

Revision of the \bar{P} ANDA Calorimeter Front-End operating parameters

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1 Motivation

2 Setup

3 Analysis

4 Outlook

Structure

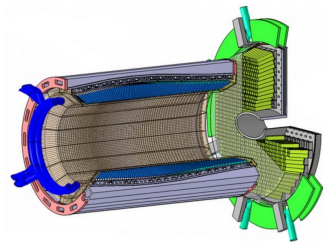
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Main focus on the front-end electronics parameters of the Barrel-EMC, due to them having a major impact on the signal quality and signal-to-noise-ratio.



Technical Design Report for: PANDA
Electromagnetic Calorimeter (EMC)



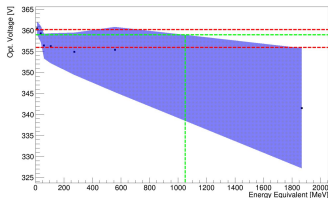
Large Area Avalanche Photodiode Gain
Optimization for the APFEL ASIC
Preamplifiers of the PANDA Calorimeter

BACHELOR THESIS
by Kim Tabea Giebenhain
from Biobertal
August 22, 2019

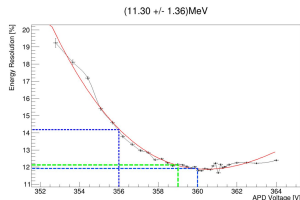
JUSTUS-LIEBIG UNIVERSITY, GIESSEN
II. INSTITUTE OF PHYSICS

INITIAL CORRECTOR: PROF. DR. KAI-THOMAS BRÜCKMANN
SECONDARY CORRECTOR: PROF. DR. CLAUDIA HEINE
TECHNICAL SUPERVISORS: DR. MARKUS MORITZ & DR. HANS-GEORG ZAUBICK

Kim Giebenhain determined an optimal gain for a high energy resolution and therefore a good signal-to-noise-ratio. The results were generated with a light pulser.



"Optimum energy resolution. The shaded area represents a 0.2% resolution loss tolerance. The red lines show the chosen limits for the optimal voltage and the green line shows the selected optimal voltage."



"Optimal energy resolution measurement at 11.3 MeV. The optimal energy resolution at ≈ 360 V is indicated as well as the energy resolution for the lower limit at 357 V."

The APD was most efficient while operated at 359 V (\rightarrow APD-Gain: 500).

The energy resolution for light pulser pulses

- improved from 43 % to 12 % for low energies (11.3 MeV)
- remained mainly constant for high energies (>200 MeV)

compared to an APD-Gain of 150.

Results still had to be proven in an actual accelerator experiment.
More questions arose in the preparations:

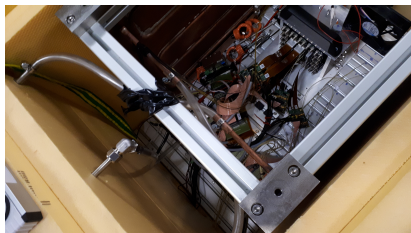
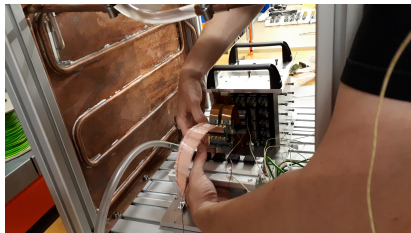
- Improvement through usage of two APDs instead of one?
- Energy resolution wished for the prototypes achievable?
- Both channels of the APFEL needed?

Structure

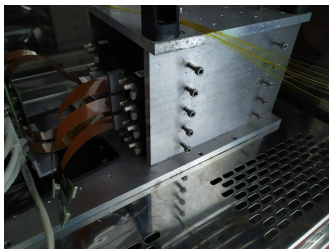
- 1 Motivation
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The setup for the beamtime at MAMI A2 tagger hall in Mainz contained the following components:

- 3 x 3 PWO-II Scintillator Matrix
- Complete readout-chain:
 - LAAPDs (2 per crystal)
 - APFEL-ASICs (1 per detector)
 - Apfelmatices (Backplanes, 1 for 4 ASICs)
 - Bufferboards
 - PANDA-SADC
- Cooling box with a refrigerated circulator for holding a temperature of $-25\text{ }^{\circ}\text{C}$
- Light pulser with glass fibres



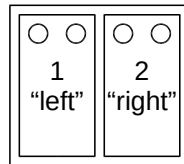
PWO-Matrix



Nine square PWO-II crystals wrapped in reflective foil, put in shrinking tube and equipped with two LAAPDs each, were arranged in a 3 by 3 matrix.



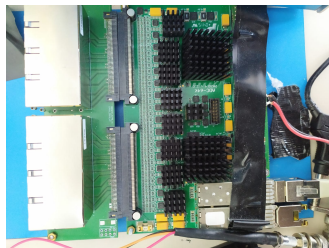
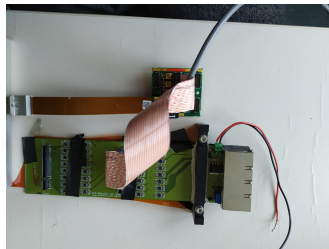
1	4	7
2	5	8
3	6	9



Readout

One APFEL-ASIC per detector gives output to an APFEL-Matrix and a Bufferboard.

This in turn delivers the signal to the PANDA SADC (v2.0) designed by Pawel Marciniewski.



Cooling



Setup was installed in an aluminium box:

- Thermally insulated
- Copper sheets and pipes
- Fans for homogeneity
- Temperature sensors for monitoring



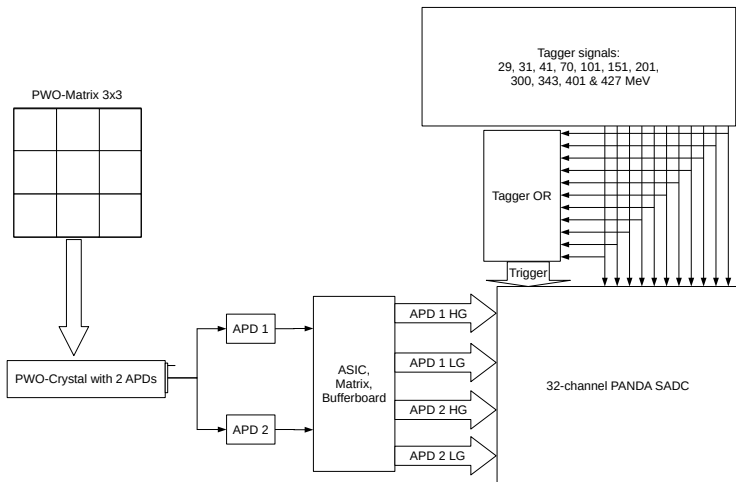
Light pulser

Light-fibre at every crystal, for later calibration.

Response to light pulser was monitored for calibration of the individual detector settings, as well as the response of a reference detector to the same signal.



Overview



ASIC\APD	150	300	400	500
LG($\times 1$)	Beam&LP	Beam&LP	Beam&LP	Beam&LP
HG($\times 16$)	Beam&LP	Beam&LP	Beam&LP	Beam&LP
HG($\times 32$)	Beam&LP	Beam&LP	Beam&LP	Beam&LP
HG($\times 16$)				Cosmic

Structure

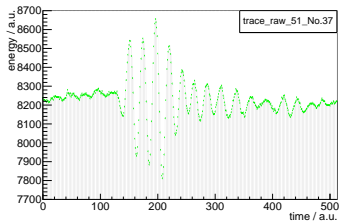
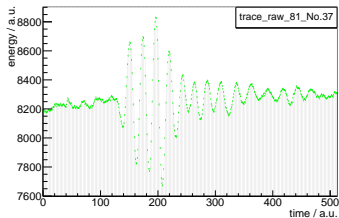
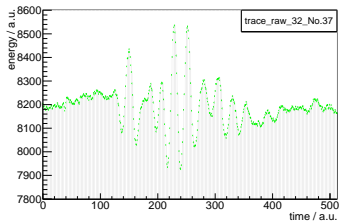
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Pick-up



The pick-up signal appeared coincident in all channels.
The source is still not known.

Feature extraction

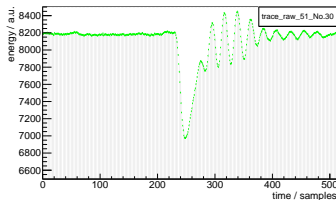
Simple peak-sensing algorithm:

1. find highest/lowest energy value and corresponding time
2. average other data points to construct baseline
3. Subtraction gives the energy in ADC-channels

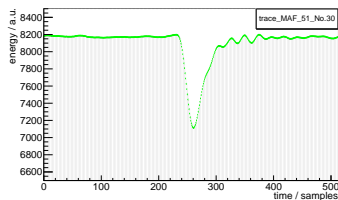
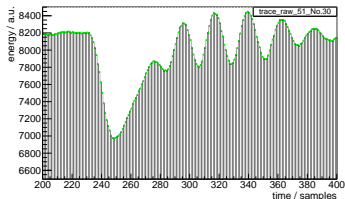
Gives a coarse idea of the signal height and is used for sorting of events.

Moving average filter

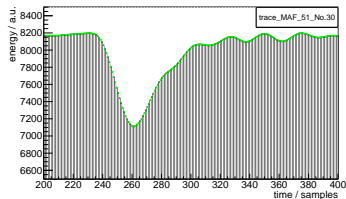
$$y(n) = \frac{1}{L} \cdot \sum_{k=0}^{L-1} x(n - k)$$



raw data

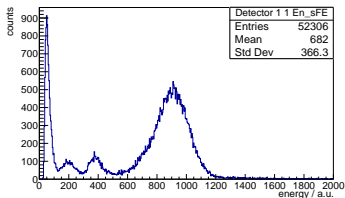
Moving Average $L = 20$ 

raw data zoom

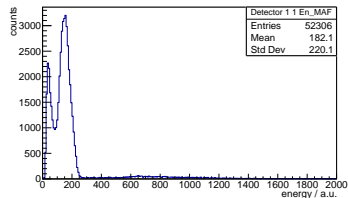
Moving Average $L = 20$ zoom

Event selection on cosmic data

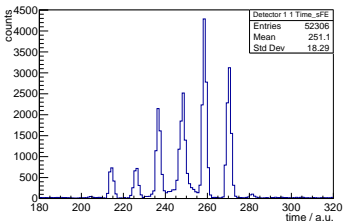
Energy Distribution simple FE



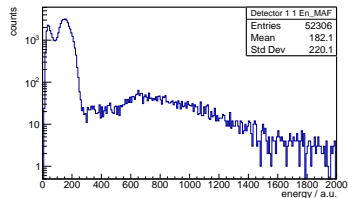
Energy Distribution after MA-Filter



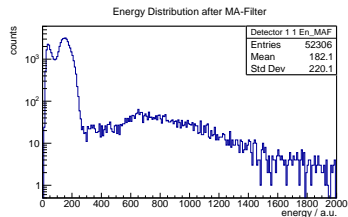
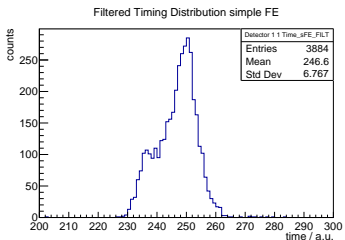
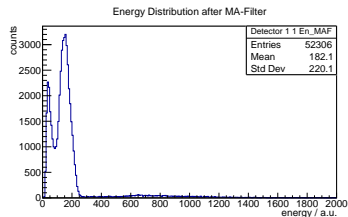
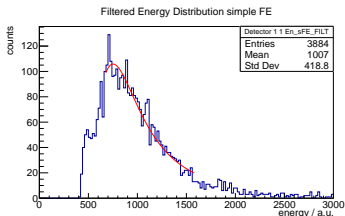
Timing Distribution simple FE



Energy Distribution after MA-Filter



Event selection on cosmic data

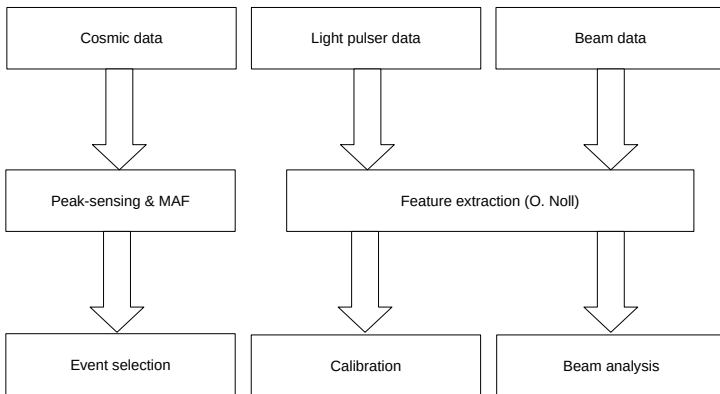


Feature extraction

Actual feature extraction by Oliver Noll's algorithm.

It's used for determining the timing and energy information of the events.

All following results are generated using this feature extraction.

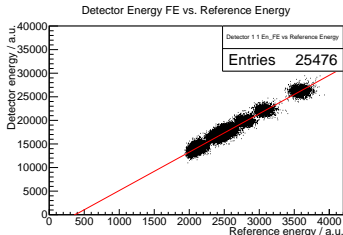


Calibration

Most probable value of deposited energy in one crystal: 28.14 MeV
(derived from GEANT4 calculations)

Calibration factor: $M(ch) = \frac{28.14 \text{ MeV}}{MPV(ch)}$

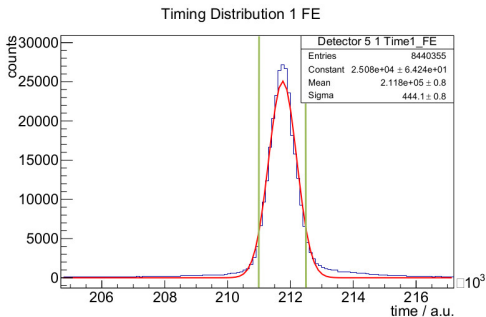
Only valid for APD gain 500 and ASIC gain x16, other settings are calibrated by comparing slopes of the light pulser measurements.



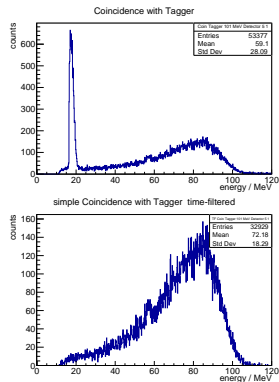
$$M_{ASIC,APD}(ch) = \frac{28.14 \text{ MeV}}{MPV(ch)} \cdot \frac{m_{16,500}(ch)}{m_{ASIC,APD}(ch)}$$

Timecuts

Several timecuts have been applied, to further sort out noise signals. The first cut was applied by the time of the central detector. The chosen window was 211000 to 212500 a.u., being roughly 2σ .



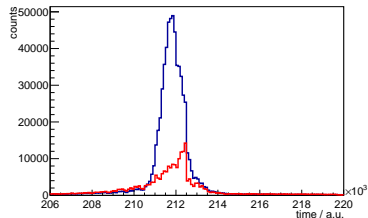
Timing and energy data of central detector coincident with Tagger
101 MeV with and without time filter



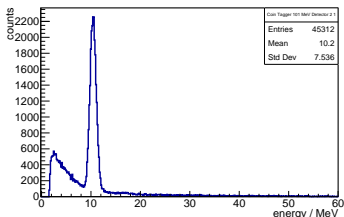
Timecuts

Another timecut was applied on the non-central detectors by the time difference between central and non-central signal.

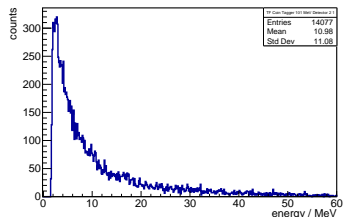
Timing Distribution 1 FE



Coincidence with Tagger

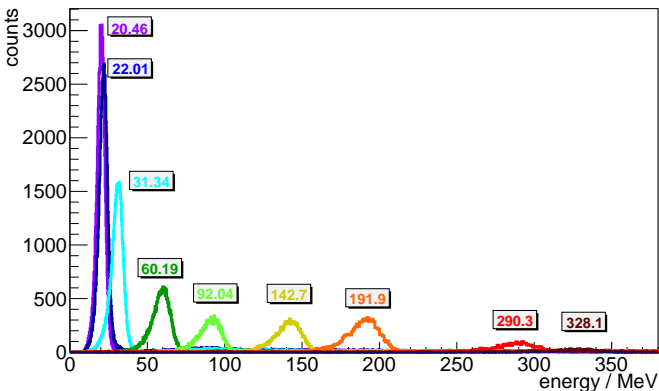


simple Coincidence with Tagger time-filtered



Lineshapes

Energy sum all detectors with energy filter



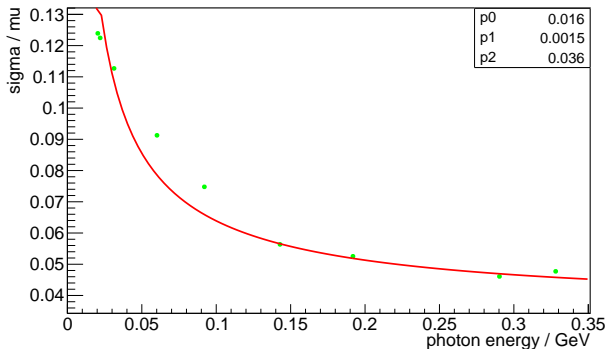
Lineshapes for HG16, APD 500

Additional to the lineshape, the energy resolution $\frac{\sigma}{E}$ is determined. Plotting the energy resolution against the energy results in the response function.

Response function

The response function follows a dependence like the following:

$$\frac{\sigma}{E} = \sqrt{\left(\frac{p_0}{\sqrt{E}}\right)^2 + \left(\frac{p_1}{E}\right)^2 + p_2^2}$$



Response function for ASIC-Gain 16 and APD-Gain 500

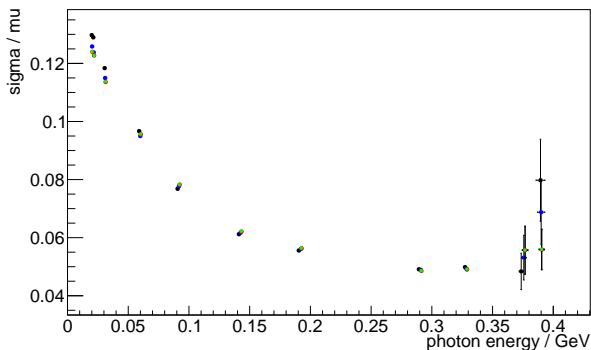
Energywise event selection

Up to this point, three different energy filters have been used for event-selection:

- Central detector has to detect signals above 1 MeV
- Feature extraction applies filter after FIR and only handles signals above 0.3 MeV
- When adding up detectors to the total deposited energy, a filter is applied, the single-crystal (SC) threshold

The effect of the SC-threshold especially, has been further analysed.

SC-threshold

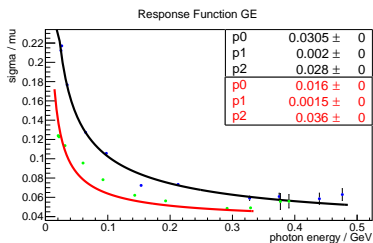


Response function for ASIC-gain 16 and APD-gain 500
SC-thresholds: black 3 MeV, blue 2 MeV, red 1 MeV and green 0.5 MeV

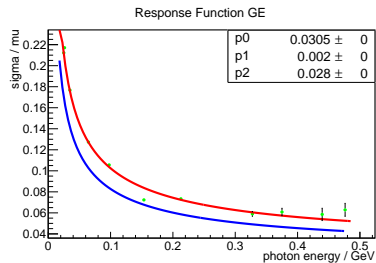
Optimal single-crystal threshold for APD-gain 500: 0.5 MeV

Optimal single-crystal threshold for APD-gain 150: 1-2 MeV

Comparison

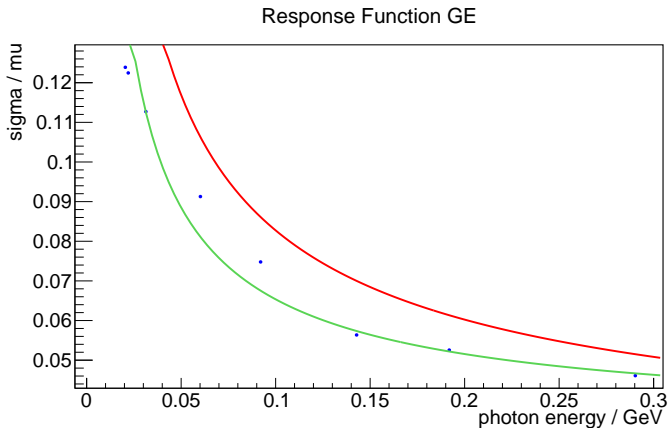


Response functions: black curve: APD-gain 150, red curve: APD-gain 500



Response functions: red curve: APD-gain 150, blue curve: PROTO120

Comparison



Response functions: green curve: APD-gain 500, red curve: PROTO120

Lowest energy data points at ≈ 30 MeV show an improvement of more than 3%

Comparison

Parameter	PROTO120 / %	APD 150 / %	APD 500 / %
p_0	2.46		
p_1	0.16		
p_2	2.32		

p_0 : photonstatistic term

p_1 : electronic term

p_2 : constant term

Comparison

Parameter	PROTO120 / %	APD 150 / %	APD 500 / %
p_0	2.46		
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Results for Gain 150 expected to be worse because of lower light yield due to crystal geometry.

$\Rightarrow p_0 \approx 3.0\%$

Comparison

Parameter	PROTO120 / %	APD 150 / %	APD 500 / %
p_0	2.46	3.05	
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Results for Gain 150 expected to be worse because of lower light yield due to crystal geometry.

$\Rightarrow p_0 \approx 3.0\%$

Results match expectations.

Comparison

Parameter	PROTO120 / %	APD 150 / %	APD 500 / %
p_0	2.46	3.05	1.6
p_1	0.16	0.2	0.15
p_2	2.32	2.8	3.6

p_0 : photonstatistic term

p_1 : electronic term

p_2 : constant term

Results for Gain 150 expected to be worse because of lower light yield due to crystal geometry.

$\Rightarrow p_0 \approx 3.0\%$

Results match expectations.

Significantly reduced statistic term!
 \rightarrow Signal-to-noise-ratio of APDs

Dynamic range

Dynamic range for the low gain ASIC-branch, assuming that the max. detectable energy is roughly 12 times higher than the range of the high gain branch: 2.2 GeV.

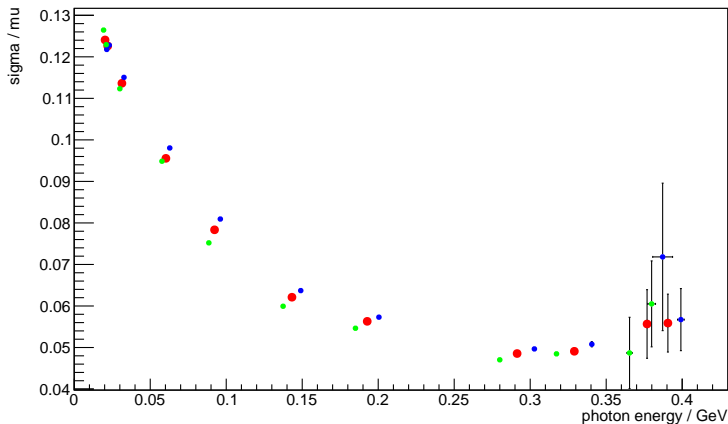
Usage of APD-gain 500 up to 40° , respectively 50° possible, without losing information.

Structure

- 1 Motivation
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Optimization of calibration factors

Response Function GE



Comparison of energy resolutions with different calibration factors on the central detector:
 red: initial data, green: central detector times 0.95, blue central detector times 1.05.
 No correction of the absolute calibration!

Tasks ahead

Currently, the calibration factors are optimized by a genetic algorithm.

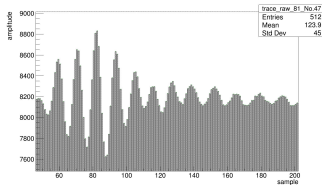
The remaining tasks for the analysis include:

- Energy resolution for remaining settings
- Improve existing fits as possible
- Timing resolution for different settings
- Position resolution for different settings
- Determination of noise level

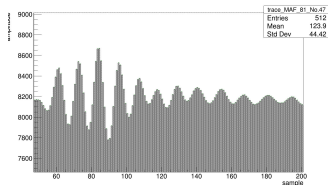
Structure

Addendum

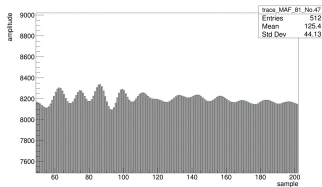
$$y(n) = \frac{1}{L} \cdot \sum_{k=0}^{L-1} x(n-k)$$



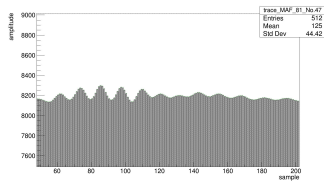
raw data



Moving Average $L = 5$



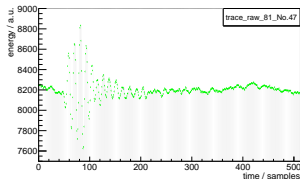
Moving Average $L = 10$



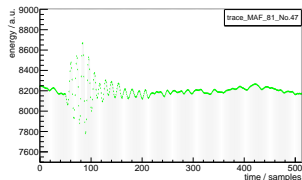
Moving Average $L = 20$

Addendum

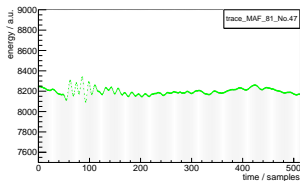
$$y(n) = \frac{1}{L} \cdot \sum_{k=0}^{L-1} x(n-k)$$



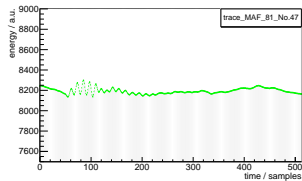
raw data



Moving Average $L = 5$



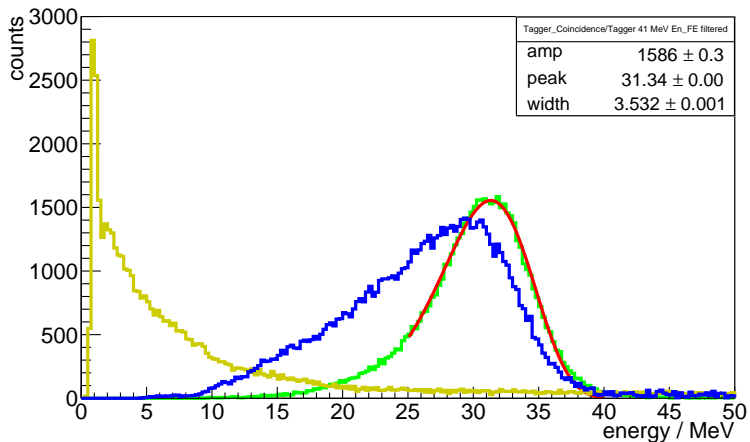
Moving Average $L = 10$



Moving Average $L = 20$

Addendum

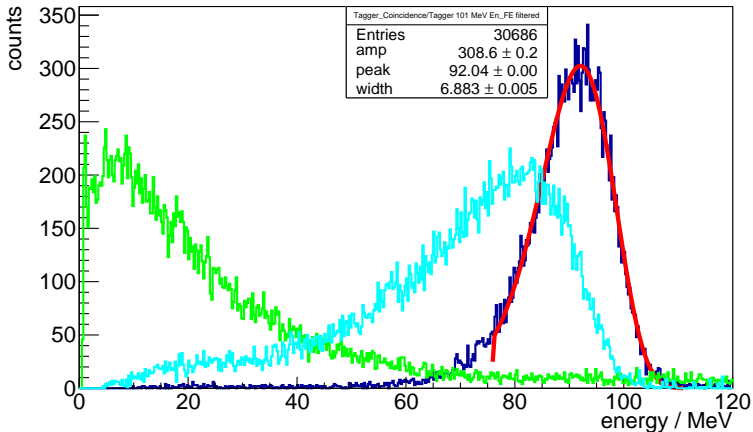
Energy sum all detectors with energy filter



Lineshape for the 41 MeV tagger channel

Addendum

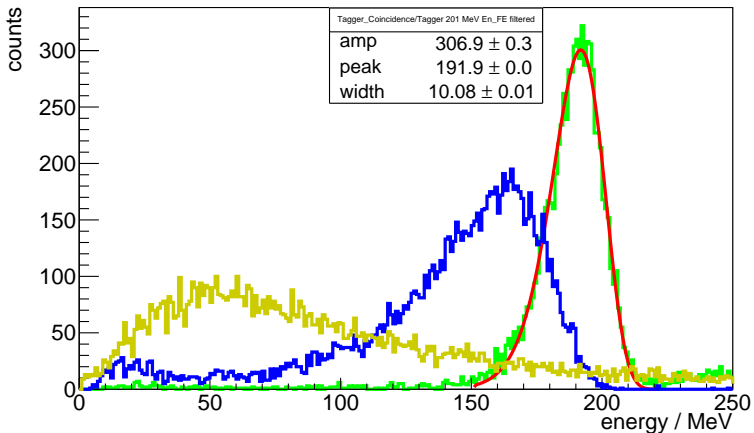
Energy sum all detectors with energy filter



Lineshape for the 101 MeV tagger channel

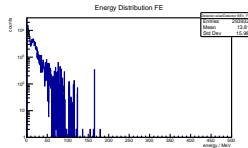
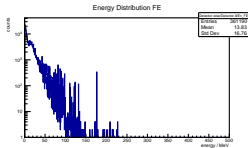
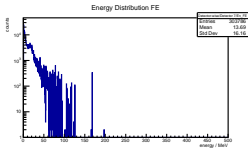
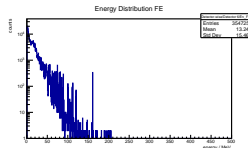
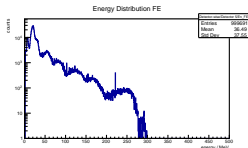
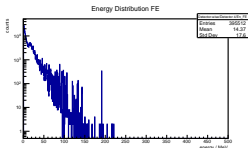
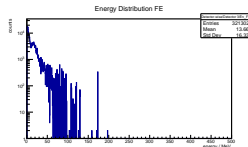
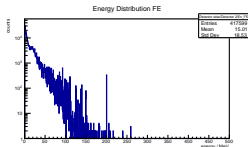
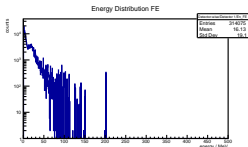
Addendum

Energy sum all detectors with energy filter



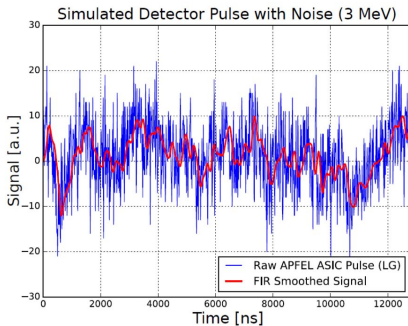
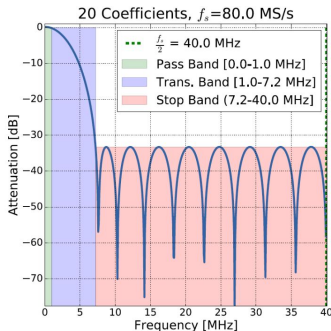
Lineshape for the 201 MeV tagger channel

Addendum



Addendum

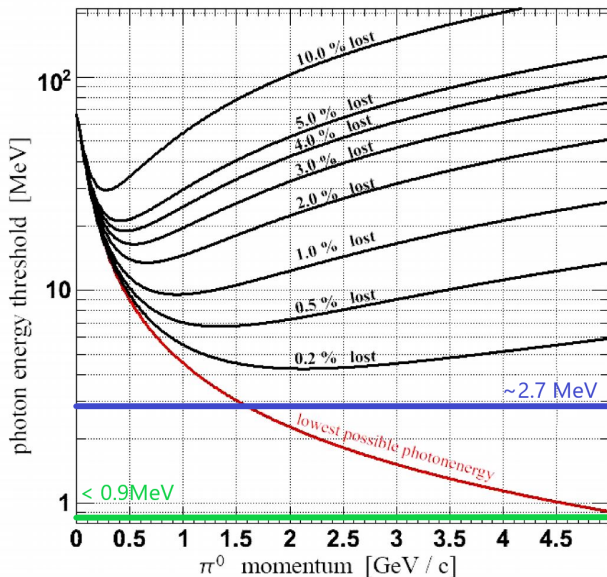
” $\overline{\text{PANDA}}$ Feature Extraction” by Oliver Noll:



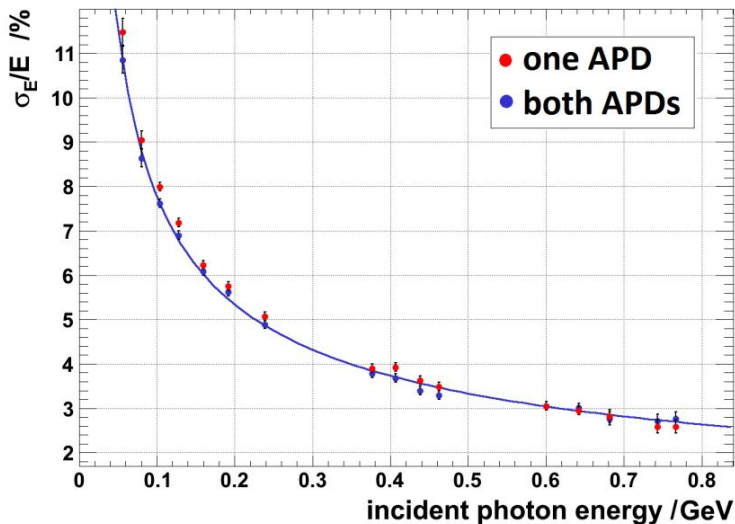
Digital Signal Processing for the EMC PANDA Collaboration Meeting 19/3, Talk by Oliver Noll

Specially tailored for use in $\overline{\text{PANDA}}$, with the APFEL-ASICs.
 ”Low-pass filter” through finite impulse responds filtering.

Addendum



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