unified** fashion
- ... founded in **microscopic** physics
- ... despite **enormous extrapolations**

II. The Skyrme EDF: **so you want to describe nuclei?**



$$\hat{H}|\Psi\rangle = E|\Psi\rangle.$$

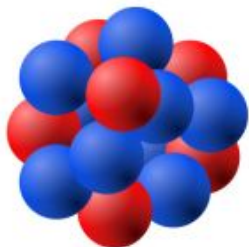
Interaction?

- Look at QCD, n-n scattering?
- Phenomenological?
- Global or local?

Model space?

- Include all nucleons?
- Numerical representation?
- Local or global?

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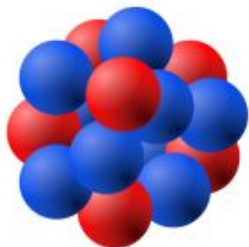
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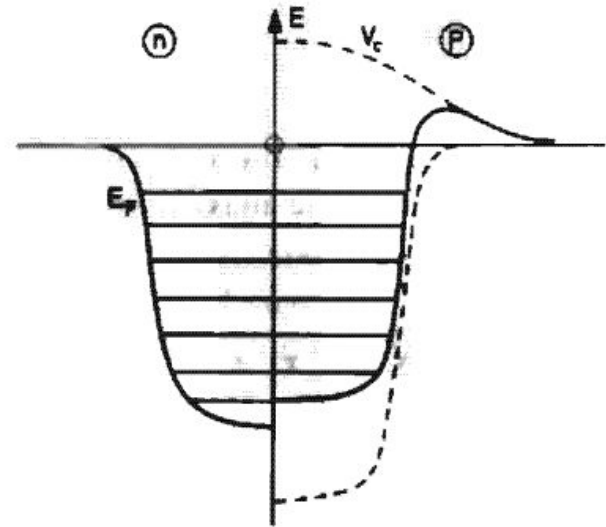
II. The Skyrme EDF: the Skyrme interaction

$$T = \sum_{i < j} \sum t_{ij} + \sum_{i < j < k} \sum t_{ijk}$$

$$t_{12} = \delta(\mathbf{r}_1 - \mathbf{r}_2) t(\mathbf{k}', \mathbf{k})$$

$$\begin{aligned}
 t(\mathbf{k}', \mathbf{k}) = & t_0(1+x_0 P^\sigma) + \frac{1}{2}t_1(1+x_1 P^\sigma)(\mathbf{k}'^2 + \mathbf{k}^2) \\
 & + t_2[1+x_2(P^\sigma - \frac{4}{5})]\mathbf{k}' \cdot \mathbf{k} \\
 & + \frac{1}{2}T[\boldsymbol{\sigma}_1 \cdot \mathbf{k}\boldsymbol{\sigma}_2 \cdot \mathbf{k} - \frac{1}{3}\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 \mathbf{k}^2 + \text{conj.}] \\
 & + \frac{1}{2}U[\boldsymbol{\sigma}_1 \cdot \mathbf{k}'\boldsymbol{\sigma}_2 \cdot \mathbf{k} - \frac{1}{3}\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 \mathbf{k}' \cdot \mathbf{k} + \text{conj.}] \\
 & + V[i(\boldsymbol{\sigma}_1 + \boldsymbol{\sigma}_2) \cdot \mathbf{k}' \times \mathbf{k}],
 \end{aligned}$$

= Central + Tensor + Spin-orbit interaction



“It is now clear that such interaction is **very different** from the interaction between two (free) nucleons.”

T. H. R. Skyrme, Nuclear Physics 9, 615–634 (1958).

P. Ring and P. Schuck, the nuclear many-body problem (1984).

II. The Skyrme EDF

D. Vautherin and D. M. Brink, PRC 5, 626-647 (1972).

$$E = \langle HF | \hat{H}_{\text{Sk.}} | HF \rangle = \int H(\mathbf{r}) d^3 \mathbf{r}.$$

$$H(\vec{\mathbf{r}}) = \frac{\hbar^2}{2m} \tau + \frac{3}{8} t_0 \rho^2 + \frac{1}{16} t_3 \rho^3 + \frac{1}{16} (3t_1 + 5t_2) \rho \tau$$

$$+ \frac{1}{64} (9t_1 - 5t_2) (\vec{\nabla} \rho)^2 - \frac{3}{4} W_0 \rho \vec{\nabla} \cdot \vec{\mathbf{J}}.$$

= Central + Tensor + Spin-orbit interaction

$$E_{\text{Sk}} = \int d^3 \mathbf{r} \sum_{t=0,1} \mathcal{E}_t(\mathbf{r}),$$

$$\mathcal{E}_t(\mathbf{r}) = C_t^{\rho\rho} \rho_t^2(\mathbf{r}) + C_t^{\rho\rho\rho\gamma} \rho_0^\gamma(\mathbf{r}) \rho_t^2(\mathbf{r})$$

$$+ C_t^{\rho\tau} \rho_t(\mathbf{r}) \tau_t(\mathbf{r}) + C_t^{\rho\Delta\rho} \rho_t(\mathbf{r}) \Delta\rho_t(\mathbf{r})$$

$$+ C_t^{\rho\nabla\cdot\mathbf{J}} \rho_t(\mathbf{r}) \nabla \cdot \mathbf{J}_t(\mathbf{r}),$$

= Central + Tensor + Spin-orbit functional

Some remarks:

Energy Density Functional = analytical form of the energy density, **can be** linked to an interaction
 Parameterization = values of the coupling constants
 Nuclear DFT is significantly more **ad-hoc** than electronic DFT.

II. The Skyrme EDF: a comparison

		Pros
Skyrme	standard	Simple zero-range interaction => parameter adjustment pushed farthest => most applications (linear response, time-dependent MF, ...)
	Fayans	Gradients in the pairing channel
	N2/3LO	Higher-number-of-gradients expansion
	SEAL	Minimal number of parameters
	Ab initio	Forms inspired by different kinds of ab initio calculations
Gogny		Finite-range interaction (with identical interaction in pairing channel)
RMF		Lorentz-invariant (with implications for coupling constants)

III. Modelspace: mean-field theory

1. Guess some initial product wavefunction with **single-particle wavefunctions**
2. Variationally minimize the energy through variation of the spwfs
3. Restore symmetries if needed
4. Mix different solutions: **GCM**

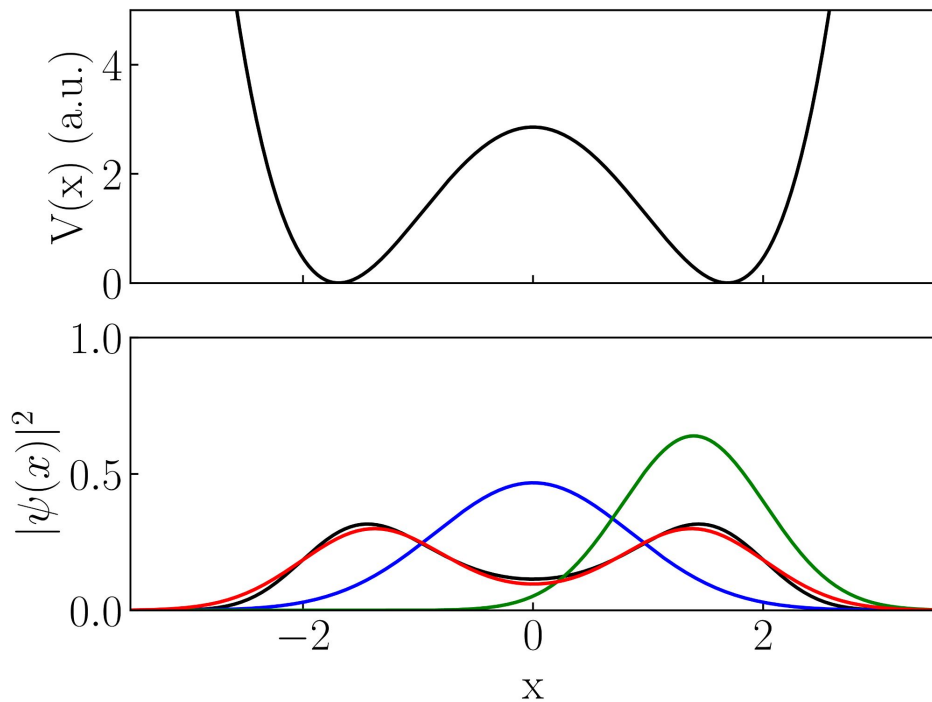
$$|\Psi\rangle = c_1^\dagger \dots c_A^\dagger |0\rangle.$$

$$|\Psi_{\text{opt.}}\rangle = \arg \min \left\{ \langle \Psi | \hat{H} | \Psi \rangle \mid \Psi = \text{simple} \right\}.$$

$$|\Psi_{\text{sym}}\rangle = \frac{1}{\sqrt{2}} \left[|\Psi_{\text{opt.}}\rangle + \hat{U} |\Psi_{\text{opt.}}\rangle \right].$$

$$|\Psi_{\text{final}}\rangle = \sum_i c_i |\Psi_{\text{sym}}\rangle.$$

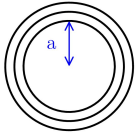
III. Modelspace: symmetry breaking and restoration



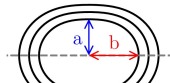
	Energy	$\langle X^2 \rangle$
Exact	1.625	1.98
Gaussian	2.152	0.727
Asym. G.	1.844	2.331
Projected	1.735	2.190

III. Modelspace: nuclear deformation

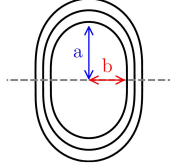
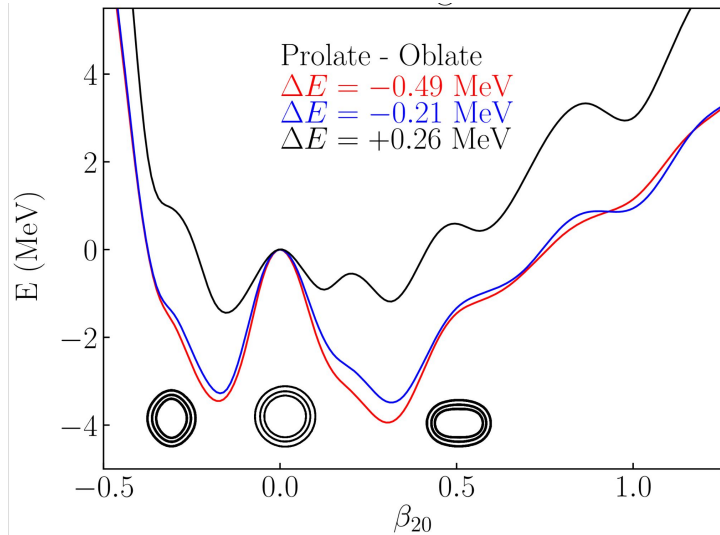
Spherical



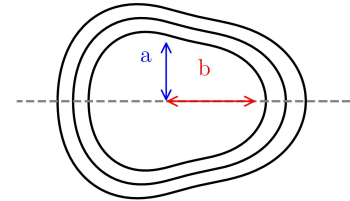
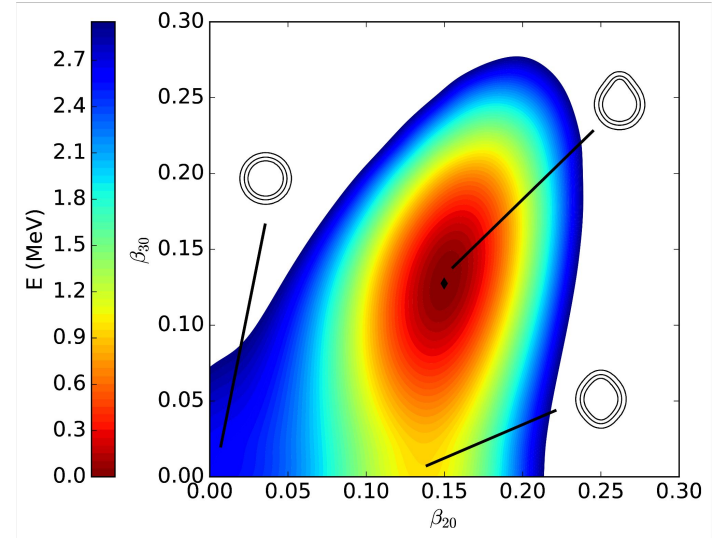
Prolate



Oblate

One DOF: β_{20} 

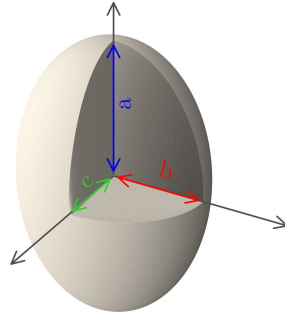
Reflection-asymmetric (RA)

Two DOF: β_{20}, β_{30} 

III. Modelspace: nuclear deformation

In general:

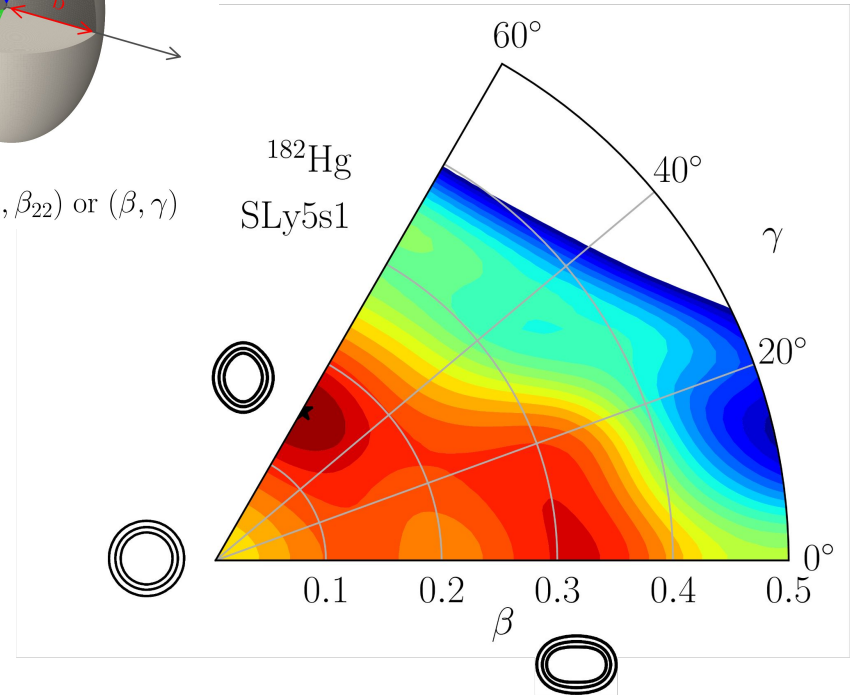
$$\beta_{\ell m} = \frac{4\pi}{3(r_0 A^{1/3})^\ell A} \int d^3\mathbf{r} \rho(\mathbf{r}) r^\ell \operatorname{Re} Y_{\ell m}(\theta, \phi),$$



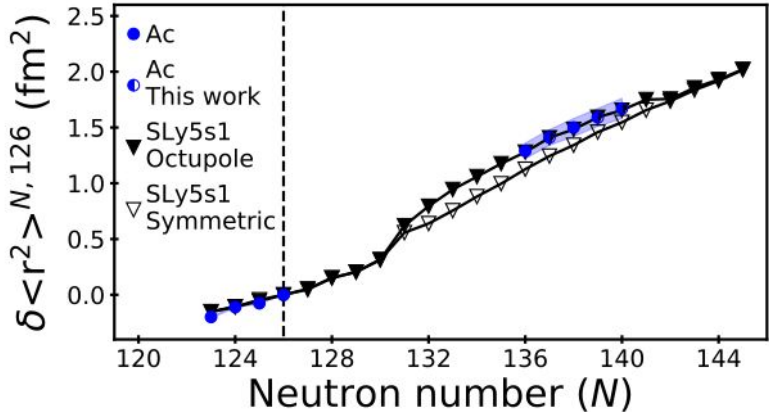
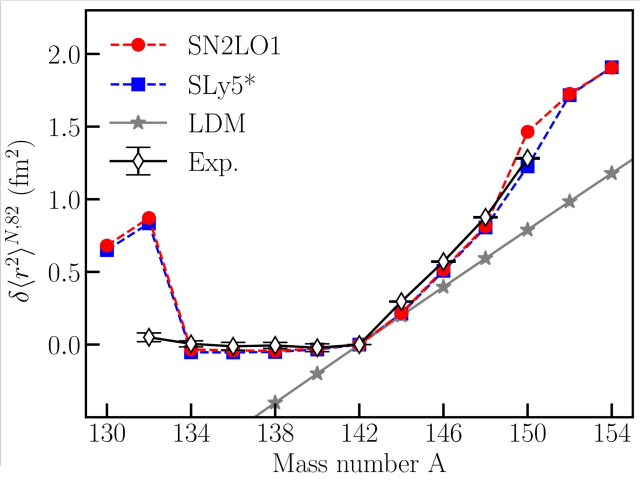
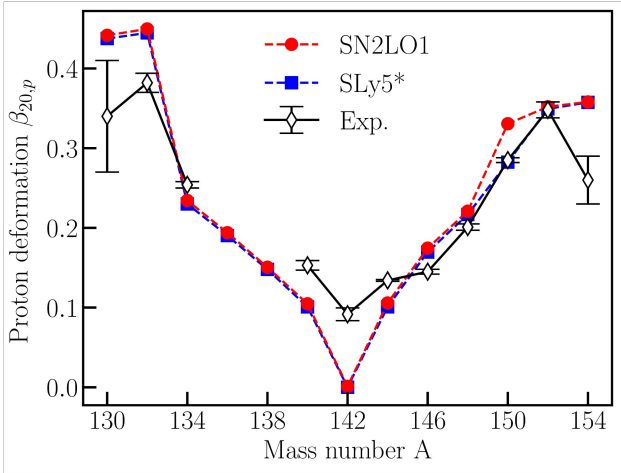
Two DOF: (β_{20}, β_{22}) or (β, γ)

Some remarks:

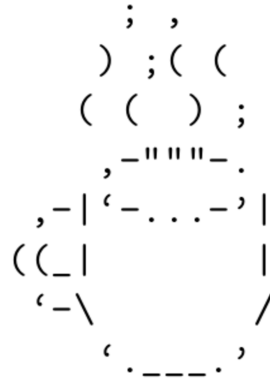
- multipole expansion of the nuclear **volume**
- a priori **all** are non-zero
- these are **outputs** of the models
- but they are **inputs** for mic-mac!



III. Modelspace: deformation and radii



III. Modelspace: MOCCa



W. R. PhD Thesis, ULB (2016).
 W. R. et. al., PRC 92, 064318 (2015).
 W. R. et. al., EPJA 55, 93 (2019).
 W.R. and M. Herbst, in preparation.
 W.R., in preparation.

Physics

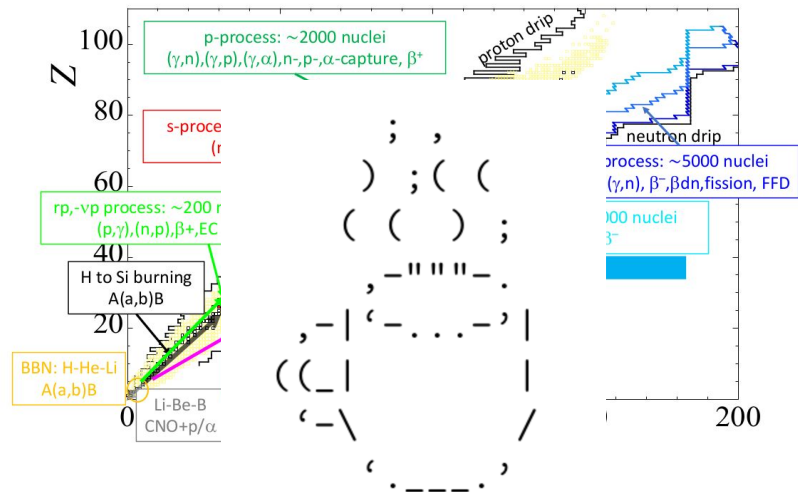
- flexibility w.r.t. **imposed symmetries**
- **unique tool** for the study of:
 - **exotic** shapes
 - **fission** properties
 - **rotational** bands
 - **electromagnetic** form factors
 - **odd-mass and odd-odd** nuclei

Technical aspects

- **3D coordinate space** representation
- developed algorithms
 - speed
 - stability (constraints, blocked states, ...)
 - numerical precision
- **automated implementation** of Skyrme EDFs

IV. The view from Brussels

- Skyrme EDF with:
 - no symmetry-restoration
 - **rotational** correction
 - **finite-size** of nucleons
- 20-25 parameters
- Fitting protocol
 - 2457 known masses
 - 884 charge radii
 - infinite nuclear matter properties
- **global** description (< 0.8 MeV rms)
- Systematic tables of 1000's of nuclei:
 - **masses, radii, deformations,**
 - **fission** properties
 - nuclear **level densities**
 - **strength** functions
- Input for:
 - r-process calculations (S. Goriely)
 - neutron star structure (N. Chamel)

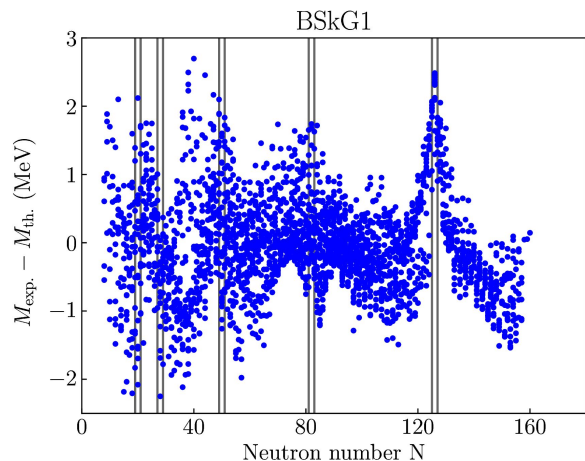


$$E_{\text{tot}} = E_{\text{HFB}} + E_{\text{corr}},$$

$$E_{\text{HFB}} = E_{\text{kin}} + E_{\text{Sk}} + E_{\text{pair}} + E_{\text{Coul}} + E_{\text{cm}}^{(1)},$$

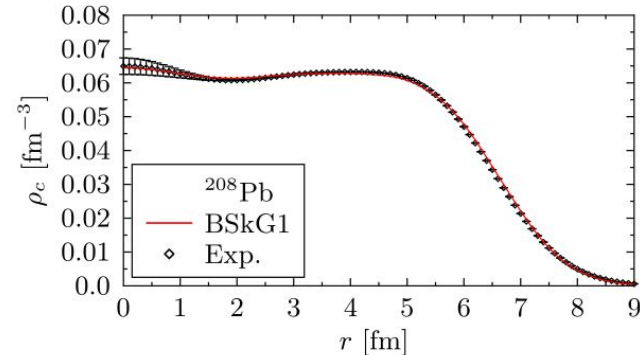
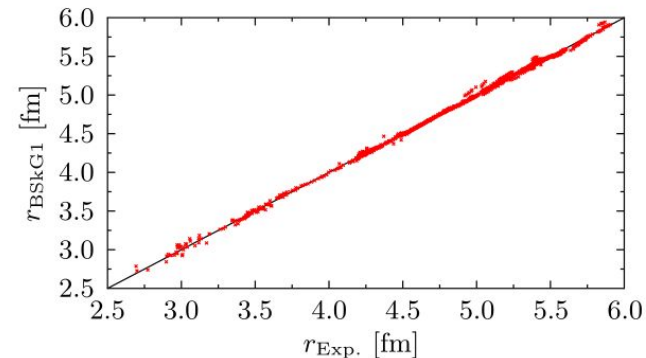
$$E_{\text{corr}} = E_{\text{rot}} + E_{\text{cm}}^{(2)} + E_{\text{W}}.$$

IV. The view from Brussels: **BSkG1/2**



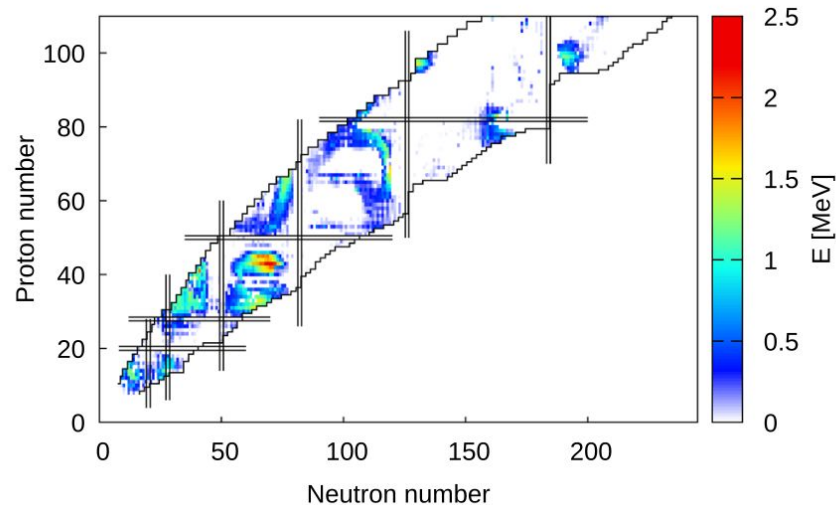
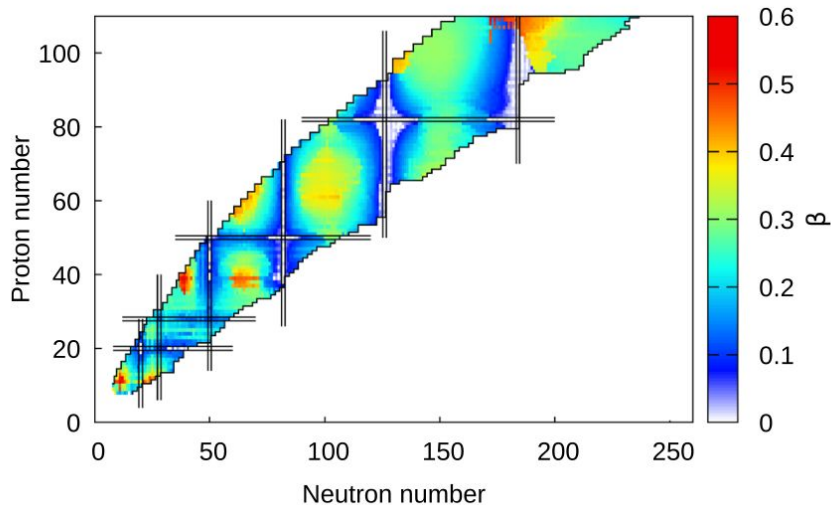
Global performance

$\sigma^2(\text{MeV})$	BSkG1	BSkG2	HFB-14	FR(L)DM
Masses	0.741	0.678	0.729	0.560
Charge radii (fm)	0.027	0.027	0.039	0.038
Primary barriers	0.87	0.45	0.61	0.79
Secondary barriers	0.86	0.45	0.70	1.35
Isomers	0.45	0.48	0.93	1.04



BSkG1: G. Scamps et al., EPJA 57, 333 (2021);
HFB-14: S. Goriely et al., PRC **75**, 064312 (2007).
FRDM: P. Möller et al., At. Data Nucl. Data Tables, **109-110** (2016).

IV. The view from Brussels: **BSkG1/2**



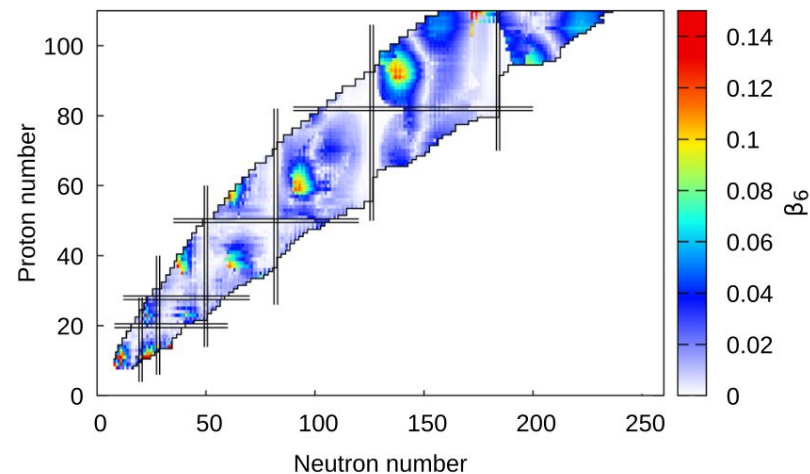
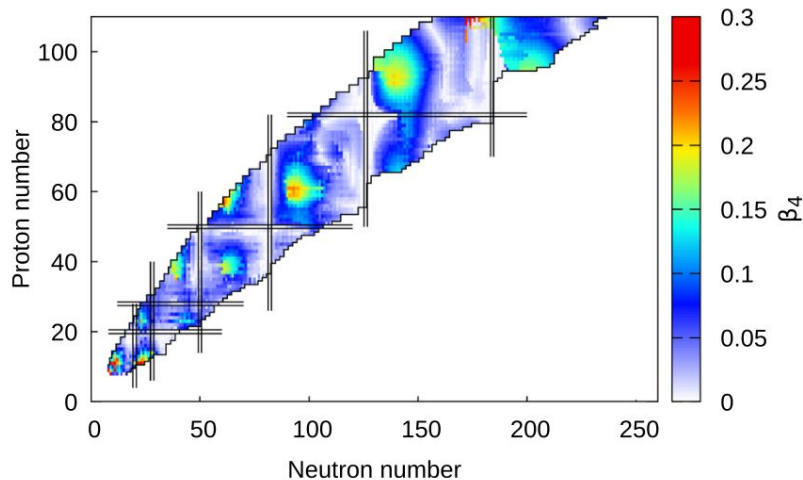
Quadrupole deformation

- quadrupole is everywhere
- and can get rather large!

Triaxial deformation

- lots of nuclei
- significantly **more** than in mic-mac

IV. The view from Brussels: **BSkG1/2**



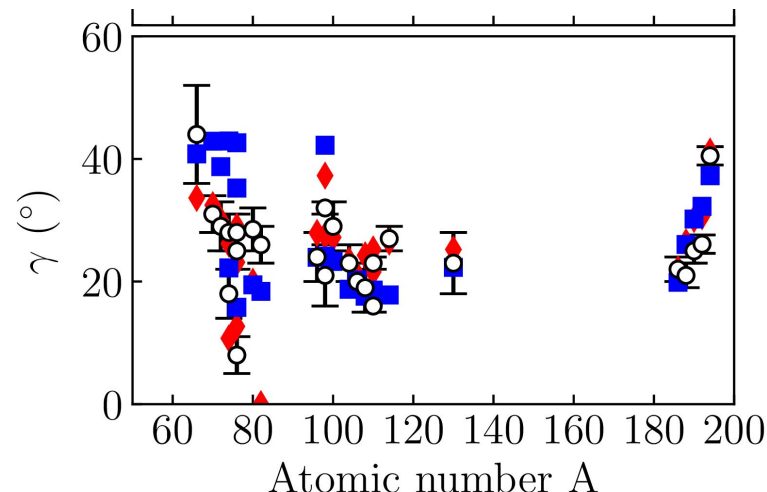
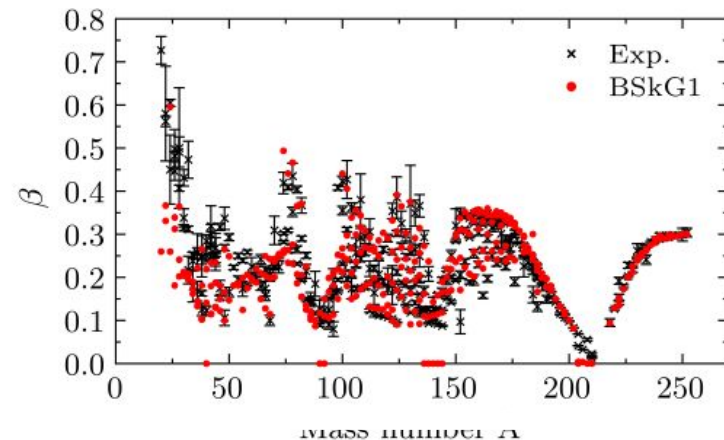
Higher order multipoles

- **always** present
- though less important
- but not necessarily for **fission!**

IV. The view from Brussels: **BSkG1/2**

- average β from B(E2) systematics
- average γ from COULEX

	Exp.		BSkG1	
	β	γ	β	γ
^{96}Mo	0.17	24(4)	0.19	24
^{98}Mo	0.17	32(1)	0.22	24
^{100}Mo	0.23	29(4)	0.25	23



IV. The view from Brussels: **the rotational correction**

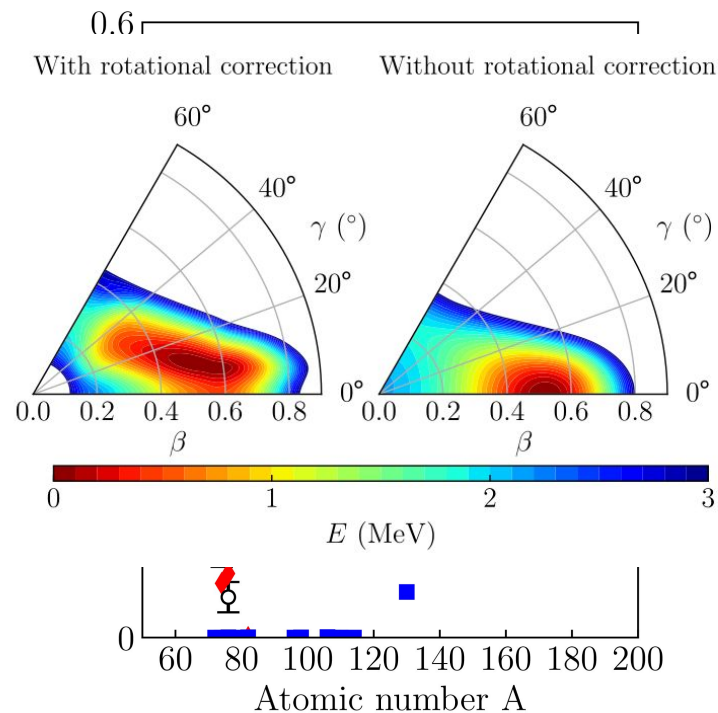
Rotational correction

- aims to mimic beyond-MF effects
- assumes rotational motion
- with consistent moments of inertia
- some parameters

Semi-variational approach:

1. construct surface in (β, γ)
2. add rotational energy
3. select minimum

$$E_{\text{tot}} = E_{\text{HFBD}} + E_{\text{corr}}, \quad E_{\text{rot}} \sim - \sum_{\mu=x,y,z} \frac{\langle J_{\mu}^2 \rangle}{2\mathcal{I}_{\mu}}.$$



V. Some Ru's, Rh's, Pd's: masses and radii

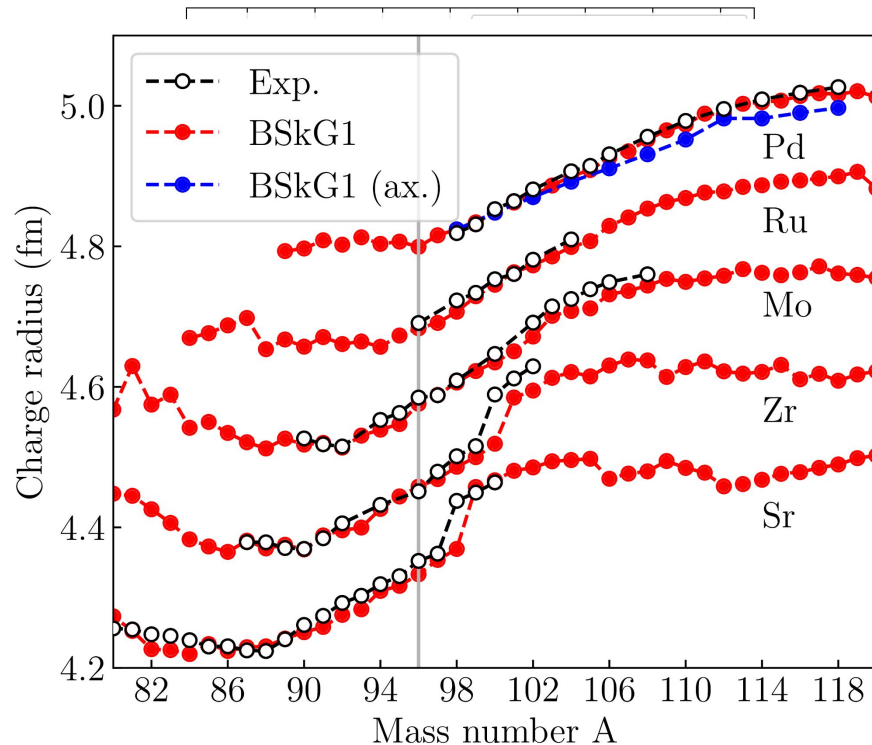
Performance in the region

- **absolute** E about 700 keV
- triaxial model agrees **well** for S2ns
- all older **axial** models are less good
- **absolute** charge radii well-reproduced

M. Hukkanen et al., in preparation.

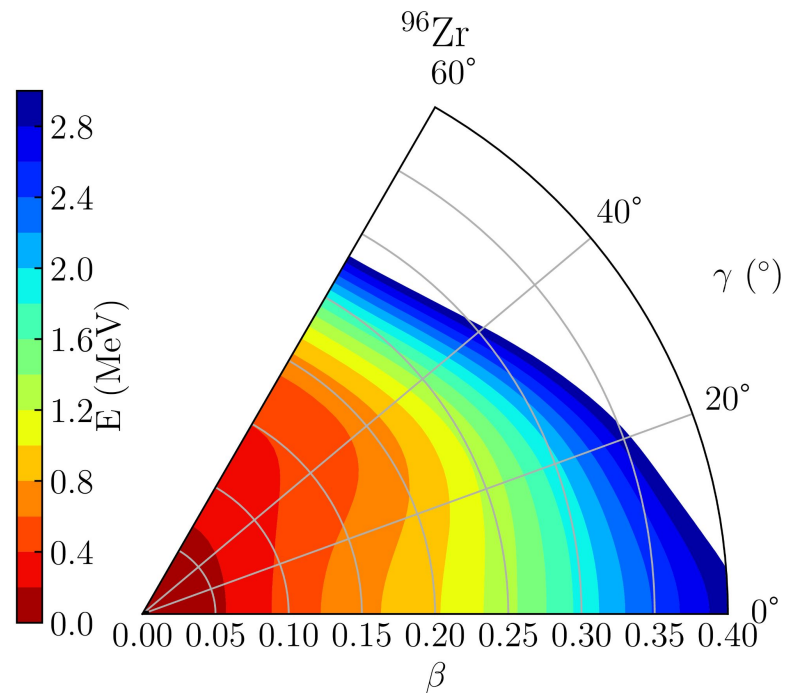
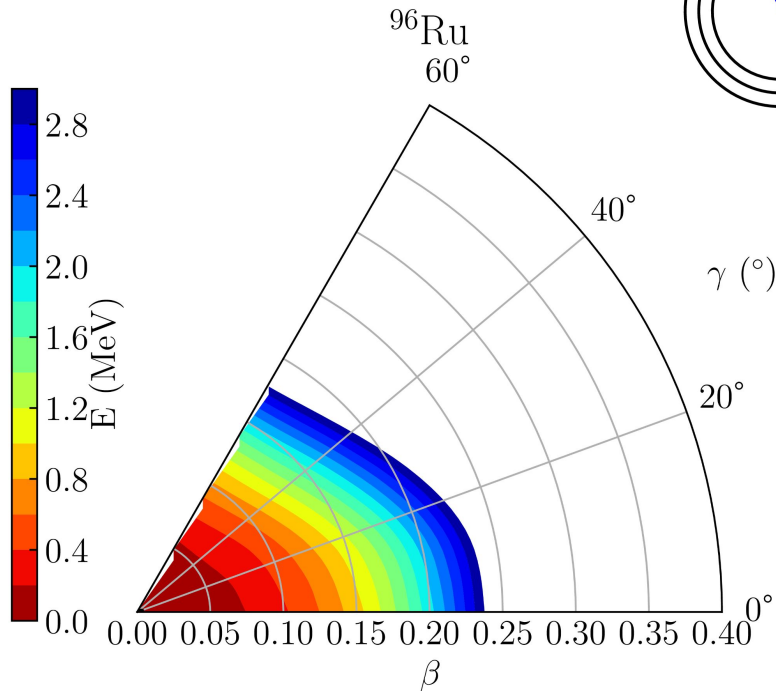
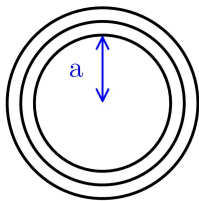
I. Angeli and K. P. Marinova, *At. Data Nuc. Data Tables*, **99** (2015).

S. Geldhof et al., *PRL* **128**, 152501 (2022).



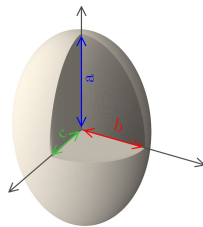
V. and A=96

Spherical

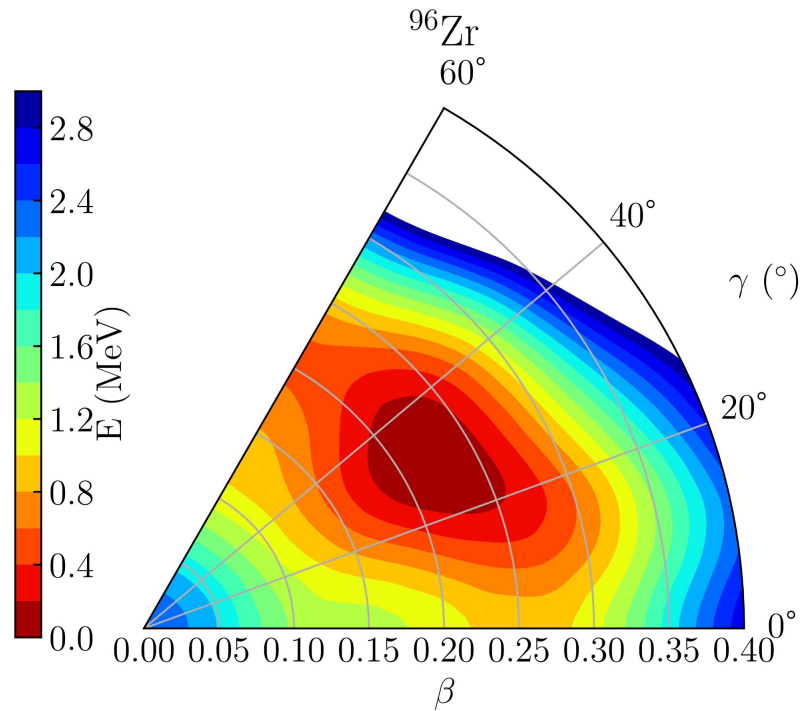
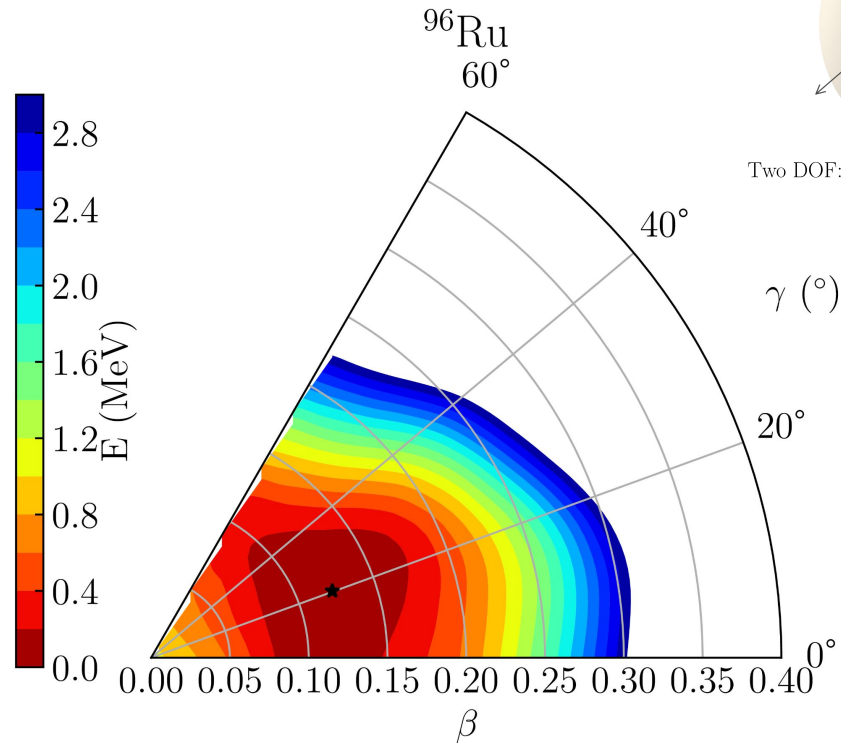


without rotational correction.

V. and A=96



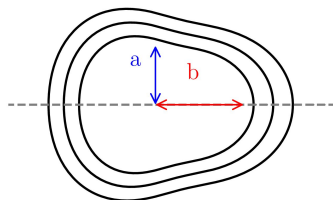
Two DOF: (β_{20}, β_{22}) or (β, γ)



with rotational correction.

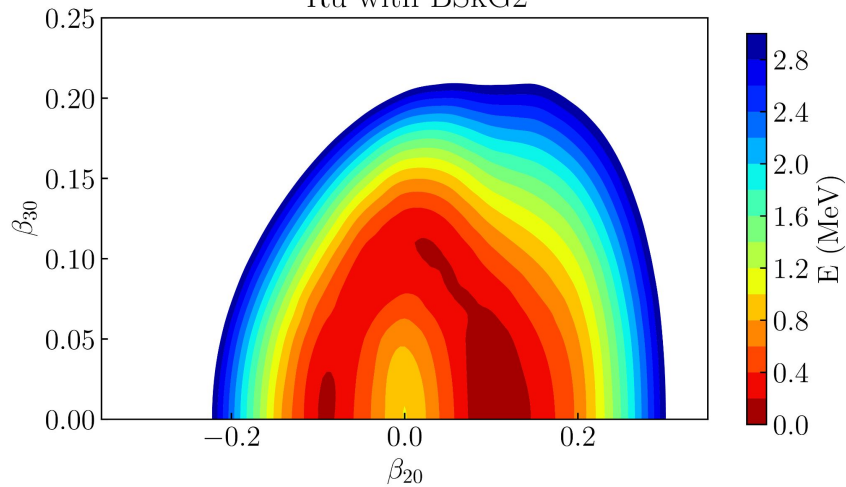
V. and A=96

Reflection-asymmetric (RA)

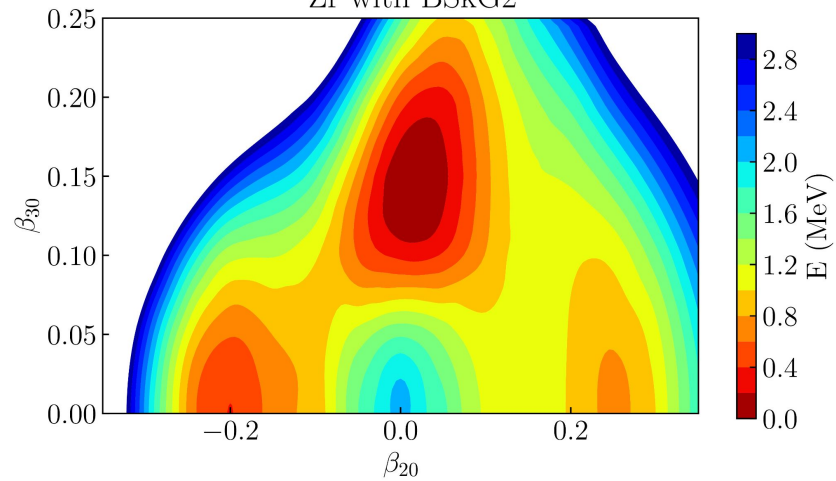


Two DOF: β_{20} , β_{30}

^{96}Ru with BSkG2

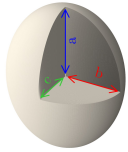


^{96}Zr with BSkG2



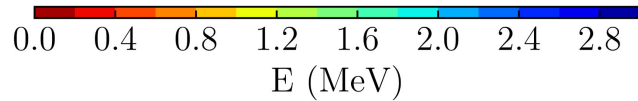
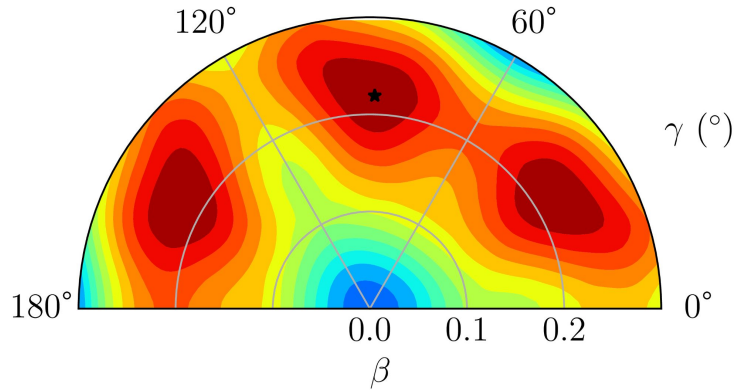
V. and A=96

Triaxial

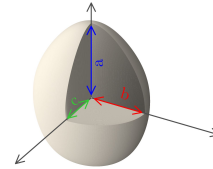


$^{96}\text{Zr}, \beta_{30} = +0.000$

Two DOF: (β_{20}, β_{22}) or (β, γ)

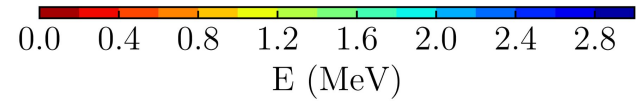
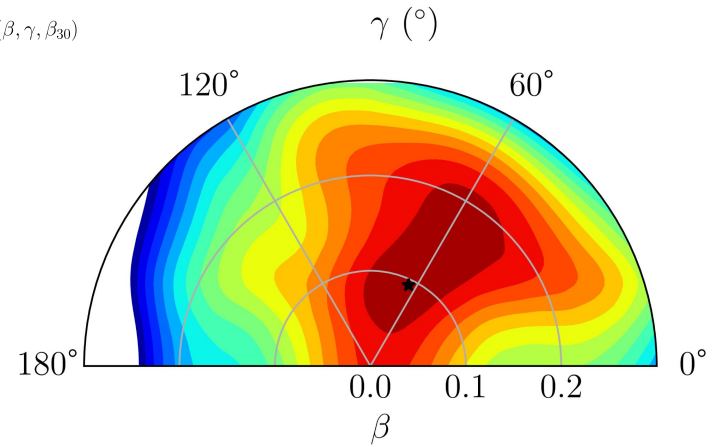


Triaxial, RA



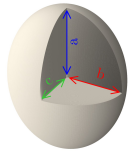
$^{96}\text{Zr}, \beta_{30} = +0.125$

Three DOF: $(\beta, \gamma, \beta_{30})$



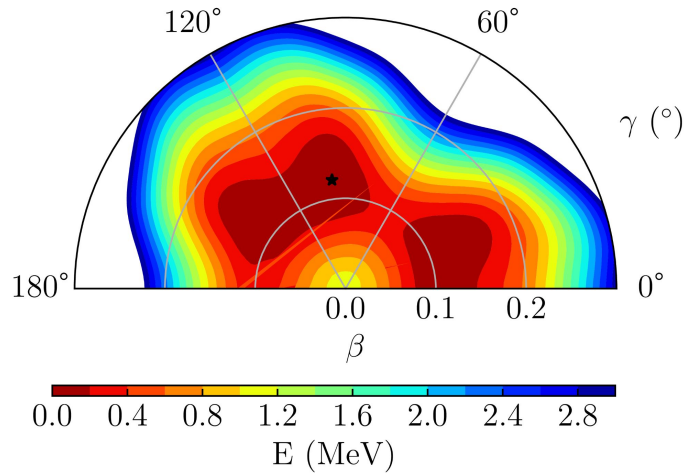
V. and A=96

Triaxial

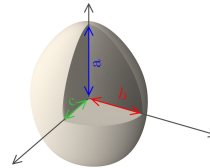


$^{96}\text{Ru}, \beta_{30} = +0.000$

Two DOF: (β_{20}, β_{22}) or (β, γ)

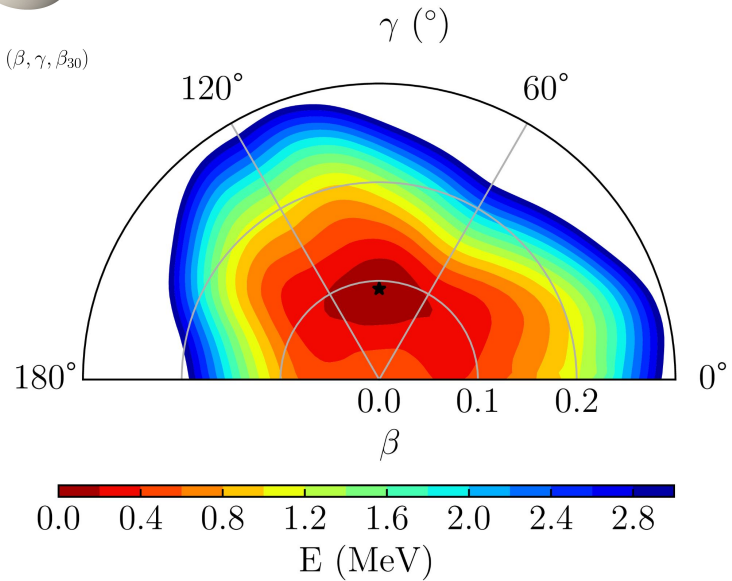


Triaxial, RA

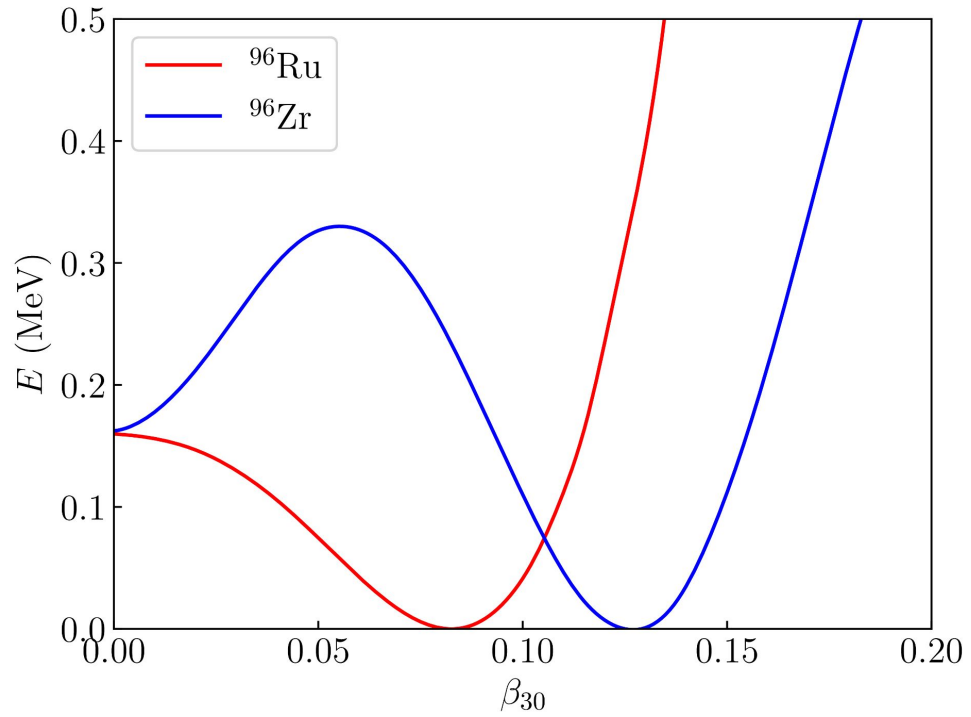


$^{96}\text{Ru}, \beta_{30} = +0.075$

Three DOF: $(\beta, \gamma, \beta_{30})$



V. and A=96



Conclusions

Skyrme Energy Density Functional

- simplicity as **prime quality**
- has been pushed very far
- many variants, many extensions sought for

Symmetries & deformations

- symmetry breaking captures correlations
- with simple wave-functions
- symmetry restoration highly desirable
- but many things can be studied without it

The Brussels Models

- excellent description: masses, radii, INM, ...
- geared for large-scale applications
- recent push towards exotic deformations
- pheno. corrections beyond mean-field

Ru, Rh, Pds and A=96

- overall good description of region
- unexpected ground-state deformations for ^{96}Ru and ^{96}Zr

