Laser-driven relativistic electron layers for coherent Thomson scattering

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A. Einstein, Annalen der Physik 17, 891 (1905)



Challenge: Ultrathin foils - high contrast laser pulses

Charge separation



lon acceleration regime $(s \sim d)$

For laser fields

$$a_0 \approx \varepsilon_0 = (n_e / n_{crit}) k_L d$$

the electrostatic field just balances the light pressure and the whole foil is accelerated.

This requires ultra-thin foils in the order of 10 - 100 nm thick.

Also circular polarized light is needed to keep electrons cold!

Radiation-Pressure Acceleration of Ion Beams Driven by Circularly Polarized Laser Pulses

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Collection and focusing of laser accelerated ion beams for therapy applications

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620 J, 62 fs, 10 PW laser pulse 400 nm foil

Electron blow-out regime ($a_0 > \varepsilon_0$)

J. Meyer-ter-Vehn and Huichun Wu, Eur. Phys. J. D55, 455 (2009)



 $\gamma_{\rm max} \approx 19$

Electron density and γ evolution

electron layer after1 cycle of interaction, wave front depleted due to electron acceleration.

 γ evolution of electrons at front peak



Expected backscattered probe light upshift $4\gamma_{\text{max}}^2 \approx 4 \times 19^2 \approx 1444$



Great disappointment! Spectral cutoff at 36 ω_0 , not at 1444 ω_0 !

For coherent Thomson scattering, transverse momentum p_{\perp} degrades Doppler factor



Uniform Laser-Driven Relativistic Electron Layer for Coherent Thomson Scattering



H.-C. Wu (武慧春),^{1,*} J. Meyer-ter-Vehn,² J. Fernández,¹ and B. M. Hegelich^{1,3}

New idea how to suppress transverse electron momentum p_{\perp} : Use additional reflector reflecting pump light, but let relativistic electrons (REM) pass unperturbed!



Reflector (about 20 nm thick ~ skin depth)

Reflected light turns $p_{\perp} \rightarrow 0$, while changing p_x only marginally!



Relativistic electrons emerge from reflector with $p_{\perp} = 0$.



Simple argument to understand

Conservation of canonical momentum

$$p_{\perp} - a = \text{const}$$

 $p_{\perp} \rightarrow 0$





2D-PIC simulations H.C. Wu (2010)

$$a_0 = 3.5, n_e / n_{crit} = 1, d / \lambda = 0.001$$

without reflector

with reflector





Signal and spectrum from electron mirror





with reflector







Case for experiments now

 $\gamma \gamma \gamma_x$

Last picture of movie in previous viewgraph



Reflectivity of relativistic mirror

H.C. Wu, J. Meyer-ter-Vehn, et al. PRL104, 234801 (2010)

Electron density:

$$n_e(x) = n_0 S(x/d)$$

Coherently backscattered amplitude:

$$a_{\rm refl} / a_{\rm inc} = \gamma \frac{n_0 k_{\rm L} d}{n_{\rm crit}} F(\xi)$$

Form factor:

$$F(\xi) = \int_{-\infty}^{\infty} S(\chi) \, \cos(\chi \xi) \, d\chi$$

in rest frame of mirror $\chi = x'/d'$ $\xi = 2k'_Ld'$

in lab frame

$$\xi = 2\gamma^2 (1 + \beta) k_L d \approx 2\pi d / \lambda_{\text{refl}}$$

for Gaussian profile :

$$S(\chi) = \exp(-\pi\chi^2)$$

$$F(\xi) = \exp\left(-\frac{\xi^2}{4\pi}\right) = \exp\left(-\left(\frac{d}{2\lambda_{\text{refl}}}\right)^2\right)$$

Reflected amplitude decays exponentially for
$$\lambda_{
m refl} < d$$

Non-linear coherent Thomson scattering

H.C. Wu, J. Meyer-ter-Vehn, et al. PRSTAB (submitted 2011)

Doppler factor :

$$D = 4\gamma_x^2 / (1 + a_{probe}^2)$$



Reflected amplitude :

$$\frac{\mathrm{E}_{\mathrm{refl}}}{\mathrm{E}_{\mathrm{inc}}} = \frac{\gamma n_0 k_L d}{n_{crit}} \cdot \frac{\mathrm{F}(2 \pi d / \lambda_{\mathrm{refl}})}{(1 + a_{probe}^2)}$$



Ultimate goal: medical application

- medical applications require photons above 20 keV
- phase-contrast imaging :

phase shifts much stronger than absorption,

breast cancer tissue, photographs by ESRF



X-ray absorption



same with phase-contrast

Conclusions

Relativistic electron mirrors can be used to compress probe pulses and to upshift frequency by factors $4\gamma^2$.

Coherent backscattering requires dense electon sheets. They can be produced by blowing out all electrons from ultrathin foils with a strong drive laser pulse.

Transverse electron momentum degrading the Doppler-shift factor can be removed by reflecting the drive laser pulse by a reflector foil That lets the relativistic electron layer pass almost unperturbed.