

# Revealing a strong phase transition in neutron star mergers with gravitational waves

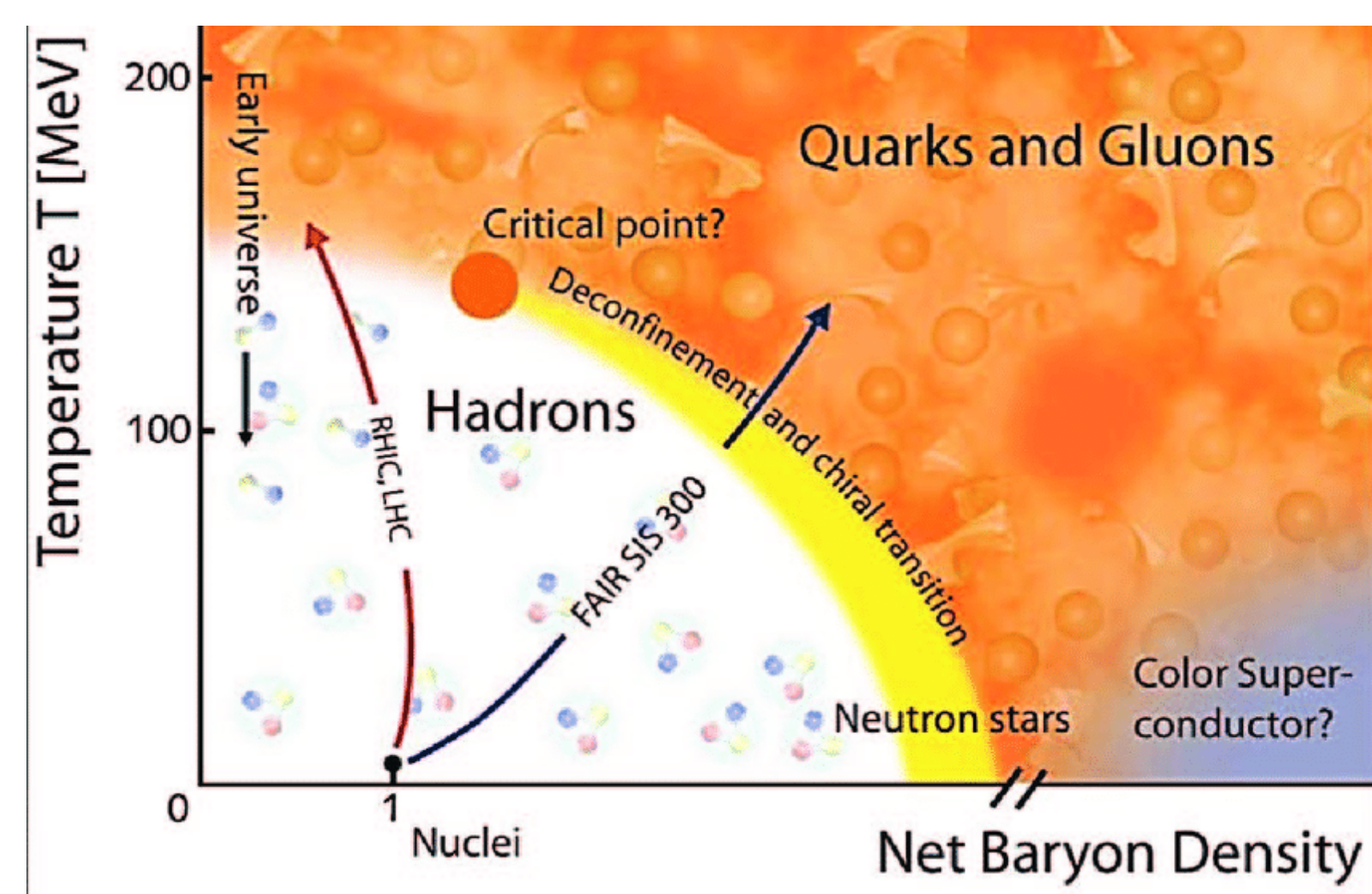


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Sebastian Blacker<sup>1,2</sup>, Niels-Uwe F. Bastian<sup>3</sup>, Andreas Bauswein<sup>2,4</sup>, David B. Blaschke<sup>3</sup>, Tobias Fischer<sup>3</sup>, Micaela Oertel<sup>5</sup>, Theodoros Soultanis<sup>6,7</sup> & Stefan Typel<sup>1,2</sup>

<sup>1</sup>TU Darmstadt, Germany; <sup>2</sup>GSI Helmholtzzentrum für Schwerionenforschung, Germany; <sup>3</sup>University of Wrocław, Poland; <sup>4</sup>Helmholtz Research Academy Hesse for FAIR, Germany; <sup>5</sup>Université de Paris, France; <sup>6</sup>Heidelberg Institute for Theoretical Studies, Germany; <sup>7</sup>Max-Planck-Institut für Astronomie, Heidelberg, Germany

## Phase diagram of matter

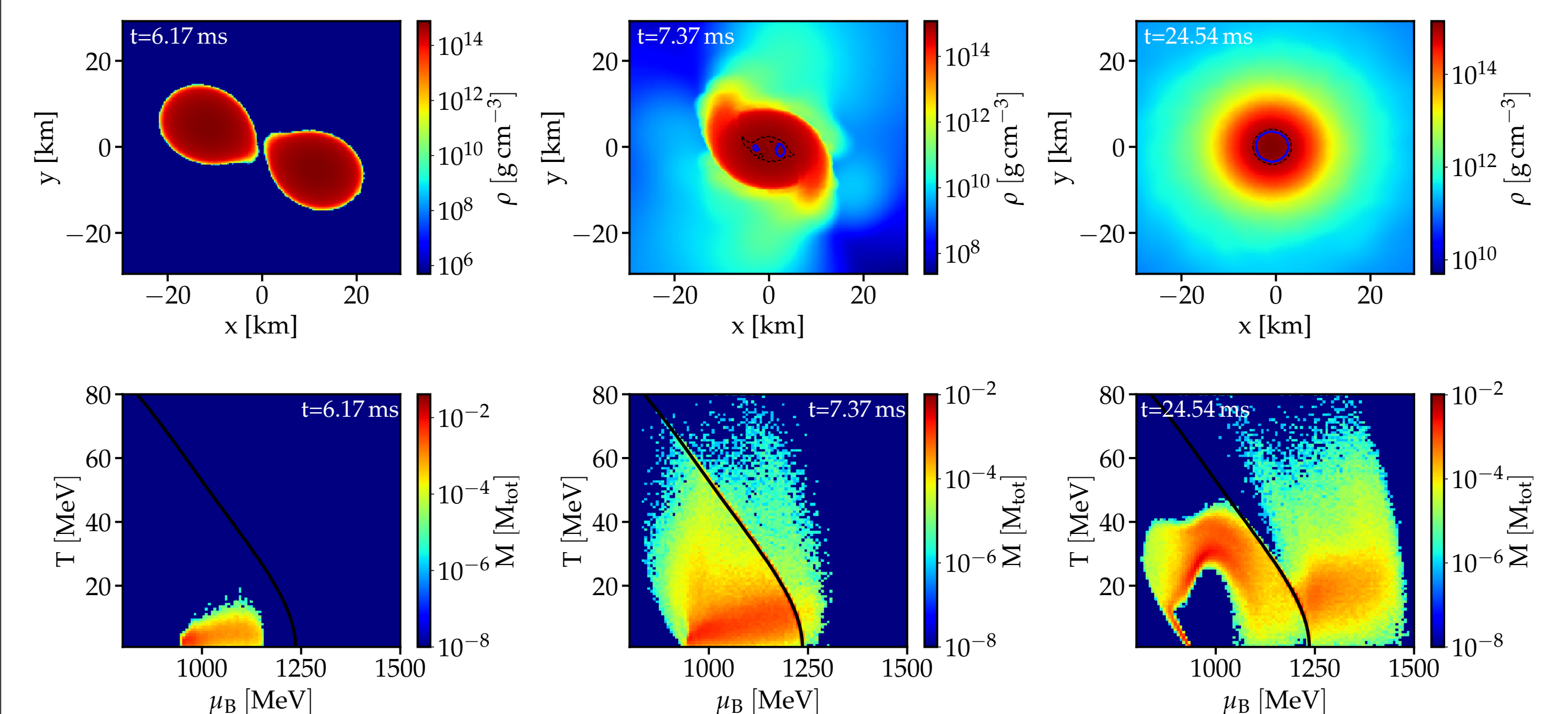


QCD predicts a phase transition from hadronic to deconfined quark matter, but at which density?

Neutron star mergers probe the high density and low temperature region inaccessible to experiments.

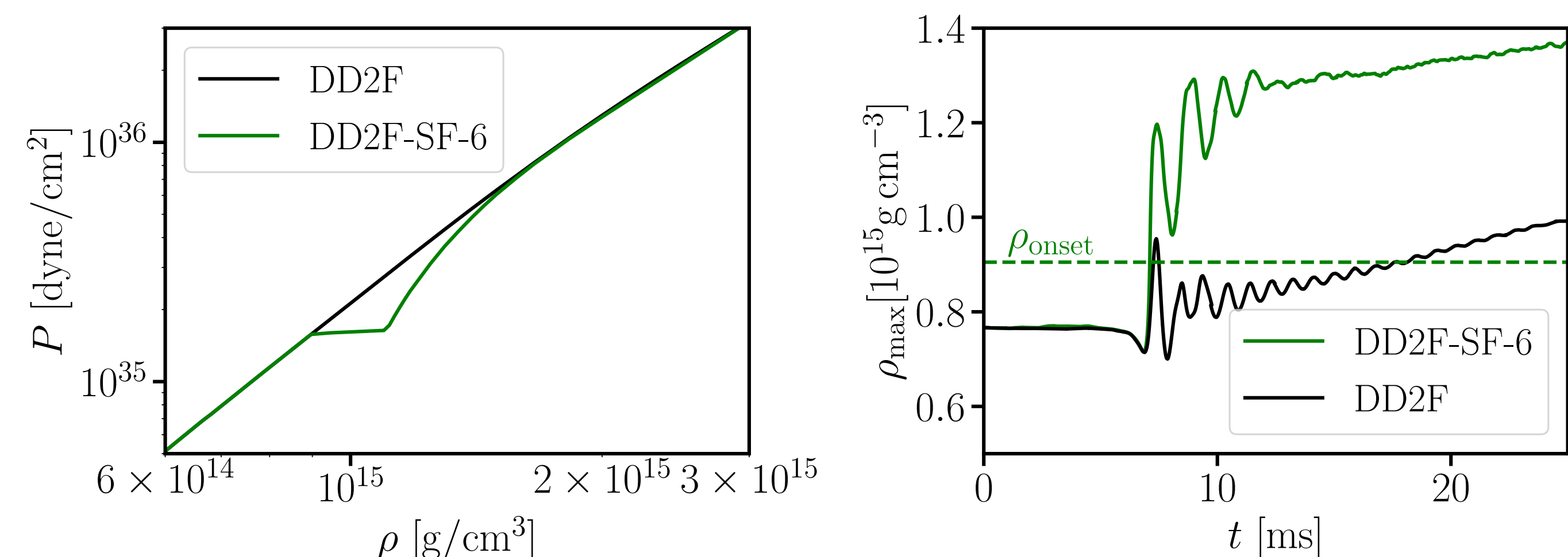
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## Merger Snapshots



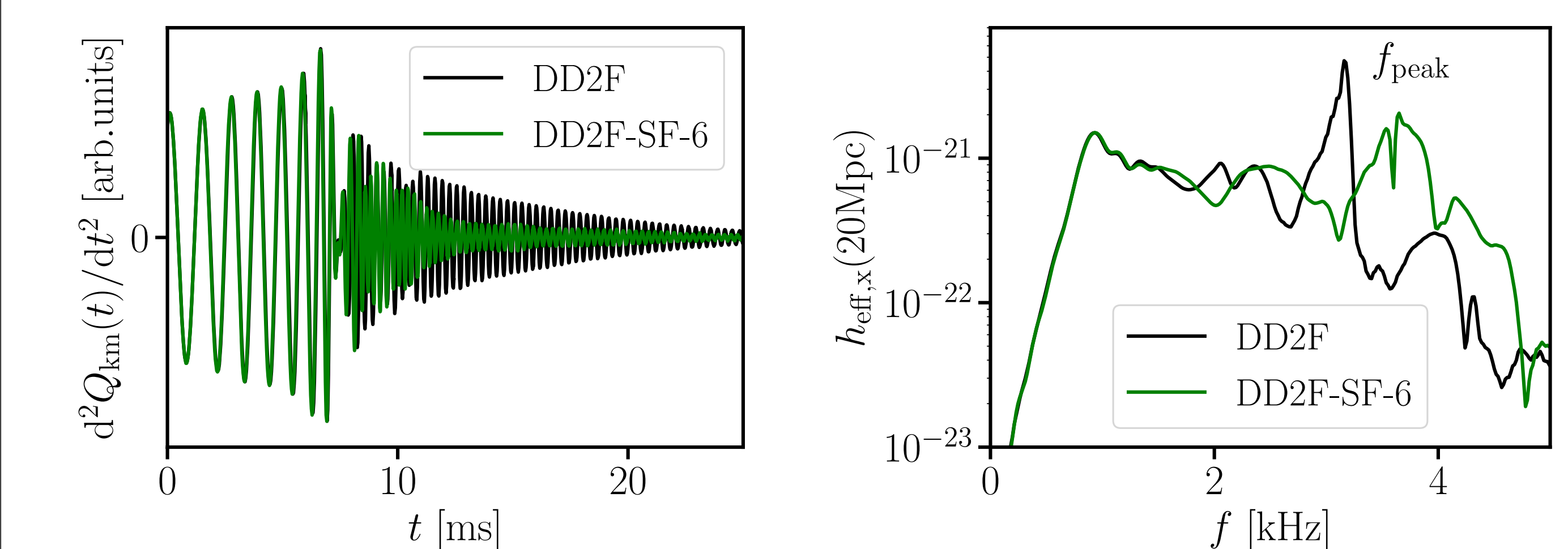
Plots taken from [1], contours highlight regions with mixed and pure quark phases.

## Impact of a phase transition



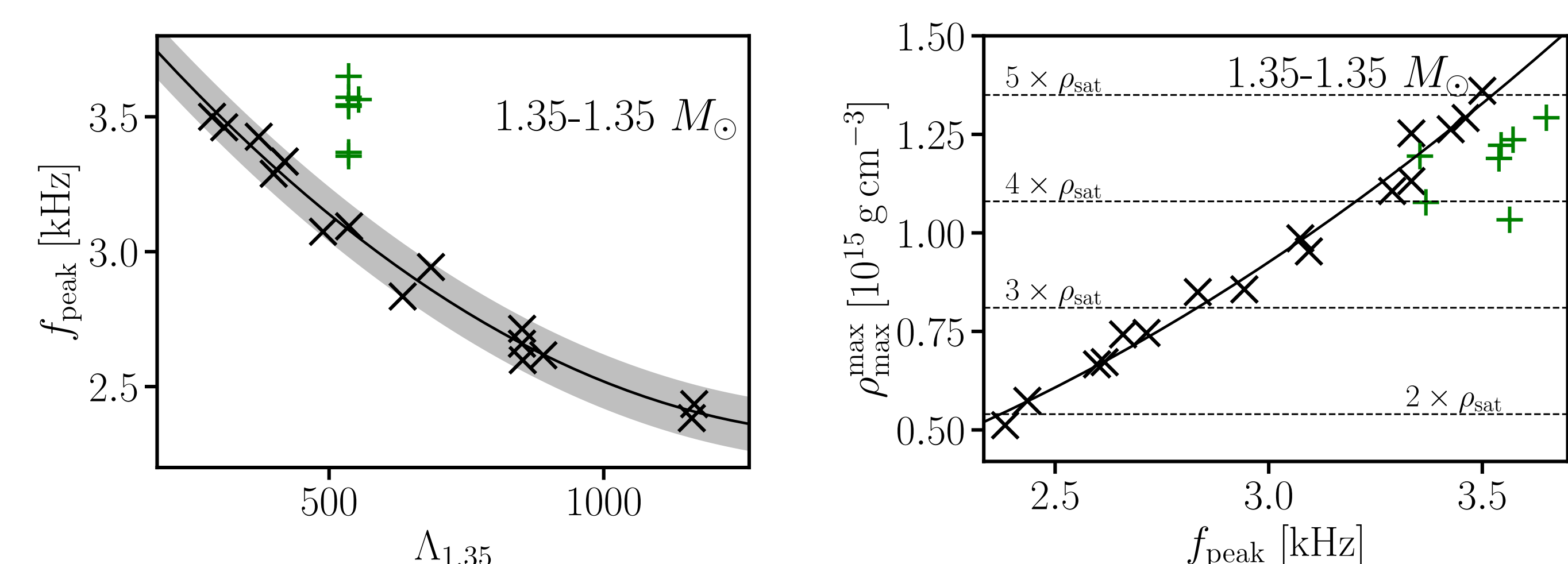
The quark-hadron phase transition strongly softens the equation of state. This softening leads to higher densities in merger remnants compared to a model with no phase transition. See [2] for details on the used model.

## Gravitational wave signal



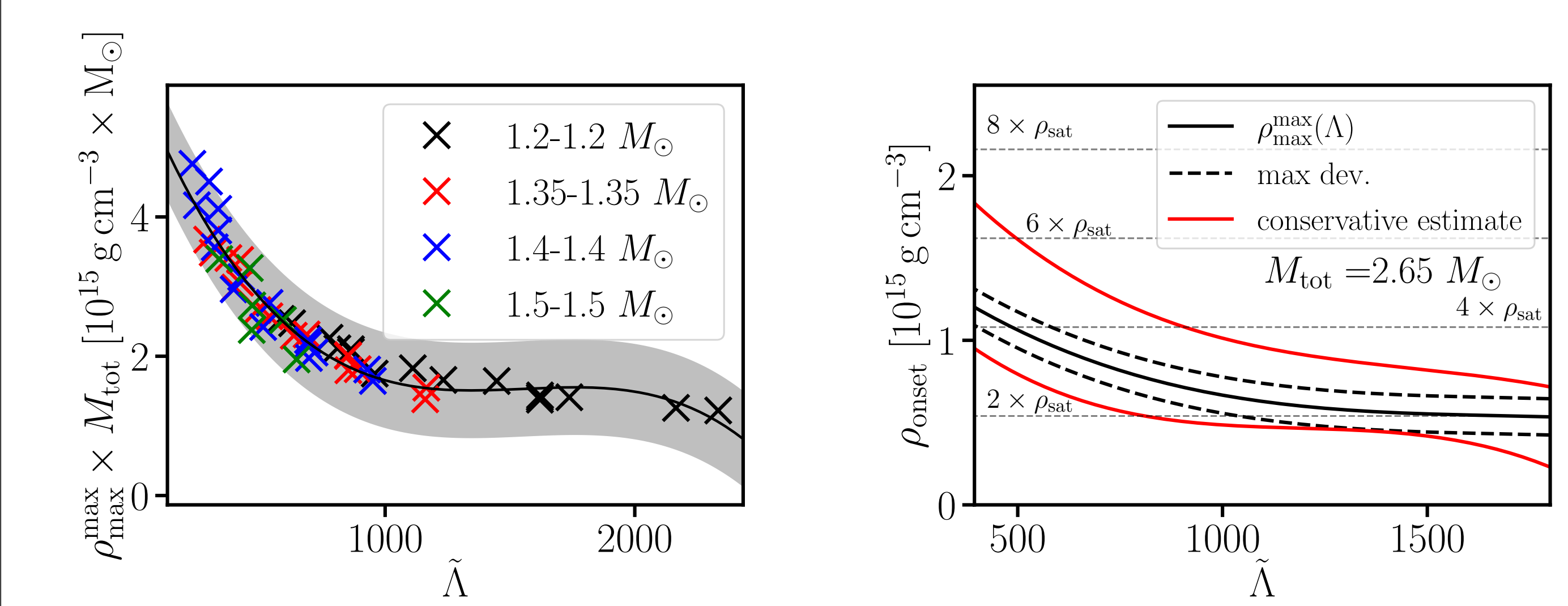
The phase transition impacts the postmerger GW signal and leads to larger frequencies. Since deconfined quarks appear after the merger the inspiral signal is unaffected.

## Identifying a phase transition



The tidal deformability  $\Lambda$  and the dominant GW frequency  $f_{\text{peak}}$  are encoded in the pre- and the postmerger, respectively. Outliers from empirical relations can be used to detect a strong phase transition and constrain the onset density. See [3] for details. Figures taken from [1].

## Constraining the onset density



The relations between  $\Lambda$ ,  $f_{\text{peak}}$  and maximum postmerger density  $\rho_{\text{max}}^{\text{max}}$  can be combined to estimate the onset density of a strong phase transition  $\rho_{\text{onset}}$ . If no sign of a transition is present a lower limit on  $\rho_{\text{onset}}$  can be inferred. A clear indication of a transition provides an upper limit on  $\rho_{\text{onset}}$ . Figures taken from [1].

## Acknowledgment

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## References

- [1] Blacker et al., PRD **102**, 123023 (2020)
- [2] Bastian, PRD **103**, 023001 (2021)
- [3] Bauswein et al., PRL **122**, 061102 (2019)