

Analytic models of the spectral properties of gravitational waves from neutron star merger remnants

T. Soultanis (HITS; MPIA, Heidelberg)

A. Bauswein (GSI, Darmstadt)

N. Stergioulas (AUTH, Thessaloniki)



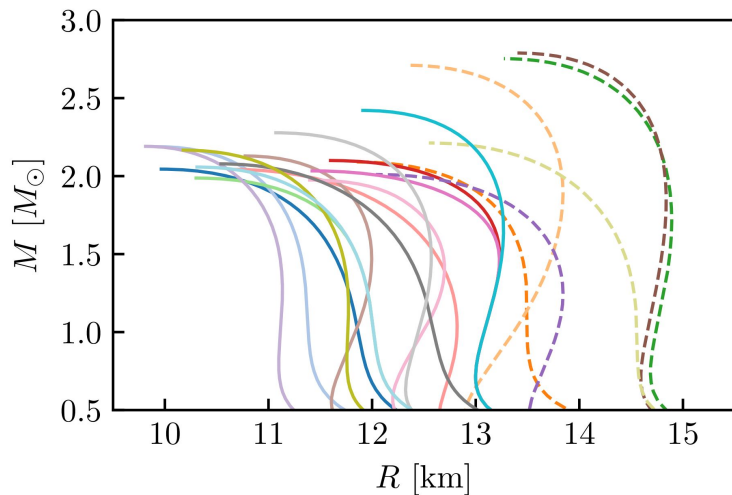
UNIVERSITÄT
HEIDELBERG
ZUKUNFT
SEIT 1386



Equation of state

- The equation of state (EoS) of a neutron star is incompletely known but theoretical models can be computed (with some uncertainties)

Neutron star mergers and
post-merger phase \Rightarrow High density regime



Adapted from Lioutas et al. (2021)

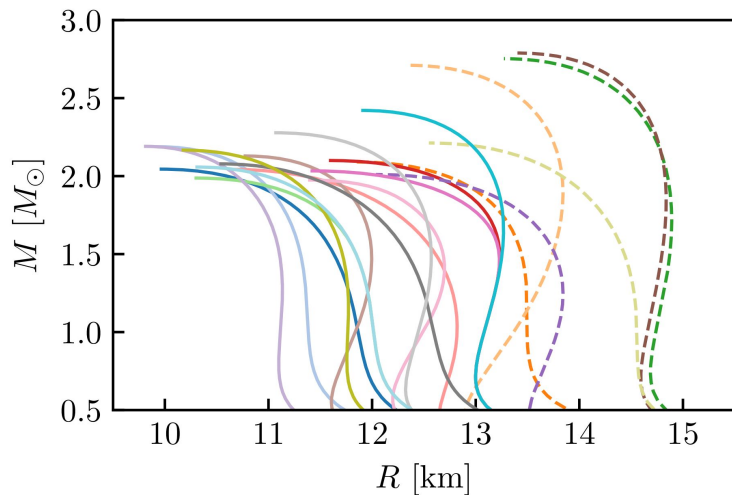
Equation of state

- The equation of state (EoS) of a neutron star is incompletely known but theoretical models can be computed (with some uncertainties)

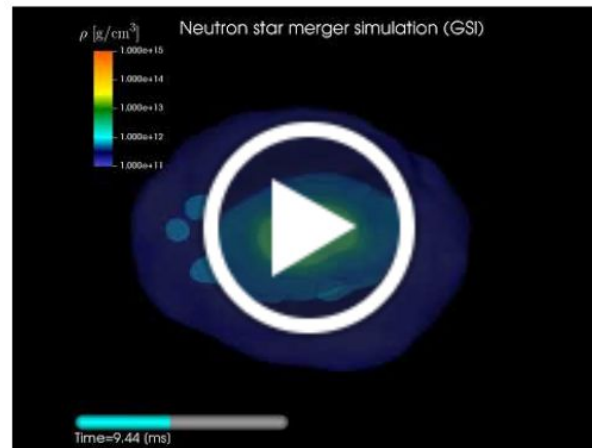
Neutron star mergers and
post-merger phase



High density regime



Adapted from Lioutas et al. (2021)



Credit: S. Blacker

Motivation

- To detect gravitational waves (GWs) we need to know how GWs look like

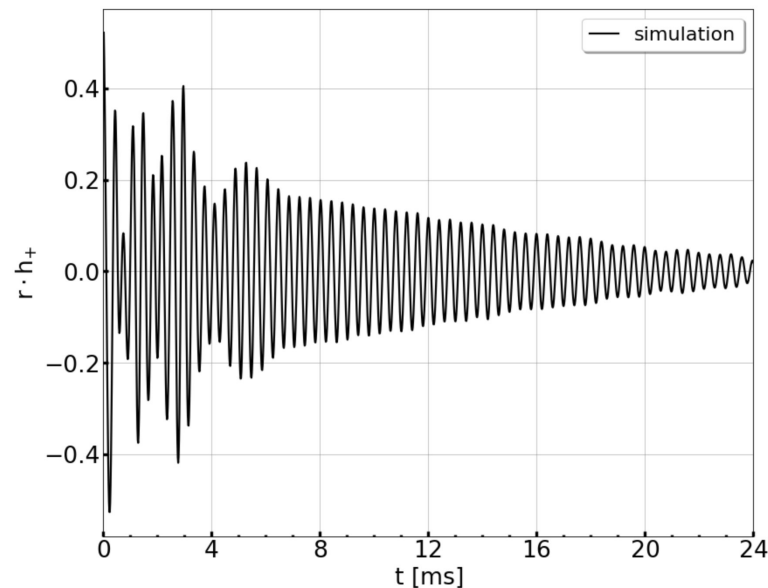
Solutions for GWs from neutron star merger
remnants



Matched-filtered search on LVC
measurements

- Analysis for potential candidates is long and computationally expensive

Fast and accurate analytic models



Simulation setup

- Simulations of binary neutron star mergers using the Einstein Toolkit (full GR and grid-based hydrodynamics)
- Sequence of models with increasing total mass using the MPA1 EOS (compatible with current constraints)

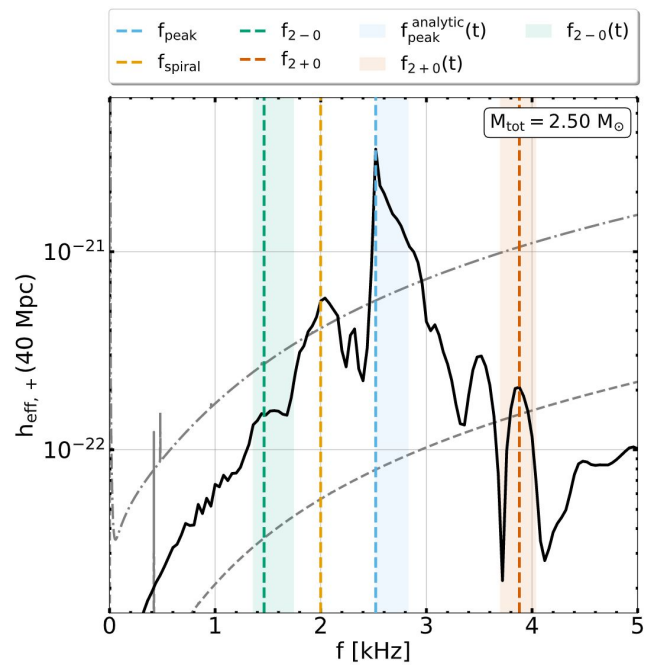
- total masses of $M_{\text{tot}} = [2.4, 2.5, 2.6, 2.7, 2.8, 2.9, 3.0, 3.1] M_{\odot}$ and ratio $q=1$

reference model: $M_{\text{tot}} = 2.5 M_{\odot}$

- Numerics:
 - spacetime evolution with Z4c formalism
 - WENO reconstruction, HLLE Riemann solver
 - $dx = 0.1875 \text{ GU} = 277 \text{ m}$, $x_{\text{max}} = 1440 \text{ GU} = 2126.28 \text{ km}$

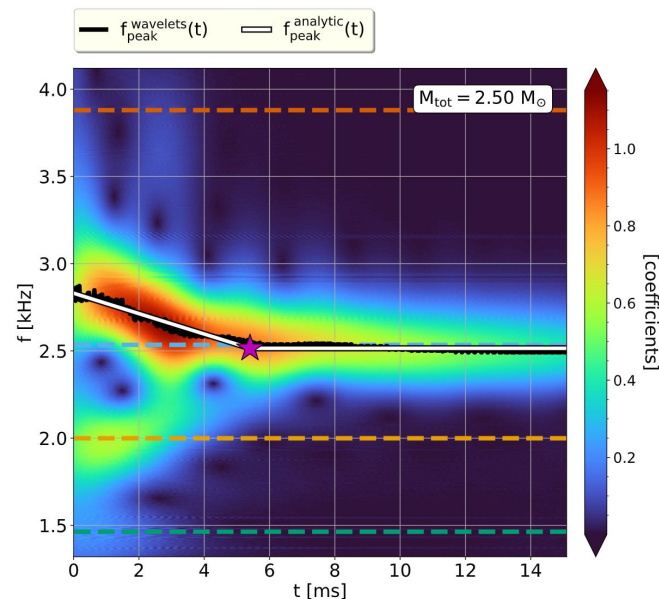
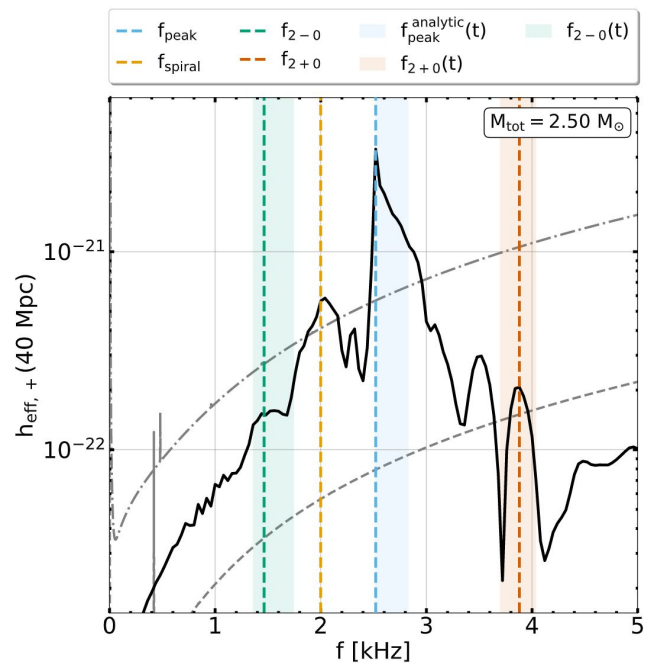
Spectral analysis

Evolution of f_{peak}



Evolution of f_{peak}

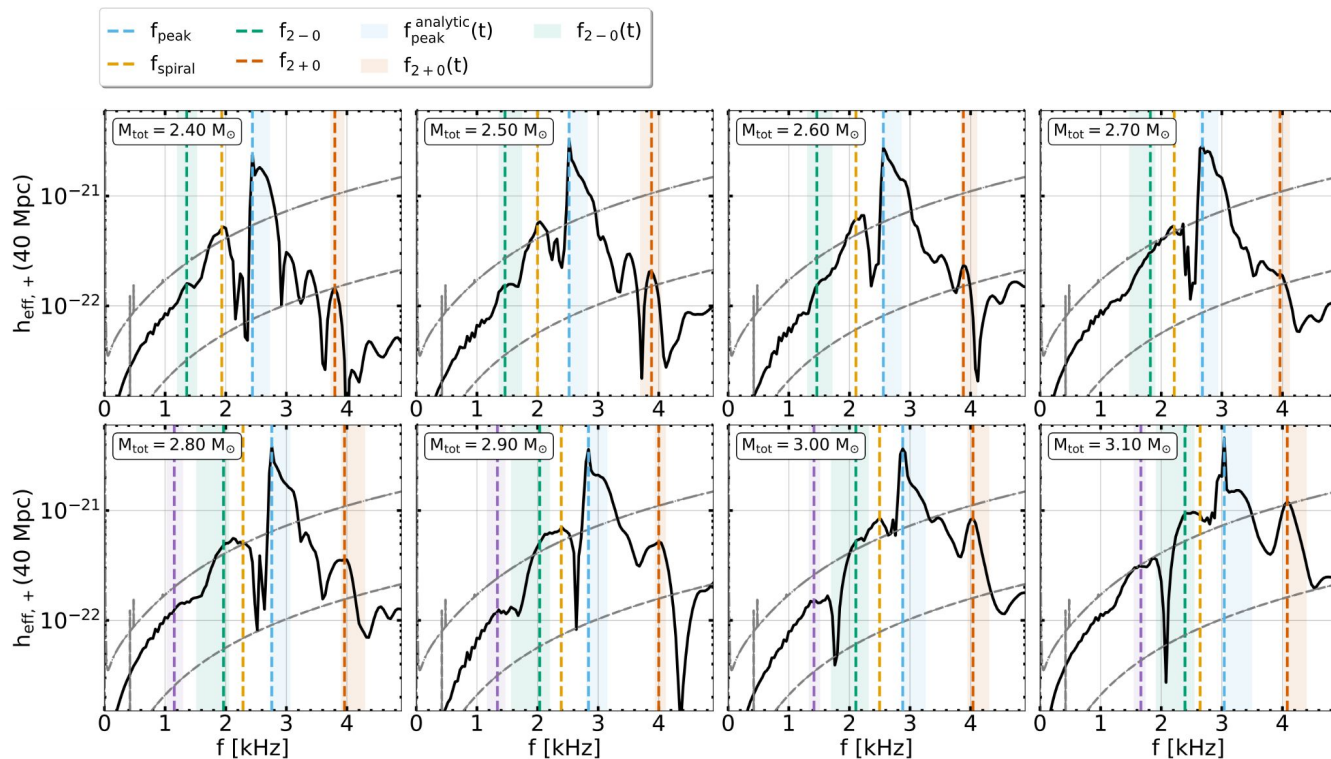
- We use spectrograms and parametrize the instantaneous f_{peak}
- 3-parameters: $\zeta_{\text{drift}}, t_*, f_{\text{peak},0}$



$$f_{\text{peak}}(t) = \begin{cases} \zeta_{\text{drift}} \cdot t + f_{\text{peak},0} & \text{for } t \leq t_* \\ f_{\text{peak}}(t_*) & \text{for } t > t_* \end{cases}$$

Sequence of models of increasing total binary mass

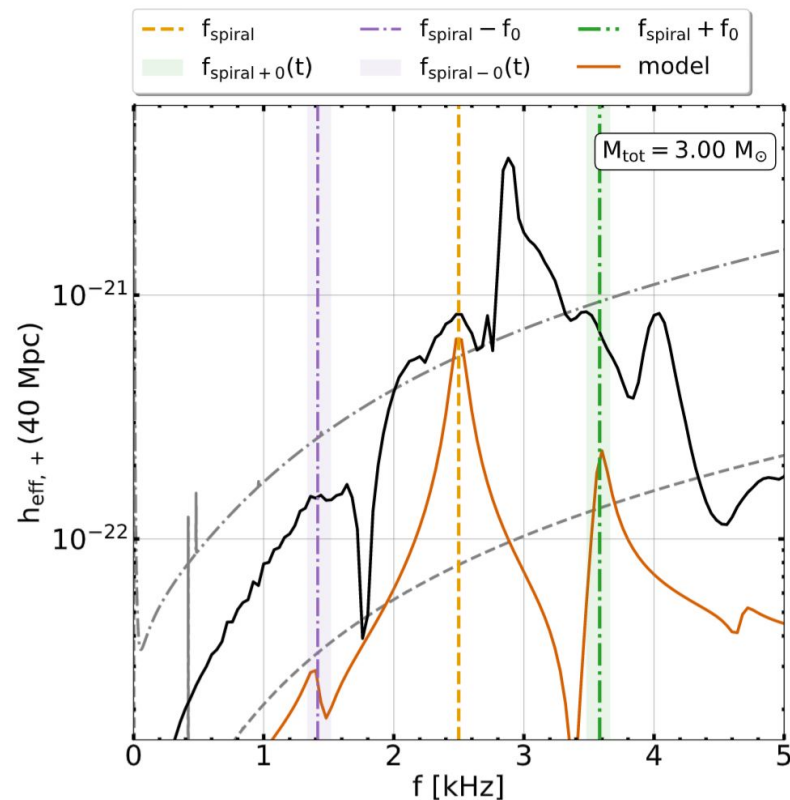
- Evolution of the spectral features depending on total binary mass



New coupling mechanism

- coupling between the ***tidal bulges*** and the ***radial*** oscillation mode

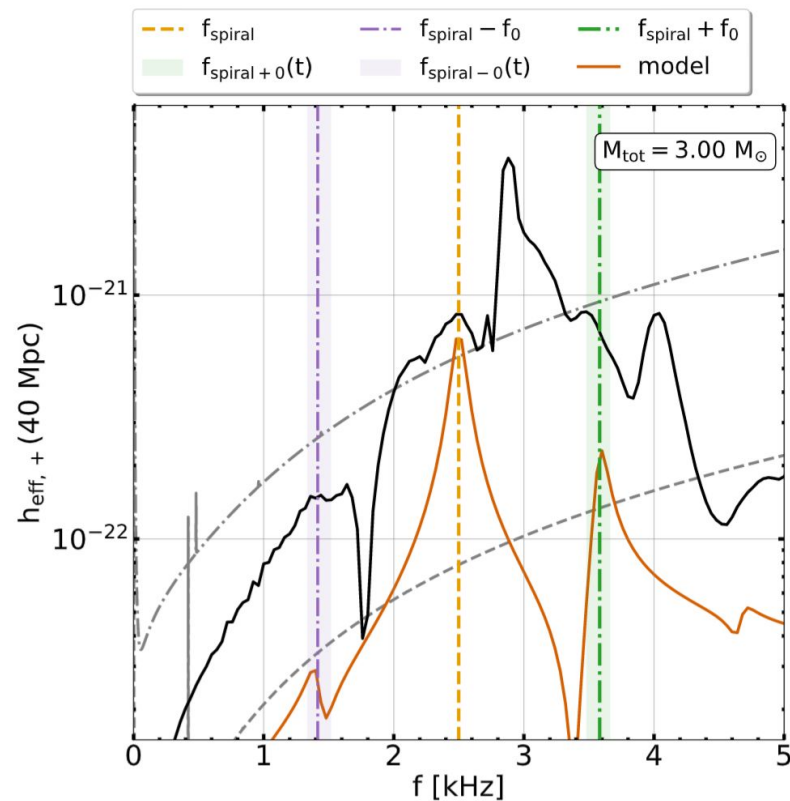
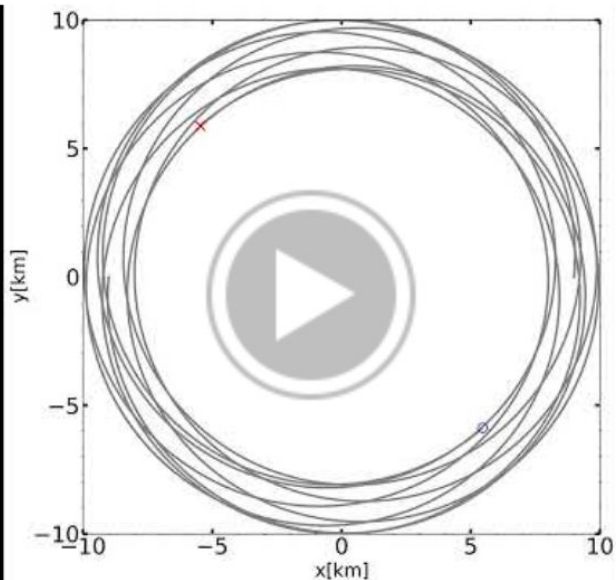
$$f_{\text{spiral} \pm 0} = f_{\text{spiral}} \pm f_0$$



New coupling mechanism

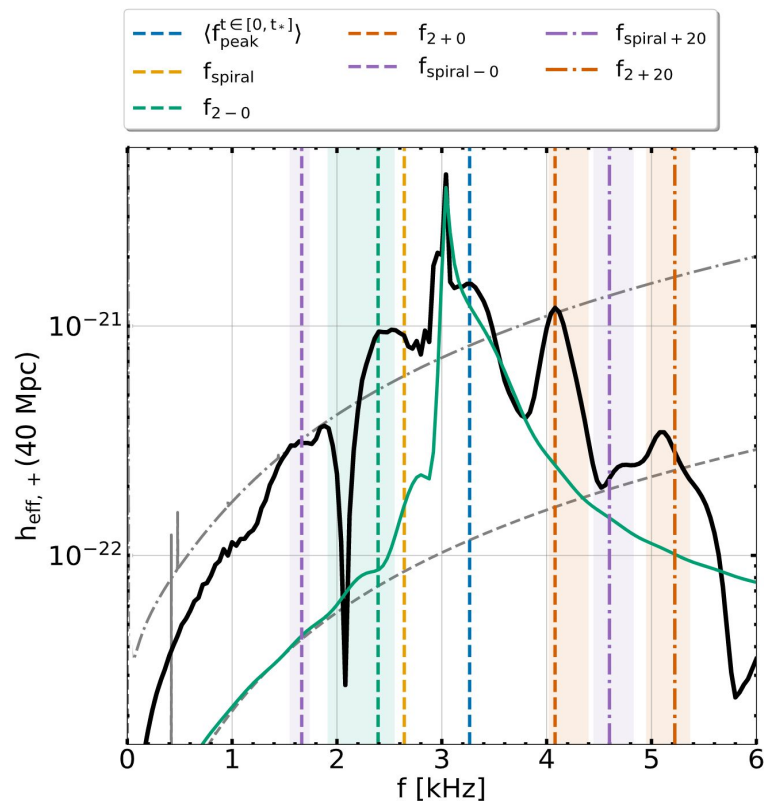
- coupling between the **tidal bulges** and the **radial** oscillation mode

$$f_{\text{spiral} \pm 0} = f_{\text{spiral}} \pm f_0$$



Models close to prompt collapse

- Different components and their couplings explain most features up to 6 kHz

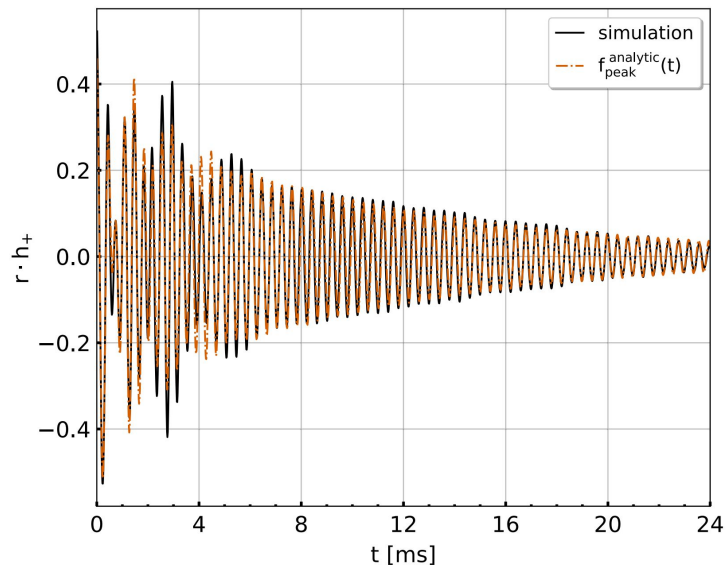


Analytic model

Analytic model: reference model

- Exponentially decaying sinusoids
 - time-dependent f_{peak}
 - 3 secondary components: $f_{\text{spiral}}, f_{2\pm 0}$

$$\begin{aligned}
 h_+(t) = & A_{\text{peak}} e^{(-t/\tau_{\text{peak}})} \cdot \sin(\phi_{\text{peak}}(t)) \\
 & + A_{\text{spiral}} e^{(-t/\tau_{\text{spiral}})} \cdot \sin(2\pi f_{\text{spiral}} \cdot t + \phi_{\text{spiral}}) \\
 & + A_{2-0} e^{(-t/\tau_{2-0})} \cdot \sin(2\pi f_{2-0} \cdot t + \phi_{2-0}) \\
 & + A_{2+0} e^{(-t/\tau_{2+0})} \cdot \sin(2\pi f_{2+0} \cdot t + \phi_{2+0})
 \end{aligned}$$

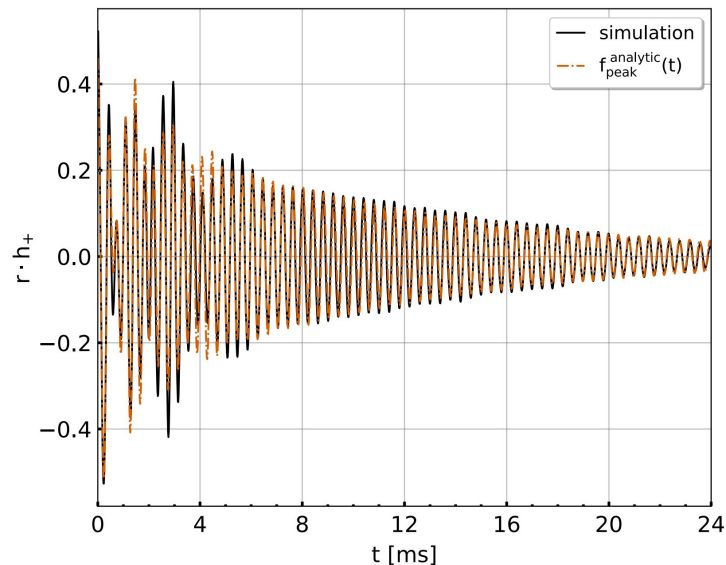
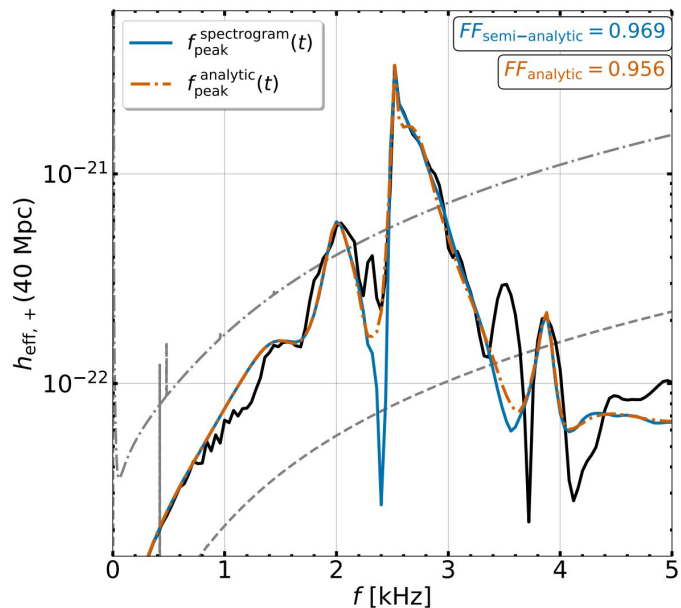


where

$$\phi_{\text{peak}}(t) = \begin{cases} 2\pi \left(f_{\text{peak},0} + \frac{\zeta_{\text{drift}}}{2} t \right) t + \phi_{\text{peak}} & \text{for } t \leq t_* \\ 2\pi f_{\text{peak}}(t_*) (t - t_*) + \phi_{\text{peak}}(t_*) & \text{for } t > t_* \end{cases}$$

Analytic model: reference model

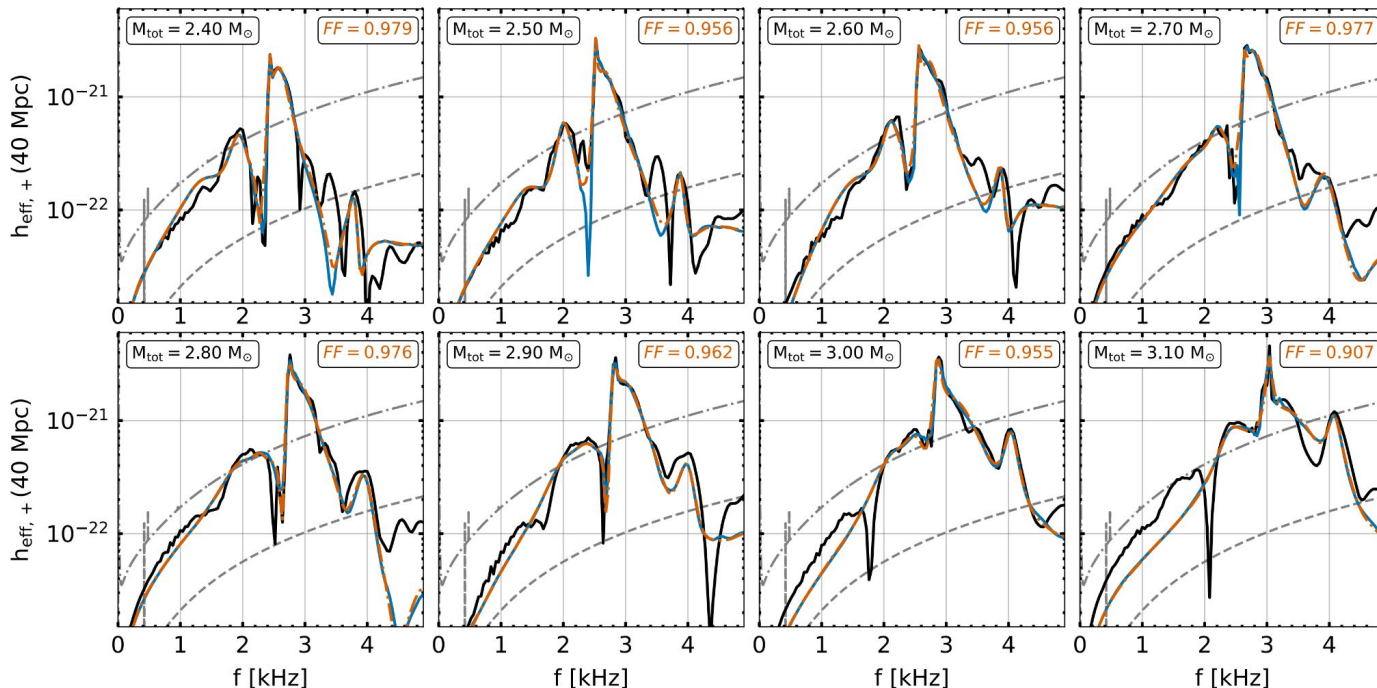
- Exponentially decaying sinusoids
 - time-dependent f_{peak}
 - 3 secondary components: $f_{\text{spiral}}, f_{2\pm0}$



Analytic model: mass sequence

- The analytic model reproduces well the f_{peak} peak for the mass sequence models

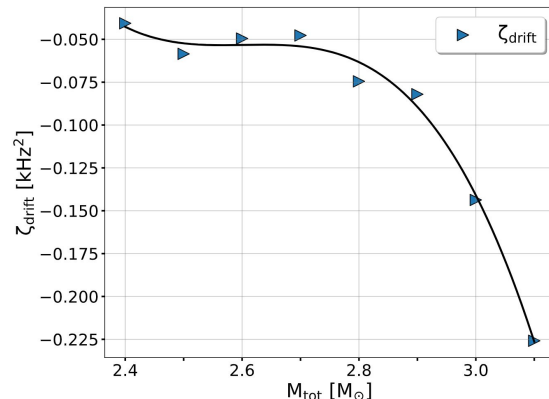
- FFs >0.95 for the majority of models



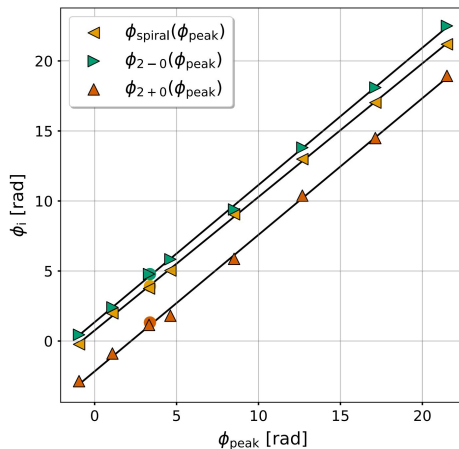
Correlations between model's parameters

- Specific trends for the description of $f_{\text{peak}}(t)$

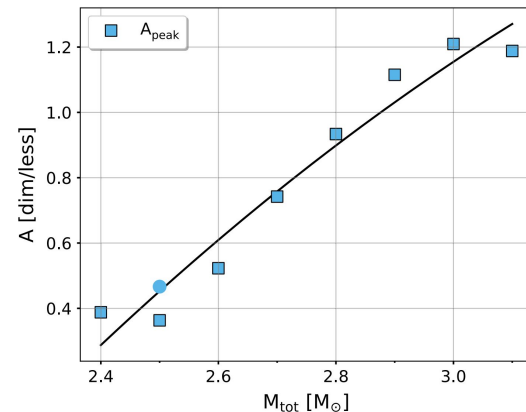
$$f_{\text{peak}}(t) = \begin{cases} \zeta_{\text{drift}} \cdot t + f_{\text{peak},0} & \text{for } t \leq t_* \\ f_{\text{peak}}(t_*) & \text{for } t > t_* \end{cases}$$



- Tight dependence of ϕ_{spiral} , $\phi_{2\pm 0}$ on ϕ_{peak}
 - slope close to 1



- Dominant mode dependence of total mass



All parameters correlate (within some uncertainty) with total binary mass M_{tot}

Summary

- Smooth transition of the spectral features depending on the total mass
- New coupling between the *antipodal bulges* and the *quasi-radial mode*: $f_{\text{spiral}\pm 0}$
- Analytic models with *exponentially damped sinusoids* performs well (FFs > 0.95)
 - 2-segment analytic linear description of $f_{\text{peak}}(t)$
- All parameters of the analytic model correlate with the total binary mass M_{tot}

[Link:](#)

