



KTH Teknikvetenskap

Astroparticle physics at KTH, Stockholm

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KTH Teknikvetenskap

Strategy of group and science goals

The main task of the group is to study the high-energy universe through X- and gamma-radiation and cosmic rays.

The fundamental scientific questions addressed concern particle acceleration and radiation processes in cosmic plasmas.

Specific topics:

- Search for dark matter
- Physics of gamma-ray bursts
- The origin of cosmic rays
- High energy emission from compact objects

The focus is on design and development of strategic satellite- and balloon-borne instrumentation. In particular, X/gamma-ray polarimeters.

Astrophysics @ KTH



CAPRICE (1990's)
Cosmic antimatter



Fermi (2008 -)
Cosmic gamma-rays



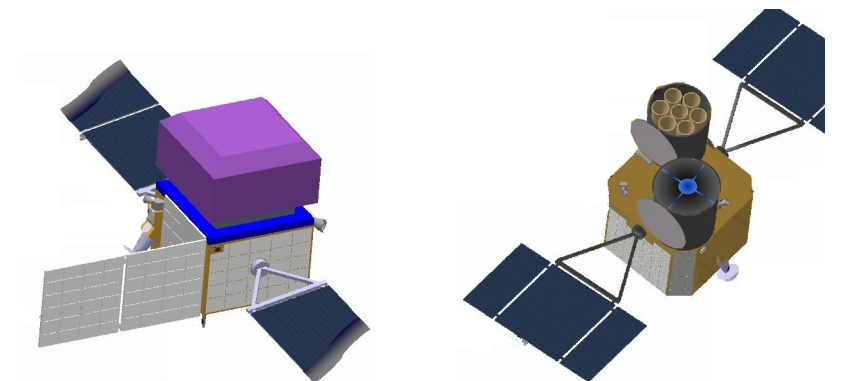
PoGO Lite (now!)
Polarised X-rays



PAMELA (2006 -)
Cosmic antimatter



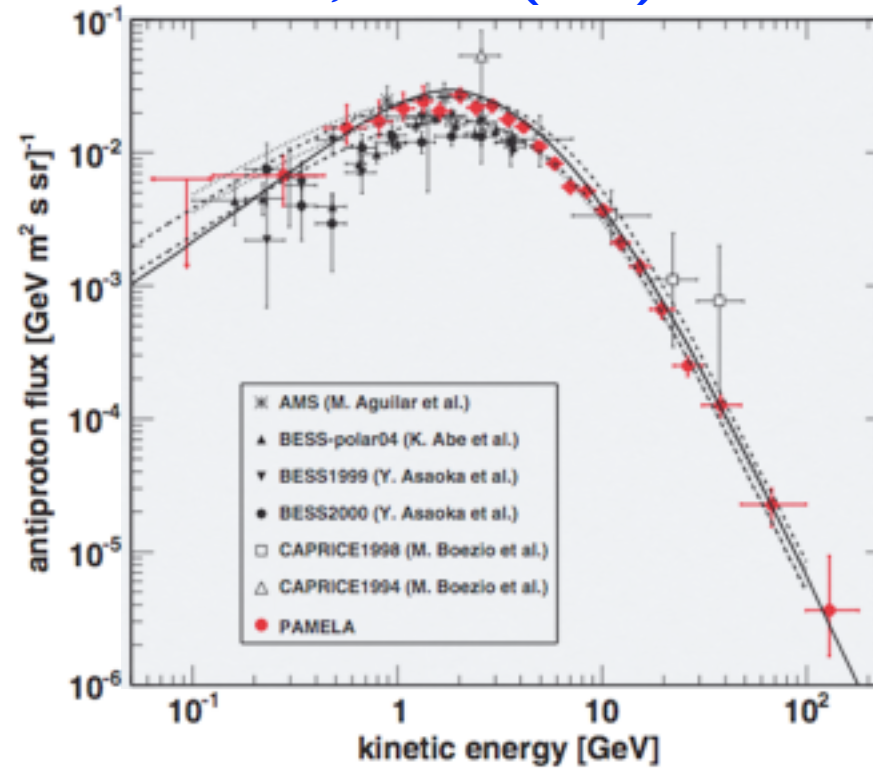
PFD / ISS (2009, 2011)
Education



GRIPS
Proposal

Main recent results achieved from platforms that we have been involved in

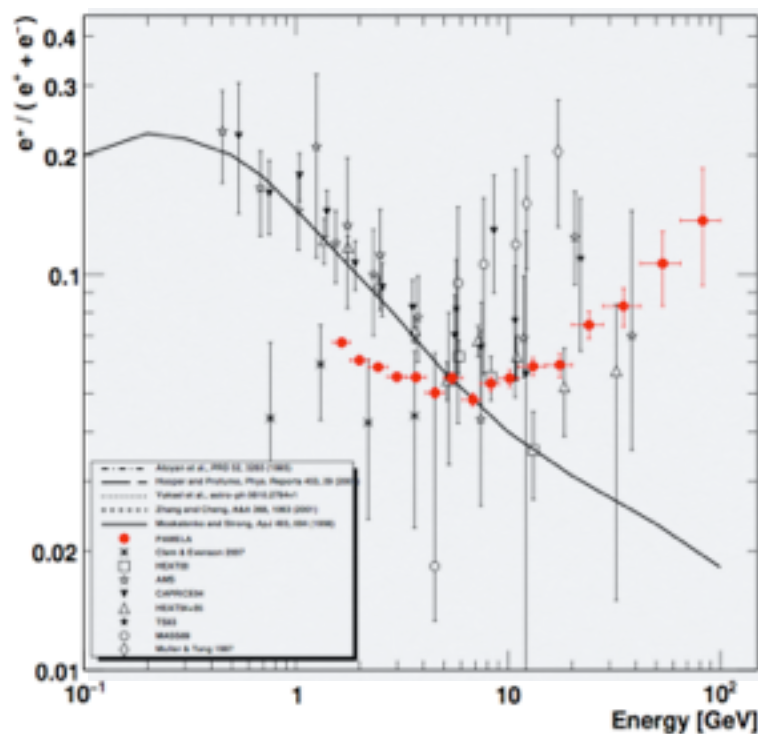
PRL 105, 121101 (2010)



Precision measurement of the cosmic-ray antiproton flux with the PAMELA satellite.

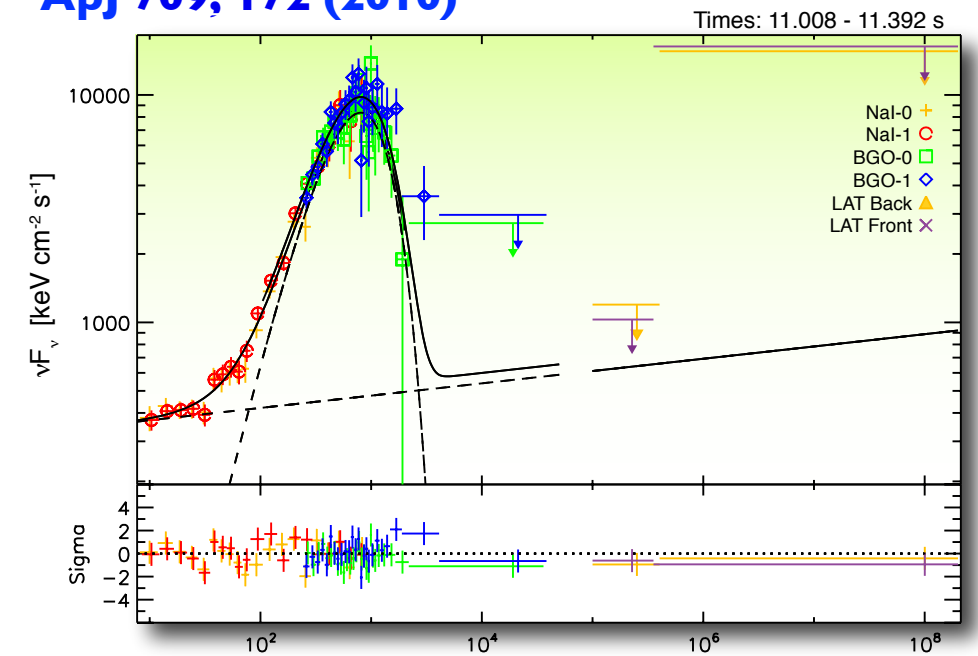
Discovery of an anomalous positron fraction with the PAMELA satellite.

Nature 458, 607 (2009)



Identification of the jet photosphere in GRBs using the *Fermi* satellite

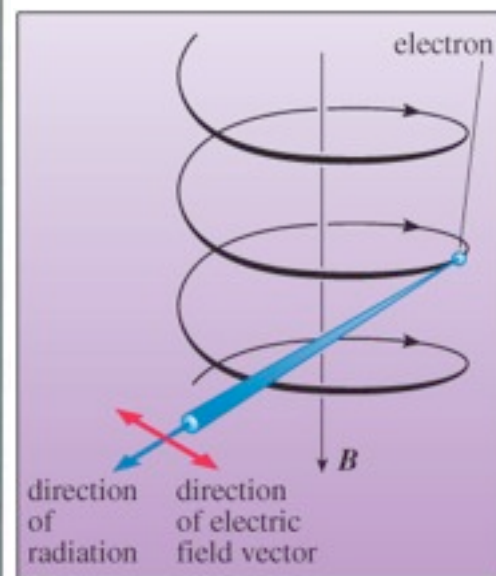
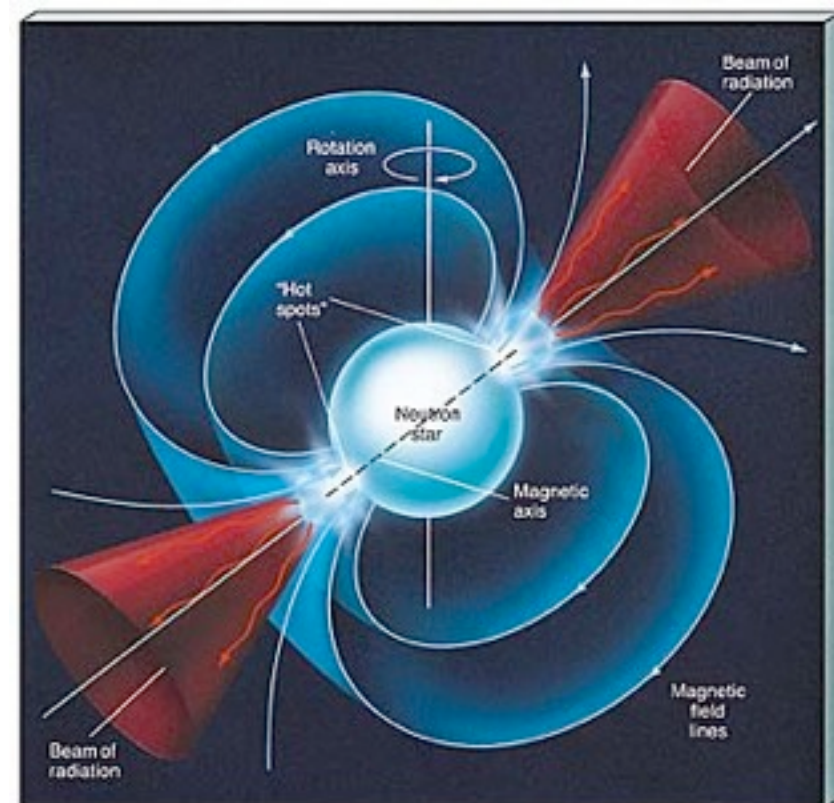
Apj 709, 172 (2010)



PoGO Lite: Targets for a X/γ polarimeter

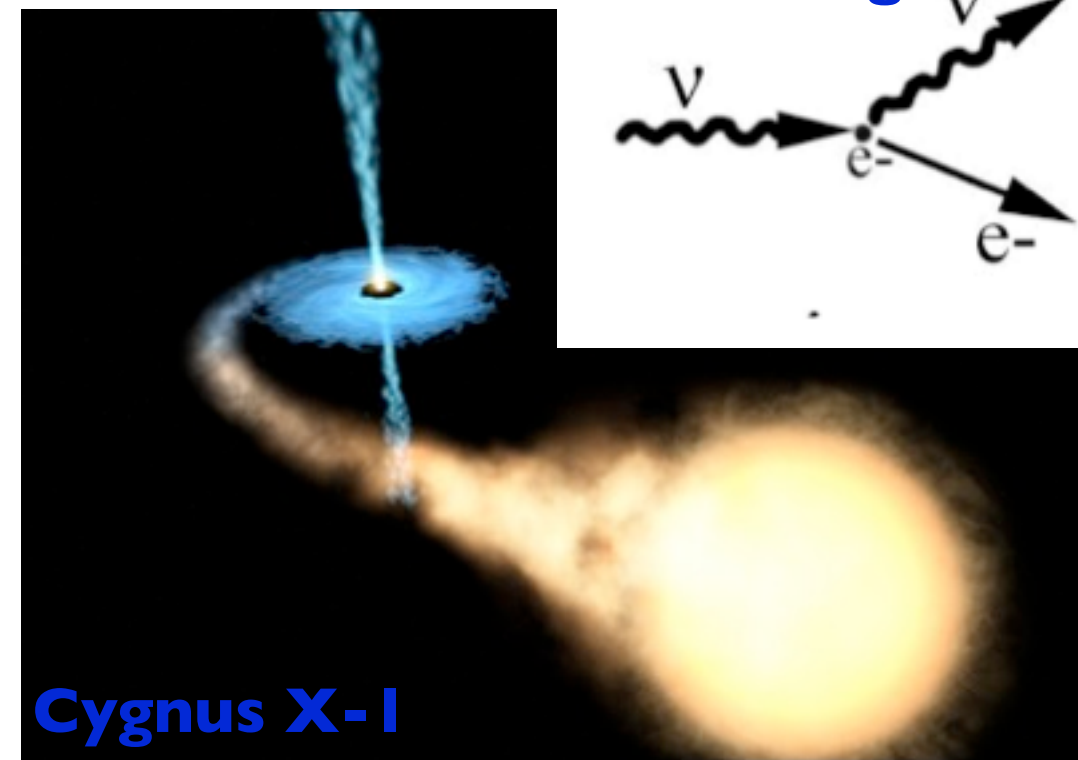
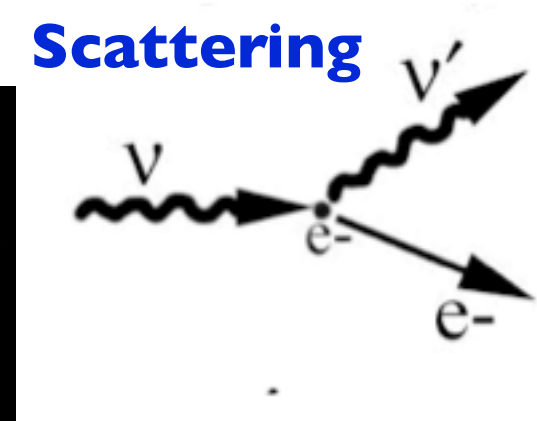
- **Synchrotron emission:**
 - **Rotation-powered neutron stars** (e.g. the Crab pulsar)
 - **Pulsar wind nebulae** (e.g. the Crab nebula)
 - **Jets in active galactic nuclei**
- **Compton scattering:**
 - **Accretion disk around black holes** (e.g. Cygnus X-1)
- **Propagation in strong magnetic field:**
 - **Highly magnetised neutron stars** (e.g. Hercules X-1)

Crab



Synchrotron emission

Scattering



Cygnus X-1

PoGO Lite

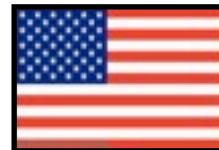


KTH

Stockholm University
DST Control
SSC Esrang



Hiroshima University,
Tokyo Inst. of Tech,
ISAS/JAXA, Waseda Uni.



SLAC - KIPAC
Uni. of Hawaii

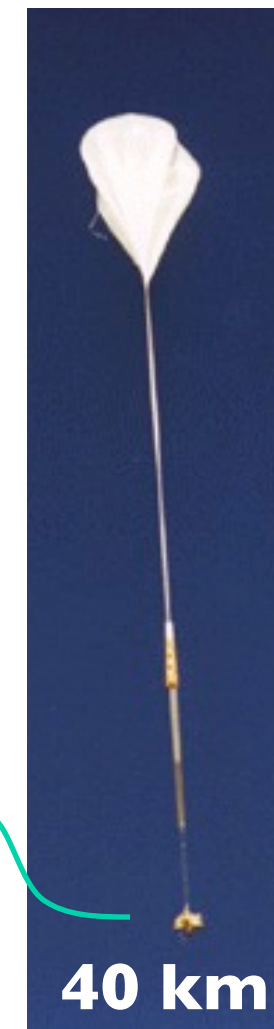


X-rays (Chandra),
optical (HST),
radio (VLA)

- **PoGO Lite** is optimised for **point sources** (e.g. Crab pulsar, Cygnus X-1).
- Measures **10% polarisation** in **200 mCrab sources** in a **6 hour** balloon observation.
- **Maiden flight** of scaled down (less effective area) **'pathfinder'** took place in **July 2011** from Esrang Space Centre, Sweden.
- **Flight terminated** due to a leaking balloon after a few hours. Payload safely recovered. Reflight in summer 2012.

γ

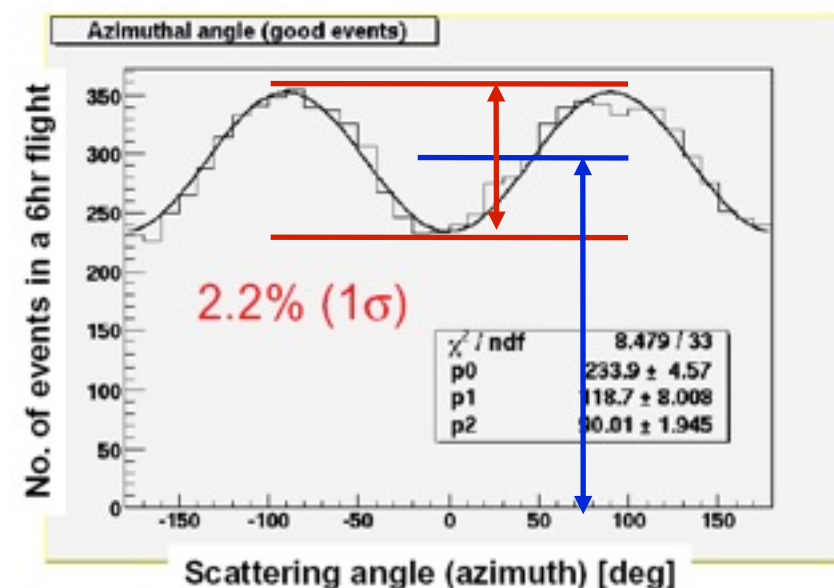
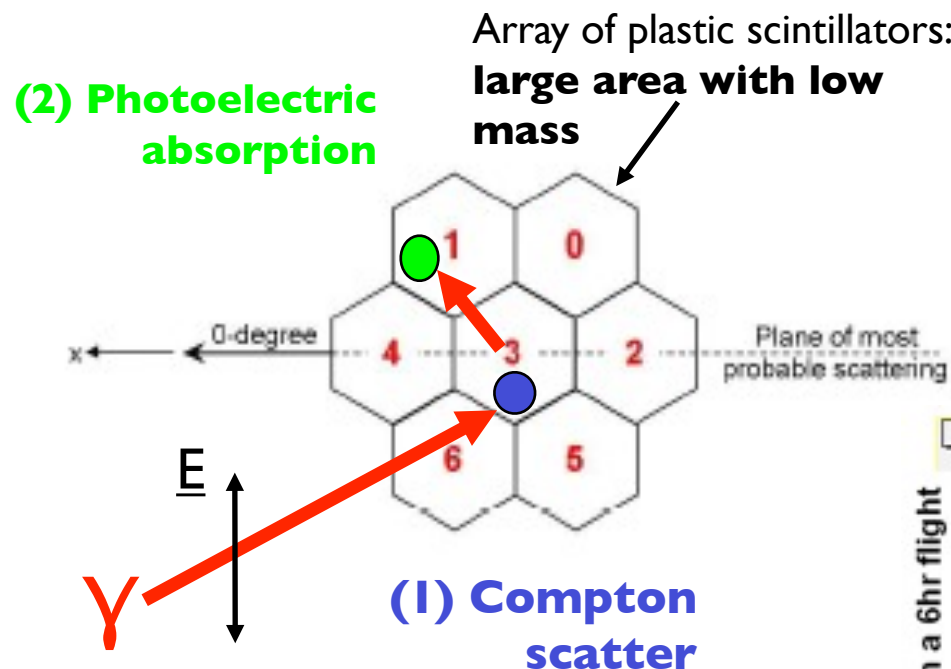
25 – ~100 keV



40 km

PoGOLite Principles

- γ from a **polarised** source undergo **Compton scattering** in segmented detector material
- Higher probability of being **scattered perpendicular** to the **electric field vector**
- Observed **azimuthal scattering angles** are **modulated by polarisation**
- Incident γ deposits little energy at Compton site
- Most deposited at photoelectric absorption site
- Large energy difference
- Can be distinguished by simple plastic scintillators (despite intrinsic poor energy resolution)



MF = difference / average

$$P_{source} = \frac{M_{obs}}{M_{100}}$$

minimum detectable polarisation

signal rate

$$MDP_{99\%} = \frac{4.29}{M_{100} R_S} \sqrt{\frac{R_S + R_B}{T}}$$

background rate

modulation for 100% polarised source

observation time



PoGOLite: Scintillator array

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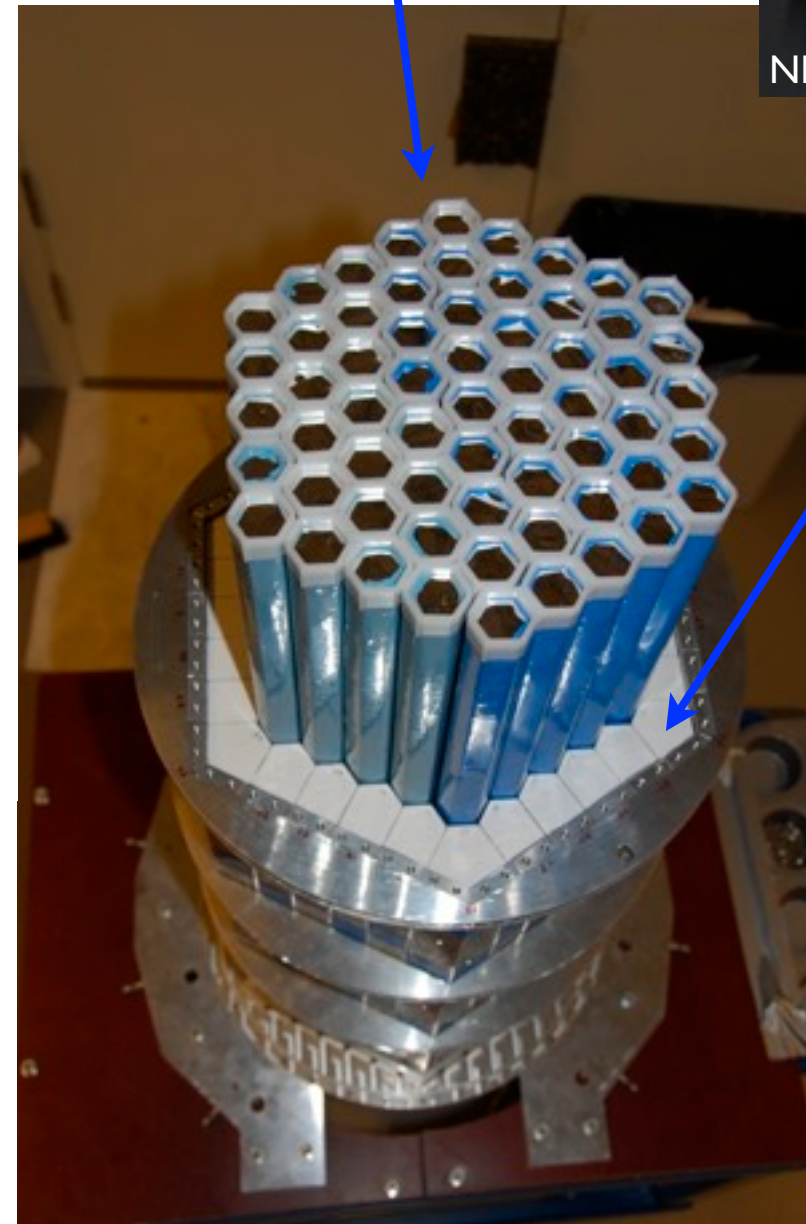
Reflective wrappings:
VM2000 (plastic) / BaSO₄ (BGO)



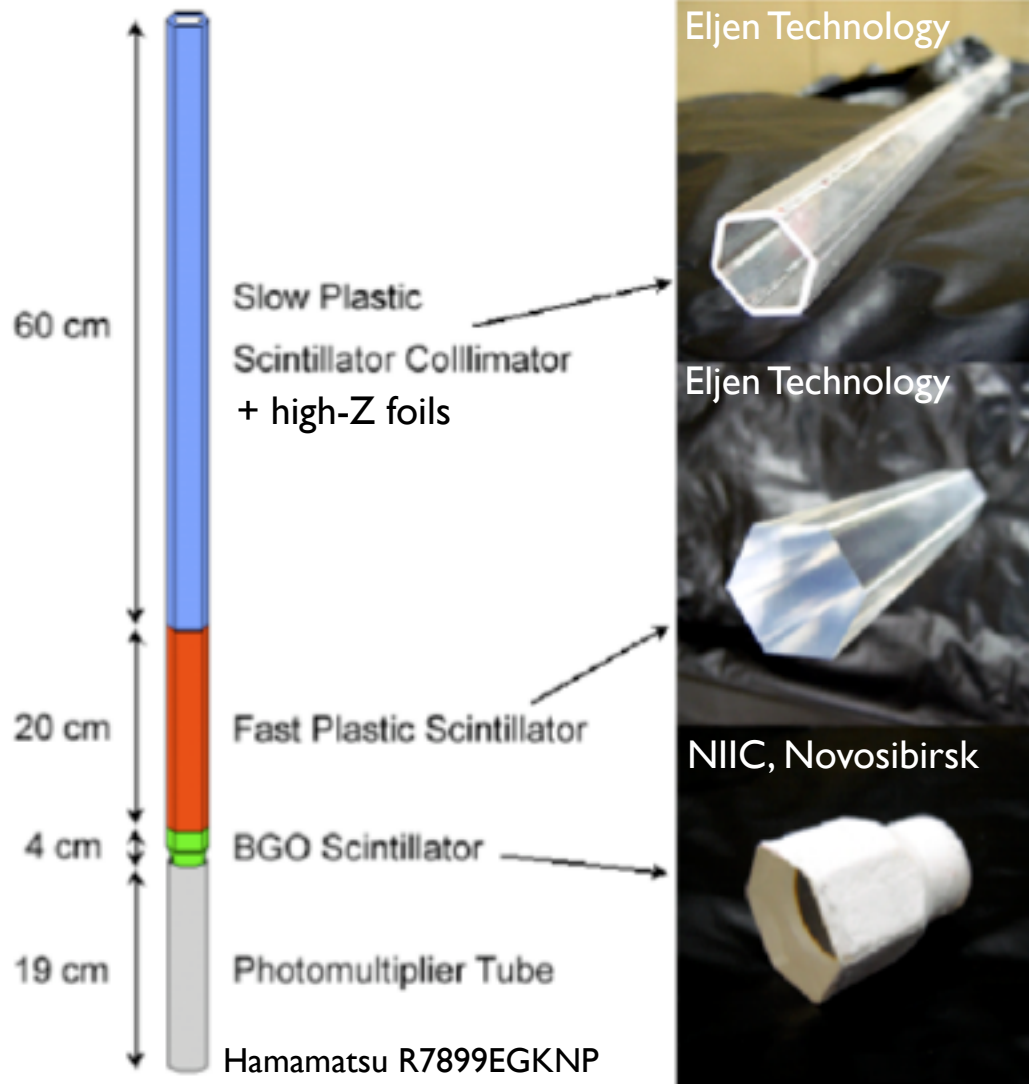
NIIC, Novosibirsk

Side anticoincidence, BGO
~ 150 kg

Phoswich Detector Cells
(~30 kg)



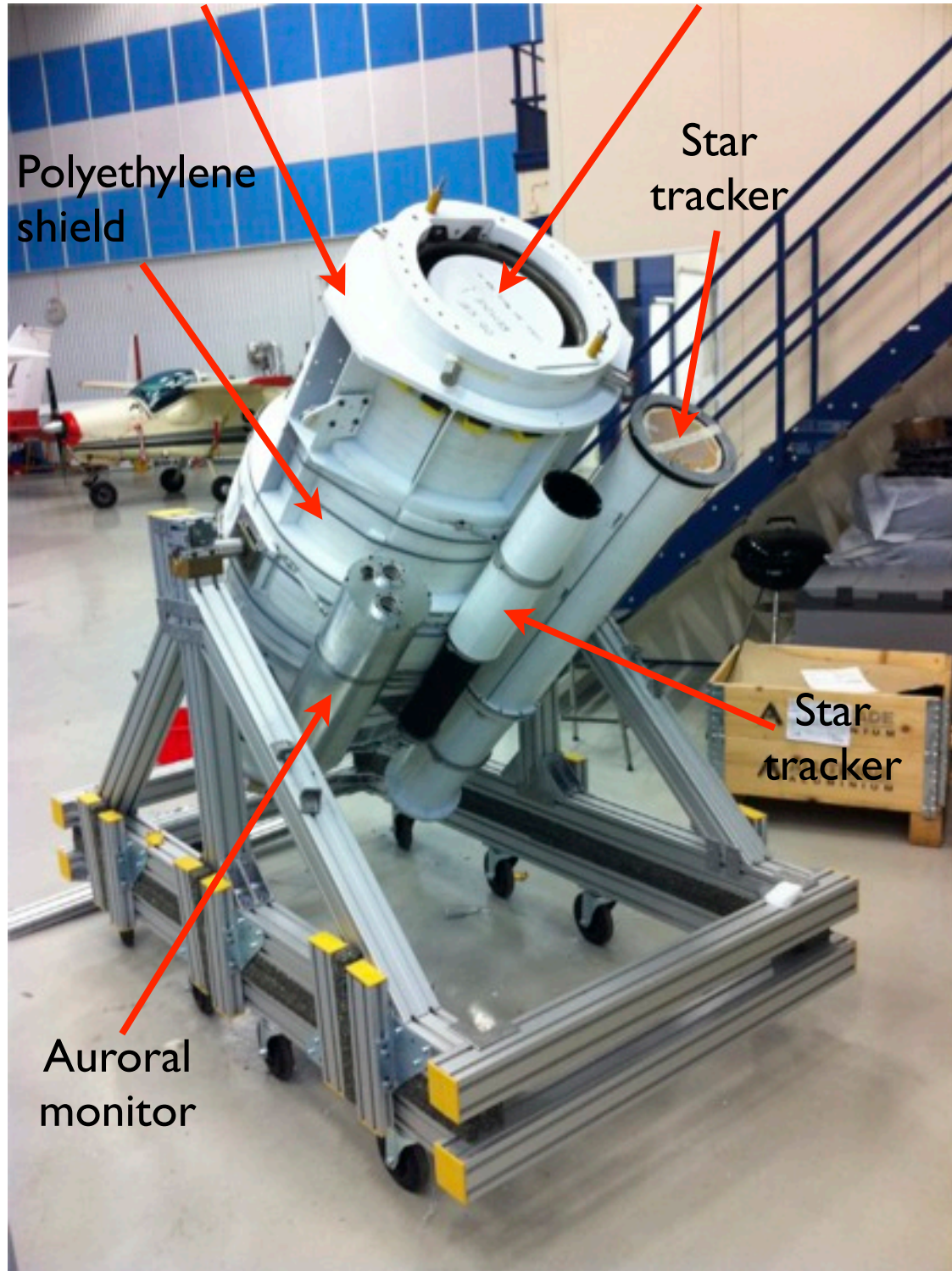
Pathfinder instrument
(61 units)



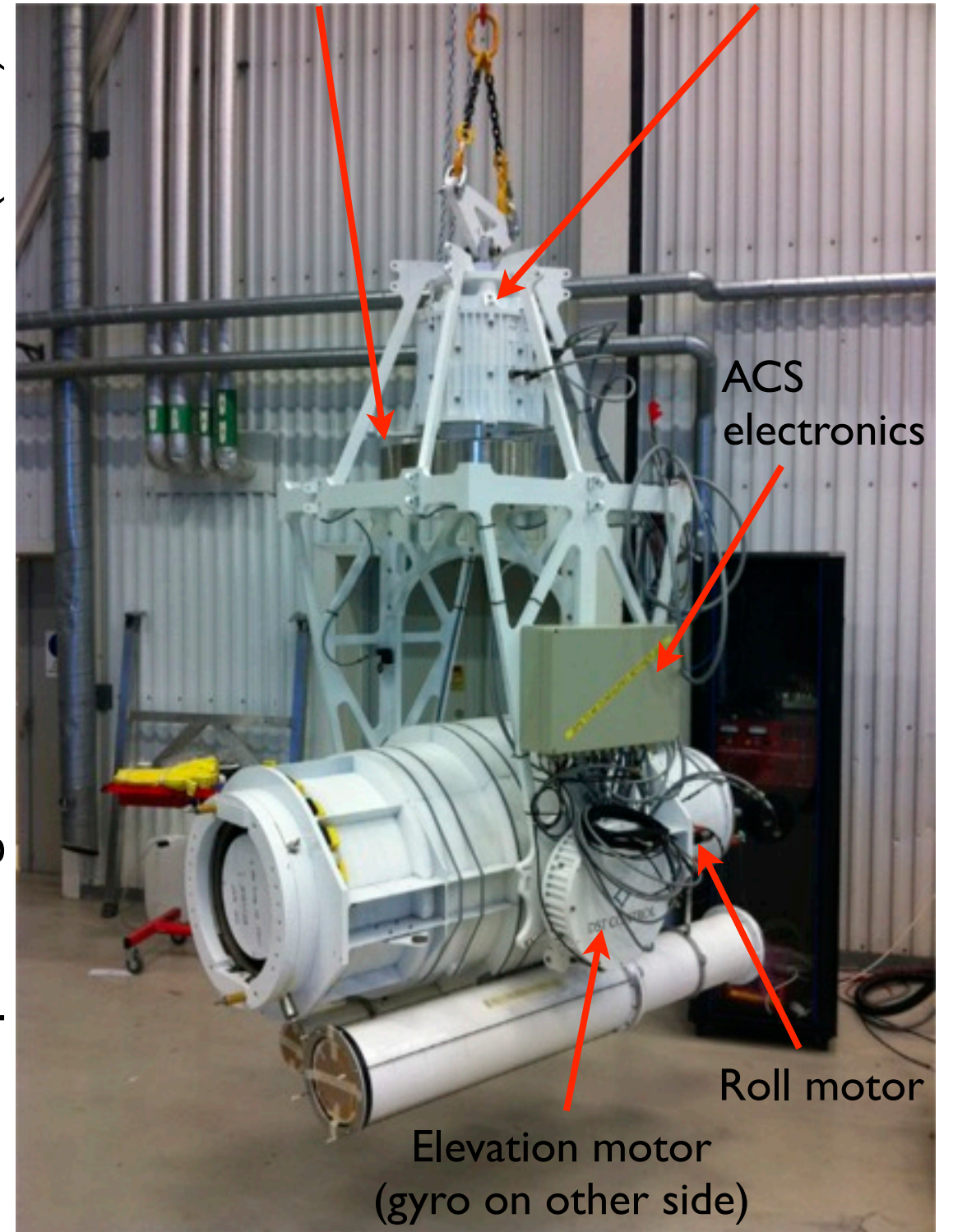
@ 25 keV only 1-3 keV Compton deposited - single p.e. detection
(PMT has 0.05 p.e. ripple @ 10⁶ gain)

PoGOLite: Polarimeter + ACS

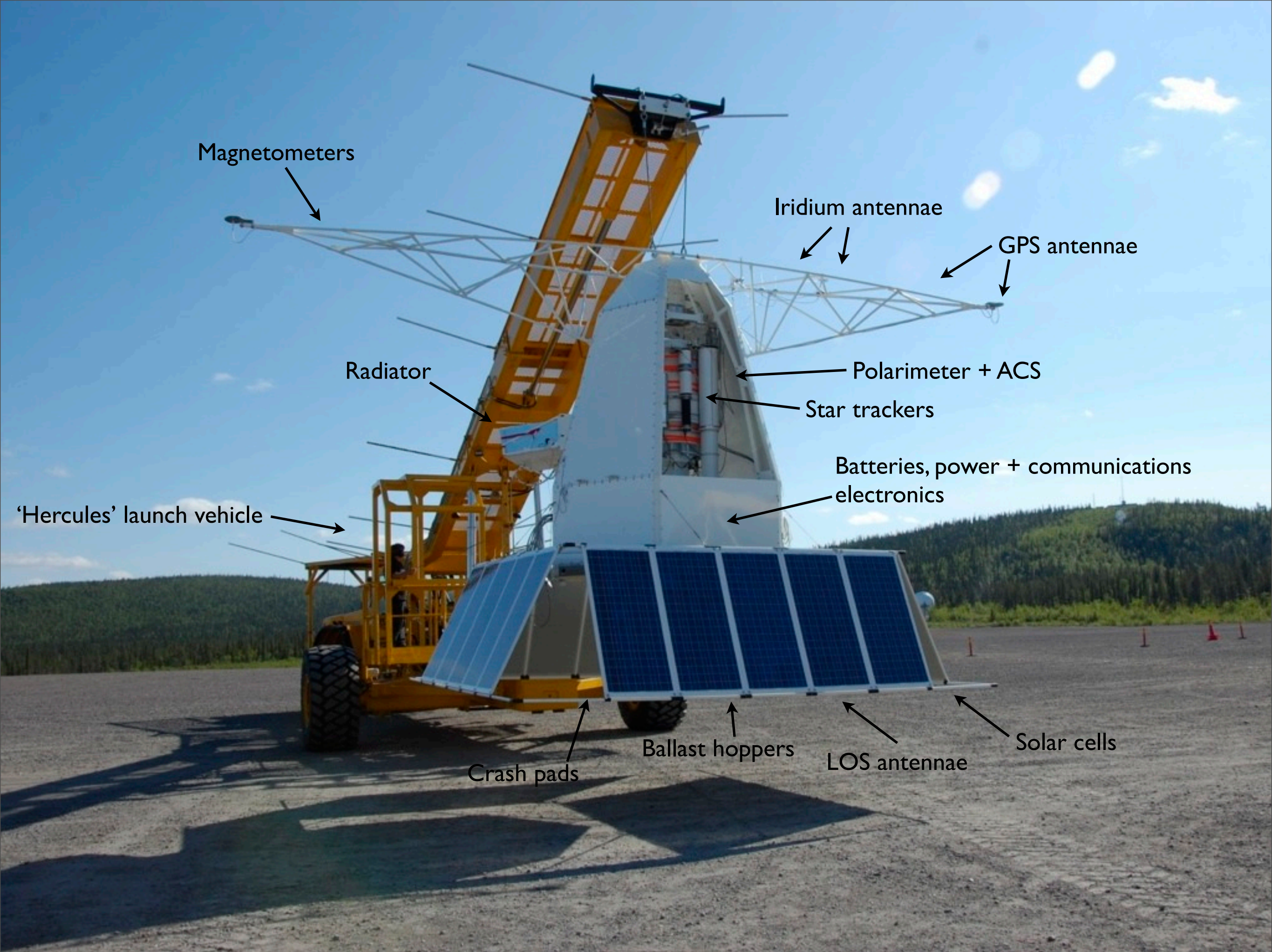
Support frame (stationary) Polarimeter (rotates)



Flywheel Momentum dump



Aim: pointing to better than 5% of FOV ($\sim 0.1^\circ$)



Magnetometers

Iridium antennae

GPS antennae

Radiator

Polarimeter + ACS

Star trackers

Batteries, power + communications electronics

'Hercules' launch vehicle

Crash pads

Ballast hoppers

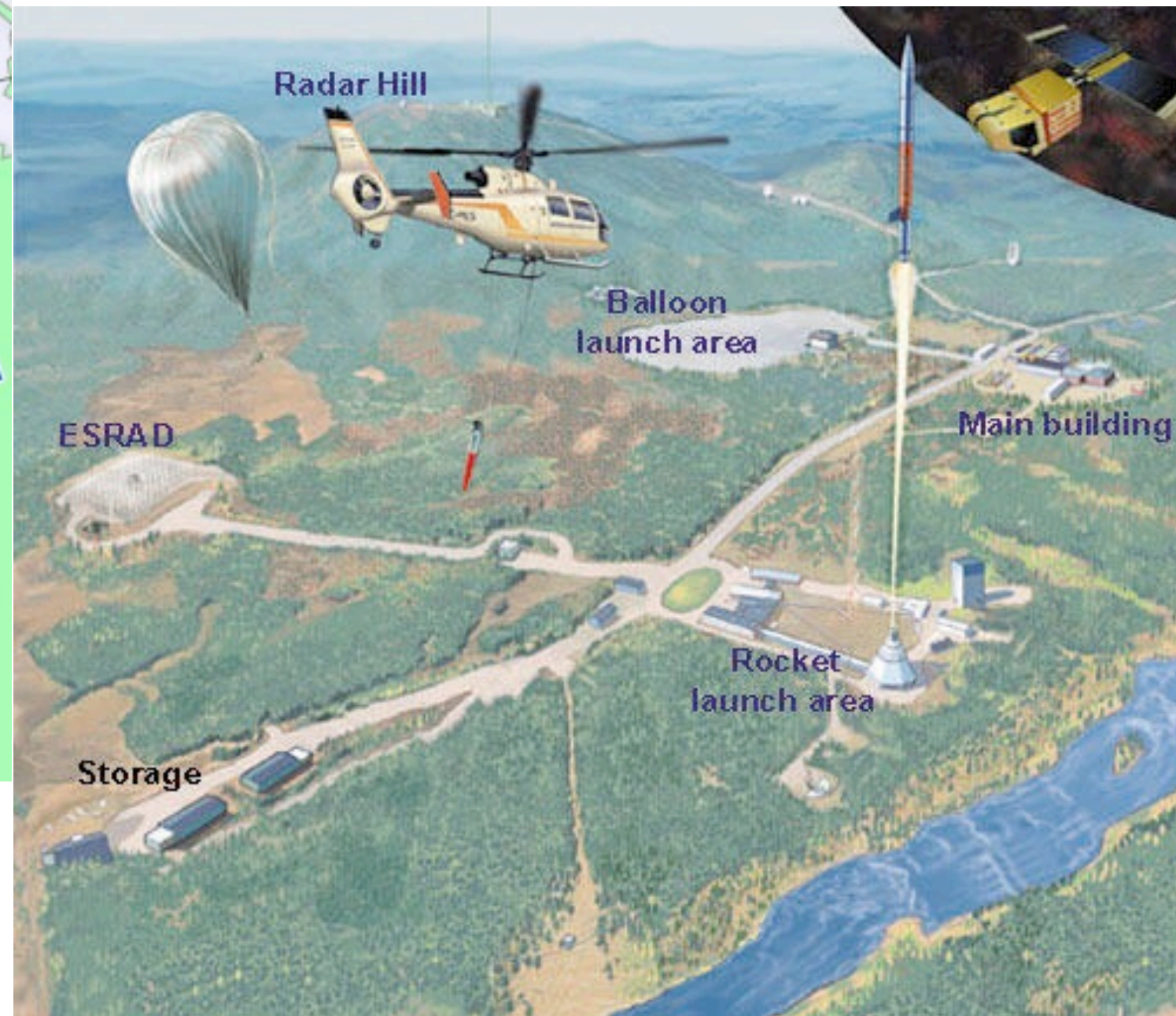
LOS antennae

Solar cells



Monday, 5 December 2011

ESRANGE: European Space Range





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Future work and collaboration

We are interested in a collaborative effort for a future high energy mission, in particular involving a polarimeter

We have valuable experience from balloon missions, in particular from the recent PoGO Lite mission campaigns

We have large experience in balloon launching from ESRANGE, such as planning, design, and logistics