

arXiv:2103.10706

New effective interactions for hypernuclei in density dependent relativistic mean field model

Collaborators: Zhong-Hao Tu (ITP/CAS), Shan-Gui Zhou (ITP/CAS, Supervisor)

Speaker :Yu-Ting RongDate:2021-04-15



Motivation

- Density dependent relativistic mean-field model for hypernuclei & effective AN interaction
- Results and discussions
- Summary

Motivation

The known hypernuclei:

- > S=-1: About 40 single Λ hypernuclei (about 20 of them with mass number A ≥ 12)
- > S=-2: A few Ξ hypernuclei and double Λ hypernuclei
- S=+1: Antihypertriton
- (Bare or effective) YN and YY interation:
 - Lattice QCD
 - Chiral effective field theory
 - Nijmegen soft-core model
 - Juelich hyperon-nucleon model
 - Skyrme Hartree Fock model
 - Relativistic mean-field model



H. Tamura. Prog. Theor. Exp. Phys. 2012 (2012) 02B012

Motivation

4 classes of effective interactions for RMF model :

DD: density dependent

Model	NN	ΛN
NL-ME	TM1,TM2,NL1,NL2,NL3,NLSH,P K1,NL-Z,	NLSH-A, PK1-Y1, et al.
DD-ME	TW99, DD-ME1, DD-ME2, PKDD, DD2, DDME-X, DD-LZ1,	DD-ME2-a, DD- ME2D-a
NL-PC	PC-PK1, PC-F1, PC-X	PCY-Si (i=1,2,3,4)
DD-PC	DD-PC1	×
NL: nonlinear coupling ME: meson exchange		

PC: point coupling



Motivation

Effective density dependent interaction in MF approach to study hypernuclei and hypernuclear matter

DD hadron field theory:

- C.M. Keil, F. Hofmann and H. Lenske. Phys. Rev. C 61 (2000) 064309 (A density)
- F. Hofmann, C.M. Keil, H. Lenske. Phys. Rev. C 64 (2001) 025804 (A density / total density)

RMF model:

- P. Finelli, N. Kaiser, D. Vretenar and W. Weise. Nucl. Phys. A 831 (2009) 163 (RMF+chiral perturbation theory; total density, B_A)
- G. Colucci and A. Sedrakian. Phys. Rev. C 87 (2013) 055806 (total denity, DD-ME2+potential) .
- S. Banik, M. Hempel and D. Bandyopadhyay. Astrophys. J. Suppl. Ser. 214 (2014) 22 (total density; DD2+potential at saturation density)
- E.N.E. van Dalen, G. Colucci and A. Sedrakian. Phys. Lett. B 734 (2014) 383 (total density; DD-ME2+mass formula)
- M. Fortin, S.S. Avancini, C. Providencia and I. Vidana. Phys. Rev. C 95 (2017) 065803 (total density / a constant, DD-ME2+B_λ, R_ω is fixed)
- C. Providencia, M. Fortin, H. Pais and A. Rabhi, Frontiers Astron. Space Sci. 6 (2019) 13 (total density)

 - $g_{\sigma\Lambda}$, $g_{\omega\Lambda}$ depend on: Λ density / total density / a constant Fitting procedure: (hyper)nuclei are assume to be spherical + B_{Λ} from experiment or other prediction / Λ potential

Motivation



A. Gal, E.V. Hungerford and D.J. Millener, Rev. Mod. Phys. 88 (2016) 035004. Open shell nuclear core deformation (Multidimensionally constrained RMF model)

Small spin-orbit splitting tensor coupling between Λ and ω

In-medium effect total density dependent coupling constant

Unpaired baryon equal filling approximation

DDRMF model

Effective Lagrangian:

$$\begin{aligned} \mathcal{L}_{B} &= \sum_{B=n,p,\Lambda} \bar{\psi}_{B} \Big(i\gamma_{\mu} \partial^{\mu} - M_{B} - g_{\sigma B} \sigma - g_{\sigma^{*}B} \sigma^{*} - g_{\omega B} \gamma_{\mu} \omega^{\mu} - g_{\phi B} \gamma_{\mu} \phi^{\mu} \\ &- g_{\rho B} \gamma_{\mu} \vec{\tau} \cdot \vec{\rho}^{\mu} - e \gamma_{\mu} \frac{1 - \tau_{3}}{2} A^{\mu} + \frac{f_{\omega \Lambda \Lambda}}{4M_{\Lambda}} \sigma_{\mu\nu} \Omega^{\mu\nu} \Big) \psi_{B} + \frac{1}{2} \partial^{\mu} \sigma \partial_{\mu} \sigma - \frac{1}{2} m_{\sigma}^{2} \sigma^{2} \\ &- \frac{1}{4} \Omega^{\mu\nu} \Omega_{\mu\nu} + \frac{1}{2} m_{\omega}^{2} \omega^{\mu} \omega_{\mu} - \frac{1}{4} \vec{R}^{\mu\nu} \vec{R}_{\mu\nu} + \frac{1}{2} m_{\rho}^{2} \vec{\rho}^{\mu} \vec{\rho}_{\mu} - \frac{1}{4} F^{\mu\nu} F_{\mu\nu} \\ &+ \frac{1}{2} \partial_{\mu} \sigma^{*} \partial^{\mu} \sigma^{*} - \frac{1}{2} m_{\sigma^{*}} \sigma^{*2} - \frac{1}{4} S^{\mu\nu} S_{\mu\nu} + \frac{1}{2} m_{\phi}^{2} \phi^{\mu} \phi_{\mu}, \end{aligned}$$

Formation for density-dependent coupling constants:

$$g_{mB}(\rho^{\nu}) = g_{mB}(\rho_{\text{sat}}) f_{mB}(x), \quad x = \rho^{\nu} / \rho_{\text{sat}},$$
$$f_{mN}(x) = \begin{cases} a_m \frac{1 + b_m (x + d_m)^2}{1 + c_m (x + d_m)^2}, & m = \sigma \text{ or } \omega, \\ e^{-a_\rho (x - 1)}, & m = \rho, \end{cases}$$

S. Typel and H.H. Wolter. Nucl. Phys. A 656 (1999) 331

Relation between NN and ΛN

$$g_{\sigma\Lambda}(\rho) = R_{\sigma}g_{\sigma N}(\rho),$$

$$g_{\omega\Lambda}(\rho) = R_{\omega}g_{\omega N}(\rho),$$

$$g_{\sigma^*\Lambda}(\rho) = R_{\sigma^*}g_{\sigma N}(\rho),$$

$$g_{\phi\Lambda}(\rho) = R_{\phi}g_{\omega N}(\rho),$$

$$f_{\omega\Lambda\Lambda}(\rho) = R_{\omega\Lambda\Lambda}R_{\omega}g_{\omega N}(\rho).$$

DDRMF model

Effective Lagrangian:

$$\begin{split} \mathcal{L}_{B} &= \sum_{B=n,p,\Lambda} \bar{\psi}_{B} \Big(i\gamma_{\mu} \partial^{\mu} - M_{B} - g_{\sigma B} \sigma - g_{\sigma^{*}B} \sigma^{*} - g_{\omega B} \gamma_{\mu} \omega^{\mu} - g_{\phi B} \gamma_{\mu} \phi^{\mu} \\ &- g_{\rho B} \gamma_{\mu} \vec{\tau} \cdot \vec{\rho}^{\mu} - e \gamma_{\mu} \frac{1 - \tau_{3}}{2} A^{\mu} + \frac{f_{\omega \Lambda \Lambda}}{4M_{\Lambda}} \sigma_{\mu\nu} \Omega^{\mu\nu} \Big) \psi_{B} + \frac{1}{2} \partial^{\mu} \sigma \partial_{\mu} \sigma - \frac{1}{2} m_{\sigma}^{2} \sigma^{2} \\ &- \frac{1}{4} \Omega^{\mu\nu} \Omega_{\mu\nu} + \frac{1}{2} m_{\omega}^{2} \omega^{\mu} \omega_{\mu} - \frac{1}{4} \vec{R}^{\mu\nu} \vec{R}_{\mu\nu} + \frac{1}{2} m_{\rho}^{2} \vec{\rho}^{\mu} \vec{\rho}_{\mu} - \frac{1}{4} F^{\mu\nu} F_{\mu\nu} \\ &+ \frac{1}{2} \partial_{\mu} \sigma^{*} \partial^{\mu} \sigma^{*} - \frac{1}{2} m_{\sigma^{*}} \sigma^{*2} - \frac{1}{4} S^{\mu\nu} S_{\mu\nu} + \frac{1}{2} m_{\phi}^{2} \phi^{\mu} \phi_{\mu}, \end{split}$$

Formation for density-dependent coupling constants:

$$g_{mB}(\rho^{\nu}) = g_{mB}(\rho_{\text{sat}}) f_{mB}(x), \quad x = \rho^{\nu} / \rho_{\text{sat}},$$
$$f_{mN}(x) = \begin{cases} a_m \frac{1 + b_m (x + d_m)^2}{1 + c_m (x + d_m)^2}, & m = \sigma \text{ or } \omega, \\ e^{-a_\rho (x - 1)}, & m = \rho, \end{cases}$$

S. Typel and H.H. Wolter. Nucl. Phys. A 656 (1999) 331

Relation between NN and ΛN

$$g_{\sigma\Lambda}(\rho) = R_{\sigma}g_{\sigma N}(\rho),$$

$$g_{\omega\Lambda}(\rho) = R_{\omega}g_{\omega N}(\rho),$$

$$g_{\sigma^*\Lambda}(\rho) = R_{\sigma^*}g_{\sigma N}(\rho),$$

$$g_{\phi\Lambda}(\rho) = R_{\phi}g_{\omega N}(\rho),$$

$$f_{\omega\Lambda\Lambda}(\rho) = R_{\omega\Lambda\Lambda}R_{\omega}g_{\omega N}(\rho).$$
-1
Ainimization of the least-square

Minimization of the least-square deviation

$$\chi^{2}(\boldsymbol{a}) = \sum_{i}^{N} \left(\frac{B_{\Lambda,i}^{\text{exp.}} - B_{\Lambda,i}^{\text{cal.}}(x_{i};\boldsymbol{a})}{\Delta B_{\Lambda,i}^{\text{exp.}}} \right)^{2},$$

Results: Selections for fitting

NN interaction: DD-ME2 (G. A. Lalazissis, T. Niksic, D. Vretenar et al. Phys. Rev. C 71 (2005) 024312) PKDD (W. H. Long, J. Meng, N. Van Giai et al. Phys. Rev. C 69 (2004) 034319)

Experimental data for fitting:

- \ /-	15 00.00 00 000 000 000 000 000 000 000 0
NU	$B_{\Lambda, exp}$
$^{12}_{\Lambda}$ C	11.36 ± 0.20
$^{13}_{\Lambda}$ C	12.0 ± 0.2
$^{16}_{\Lambda}O$	13.0 ± 0.2
²⁸ Si	17.2 ± 0.2
$^{32}_{\Lambda}S$	17.5 ± 0.5
$^{40}_{\Lambda}$ Ca	18.7 ± 1.1
51 V	21.5 ± 0.6
$^{52}_{\Lambda}$ V	21.8 ± 0.3
⁸⁹ Y	23.6 ± 0.5
¹³⁹ La	25.1 ± 1.2
$^{208}_{\Lambda} Pb$	26.9 ± 0.8

A. Gal, E.V. Hungerford and D.J. Millener. Rev. Mod. Phys. 88 (2016) 035004. P.H. Pile, S. Bart, et al. Phys. Rev. Lett. 66 (1991) 2585.

¹⁶_{Λ}N: B_{Λ}=13.76(16) MeV, CSB

 ${}^{12}{}_{\Lambda}B: B_{\Lambda} = 11.52(2) \text{ MeV, CSB}$

 13,14 _AC: only emulsion data

Root mean square (rms) deviation:

$$\Delta = \sqrt{\frac{1}{N} \sum_{i}^{N} \left(B_{\Lambda,i}^{\text{exp.}} - B_{\Lambda,i}^{\text{cal.}} \right)^2},$$

Root of relative square (rrs) deviation:

$$\delta = \sqrt{\frac{1}{N} \sum_{i}^{N} \left(\frac{B_{\Lambda,i}^{\text{exp.}} - B_{\Lambda,i}^{\text{cal.}}}{B_{\Lambda,i}^{\text{exp.}}} \right)^{2}},$$

Results: Properties of normal core nuclei

- > Nuclear cores: ¹¹C, ¹²C, ¹⁵O, ²⁷Si, ³¹S, ³⁹Ca, ⁵⁰V, ⁵¹V, ⁸⁸Y, ¹³⁸La and ²⁰⁷Pb
- Binding energies calculated with DD-ME2 and PKDD are close to AME2016 (relative deviations are within 5%)
- > Axially deformed shape are obtained except ¹⁵O and ²⁰⁷Pb



Results: Λ separation energies & deformation

NN interaction: DD-ME2



Results: Λ separation energies & deformation

NN interaction: DD-ME2

Reference hypernuclei: ²⁰⁸_APb, ¹³⁹_ALa, ⁸⁹_AY, ^{51,52}_AV, ⁴⁰_ACa (Group 3)



Results: Λ separation energies & deformation

NN interaction: DD-ME2

Reference hypernuclei: ²⁰⁸_APb, ¹³⁹_ALa, ⁸⁹_AY, ^{51,52}_AV, ⁴⁰_ACa, ³²_ASi, ²⁸_ASi, ¹⁶_AO (Group 2)



Results: Λ separation energies & deformation

NN interaction: DD-ME2

Reference hypernuclei: ²⁰⁸_APb, ¹³⁹_ALa, ⁸⁹_AY, ^{51,52}_AV, ⁴⁰_ACa, ³²_AS, ²⁸_ASi, ¹⁶_AO, ¹³_AC, ¹²_AC (Group 1)



Results: Λ separation energies & deformation

NN interaction: PKDD



Results: Λ separation energies & deformation

NN interaction: PKDD

Reference hypernuclei: ${}^{208}_{\Lambda}$ Pb, ${}^{139}_{\Lambda}$ La, ${}^{89}_{\Lambda}$ Y, ${}^{51,52}_{\Lambda}$ V, ${}^{40}_{\Lambda}$ Ca (Group 3)



Results: Λ separation energies & deformation

NN interaction: PKDD

Reference hypernuclei: ²⁰⁸_APb, ¹³⁹_ALa, ⁸⁹_AY, ^{51,52}_AV, ⁴⁰_ACa, ³²_AS, ²⁸_ASi, ¹⁶_AO (Group 2)



Results: Λ separation energies & deformation

NN interaction: PKDD

Reference hypernuclei: ²⁰⁸_APb, ¹³⁹_ALa, ⁸⁹_AY, ^{51,52}_AV, ⁴⁰_ACa, ³²_AS, ²⁸_ASi, ¹⁶_AO, ¹³_AC, ¹²_AC (Group 1)



Results: Parameter correlation

 $R\sigma$ and $R\omega$ are correlated —> Λ potential B_{Λ} of light hypernuclei —> small $R\sigma$ —> mass dependent? Structure effect?



$$R_{\omega} = 1.228R_{\sigma} - 0.097$$

Results: Theoretical uncertainty & parameter correlation



Results: Theoretical uncertainty & parameter correlation





Results: Hypernuclear density

Impact on density —> change the Λ density in light hyernuclei



Results: Hyperon star

- Baryons: $p, n, \Lambda, \Sigma^+, \Sigma^0, \Sigma^-, \Xi^+, \Xi^+$ leptons: e^-, μ^-
- mesons: σ, ω, ρ
- Coupling constants:

$$\begin{split} &2g_{\omega\Xi} = g_{\omega\Sigma} = 2g_{\omega N}/3, \\ &g_{\rho\Sigma} = 2g_{\rho\Xi} = 2g_{\rho N}, \\ &U_Y^{(N)} = g_{\sigma Y}\sigma + g_{\omega Y}\omega + g_{\rho Y}\rho + \Sigma_R, \\ &U_{\Sigma}^{(N)} = +30 \text{ MeV}, \ U_{\Xi}^{(N)} = -15 \text{ MeV}. \end{split}$$

Equation of state:

$$\begin{split} \varepsilon &= \frac{1}{2}m_{\sigma}^{2}\sigma^{2} + \frac{1}{2}m_{\omega}^{2}\omega^{2} + \frac{1}{2}m_{\rho}^{2}\rho^{2} + \sum_{B}\varepsilon_{\mathrm{kin}}^{B}, \\ P &= -\frac{1}{2}m_{\sigma}^{2}\sigma^{2} + \frac{1}{2}m_{\omega}^{2}\omega^{2} + \frac{1}{2}m_{\rho}^{2}\rho^{2} + \rho^{\nu}\Sigma_{R} + \sum_{B}P_{\mathrm{kin}}^{B}. \end{split}$$

TOV equation:

$$\begin{split} \frac{dP}{dr} &= -\frac{[P(r) + \varepsilon(r)][M(r) + 4\pi r^3 P(r)]}{r(r - 2M(r))},\\ \frac{dM}{dr} &= 4\pi^2 \varepsilon(r). \end{split}$$

Results: Hyperon star

Impact on hyperon star —>The lighter R_{σ} , the softer EoS and smaller maximum mass of neutron star.



Results: Excited states and single particle levels

- > Excited energies in p_{Λ} , d_{Λ} and f_{Λ} states are close to experimental values
- > Energy splitting for spin partners are close to the experimental value with tensor coupling





S. Ajimura, H. Hayakawa, T. Kishimoto, et al. Phys. Rev. Lett. 86 (2001) 4255.

Summary

- The effective ∧N interactions for hypernuclei are investigated by a deformed DDRMF model and new parameter sets are proposed.
- Shape change effect is obvious in the A \leq 40 mass region
- Data from light hypernuclei make the ΛN coupling constants smaller
- R_{σ} and R_{ω} are strongly correlated and uncertainties for R_{σ} are given
- Density for Λ in light mass region and EoS of hypernuclear matter are sensitive to the parameters
- Single particle excited energies are close to the experimental values
- Spin-orbit splitting in *p* state can be well described by including tensor coupling

Thanks for discussions with Johann Haidenbauer, Hoai Le, Andreas Nogga, Xiang-Xiang Sun and Kun Wang.





Thanks for your attention!

Yu-Ting Rong Supervisor: Shan-Gui Zhou Institute of Theoretical Physics, Chinese Academy of Sciences