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Coherent Short Wavelength Radiation: Relativistic High Harmonics from Overdense Plasmas

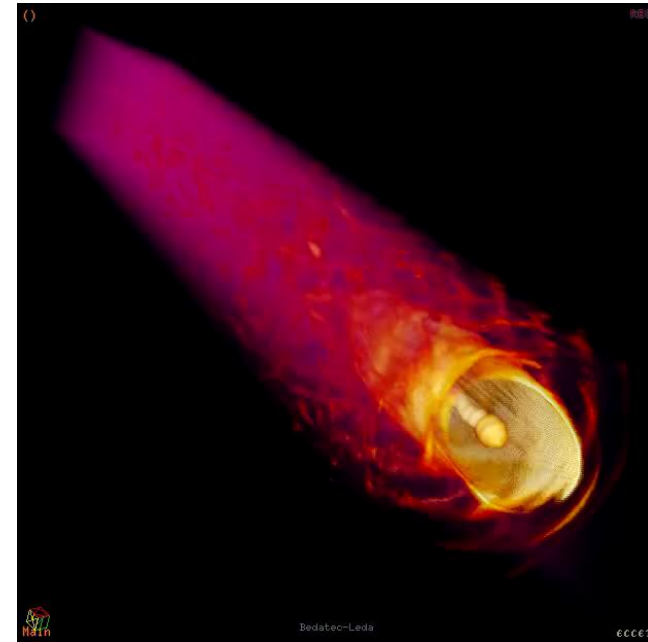
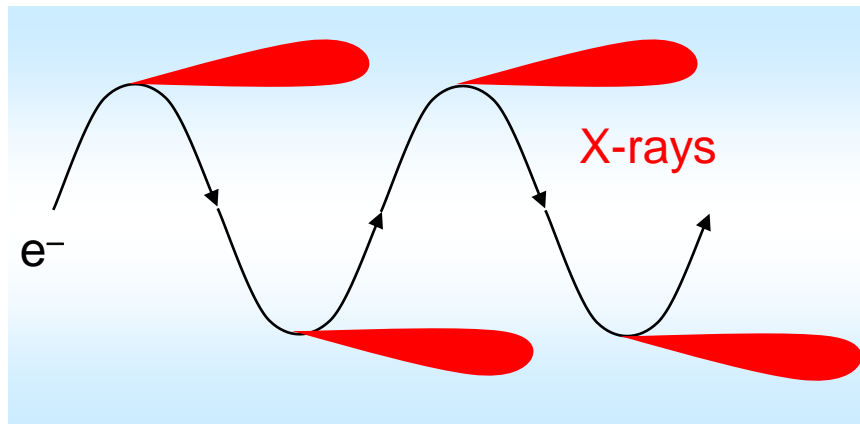
EMMI-Workshop, GSI Darmstadt, 2011

Outline

- **X-ray emission by relativistic electrons**
 - **Relativistic harmonics and attosecond pulses from overdense plasmas, spectrum $n^{-8/3}$**
 - **Coherent Synchrotron Emission (CSE) spectrum $n^{-6/5}$**
 - **Spectral modulations and the femtosecond plasma surface dynamics**
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Synchrotron radiation from the Laser-Plasma Bubble

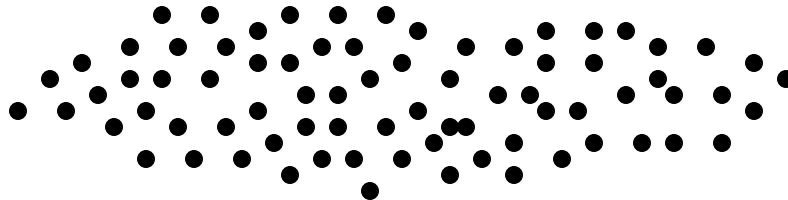
see Kiselev, Pukhov, Kostyukov, Phys. Rev. Lett. **93**, 135004 (2004)



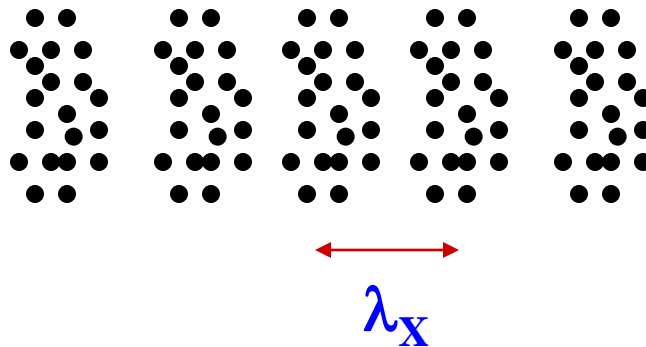
Electron moving on a circular orbit in the ultrarelativistic limit emits synchrotron radiation (high harmonics) when $\Delta\theta > 1/\gamma$.

Incoherent vs coherent emission

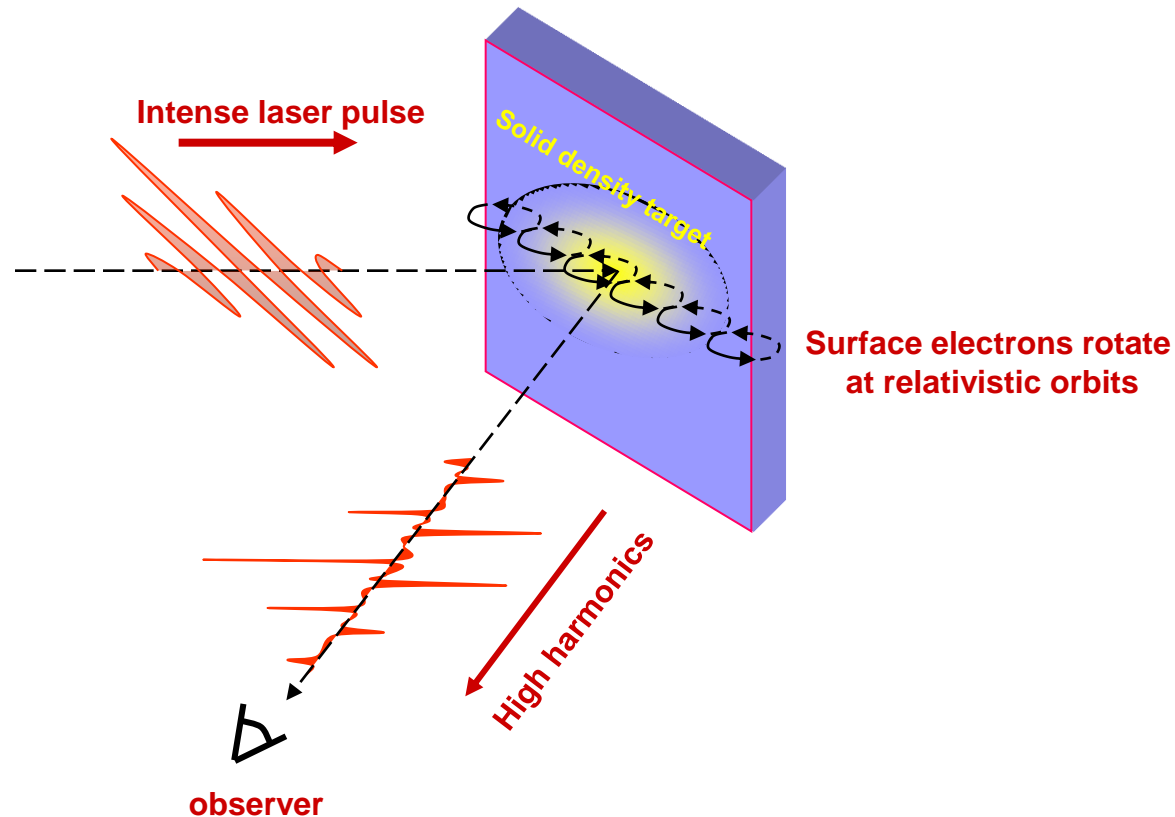
Unbunched electrons: incoherent emission: $N_{\text{ph}} \sim N_e$.



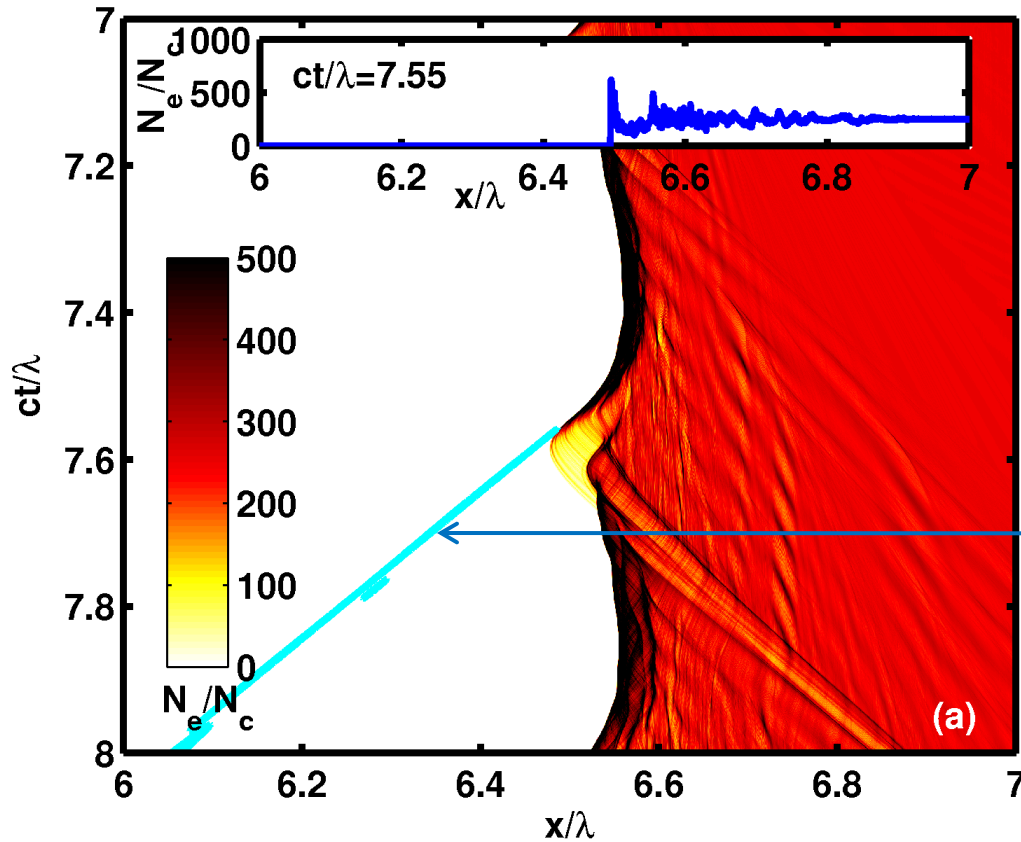
Electrons bunched at λ_x : coherent emission: $N_{\text{ph}} \sim N_e^2$.



Relativistic High Harmonics from solid state target surfaces



Relativistic Harmonics



Parameters:
 $N_e = 250 N_c$
 step profile
 $a_0 = 60$

Harmonic emission

Analytic description of HHG

Wave equation
$$-\frac{1}{c^2} \partial_t^2 \mathbf{A}(t, x) + \partial_x^2 \mathbf{A}(t, x) = -\frac{4\pi}{c} \mathbf{j}_\perp(t, x)$$

Solution
$$\mathbf{A}(t, x) = 4\pi \int_{-\infty}^{+\infty} \mathbf{j}_\perp(t', x') G(t, x, t', x') dt' dx'$$

Green function satisfying boundary conditions

$$G(t, x, t', x') = \frac{1}{2} \left[\theta \left(t - t' - \frac{|x - x'|}{c} \right) - \theta \left(t - t' - \frac{x - x'}{c} \right) \right]$$

Analytic description of HHG

Electric field

$$\mathbf{E}_{\perp}(t, x) = \frac{2\pi}{c} \int_x^{+\infty} \left[\mathbf{j}_{\perp} \left(t + \frac{x - x'}{c}, x' \right) - \mathbf{j}_{\perp} \left(t - \frac{x - x'}{c}, x' \right) \right] dx'$$

↑

Reflected
wave

↑

Incoming
wave

So that at the left of plasma

$$\mathbf{E}_{\perp}(t, x) = \mathbf{E}_i \left(t - \frac{x}{c} \right) + \mathbf{E}_r \left(t + \frac{x}{c} \right)$$

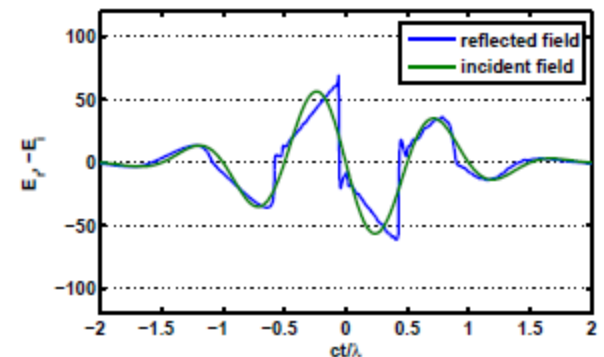
The ROM boundary condition

Baeva, Gordienko, Pukhov, *Phys. Rev. E* **74**, 046404 (2006)

- Skin layer evolution time long compared to skin length (e.g. step-like profile), then follows ARP boundary condition

$$E_i(x_{ARP}(t) - ct) + E_r(x_{ARP}(t) + ct) = 0$$

- Reflected field phase modulation of incident field
 - same maxima and minima
 - same sequence of monotonic intervals



Normal incidence, $N_e = 250 N_c$, sharp edged profile, $a_0 = 60$

Analytical derivation of the BGP spectrum

Fourier transform of the reflected wave

$$E_r(\omega) = -\int E_i\left(t - \frac{x_{ARP}}{c}\right) e^{i\omega(t+x/c)} \left(1 + \frac{\dot{x}_{ARP}}{c}\right) dt$$

We assume for the incident wave $E_i(t) = a(t) \left[e^{i\omega_0 t} + e^{-i\omega_0 t} \right] / 2$

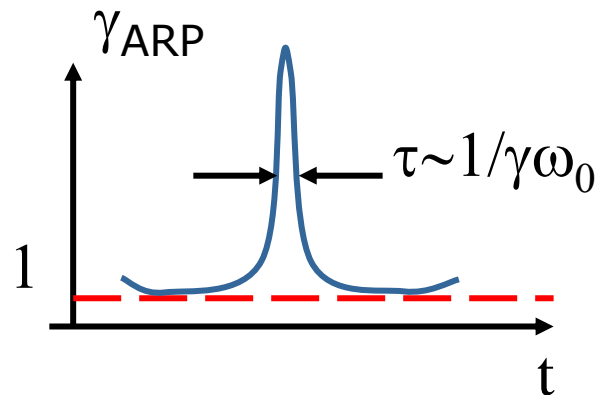
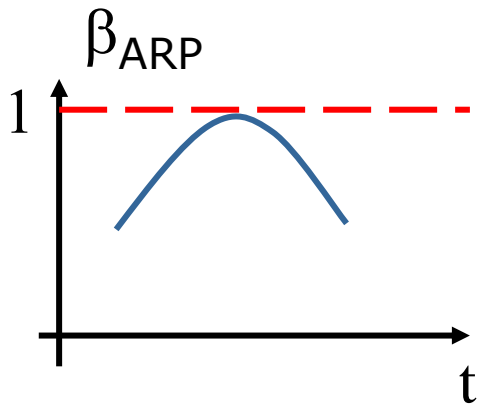
The reflected wave is thus $E_r(\omega) = E_+ + E_-$

$$E_{\pm}(\omega) = -\int a\left(t - \frac{x_{ARP}}{c}\right) e^{i\left[\omega\left(t + \frac{x_{ARP}}{c}\right) \pm \omega_0\left(t - \frac{x_{ARP}}{c}\right)\right]} \left(1 + \frac{\dot{x}_{ARP}}{c}\right) dt$$

Stationary points and the γ -spikes

The stationary phase points correspond to the instants when the apparent reflection point (ARP) moves towards the observer with maximum velocity.

The corresponding ARP gamma factor $\gamma_{\text{ARP}} = \left(1 - \dot{x}_{\text{ARP}}^2\right)^{-1/2}$

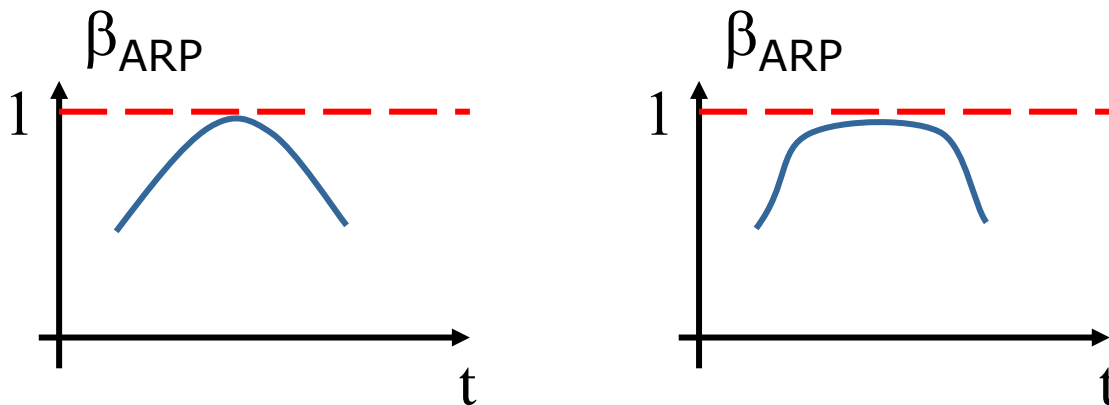


Order of the γ -spikes

The spectrum depends on the exact behavior of the ARP in the neighborhood of the γ -spike.

The $2n$ -th order maximum of the surface velocity

$$\frac{d^k}{dt^k} x_{\text{ARP}} = 0, \quad \text{for all } 1 < k < 2n - 1$$



General form of the spectrum

The spectrum can be calculated for an arbitrary order n of the surface velocity maximum

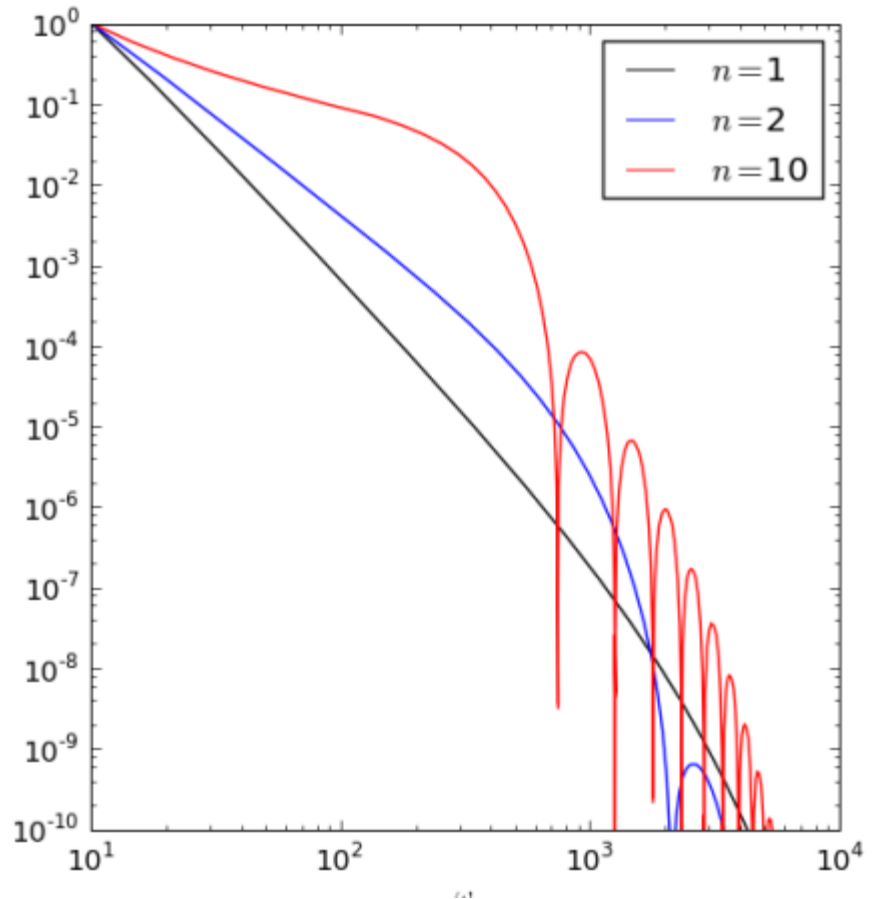
$$I_n(\omega) \propto \omega^{-\frac{4n+4}{2n+1}} \left[\text{Gai}_n \left(\frac{\omega\gamma^{-2} - 4\omega_0}{2(\alpha\omega)^{1/(2n+1)}} \right) - \text{Gai}_n \left(\frac{\omega\gamma^{-2} + 4\omega_0}{2(\alpha\omega)^{1/(2n+1)}} \right) \right]^2$$

Where Gai_n is the generalized Airy function:

$$\text{Gai}_n(x) = \frac{1}{2\pi} \int_{-\infty}^{\infty} \exp\left[i\left(xt + \frac{t^{2n+1}}{(2n+1)}\right)\right] dt$$

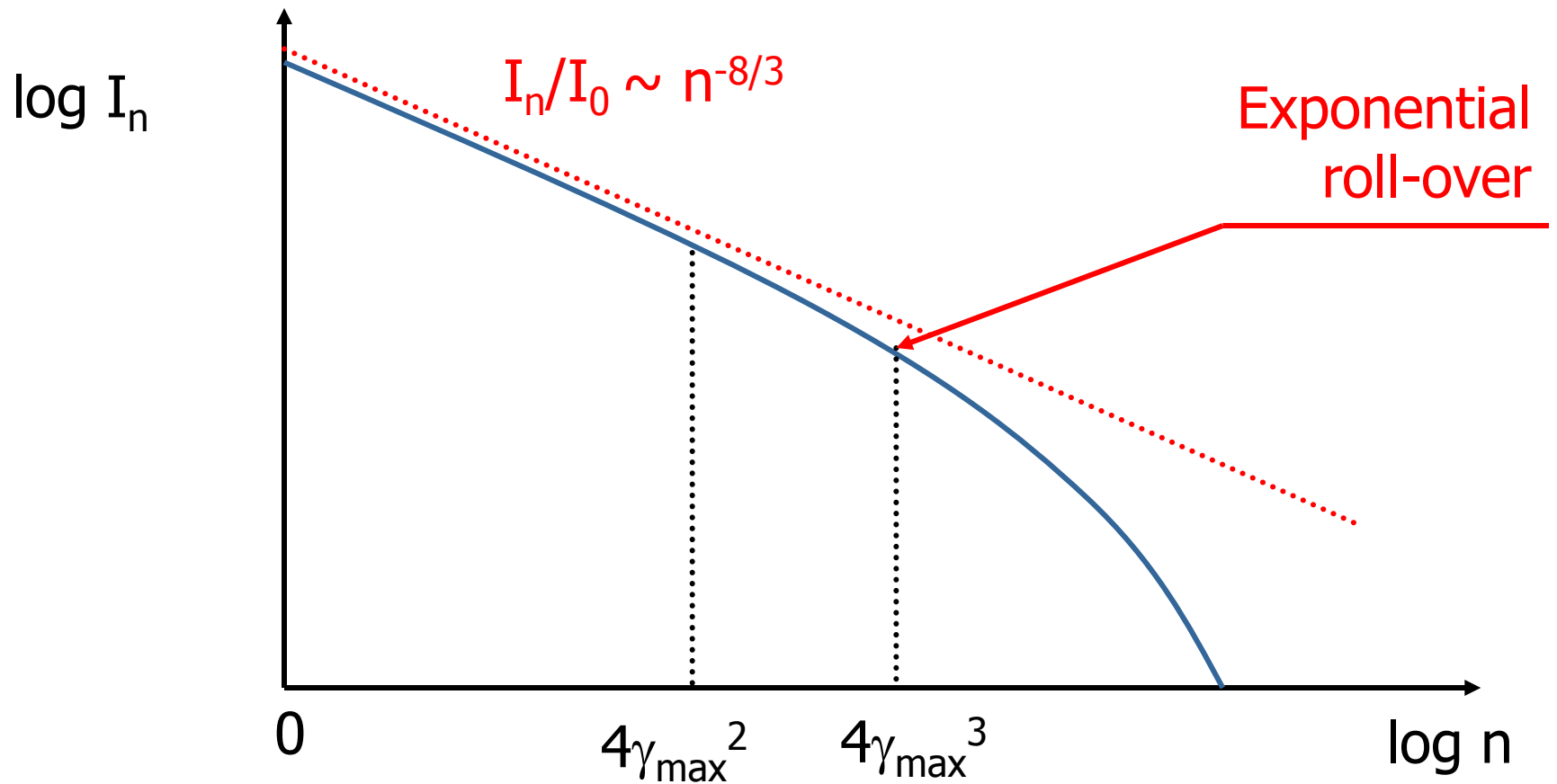
General form of the spectrum

The spectrum shape
for orders $n=1,2,10$
of the surface
velocity maximum.
 $\alpha=1, \gamma=8$ have been used



The BGP Spectrum

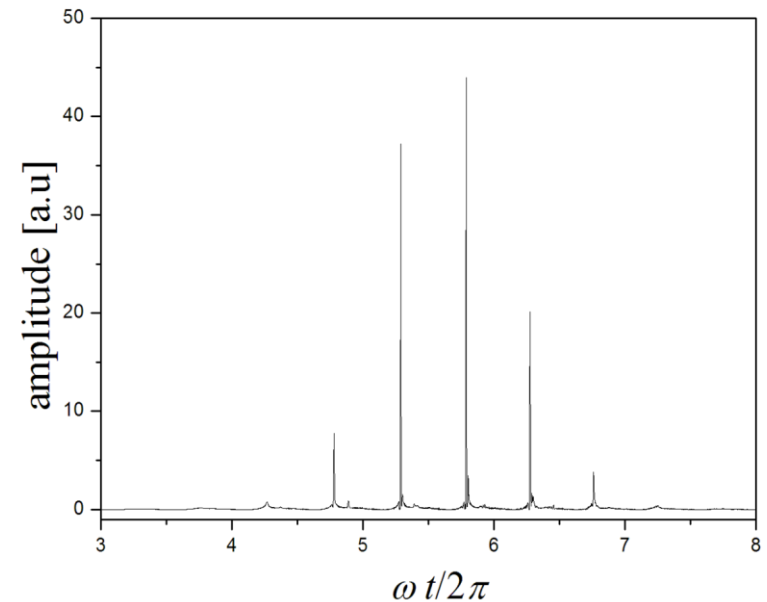
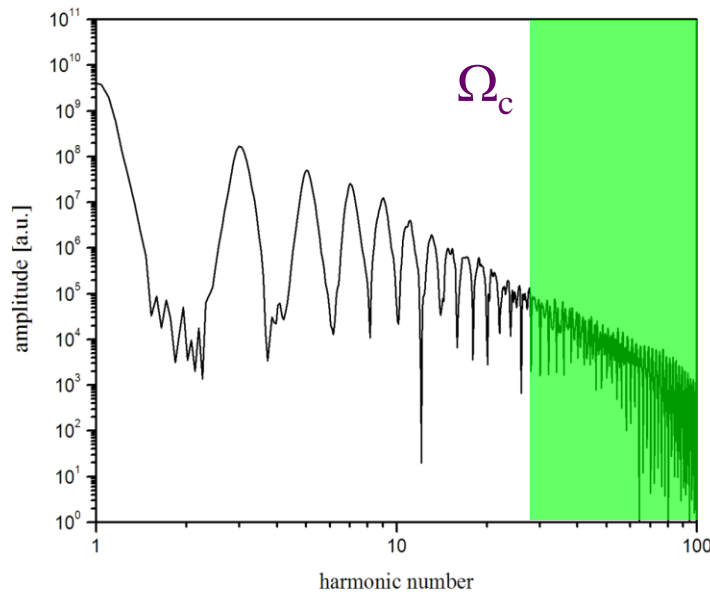
Baeva, Gordienko, Pukhov, *Phys. Rev. E* **74**, 046404 (2006)



(Sub-)Attosecond Pulses

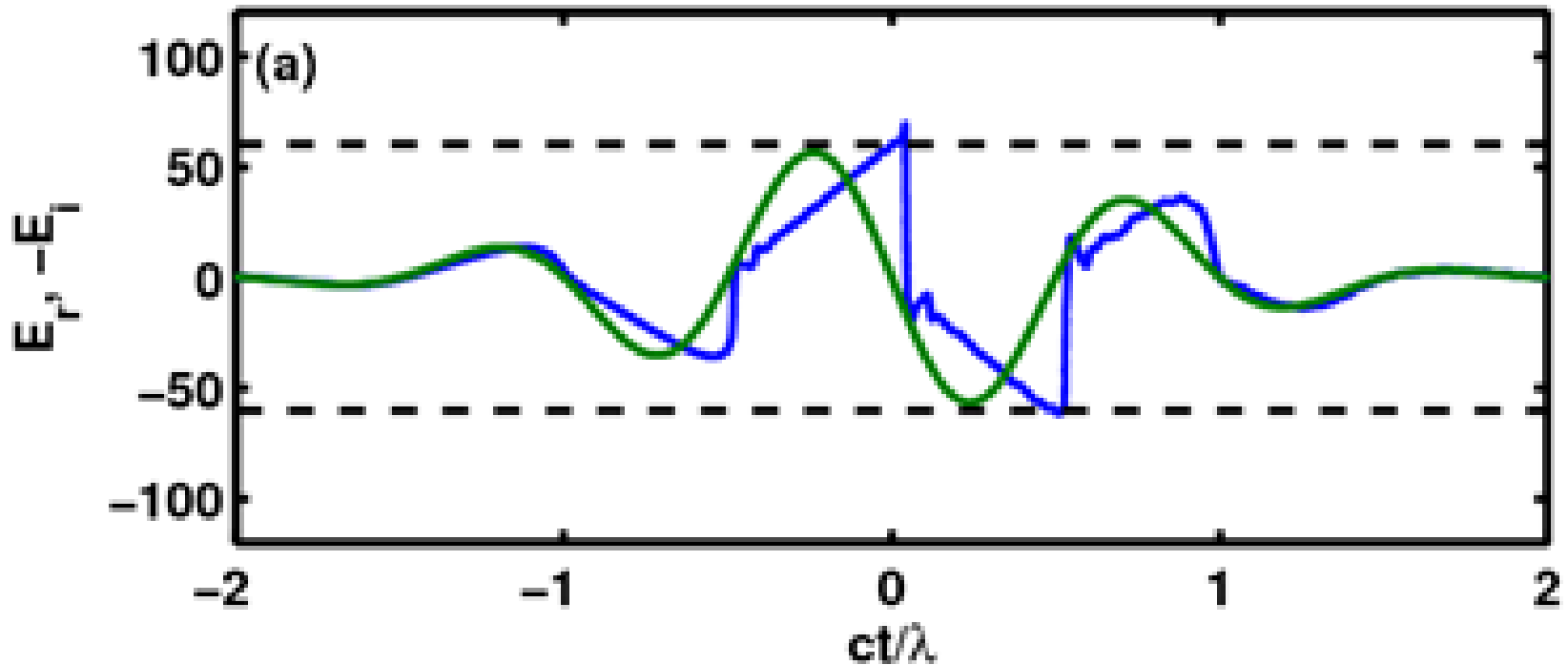
Baeva, Gordienko, Pukhov, *Phys. Rev. E* **74**, 046404 (2006)

- After proper filtering of RHHG one obtains a train of (sub-)attosecond pulses



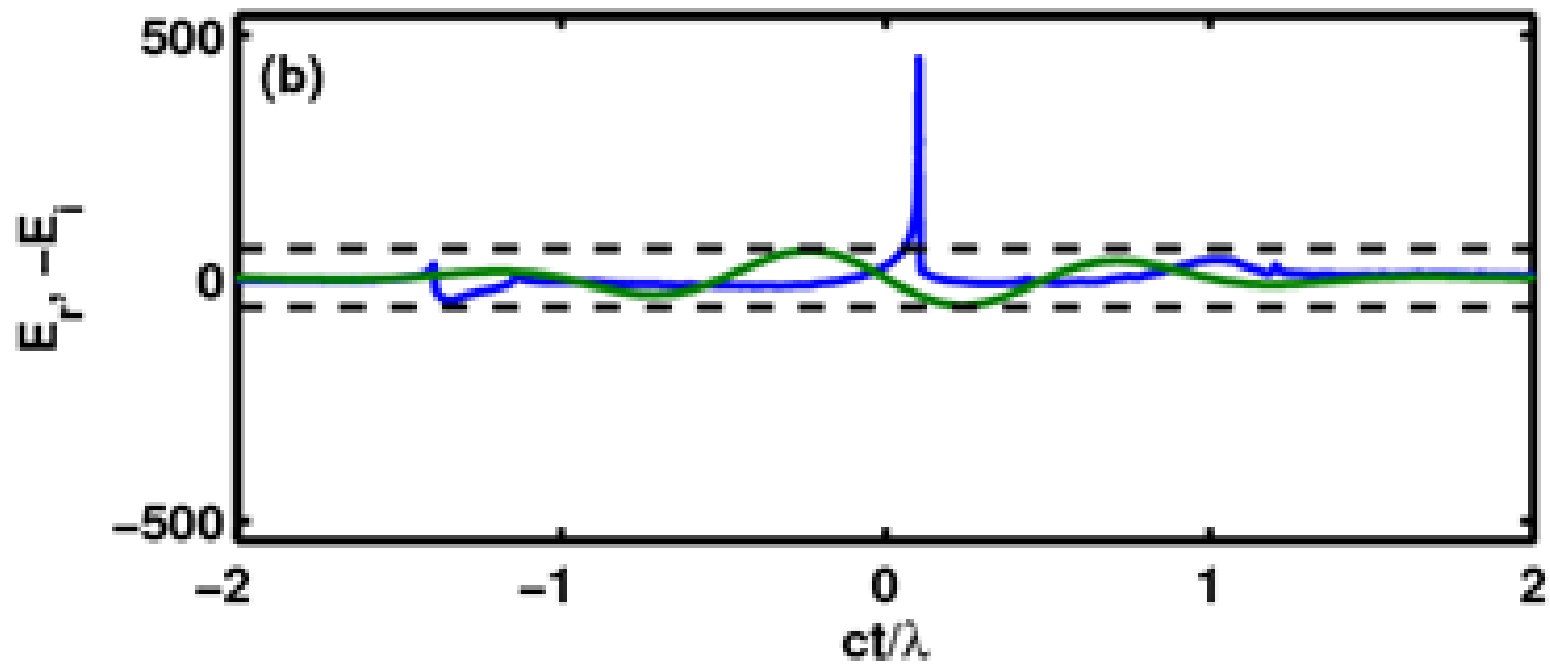
How universal is the $-8/3$ spectrum?

BGP case: pure phase modulation of the laser field



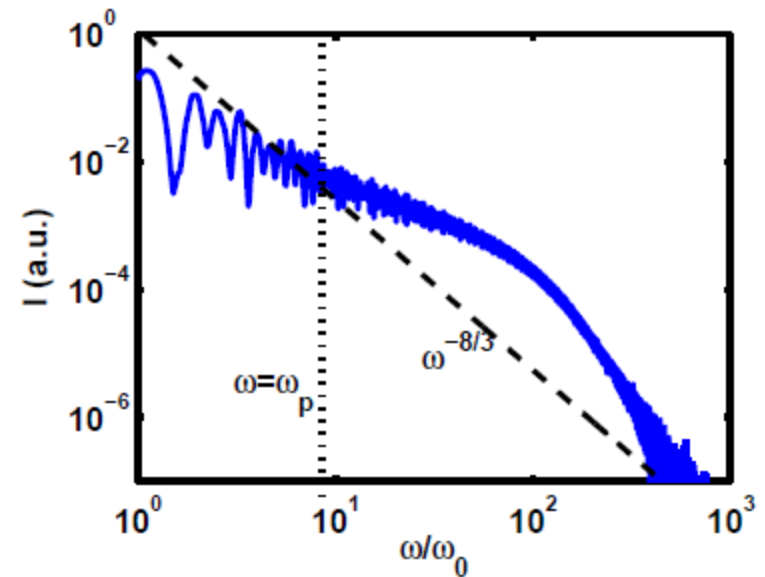
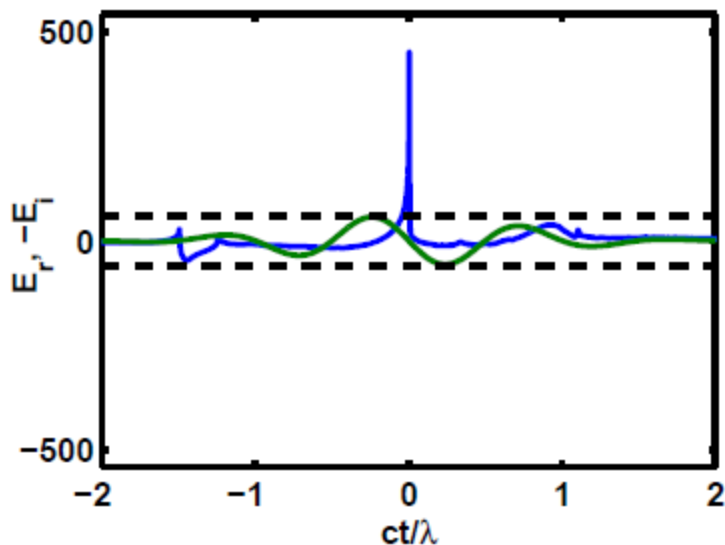
Giant single attosecond pulse

63° oblique incidence, selected CEP and density gradient.



Violation of the boundary condition. Much flatter spectrum, super intense isolated pulse

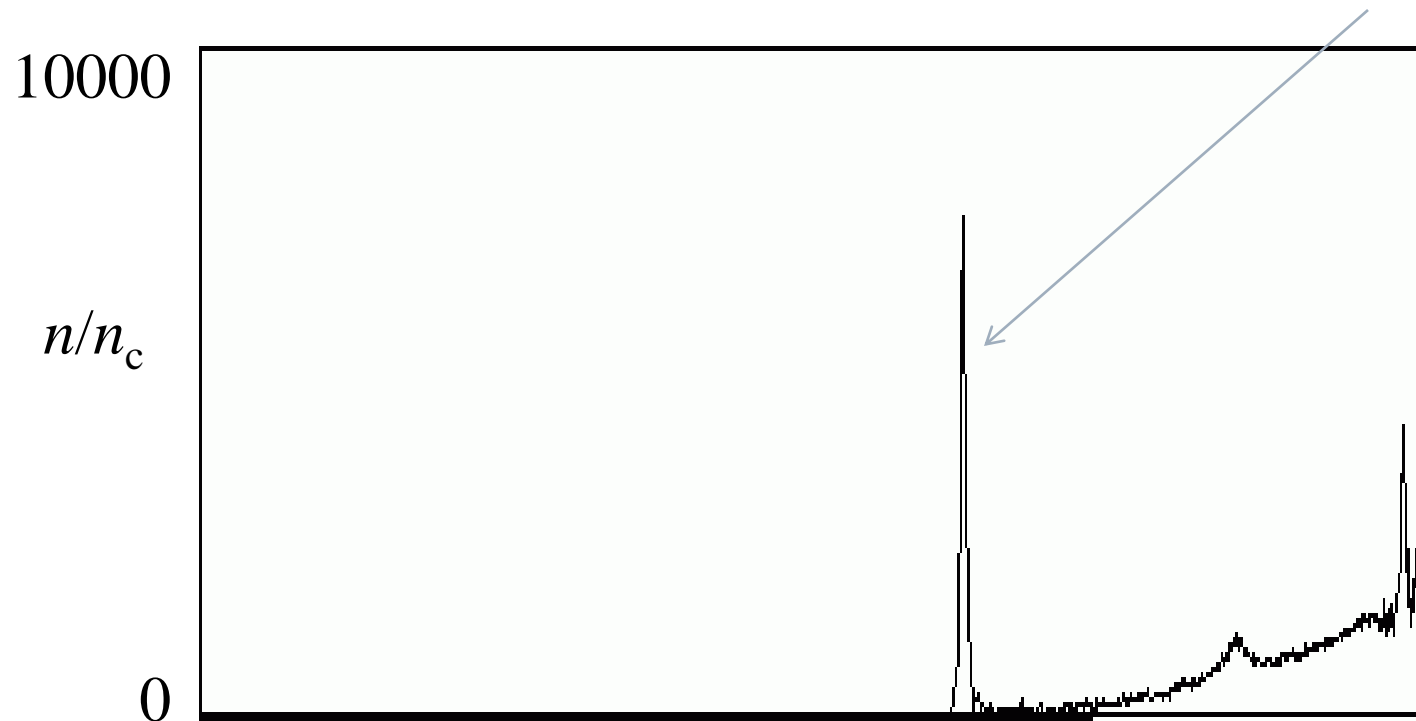
Simulation parameters: plasma density ramp $\propto \exp(x/(0.33\lambda))$ up to a maximum density of $N_e = 95 N_c$ (lab frame), oblique incidence at 63° angle (p-polarised). Laser field amplitude is $a_0 = 60$.

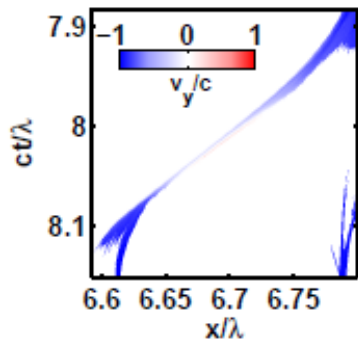


- Violation of ARP boundary condition \implies no ROM
- No spectral cut-off at plasma frequency \implies no CWE

Nanobunching in electron density

Coherent synchrotron emission from the density peak





Simulation parameters: $a_0 = 60$, plasma density ramp $\propto \exp(x/(0.33\lambda))$, maximum density $N_e = 95 N_c$, p-polarised incidence at 63°

- 2nd order zero in v_y
- Bunch width: $\delta_{FWHM} \approx 0.0015\lambda$, roughly Gaussian profile

Excellent Agreement between Theory and Simulation

- Theoretical Spectrum:

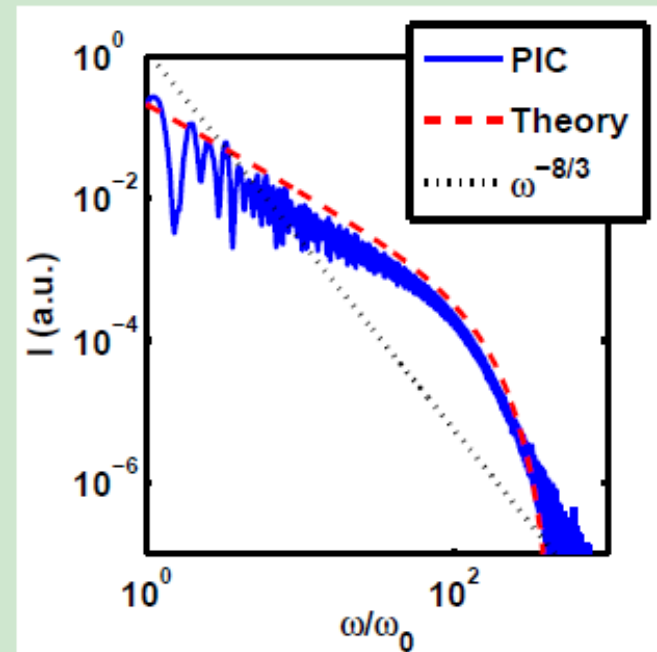
$$I(\omega) \propto \omega^{-6/5} \times \left[S'' \left(\left(\frac{\omega}{\omega_{rs}} \right)^{4/5} \right) \right]^2 \times \exp \left(- \left(\frac{\omega}{\omega_{rf}} \right)^2 \right)$$

using

$$\omega_{rs} = 800$$

$$\omega_{rf} = 225$$

compatible with PIC data

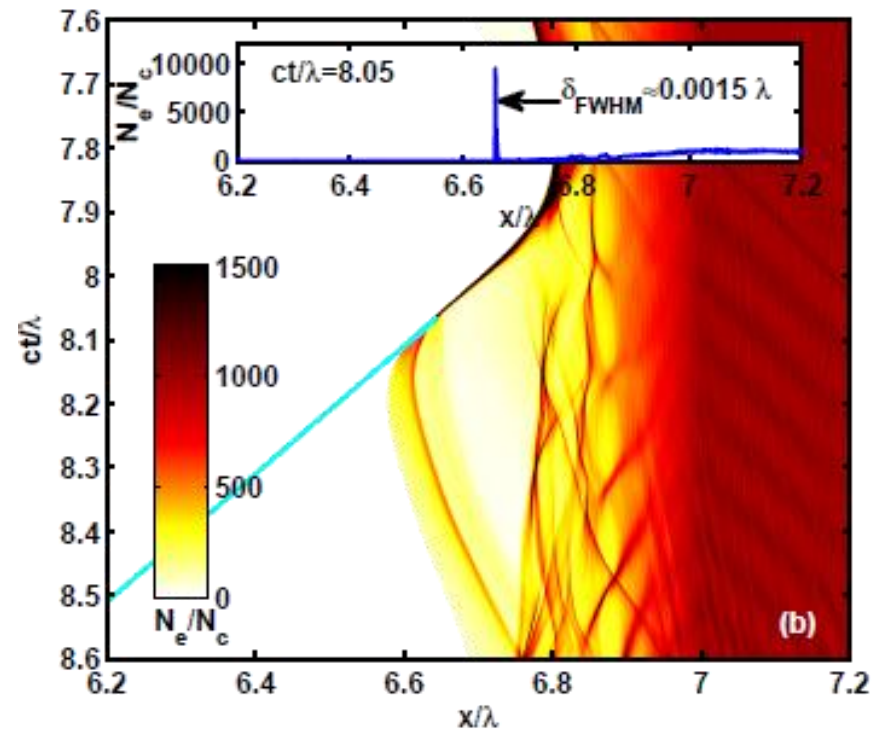
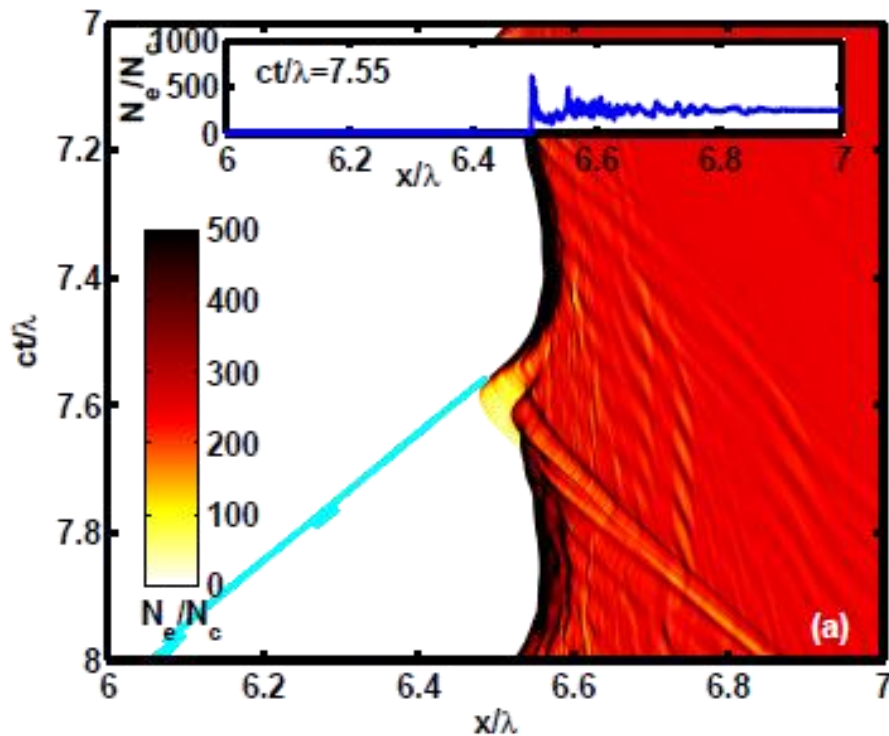


Two relativistic regimes of HHG

1. BGP case:
plasma boundary stays “conjunct”
the skin layer emits as a whole
→ the universal $n^{-8/3}$ spectrum
2. Nanobunching of plasma electrons,
Coherent Synchrotron Emission (**CSE**)
→ much flatter spectra, $n^{-4/3}$ or $n^{-6/5}$

DadB, Pukhov, Phys. Plasmas 17, 033110 (2010)

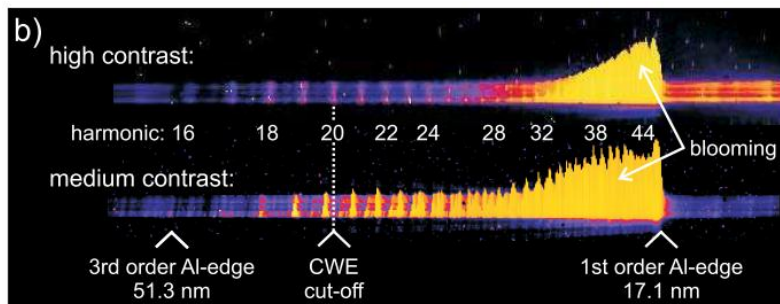
BGP vs CSE case



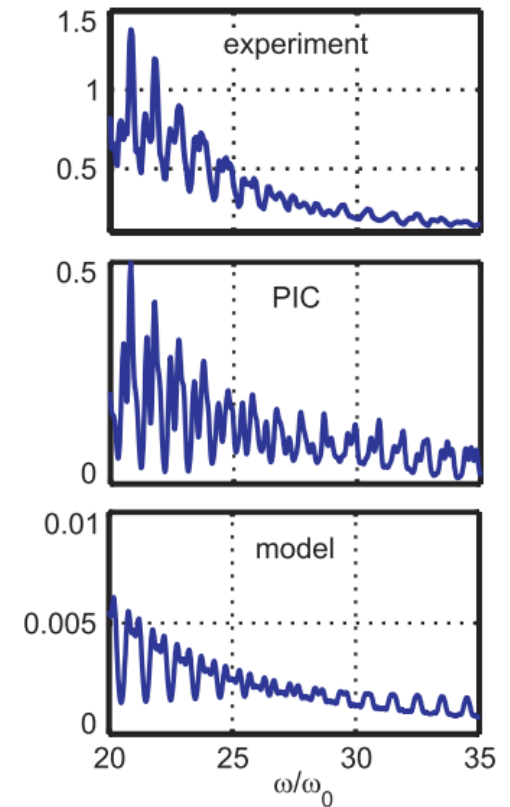
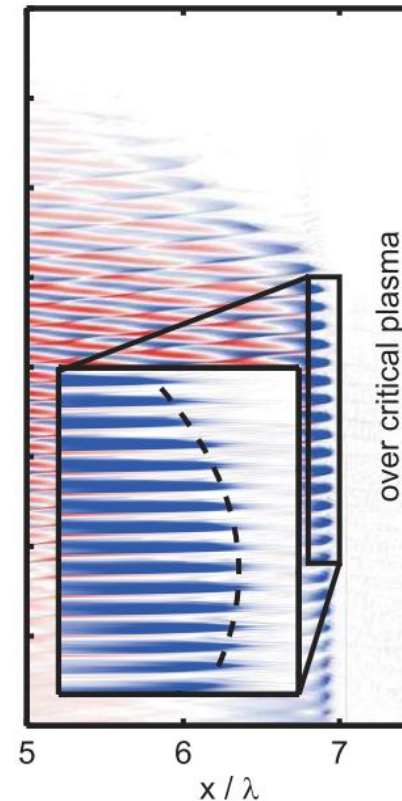
Pictures: electron density and contour lines (cyan) of the emitted harmonics radiation ($\omega > 4.5\omega_0$).
 Simulation parameters: $a_0 = 60$ in both cases, (a) normal incidence, $N_e = 250 N_c$, sharp edged profile;
 (b) plasma density ramp $\propto \exp(x/(0.33\lambda))$, maximum density $N_e = 95 N_c$, p-polarised incidence at 63°

HHG experiment on D'Arcturus

M. Behmke et al., *Phys. Rev. Lett.*, accepted (2011)



The harmonic spectrum contains information on the femtosecond dynamics of relativistic plasma



Summary

- High harmonics from solid surfaces may become the very bright coherent source of XUV and X-rays
- The spectrum is a power law with an exponential roll over. The exponent is
 $p = -8/3$ for BGP spectrum
 $p = -6/5$ for CSE spectrum
- Giant intensity attosecond pulse is generated due to CSE.
- Spectral modulations encode the femtosecond plasma surface dynamics