



Isomeric states of $^{113,115}\text{In}$ in radiative proton–capture reactions at energies of astrophysical interest

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Introduction & Motivation

- Isomeric states in the neutron–deficient nuclei around $A \sim 100$ are important for astrophysical processes, as these states are often involved in vast reaction networks taking place at astrophysical sites
- ^{113}In is generally considered a p nucleus known to be significantly underproduced in most astrophysical models
- For ^{115}In an older set of isomeric cross sections is reported in [4], while a more recent work was focused on its production for medical applications [5], at beam energies well outside the Gamow energy window for the reaction
- Measurements of reaction rates and cross sections in this mass regime provide stringent tests to the theoretical models
- Measurements of cross sections inside the Gamow window ($\sim 1.6\text{--}4.8$ MeV, corresponding to $T_9 \sim 1.7\text{--}3.3$ GK) are expected to provide data for better understanding the p process in this mass region

Experimental Details

- Experiments were carried out at the 5.5 MV T11 Tandem Van de Graaff accelerator of NCSR “Demokritos”, Greece
- Radiative proton capture reactions were studied at proton beam energies $E_p = 3.0\text{--}4.0$ MeV
- Array of 3 HPGe detectors, placed at 55° , 90° and 155° , respectively
- Isotopically enriched $^{112,114}\text{Cd}$ targets were used
- De–excitations to the g.s. were measured in–beam
- The activation method was employed for the measurement of the isomeric cross section, σ_{is}
- σ_{is} was also measured via the in–beam method, by measuring decays to the isomeric state in the in–beam spectra

Results

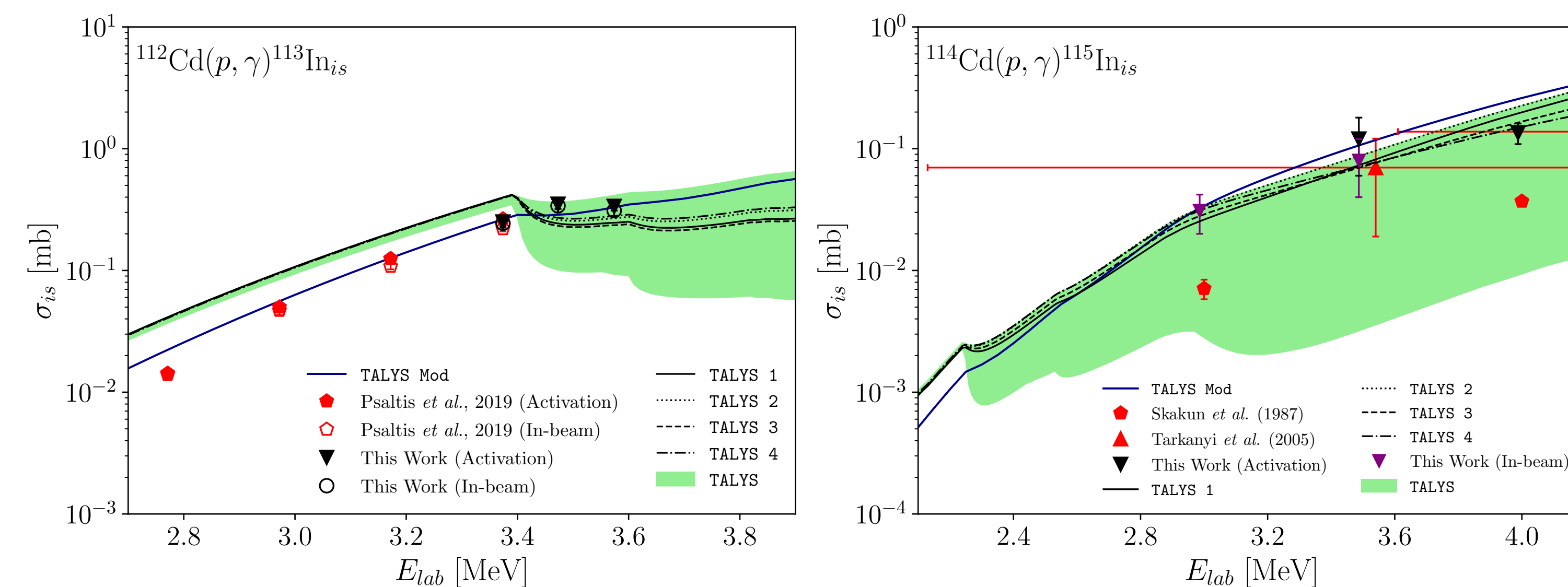


Figure 1: Isomeric cross sections for the reactions $^{112,114}\text{Cd}(p, \gamma)^{113m,115m}\text{In}$. The shaded areas correspond to the full range of calculated values with every combination of models employed [6]. The lines correspond to the best data matching calculations [1]. The results are compared to experimental data, previously published from our group [3].

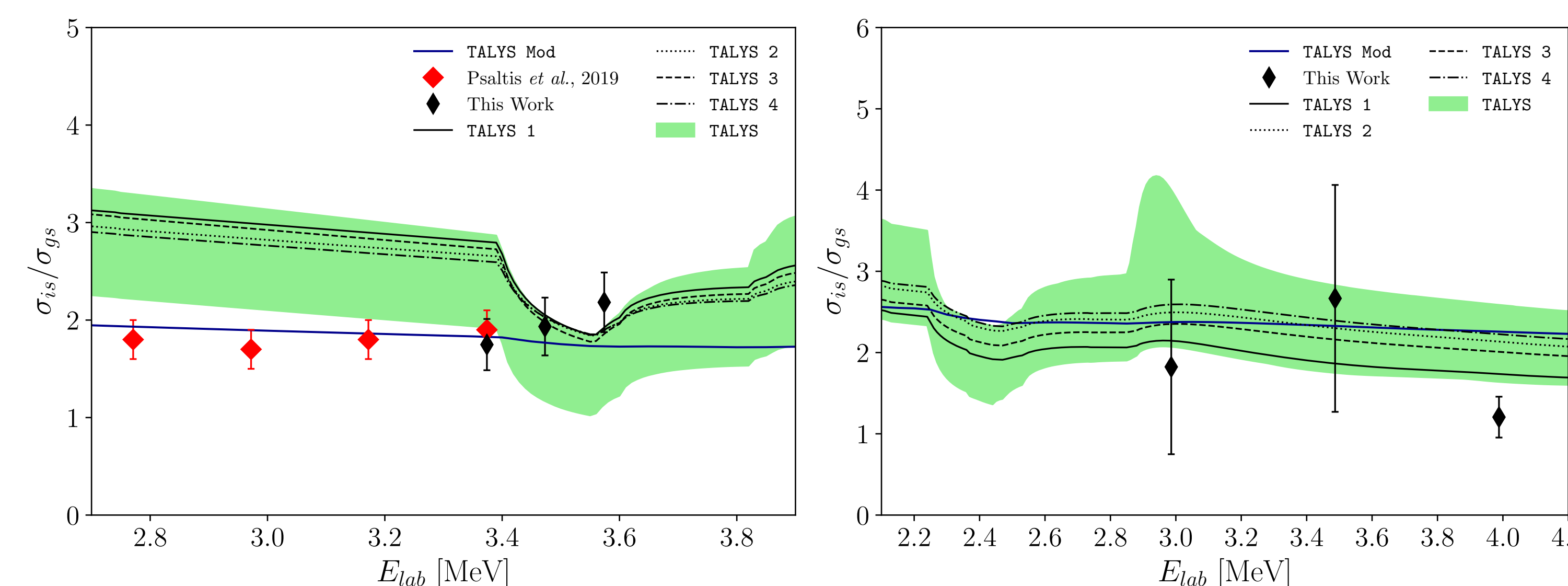


Figure 2: Isomeric to ground state ratios for the reactions $^{112,114}\text{Cd}(p, \gamma)^{113m,115m}\text{In}$

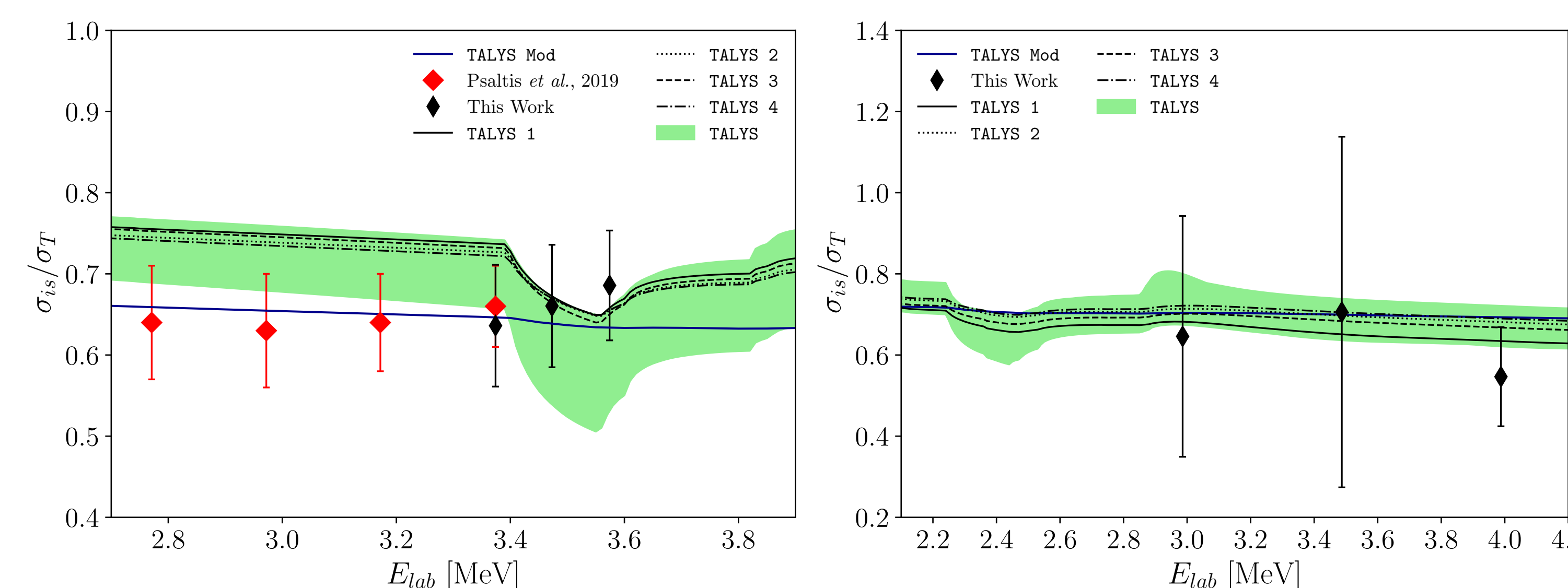


Figure 3: Same as in Fig. 2, for the isomeric to total cross section ratios.

Overview & Conclusions

- Excellent agreement with the data from the earlier work of Ref. [3], and extension to energies above the neutron emission threshold ($E_{th} = 3.397$ MeV), while still inside the Gamow energy window of astrophysical importance [1, 2]
- Better agreement with the data of Ref. [5] than Ref. [4], improving on the large uncertainties.
- Experimental results are compared to theoretical predictions using every possible combination of OMP+NLD+ γ SF provided by the TALYS v1.95 code [6] (green shaded area in figures) [1, 2]
- The best set of OMP+NLD+ γ SF model parameters was determined, in an effort to achieve a good description of the experimental data for each reaction channel in a simultaneous fashion [1, 2]
- Further experimental and theoretical work required for a firm insight at the driving mechanisms behind the p process nucleosynthesis, in an energy region where the experimental data are scarce

Acknowledgements

AZ and AC acknowledge support by the Hellenic Foundation for Research and Innovation (HFRI) and the General Secretariat for Research and Technology (GSRT) under the HFRI PhD Fellowship grant (GA. No. 101742/2019 and 74117/2017, respectively).



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