



# Zoé Favier

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EMMI Workshop/NUSTAR program  
Online/GSI, September 16<sup>th</sup> 2021



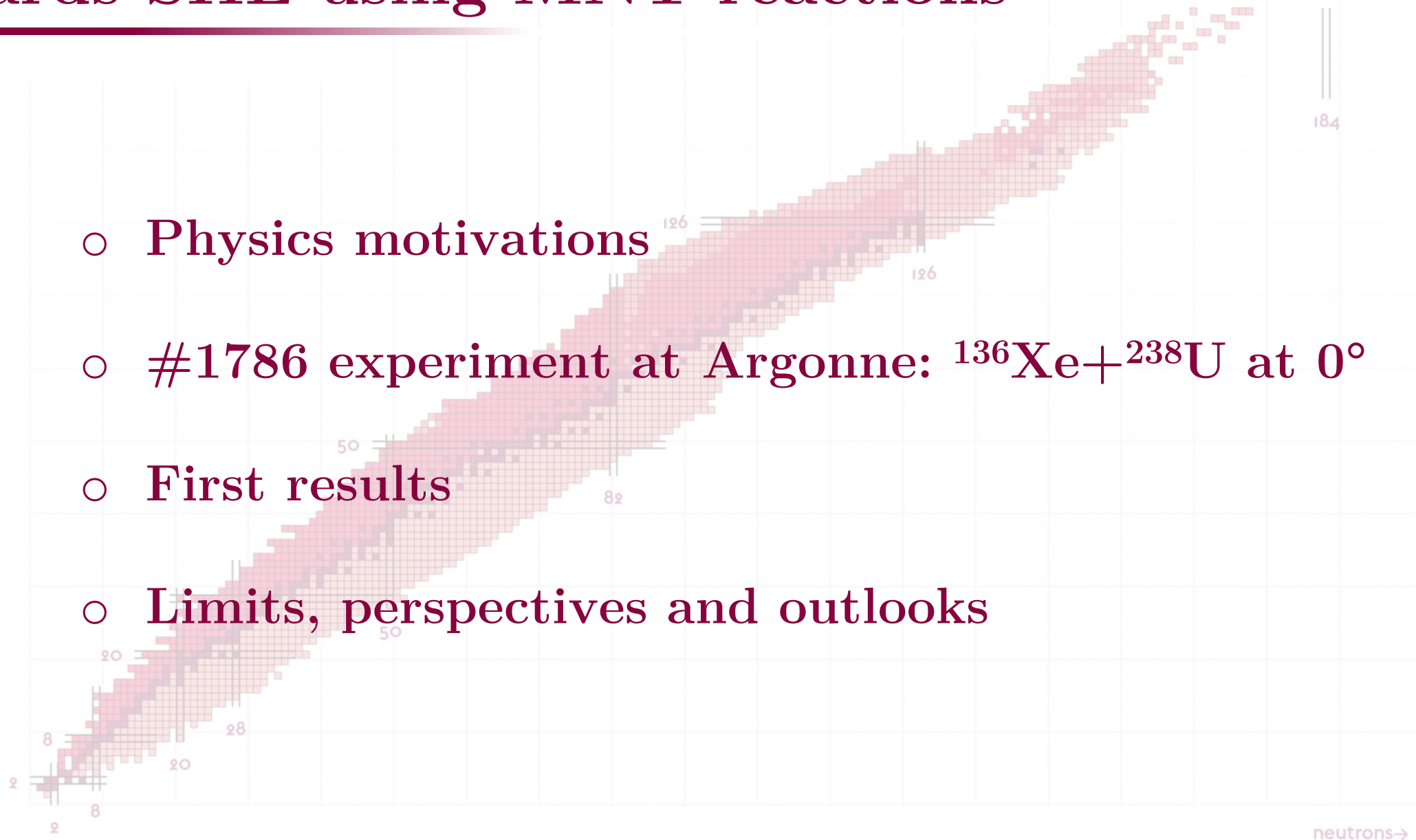
## Towards the superheavy elements:

MNT experiment carried out at Argonne:  $^{136}\text{Xe} + ^{238}\text{U}$  at  $0^\circ$



# Towards SHE using MNT reactions

- Physics motivations
- #1786 experiment at Argonne:  $^{136}\text{Xe} + ^{238}\text{U}$  at  $0^\circ$
- First results
- Limits, perspectives and outlooks



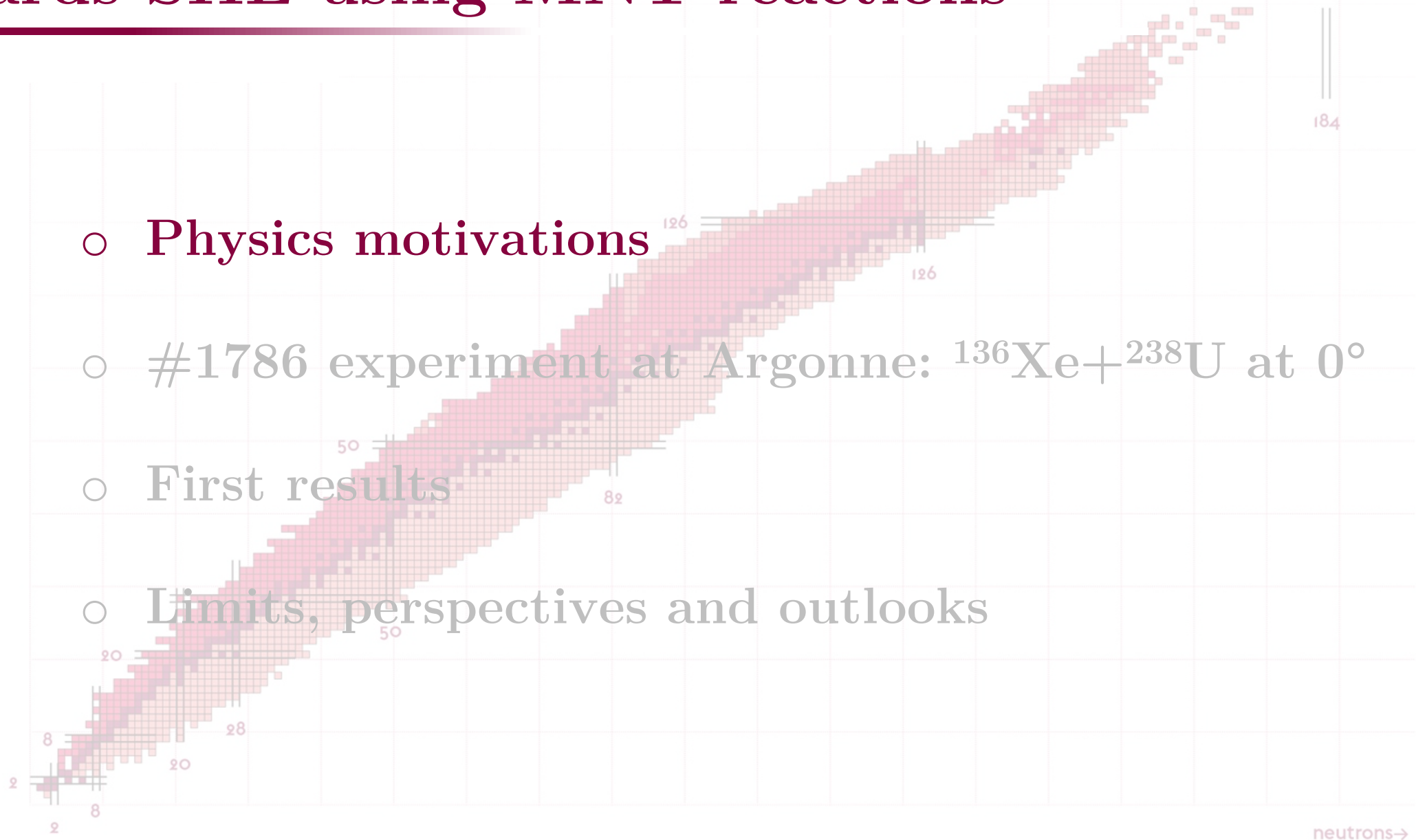
# Towards SHE using MNT reactions

- Physics motivations

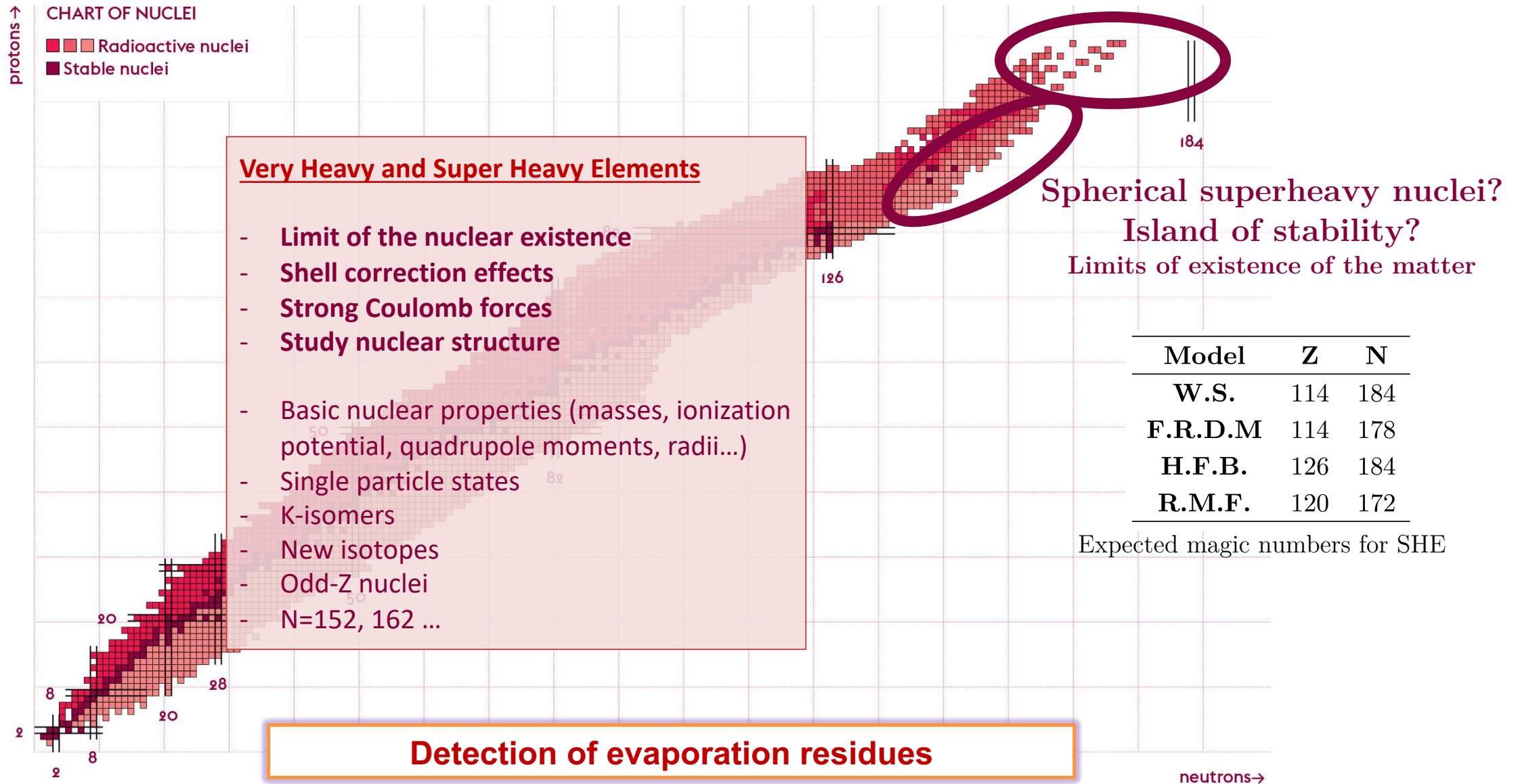
- #1786 experiment at Argonne:  $^{136}\text{Xe} + ^{238}\text{U}$  at  $0^\circ$

- First results

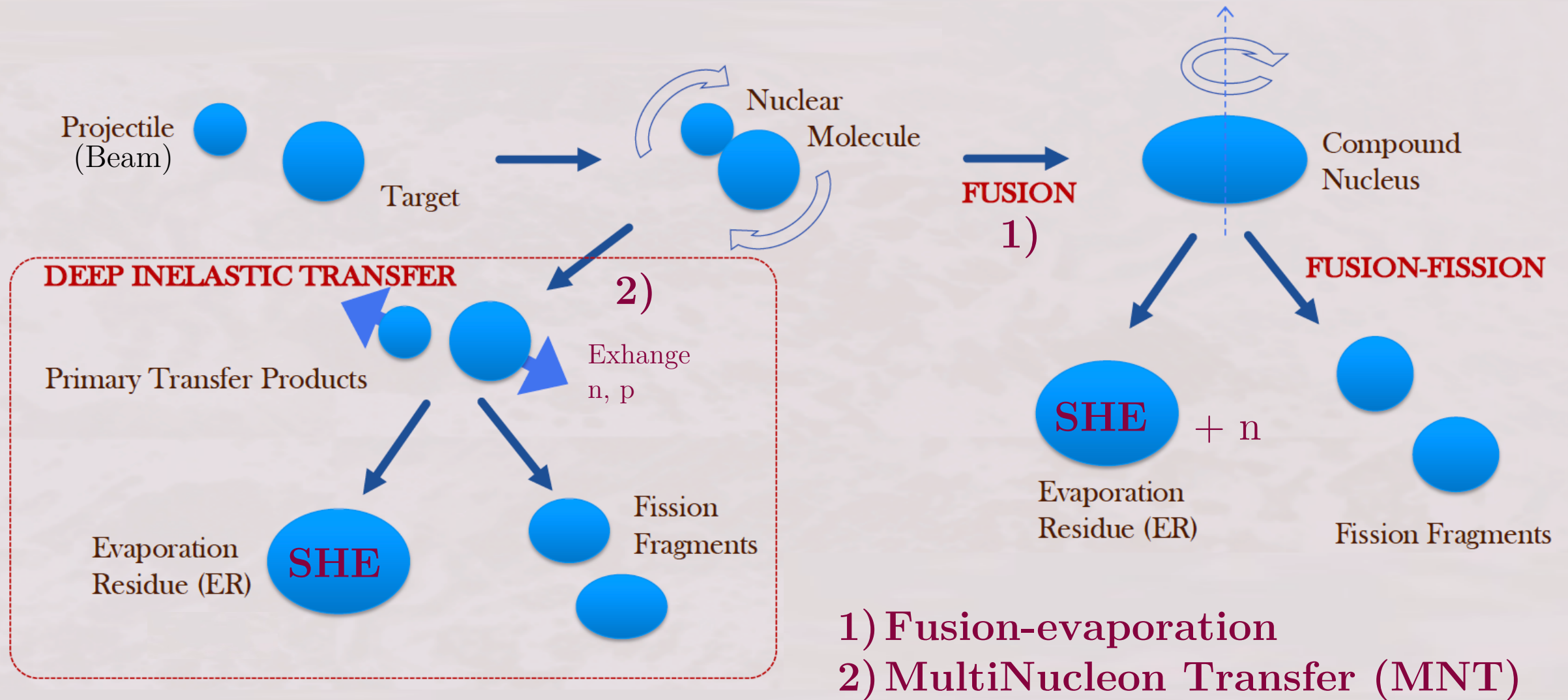
- Limits, perspectives and outlooks



# Physics objectives: SHE, a perfect laboratory

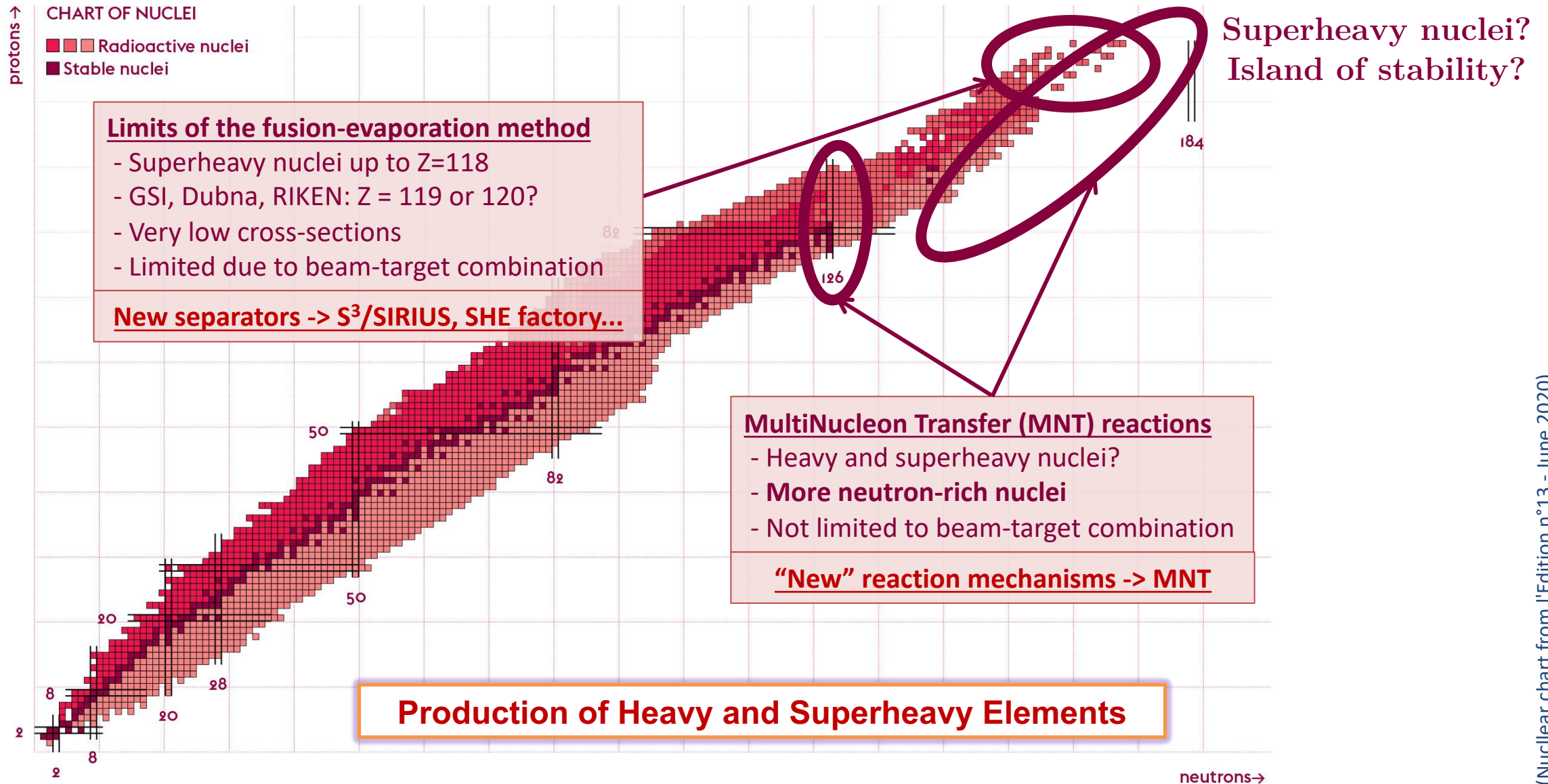


# Reaction mechanisms to produce SHE?

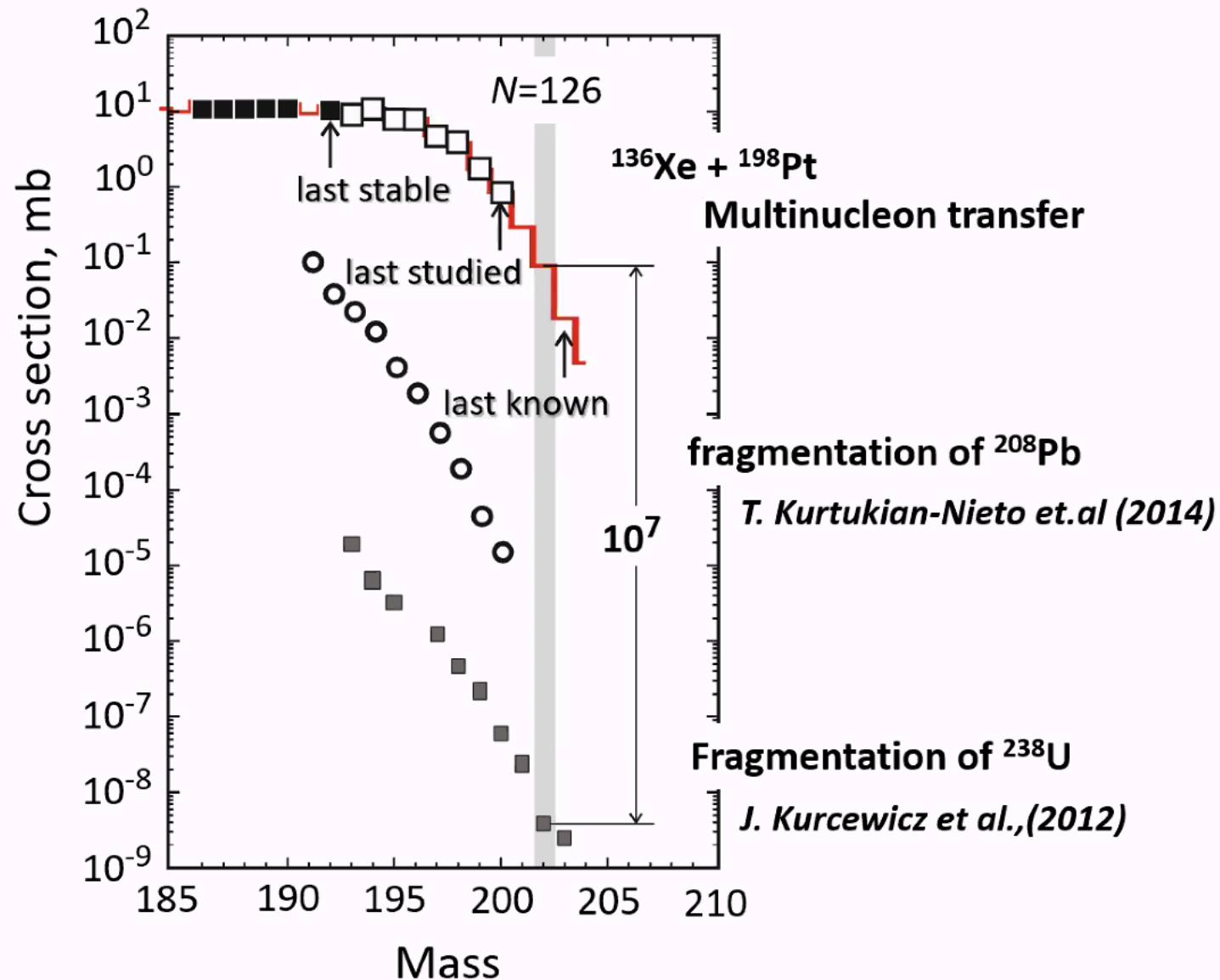




# Physics objectives: VHE and SHE

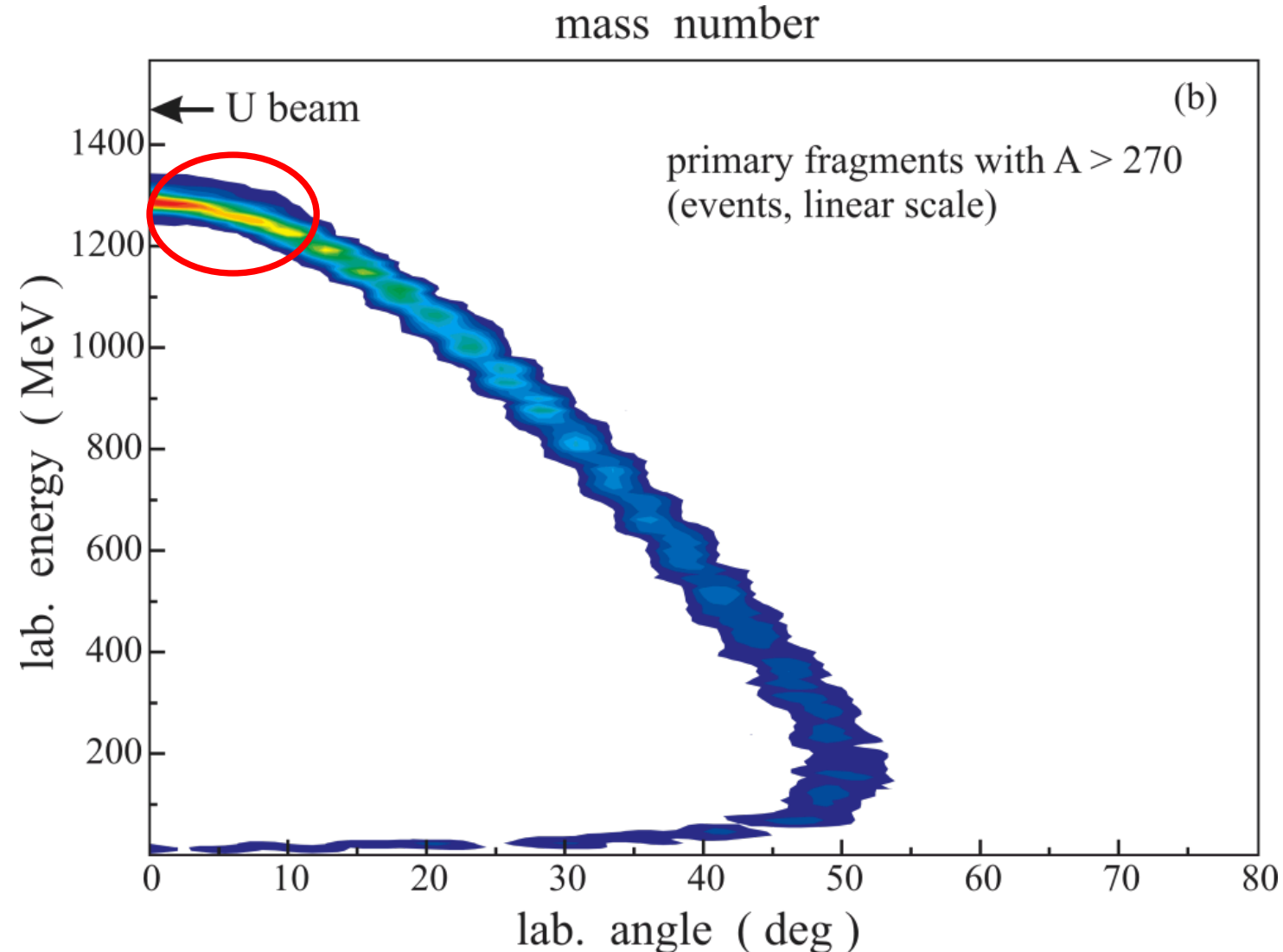


# Production of neutron-rich Os isotopes



# Deep-inelastic studies at $0^\circ$

V.I. Zagrebaev and W. Greiner, PRC 83 (2011) 044618

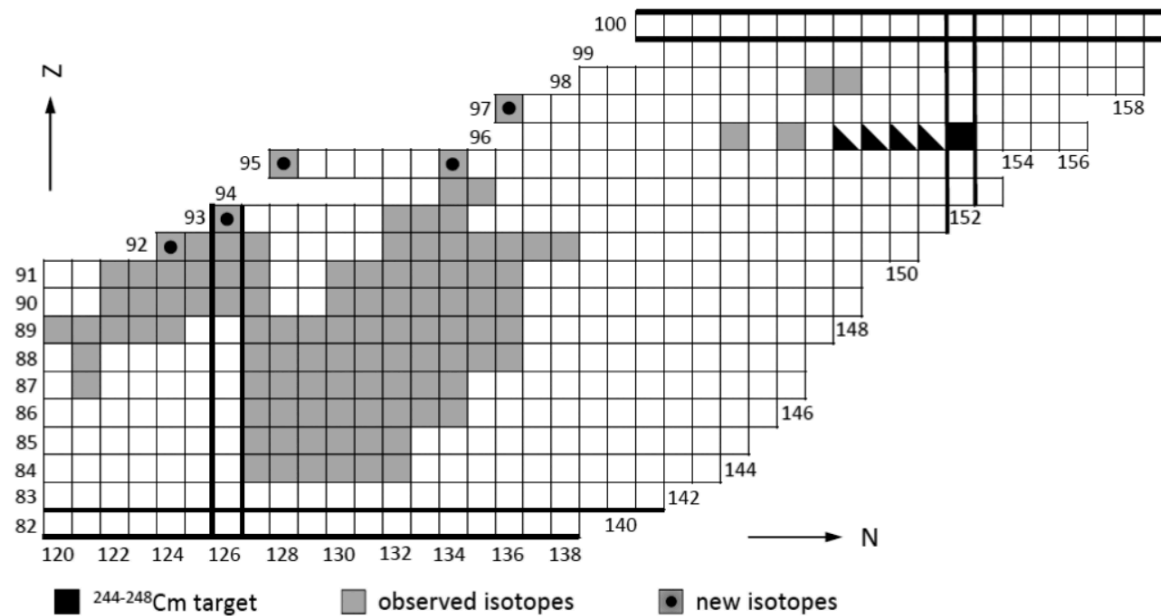


The angular distribution of the superheavy nuclei does not reveal any grazing feature, *it is forward directed.*

# First MNT studies in the SHE region at 0°

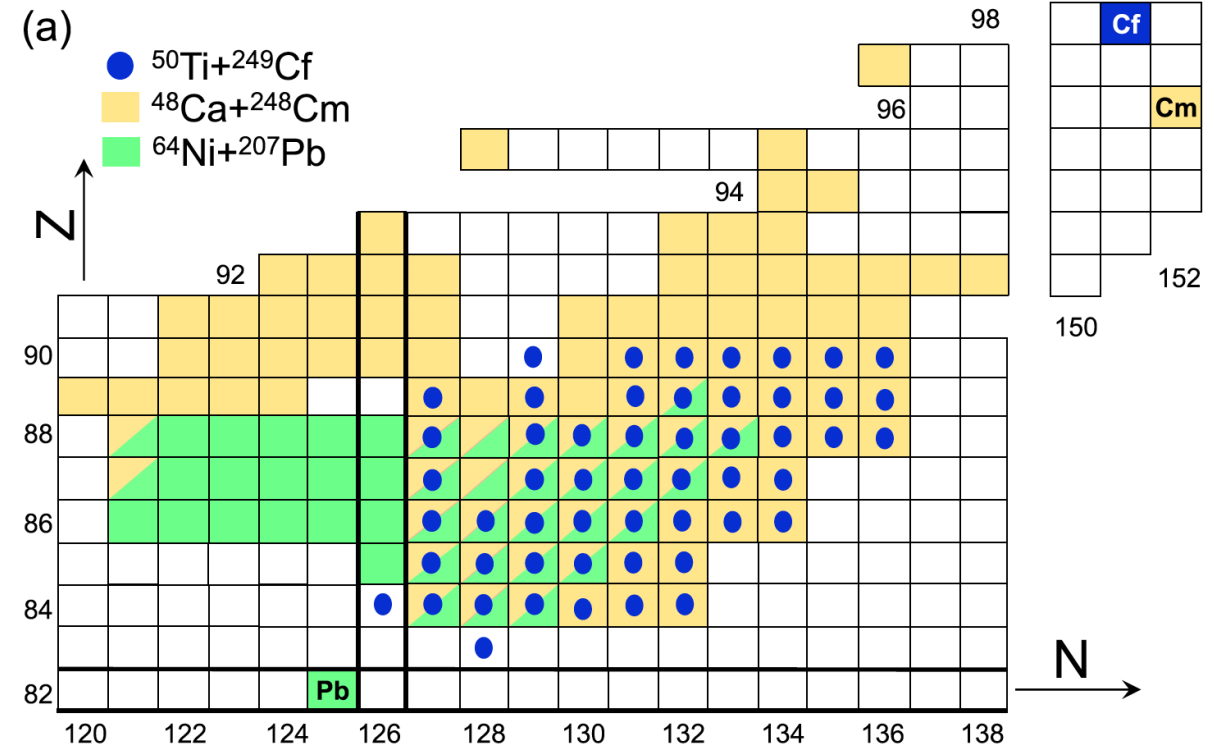
$^{48}\text{Ca}+^{248}\text{Cm}$  @ (5.3 MeV/n) At SHIP (GSI) 0°

S. Heinz, et al. Eur. Phys. J. A (2016) 52: 278



$^{50}\text{Ti}+^{249}\text{Cf}$  @ (6.1 MeV/n) At TASCA (GSI) 0°

<sup>1</sup>A. Di Nitto, et al. PLB 784 (2018)199–205

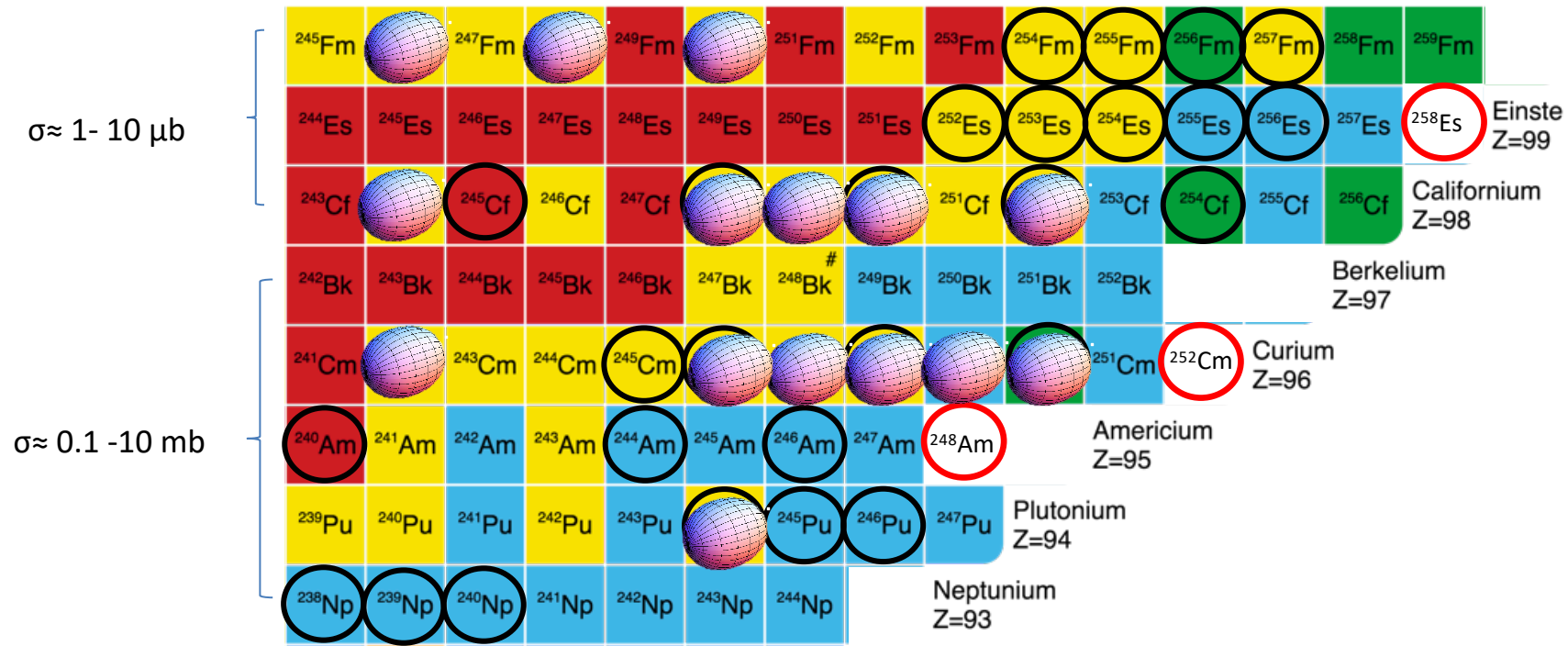


The nuclei observed in the  $^{50}\text{Ti}+^{249}\text{Cf}$  and  $^{48}\text{Ca}+^{248}\text{Cm}$  reactions originate from a process where a large number of nucleons flow in the direction of the projectile nucleus with cross-sections significantly higher than the fusion-evaporation ones.<sup>1</sup>



# Towards superheavy elements

$^{136}\text{Xe} + ^{248}\text{Cm} @ 5 \text{ MeV/n around } 0^\circ$



Prompt spectroscopy allows to measure:

- Highly excited states not populated via fusion-evaporation reaction
- $\gamma$ -rays spectroscopy will provide crucial information for theoretical models

E.g.: Via measurement of excitation energy and spin distribution

# Towards SHE using MNT reactions

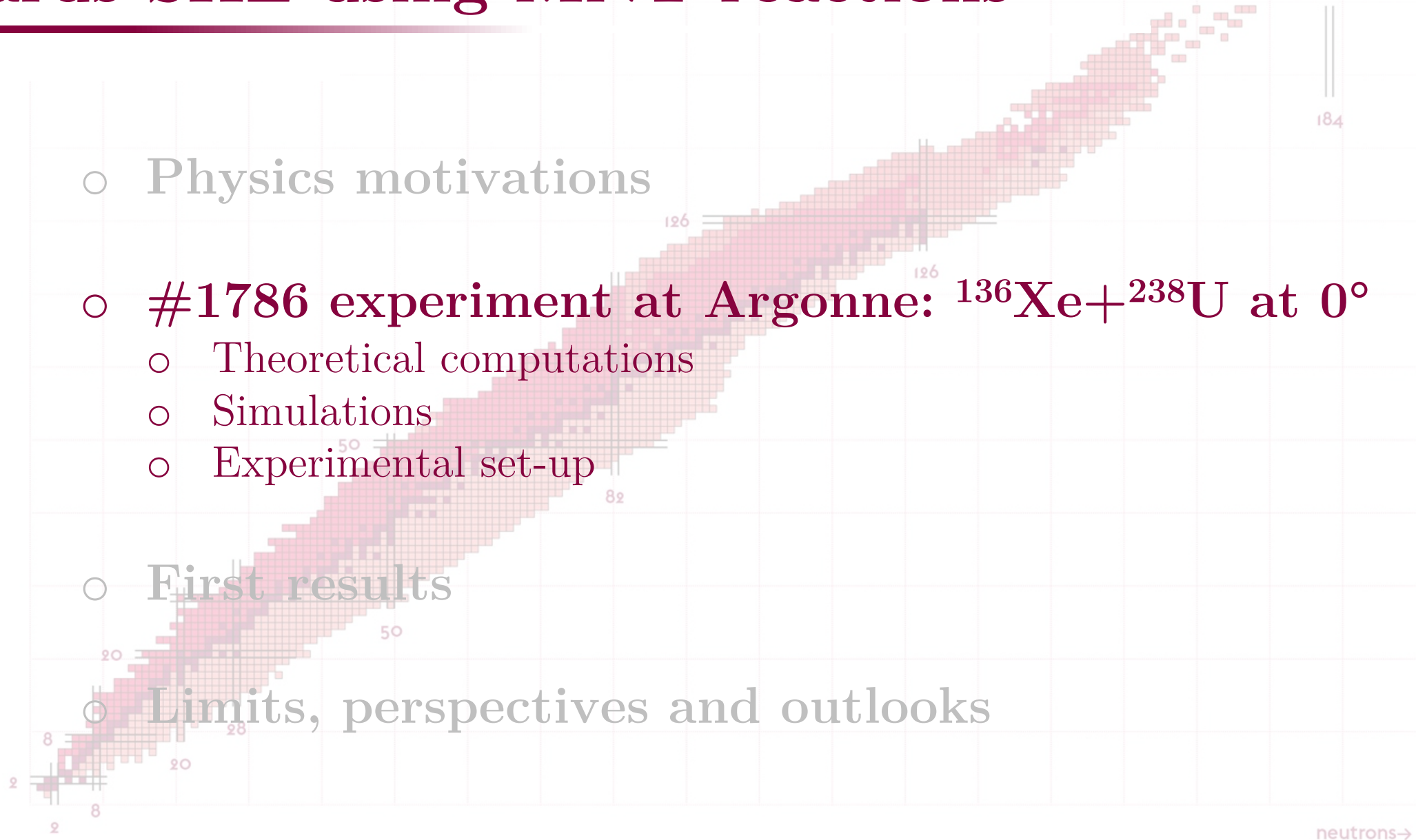
- Physics motivations

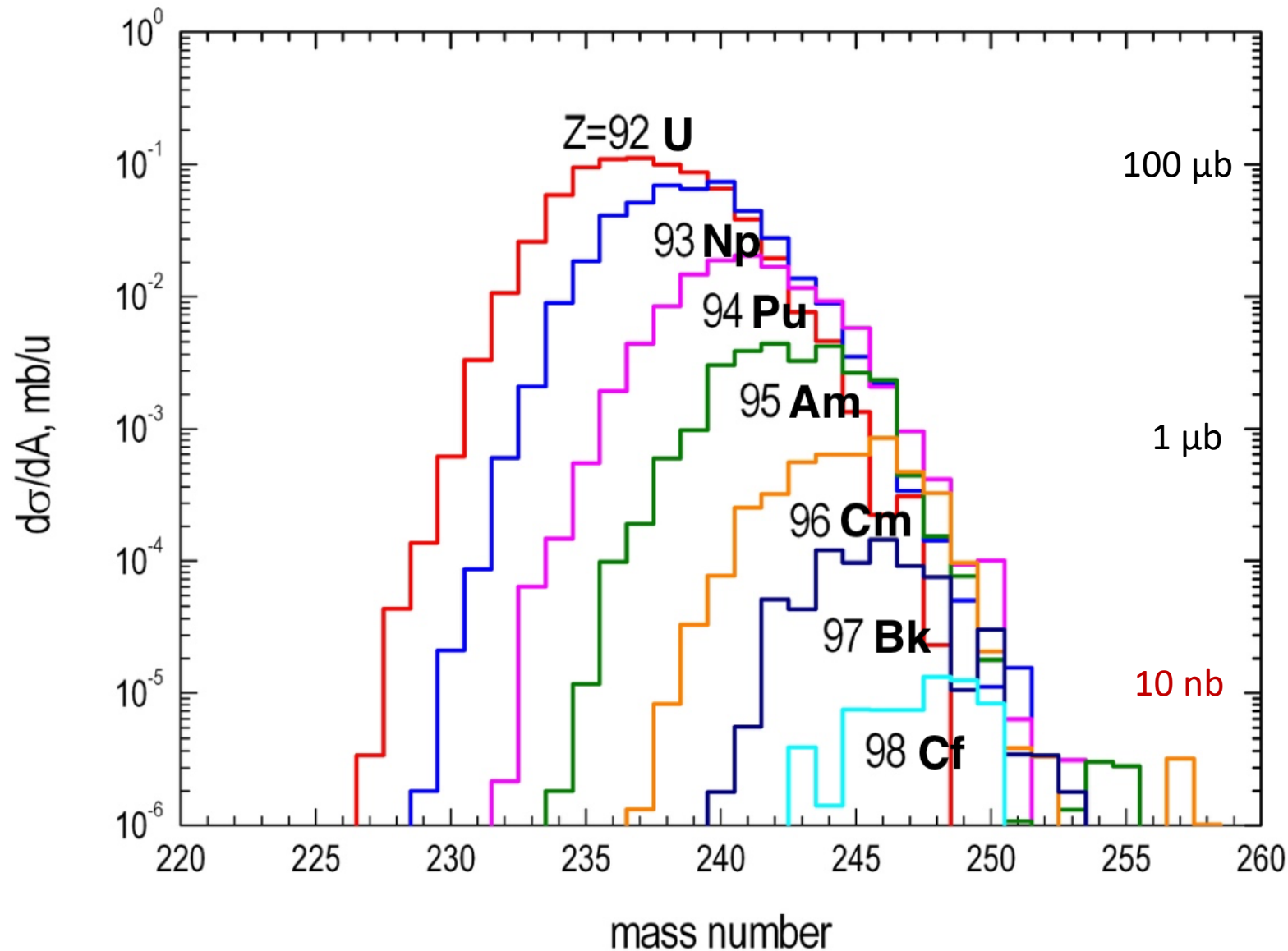
- **#1786 experiment at Argonne:  $^{136}\text{Xe} + ^{238}\text{U}$  at  $0^\circ$**

- Theoretical computations
- Simulations
- Experimental set-up

- First results

- Limits, perspectives and outlooks



The experiment:

ZF, D. Seweryniak et al:

« Synthesis of heavy and superheavy neutron-rich nuclei in multinucleon transfer reactions close to  $0^\circ$  »

The reaction:

$^{136}\text{Xe} + ^{238}\text{U}$  at  $0^\circ$

@3.6, 4.4 and 5.9 MeV/u, 20 pnA  
Gammasphere-AGFA-XArray

← Courtesy of A. Karpov  
(using the Langevin's model)

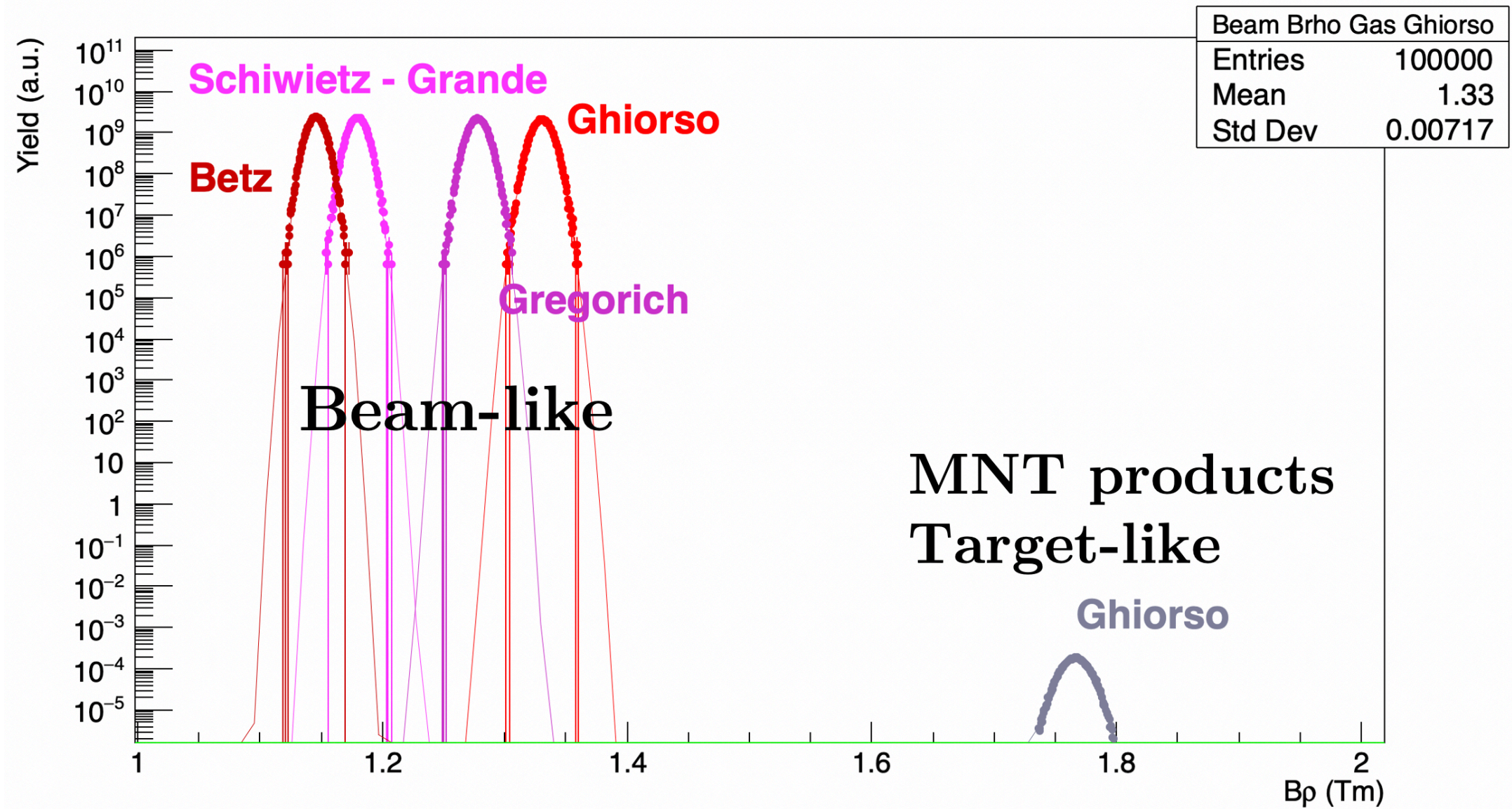
$E_{\text{cm}} = \sim 490 \text{ MeV}$

Integrated between  $0-5^\circ$



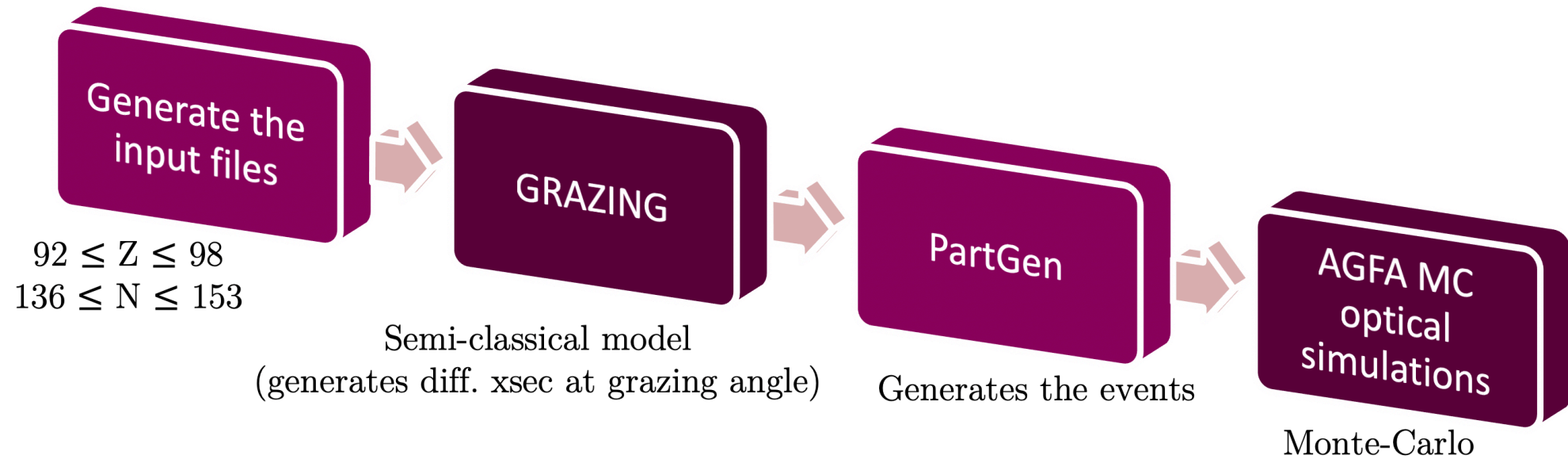
# Kinetic simulations (VGS code)

Developed by Ch. Theisen





# Optical simulations (transmission of AGFA)



## ➤ GRAZING Code:

(greetings to M. Siciliano)

- Aim: to produce an event generator to feed the AGFA MC simulation
- Estimates the Q-value (with ame2012.cal)
- Generates the output files of the event generator

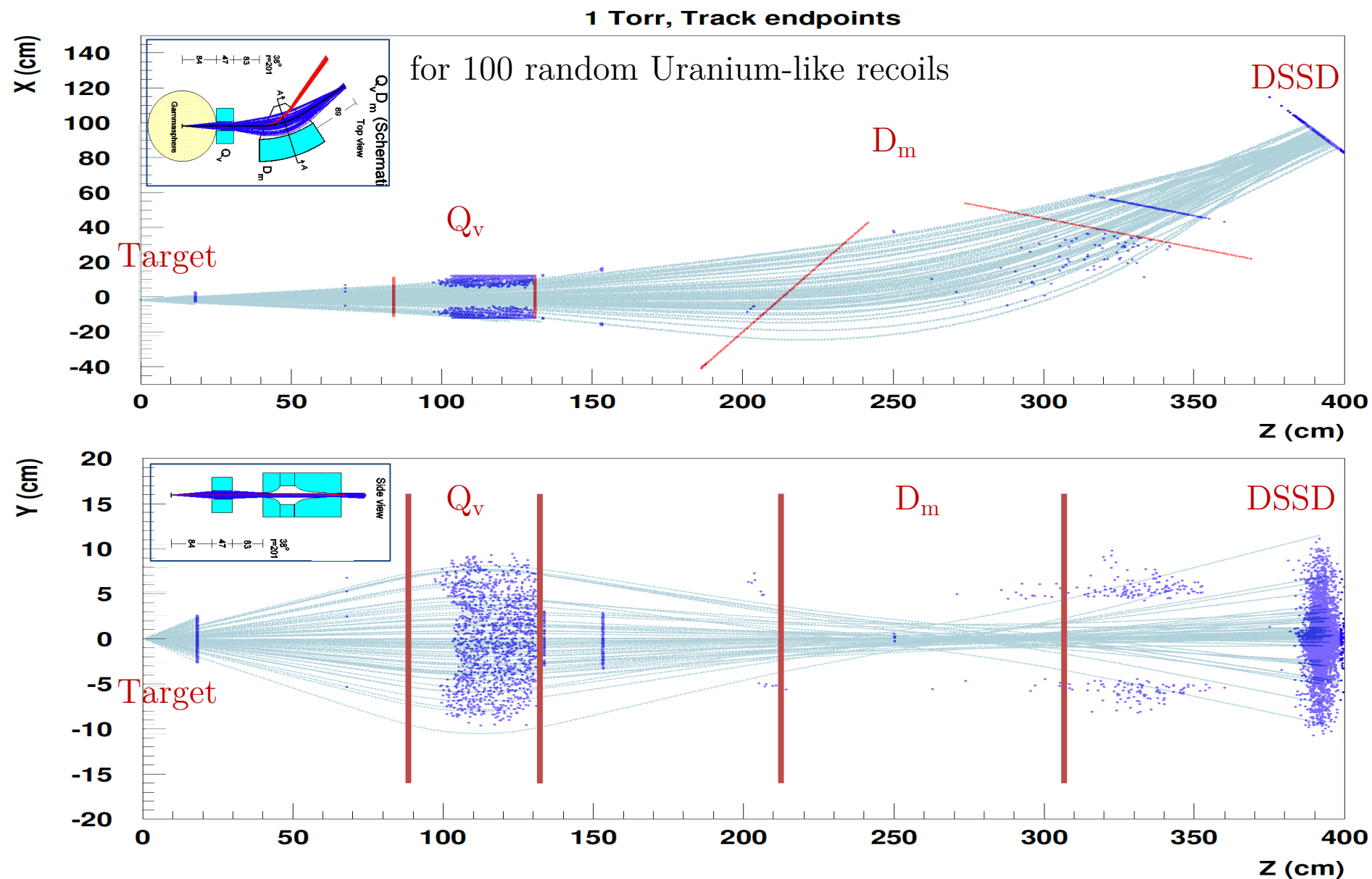
$Z$     $A$     $Q$     $\bar{Q}$     $E$     $\vec{x}$     $\frac{\vec{p}}{\|\vec{p}\|}$     $B\rho$

## ➤ AGFA Monte-Carlo simulation:

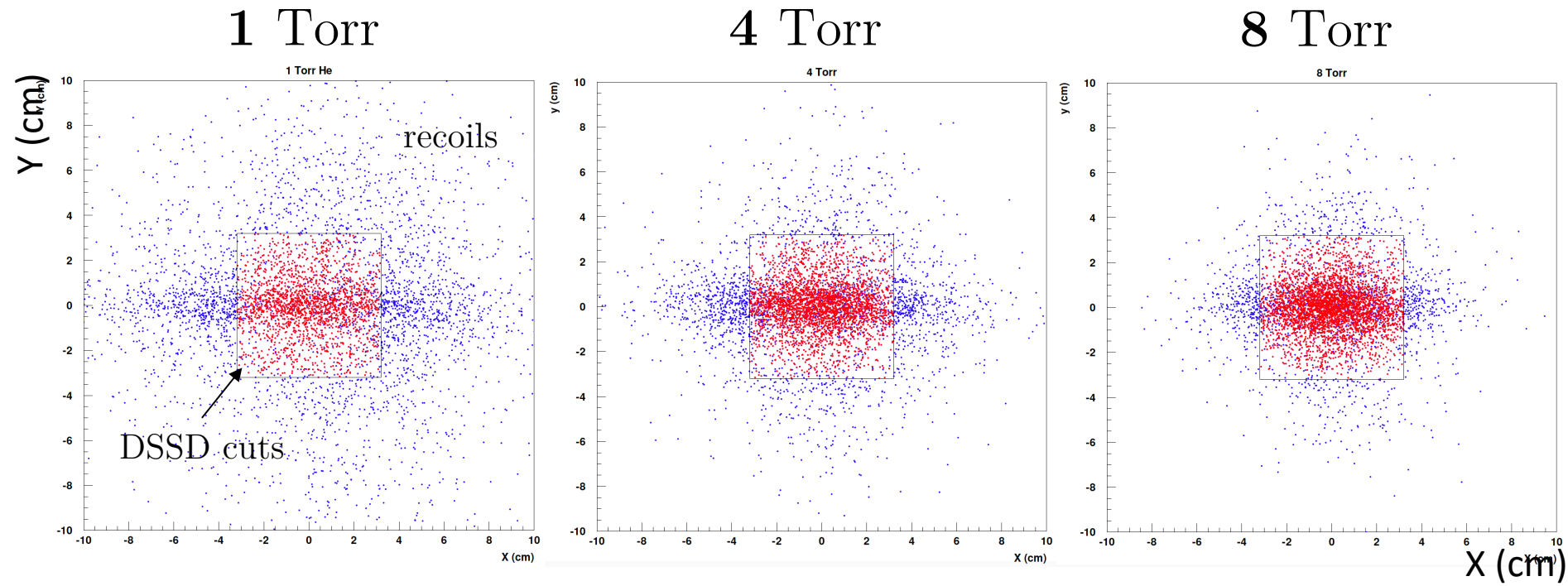
(greetings to D.H. Potterveld)

- Charge state model: A. Ghiorso et al. NIM in PRA 269 (1988) 192-201
- Optical simulations of AGFA

# AGFA Monte-Carlo simulations



# AGFA Monte-Carlo simulations

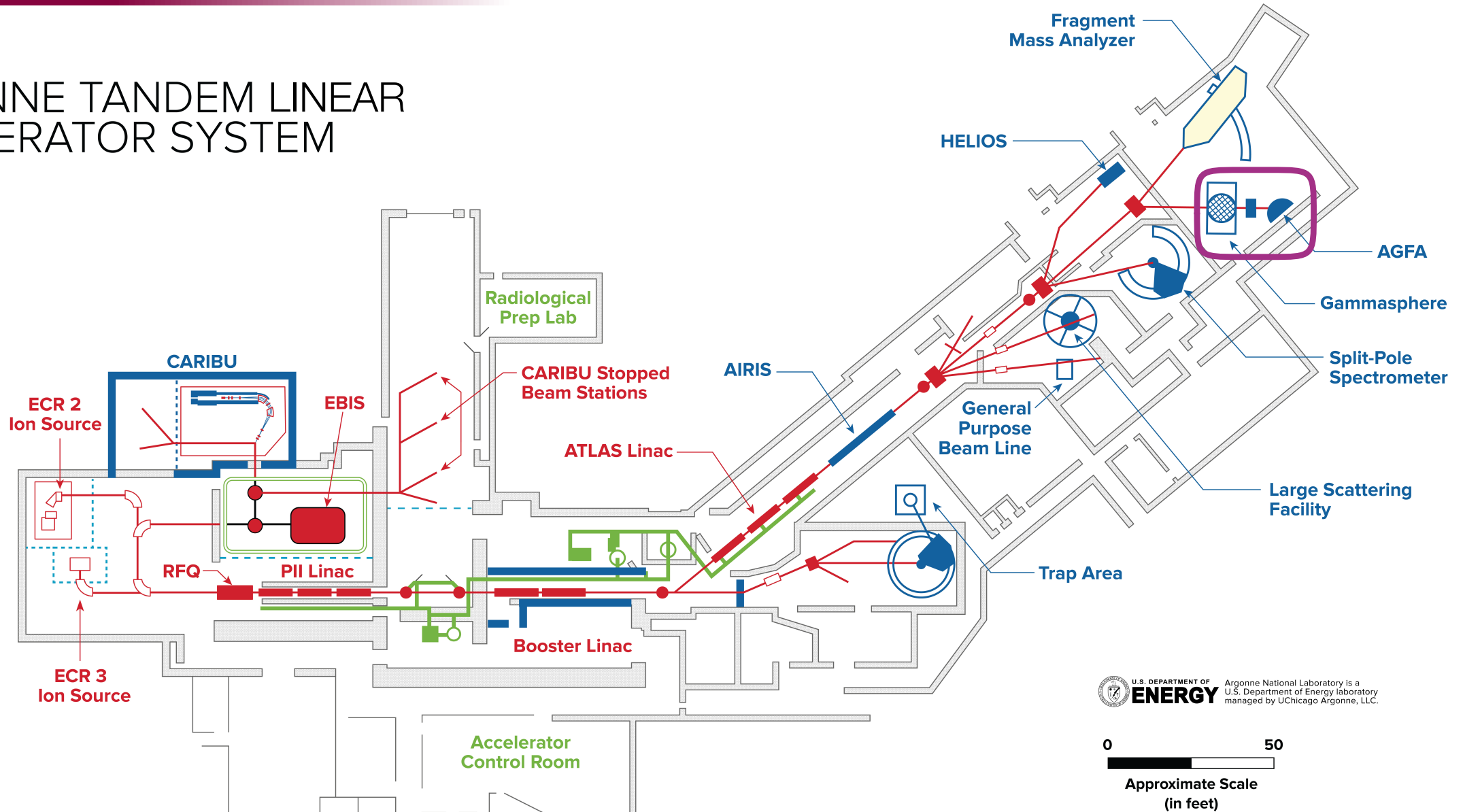


Pressure (Torr)	$N_{AGFA}$	$N_{DSSD}$	Transmission to the DSSD (%)
0.44	4735	1275	13.6
1.0	5159	1936	20.7
2.0	5215	2229	23.8
4.0	5384	3106	33.2
8.0	5114	3827	40.9

**~28% at 3 Torr**

# ATLAS at Argonne

## ATLAS ARGONNE TANDEM LINEAR ACCELERATOR SYSTEM



U.S. DEPARTMENT OF  
**ENERGY**

Argonne National Laboratory is a  
U.S. Department of Energy laboratory  
managed by UChicago Argonne, LLC.

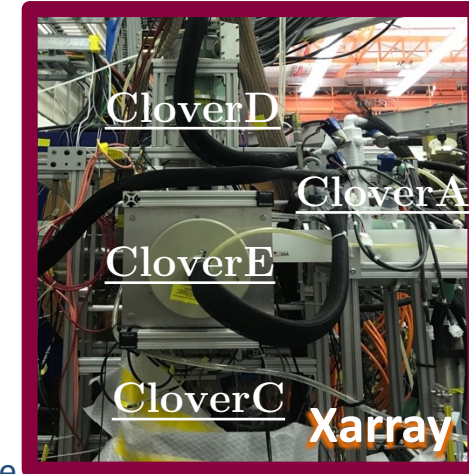
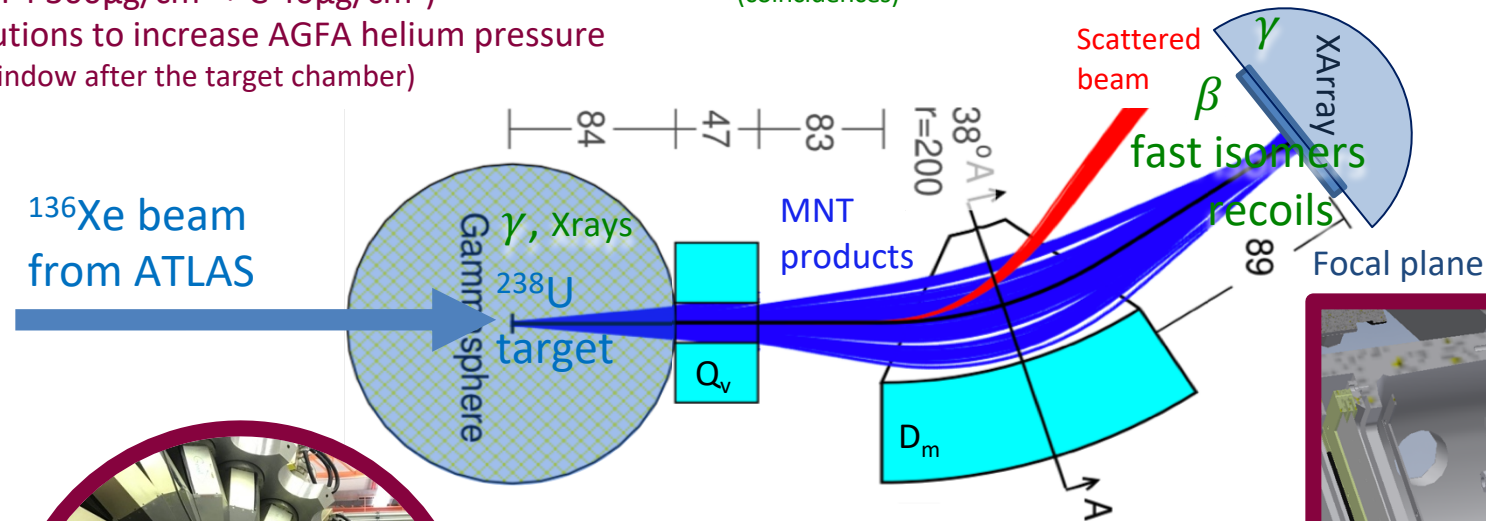
0 50  
Approximate Scale  
(in feet)



# Multinucleon transfer reactions (MNT) at Argonne

- Proof-of-principle experiment to study neutron-rich heavy nuclei using MNT reactions
- 1<sup>st</sup> MNT reaction using AGFA @ANL to produce heavy nuclei
- **Beam:**  $^{136}\text{Xe}$  @605, 705 and 809MeV
- **Target:**  $^{238}\text{U}$  (UF<sub>4</sub> 300μg/cm<sup>2</sup> + C 40μg/cm<sup>2</sup>)
- Innovative solutions to increase AGFA helium pressure up to 3Torr (Ti window after the target chamber)

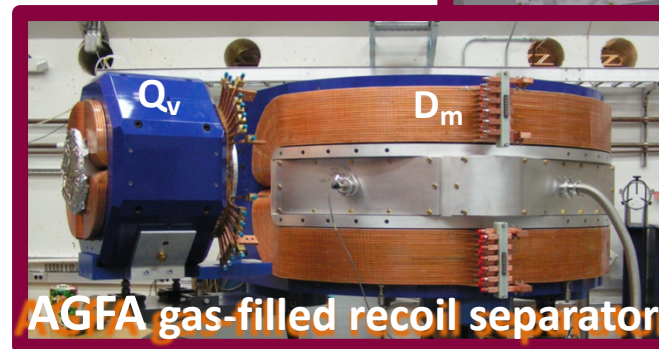
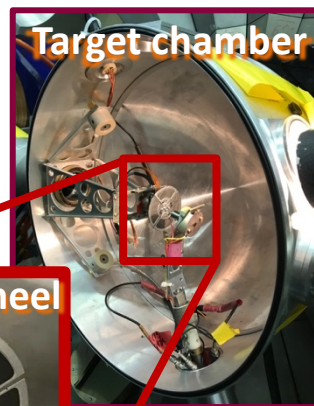
## Target-like products identification (coincidences)



Decay spectroscopy



Prompt spectroscopy



# Experimental conditions

5  $B\rho$

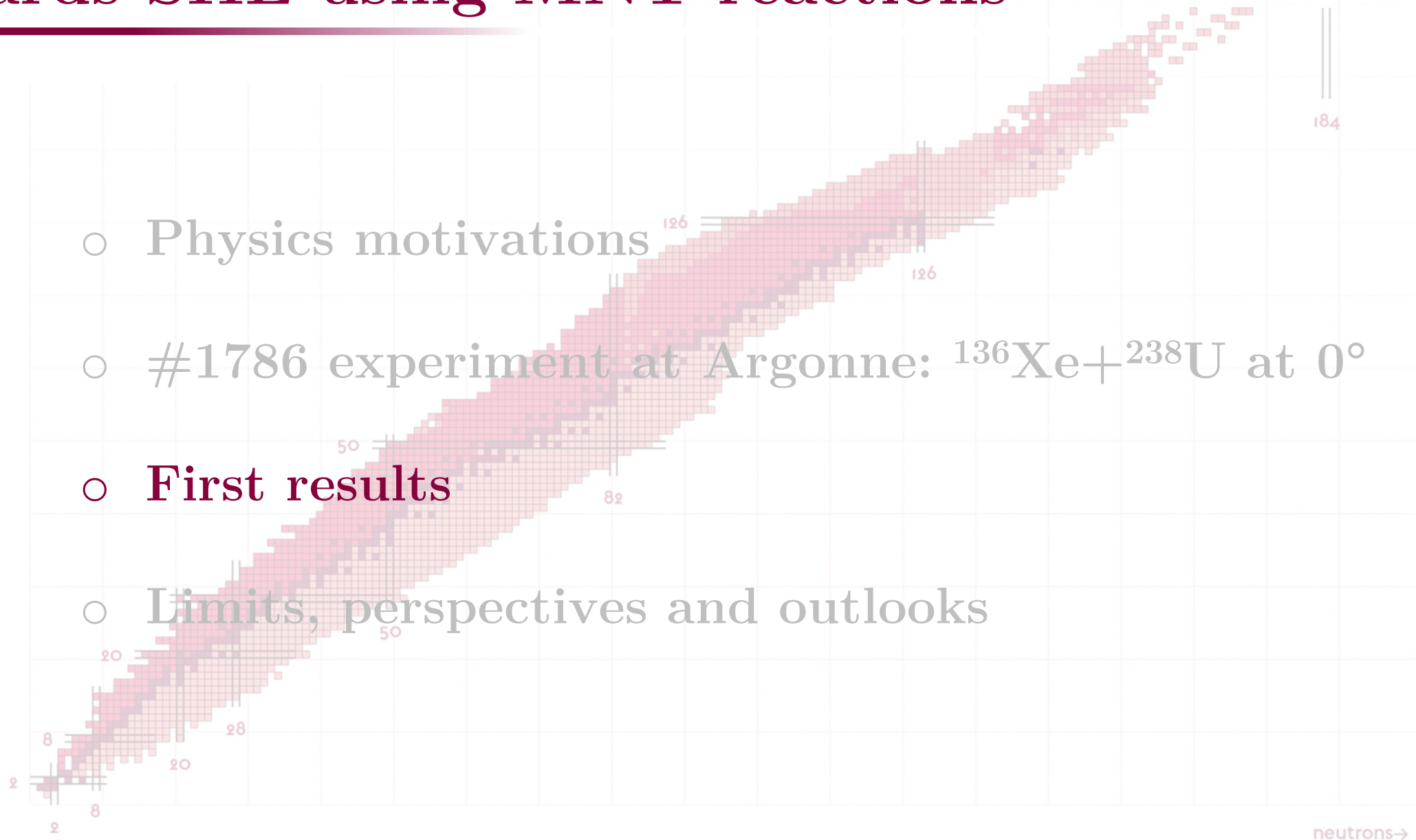
3 energies

Beam energy (MeV)	Run numbers	Theor. $\beta=v/c$ (%)
No beam	1->19 ans 117->129	0
605	20->50	6.91
705	<b>73-&gt;116</b>	<b>7.55</b>
809	51->72	8.16

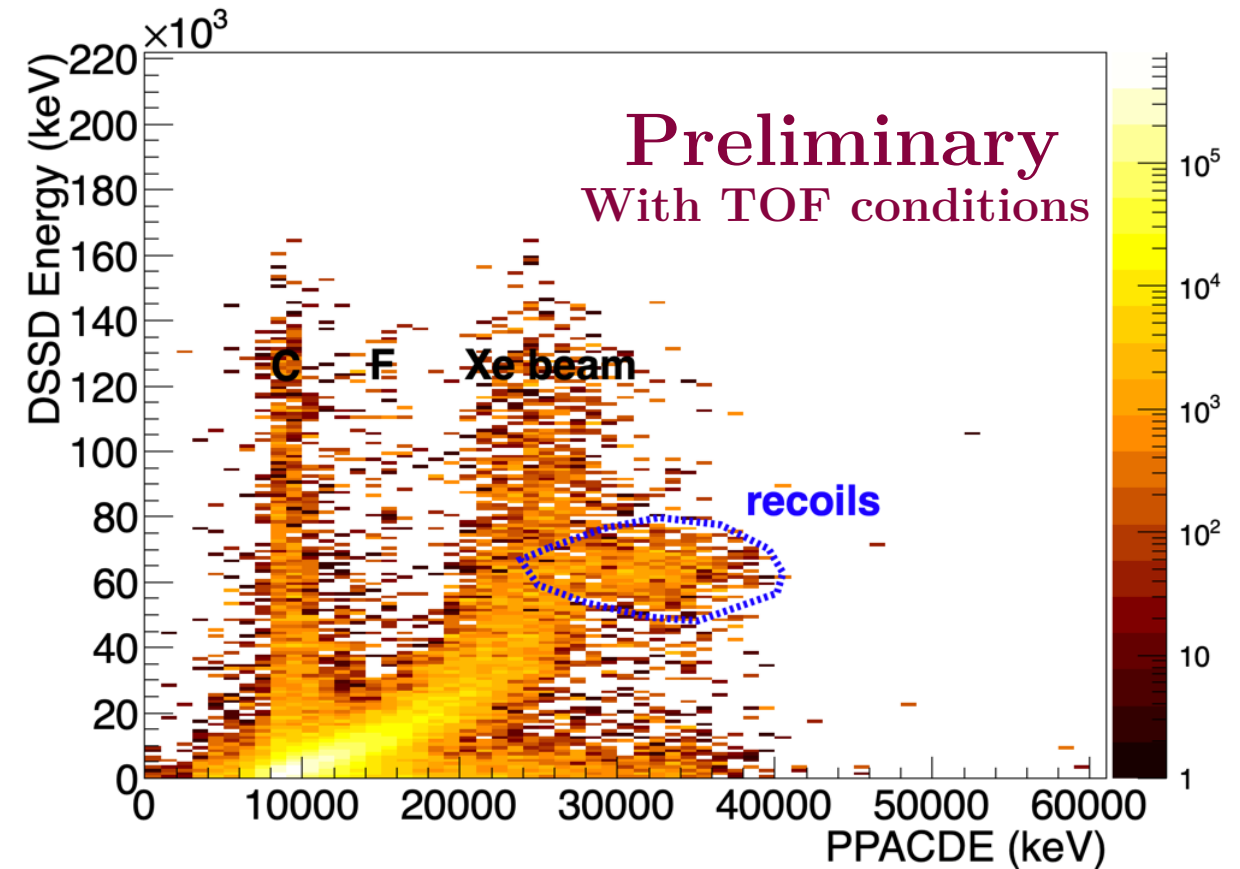
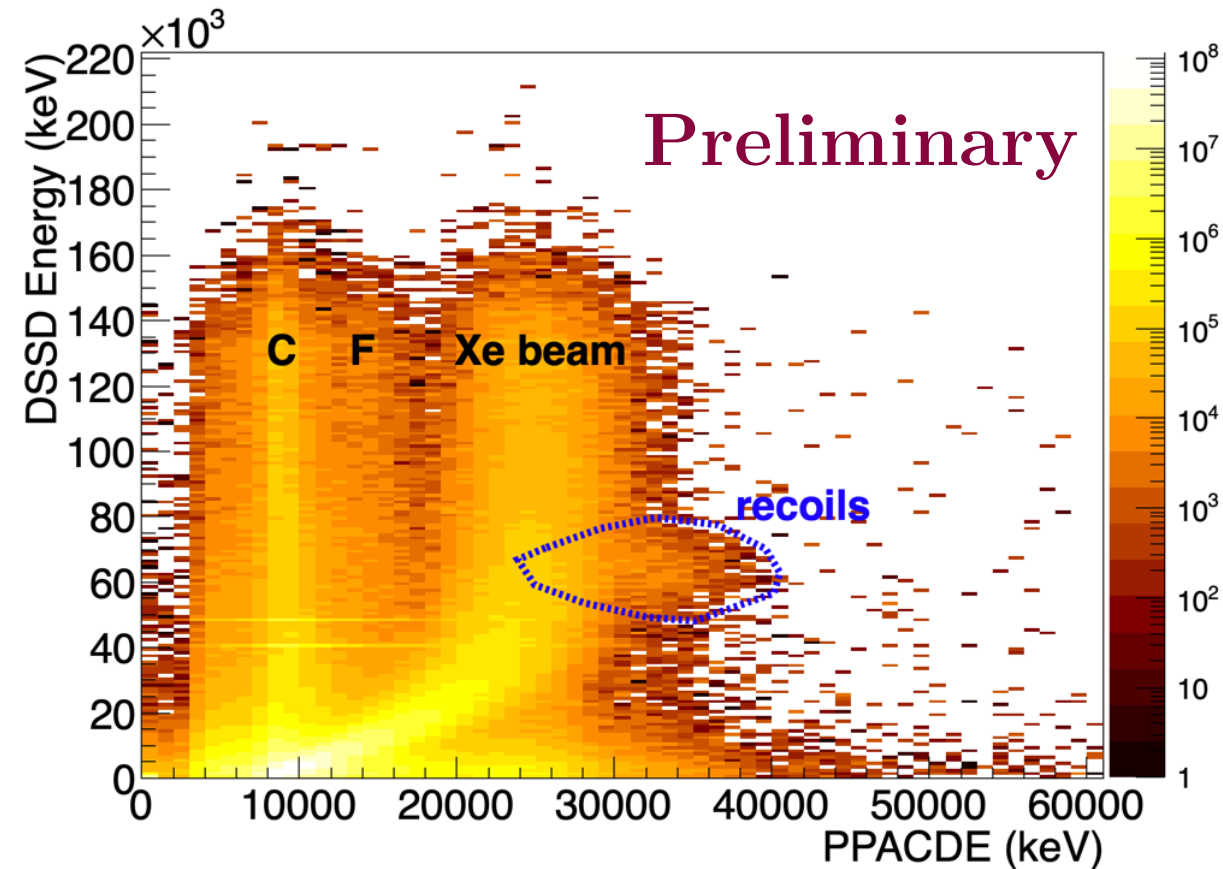
Beam energy (MeV)	Run numbers	$B\rho$ (Tm)
605	20	2.4
	21->27	1.8
	28	2.0
	29	2.2
	30	2.4
	21->50	1.8
705	<b>73-&gt;116</b>	<b>1.8</b>
809	51->66	1.8
	67->72	1.9

# Towards SHE using MNT reactions

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- #1786 experiment at Argonne:  $^{136}\text{Xe} + ^{238}\text{U}$  at  $0^\circ$
- **First results**
- Limits, perspectives and outlooks

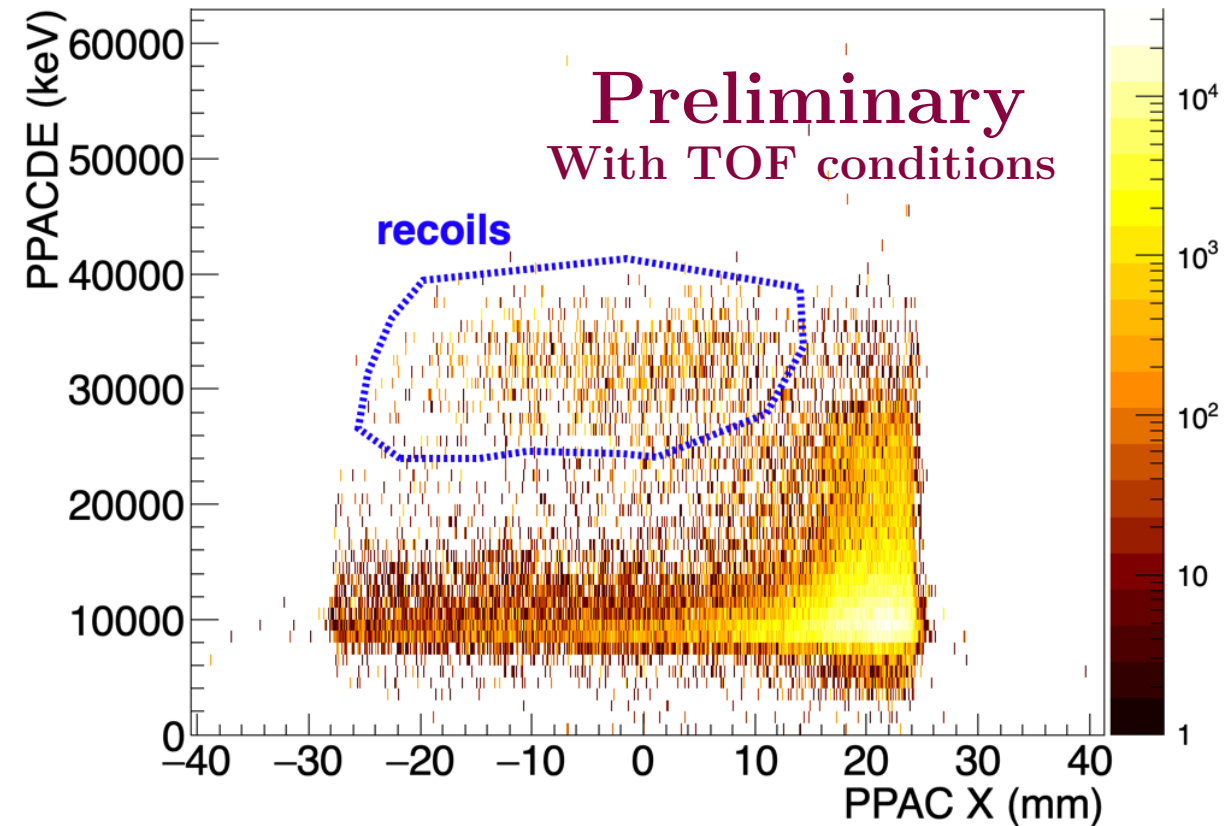
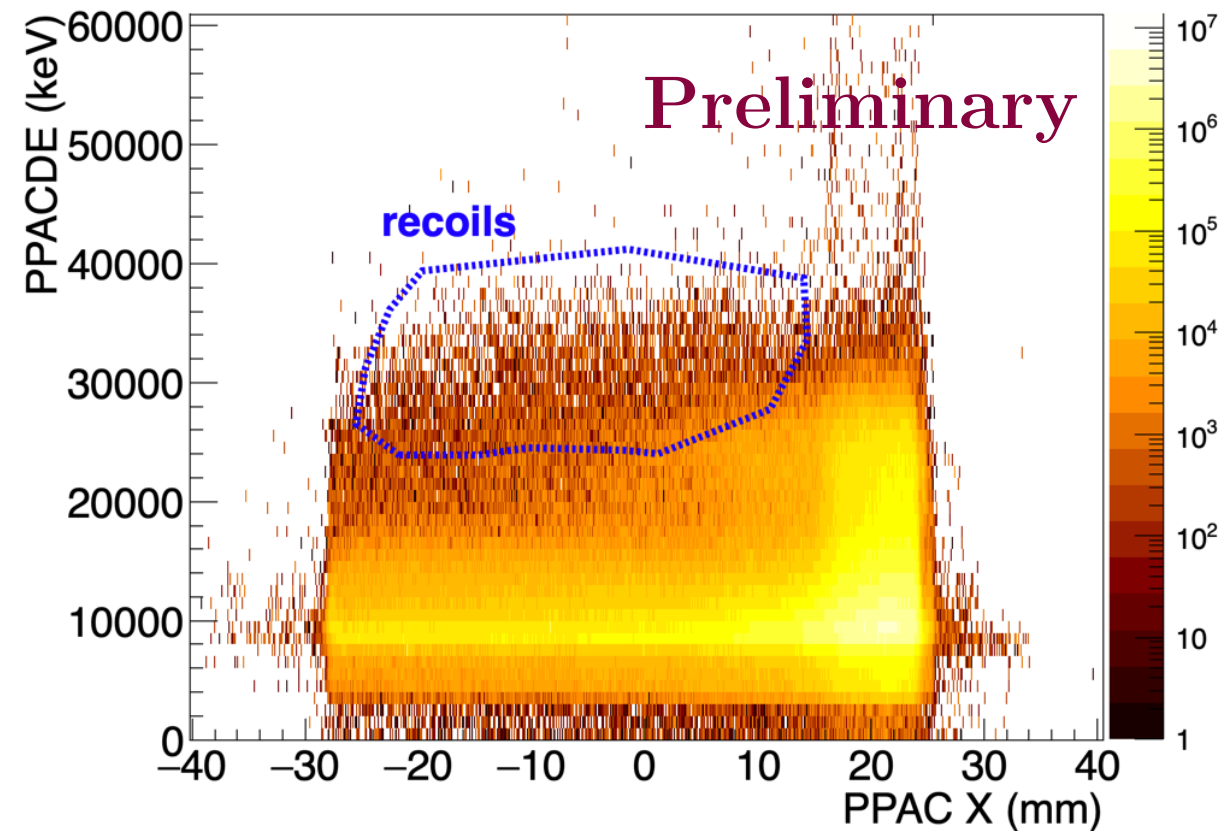


# Identification of neutron-rich nuclei (recoils)

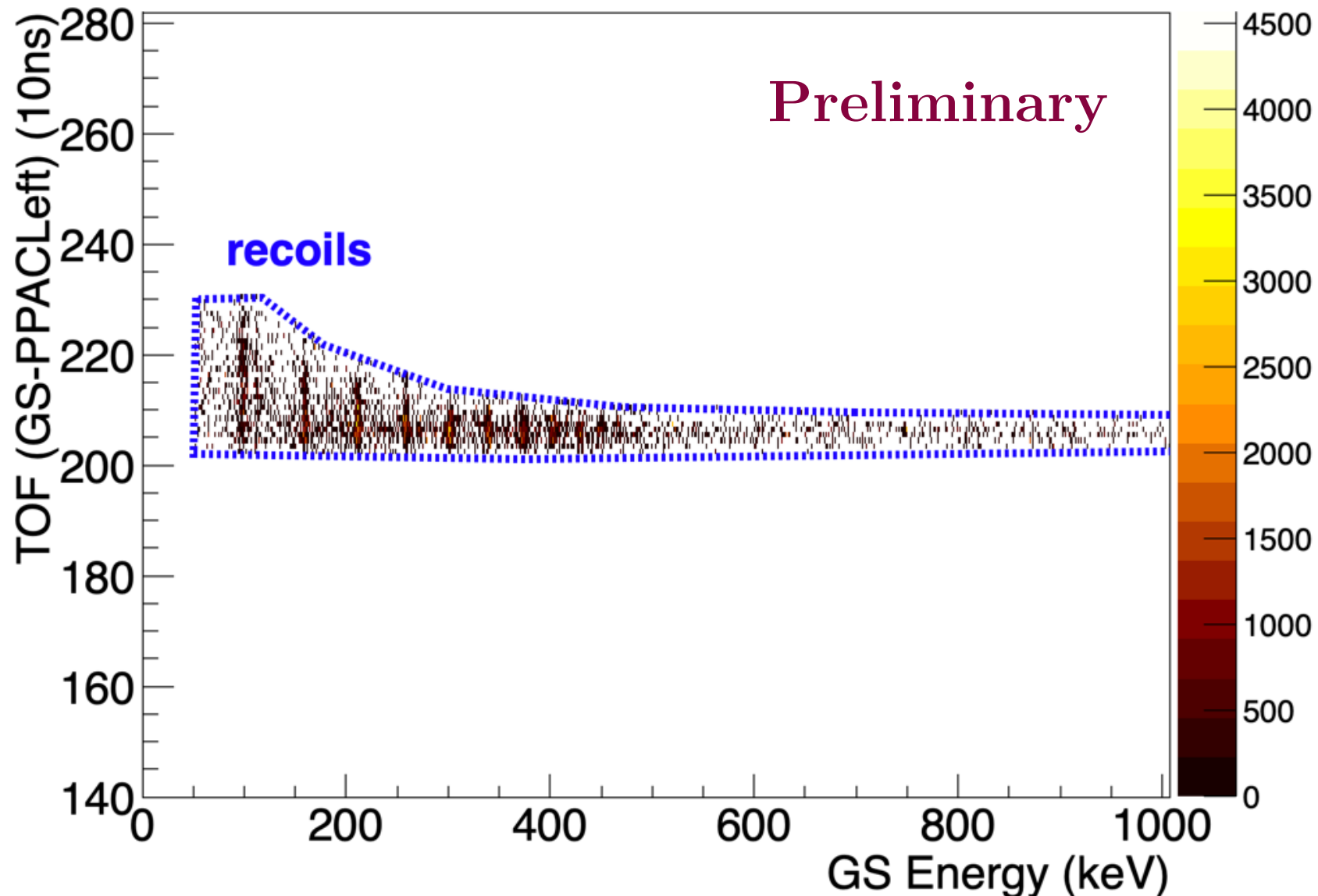




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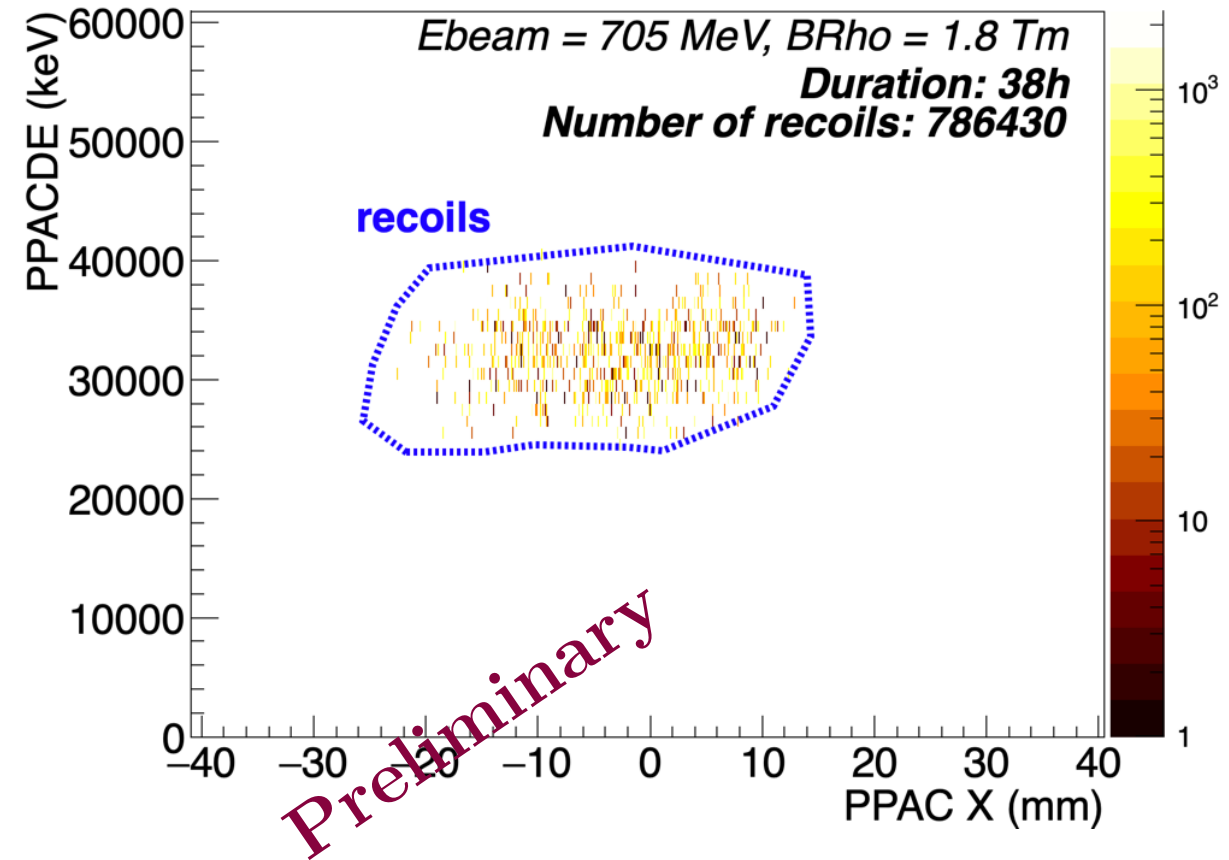
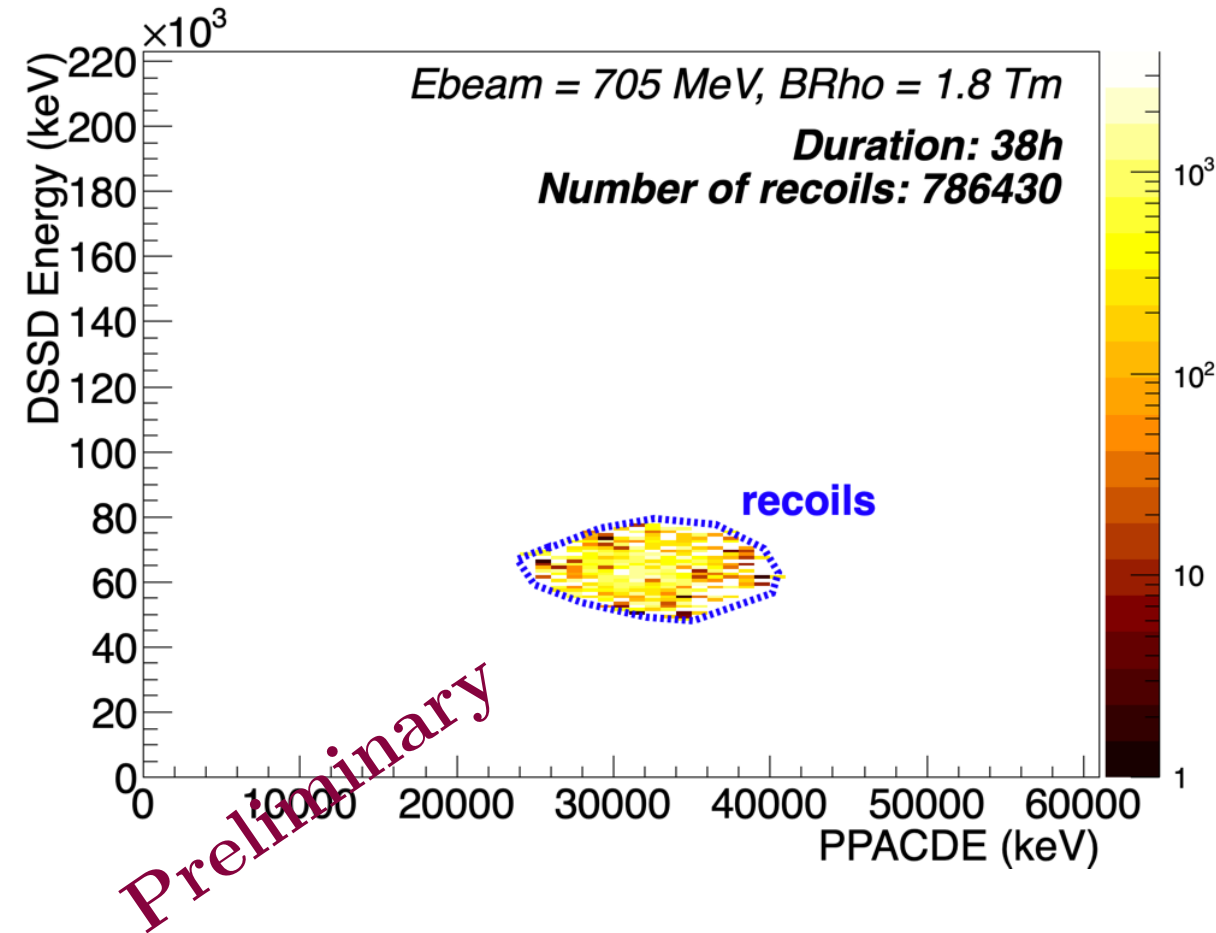


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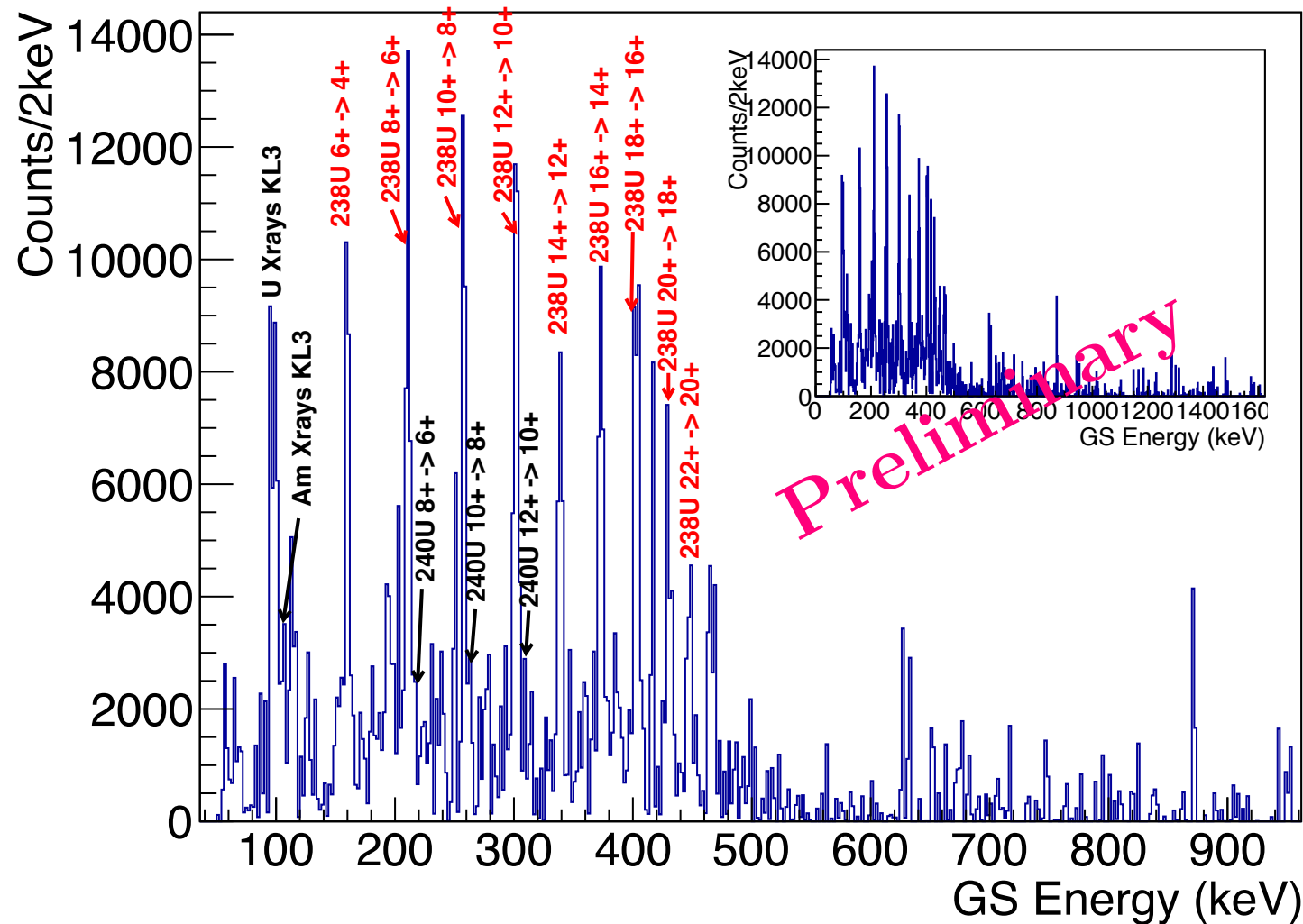


# Identification of neutron-rich nuclei (recoils)

## Correlations in time and position



# Identification of neutron-rich nuclei (recoils)

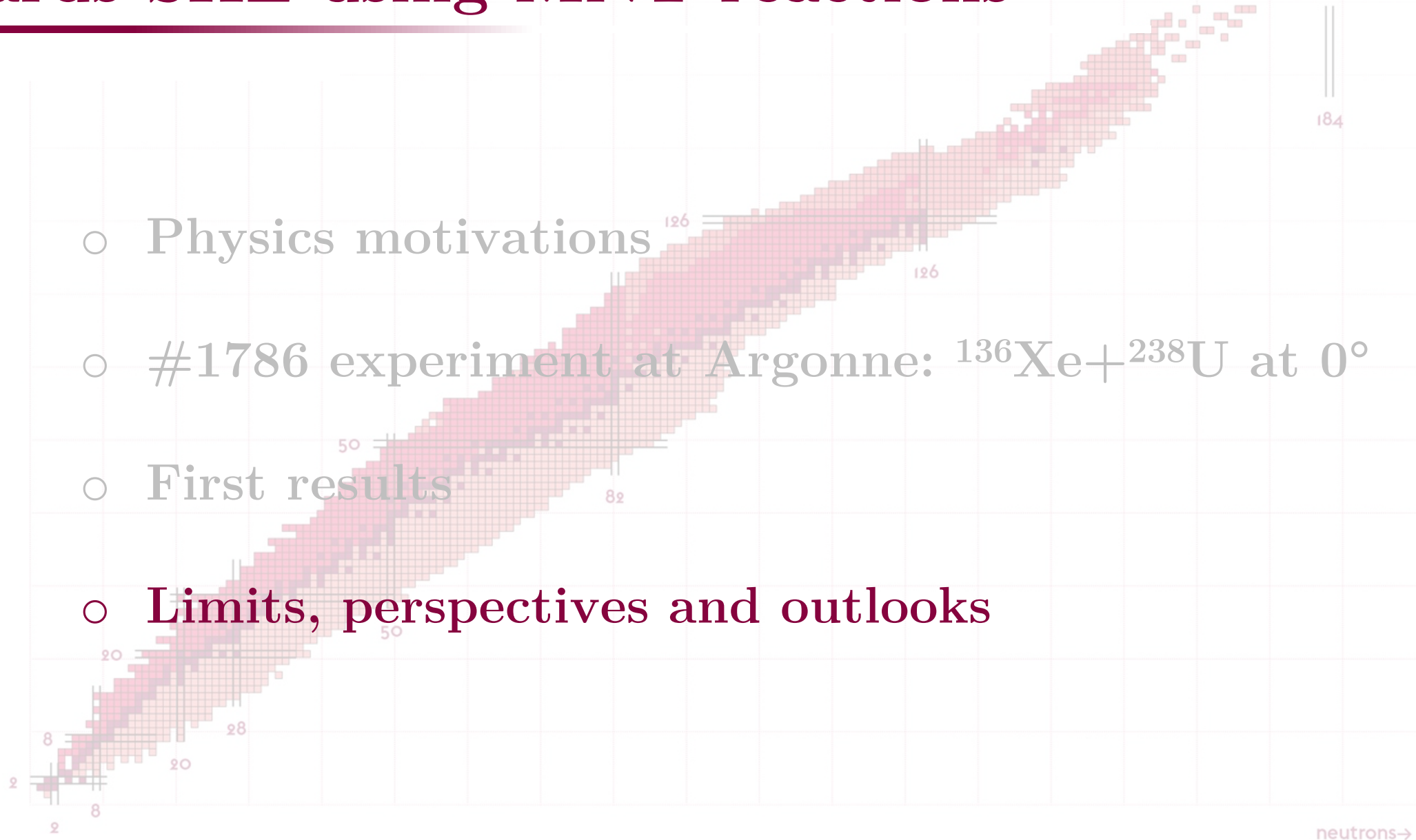


- $\gamma - \gamma$  correlations
- Transfer of 2n ( $^{240}\text{U}$ )
- Xrays of Am, Np, U (transfer of 1p, 3p)?
- Correlations with GS and XArray
- New accepted experiment at PAC2021 to get more statistics!

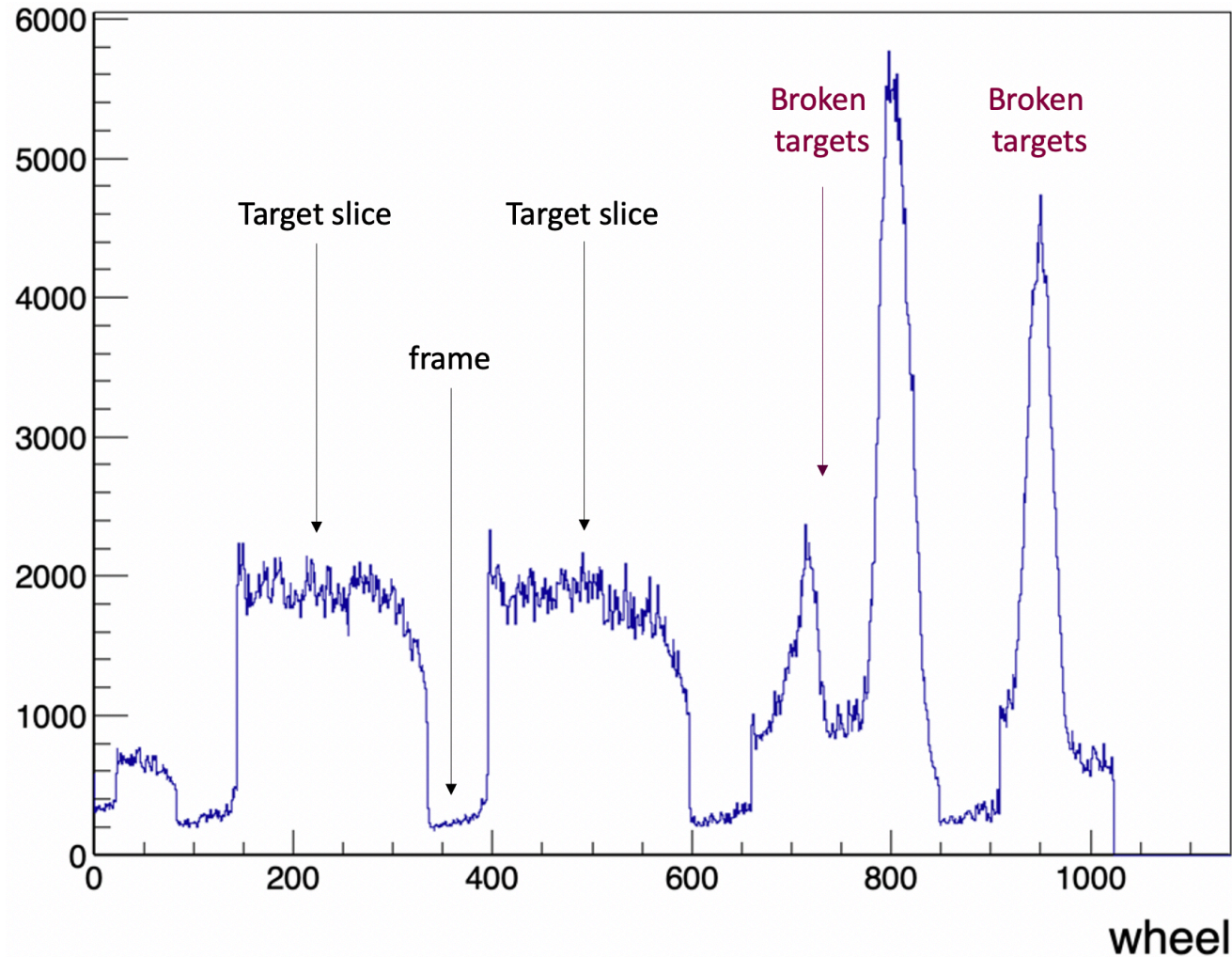


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# Main limits of the #1786 experiment



## Lack of statistics

### 6.5 Shift losses due to:

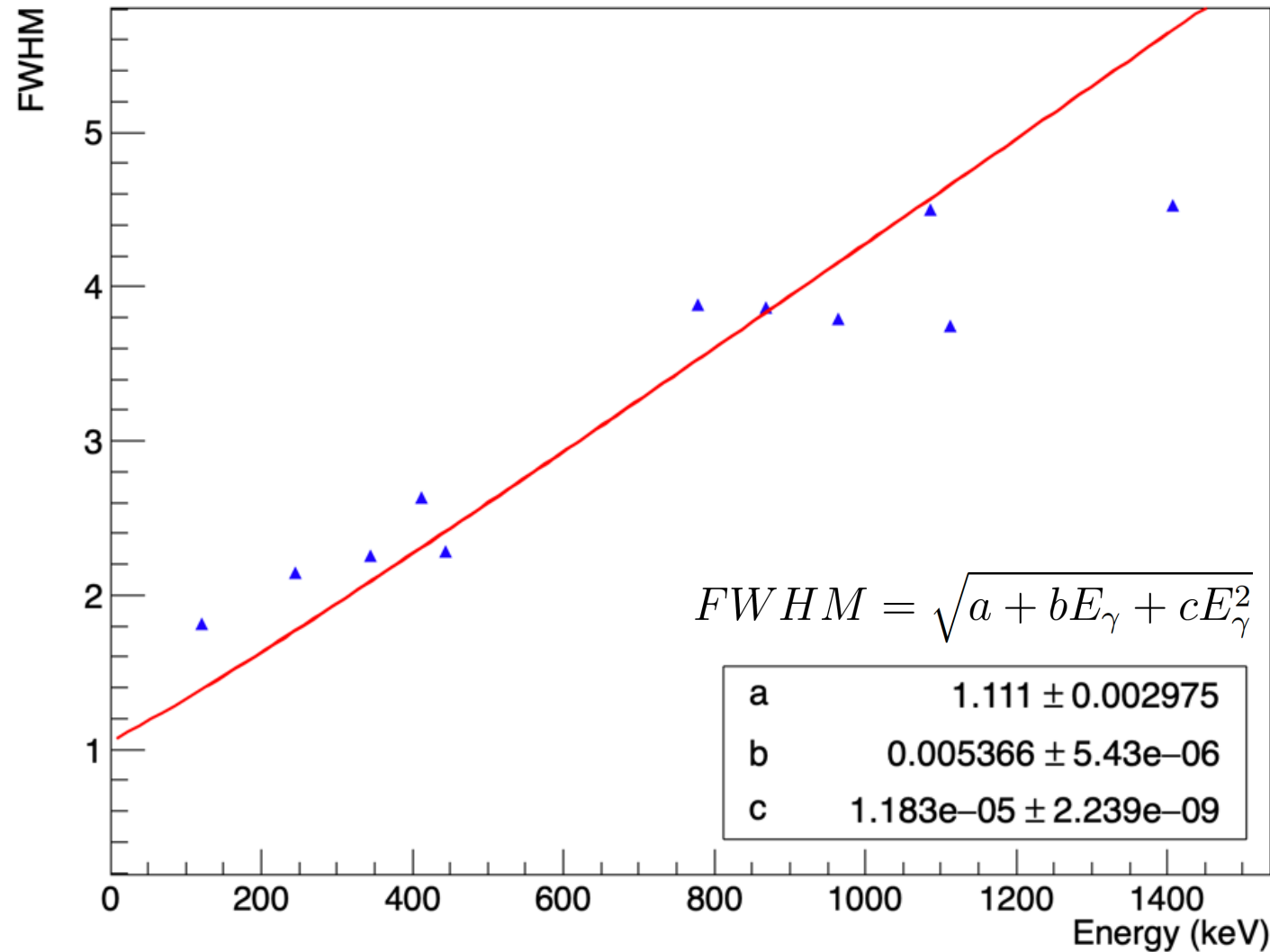
- DAQ issues (reboot IOTs)
- broken targets (mechanical constraints) had to vent, open the target chamber...

### Limitations :

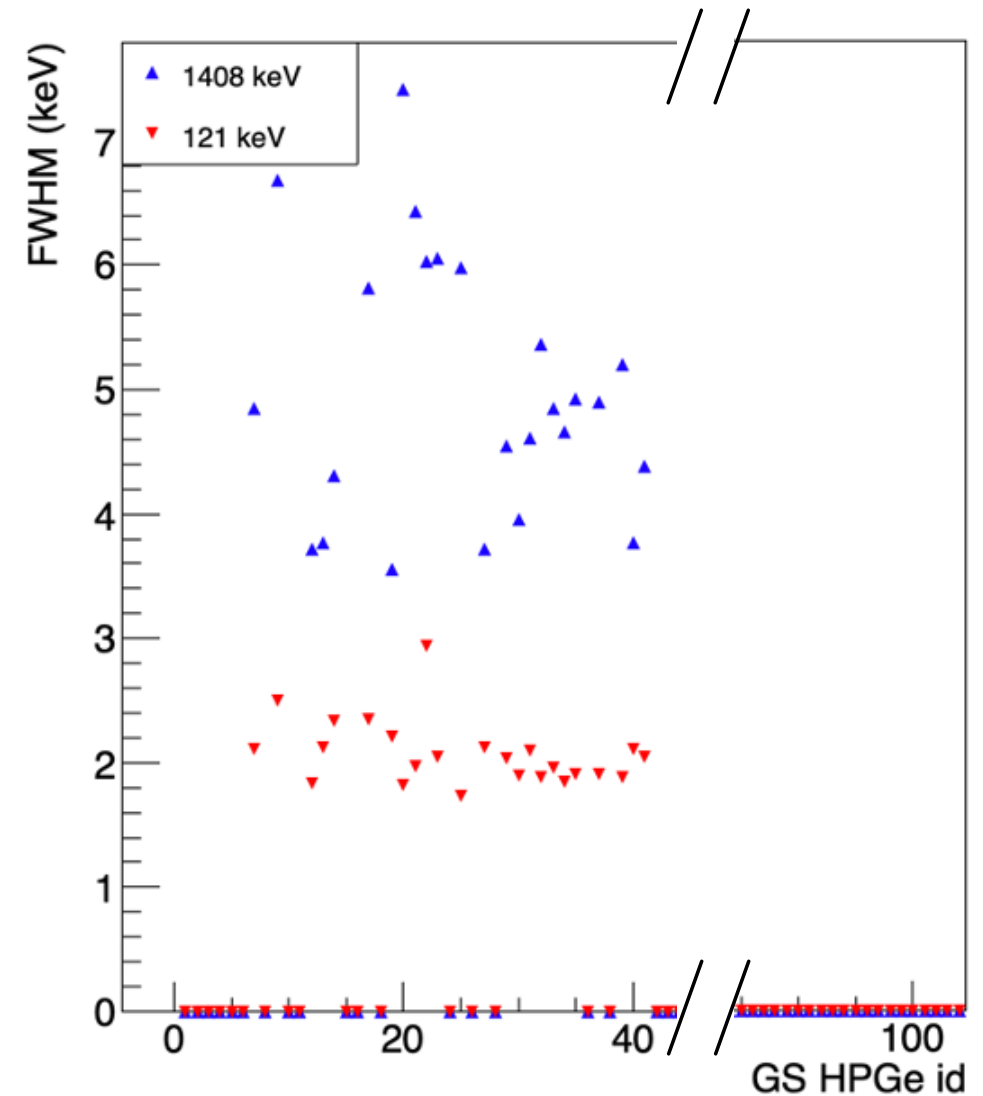
- APEX  $\text{UF}_4$  targets, lots of fission events... had to decrease the beam intensity
- Frequency in the GS crystals
- 70/110 working detectors of GS
- bad resolution of the GS crystals
- can not rely on alphas
- no mass identification

# GS Calibration with a $^{152}\text{Eu}$ source

FWHM(Energy) of the GS HPGe det8



$^{152}\text{Eu}$  calibration of Gammasphere



# New #1930 experiment accepted PAC 2021

- $^{136}\text{Xe} + ^{238}\text{U}$  at  $0^\circ$
- **New metallic targets manufactured at GSI**  
(thanks to B. Lomel, B. Kindler, and their team)  
Not the old APEC  $\text{UF}_4$  ones
- **Annealing of some GS detectors: 70  $\rightarrow$  110?**
- **More beam intensity (from 2.25 to 20pnA)**
- **More gas pressure from 3 to 5 Torr in AGFA**
  
- **PAC2019 (#1786) For the  $E_{\text{beam}}=705$  MeV:**  
We collected 786 430 recoils, within 38h at 2.25pnA  
i.e a rate of 20755 recoils/h  
These recoils emitted  $4.7 \cdot 10^7$  single-gammas.
  
- **PAC 2021 (#1930) For the  $E_{\text{beam}}=705$  MeV:**  
Running at 20 pnA with the same beam for 110 working detectors  
in Gammasphere (instead of 70), during 4 days (96 hours)  
we expect  $1.7 \cdot 10^7$  recoils (**gain of factor 20**)  
emitting  $1.7 \cdot 10^8$  single-gammas (**gain of factor 35**).



# Conclusions and perspectives

- Simulations and particle generator
  - Kinetic and optical simulations
  - Helped us predict the separation power of AGFA for MNT reactions
- MNT at Argonne
  - First proof-of-principle experiment of MNT using AGFA
  - Good separation of the BLFs and TLFs with AGFA
  - First transfers observed to be confirmed with the new experiment (PAC2021)
- MNT at GSI/FAIR
  - First experiments with TASCA and SHIP for SHE with MNT
  - See the talk of Timo Dickel
- MNT for VHE and SHE
  - A new path for neutron-rich VHE and SHE
  - Towards the “Island of Stability”? Mass identifications with gas-cells?



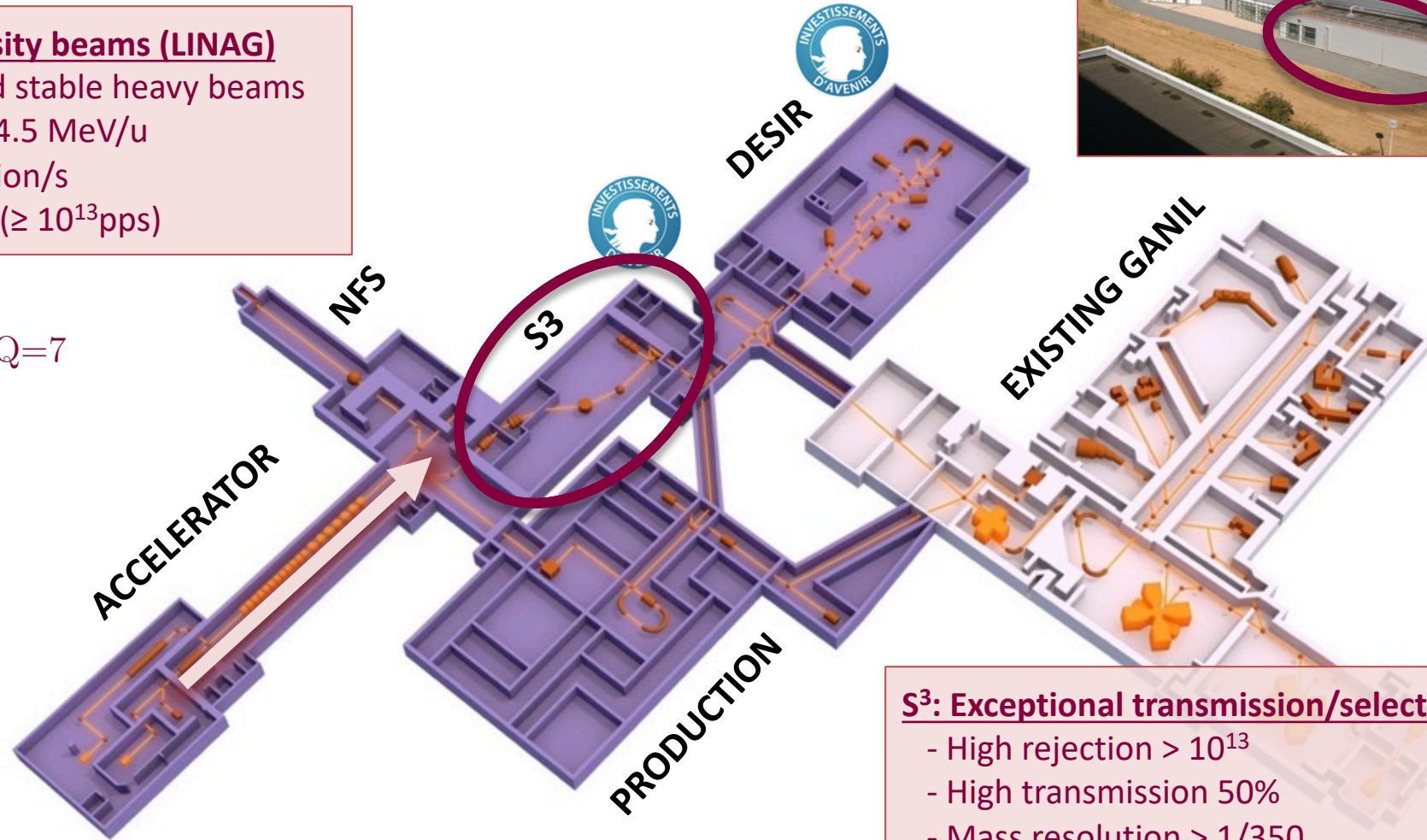
# S<sup>3</sup> in the SPIRAL2 project

## Very high intensity beams (LINAG)

- Deuterons and stable heavy beams
- $E_{\text{beam}}$  = up to 14.5 MeV/u
- Intensity  $10^{14}$  ion/s  
beyond  $1\mu\text{A}$  ( $\geq 10^{13}$ pps)

A/Q=3

NEWGAIN A/Q=7



## S<sup>3</sup>: Exceptional transmission/selection combination

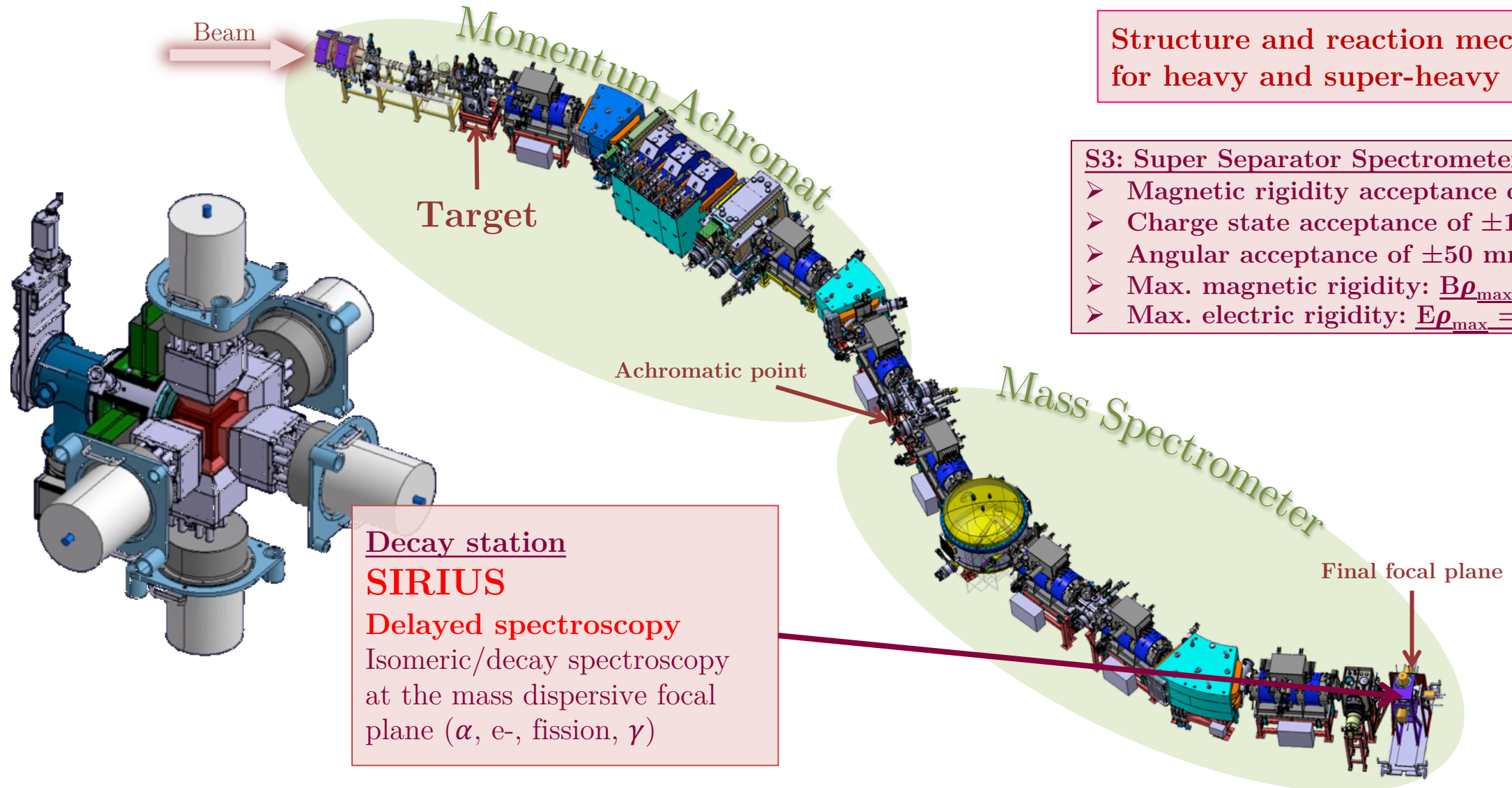
- High rejection  $> 10^{13}$
- High transmission 50%
- Mass resolution  $> 1/350$

A cutting-edge instrumentation for S<sup>3</sup>  
(SIRIUS, LEB, FISIC...)





# S<sup>3</sup>: Super Separator Spectrometer



Structure and reaction mechanism  
for heavy and super-heavy nuclei

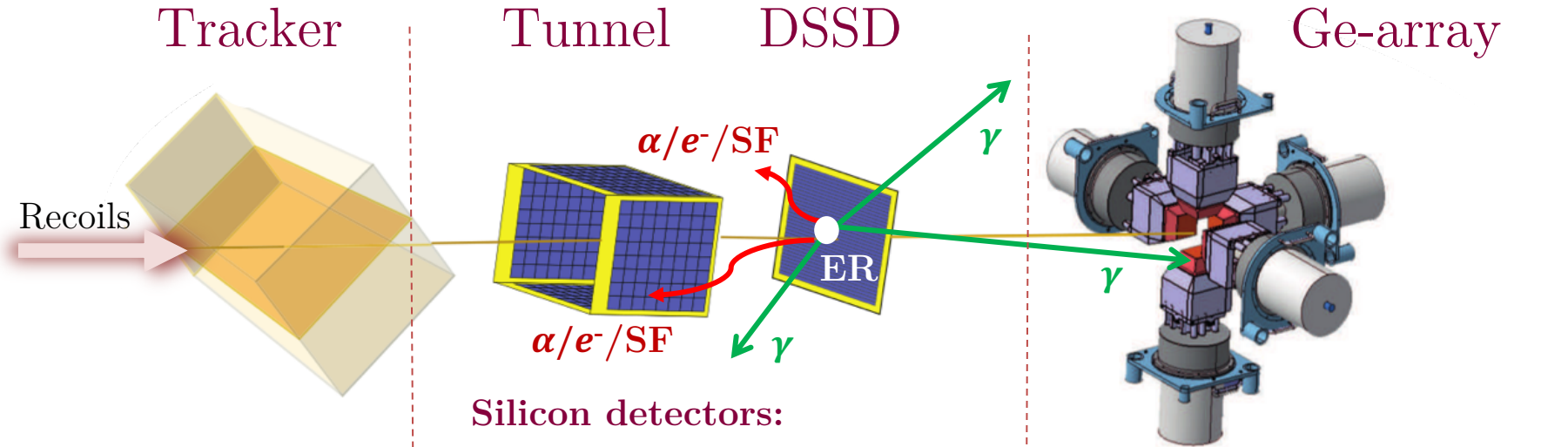
## S3: Super Separator Spectrometer

- Magnetic rigidity acceptance of  $\pm 7\%$
- Charge state acceptance of  $\pm 10\%$
- Angular acceptance of  $\pm 50$  mrad
- Max. magnetic rigidity:  $B\rho_{\max} = 1.8\text{Tm}$
- Max. electric rigidity:  $E\rho_{\max} = 12\text{MV}$



# SIRIUS decay station

SIRIUS: System for the Investigation of Recoiling Ions Using S3



## Time of Flight:

- Emissive foil
  - Thin windows
  - High Time resolution
  - Mass Identification
- $A/\Delta A \sim 300$

## Silicon detectors:

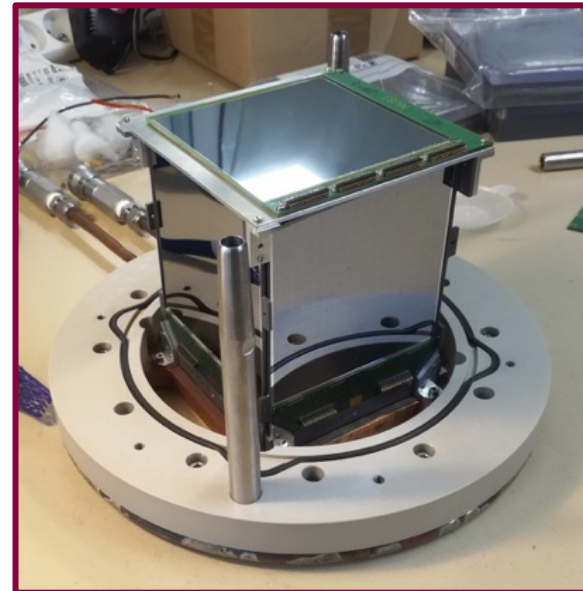
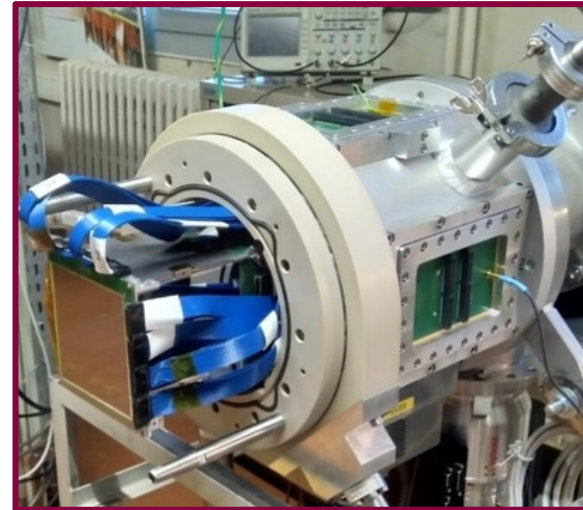
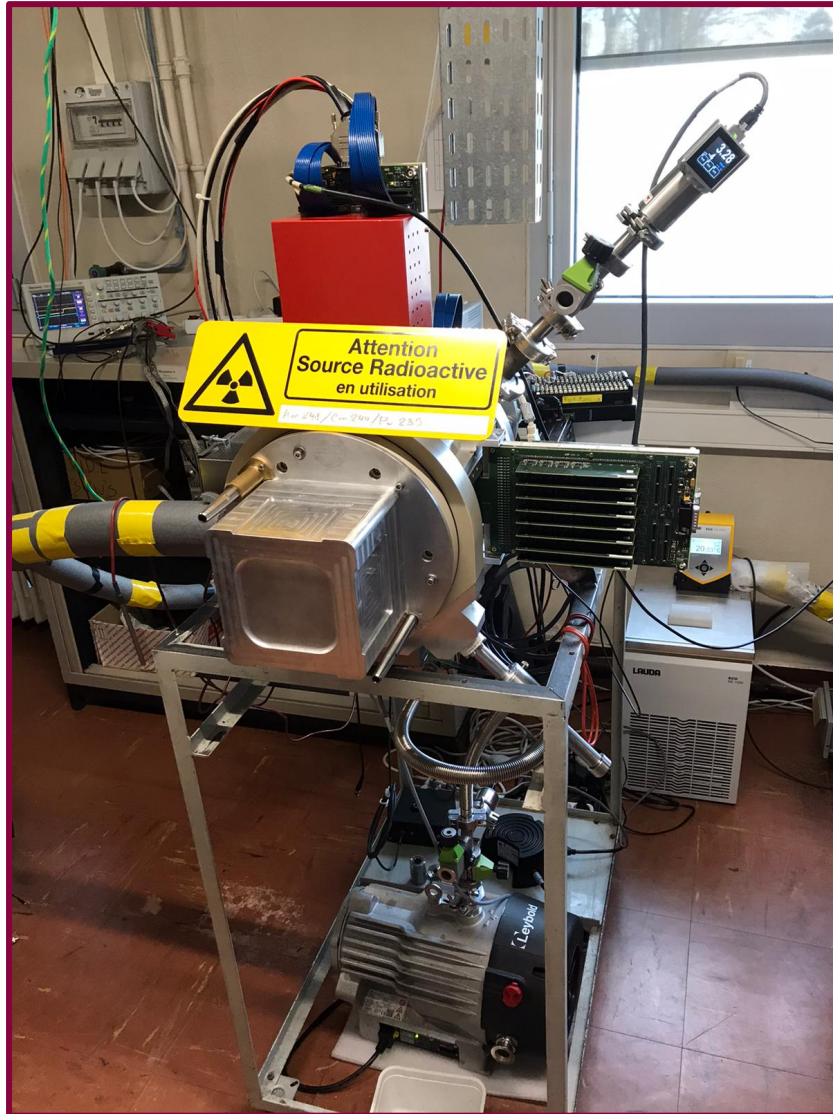
- Charged particle discrimination for recoil, beta and decay alpha
- High resolution alpha and conversion electron spectroscopy
- Measurement of TKE for spontaneous fission
- Access to short decay times

## $\gamma$ -ray detection :

- 5 EXOGAM clover detectors
- Efficiency of 40% at 121 keV

**Digital electronics:**  
Digital signal processing

# SIRIUS detectors



## Decay station

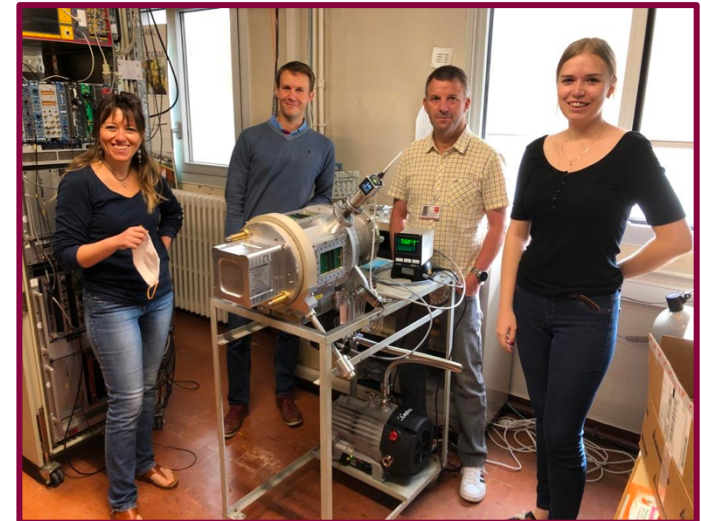
## SIRIUS

### Delayed spectroscopy

Since July 2020 at CEA/Irfu

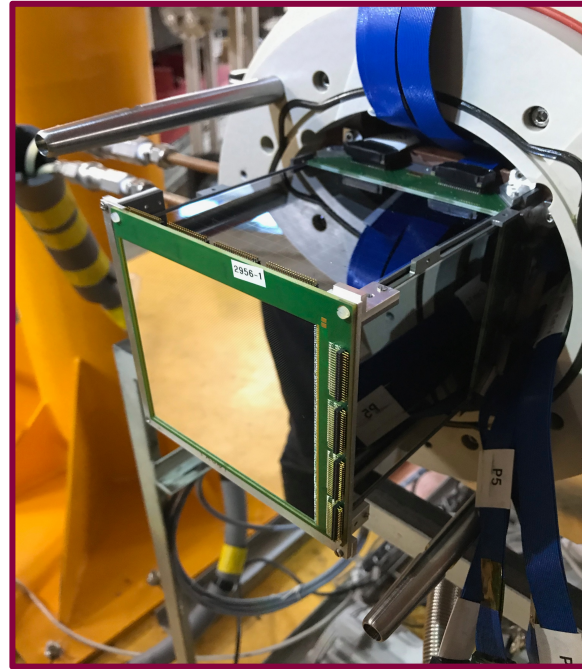
Complete chamber from Strasbourg.  
Tunnels from IJClab.

DSSD and electronics from CEA/Irfu.  
Offline commissioning (sources  $\alpha$ ,  $\beta^-$ )





# SIRIUS detectors



Decay station

**SIRIUS**

**Delayed spectroscopy**

Since March 2021 at **GANIL**

Complete chamber from Strasbourg.

Tunnels from IJClab.

DSSD and electronics from CEA/Irfu.

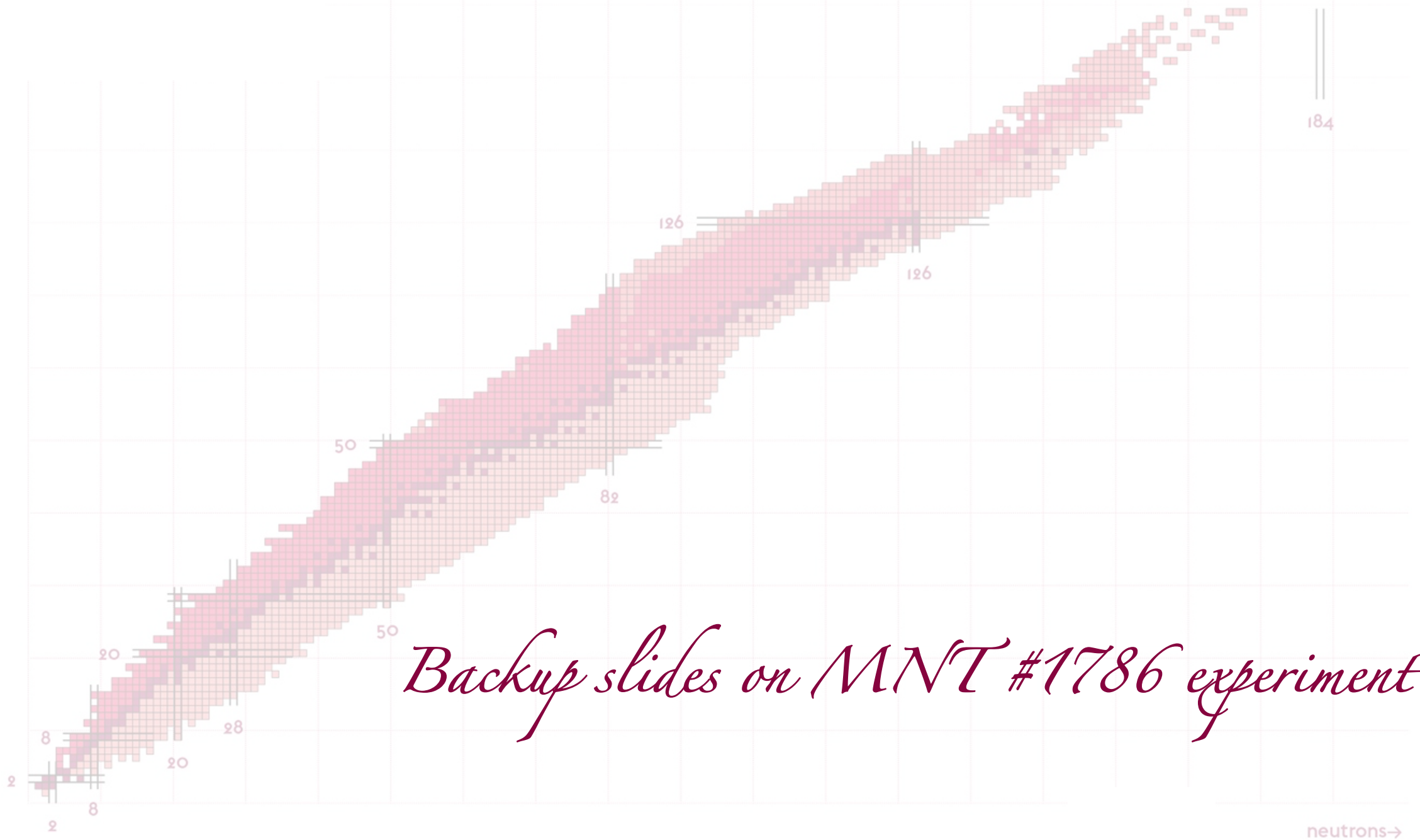
Offline commissioning (sources  $\alpha$ ,  $\beta^-$ )





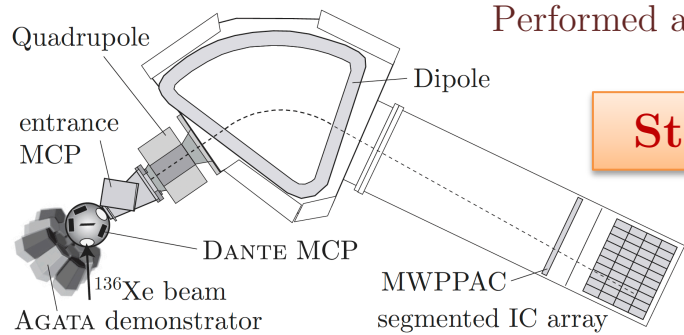
*Thank you  
for your kind attention*





*Backup slides on MNT #1786 experiment*

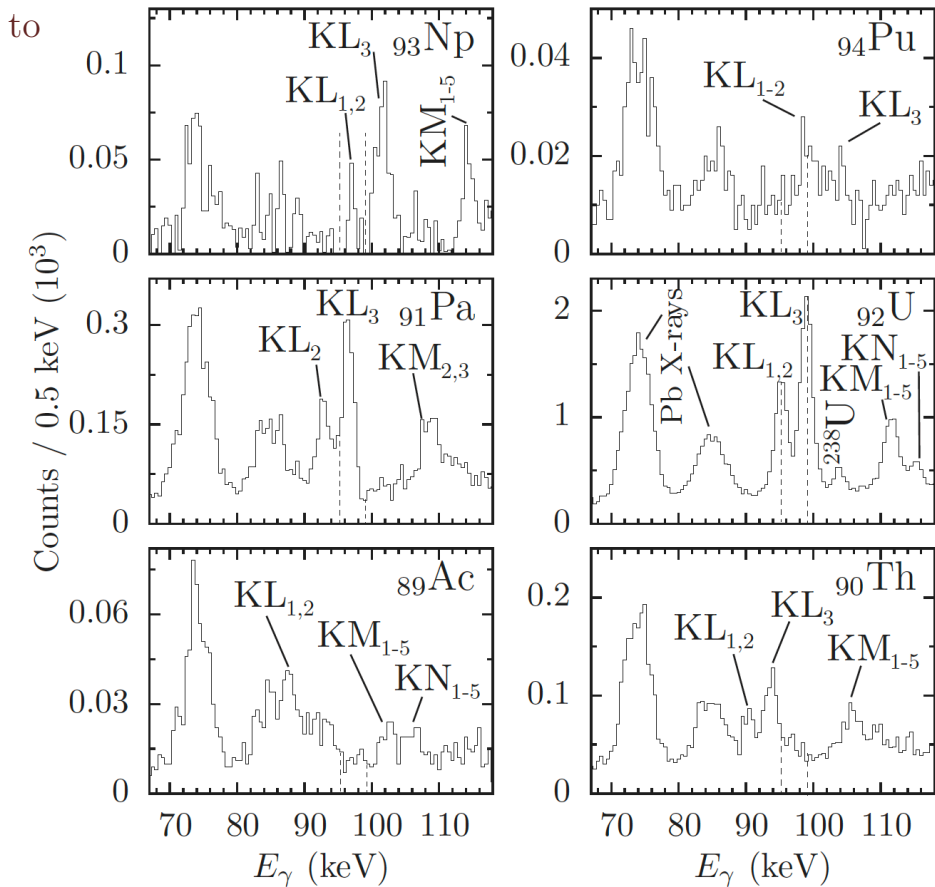
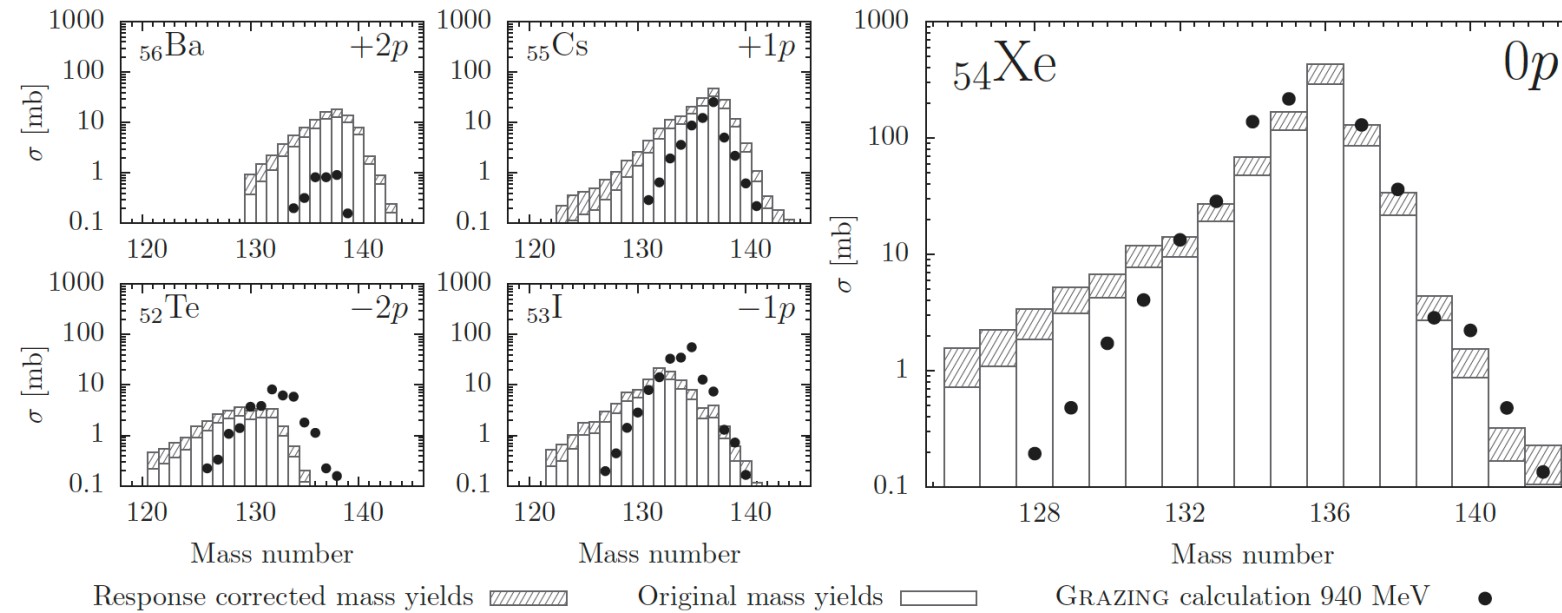
# MNT $^{136}\text{Xe}+^{238}\text{U}$ at 1 GeV and $\theta_{\text{lab}} \sim 50^\circ$



Performed at INFN LNL

Study of the beam-like (light fragments) and target-like (heavy fragments)

Cross sections of GRAZING calculations and experimental yields (histogram bars) normalized to the calculated cross section of the +1 channel of  $^{137}\text{Xe}$ .

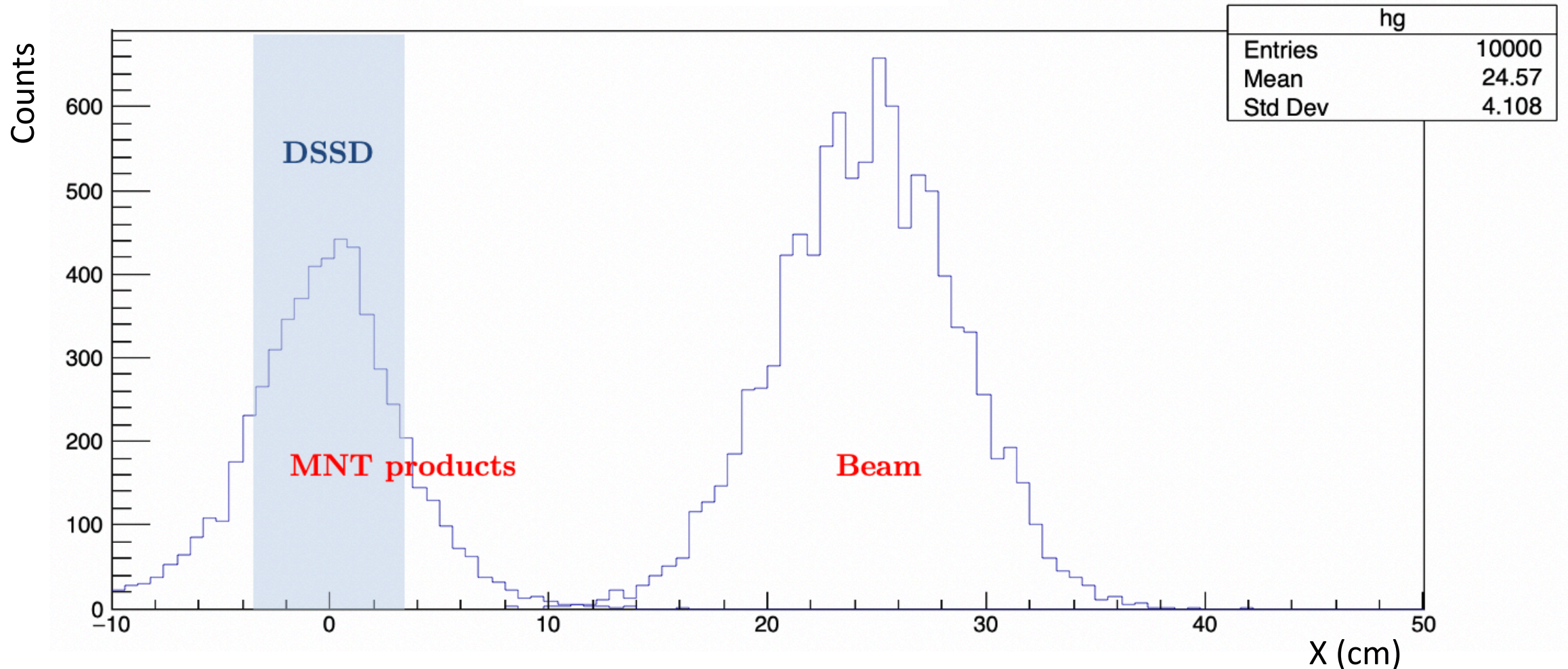


A. Vogt, et al. PRC 92, 024619 (2015)

# AGFA Monte-Carlo simulations

At 2 Torr

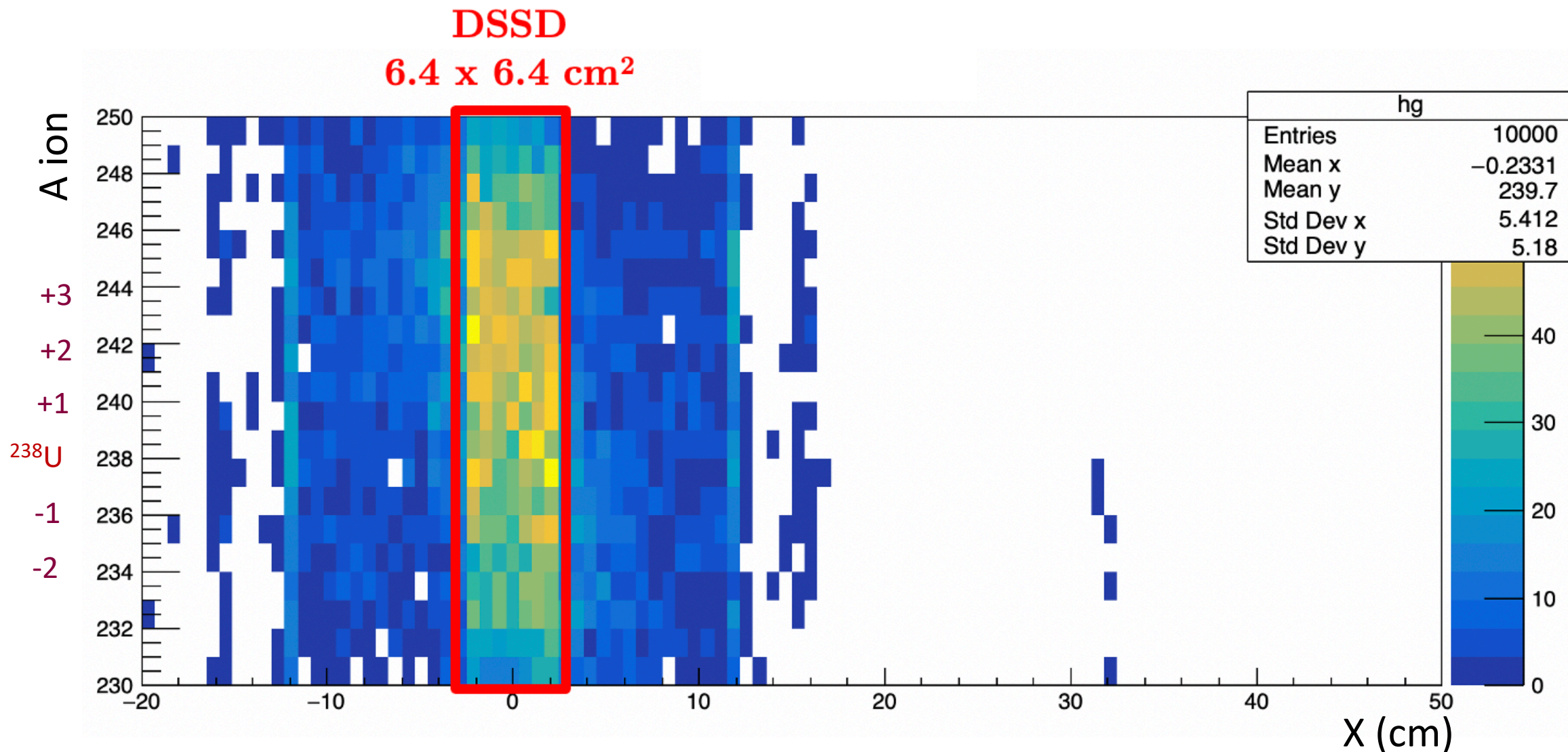
Particles reaching the focal plane





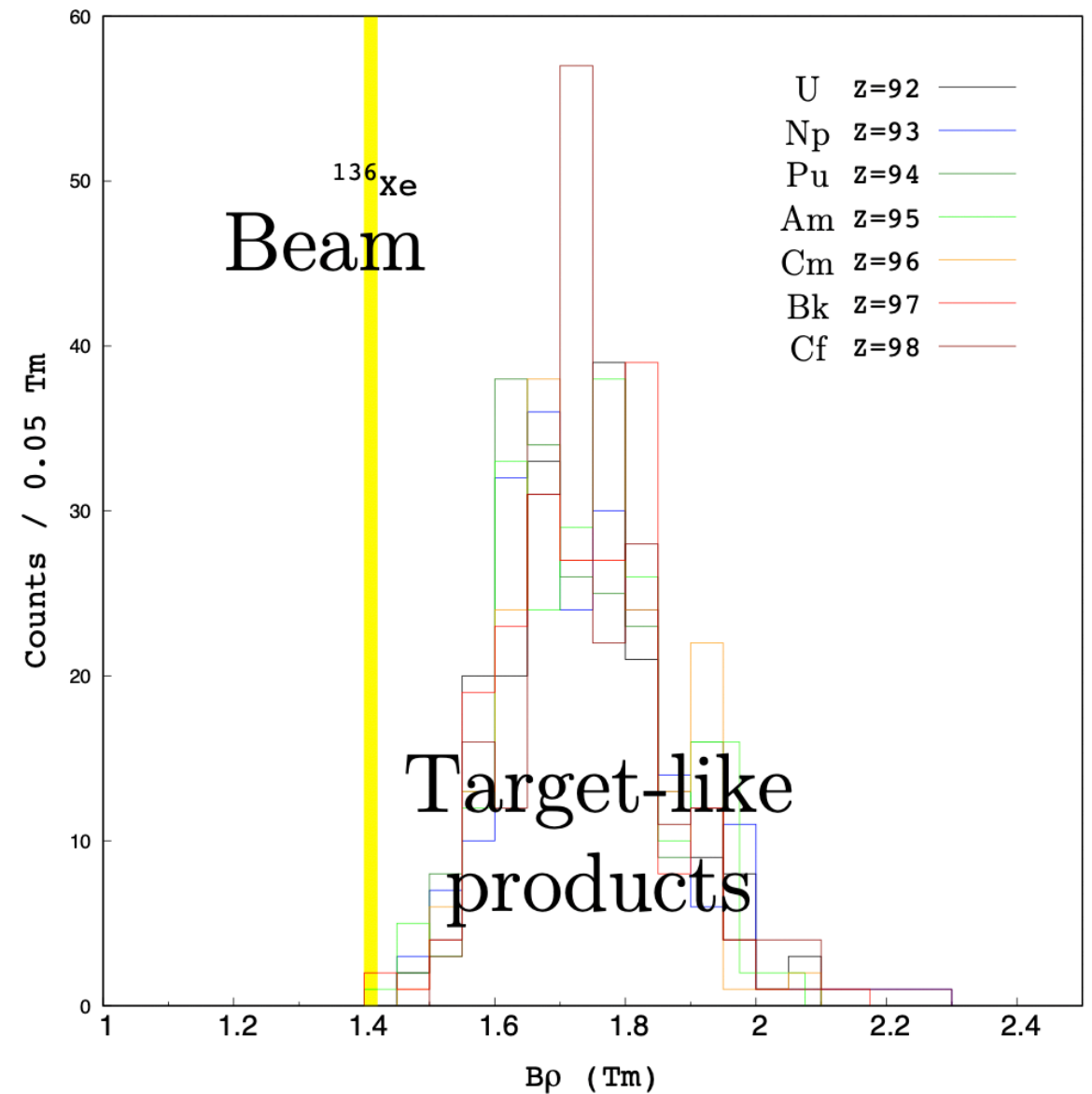
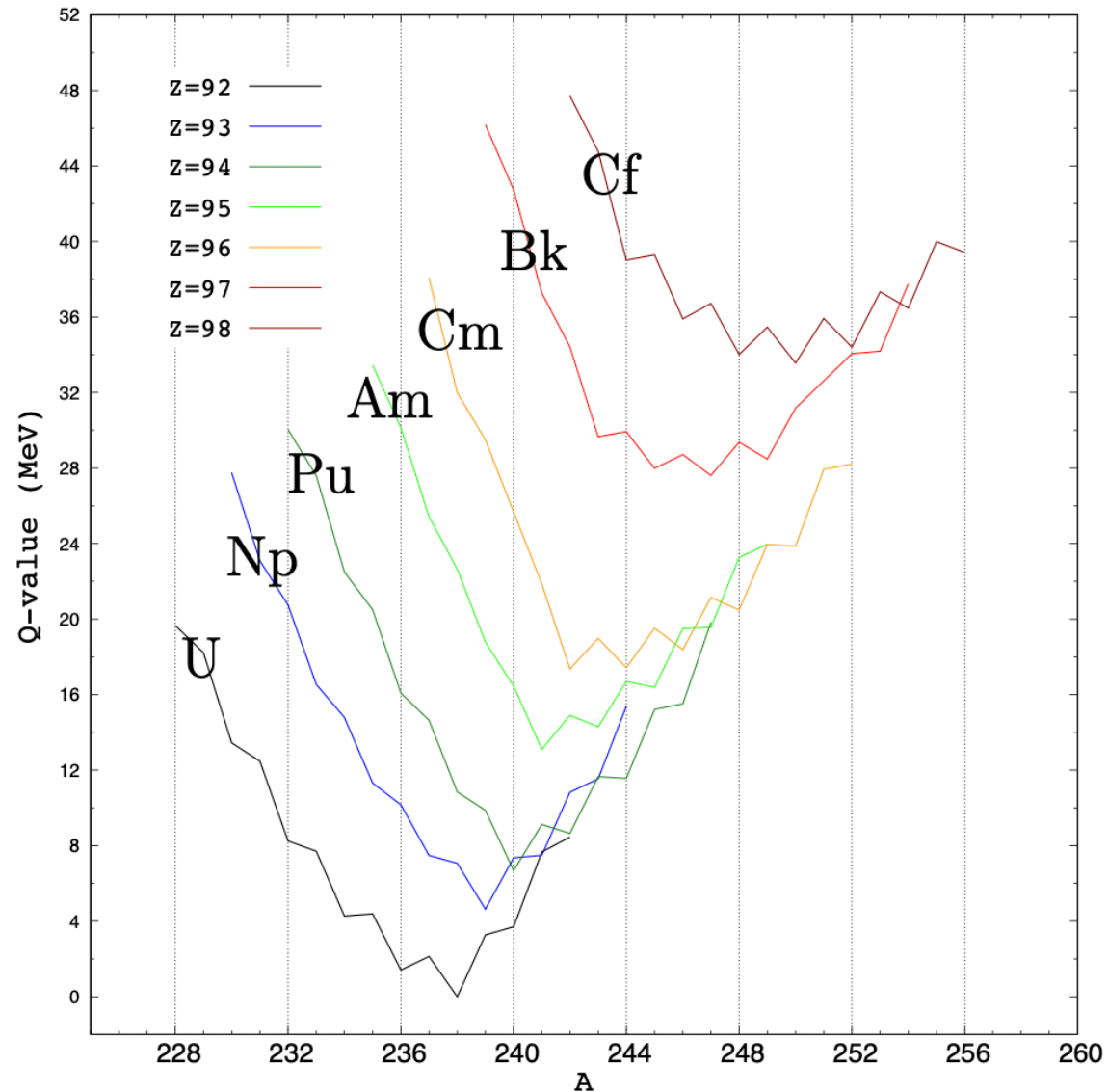
# AGFA Monte-Carlo simulations

At 2 Torr

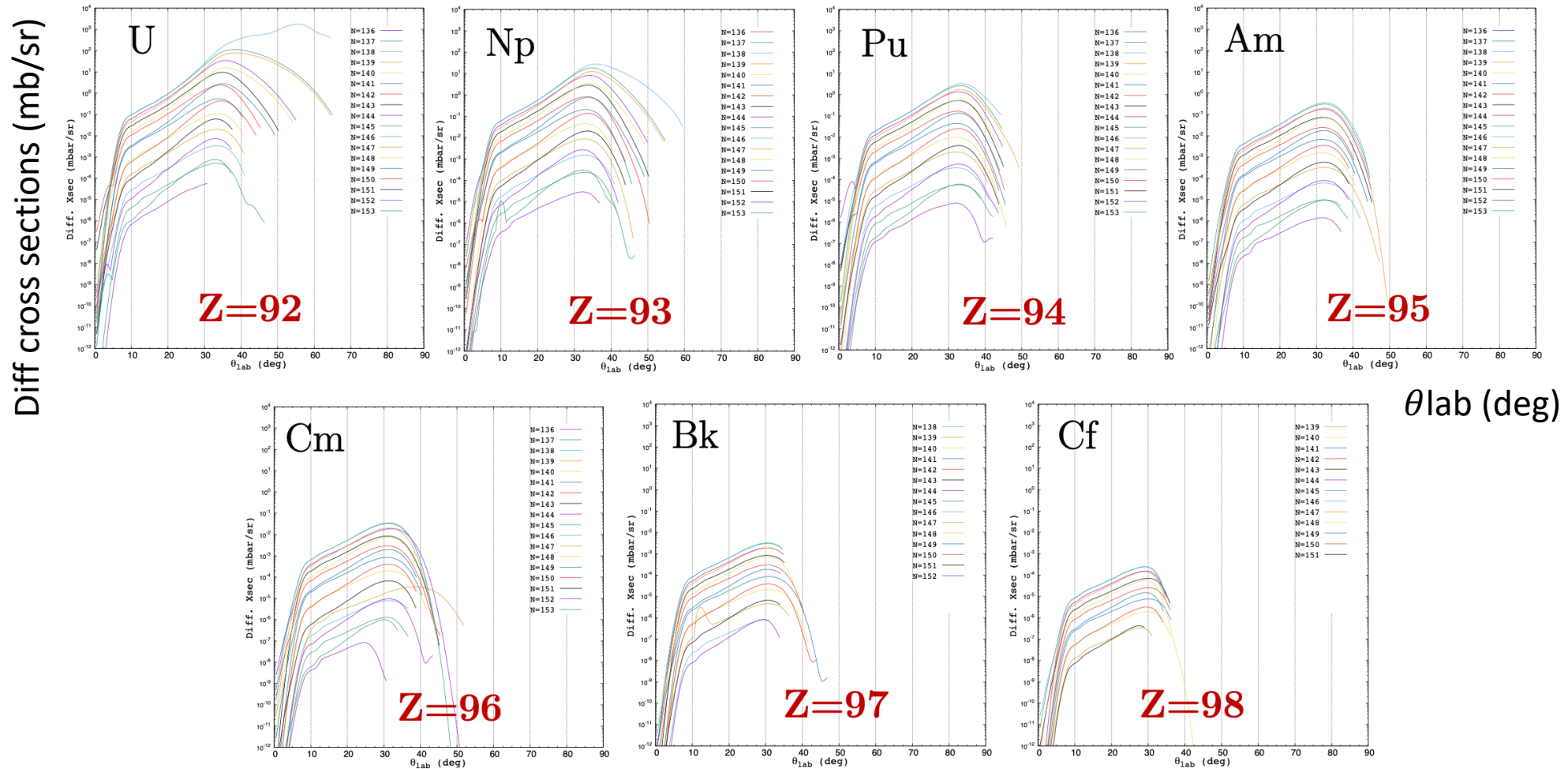




# From the particle generator



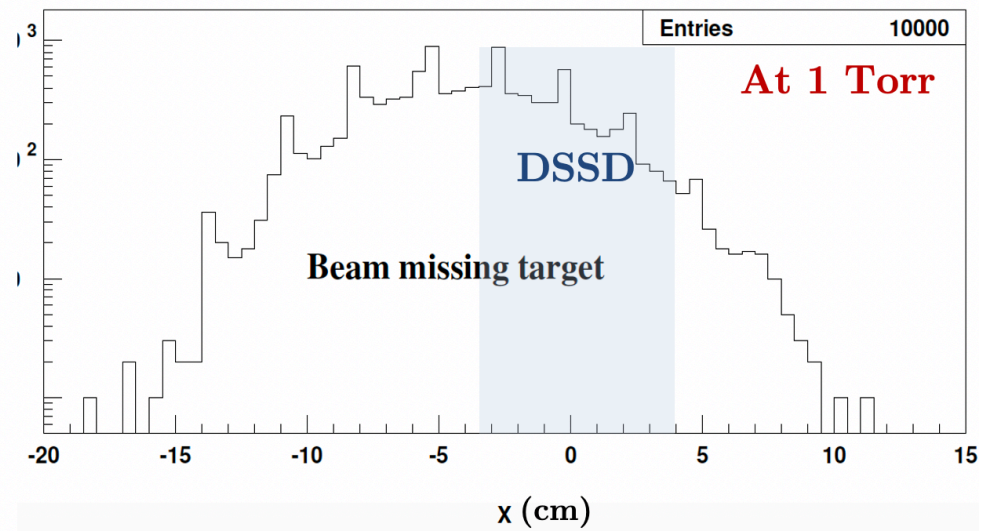
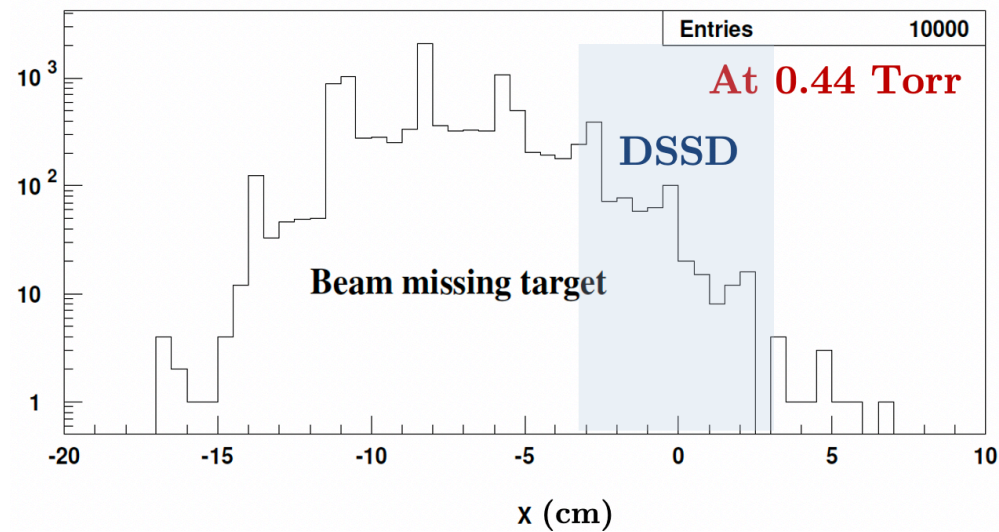
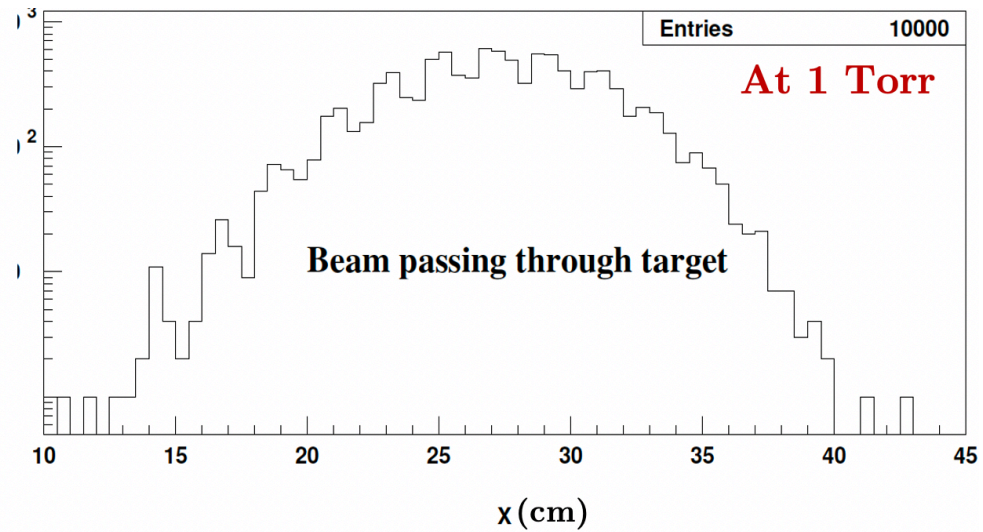
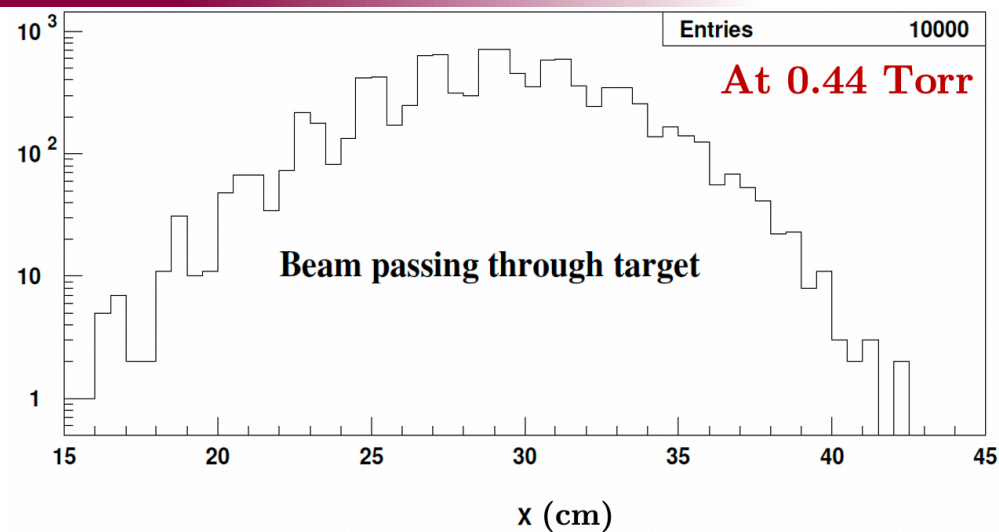
# GRAZING: angular distributions



Angular distribution  $\theta_{lab}$  (deg) of the differential cross-sections (mbar/sr) of the target-like nuclei generated by GRAZING.

At the grazing angle, GRAZING is giving the nice reproduction of the reactions but between 0 and 5° we are above the limits of the model.

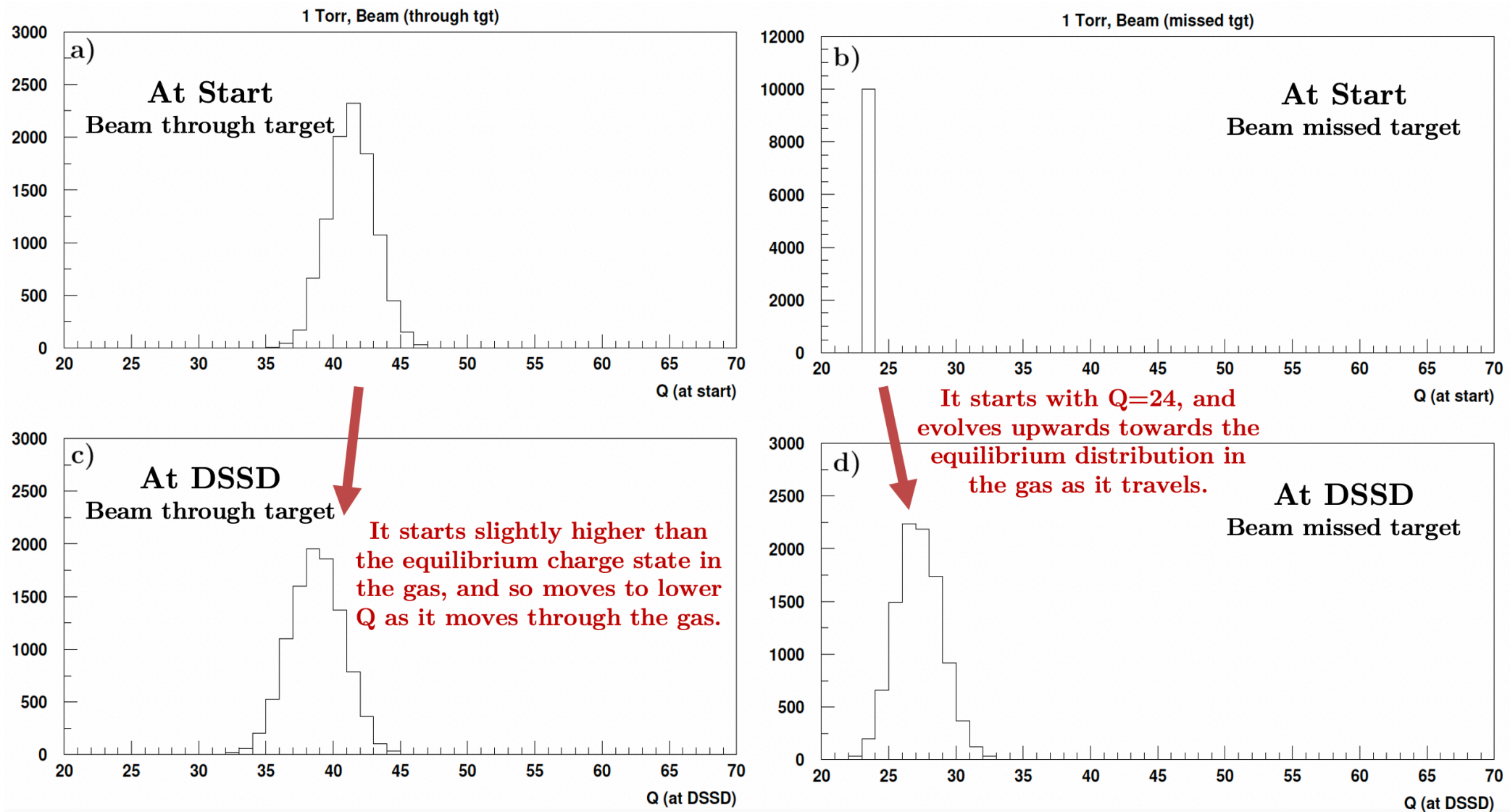
# Monte-Carlo AGFA simulations



**X (at the DSSD, in cm) of the  $^{136}\text{Xe}$  beam**

If the beam misses the target, there is significant amount that will hit the DSSD.

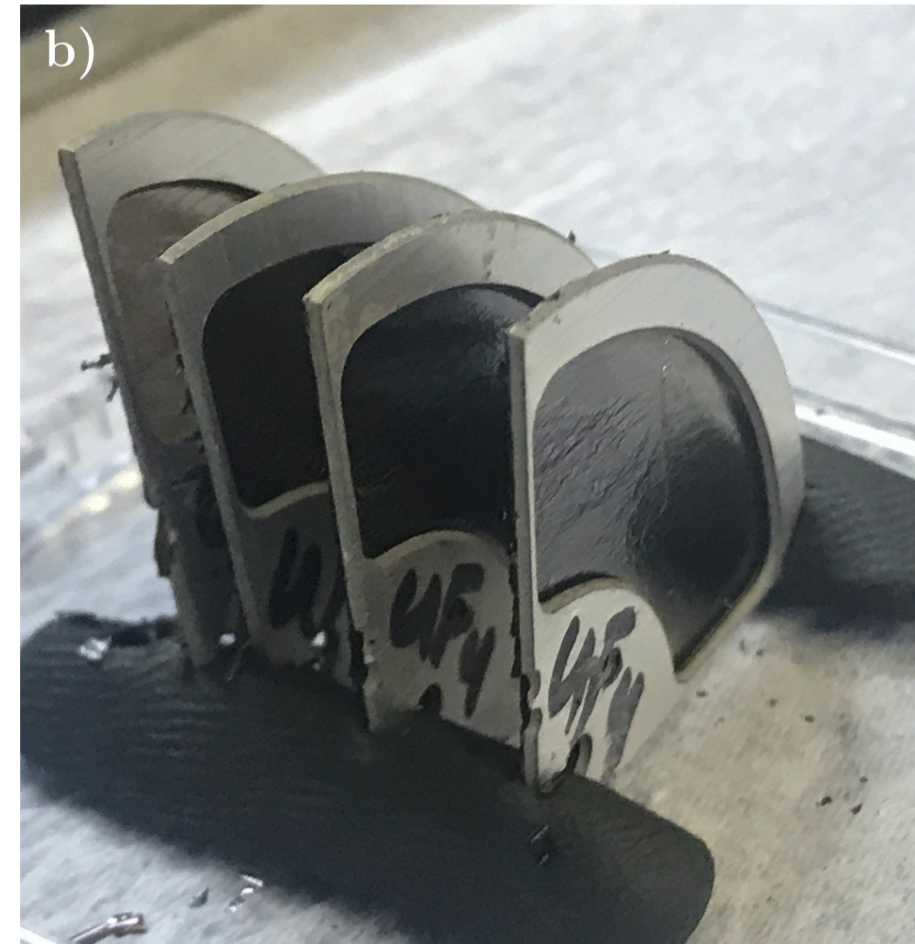
# Monte-Carlo AGFA simulations



Charge states distributions



# $^{238}\text{U}$ Targets





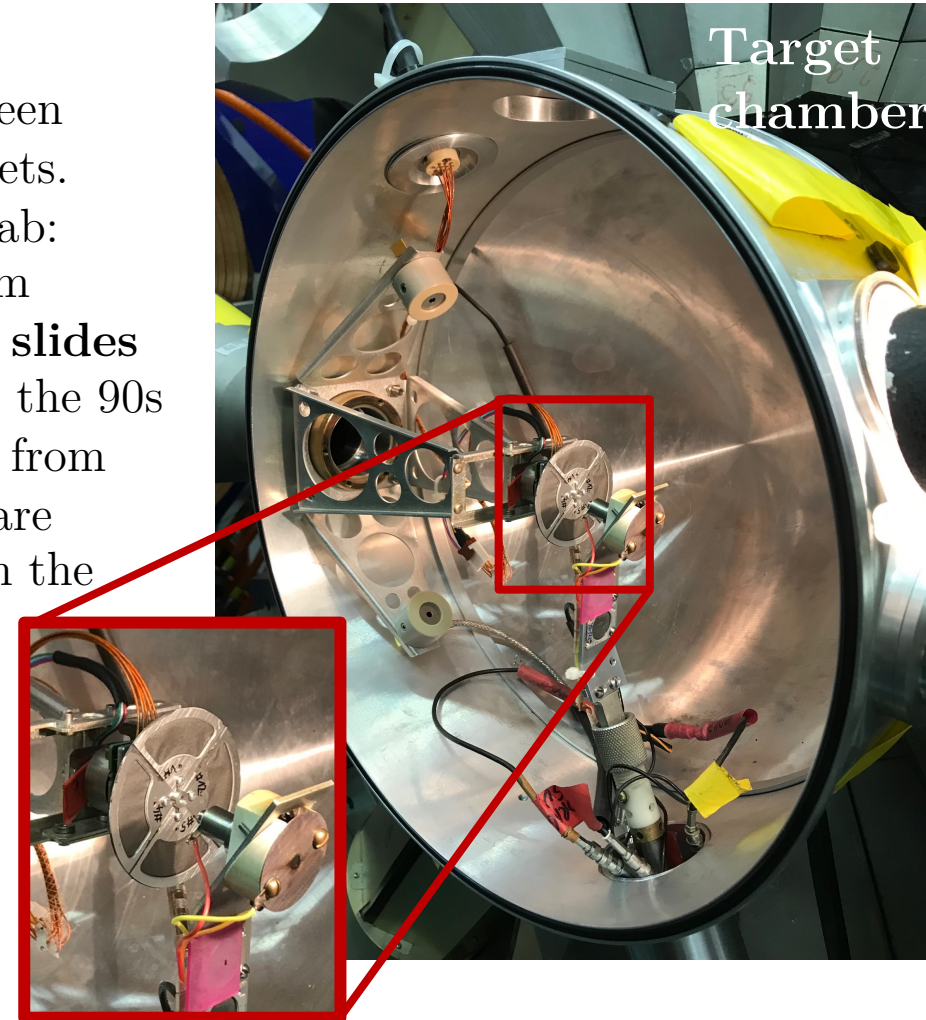
# $^{238}\text{U}$ Targets

## ➤ Up-to-now targets

- John P. Greene and his team have been striving on the preparation of the targets.
- Old  $\text{UF}_4$  thin targets are the target lab:  
 **$300\text{ }\mu\text{g}/\text{cm}^2$   $\text{UF}_4$**  prepared by vacuum evaporation onto  **$40\text{ }\mu\text{g}/\text{cm}^2$  carbon slides**
  - from the APEX<sup>1</sup> experiment in the 90s
  - they also prepared many slides from this era as well. All these slides are now exhausted and are floated in the Hot Lab.

## ➤ Wheel with slits

- Small target wheel quadrants



<sup>1</sup>J.P. Greene, et al. NIM in Physics Research A 362 (1995) 81-89

<sup>1</sup>G.E. Thomas, et al. NIM in Physics Research A 362 (1995) 201-204

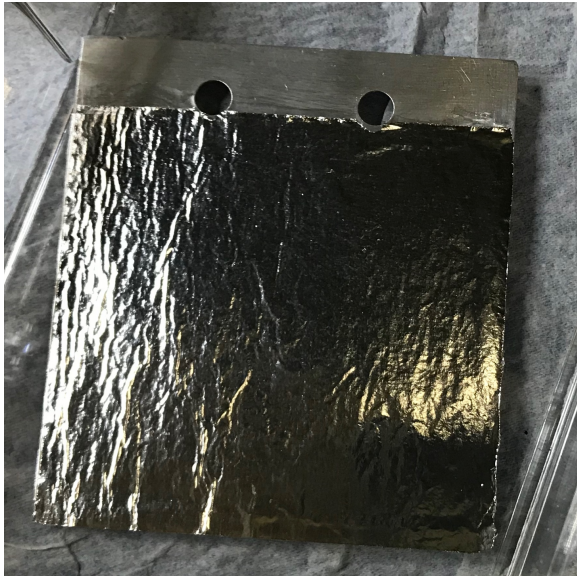
# Windows and degraders

## ➤ Windows

- John P. Greene made a Titanium Window  
Thickness is  $0.451 \text{ mg/cm}^2$

## ➤ Degraders for the DSSD

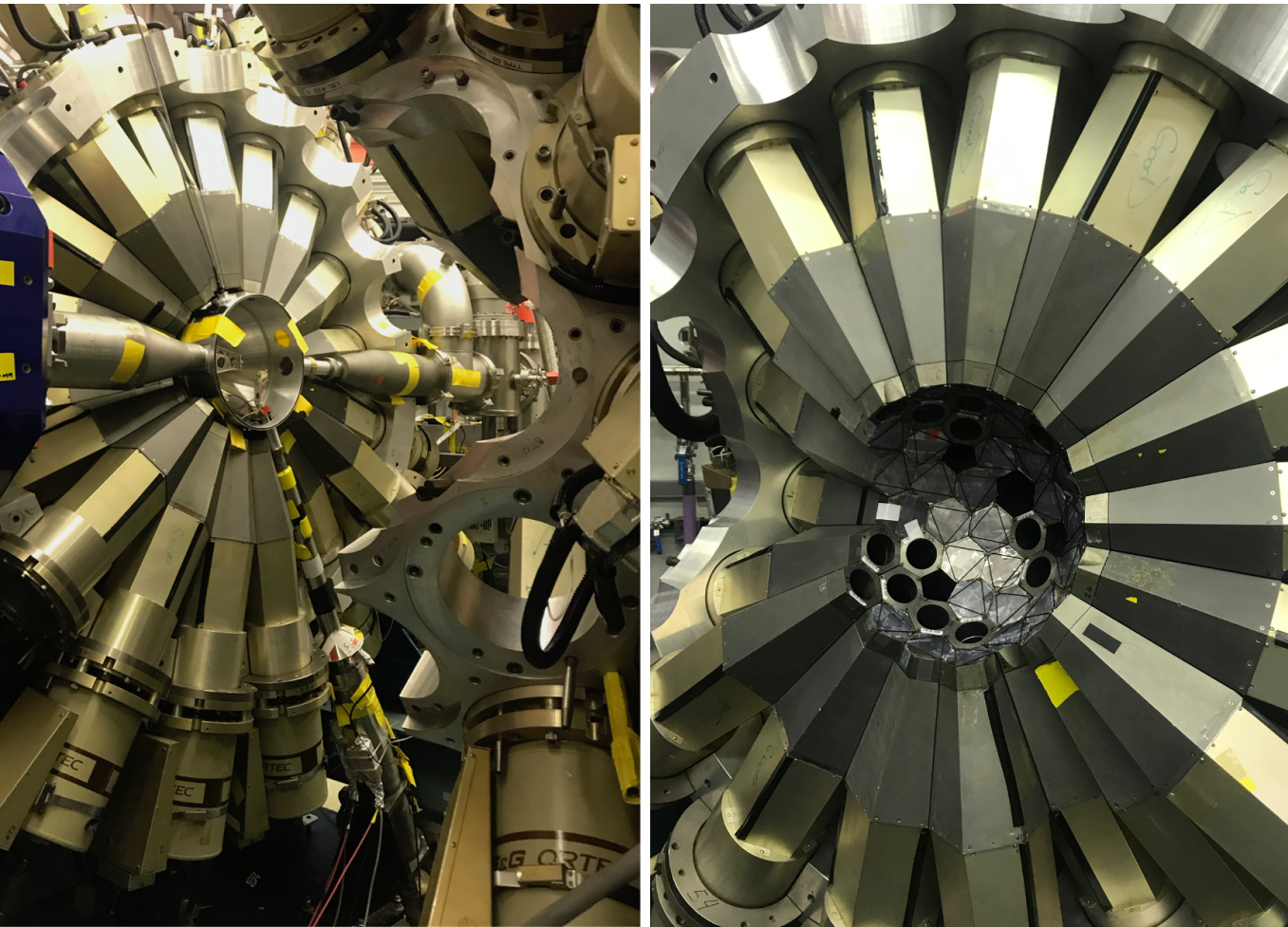
- Different thicknesses of Aluminum foils



At 2 Torr		
	Ti window	
E_beam (MeV)	Al degrader thickness (mg/cm2)	E_remain (MeV)
605	4,26	115
	5,42	60
	6,44	30
	7,36	5
809	4,26	130
	5,42	100
	6,44	65
	7,36	37
	all energies +/- 15 MeV	

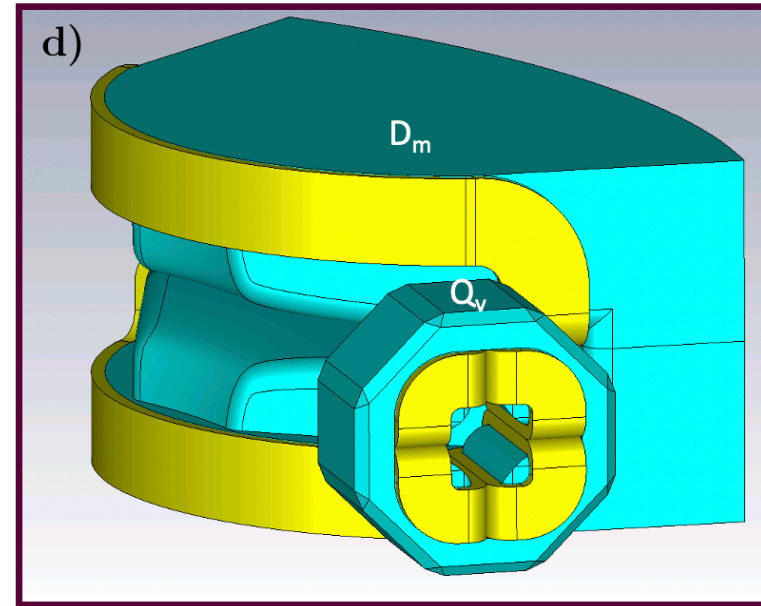
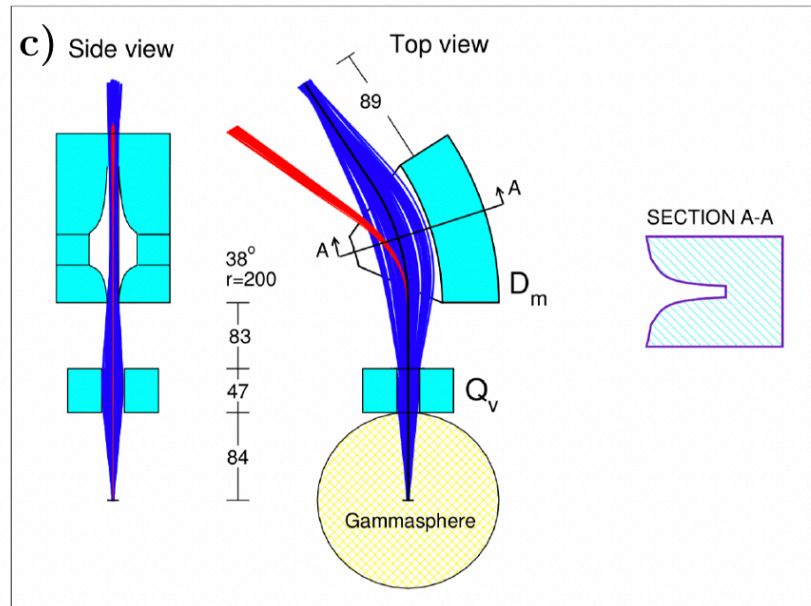
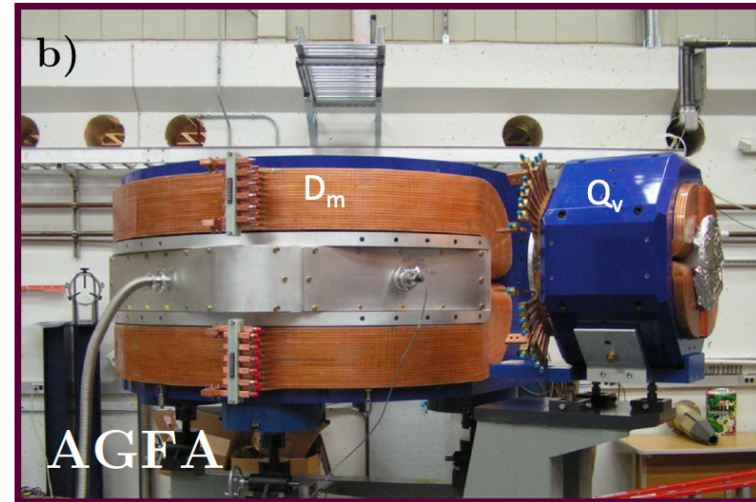


# GammaSphere





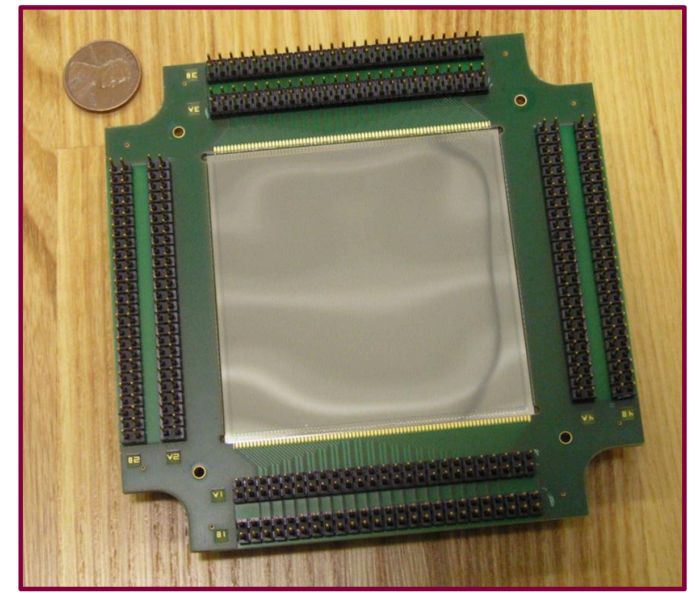
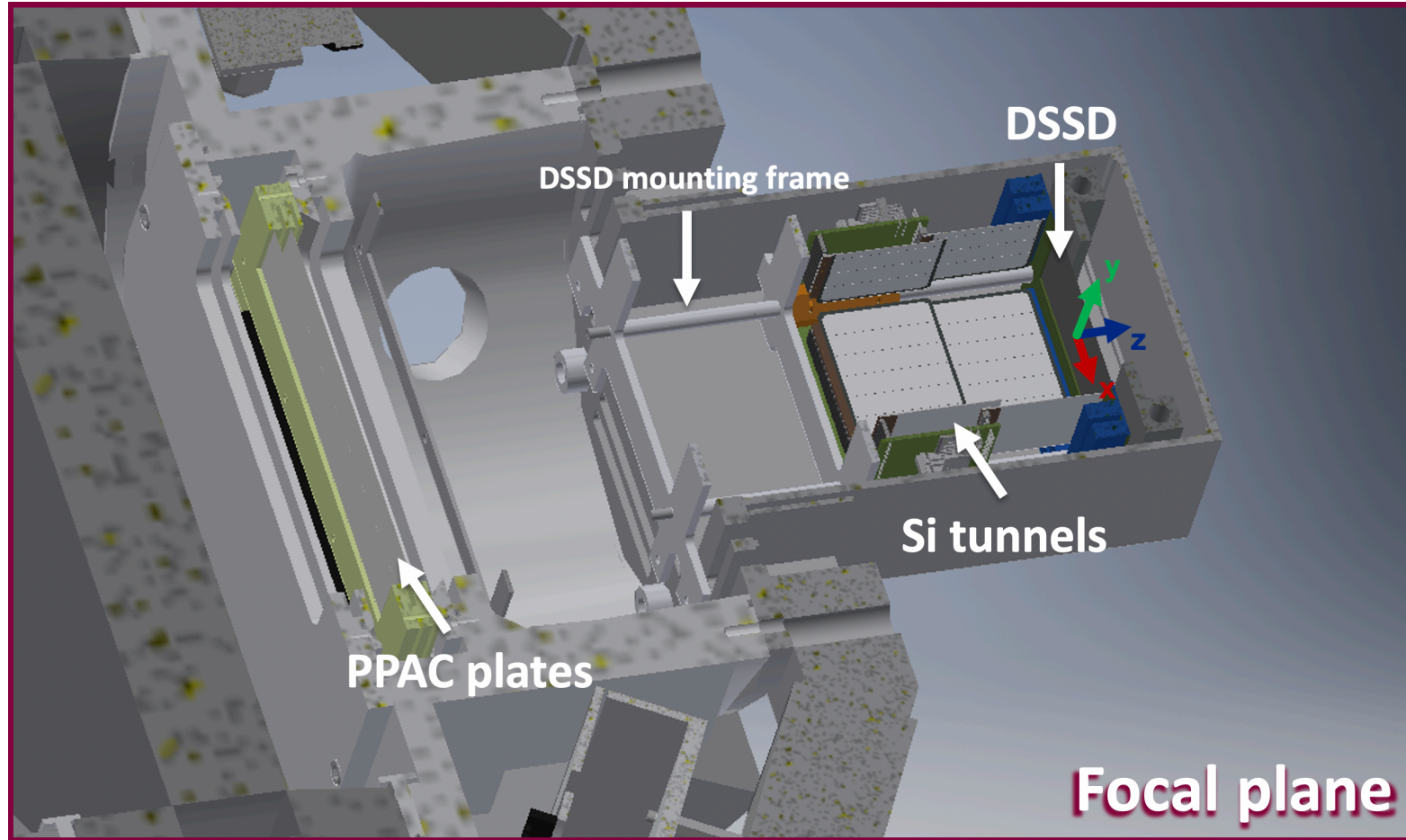
a) Parameter	Value	
Configuration	$Q_v D_m$	
Maximum bending power, $B\rho$	2.5 Tesla-m	
Maximum field at $Q_v$ pole tip	1.24 Tesla	
Maximum field at $D_m$ pole tip	1.7 Tesla	
Bend angle	38 degrees	
Target to $Q_v$ distance	40 cm	84 cm
Solid angle, $\Omega$	44 msr	22 msr
Target to focal plane distance	3.7 m	4.3 m



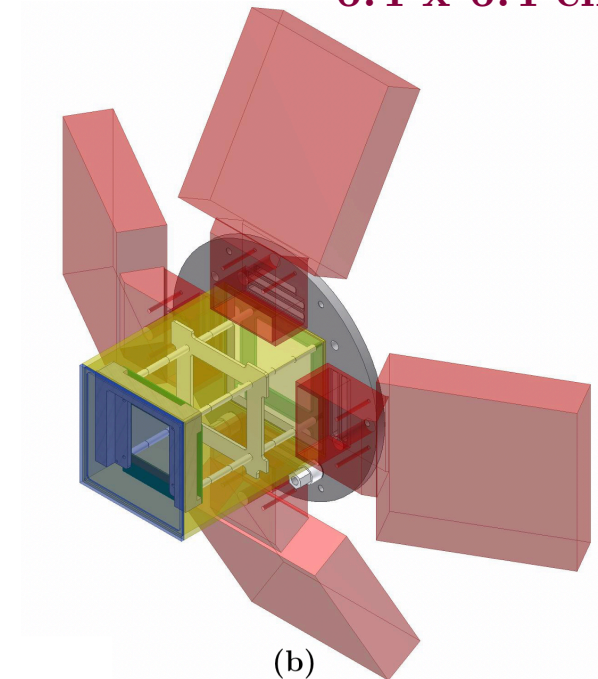
B.B. Black. EPJ Web of Conferences 163, 00003 (2017)



# DSSD, PPAC and Si tunnels



(a)  $6.4 \times 6.4 \text{ cm}^2$

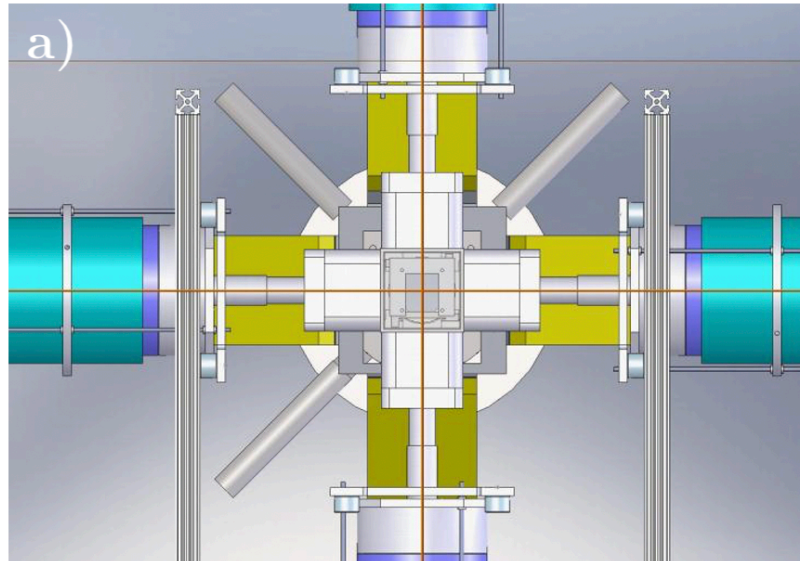


(b)

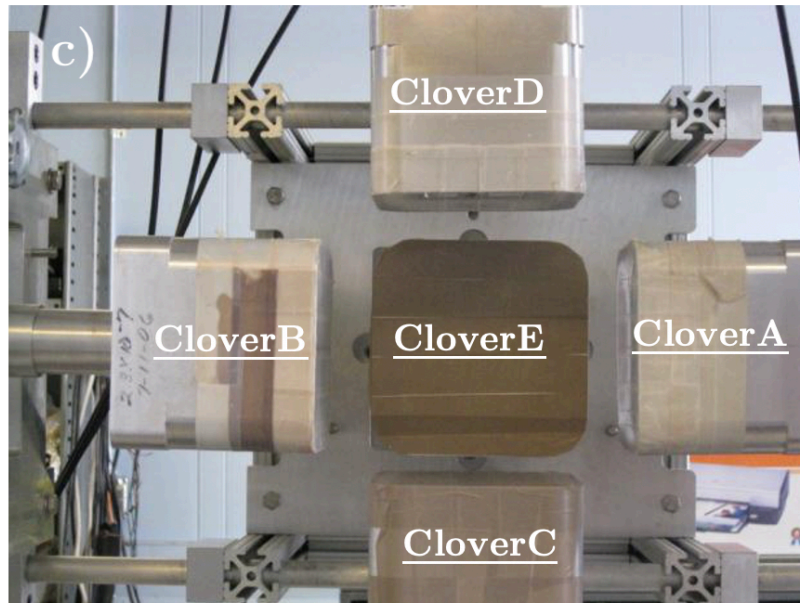


# XArray

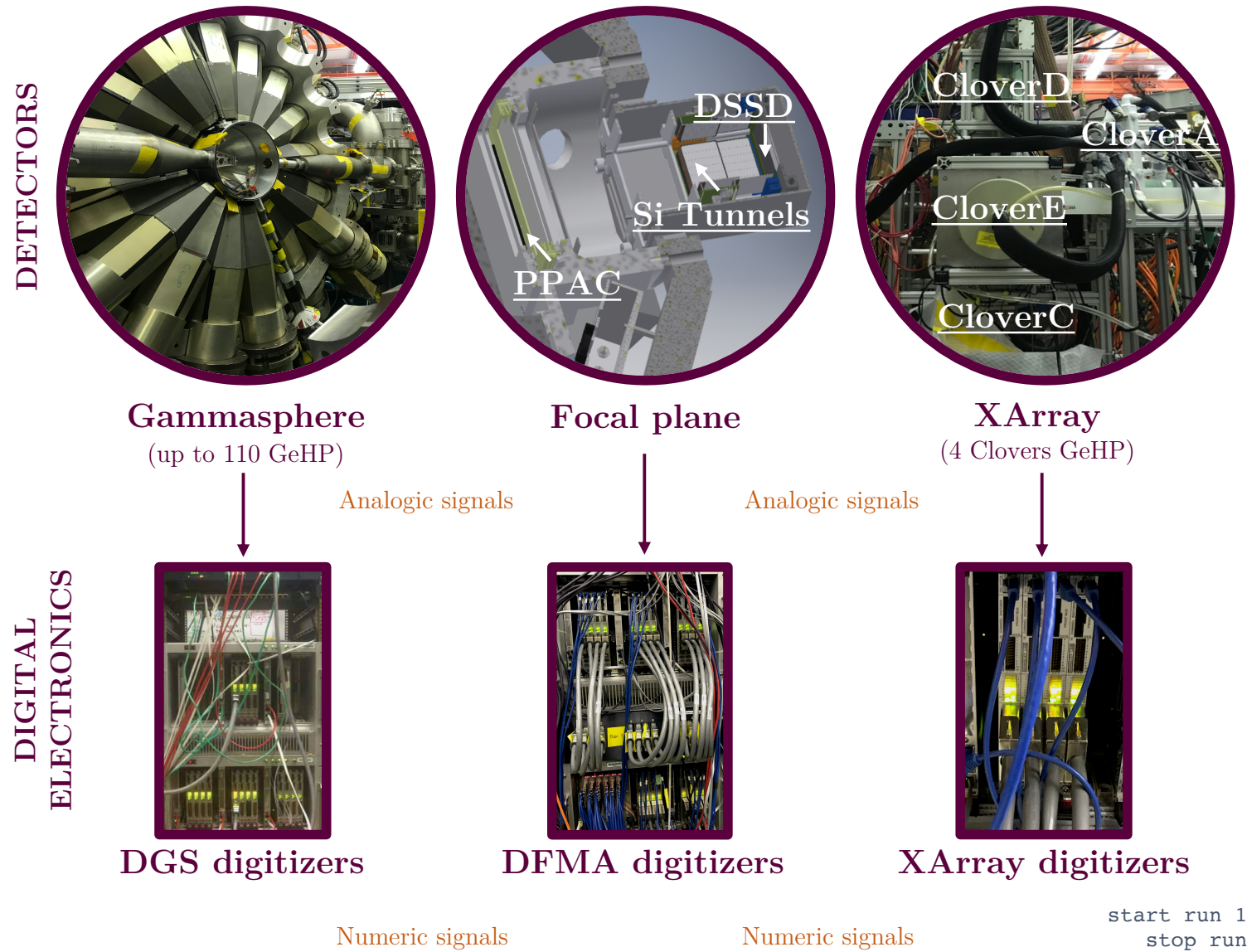
EXOGRAM  
Germanium  
clovers



Clover B  
was missing

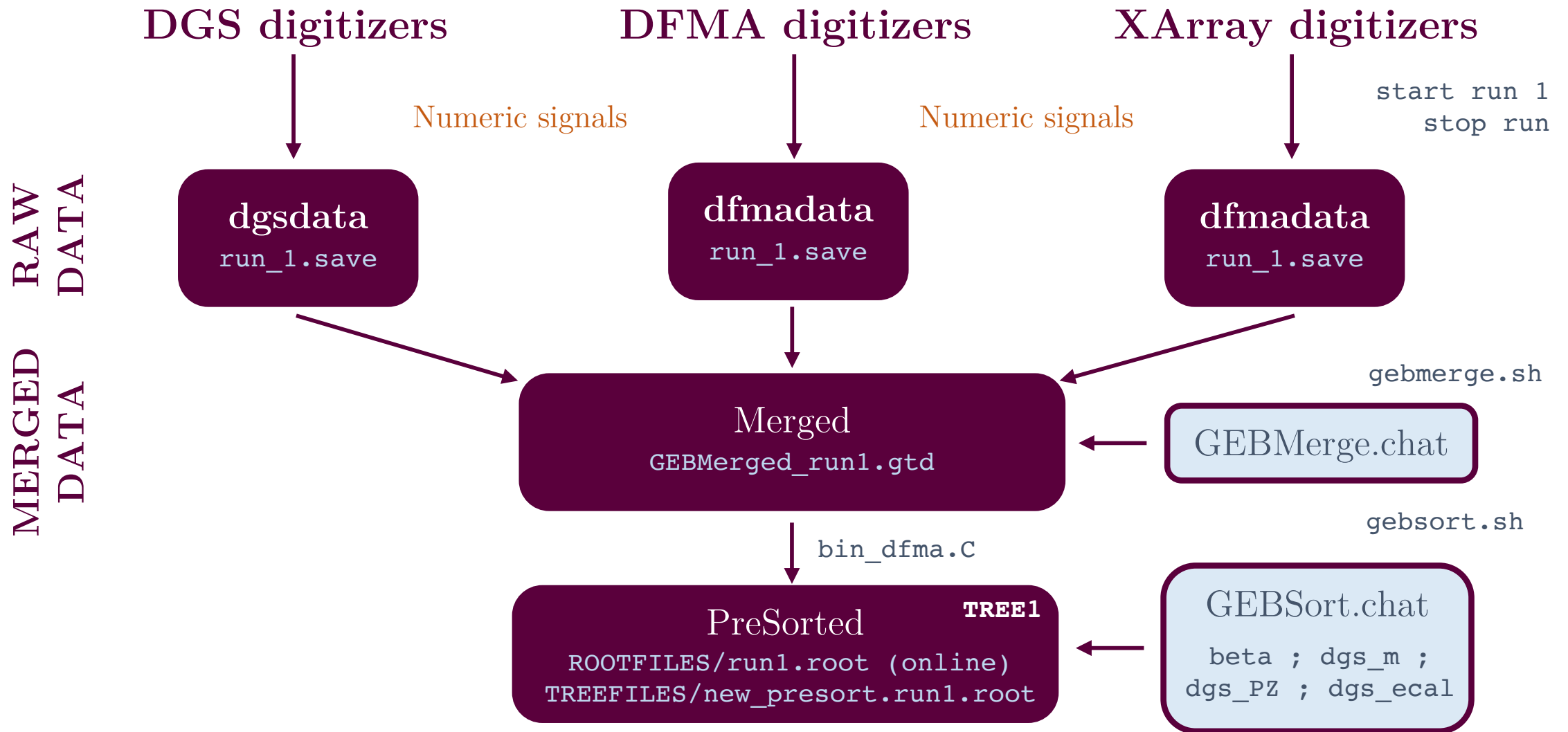


# Digital electronics

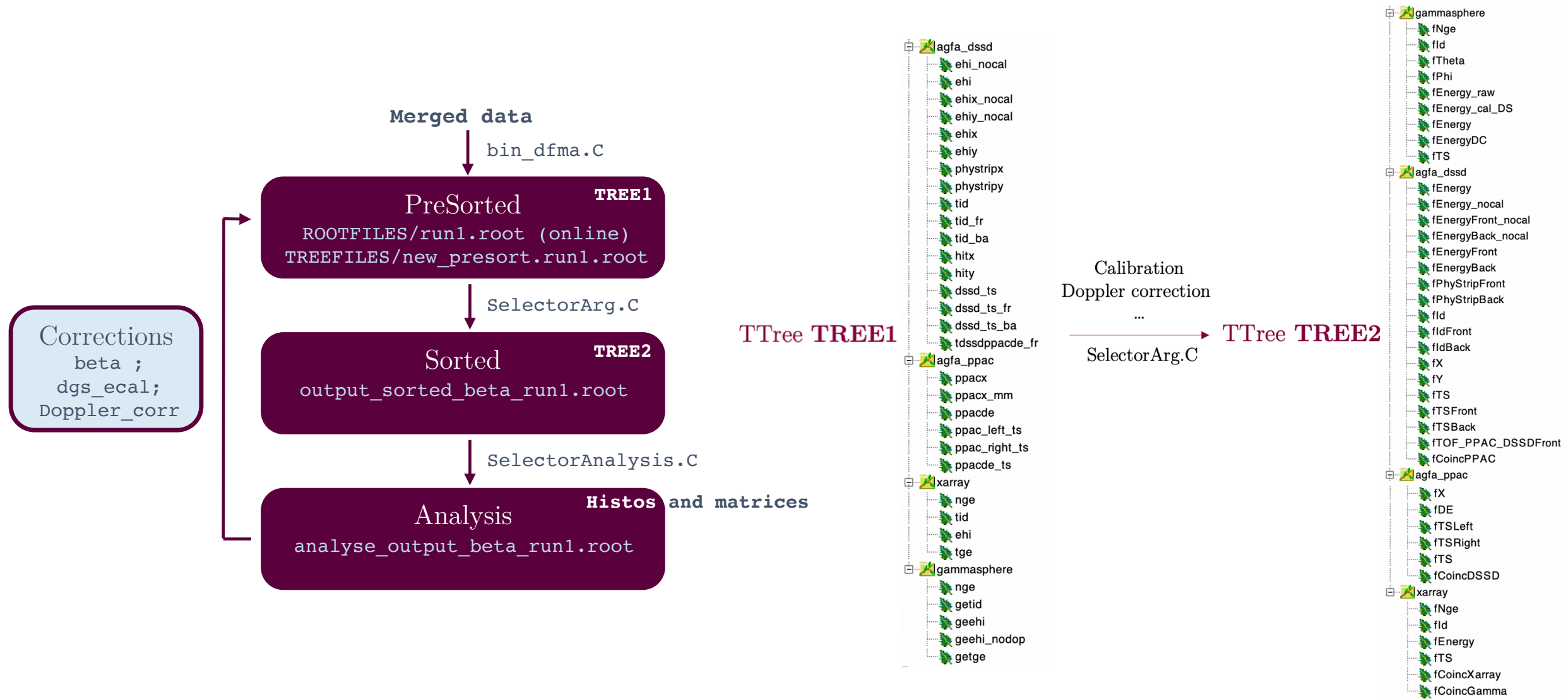




# Presorting scheme



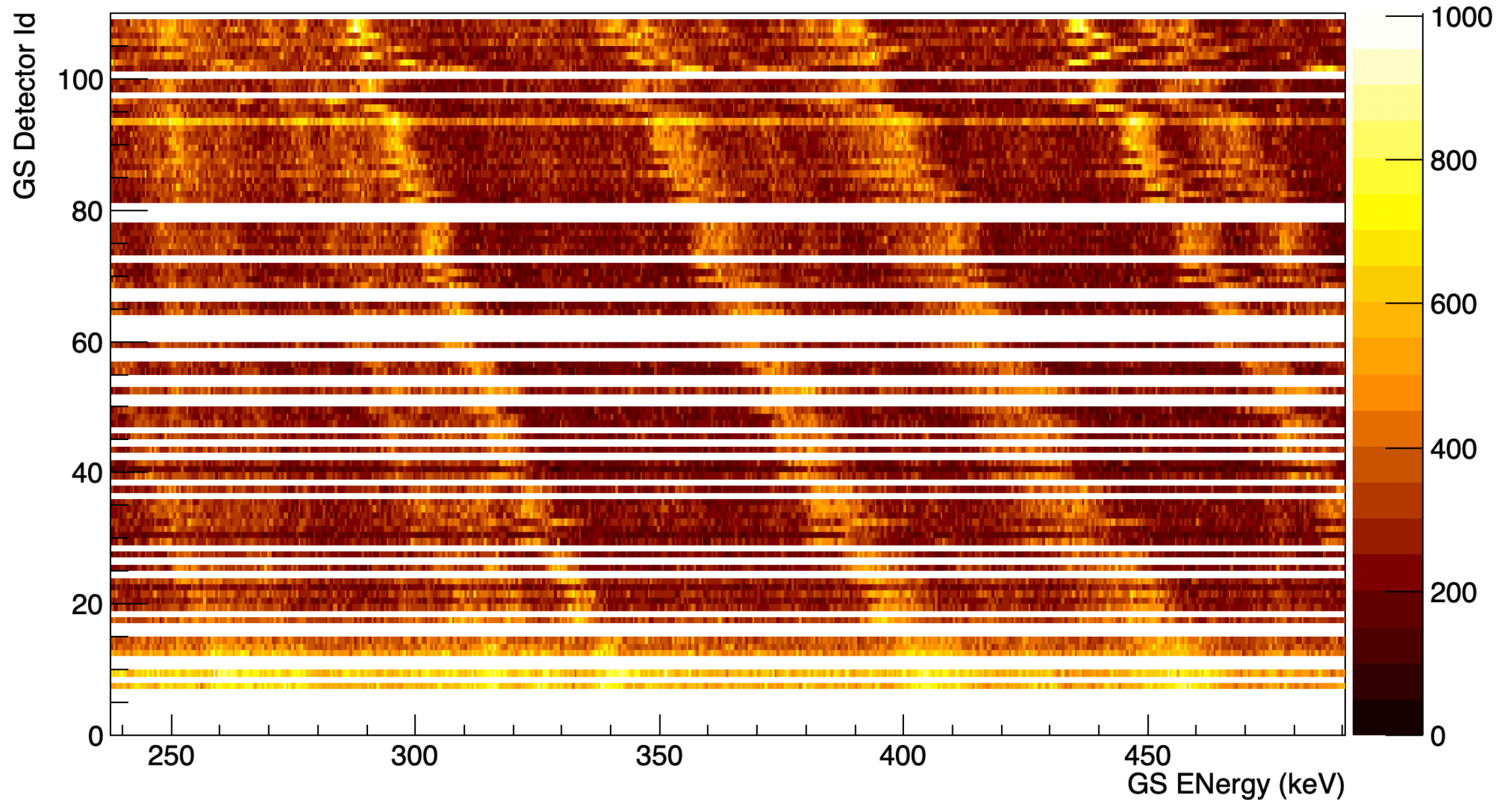
# Sorting and analysis scheme



# Calibrations

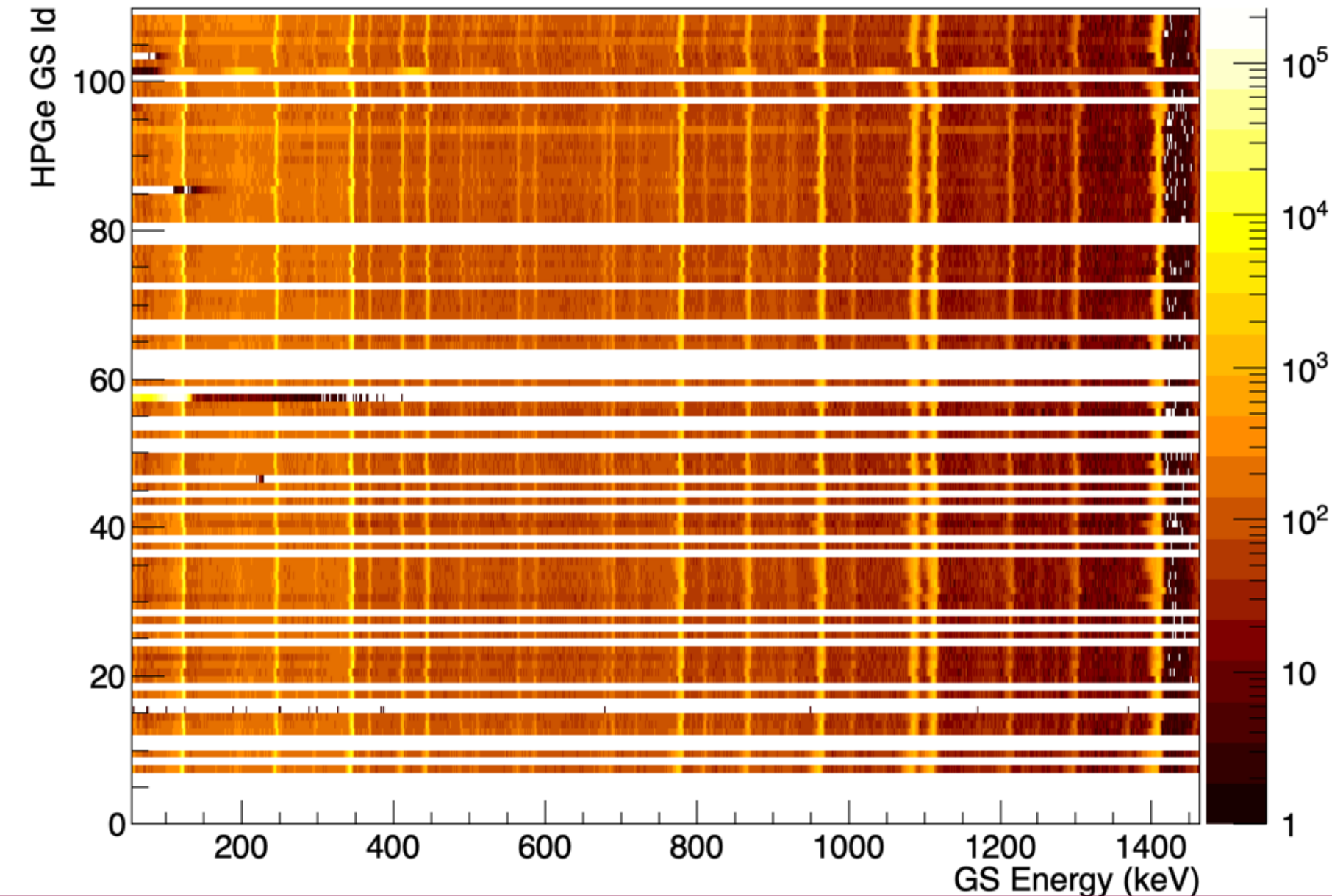
Detector	Source	Energy range [keV]	Activity [kBq]	Run number
Gammasphere	$^{182}\text{Ta}$	0-2000	85 (17/09/18)	127
	$^{152}\text{Eu}$	121-1410	?	126
	Mixed gamma	0-2000	115 (01/10/09)	125
Xarray	$^{182}\text{Ta}$	0-2000	85 (17/09/18)	123
	$^{152}\text{Eu}$	121-1410	?	124
	Mixed gamma	0-2000	115 (01/10/09)	129
	$^{243}\text{Am}$	5000-5400	148 (15/03/06)	128
DSSD	$^{239}\text{Pu}$ , $^{244}\text{Cm}$ , ( $^{241}\text{Am}$ )	5000-6000	?	120

# Without Doppler correction / Calibration





# Doppler correction (7.55% at $E_{\text{beam}}=705$ MeV)

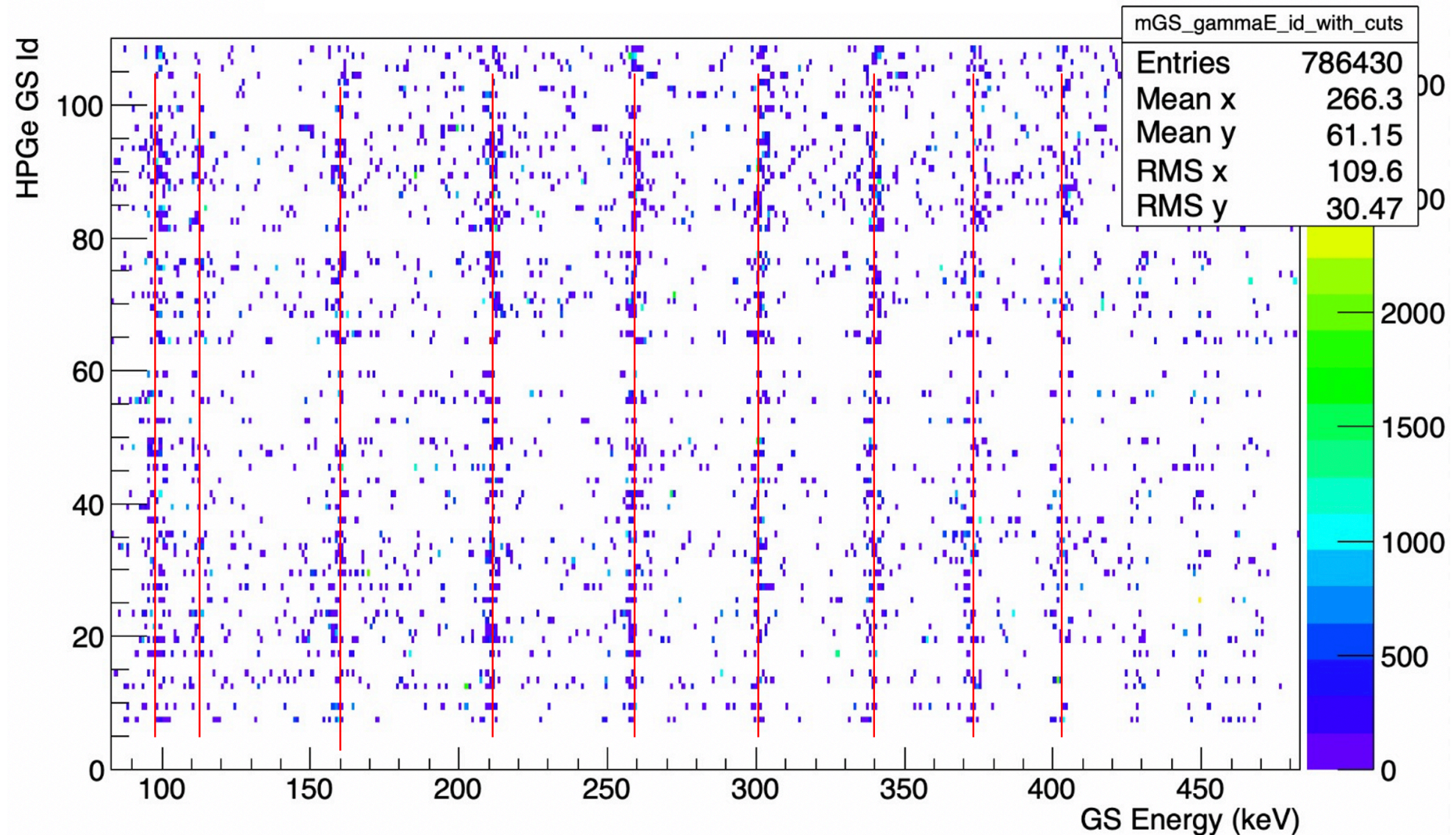


Doppler correction  
function:

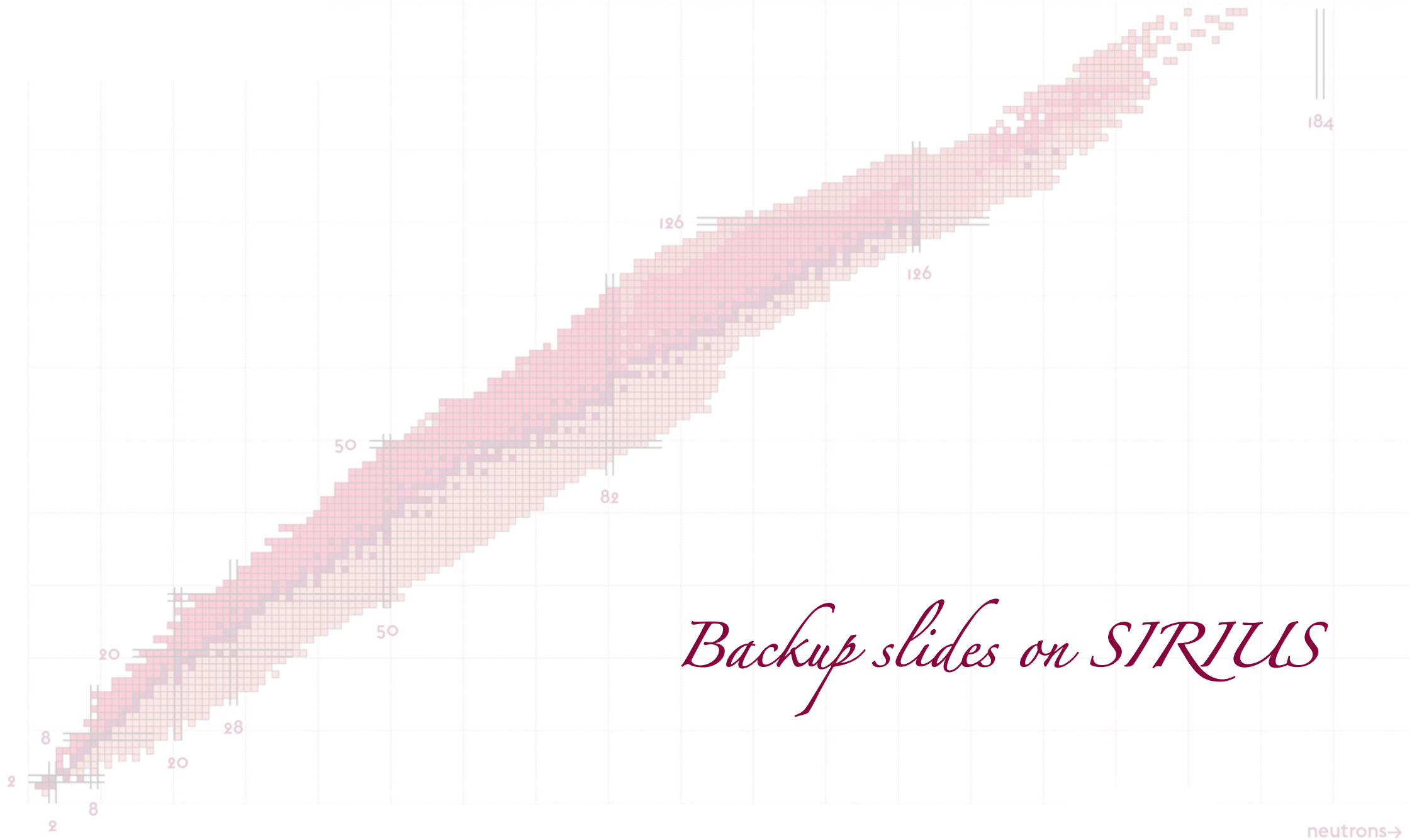
$$E_{\gamma} = E_0 \frac{\sqrt{1 - \beta^2}}{1 - \beta \cos \theta}$$

# Doppler correction (7.55% at $E_{\text{beam}}=705$ MeV)

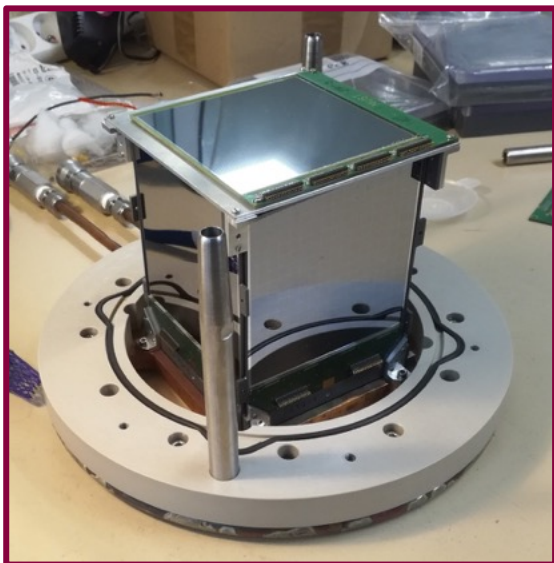
With cuts on the recoils







# DSSD test bench @CEA/Irfu



- Testing of DSSD, full electronics chain, with tri-alpha and electron sources
- Digital Signal Processing with Jordanov Filters (optimization of m and k)
- Characterization of the DSSD front-end electronics
- Write online and offline codes for data acquisition (NUMEXO2 cards)
- Different temperatures (-20, 0, +20°C)
- Test of low gain and high gain (0.5pF/1pF and 9pF)

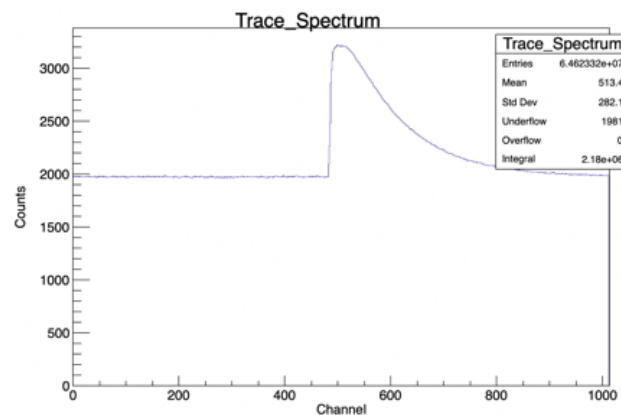
## DSSD

128 strips x  
128 strips y

## Conditions:

U=55V  
I=1.4μA  
Temp. 20°C

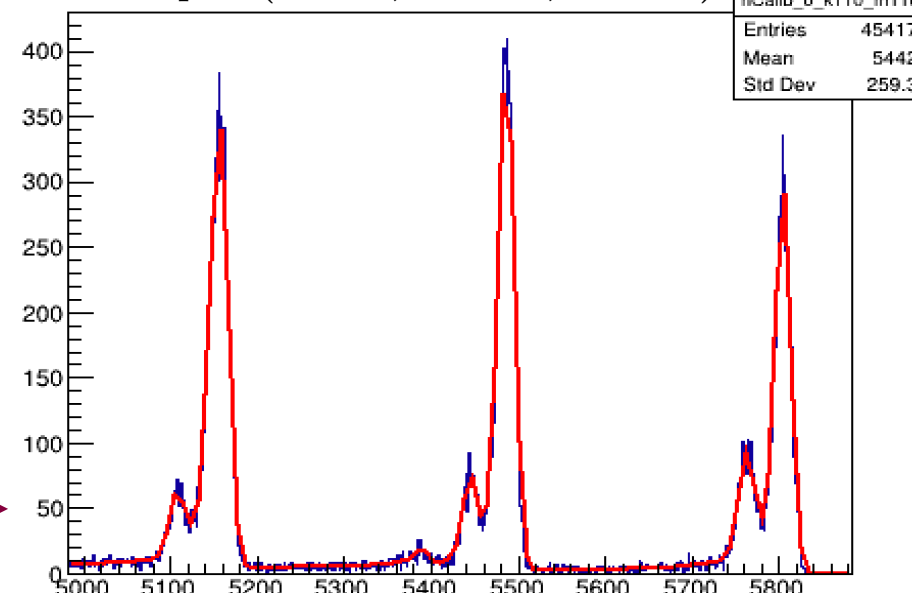
992 samples  
Sampling period 5ns



Trace

Jordanov  
Filter

## Tri alpha (239Pu, 241Am, 244Cm)

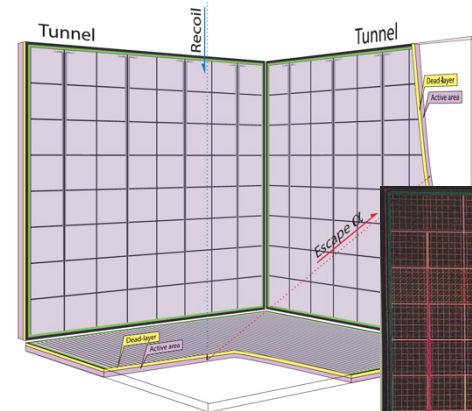
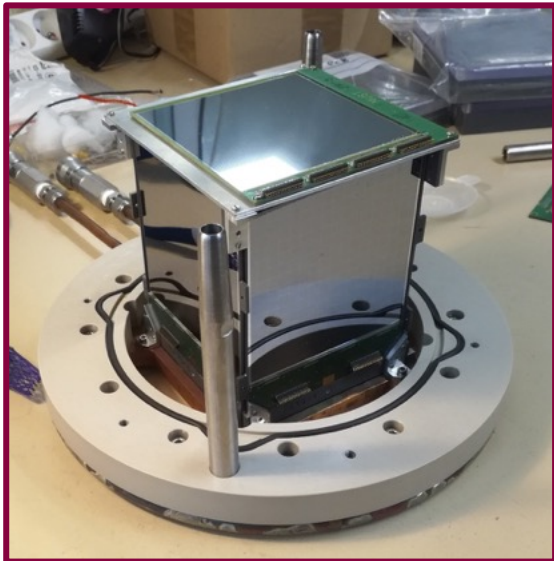


18 keV resolution



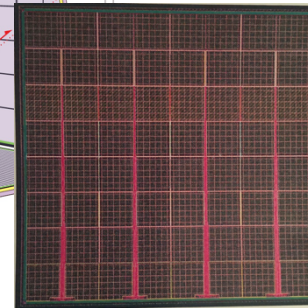
# Tunnels test bench @IPHC/IJClab

*Courtesy of O. Dorvaux*

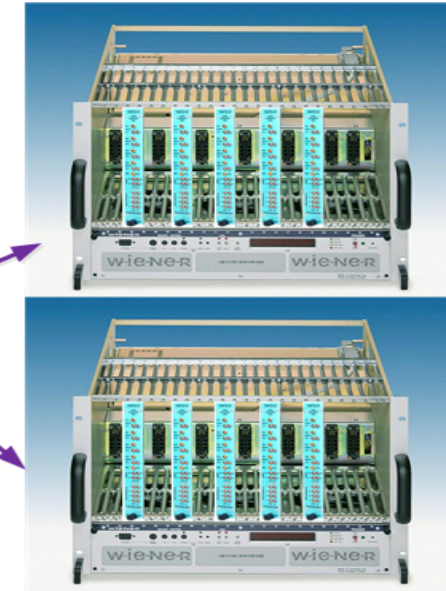


4 x 64 pixels  
Si Pad Detector

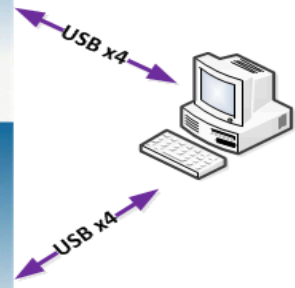
*@A Lopez-Martens*



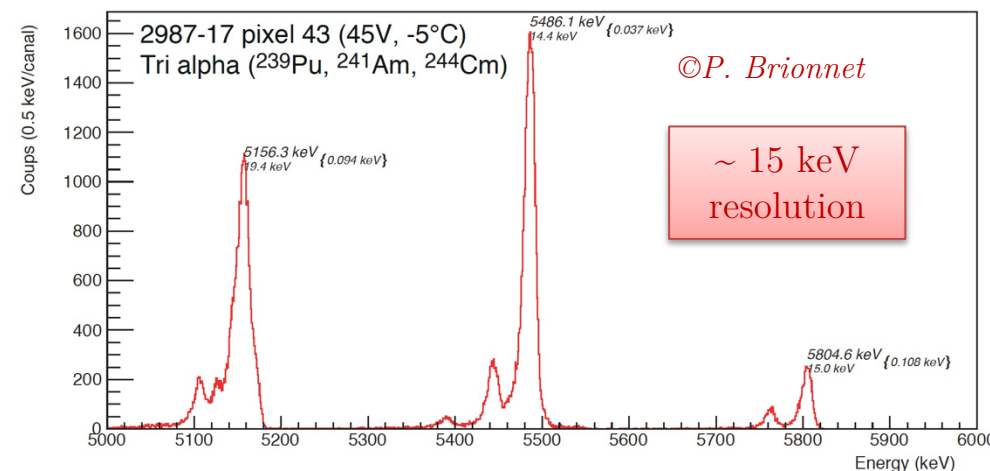
32 voies préampli  
CR110 custom



2 Crates with 4 TNT2D  
Total of 8 TNT2D cards = 32 channels



Still waiting for tests  
with the Numexo2 electronics,  
it will be finished at IJClab/GANIL



*©P. Brionnet*

With CREMAT electronics

# SIRIUS online commissioning @LISE2000?

## Objectives:

Testing the rejection of the line LISE2000 @GANIL with a small DSSD to study the feasibility of the SIRIUS commissioning at LISE2000.

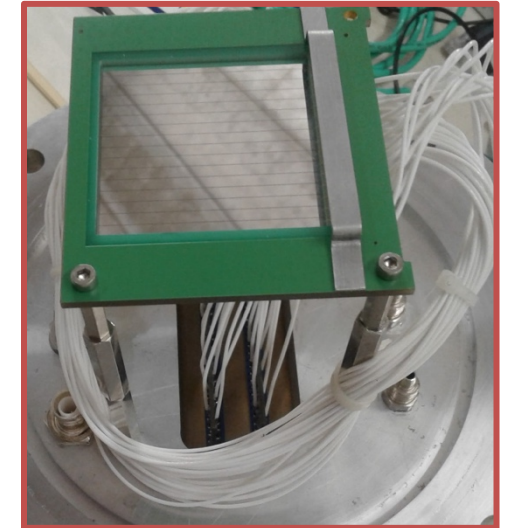
## Chosen reaction:

$^{40}\text{Ar} + ^{174}\text{Yb} \rightarrow ^{209,210}\text{Ra}$ . 4,62 MeV/u

Cross section: 1.4 mb

## Set-up:

- 1 DSSD 16x16 strips 300  $\mu\text{m}$   
(50x50 mm<sup>2</sup> active area with PAC 15mV/MeV)
- 2 amplifiers CAEN N568B  
(OUT1: GANIL ADC,  
FOUT1: 16 strips CFD CAEN N843)
- 1 MWPC (50x50 mm<sup>2</sup>) for TOF measurements

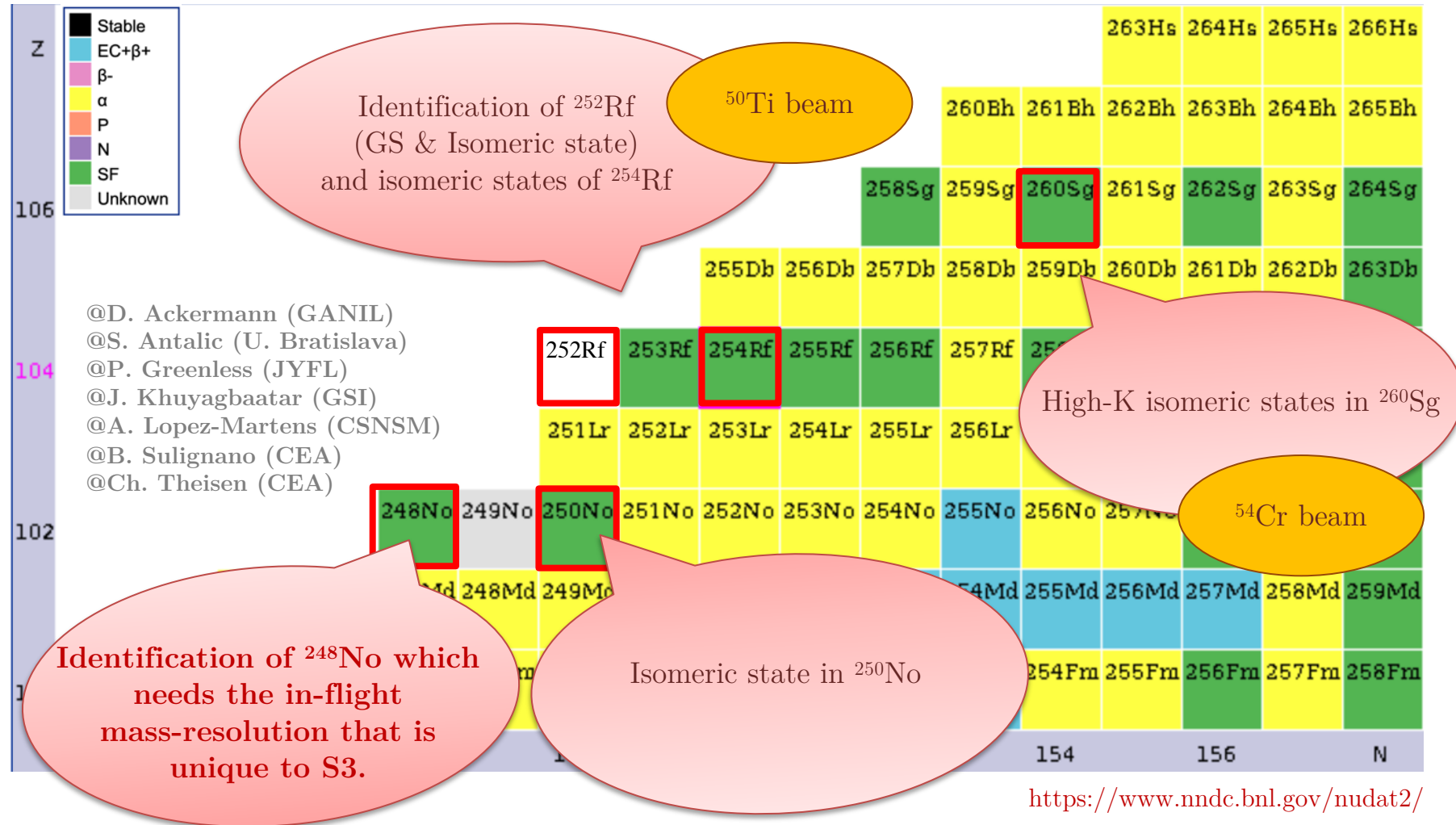


First test in June 2019  
→ New test in April 2021

Work has been done to improve the transmission of LISE2000!

# SIRIUS Day1 experimental plan (A/Q=3)

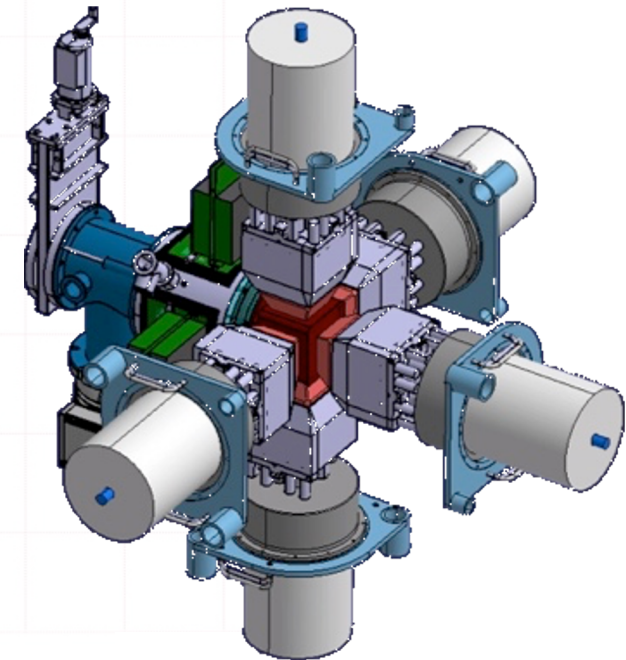
The selection of day one experiments for  $S^3$  was made to determine out the list of pre-proposals with “high impact” (discovery & unique proposals for  $S^3$ ) and feasibility.





# Conclusions and perspectives

- Developing a cutting-edge technology
  - SIRIUS in the SPIRAL2/S<sup>3</sup> framework
  - SIRIUS: “Spectroscopy and Identification of Rare Isotopes Using S3”
  - All the different parts of SIRIUS have been tested offline @IPHC, IJClab, GANIL and IRFU test benches
  - Since July 2020, SIRIUS @CEA/Irfu (DSSD+Tunnels)
  - **Offline commissioning and tests of the complete acquisition chain!**
- March 23-24, 2021:
  - Move SIRIUS from CEA-Saclay to GANIL!
- March-April 2021:
  - Continuation of the tests of SIRIUS @GANIL (Trackers, HPGe, electronics...)
- April 23, 2021:
  - T21-01 new test with LISE2000
  - $^{40}\text{Ar}^{6+}$  4,6 MeV/u
  - $^{40}\text{Ar} + ^{174}\text{Yb} \rightarrow ^{209,210}\text{Ra}$
  - Online commissioning of SIRIUS with LISE2000?



# SIRIUS Collaboration

**CEA/IRFU** : ZF, Th. Chaminade, B. Sulignano, M. Authier,, E. Delagnes, D. Desforge, A. Drouart, A. Grabas, W. Korten, H. Le Provost, Ch. Theisen, M. Vandebrouck, M. Zielinska

**GANIL** : D. Ackermann, M. Blaizot, A. Boujrad, R. Chakma, E. Clément, S. Coudert, S. Herlant, G. Lebertre, C. Maugeais, J. Piot, F. Saillant, H. Savajols, G. Wittwer

**IJClab** : V. Alaphilipe, L. Gibelin, K. Hauschild, N. Karkour, X. Lafay, L. LeBlanc, D. Linget, A. Lopez-Martens

**IPHC** : P. Brionnet, O. Dorvaux, B. Gall, Th. Goeltzenlichter, C. Mathieu

