The Role of Baryon Structure in Dense Nuclear Matter

Theo F. Motta (JLU Gießen)

@ EMMI Workshop, Kitzbühel

September 14, 2022

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The Quark-Meson Coupling Model

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• The QMC model is a relativistic phenomenological models. What we want, fundamentally, is

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- However, we also want to retain some information on the *baryon structure* without solving QCD.
- We choose to model baryon-baryon interactions as a quark-meson interaction in the subhadronic level.



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Bag Model

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• Solve Dirac with the appropriate boundary conditions

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• In MFA we have:

$$(i\partial - m^{\star})\psi = 0$$

where $m^* = m - g^q_\sigma \bar{\sigma}$

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$$\mathcal{L} = \bar{\Psi}_B (i\partial \!\!\!/ - \mathcal{M}_B) \Psi_B + g^B_\sigma \sigma \bar{\Psi} \Psi$$

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$$\mathcal{L} = \bar{\Psi}_B (i\partial \!\!\!/ - M_B) \Psi_B + g^B_\sigma \sigma \bar{\Psi} \Psi - \frac{d}{2} (g^B_\sigma \sigma)^2 \bar{\Psi} \Psi + \cdots$$

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Some QMC Highlights

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Some QMC Highlights



Figure: Nuclei binding energies from (Martinez et al 2019)

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Some QMC Highlights



Figure: Stellar structure results from (Motta et al 2019) plus recent NICER results

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Content

About the particle content, the QMC model makes a few interesting claims.



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Delta Isobars



Chemical potentials in equilibrium. Say, after $\mu_n = \mu_\Lambda$ they merge and the Λ s start to populate the system.

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Delta Isobars



It so happens that, due to the natural repulsion that arises from the QMC model, the $\mu_{n} + \mu_{e}$ combination never reaches μ_{Λ} .



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Detla Meson

$$\begin{split} M_B^{\star}(\bar{\sigma},\bar{\delta}) = & M_B - g_{\sigma}\bar{\sigma} + \frac{d}{2} (g_{\sigma}\bar{\sigma})^2 \\ & - t_B^{\delta} g_{\delta} I_B \bar{\delta} + \tilde{t}_B^{\delta} (g_{\delta}\bar{\delta})^2 + \tilde{d} g_{\sigma} g_{\delta} \bar{\sigma} I_B \bar{\delta} \end{split}$$

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Detla Meson

$$\begin{split} \mathcal{M}_{B}^{\star}(\bar{\sigma},\bar{\delta}) = & \mathcal{M}_{B} - g_{\sigma}\bar{\sigma} + \frac{d}{2} \left(g_{\sigma}\bar{\sigma}\right)^{2} \\ & - t_{B}^{\delta}g_{\delta}I_{B}\bar{\delta} + \tilde{t}_{B}^{\delta} \left(g_{\delta}\bar{\delta}\right)^{2} + \tilde{d}g_{\sigma}g_{\delta}\bar{\sigma}I_{B}\bar{\delta} \\ = & \mathcal{M}_{B} - g_{\sigma}^{B}(\bar{\sigma},\bar{\delta})\sigma - g_{\delta}^{B}(\bar{\sigma},\bar{\delta})I_{B}\bar{\delta}. \end{split}$$



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DM Capture

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• Assume DM particles collide with NS matter and, in doing so, become gravitationally trapped (DM Capture).

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- Assume DM particles collide with NS matter and, in doing so, become gravitationally trapped (DM Capture).
- The N-DM form factor is usually taken to be a constant; calculated at zero momentum transfer.
- However, given the density of a NS, the neutrons have very high momenta. A constant form factor is a poor approximation.
- Furthermore, the in-structure of the nucleon is not only relevant at high momentum transfer, but it's also *modified* by the medium.

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Form Factors

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Form Factors

Including the effective mass dependence we already have

$$c_{n}^{S}\left(m_{n}^{\text{eff}}\right) = \frac{2(m_{n}^{\text{eff}})^{2}}{\Lambda^{4}v^{2}} \left[\sum_{q=u,d,s} f_{T_{q}}^{(n)} + \frac{2}{9}f_{T_{G}}^{(n)}\right]^{2} \\ c_{n}^{P}\left(m_{n}^{\text{eff}}\right) = \frac{2(m_{n}^{\text{eff}})^{2}}{\Lambda^{4}v^{2}} \left[\sum_{q=u,d,s} \left(1 - 3\frac{\bar{m}}{m_{q}}\right)\Delta_{q}^{(n)}\right]^{2} \\ c_{n}^{V} = \frac{9}{\Lambda^{4}}, \quad c_{n}^{A} = \frac{1}{\Lambda^{4}} \left[\sum_{q=u,d,s} \Delta_{q}^{(n)}\right]^{2}$$
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(1)

• On top of that, considering the energy scale of the interaction, we introduce the *t* momentum dependence in the standard way

$$c_n^l(t) = \frac{c_n^l}{\left(1 - t/Q_0^2\right)^2}, \quad l \in \{S, P, V, A, T\}$$
 (2)

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Interaction rate



Figure: Normalised differential DM-neutron interaction rate per DM energy loss. Constant neutron coupling $c_n^S(0)$ (light blue line) and w/ transferred momentum form factor dependence (magenta line), $m_{\chi} = 1$ TeV, B = 0.5, $\mu_{F,n} = 0.4$ GeV and $Q_0 = 1$ GeV.

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Capture rate



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